Bimetric Theories: Ghost-Free Spin-2 Interactions

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Collaborators:

SFH, A. Schmidt-May, M. von Strauss

arXiv:1109.3230 (with J. Enander, E. Mörtsell), 1203.5283, 1204.5202, 1208:1515, 1208:1797, 1212:4525, 1303.6940, 1406.xxxx (to appear)

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▶ SFH, R. A. Rosen,

arXiv:1103.6055, 1106.3344, 1109.3515 (with Schmidt-May), 1109.3230, 1111.2070

Outline of the talk

Motivation

Generic massive & interacting spin-2 fields

The Ghost Problem

Ghost-free bimetric theory

Bimetric Theory & Conformal Gravity

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Motivation

New physics beyond the Standard Model (SM) and General Relativity (GR):

> Dark matter, Dark energy (the cosmological constant problem), Matter-Antimatter asymmetry, Neutrinos, Inflation, Quantum gravity, ···

- New physics needs new tools (theoretical models): Supersymmetry, String Theory, Something yet unknown ?
- Spin-2 physics as a tool for new physics:

Modifies gravity, orthogonal to other approaches

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The Spin Inventory

Spin 0, 1/2, 1: Theories exist for massless, massive fields. Interactions are known. Building blocks of SM.

$$\exists \mathsf{x}: \qquad \sqrt{|\det g|}(F_{\mu\nu}F^{\mu\nu}-m^2\,g^{\mu\nu}A_{\mu}A_{\nu})$$

Spin 2: Theory for a single, massless field $g_{\mu\nu}$:

Einstein-Hilbert action: $\sqrt{|\det g|} R(g)$

Are there consistent theories for massive or multiple spin-2 fields? Neutral? Charged? Higher Spins?

New theories beyond SM & GR, but the ghost problem!

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Linear massive spin-2 fields

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

Massless spin-2:

Non - linear :
$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 0$$
Linear : $\mathcal{E}^{\rho\sigma}_{\mu\nu}h_{\rho\sigma} \equiv \Box h_{\mu\nu} + \cdots = 0$ 2 polarization modes ($\pm 2, \pm X, \emptyset$)

Massive spin-2:

Non-linear extension: The Boulware-Deser Ghost (1972)

Nonlinear massive spin-2 fields

• "Massive gravity" (needs a new field $f_{\mu\nu}(x)$)

$$\mathcal{L}=m_
ho^2\sqrt{-g}\left[R-m^2\,V(f^{-1}g)
ight],\qquad V\sim f^{\mu
ho}\,f^{
u\sigma}\,g_{\mu
u}\,g_{
ho\sigma}+\cdots$$

Interacting spin-2 fields (dynamical g_{μν} and f_{μν})

$$\mathcal{L} = m_p^2 \sqrt{-g} \left[R - m^2 V(f^{-1}g) \right] + \mathcal{L}(f, \nabla f)(?)$$

Bimetric: $\mathcal{L}(f, \nabla f) = m_f^2 \sqrt{-f} R_f(?)$ [Isham-Salam-Strathdee, 1971]

Generically, both contain a nonlinear \mathcal{GHOST}

[Boulware-Deser, 1972]

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The Ghost Problem

Ghost: A field with negative kinetic energy

Example:

$$\mathcal{L} = \mathbf{T} - \mathbf{V} = (\partial_t \phi)^2 \cdots$$
 (healthy)

But

$$\mathcal{L} = T - V = -(\partial_t \phi)^2 \cdots$$
 (ghost)

Consequences:

- Negative probabilities, violation of unitarity
- Instability: unlimited energy transfer from ghost to other fields

Ghost in Spin-2 Theories:

Generic massive gravity:

- Linear: 5 modes
- Non-linear: 5 + 1 (ghost)

Generic bimetric theory:

- Linear: 5 (massive) + 2 (massless) modes
- Non-linear: 7 + 1 (ghost)

Do ghost-free massive gravity & bimetric theories exist?

