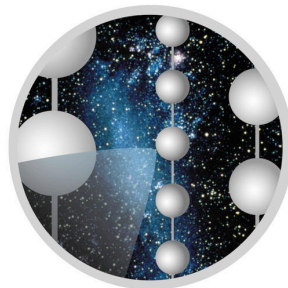




Cosmic-ray Spectrum, Composition, and Anisotropy Measured with IceCube/IceTop

Alessio Tamburro for the IceCube Collaboration
atamburro@idecube.wisc.edu

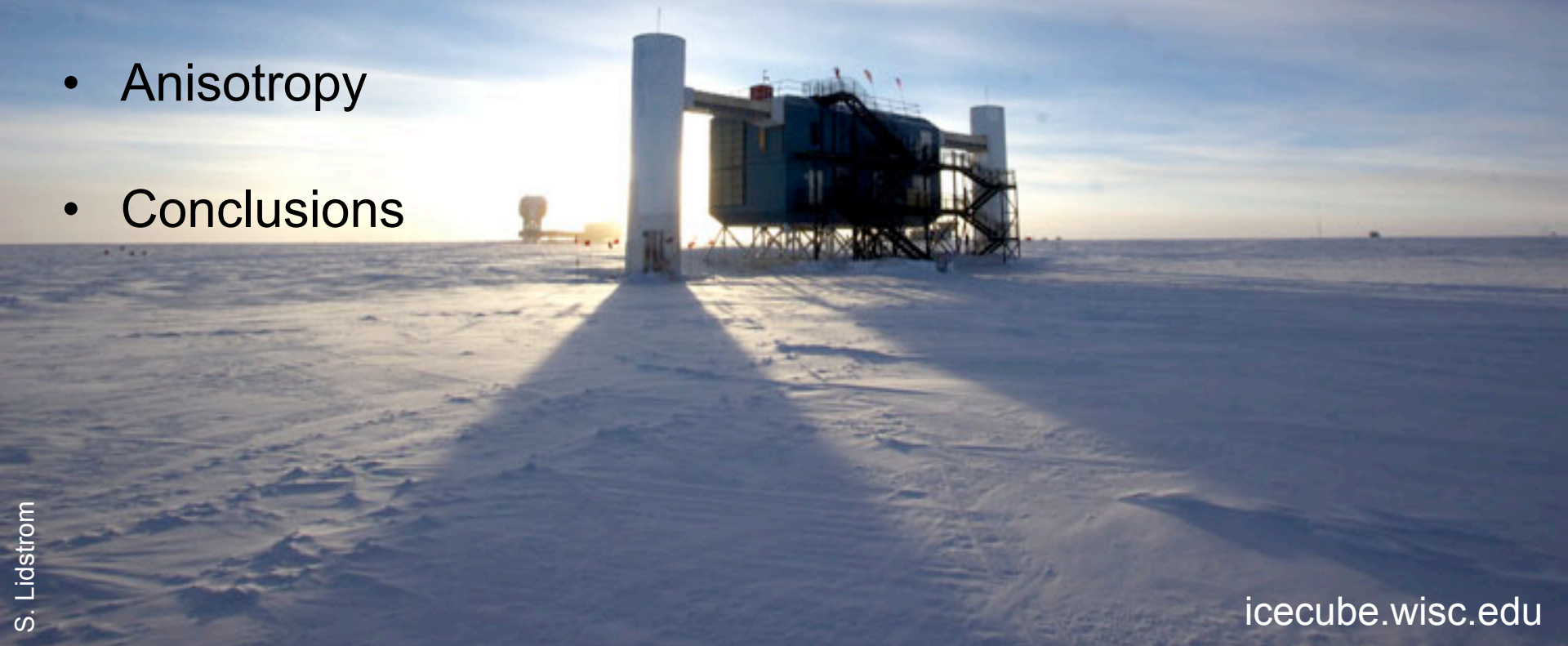


ICECUBE

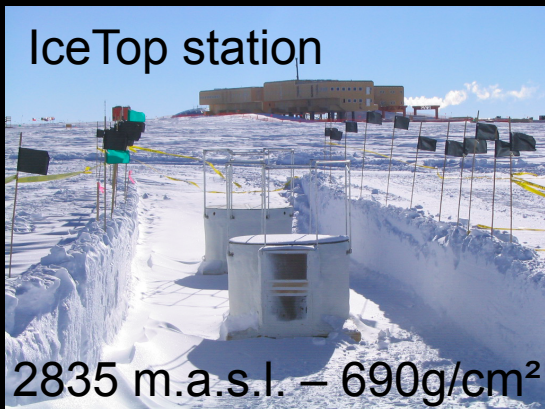


RICAP 2013, Rome – May 22, 2013

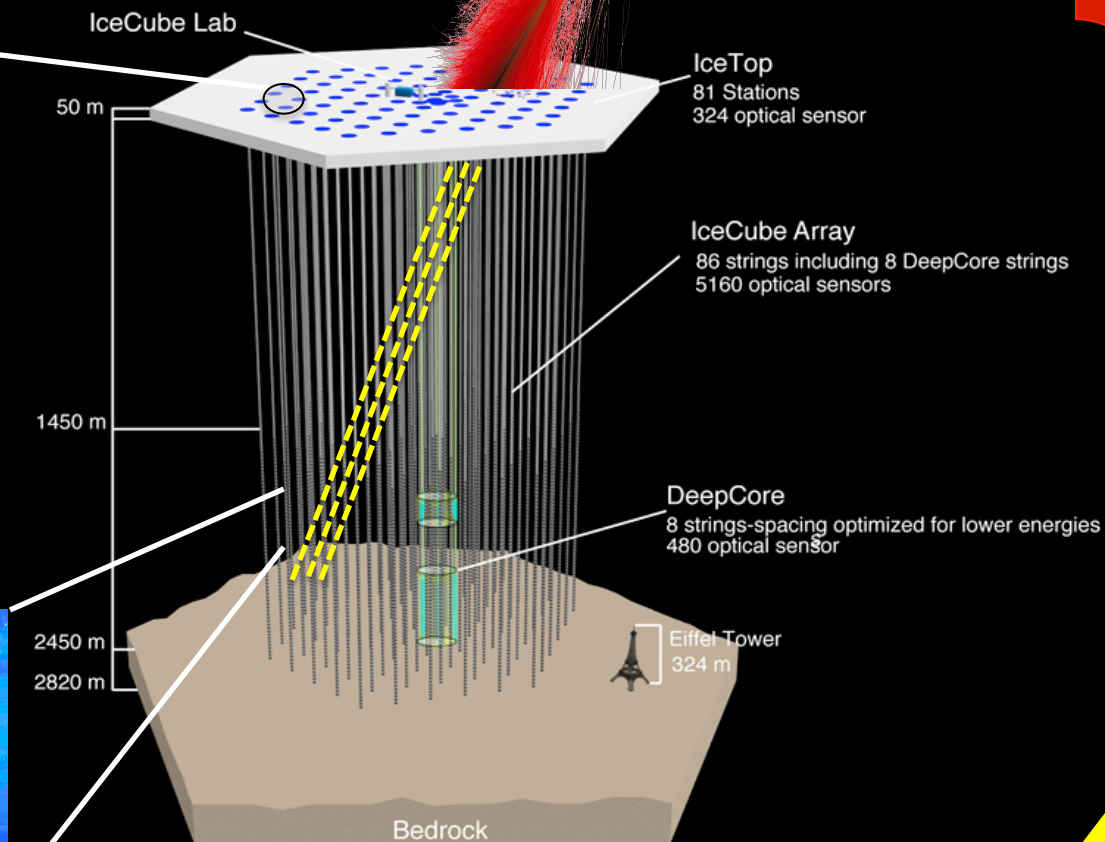
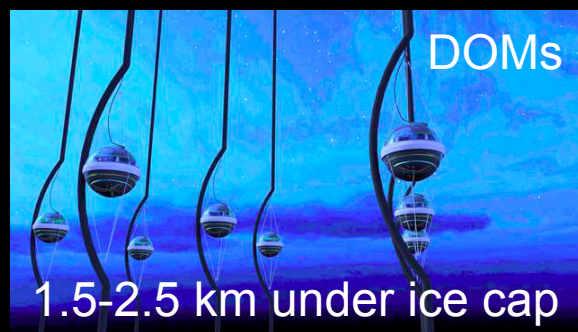
- IceCube and Cosmic-ray Physics
- IceTop-73 Spectrum Analysis
- IceTop-73/IceCube-79 Composition and Spectrum Analysis
- Anisotropy
- Conclusions



IceTop station

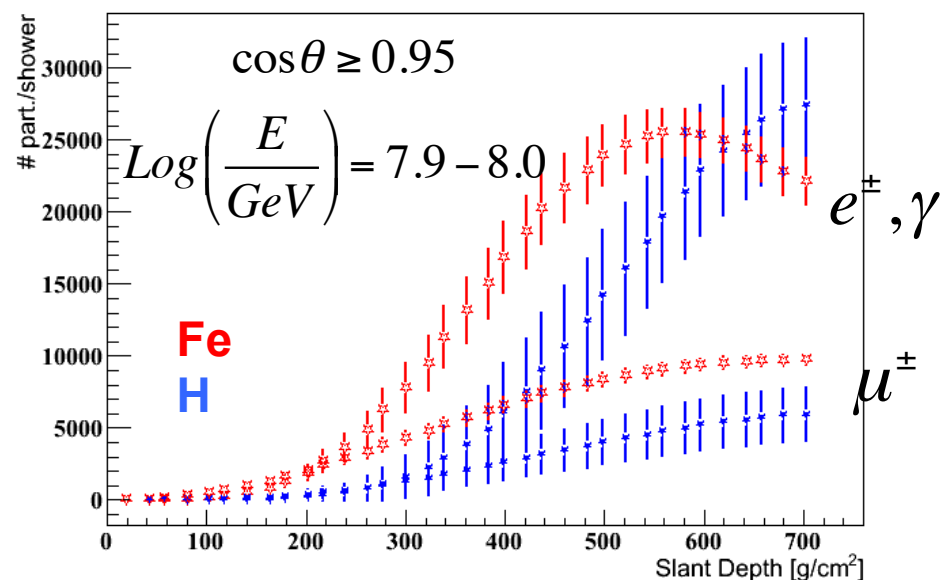
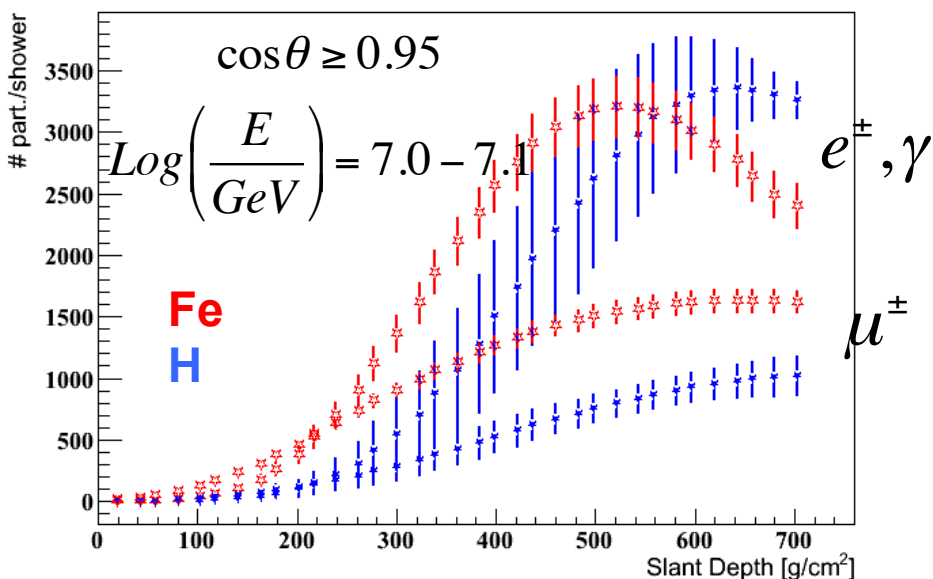


Close to shower max



@ $E \sim 3 \times 10^{16}$ eV or 30 PeV:
Over 2,000 events/yr
(Max $E \sim 3$ EeV due to km² size)

650 coinc. muon bundles/yr
(Max $E \sim$ EeV)



CR spectrum (PeV – EeV):

