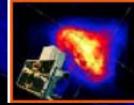


# The AGILE gamma-ray sky and the AGILE data center at ASDC

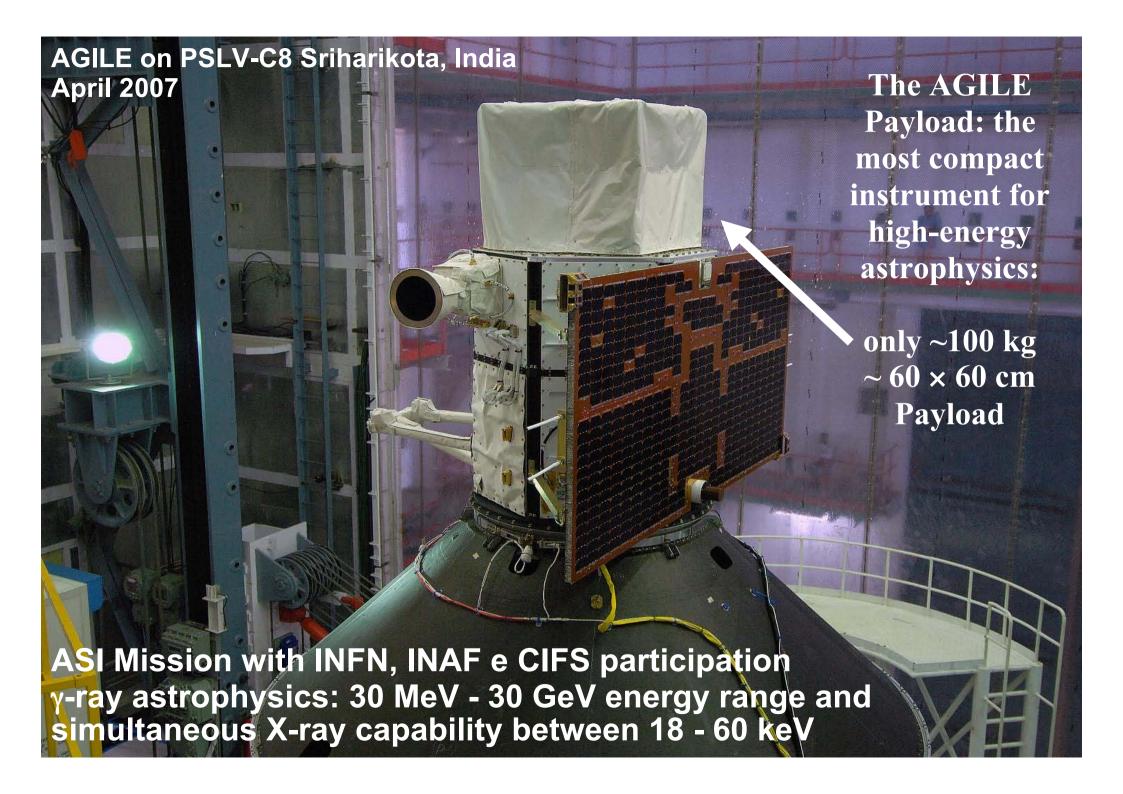


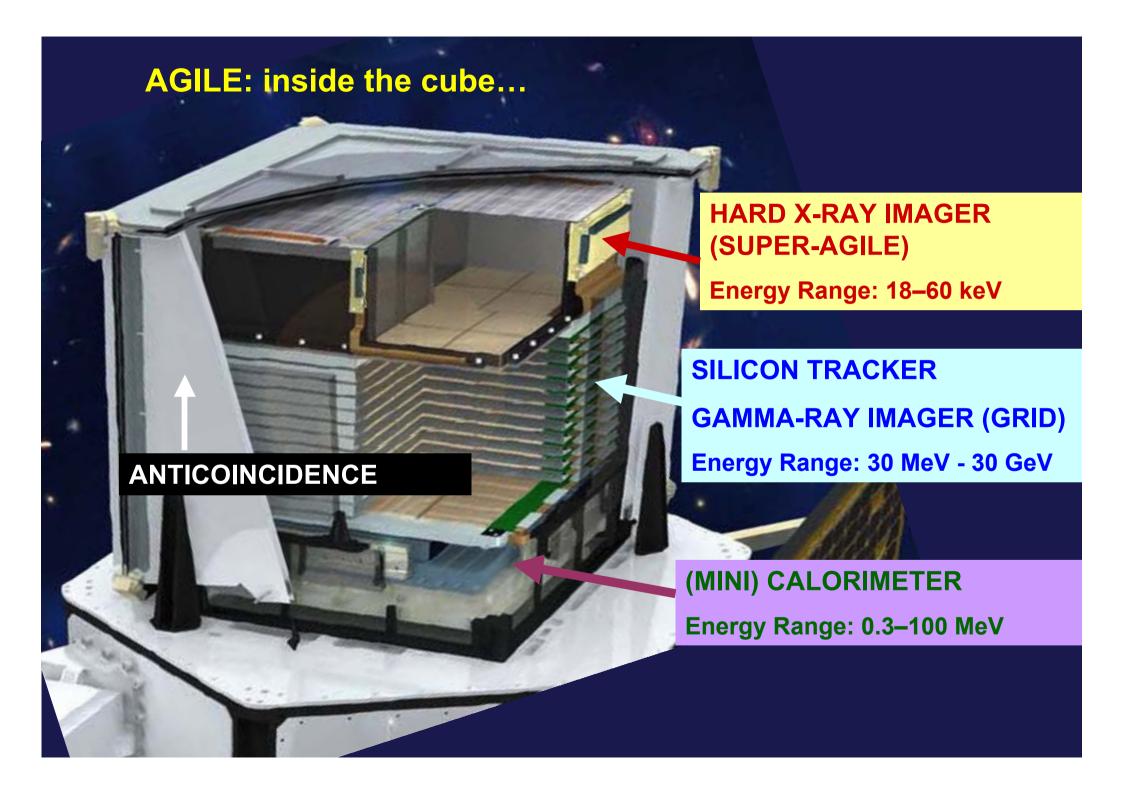
Bruno Rossi Prize 2012

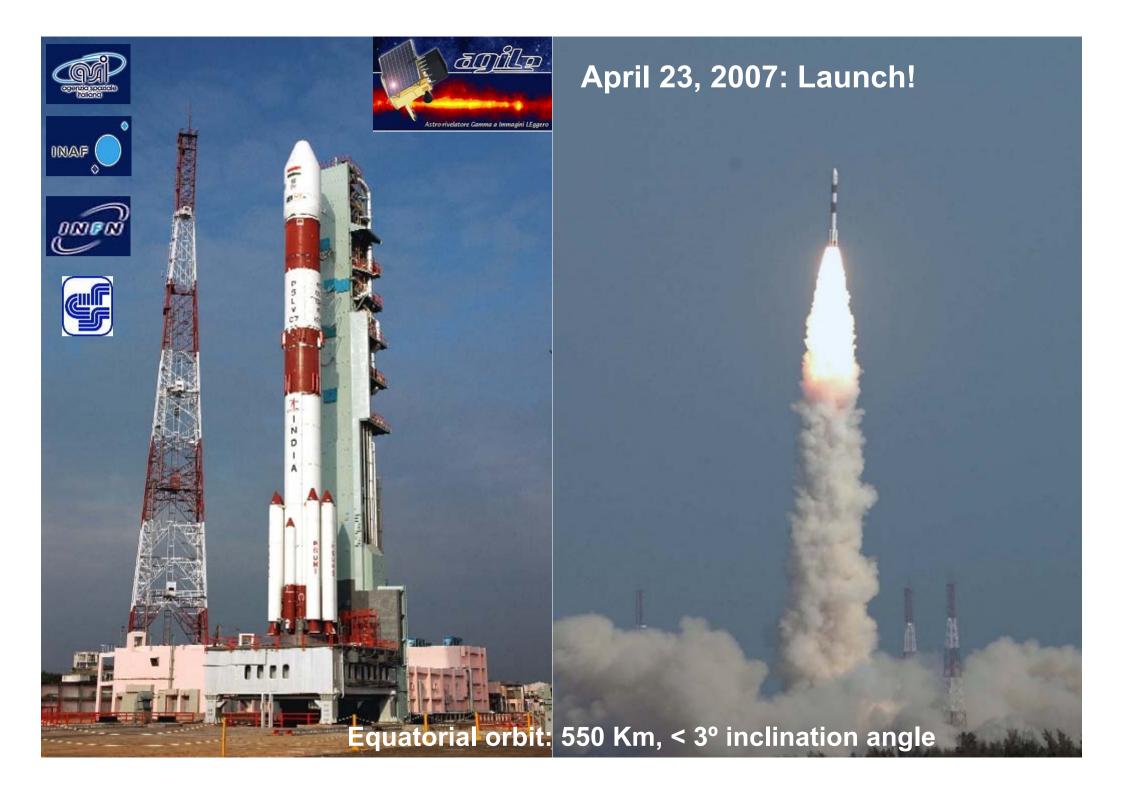
Marco Tavani and the AGILE team



Carlotta Pittori - ADC, on behalf of the AGILE Collaboration







### **AGILE** orbital parameters

Baseline equatorial orbit: 550 Km, 3° inclination

**Semi-major axis: 6922.5 km (± 0.1 km)** 

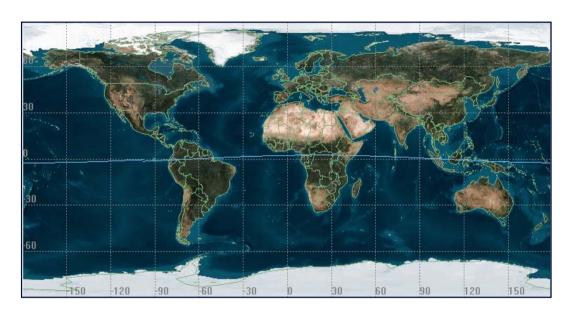
