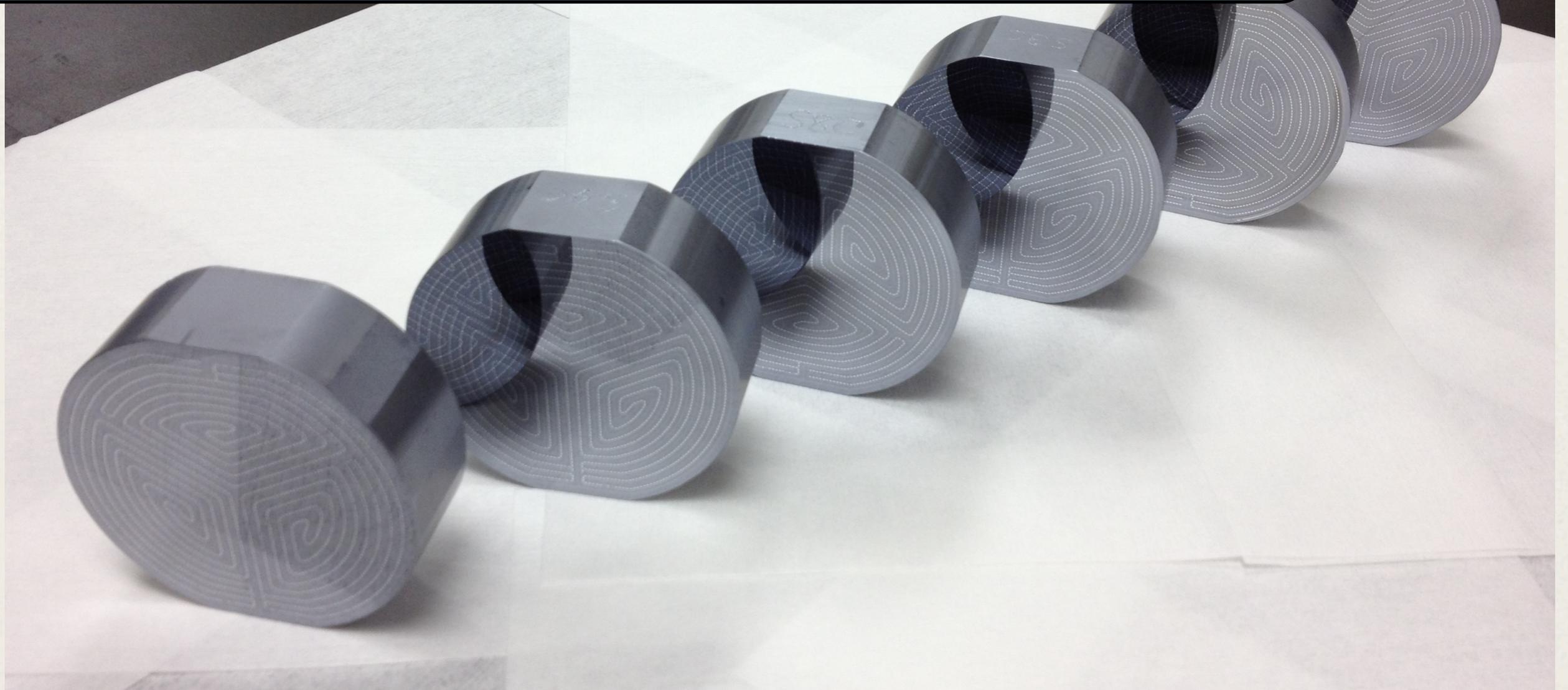


# Background Considerations for SuperCDMS



Jodi Cooley - SuperCDMS  
Southern Methodist University

# SuperCDMS Collaboration

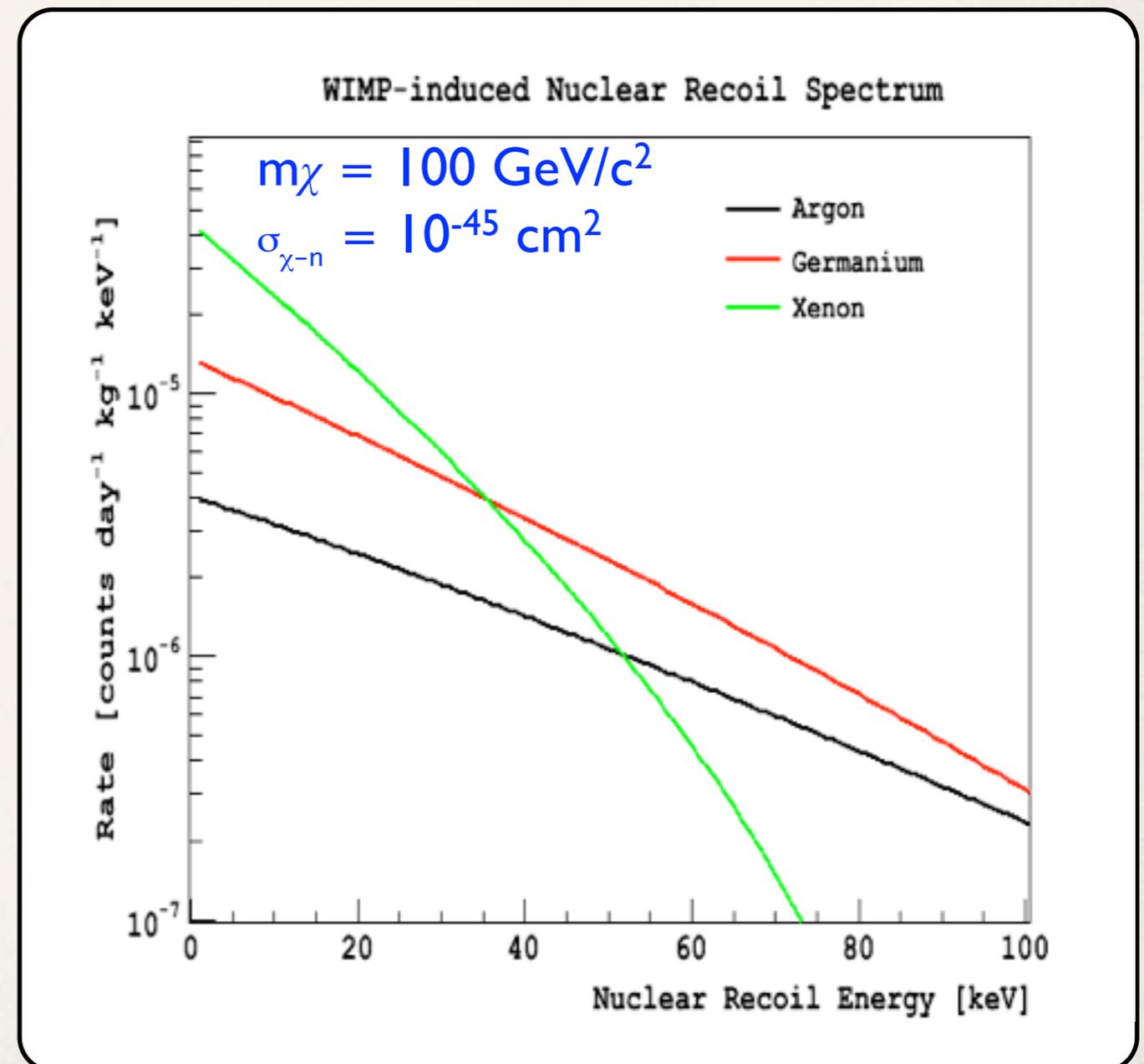


- ❖ California Institute of Technology
- ❖ Fermi National Accelerator Laboratory
- ❖ Massachusetts Institute of Technology
- ❖ NIST
- ❖ Pacific Northwest National Laboratory
- ❖ Queen's University
- ❖ Santa Clara University
- ❖ SLAC/KIPAC
- ❖ Southern Methodist University\*
- ❖ Stanford University
- ❖ Syracuse University
- ❖ Texas A&M
- ❖ University of British Columbia
- ❖ University of California, Berkeley
- ❖ University of California, Santa Barbara
- ❖ University of Colorado, Denver
- ❖ University of Evansville
- ❖ University of Florida
- ❖ FT-UAM/CSIC and Universidad Autonoma de Madrid
- ❖ University of Minnesota

\*The SMU SuperCDMS group is supported by the NSF under grant number 1151869.

# Direct Detection Event Rates

- ❖ **Elastic scattering** of a WIMP deposits small amounts of energy into recoiling nucleus (~ few 10s of keV)
- ❖ Featureless **exponential spectrum**
- ❖ **Expected rate:**  
**< 5 interaction per ton per day**  
**( $3.8 \times 10^{-44} \text{ cm}^2$  for  $m_\chi = 70 \text{ GeV}$ )**
- ❖ Background caused by the radioactivity of most materials is higher than this rate!

