NIC2018

Experimental Challenge to the Cosmological Li Problem

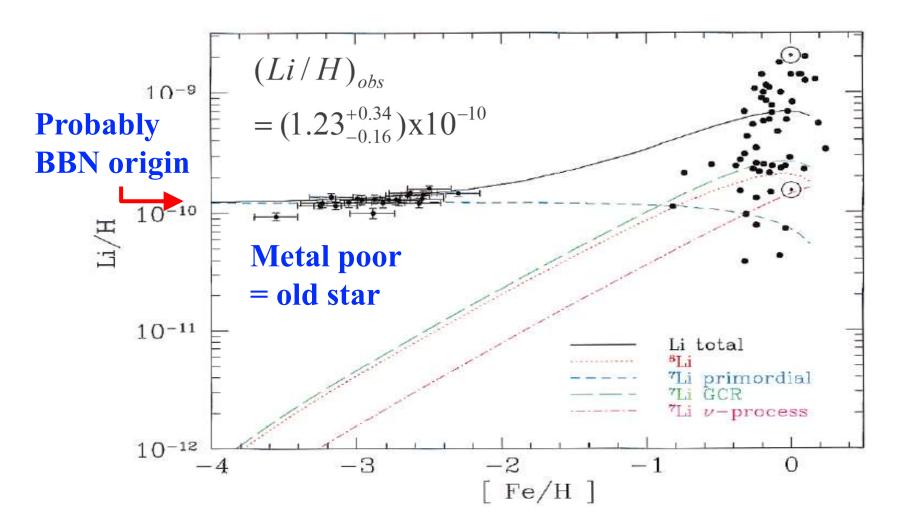
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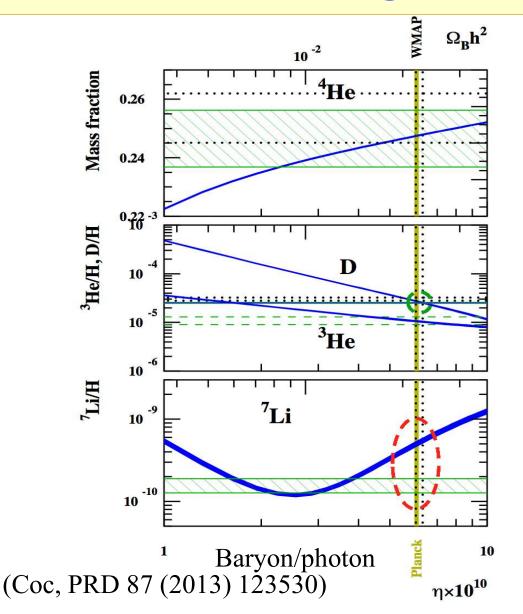
- **1. Cosmological Li Problem in BBN**
- 2. ⁷Be + n destruction channels
- 3. Other destruction channels of ⁷Be
- 4. Summary

Observation of Primordial ⁷Li Abundance

S. G. Ryan *et al*., Astrophys. J. 600 (2000) L57.



Primordial Light Nuclides – Cosmological Li Problem

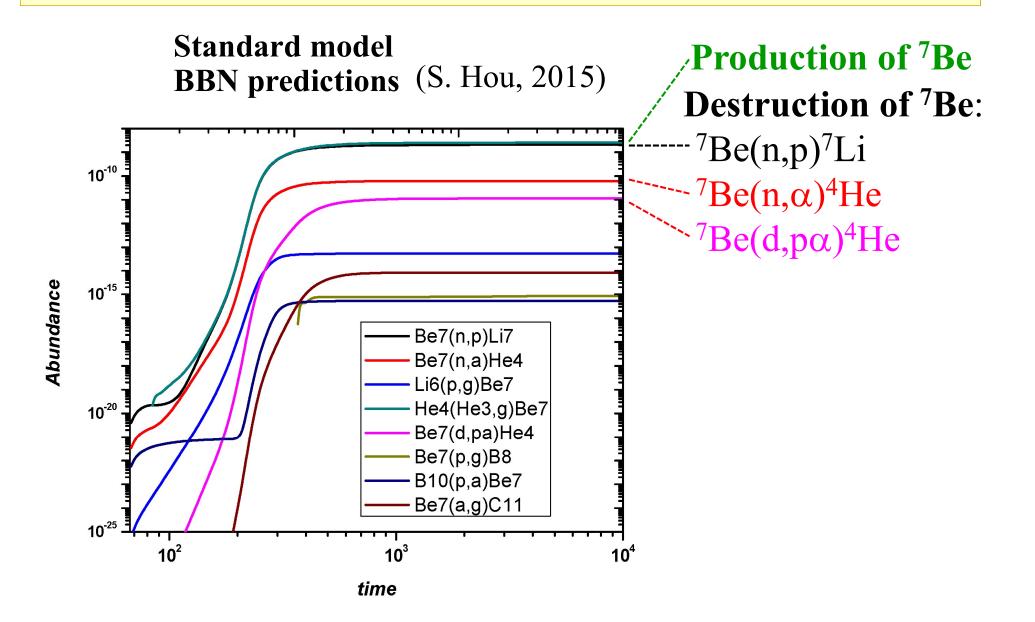


Microwave background Measurement: Plank observation: $\Omega_b h^2 = (2.23 \pm 0.02) \times 10^{-2}$

