

# HIGGS PHYSICS: WHAT WE LEARNED WHAT WE CAN STILL HOPE TO LEARN

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les Rencontres de Physique  
de la Vallée d'Aoste  
La Thuile, March 1st, 2013

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ICREA, IFAE  
Barcelona

# FIRST GLIMPSES OF HIGGS' FACE

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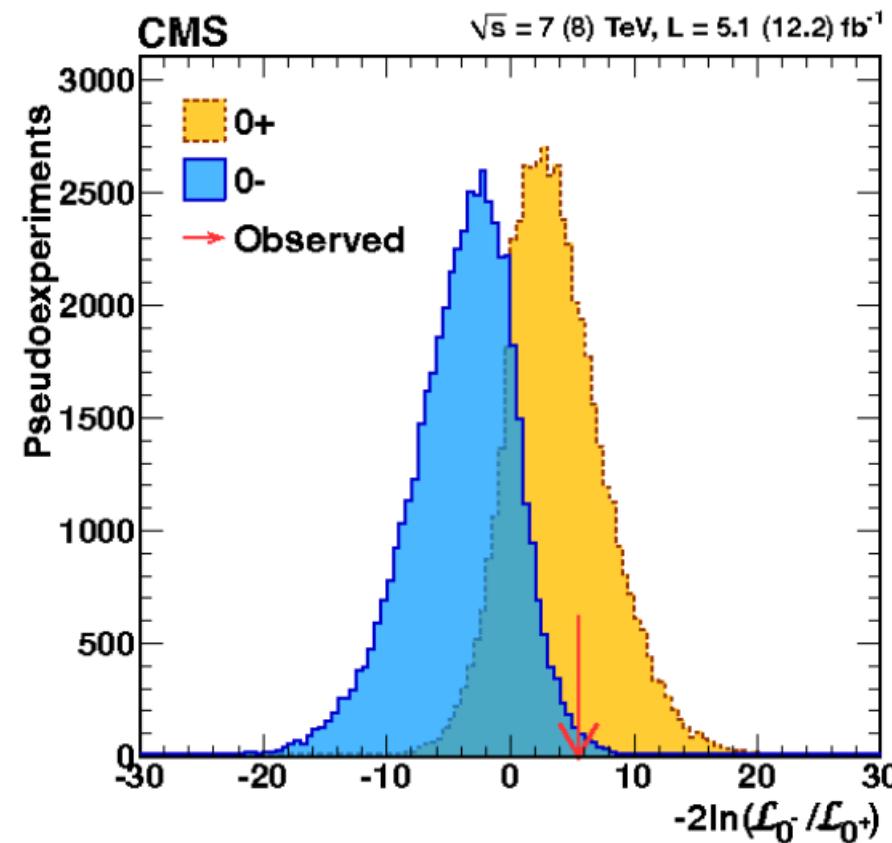
# HIGGS BOSON

$h^0$

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$h^0$

$J^P = 0^+$



$0^-$  excluded @  
97% CL

# HIGGS BOSON

$h^0$

$J^P = 0^+$

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## $h^0$ MASS

VALUE (GeV)

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*DOCUMENT ID*

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$125.2 + 0.3(\text{stat}) + 0.6 (\text{syst})$

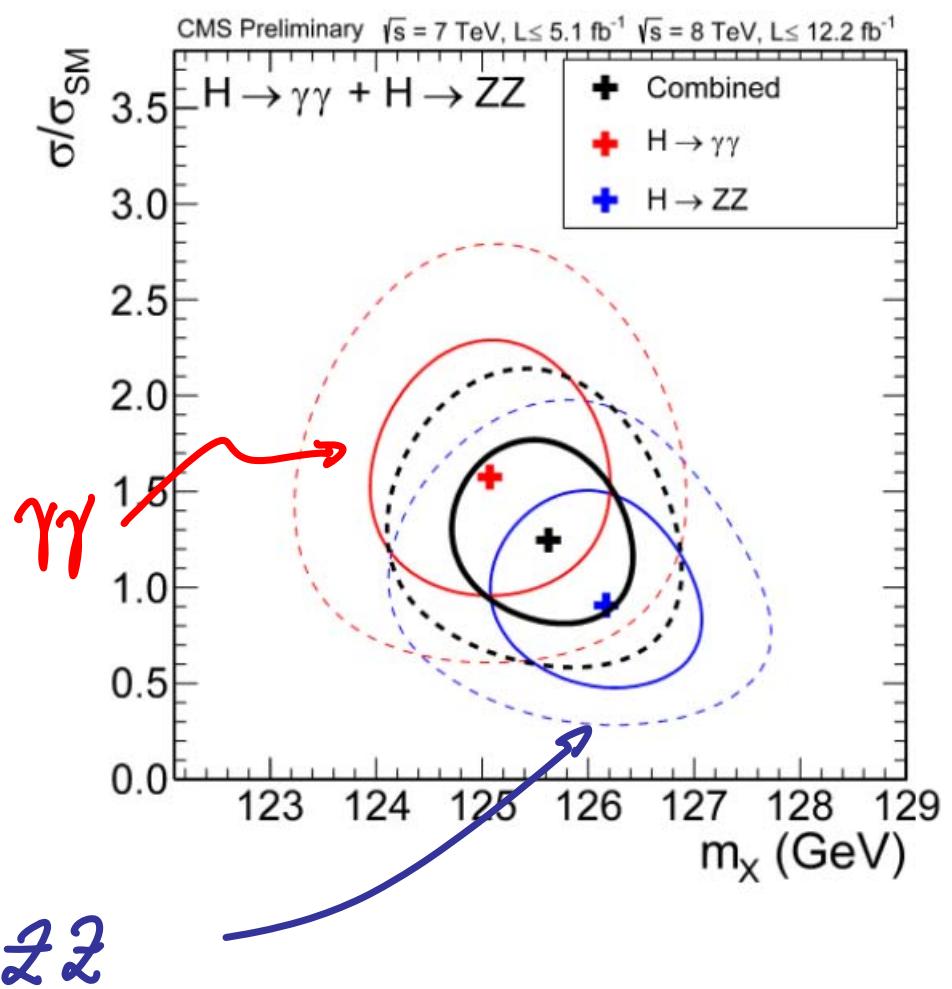
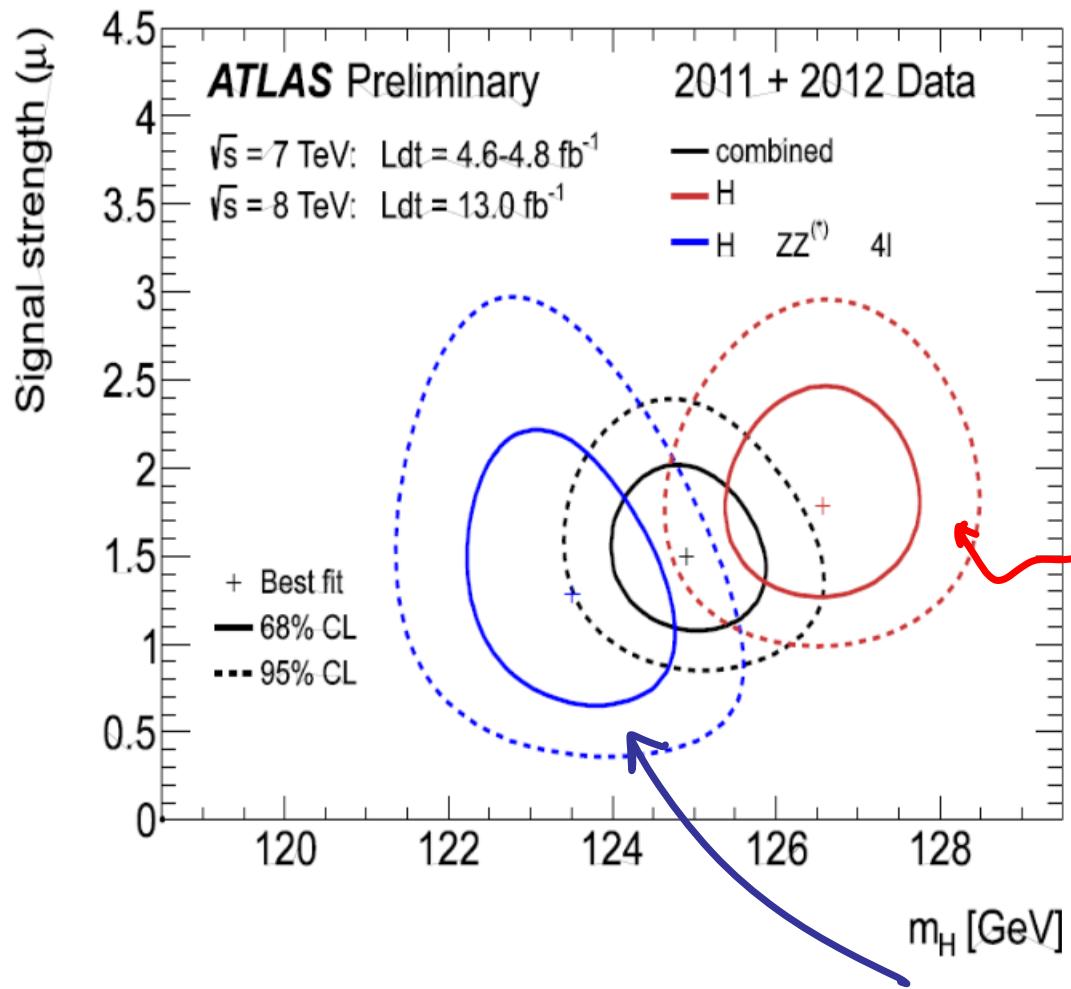
ATLAS-CONF-2012-170

$125.3 + 0.4(\text{stat}) + 0.5 (\text{syst})$

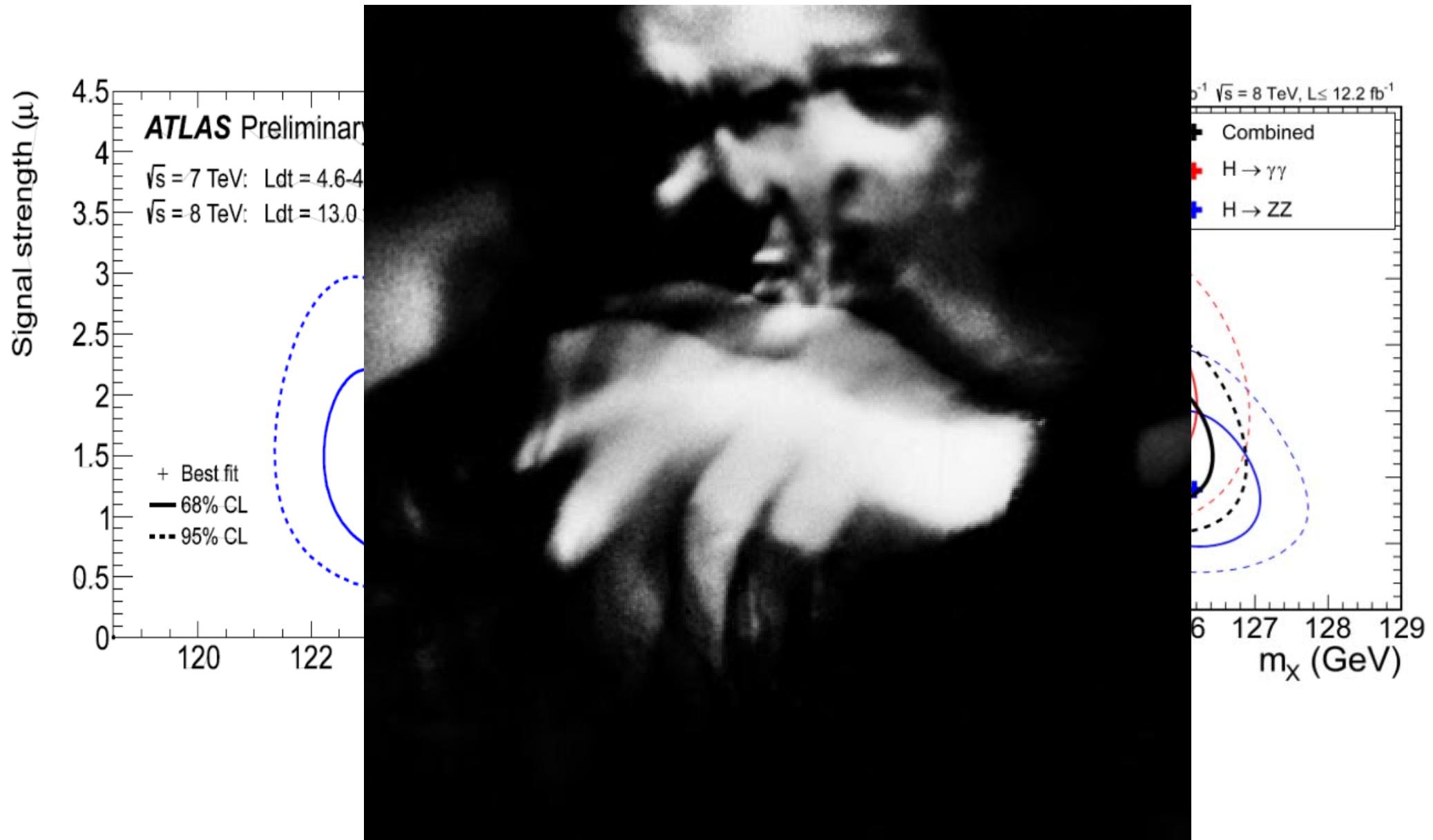
CMS-HIG-12-028

# MASS DETERMINATION

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# HIGGS BOSON

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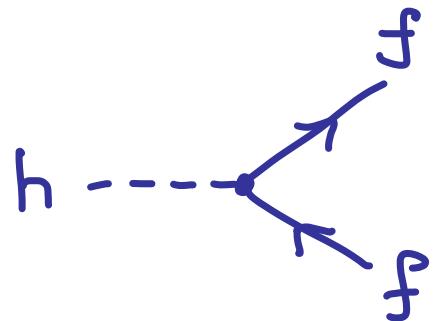
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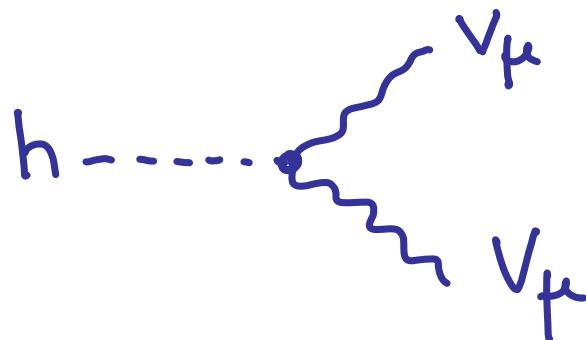
## $h^0$ couplings

# HIGGS COUPLINGS SEEN SO FAR

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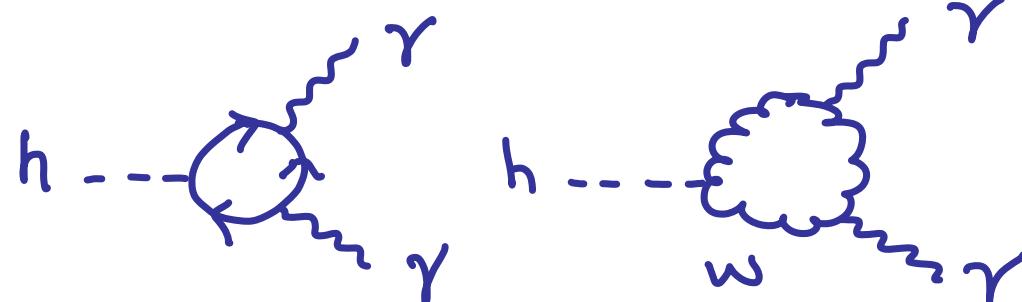
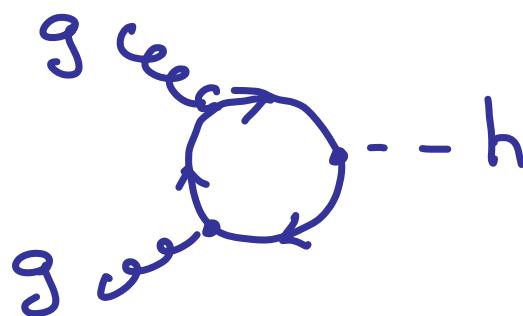


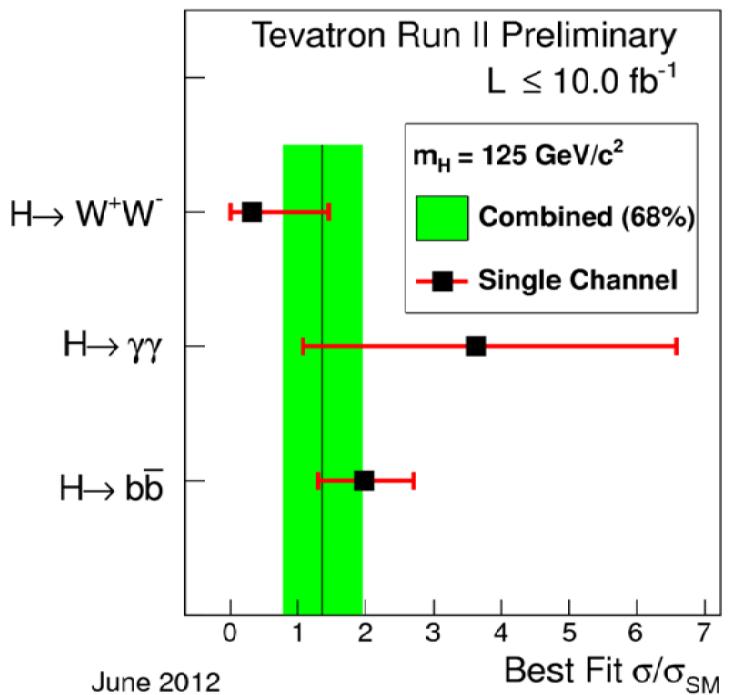
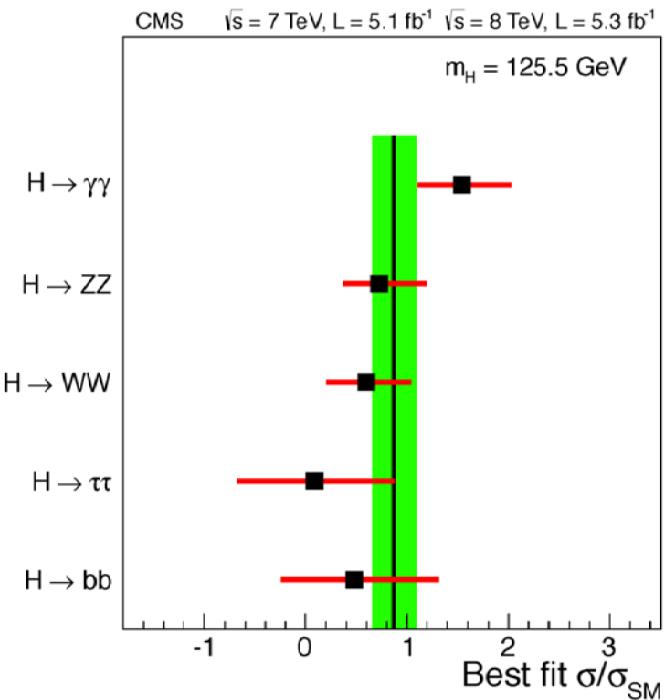
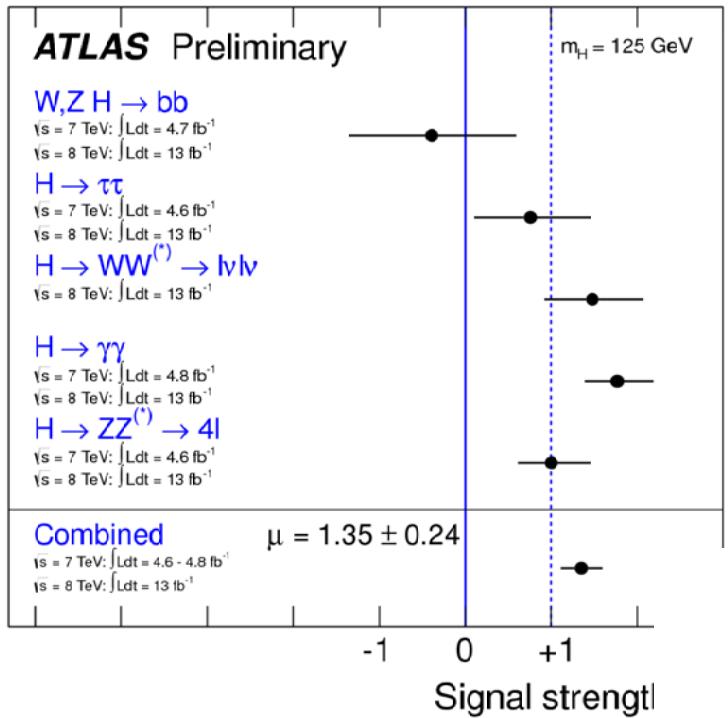
$$f = \begin{cases} \text{top (indirectly)} \\ \text{bottom} \\ \text{tau} \end{cases}$$



$$\nu_\mu = \begin{cases} w \\ z \end{cases}$$

LOOP INDUCED





See next talks  
by Rescigno &  
Azzurri.

Values close to SM-like .

But wait, The Higgs is no ordinary particle !

- \* We're seeing the first non-gauge interactions
- \* We might be seeing the first spin  $\phi$  fundamental particle !
- \* We want to learn about the mechanism behind electroweak symmetry breaking !
- \* From that perspective, some quantities are more important than others:
  - Mass value
  - Important
  - Determining  $J^P$ /Discarding  $J=2$  Less so
  - Precise measurement of couplings
  - Crucial

# WHY COUPLINGS MATTER

SM Higgs sector is the less tested and more problematic



Affected by  
hierarchy problem

Calls for  
new physics at  
the TeV scale

It's very likely that the Higgs will  
depart from its SM properties

The importance of measuring Higgs couplings:  
window to natural new physics

# LEARNING ABOUT THE HIGGS BOSON

EWSB

1. Implications from  $M_h \sim 125$  GeV  
→ Potential Instability in SM

2. Global fits to Higgs Data  
BSM Hopes

→ Deviations from SM couplings

For recent work see :

- Stability.