Complication: Since the ghost shows up nonlinearly, its absence needs to be established nonlinearly

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Construction of ghost-free massive gravity

Development of "Decoupling limit" analysis

[Creminelli, Nicolis, Papucci, Trincherini, (hep-th/0505147)]

Ghost-free massive gravity proposal

[de Rham, Gabadadze, (1007.0443); de Rham, Gabadadze, Tolley, (1011.1232)]

$$V_{dRGT}\left(\sqrt{g^{-1}\,\eta}
ight)\,,$$

 Proof of absence of ghost, generalization to V(\sqrt{g^{-1} f}) [SFH, Rosen (1106.3344, 1111.2070)]
 [SFH, Rosen, Schmidt-May (1109.3230)]

Interacting spin-2 fields g & f (Bimetric theory)

[SFH, Rosen (1109.3515)]

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Multiple spin-2 fields

[Hinterbichler, Rosen (1203.5783)]

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Ghost-free bimetric theory

A dynamical theory for spin-2 fields $g_{\mu
u}$ & $f_{\mu
u}$

$$\mathcal{L} = m_{\rho}^2 \sqrt{-g} \left[R - m^2 V(g^{-1}f) \right] + \mathcal{L}(f, \nabla f)$$

- What is $V(g^{-1}f)$?
- what is $\mathcal{L}(f, \nabla f)$?
- Proof of absence of ghost
- What are the implications of the theory?

Digression: Elementary symmetric polynomials of X with eigenvalues $\lambda_1, \dots, \lambda_4$:

$$\begin{split} e_0(\mathbb{X}) &= 1, \qquad e_1(\mathbb{X}) = \lambda_1 + \lambda_2 + \lambda_3 + \lambda_4, \\ e_2(\mathbb{X}) &= \lambda_1 \lambda_2 + \lambda_1 \lambda_3 + \lambda_1 \lambda_4 + \lambda_2 \lambda_3 + \lambda_2 \lambda_4 + \lambda_3 \lambda_4, \\ e_3(\mathbb{X}) &= \lambda_1 \lambda_2 \lambda_3 + \lambda_1 \lambda_2 \lambda_4 + \lambda_1 \lambda_3 \lambda_4 + \lambda_2 \lambda_3 \lambda_4, \\ e_4(\mathbb{X}) &= \lambda_1 \lambda_2 \lambda_3 \lambda_4 = \det \mathbb{X}. \end{split}$$

$$\begin{split} e_{0}(\mathbb{X}) &= 1, \qquad e_{1}(\mathbb{X}) = [\mathbb{X}], \\ e_{2}(\mathbb{X}) &= \frac{1}{2}([\mathbb{X}]^{2} - [\mathbb{X}^{2}]), \\ e_{3}(\mathbb{X}) &= \frac{1}{6}([\mathbb{X}]^{3} - 3[\mathbb{X}][\mathbb{X}^{2}] + 2[\mathbb{X}^{3}]), \\ e_{4}(\mathbb{X}) &= \frac{1}{24}([\mathbb{X}]^{4} - 6[\mathbb{X}]^{2}[\mathbb{X}^{2}] + 3[\mathbb{X}^{2}]^{2} + 8[\mathbb{X}][\mathbb{X}^{3}] - 6[\mathbb{X}^{4}]), \\ e_{k}(\mathbb{X}) &= 0 \qquad \text{for} \quad k > 4, \\ & [\mathbb{X}] = \text{Tr}(\mathbb{X}), \qquad e_{n}(\mathbb{X}) \sim (\mathbb{X})^{n} \end{split}$$

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• The $e_n(\mathbb{X})$'s and $det(\mathbb{1} + \mathbb{X})$:

$$\det(\mathbb{1}+\mathbb{X})=\sum\nolimits_{n=0}^{4}e_{n}(\mathbb{X})$$

Introduce "deformed determinant" :

$$\widehat{\det}(\mathbb{1}+\mathbb{X})=\sum\nolimits_{n=0}^{4}\beta_{n}\,e_{n}(\mathbb{X})$$

