

"Hot" electron currents in ultra-intense laser-solid interactions



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hzdr

HELMHOLTZ
ZENTRUM DRESDEN
ROSSENDORF



Laser particle acceleration at HZDR

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HZDR

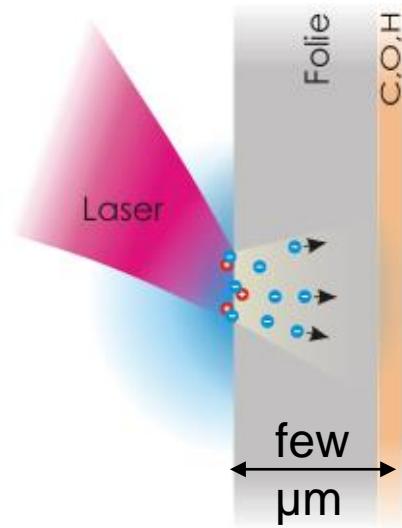
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Why are we interested in electron currents?

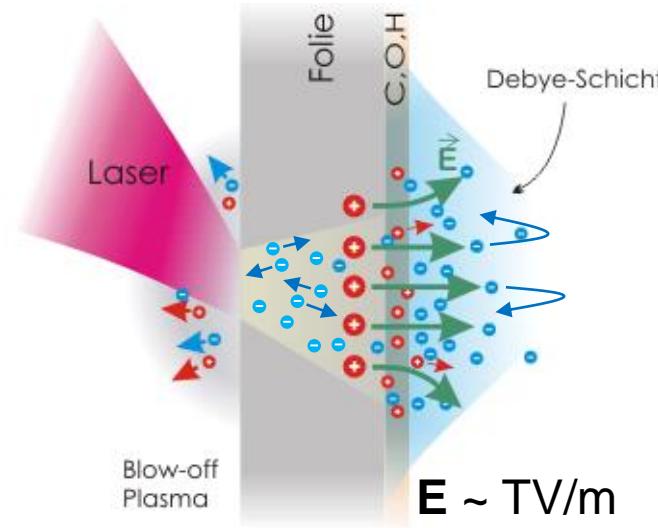
- self generated **magnetic fields** for guiding/divergence control
e.g. Leblanc, Sentoku PRE **89** 023109 (2014)
- TNSA
- diagnostics
- bulk heating, e.g. L. Huang et al. POP **20** 093109 (2013)

Idealized setup

absorption and electron acceleration



electron transport, creation of Debye sheath, recirculation



accelerated electrons: density n, temperature T

Why often Maxwellian-like distributions are observed

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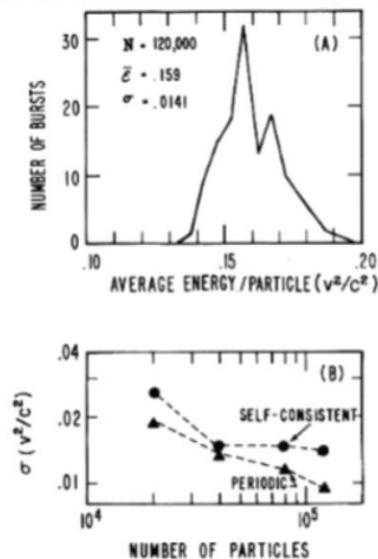


FIG. 4. (a) Pulse-height analysis of Fig. 3(b) data. Number of pulses vs pulse height. (b) Standard deviation of data from (a) vs number of simulation particles for self-consistent simulation calculations.

The generation of high-intensity laser pulses has been studied in the last few years.

tion should be compared with the Monte Carlo particle approach in which particles move randomly from the thermal bath. This steady injection removes the need for steady-state calculations and allows a systematic study of the characteristics of the collected, after passing through the interaction region, at a spatial point in time. By avoiding collecting particles originating from a later time, one can construct a distribution function for a given beam velocity during each cycle. This allows us to obtain a by-cycle composition of the distribution function.

In Fig. 5 we show (a) the distribution of hot par-

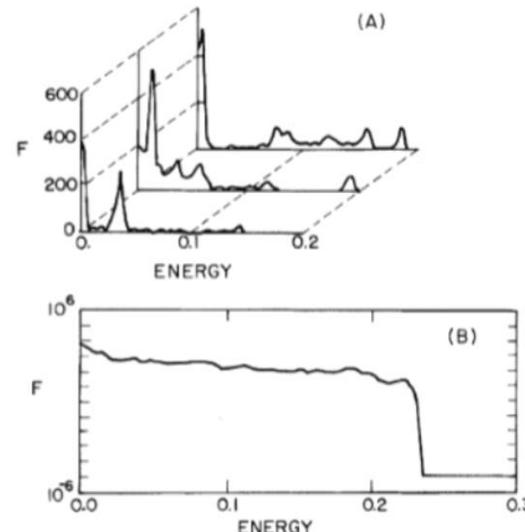


FIG. 5. Hot-electron distributions due to an injected beam with velocity $v/c = -0.08$. (a) Distributions

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free modes of

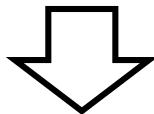
each weighted by the initial Maxwellian. From this result it is tempting to speculate that the Maxwellian character of the hot-electron distribution observed in self-consistent simulations of resonant absorption is directly related to the random fluctuations in the distribution of hot particles produced from cycle to cycle. Certainly, without this element of randomness, it would be very difficult to understand this observation.

excitations are self-consistently related to the

Picosecond pulses vs. few-femtosecond pulses

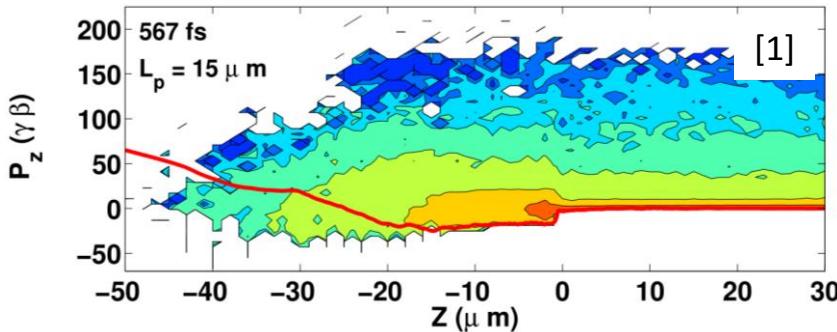
picosecond pulses

- expansion
- preplasma generation
- bulk heating
- recirculation



- relativistic oscillation of critical density surface
- electrons enter laser wave with different initial velocities in different phases