M. Holthausen, K.S. Lim, M. Lindner [hep-ph/1112.2415]

F. Bezrukov, M.Y. Kalmykov, B.A. Kniehl, M. Shaposhnikov [hepph/1205.2893]

J. Elias-Miró, J.R.E., G.F. Giudice, G. Isidori, A. Riotto, A. Strumia  
+ H.M. Lee + G. Degrassi, S. Di Vita

[hep-ph/1112.3022], [hep-ph/1203.0237], [hep-ph/1205.6497]

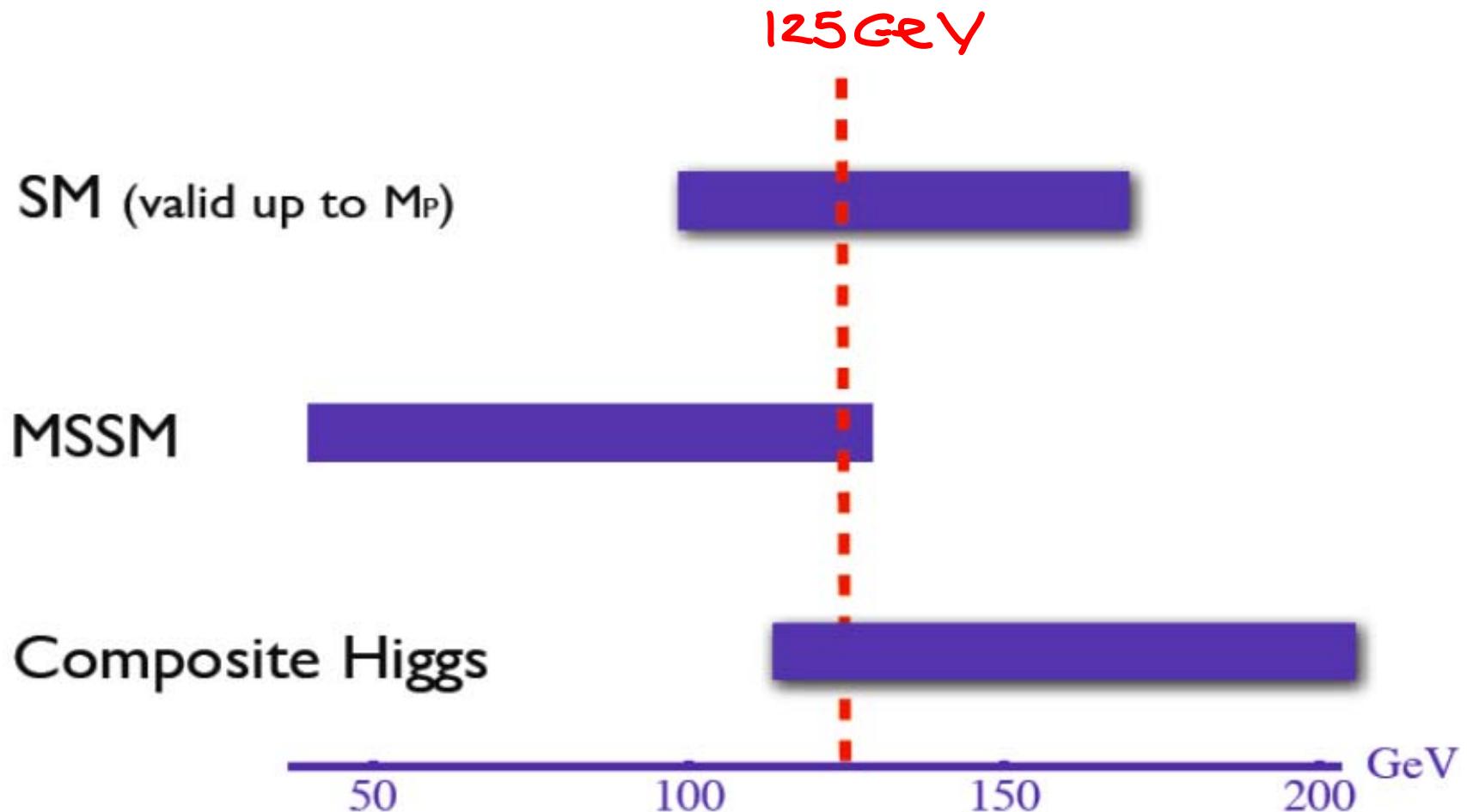
For recent work see :

- Global Fits to Higgs Data

- Carmi, Falkowski, Kuflik, Volansky [1202.3144][1206.4201]  
+ Zupan [1207.1718] Azatov, Contino, Galloway [1202.3415][1206.3171]  
Giardino, Kannike, Raidal, Strumia [1203.4254][1207.1347]  
Ellis, Tso [1204.0464][1207.1693]  
Klute, Lafaye, Plehn, Rauch, Ternas [1205.2699]  
Plehn, Rauch [1207.6108] Low, Lykken, Shaughnessy [1207.1093]  
Corbett, Eboli, González-Fraile, González-García [1207.1344][1211.4580]  
Montull, Riva [1207.1716], Belanger et al. [1212.5244][1302.5694]  
Cacciapaglia et al. [1211.4580] +...  
J.R.E., C.Grojean, M. Muhlleitner, M.Trott  
[hep-ph/1202.3697], [hep-ph/1205.6790], [hep-ph/1207.1717]

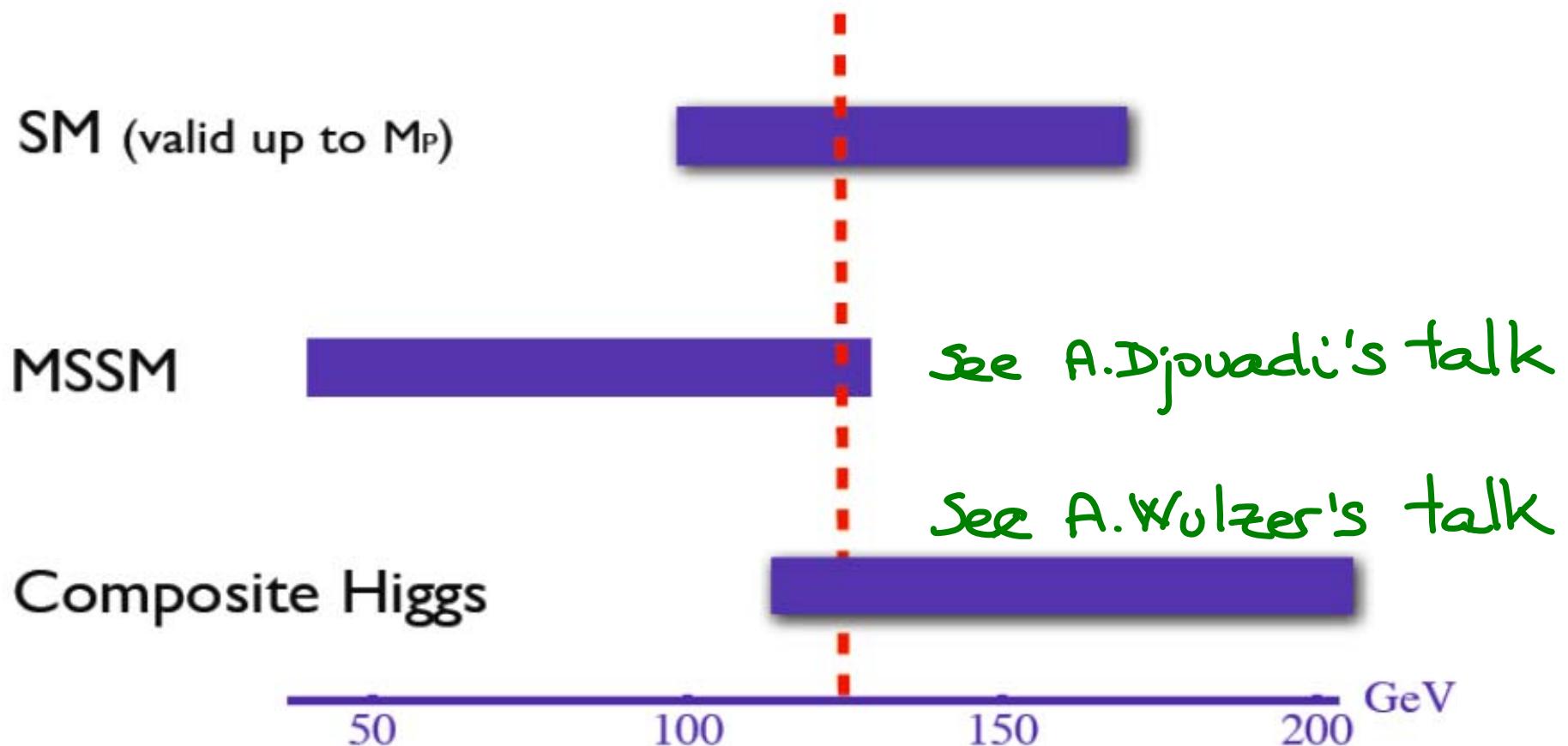
# $M_h$ AS MODEL DISCRIMINATOR

## Higgs mass range



# $M_h$ AS MODEL DISCRIMINATOR

## Higgs mass range



# ① $M_H \approx 125$ GeV. IMPLICATIONS FOR STABILITY

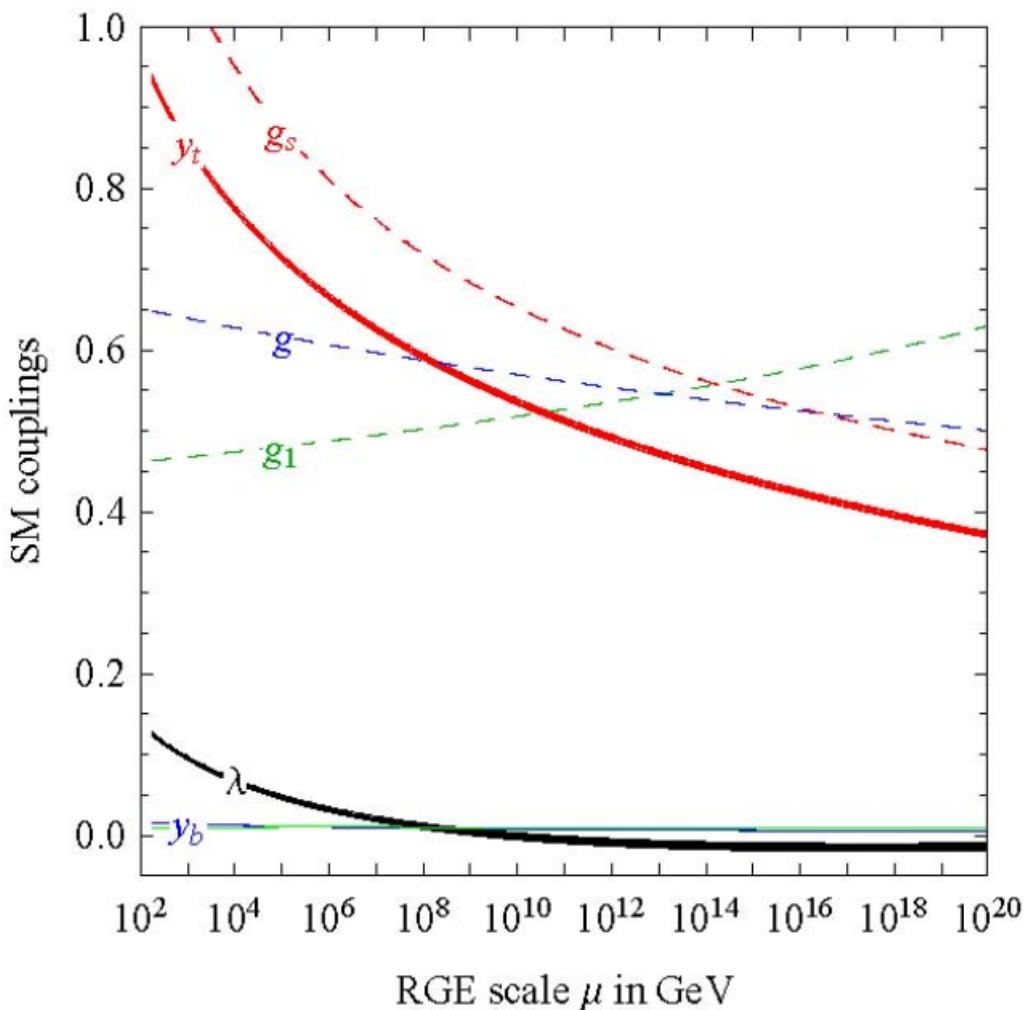
Assume scalar has SM props. and no BSM Physics

All SM parameters known

$$M_h \rightarrow \lambda(\text{EW})$$

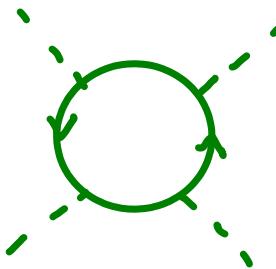
forgetting naturalness, can  
the pure SM be valid  
up to  $M_{\text{Pl}}$ ?

Weakly coupled up to  $M_{\text{Pl}}$



# VACUUM INSTABILITY

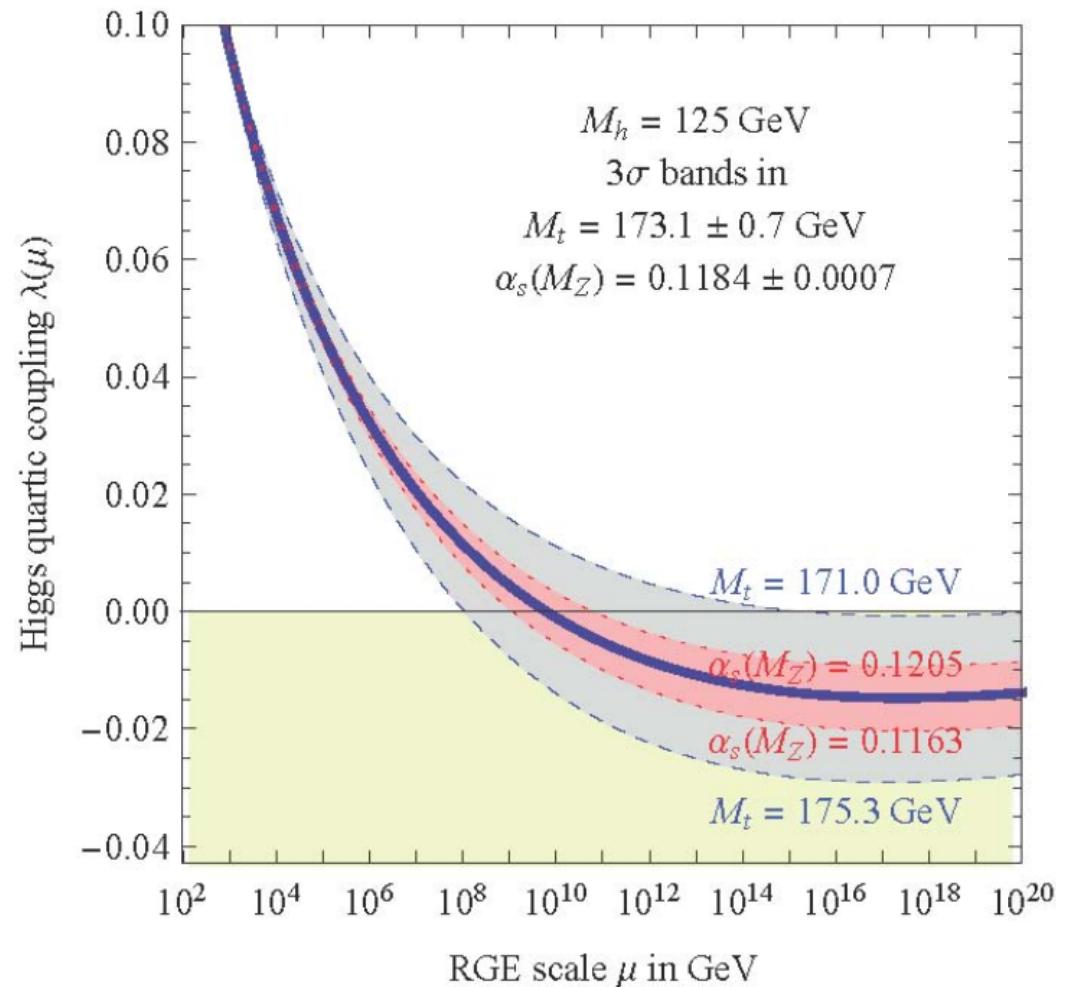
$$\frac{d\lambda}{d\ln Q} \sim -\frac{h_t^4}{16\pi^2}$$



$\lambda < 0$  at  $\Lambda_I \sim 10^{10}$  GeV

Higgs potential instability

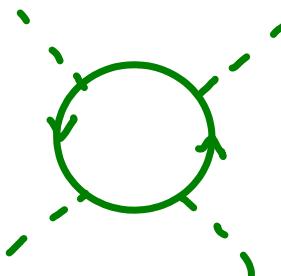
$$V(\phi \gg M_t) \approx \frac{1}{4} \lambda(Q \approx h) h^4$$



Cabibbo et al'79, Hung'79, Lindner'86

# VACUUM INSTABILITY

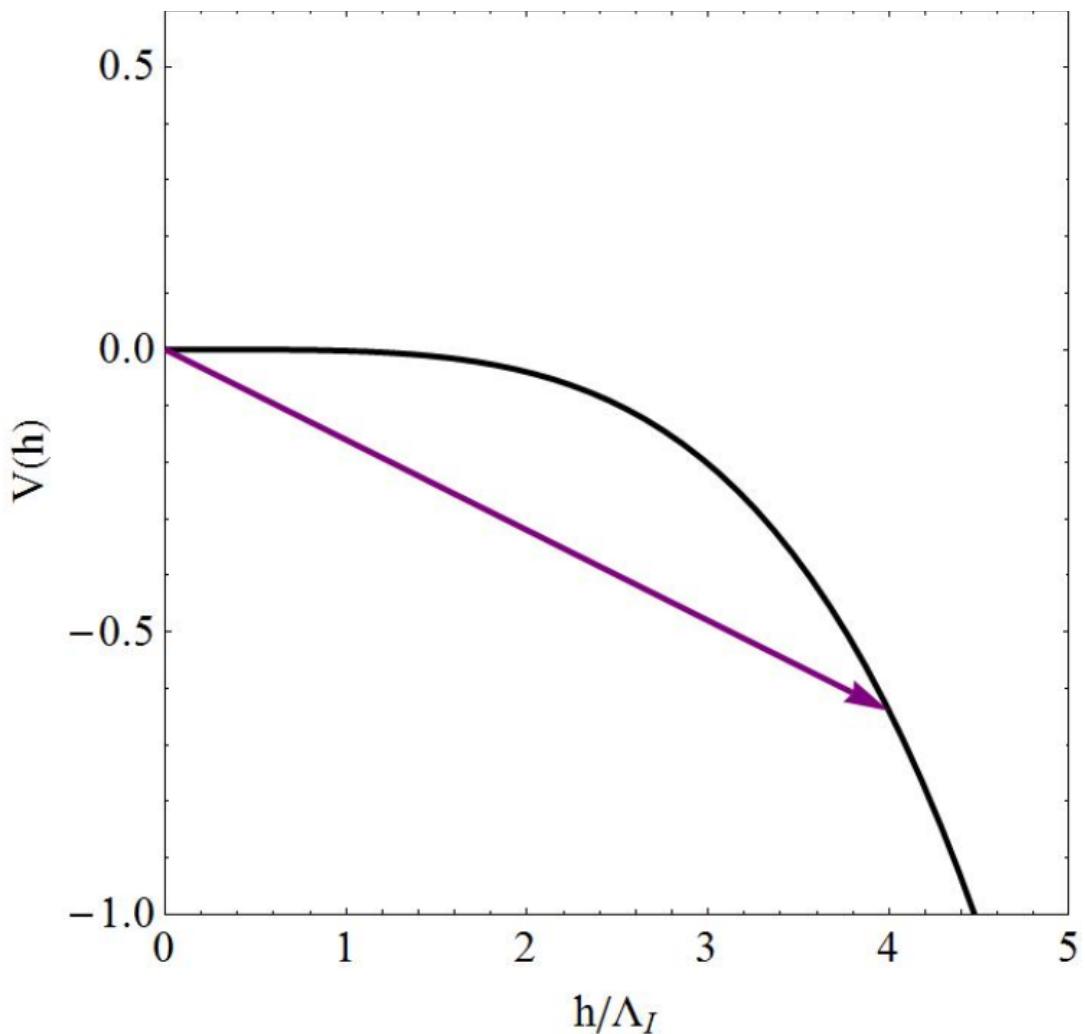
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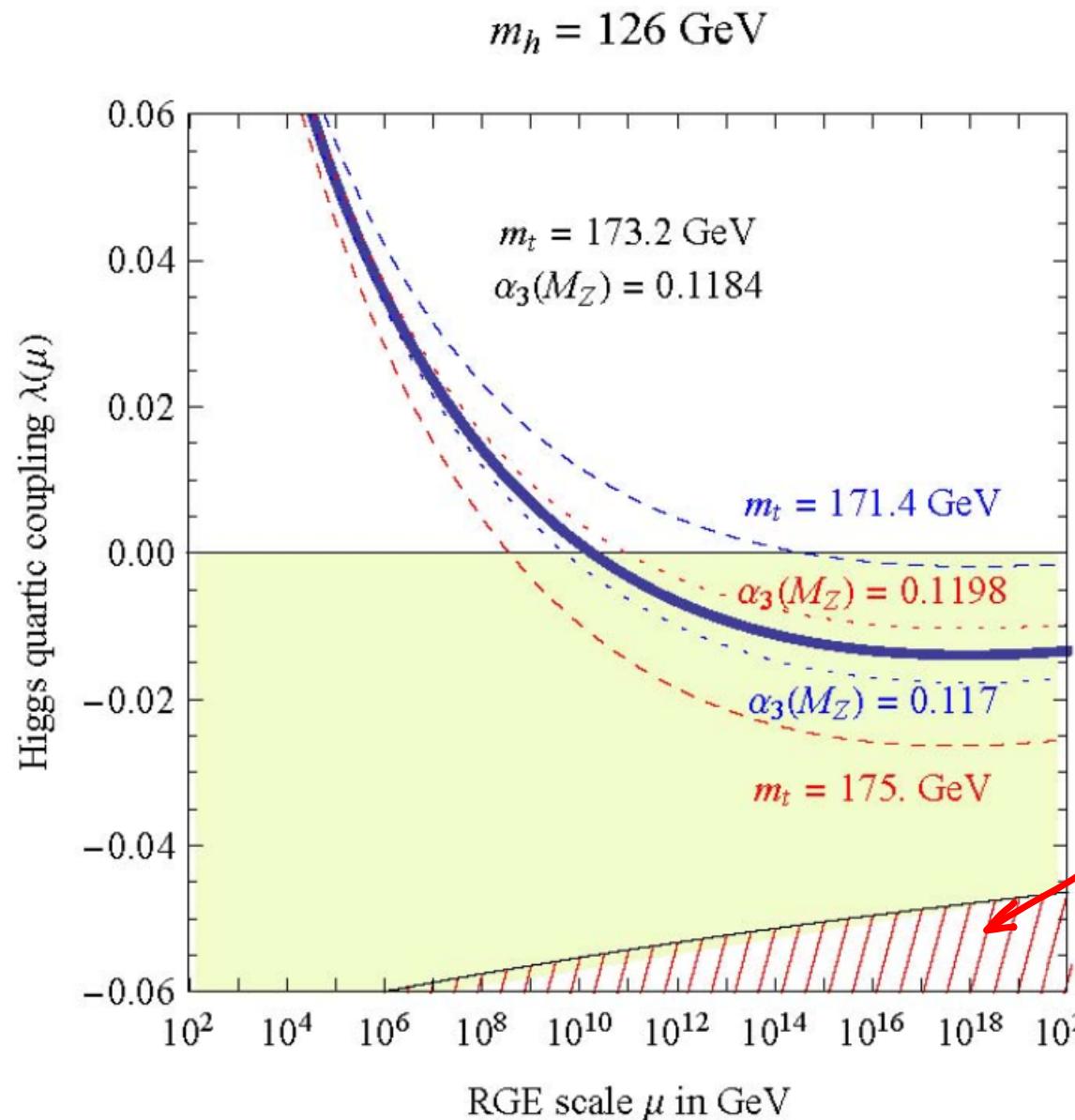
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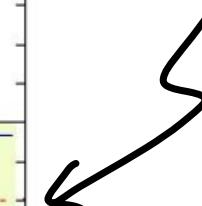
$$V(\phi \gg M_t) \simeq \frac{1}{4} \lambda(Q \approx h) h^4$$