Requirement: 6928.0 ± 10 km

Inclination angle: 2.48° (±0.04°)

Requirement: < 3°

**Eccentricity:** 0.002 (±0.0015)

Requirement: < 0.1°



### **TPZ** orbital decay estimate:

Height < 400Km on 20/04/2017

(A/M=0.009 sqm/Kg)

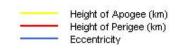
Worst case (A/M=0.012 sqm/Kg): 02/11/2015 -> 450

Best case (A/M=0.006 sqm/Kg): 29/04/2023

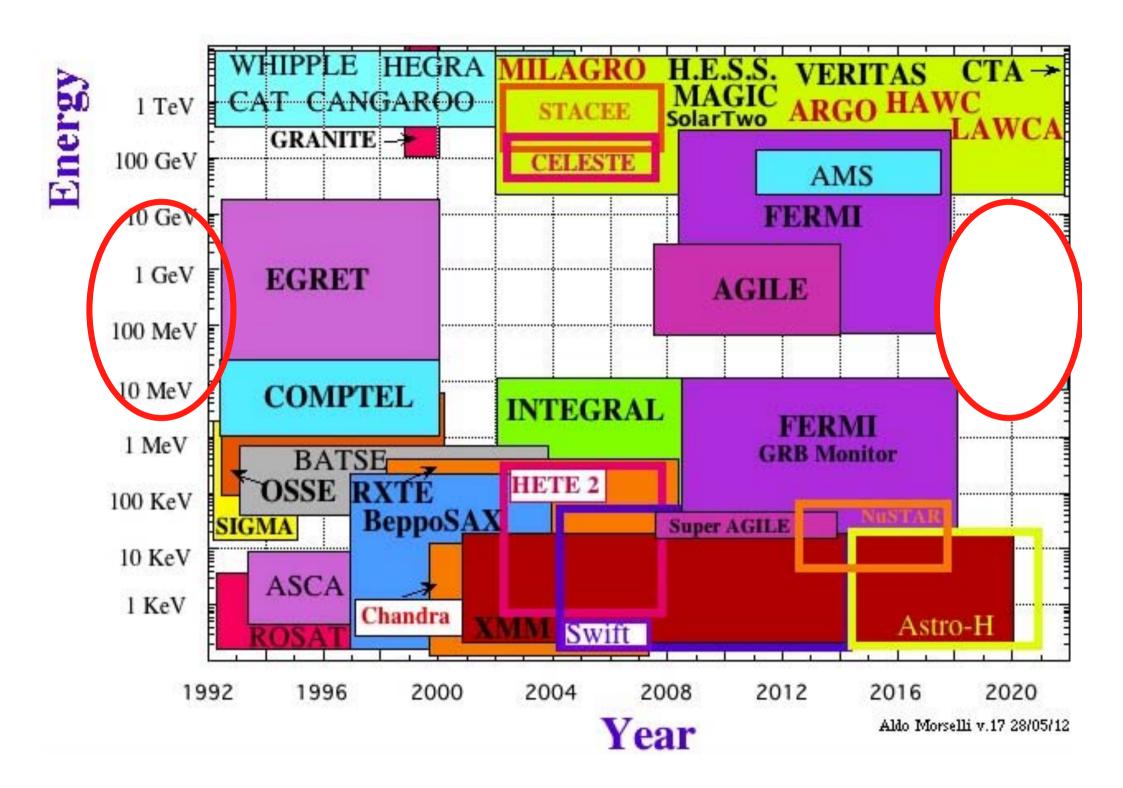
0.0026 0.0024 0.0022 0.0020 0.0020 0.0018 0.0018 0.0016 15 Mar 2013 00:00:00.000 29 Jul 2014 06:54:49.710 12 Dec 2015 13:49:39.420 Time (UTCG)