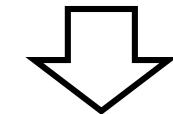
→ randomization



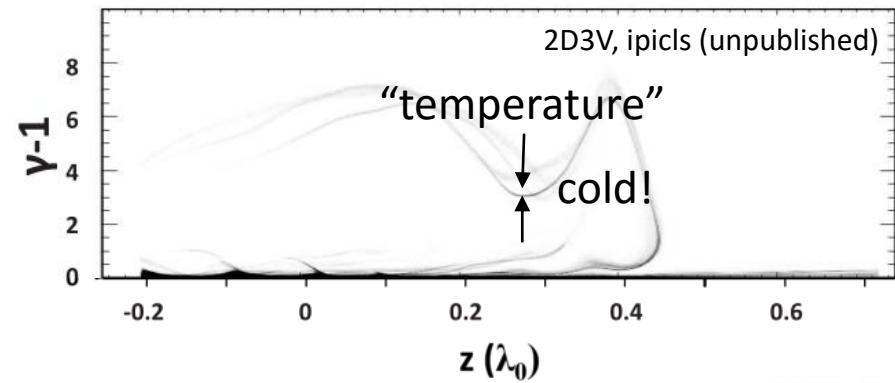
[1] B.S. Paradkar, Thesis "The effects of pre-formed plasma on the generation and transport of fast electrons in relativistic laser-solid interactions", University of California (2012)

few-femtosecond pulses

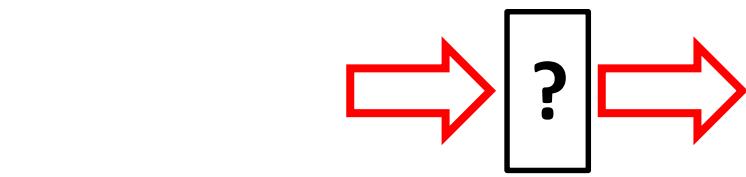
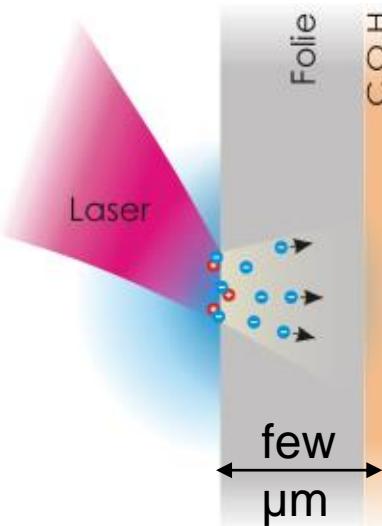
- flat target remains virtually flat
- only few 10s of nm preplasma
- no significant bulk heating
- no recirculation



- electrons enter laser wave at rest
→ problem is fully **deterministic**



Energy flux conservation for long pulses laser pulses



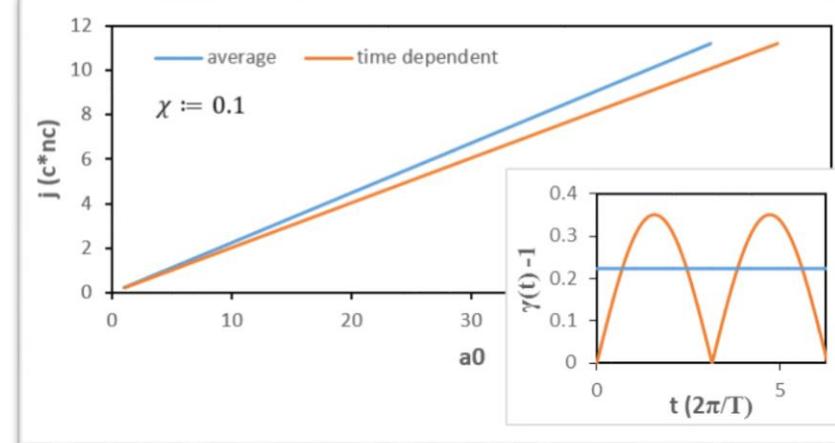
*absorbed energy
(per unit area and time)*
 χI