# LIFE IN A METASTABLE VACUUM

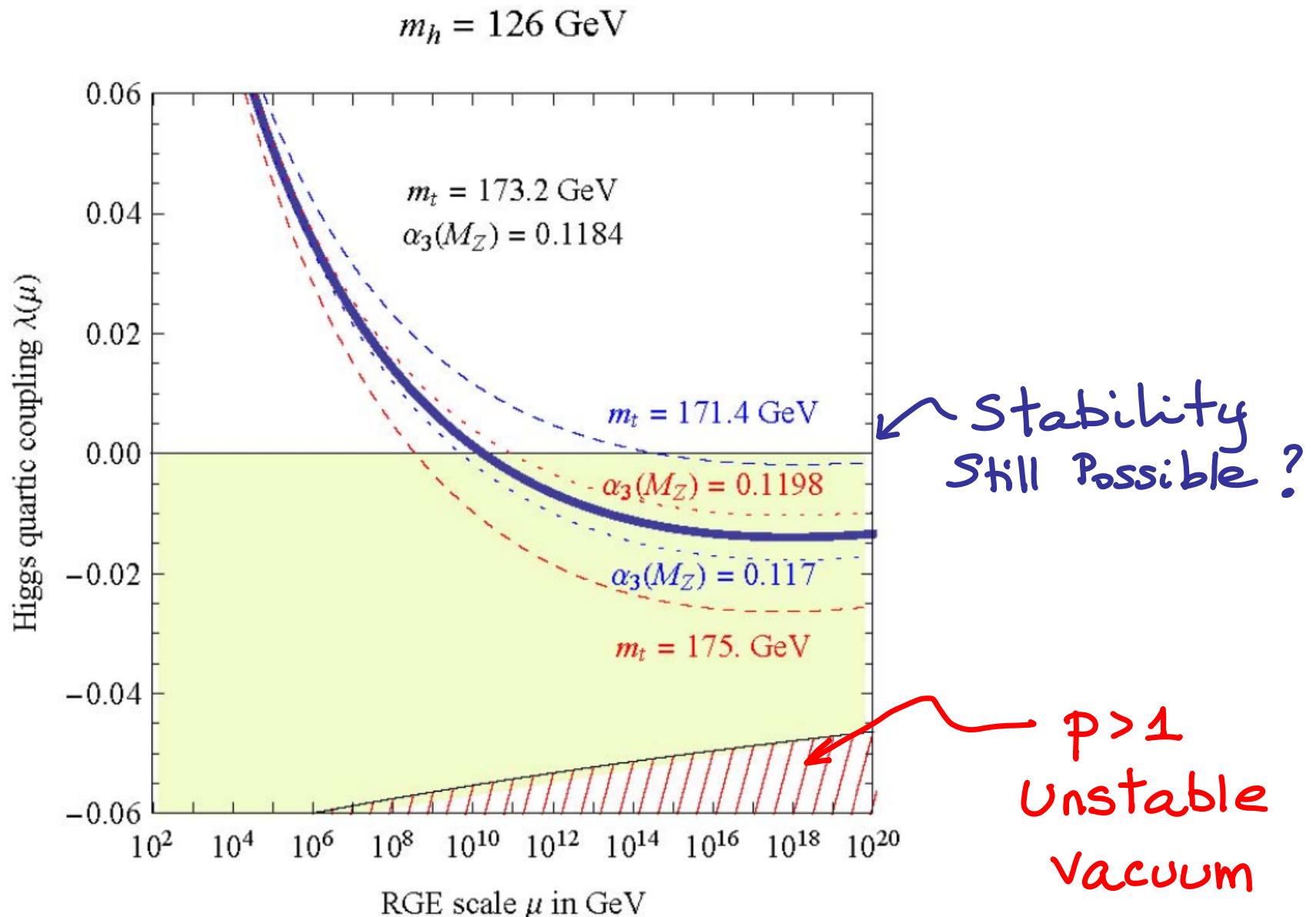


Lifetime  $\propto \exp \frac{1}{|\lambda|}$   
 $\gg \text{age of Universe}$



$p > 1$   
Unstable  
vacuum

# LIFE IN A METASTABLE VACUUM



## NNLO STABILITY BOUND

For stability up to  $M_{Pl}$ :

$$M_h [\text{GeV}] > 129.4 + 1.4 \left( \frac{M_t (\text{GeV}) - 173.1}{0.7} \right) - 0.5 \left( \frac{\alpha_s(M_Z) - 0.1184}{0.0007} \right) \pm 1.0_{\text{th}}$$

State-of-the-art NNLO calculation:

- 2-loop Veff (Ford, Jack, Jones [hep-ph/0111190])
- 3-loop RGES (... , Chetyrkin, Zoller [hep-ph/1205.2892])
- 2-loop matching in  $\lambda \leftrightarrow M_h^2$ ;  $h_t \leftrightarrow M_t$   
(..., Shaposhnikov et al [hep-ph/1205.2893],  
, Degrassi et al [hep-ph/1205.6497])

## TOP MASS CAVEATS

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Have assumed

$$M_t = 173.1 \pm 0.7 \text{ GeV}$$

from Tevatron + LHC is the top pole mass.

Mitov's & Uwer's talks

Theoretically cleaner determination from  $\sigma(t\bar{t})$   
but larger error

$$M_t = 173.3 \pm 2.8 \text{ GeV}$$

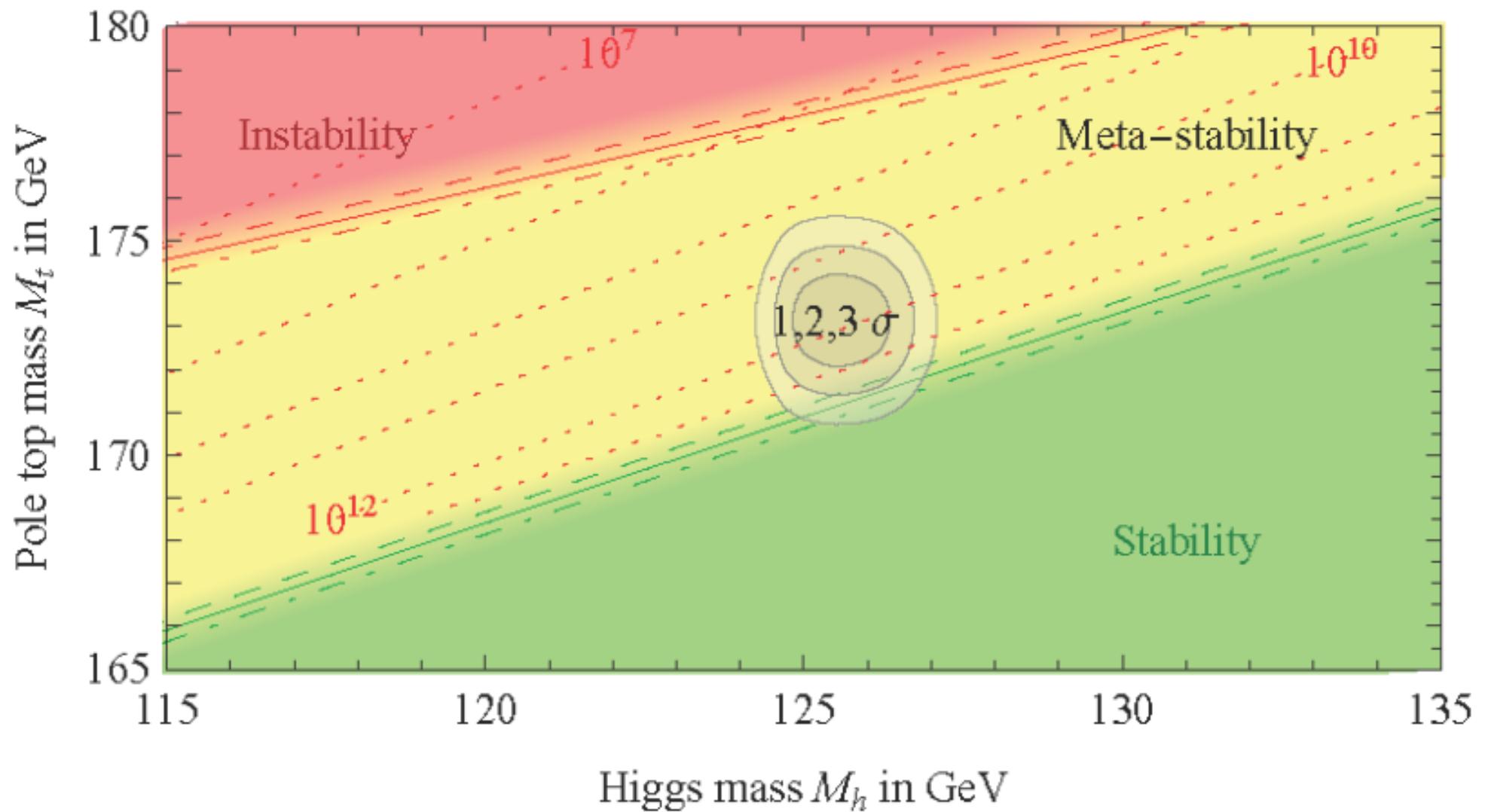
would still allow for stability

Alekhin, Djouadi, Moch [hep-ph/1207.0980]

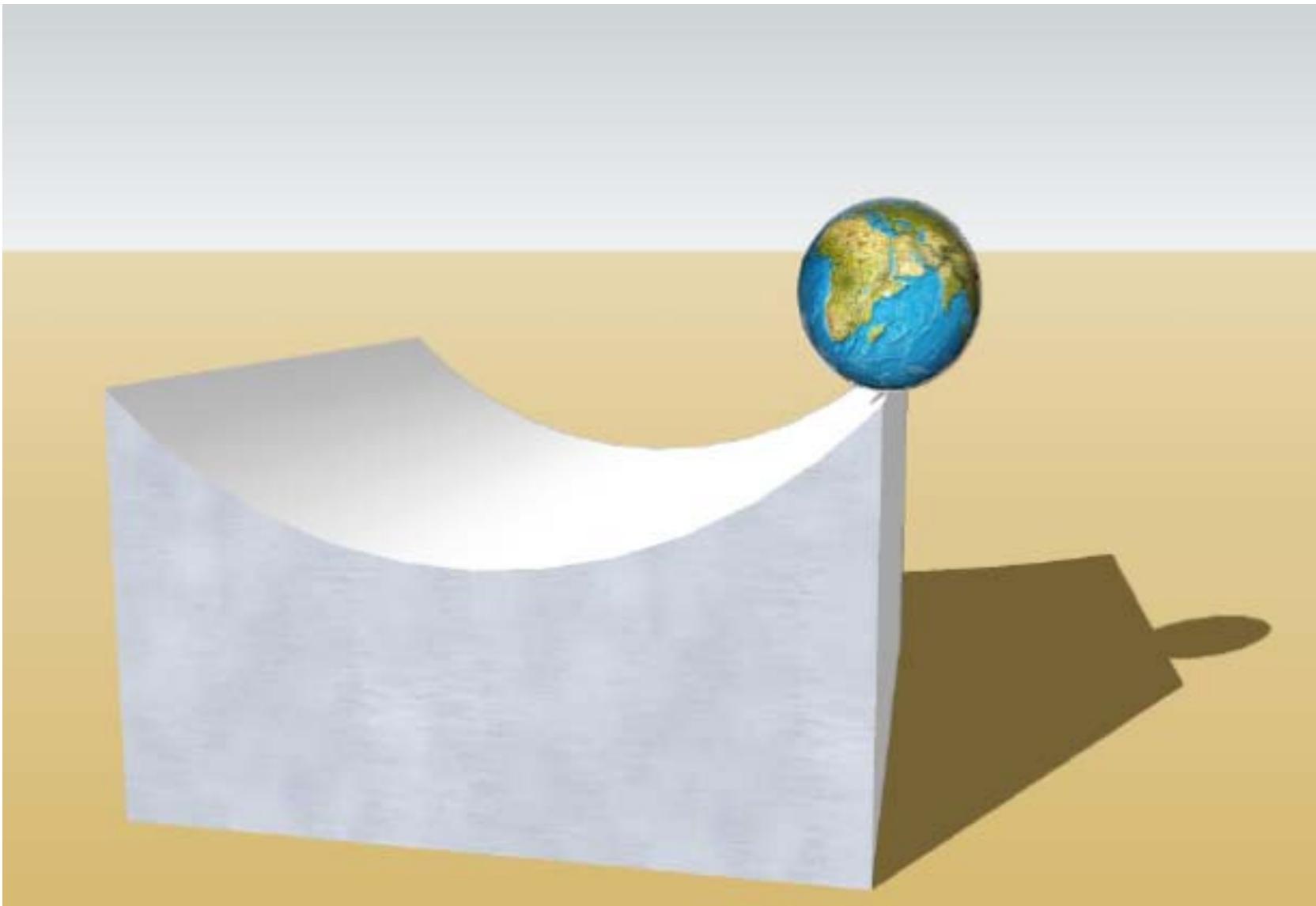
Too conservative given the good agreement...

# LIVING AT THE EDGE

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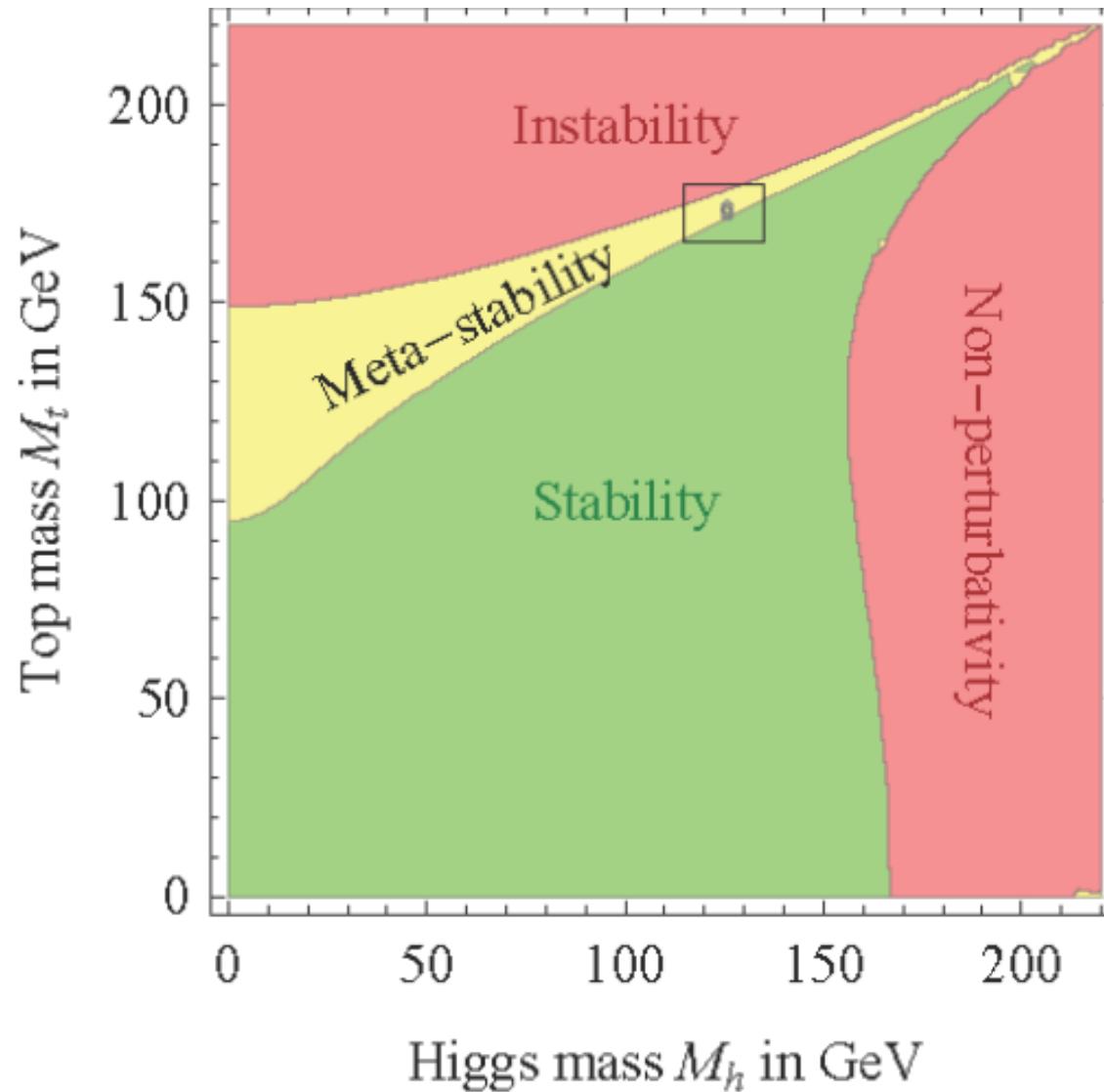


# LIVING AT THE EDGE



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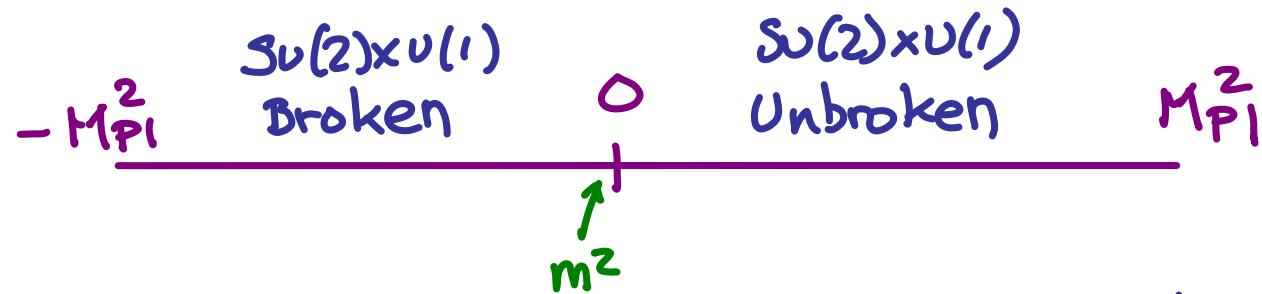


# NEW KNOWLEDGE BRINGS NEW QUESTIONS

- ★ Why do we live near the critical boundary for stability?

$$\lambda(M_{Pl}) \approx 0$$

- ★ Is this related to our living near the phase boundary  $m^2/M_{Pl}^2 \approx 0$ ?

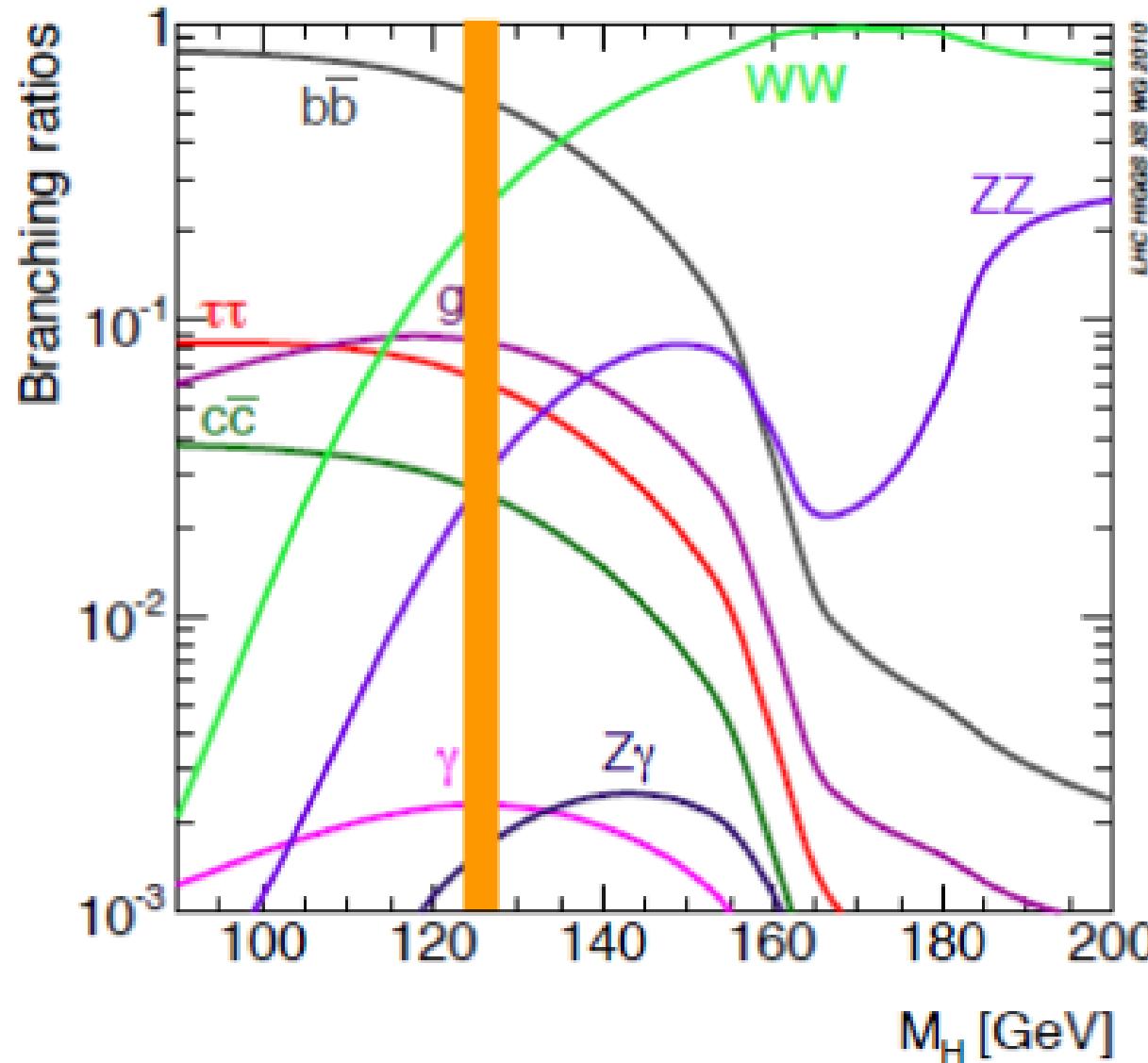


- ★ Is the EW scale determined by Planck scale physics?
- ★ Or is this just a coincidence? BSM...

