PTOLEMY: A Proposal for Thermal Relic Detection of Massive Neutrinos and Directional Detection of MeV Dark Matter

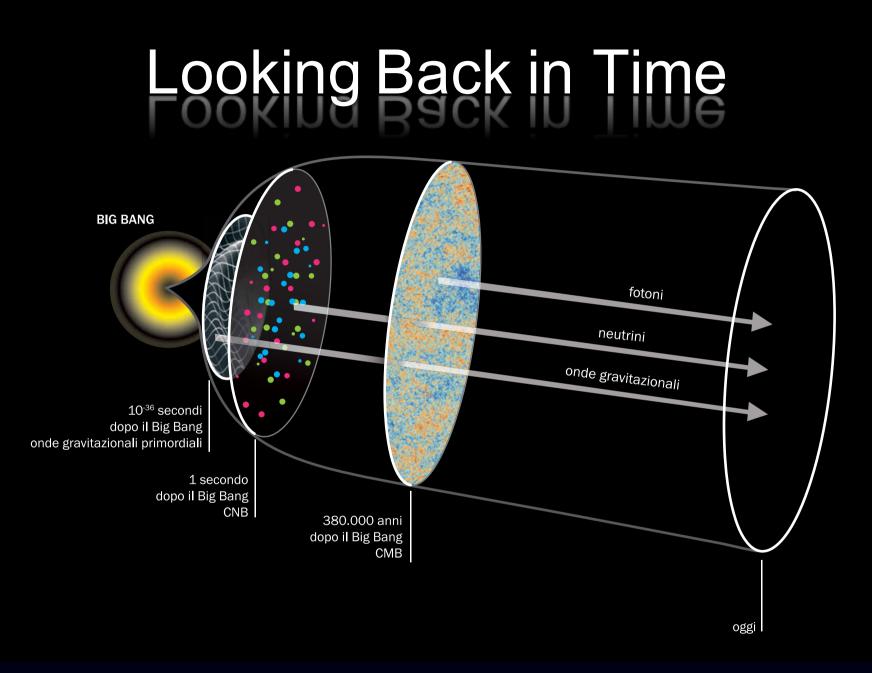
Alfredo Cocco, <u>Chris Tully</u>, Marcello Messina for the PTOLEMY Collaboration

LNGS SCIENTIFIC COMMITTEE OPEN SESSION 26 MARCH 2018

PTOLEMY: A Proposal for Thermal Relic Detection of Massive Neutrinos and Directional Detection of MeV Dark Matter

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Y. Raitses²⁵, N. Rossi¹⁰, F. Zhao²⁶, K.M. Zurek^{21,22}

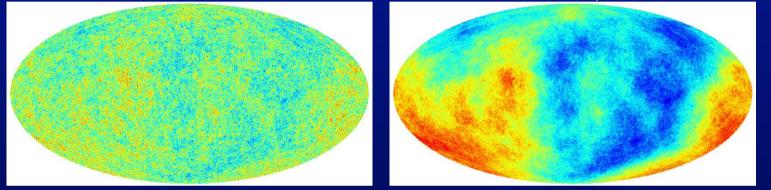
¹Racah Institute of Physics, Hebrew University of Jerusalem, Jerusalem, Israel ²INFN Laboratori Nazionali del Gran Sasso, L'Aquila, Italy ³Gran Sasso Science Institute (GSSI), L'Aquila, Italy ⁴INFN Laboratori Nazionali di Frascati, Frascati, Italy ⁵Università degli Studi di Genova e INFN Sezione di Genova, Genova, Italy ⁶Università degli Studi di Milano-Bicocca e INFN Sezione di Milano-Bicocca, Milano, Italy ⁷INFN Sezione di Napoli, Napoli, Italy ⁸Università degli Studi di Napoli Federico II, Napoli, Italy ⁹Università degli Studi di Pisa e INFN Sezione di Pisa, Pisa, Italy ¹⁰INFN Sezione di Roma, Roma, Italy ¹¹Università degli Studi di Roma La Sapienza, Roma, Italy ¹²Istituto Nazionale di Ricerca Metrologica (INRiM), Torino, Italy ¹³Nationaal instituut voor subatomaire fysica (NIKHEF), Amsterdam, Netherlands ¹⁴Centro de Investigaciones Energéticas, Medioambientales y Tecnológicas (CIEMAT), Madrid, Spain ¹⁵Consejo Superior de Investigaciones Científicas (CSIC), Madrid, Spain ¹⁶Universidad Politécnica de Madrid, Madrid, Spain ¹⁷Instituto de Física Corpuscular (IFIC), Valencia, Spain ¹⁸Stockholm University, Stockholm, Sweden ¹⁹Uppsala University, Uppsala, Sweden ²⁰New York University Abu Dhabi, Abu Dhabi, UAE ²¹Lawrence Berkeley National Laboratory, University of California, Berkeley, CA, USA ²²Department of Physics, University of California, Berkeley, CA, USA ²³Argonne National Laboratory, Chicago, IL, USA ²⁴Kavli Institute for Cosmological Physics, University of Chicago, Chicago, IL, USA ²⁵Princeton Plasma Physics Laboratory, Princeton, NJ, USA ²⁶Department of Physics, Princeton University, Princeton, NJ, USA



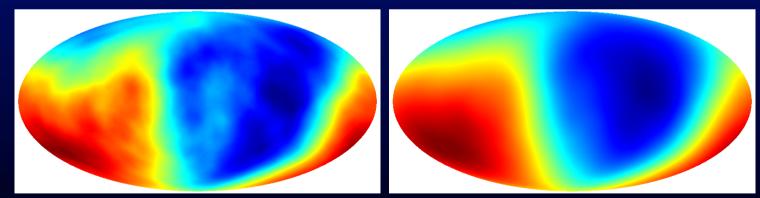
The Neutrino Sky

$m_v < 0.00001 \text{ eV}$

$m_v \sim 0.001 \text{ eV}$



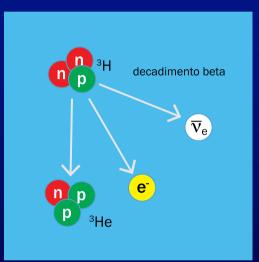
Hannestad, Brandbyge (2009)

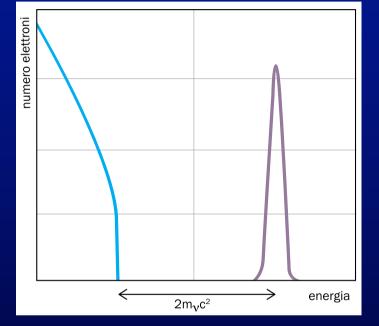


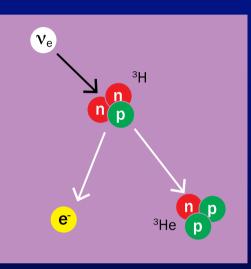
 $m_v \sim 0.01 \text{ eV}$

 $m_v \sim 0.1 \text{ eV}$

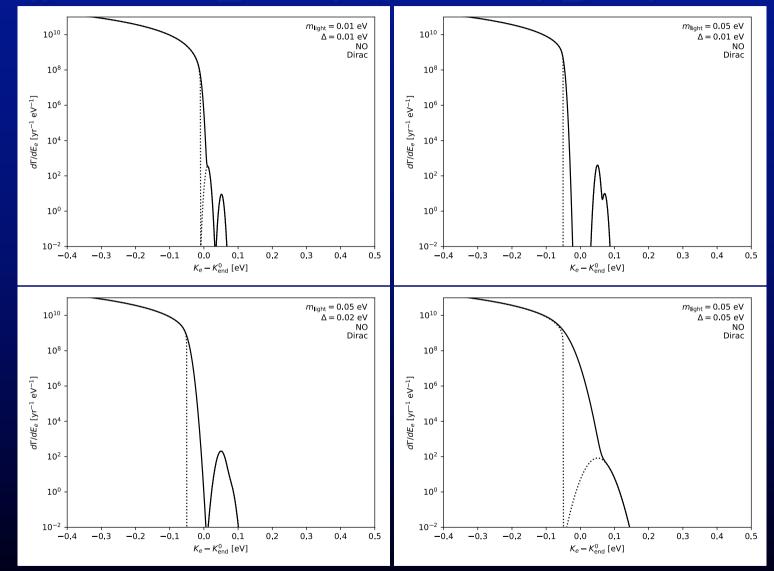
Detection Concept: Neutrino Capture







Challenges: Resolution and Backgrounds

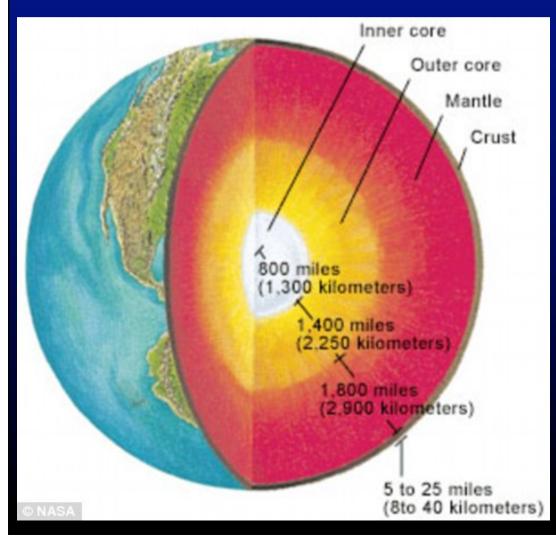


Normal Ordering

High Radio-Purity Carbon

Thumb radioactivity (1 per second \rightarrow 1 per 100 years)

