

Phenomenological analysis related to schemes with sterile neutrinos

Workshop on Beyond Three Family Neutrino Oscillations

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Outline

SBL phenomenology

Active-sterile mixing

$$U_{e4}$$

$$U_{\mu 4}$$

$$U_{\tau 4}$$

Global data

More exotic proposals

Sterile neutrino + NSI

Soft decoherence

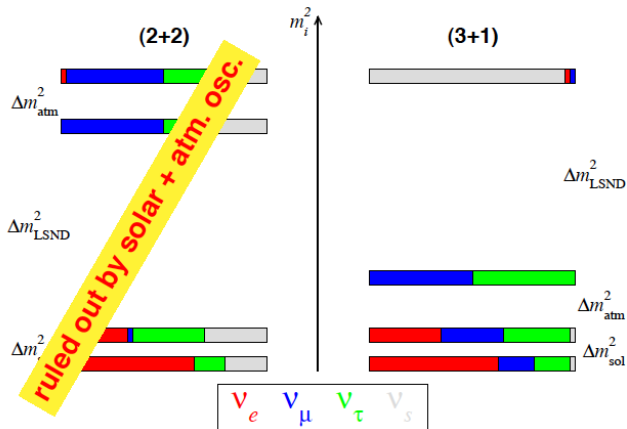
Conclusions

Short-baseline (SBL) experiments

- ▶ $\nu_\mu \rightarrow \nu_e, \bar{\nu}_\mu \rightarrow \bar{\nu}_e$ appearance searches
LSND, MiniBooNE, KARMEN, NOMAD
- ▶ $\bar{\nu}_e$ disappearance experiments (reactors)
Bugey-3 (15, 40, 95), Bugey-4, Goessgen, Krasnoyarsk, Rovno, ILL, Chooz, Palo Verde
- ▶ ν_μ disappearance
CDHS, MiniBooNE, atmospheric neutrinos, MINOS NC

SBL: $E/L \sim \text{eV}^2 \rightarrow \text{set } \Delta m_{21}^2 \approx \Delta m_{31}^2 \approx 0$
 exceptions: Chooz, Palo Verde, atmospheric, MINOS NC

4-neutrino schemes



3+1: small perturbation to 3-active case

3+1 SBL oscillations

appearance

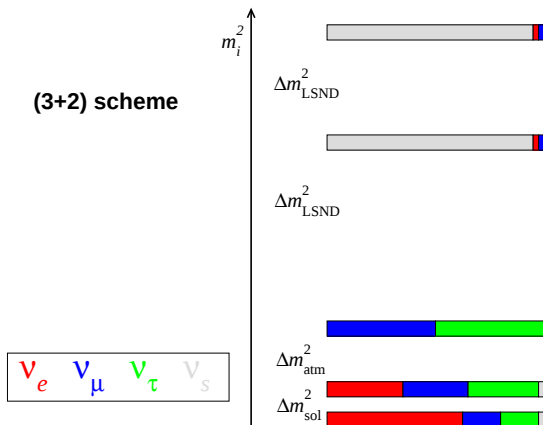
$$P_{\mu e} = \sin^2 2\theta_{\text{app}} \sin^2 \frac{\Delta m_{41}^2 L}{4E} \quad \sin^2 2\theta_{\text{app}} = 4|U_{e4}|^2|U_{\mu 4}|^2$$

disappearance

$$P_{\alpha\alpha} = 1 - \sin^2 2\theta_{\text{dis}} \sin^2 \frac{\Delta m_{41}^2 L}{4E} \quad \sin^2 2\theta_{\text{dis}} = 4|U_{\alpha 4}|^2(1 - |U_{\alpha 4}|^2)$$

- ▶ effective 2-flavour oscillations
- ▶ no CP violation $\rightarrow$ can't reconcile $\bar{\nu}$ (LSND, MB) and ν (MB) data
- ▶ constraints from ν_e (ν_μ) disappearance experiments on U_{e4} ($U_{\mu 4}$)
appearance mixing angle quadratically suppressed