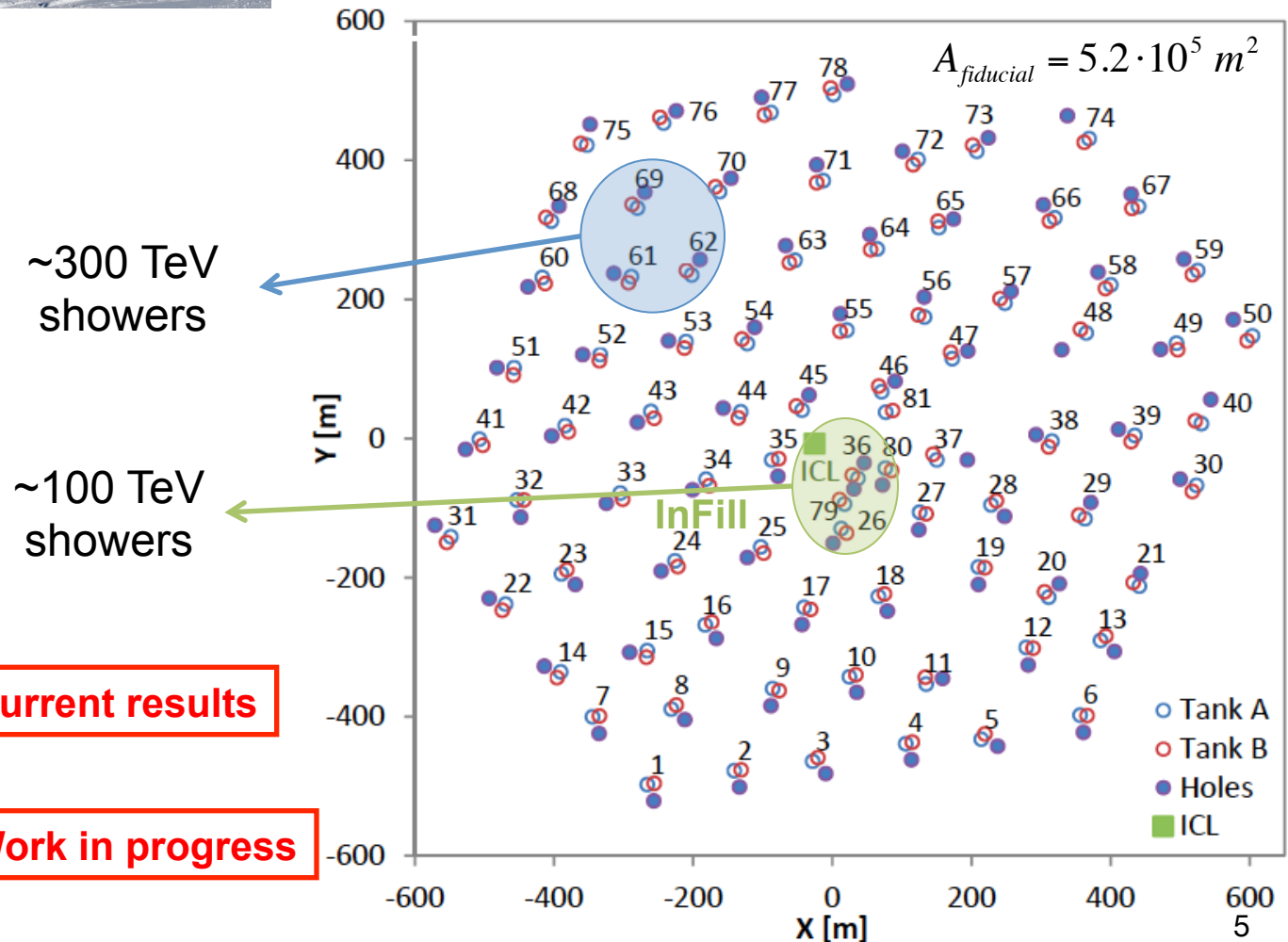
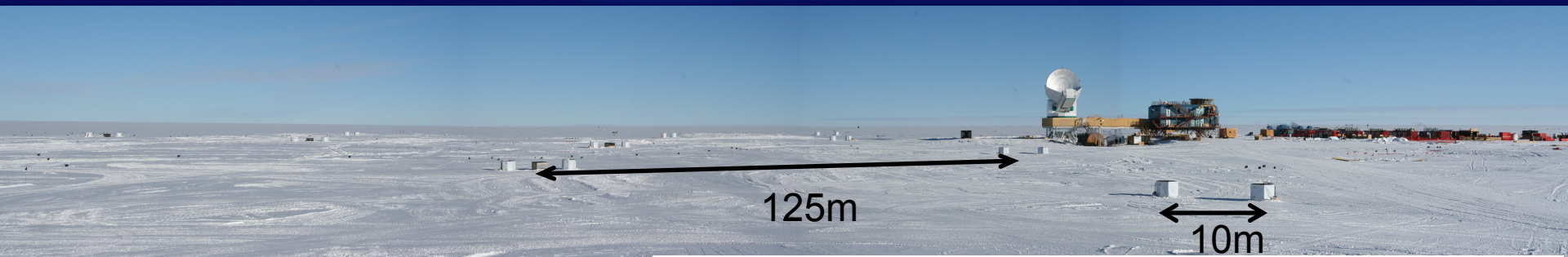
- less composition-dependent @ $10^{7.5}$ GeV and above
- precisely measured (low fluctuations at IceTop altitude)

CR composition (PeV – EeV):

- in-ice info allows for studying muon content of showers
- will see transition from galactic to extra-galactic?

CR anisotropy with IceCube/IceTop (10 TeV – 2 PeV, 10^{-3} level):

- Large statistics 10^{11} muons/yr – 10^8 showers/yr
- Angular resolution better than 10°



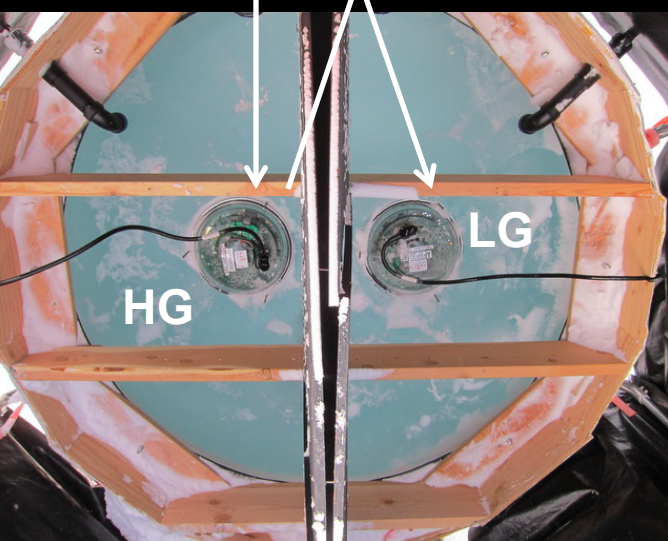
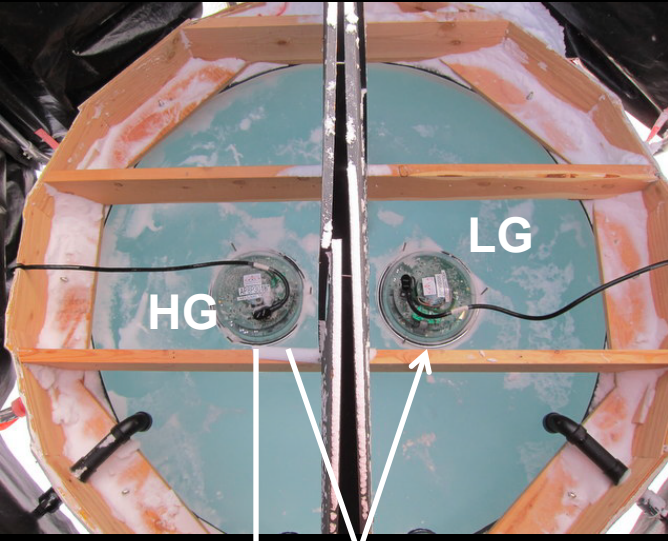
5+ stations
 $E \sim \text{PeV} - \text{EeV}$

Current results

3+ stations
 $E > 100 \text{ TeV}$

Work in progress

Tank A



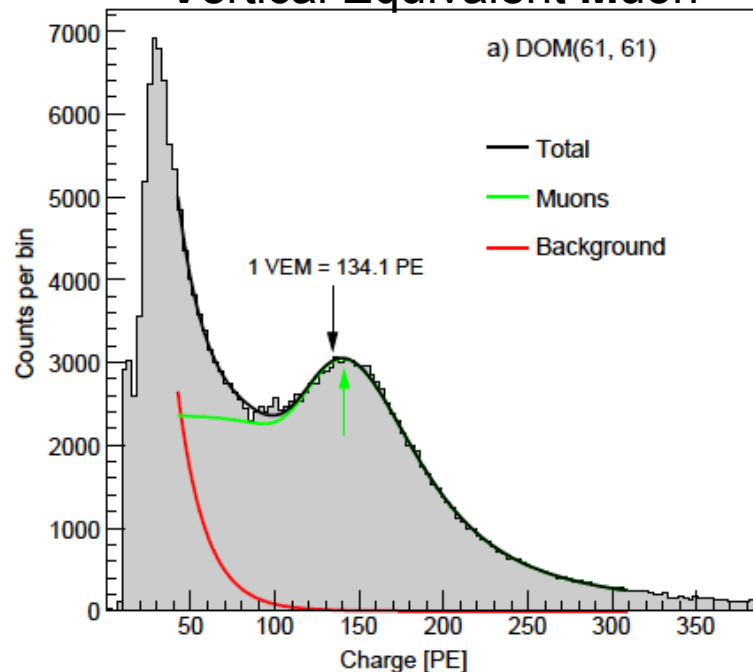
Tank B

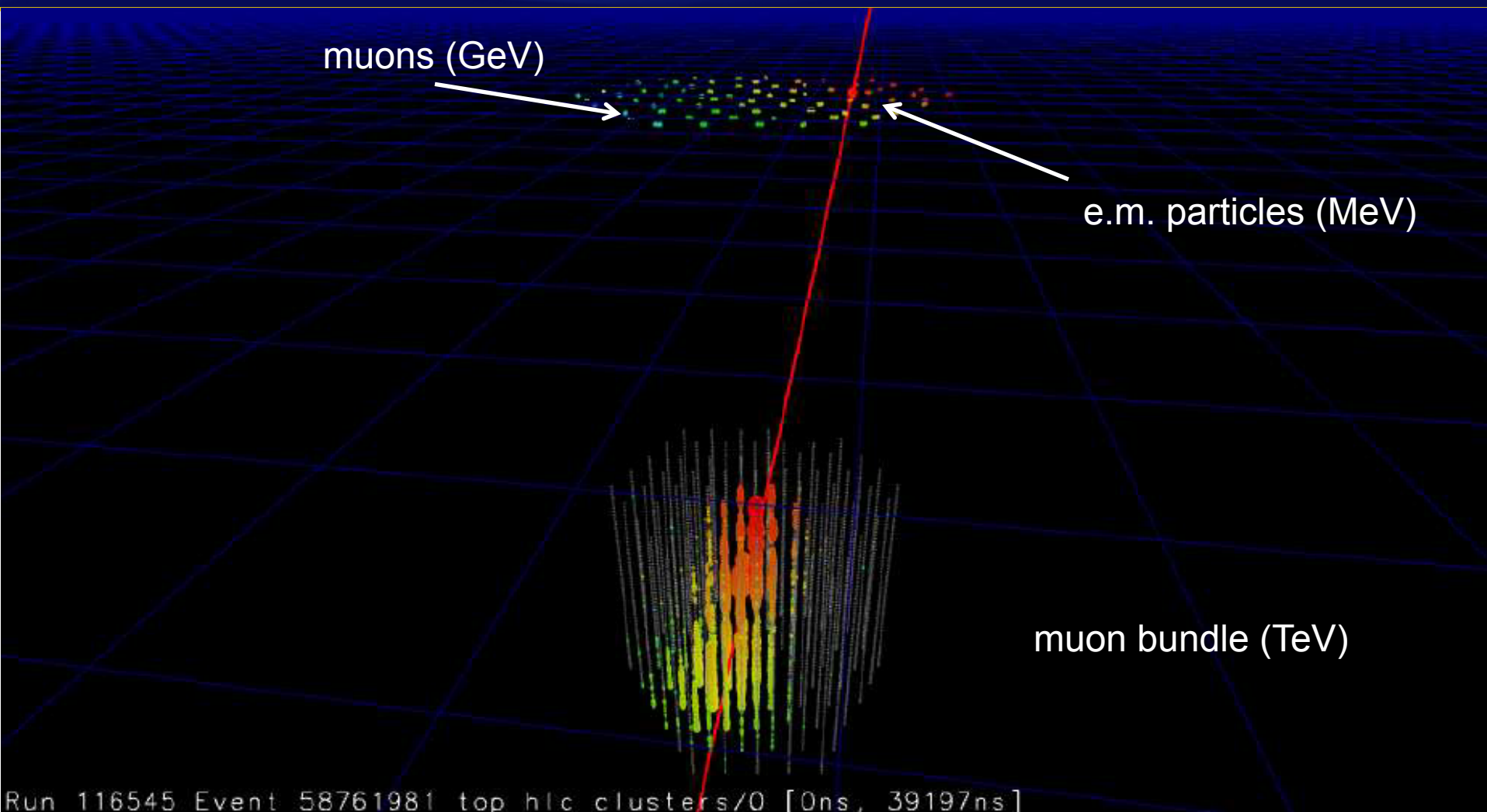
Single hits (1 DOM) or SLC
(muon detection + veto of air showers for in-ice studies)
(**2,000 Hz**)

Station trigger
HG-HG or HG-LG coincidence in 1 μ s (**30 Hz**)

IceTop trigger
3+ stations in 6 μ s (**20 Hz**)