Graphene fabrication from $CO_2 \rightarrow CH_3OH \rightarrow CH_4$



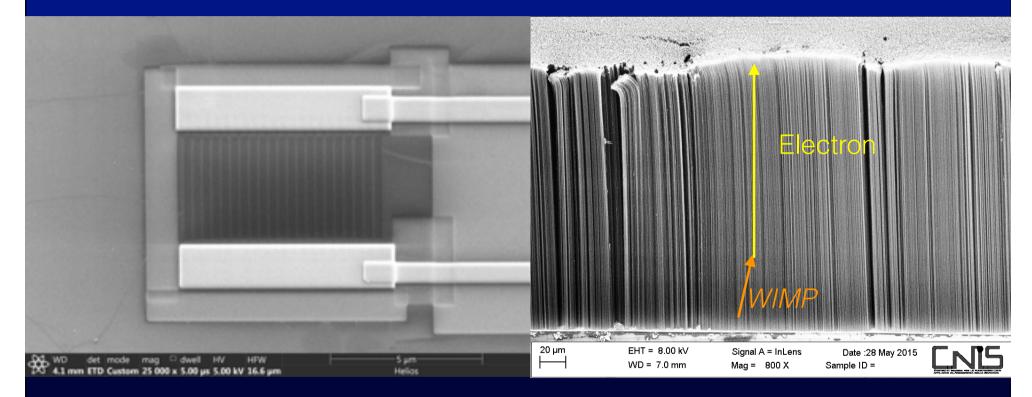


Kinder Morgan Doe Canyon CO₂ facility in southwestern Colorado

Graphene Targets: Two Concepts

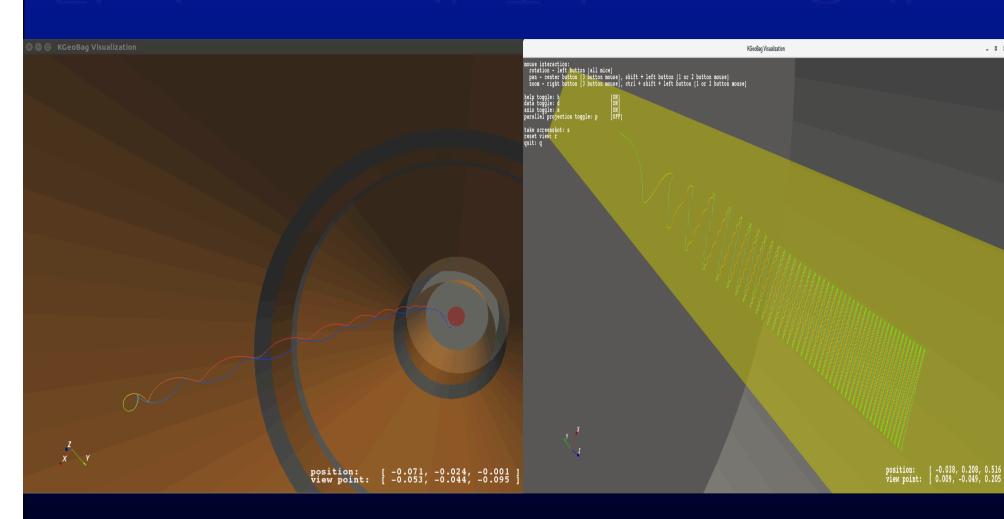
PTOLEMY-G³

PTOLEMY-CNT



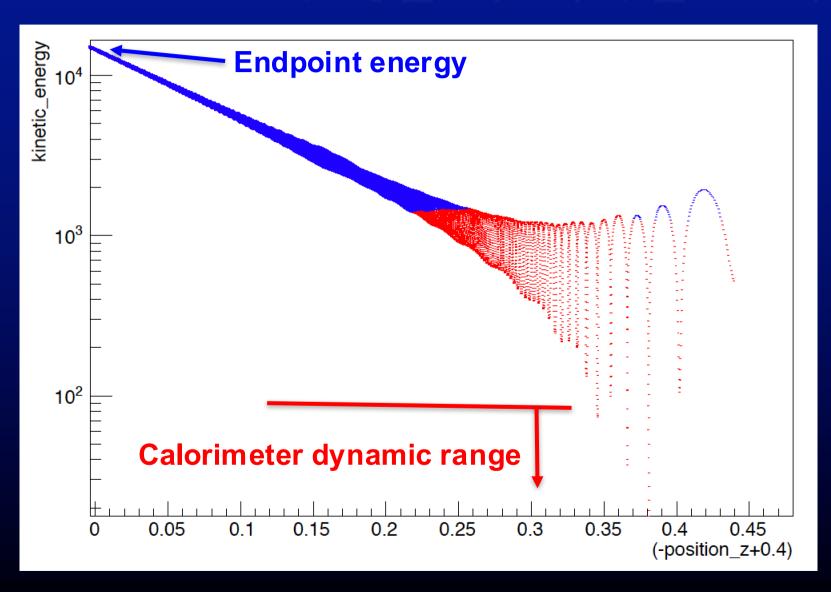
Self-instrumented with G-FETs Anisotropy of aligned CNTs

Electromagnetic Telescope Optics



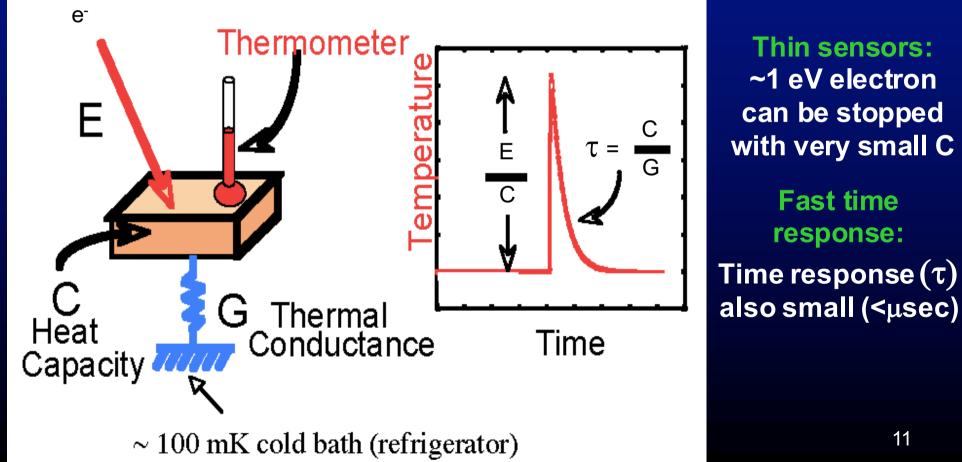
_ D X

Measurement of Endpoint Energies



Microcalorimetry

Optimize Transition-Edge Sensors for low energy ightarrowelectron calorimetry with an energy resolution sufficient to resolve the neutrino mass



Thin sensors: ~1 eV electron can be stopped with very small C

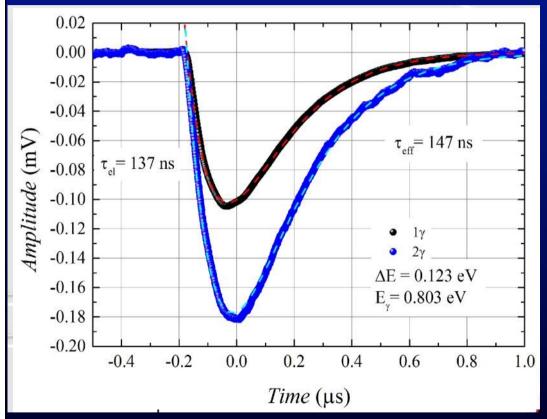
Fast time

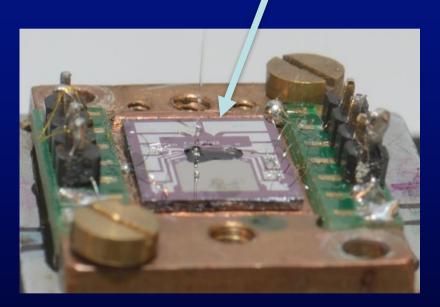
response:

Single Infrared Photon Detectors

Results from INRIM (Torino) -Istituto Nazionale di Ricerca Metrologica

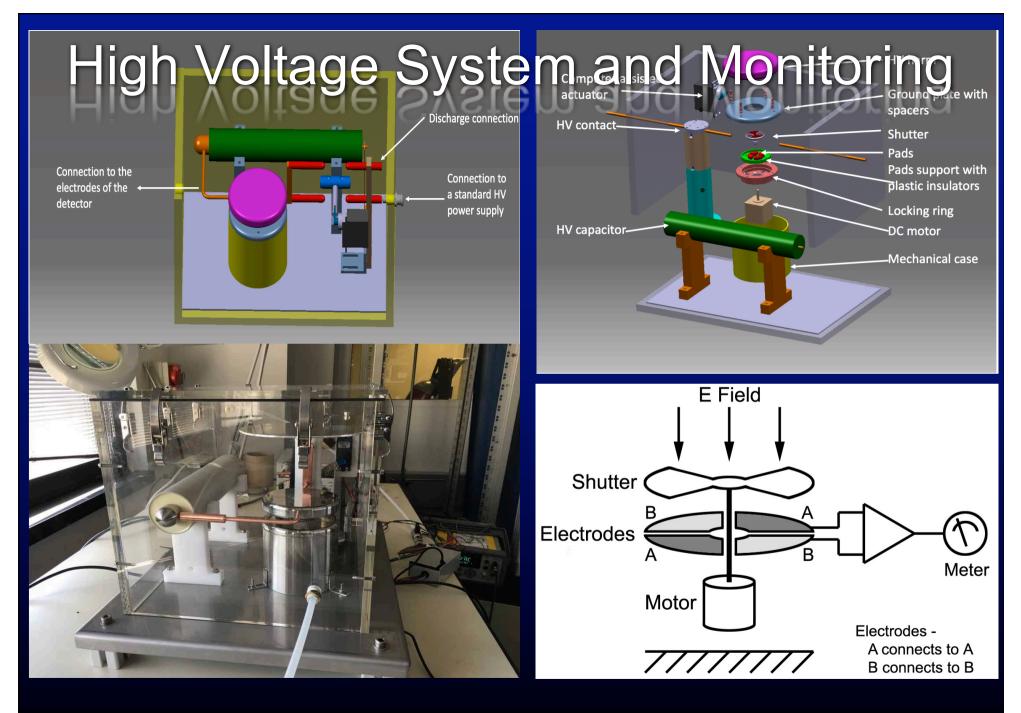
Transition-Edge Sensor





Infrared Photons E=0.8eV $\sigma_{\rm E}$ = 0.05 eV @ 300mK

→ Exceeding goals for energy resolution



PTOLEMY Prototype

PonTecorvo Observatory

for Light, Early-universe, Massive-neutrino Yield

 \rightarrow

R&D Prototype @ PU (June 7, 2017)