5-neutrino schemes



Peres, Smirnov, hep-ph/0011054; Sorel, Conrad, Shaevitz, hep-ph/0305255

3+2 SBL oscillations

appearance:

$$\begin{aligned}
 P_{\nu_{\mu} \rightarrow \nu_e} &= 4 |U_{e4}|^2 |U_{\mu 4}|^2 \sin^2 \phi_{41} + 4 |U_{e5}|^2 |U_{\mu 5}|^2 \sin^2 \phi_{51} \\
 &+ 8 |U_{e4} U_{\mu 4} U_{e5} U_{\mu 5}| \sin \phi_{41} \sin \phi_{51} \cos(\phi_{54} - \delta)
 \end{aligned}$$

disappearance:

$$P_{\nu_{\alpha} \rightarrow \nu_{\alpha}} \approx 1 - 4 \sum_{i=4,5} |U_{\alpha i}|^2 \sin^2 \phi_{i1} - 4 |U_{\alpha 4}|^2 |U_{\alpha 5}|^2 \sin^2 \phi_{54}$$

$$[\phi_{ij} \equiv \Delta m_{ij}^2 L / 4E]$$

- ▶ phase $\delta \equiv \arg \left(U_{e4}^* U_{\mu 4} U_{e5} U_{\mu 5}^* \right) \rightarrow$ CP violation Karagiorgi et al. 06; Maltoni, TS 07
- ▶ BUT: constrain $|U_{ei}|$ and $|U_{\mu i}|$ ($i = 4, 5$) from disappearance to be reconciled with appearance amplitudes $|U_{ei} U_{\mu i}|$

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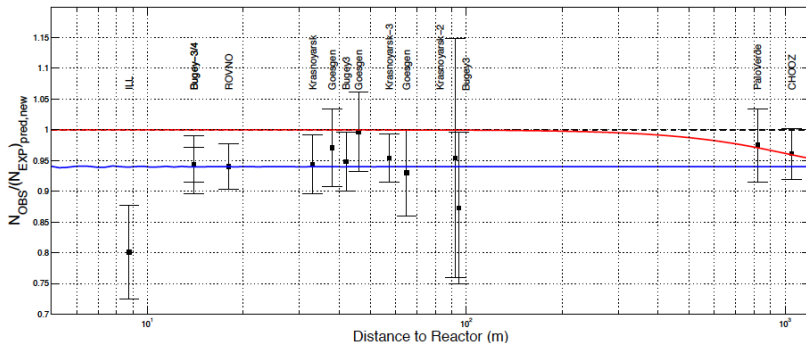
What do we know about active-sterile mixing?

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

independent information on $U_{\alpha 4}$ mainly from disappearance exps

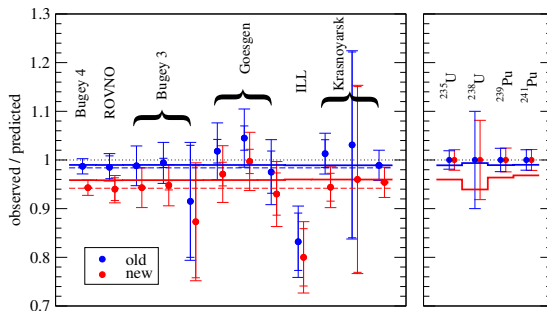
The “reactor anomaly”

Mention et al., 1101.2755, talk by D. Lhuillier



oscillations with $\Delta m^2 \sim 1 \text{ eV}^2$ may account for disappearance
at baselines $L \lesssim 100 \text{ m}$

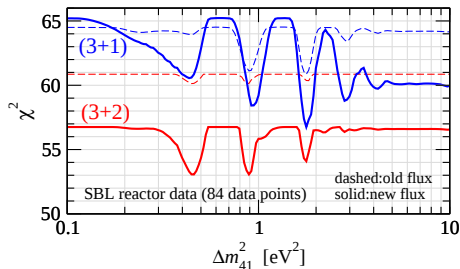
The “reactor anomaly”



TS, Tortola, Valle 11

- ▶ $\chi^2 = 8.1(13.0)/12$ dof for old (new) fluxes $\Rightarrow$ still acceptable fit within errors (P-value: 37%)
- ▶ flux free analysis with new fluxes: $f = 0.942 \pm 0.024$, $f = 1$ disfavored with $\Delta\chi^2 = 6.2$ (2.5σ)

