





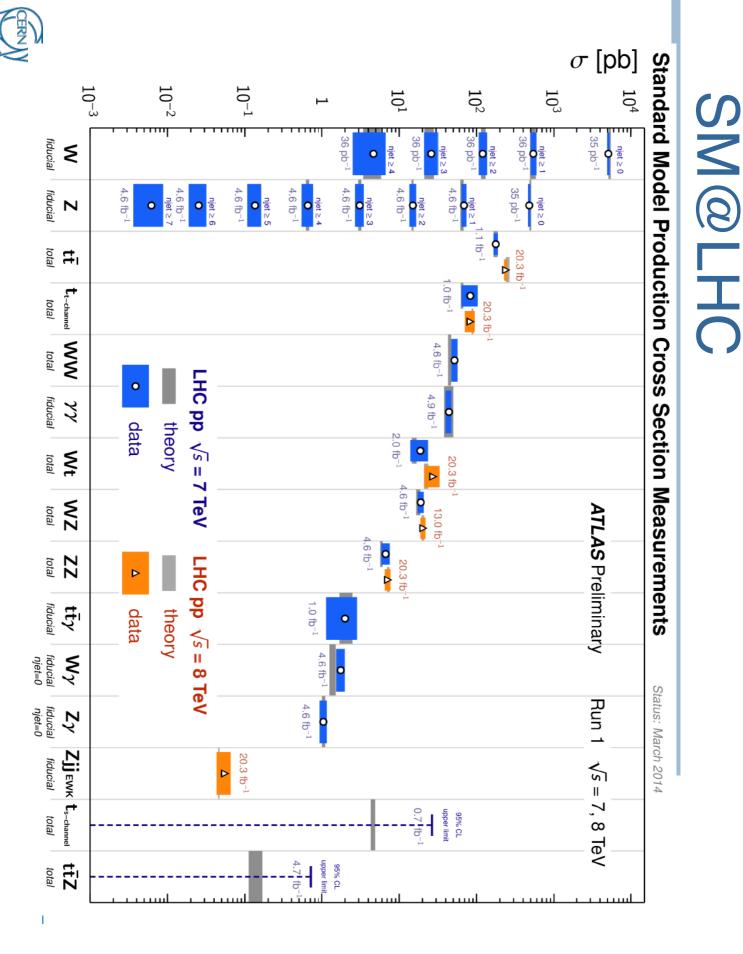


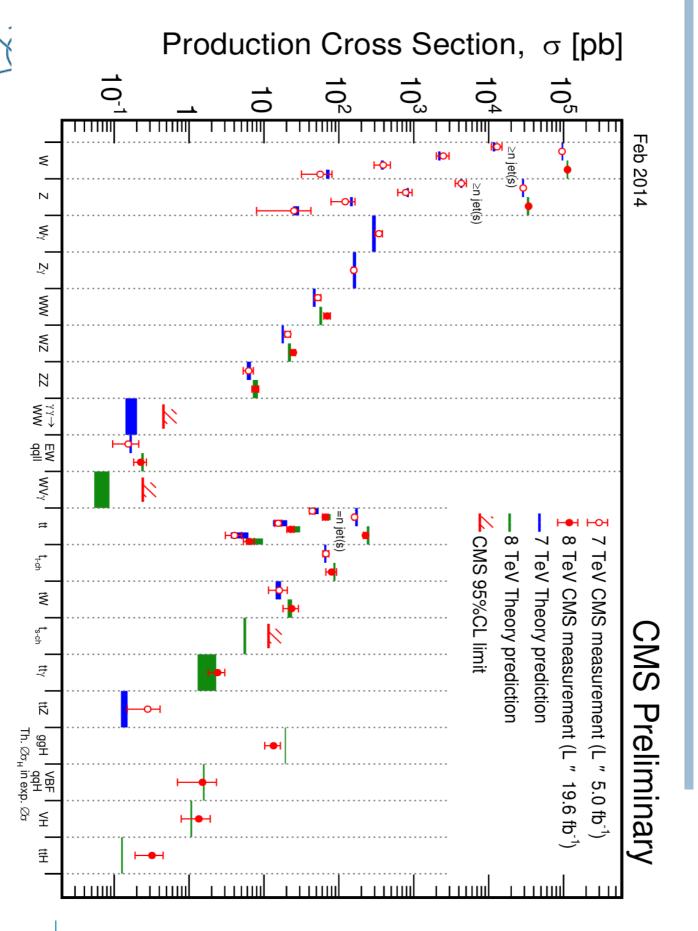
Arcetri, May 17th 2016 Sergio Bertolucci INFN

After LHC Run 1 (2010-2012):

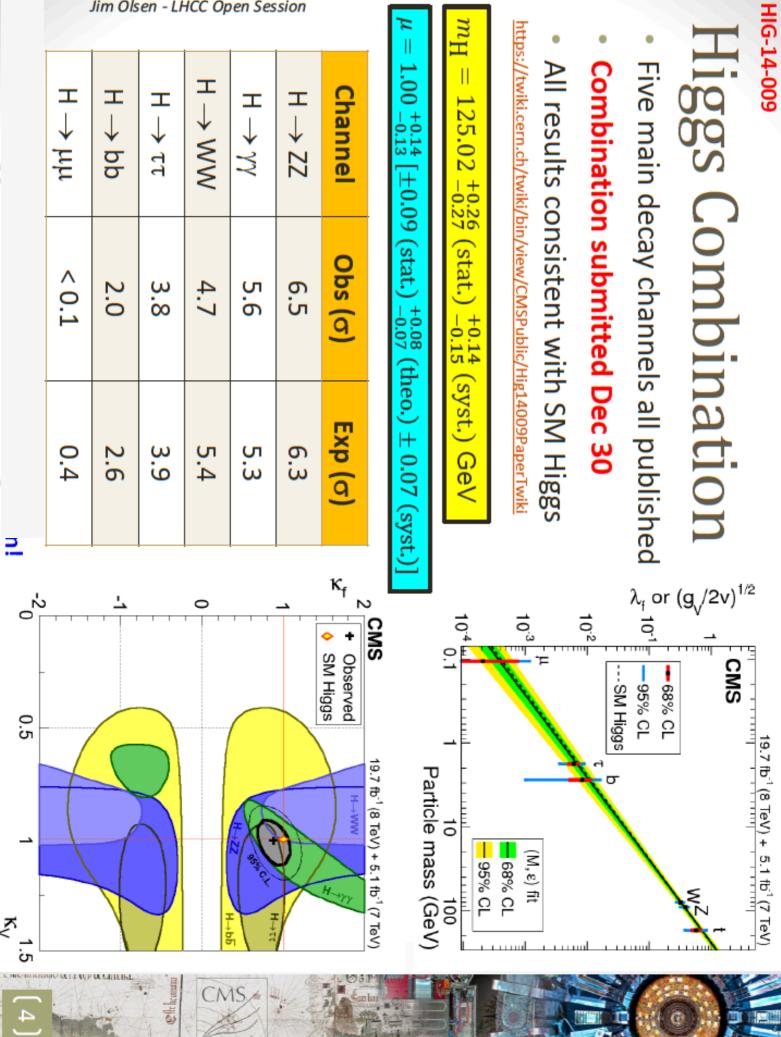
- We have consolidated the Standard Model and very sensitive to New Physics, $B_s \rightarrow \mu\mu$ decay) (a wealth of measurements at 7-8 TeV, including the rare,
- We have completed the Standard Model: discovery of the messenger of the BEH-field, the Higgs boson discovery
- We have found interesting properties of the hot dense matter
- scrutiny although tantalizing hints have survived We have NO evidence of New Physics,





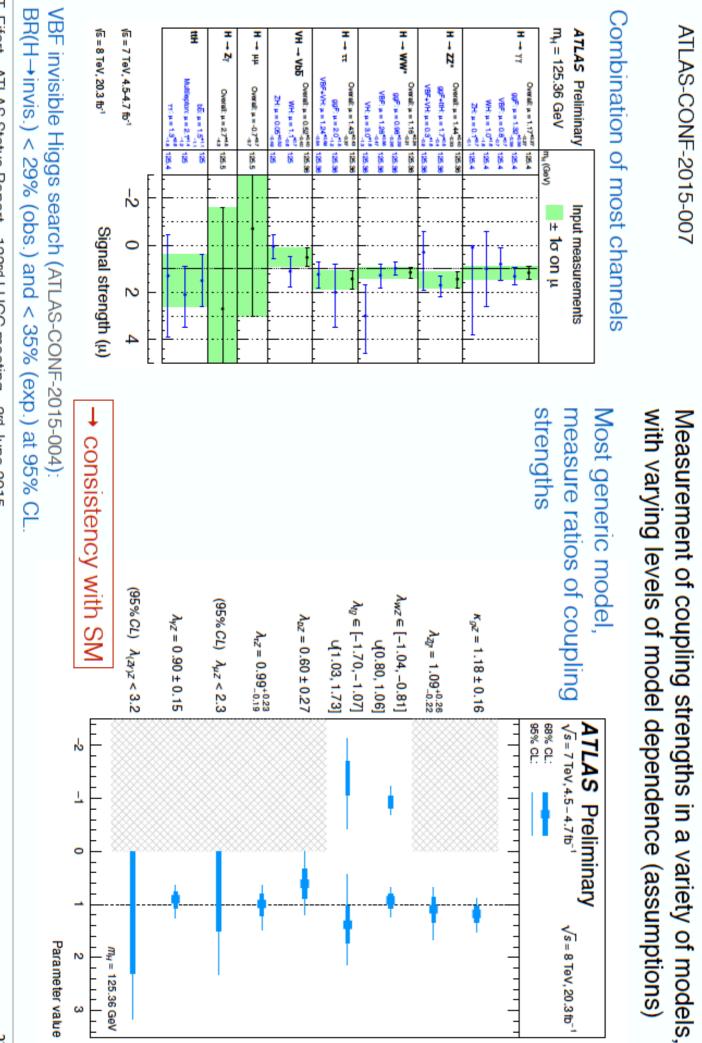


SM@LHC

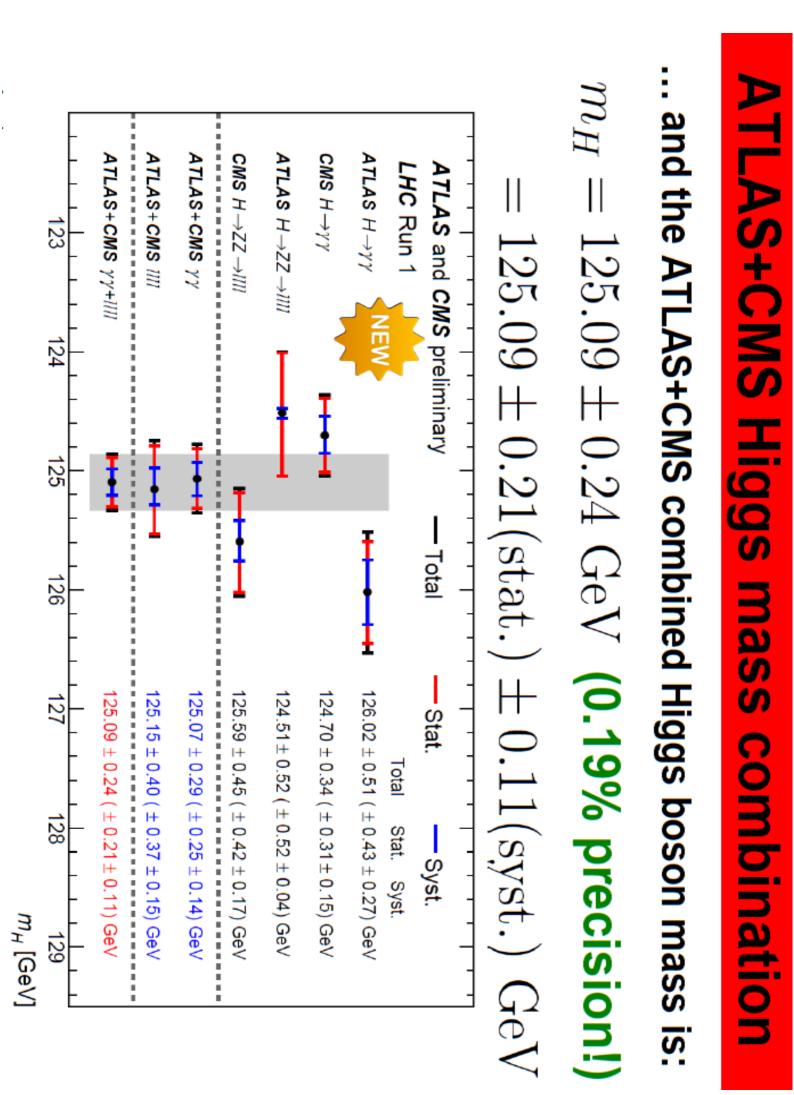


Jim Olsen - LHCC Open Session

T. Eifert - ATLAS Status Report - 122nd LHCC meeting - 3rd June 2015



Higgs production, rates, couplings



ATLAS Status

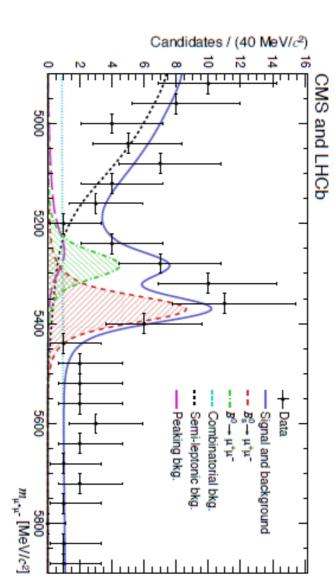
"Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1*rr* theoretical signal cross section uncertainty.