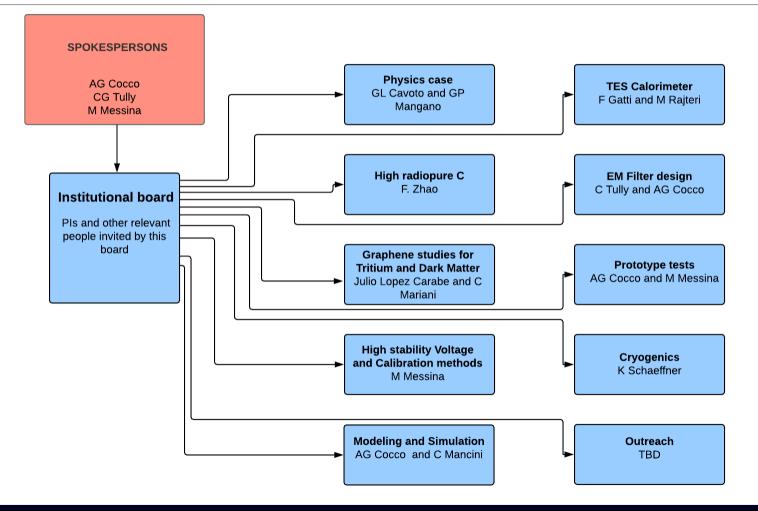
(Princeton T

Supported by: The Simons Foundation The John Templeton Foundation

PTOLEMY Working Groups

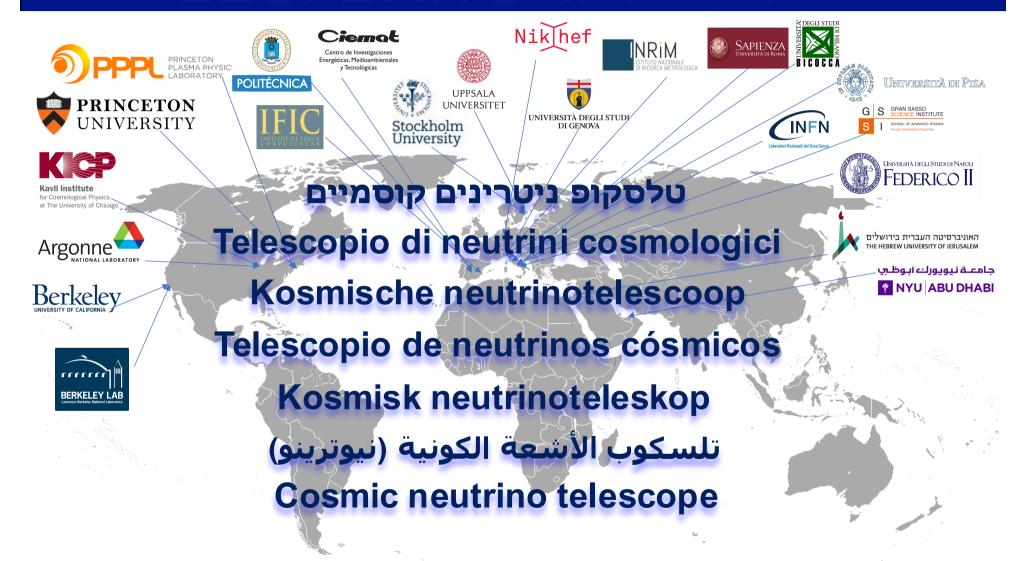
PTOLEMY ORGANIZATION CHART

| March 17, 2018



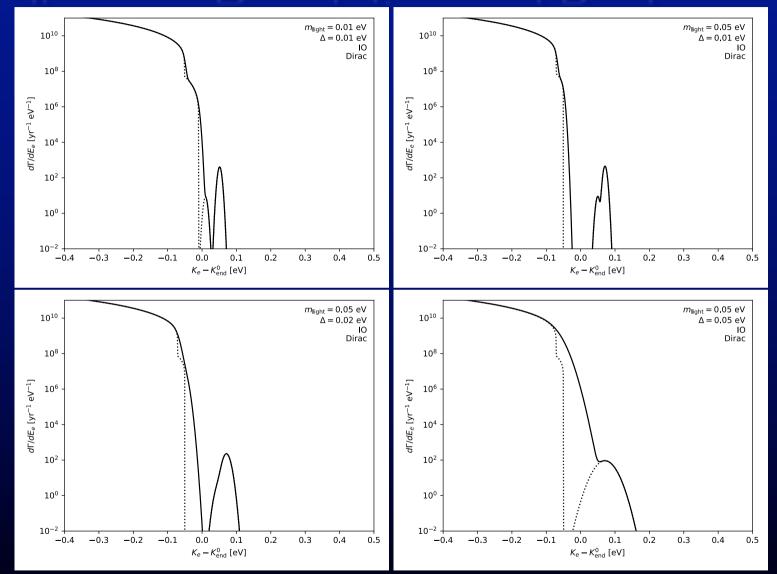
15

PTOLEMY Collaboration



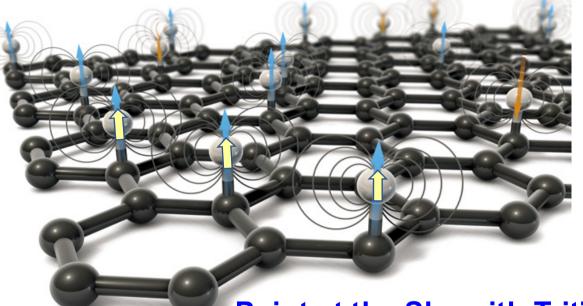
Commentary

Challenges: Resolution and Backgrounds



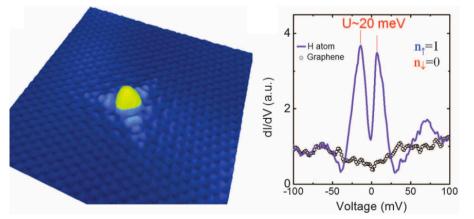
Inverted Ordering

Polarized Tritium Target



Lisanti, Safdi, CGT, 2014.

Point at the Sky with Tritium Nuclear Spin 👔



Hydrogen doping on graphene reveals magnetism

Gonzalez-Herrero, H. *et al.* Atomic-scale control of graphene magnetism by using hydrogen atoms. *Science (80)*. **352**, 437–441 (2016).

Polarized ³H Decay

$$\frac{d^5\omega}{dE_e d\Omega_e d\Omega_\nu} = \frac{G_F^2}{(2\pi)^5} p_e E_e (\Delta m - E_e)^2 \xi [1 + a\boldsymbol{\beta} \cdot \hat{\boldsymbol{\nu}} + \hat{\mathbf{P}} \cdot (A\boldsymbol{\beta} + B\hat{\boldsymbol{\nu}})] , \qquad (1)$$

where G_F is the Fermi constant, Δm is the difference between the ³H and ³He mass, p_e (E_e) is the electron impulse (energy), β (ν) is the electron (neutrino) three-velocity, and $\hat{\mathbf{P}}$ is the ³H polarization versor. The quantities ξ , a, A and B contain the nuclear matrix elements, and and can be written in terms of the "standard" Fermi (F) and Gamow-Teller (GT) matrix elements as

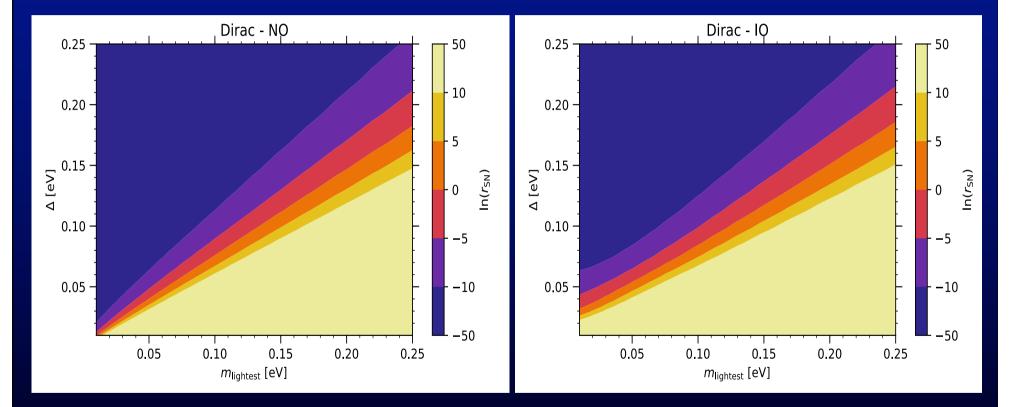
$$\xi = |F|^2 + g_A^2 |GT|^2 , \qquad (2)$$

$$a\xi = |F|^2 - \frac{g_A^2}{3}|GT|^2 , \qquad (3)$$

$$A\xi = -\frac{2}{3}g_A^2|GT|^2 + \frac{2}{\sqrt{3}}|GT||F| , \qquad (4)$$

$$B\xi = +\frac{2}{3}g_A^2|GT|^2 + \frac{2}{\sqrt{3}}|GT||F| .$$
(5)