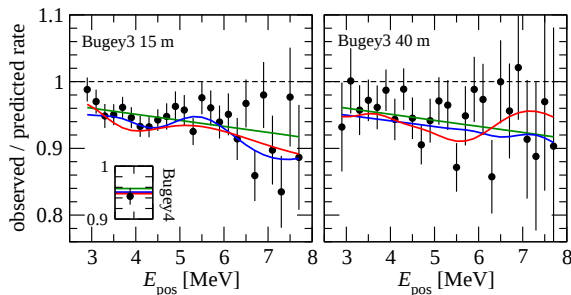
3+1 and 3+2 fit to SBL reactor data



preference for sterile neutrino oscillations with new fluxes

- ▶ $\Delta\chi^2(3 + 1/\text{no-osc}) = 8.5 \rightarrow 98.6\% \text{ CL (2 dof)}$
- ▶ $\Delta\chi^2(3 + 2/\text{no-osc}) = 12.1 \rightarrow 98.3\% \text{ CL (4 dof)}$
- ▶ old flux: $\Delta\chi^2 = 3.6(4.4)$ for 3+1 (3+2)
- ▶ further improvement of fit for 3+2 wrt 3+1 with new fluxes

3+1 and 3+2 fit to SBL reactor data



$$\chi^2_{\text{no osc}} = 59.0/69$$

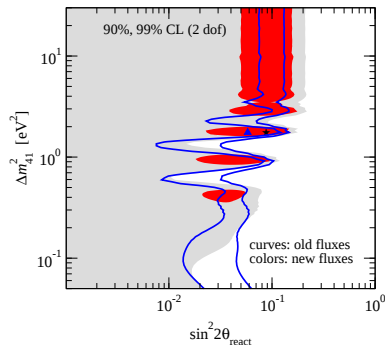
$$\chi^2_{3+1} = 50.1/67$$

$$\chi^2_{3+2} = 46.5/65$$

	Δm_{41}^2 [eV ²]	$ U_{e4} $	Δm_{51}^2 [eV ²]	$ U_{e5} $
3+1	1.78	0.151		
3+2	0.46	0.108	0.89	0.124

Kopp, Maltoni, TS, 11

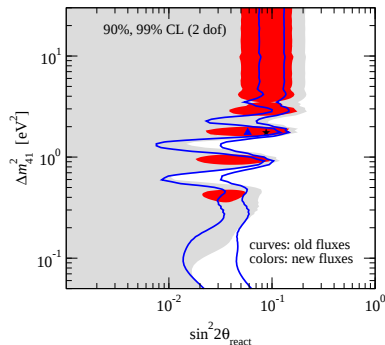
3+1 and U_{e4} from SBL reactor data



$$\begin{aligned}\sin^2 2\theta_{\text{react}} &= 4|U_{e4}^2|(1 - |U_{e4}^2|) \\ &\approx 4|U_{e4}^2|\end{aligned}$$

- ▶ previous constraints become now “preferred regions” with $U_{e4} > 0$

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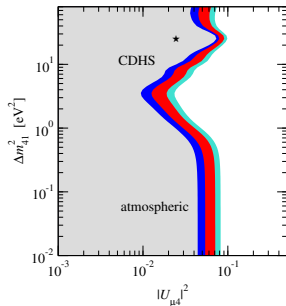


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- ▶ “Gallium anomaly” gives additional hint for $U_{e4} > 0$

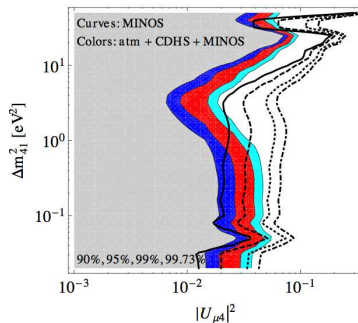
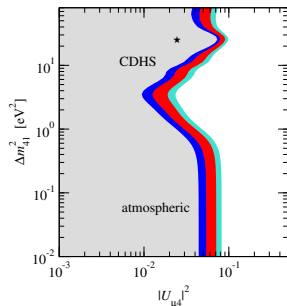
Acero, Giunti, Laveder, 0711.4222; Giunti, Laveder, 1006.3244; talks by E. Bellotti and C. Giunti

3+1 and limits on $U_{\mu 4}$



- ▶ ν_μ disappearance CDHS (and also MiniBooNE)
- ▶ atmospheric neutrinos Bilenky, Giunti, Grimus, TS 99; Maltoni, TS, Valle 01