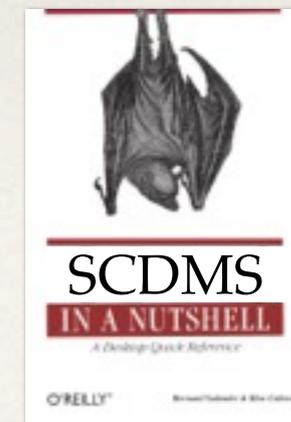


# Challenges

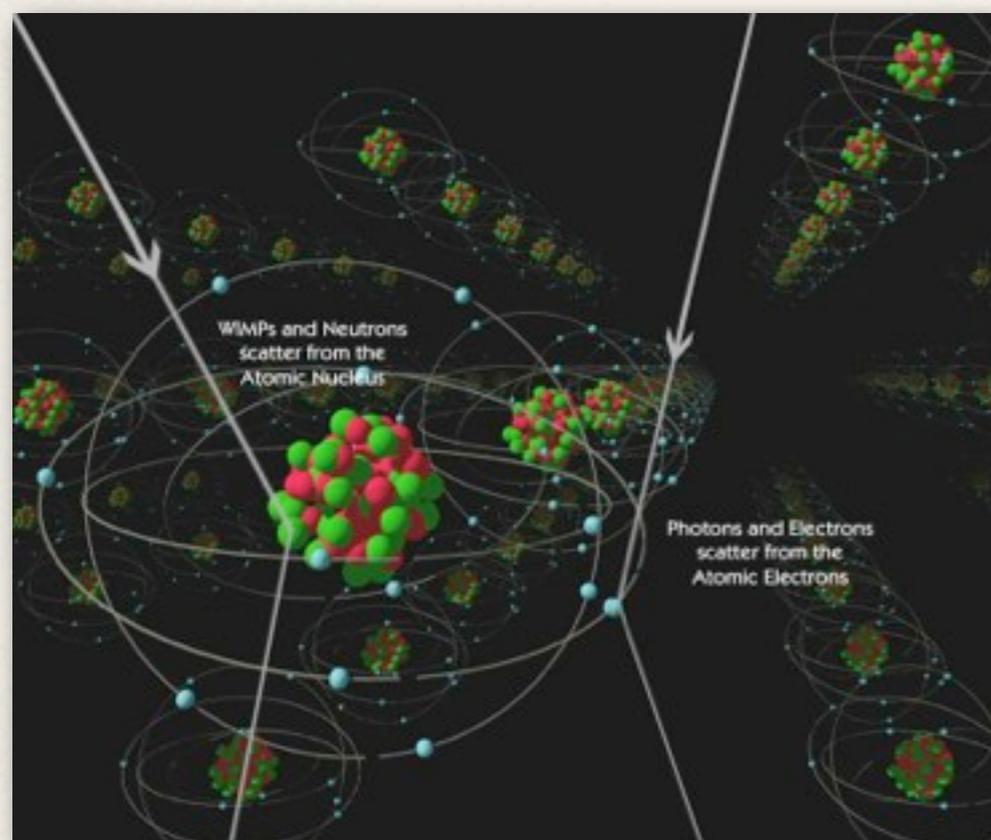
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- ✓ **Low energy thresholds** (~10 keV)
- ✓ **Rigid background controls**
  - ➔ Clean materials
  - ➔ shielding
  - ➔ discrimination power
- ✓ **Substantial Depth**
  - ➔ neutrons look like WIMPS
- ✓ **Long exposures**
  - ➔ large masses, long term stability

# The Big Picture



Use a combination of **discrimination** and **shielding** to maintain a “**< 1 event expected background**” experiment with **low temperature** semiconductor detectors



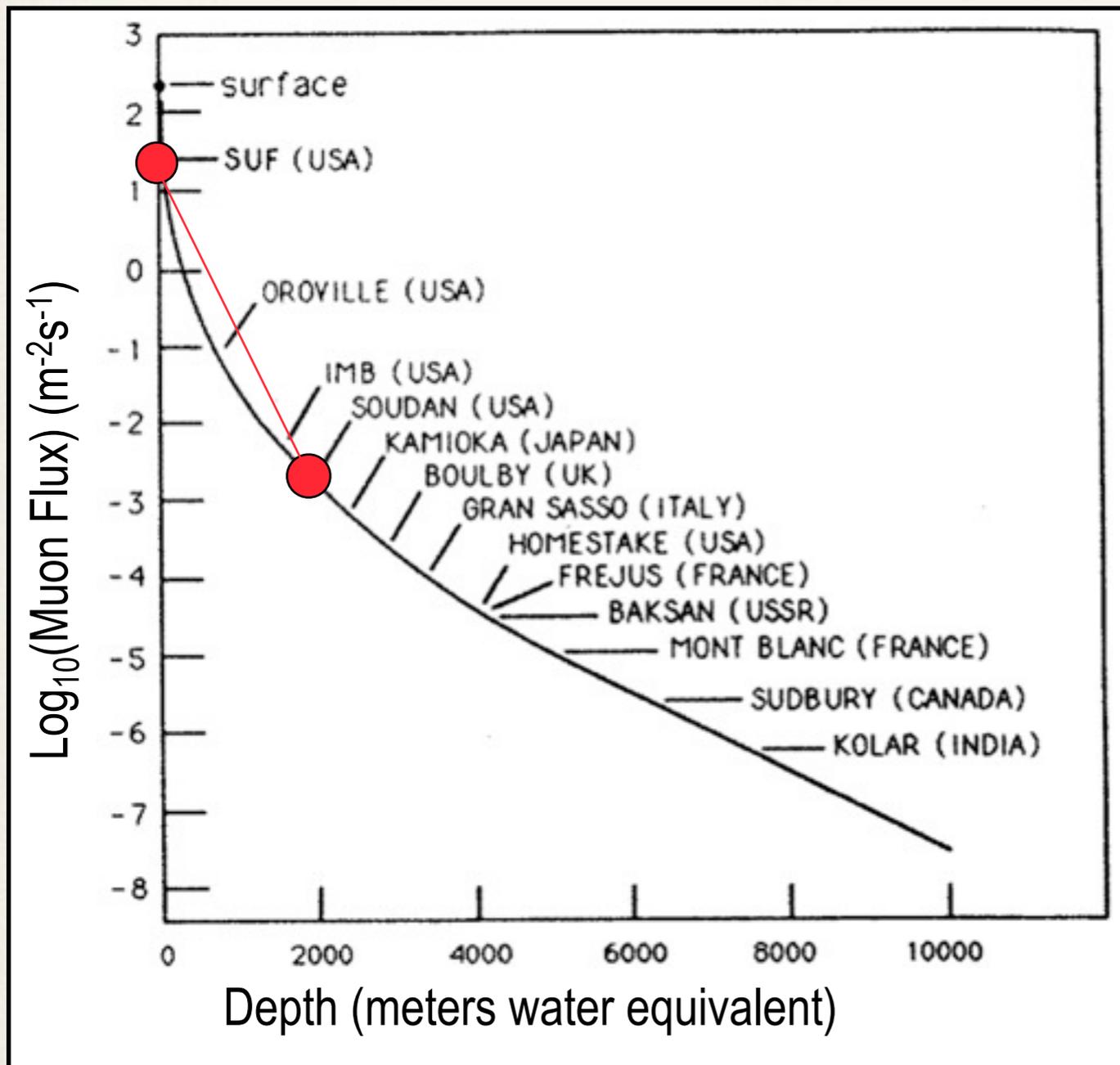
Primary Backgrounds:  
Photons, Electrons and Neutrons.

Discrimination from measurements of  
**ionization** and **phonon energy**.

Keep backgrounds low as possible  
through shielding and material selection.



# Depth is Important!



mwe = meters water equivalent

**SUF (CDMS I)**

**17 mwe**

**0.5 n/d/kg**

**(182.5 n/y/kg)**

**Soudan (CDMS II,  
SuperCDMS)**

**2090 mwe**

**0.05 n/y/kg**

# Shielding

---

**Active Muon Veto:**  
rejects events from cosmic rays



CDMS active muon veto

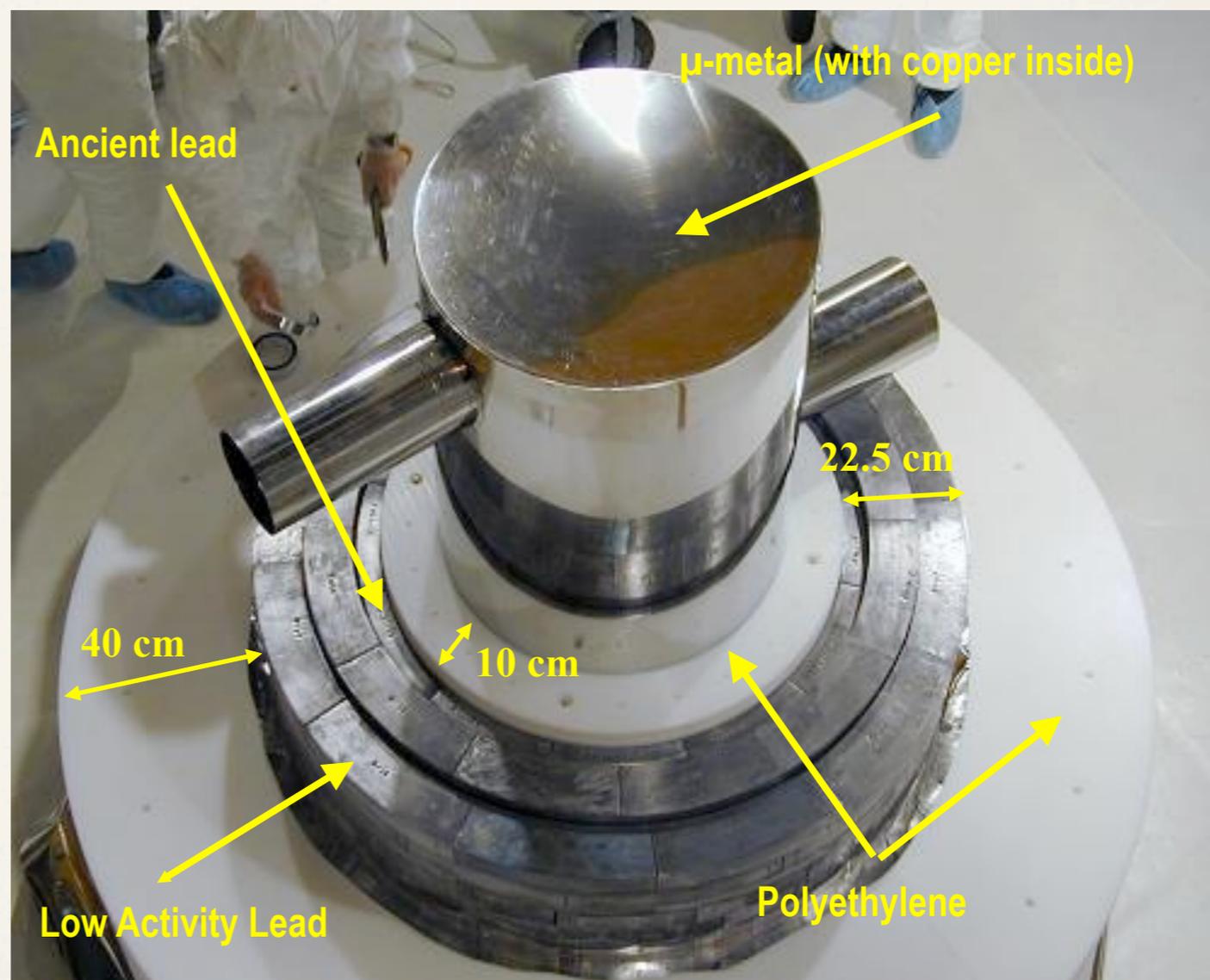
# Shielding - An Example

## **Active Muon Veto:**

rejects events from cosmic rays

**Pb:** shielding from gammas resulting from radioactivity

**Polyethylene:** moderate neutrons produced from fission decays and from  $(\alpha, n)$  interactions resulting from U/Th decays