CNB Signal-to-Noise



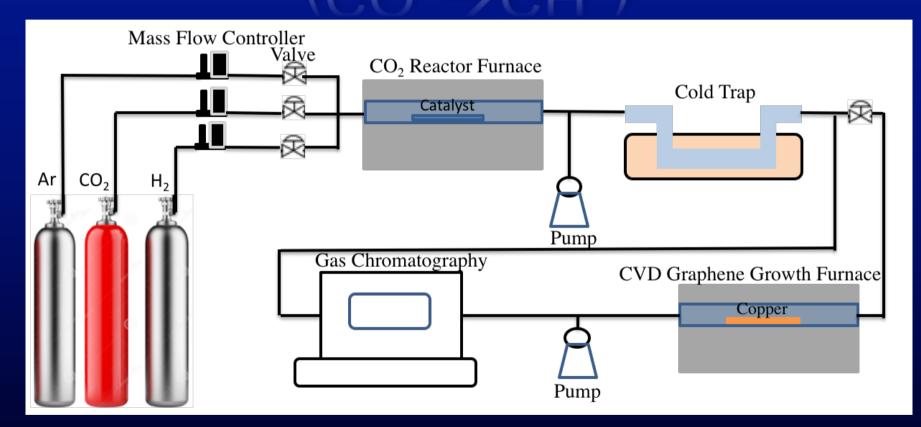
Normal Ordering

Inverted Ordering

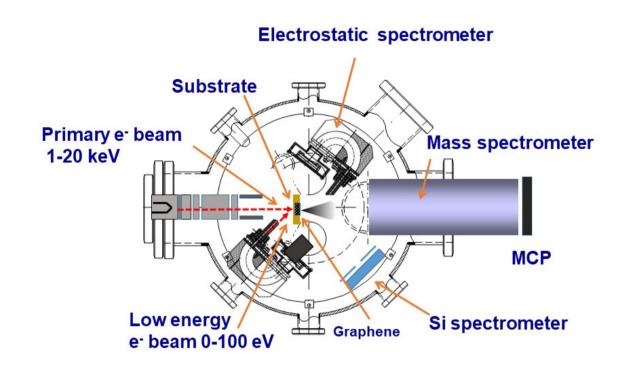
Cryogenic System of CRESST



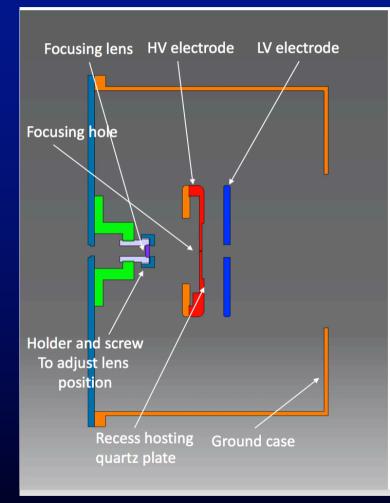
High Radio-Pure ¹²C ($CO_2 \rightarrow CH_4$)

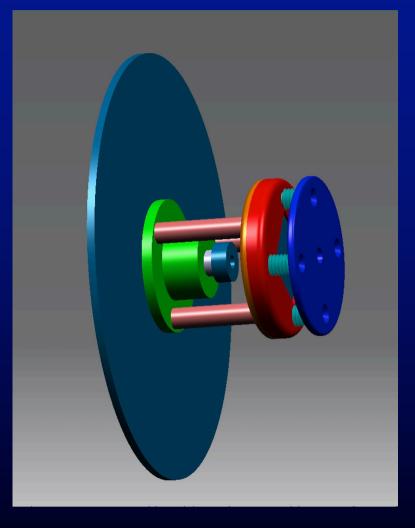


Electron-Graphene Interaction Chamber



Electron Gun

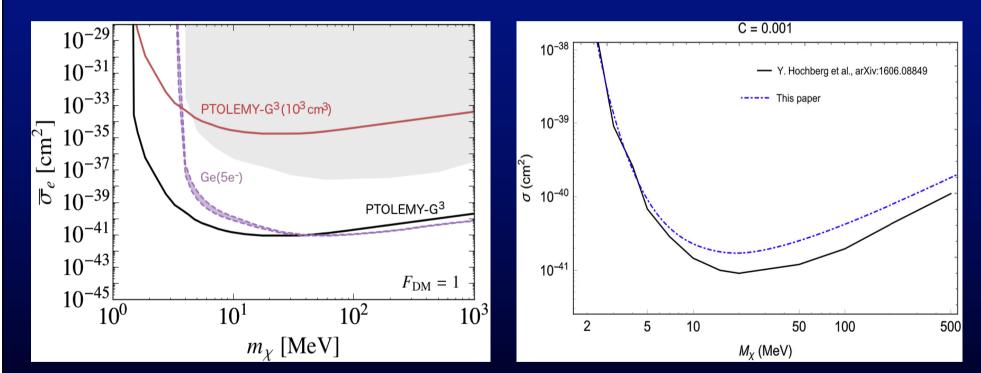




Direction Detection MeV Dark Matter Searches

PTOLEMY-G³

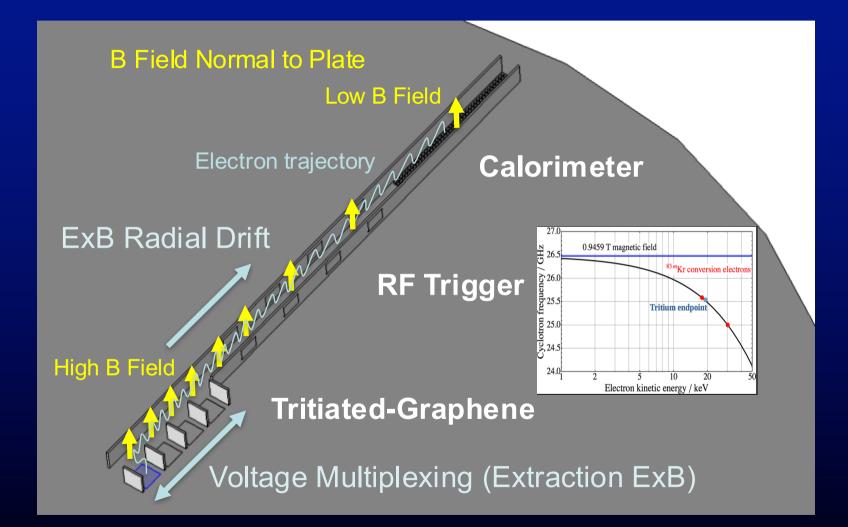
PTOLEMY-CNT



Anisotropy of aligned CNTs

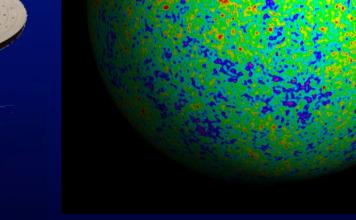
Self-instrumented with G-FETs

Scalable Underground Design



Celestial Globes



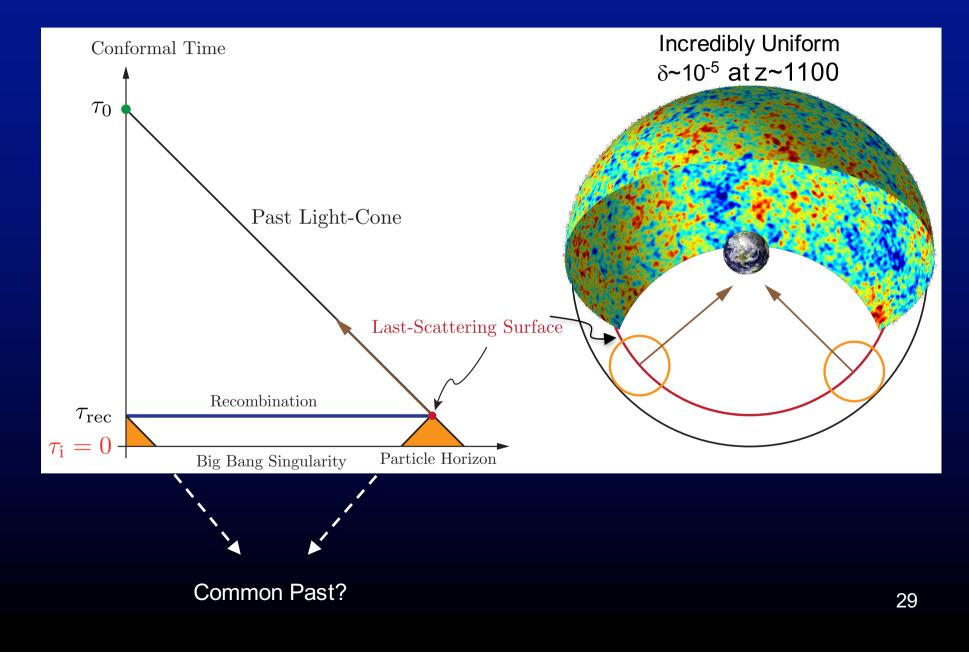


Adiabatic Density Anisotropies δ ~10⁻⁵ at z~1100

Johann Schöner, c.1534

WMAP, c.2009

Big Bang Cosmology



Our Home in the Universe

Ostriker, J.P. & Peebles, P.J.E. 1973, Ap. J. 186, 467.