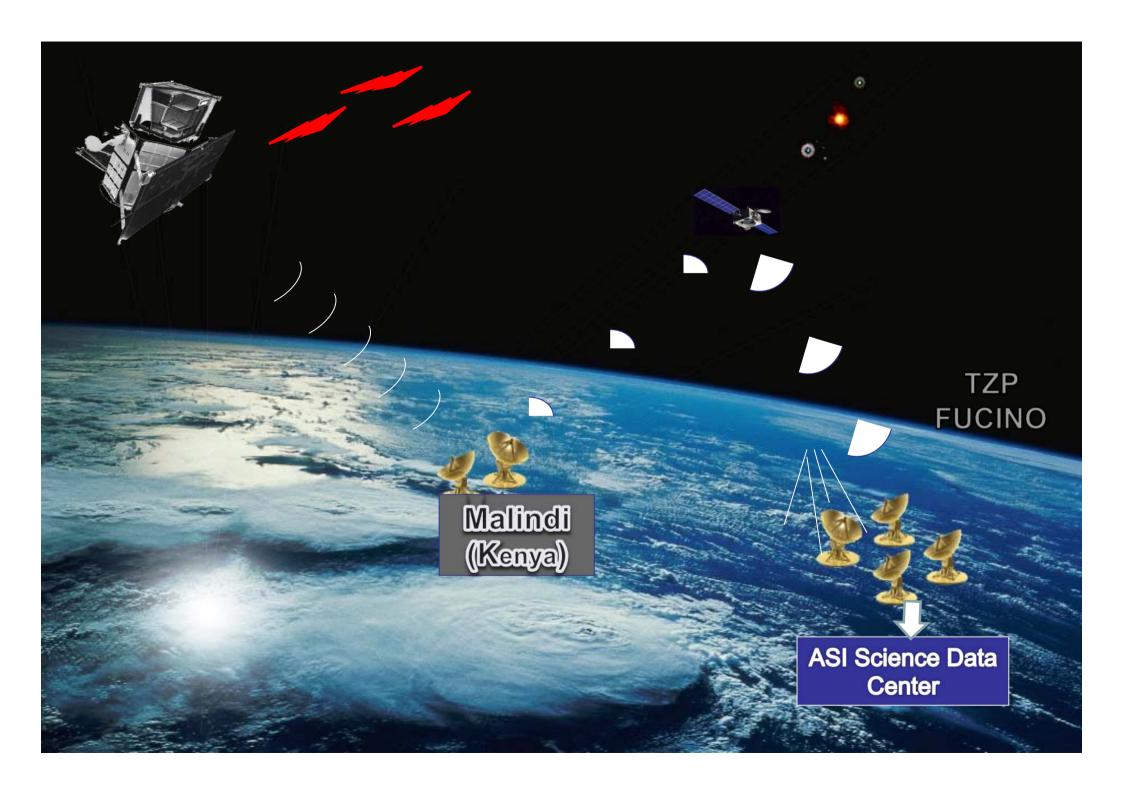
Satellite-AGILE - 28 Mar 2013 10:08:13

(March 2013 updated estimate, using recent solar flux "Schatten" forecasts  $+ 2\sigma$ )



0 0028





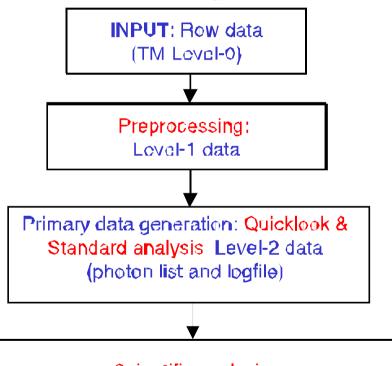


• The ADC, based at ASDC-ESRIN, is in charge of all the scientific oriented activities related to the analysis and archiving of AGILE

data:

From scientific telemetry (TM) Level–0:

- ✓ Preprocessing → Level-1 data
- ✓ Quick-Look Analysis (transient detection)
- ✓ Standard analysis → Level-2 data (photon list)
- ✓ Scientific analysis (source detection, diffuse gamma-ray background)
- ✓ Archiving and distributing all scientific AGILE data



### Scientific analysis:

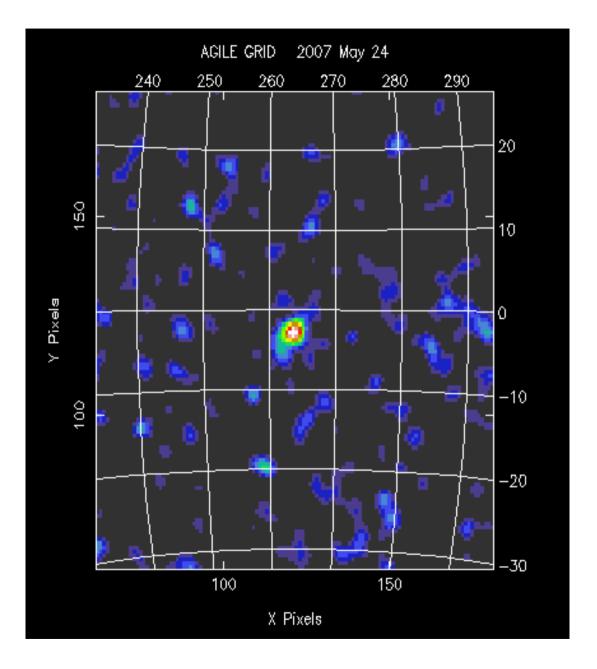
Level-3 data

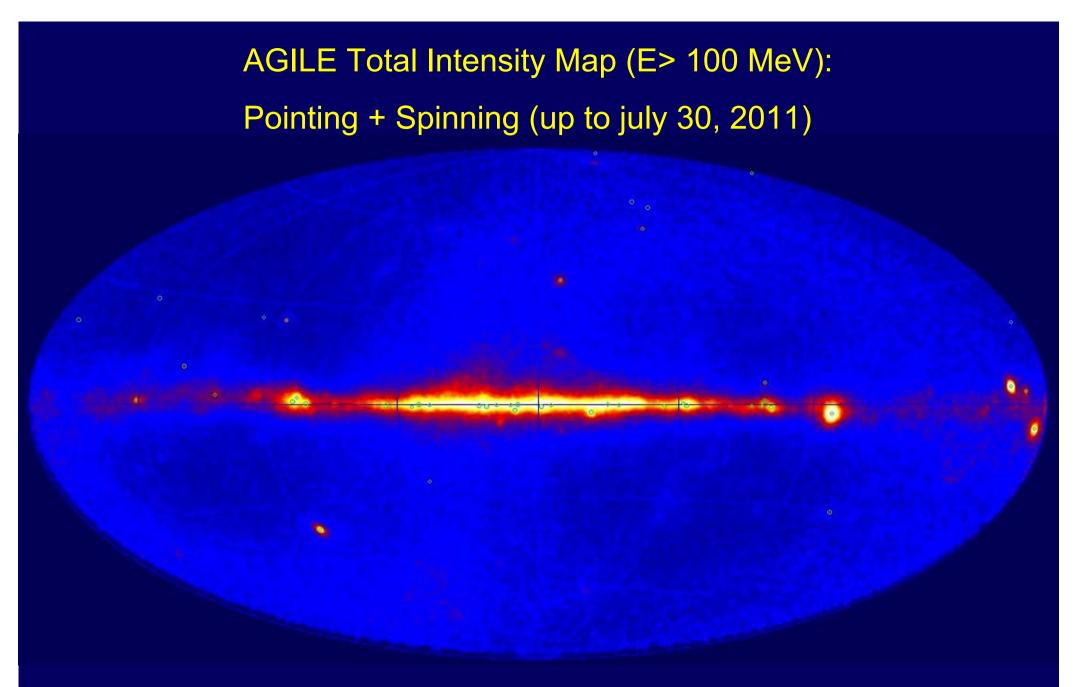
OUTPUT: High level data products
(count maps, spectra, light curves...)

## First AGILE GRID light ADC 24/5/2007

Commissioning Phase: AGILE Vela PSR Count Map

(~ 20000 s)





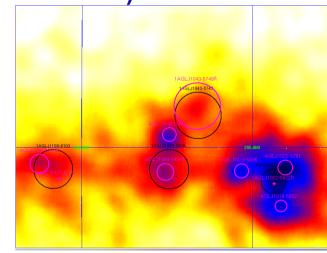
"The First AGILE-GRID Catalog of High Confidence Gamma-Ray Sources" C. Pittori et al., A&A 506, 2009 (green circles, first year of operations)

### **WORK IN PROGRESS:**

• A revised AGILE bright gamma--ray sources list and its variability study in pointing mode (F. Verrecchia et al., submitted to A&A)

Variability study of an improved 1AGL source list (54 sources) on the timescale of the AGILE pointed observations (Observation Blocks)

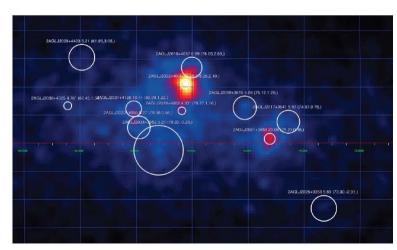
Refined positioning of some 1AGL sources: the Carina region →