BBN calculation overestimate by a factor of ~3!

Primordial Nucleosynthesis (BBN) ъ 6 ō (Wagoner, APJ 179, 343, 1973) 14 ⁷Li production N N mainly by EC of ⁷Be ²C after BBN в B * ⁷Li can be easily destroyed by ⁷Li(p,α) Be during BBN time. œ.n α,γ (³He,2p)-³He a,, ⁴He α. D p,n $-n,\gamma$ 4 n.a

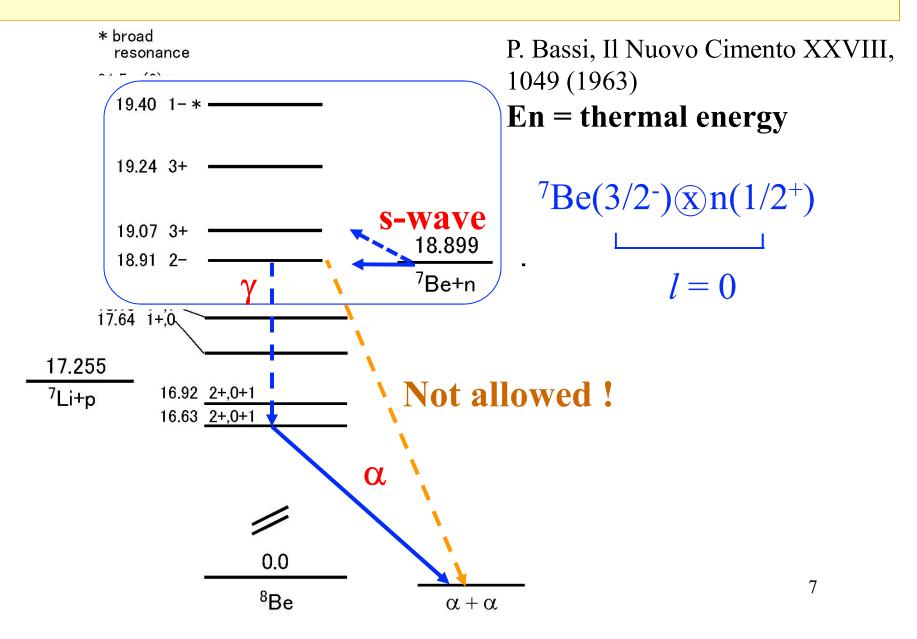
Primordial Light Nuclides -7Li Problem-



1. Fate of ⁷Be by (n,α)

- Most unknown reaction for ⁷Be destruction -

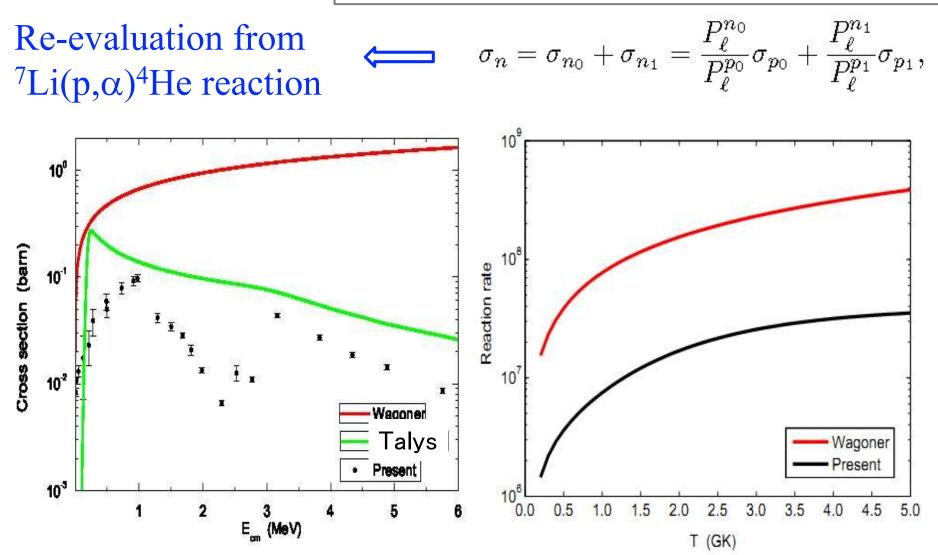
⁷Be(n,α)⁴He Reaction



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New estimate of ⁷Be(n,α)⁴He cross sections