• The $e_n(\mathbb{X})$'s and $det(\mathbb{1} + \mathbb{X})$:

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Introduce "deformed determinant" :

$$\widehat{\det}(\mathbb{1}+\mathbb{X})=\sum_{n=0}^{4}\frac{\beta_{n}}{\rho_{n}}e_{n}(\mathbb{X})$$

Observation:

$$V\left(\sqrt{g^{-1}f}\right) = \sum_{n=0}^{4} \beta_n \, e_n\left(\sqrt{g^{-1}f}\right)$$

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Ghost-free "bi-metric" theory

Ghost-free combination of *kinetic* and *potential* terms for *g* & *f*:

$$\mathcal{L} = m_g^2 \sqrt{-g} R_g - 2m^4 \sqrt{-g} \sum_{n=0}^4 \beta_n \, e_n(\sqrt{g^{-1}f}) + m_f^2 \sqrt{-f} \, R_f$$

[SFH, Rosen (1109.3515,1111.2070)]

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Symmetry under $f \leftrightarrow g$,

$$\sqrt{-g}\sum_{n=0}^{4}\beta_{n}\,e_{n}(\sqrt{g^{-1}f})=\sqrt{-f}\sum_{n=0}^{4}\beta_{4-n}\,e_{n}(\sqrt{f^{-1}g})$$

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Symmetry under $f \leftrightarrow g$,

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Hamiltonian analysis: 7 nonlinear propagating modes, no ghost!

$$C(\gamma,\pi)=0$$
, $C_2(\gamma,\pi)=\frac{d}{dt}C(x)=\{H,C\}=0$

Ghost-free Matter couplings

Similar to GR minimal couplings:

 $\mathcal{L}_{g}(g,\phi) + \mathcal{L}_{f}(f,\tilde{\phi})$

Convention:

 $g_{\mu
u}$: the gravitational metric, ϕ : observed matter. Then,

 $m_g \sim M_P$ (Planck mass)

Equations of motion:

$$\begin{split} m_g^2 \left[R_{\mu\nu}(g) - \frac{1}{2} g_{\mu\nu} R(g) \right] + V_{\mu\nu}^g(g,f) &= T_{\mu\nu}^g(g,\phi) \\ m_f^2 \left[R_{\mu\nu}(f) - \frac{1}{2} f_{\mu\nu} R(f) \right] + V_{\mu\nu}^f(g,f) &= T_{\mu\nu}^f(f,\tilde{\phi}) \end{split}$$

Matter equations of motion are unchanged (geodesic motion)

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Physical content (Mass spectrum)

The theory has 7 propagating modes. Consider

$$g_{\mu
u} = ar{g}_{\mu
u} + \delta g_{\mu
u} , \quad f_{\mu
u} = ar{f}_{\mu
u} + \delta f_{\mu
u}$$

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Well defined mass spectrum (FP masses) exists for $\overline{f} = c^2 \overline{g}$, Linear modes:

Physical content (Mass spectrum)

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Well defined mass spectrum (FP masses) exists for $\overline{f} = c^2 \overline{g}$, Linear modes:

Massless spin-2 (2): $\delta G_{\mu\nu} = \left(\delta g_{\mu\nu} + \frac{m_f^2}{m_g^2} \delta f_{\mu\nu}\right)$ Massive spin-2 (5): $\delta M_{\mu\nu} = \left(\delta f_{\mu\nu} - c^2 \delta g_{\mu\nu}\right)$,

• m_{FP}^2 , c^2 and Λ are given in terms of the β_0, \dots, β_4 .

 g_{μν}, f_{μν} are mixtures of massless and massive modes [SFH, Schmidt-May, von Strauss 1208:1515, 1212:4525]

Limits of bimetric theory (m_g vs m_f)

dRGT Massive gravity limit (fixed $f_{\mu\nu}$):

$$m_g = M_P \,, \qquad m_f^2 \to \infty$$

Most of the recent work has been done for this theory.