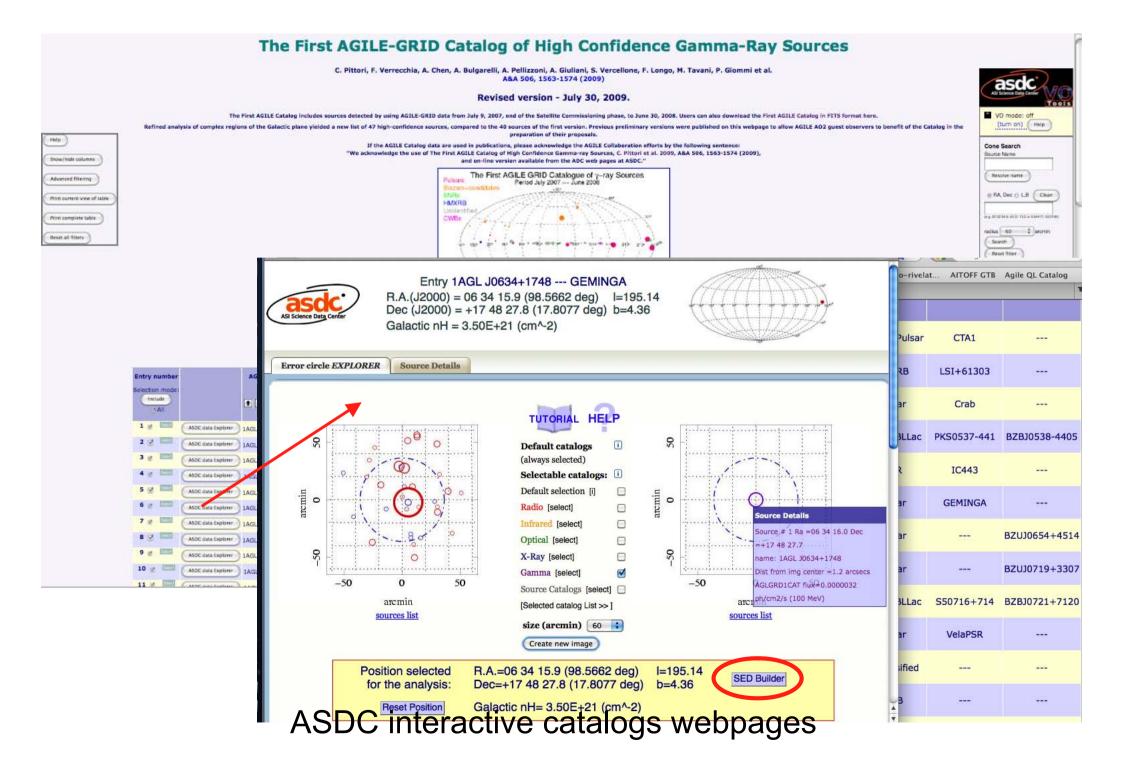


• The second AGILE Catalog (A. Bulgarelli et al.)

New AGILE-GRID source catalog over the whole period of AGILE pointed observations (2.3 years), with improved event filter and updated calibrations.

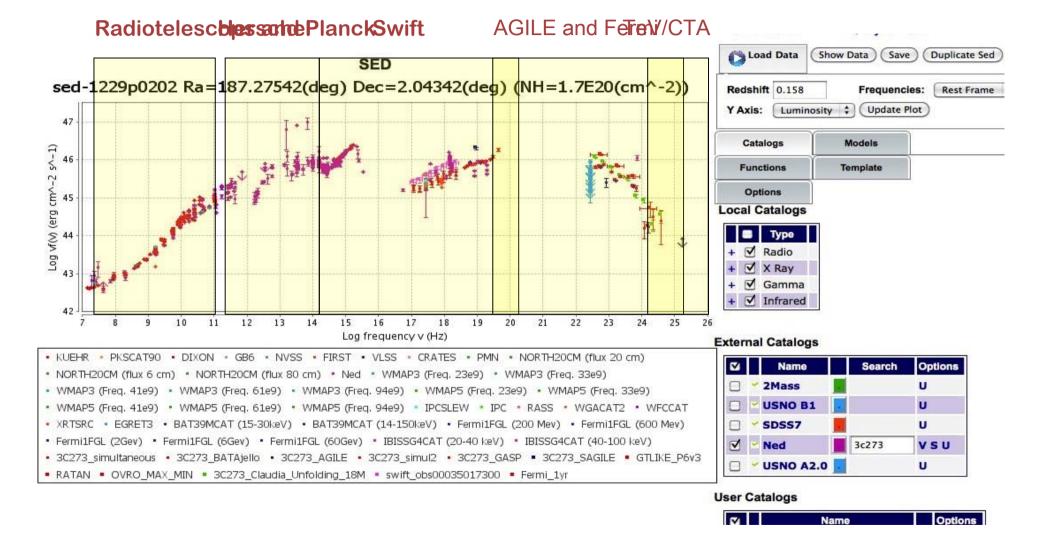
More than 180 sources on the galactic plane only: the Cygnus region →



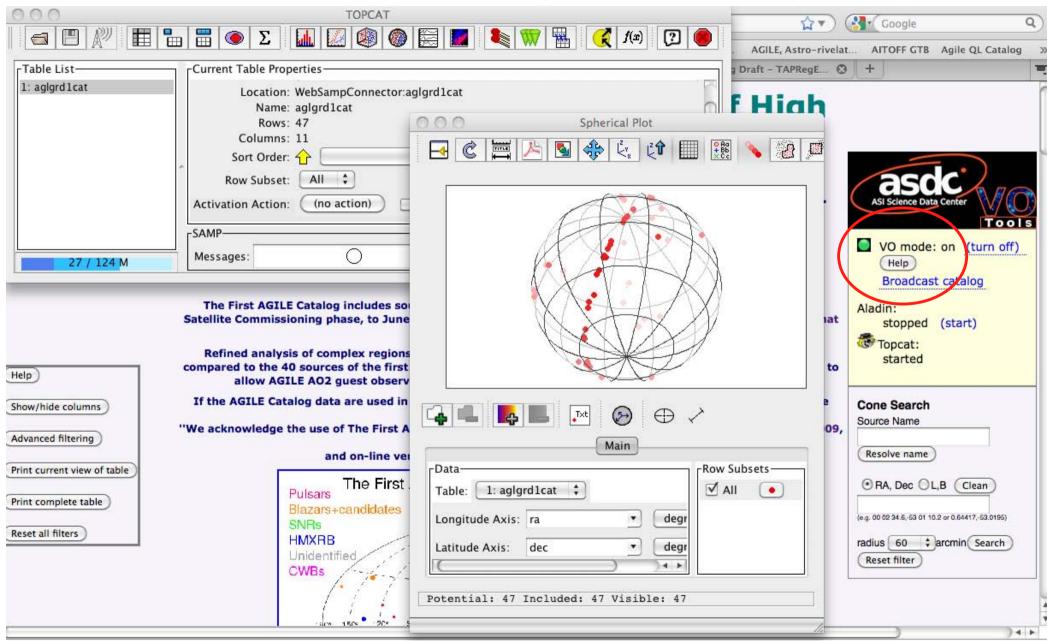


### The ASDC SED Builder

## New SED(t) v2.2: VO tools and TIME domain



# Virtual Observatory Standards (*in progress*) and Tool for OPerations on Catalogues And Tables (**Topcat**)



### The AGILE MCAL Gar

**NEW: MCAL GRB Catalog** 

(M. Galli et al., 2013)

### Swift-XRT light curves of GRB 090510

Last updated after receiving ObsID 00351588001, version 19

Related pages: Burst Analyser | Enhanced position | Spectrum | GRB Region information | XRT Catalogue entry | Download obs data | GCN Notices | GCN Circulars