②

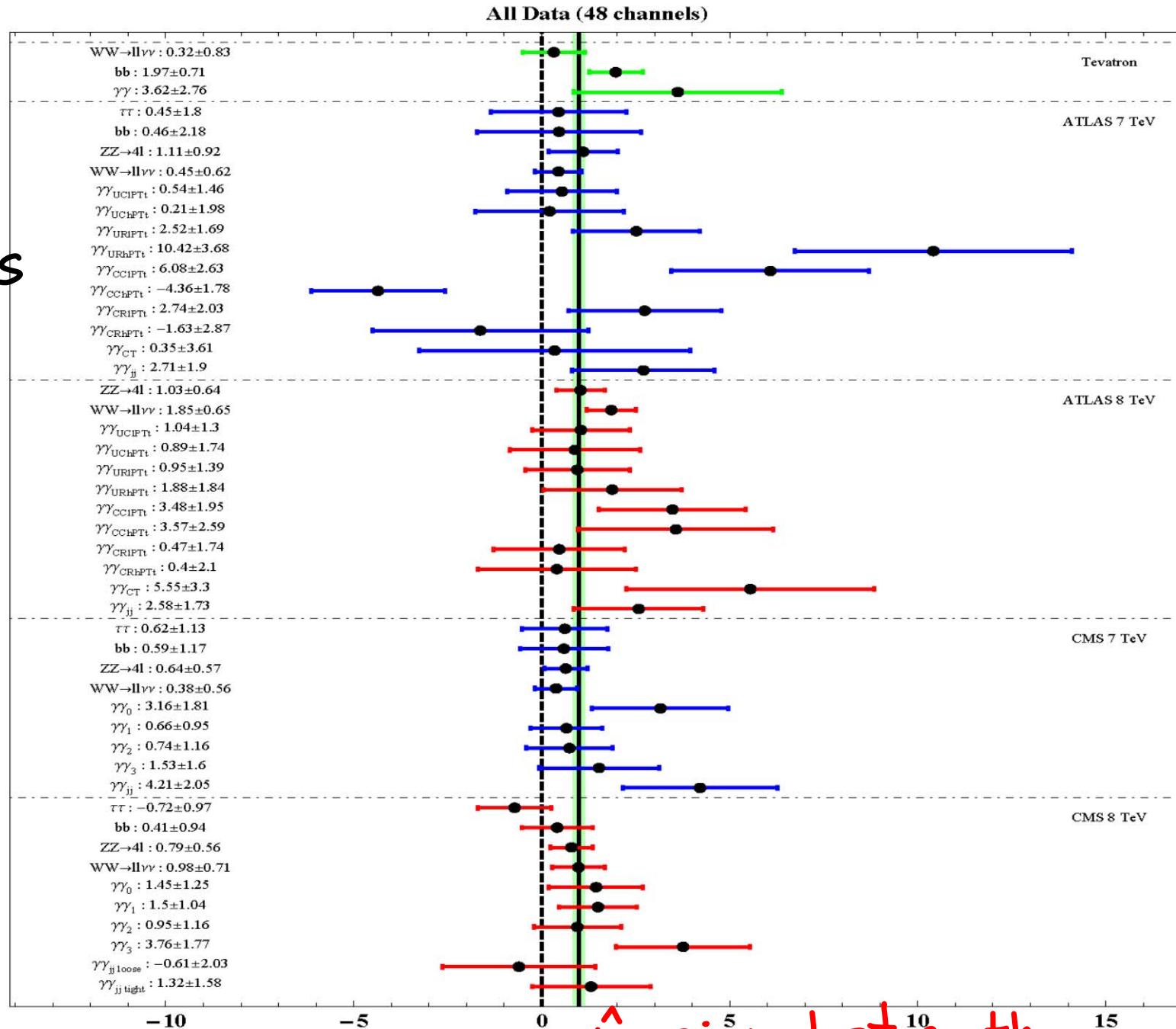
## HIGGS COUPLINGS

$m_h \sim 125$  GeV: many channels!



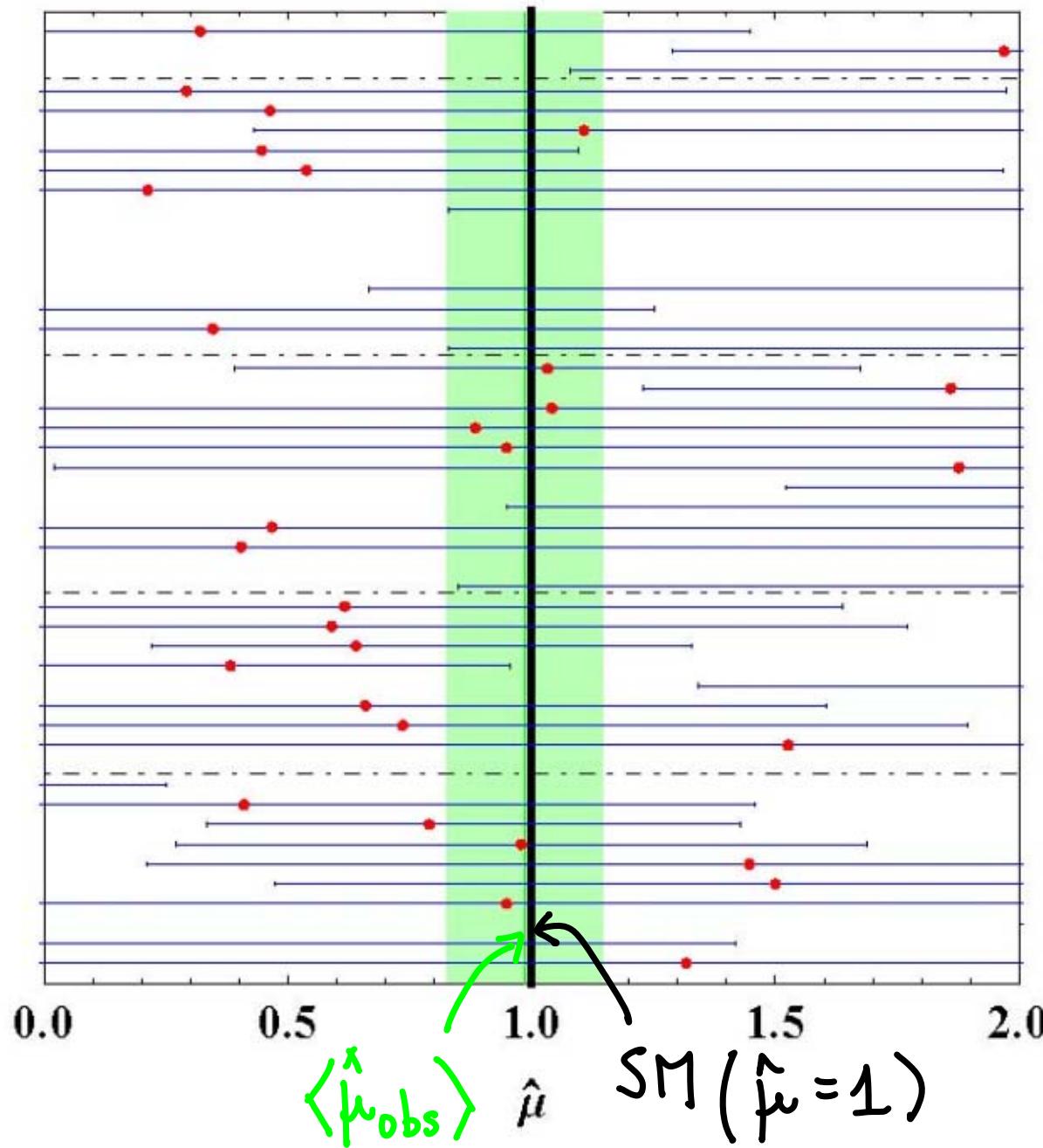
# DATA USED IN OUR FITS

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# DATA USED IN OUR FITS

All Data



Zoom of  
previous plot

## EFFECTIVE LAGRANGIAN APPROACH

↙ valid at  $E \sim M_t$ .

Field content : SM + scalar  $h$  (no extra light states)

$$\mathcal{L} = \mathcal{L}[h] - (M_W^2 W_\mu^+ W^\mu_- + \frac{1}{2} M_Z^2 Z_\mu Z^\mu) \left[ 1 + 2a \frac{h}{v} + O(h^2) \right]$$
$$- m_i \bar{\psi}_i \psi_i \left[ 1 + c \frac{h}{v} + O(h^2) \right] + \dots$$

Contino et al  
'10 '12

Incorporates  $SU(2) \times U(1)_Y \rightarrow U(1)_{em}$  breaking

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$$- m_i \bar{\psi}_i \Gamma_i \psi_i \left[ 1 + c \frac{h}{v} + O(h^2) \right] + \dots$$

Incorporates  $SU(2) \times U(1)_Y \rightarrow U(1)_{em}$  breaking  
+ custodial symmetry

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$$- m_i \bar{\psi}_i \tilde{\psi}_i \left[ 1 + c \frac{h}{v} + O(h^2) \right] + \dots$$

Incorporates  $SU(2) \times U(1)_Y \rightarrow U(1)_{em}$  breaking

+ custodial symmetry

+ no tree-level FCNC from  $h$  exchange

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Incorporates  $SU(2) \times U(1)_Y \rightarrow U(1)_{em}$  breaking

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First terms in a "chiral lagrangian" with cutoff

$$\lambda \gtrsim 4\pi v$$

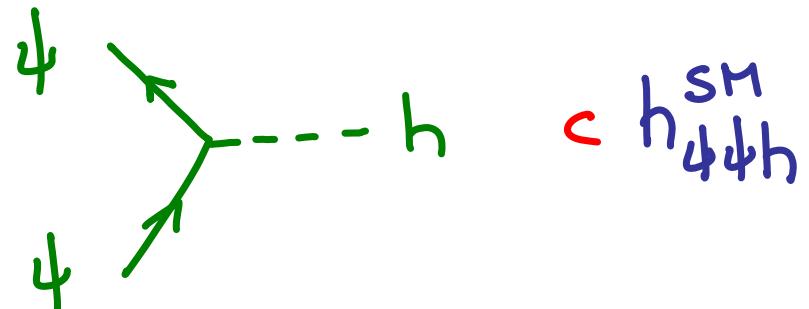
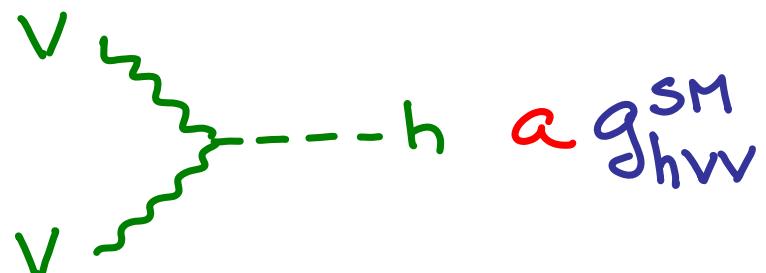
# EFFECTIVE LAGRANGIAN APPROACH

↙ valid at  $E \sim M_T$ .

Field content : SM + scalar  $h$  (no extra light states)

$$\mathcal{L} = \mathcal{L}[h] - (M_W^2 W_\mu^+ W^\mu_- + \frac{1}{2} M_Z^2 Z_\mu Z^\mu) \left[ 1 + 2a \frac{h}{v} + O(h^2) \right] - m_i \bar{\psi}_i \psi_i \left[ 1 + c \frac{h}{v} + O(h^2) \right] + \dots$$

2-parameter extension of the SM with



$$SM \equiv (a, c) = (1, 1)$$

## EFFECTIVE LAGRANGIAN APPROACH

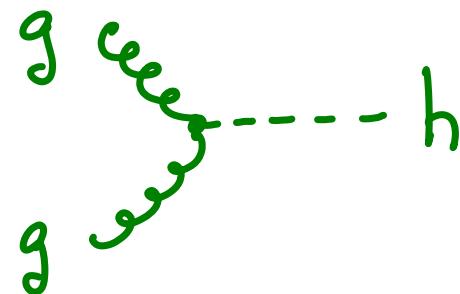
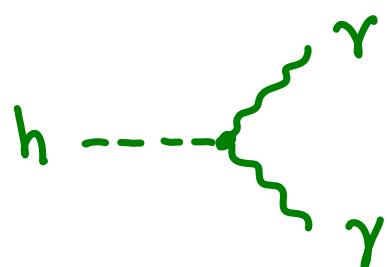
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Other operators very relevant for Higgs searches:

$$\frac{g^2}{16\pi^2} \left[ c_\gamma A_{\mu\nu} A^{\mu\nu} + c_g G_{\mu\nu}^A G^{A\mu\nu} \right] \frac{h}{v}$$



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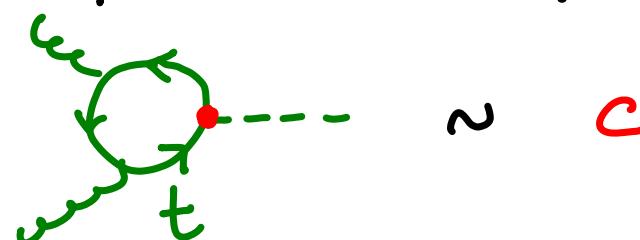
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are assumed to come from SM loops only:

$gg \rightarrow h$



## EFFECTIVE LAGRANGIAN APPROACH

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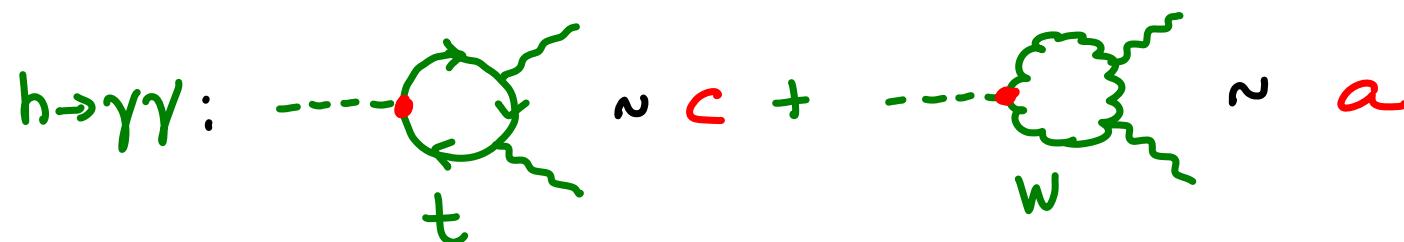
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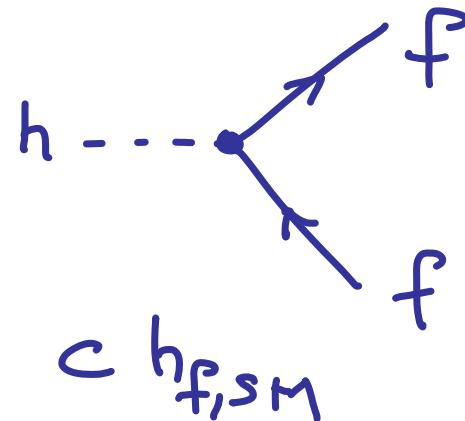
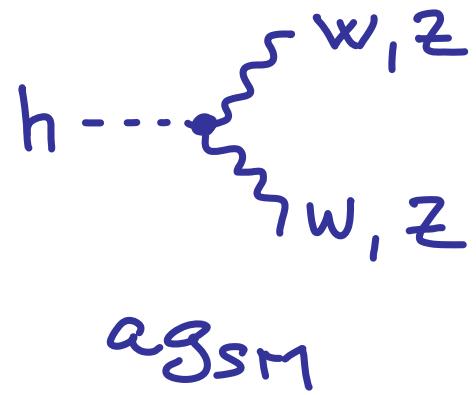
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are assumed to come from SM loops only.

Motivated by composite PNGB Higgs scenarios  
Will relax this assumption later on.

## PHYSICAL RANGE OF $(a, c)$



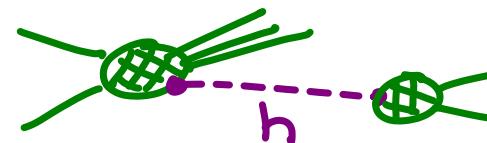
$a g_{SM} h \nu_\mu \nu^\mu$  set  $a > 0$  by choosing sign of  $h$

indep.  $\left\{ \begin{array}{ll} m_f \bar{f} f & \text{set } m_f > 0 \text{ by choosing phase of } \bar{f} f \\ c h_{f,SM} h \bar{f} f & c \text{ can be of either sign.} \end{array} \right.$

# RATES IN SM( $a, c$ )

★ Only  $h$  couplings are modified :

Signal rate modified  
kinematics unchanged



★ Different production mechanisms change differently.

$$\sigma \left[ \text{loop} \dots, \text{gg-fusion} \right] \sim c^2 \quad \sigma \left[ \text{gg-fusion}, \text{loop} \dots \right] \sim a^2$$

★ Expected signal strengths :

$$\mu_i = \left[ \frac{\sigma_{pp \rightarrow h \rightarrow X}^{SM(a,c)}}{\sigma_{pp \rightarrow h \rightarrow X}^{SM}} \right]_i = \frac{\sum_{p_i} \epsilon_{p_i} \sigma_{p_i}(a,c) \times BR_{h \rightarrow X}^{SM(a,c)}}{\sum_{p_i} \epsilon_{p_i} \sigma_{p_i} \times BR_{h \rightarrow X}^{SM}}$$

Calculate  $\mu_i(a,c)$  and confront with the measured  $\mu_i$ 's using a  $\chi^2$  fit.

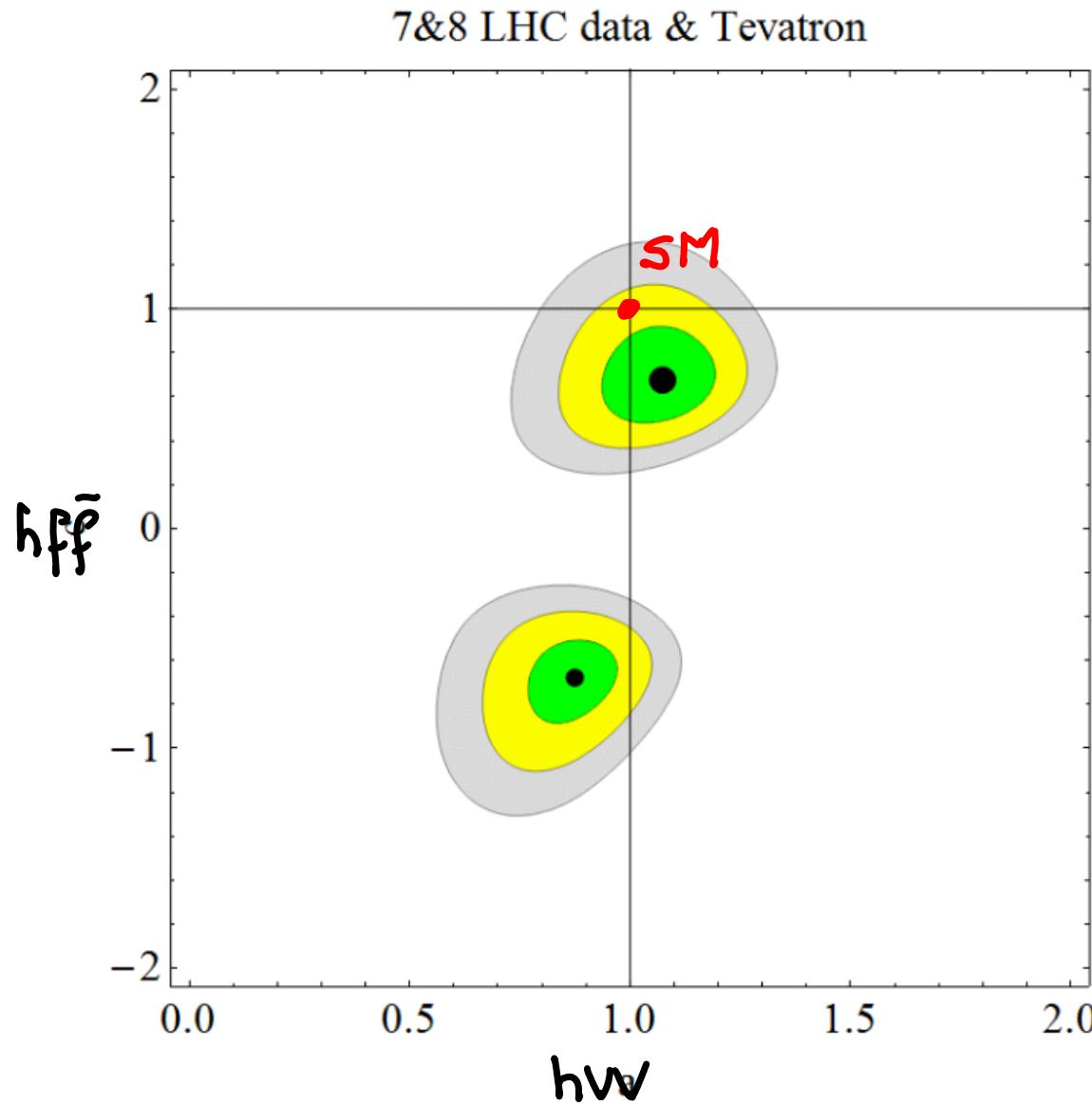
# FITS IN SM( $a, c$ )

$\chi^2$  fit to  $\hat{\mu}_i \pm \sigma_i$  from 48 channels (ATLAS+CMS+Tevatron)