Vertical Equivalent Muon



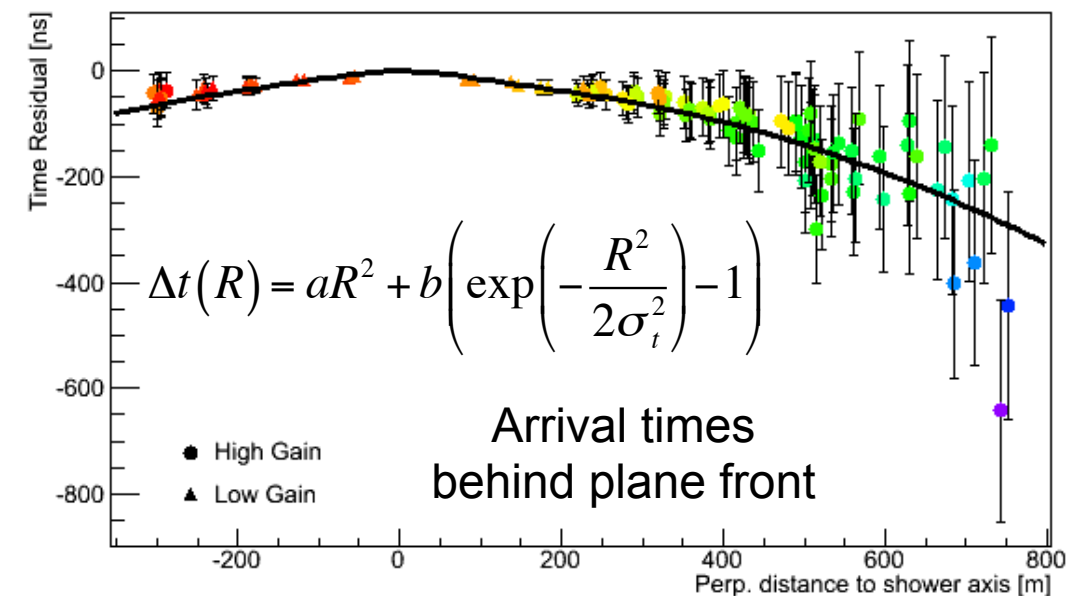
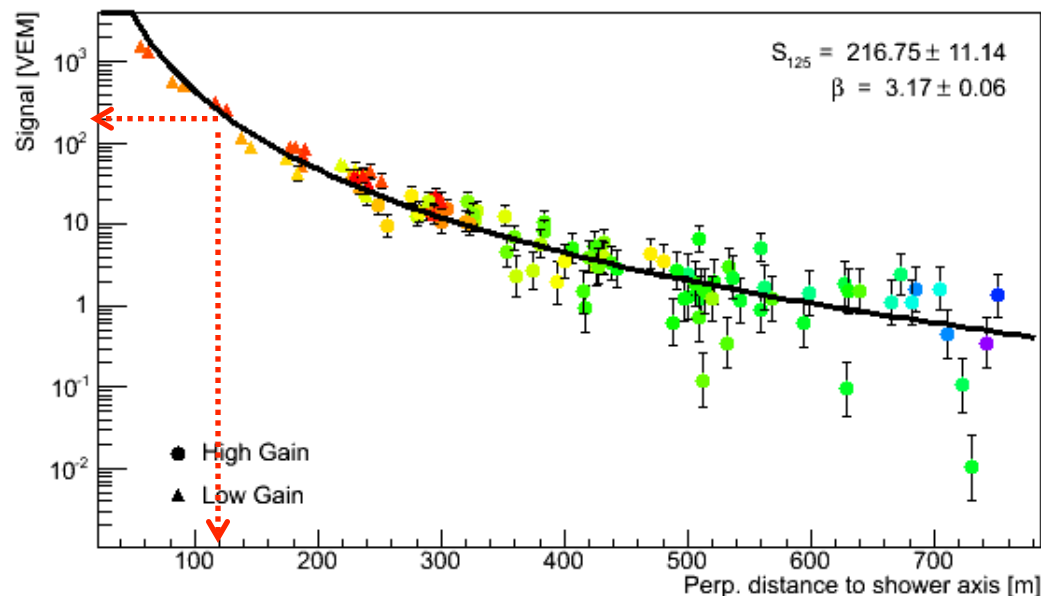
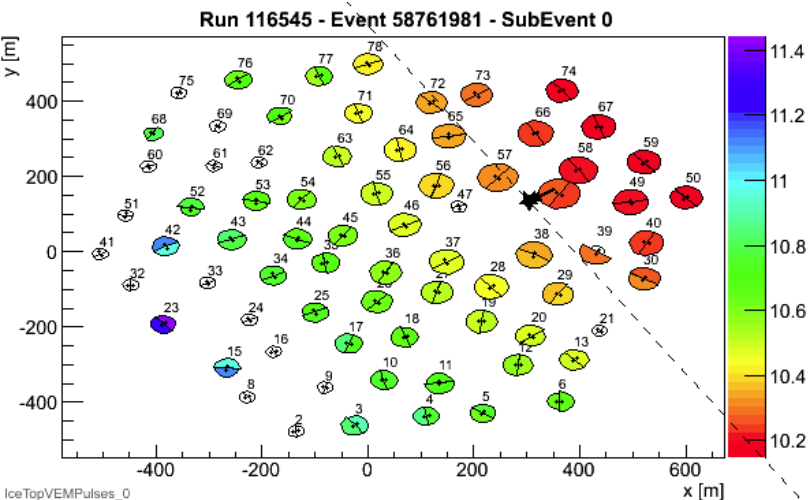


Coincident events:

- Shower core contained within IceTop array
- Muon bundle track through in-ice array

$$\theta = 11.5^\circ, \phi = 33.0^\circ$$

$$E = 2.90 \cdot 10^{17} \text{ eV (H assumption)}$$

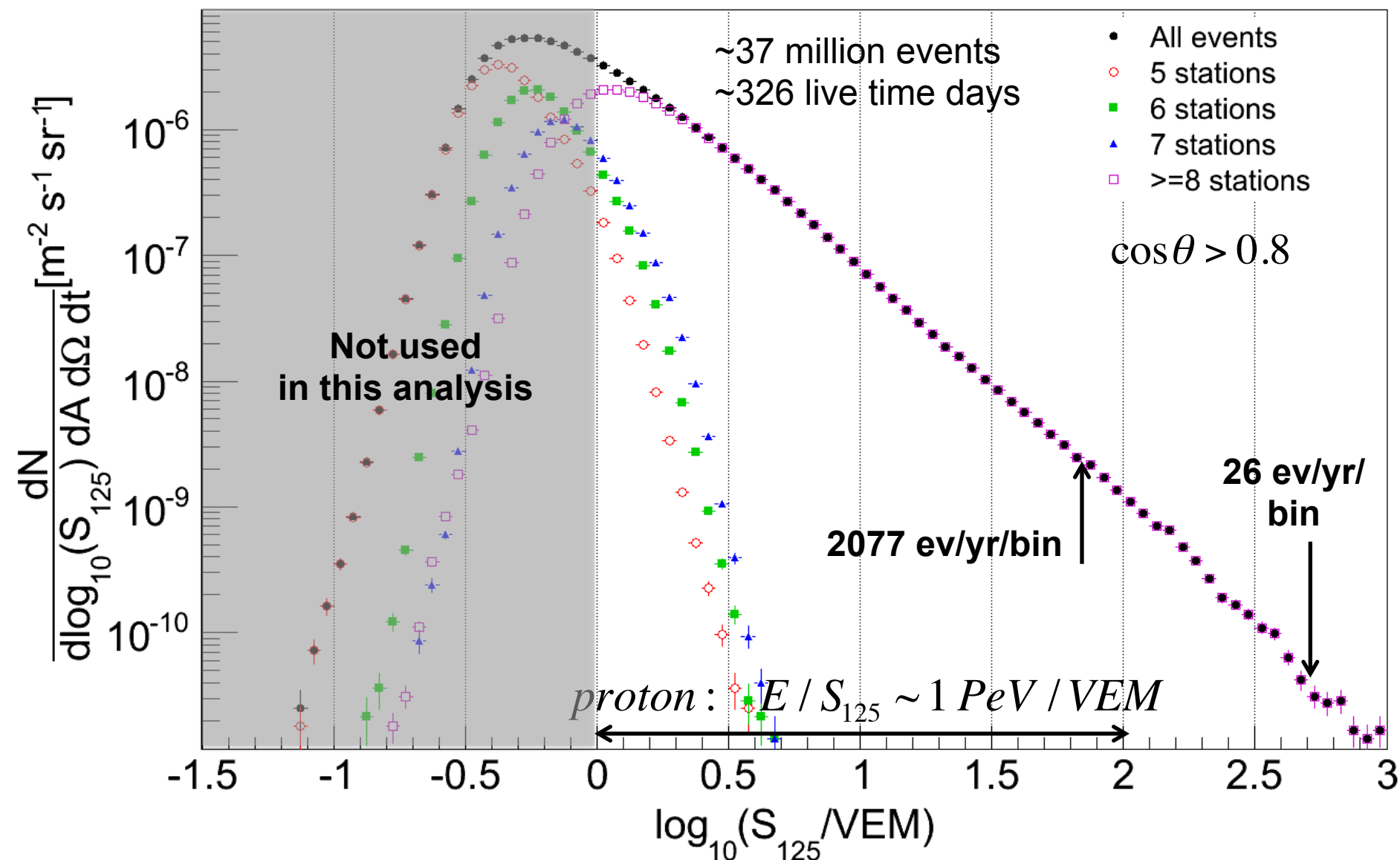


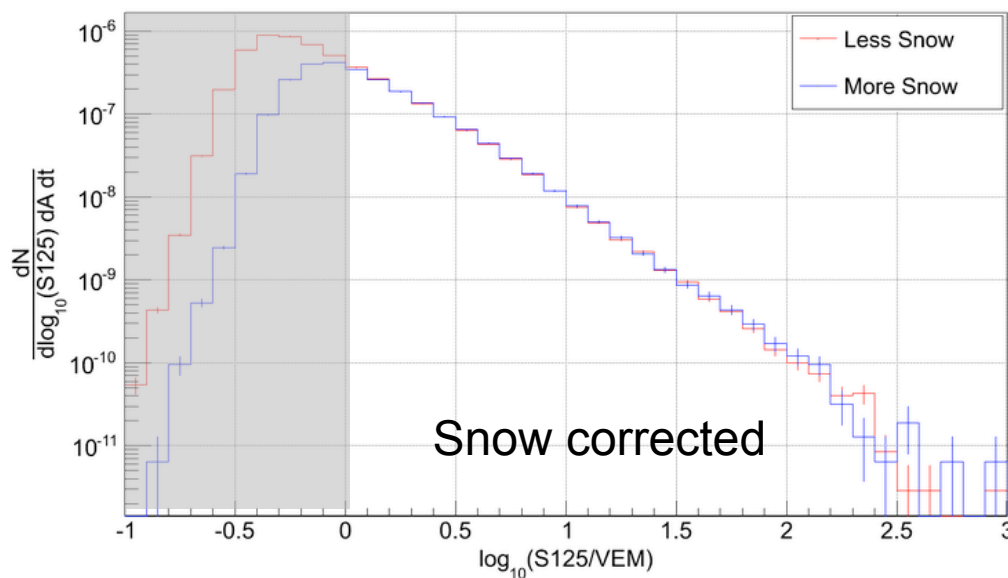
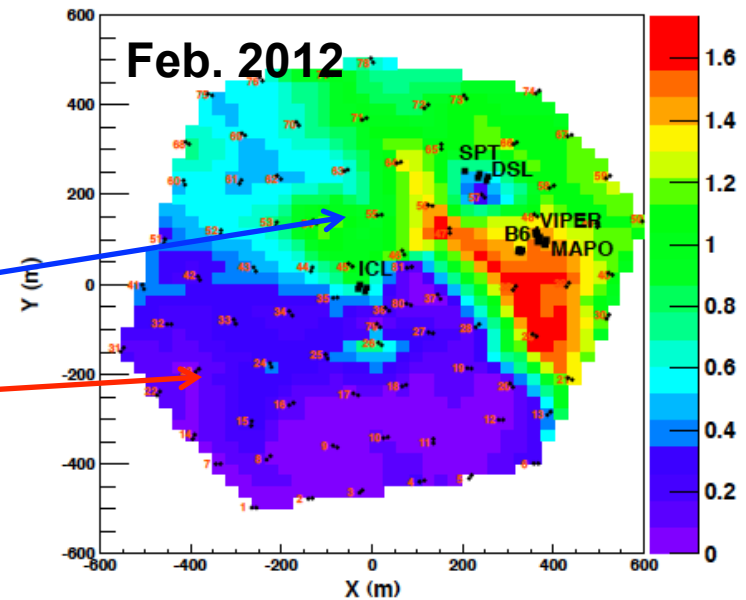
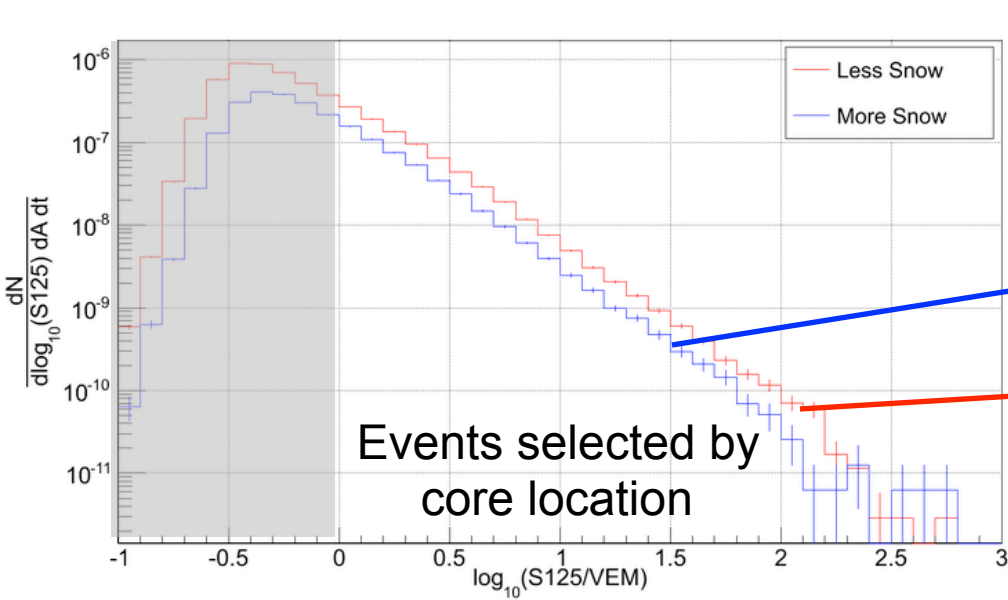
$$S(R) = S_{125} \left(\frac{R}{125m} \right)^{-\beta - \kappa \log(R/125m)}$$

