

Neutrino velocity measurement with ICARUS

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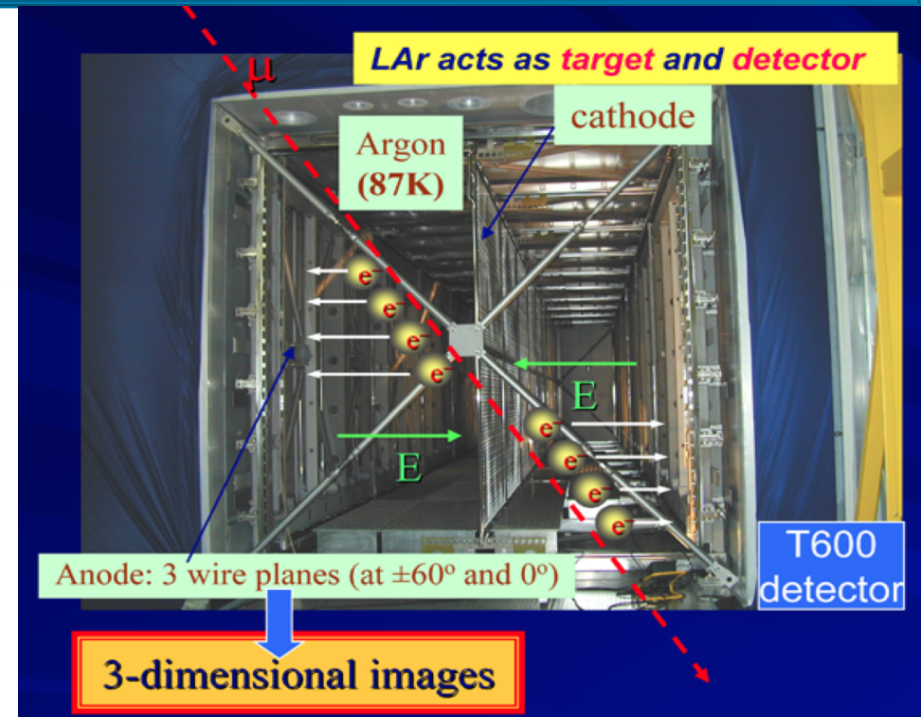
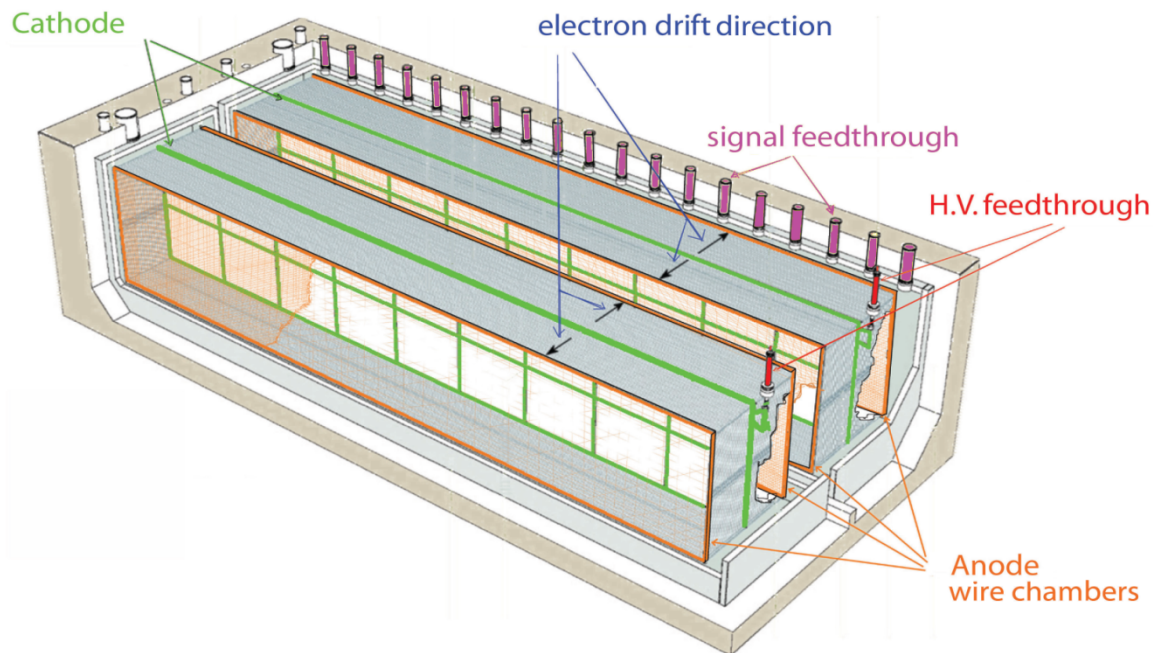
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ICARUS contribution to the superluminal neutrino problem

- ❑ A search for the analogue to Cherenkov radiation by high energy neutrinos at superluminal speeds in ICARUS.
(ICARUS Coll. and A. Cohen (Boston Univ.), arXiv:1109.6562)
- ❑ Measurement of the neutrino time-of-flight with the ICARUS detector at the CNGS beam.
(ICARUS Coll. and P. Alvarez Sanchez, J. Serrano (CERN), arXiv:1203.3433)
- ❑ Participation in the new measurement campaign in 2012.

The ICARUS T600 detector



- Two identical modules
 - $3.6 \times 3.9 \times 19.6 \approx 275 \text{ m}^3$ each
 - Liquid Ar active mass: $\approx 476 \text{ t}$
 - Drift length = 1.5 m (1 ms)
 - HV = -75 kV E = 0.5 kV/cm
 - v-drift = 1.55 mm/ μs
- 4 wire chambers:
 - 2 chambers per module
 - 3 readout wire planes per chamber, wires at 0, $\pm 60^\circ$
 - ≈ 54000 wires, 3 mm pitch, 3 mm plane spacing
- 20+54 PMTs, 8" \varnothing , for scintillation light:
 - VUV sensitive (128nm) with wave shifter (TPB)

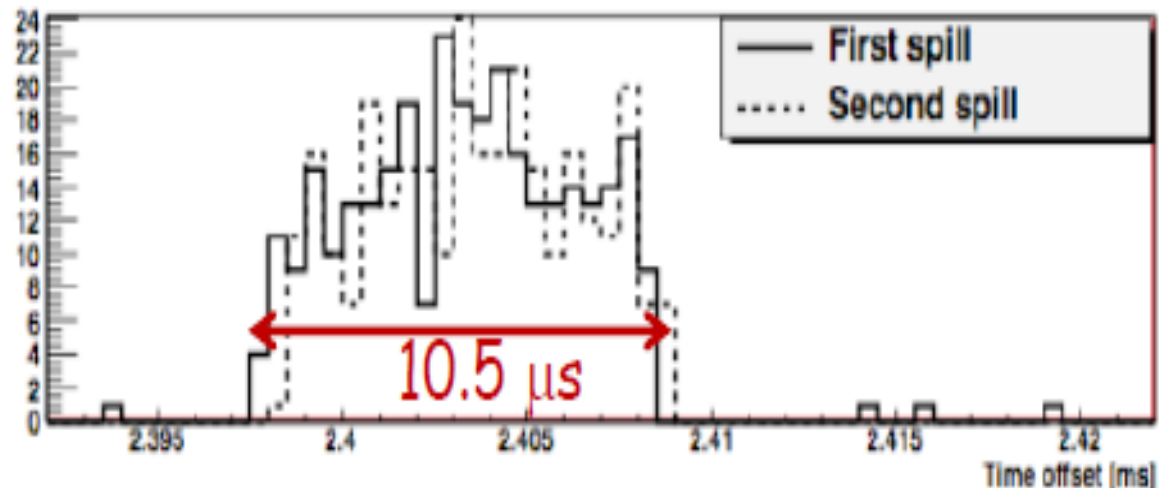
Key feature: LAr purity from electro-negative molecules (O_2 , H_2O , CO_2).
Now: 0.05 ppb (O_2 equivalent) \rightarrow 6 ms lifetime.

The ICARUS T600 light detection system

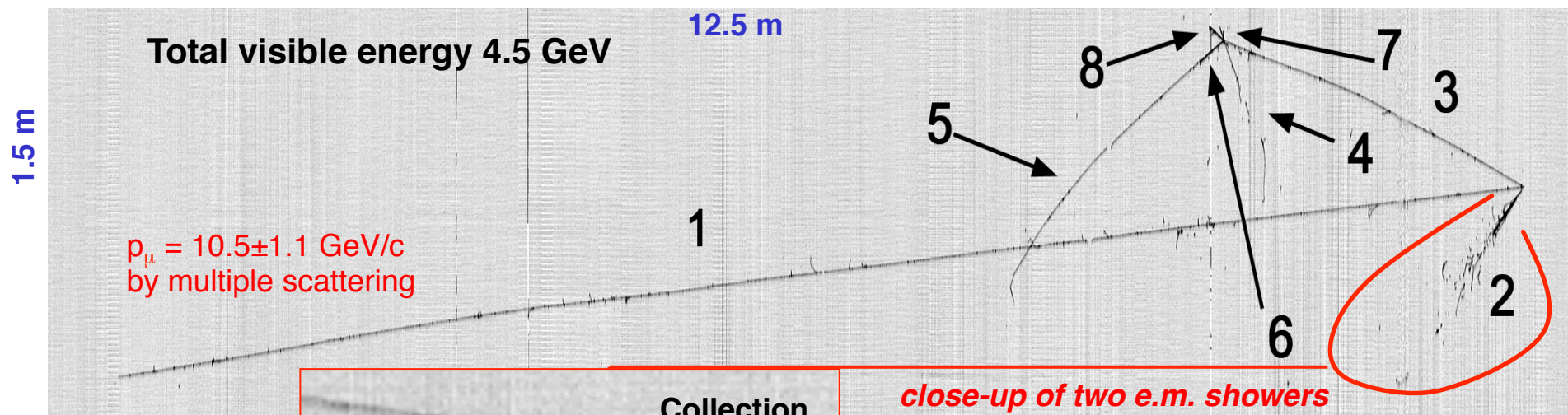
- Scintillation light produced by ionizing events ($\sim 2.5 \cdot 10^7 \gamma/\text{GeV}$) in the LAr-TPC is recorded with 74 (20+54) 8" PMT's.
- The PMT's in ICARUS T600 are organized in arrays behind the 4 wire chambers (1L, 1R, 2L, 2R). The 4 sums of the signals from the PMT arrays are used for triggering and to locate each event within the drift direction ($T=0$).
- The trigger threshold at about 100 phe allows detecting events with energy deposition as low as few hundred MeV with full efficiency.
- The scintillation in LAr has two distinct components with 6 ns (1/4 of the total) and 1.5 (3/4) μs decay time. About 4% of the γ ($10^6/\text{GeV}$) are emitted within 1 ns from the ionization process: 50-100 phe/GeV/D²[m] are recorded on the PMT's (D = distance in m from ionizing track to PMT).