68%

95%

99%



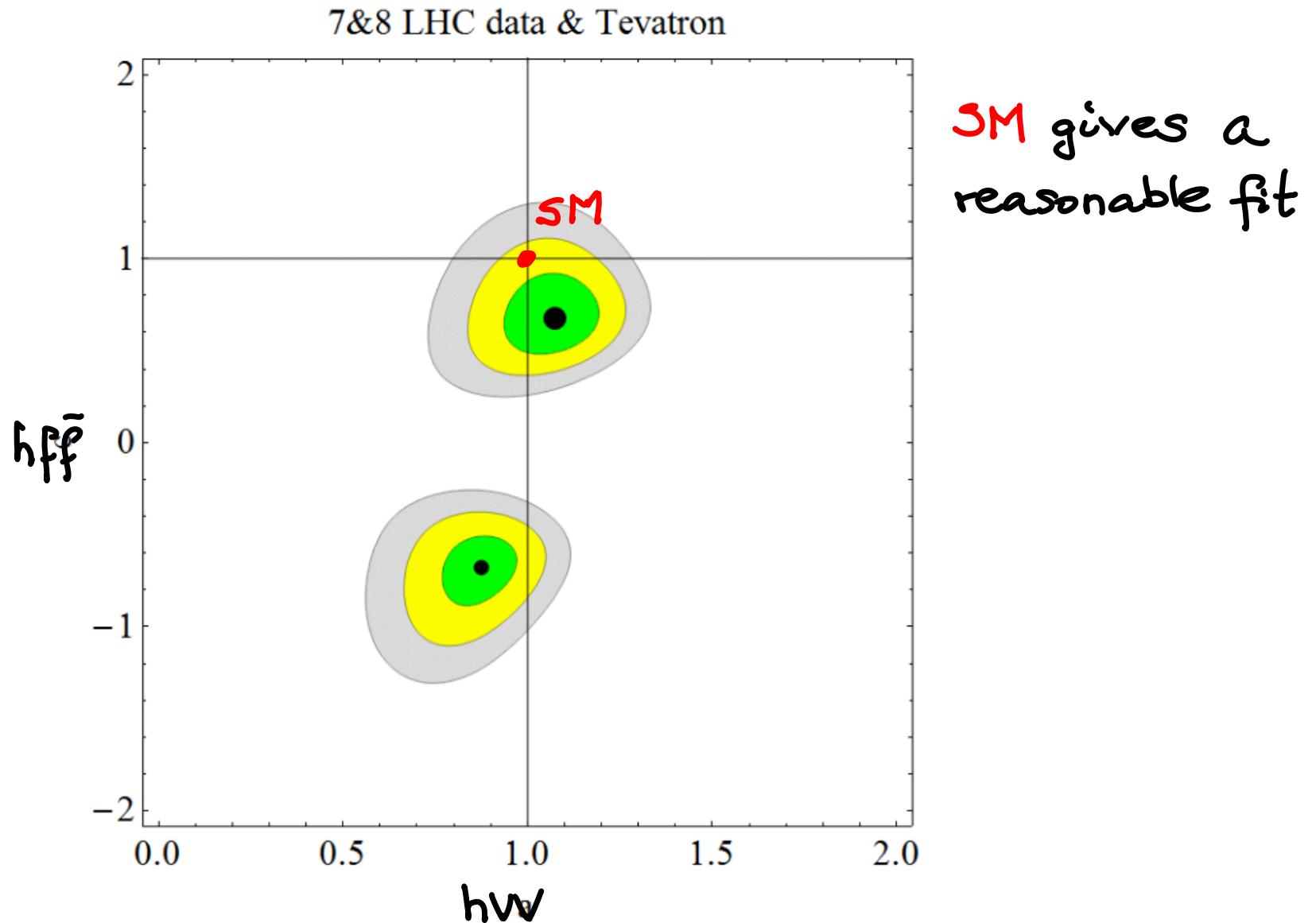
# FITS IN SM( $a, c$ )

$\chi^2$  fit to  $\hat{\mu}_i \pm \sigma_i$  from 48 channels (ATLAS+CMS+Tevatron)

68%

95%

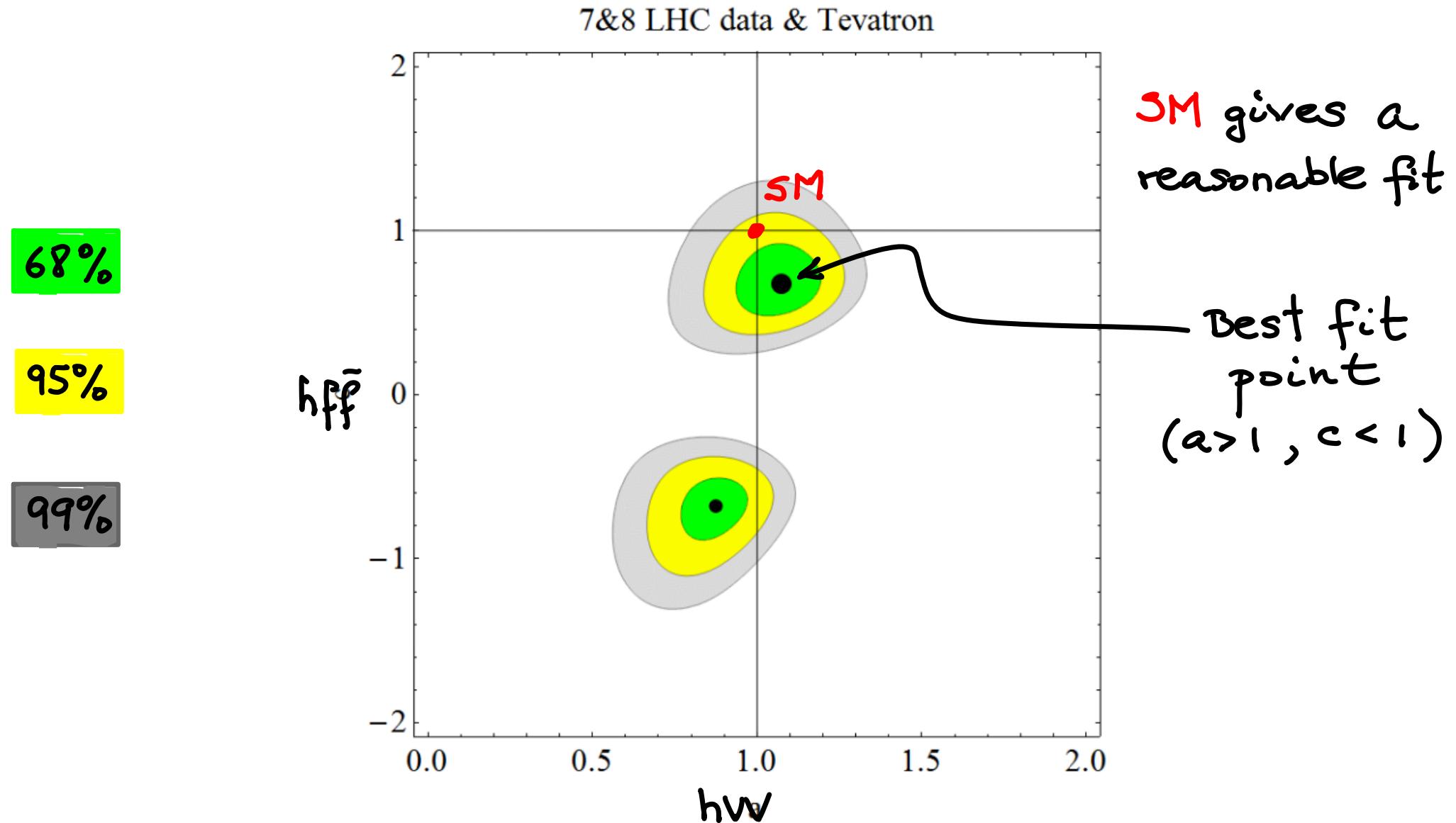
99%



SM gives a reasonable fit

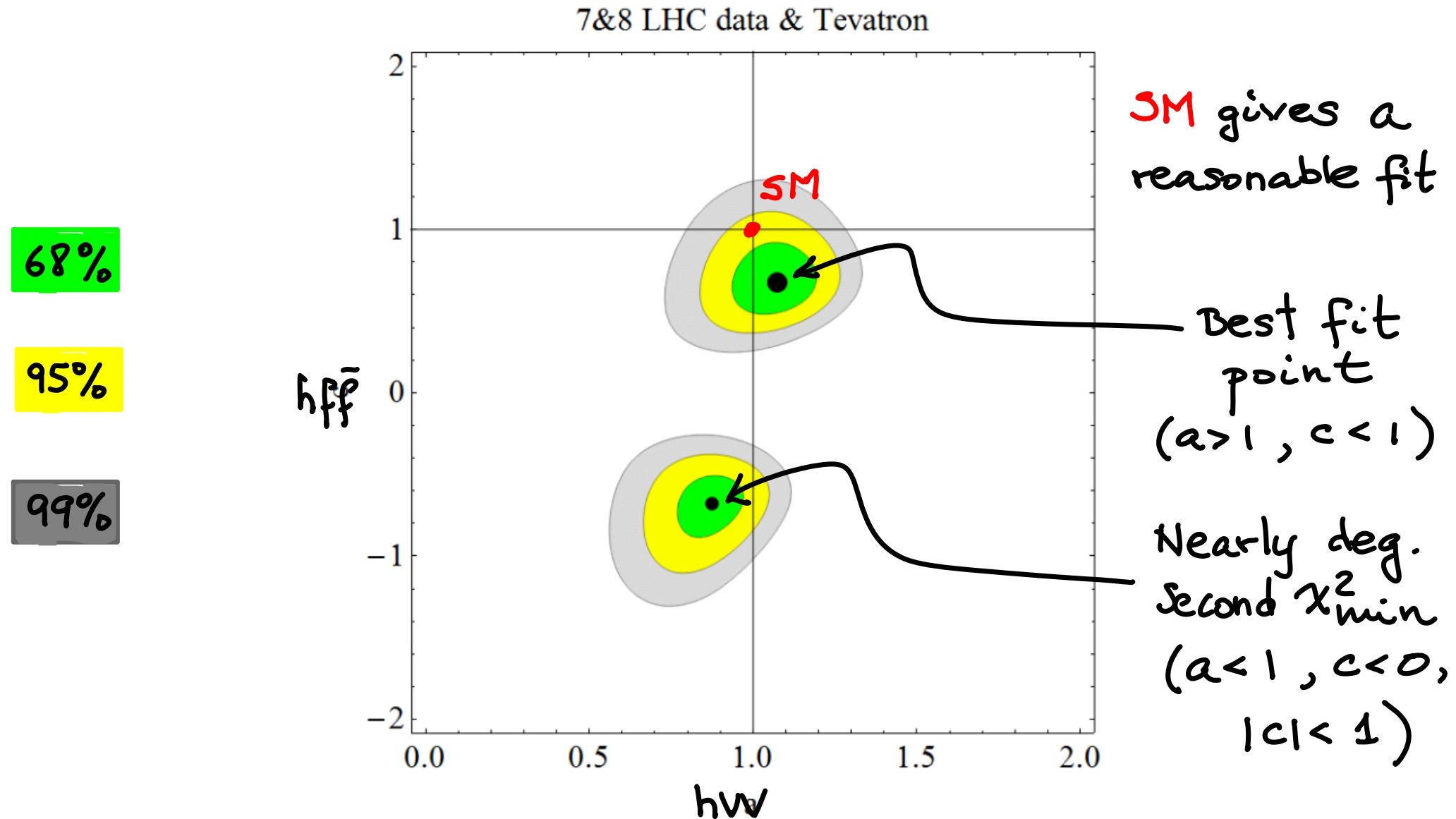
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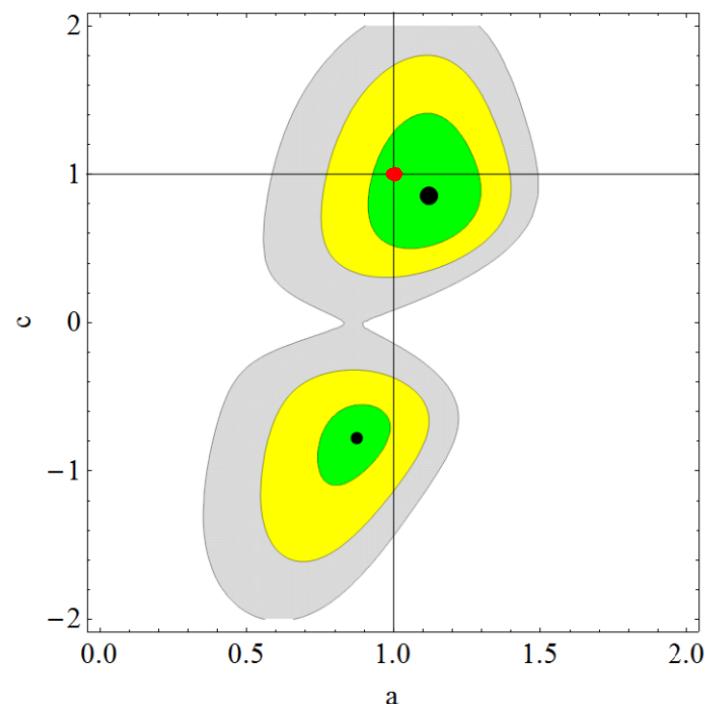
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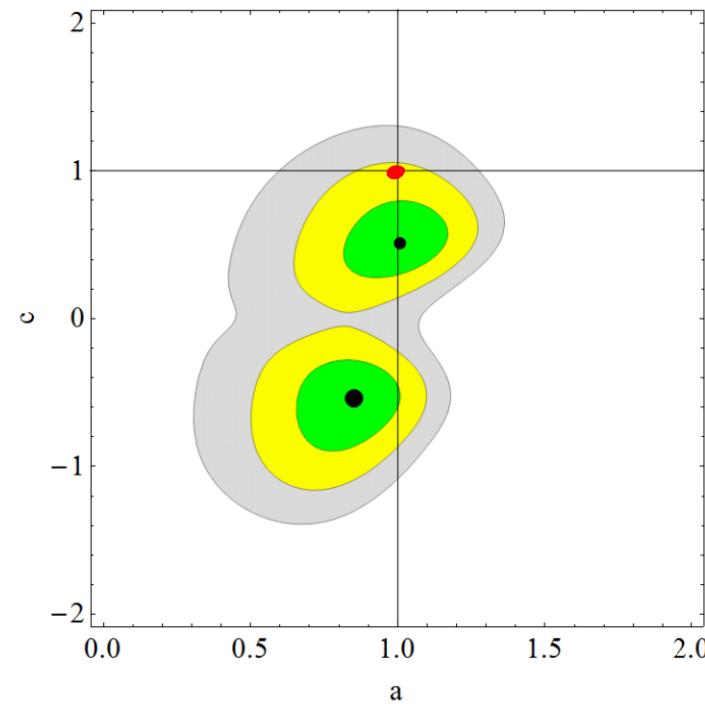
# FITS TO SEPARATE EXP.

7&8 ATLAS data ( $m_h=126.5$  GeV)



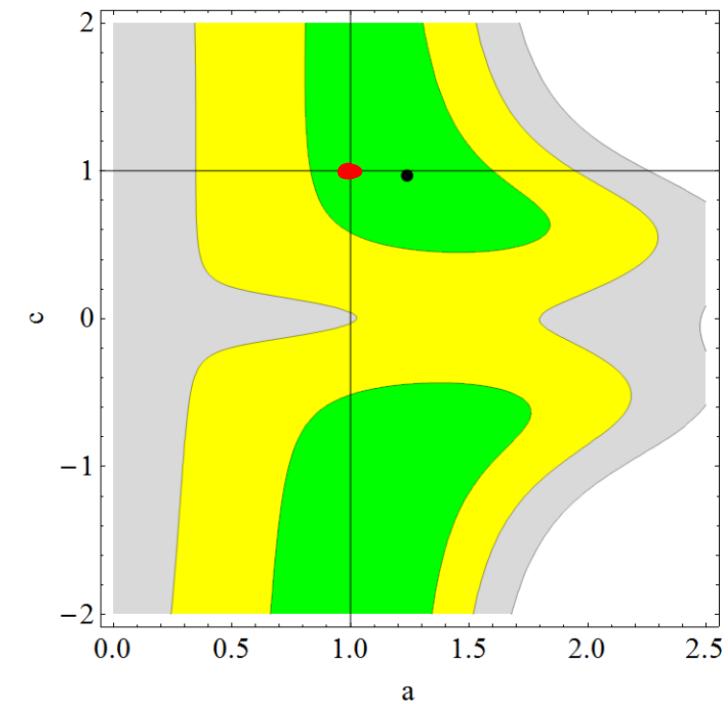
ATLAS

7&8 CMS data ( $m_h=125$  GeV)



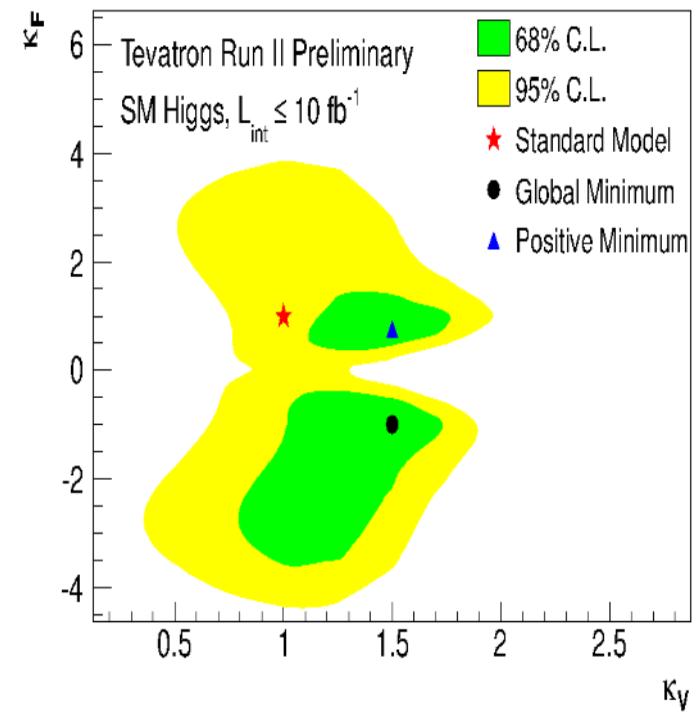
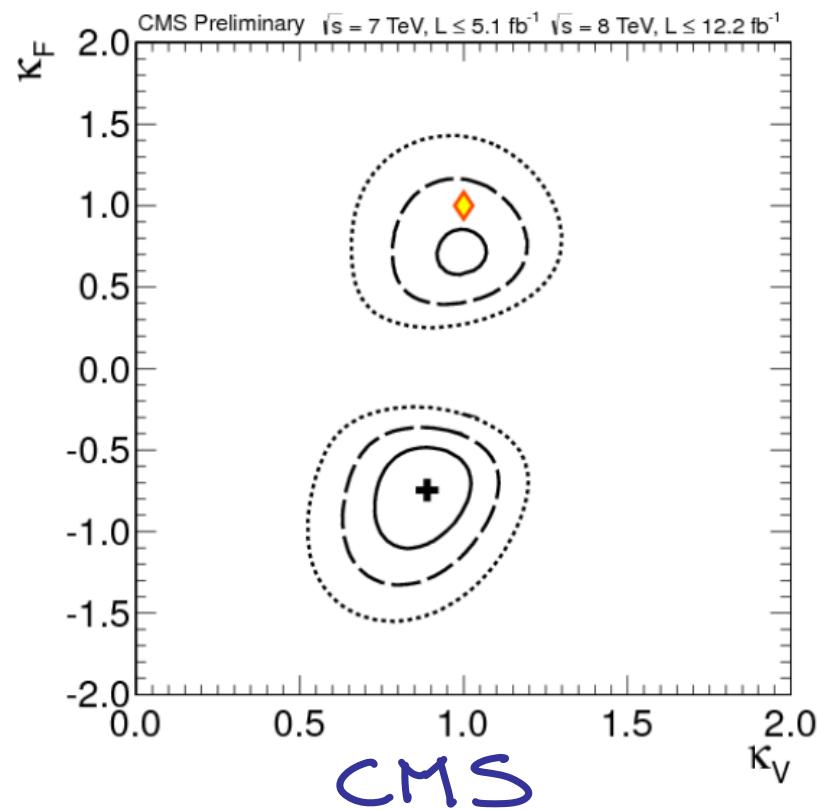
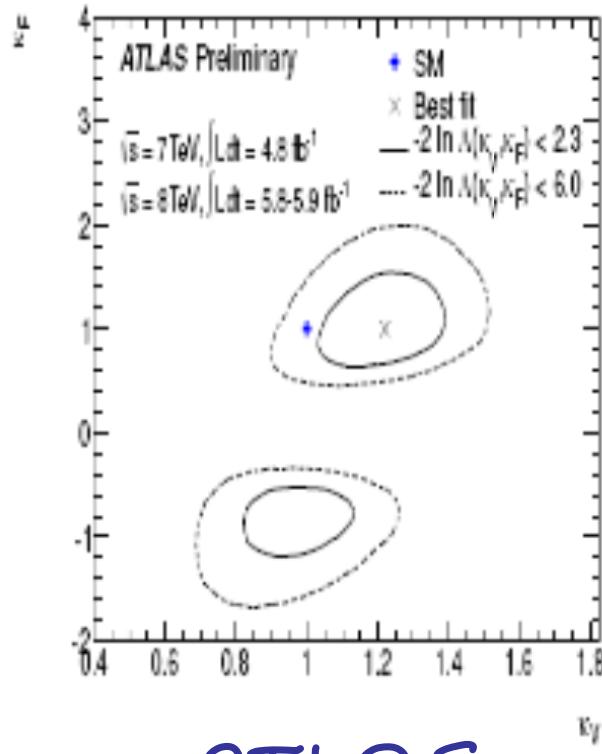
CMS

Tevatron data ( $m_h=125$  GeV)