*energy carried
away by electrons*
 $n_{e,hot}(\gamma_h - 1)m_e v_z$

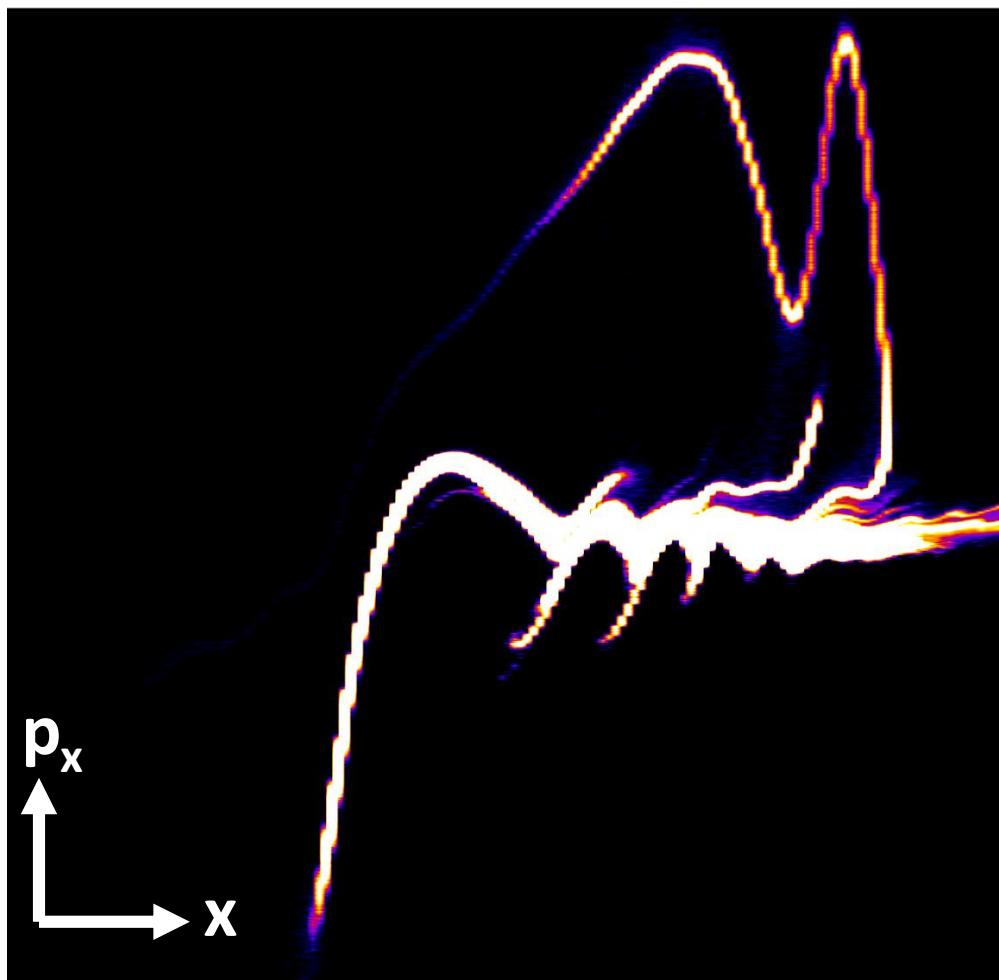
$$\text{assuming } p_z = \gamma_h - 1$$

$$\text{and } \gamma_h \cong \frac{\chi a_0}{\sqrt{2}} + 1$$

$$\Rightarrow j_h = n_{e,hot} v_z \cong \gamma_h n_c c$$



Short laser pulse (first period)

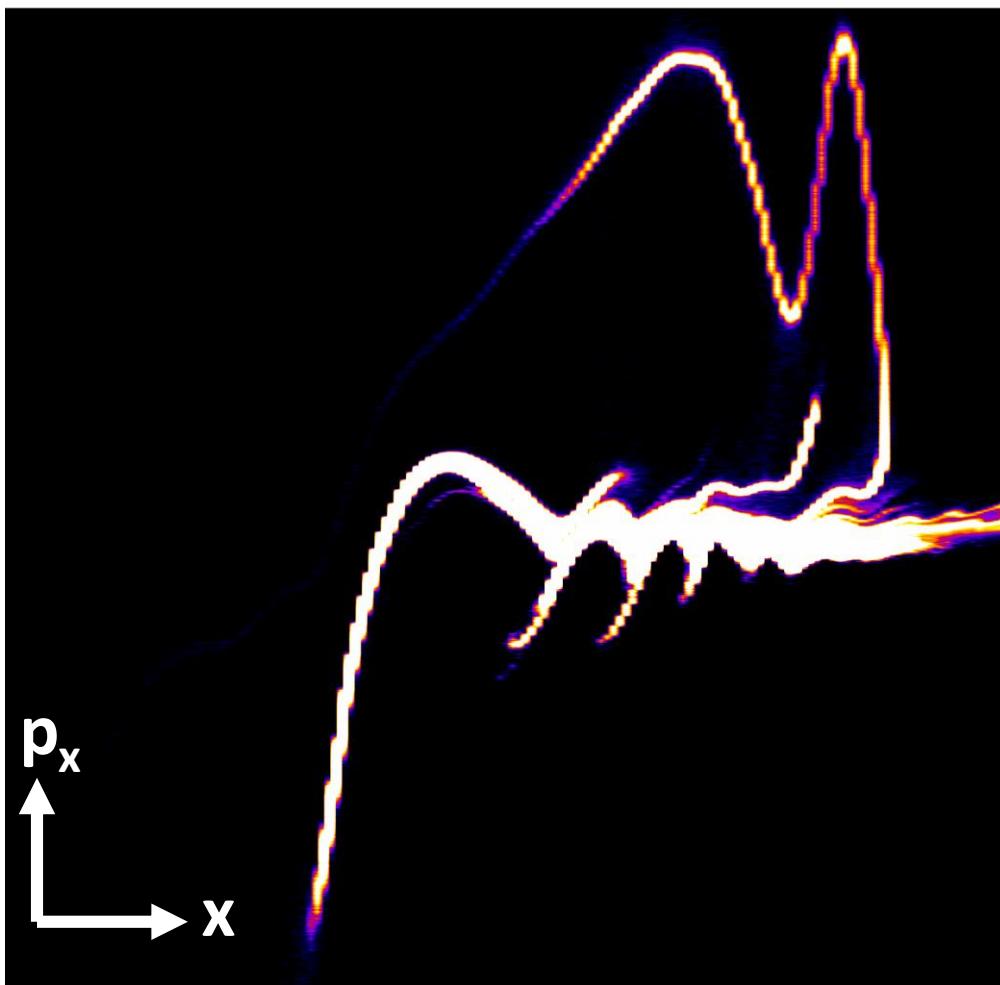


blue: E_x
green: E_y
red: B_z

PIC simulations

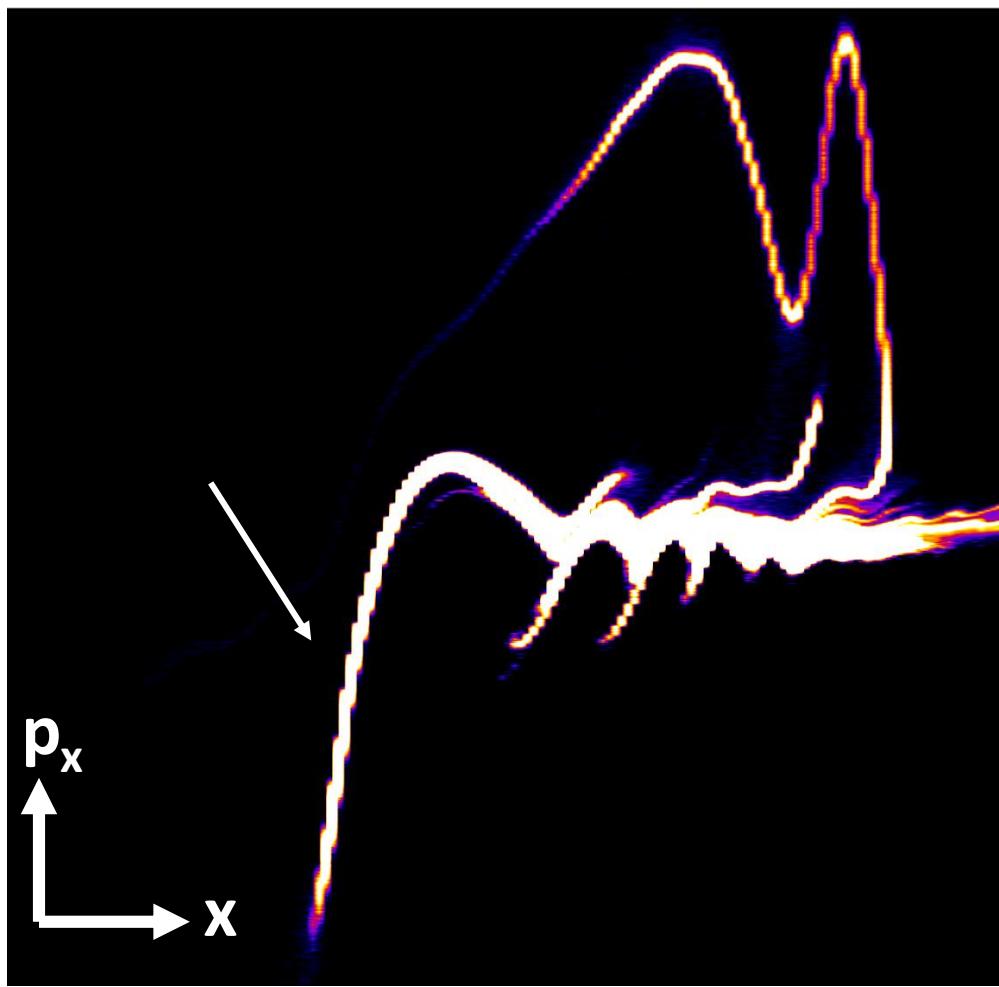
- ipicls2d (Y. Sentoku)
- plane laser, no ramp-up of intensity
- plane target, $400 n_c$
- fully ionized, $Q/A = 1/2$
- small preplasma, scale length 1 micron
- $a_0=1\dots 60$