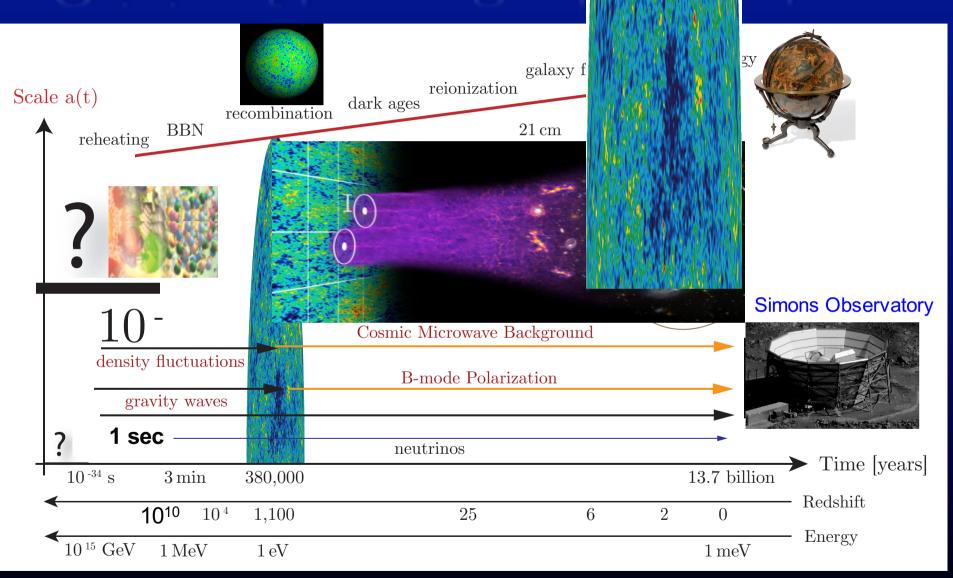
To study the stability of flattened galaxies, we have followed the evolution of simulated galaxies containing 150 to 500 mass points. Models which begin with characteristics similar to the disk of our Galaxy (except for increased velocity dispersion and thickness to assure local stability) were found to be rapidly and grossly unstable to barlike modes. These modes cause an increase in random kinetic energy, with approximate stability being reached when the ratio of kinetic energy of rotation to total gravitational energy, designated t, is reduced to the value of 0.14 ± 0.02 . Parameter studies indicate that the result probably is not due to inadequacies of the numerical N-body simulation method. A survey of the literature shows that a critical value for limiting stability $t \simeq 0.14$ has been found by a variety of methods.

-25.000 light years

Models with added spherical (halo) component are more stable. It appears that halo-to-disk mass ratios of 1 to $2\frac{1}{2}$, and an initial value of $t \simeq 0.14 \pm 0.03$, are required for stability. If our Galaxy (and other spirals) do not have a substantial unobserved mass in a hot disk component, then apparently the halo (spherical) mass *interior* to the disk must be comparable to the disk mass. Thus normalized, the halo masses of our Galaxy and of other spiral galaxies *exterior* to the observed disks may be extremely large.

Subject headings: galactic structure --- stellar dynamics

Origin of Large Scal

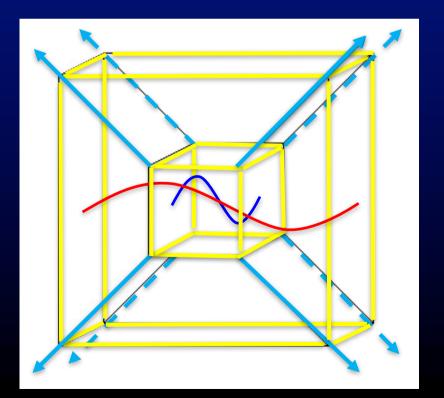


Baumann (TASI 2012) icture

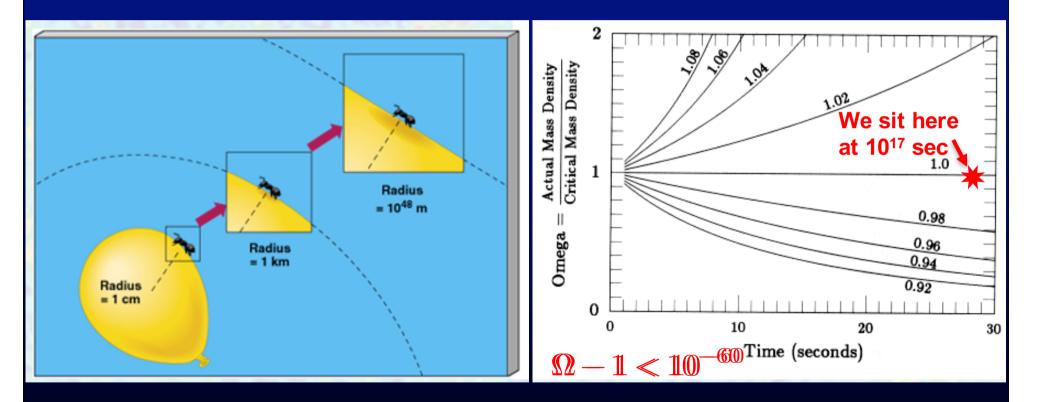
Expanding Universe

Expansion rate of the Universe: \dot{a} \rightarrow Kinetic Energy $\propto \dot{a}^2$

Energy density of the Universe: \rightarrow Potential Energy $\propto \rho$ $\rho_{matter} \propto 1/a^3$ sum from all matter, radiation and vacuum energy $\rho_{radiation} \propto 1/a^4$ $\rho_{\Lambda} \propto constant$

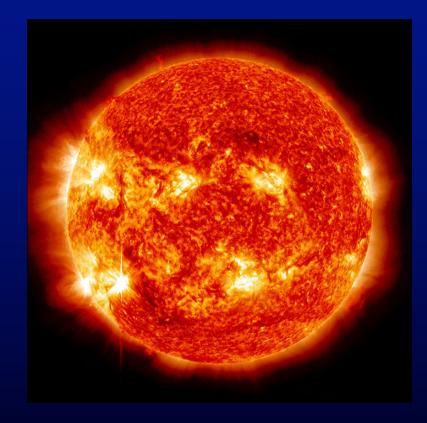


Balance of Kinetic and Potential Energy (ratio) $\Omega = 1.000(3)^{(m-1)}$ (known to better than 0.3%)



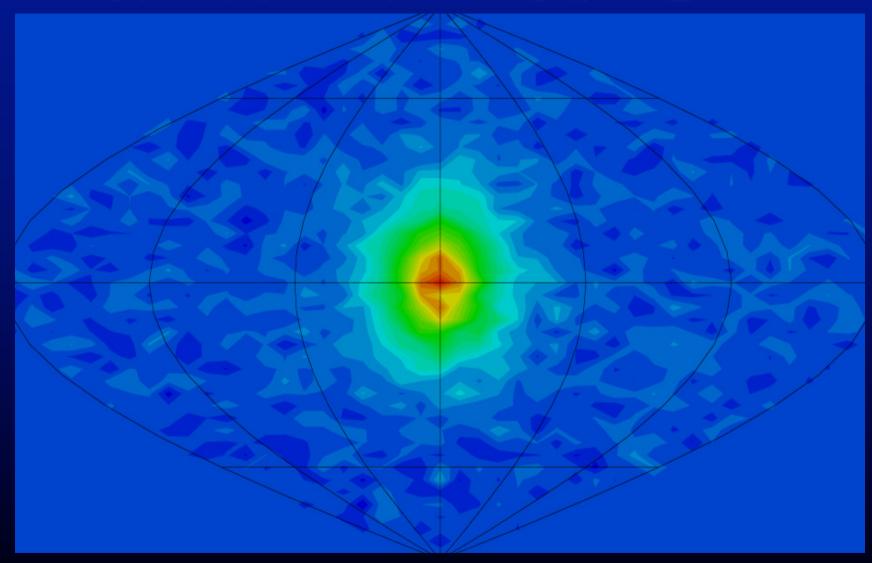
Expansion in a dark energy (cosmological constant) dominated Universe

View of the Sun

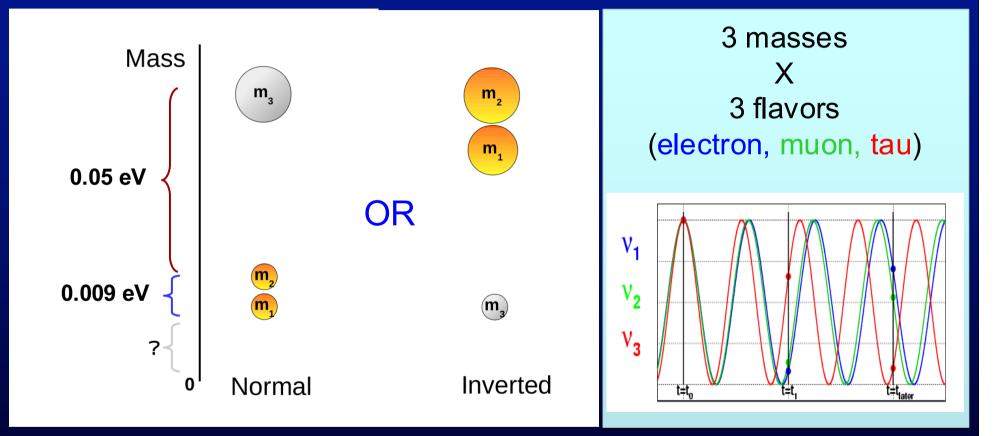


~8 min. away

Neutrino view of the Sun



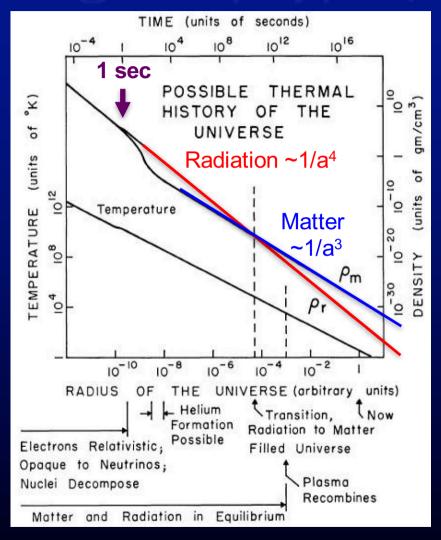
Neutrino Masses from Oscillations



The absolute neutrino masses are not known.