Tevatron

# FITS TO SEPARATE EXP.

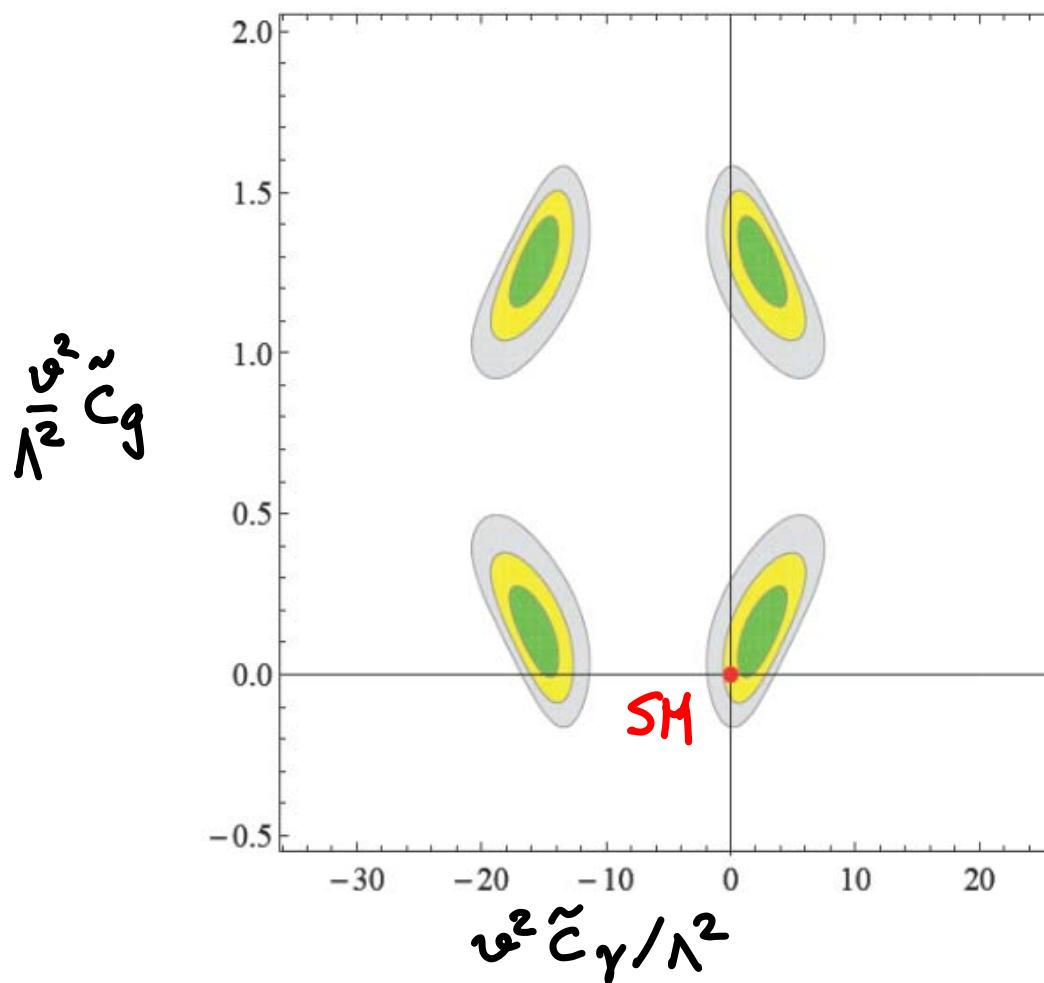


OFFICIAL VERSION

# BSM EFFECTS IN $h\gamma\gamma$ , $hgg$

Take  $a=c=1$ , add  $\frac{-1}{32\pi^2} \left[ e^2 \tilde{c}_\gamma F_{\mu\nu} F^{\mu\nu} + g_s^2 \tilde{c}_g G_{\mu\nu}^A G^{A\mu\nu} \right] \frac{H^\dagger H}{\Lambda^2}$

e.g. from loops of heavy charged particles  $\Rightarrow$  SM ( $\tilde{c}_\gamma, \tilde{c}_g$ )

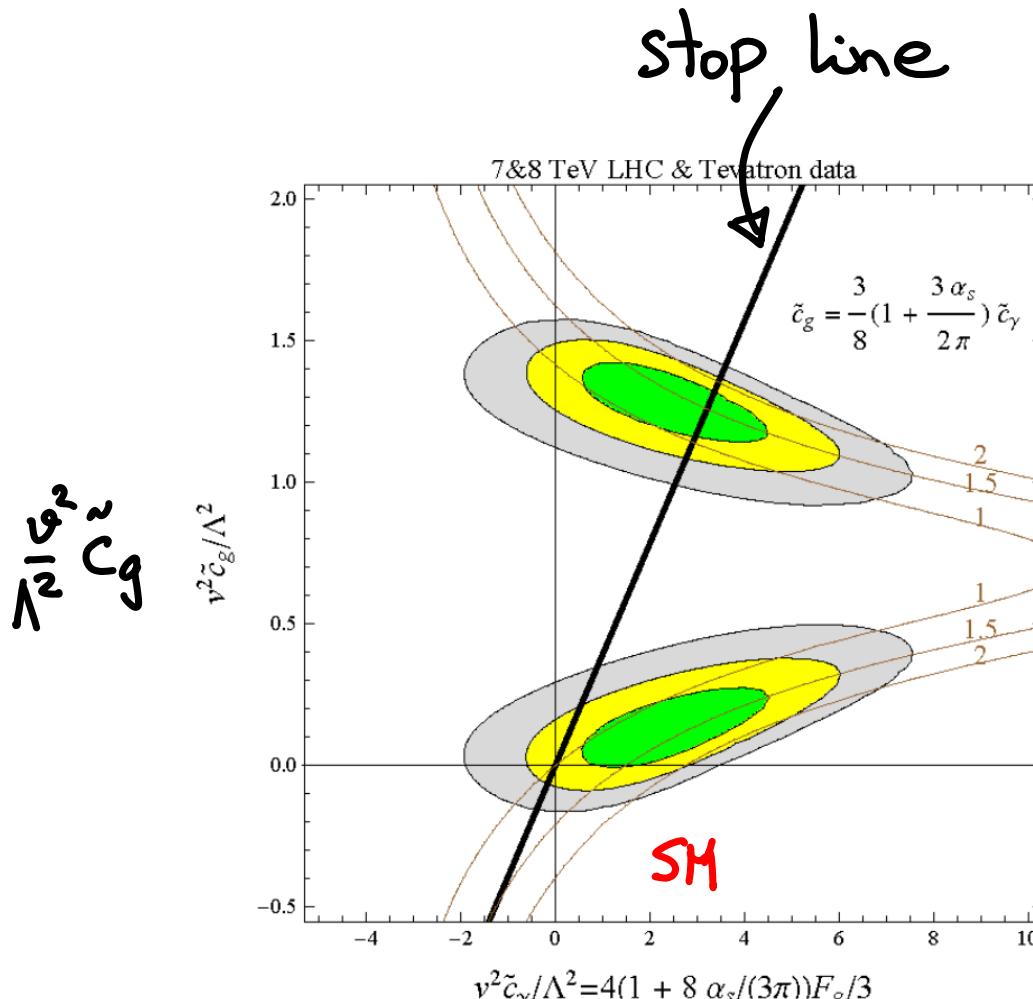


Model-independent result:  
Map particular models to this plane

# BSM EFFECTS IN $h\gamma\gamma$ , $hgg$

---

Example : Stop, which gives  $\tilde{c}_g = 2N_c Q_t^2 \tilde{c}_g$



See JHEP, Gogoberidze,  
Sanz, Trott '12

$$v^2 \tilde{c}_\gamma / \Lambda^2$$

## SOME REMARKS

- ★ With more data, fits to more parameters will be possible.
- ★ Higher order ( $p$ -dep) operators in  $\mathcal{L}_{\text{eff}}$  will be required for highly boosted Higgs.
- ★ Current Higgs data can be used to constrain BSM models either for giving two large signal rates or too small (eg. limits on Higgs invisible BR)
- ★ Higgs data can be used to probe indirectly regions of parameter space hard for direct searches (eg. light stops)

# CONCLUSIONS

We finally have data to explore the physics of electroweak symmetry breaking!

★  $M_h \simeq 125 \text{ GeV} \Rightarrow$  Unstable EW vacuum w/o BSM

Long-lived and intriguingly close to stability boundary

Deviations of Higgs properties from the SM are a must of natural BSM: extremely important measurements.

★ Couplings compatible with SM

Some interesting deviations ( $\gamma\gamma$ ), not statistically signif (yet?)

Future deviations can be crucial to steer the search for physics beyond the SM.

Looking forward to Moriond'13 !

# LIFE IN A METASTABLE VACUUM

$$p = \text{Decay prob.} = \underbrace{\frac{\text{Decay rate}}{\Delta t \cdot \Delta V}}_{h^4 e^{-S_4}} \tau_0^4 \quad \text{with} \quad \tau_0^4 \sim (e^{140}/M_{Pl})^4$$

$$h^4 e^{-S_4} \sim h^4 \exp\left(-\frac{8\pi^2}{3|\lambda(h)|}\right) \sim h^4 \exp\left[-\frac{2600}{|21/0.01|}\right]$$

easily wins over  $\tau_0^4$

$p \ll 1$  : Lifetime of EW vacuum much longer than  $\tau_0$

# ERROR BUDGET OF STAB. BOUND

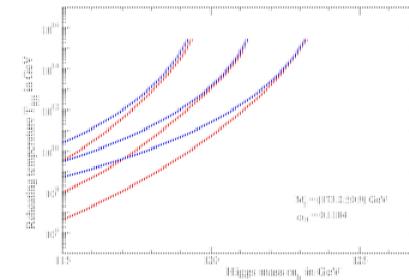
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Type of error	Estimate of the error	Impact on $M_h$
$M_t$	experimental uncertainty in $M_t$	$\pm 1.4$ GeV
$\alpha_s$	experimental uncertainty in $\alpha_s$	$\pm 0.5$ GeV
Experiment	<b>Total combined in quadrature</b>	$\pm 1.5$ GeV
$\lambda$	scale variation in $\lambda$	$\pm 0.7$ GeV
$y_t$	$\mathcal{O}(\Lambda_{\text{QCD}})$ correction to $M_t$	$\pm 0.6$ GeV
$y_t$	QCD threshold at 4 loops	$\pm 0.3$ GeV
RGE	EW at 3 loops + QCD at 4 loops	$\pm 0.2$ GeV
Theory	<b>Total combined in quadrature</b>	$\pm 1.0$ GeV

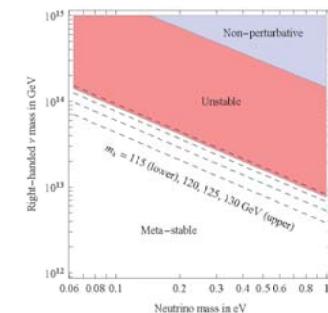
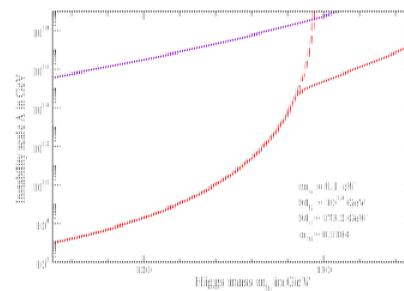
Table 1: Dominant sources of experimental and theoretical errors in the computation of the SM stability bound on the Higgs mass, eq. (2).

# OTHER IMPLICATIONS NOT DISCUSSED

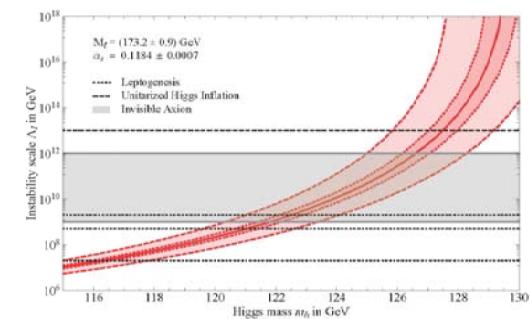
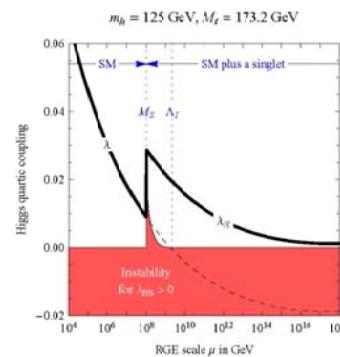
- Cosmology :  
Bound on  $T_{RH}$ ; inflation,...



- See-saw neutrinos :  
Bound on  $M_{NR}$

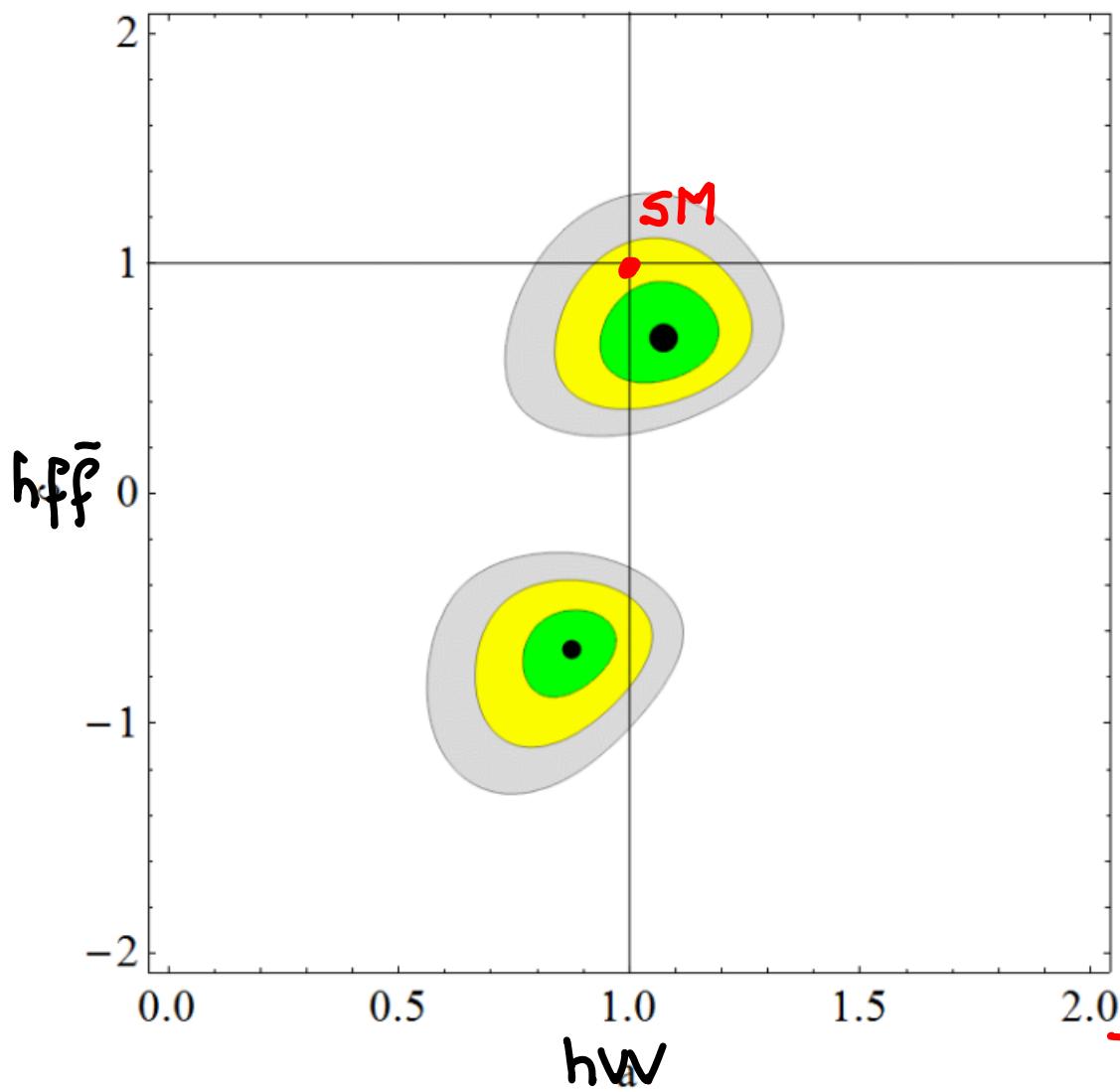


- Singlet fix of instability  
(from see-saw; axion;  
Higgs inflation,...)



# FITS IN SM( $a, c$ )

7&8 LHC data & Tevatron



How to improve SM's fit:

$$|c| < 1 \Rightarrow BR_{b\bar{b}} \downarrow \sim 20\%$$



$$BR_{\gamma\gamma} \uparrow \uparrow \times 2-3$$

$$BR_{WW} \uparrow$$

$$\text{compensating } \sigma_{gg} \downarrow \sim 50\%$$

Two solutions due to approx.

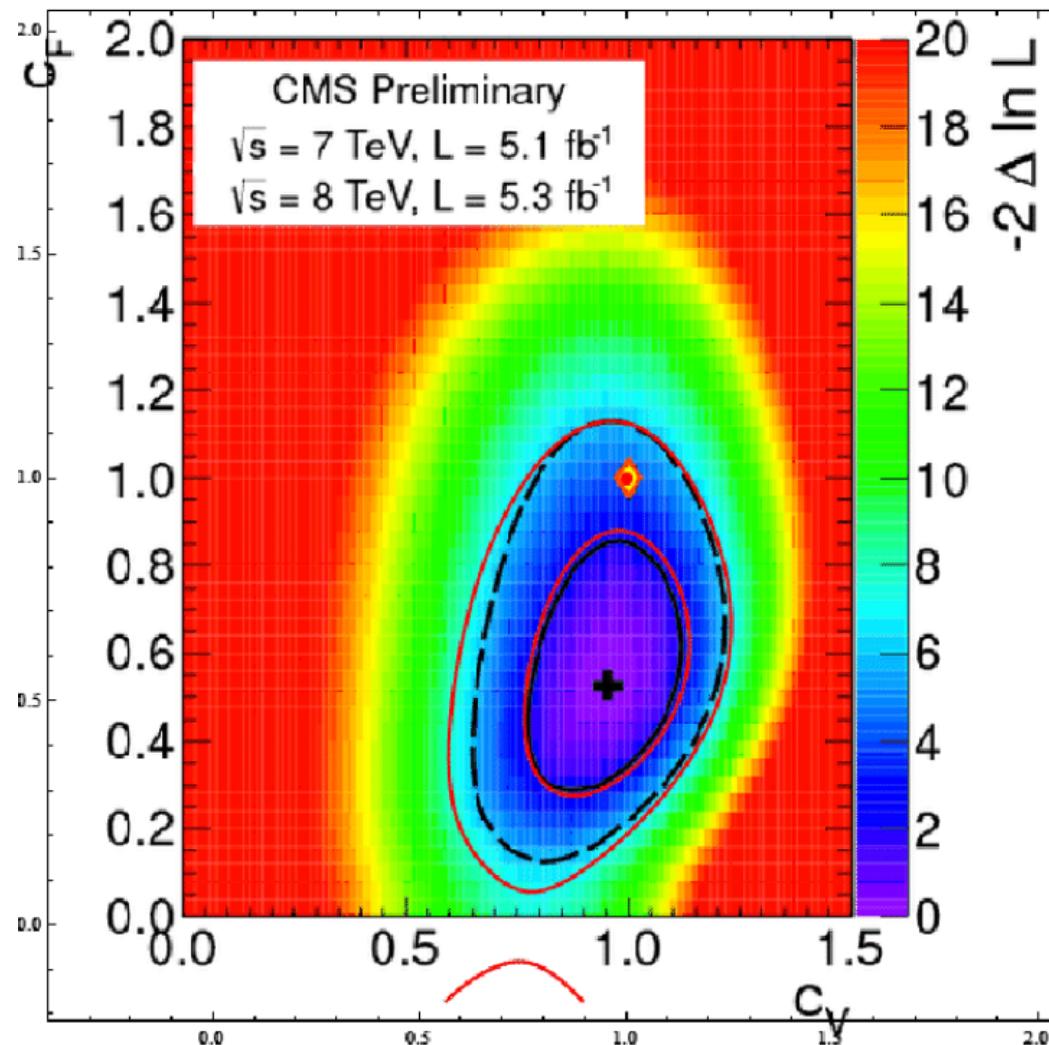
$$a \leftrightarrow -a \quad c \leftrightarrow -c \text{ sym.}$$

Broken by  $\gamma\gamma\gamma$  amplitude

$$\sim |1.26a - 0.26c|^2$$

Constr. interference for  $c < 0$

## COMPARISON WITH CMS

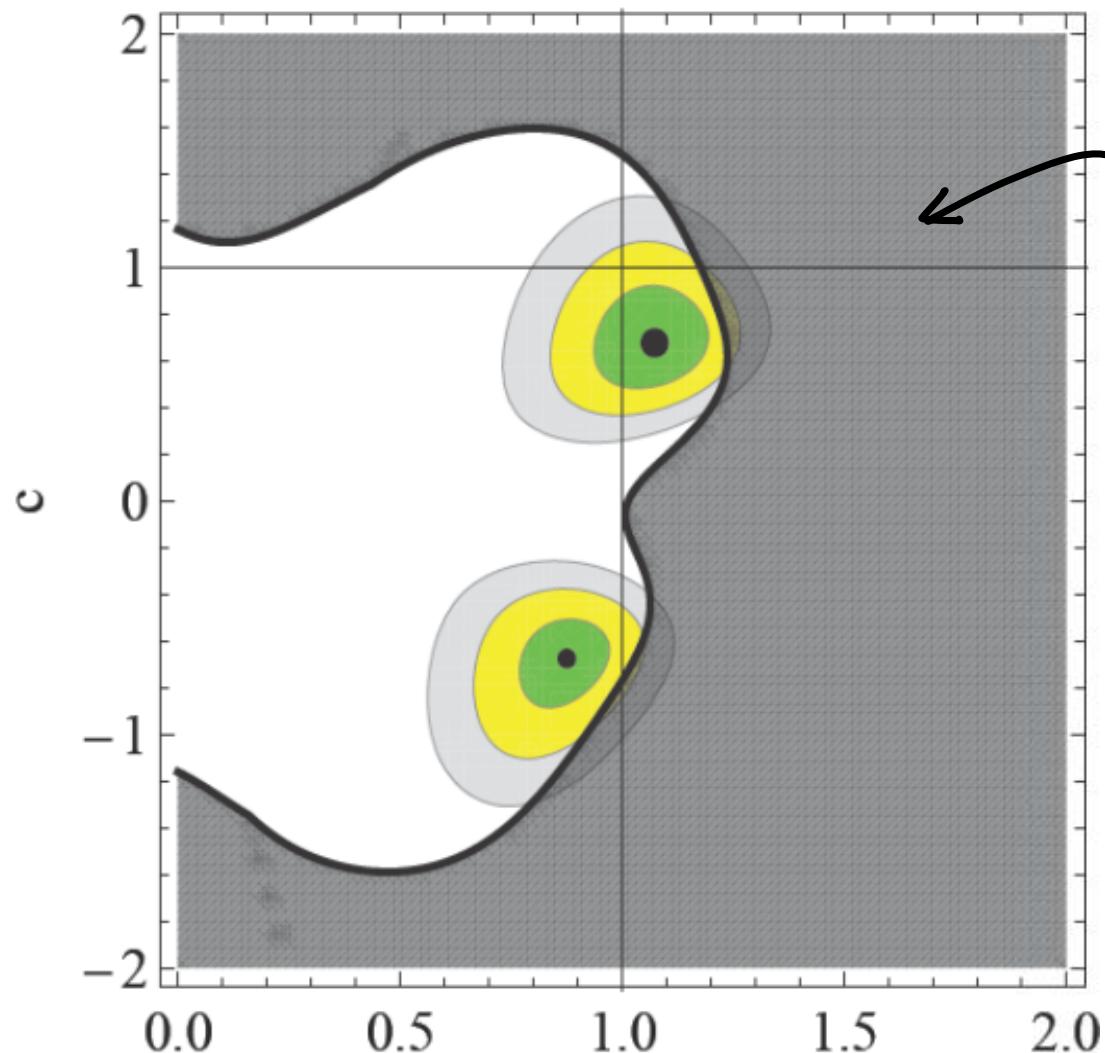


Good agreement

Had to impose prior  $c > 0$ . (Best fit point for  $c < 0$ ).