	Mass scale [TeV]	1 dilative autodana dana dana dana dina ta theoretic	10 ⁻¹	- 	full data	tu V	vs = 8 1eV partial data	vs = 7 lev full data	
20 Coentrinoi	m(r)>200 GeV	480 GeV	-	e-112	100			Scalar charm, c→citi	Other
1011 01205	mill ⁰ L-montheast			32	Koo	3	-	Carlos abases 5	
1404.250		850 GeV	*	20.3	1 See	4.8-0	2 e. µ (SS)	$\tilde{s} \rightarrow \tilde{t}_1 s, \tilde{t}_1 \rightarrow bs$	
ATLAS-CONF-2013-091	8P(r)=8P(r)=8P(r)=0%	916 GeW	241	20.3		6-7 jets		2-000 2-101100 - 0001001 - 000000	
1405.5088	medication of the second states of the second state	450 G#V	<u>.</u>	20.3	ti a		ω	Pty: St_NPP Pt_res are	R
1405.5085	$m(\tilde{x}^0) > 0.2 \times m(\tilde{x}^0)$. $J \rightarrow \infty$	750 GeV		20.3				$E^+X^-, X^+ \rightarrow WE^0, X^0 \rightarrow ee\bar{e} \rightarrow ee\bar{e}$	PV
1404.2500	m(d)=m(d), etasa<1 mm	1.35 TeV	d, to	2013	_	0-3 /	2 e. u (SS)	Bilmagr RPV CMSSM	′
1919 1979	J	1 1 TeV	8- 3	4.0			10.0 + 7	LFV $pp \rightarrow p_1 + X, p_1 \rightarrow e(u) + r$	
10101010			D				0	EV market V & Lot La	
ATLAS-CONF-2013-082	1.5 <cr<156 br(µ ="1," m(t<sup="" mm,="">2)=108 GeV</cr<156>	1.0 TeV	ā	20.3		×	Ŀ.	q̃q̃, k_1^0 →qqµ (RPV)	
1409.5542	2 <r(2)<3 model<="" ns,="" sps8="" th=""><th>435 GeV</th><th>λų.</th><th>20.3</th><th>á</th><th></th><th></th><th>GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_1^0$</th><th>Lo pi</th></r(2)<3>	435 GeV	λų.	20.3	á			GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_1^0$	Lo pi
1411.6795	10 <tan 60<="" th=""><th>537 GeV</th><th>×.</th><th>19.1</th><th></th><th></th><th>(e,µ) 1-2µ</th><th>GMSB, stable $\overline{\tau}, \overline{X}_1^0 \rightarrow \overline{\tau}(\overline{e}, \overline{\mu}) + \tau(e, \mu)$</th><th>ng arti</th></tan>	537 GeV	×.	19.1			(e,µ) 1-2µ	GMSB, stable $\overline{\tau}, \overline{X}_1^0 \rightarrow \overline{\tau}(\overline{e}, \overline{\mu}) + \tau(e, \mu)$	ng arti
1411.6795		1.27 TeV	28.	19.1	,	,	trk	Stable ji R-hadron	-liv icle
1310.6584	m(t ²)=100 GeV, 10 µs <r(j)<1000 s<="" th=""><th>832 GeV</th><th>ati</th><th>27.9</th><th>s Yes</th><th>1-5 jets</th><th>0</th><th>Stable, stopped § R-hadron</th><th>90 75</th></r(j)<1000>	832 GeV	ati	27.9	s Yes	1-5 jets	0	Stable, stopped § R-hadron	90 75
1310.3675	$m(\tilde{x}_1^2) - m(\tilde{x}_1^2) = 100 \text{ MeV}, \pi(\tilde{x}_1^2) = 0.2 \text{ ns}$	~	X ⁺ 270 GeV	20.3	Ŕ	1 jet	Disapp. trik	Direct $\hat{\mathcal{X}}_1^* \hat{\mathcal{X}}_1^-$ prod., long-lived $\hat{\mathcal{X}}_1^*$	í
	(i) A straight of a straight of the straigh	100.000	ê		100			A3A3, A23 - 1686	
	man and a second s			20.2		0 1		$F_{0}F_{0} = F_{0}$	
1501.07110	michaente michaente destructure		3 ² . 3 ² 250 GeV	20.3		0-2 5		222 WDB 22 22 22 22 22 22 22 22 22 22 22 22 22	
1403.5294, 1402.7029	m071 am073, m071 and steptons decoupled			20.3		0-2 jets	2-3 c. µ	$\overline{F}^{\pm}\overline{F}^{0} \rightarrow W\overline{F}^{0}Z\overline{F}^{0}$	di
1402.7029	mk? \=mk?), mk?)=0, mk?, H=0.5(mk?)+mk?))	700 GeV		20.3	đ,	0	0 n.1	$\overline{X}^{\pm} \overline{X}^{\pm} \rightarrow \overline{h} v h f(\overline{v} v), f \overline{v} h f(\overline{v} v)$	EW rec
1407.0350	$m(2^0) \rightarrow 0$ GeV $m(2, 2) \rightarrow 0$ S($m(2^0) \rightarrow m(2^0)$)	100-350 GeV		20.3	Ň		NJ I	FTF Ft + TW(+0)	
1400.5294	m(2)=0 GaV, m(2, 9)=0.5(m(2))+m(2))	140-465 GeV	1	20.3	đ i	0	20.1	$\tilde{X}^{\dagger}\tilde{X}^{\dagger}$, $\tilde{X}^{\dagger} \rightarrow \tilde{U}U(\tilde{Y})$	
1403.5294	m(3 ⁰)=0 GeV	90-325 GeV	06 2	20.3	Yes	0	20,14	$\tilde{L}_1 = \tilde{L}_1 = . \tilde{L} \rightarrow \ell \tilde{X}_1^0$	
1403.5222	m(t ²) <200 GeV	290-600 GeV	12	20.3	Yes	1.0	3 e, µ (Z)	$i_2 i_3, i_2 \rightarrow i_1 + Z$	
1400.5222	m(R ²)>150 GeV	150-580 GeV	7	20.3	(je	1 b	$2 e_{\tau} \mu(Z)$	7(f) (natural GMSB)	
1407.0608	m(i) - m(ii) <85 GeV		7, 90-240 GeV	20.3	-tag Yes	mono-jet/c-tag Yes	0	$\tilde{i}_1 \tilde{i}_1, \tilde{i}_1 \rightarrow c \tilde{k}_1^{-1}$	ect
1407.0583,1406.1122	m(R ₁)=1 GeV	210-640 GeV	ñ.	28	Ť.	1-2 6	0-1 e.,µ	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c \tilde{k}_1^T$	t pi
1403.4853, 1412.4742	m(t ²)=1 GeV	215-530 GeV	71 90-191 GeV	20.3		0-2 jets	20,1	$\tilde{i}_1 \tilde{i}_1, \tilde{i}_1 \rightarrow Wb\tilde{\mathcal{K}}_1^{\prime}$ or $d\tilde{\mathcal{K}}_1^{\prime}$	so
1209.2102, 1407.0583	$m(\tilde{k}_{1}^{*}) = 2m(\tilde{k}_{1}^{*}), m(\tilde{k}_{1}^{*})=55 \text{ GeV}$	230-460 GeV	7, 110-167 GeV	4.7	í,	1-2 5	1-2 e.,u	$\tilde{i}_1 \tilde{i}_1, \tilde{i}_1 \rightarrow b \tilde{k}_1^{-}$	qua luc
1404.2500	m(k ²)-2 m(k ²)	275-440 GeV	5,	20.3	Yes	0-3 /2	2 e, µ (SS)	$b_1b_1, b_1 \rightarrow \delta \tilde{k}_1$	ark tio
1308.2631	m(t ⁰)=90 GeV	100-620 GeV	ð,	20.1) Tes	2.6	0	$b_1b_1, b_1 \rightarrow bk_1^0$	s n
	. and a set of the set		•					0	
1407.0600	m(2)=300 GeV	1.3 TeV	241 1	20	đ i	35	0-1 c. x	2	31 8
1407.0600	m0 ²⁰ k400 GeV	1.34 TeV	240 3	20.1		а Э	0-1 e.u		f g ma
1308.1841	m[2]1 - 350 GaV	1.1 TeV	Ny 2	20.3	n Fi	7-10 jets		$\frac{2}{N-2}$	en. ed.
1407 0500	mil ²⁰ Loann /Savi	.		3	s	4	-	0.5247 =	
1502.01518	m(G)>1.8 × 10 ⁻⁴ eV, m(g)-m(g)=1.5 TeV	865 GeV	F ^{U2} scale	20.3	99. 1688	mono-jet	0	Gravitino LSP	
ATLAS-CONF-2012-152	m(NLSP)>200 GeV	690 GeV	341	5.8	s Mes	0-3 jets	$2 e, \mu(Z)$	GGM (higgsino NLSP)	
1211.1167	m(R ⁰)>220 GeV	900 GeW	24	4.8	1	1.6	Ŷ	GGM (higgsino-bino NLSP)	In
ATLAS-CONF-2012-144	m(2)>50 GeV	619 GeV	941	4.8	đ	,	$1 e, \mu + \gamma$	GGM (wine NLSP)	ciu
ATLAS-CONF-2014-001	m(t ⁰)>50 GeV	1.28 TeV	24	20.3	1és	,	27	GGM (bino NLSP)	isi
1407.0603	TeV tanβ≥20	1.6 TeV	991	20.3	s Yes		1-2 7 + 0-1 (GMSB (Z NLSP)	ve
1501.03555	m(t ⁰)=0 GeV	1.32 TeV	24	20	,	0-3 jets	20,4	$\overline{g}\overline{g}, \overline{g} \rightarrow gg(\ell\ell/(\nu/\nu\nu)\overline{k})$	Se
1501.03555	$m(\tilde{x}_1^0 <300 \text{ GeV}, m(\tilde{x}_1^1)=0.5(m \tilde{x}_1^0 +m(\tilde{x}_1^0))$	1.2 TeV	эh	20	s Mes	3-6 jots	10,4	$\tilde{X}\tilde{X}, \tilde{X} \rightarrow qq\tilde{X}_{1}^{+} \rightarrow qqW^{\pm}\tilde{X}_{1}^{0}$	
1405.7875	m(R ₁)=0 GeV	1.33 TeV	24	20.3	8 168	2-6 jets	0	$\overline{x}\overline{x}, \overline{x} \rightarrow 0\overline{0}\overline{x}_{1}^{(i)}$	
1411.1559	$m(\hat{q}) - m(\hat{e}_{1}^{0}) = m(c)$		q	20.3		0-1 jet	1γ	$\tilde{q}\tilde{q}\gamma, \tilde{q} \rightarrow q\tilde{t}_{1}^{0}$ (compressed)	
1405.7875	m(R ₁)=0 GeV, m(1 st gen. 4)=m(2 nd gen. 4)	850 GeV	ð	20.3		2-6 jets	0	$\bar{q}\bar{q}, \bar{q} \rightarrow q\bar{q}^0$	
1405,7875	1.7 TeV m(\$)-m(\$)	C1	4.2	20.3	s Yes	2-6 jets	0	MSUGRACMSSM	
	-				-				1
Reference		Mass limit		[<i>L dr</i> (m ⁻¹]	E_{T}^{miss}	Jets	e, μ, τ, γ	Model	
$\sqrt{s} = 7, 8 \text{ TeV}$								Status: Feb 2015	Sta
					2		carcie	-	2
C Preliminary	ATI AS		Saarchaet - 05% CI I owar I imite	2	70	*	archa	ATI AS SIISV S	2

Run-1 SUSY program completing https://twiki.cern.ch/twiki/bin/view/AtlasPublic/SupersymmetryPublicResults



projection of invariant mass in most sensitive bins



6.2 σ for the $B^0_{s} \rightarrow \mu^+ \mu^-$ (Expected SM 7.6 σ) ★ First observation

 $\mathcal{B}(B_s^0 \to \mu^+ \mu^-) = 2.8^{+0.7}_{-0.6} \times 10^{-9}$

Result demonstrates power of combing data from >1 experiment (an LHC first!) Fit to full run I data sets of both experiments, sharing parameters

CMS and LHCb $B^0_{s,d} \rightarrow$

 $* \mu \mu$ combination

>

Where we stand

- We have exhausted the number of "known unknowns" within the current paradigm.
- Although the SM enjoys an enviable state of observations. supported in most cases by experimental cannot explain several outstanding questions, health, we know it is incomplete, because it



Looking for "unknown unknowns"

Needs a synergic use of:

- High-Energy colliders
- neutrino experiments (solar, short/long baseline, reactors, 0vββ decays),
- cosmic surveys (CMB, Supernovae, BAO, Dark E)
- gravitational waves
- dark matter direct and indirect detection
- precision measurements of rare decays and phenomena
- dedicated searches (WIMPS, axions, dark-sector particles)



From the Update of the European Strategy for Particle Physics

CERN Member States and of the national institutes, laboratories and universities closely collaborating with the European organizational model for particle physics founded on the sustained long-term commitment of the CERN The success of the LHC is proof of the effectiveness of

Europe should preserve this model in order to keep its leading role, sustaining the success of particle physics and the benefits it brings to the wider society.