CDMS - Layers of Polyethylene and Lead

# Shielding

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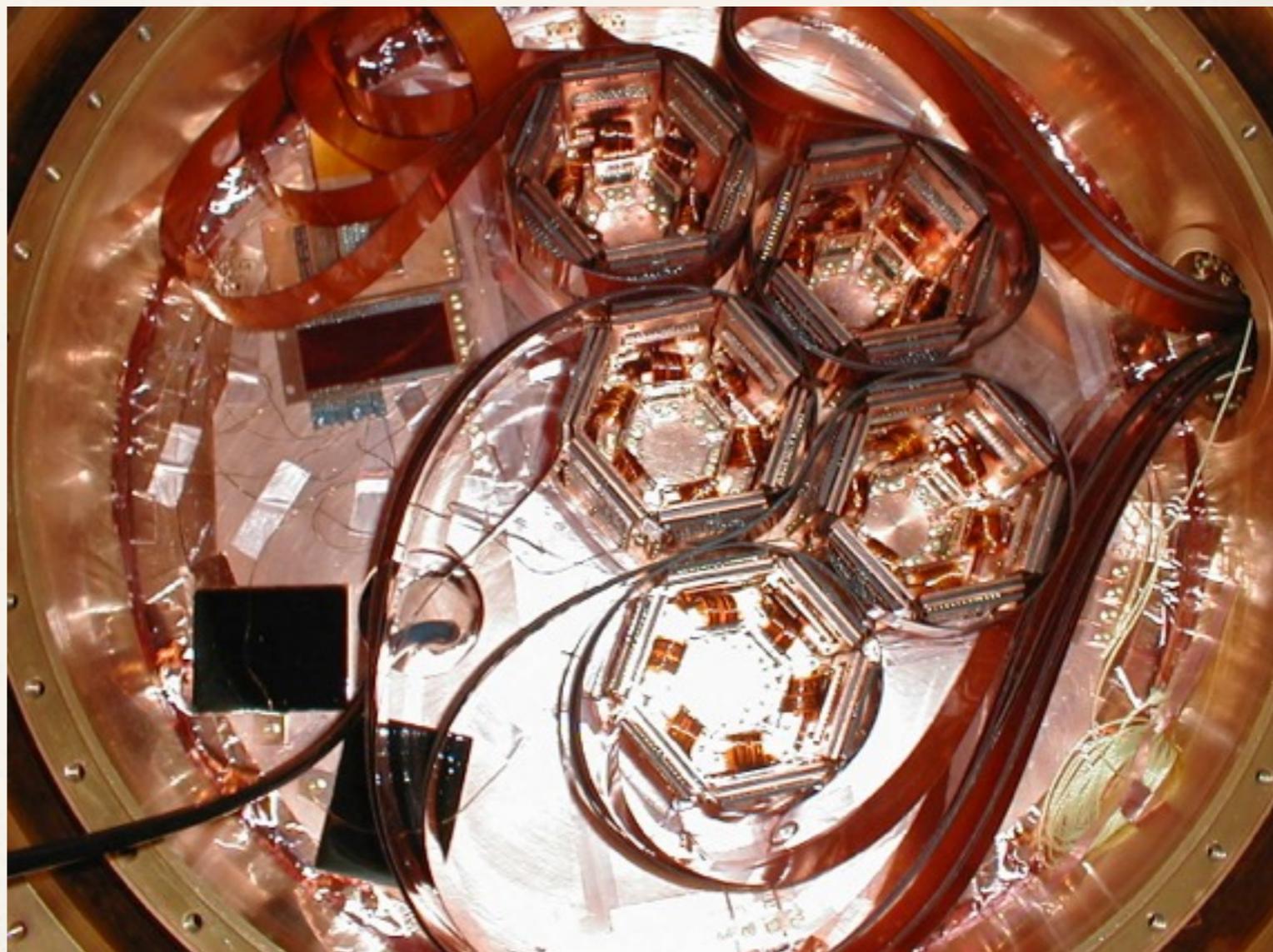
## **Active Muon Veto:**

rejects events from cosmic rays

**Pb:** shielding from gammas  
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**Polyethylene:** moderate  
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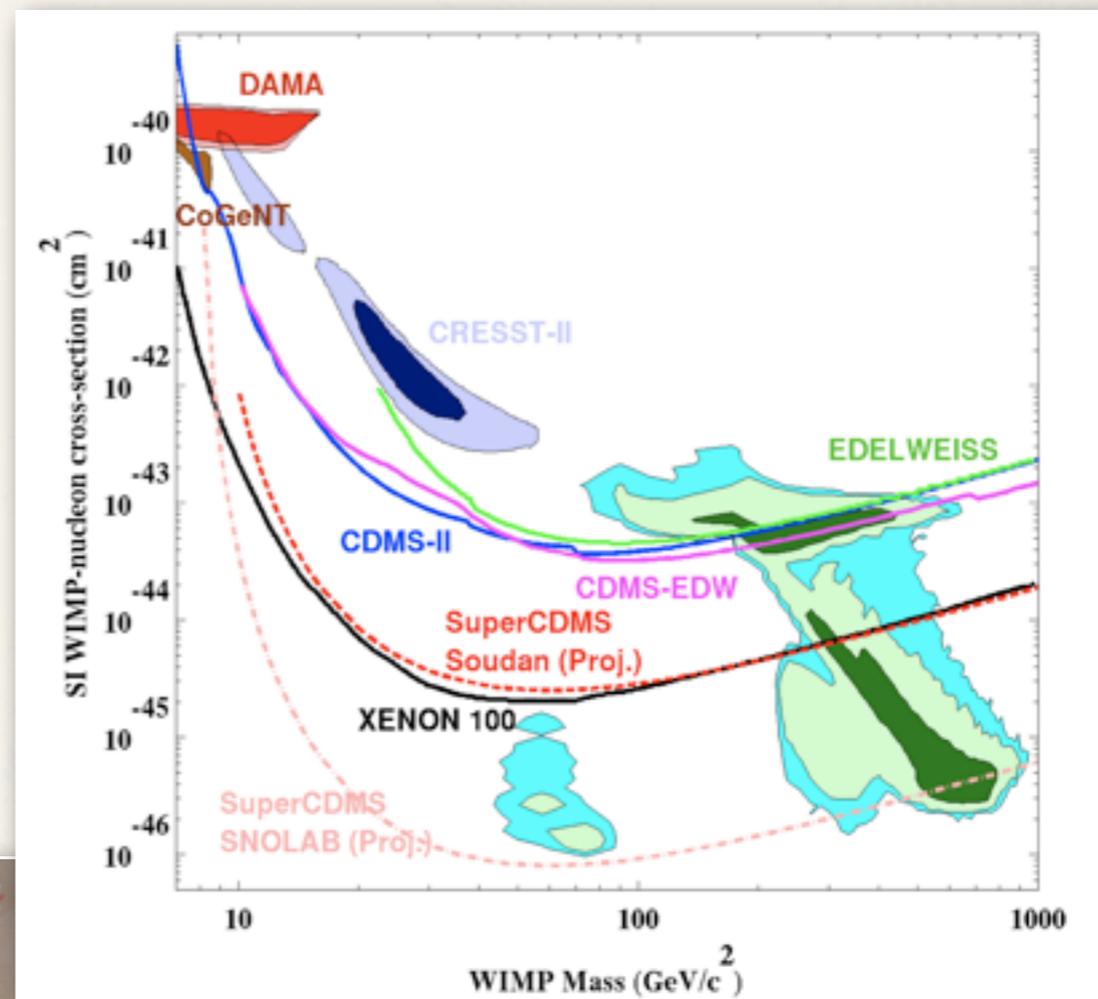
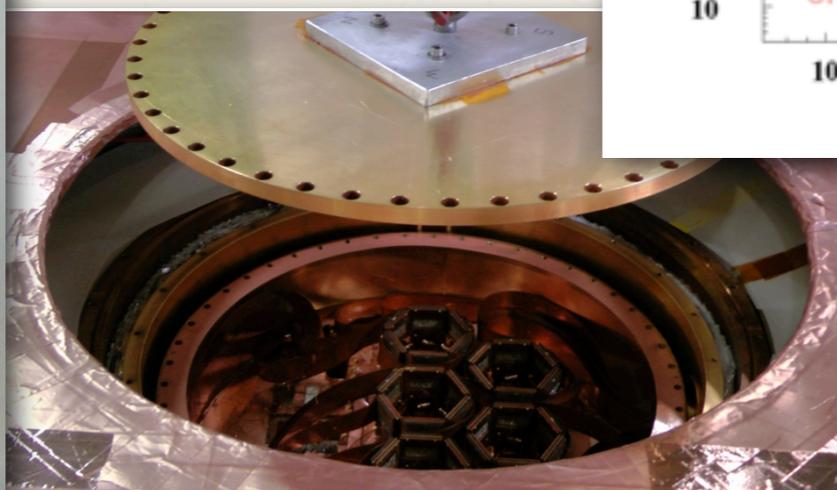
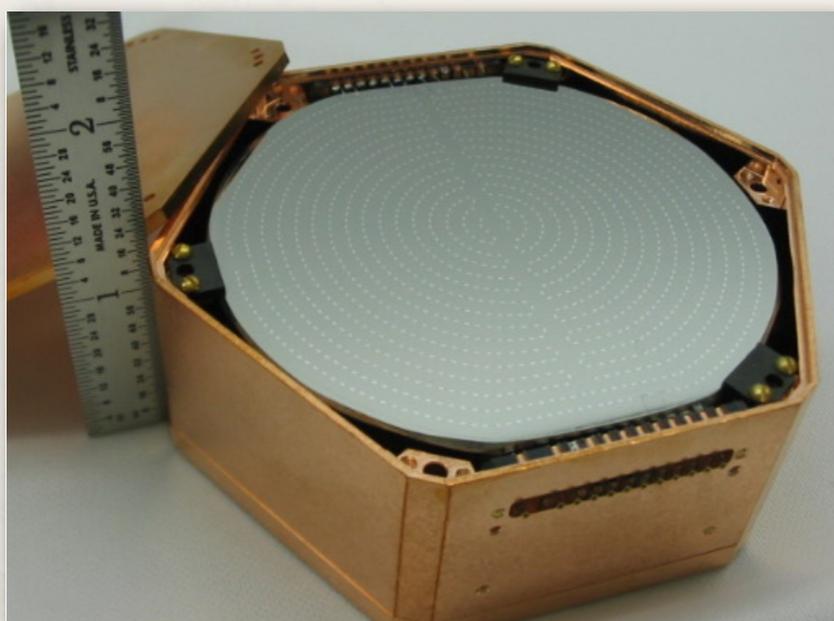
**Cu:** shielding from gammas