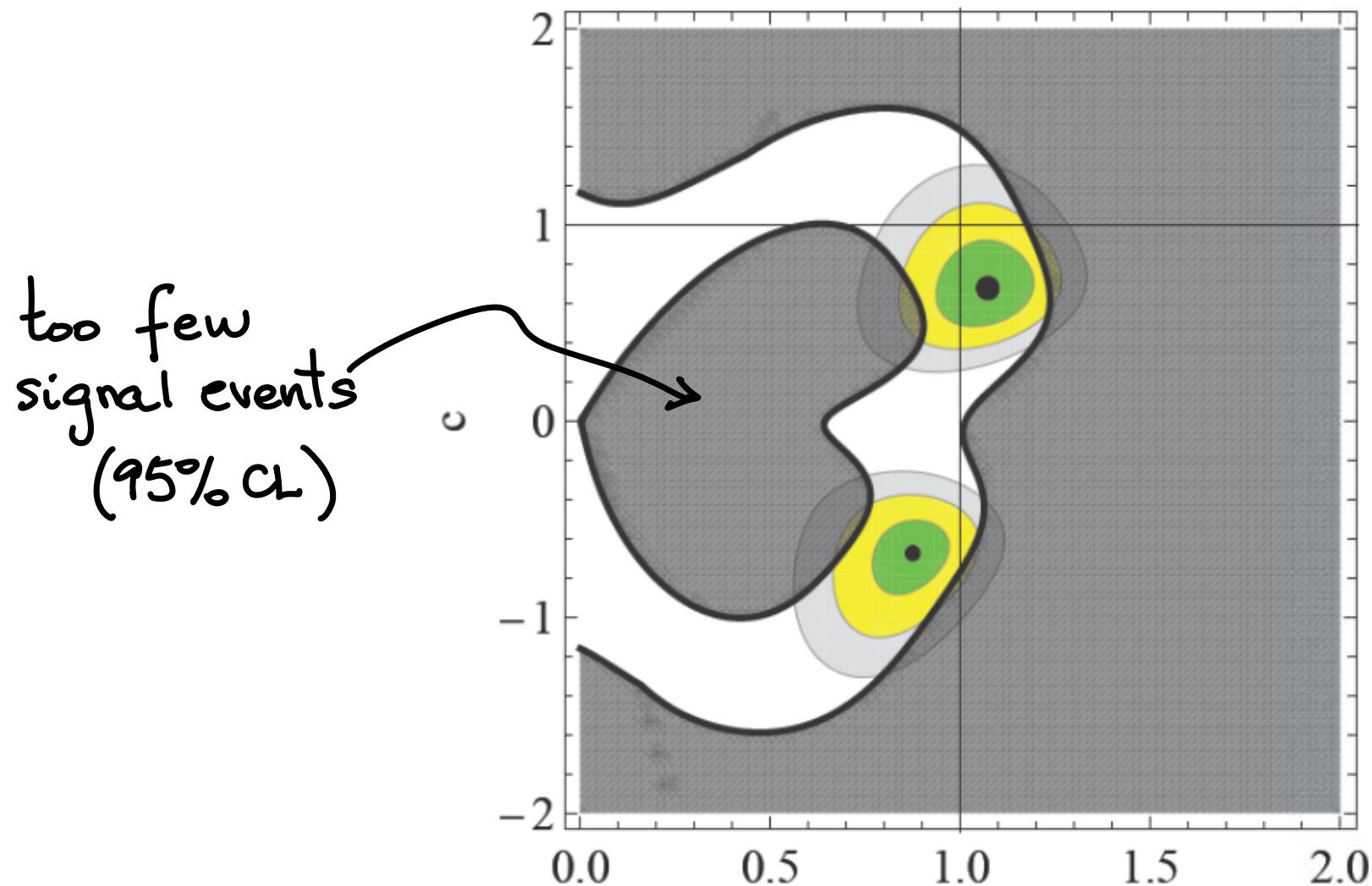
## EXCLUDING SM ( $a, c$ )

7&8 LHC data & Tevatron



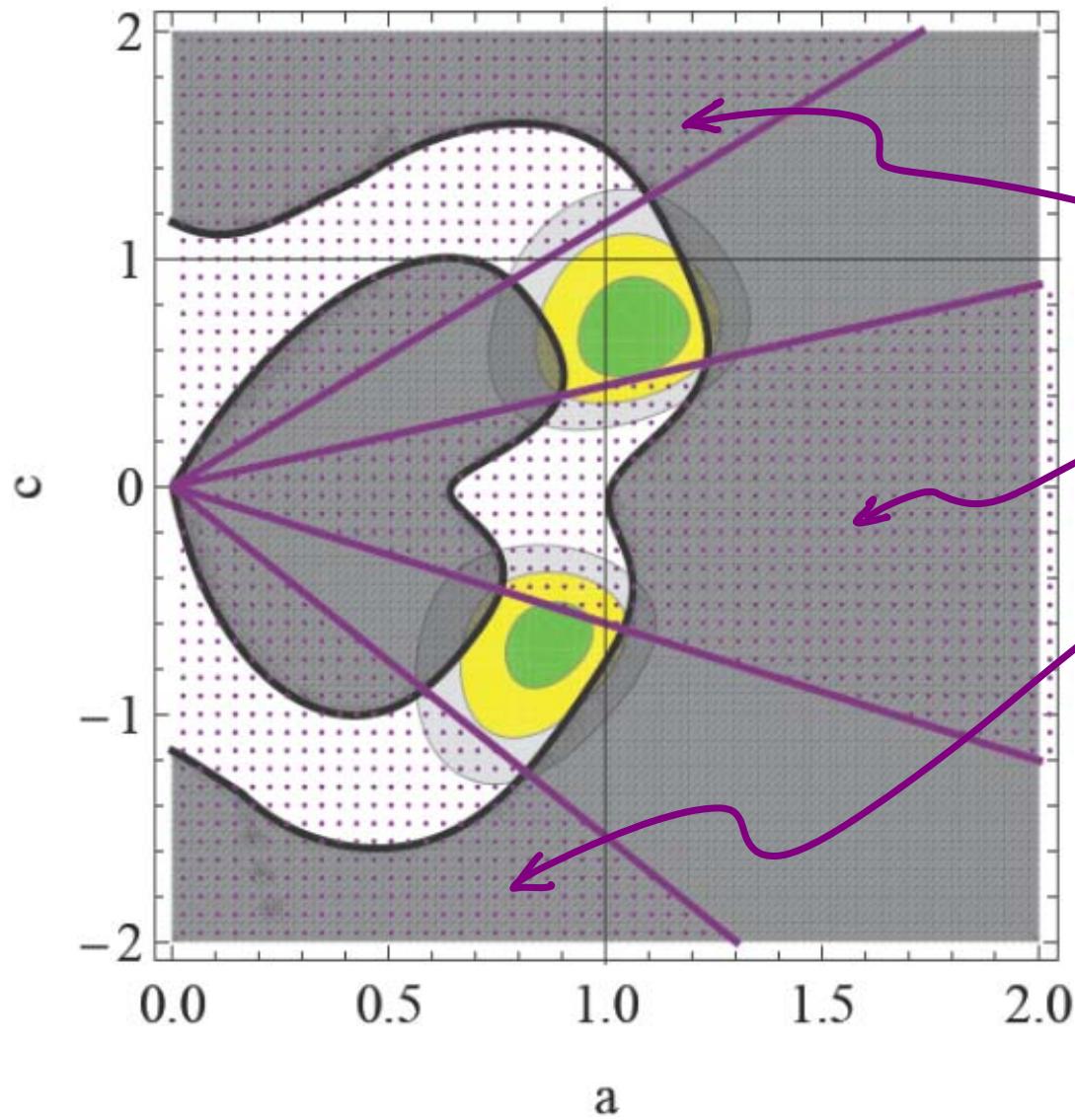
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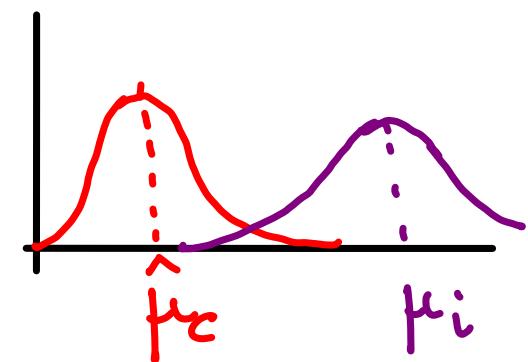


# EXCLUDING SM( $a, c$ )

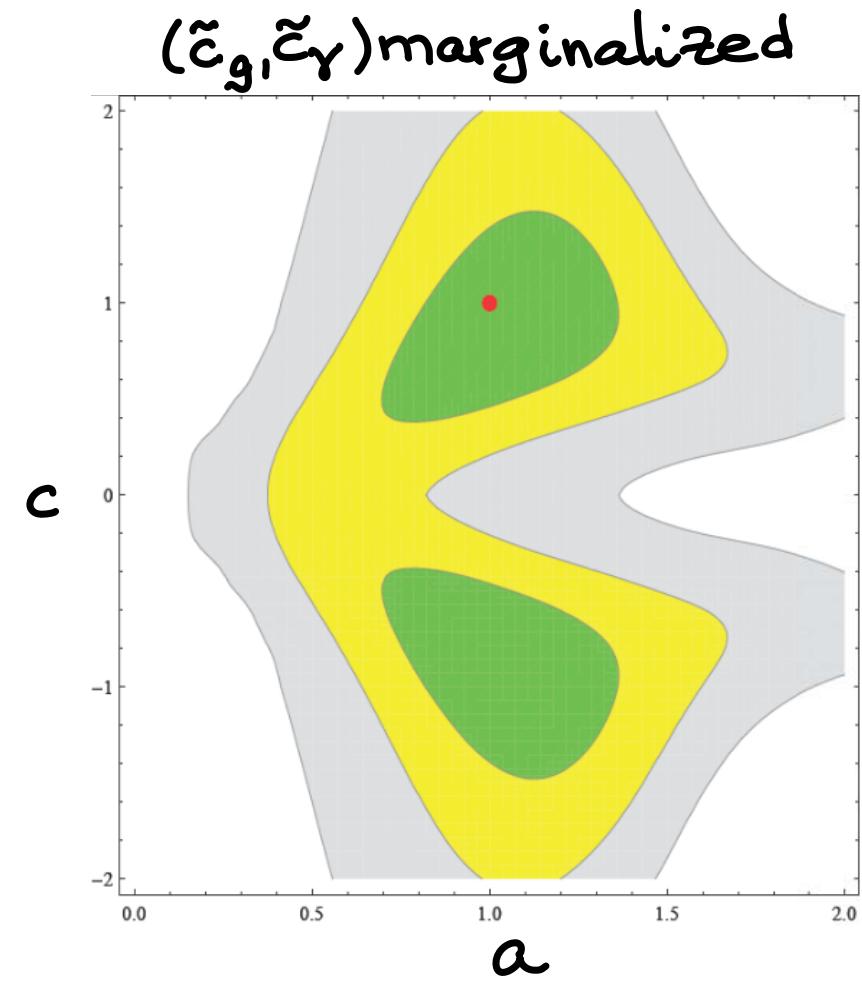
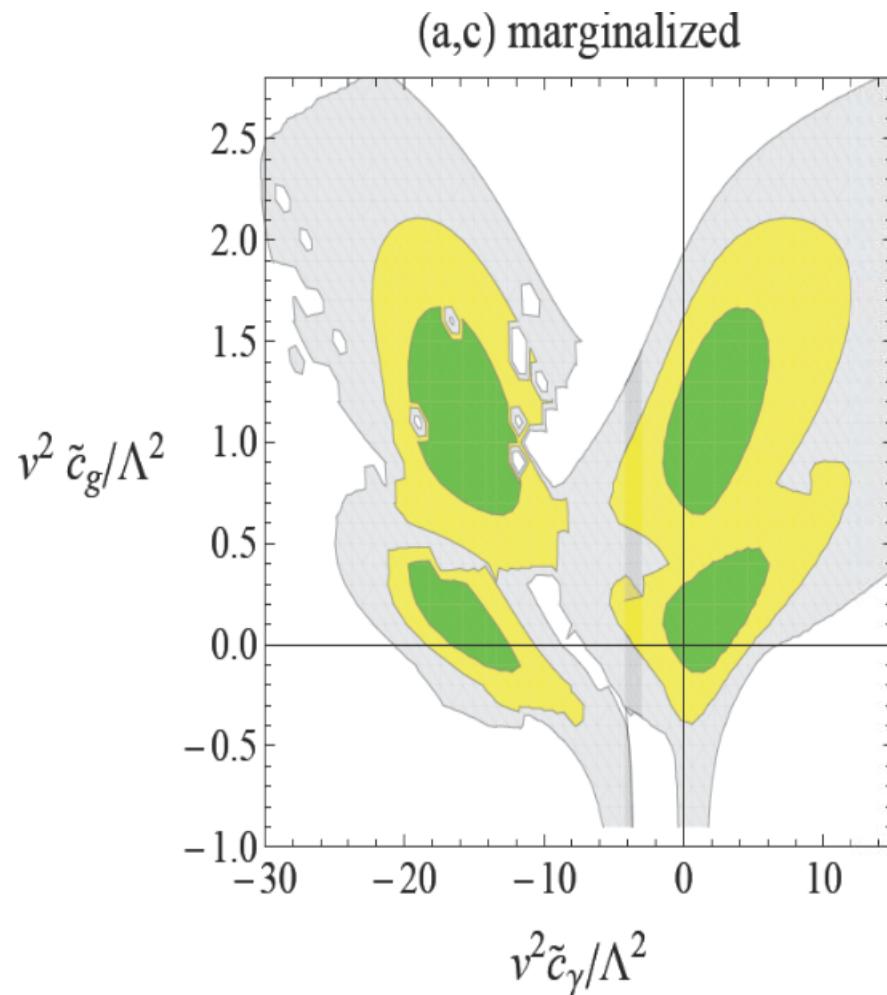
7&8 LHC data & Tevatron



Too much  
tension  
in some  
channels



ANOMALOUS h COUPLINGS  $(a,c) \neq (1,1)$   
 OR LOOP INDUCED EFFECTS  $(\tilde{c}_g, \tilde{c}_\gamma) \neq (0,0)$  ?



Need more data !

(4b)

## HIGGS INVISIBLE WIDTH?

$h \rightarrow XX$  for light (SM singlet)  $\times$  (Higgs portal, DM, ...)

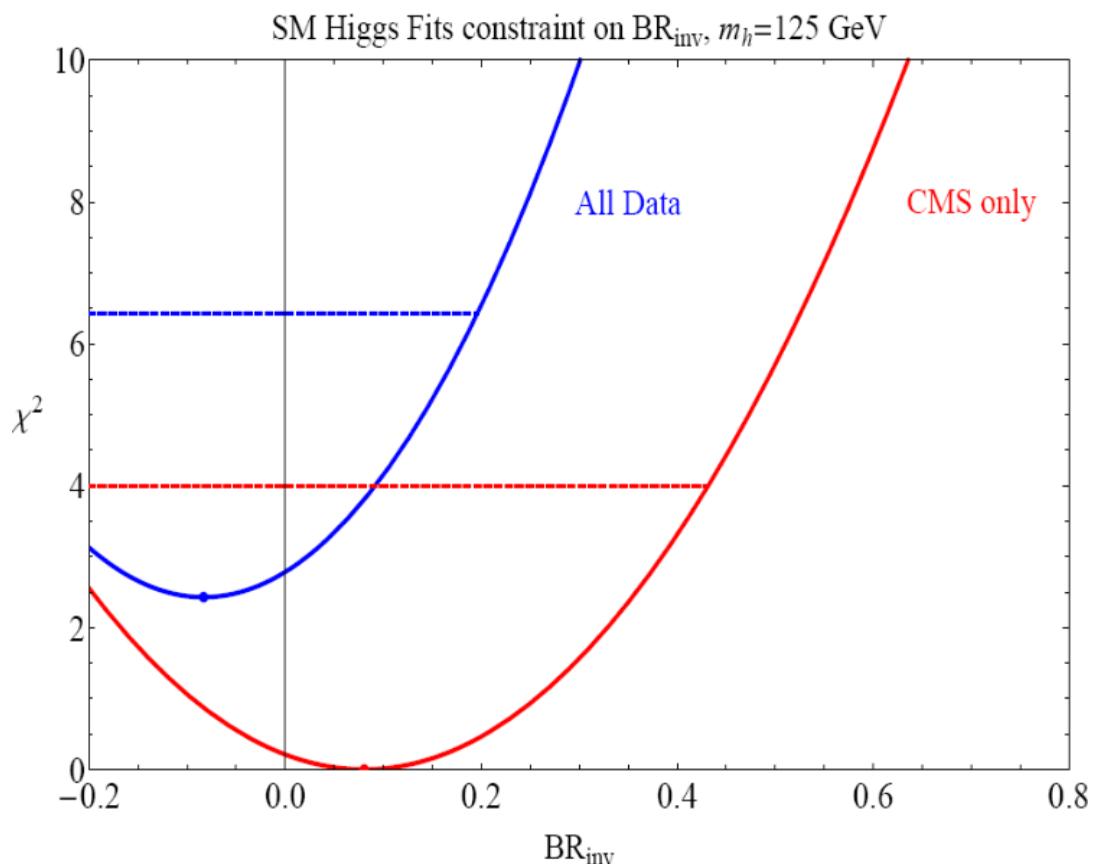
Universal reduction of  $\text{Br}_j = \frac{\mathcal{I}(h \rightarrow j)}{\mathcal{I}_{\text{SM}} + \mathcal{I}_{\text{inv}}} = (1 - \text{BR}_{\text{inv}}) \text{Br}_{\text{j}}^{\text{SM}}$

$\chi^2$  fit to combined  $\hat{\mu}_c \pm \sigma_c$

$\chi^2_{\text{min}}$  at  $\text{BR}_{\text{inv}} \sim (-0.1, 0.1)$

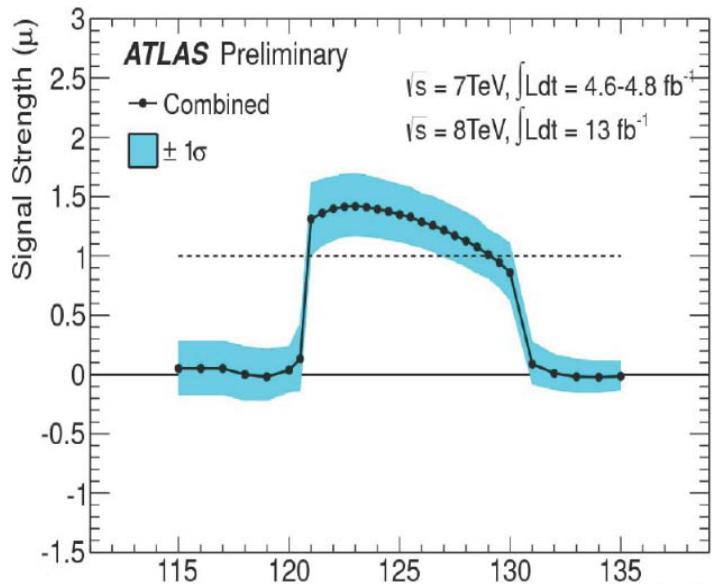
SM with  $\text{BR}_{\text{inv}} = 0$  OK

Still there is room for  
large  $\text{BR}_{\text{inv}}$  ( $\lesssim \frac{0.2}{0.45}$  @ 95% CL)

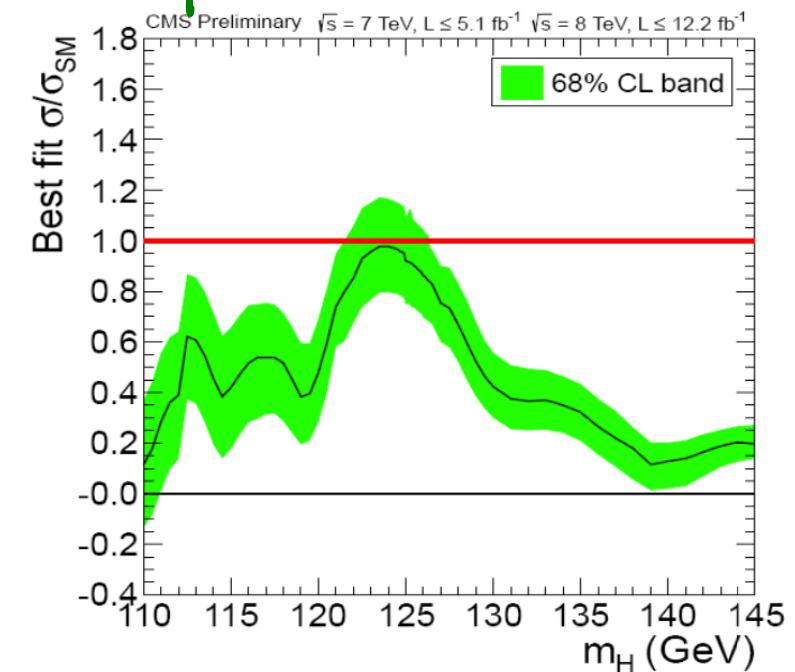


# HIGGS INVISIBLE WIDTH?

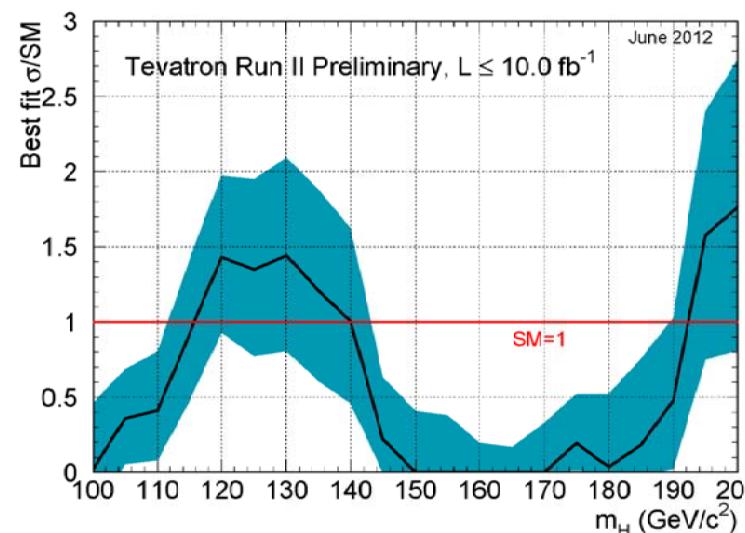
Indirect evidence. Interpret  $\hat{\mu} < 1$  as  $\hat{\mu} = 1 - BR_{inv}$