S_{125} “Shower size” at 125m

β Slope at 125m

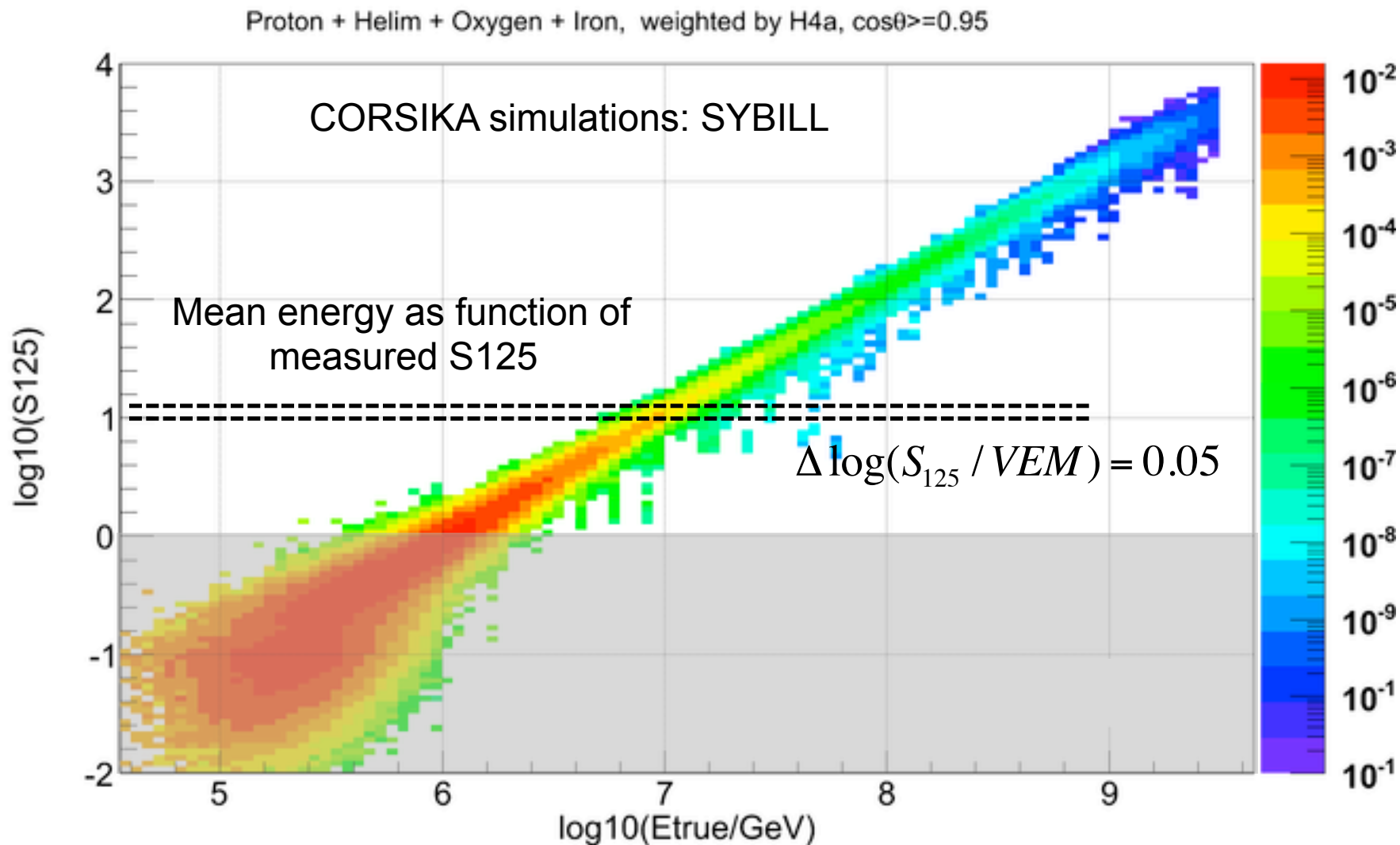
Tank signals are mostly
from shower e.m. component.

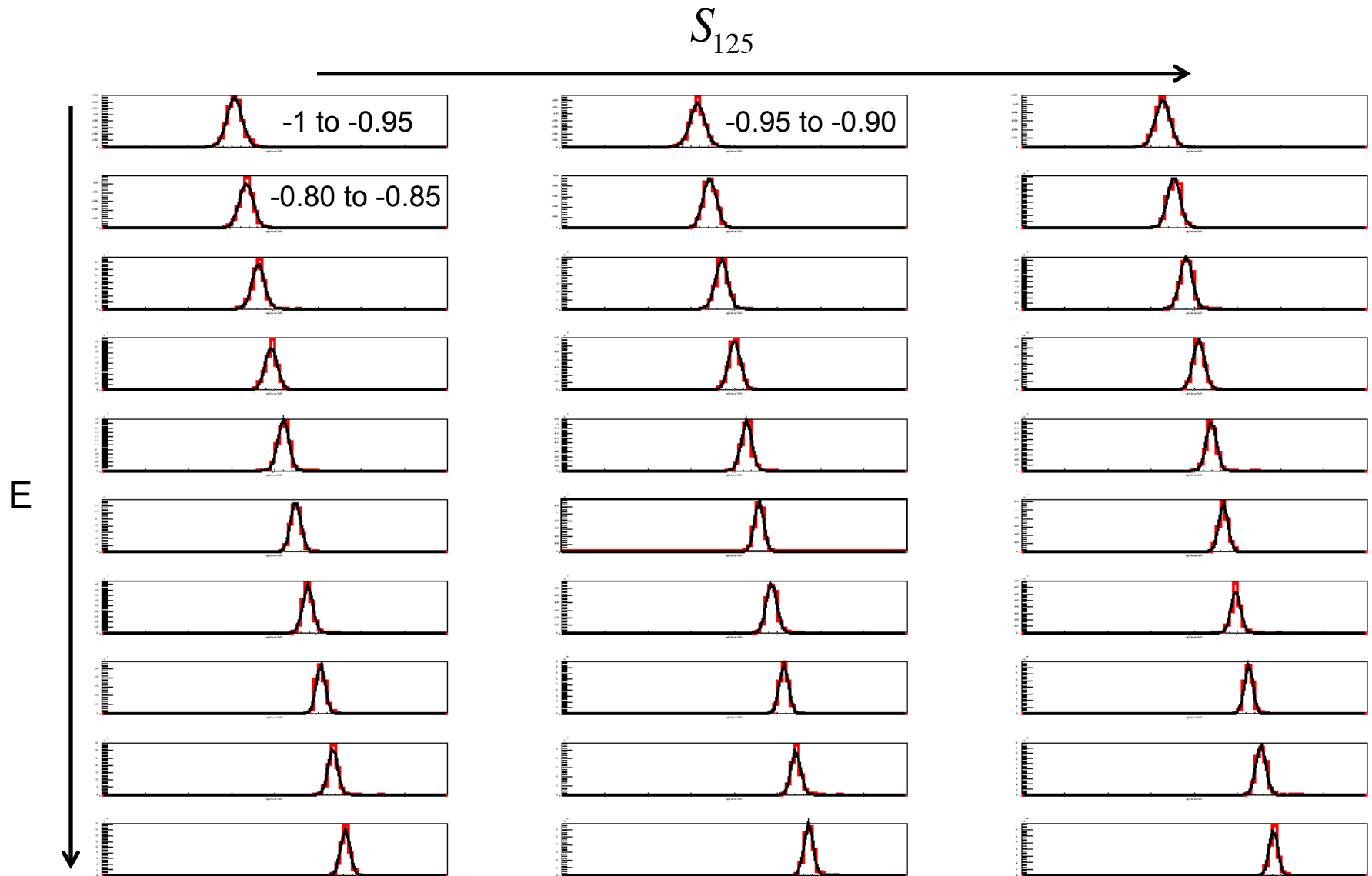


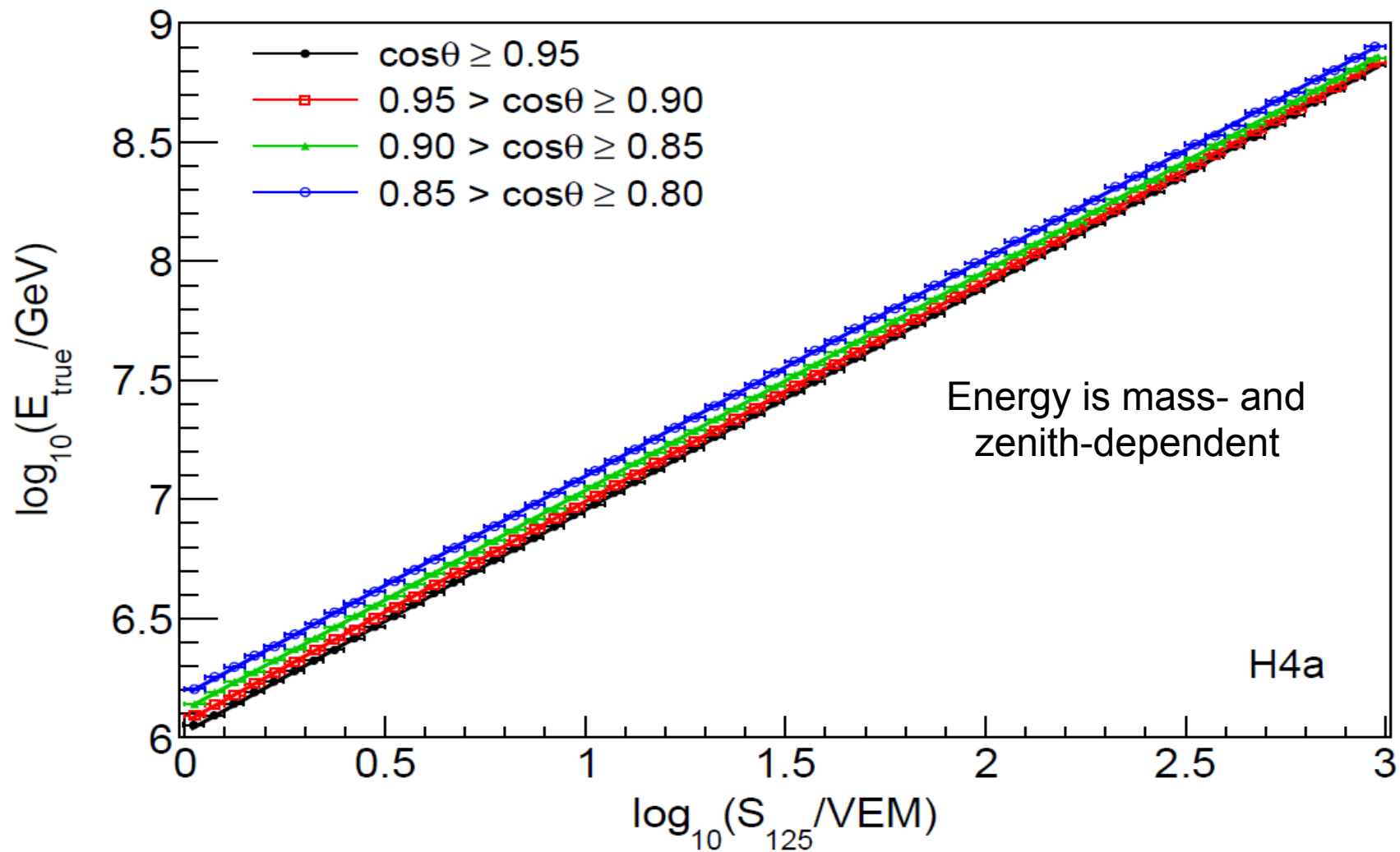


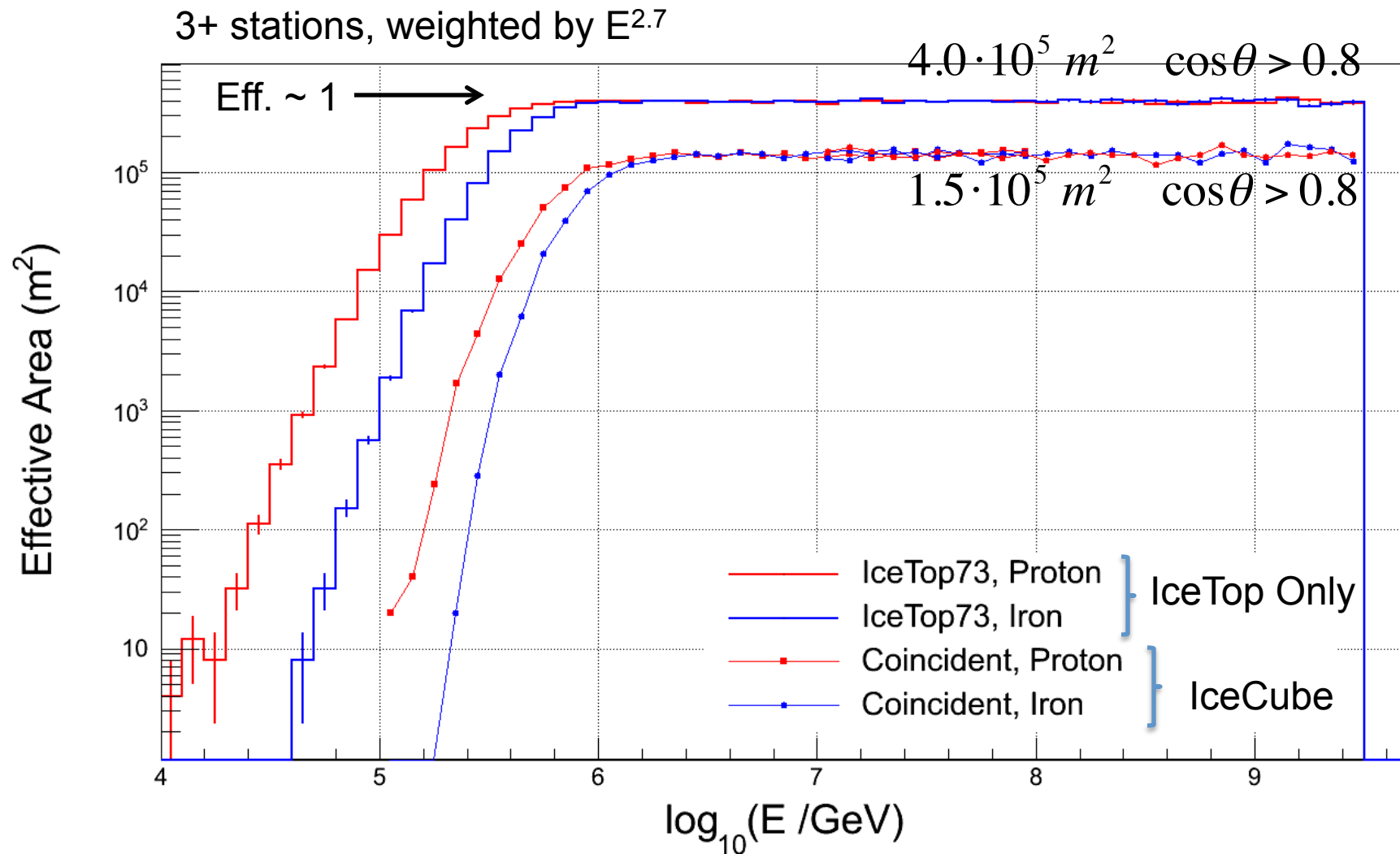
$$S_{corr} = S_{meas.} e^{\frac{z \cos \theta}{\lambda}}$$