The scale of the facilities required by particle physics is **resulting in the globalization of the field**. The European Strategy takes into account the worldwide fields and should continue to do so. particle physics landscape and developments in related



Particle physics is global.

The United States and major players in other regions can together address the full breadth of the field's most priority facilities hosted elsewhere. world-class facility at home and partners in highurgent scientific questions if each hosts a unique

as an example of a successful large international with the Large Hadron Collider (LHC) at CERN serving science project. Strong foundations of international cooperation exist,

intensely competitive field. this perspective is finding worldwide resonance in an cooperation is an important theme of this report, and of international projects. Building further international Reliable partnerships are essential for the success



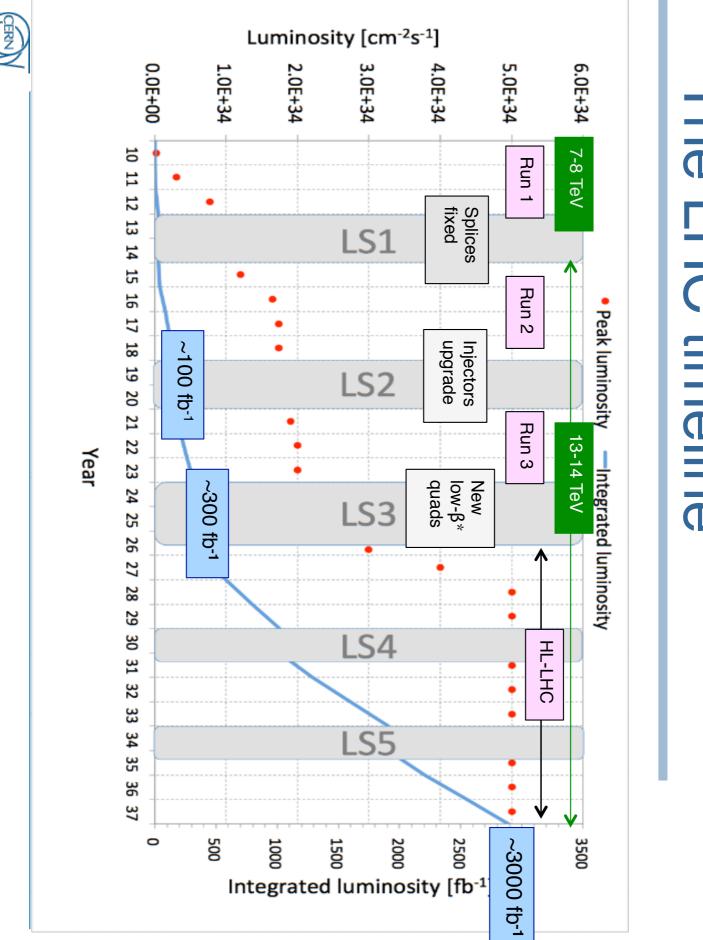
From Japan HEP Community

projects, which comprise the core of future high energy physics research in Japan. The committee makes the following recommendations concerning large-scale

higher, accelerator R&D should be strengthened in order to realize the necessary collision energy. the upgraded LHC version. Should the energy scale of new particles/physics be In parallel, continuous studies on new physics should be pursued for both LHC and experiments at low collision energy should be started at the earliest possible time Should a new particle such as a Higgs boson with a mass below approximately realization of an e+e- linear collider. In particular, if the particle is 1 TeV be confirmed at LHC, Japan should take the leadership role in an early light,

studies on CP symmetry through neutrino oscillations accompanied by the necessary reinforcement of accelerator intensity, so allowing search for proton decays, which would be direct evidence of Grand Unified Should the neutrino mixing angle θ_{13} be confirmed as large, Japan should aim to realize a large-scale neutrino detector through international cooperation, This new large-scale neutrino detector should have sufficient sensitivity to allow the

I heories.



The LHC timeline

The question

- Is the mass scale beyond the LHC reach ?
- Is the mass scale within LHC's reach, but final states are elusive ?

through: We should be prepared to exploit both scenarios,

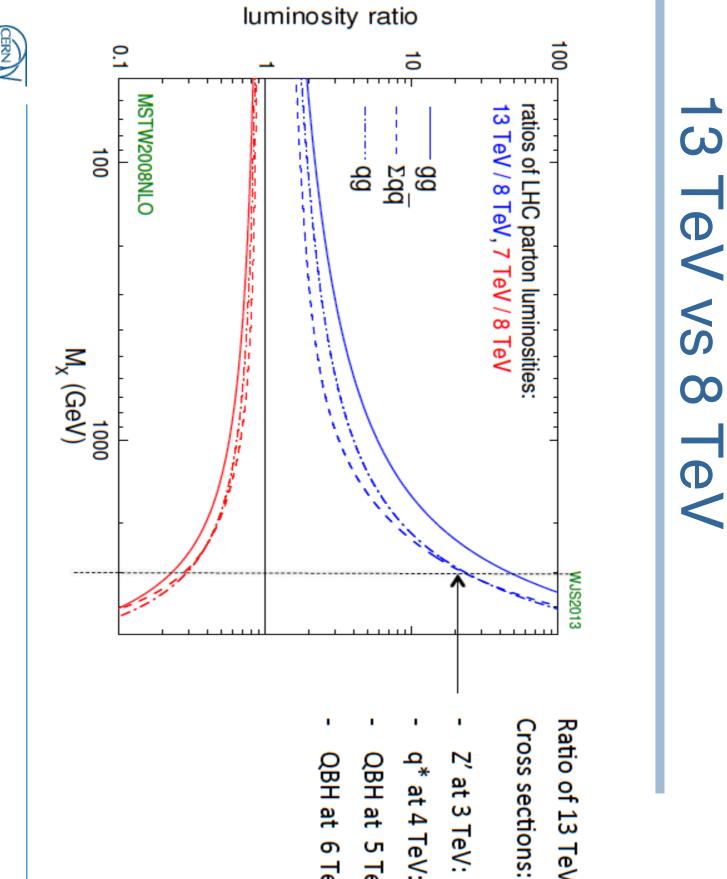
- Precision
- Sensitivity (to elusive signatures)
- Extended energy/mass reach



Extending the reach...

- Weak boson scattering
- Higgs properties
- Supersymmetry searches and measurements
- Exotics
- t properties
- Rare decays
- CPV
- ..etc





Ratio of 13 TeV / 8 TeV

- Z' at 3 TeV: 20
- q* at 4 TeV: 56
- QBH at 5 TeV: 370
- QBH at 6 TeV: 9000



1695 Openings and

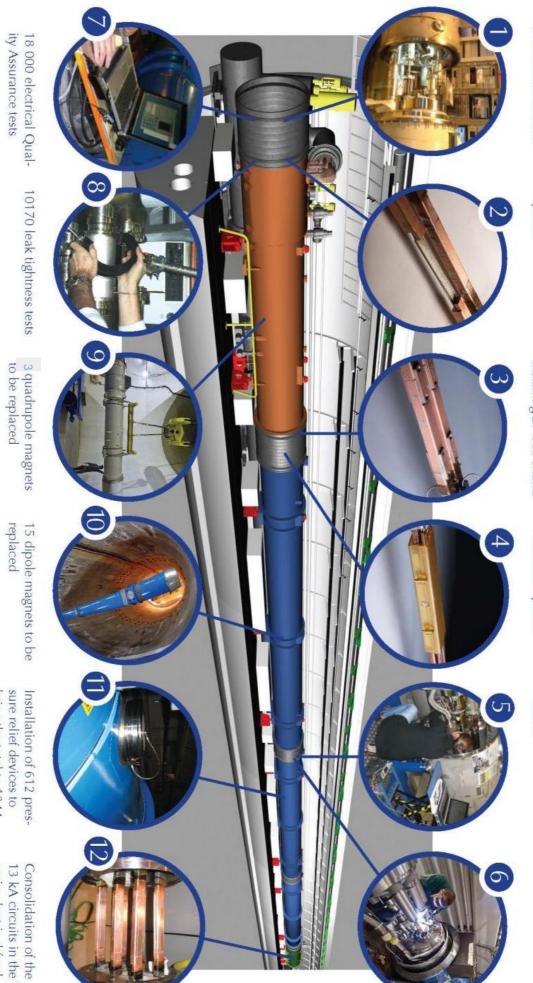
The main 2013-14 LHC consolidations



300 000 electrical ments resistance measure-

Installation of 5000

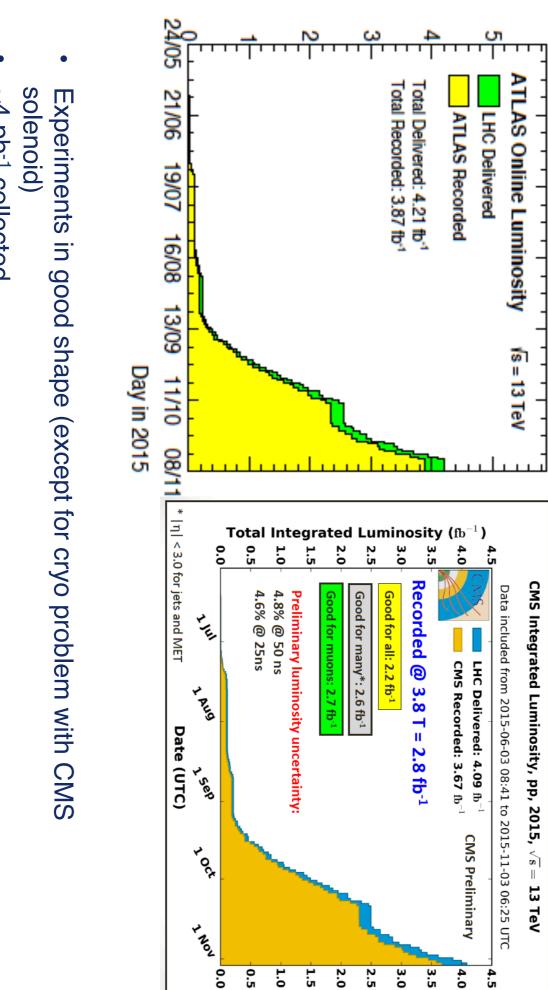
of stainless steel lines 10170 orbital welding





bring the total to 1344

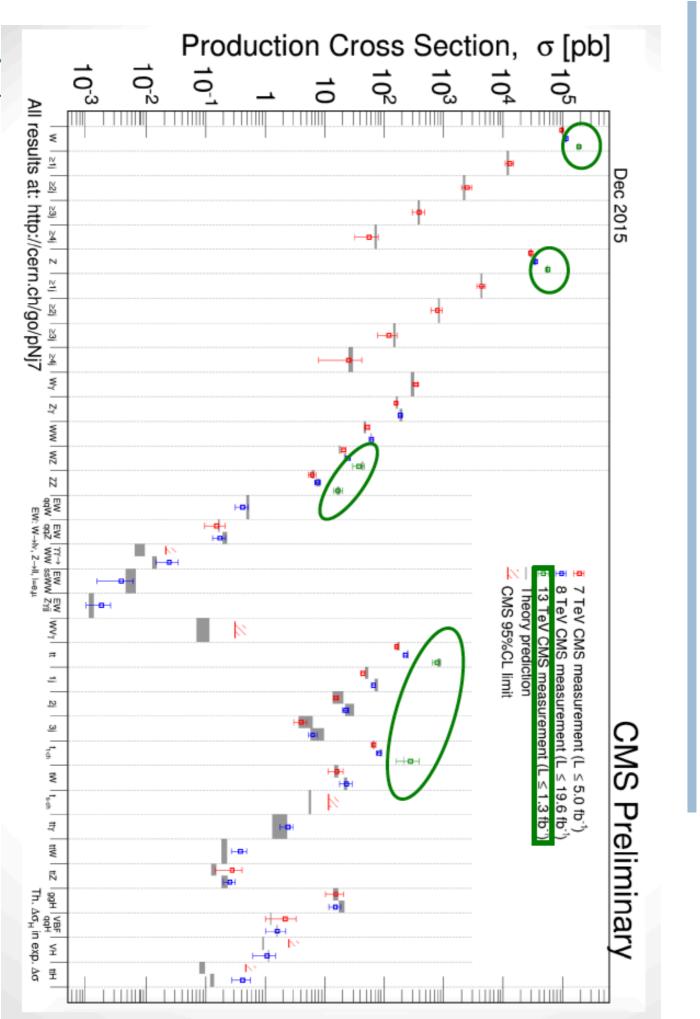
main electrical feed-13 kA circuits in the 16 boxes