ICARUS CNGS trigger

- CNGS “Early Warning” signal sent 80 ms before SPS proton extraction, containing the predicted extraction time of the 2 consecutive CNGS spills ($10.5\ \mu\text{s}$).
- Trigger: photomultiplier sum signal for each chamber within $60\ \mu\text{s}$ wide beam gate, open in coincidence with SPS proton extraction.
- 80 triggers per day recorded at a rate of about 1 mHz, with a time distribution in agreement with spill duration.
- 2.40 ms offset value in agreement with $\sim 2.44\ \text{ms}$ ν -tof ($\sim 40\ \mu\text{s}$ fiber transit time from external lab to Hall B).



LAr-TPC: powerful technique. Run 9927 Event 572

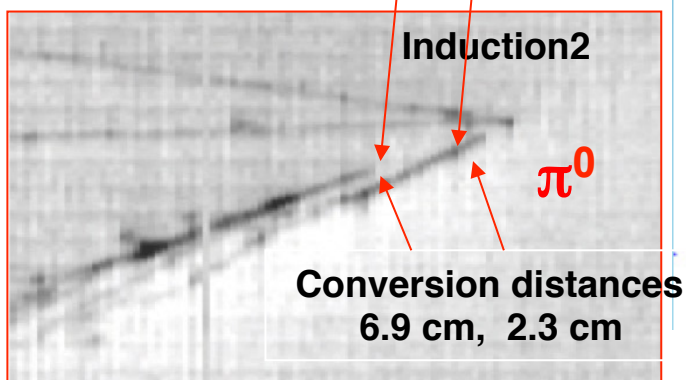
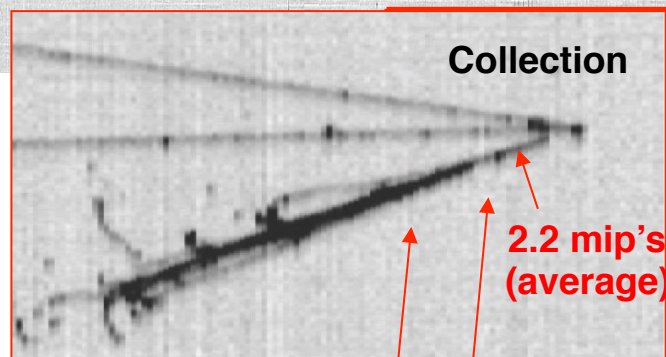


Primary vertex:

- very long μ (1),
- e.m. cascade (2),
- π (3).

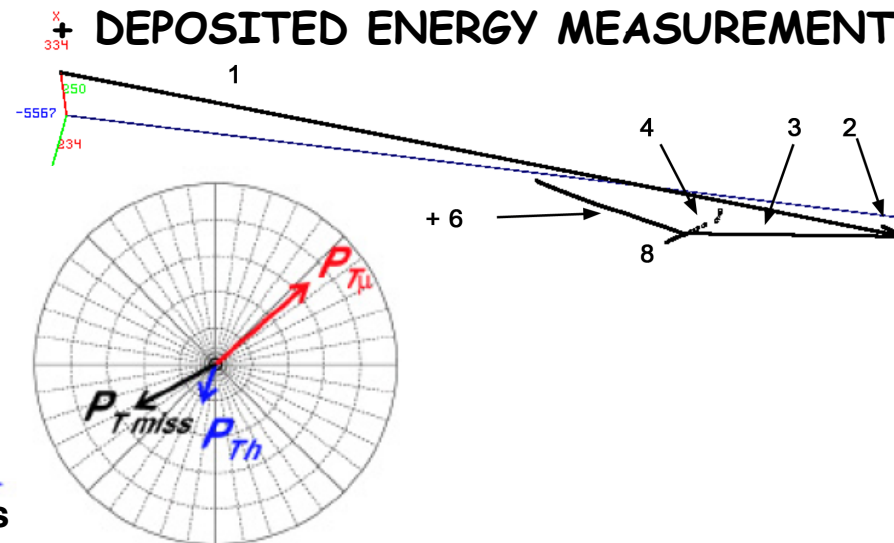
Secondary vertex:

Longest track (5) is μ coming from stopping K (6), μ decay is observed.



close-up of two e.m. showers

ALL PARTICLES RECONSTRUCTED IN 3D + DEPOSITED ENERGY MEASUREMENT



Total transverse momentum $\sim 250 \text{ MeV}$
(consistent with Fermi momentum)

$$M_{\gamma\gamma}^* = 125 \pm 15 \text{ MeV}/c^2$$

Search for the analogue of Cherenkov radiation by CNGS neutrinos at superluminal speeds in ICARUS

Search for superluminal ν 's radiative processes in ICARUS

- Cohen and Glashow [Phys. Rev. Lett., 107 (2011) 181803] argued that superluminal ν should lose energy mainly via e^+e^- brehmstrahlung:

$$\frac{dE}{dx} = \frac{5}{112} \frac{G_F^2}{192\pi^3} E_\nu^6 \delta^3 \quad \Gamma = \frac{2}{35} \frac{G_F^2}{192\pi^3} E_\nu^5 \delta^3 \quad (\Gamma: \text{Emission rate})$$

resulting in $0.78 \cdot E_\nu$ energy lost /emission on average where $\delta = (v_\nu^2 - c^2)/c^2$

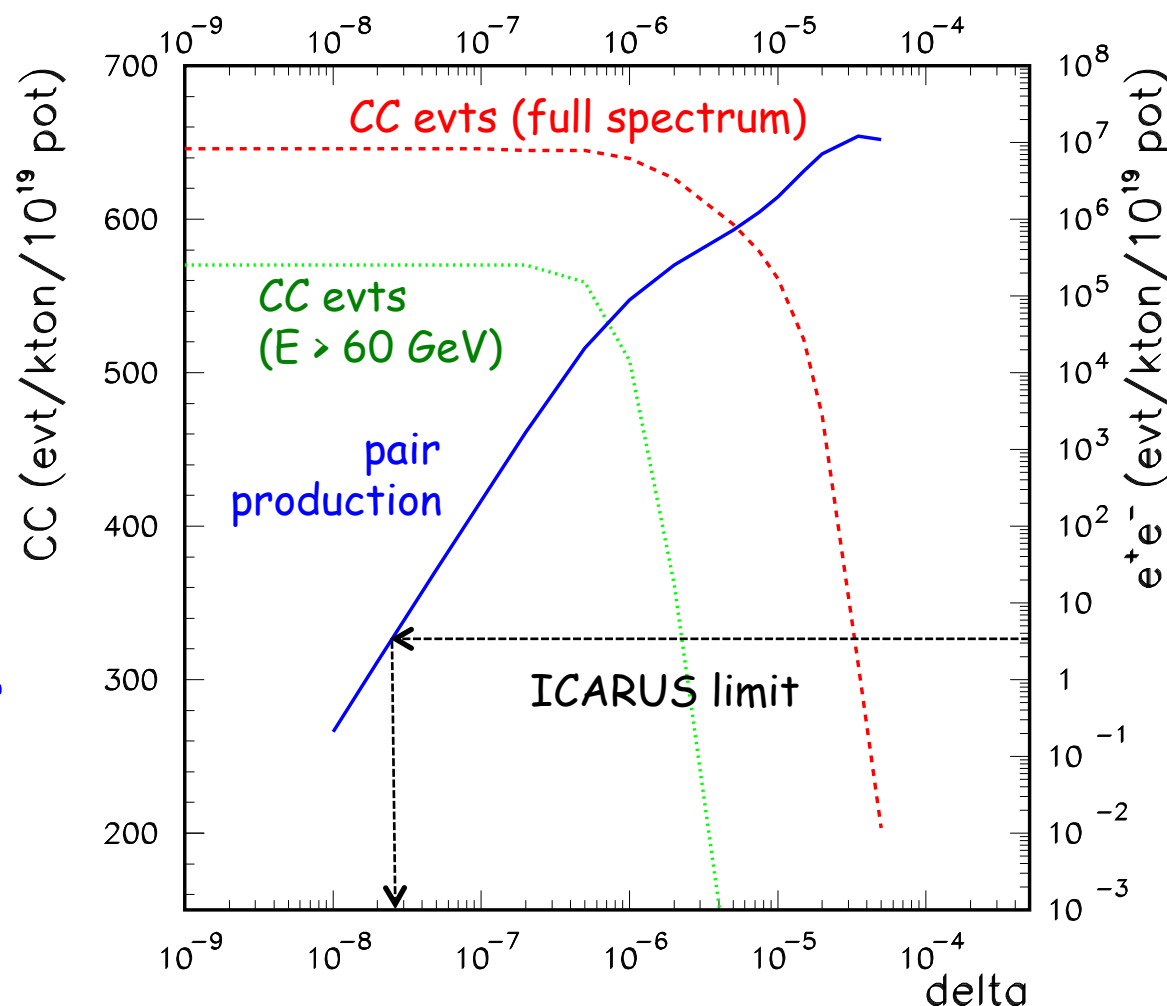
- The process kinematics, folded in the CNGS beam, has been studied with a full FLUKA simulation, for several values of δ parameter.
 - for $\delta = 5 \cdot 10^{-5}$ (OPERA) full ν event suppression for $E_\nu > 30 \text{ GeV}$ and
 - $\sim 10^7$ e^+e^- pairs / 10^{19} pot \cdot kt expected in ICARUS, appearing as isolated e.m. shower with $E_{\text{dep}} > 200 \text{ MeV}$, pointing to CNGS beam axis within 150 mrad and without hadronic activity.
- Effects searched in $6.7 \cdot 10^{18}$ pot \cdot kt CNGS ICARUS exposure (2010/11).
 - No spectrum suppression found in both NC and CC data (~ 400 events).
 - No e^+e^- pair event candidate found.