It's not known at this time whether neutrinos masses are "Normal" or "Inverted".

Cosmic Neutrino Background



Dicke, Peebles, Roll, Wilkinson (1965)

Number density: $n_{y} = 112/cm^{3}$ **Temperature:** T_v~1.95K Time of decoupling: $t_v \sim 1$ second neutron/proton ratio @start of nucleosynthesis Velocity distribution: $<v_{v}> ~ T_{v}/m_{v}$

Non-linear distortions Villaescusa-Navarro et al (2013) 37

Cosmic Elements

3 element theory

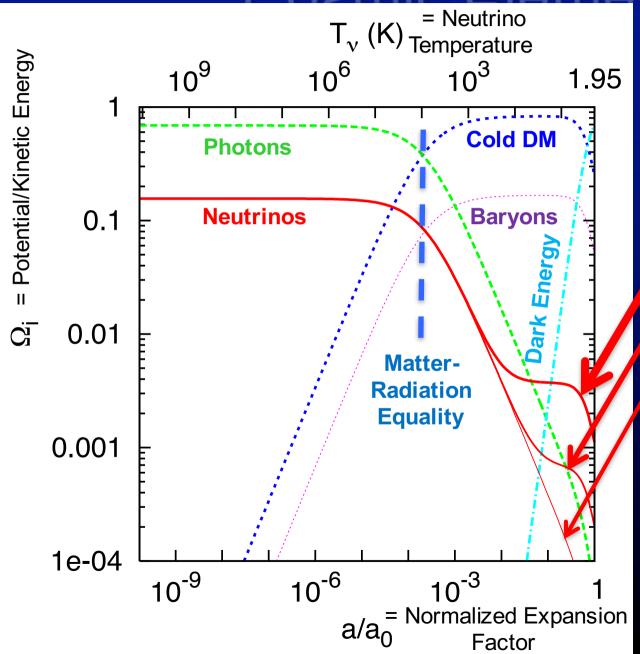
γ (photons)
ν (neutrinos)
p,n (baryons)

4 element theory

 χ (cold dark matter)

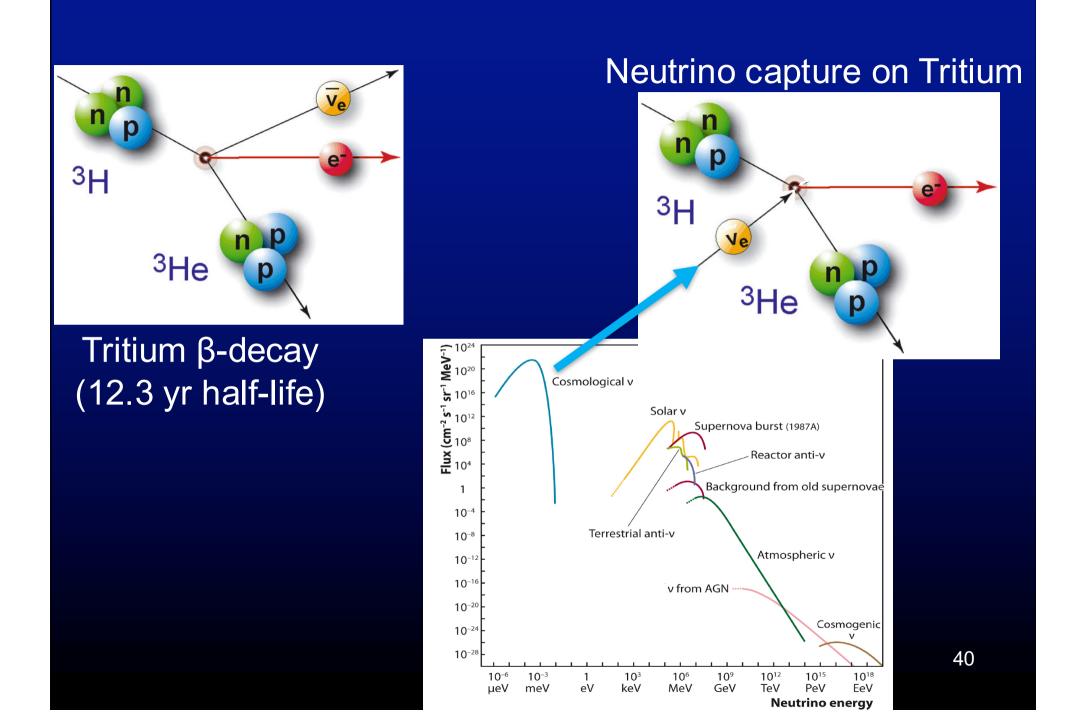
5 element theory (+Aether/Void) Λ (dark energy)

Cosmic Elements



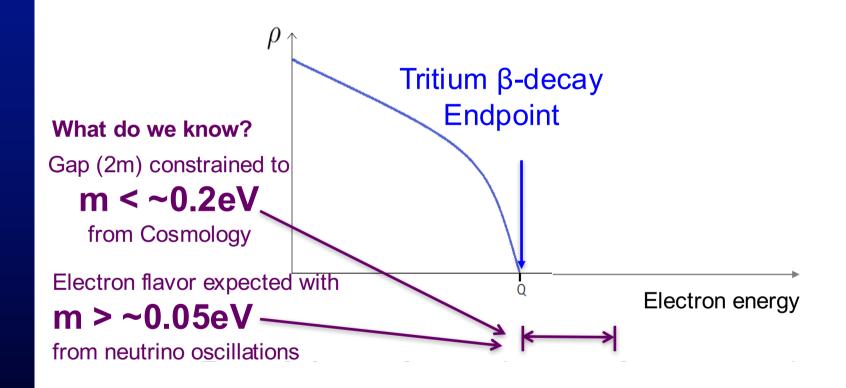
J. Lesgourgues

Individual neutrino contributions assuming Normal Hierarchy and $m_3 = 0.05 \text{ eV},$ $m_2 = 0.009 \text{ eV},$ $m_1 = 0$



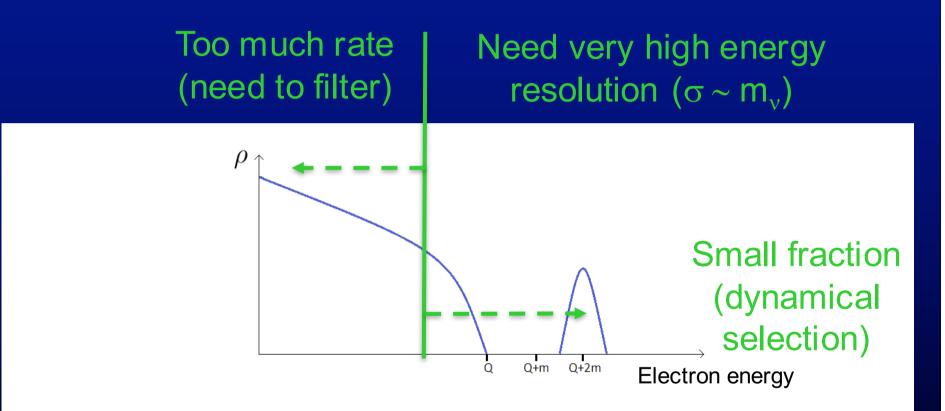
Relic Neutrino Detection

 Basic concepts for relic neutrino detection were laid out in a paper by Steven Weinberg in 1962 [Phys. Rev. 128:3, 1457]



Tritium and other isotopes studied for relic neutrino capture in this paper: JCAP 0706 (2007)015, hep-ph/0703075 by Cocco, Mangano, Messina

Experimental Perspective



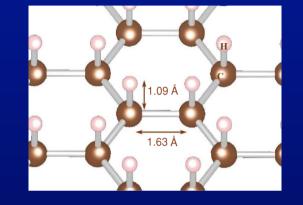
Emitted electron density of states vs kinetic energy for neutrino capture on beta decaying nuclei. The spike at Q + 2m is the CNB signal

Graphene (2-D Material)