The General Relativity limit:

$$m_g = M_P \,, \qquad m_f o 0$$

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More natural and phenomenologically viable.

Bimetric as "gravity + a massive spin-2 field"

► Gravity:

 $g_{\mu\nu}$ coupled to observed matter (not purely massless)

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• Massive spin-2 field: $M_{\mu\nu} = g_{\mu\rho} (\sqrt{g^{-1}f})^{\rho}_{\ \nu} - c g_{\mu\nu}$ coupled to gravity Bimetric as "gravity + a massive spin-2 field"

► Gravity:

 $g_{\mu\nu}$ coupled to observed matter (not purely massless)

• Massive spin-2 field: $M_{\mu\nu} = g_{\mu\rho} (\sqrt{g^{-1}f})^{\rho}_{\ \nu} - c g_{\mu\nu}$ coupled to gravity

For $M_P = m_g >> m_f$: $g_{\mu\nu}$ is mostly massless!

In this setup, the *long range* of gravitational force is correlated with gravity being the *weakest* force in the spin-2 sector

What next?

Observational considerations:

Viability of applications to cosmology and gravity, classical solutions, cosmological perturbations, gravitationsl waves, etc. (Old motivations may not be met)

[See talk by Luigi Pilo]

Theoretical considerations:

Deeper understanding of structure, origin of masses, dynamics, generalizations, the charged case, limitations, etc. Relation to "partial masslessness" and Conformal gravity

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Digression I: Conformal gravity

Higher Derivative gravity:

$$S^{ ext{HD}}_{(2)}[g]=m_g^2\int d^4x\sqrt{g}\left[\Lambda+c_1R(g)-rac{c_2}{m^2}\left(R^{\mu
u}R_{\mu
u}-rac{1}{3}R^2
ight)
ight]$$

7 modes: massless spin-2 + massive spin-2 (ghost) [Stelle (1977)]

Set:
$$\Lambda = 0, c_1 = 0 \implies$$
 Conformal gravity

Digression I: Conformal gravity Conformal Gravity:

$$\mathcal{S}^{ ext{CG}}[g] = c \int d^4x \sqrt{g} \left[R^{\mu
u} R_{\mu
u} - rac{1}{3} R^2
ight] = c \int d^4x \sqrt{g} \mathcal{W}^2 \,,$$

(similar to the Yang-Mills action)

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Invariance:

$$g_{\mu
u}
ightarrow oldsymbol{e}^{\phi} oldsymbol{g}_{\mu
u}$$

EoM (Bach tensor):

$$egin{aligned} B_{\mu
u} &\equiv -
abla^2 P_{\mu
u} -
abla_\mu
abla
u P_{\mu
u} &= R_{\mu
u} - rac{1}{6}g_{\mu
u}R \end{aligned}$$

Spectrum: 6 modes: 2 (massless spin-2) + 4 ghosts [Riegert (1984), Maldacena (2011)] Digression II: Partially massless FP theory Back to the Fierz-Pauli equation:

$$ar{\mathcal{E}}^{
ho\sigma}_{\mu
u}\,h_{
ho\sigma}-oldsymbol{\Lambda}ig(h_{\mu
u}-rac{1}{2}ar{g}_{\mu
u}h_{
ho}^{
ho}ig)+rac{m_{
m Fp}^2}{2}ig(h_{\mu
u}-ar{g}_{\mu
u}h_{
ho}^{
ho}ig)=0$$

 $\bar{g}_{\mu\nu}$: dS/Einstein background.