ICARUS result on superluminal δ parameter

- The lack of pair bremsstrahlung events in CNGS ICARUS 2010/2011 data, allows to set the limit:

$$\delta = (v_\nu^2 - c^2)/c^2 < 2.5 \cdot 10^{-8} \text{ 90\% CL}$$

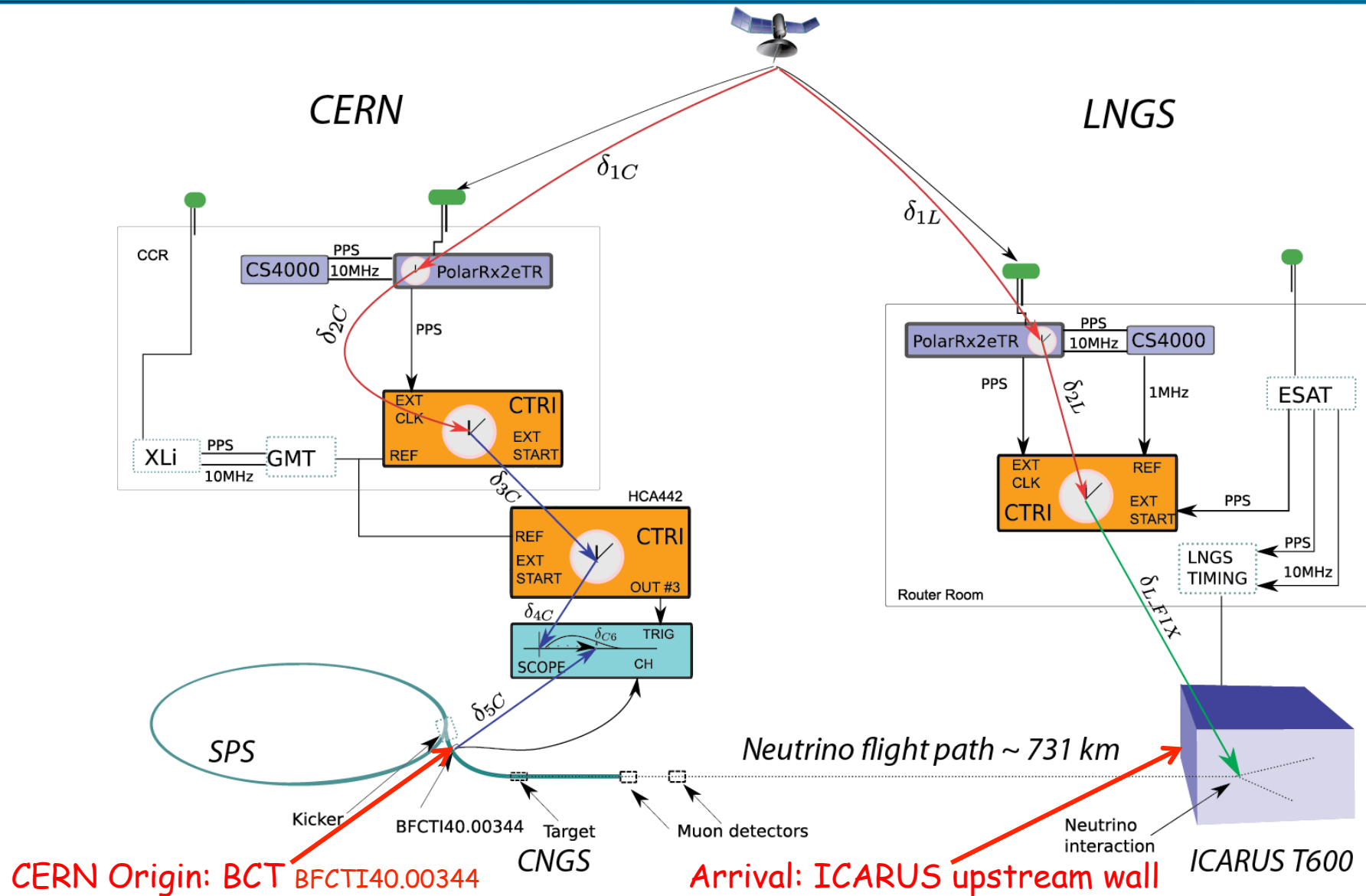
- comparable to the SuperK limit $\delta < 1.4 \cdot 10^{-8}$ from the lack of depletion of atmospheric ν 's,
- somewhat larger than the lower energy velocity constraint $\delta < 4 \cdot 10^{-9}$ from SN1987A.



- The remaining ICARUS 2011 events (3 x more statistics) and 2012 data will be available but, due to δ^3 dependence of pair bremsstrahlung cross section, no major improvement on δ sensitivity is expected.

Measurement of neutrino time-of-flight with the ICARUS detector in the CNGS beam

CERN-LNGS time-link for neutrino tof measurement

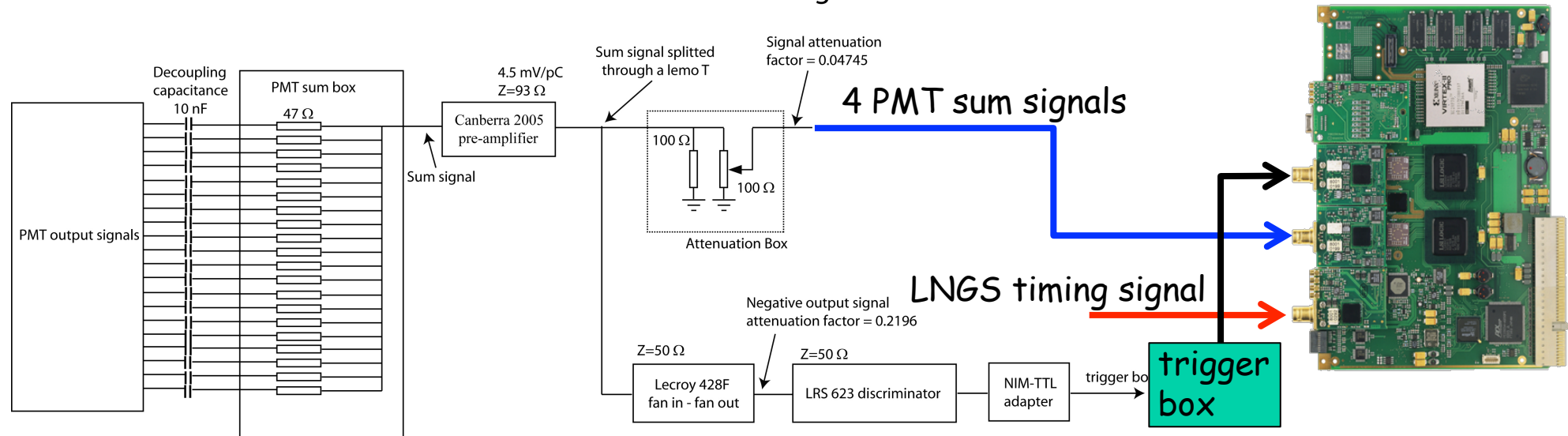


Proton extr. time and BCT waveforms, recorded with ACQIRIS 1 GS/s DP110, logged into CNGS DB at CERN, replicated at LNGS thanks to "Servizio di Calcolo"

✓ t.o.f measurement with ICARUS in CNGS bunched beam

- For the 2011 CNGS bunched beam run, the ICARUS PMT readout was equipped with an **independent DAQ system derived from that of the WArP experiment**, based on three "2-channel, 8-bit, 1-GHz ACQIRIS AC240" digitizer boards.
- The 4 PMT-Sum signals, integrated with a fast CANBERRA 2005 preamplifier, are recorded together with the absolute time signal distributed by the LNGS laboratory (PPmS).

Block scheme for the PMT sum signal



- The PMT-DAQ is triggered by the ICARUS-CNGS trigger.

PMT-DAQ scheme

- The dynamic range of AC240 boards can be adjusted to signal pulse height, and ranges from a minimum of 50 mV up to a maximum of 1 V, easily matching the ICARUS PMTs signal range (~ 3 phe/ADC #).
- Channel-to-channel synchronization: < 10 ps; Board-to-boards synchronization: 16 ns (with special jumpers ~ 0.1 ns). Stability: < 1 ppm.
- Similarly to ICARUS, in the new PMT-DAQ signals are continuously read out/stored in a circular memory (24 KB/ch depth corresponding to $24 \mu\text{s}$ time interval, sampling at 1 GHz).
- When a CNGS trigger occurs, additional data are readout and stored in an additional memory buffer until the required time interval is fully readout ($1400 \mu\text{s}$, ~ 1.4 MB/channel). The content of both buffers is then sent to the ICARUS data storage.