~4 pb⁻¹ collected

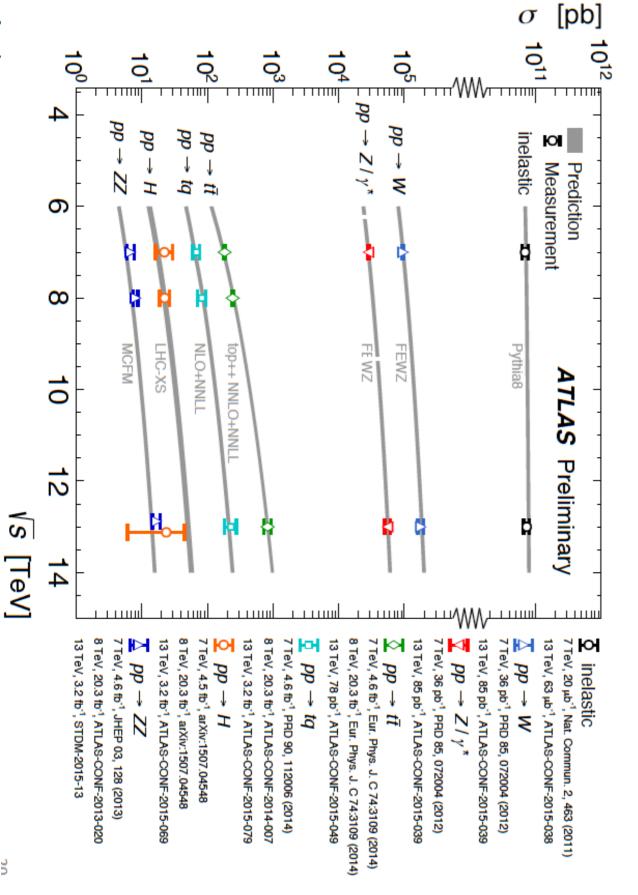
Run2 @13 Tev in 2015

Total Integrated Luminosity [fb⁻]



SM Physics at 13 TeV

SM Physics at 13 TeV



The tantalizing diphotons: ATLAS

Search for a Two Photons Resonance (I)

ATLAS-CONF-2015-081

Inclusive search for two photon resonance

(optimized for a scalar resonance)

- Selection of two photons with pT/m thresholds of 0.3 and 0.4 and pT dependent calorimeter and track isolation criteria
- Typical prompt photon purity 90%

Background from a functional

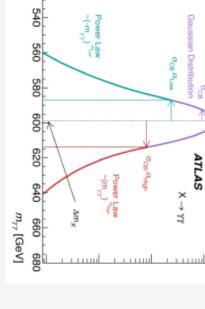
Similar to the dijet search but chosen using the Fisher F-test and the spurious signal method measured in events from Sherpa, Diphox and Jetphox:

$$f_{bkg}(x; b, \{a_k\}) = (1 - x^{1/3})^b x^{\sum_{j=0}^{k} a_j \log(x)^j}$$

$$\frac{n_{\gamma\gamma}}{\sqrt{s}}$$
 Here a simple form \sqrt{s} with k=0 is used

א ≡ -

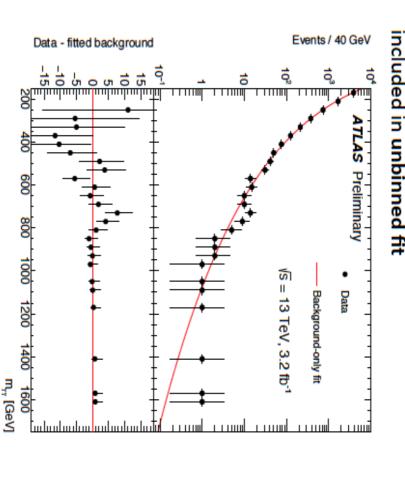
Signal Model NWA: Use Double Sided Crystal Ball 25% of the resonance mass samples with different widths with up to function LW: Use DSCB fitted from simulated Arbitrary Units ನ್ನ ನ್ನೆ 10 520 540 560 "(-m_{pp})").....



The tantalizing diphotons: ATLAS

Search for a Two Photons Resonance (II)

Results: Events with mass in excess of 200 GeV are



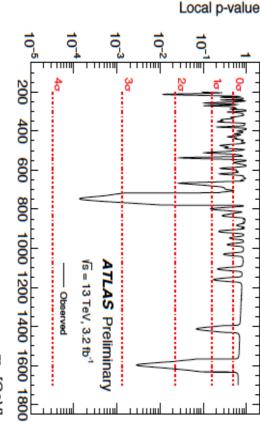
unblinding) of 200 GeV to 2.0 TeV the global Taking a LEE in a mass range (fixed before

significance of the excess is 2.00

750 GeV

3.9o

observed at a mass hypothesis of minimal p₀ of In the NWA search, an excess of 3.6σ (local) is



yielding a width of approximately 45 GeV The data was then fit under a LW hypothesis

(Approx. 6% of the best fit mass of approximately 750 GeV)

As expected the local significance increases to

profiled in the NWA fit and is pulled by 1.5σ

In the NWA fit the resolution uncertainty is

m_x [GeV]

upper range in resolution fixed after unblinding) up to 10% of the mass hypothesis of 2.3σ (Note: Taking into account a LEE in mass and width of

..and CMS

Search for diphoton resonances EXO-15-004

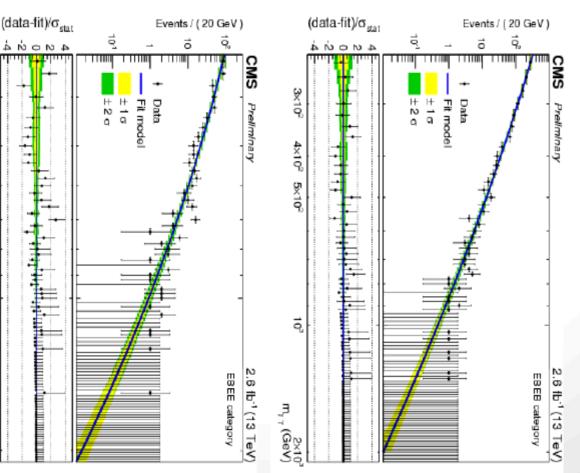
- Two categories: barrel-barrel (EBEB), barrel-endcap (EBEE)
- p_T(γ) > 75 GeV, I_{ch} < 5 GeV (in 0.3 cone around photon direction)
- Efficiency, scale and resolution calibrated on Z → ee and high-mass DY events
- Search for RS graviton with three assumptions on coupling: $\tilde{\kappa} = 0.01$ (narrow), 0.1, 0.2 (wide)
- Blind analysis, no changes have been made to the analysis since unblinding data in the signal region

1×10%

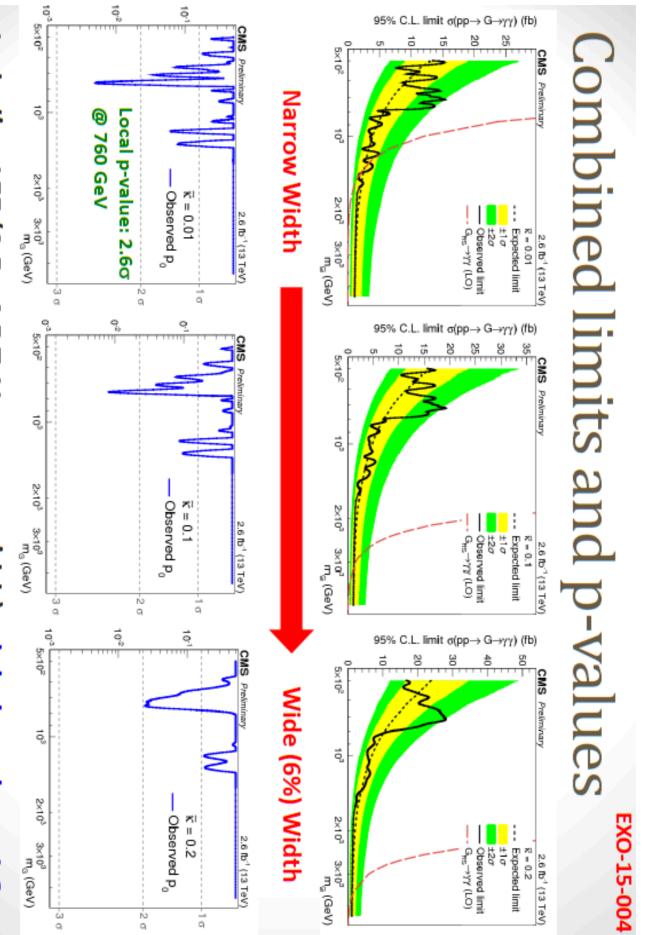
5×10° 6×10°

ą

m, , (GeV)



Including LEE (0.5 - 4.5 TeV; narrow width), global p-value < 1.2σ



..and CMS

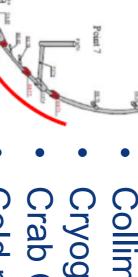
Only time will tell...



Project leadership: L. Rossi and O. Brüning Major intervention on more than 1.2 km of the LHC

ATLAS

- Machine protection
- Cold powering
- Crab Cavities



New IR-quads Nb₃Sn (inner triplets)

