# Equilibration rates in a strongly coupled nonconformal quark-gluon plasma

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<u>Based on</u> arXiv:

1503.07114 with M.Heller and R.Myers; also on earlier work on quantum quenches with L.Lehner and R.Myers

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## Why QGP thermalization is fast?

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 $\Longrightarrow$  So, we like to refer to AdS/CFT correspondence, and use large-N SYM as a proxy for a real QGP

The motivations:

• QCD thermodynamics from lattice; (Karsch, Laermann, hep-lat/0305025). The plateau is  $\sim 80\%$  of the SB result — close to 3/4 in SYM thermodynamics



• The small shear viscosity ratio (Policastro, Son, Starinets, hep-th/0104066)

$$\frac{\eta}{s} = \frac{1}{4\pi}$$

 $\implies$  From A.Bazarov et.al (HotQCD Collaboration), arXiv:1407.6387:



 $\implies$  The violation of the conformality,

$$\frac{\epsilon - 3p}{\epsilon} \sim 50\%$$

at the maximum

So, should be use nonconformal modes of gauge/gravity correspondence to model equilibration of QGP?

 $\implies$  I am going to use top-down holographic model to address this question

# Outline of the talk:

- A toy model for holographic equilibration
  - where the relaxation time(s) is(are) encoded?
- $\mathcal{N} = 2^*$  gauge theory/supergravity holography
  - Gauge theory perspective
  - Holographic Pilch-Warner RG flow
  - Matrix model and localization results
- Spectra of quasinormal modes
  - Relaxation rates for homogeneous/isotropic perturbations:  $\{\mathcal{O}_2, \mathcal{O}_3, \mathcal{O}_4, T_{\mu\nu}, \cdots\}$
  - Relaxation rates for generic transverse/traceless perturbations of  $T_{\mu\nu}$
- Conclusion beyond  $\mathcal{N} = 2^*$  holography
  - relaxation with chemical potential in  $\mathcal{N} = 4$  SYM
  - relaxation in bottom-up holographic models

• Consider  $\mathcal{N} = 4$  large-N SU(N) SYM theory at strong coupling in thermal state:

$$\epsilon = \frac{3}{8}\pi^2 N^2 T^4$$
,  $p = \frac{1}{8}\pi^2 N^2 T^4$ ,  $s = \frac{1}{2}\pi^2 N^2 T^3$ 

- The holographic dual to this state is a Schwarzschild black hole in Poincare-slice  $AdS_5$ . It has the thermodynamic properties (Hawking temperature, Bekenstein-Hawking entropy,...) as above
- SYM has gauge invariant fermion bi-linear operators of dimension Δ = 3:
  O<sub>3</sub>, and gauge invariant scalar bi-linears of dimension Δ = 2: O<sub>2</sub>. In thermal equilibrium,

$$\langle \mathcal{O}_3 \rangle_{T \neq 0} = 0, \qquad \langle \mathcal{O}_2 \rangle_{T \neq 0} = 0$$

• We can 'prepare' a non-equilibrium states of the  $\mathcal{N} = 4$  plasma (thus inducing a non-trivial time-dependence of  $\mathcal{O}_{\Delta}$ ) by quenching the coupling constants of these relevant operators:

$$H_{SYM} \rightarrow H_{SYM} + \lambda_{\Delta} \mathcal{O}_{\Delta}$$
  
 $\lambda_{\Delta} = \lambda_{\Delta}(t), \qquad \lambda_{\Delta}(-\infty) = 0$ 

• Specifically, we assume

$$\lambda_{\Delta}(t) = \lambda_{\Delta}^{0} \left( \frac{1}{2} + \frac{1}{2} \tanh \frac{t}{\mathcal{T}} \right), \qquad \mathcal{T} = \frac{\alpha}{T_{i}},$$

where:

- $T_i$  is the temperature of the thermal state at  $t \to -\infty$ ,
- $\lambda_{\Delta}^{0}$  is the amplitude of the quench, taken to be small compare to the initial temperature,

$$\frac{|\lambda_{\Delta}|}{T_i^{4-\Delta}} \ll 1$$

- $\alpha$  is the rate of quench, measure in units of inverse temperature.
- Note that  $\alpha$  can be arbitrarily small/large, corresponding to abrupt/adiabatic quenches
- $\alpha \to 0$  limit (infinitely sharp step-function quench) can be thought as preparing a system in an excited state at t = 0 and allowing it to relax

 $\implies$  I am not going to explain how to set up about quench holographically, and rather move to to discuss the results

 $\implies$  Our primary observable is the expectation value of the quenching operator:

$$\mathcal{O}_{\Delta} = \mathcal{O}_{\Delta}(t)$$



- Evolution of the normalizable component  $\mathcal{O}_3$  (left panel) and  $\mathcal{O}_2$  (right panel) during the quenches with  $\alpha = 1$ . The dashed red lines represent the adiabatic response.
- As  $\tau \to +\infty$  the expectation values approach their equilibrium values in a damped-oscillatory manner (More on this later).