Hou, He, Kubono, and Chen, PRC 91 (2015) 055802



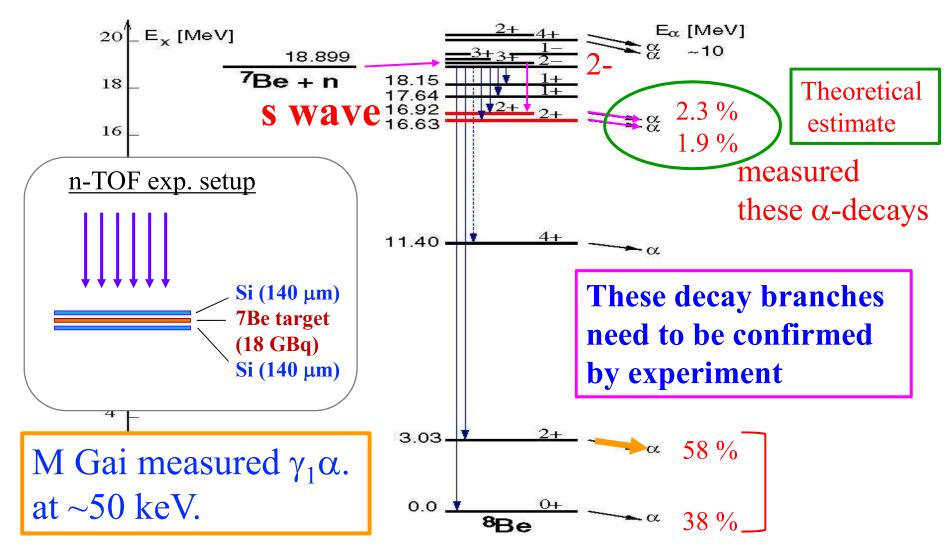
Several experiments made and being made

- 1. RCNP time-rev. method (Kyoto-RIKEN-RCNP-Tohoku)
- 2. ISRAEL direct measurement
- 3. n-TOF direct measurement
- 4. Legnaro Trojan Horse Method
- 5. CRIB/RIBF Trojan Horse Method
- 6. JAEA Branching ratio
- 7. FSU 7Be(d,a)
- 8. . . .

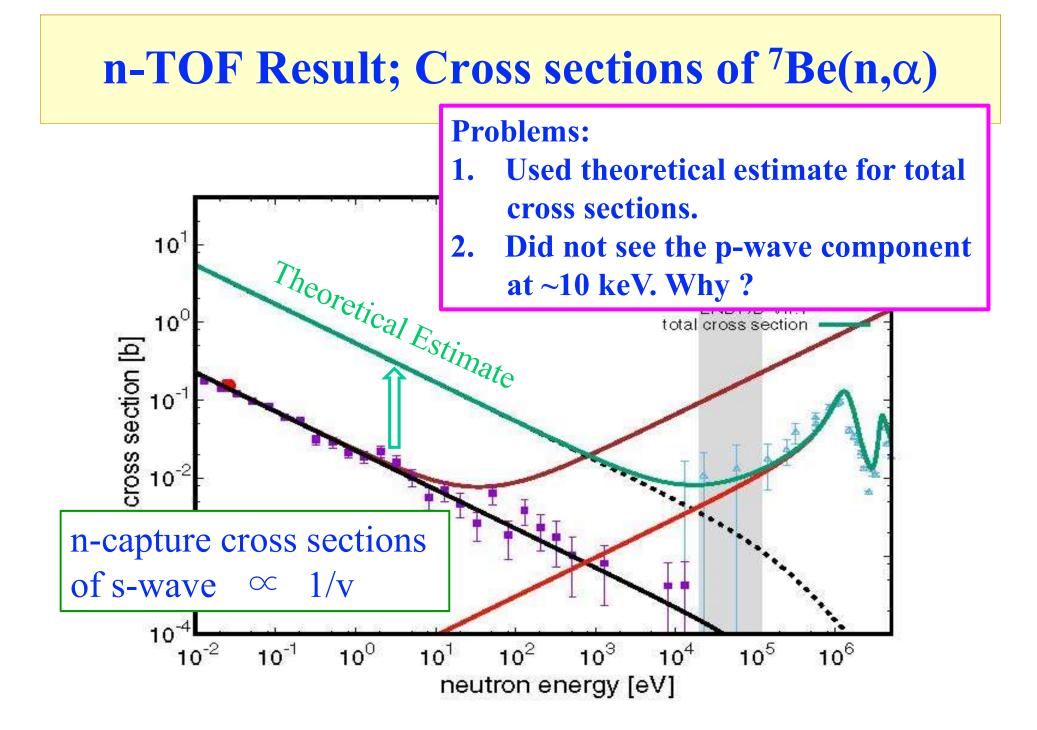
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⁷Be(n,α) experiment at n-TOF