IC PROJECT

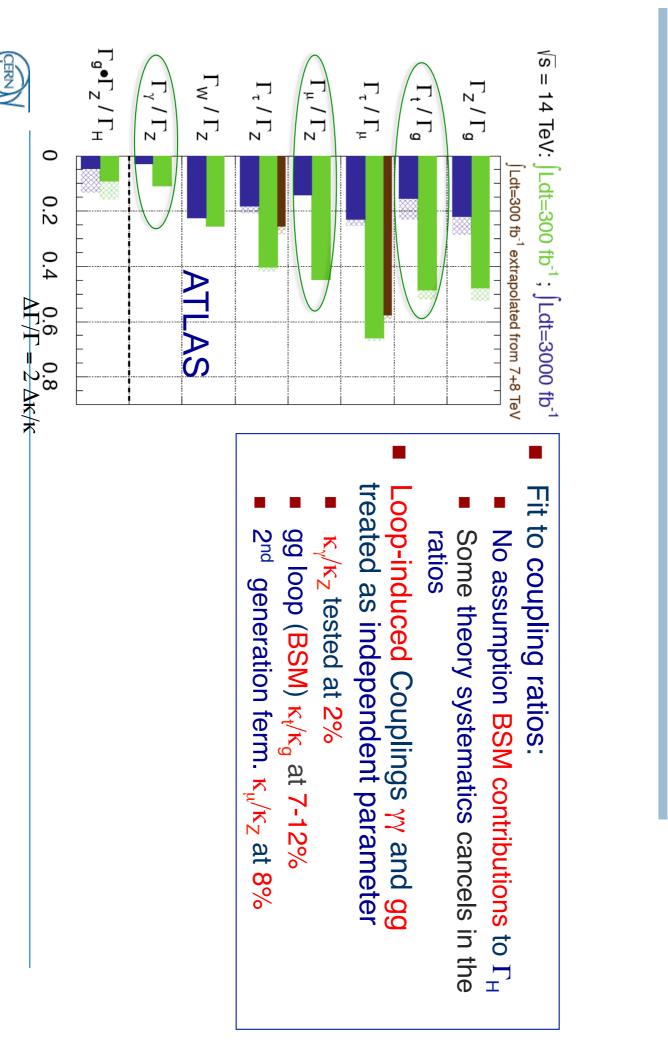
UNDERGROUND

The HL-LHC Project

- New 11 T Nb₃Sn (short) dipoles
- Collimation upgrade
- Cryogenics upgrade

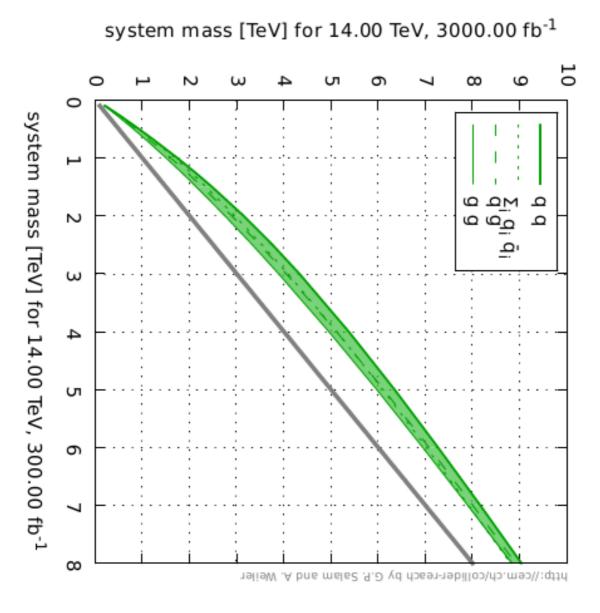


Higgs couplings fit at	cou	oling	s fit a	IT HL-LHC
			Uncertainty	ainty (%)
	Coupling	300	$300 {\rm ~fb^{-1}}$	3000 fb^{-1}
		Scenario 1	Scenario 2	Scenario 1 Scenario 2
CMS	κ_{γ}	6.5	5.1	5.4 1.5
	κV	5.7	2.7	4.5 1.0
	κ_{g}	11	5.7	7.5 2.7
	κ_b	15	6.9	11 2.7
	κ_t	14	8.7	8.0 3.9
	$\kappa_{ au}$	8.5	5.1	5.4 2.0
Assur	nption	NO invisi	CMS Projection visible/undetectabl	CMS Projection Assumption NO invisible/undetectable contribution to Γ_{H} :
- Scenario	1: syste	m./Theory	err. unch	Scenario 1: system./Theory err. unchanged w.r.t. current analysis
- scenario ∠: systemati ✓ γγ loop at 2-5% level	∠: system at 2-5%	evel	aled by T/S	Scenario 2: systematics scaled by T/sqrt(L), theory errors scaled by ½ / γγ loop at 2-5% level
🗸 down-ty	pe fermi	on couplin	down-type fermion couplings at 2-10% level	0% level
✓ direct to	<mark>p</mark> couplir	direct top coupling at 4-8% level	level	
✓ gg loop	gg loop at 3-8% level	level		
CERN				

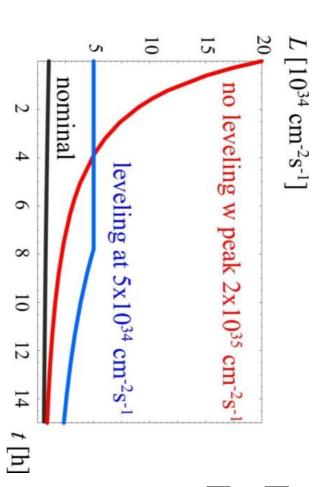


Coupling Ratios Fit at HL-LHC



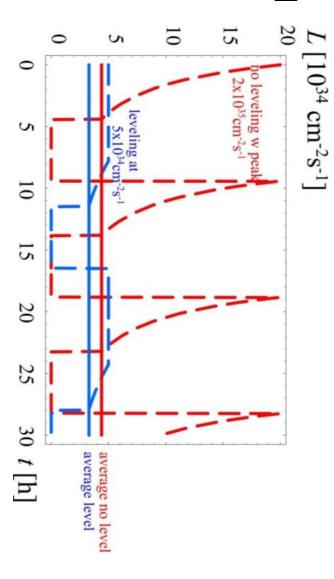


Luminosity Levelling, a key to success



- About 250 to 300 fb⁻¹/year (40% stable beams) Obtain about 3 - 4 fb⁻¹/day

High peak luminosity Minimize pile-up in "constant" luminosity experiments and provide





Baseline parameters of HL for reaching 250 -300 fb⁻¹/year

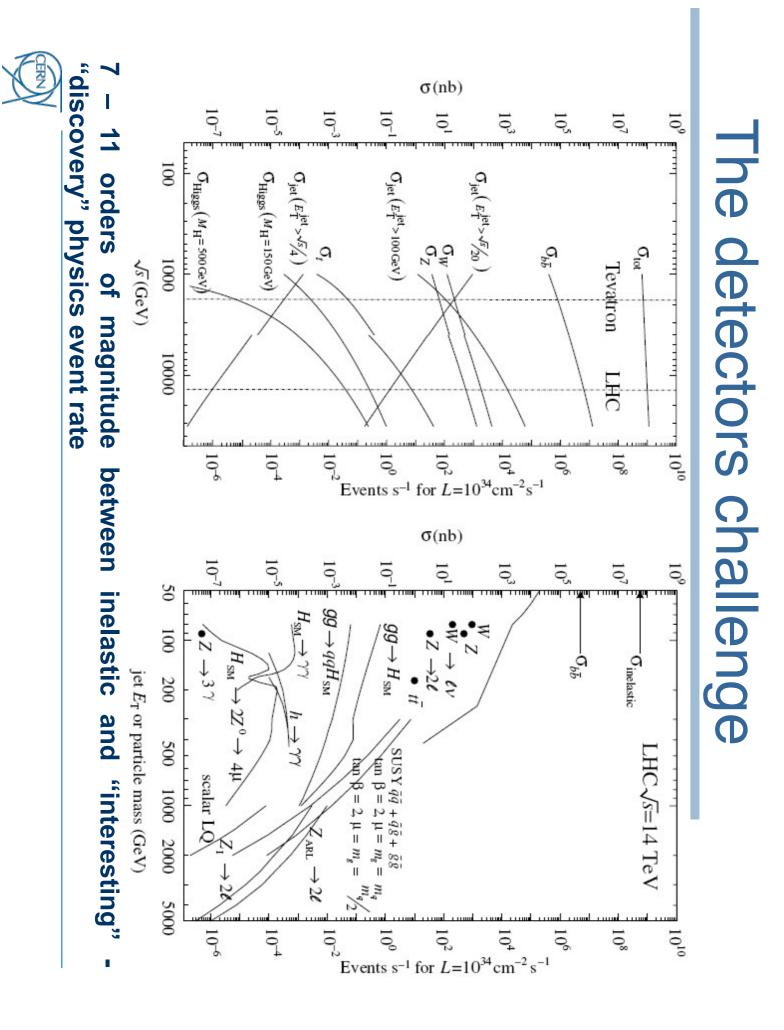
25 ns is the option However: 50 ns should be kept as alive and possible because we DO NOT hav

possible because we DO NOT have enough experience on the actual limit (e-clouds, l_{beam})

Continuous global optimisation with LIU

#Pile up @5E34	T _{leveling} [h] @ 5E34	Virtual lumi [10 ³⁴]	Peak lumi [10 ³⁴]	Loss factor	X-angle [µrad]	β^* [cm] (baseline)	γε _{x,y} [μm]	σ _{δp/p} [10- ³]	σ _z [cm]	ε _L [eV.s]	p/bunch [10 ¹¹]	# Bunches	
123	7.8	20.0	6.0	0.30	590 (12.5 σ)	15	2.5	0.1	7.5	2.5	2.0 (1.01 A)	2808	25 ns
247	6.8	22.7	7.4	0.33	590 (11.4 σ)	15	3.0	0.1	7.5	2.5	3.3 (0.83 A)	1404	50 ns



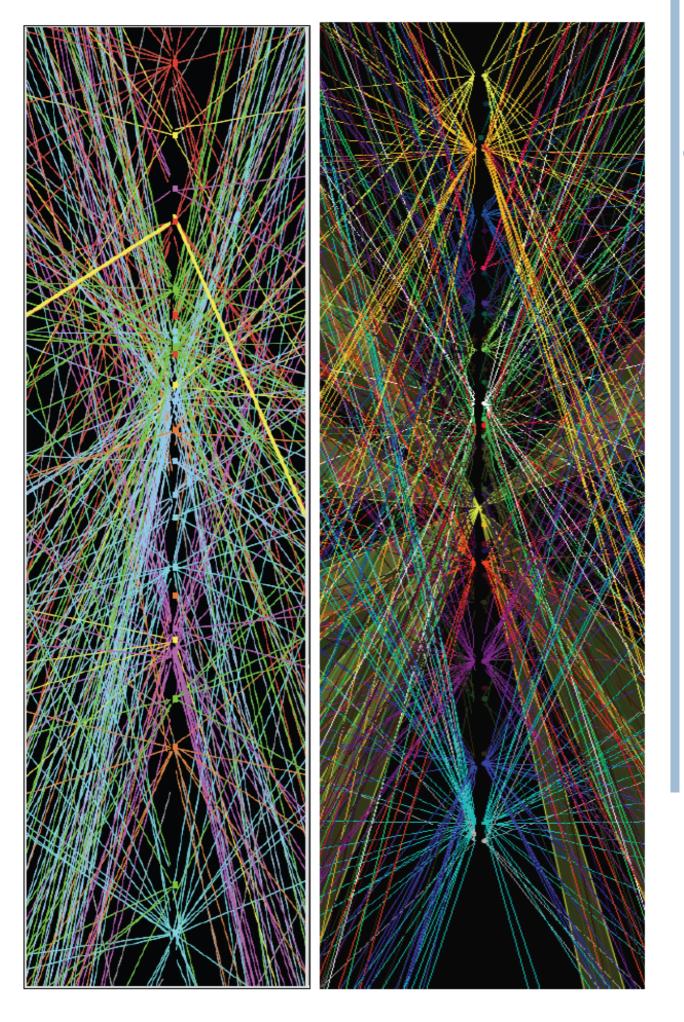


The detectors challenge

their capabilities to perform precision measurements at to maintain full sensitivity for discovery, while keeping low p_{T} , in the presence of: In order to exploit the LHC potential, experiments have

- Pileup
- <PU> ≈ 50 events per crossing by LS2
- <PU> \approx 60 events per crossing by LS3
- <PU> ≈ 140 events per crossing by HL-LHC
- Radiation damage
- Requires work to maintain calibration
- Limits performance-lifetime of the detectors
- Light loss (calorimeters)
- Increased leakage current (silicon detectors)