PMT-DAQ layout

1-L 1-R 2-R



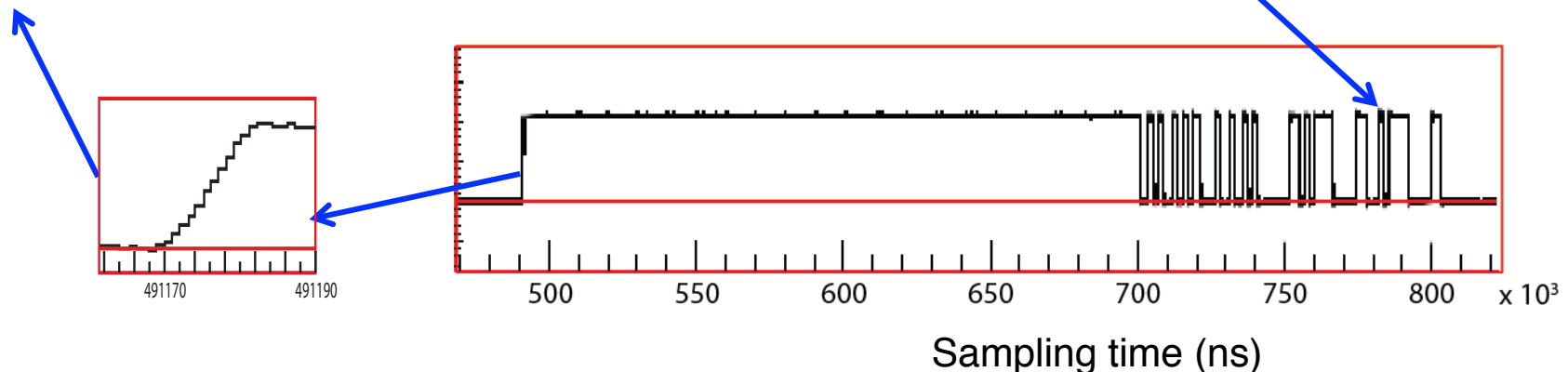
timing 2-L timing

LNGS timing signal

- LNGS timing signal (**PPmS**) consists of a TTL 15 ns positive edge sent out every ms and followed, after 200 μ s, by a pulsed structure whose logical states contain the information on the absolute time of the first edge. This signal generated in the external lab is sent to underground Halls via ~ 8 km of dedicated optical fiber.
- *This introduces a 42036.6 ± 1.3 ns delay accurately calibrated in Dec. 2011 in collaboration with the LNGS "Technical and Research divisions (Computing and Electronics Services)".*

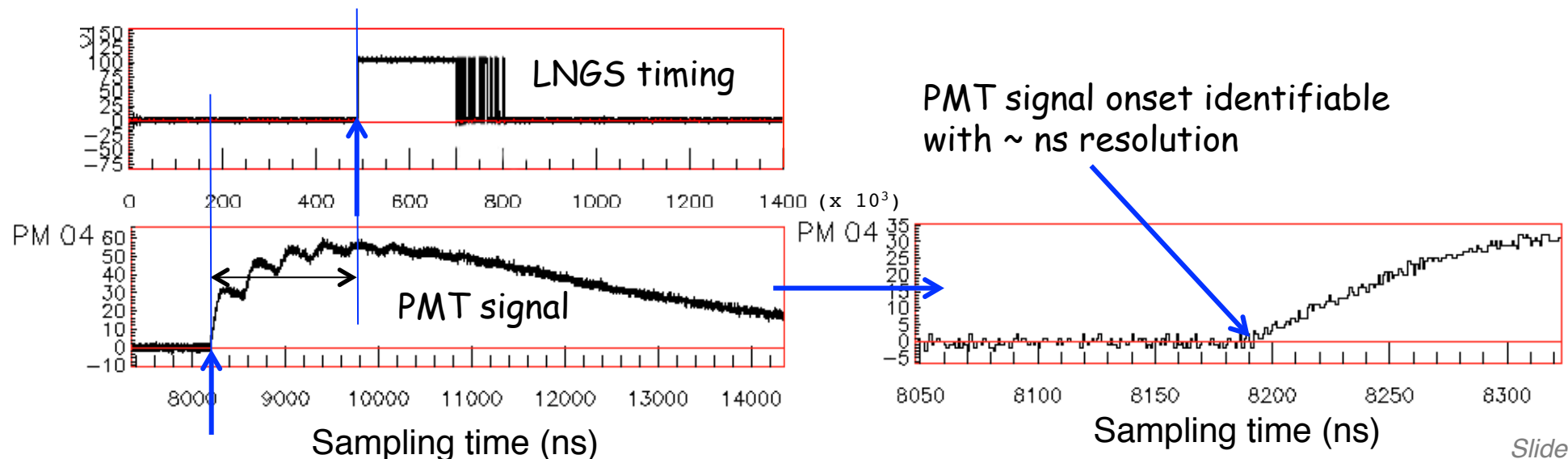
Stable and reproducible 15 ns leading edge
(few ns precision)

64 bits carrying information
about date and time (up to ns)



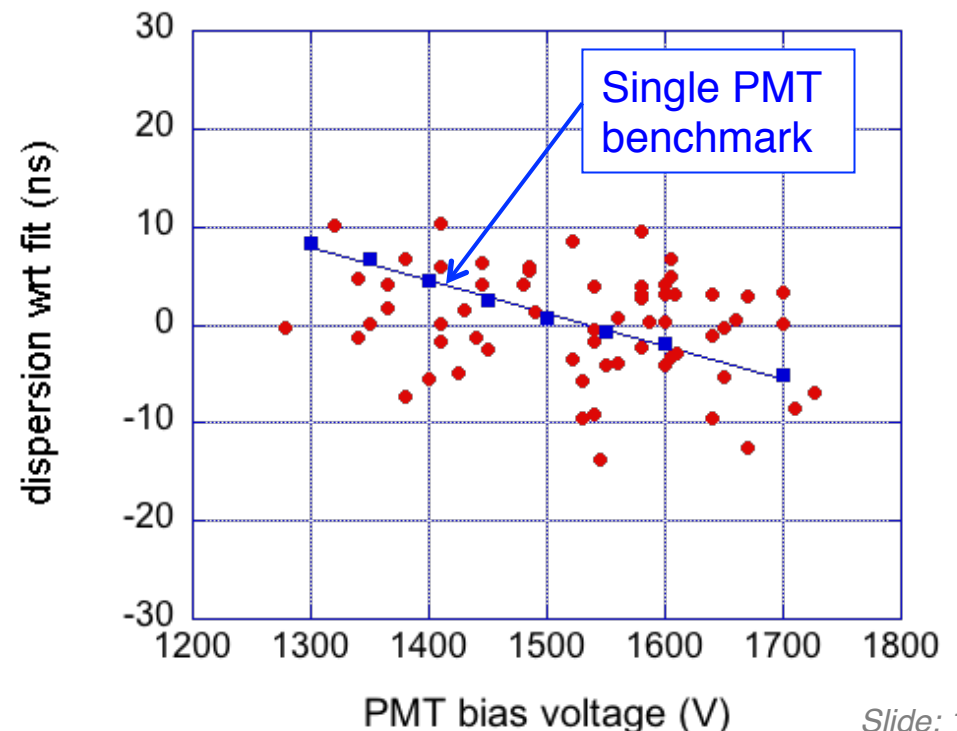
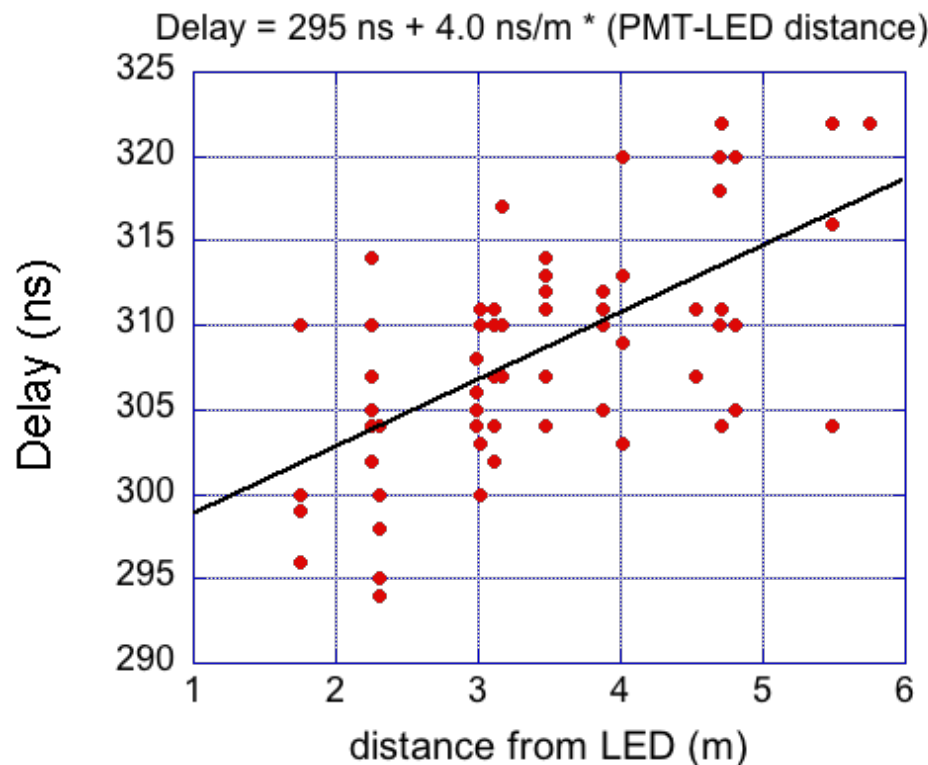
ICARUS PMT signal timing

- PMT and timing signals are recorded on the same AC240 boards. The time interval between these 2 signals, **measurable with few ns precision**, allows the determination of the absolute time of the scintillation light pulse in the T600 detector.
- Propagation of signals from PMT's to the AC240 boards including transit time in PMT (nominal value: $\sim 75 \pm 2$ ns) and CANBERRA (~ 5 ns) **has also been accurately calibrated** (with LED's in LAr and signal reflections): **299 (1L), 293 (1R) 295 (2L) 295 (2R) ± 5.5 ns** (mainly due to PMT transit time spread vs bias voltage).