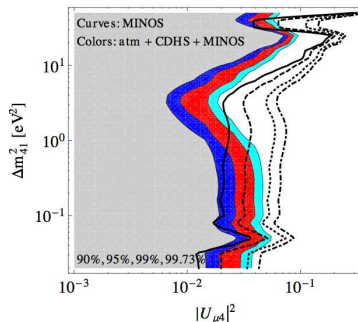
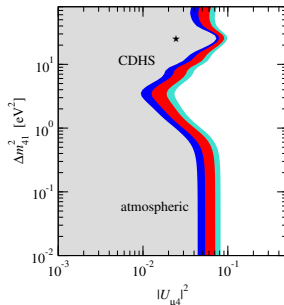
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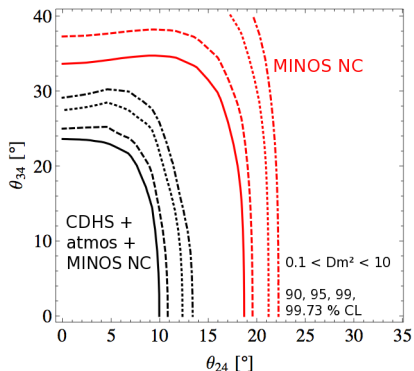
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- ▶ MINOS NC data [1001.0336, 1104.3922](#); talk by R. Nichol
- ▶ additional constraints from MiniB $\nu_\mu(\bar{\nu}_\mu)$ at $\Delta m^2 \gtrsim 10 \text{ eV}^2$ [talk by M. Sorel](#)

3+1 and limits on $U_{\tau 4}$



$$|U_{\mu 4}^2| = \sin^2 \theta_{24} < 0.03$$

$$|U_{\tau 4}^2| = \cos^2 \theta_{24} \sin^2 \theta_{34} < 0.16$$

What do we know about active-sterile mixing?

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & \approx 0.15 \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & < 0.17 \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & < 0.4 \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \quad \begin{aligned} |U_{e4}^2| &\approx 0.025 \\ |U_{\mu 4}^2| &< 0.03 \\ |U_{\tau 4}^2| &< 0.16 \end{aligned}$$

for $3+n$:

$$\text{limit on } |U_{\alpha 4}^2| \rightarrow \text{limit on } \sum_{i=4}^{3+n} |U_{\alpha i}^2|$$

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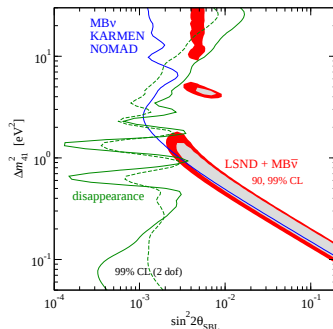
Conclusions

$\nu_\mu \rightarrow \nu_e$ data at the $E/L \sim 1 \text{ eV}^2$ scale

- ▶ **LSND** $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$, $87.9 \pm 22.4 \pm 6.0$ excess events
 $P = (0.264 \pm 0.067 \pm 0.045)\% \sim 3.8\sigma$ away from zero
- ▶ **MiniBooNE** $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$, $\sim 2\sigma$ excess
consistent with LSND in 2ν framework
- ▶ **MiniBooNE** $\nu_\mu \rightarrow \nu_e$
 $E > 475$: no excess, $E < 475$: $\sim 3\sigma$ excess
- ▶ **KARMEN** $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$, tight constraint on LSND region
(slightly smaller L/E than LSND)

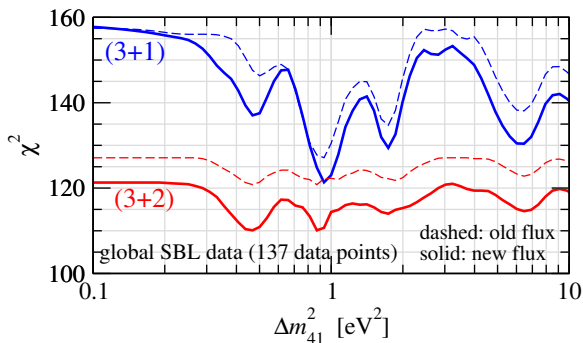
talk by M. Sorel

3+1 global fit



- ▶ no CP violation $\rightarrow$ LSND+MB $\bar{\nu}$ versus MB ν
- ▶ despite relaxed constraints on U_{e4} no improvement of global 3+1 fit
- ▶ LSND+MB $\bar{\nu}$ versus rest: $\chi^2_{PG} = 21.5(24.2)$ for new (old) flux
 $\rightarrow$ compatibility of less than 10^{-5}