CDMS - Top view of inner most shield layer

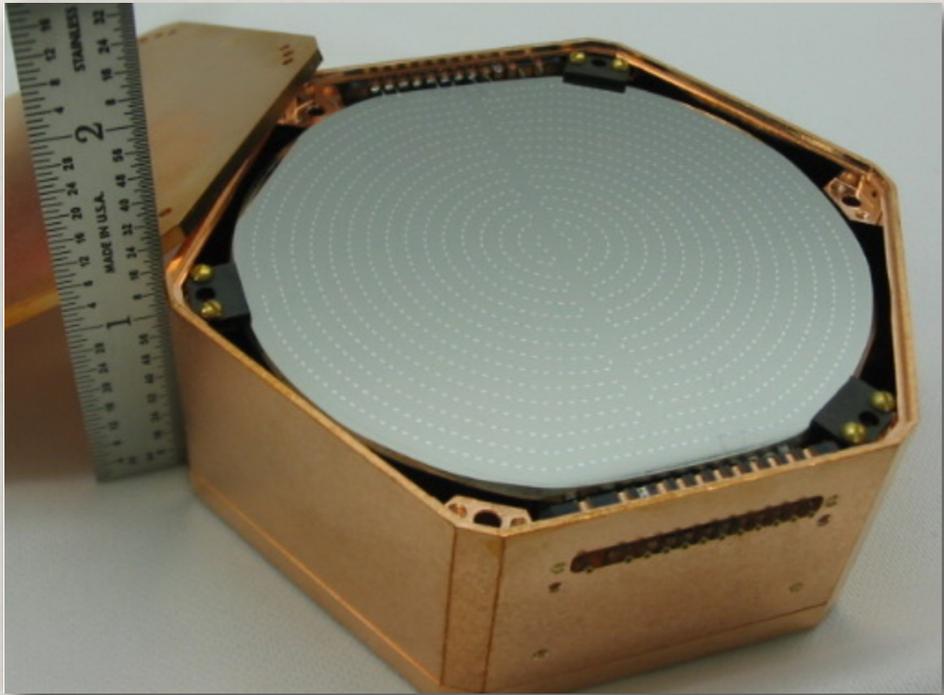
# SuperCDMS at Soudan

- ❖ Currently operating 5 towers of advanced iZIP detectors (~9 kg Ge) in the existing cryostat at the Soudan Underground Laboratory.
- ❖ After 3 years of operation, expected to improve sensitivity to spin-independent WIMP-nucleon interactions by a factor of ~10 over existing CDMS II results.

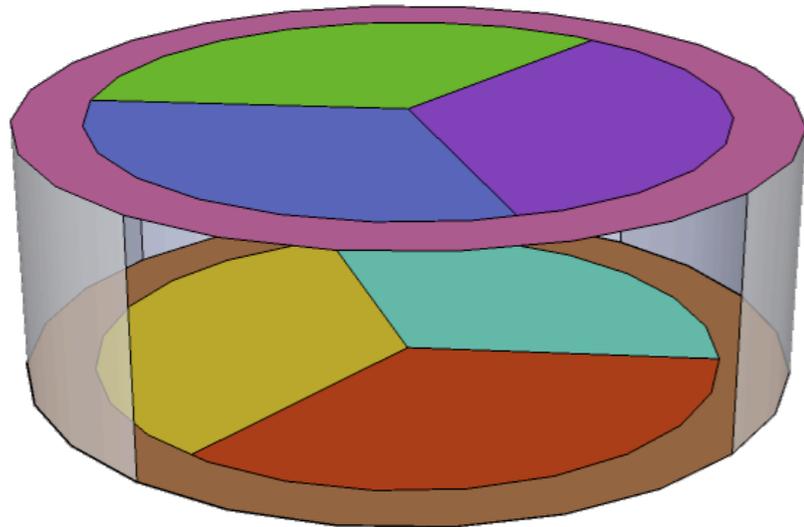


Installation complete Nov. 8, 2011.  
Operating with final detector settings since Mar. 2012.

# The Detectors



**Phonon sensor layout:**



- \* 76 x 25 mm interleaved ZIP (iZIP) double sided detectors
  - \* (2.5x thicker than CDMS II)
- \* Ionization electrodes are interleaved with narrow strips of phonon sensors.
  - \* Phonon sensors optimized to enhance phonon signal to noise ratio
- \* Optimized sensor layout
  - \* Each side has one outer channel to reject zero charge events and 3 inner channels to reject surface events.
- \* Ionization channels can be used to reject surface events

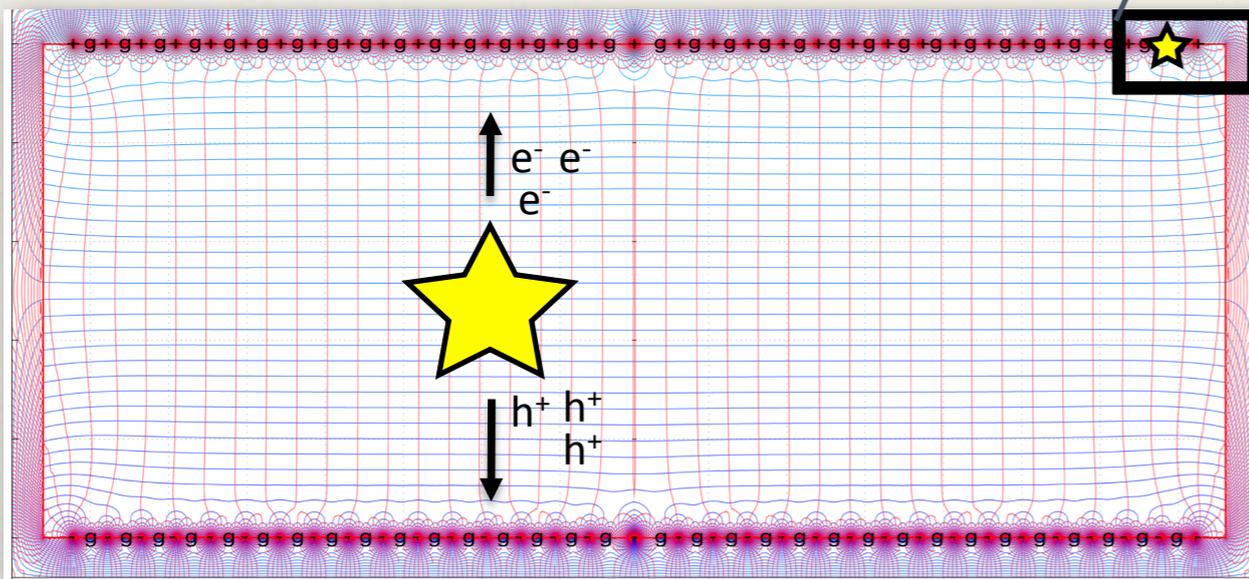
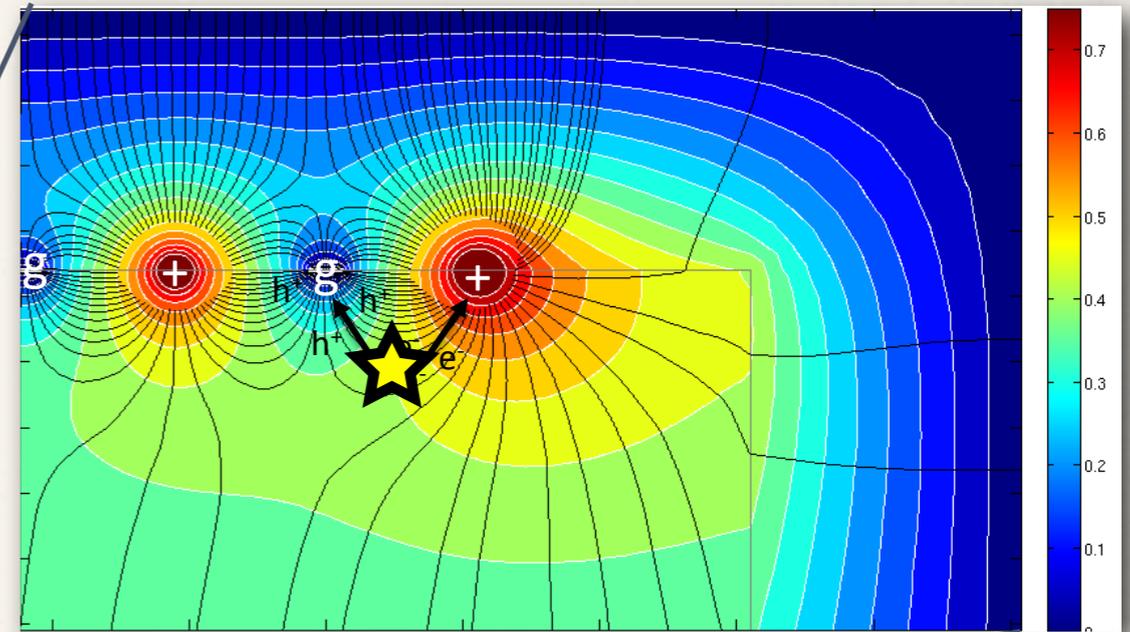
# SuperCDMS iZIPs: Charge Signal