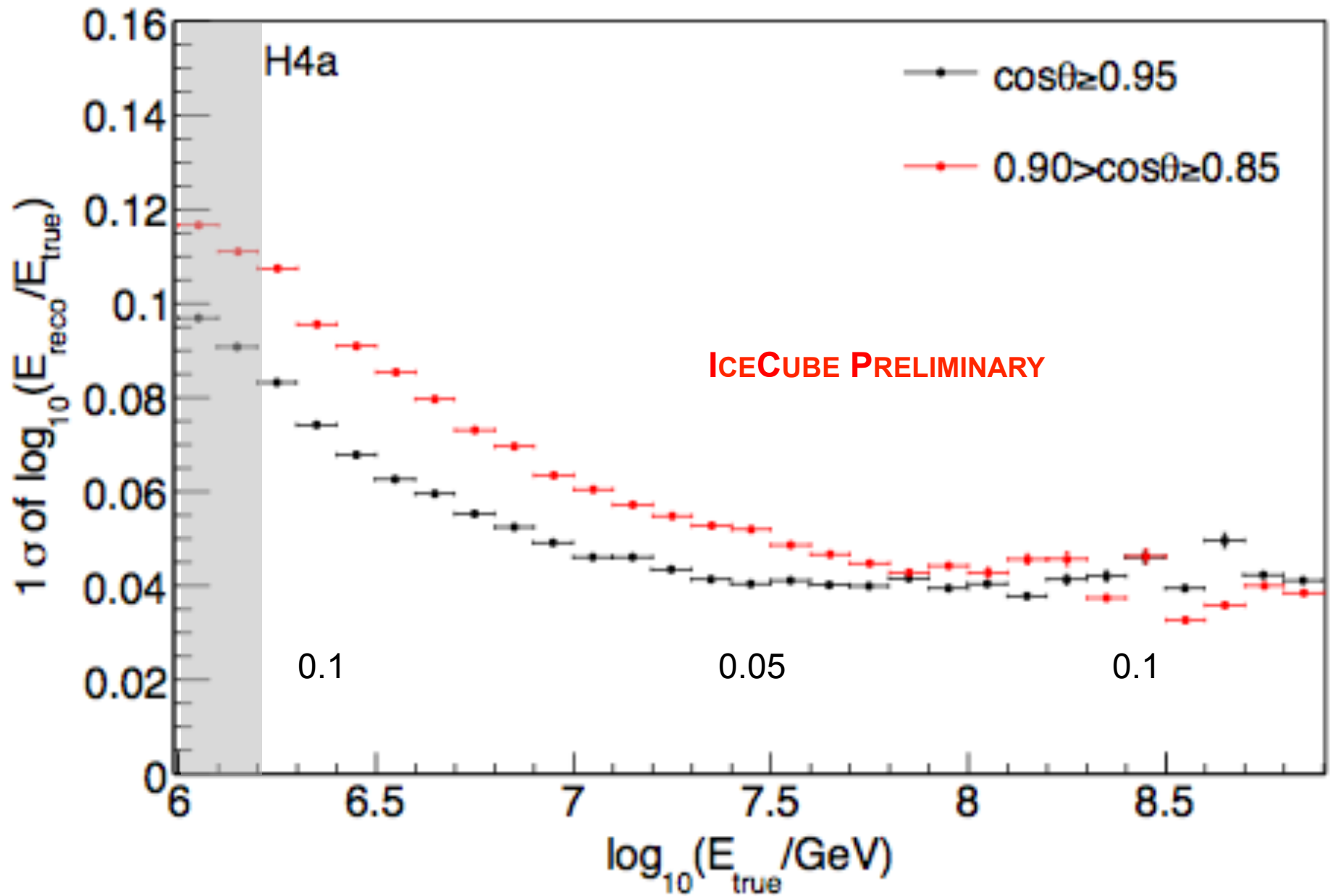
Need to assume reasonable composition to derive energy spectrum



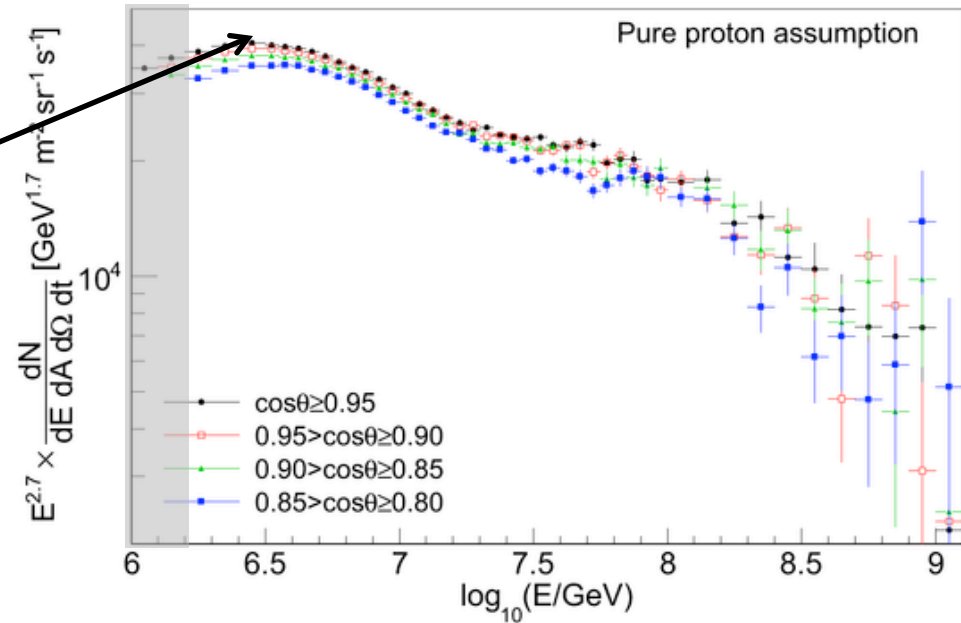




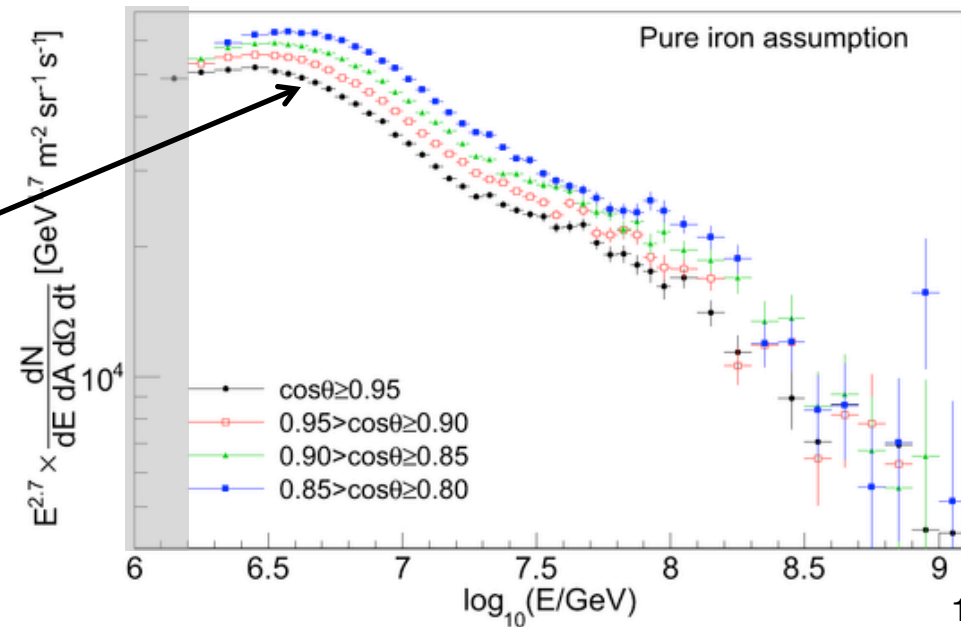


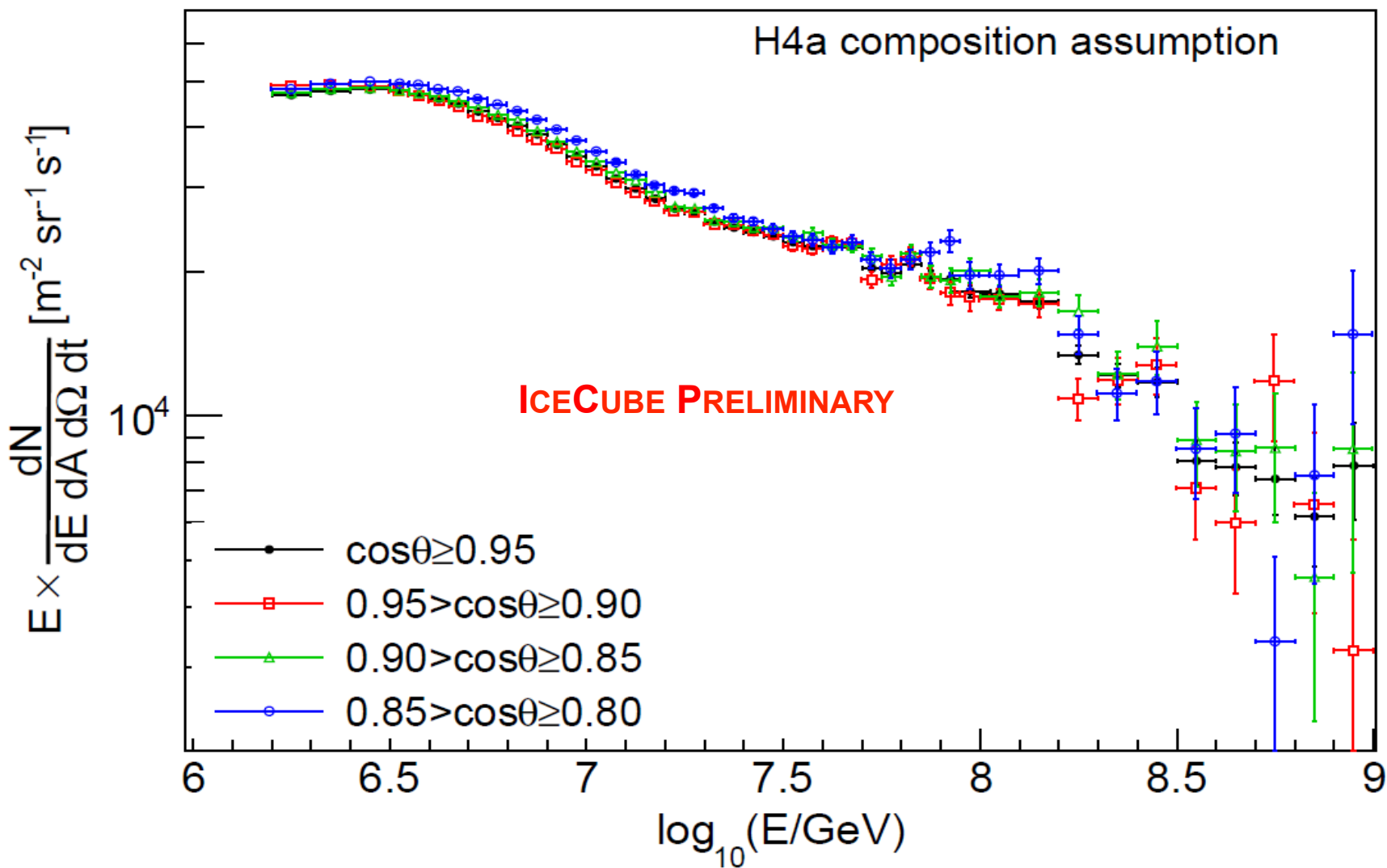


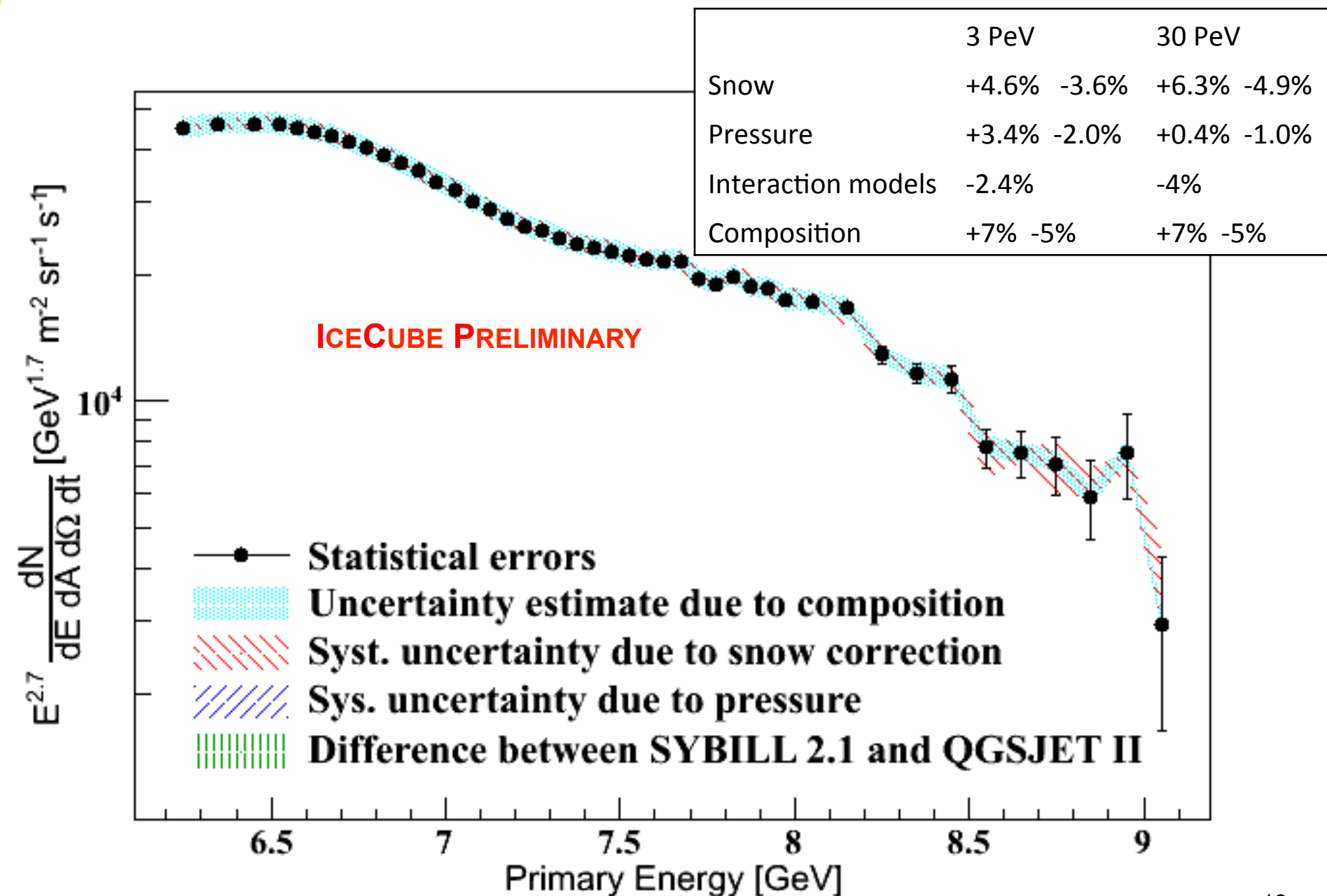
Inferred proton flux is **higher**
near the vertical