Molecular Broadening

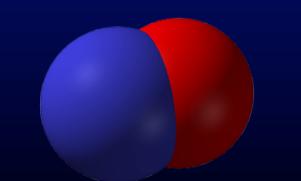
$T-T \rightarrow (T-He^3)^*$

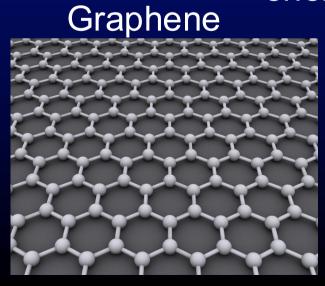


<3eV binding energy

~3eV He³ recoil at endpoint

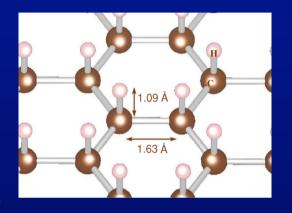


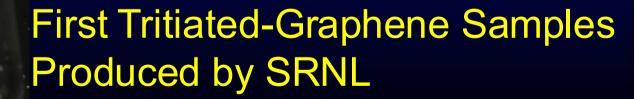




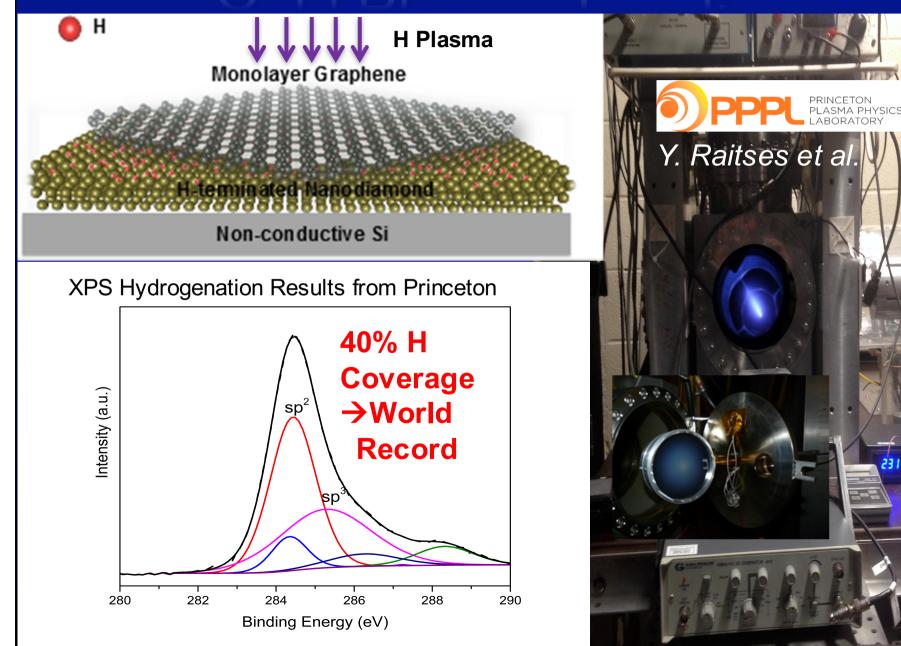
Molecular Broadening

SRNL



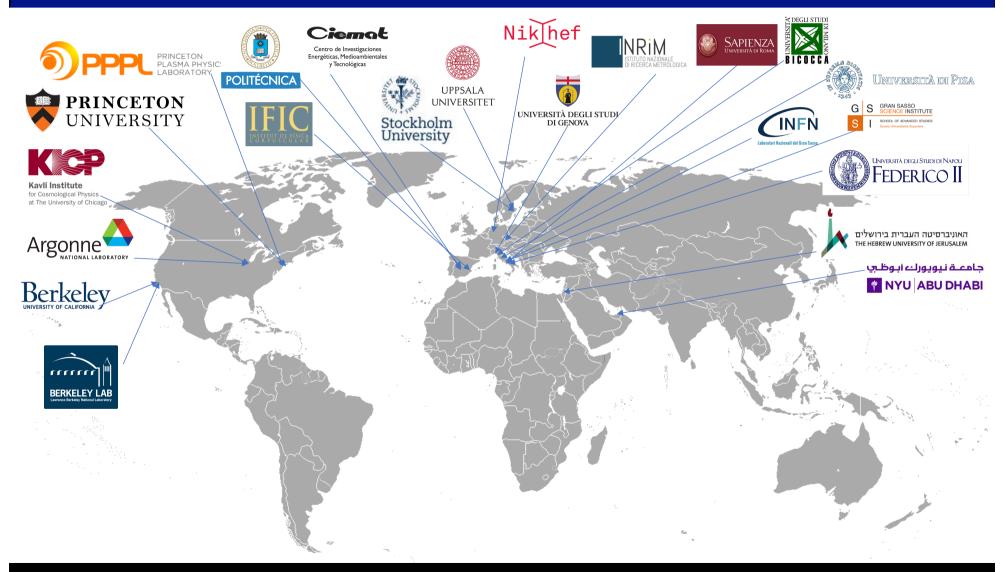


Cold Plasma Loading



PTOLEMY World-Wide Collaboration

2015 Targeted Grant Award from the SIMONS FOUNDATION

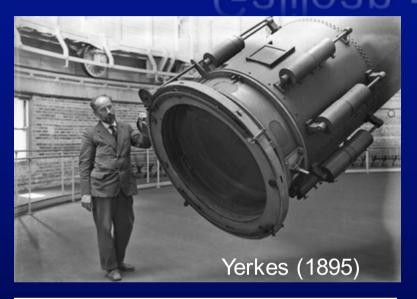


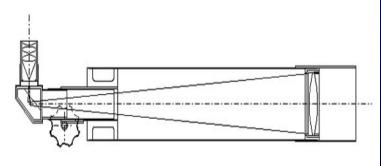
Underground Environment

Gran Sasso National Laboratory, Italy

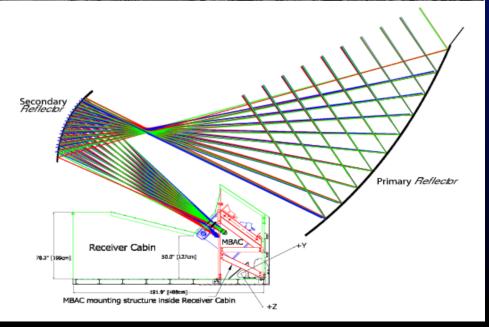
PTOLEMY kick-off meeting 11-12 December 2017 http://ptolemy.lngs.infn.it

Refractor \rightarrow Reflector Telescopes Galilean \rightarrow Newtonian







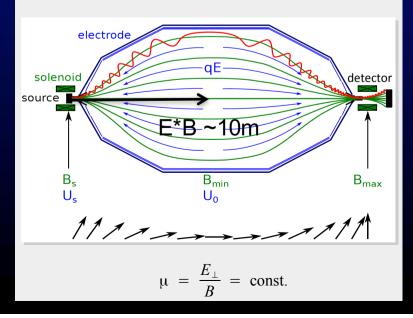


MAC-E "Telescope"

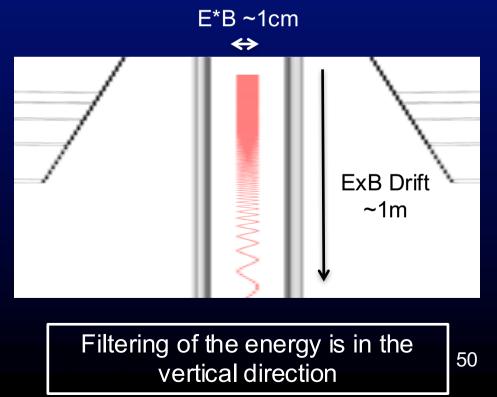


MAC-E filter technique

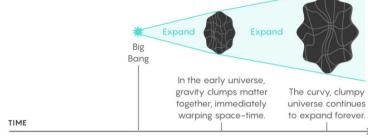
Magnetic Adiabatic Collimation with Electrostatic filter Picard et al., NIM B63 (1992) 345



PTOLEMY implements a "reflector" method that is four orders of magnitude more compact along the direction of the B field



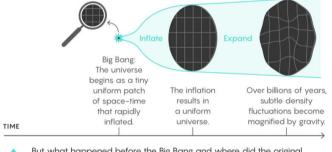
To explain why the universe was expanding, cosmologists began theorizing in the 1920s that a Big Bang event birthed the universe from an infinitely dense point 13.8 billion years ago.



But cosmologists observe a uniform early universe, not a clumpy crumpled one. Something was missing.

Cosmic Inflation

About 30 years ago, cosmologists proposed an updated Big Bang theory called "cosmic inflation" to explain our smooth, flat universe.



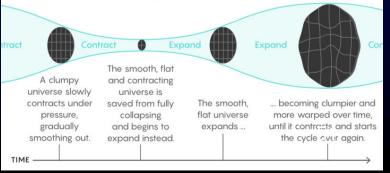
But what happened before the Big Bang and where did the original patch of space-time come from?



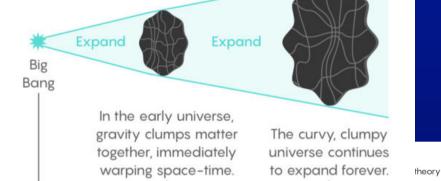
https://www.quantamagazine.org/ big-bounce-models-reignite-bigbang-debate-20180131

The Big Bounce

Recently, researchers have been taking a new look at the possibility of an expanding and contracting universe that could cycle forever.



To explain why the universe was expanding, cosmologists began theorizing in the 1920s that a Big Bang event birthed the universe from an infinitely dense point 13.8 billion years ago.



TIME

But cosmologists observe a uniform early universe, not a clumpy crumpled one. Something was missing.

TIME



The inflation Over billions of years, results in subtle density a uniform fluctuations become universe. magnified by gravity.