## Bulk Events:

Equal but opposite ionization signal appears on both sides of detector (symmetric)

## Surface Events:

Ionization signal appears on one detector side (asymmetric)



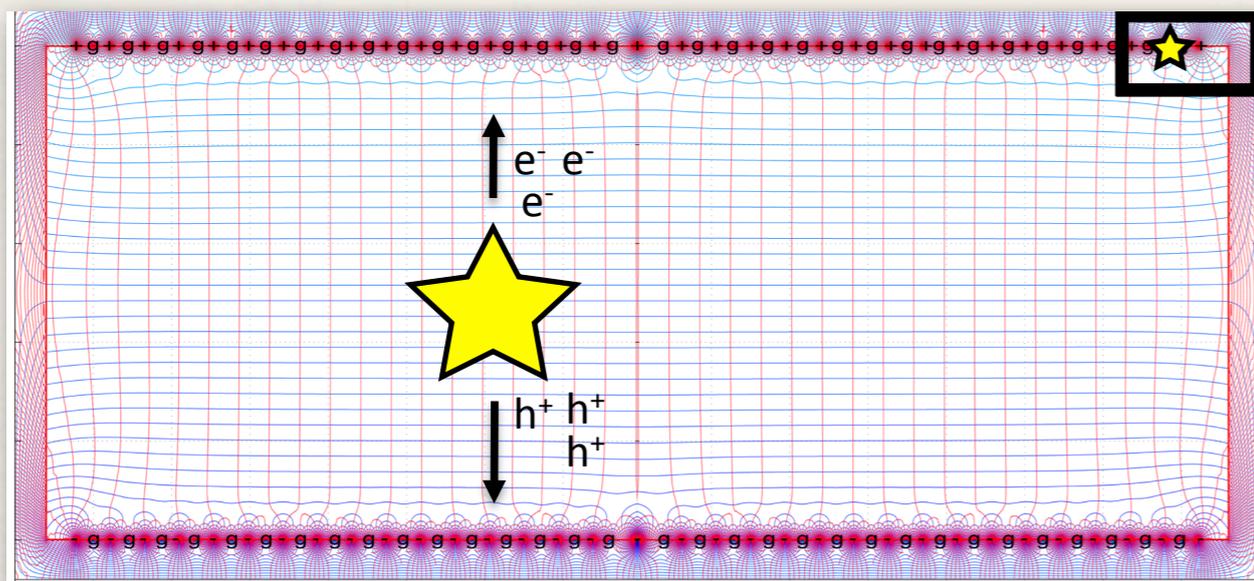
# SuperCDMS iZIPs: Ionization Signal

## Bulk Events:

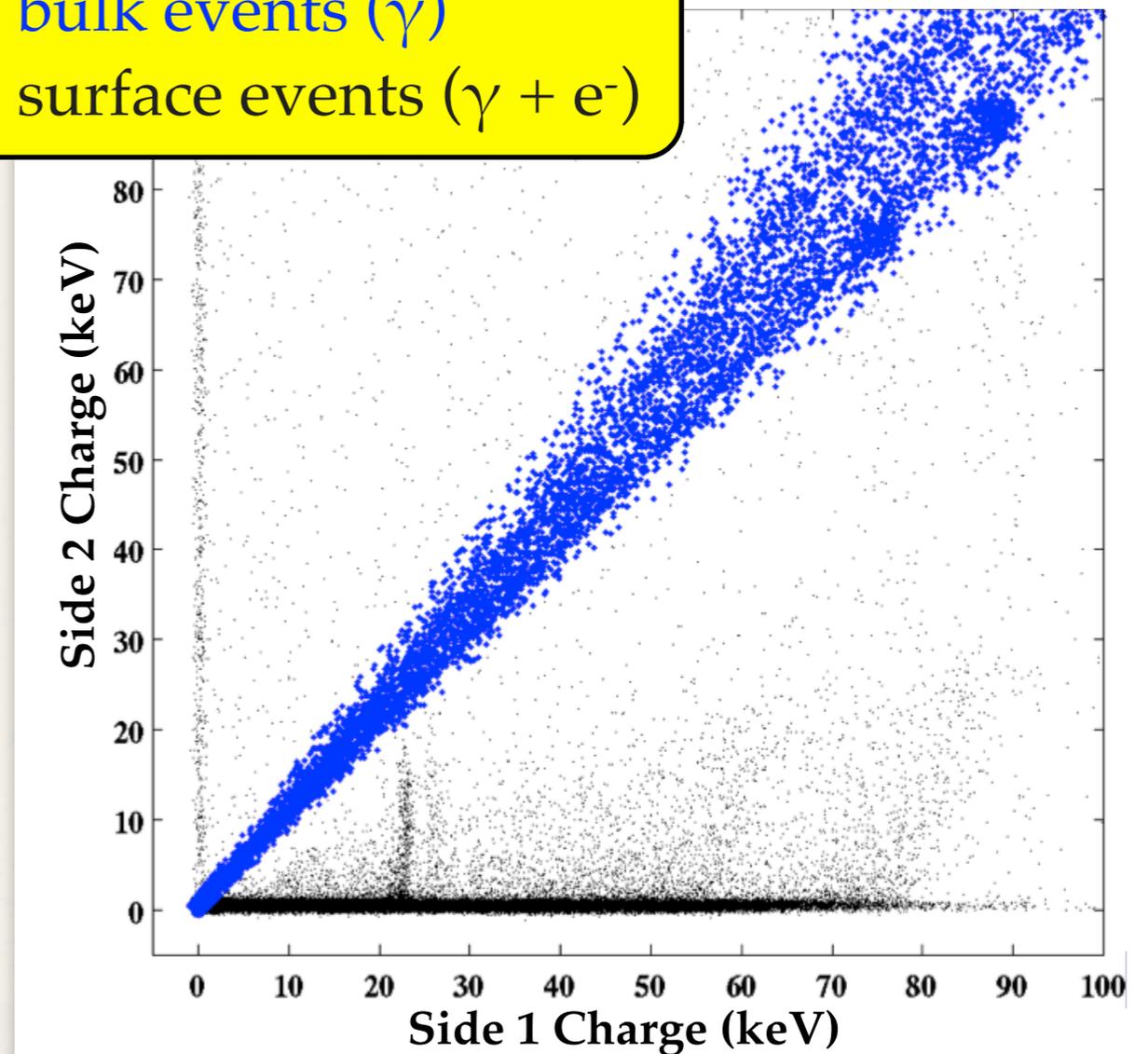
Equal but opposite ionization signal appears on both sides of detector (symmetric)