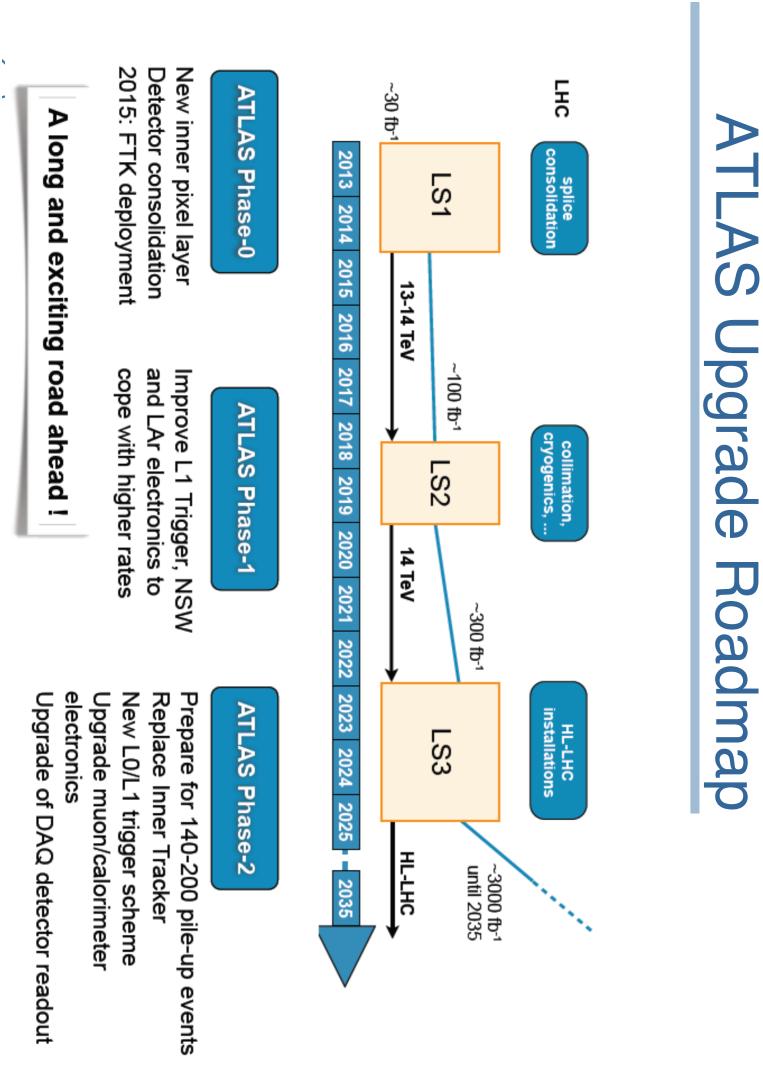


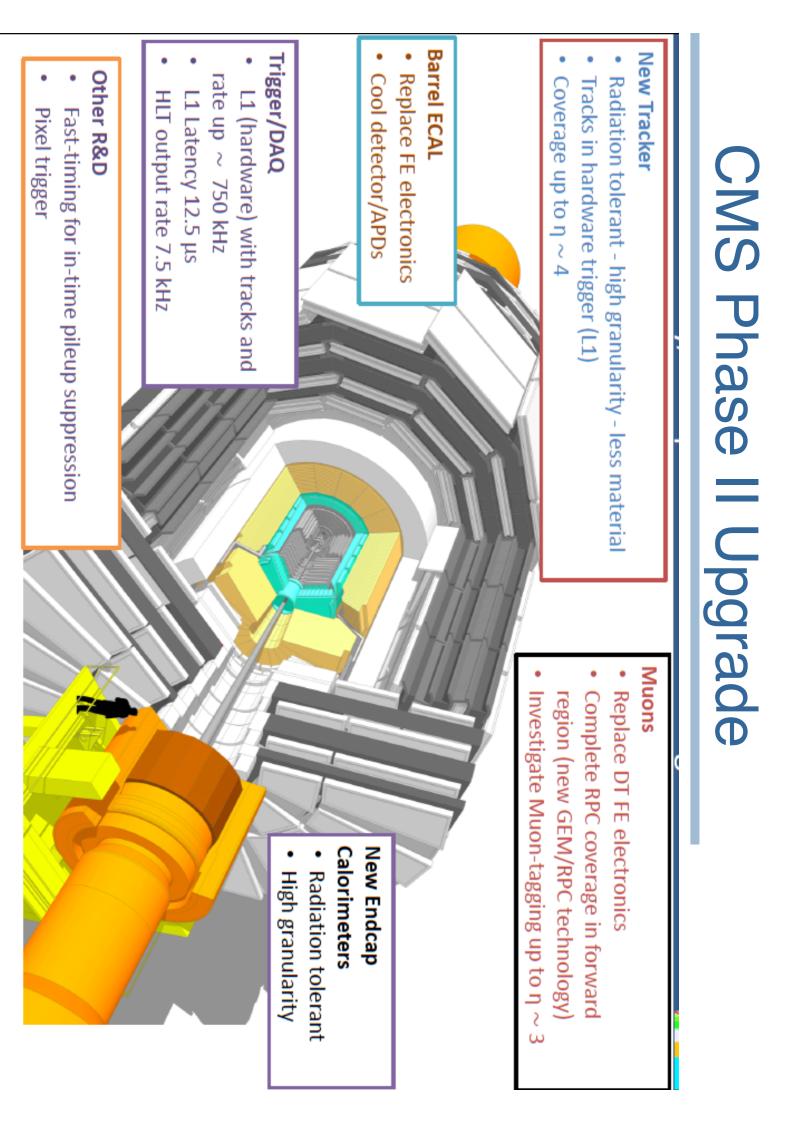


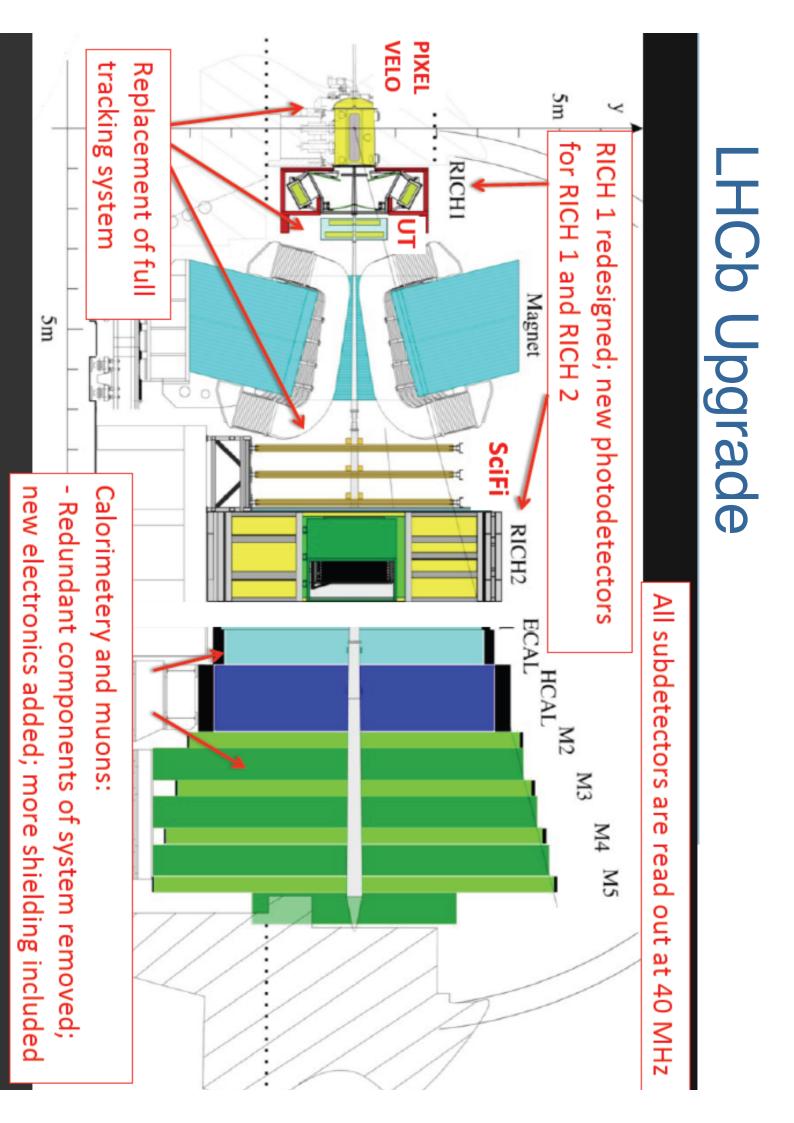
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Try to visualize x5!







ALICE Upgrade

New Inner Tracking System (ITS)

- improved pointing precision
- less material -> thinnest tracker at the LHC

Time Projection Chamber (TPC)

continuous readout New Micropattern gas detector technology

New Central Trigger

Processor (CTP)

Data Acquisition (DAQ)/ High Level Trigger (HLT)

- new architecture

- compression

50kHz Pbb event rate

TOF, TRD

New Trigger

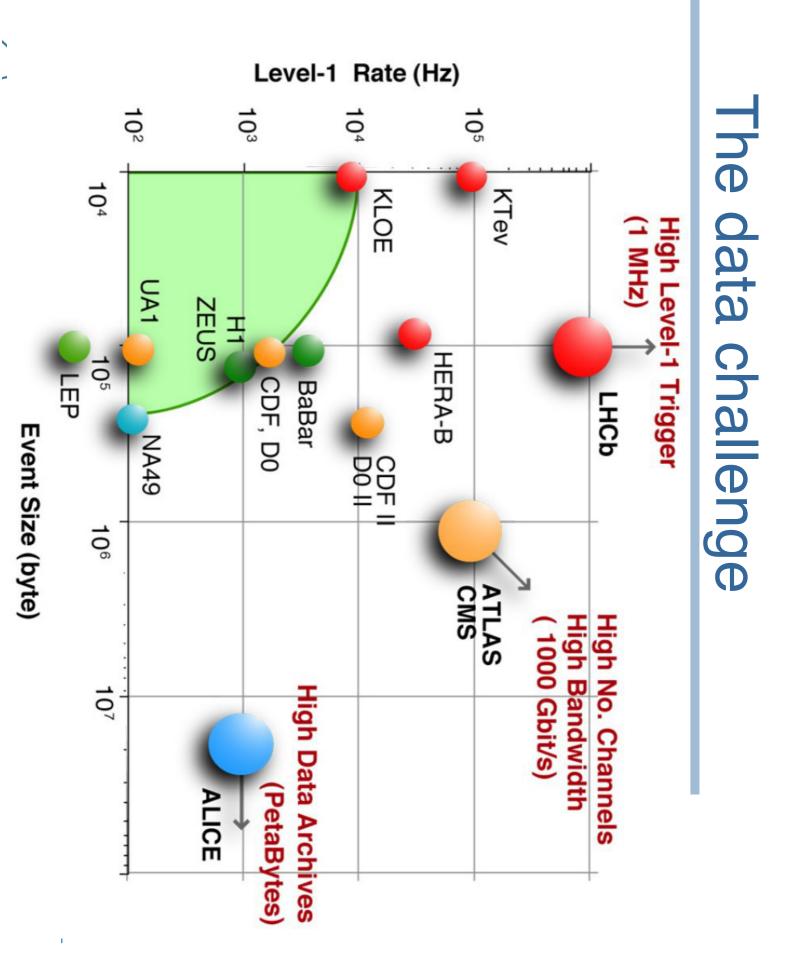
(c) by St. Rossegger

Detectors (FIT)

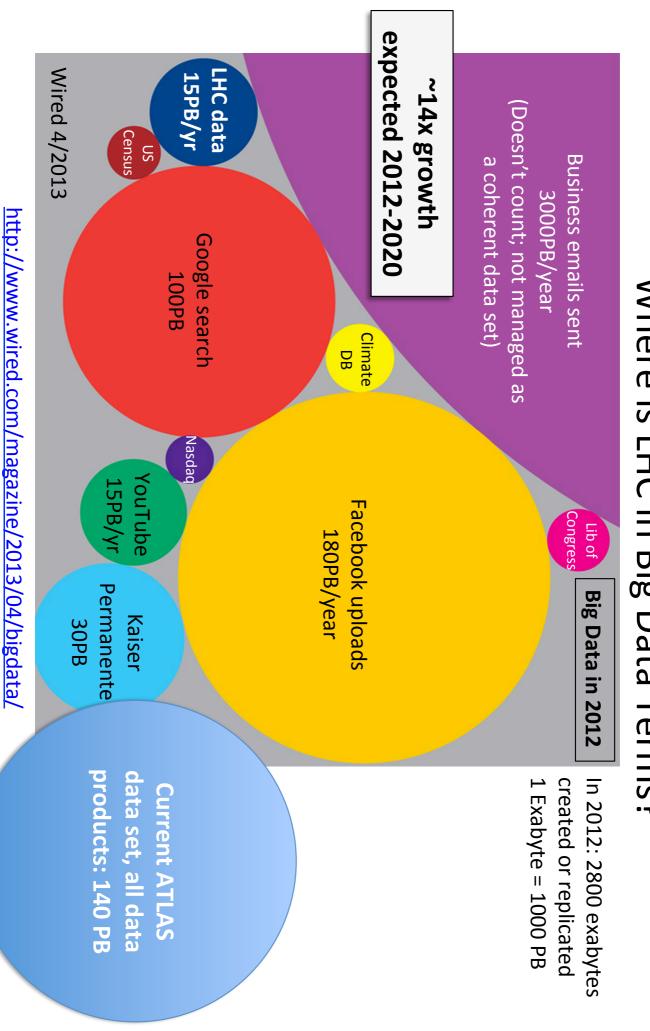
Faster readout

- on line tracking & data

- Muon Forward Tracker (MFT)
- new Si tracker
- Improved MUON pointing precision
- MUON ARM
- continuous readout
- electronics



Data Management Where is LHC in Big Data Terms?



BROOKH&VEN

October 15, 2013

100 km tunnel infrastructure in Geneva area with possibility of e+-e- (TLEP) and p-e (VLHeC) design driven by pp-collider requirements

Conceptual Design Report and cost review for the next ESU (≥2018)

FCC Design Study Kick-off Meeting: 12-14. February 2014 in Geneva international collaboration established, design study proceeding fast

Geneva



FCC-hh: 100 TeV

- explore directly the 10-50 TeV E-scale
- provide conclusive exploration of EWSB dynamics
- study nature the Higgs potential and EW phase transition
- say final word about heavy WIMP dark matter
- etc.