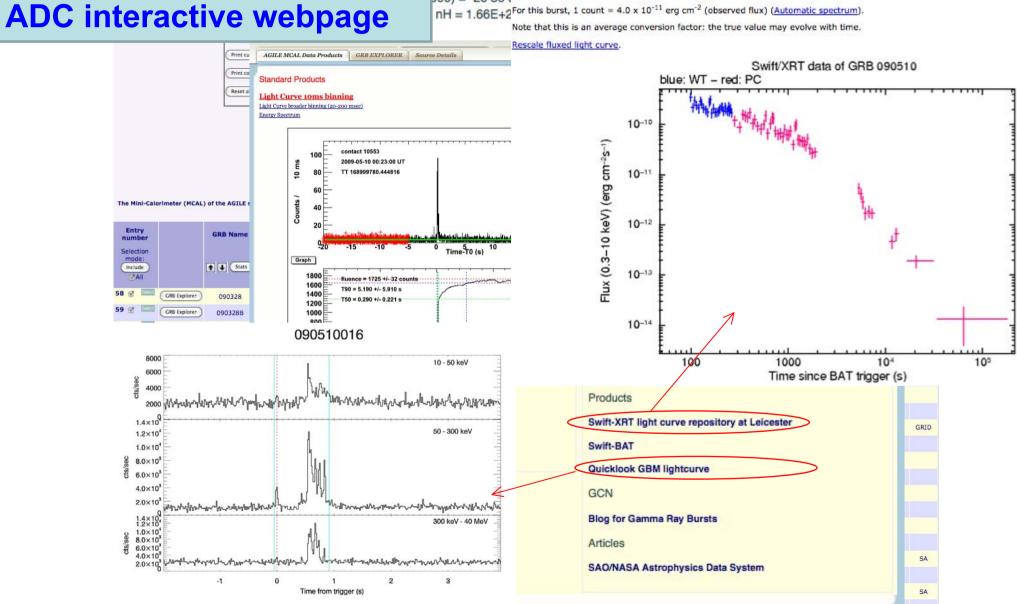
Rebin this light curve | About these products.

Ent

)00) = 22 14 1: Flux Light Curve

000) = -2636(

nH = 1.66E+2For this burst, 1 count = 4.0 x 10<sup>-11</sup> erg cm<sup>-2</sup> (observed flux) (Automatic spectrum).

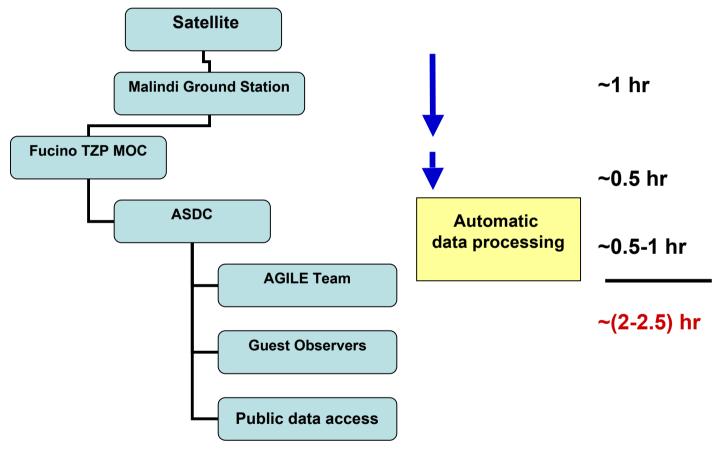


### **AGILE:** 6th year in orbit

- AGILE demonstrates for the first time the covering of  $\sim$  1/5 of the entire gamma-ray sky (FoV  $\sim$  2.5 sr) with excellent angular resolution and competitive sensitivity.
- AGILE shows for the first time an optimal performance of its gamma-ray and hard X-ray imagers.
- > 31280 orbits, May 15, 2013
- Pointing observation mode up to October 18, 2009 and spinning observation mode since October 2009.
- Very good scientific performance, especially at ~ 100 MeV
- Guest Observer Program open to the scientific community:

```
Cycle-1: completed, Dec. 1, 2007 – Nov 30, 2008
Cycle-2: completed, Dec. 1, 2008 – Nov 30, 2009
Cycle-3: completed, Dec. 1, 2009 – Nov 30, 2010
Cycle-4: completed, Dec. 1, 2000 – Nov 30, 2011
Cycle-5 and Cycle-6: on-going data taking
```

# AGILE: "very fast" Ground Segment (with contained costs)

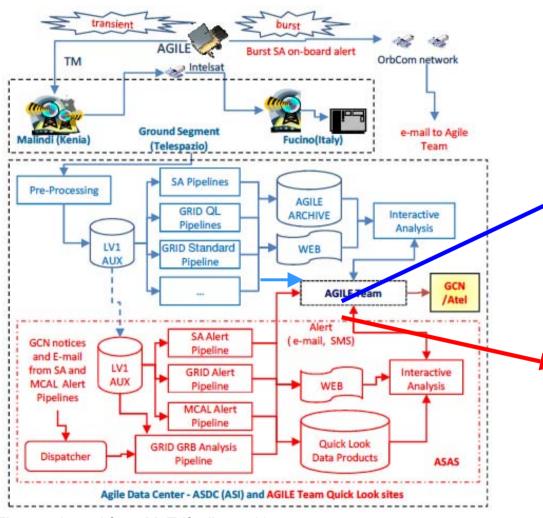


Record for a gamma-ray mission!

### **AGILE Science Alert System**

- The system is distributed among the ADC @ ASDC and the AGILE Team Institutes (Trifoglio, Bulgarelli, Gianotti et al.)
- Automatic Alerts to the AGILE Team are generated within T<sub>0</sub> + 45 min (SA) and T<sub>0</sub> + 100 min (GRID)
- GRID Alerts are sent via email (and sms) both on a contact-by-contact basis and on a daily timescale
- Refined manual analysis on most interesting alerts performed every day (daily monitoring)
- 104 ATel (48 in pointing + 56 in spinning) and 40 GCN published up to May, 2013

### Selected alerts sent via email, sms

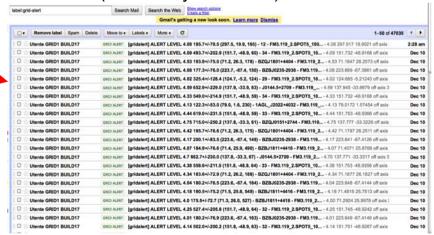


(Figure adapted from M. Trifoglio et al.)

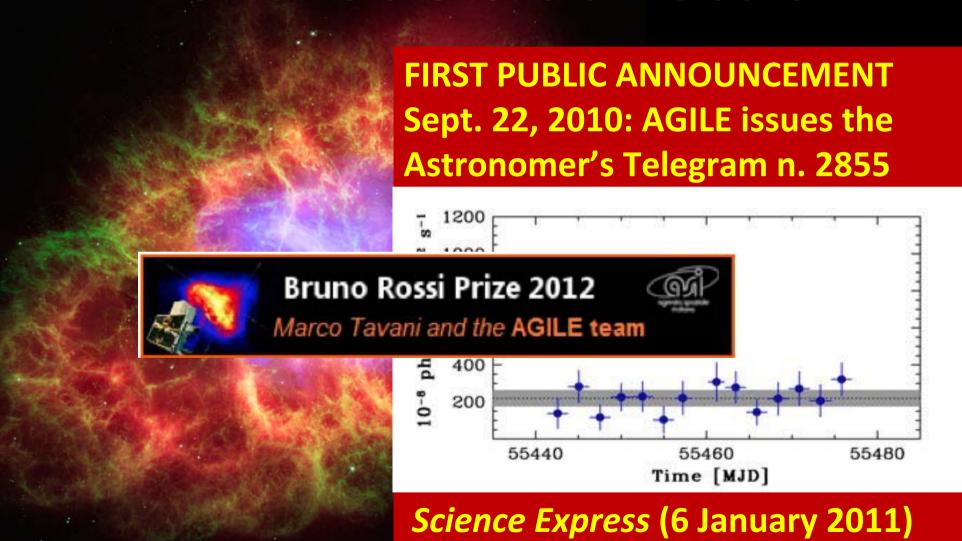


Daily reports on a 48h time scale (sent twice a day)