## Surface Events:

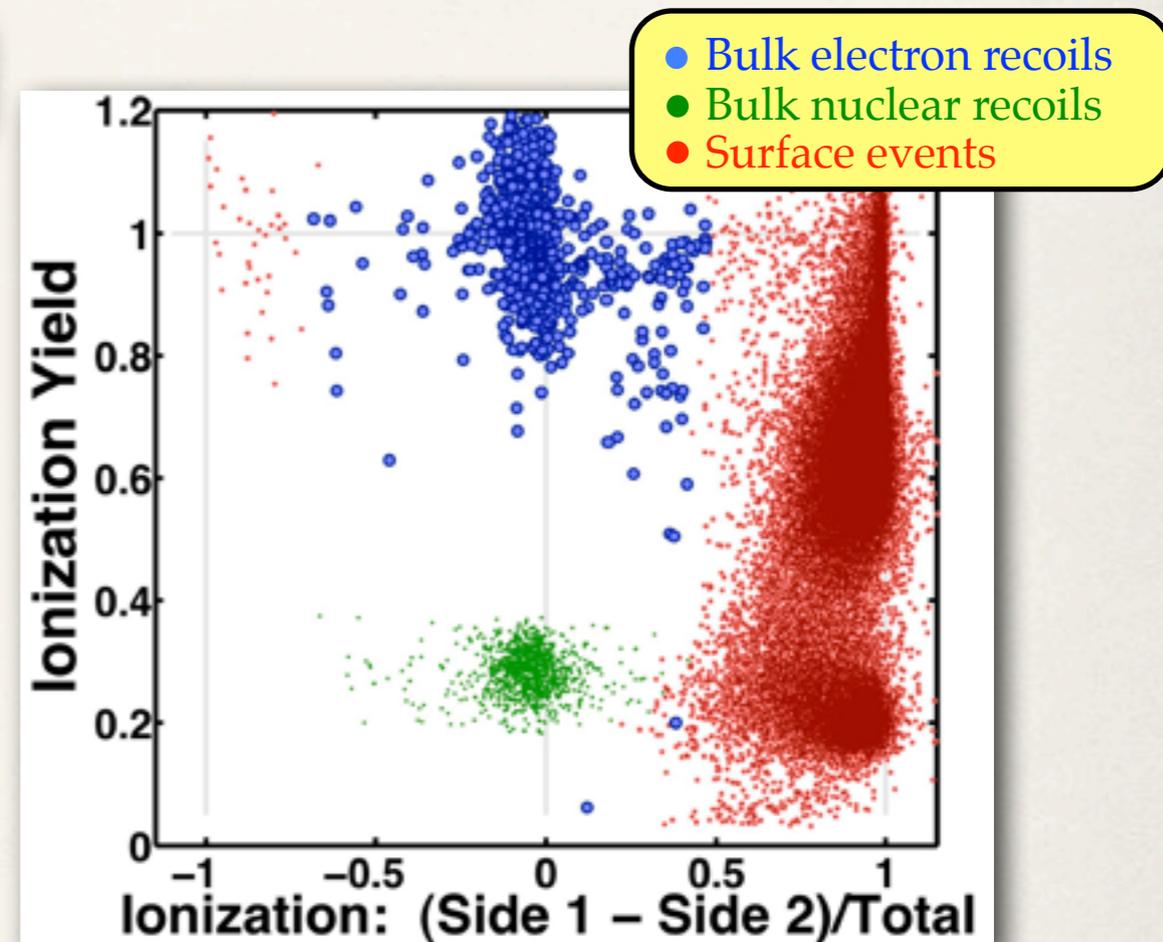
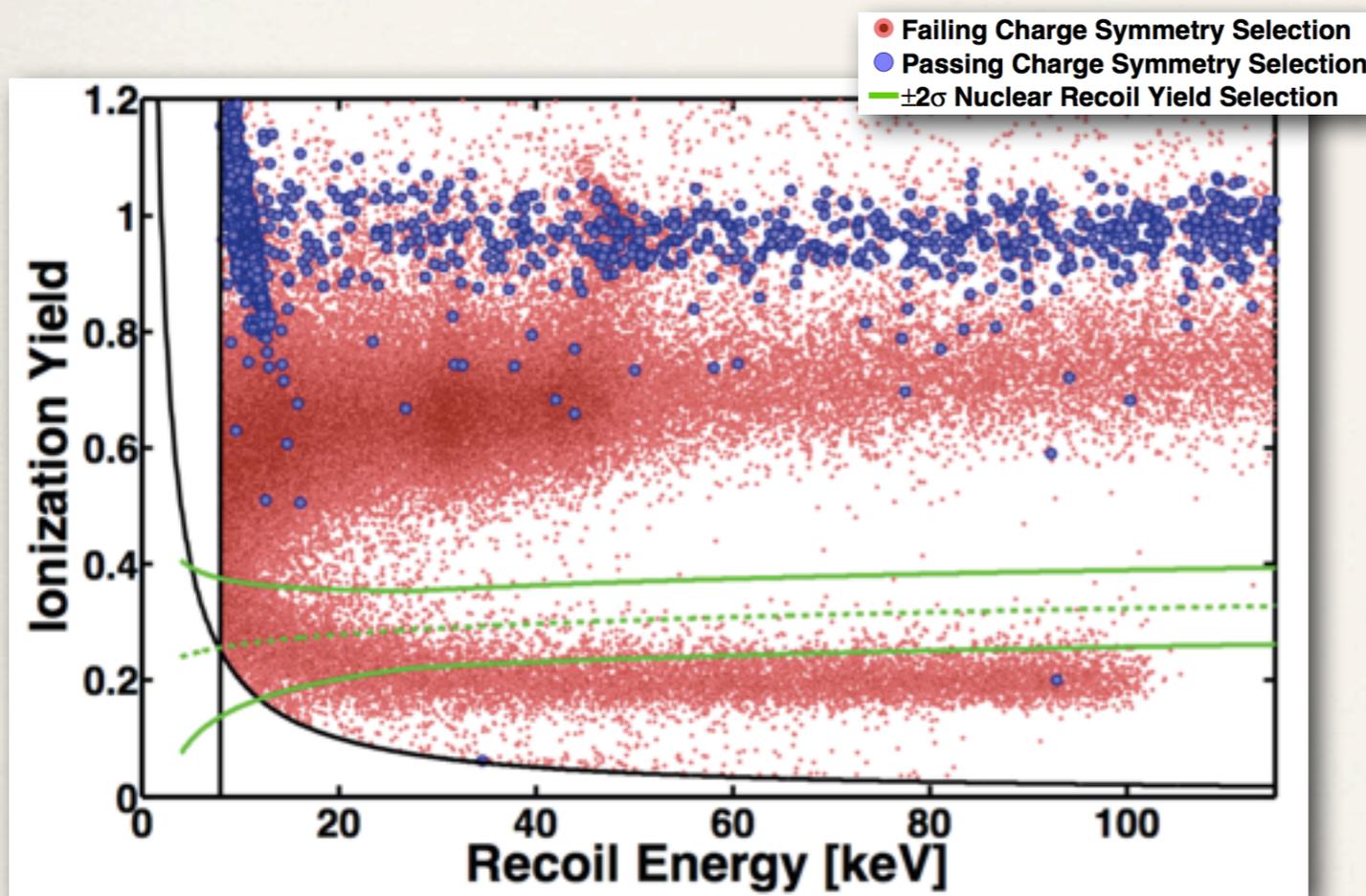
Ionization signal appears on one detector side (asymmetric)



- bulk events ( $\gamma$ )
- surface events ( $\gamma + e^-$ )



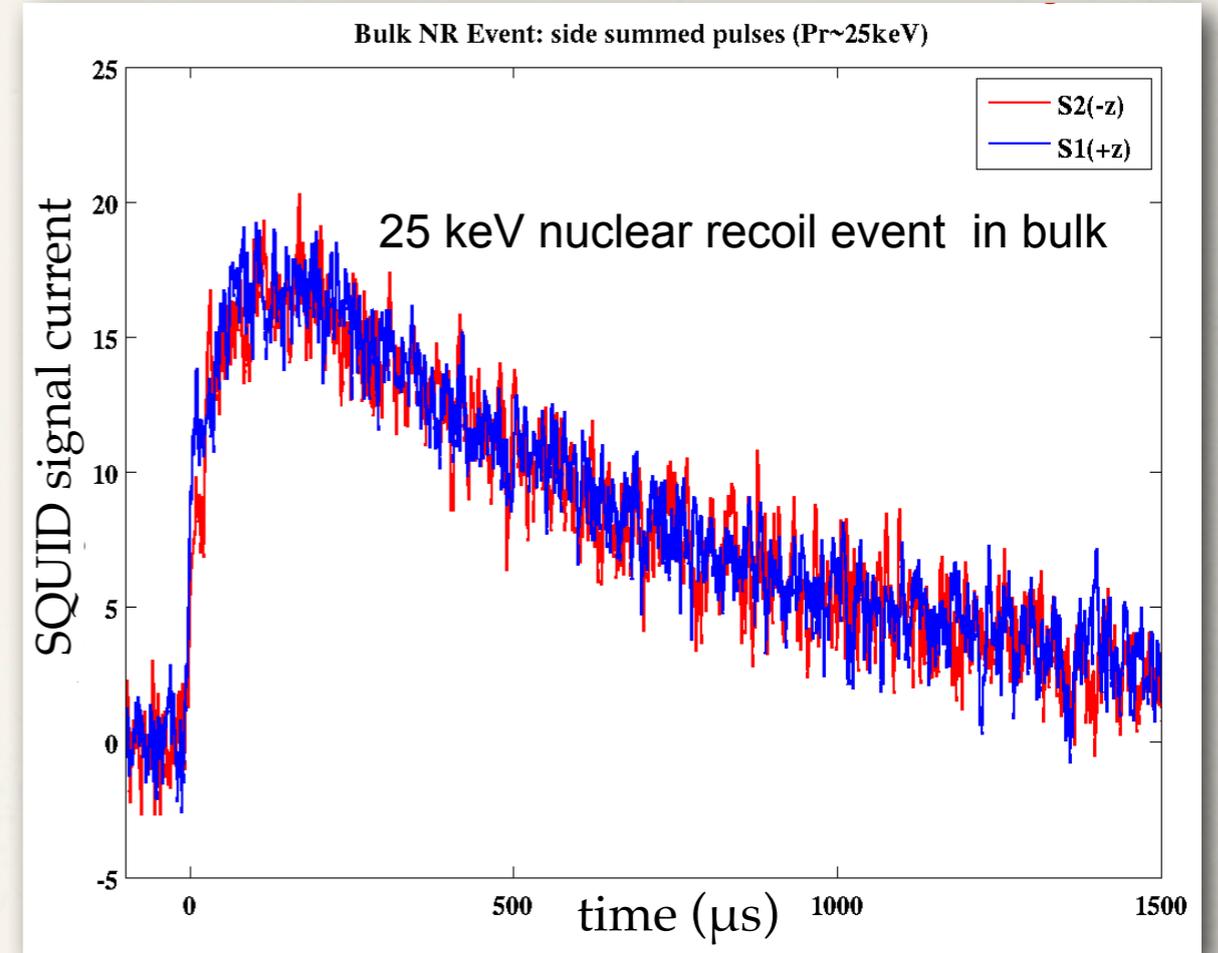
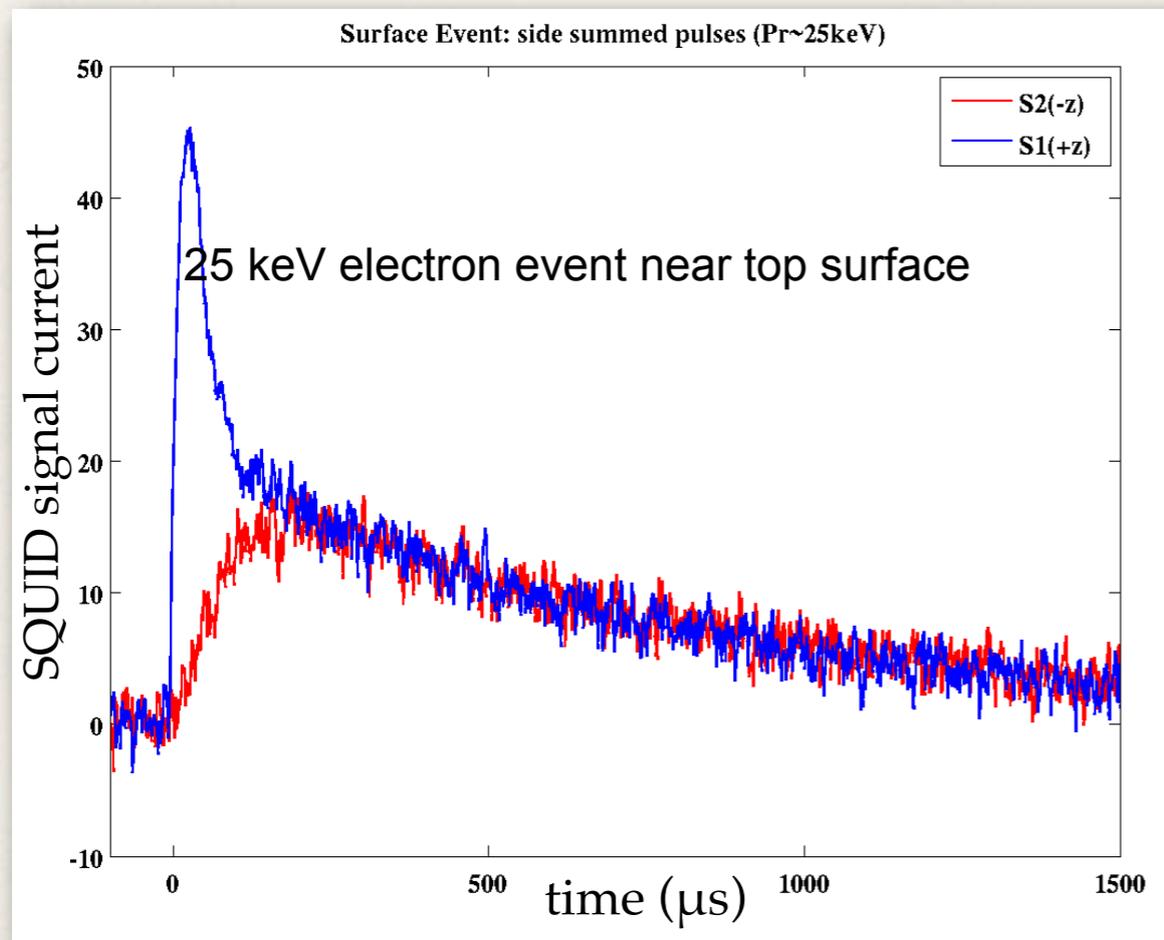
# SuperCDMS Soudan: $^{210}\text{Pb}$ Test