Short laser pulse (first period)



blue: E_x
green: E_y
red: B_z

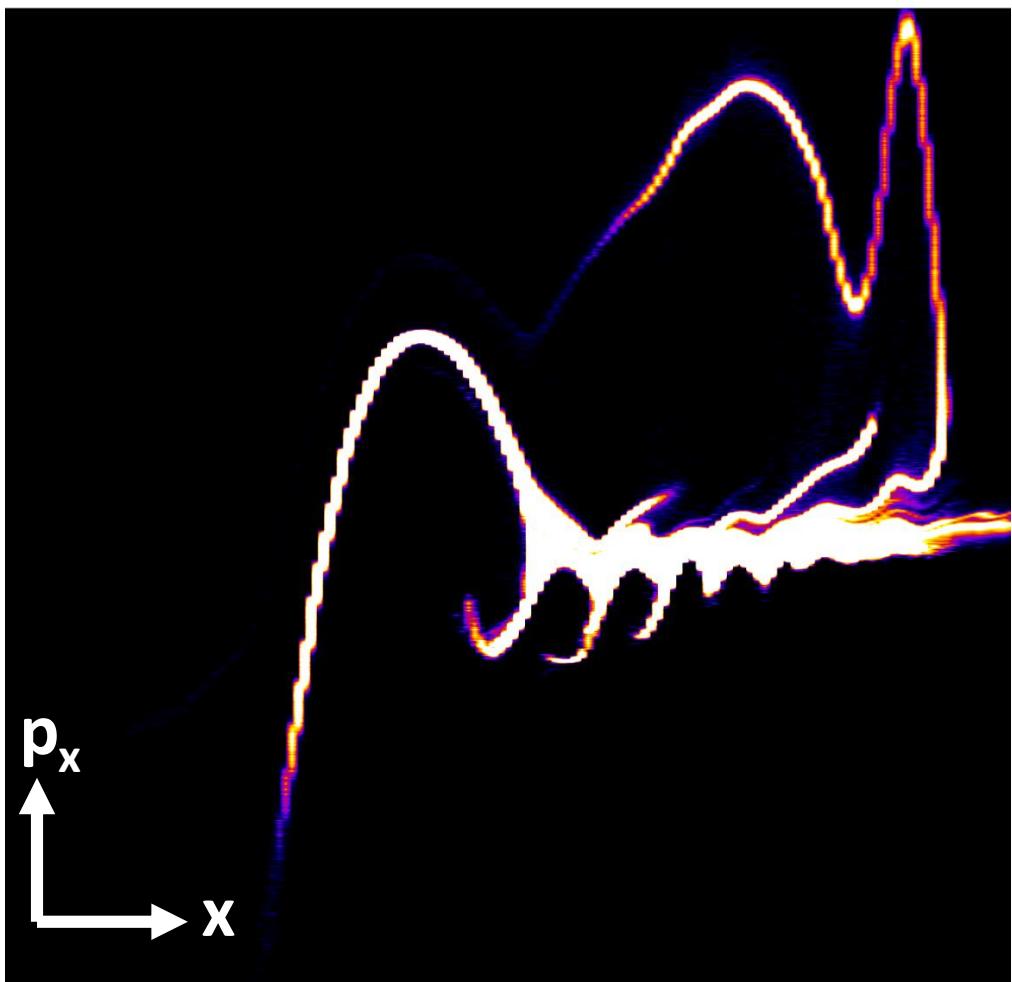
Short laser pulse (first period)



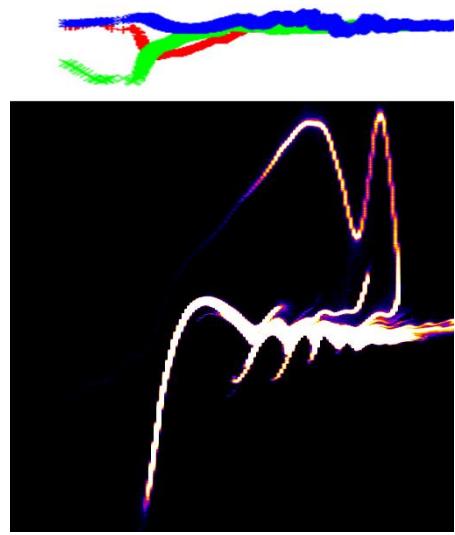
blue: Ex
green: Ey
red: Bz

- extraction

Short laser pulse (first period)