Contact-by-contact alerts on a 48h time scale (sent every ~100 min)



## The variable Crab Nebula!





#### First AGILE catalog of high-confidence gamma-ray sources

C. Pittori<sup>1</sup>, F. Verrecchia<sup>1</sup>, A. W. Chen<sup>2,3</sup>, A. Bulgarelli<sup>4</sup>, A. Pellizzzoni<sup>5</sup>, A. Giuliani<sup>2,3</sup>, S. Vercellone<sup>6</sup>, F. Longo<sup>7,8</sup>, M. Tavani<sup>9,10,11,3</sup>, P. Giommi<sup>1,12</sup>, G. Barbiellini<sup>7,8,3</sup>, M. Trifoglio<sup>4</sup>, F. Gianotti<sup>4</sup>, A. Argan<sup>9</sup>, A. Antonelli<sup>4</sup>, F. Boffelli<sup>4</sup>, P. Caraveo<sup>2</sup>, P. W. Cattaneco<sup>4</sup>, V. Cocco<sup>6</sup>, S. Colafrancesco<sup>4,12</sup>, T. Contess<sup>2</sup>, E. Costa<sup>8</sup>, Z. E. Costa<sup>8</sup>, A. Farmando<sup>9,10</sup>, E. Del Monte<sup>6</sup>, G. De Paris<sup>6</sup>, G. Di Cocco<sup>6</sup>, G. Di Persic<sup>6</sup>, I. Donnarumma<sup>8</sup>, Y. Evangellista<sup>8</sup>, G. Fanari<sup>1</sup>, M. Feroci<sup>9</sup>, A. Ferrari<sup>1,15</sup>, M. Forini<sup>2</sup>, F. Fornari<sup>2</sup>, F. Fuschino<sup>4</sup>, T. Froysland<sup>4,11</sup>, M. Frutti<sup>9</sup>, M. Britti<sup>1</sup>, M. Garbielli<sup>8</sup>, D. Gasparimi<sup>1</sup>, C. Labani<sup>4</sup>, I. Lapshov<sup>9,17</sup>, F. Lazzarotto<sup>9</sup>, F. Liello<sup>9</sup>, P. Lipila<sup>1,8,19</sup>, E. Mattanin<sup>2</sup>, M. Marstopietro<sup>2,2</sup>, A. Morselli<sup>11</sup>, L. Pacciami<sup>9</sup>, F. Perotti<sup>2</sup>, G. Piano<sup>9,10,11</sup>, P. Piccozzal<sup>0,11</sup>, M. Pilia<sup>2,2,2</sup>, C. Pontoni<sup>3,8</sup>, G. Porrovecchio<sup>9</sup>, B. Preger<sup>1</sup>, M. Press<sup>2,2</sup>, R. Primwera<sup>1</sup>, G. Puccella<sup>9</sup>, M. Agpisarda<sup>2,0</sup>, A. Rappoldi<sup>11</sup>, F. Rossi<sup>4</sup>, A. Rubin<sup>9</sup>, S. Sabatini<sup>10</sup>, P. Santolamazza<sup>1</sup>, E. Scalise<sup>9</sup>, P. Soffitta<sup>9</sup>, S. Stellato<sup>1</sup>, E. Striani<sup>10</sup>, F. Tamburelli<sup>1</sup>, A. Traci<sup>4</sup>, A. Traci<sup>8</sup>, E. Vallazza<sup>2,8</sup>, V. Vittorio<sup>13</sup>, A. Zambara<sup>3</sup>, D. Zanellolla<sup>18,9</sup>, and Labotti<sup>12</sup>

### Sect. 6.1 Notes on individual

1AGL J0535+2205 and 1AGL J0634+1748 (Crab and Geminga). These two well known strong  $\gamma$ -ray pulsars, together with the Vela pulsar, were used for in-flight AGILE calibrations. We report the flux values obtained during calibration subperiods. These values agree with pulsed flux values reported in (Pellizzoni et al. 2009). We note, however, that we observed higher flux values, over  $1\sigma$  from the reported mean flux, for both sources when merging all the data, including shorter (1 day) integration periods during 2007. This point is under investigation.



IIIaiviaaai		ω -			-
Flare date	Duration	Peak γ-ray flux	Instruments		1
October 2007	~ 15 days	~ 6·10 <sup>-6</sup> ph cm <sup>-2</sup> s <sup>-1</sup>	AGILE	1 11	1 1
February 2009	~ 15 days	~ 4·10 <sup>-6</sup> ph cm <sup>-2</sup> s <sup>-1</sup>	Fermi	1 1	† ‡
September 2010	~ 4 days	~ 5·10 <sup>-6</sup> ph cm <sup>-2</sup> s <sup>-1</sup>	AGILE, Fermi		
April 2011	~ 2 days	~ 30·10 <sup>-6</sup> ph cm <sup>-2</sup> s <sup>-1</sup>	Fermi, AGILE	4380	54400

March 2013: **new** γ-ray flaring state detected by Fermi and AGILE

 a big theoretical challenge: the Crab Nebula is not a standard candle in gamma-rays!



# AGILE: Gamma-Rays MAIN DISCOVERIES AND SURPRISES!

### **UNEXPECTED NEWS FROM THE γ-RAY SKY:**

# AGILE DISCOVERY OF THE CRAB NEBULA VARIABILITY IN γ-RAYS

Tavani et al., <u>Science</u>, 331, 736 (2011)

### Fermi confirmation:

AGILE updates: see Edoardo Striani TALK "The Crab Nebula: observations and simulations" on Thu, May 23, Parallel Session E at 15:55

 Carina region: γ-ray detection of the colliding wind massive binary system η-Car with AGILE

Tavani et al., ApJ, 698, L142, 2009 (arXiv:0904.2736)

Cygnus region microquasars:

Cyg X-1 updates: see Sabina Sabatini TALK on Thu, May 23, Parallel Session G at 16:45

Cyg X-3 updates: see Giovanni Piano TALK on Thu, May 23, Parallel Session G at 17:00

Tavani et al., Nature 462, 620, 2009 (arXiv:0910.5344)

 Detection of Gamma-Ray Emission from the Vela Pulsar Wind Nebula with AGILE

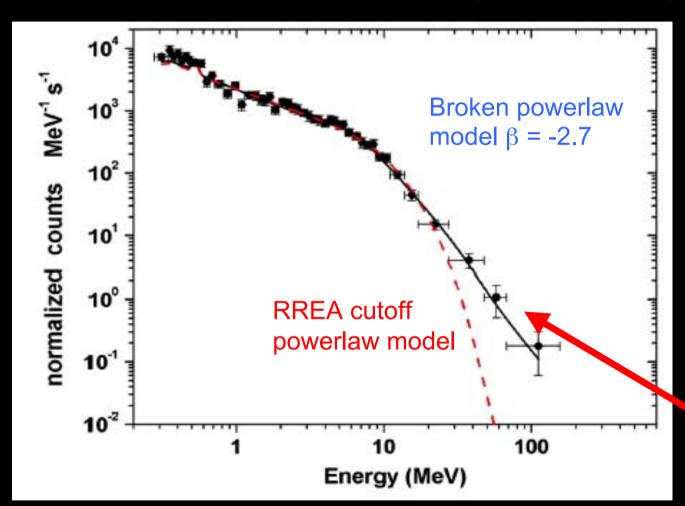
W44 updates: see Martina Cardillo TALK on Wed, May 22, Parallel Session B at 15:15

CITO OTHER TOTAL

Giuliani et al., ApJ, 742, 2011

## TGF Cumulative spectrum

110 TGFs 1806 photons 142  $\gamma$  E> 10 MeV 26  $\gamma$  E> 20 MeV



Significant detection of γ >40 MeV!!
Uneplained by standard RREA model: challenge for emission models

Tavani et al., Phys. Rev. Letters 106, 018501 (2011)

Slide adapted from M. Marisaldi, 10th AGILE WS

AGILE-MCAL crucial spectral contribution up to 100 MeV!!