- \* March - July 2012: ~65,000 electrons and ~15,000  $^{210}\text{Pb}$  recoil surface event collected from  $^{210}\text{Pb}$  source.
- \* 0 events leaking into the signal region ( $< 2.5 \times 10^{-5}$  @90% C.L. misID)

- \* 63% fiducial volume (8-100 keVnr)
- \* Good enough for a 200 kg experiment run for 4 years at SNOLAB!

# SuperCDMS iZIP: Phonon Signal

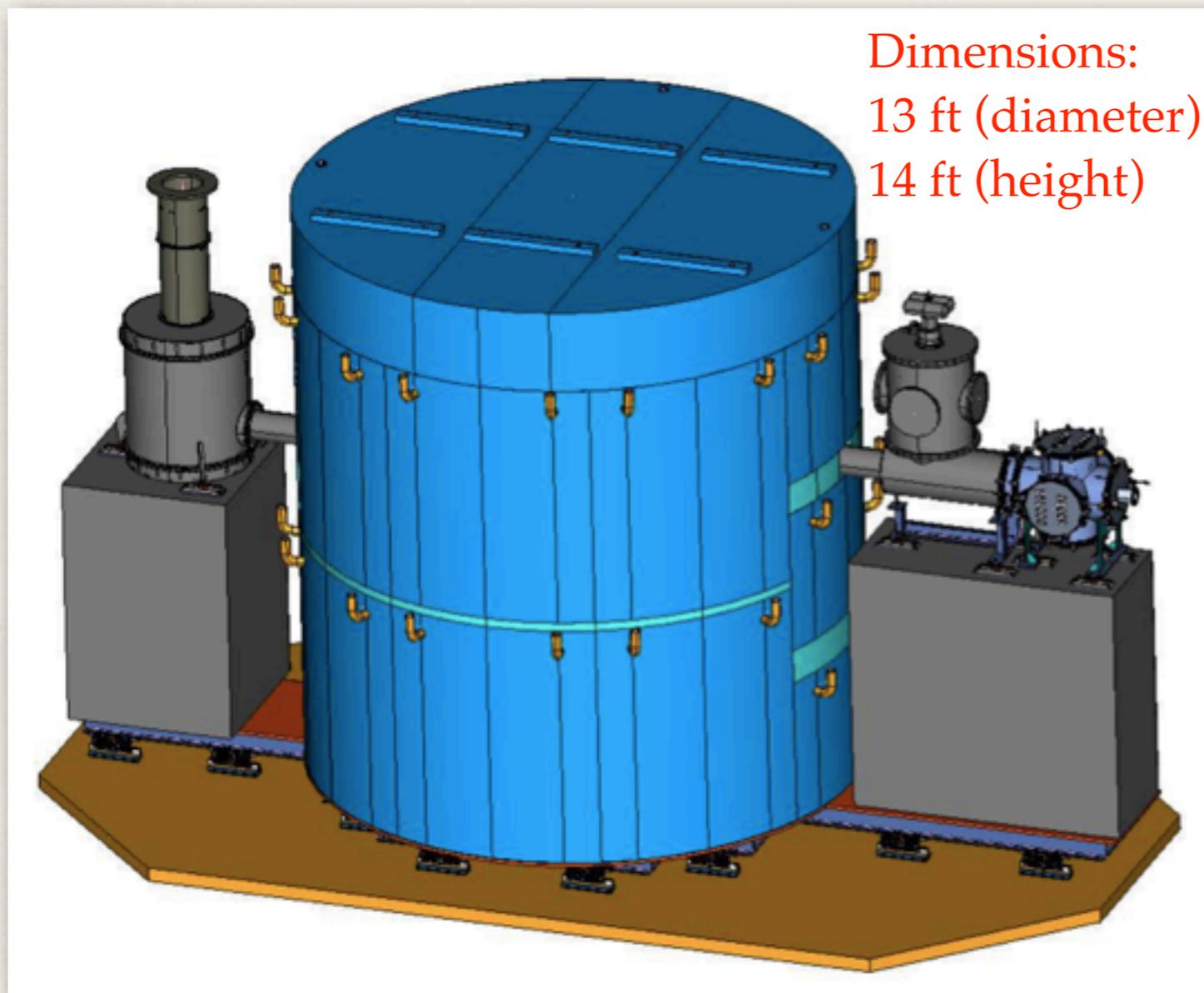


- \* Phonon timing pulse information still possible.
- \* Surface electron vs bulk nuclear recoil event discrimination
- \* PULSE SHAPE HAS NOT YET BEEN USED! (It's not needed.)

# SuperCDMS SNOLAB: G2 Experiment at the Ladder Lab



# SuperCDMS at SNOLAB

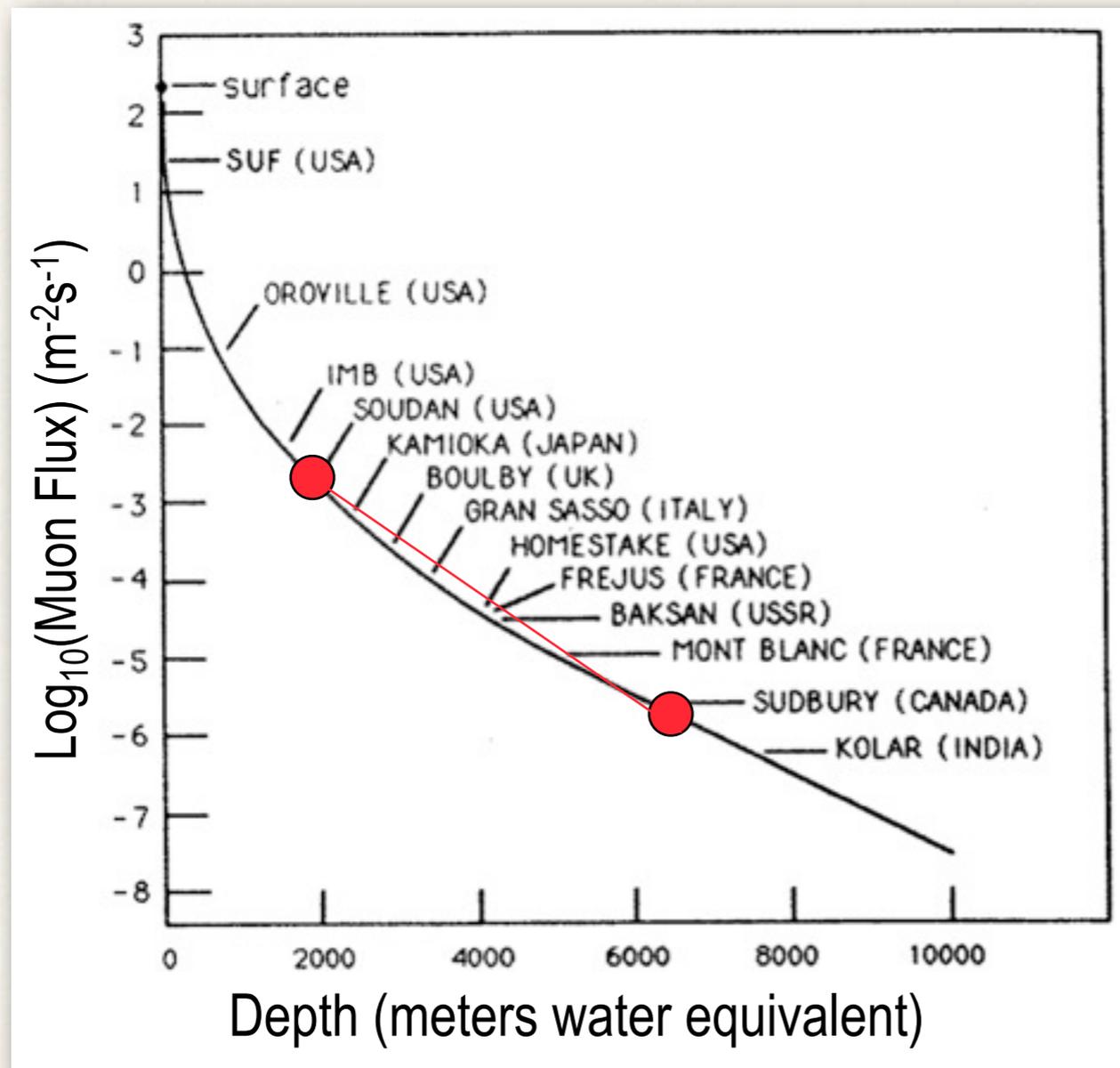


## Planned Setup

- \* cryostat volume of up to 400 kg target
  - \* 200 kg experiment with sensitivity of  $8 \times 10^{-47} \text{ cm}^2$  at  $60 \text{ GeV}/c^2$
- \* Pb/Cu shielding for external radiation
- \* increased PE shielding (neutrons)
- \* possible neutron veto