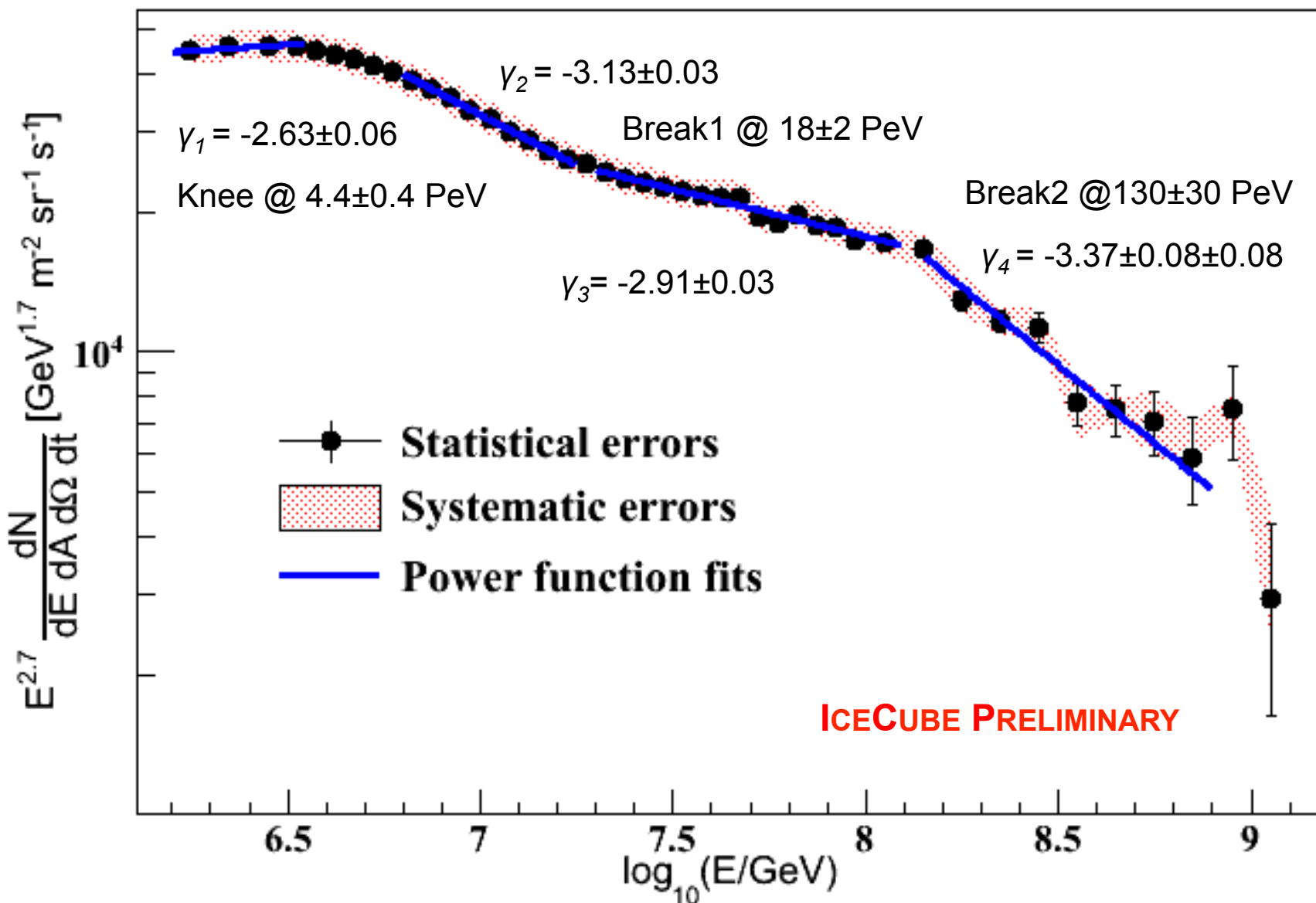
Inferred iron flux is **suppressed**
near the vertical