- Normal lightnings involve a potential difference ~ 500 kVolts
- Terrestrial Gamma-Ray Flashes (TGF) involve DV > 100 MVolts!
- Models??: Relativistic Runaway Electron Avalanche (RREA) with relativistic feedback (Dwyer 2008). Bremsstrahlung + Compton scattering. *Much theoretical work in progress*
- RHESSI cumulative spectrum compatible with a production altitude of 15-21 km (just above tropical thunderstorms)

### **AGILE MCAL:** an optimal detector for TGF

- MCAL energy range is extended up to 100 MeV
- Efficient trigger at ms and sub-ms time scale (the TGF time scale)
- AGILE equatorial orbit at 2.5° inclination is optimal for mapping the equatorial region, where most of the events take place
- A real-time monitoring and alert system can be implemented for correlation with other meteo resources (work in progress)

### Agile Services: user pittori



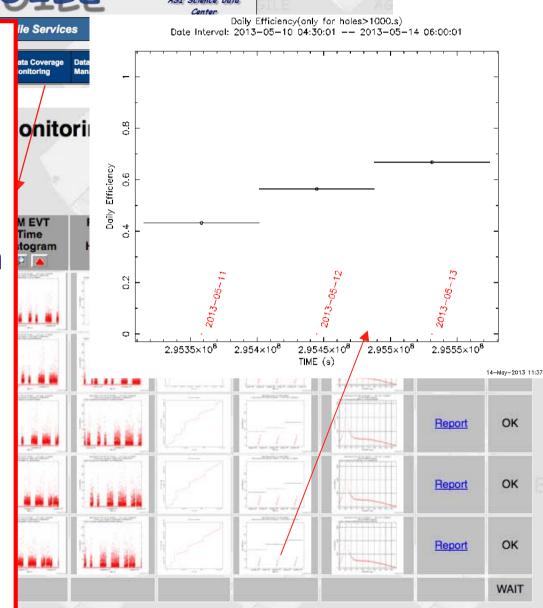




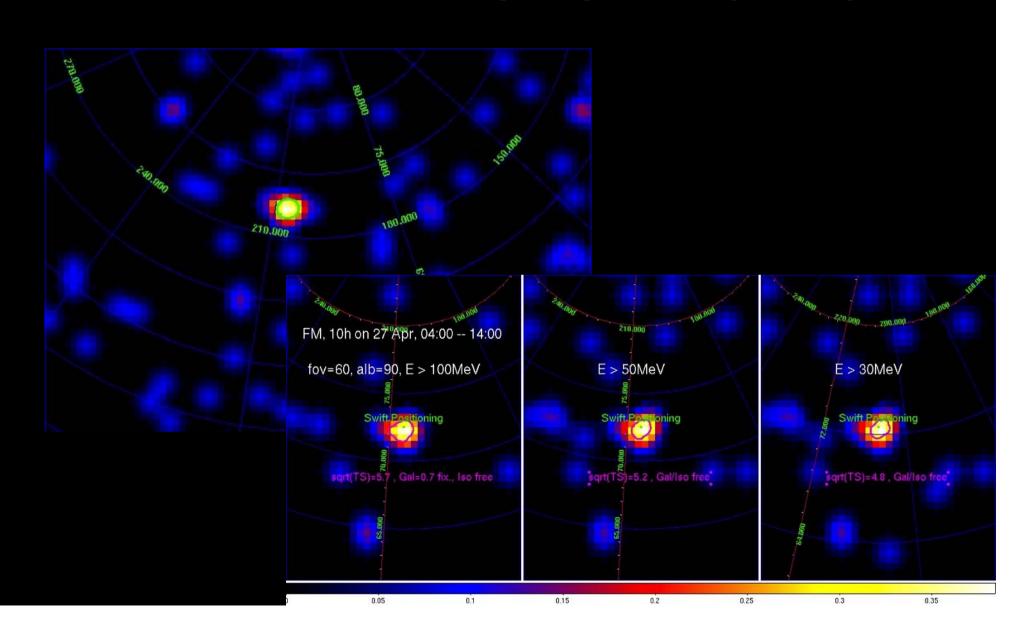
# Malindi Ground station problems impact on AGILE telemetry download

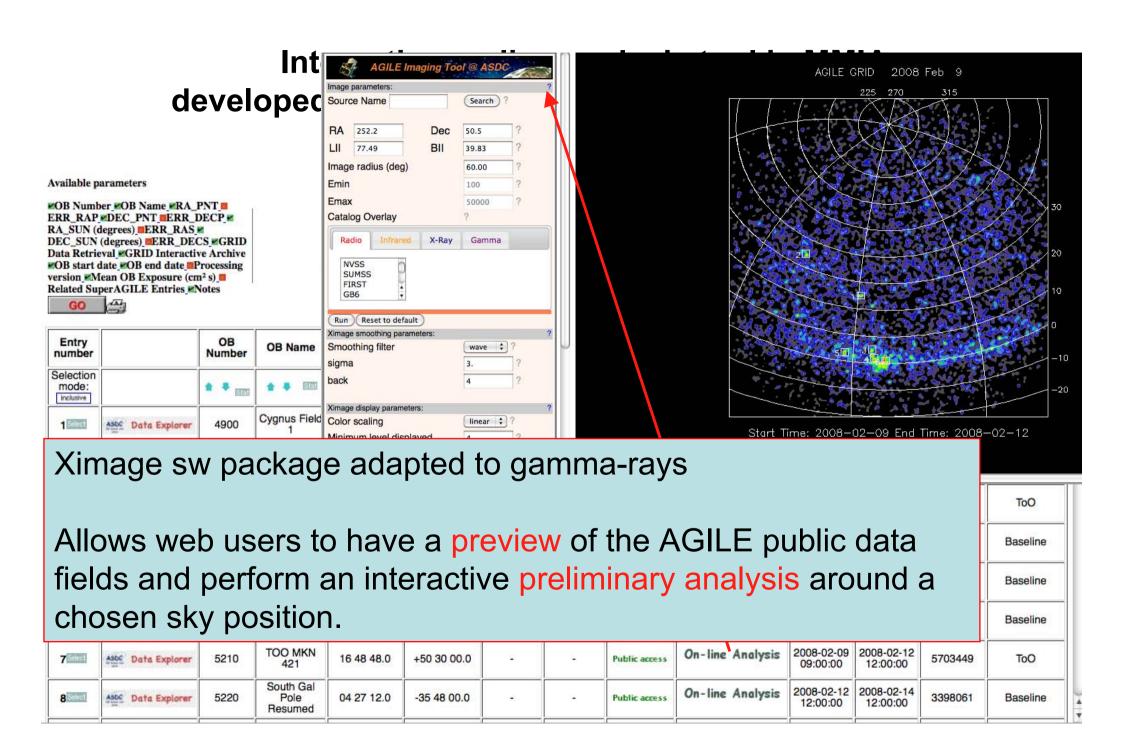
Due to ASI's Malindi ground station technical problems, the acquisition of telemetry data from the AGILE satellite has been significantly reduced since December 21, 2012.