# Why SNOLAB?

## Depth is Important



*Soudan*

*2090 mwe*

*0.05 n/y/kg*

*SNOLab*

*6060 mwe*

*0.2 n/y/ton*

*(0.0002 n/y/kg)*

❖ Only need to worry about radiogenic neutrons!

# Radiogenic Neutrons

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- ❖ External Radiogenic Neutrons
  - ❖ Resulting from fission and alpha-n interactions from U, Th in cavern rock
  - ❖ Expected to be negligible with passive shielding
- ❖ Internal Radiogenic Neutrons
  - ❖ Resulting from fission and alpha-n interactions from U, Th in copper cans, shielding and supports.
  - ❖ Expected to be ~1 event, depending on material cleanliness

For these reasons we are considering a neutron veto in the shield design.

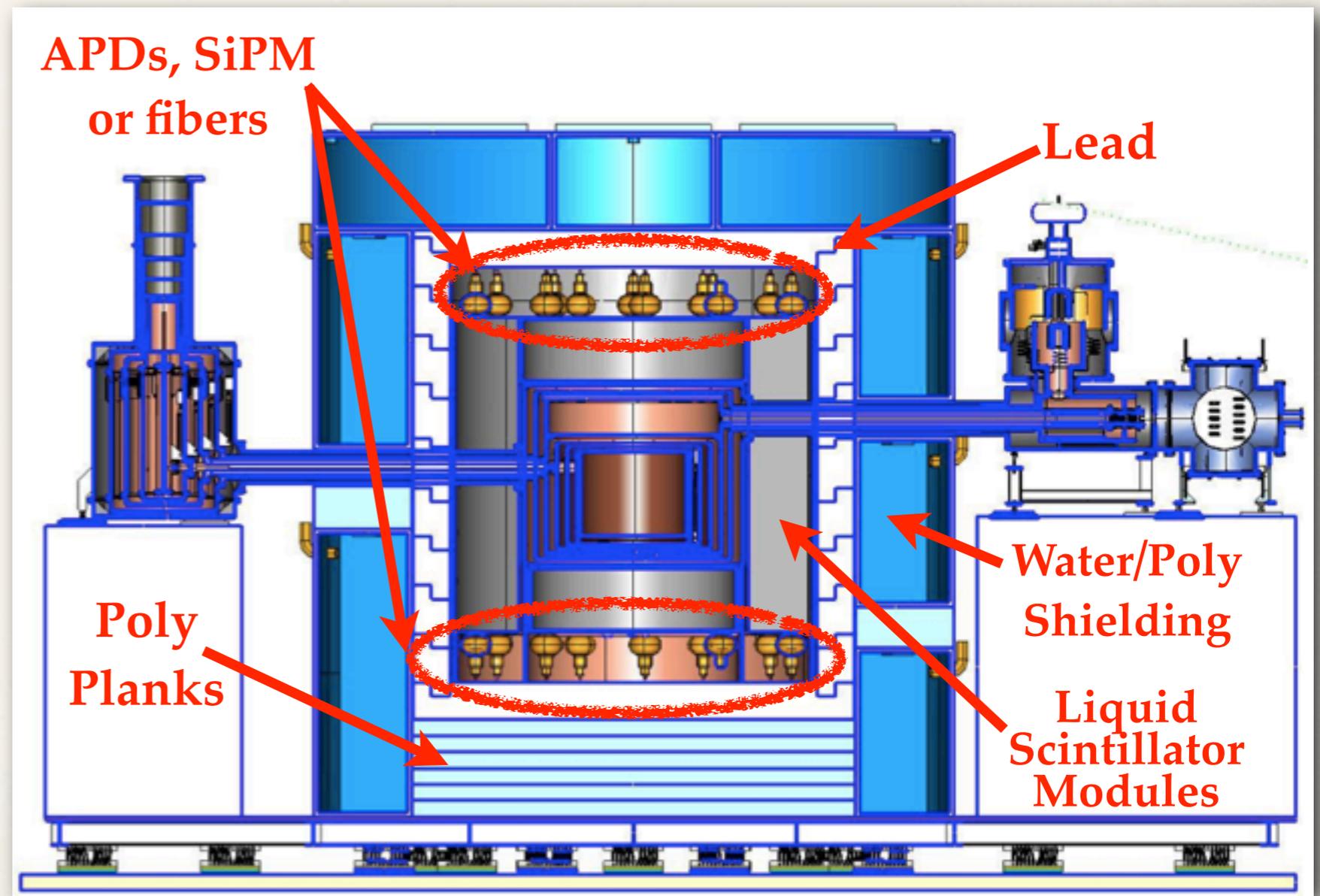
# Physics Requirements for a Neutron Veto

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- ❖ Total unvetoes background in iZIP detectors:  $< 1$  event in the 200 kg phase.
- ❖ Total background rate (gamma + neutrons) must not generate excessive background rates in the iZIP detectors
  - ❖ Implies radio-clean construction
- ❖ Negligible contribution to dead-time
  - ❖ Implies low ( $< \text{kHz}$ ) non-coincident trigger rate
- ❖ High efficiency to detect neutron capture
  - ❖ Aim for 90% or better

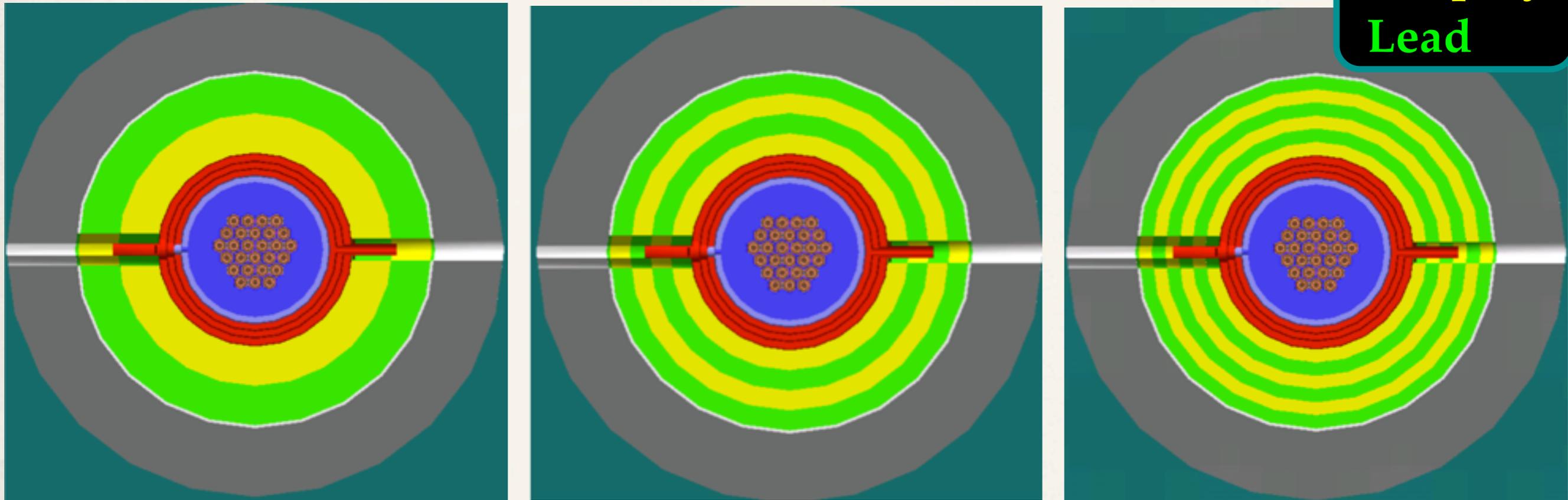
# Design Details

- ❖ Surround the cryostat with a high efficiency neutron detector to tag neutrons.
- ❖ Modular tanks of liquid scintillator, with radial thickness 0.4 m, viewed by phototubes.
  - ❖ Details of scintillator to use (Gd or B loaded) under consideration.



# Alternate Design

Copper  
Gd-poly  
Lead



- ❖ Alternating layers of Gd-loaded poly / scintillator and lead.
- ❖ Preliminary studies underway.

# Summary

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- ❖ SuperCDMS at Soudan (~9 kg) is taking data with iZIP detectors and expects to reach a WIMP-nucleon sensitivity of  $2 \times 10^{-45} \text{ cm}^2$  for spin-independent interactions.
- ❖ We have demonstrated surface event rejection with the new iZIP detector design using  $^{210}\text{Pb}$  sources which paves the way for better than  $10^{-46} \text{ cm}^2$  sensitivity at SNOLAB.
- ❖ Ongoing studies are assessing the necessity and feasibility of including a neutron veto in the SuperCDMS at SNOLAB design.