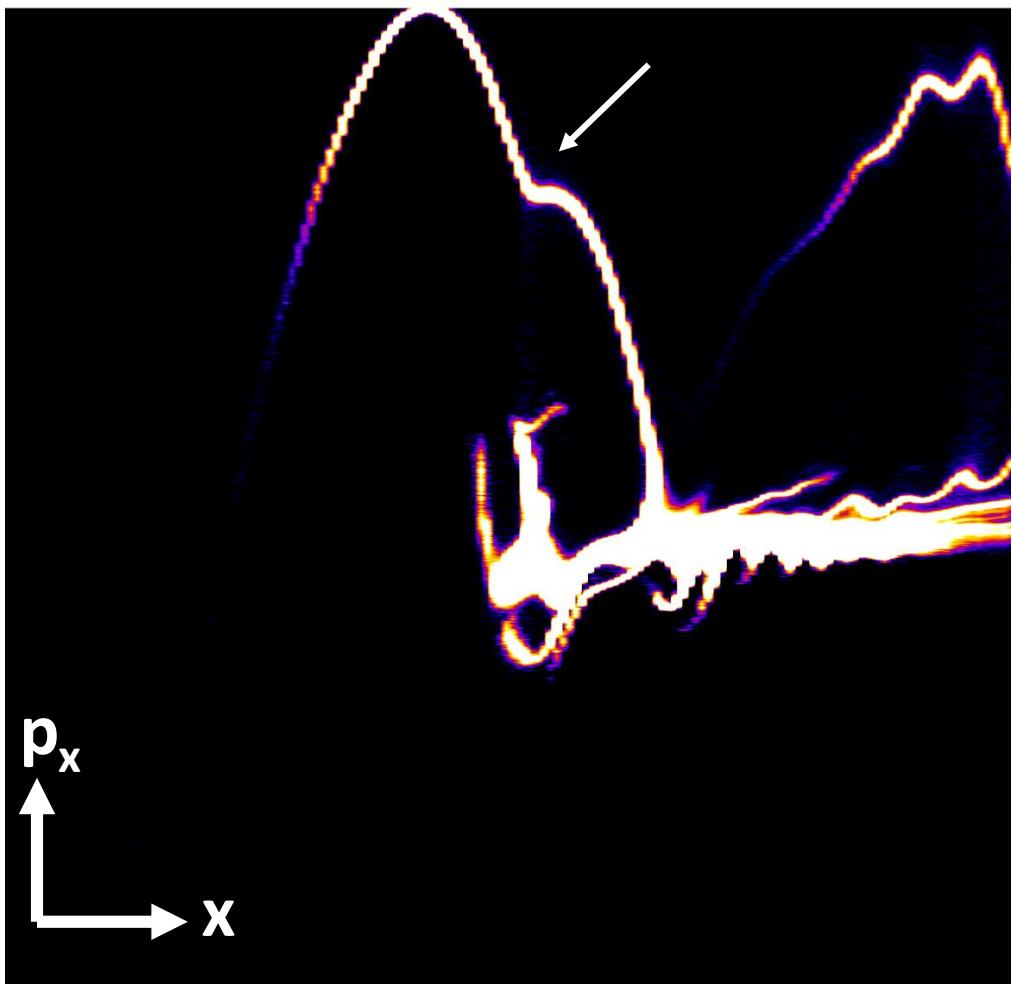


blue: E_x
green: E_y
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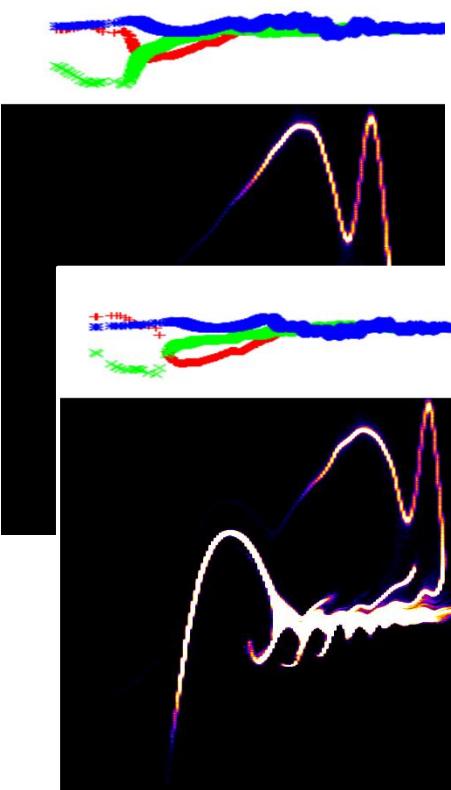
- extraction
- reflection
in B field

Short laser pulse (first period)

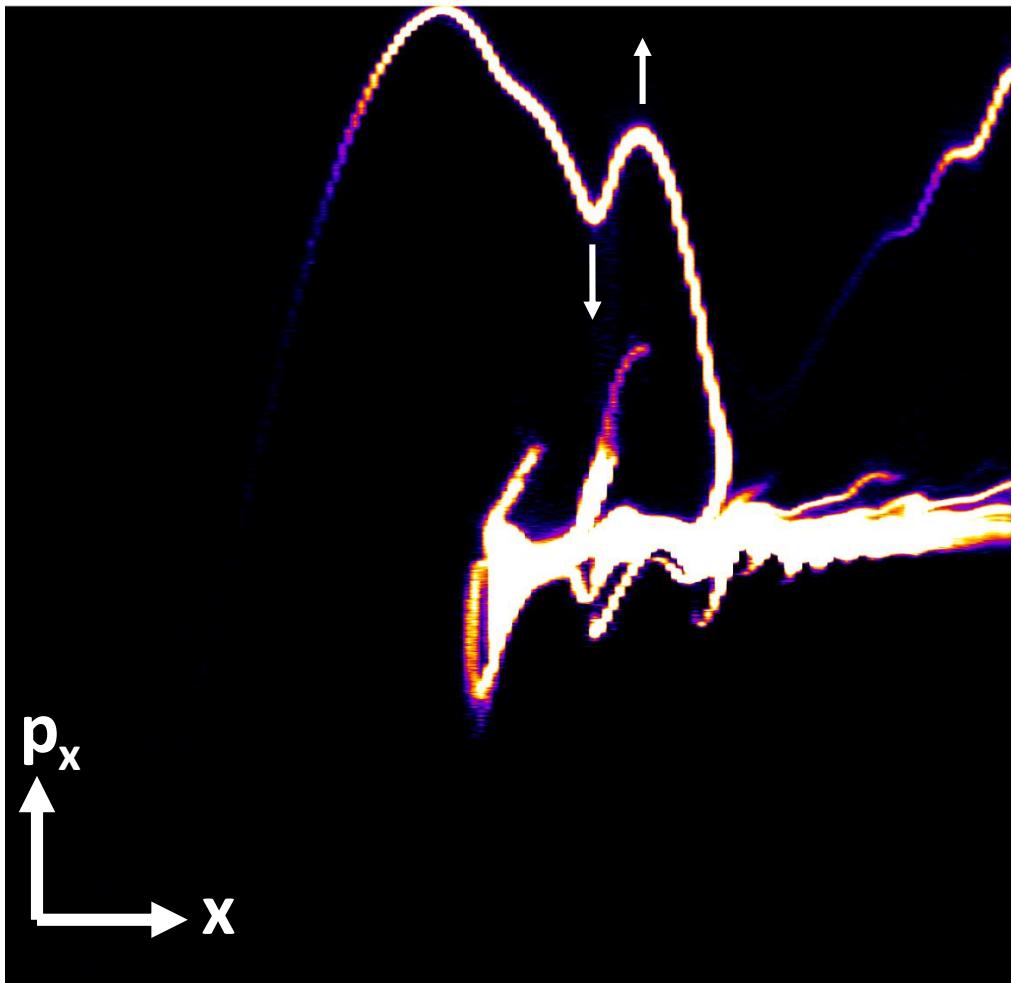
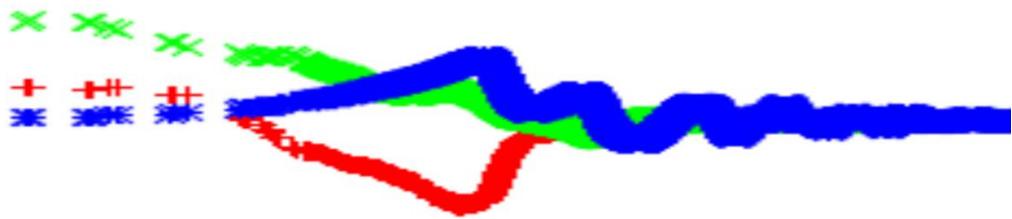