The response of  $\mathcal{O}_{\Delta}$  depends on  $\Delta$ :

• for fast quenches,  $\alpha$  is small,



 $\implies$  The response is quite different!

 $\implies$  How do we characterize equilibration time?

Introduce

$$\delta_{neq}(\tau) \equiv \left| \frac{\mathcal{O}_{\Delta}(\tau) - [\mathcal{O}_{\Delta}(\tau)]_{adiabatic}}{[\mathcal{O}_{\Delta}(\tau)]_{equilibrium}} \right| \,,$$

where  $[\mathcal{O}_{\Delta}(\tau)]_{equilibrium}$  is the adiabatic response that can be computed analytically.

■ Note,

$$\lim_{\tau \to \pm \infty} \delta_{neq}(\tau) \to 0$$

as at early/late times the system is in equilibrium.

#### $\implies$ In practice,



Extraction of the excitation/equilibration rates for  $\alpha = 1$  quench. The horizontal green line is the threshold for excitation/equilibration which we define to be 5% away from local equilibrium as determined by  $\delta_{neq}$ . The dashed red lines indicate the earliest and latest times of crossing this threshold, which we denote as  $\tau_{ex}$  (for excitation time) and  $\tau_{relax}$  (for equilibration time), respectively.



 $\implies$  Going to small  $\alpha$  (ln  $\alpha \rightarrow -\infty$ ) corresponds to preparing the state with an abrupt quench of a dim- $\Delta$  operator. The dashed scaling line translates into a **universal relaxation time**:

$$t_{relax} \sim \frac{1}{T}$$

independent of  $\alpha$ !

#### $\implies$ We can do more:



Behavior of the response coefficients versus time for representative fast quenches. As is evident in the picture, the same quasinormal mode governs the dynamics very quickly after the quench:

$$\Delta = 3: \qquad \frac{\omega}{2\pi T}\Big|_{fit} \simeq (1.095 - i\,0.87), \qquad \frac{\omega}{2\pi T}\Big|_{BH} \simeq (1.099 - i\,0.879)$$
$$\Delta = 2: \qquad \frac{\omega}{2\pi T}\Big|_{fit} \simeq (0.64 - i\,0.4), \qquad \frac{\omega}{2\pi T}\Big|_{BH} \simeq (0.644 - i\,0.411)$$

Lowest quasinormal modes of the black hole in the gravitational dual control the relaxation in strongly coupled gauge theory plasma

- Such feature was also observed in various other holographic examples
- It probably should not be a surprise that the relaxation rate is  $\frac{1}{T}$ , as after preparing the state in a high-temperature plasma after an abrupt quench temperature is **the only scale**

 $\implies$  This also motivates to look at non-conformal examples of gauge/gravity correspondence.

### $\mathcal{N} = 2^*$ gauge theory (a QFT story)

 $\implies$  Start with  $\mathcal{N} = 4 SU(N)$  SYM. In  $\mathcal{N} = 1$  4d susy language, it is a gauge theory of a vector multiplet V, an adjoint chiral superfield  $\Phi$  (related by  $\mathcal{N} = 2$  susy to V) and an adjoint pair  $\{Q, \tilde{Q}\}$  of chiral multiplets, forming an  $\mathcal{N} = 2$  hypermultiplet. The theory has a superpotential:

$$W = \frac{2\sqrt{2}}{g_{YM}^2} \operatorname{Tr}\left(\left[Q, \tilde{Q}\right]\Phi\right)$$

We can break susy down to  $\mathcal{N} = 2$ , by giving a mass for  $\mathcal{N} = 2$  hypermultiplet:

$$W = \frac{2\sqrt{2}}{g_{YM}^2} \operatorname{Tr}\left(\left[Q,\tilde{Q}\right]\Phi\right) + \frac{m}{g_{YM}^2}\left(\operatorname{Tr}Q^2 + \operatorname{Tr}\tilde{Q}^2\right)$$