Calibration of PMT's signal propagation

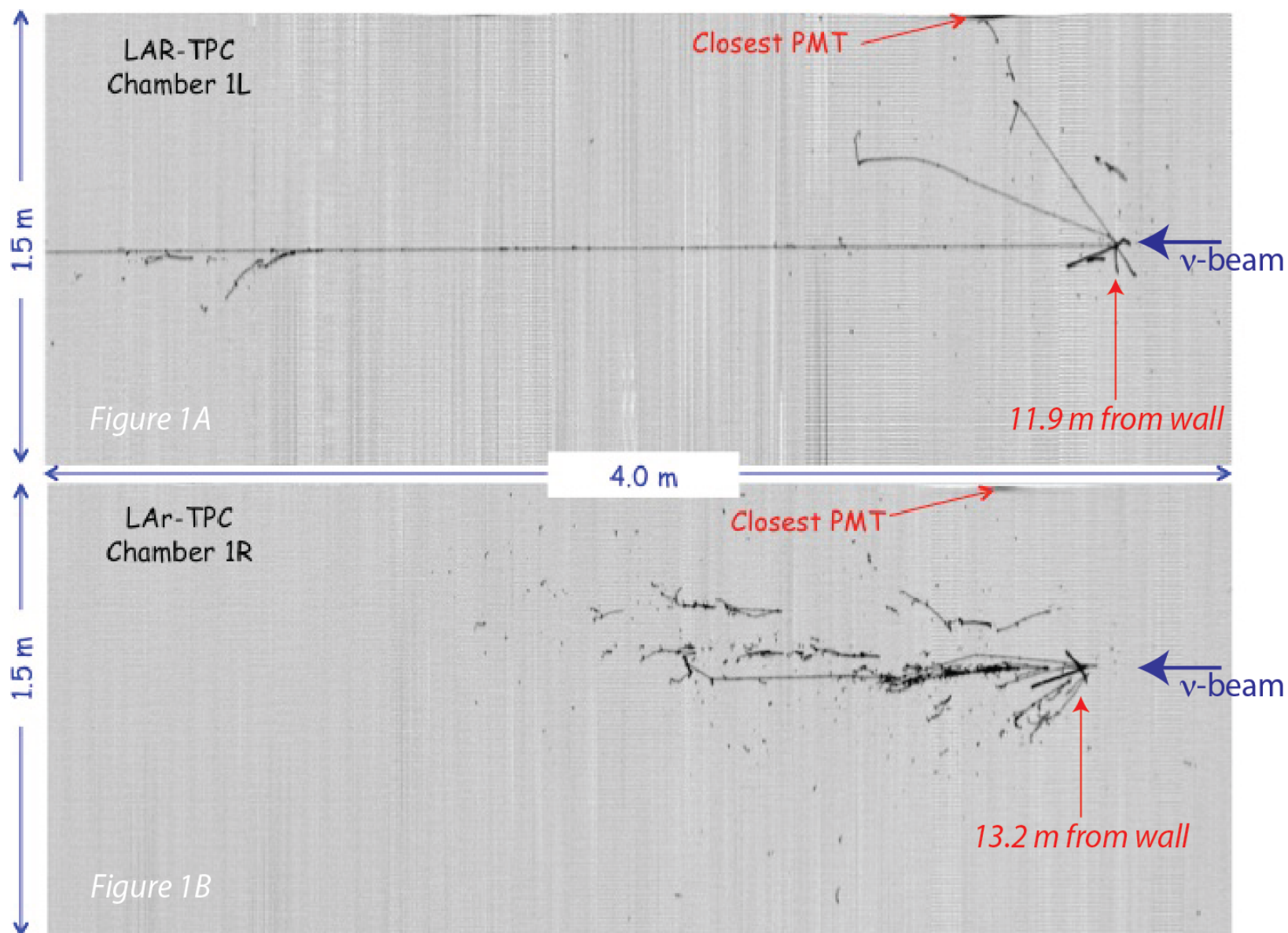
- Blue LED's immersed in LAr on top of Cathode
- Measurement of individual PMT's signals delay in agreement with measurement PMT-Sum with full cabling chain and measurement of each cable section with test pulse.
- Residual delay fluctuation (± 5.5 ns) compatible with variation of PMT transit time w.r.t biasing voltage.



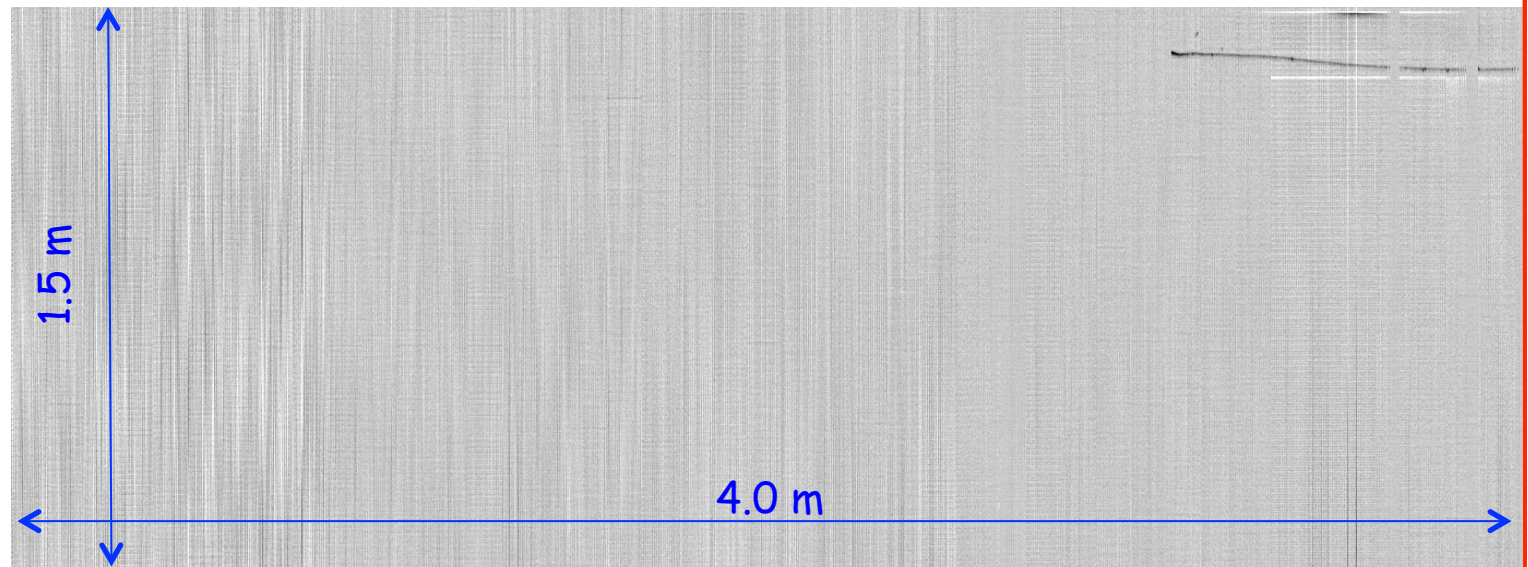
Data taking with bunched CNGS

- ICARUS started data taking with the low intensity bunched beam (4 bunches/spill, 3 ns FWHM, 524 ns separation) on October 27th.
- Unfortunately, due to an undetected time shift of 2 ms of the Early Warning Signal w.r.t. the actual CERN-SPS proton extraction time, ICARUS DAQ trigger gate was opened after the ν arrival at LNGS, resulting in empty events.
- The problem was spotted out and settled on October 31st. Until November 5th, ICARUS collected $\sim 2.2 \cdot 10^{16}$ pot.
- In this period ICARUS observed 7 beam-associated events, consistent with the expected overall event rate:
 - 2 CC ν_μ events
 - 1 NC ν event
 - 1 stopping + 3 crossing μ 's from ν interaction in upstream rock

CC and NC neutrino events

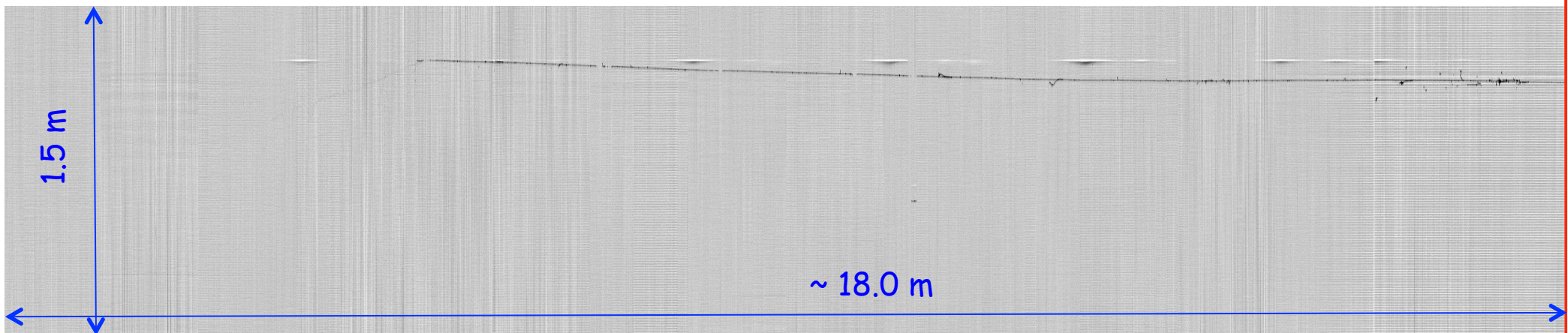


Stopping and through-going muons



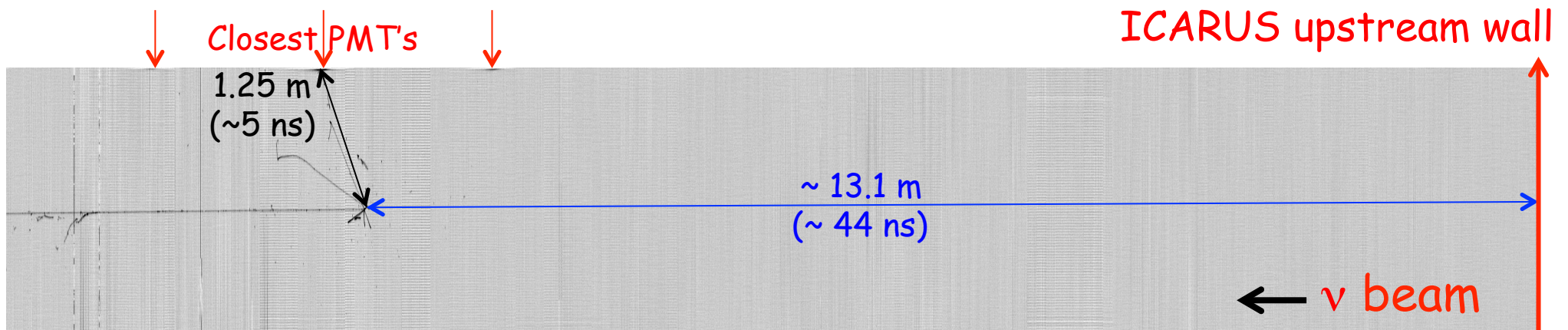
NOTE THE DIFFERENT ASPECT RATIOS

← ν beam



Neutrino arrival time

- To use the upstream wall position of ICARUS active volume as reference point for neutrino timing, some additional corrections have to be included namely:
 - the position of the interaction vertex along the 18 m of the detector length
 - the distance of the event vertex from the closest PMT.
- Both corrections can be precisely deduced from the event topology in the LAr-TPC through visual scanning. Events in the standard ICARUS DAQ and the new AC240 based DAQ are associated through their common acquisition time.