(Barbagallo, PRL 117, 152701, 2016)

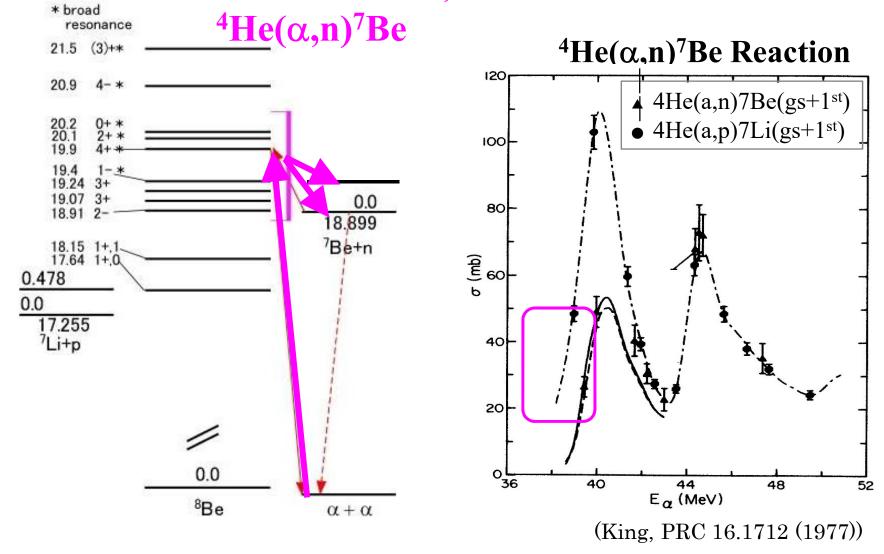


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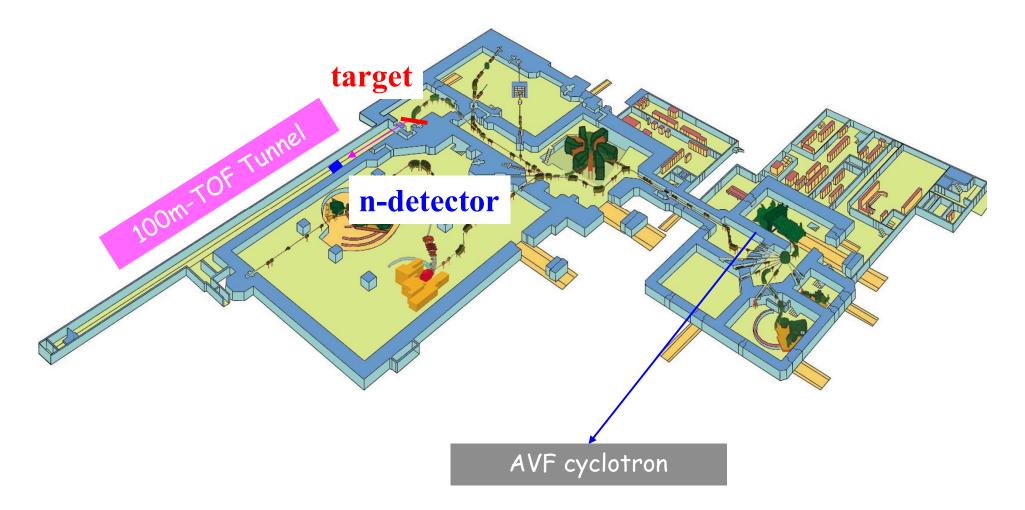
⁷Be(n, α)⁴He Study by ⁴He(α ,n)⁷Be