FCC-ee: 90-350 GeV

- indirect sensitivity to E scales up to O(100 TeV) by measuring most Higgs couplings to O(0.1%), improving the precision of EW parameters measurements by ~20-200 $\Delta M_W < 1$ MeV, $\Delta m_{top} \sim 10$ MeV, etc.
- sensitivity to very-weakly coupled physics (e.g. light, weakly-coupled dark matter)
- etc.
- FCC-ep: ~ 3.5 TeV
- \Box unprecedented measurements of PDF and α_s
- new physics: leptoquarks, eeqq contact interactions, etc.
- \Box Higgs couplings (e.g. Hbb to ~ 1%)
- etc.

Machines are complementary and synergetic, e.g. from measurement of ttH/ttZ ratio, and using ttZ coupling and H branching ratio from FCC-ee, FCC-hh can measure ttH to $\sim 1\%$

The challenge is not only the machine...

Detectors R&D :

- trackers Ultra-light, ultra-fast, ultra-granular, rad-hard, low-power Si
- at high-rate machines,..) 10⁸ channel imaging calorimeters (power consumption and cooling
- resolutions of ~10% for p~20 TeV muons big-volume 5-6 T magnets (~2 x magnetic length and bore of ATLAS and CMS, ~50 GJ stored energy) to reach momentum

Theory:

improved theoretical calculations (higher-order EW and QCD Work together with experiments on model-independent analyses in precision on EW observables, Higgs mass and branching ratios. corrections) needed to match present and future experimental

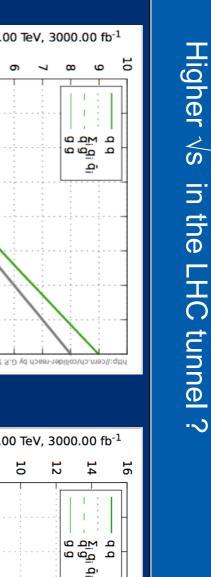
CERN

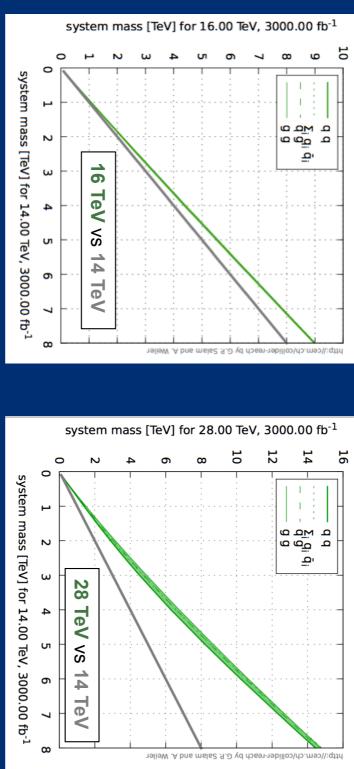
the framework of Effective Field Theory



FCC Week 2016 Fabiola Gianotti,

amount of HW and physics reach changes, technical challenges, cost, with increasing Various options





WG set up to explore technical feasibility of pushing LHC energy to: 1) design value: 14 IeV

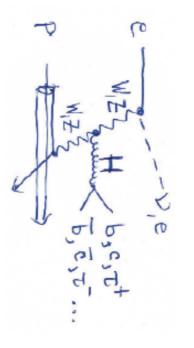
- ultimate value: 15 TeV (corresponding to max dipole field of 9 T)
- 3) beyond (e.g. by replacing 1/3 of dipoles with 11 T Nb₃Sn magnets) \rightarrow Identify open risks, needed tests and technical developments, trade-off
- between energy and machine efficiency/availability

 \rightarrow Report on 1) end 2016, 2) end 2017, 3) end 2018 (in time for ES)

HE-LHC (part of FCC study): ~16 T magnets in LHC tunnel ($\rightarrow \sqrt{s}$ ~ 30 TeV) I uses existing tunnel and infrastructure; can be built at fixed budget strong physics case if new physics from LHC/HL-LHC powerful demonstration of the FCC-hh magnet technology

LHeC, not only PDFs

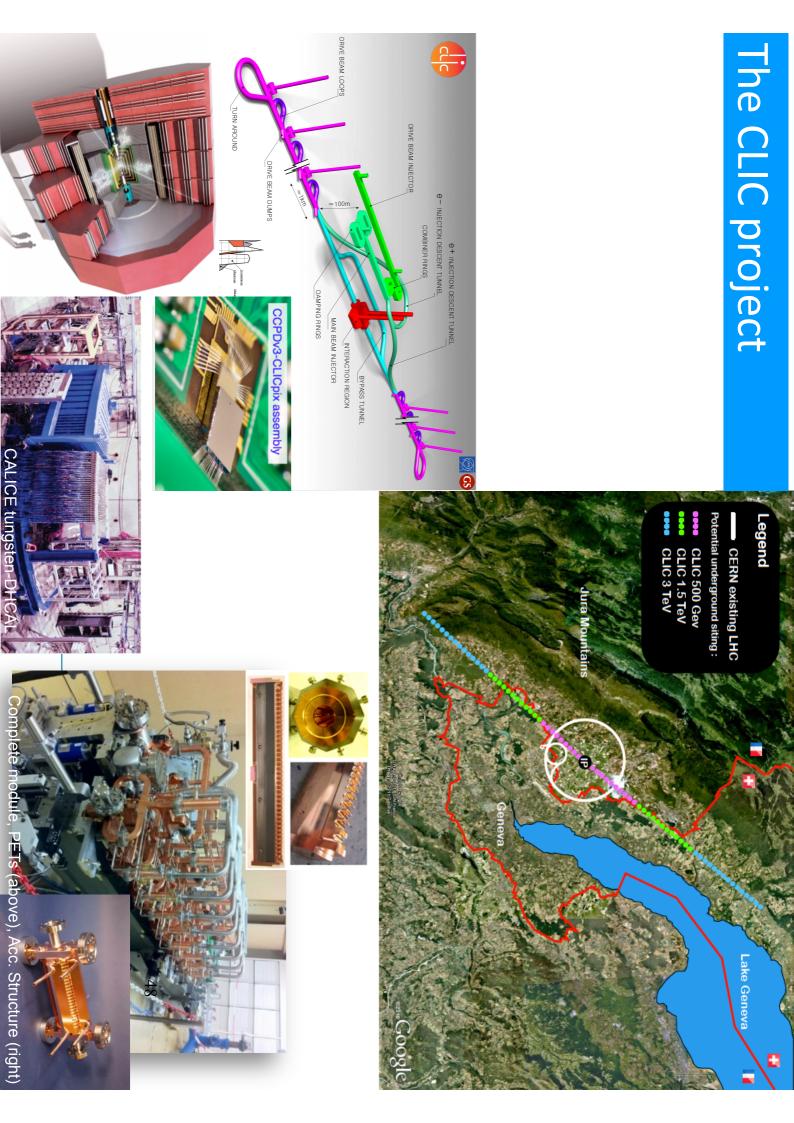




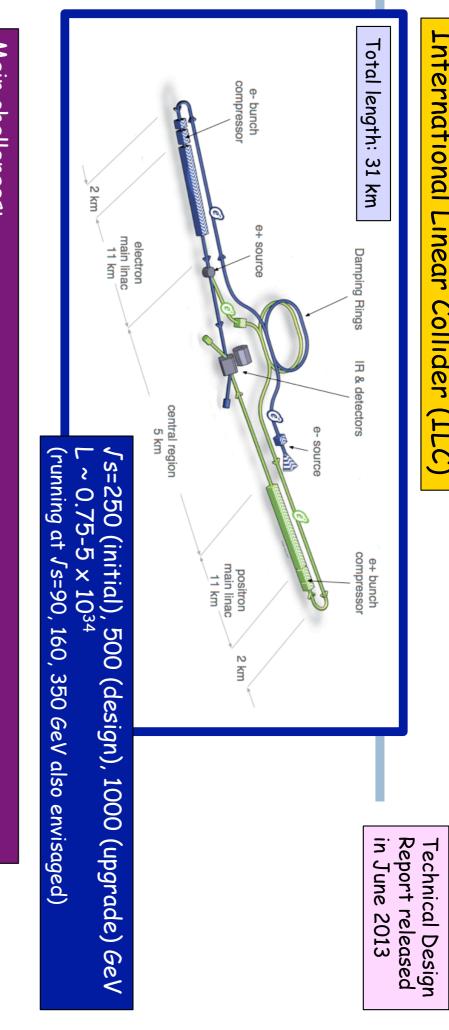
Continuing activity on Physics Detector ERL

Goal: L~10³⁴ cm⁻²s⁻¹









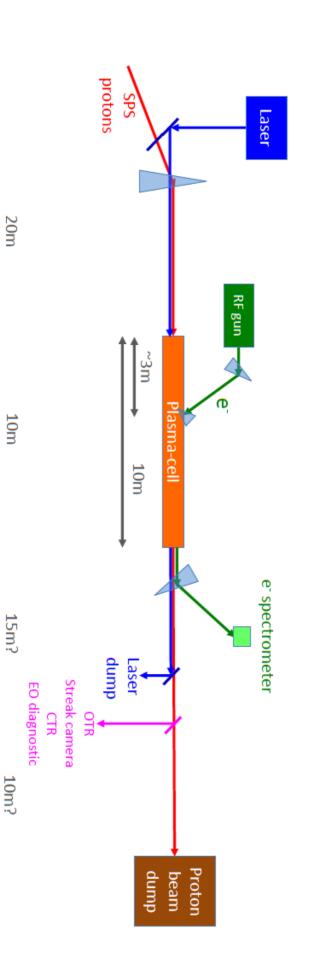
Main challenges:

- ~ 15000 SCRF cavities (1700 cryomodules), 31.5 MV/m gradient
- 1 TeV machine requires extension of main Linacs (50 km) and 45 MV/m
- Positron source; suppression of electron-cloud in positron damping ring
- Final focus: squeeze and collide nm-size beams
- ightarrow Construction could technically start ~2019, duration ~10 years ightarrow physics could start ~2030 Japan interested to host → decision ~2018 based also on ongoing international discussions (e.g. European xFEL at DESY is 5% of ILC, gradient 24 MV/m, some cavities achieved 29.6 MV/m) Mature technology: 20 years of R&D experience worldwide

Disruptive Technologies: Wakefield Acceleration

ATVAKE

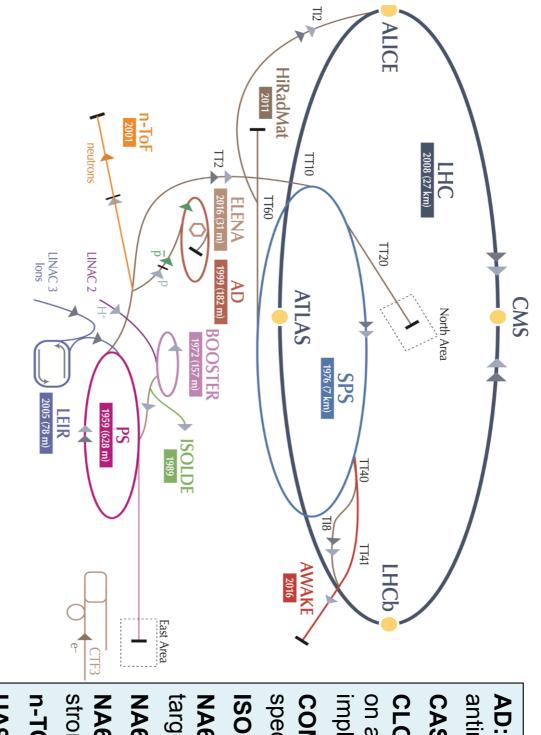
Experimental Layout



E. Gschwendtner, 4/9/2012



A compelling scientific programme beyond the LHC



~20 experiments > 1200 physicists

AD: Antiproton Decelerator for antimatter studies CAST, OSQAR: axions

CLOUD: impact of cosmic rays on aeorosols and clouds → implications on climate

COMPASS: hadron structure and spectroscopy ISOLDE: radioactive nuclei facility

NA61/Shine: ions and neutrino targets

NA62: rare kaon decays

NA63: radiation processes in strong EM fields

n-TOF: n-induced cross-sections

UA9: crystal collimation

Neutrino Platform: collaborating with experiments in US and lapan → see later

accelerator neutrinos \rightarrow imply neutrinos have masses and mix Neutrino oscillations (e.g. $\nu_{\mu} \rightarrow \nu_{e}$) established (since 1998) with solar, atmospheric, reactor and

Since then: great progress in understanding ν properties at various facilities all over the world