Neutrino event timing ($\text{tof}_\nu = T_{\text{stop}} - T_{\text{start}}$)

Starting Time @ CERN BCT:

$$T_{\text{START}} = T_{\text{CERN}} + \Delta_{\text{FIX}_C} + \Delta_{\text{VAR}_C}$$

- T_{CERN} : CERN SPS p-extraction time.
- Fixed delay: $\Delta_{\text{FIX}_C} = D_1 + \delta_{3C} + \delta_{4C} + \delta_{5C}$
 $D_1 = -99216$ ns (overall offset CERN -GPS time bases);
 $\delta_{3C} = 10078 \pm 2$ ns;
 $\delta_{4C} = 27 \pm 1$ ns;
 $\delta_{5C} = -580 \pm 5$ ns.
- Variable corr.: $\Delta_{\text{VAR}_C} = \delta_{1C} + \delta_{2C} + \delta_{6C}$
 δ_{1C}, δ_{2C} (GPS related);
 δ_{6C} p-pulse delay w.r.t. SPS extract.
(± 1.3 ns due to p-bunch width).

Arrival Time @ ICARUS entry wall:

$$T_{\text{STOP}} = T_{\text{LNGS}} + \delta_{L_FIX} + \Delta_{\text{VAR}_L}$$

- T_{LNGS} : LNGS PPmS time on AC240.
- Fixed delay: $\delta_{L_FIX} = 42037 \pm 1.3$ ns
(PPmS signal propagation).
- Variable corr.: $\Delta_{\text{VAR}_L} = \delta_{1L} + \delta_{2L} - (\delta_{3L} - \delta_{4L}) - \delta_{5L} - \delta_{6L} - \delta_{7L}$
 δ_{1L}, δ_{2L} (GPS related);
 $\delta_{3L} - \delta_{4L}$ (± 2 ns) PPmS-PMT delay;
 δ_{5L} (± 5.5 ns) PMT cabling delay;
 δ_{6L} ($< \pm 0.1$ ns) vertex position in LAr;
 δ_{7L} ($< \pm 1$ ns) vertex-PMT distance.

$\delta_{1C}, \delta_{2C}, \delta_{1L}, \delta_{2L}$: variable corrections calculable from GPS data.

Recipe, developed by OPERA/CERN, publicly available since end February 2012.

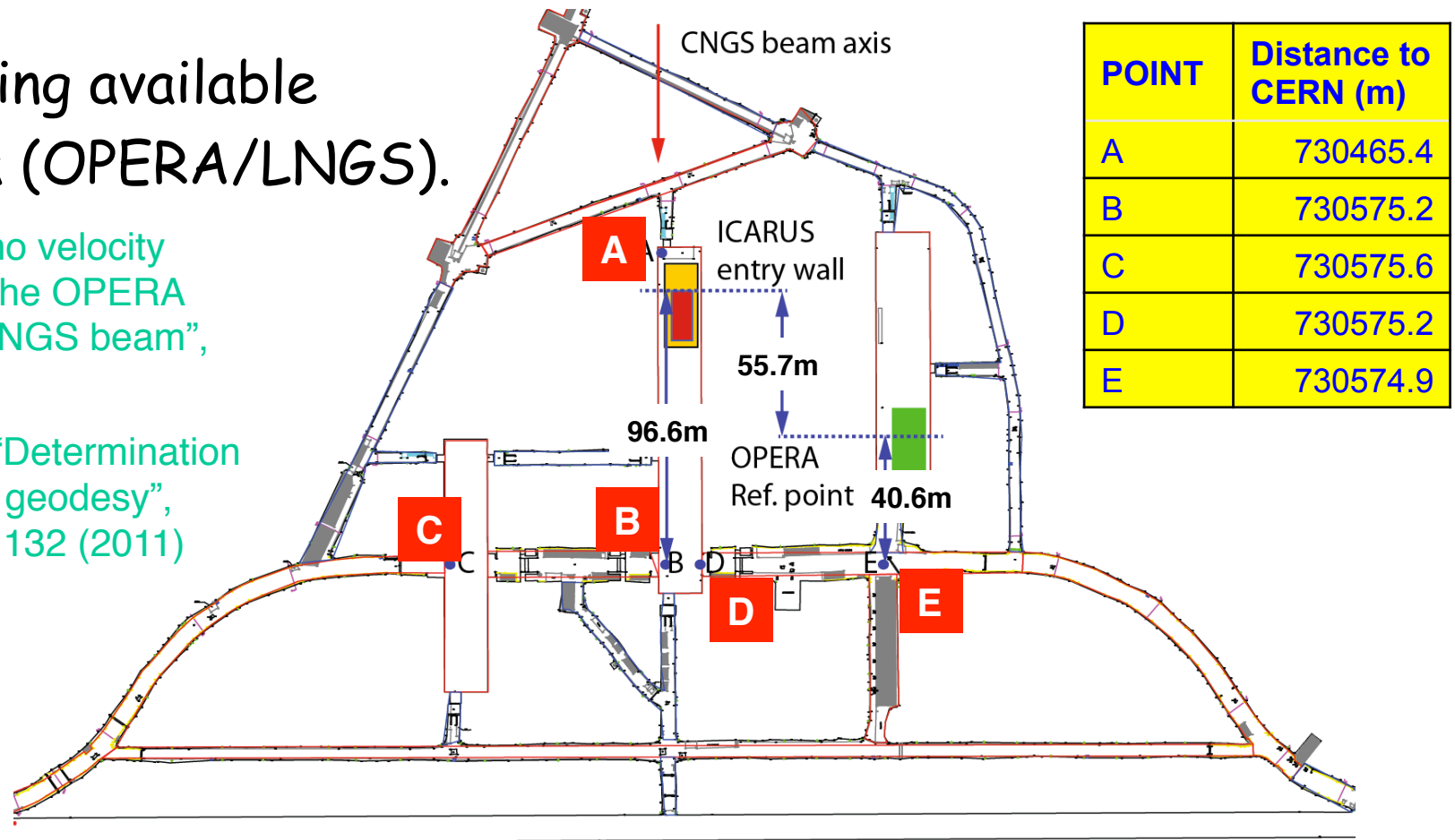
Estimated residual fluctuations on $(\delta_{1C} + \delta_{2C} - \delta_{1L} - \delta_{2L})$: ± 2 ns

CERN-ICARUS baseline estimation

Rely on existing available geodesy data (OPERA/LNGS).

G. Brunetti, "Neutrino velocity measurement with the OPERA experiment in the CNGS beam", PhD Thesis (2011)

G. Colosimo et al., "Determination of the CNGS global geodesy", OPERA public note 132 (2011)



- The distance from CERN origin (BCT) to ICARUS upstream wall is: 731222.3 ± 0.5 m. The expected time of flight for $v = c$ from the BCT position at CERN and the T600 upstream wall is 2439098 ± 1.7 ns (including 2.2 ns contribution due to earth rotation).

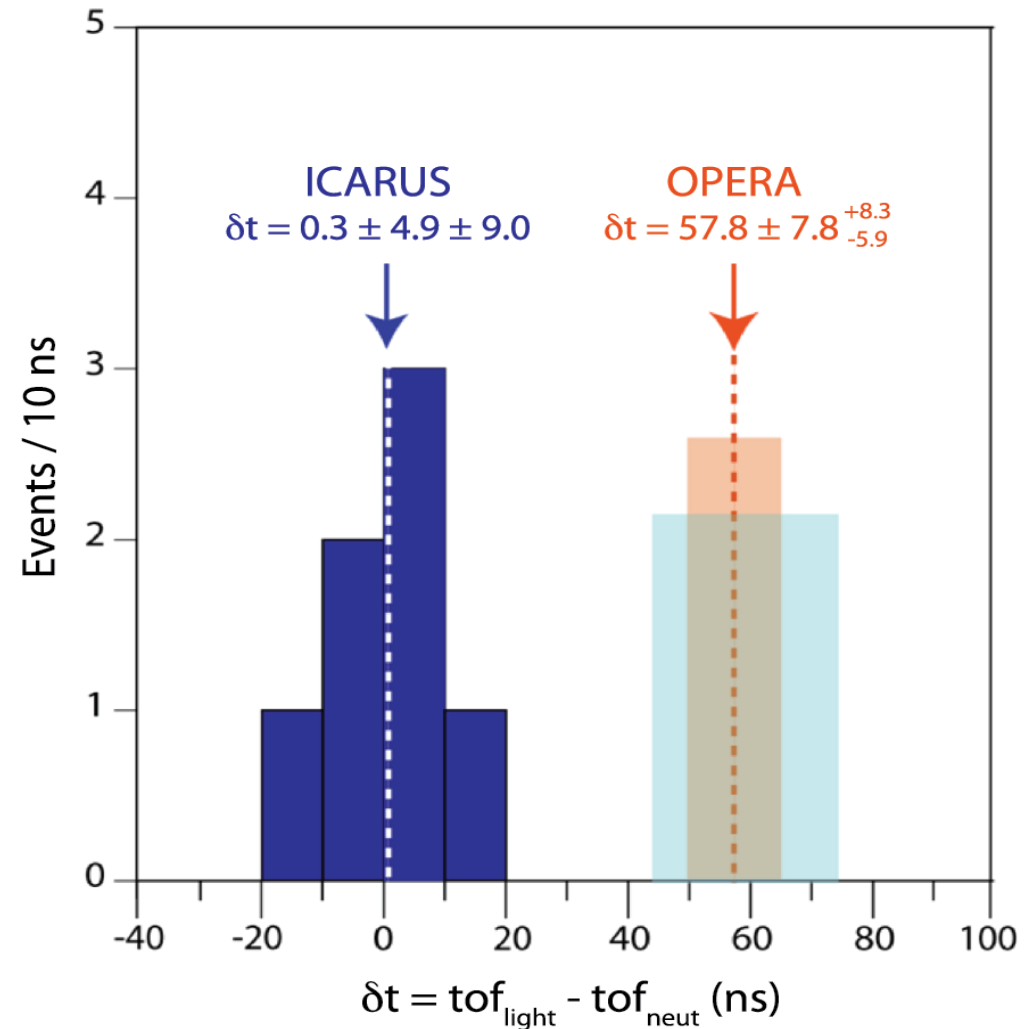
Summary of neutrino events and related corrections