blue: Ex
green: Ey
red: Bz

- extraction
- reflection in B field
- modulation in transient ambipolar field

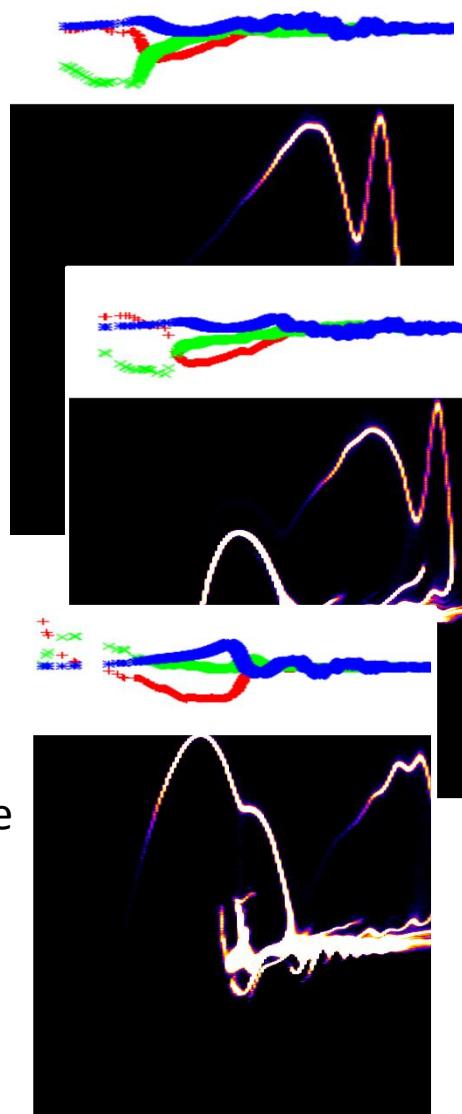


Short laser pulse (first period)

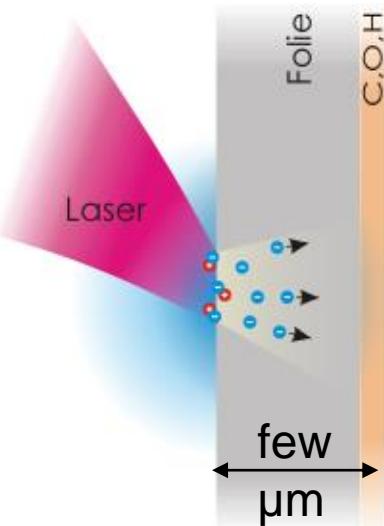


blue: Ex
green: Ey
red: Bz

- extraction
- reflection in B field
- modulation in transient ambipolar field
- plasma wave interaction



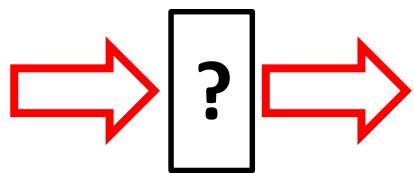
Energy flux conservation including spectro-temporal distribution



$f_t(t)$?

$p_z(t)$?

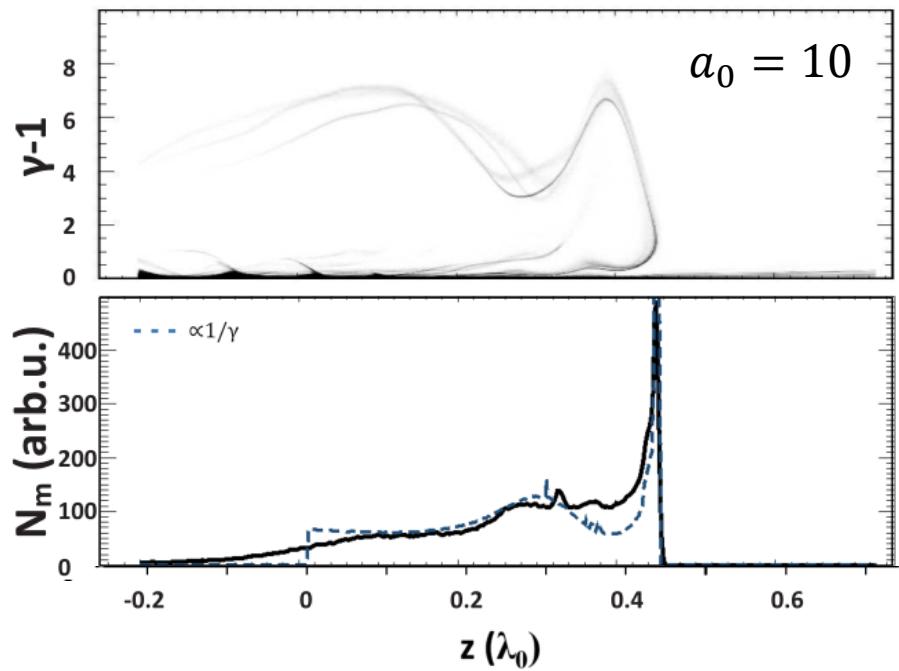
$\gamma_h(t)$?