Use the time-reverse reaction Kawabata, Kubono, PRL 118(2017) 052701

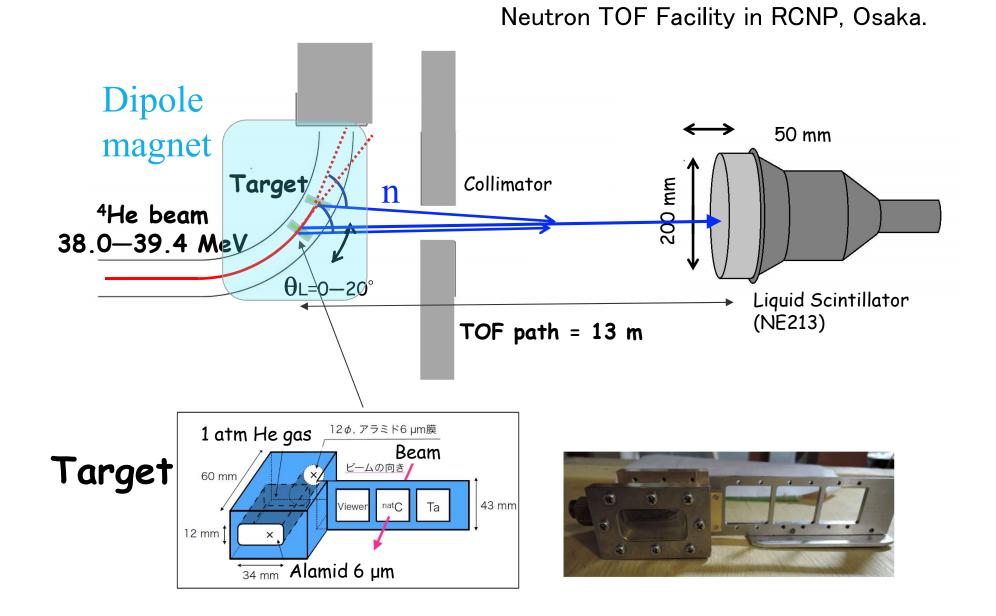


Experimental Facility

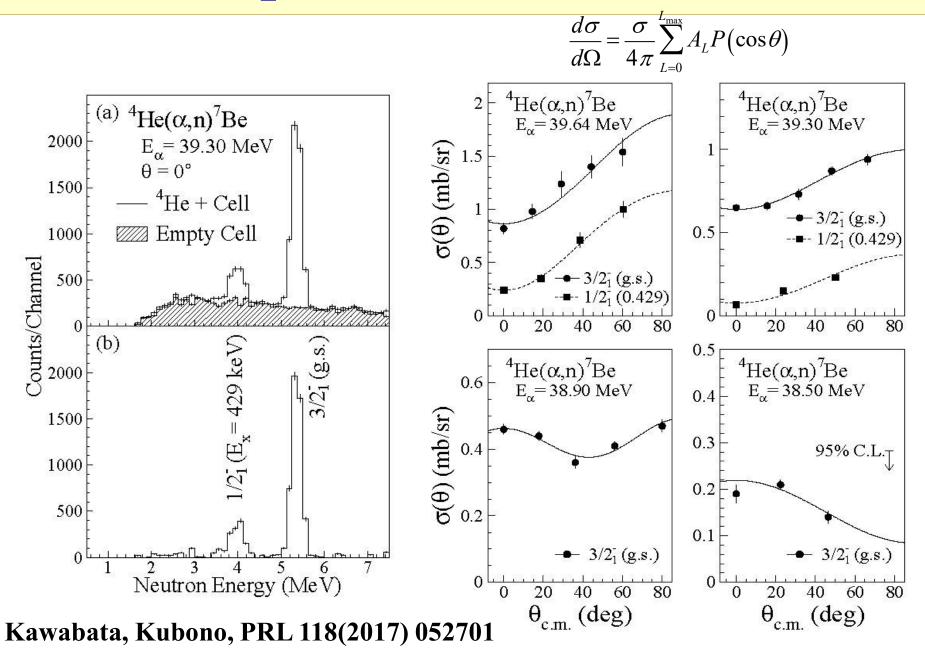
Neutron TOF Facility in RCNP, Osaka.