ATLAS



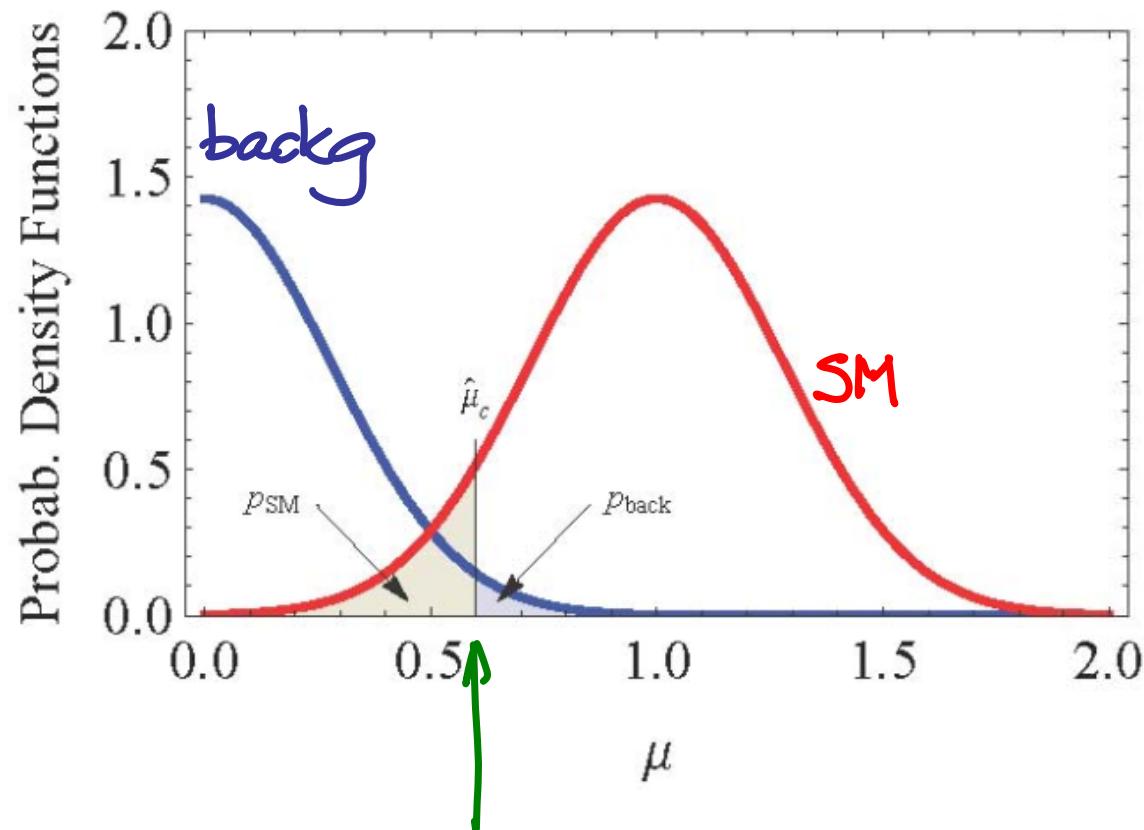
CMS



Tevatron

# HIGGS INVISIBLE WIDTH?

To trust this indirect evidence need  $p_{SM}, p_{backg} \ll 1$

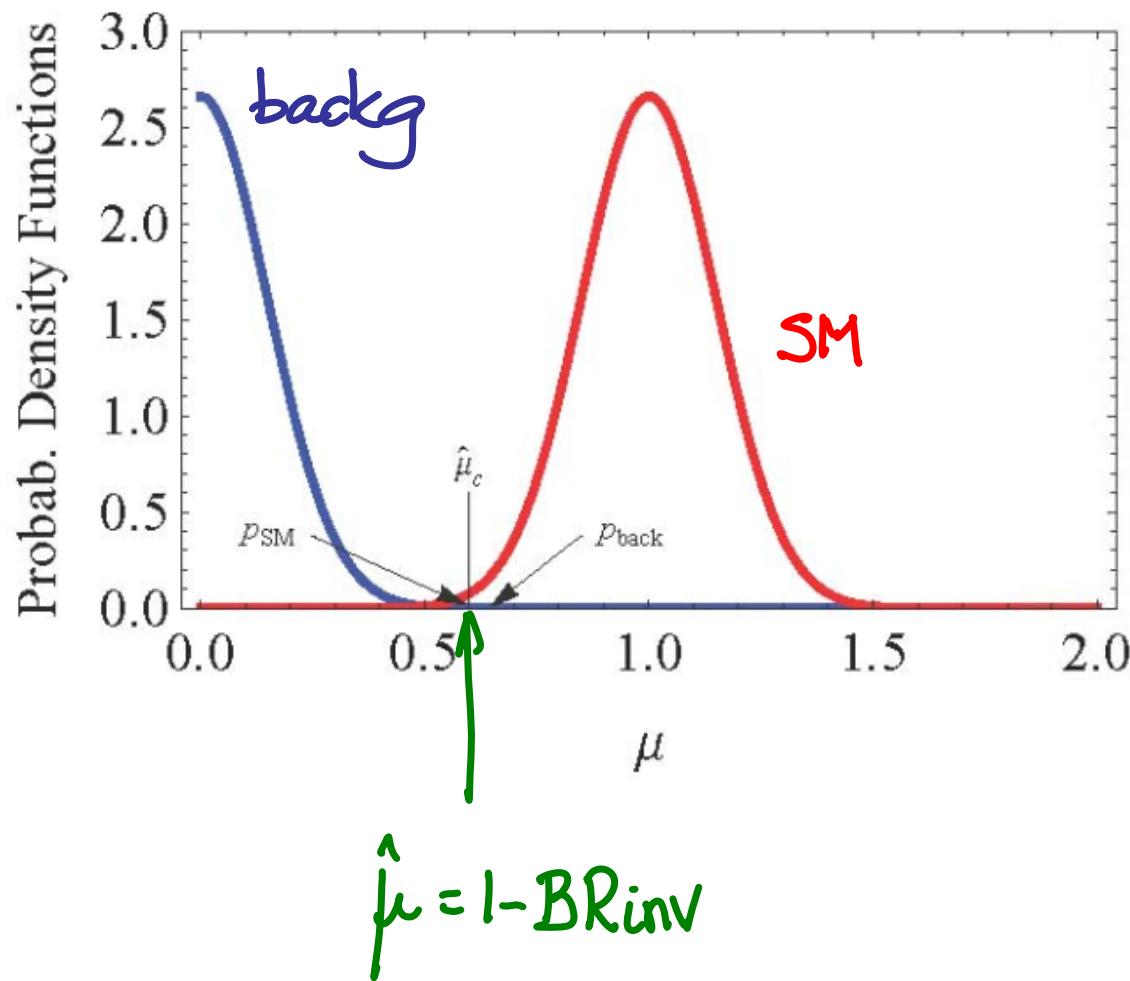


$$\hat{\mu} = 1 - BR_{inv}$$

It's crucial to reduce  
 $q$

# HIGGS INVISIBLE WIDTH?

To trust this indirect evidence need  $p_{SM}, p_{backg} \ll 1$   
Must also give small  $\chi^2_{min}$



It's crucial to reduce  
 $q$

$$\frac{1}{q^2} = \sum_i \frac{N_{ch}}{\sigma_i^2} \Rightarrow$$

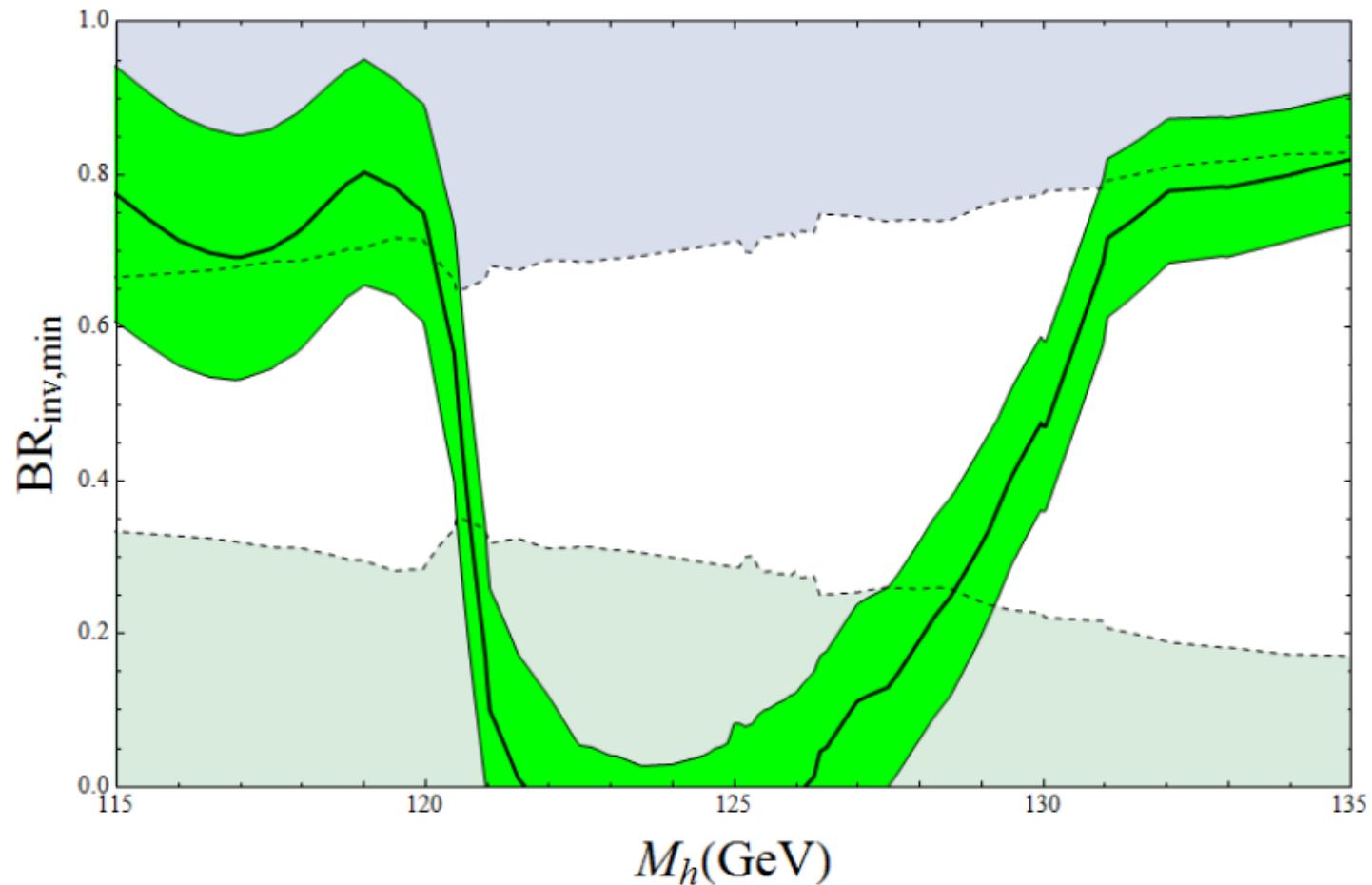
$$\sigma \sim \sqrt{\frac{q}{N_{ch}}}$$

$$\sigma_i \sim \frac{\sqrt{n_{b,i}}}{n_{S,i}^{SM}} \sim \frac{1}{\sqrt{\mathcal{L}}}$$

# HIGGS INVISIBLE WIDTH?

Current situation as a function of  $m_h$

$$\text{BR}_{\text{inv}} = 1 - \hat{\mu}_c$$

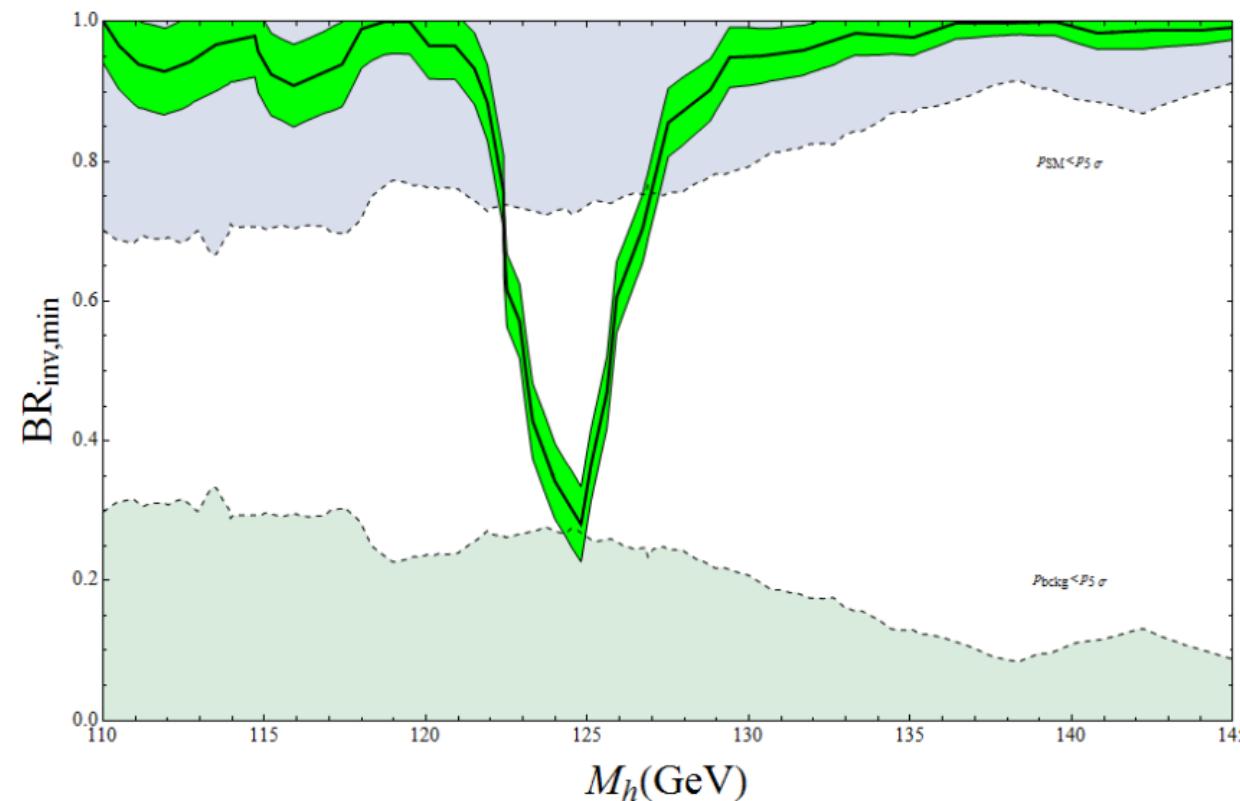


No strong conclusion possible

# HIGGS INVISIBLE WIDTH?

Future situation ( $\sigma_c \sim \frac{0.3}{5}$  for  $m_h = 125$  GeV)

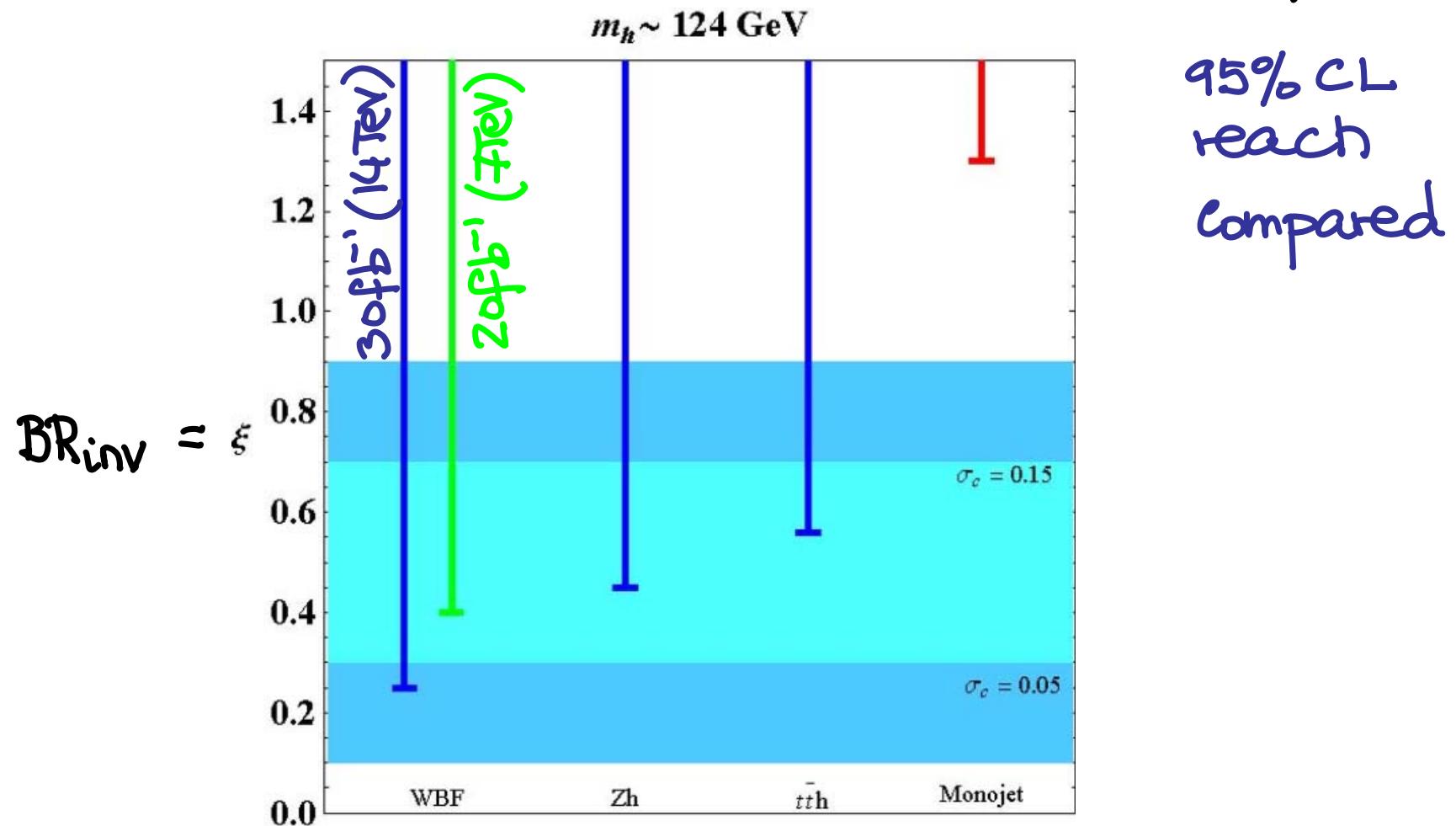
Hypothetical observed  $\hat{\mu}$ , of course...



# HIGGS INVISIBLE WIDTH?

How to tell apart  $\text{BR}_{\text{inv}} > 0$  from e.g.  $a=c < 1$  ?

Must validate with direct searches for inv. Higgs decays

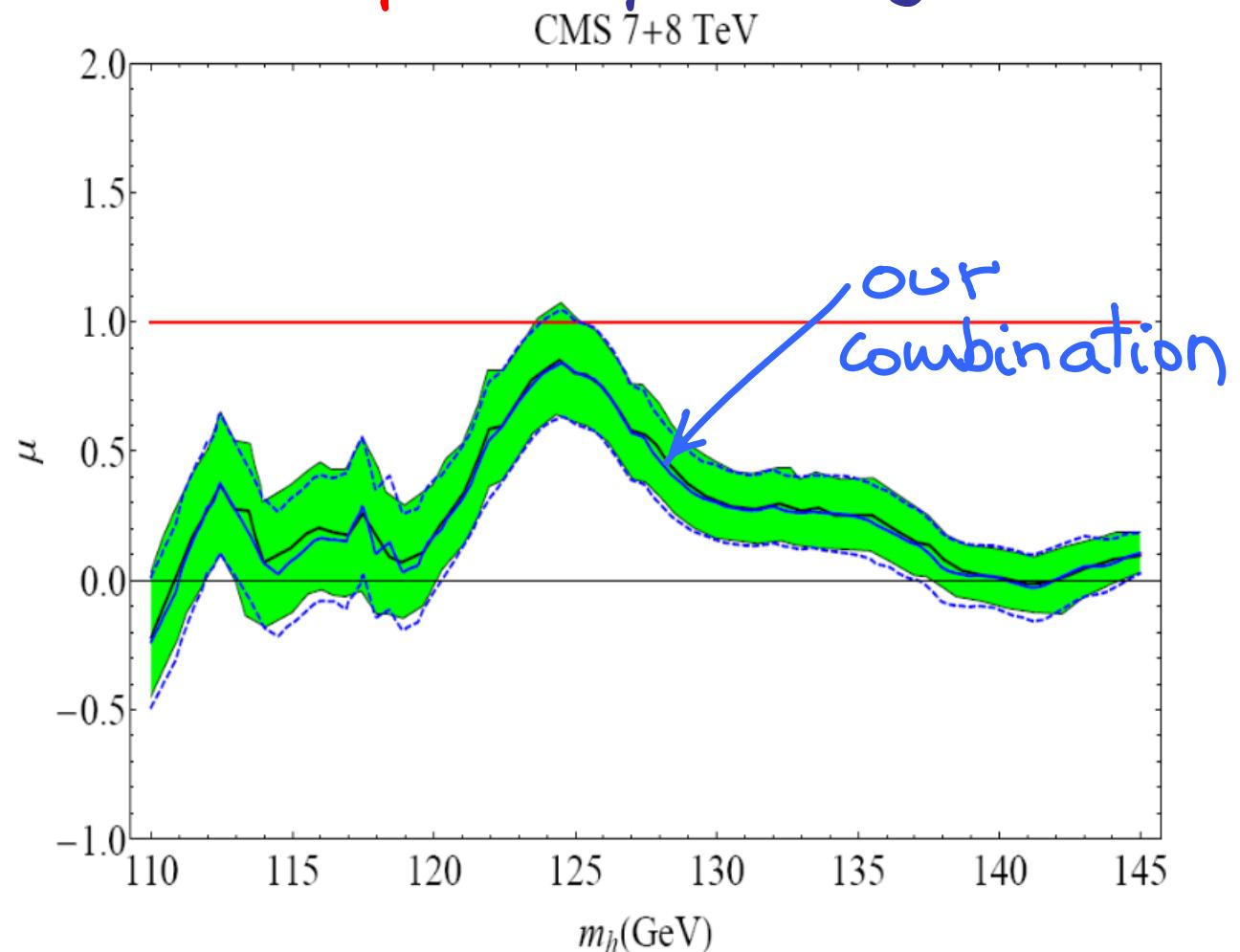


# RATES IN SM( $a, c$ )

- ★ Need to know efficiencies  $\epsilon_{p_i}$  to properly rescale  $\sigma_i$
- ★ Need to know measured  $\hat{\mu}$ 's separately at 7, 8 TeV

We checked  
this works well :

official CMS  
combination  
at 7+8 TeV



# ASKING THE HIGGS ABOUT STOPS