Let $n_{e,hot}(t) \equiv n_{e,hot}^0 f_t(t)$, then
 $\frac{\chi a_0^2}{2} = n_{e,hot}^0 \langle f_t(t)(\gamma_h(t) - 1)m_e v_z(t) \rangle_t$
and
 $j_h(t) = n_{e,hot}^0 f_t(t) v_z(t)$

a
(per

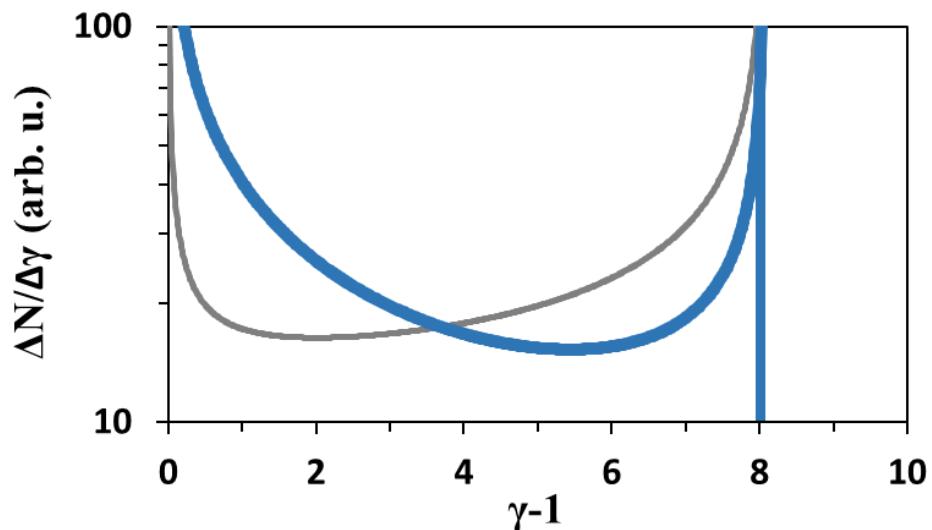
Energy flux conservation including spectro-temporal distribution



- ✓ assuming $f_t(t) \propto \gamma_h(t)^{-1}$ [1]
- ✓ and $p_z(t) = \sqrt{\gamma_h(t)^2 - 1}$
- (✓) and $\gamma_h(t) \cong \sqrt{1 + \hat{a}_0^2 \sin^2 t}$

[1] PRL 107 205003 (2011)

Energy flux conservation including spectro-temporal distribution



- **not** Maxwellian or exponential
- peak at high energies
- Shape depends on f_t

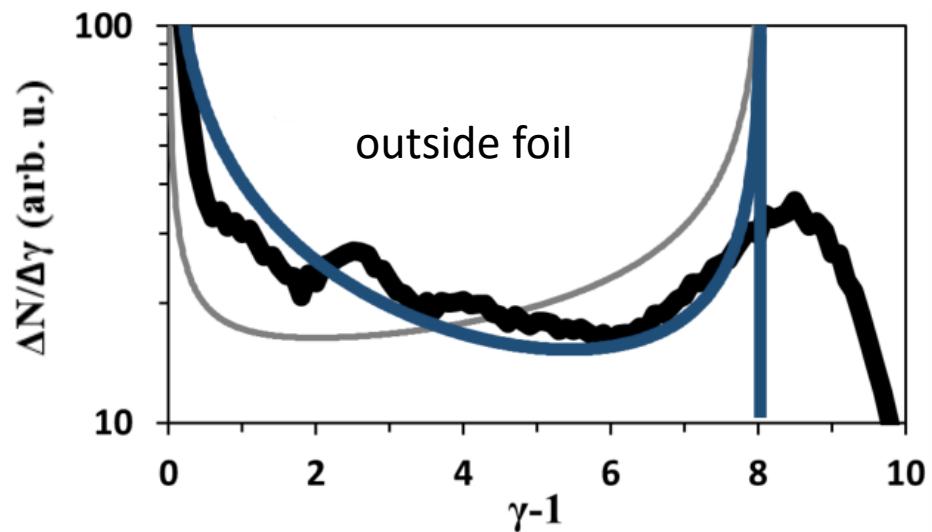
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Simple comparison - spectrum:

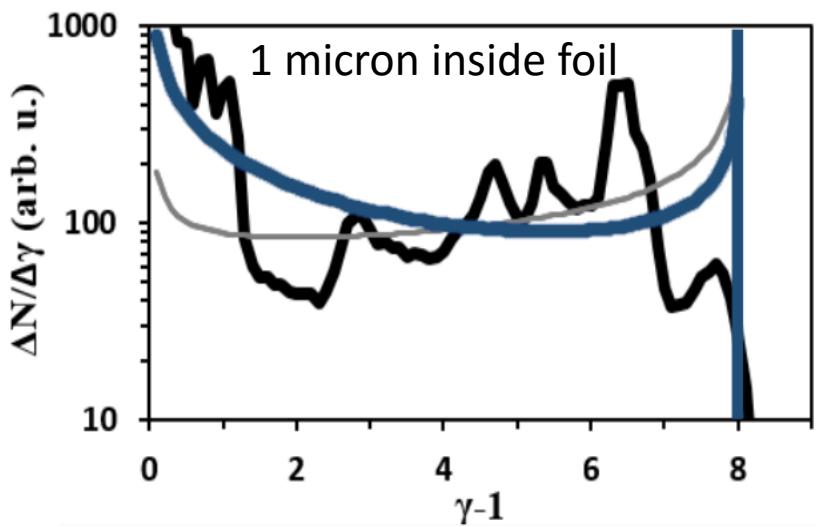
$$f_\gamma(\gamma_h) = \frac{dN}{dt} \frac{dt}{d\gamma_h} = \frac{f_t}{\dot{\gamma}_h}$$

$$f_\gamma(\gamma_h) = \frac{\pi}{2K(-a_0^2)} \frac{1}{\sqrt{(\gamma_h^2 - 1)(a_0^2 - \gamma_h^2 + 1)}}$$