Nevertheless, several open questions:

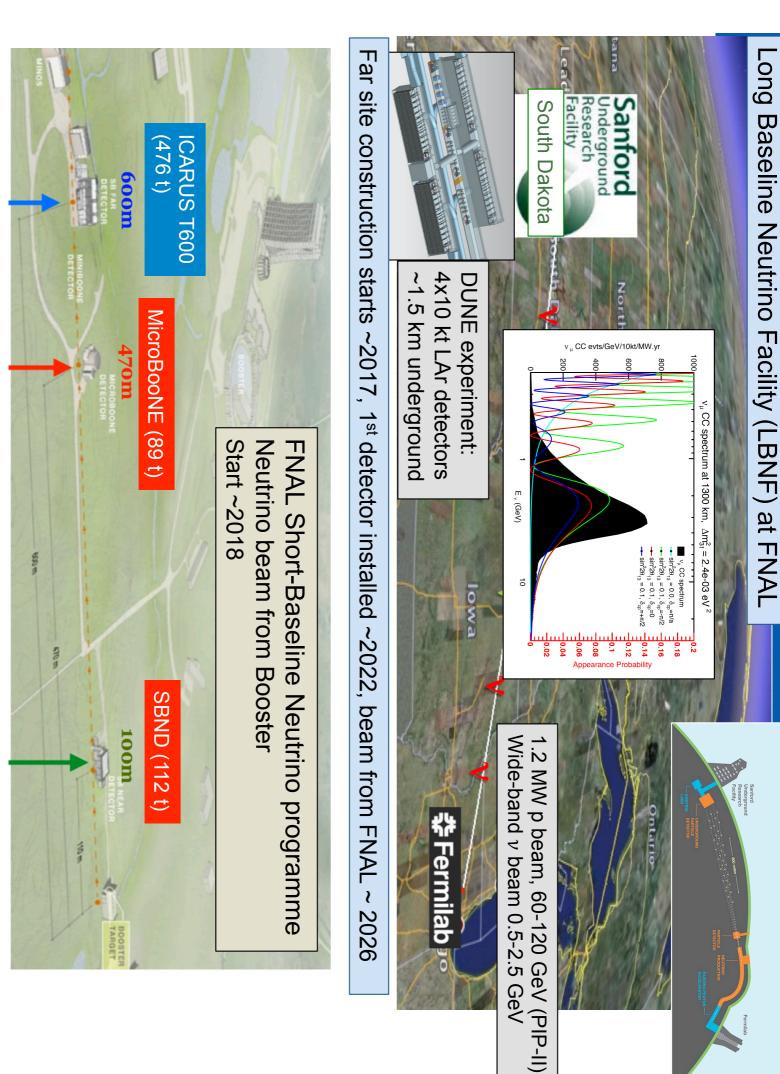
- Origin of ν masses (e.g. why so light compared to other fermions ?)
- Mass hierarchy: normal (v_3 is heaviest) or inverted (v_3 is lightest) ?
- Why mixing much larger than for quarks ?
- CP violation (observed in quark sector): do ν and anti- ν behave in the same way?
- Are there additional (sterile) ν (hints from observed anomalies)?

Accelerator experiments can address some of above questions studying $\nu_{\mu}
ightarrow \nu_{e}$ oscillations and the searched-for effects tiny \rightarrow Next-generation facilities planned in US and Japan. Need high-intensity p sources (> 1MW) and massive detectors, as v are elusive particles

f) Rapid progress in neutrino oscillation physics, with significant European involvement, has established a strong scientific case for a long-baseline neutrino programme exploring CP violation the way for a substantial European role in future long-baseline experiments. Europe should and the mass hierarchy in the neutrino sector. CERN should develop a neutrino programme to pave explore the possibility of major participation in leading long-baseline neutrino projects in the US

European Strategy 2013

and Japan.



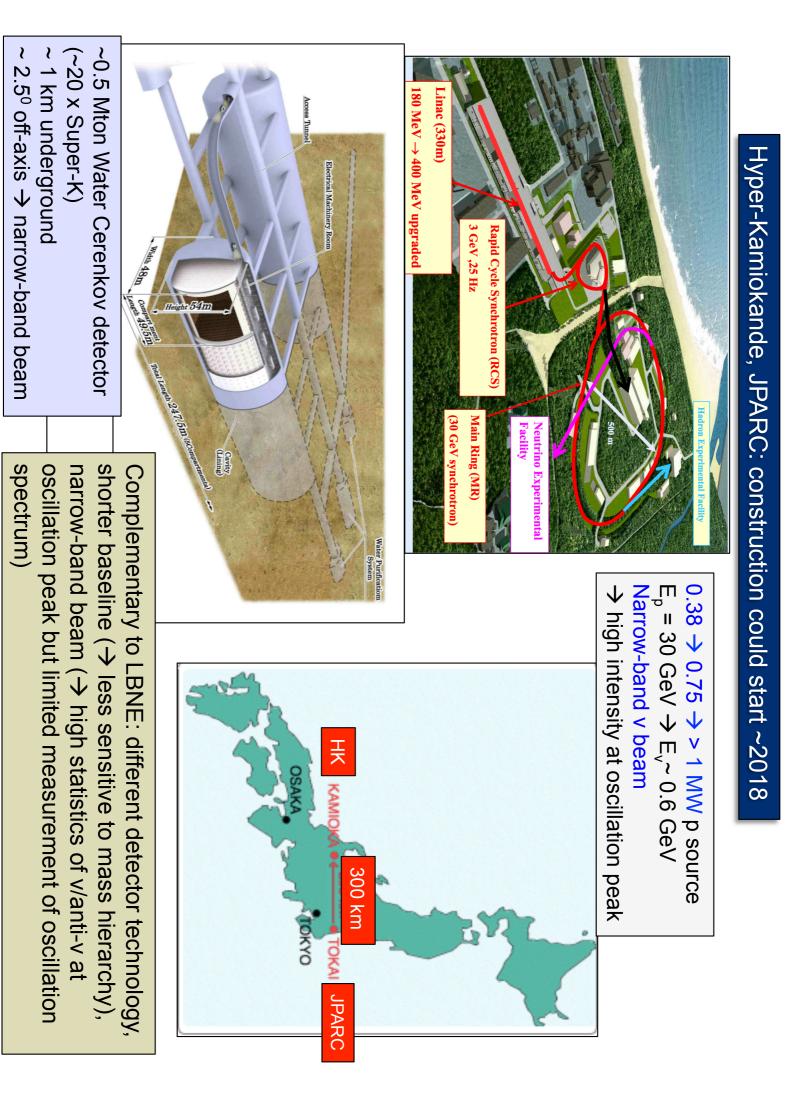
A 25+ years Physics Program

On the beam:

- Perform a comprehensive investigation of neutrino oscillations
- test CP violation in the lepton sector
- determine the ordering of the neutrino masses
- test the three-neutrino paradigm
- Perform a broad set of neutrino scattering measurements with the near detector

accelerator physics topics Exploit the large, high-resolution, underground far detector for non-

- atmospheric neutrino measurements
- searches for nucleon decay
- core-collapse supernova). measurement of astrophysical neutrinos (especially those from a





CERN Neutrino Platform

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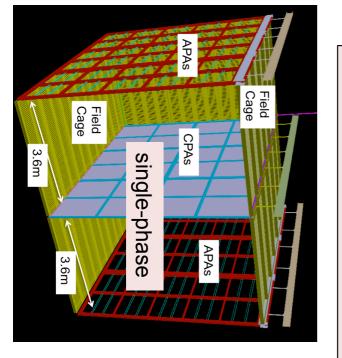
Mission:

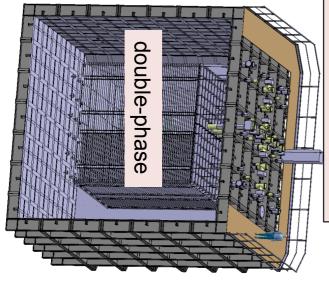
- Provide charged beams and test space to
- □ North Area extension
 □ Support European participation in accelerator
- neutrino experiments in US and Japan:
- → R&D to demonstrate large-scale LAr technology (cryostats, cryogenics, detectors)
- \checkmark Construction of one cryostat for DUNE detector modules
- ightarrow Construction of BabyMIND magnet: muon spectrometer for WAGASCI experiment at JPARC

Refurbishment of ICARUS T600 for short baseline programme → ship to FNAL beg 2017

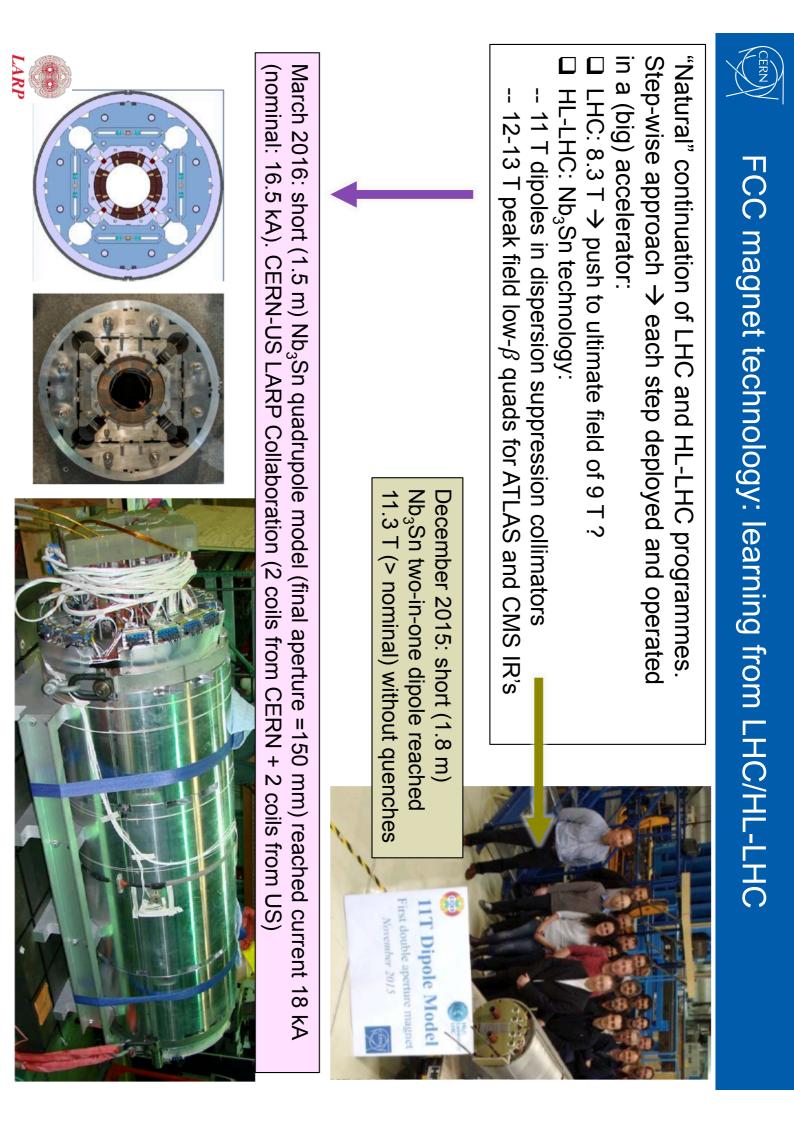


Construction and test of "full-scale" prototypes of DUNE drift cells: ~ $6x6x6 \text{ m}^3$, ~ 700 tons





ready for beam tests in 2018 (before LS2)





A "Physics Beyond Colliders" Study Group has been put in place

Mandate

outstanding questions in particle physics through projects: Explore opportunities offered by CERN accelerator complex and infrastructure to address

Examples: searches for rare processes and very-weakly interacting particles, electric dipole moments, etc. I complementary to other efforts in the world \rightarrow optimise resources of the discipline globally exploiting unique capabilities of CERN accelerator complex and infrastructure complementary to future high-energy colliders (HE-LHC, CLIC, FCC)

→ Enrich and diversify CERN's future scientific programme

Kick-off meeting in Summer 2016 Will bring together accelerator scientists, experimental and theoretical physicists

Final report end 2018 \rightarrow in time for European Strategy

with other laboratories and institutions in Europe (and beyond) One of the goals is to involve interested worldwide community, and to create synergies

In summary

- An exciting period in front of us:
- We have finished the inventory of the "known unknown"...
- few tantalizing hints to probe) ...but we have a vast space to explore (and a
- 20 years We have a solid physics program for the next 15
- In this time period we have to prepare for the political frames. next steps, setting directions, technologies and





We will need: agenda of the field. Experimental results will be dictating the

- Flexibility
- Preparedness
- Visionary global policies



THANK YOU