All AGILE payload functions are nominal, and normal operations will be resumed as soon as the Malindi antenna system problems are solved.



# Despite recent low Malindi dowlink efficiency, AGILE detected GRB 130427A: the most energetic gamma-ray burst yet!





Warning: use imaging tool only as a preview of the AGILE γ-ray field. To perform your own scientific analysis, please download data and use the official public AGILE software available at: http://agile.asdc.asi.it/public/ following the AGILE Software User Manual

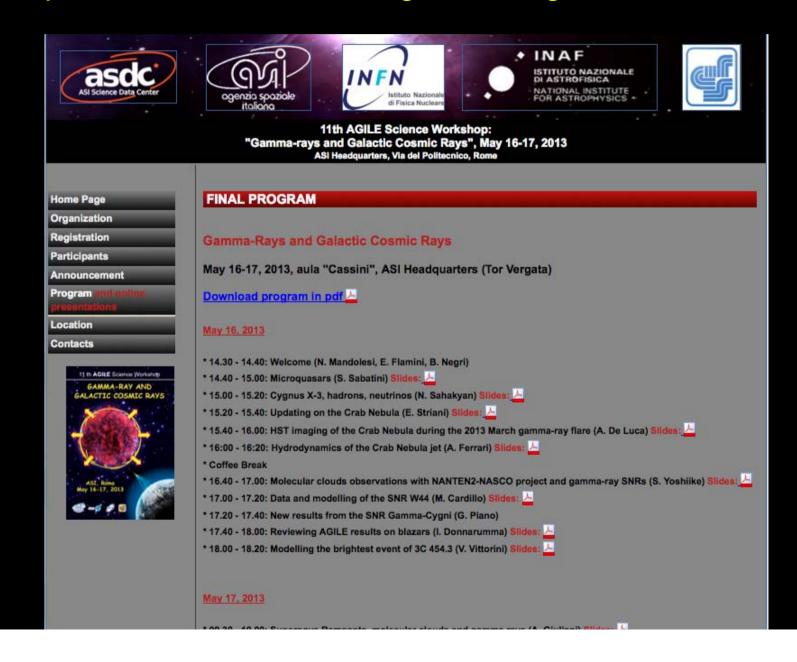
### Index of /public/AGILE\_SW\_5.0\_SourceCode

Icon	Name	Last modified	Size	Description
[DIR]	Parent Directory	\$1000 1-000000 F.CCC 1420	-	
[ ]	AGILE-IFC-OP-009 Build-21.pdf	22-Nov-2011 18:24	928K	
[ ]	BUILD GRID 5.0.tgz	22-Nov-2011 16:56	121M	
[TXT]	SoftwareReleaseNote 5.0.txt	25-Nov-2011 16:01	16K	
[TXT]	readme 5.0.txt	22-Nov-2011 16:57	5.2K	
[ ]	test dataset 5.0.tgz	22-Nov-2011 16:57	346M	

Apache Server at agile.asdc.asi.it Port 80

NEW: web interface for interactive on-line ML analysis on AGILE on legacy (LV3) data archive under construction!! Stay tuned!

# 11<sup>th</sup> AGILE Science Workshop: on-line presentations from http://www.asdc.asi.it/11thagilemeeting/



### **BACKUP SLIDES**

Table 3: AGILE Scientific Performance

Gamma-ray Imaging Detector (GRID)		
Energy Range	30 MeV - 50 GeV	96
Field of view	$\sim 3 \ \mathrm{sr}$	
Sensitivity at 100 MeV (ph cm <sup>-2</sup> s <sup>-1</sup> MeV <sup>-1</sup> )	6×10 <sup>-9</sup>	$(5\sigma \text{ in } 10^6 \text{ s})$
Sensitivity at 1 GeV (ph cm <sup>-2</sup> s <sup>-1</sup> MeV <sup>-1</sup> )	4×10-11	$(5\sigma \text{ in } 10^6 \text{ s})$
Angular Resolution at 1 GeV	36 arcmin	(68% cont. radius
Source Location Accuracy	~5-20 arcmin.	S/N~10
Energy Resolution	$\Delta E/E\sim 1$	at 300 MeV
Absolute Time Resolution	$\sim 1~\mu s$	
Deadtime	~ 200 µs	
Hard X-ray Imaging Detector (Super-AGII	LE)	
Energy Range	10 - 40  keV	
Field of view	107°×68°	FW at Zero Sens.
Sensitivity (at 15 keV)	∼5 mCrab	(5σ in 1 day)
Angular Resolution (pixel size)	∼ 6 arcmin	10.000.000.000.000
Source Location Accuracy	~2-3 arcmin	S/N~10
Energy Resolution	$\Delta E < 4 \text{ keV}$	1000 C 1000 C 1000
Absolute Time Resolution	$\sim 4  \mu s$	
Deadtime (for each of the 16 readout units)	$\sim 4\mu s$	
Mini-Calorimeter		
Energy Range	0.3 - 200 MeV	60 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5
Energy Resolution	$\sim 1~{ m MeV}$	above 1 MeV
Absolute Time Resolution	~ 3 µs	
Deadtime (for each of the 30 CsI bars)	$\sim 20\mu\mathrm{s}$	

### **WORK IN PROGRESS**

### Systematic search for GeV counterparts of TeV sources

LONGO, F., LUCARELLI, F., PITTORI, C., RAPPOLDI, A., VERRECCHIA, F.

- Automated search for E>100 MeV sources spatially correlated with known TeV sources, using GRID Pointing data (2.3 yrs) (ADC public data archive).
- Data analysis criteria: multi-source Maximum Likelihood (ML) analysis around the TeV position
- Input catalogue: revised source list built from the TeVCat catalogue (see TGevCat Catalogue @ ASDC).

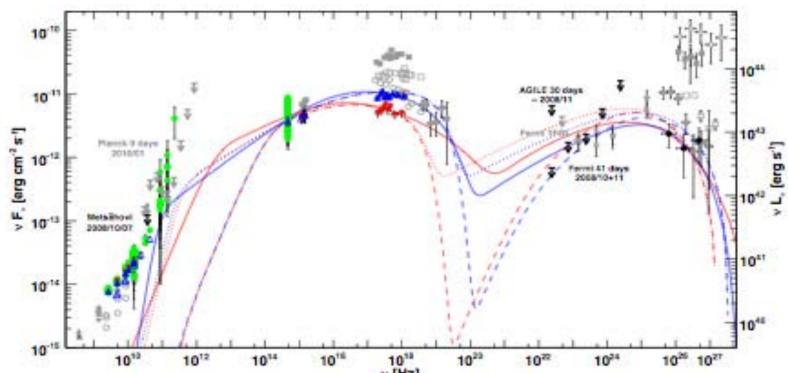
Source Class	TeV Catalog	Detections sqrt(TS>4)
AGN (HBL, LBL,)	45	10
Starburst	2	
PWN	24	8
SNR	14	5
XRB	3	1
UNID	25	6
Other Galactic	3	1

- ~30% of TeV sources show counterparts in AGILE first 2.3yrs data.
- Few GRID detections (1-2 HBL, few Galactic) not in 2FGL Cat.
- Upgrade is ongoing (latest calib. matrices (I0023) and updated TeV input list).

Rappoldi et al. (2009), Lucarelli et al. (2011)

### MWL campaigns on TeV sources

1ES2344+514 (Rugamer et al., arXiv:1211.2608v1)



More to come: PKS1510 (Feb. '12 flare), Mrk180, 1ES1011+496.

<u>GRID Level 3 Archive</u> (1-day integration EXP, COUNTS and GAS maps) helps to investigate flaring activity over > 5 years of data in a very short computing time.