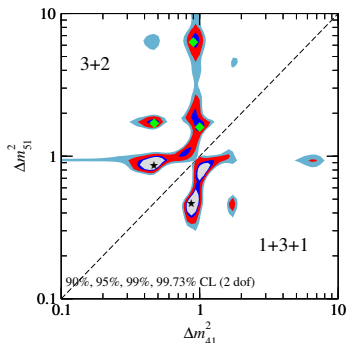
3+2 global fit



- ▶ $\Delta\chi_{3+2}^2$ (old vs new fluxes) = 11.1
- ▶ $\Delta\chi^2$ (3+1 vs 3+2) = 11.2 (97.6% CL, 4 dof)
6.3 for old flux

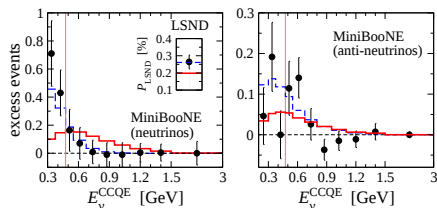
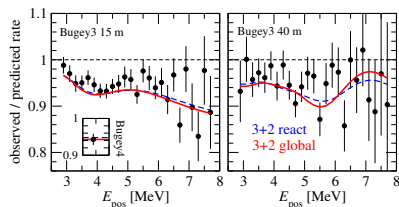
Kopp, Maltoni, TS, 11

3+2 global fit



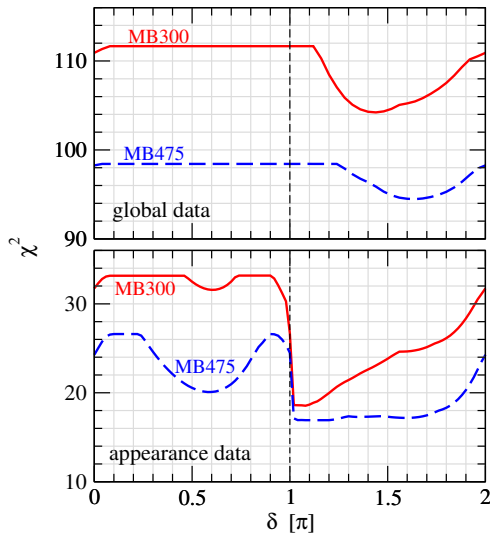
	Δm_{41}^2	$ U_{e4} $	$ U_{\mu 4} $	Δm_{51}^2	$ U_{e5} $	$ U_{\mu 5} $	δ/π	$\chi^2/130$
3+2	0.47	0.128	0.165	0.87	0.138	0.148	1.64	110.1
3+2'	0.47	0.117	0.201	1.70	0.151	0.101	1.39	114.4
3+2'	1.00	0.133	0.163	1.60	0.122	0.079	1.48	114.4
3+2'	0.90	0.123	0.163	6.30	0.135	0.091	1.67	115.0
1+3+1	0.47	0.129	0.154	0.87	0.142	0.163	0.35	106.1

3+2 global fit



- ▶ 3+2 fit to appearance data only (LSND, $MB\nu$, $MB\bar{\nu}$, KARMEN, NOMAD) allows for explanation of MB low energy excess [Maltoni, TS, 07](#)
- ▶ BUT: this requires large U_{e4} , U_{e5} , $U_{\mu4}$, $U_{\mu5}$ in sever disagreement with disappearance bounds
- ▶ MB low-E excess cannot be explained using global data

CP violation?