Event #	1	2	3	4	5	6	7
<i>General properties of the event</i>							
Date	31-Oct	1-Nov	1-Nov	2-Nov	2-Nov	2-Nov	4-Nov
Time (UTC+1)	08:41:22.554	03:57:00.954	09:11:56.154	02:49:16.556	11:18:51.356	16:53:41.756	23:31:27.356
Run number	10949	10949	10949	10950	10951	10951	10956
Event Number	338	1247	1885	1053	91	802	773
Type of event	Mu-rock	Nu-NC	Nu-CC	Mu-rock	Nu-CC	Mu-rock	Mu-rock
TCP chamber (PMT array)	1R(1L)	1R(1L)	1L(1L)	2L(2R)	1L(1L)	2L(2R)	2L(2R)
<i>CERN related data</i>							
CERN extr. time, T_{CERN} (CERN t-base, ns)	133004	113399	114172	131046	116852	117040	118340
1 st CERN time link correction, δ_{1C} (ns)	2443	2457	2464	2487	2500	2507	2558
2 nd CERN time link correction, δ_{2C} (ns)	86603	86590	86585	86564	86556	86537	86500
BCT delay (first bunch), δ_{6C} (ns)	6047	6044	6049	6055	6050	6041	6046
P-transit time @ BCT, T_{START} (GPS t-base ns)	138406	118799	119579	136461	122267	122434	123753
<i>LNGS related data</i>							
LNGS PPMs time, T_{LNGS} (LNGS t-base, ns)	3000000	3000000	3000000	3000000	3000000	3000000	3000000
1 st LNGS time link correction, δ_{1L} (ns)	-7316	-7317	-7321	-7318	-7313	-7312	-7323
2 nd LNGS time link correction, δ_{2L} (ns)	7035	7030	7062	7029	7064	7060	7041
LNGS time on AC240, δ_{3L} (ns)	471602	491167	489889	474031	487231	486570	486296
PMT time on AC240, δ_{4L} (ns)	(*) 8196	8193	8194	8190	8185	8204	8212
PMT cable delay, δ_{5L} (ns)	299	299	299	295	299	295	295
Vertex position correction, δ_{6L} (ns)	0	40	44	16	28	5	1
PMT position correction, δ_{7L} (ns)	12	11	5	13	4	13	13
Event time in ICARUS, T_{STOP} (GPS t-base, ns)	2578030	2558421	2559730	2575578	2562406	2563101	2563357
<i>Neutrino time-of-flight calculation</i>							
$T_{\text{STOP}} - T_{\text{START}}$ (ns)	2439632	2439622	2440151	2439117	2440139	2440667	2439604
Nearest proton bunch	2	2	3	1	3	4	2
Bunch related additional time (ns)	524	524	1048	0	1048	1572	524
Neutrino time of flight (tof_{nu} , ns)	2439104	2439098	2439103	2439117	2439091	2439095	2439080
$\delta t = \text{tof}_c - \text{tof}_{\text{nu}}$ (ns, $\text{tof}_c = 2439098\text{ns}$)	-6	0	-5	-19	+7	+3	+18

For clarity, all times are referred to the CERN (UTC+1) time as T=0.

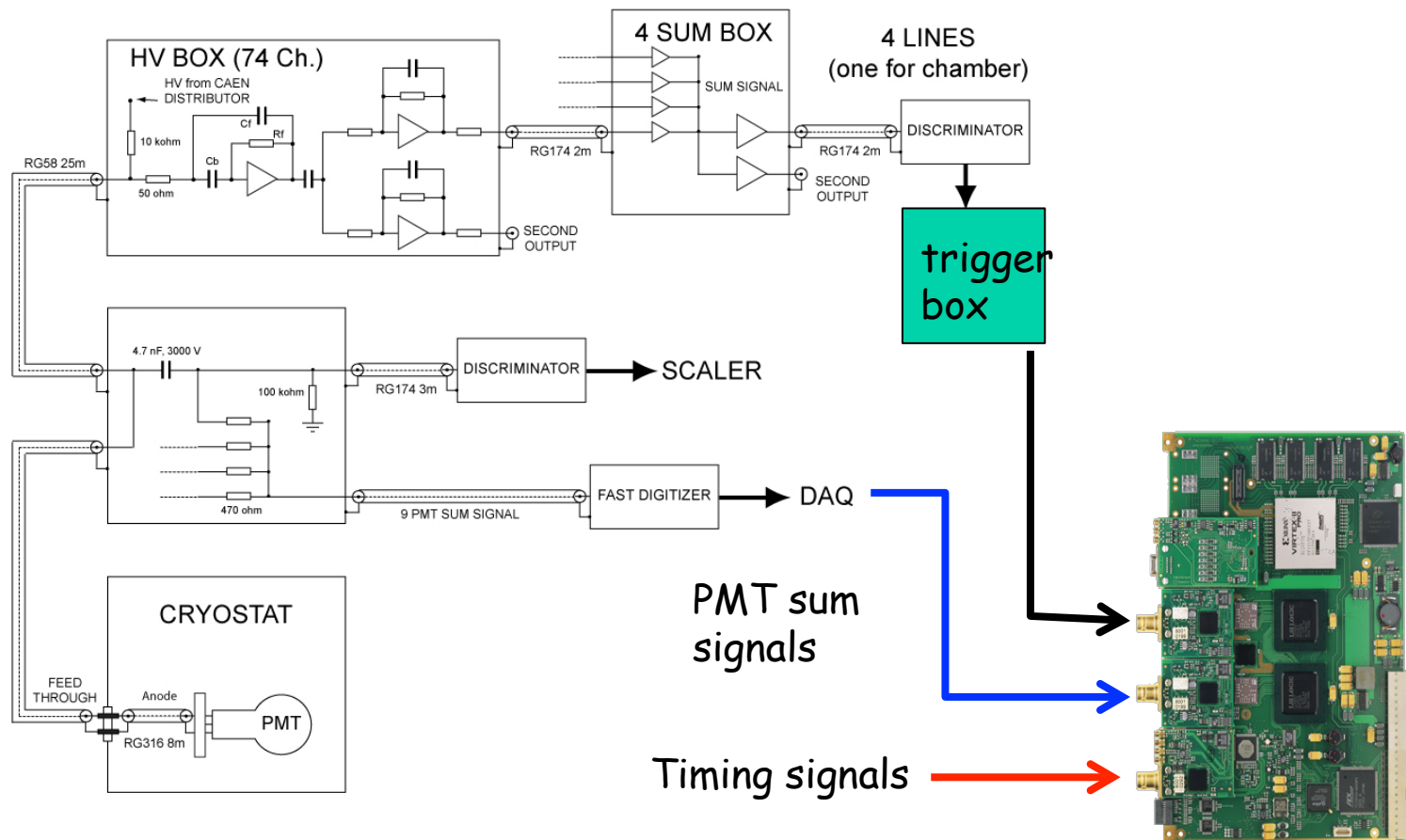
Neutrino time of flight result

- The average $\delta t = \text{tof}_c - \text{tof}_v$ of the 7 events is **+0.3 ns** with an r.m.s. of **10.5 ns** and **~9 ns** systematic error, obtained from combining in quadrature all quoted uncertainties.
- The statistical error on the average (**4.9 ns**) is estimated from a Student distribution with 6 d.o.f.
- **ν velocity compatible with speed of light**



Perspectives for 2012 run

- The new dedicated PMT-DAQ system proved to work as expected. Some improvements have been already implemented:
 - Direct PMT-sum signal recording with higher granularity (x 3)
 - All DAQ AC240 board synchronization at 0.1 ns



Timing and position measurement upgrades

- Three time distribution systems will be available :
 - In parallel with the recording of the **existing LNGS PPmS** signal, ICARUS will send a trigger signal, for time stamping, to the new **independent clock synchronization system at LNGS**, presently under installation in the external Labs by **Borexino Collaboration**.
 - Moreover ICARUS will profit of the **additional timing distribution system over the "White Rabbit"** protocol which be soon operation at LNGS under responsibility of CERN.
 - the provided PPmS signal will be recorded on PMT-DAQ
 - the available time stamping feature will be exploited.
- ICARUS will contribute to the new geodetic measurement campaign with Borexino/LVD foreseen for next months.