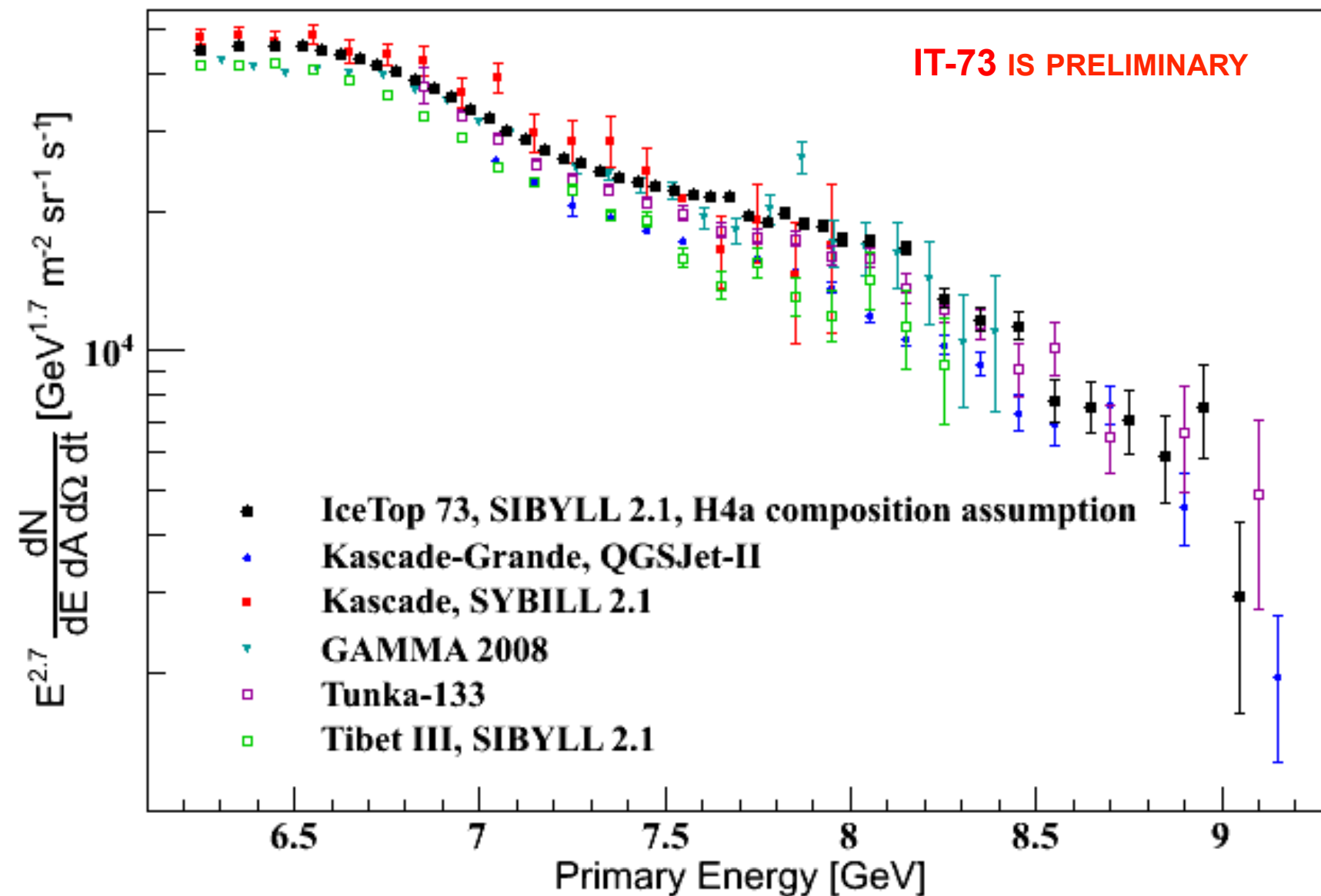


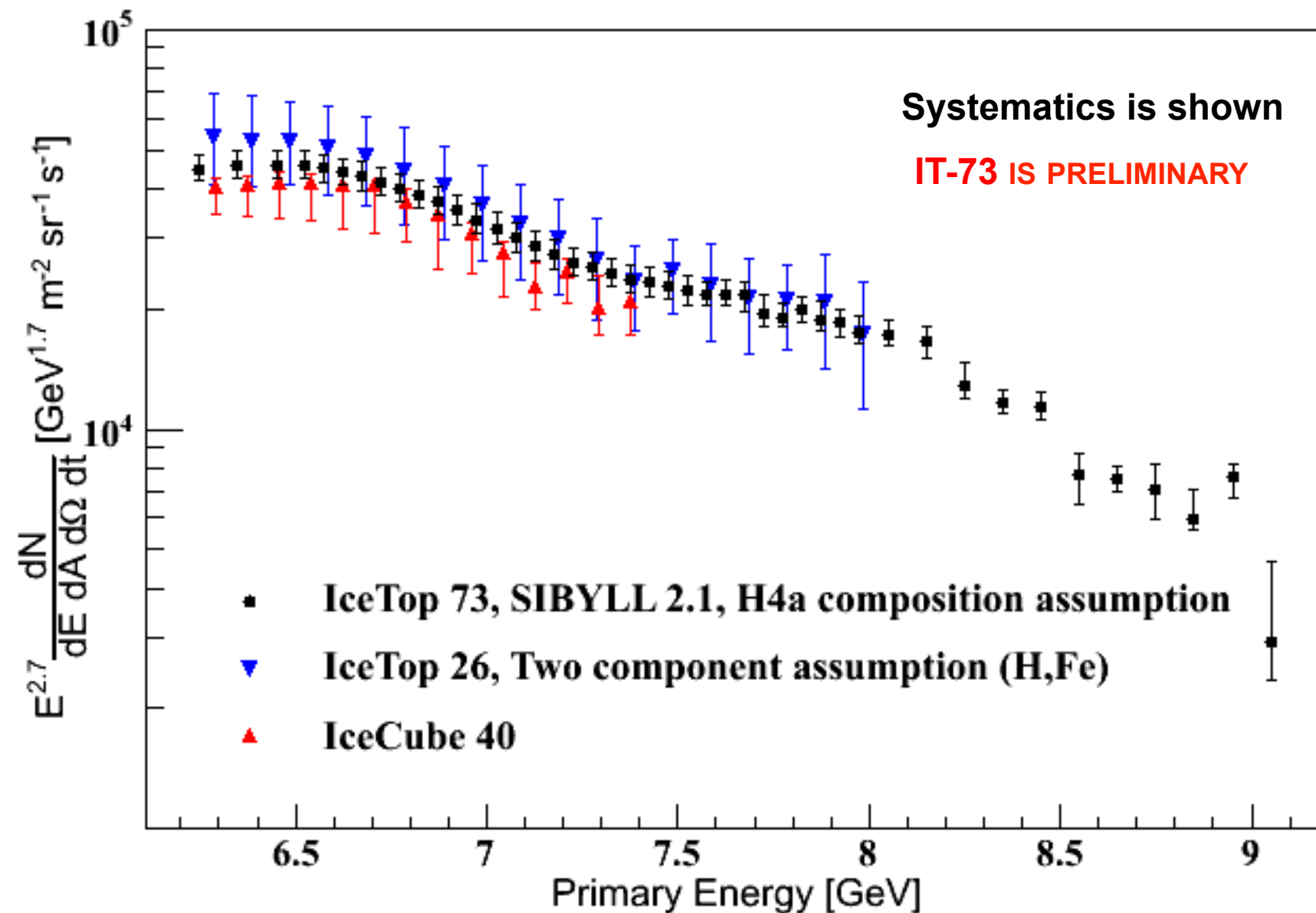


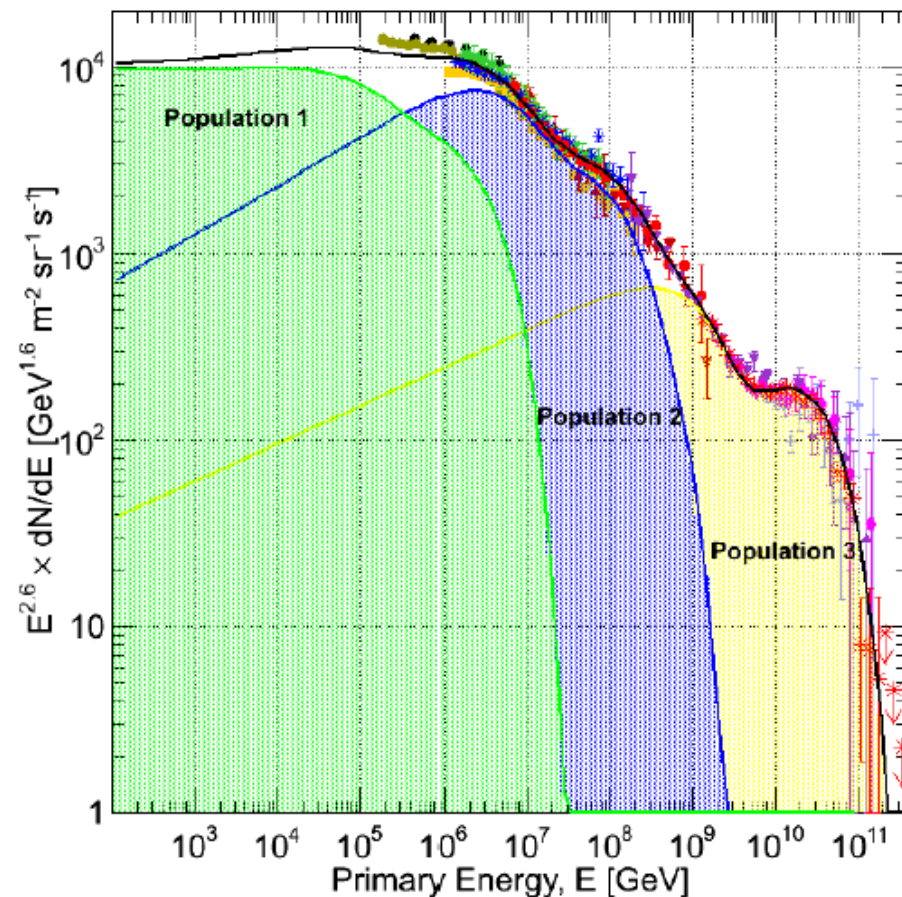
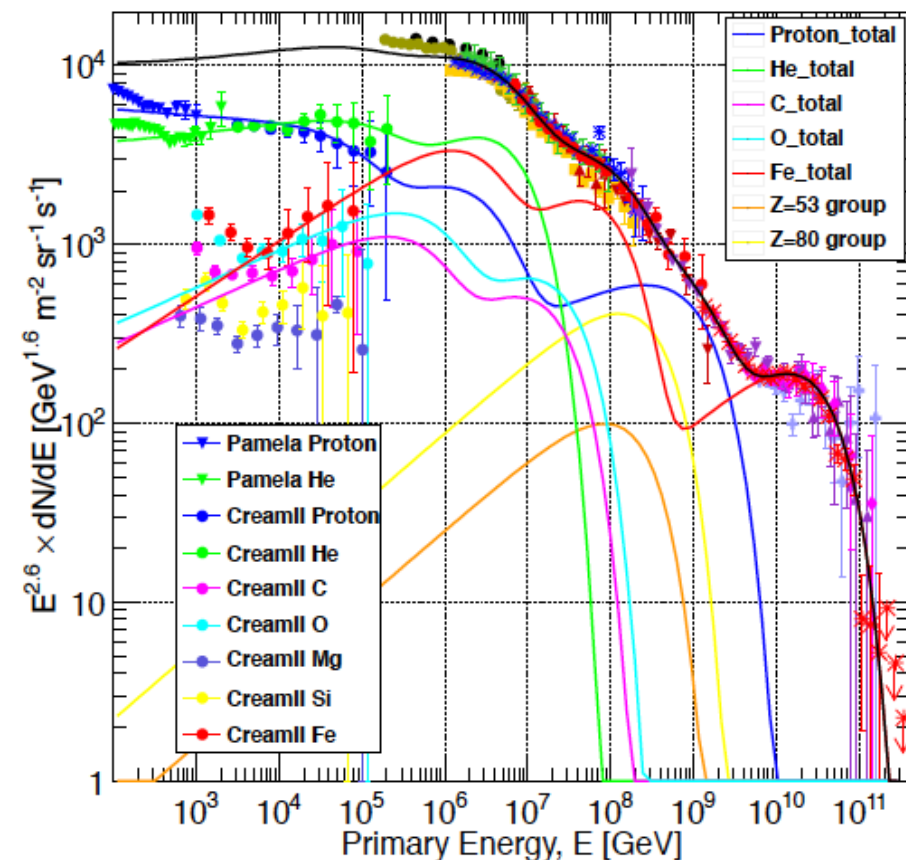


Energy “break” measured where a change in power law slope is observed









Including composition measurements when fitting Fe cutoffs identifies finer spectral structures corresponding to different populations of sources

Will IceCube see a transition from galactic to extra-galactic CRs?

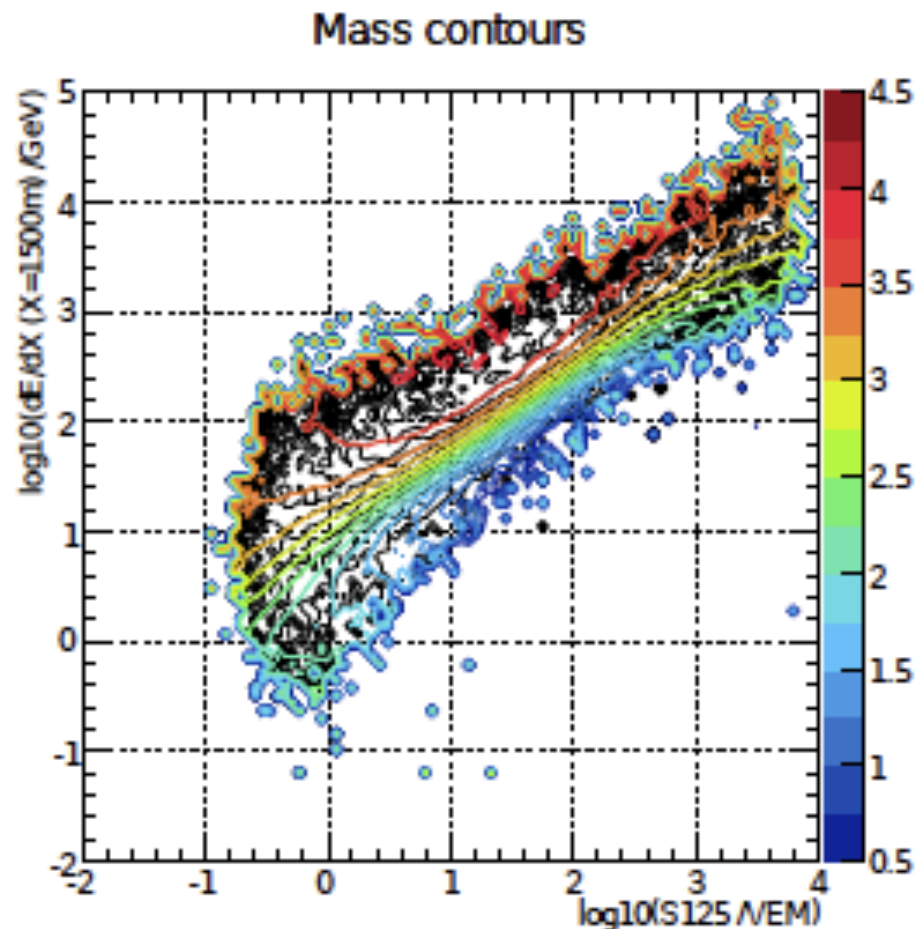
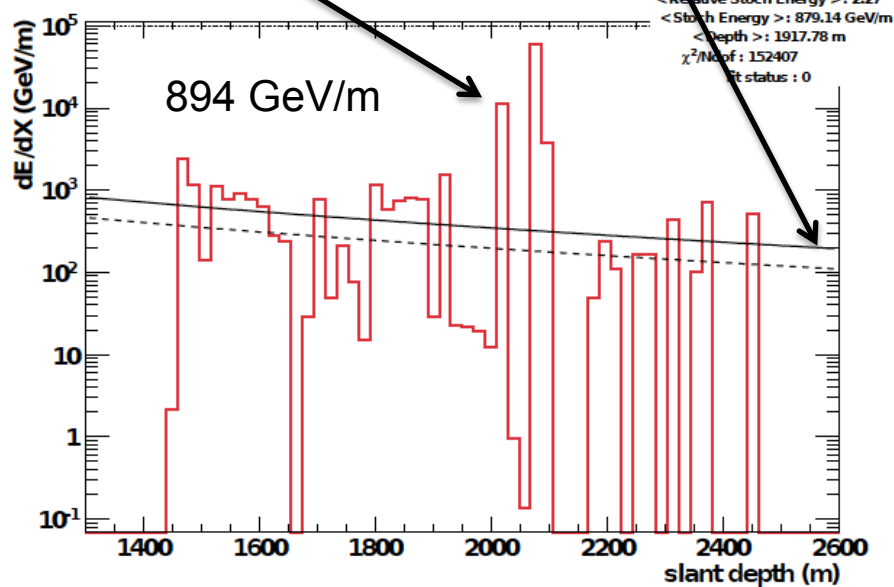
Gaisser, Stanev, Tilav, 2013
(See Stanev's talk)

Larger S125:
Higher energy/Deeper Xmax
Larger dE/dx:
More muons/Heavier primary

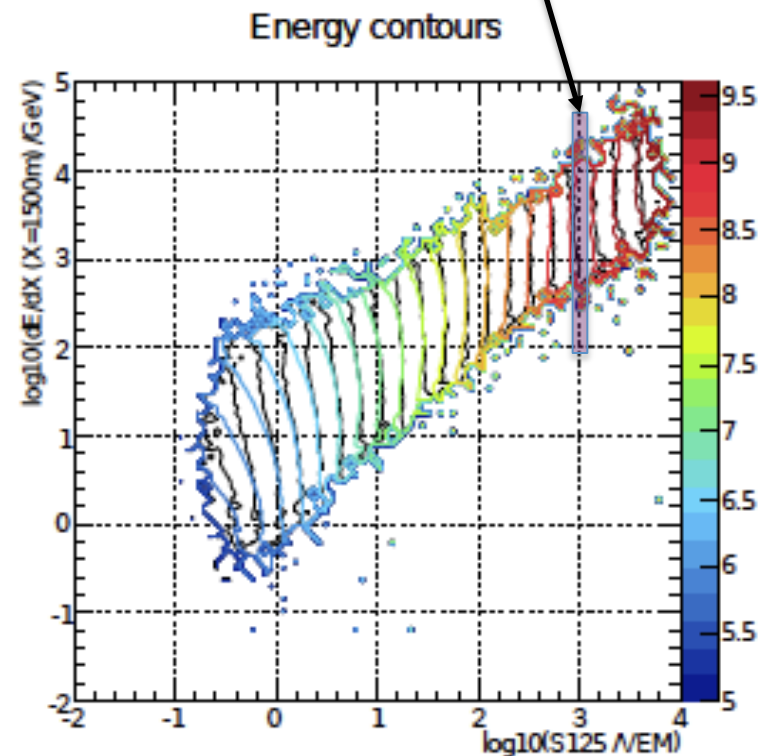
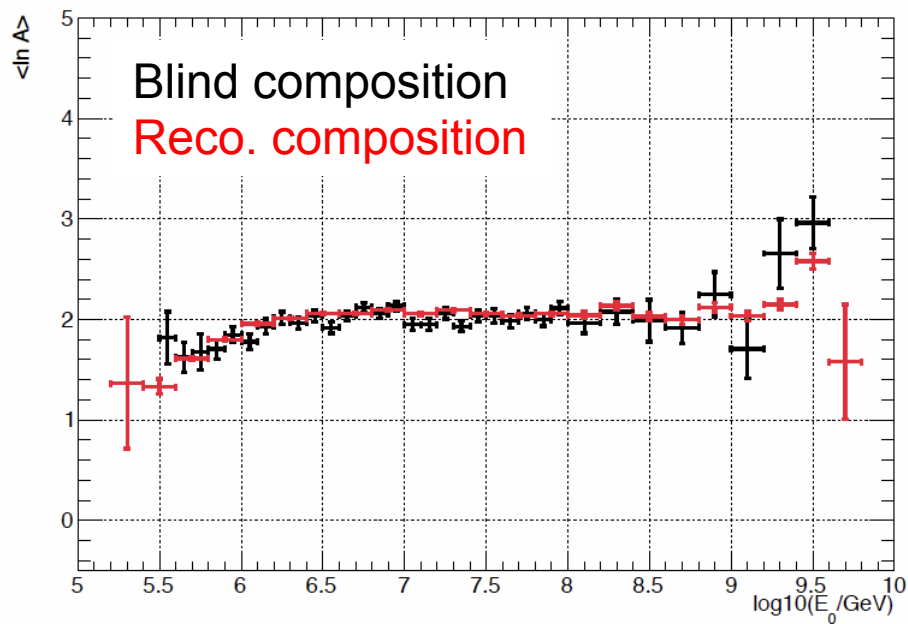
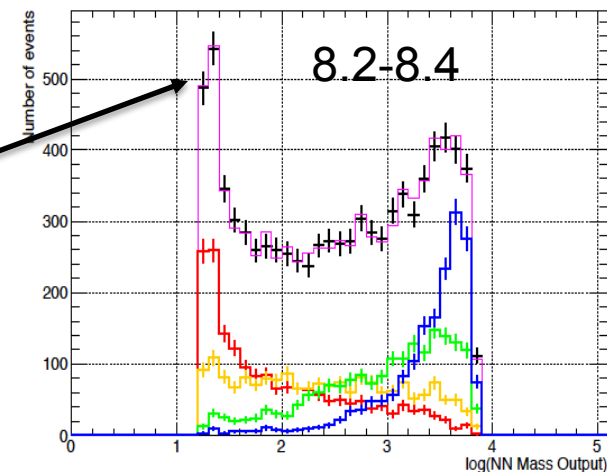
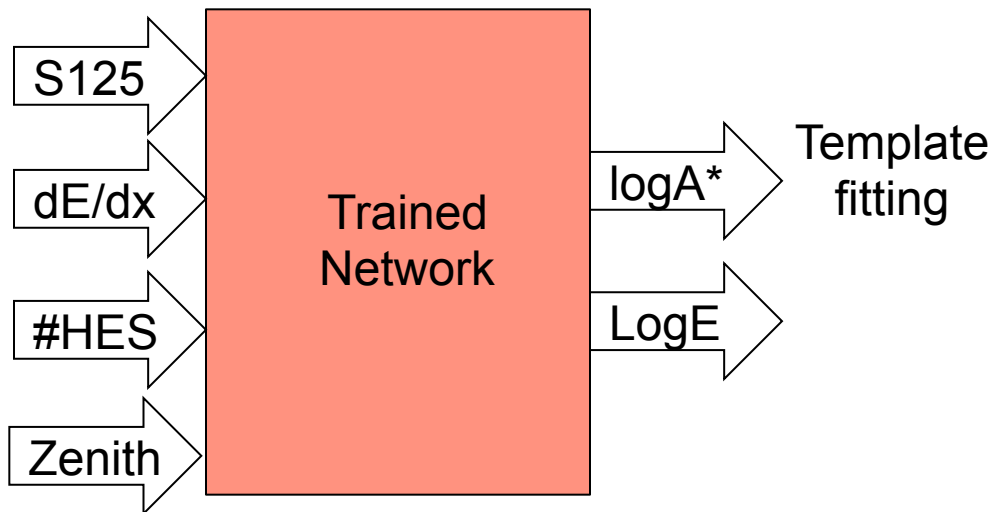
Avg. muon energy loss