Maltoni, TS 07

Still considerably tension in 3+2 fit

	LSND+MB($\bar{\nu}$) vs rest		appearance vs disapp.	
	old	new	old	new
$\chi^2_{\text{PG},3+2}/\text{dof}$	25.1/5	19.9/5	19.9/4	14.7/4
PG ₃₊₂	10^{-4}	0.13%	5×10^{-4}	0.53%

- ▶ explanation of appearance signal requires U_{ei} and $U_{\mu i}$ ($i = 4, 5$)
- ▶ no hint for oscillations seen in ν_μ disappearance
 $\Rightarrow$ constraints on $U_{\mu i}$ in tension with signals
- ▶ MINOS NC data not (yet) included in the fit, slightly change compatibility Kopp, Maltoni, TS, in preparation

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- ▶ 3-neutrinos and CPT violation Murayama, Yanagida 01;
Barenboim, Borissov, Lykken 02; Gonzalez-Garcia, Maltoni, TS 03
- ▶ 4-neutrinos and CPT violation Barger, Marfatia, Whisnant 03
- ▶ Exotic muon-decay Babu, Pakvasa 02
- ▶ CPT viol. quantum decoherence Barenboim, Mavromatos 04
- ▶ Lorentz violation Kostelecky et al., 04, 06; Gouvea, Grossman 06
- ▶ mass varying ν Kaplan, Nelson, Weiner 04; Zurek 04; Barger, Marfatia, Whisnant 05
- ▶ shortcuts of sterile ν s in extra dim Paes, Pakvasa, Weiler 05
- ▶ decaying sterile neutrino Palomares-Riu, Pascoli, TS 05; Gninenko 10
- ▶ 2 decaying sterile neutrinos with CPV
- ▶ energy dependent quantum decoherence Farzan, TS, Smirnov 07
- ▶ sterile neutrinos and new gauge boson Nelson, Walsh 07
- ▶ sterile ν with energy dep. mass or mixing TS 07
- ▶ sterile ν with nonstandard interactions Akhmedov, TS 10

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most of these proposals involve sterile neutrinos

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most of these proposals have problems with some data

3+1 + NSI

Akhmedov, TS 10 assume

- ▶ a 4th neutrino with $\Delta m_{41}^2 \sim 1 \text{ eV}^2$
- ▶ a new type of CC-like interaction

$$\mathcal{L}_{\text{NSI}} = -2\sqrt{2}G_F \sum_{\alpha,\beta} \varepsilon_{\alpha\beta}^{ff'} (\bar{f} P_{L,R} \gamma^\mu f') (\bar{l}_\alpha P_L \gamma_\mu \nu_\beta) + h.c.$$

f, f' fermions depending on production or detection process

(NC-like NSI will have no relevant effect in short-baseline experiments since matter effect is very small)

Transition probability in (3+1) NSI

- CP viol and zero-distance effects:

$$\begin{aligned}
 P_{\alpha\beta}(L) = & \underbrace{4 \left[|\alpha_{\alpha\beta}|^2 - \text{Re}(\beta_{\alpha\beta}^* \alpha_{\alpha\beta}) \right] \sin^2(\Delta/2)}_{\text{similar to standard oscillations}} \\
 & + \underbrace{|\beta_{\alpha\beta}|^2}_{\text{zero-distance effect}} + \underbrace{2\text{Im}(\beta_{\alpha\beta}^* \alpha_{\alpha\beta}) \sin \Delta}_{\text{NSI-osc interference} \rightarrow \text{CP viol}}
 \end{aligned}$$

$\alpha_{\alpha\beta}, \beta_{\alpha\beta}$: functions of $U_{\alpha 4}, \varepsilon_{\alpha\beta}$

Source and detection processes

- ▶ CP viol and zero-distance effects
- ▶ can decouple LSND/KARMEN from the rest:

	source	detection
LSND/KARMEN	μ decay	ν -nucl CC (e)
MiniB/NOMAD	π decay	ν -nucl CC (e)
CDHS (atm)	π decay	ν -nucl CC (μ)
Bugey/Chooz	ν -nucl CC (e)	ν -nucl CC (e)

$$|\alpha_{\mu e}^{\text{LK}}|, |\beta_{\mu e}^{\text{LK}}|, \delta^{\text{LK}}, \alpha_e, \alpha_\mu, |\beta_{\mu e}|, \delta, \Delta m_{41}^2$$

→ 8 parameters (7 for 3+2 osc)