Thank you

BACKUP SLIDES

Search for superluminal ν 's radiative processes in ICARUS

- Cohen and Glashow [Phys. Rev. Lett., 107 (2011) 181803] argued that superluminal ν 's should lose energy by producing γ and e^+e^- pairs, through Z_0 mediated processes, as for Cherenkov light emission:

- The neutrino energy loss and pair emission rates are :

$$\frac{dE}{dx} = \frac{5}{112} \frac{G_F^2}{192\pi^3} E_\nu^6 \delta^3 \quad \Gamma = \frac{2}{35} \frac{G_F^2}{192\pi^3} E_\nu^5 \delta^3$$

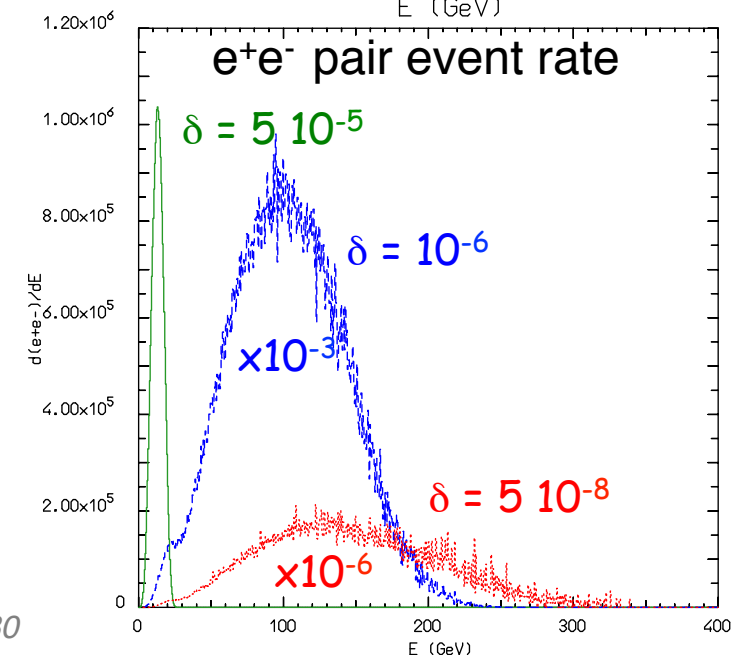
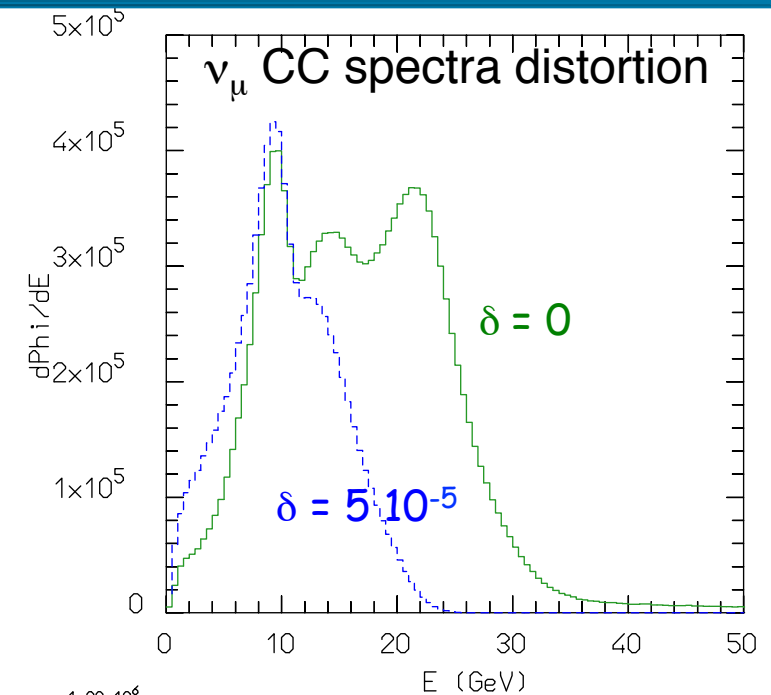
- with $\delta = (v_\nu^2 - c^2)/c^2$, resulting in $(\Gamma/E) (dE/dx) \sim 0.78 E_\nu$ energy lost per emission on average.
- To properly account for the actual process kinematics, a full FLUKA MC simulation of the effect, embedded in the CNGS beam, has been performed for several values of δ parameter:
 - for $\delta = 5 \cdot 10^{-5}$ the resulting mean free path is 490 Km at $E_\nu \sim 19$ GeV with a deflection angle $\sim \sqrt{\delta}$ comparable with CNGS beam width; full event suppression for $E_\nu > 30$ GeV expected.
 - e^+e^- emission angle w.r.t incoming ν direction is also $O(\sqrt{\delta})$.

Expectations:

- Both e^+e^- production and ν spectrum distortion have been studied to probe superluminal effects and test OPERA claim ($\delta \sim 5 \cdot 10^{-5}$). The e^+e^- signature allows to test much smaller values of δ , close to the SN1987A limit ($\delta < 4 \cdot 10^{-9}$).

Expected ν event rate and e^+e^- pair production spectra for 10^{19} pot*kt of ICARUS exposure and different δ values

δ	CC (all flavours)	NC (all flavours)	CC > 60 GeV ($\nu_\mu + \bar{\nu}_\mu$)	e^+e^-
0	644	203	57	0
$5 \cdot 10^{-8}$	644	203	57	27
$5 \cdot 10^{-7}$	643	203	56	$2.1 \cdot 10^4$
$5 \cdot 10^{-6}$	594	188	8.5	$7.2 \cdot 10^5$
$5 \cdot 10^{-5}$	203	85	$< 10^{-6}$	$1.1 \cdot 10^7$

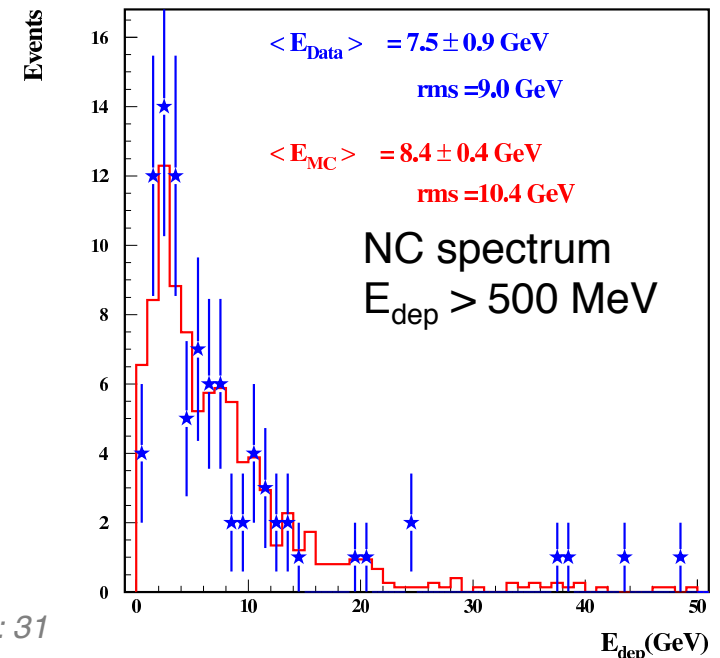
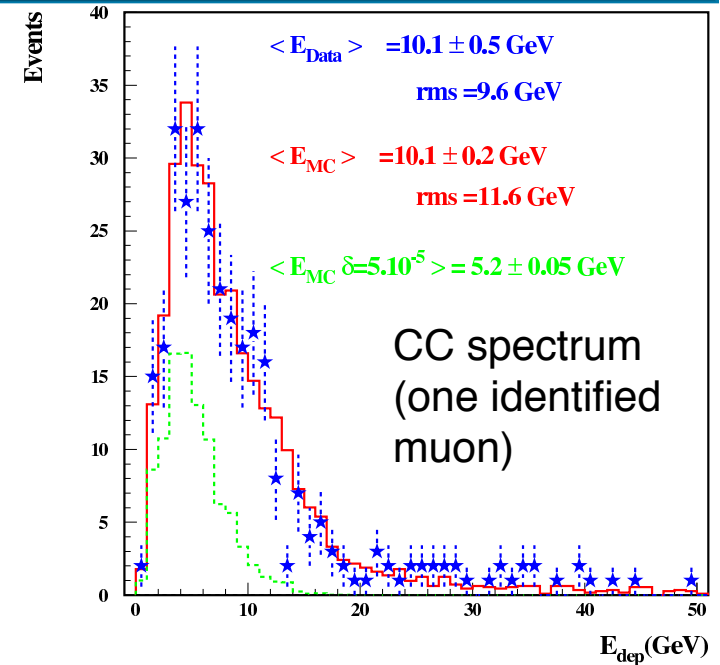


CNGS ν spectrum distortion in ICARUS

- Data from 2010 and part of 2011 CNGS runs (5.05 10^{18} pot·kt fiducial exposure).
- Bias-free analysis:
 - calorimetric measurements of raw event energy deposition in ICARUS (only correction for signal quenching in LAr applied).
 - 2.5 m downstream fiducial volume cut on interaction vertex to identify muon tracks (for CC /NC separataion).

Rates	Observed	Expected $\delta=0$	Expected $\delta=5 \cdot 10^{-5}$
CC	308	315 ± 5	98.1 ± 2
NC	89	93.1 ± 3	33.0 ± 1
ν_μ CC, $E_{\text{dep}} > 25 \text{ GeV}$	25	18 ± 1.3	$< 10^{-6}$

- Experimental E_{dep} distributions for CC and NC events in agreement with MC expectations for undistorted ($\delta = 0$) CNGS ν beam.
- $E_\nu > 30 \text{ GeV}$ full event suppression, predicted for $\delta = 5 \cdot 10^{-5}$, not observed.



Search for e^+e^- pairs in CNGS beam with ICARUS

- Both $\nu_x \rightarrow \nu_x + \gamma$ and $\nu_x \rightarrow \nu_x + e^+e^-$ events would appear as isolated narrow e^+e^- pairs:
 - no hadronic activity;
 - single e.m. shower $E_{\text{dep}} > 200 \text{ MeV}$ with vertex within fiducial volume and pointing towards the CNGS beam axis within 150 mrad.
 - Hypothetical background sources investigated through MC simulations for both NC/CC interactions also in surrounding material/rock. Residual after the selection cuts:
 - Coherent π^0 production: 0.14 ± 0.03 ; Quasi-elastic CC ν_e : 0.04 ± 0.01
- Control sample (loose cuts):
- 1 event with a single isolated π^0 detected (1.4 ± 0.05 expected).
 - 1 isolated γ event observed at 60° w.r.t. beam direction (2 ± 0.4 predicted without angular cuts).
- No real events survives the selection: $\sim 7.4 \cdot 10^6$ e^+e^- pairs expected for $\delta \approx 5 \cdot 10^{-5}$ and integrated exposure of $6.7 \cdot 10^{18} \text{ pot}\cdot\text{kt}$