But what happened before the Big Bang and where did the original patch of space-time come from?

begins as a tiny

uniform patch

of space-time

that rapidly

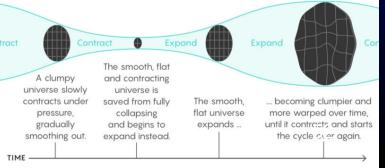
inflated.



https://www.quantamagazine.org/ big-bounce-models-reignite-bigbang-debate-20180131

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TIME

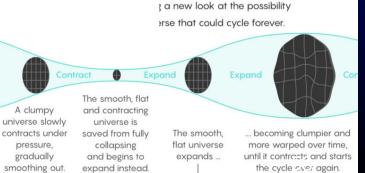
	Big Bang: The universe		Expand
TIME	begins as a tiny uniform patch of space-time that rapidly inflated.	The inflation results in a uniform universe.	Over billions of years, subtle density fluctuations become magnified by gravity.

But what happened before the Big Bang and where did the original

TIME



patch of space-time come from?

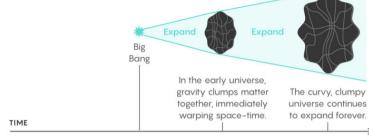


https://www.quantamagazine.org/ big-bounce-models-reignite-big-

💢 Quantamagazine

bang-debate-20180131

To explain why the universe was expanding, cosmologists began theorizing in the 1920s that a Big Bang event birthed the universe from an infinitely dense point 13.8 billion years ago.



But cosmologists observe a uniform early universe, not a clumpy crumpled one. Something was missing.

Cosmic Inflation

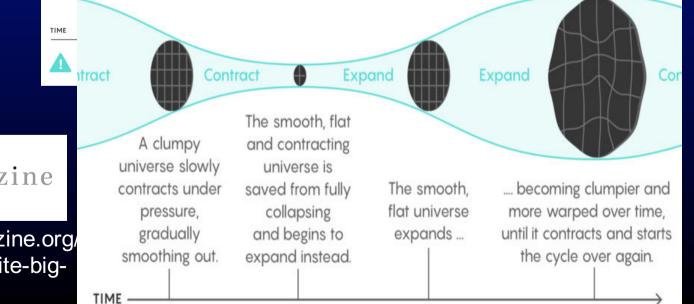
About 30 years ago, cosmologists proposed an updated Big Bang theory

called

The Big Bounce

Recently, researchers have been taking a new look at the possibility

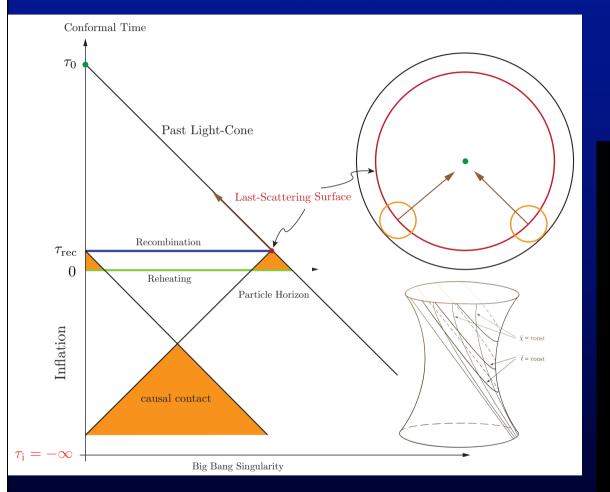
of an expanding and contracting universe that could cycle forever.



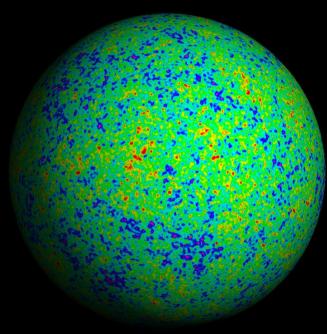


https://www.quantamagazine.org/ big-bounce-models-reignite-bigbang-debate-20180131

Big Bang Cosmology

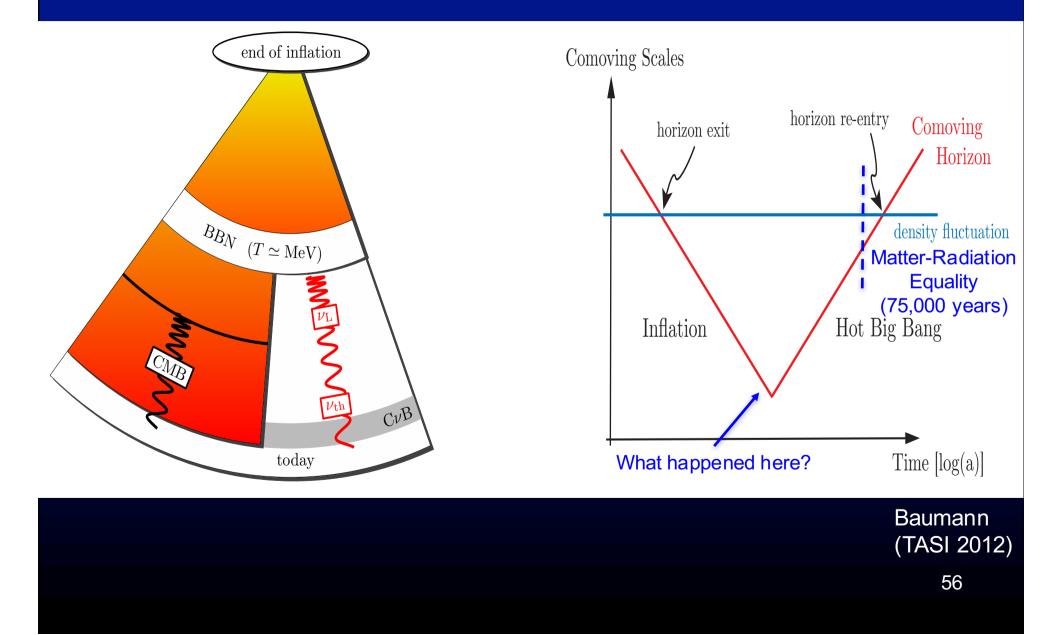


Adiabatic Density Anisotropies $\delta \sim 10^{-5}$ at $z \sim 1100$

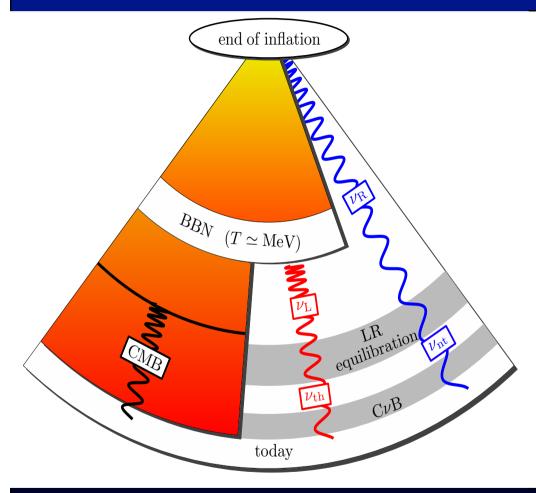


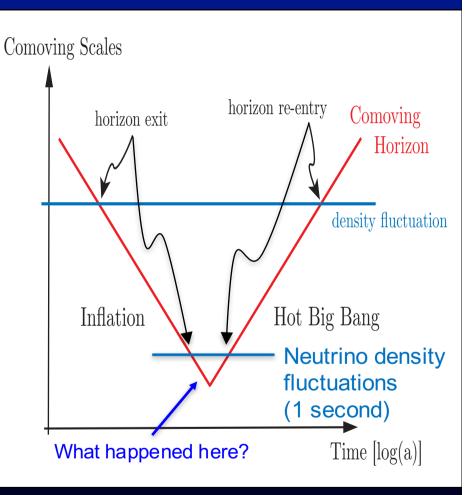
Where we think there is an initial τ_i =0 Big Bang Singularity is believed to be the "end" of an inflation period that slowly pulled out (>60 e-folds $a(\tau) \sim e^{H\tau}$) of a "de Sitter"-like spacetime

Inflation -> Hot Big Bang



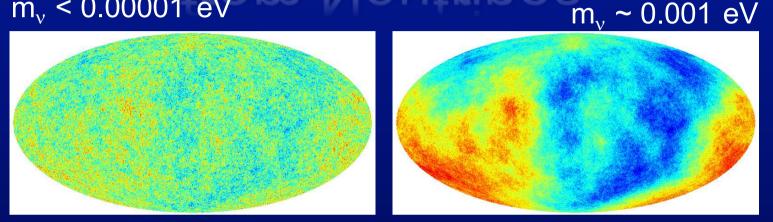
Inflation Hot Big Bang



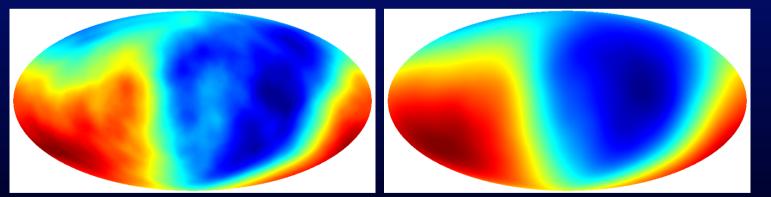


Ratz (Erice 2017)

The Future: Celestial Globes from Neutrinos Martin Rees@CMB50 m_v < 0.00001 eV m ~ 0.001 eV



Hannestad, Brandbyge (2009)



$m_v \sim 0.01 \text{ eV}$

2015 Targeted Grant Award from the SIMONS FOUNDATION

and additional support from the



 $m_v \sim 0.1 \text{ eV}$

John Templeton Foundation