Higuchi Bound:

$$m_{FP}^2 = \frac{2}{3}\Lambda$$

New gauge symmetry:

$$\Delta h_{\mu
u} = (
abla_{\mu}
abla_{
u} + rac{\Lambda}{3})\xi(x)$$

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Gives 5-1=4 propagating modes [Deser, Waldron, · · · (1983-2012)]

Non-linear extension? Relation to Conformal Gravity?

Curvature expansion of bimetric equations [SFH, Schmidt-May, von Strauss, 1303:6940]

$$m_g^2 \left[R_{\mu\nu}(g) - \frac{1}{2} g_{\mu\nu} R(g) \right] + V_{\mu\nu}^g(g, f) = 0$$
(1)
$$m_f^2 \left[R_{\mu\nu}(f) - \frac{1}{2} f_{\mu\nu} R(f) \right] + V_{\mu\nu}^f(g, f) = 0$$
(2)

From (1), (
$$P_{\mu\nu} = R_{\mu\nu} - \frac{1}{6}g_{\mu\nu}R$$
),
 $f_{\mu\nu} = a g_{\mu\nu} + \frac{b}{m^2}P^{\mu}_{\ \nu} - \frac{c}{m^4}$ " P^2 "…

Substitute in (2),

 $A g_{\mu\nu} + \frac{B}{m^2} \left[R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R \right] + \frac{C}{m^4} B_{\mu\nu} + \frac{D}{m^4} "P^2 " + \frac{"P^3"}{m^6} + \dots = 0$ Can A = B = D = 0 ? Yes!

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A ghost-free generalization of conformal gravity? When parameters $\beta_0, \dots \beta_4$ satisfy ($\alpha = m_f/m_g$),

$$\alpha^4\beta_0 = 3\alpha^2\beta_2 = \beta_4, \qquad \beta_1 = \beta_3 = 0$$

the bimetric EoM's imply

$$B_{\mu\nu}(g) + \mathcal{O}(R^3(g)/m^2) = 0$$

Invariance in low curvature limit:

 $\Delta g_{\mu\nu} = \phi g_{\mu\nu} + \cdots$

But now the parent bimetric theory is ghost-free!

Symmetry at higher orders? Always 6 propagating modes (rather then 7 modes) ?

(Note: No massive gravity limit $\alpha \to \infty$!)

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Extended Weyl invariance

Explicit symmetry to order ∂^6/m^6 (for $m_g = m_f$):

$$\Delta g_{\mu\nu} = \phi g_{\mu\nu} - \hat{m}^{-2} (P_{\mu\nu}\phi + \nabla_{\mu}\nabla_{\nu}\phi) + \hat{m}^{-4}(\partial^{4}) + \hat{m}^{-6}(\partial^{6}) + \cdots$$

$$\Delta f_{\mu\nu} = \phi f_{\mu\nu} - \hat{m}^{-2} (P'_{\mu\nu}\phi + \nabla'_{\mu}\nabla'_{\nu}\phi) + \hat{m}^{-4}(\partial^{4}) + \hat{m}^{-6}(\partial^{6}) + \cdots$$

[SFH, Schmidt-May, von Strauss, 1406:xxxx]

PM symmetry in dS/Einstein backgrounds $\bar{f} = c^2 \bar{g}$,

$$\Delta M_{\mu\nu} = (\nabla_{\mu}\nabla_{\nu} + \frac{\Lambda}{3})\xi(x)$$
$$\Delta G_{\mu\nu} = 0$$

The specific bimetric theory (or an appropriate generalization) could provide a nonlinear realization Partial Masslessness (away from Einstein backgrounds) as well as a ghost-free extension of Conformal Gravity (more work needed).

Discussion

- Ghost-free spin-2 interactions
- Relevance to CG and PM theories.
- Relation between metric and vielbein formulations
- Derivative interactions, changed spin-2 fields
- Causal structure of theories with 2 "metrics", related consistency issues.

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- Origin of the interactions (a la Higgs)
- General results on dynamics and viability

Thank you!