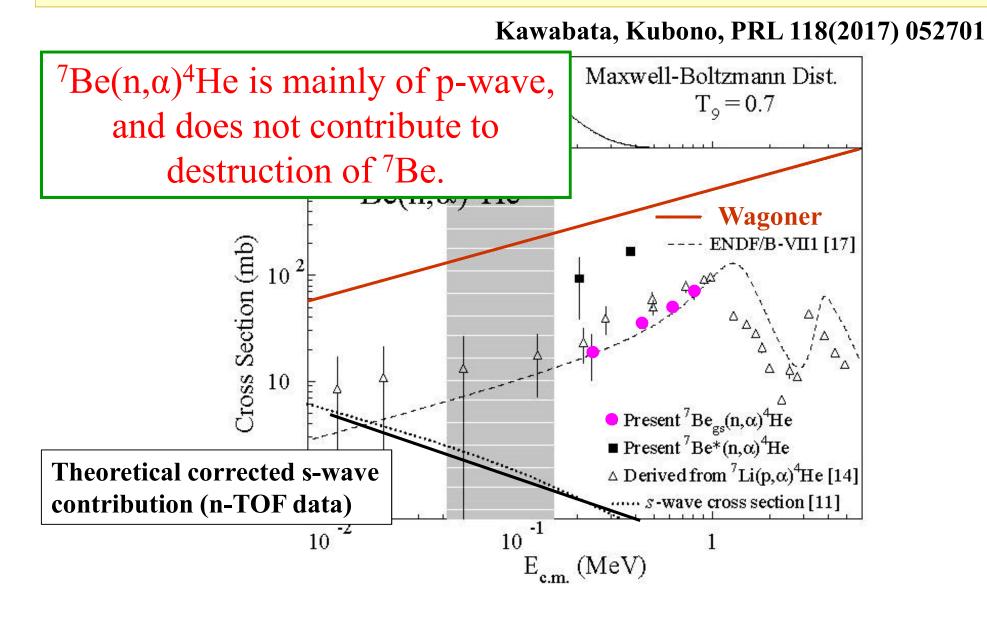
Experimental Setup II



Experimental Results

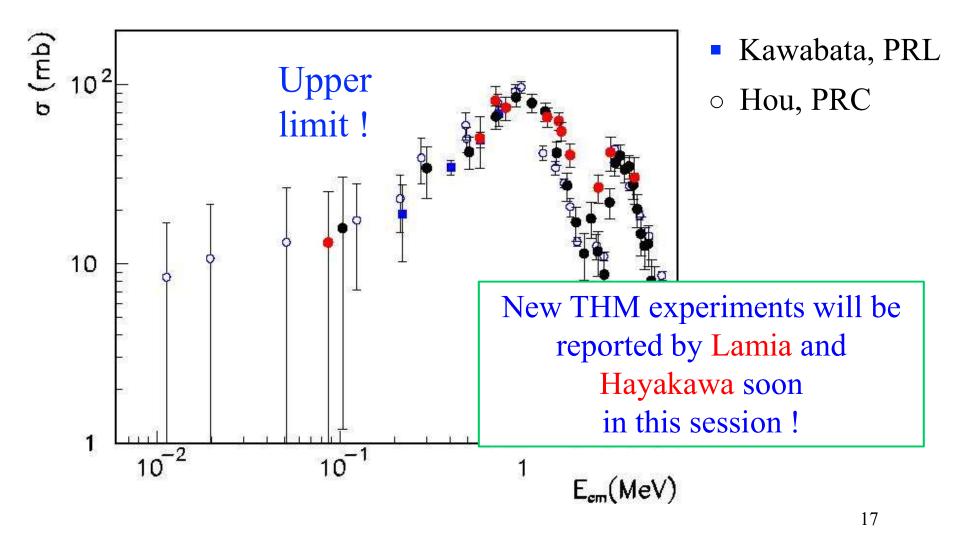


Experimental Results of ⁷**Be(n,α)**



Trojan Horse Method for ⁷Be(n,α)

L. Lamia, APJ 850, 175



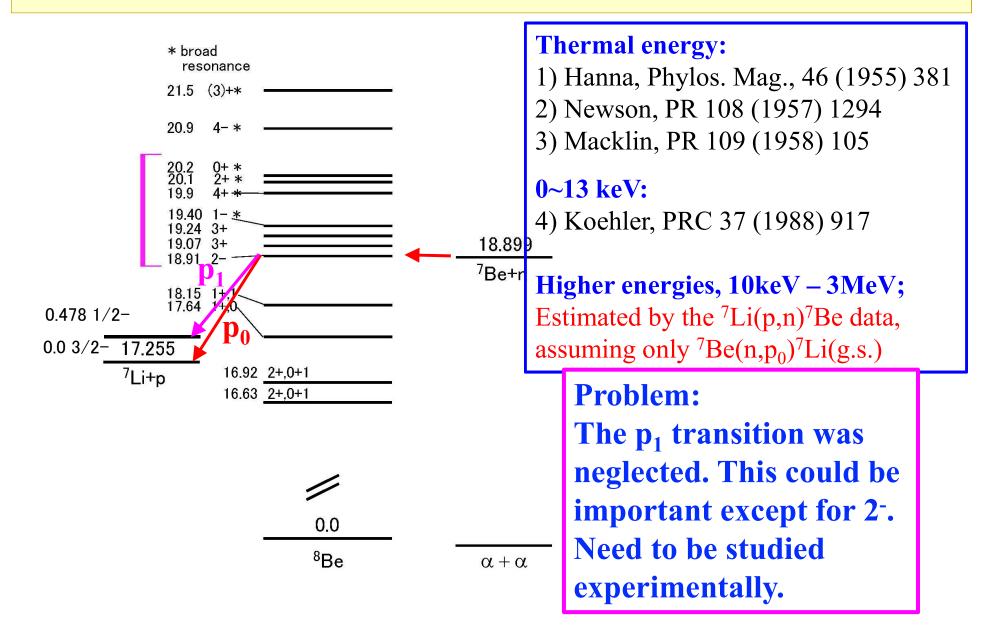
Conclusion for ⁷**Be(n,α)**⁴**He reaction in BBN**