This theory is known as  $\mathcal{N} = 2^*$  gauge theory

When  $m \neq 0$ , the mass deformation lifts the  $\{Q, \tilde{Q}\}$  hypermultiplet moduli directions; we are left with the (N-1) complex dimensional Coulomb branch, parametrized by

$$\Phi = \operatorname{diag}\left(a_1, a_2, \cdots, a_N\right), \qquad \sum_i a_i = 0$$

We will study  $\mathcal{N} = 2^*$  gauge theory at a particular point on the Coulomb branch moduli space:

$$a_i \in [-a_0, a_0], \qquad a_0^2 = \frac{m^2 g_{YM}^2 N}{\pi}$$

with the (continuous in the large N-limit) linear number density

$$\rho(a) = \frac{2}{m^2 g_{YM}^2} \sqrt{a_0^2 - a^2}, \qquad \int_{-a_0}^{a_0} da \ \rho(a) = N$$

**Reason:** we understand the dual supergravity solution only at this point on the moduli space.

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**Reason:** This moduli space point is a large-N saddle point obtained from localization (in 2 transparencies)

 $\mathcal{N} = 2^*$  gauge theory (a supergravity story — a.k.a Pilch-Warner flow) Consider 5d gauged supergravity, dual to  $\mathcal{N} = 2^*$  gauge theory. The effective five-dimensional action is

$$S = \frac{1}{4\pi G_5} \int_{\mathcal{M}_5} d\xi^5 \sqrt{-g} \left(\frac{1}{4}R - (\partial \alpha)^2 - (\partial \chi)^2 - \mathcal{P}\right) \,,$$

where the potential  $\mathcal{P}$  is

$$\mathcal{P} = \frac{1}{16} \left[ \left( \frac{\partial W}{\partial \alpha} \right)^2 + \left( \frac{\partial W}{\partial \chi} \right)^2 \right] - \frac{1}{3} W^2 \,,$$

with the superpotential

$$W = -\frac{1}{\rho^2} - \frac{1}{2}\rho^4 \cosh(2\chi), \qquad \alpha \equiv \sqrt{3}\ln\rho$$

 $\implies$  The 2 supergravity scalars  $\{\alpha, \chi\}$  are holographic dual to dim-2 and dim-3 operators which are nothing but (correspondingly) the bosonic and the fermionic mass terms of the  $\mathcal{N} = 4 \rightarrow \mathcal{N} = 2$  SYM mass deformation.

PW geometry ansatz:

$$ds_5^2 = e^{2A} \left( -dt^2 + d\vec{x}^2 \right) + dr^2$$

solving the Killing spinor equations, we find a susy flow:

$$\frac{dA}{dr} = -\frac{1}{3}W, \qquad \frac{d\alpha}{dr} = \frac{1}{4}\frac{\partial W}{\partial \alpha}, \qquad \frac{d\chi}{dr} = \frac{1}{4}\frac{\partial W}{\partial \chi}$$

Solutions to above are characterized by a single parameter k:

$$e^A = \frac{k\rho^2}{\sinh(2\chi)}, \qquad \rho^6 = \cosh(2\chi) + \sinh^2(2\chi) \ln \frac{\sinh(\chi)}{\cosh(\chi)}$$

In was found (Polchinski, Peet, AB) that

$$k = 2m$$

 $\implies$  Precision test on  $\mathcal{N} = 2^*$  holography from Pestun's localization

- Supersymmetrically compactify  $\mathcal{N} = 2^*$  gauge theory on  $S^4$
- Moduli of the theory are conformally coupled scalars, so they will all be lifted via coupling to  $S^4$  curvature
- The exact partition function of the compactified theory is known due to Pestun's localization (reduces to a matrix model):

$$Z_{\mathcal{N}=2^*} = \int d^{N-1}\hat{a} \prod_{i < j} \frac{(\hat{a}_i - \hat{a}_j)^2 H^2(\hat{a}_i - \hat{a}_j)}{H(\hat{a}_i - \hat{a}_j - mR)H(\hat{a}_i - \hat{a}_j + mR)}$$
$$\times e^{-\frac{8\pi^2 N}{\lambda} \sum_j \hat{a}_j^2} |\mathcal{Z}_{\text{inst}}|^2$$

where

$$H(x) \equiv \prod_{n=1}^{\infty} \left(1 + \frac{x^2}{n^2}\right)^n e^{-\frac{x^2}{n}}$$

• In the large-N limit the partition function is dominated by the saddle point, that can be computed analytically (AB, J.G.Russo and K.Zarembo, 1301.1597)

## $\implies$ One recovers:

- the moduli space point picked out by supergravity (as a matrix model saddle point)
- the susy Wilson loops agree both in matrix model and in holographic dual
- the matrix model free energy agree with the holographic free energy (Bobev et.al, arXiv:1311.1508)