3+1 + NSI

CP viol, decouple LSND/KARMEN $\Rightarrow$

- ▶ $\chi^2_{(3+1)\text{osc}} - \chi^2_{(3+1)\text{NSI}} = 18.5(5 \text{ dof}) \rightarrow 99.76\% \text{ CL}$
appearance versus disappearance: PG compatibility of 15%
- ▶ $|U_{e4}| \approx 0.116$, $|U_{\mu 4}| \approx 0.205$ (more freedom from new react flux)
 $|\varepsilon_{\mu s}^{ud}| \approx 0.05$, $|\varepsilon_{e\mu}^{ud}| \approx 0.011$, $|\varepsilon_{\mu s}^{e\nu}| \approx 0.03$, $|\varepsilon_{\mu e}^{e\nu}| \approx 0.01$
strength of new interactions at the level of **few %** compared to G_F
in agreement with bounds Biggio, Blennow, Fernandez-Martinez, 0907.0097
- ▶ if due to dim-6 operator $\varepsilon G_F (\bar{f} f') (\bar{\ell} \nu)$ expect a charged mediator with $m_\phi \sim m_W / \sqrt{\varepsilon} \sim \text{TeV} \rightarrow$ excellent prospect at LHC
- ▶ **BUT:** seems difficult (impossible?) to obtain such operators at dim-6 in gauge invariant way $\rightarrow$ go to dim-8 and involve some fine tuning

Gavela, Hernandez, Ota, Winter, 0809.3451; Antusch, Baumann, Fernandez, 0807.1003

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- ▶ if due to dim-6 operator $\varepsilon G_F(\bar{f}f')(\bar{\ell}\nu)$ expect a charged mediator with $m_\phi \sim m_W/\sqrt{\varepsilon} \sim \text{TeV} \rightarrow$ excellent prospect at LHC
- ▶ **BUT:** seems difficult (impossible?) to obtain such operators at dim-6 in gauge invariant way $\rightarrow$ go to dim-8 and involve some fine tuning

Gavela, Hernandez, Ota, Winter, 0809.3451; Antusch, Baumann, Fernandez, 0807.1003

3+1 + NSI

CP viol, decouple LSND/KARMEN $\Rightarrow$

- ▶ $\chi^2_{(3+1)\text{osc}} - \chi^2_{(3+1)\text{NSI}} = 18.5(5 \text{ dof}) \rightarrow 99.76\% \text{ CL}$
appearance versus disappearance: PG compatibility of 15%
- ▶ $|U_{e4}| \approx 0.116$, $|U_{\mu 4}| \approx 0.205$ (more freedom from new react flux)
 $|\varepsilon_{\mu s}^{ud}| \approx 0.05$, $|\varepsilon_{e\mu}^{ud}| \approx 0.011$, $|\varepsilon_{\mu s}^{e\nu}| \approx 0.03$, $|\varepsilon_{\mu e}^{e\nu}| \approx 0.01$
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Energy dependent quantum decoherence

Farzan, TS, Smirnov 07

- ▶ **no sterile neutrino needed**
- ▶ postulate quantum-decoherence of the mass states $\Rightarrow$ damping of the interference terms in the oscillation probabilities.
- ▶ damping affects only oscillations with Δm_{31}^2 and rapidly decreases with the neutrino energy
- ▶ reconcile LSND signal with $MB\nu$ and other null-result exps. (signal in $MB\bar{\nu}$ requires CPT violation decoherence)
- ▶ standard explanations of solar, atmos, KamL, MINOS not affected
- ▶ LSND signal controlled by the 1-3 mixing: $0.0014 < \sin^2 \theta_{13} < 0.034$
- ▶ comparison of near and far measurements at reactors: null-result positive signal for θ_{13} in long-baseline accelerator experiments
- ▶ seems to be consistent with reactor anomaly!

Outline

SBL phenomenology

Active-sterile mixing

$$U_{e4}$$

$$U_{\mu 4}$$

$$U_{\tau 4}$$

Global data

More exotic proposals

Sterile neutrino + NSI

Soft decoherence

Conclusions

Conclusions

- ▶ There are a couple of "hints" for sterile neutrinos at the eV scale
- ▶ Unfortunately none of these "hints" is convincing by itself
- ▶ It is intriguing that they all point to a similar mass range
BUT, significant tension remains in the global fit
(non-observation of ν_μ disappearance)
- ▶ Are these sterile neutrinos consistent with cosmology?
(CMB, LSS, BBN) see talks by F. Villante, A. Melchiorri

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Thank you for your attention