- 1. The p-wave cross sections dominate at the BBN temperatures.
- 2. The total cross sections (s +p) is about one order of magnitude smaller than the Wagoner rate.
- **3.** The ⁷Be(n,α)⁴He reaction does not solve the Li problem in BBN.

2. Fate of ⁷Be by (n,p)

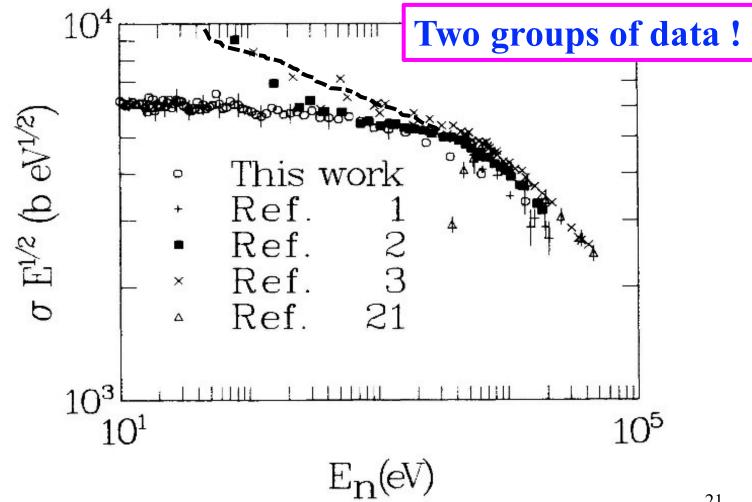
- The major destruction reaction. Any uncertainty, well known ? -