 $\implies$  All these checked are in addition to earlier agreement with the metric on the moduli space computed either in supergravity or from QFT using Seiberg-Witten techniques  $\implies$  What do we do:

• Take PW gravitational dual to  $\mathcal{N} = 2^*$  gauge theory and construct black hole solutions. The thermodynamics of the black hole has a nontrivial dependence of 2 scales: T and m. The m dependence is quite profound:

$$s(T) \propto \begin{cases} N^2 \ T^3 \ , \qquad \frac{m}{T} \ll 1 \\ N^2 \ \frac{T^4}{m} \ , \qquad \frac{m}{T} \gg 1 \end{cases}$$

- Using the standard techniques we compute quasinormal modes of the BH corresponding to
  - the stress-energy tensor  $T_{\mu\nu}$
  - operators  $\{\mathcal{O}_2, \mathcal{O}_3\}$  inducing the RG flow
  - 'passive' operators  $\{\mathcal{O}_2, \mathcal{O}_3\}$
  - also study the momentum dependence on the quasinormal frequencies to get an idea of relaxation of spatially inhomogeneous excitations



 $\Rightarrow$  (L) Trace of the energy-momentum tensor normalized to the energy density of  $\mathcal{N} = 4$  SYM ( $\epsilon_0 = \frac{3}{8}\pi^2 N_c^2 T^4$  with  $N_c$  denoting the number of colors) as a function of m/T. The results indicate that, thermodynamically, the effects of the conformal symmetry breaking are the strongest at  $m/T \approx 4.8$ .

 $\implies$  (R) Trace anomaly in deep IR — approach to a  $CFT_5$ 

 $\implies$  There is an interesting story with the IR properties of the flow:



 $\Rightarrow$ Ratio of viscosities  $\frac{\zeta}{\eta}$  versus the speed of sound in  $\mathcal{N} = 2^*$  gauge theory plasma (AB, arXiv:0708.3459). Dashed line is the bulk viscosity bound,  $\frac{\zeta}{\eta} \geq 2\left(\frac{1}{3} - c_s^2\right)$ . A single point represents extrapolation of the speed of sound and the viscosity ratio to  $T \to +0$ .

$$\implies$$
 Note, for a  $CFT_5$ ,  $\epsilon = 4p$ , so  $c_{s,CFT_5}^2 - \frac{1}{4}$ , and  
 $\frac{1}{3} - c_s^2 = \frac{1}{3} - \frac{1}{4} = \frac{1}{12} = 0.08333\cdots$ 



Real (green continuous) and minus imaginary (red dashed) parts of the lowest quasinormal mode frequencies for operators of dimensions  $\Delta = 2, 3$ and 4 (from bottom to top). The frequencies do not change significantly as a function of m/T, which leads to universal equilibration in 1/T. One can also infer from this plot that all the frequencies asymptote at low temperatures to the quasinormal mode of a massless scalar field living in the (1+5)-dimensional AdS-Schwarzschild geometry (dotted curves).



• Momentum dependence of the real (green) and minus imaginary part (red) of the QNM frequency of operators with  $\Delta = 2, 3$  and 4 (from bottom to top) for m/T = 0 ( $\mathcal{N} = 4$  SYM, dashed) and m/T = 4.8 (continuous). Surprisingly, corresponding curves are very close to each other despite of the fact that m/T = 4.8 matches the locus of the maximal deviation from conformal invariance in thermodynamics of  $\mathcal{N} = 2^*$ .

• There is **very weak** dependence on k in  $\text{Im}\,\omega$ 

## **Conclusion:**

- I argued that relaxation time in strongly coupled plasma is encoded in the spectrum of quasinormal modes of BH in the holographic dual
- One can have a controlled examples of the top-down holography where it is possible to systematically study effects of non-conformality on the relaxation time
- We found that

$$\tau_{relax} \propto \frac{1}{T}$$

 $\underline{\mathbf{universally}}$ , even though are other microscopic scales in the plasma (masses, etc)

- The spatial relaxation is <u>ultralocal</u> imaginary parts of the quasinormal modes are almost flat in momentum k
- Similar conclusions are reached in other models:
  - thermalization in  $\mathcal{N} = 4$  plasma in the presence of charge densities/magnetic fields (J.Fuini, L.Yaffe, arXiv:1503.07148)
  - Bottom-up pheno models of holography (R.Janik et.al, arXiv:1503.07148; T.Ishii et.al, arXiv:1503.07766)