

Pulse Shape Discrimination for high pressure ³He counters

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The success of low counting rate experiments strongly depends on the accurate knowledge and reduction of the backgrounds present during the measurement campaigns. For this reason, dark matter search and low energy nuclear astrophysics experiments are performed in underground laboratories, shielded from the cosmic rays [1, 2]. Two important examples of low counting rate experiments for nuclear astrophysics are the measurement of 13 C(α ,n) and 22 Ne(α ,n) cross-sections at low energies, which are considered to be the main neutron sources for the astrophysical s-process [3, 4].

These experiments require the use of high intrinsic efficiency detectors, such as the high pressure 3 He counters, sensitive to thermal neutrons through the 3 He(n,p) reactions. The intrinsic alpha-activity contained in the walls of these counters becomes a major source of background in low count rate scenarios.

To reduce this intrinsic background, over the years, many efforts were made in the pulse shape discrimination between the neutron and α event signature. Two recent examples are the use of current sensitive preamplifiers, taking advantage of the double peak structure formed by the triton and proton particles [5, 6] and low pressure 3 He counters, for which the rise-time differences between α and 3 He(n,p) events permit pulse shape discrimination [7].

However, most experiments are performed with charge sensitive preamplifiers in the electronic chain, thus losing the double peak structure of ${}^{3}\text{He}(n,p)$ reactions, and the usage of higher-pressure (and thus higher-efficiency) counters limits the usefulness of discrimination by rise time.

We will present an improved pulse shape discrimination methodology developed for high pressure ³He counters read out through charge sensitive preamplifiers electronics using standard digitizing methods, and a Monte Carlo simulation of the detector response to the different particle interactions.

References

- [1] A. Best et al. NIM A 812(2016)16
- [2] D. Jordan et al. Astroparticle Physics 42 (2013) 16
- [3] F. Käppler et al. Nuclear Physics A 777 (2006) 219-310
- [4] U. Greife et al. The Astrophysical Journal 414 (1993) 735-739
- [5] A .Amsbaugh et al., NIM A 579 (2007) 10541080
- [6] Z.M. Zeng et al., NIM A 866(2017)242247

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[7] T.J.Langford et al., NIM A 717(2013)5157