⁷Be(n,p), the major destruction reaction



⁷Be(n,p)Li reaction cross sections

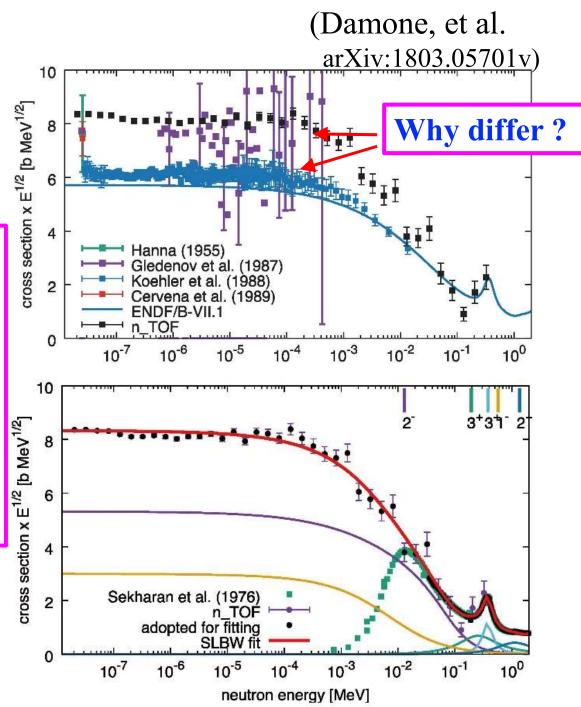
Koehler, PRC 37 (1988) 917



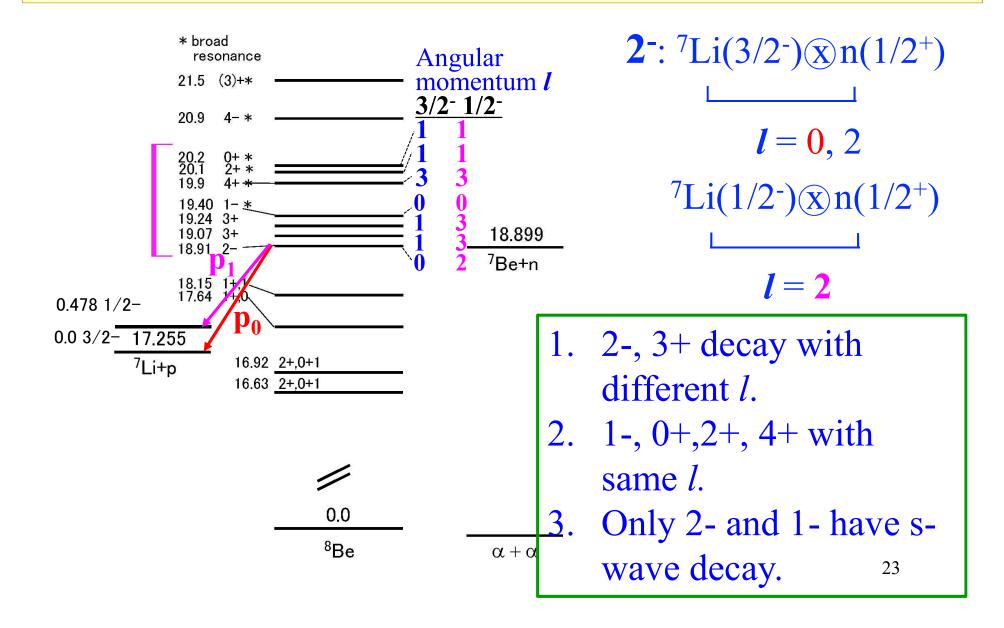
⁷Be(n,p)⁷Li Measured (a)n-TOF

Problem: The total cross sections were obtained assuming s-wave transitions for all. But, they are mostly non <u>s-wave</u>.

-> Need to measure angular distributions



⁷Be(n,p) reaction and the spin selection rule



Measurement of the branching ratio for the ⁷Be(n,p)⁷Li reaction

(Iwasa,SK, ..)

The branching ratios p_1/p_0 were measured for these resonances at JAEA tandem facility, using ENMA spectrometer.

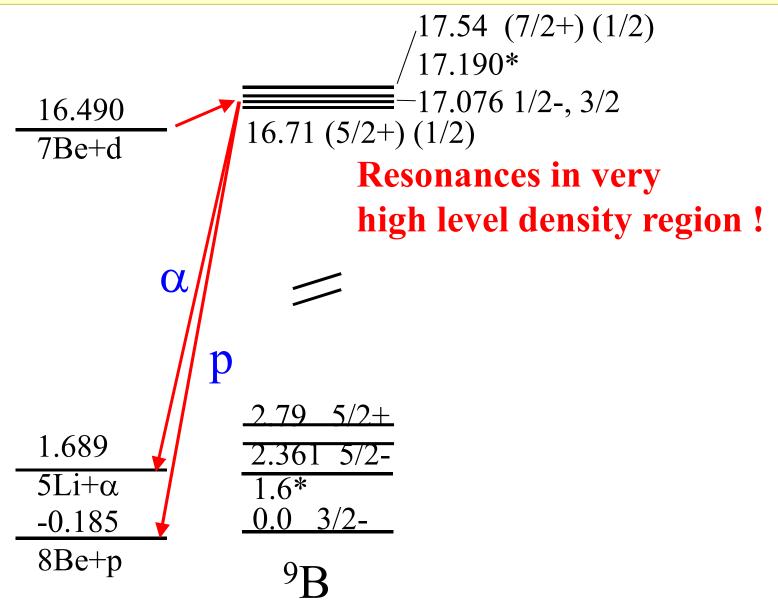
Significant decay branches for p_1 were observed for the resonances except for the lowest 2- resonance.

Similar result was obtained by THM. Waite for Hayakawa talk.

3. Fate of ⁷Be by other channels

- d, 3He, t induced reactions? -

⁷Be(d,p) and (d, α) reactions



⁷Be(d,α)⁵Li Reaction

Ingo Wiedenhoever reported (first day of this symposium):

• The excitation function of ⁷Be(d,p) and ⁷Be(d,alpha) with good sensitivity into the Gamow window of BBN.

• Some new resonances in the (d,alpha) cross sections were discovered, and this reaction is dominating over the (d,p) reaction. This result reduce the Li problem by $\sim 17\%$.

4. Other possibilities for the Li Problem

Other reactions in BBN
; ³He(α,γ), ²H(p,γ),.

Summary on Li Problem in BBN

- 1. $^{7}Be(n,\alpha)$ reaction is now reasonably well studied, and would not contribute much to the ^{7}Be destruction.
- 2. ⁷Be(n,p) reaction needs to be refined better by experiments considering the spin selection.
- 3. ⁷Be+d channels may influence to the ⁷Be destruction.
- 4. $^{7}Be+^{3}He$, t, α channels still are not well investigated yet. They should be studied carefully.
- 1. Further experiments are required for reaction rates.
- 2. Other possibilities of **new unknown physics** for BBN.
- 3. **Further work in observation** is also highly welcome.