Muon stochastic loss

Run 116545 event 58761981



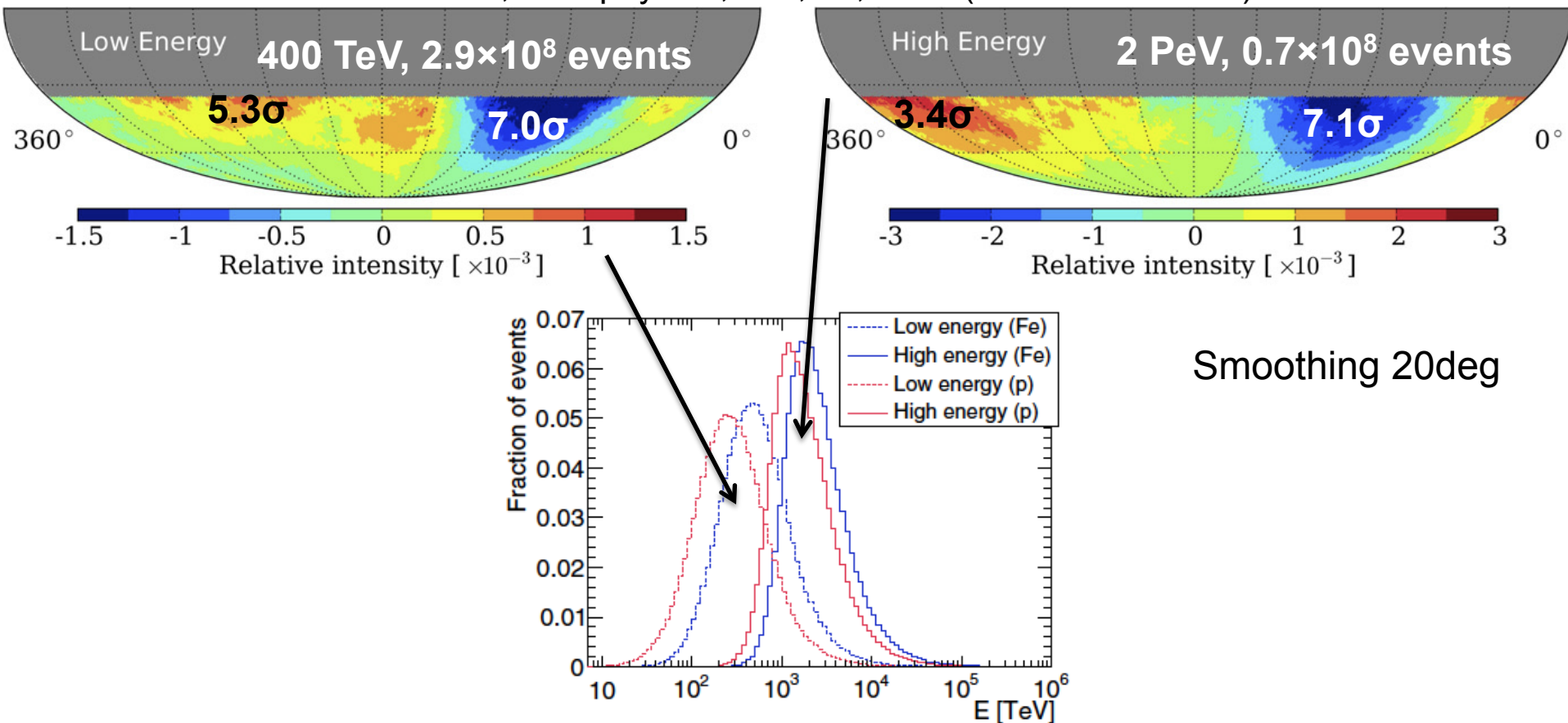
Neural Network Approach



Results not formally approved in time to be shown here

- We observe same trend observed with IC40 data up to $10^{8.8}$ GeV
- Systematics studies on-going
- **We can discriminate different masses**
- **We can reach 1 EeV**
- Full year results will be available at ICRC

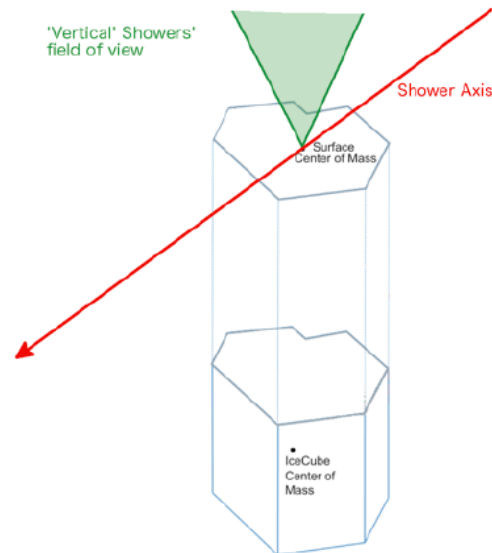
IceCube Coll., Astrophys. J., 765, 55, 2013 (See Desiati's talk)



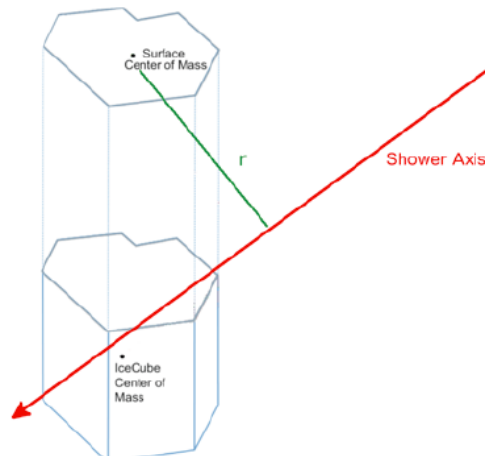
IceTop confirms a “dipole-like” large scale anisotropy, inconsistent with CG effect

- IceTop spectrum:
 - Unprecedented precision ($0.05 \log(E/\text{GeV})$; 3PeV-0.1EeV)
 - 4 power laws identified in PeV – EeV
 - Beyond systematics we observe significant hardening (-2.91) above 18 PeV and steepening (-3.37) above 130 PeV
- IceTop/IceCube composition and spectrum:
 - $\langle \log A \rangle$ will be measured up to 1 EeV
 - Different masses will be discriminated
- Anisotropy:
 - Significant dipole-like structure confirmed in Southern Sky



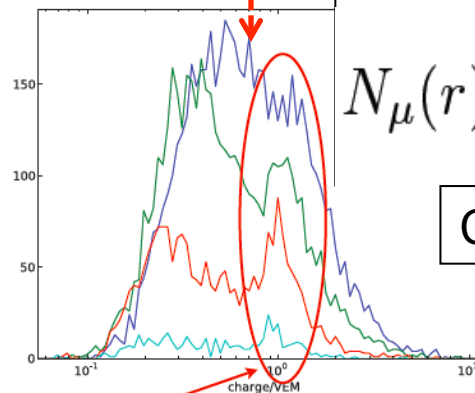
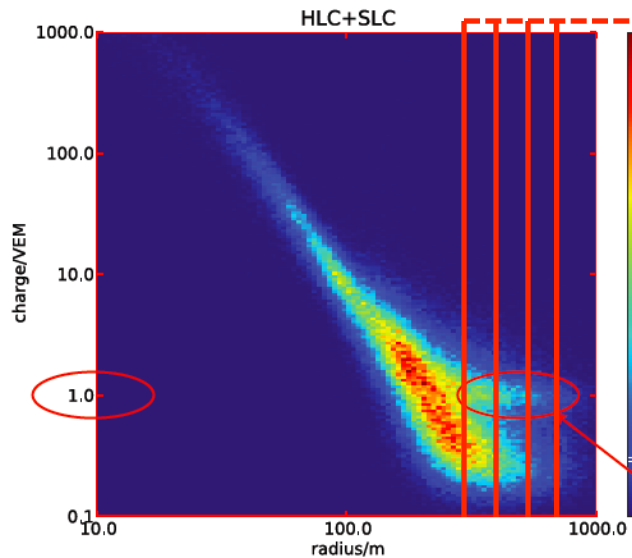


On-axis
+100% exposure (<60°)



Off-axis (<60°)
poorer reco., SLC-dominated

Increase sensitivity to
muon content of showers
Reach out for GC

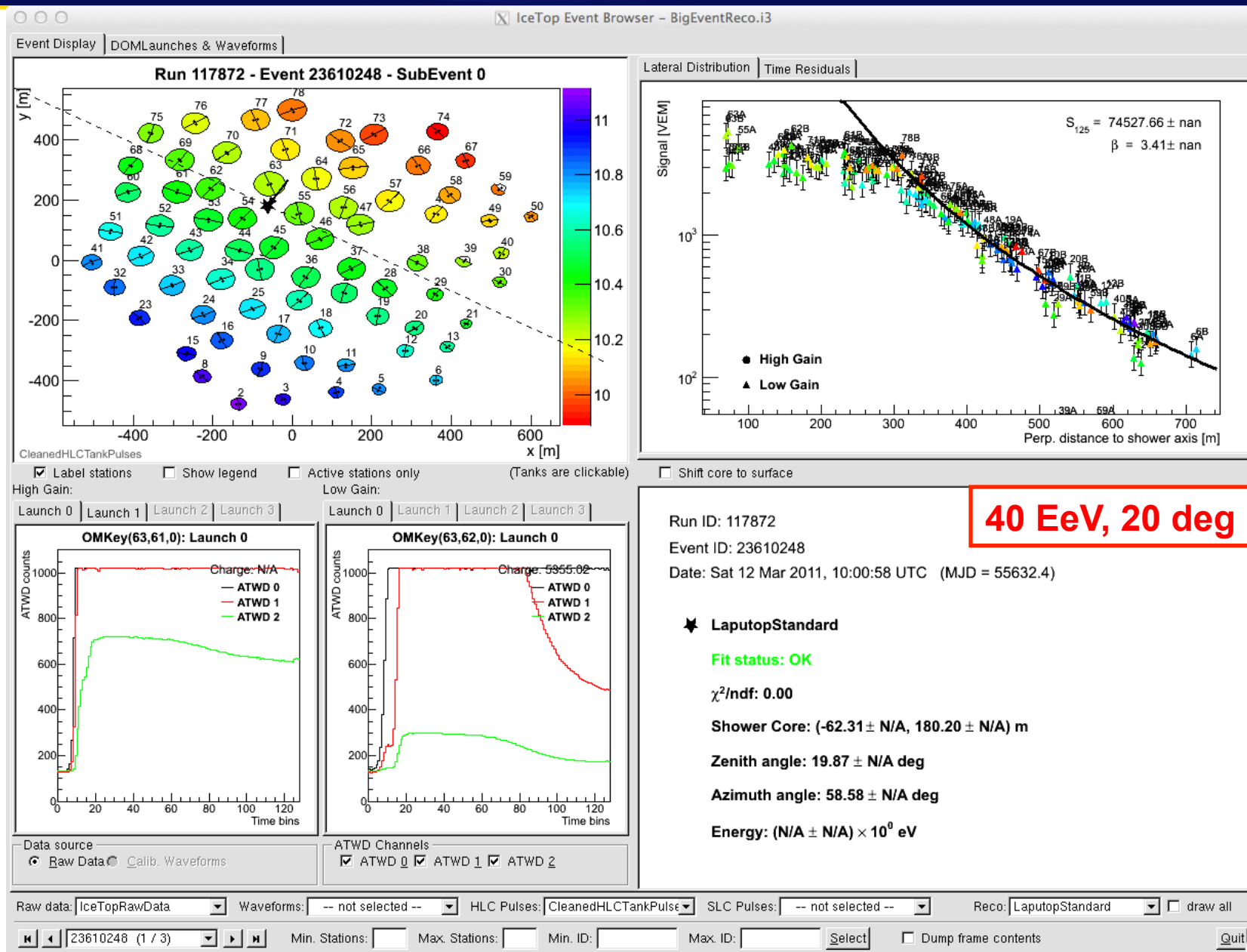


Muon LDF

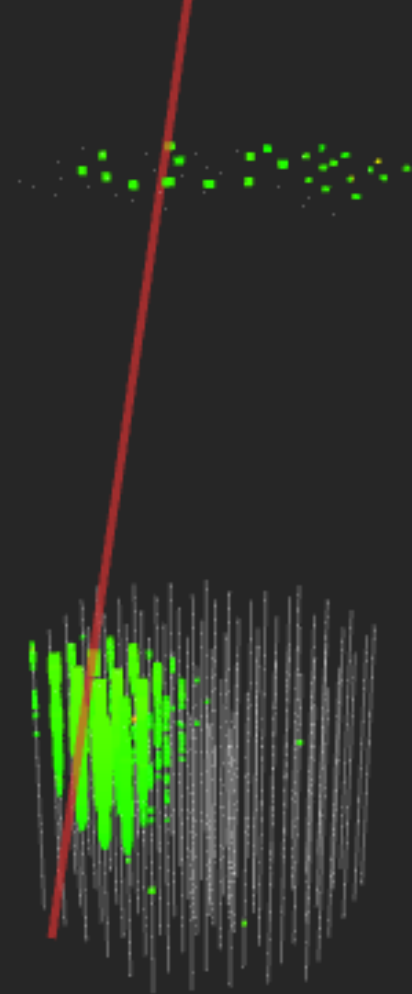
$$N_{\mu}(r) = A r^{-0.75} \left(1 + \frac{r}{320\text{m}}\right)^{-\gamma}$$

Compare with Greisen prediction

Muons



```
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  SubEventID : 0  
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```



```
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