Energy flux conservation including spectro-temporal distribution



- assuming $f_t(t) \propto \gamma_h(t)^{-1}$ [1]
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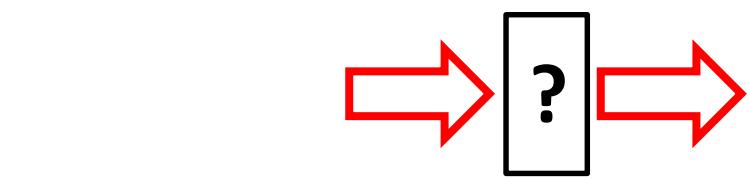
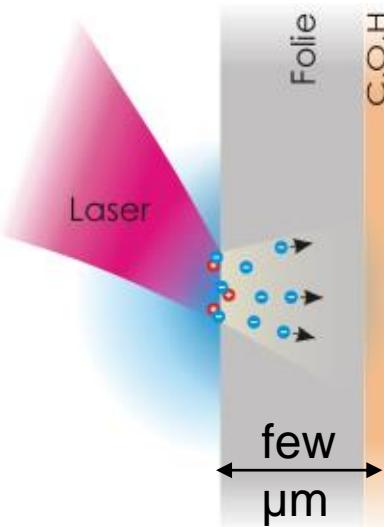


Simple comparison - spectrum:

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$$f_\gamma(\gamma_h) = \frac{\pi}{2K(-a_0^2)} \frac{1}{\sqrt{(\gamma_h^2 - 1)(a_0^2 - \gamma_h^2 + 1)}}$$

Temporal variation of current (within a bunch)



*absorbed energy
(per unit area and time)*

$$\chi I$$

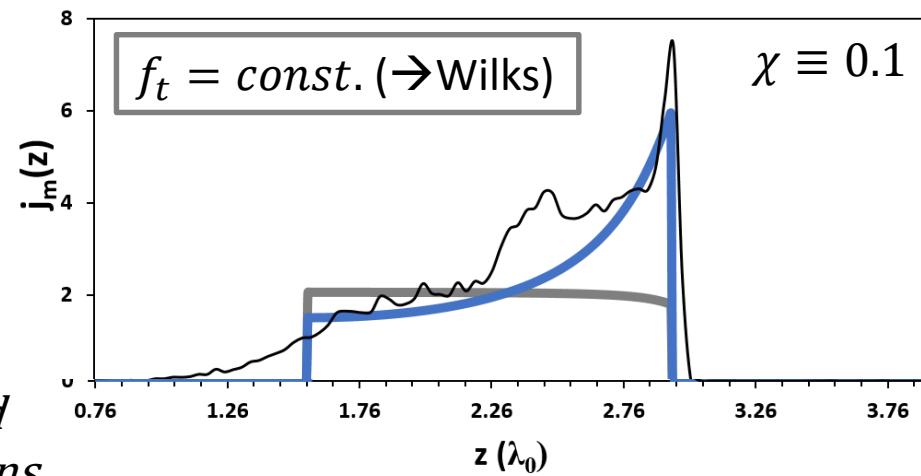
*energy carried
away by electrons*

$$\langle n_{e,hot}(t)(\gamma_h(t) - 1)m_e v_z(t) \rangle$$

$$\chi \frac{a_0^2}{2} = n_{e,hot}^0 \langle f_t(t)(\gamma_h(t) - 1)v_h(t) \rangle$$

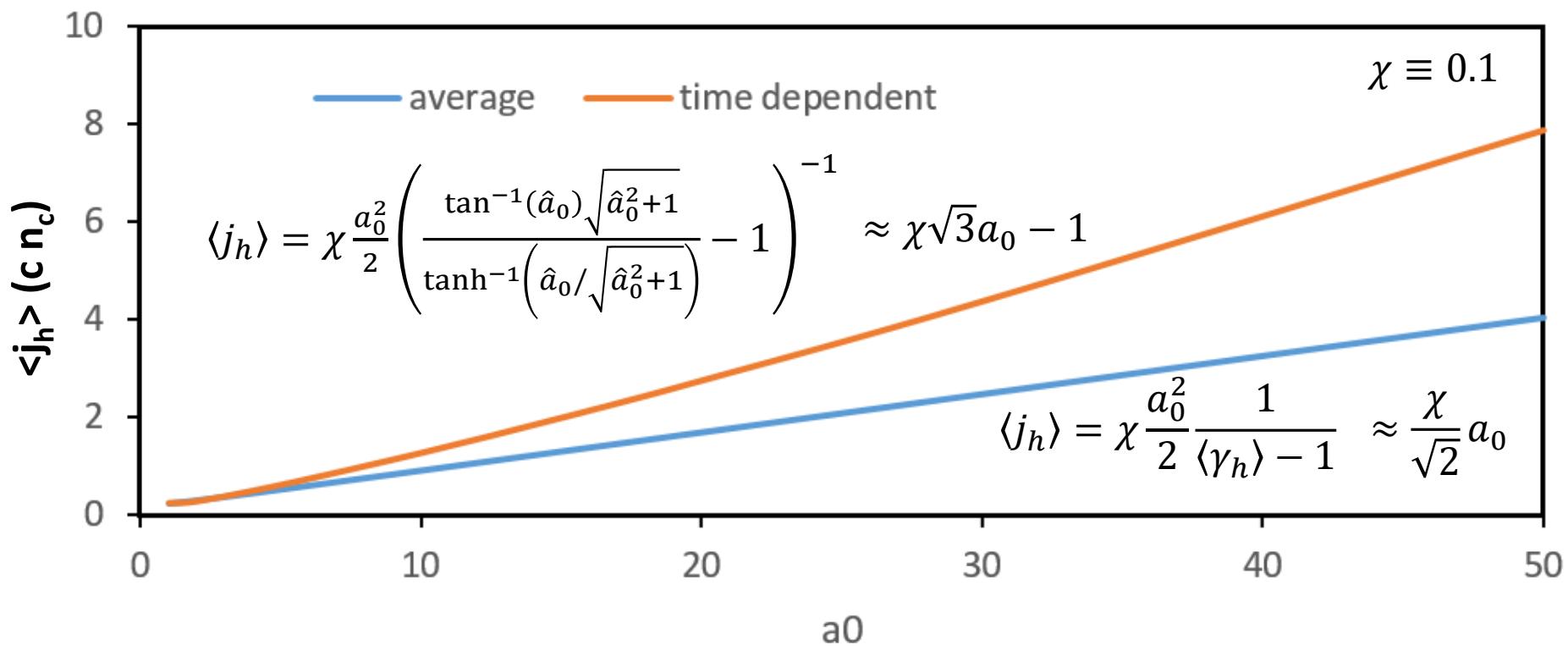
$$j_h(t) = n_{e,hot}^0 f_t(t) v_h(t)$$

- ✓ assuming $f_t(t) \propto \gamma_h(t)^{-1}$
- ✓ and $p_z(t) = \sqrt{\gamma_h(t)^2 - 1}$
- (✓) and $\gamma_h(t) \cong \sqrt{1 + \hat{a}_0^2 \sin^2 t}$

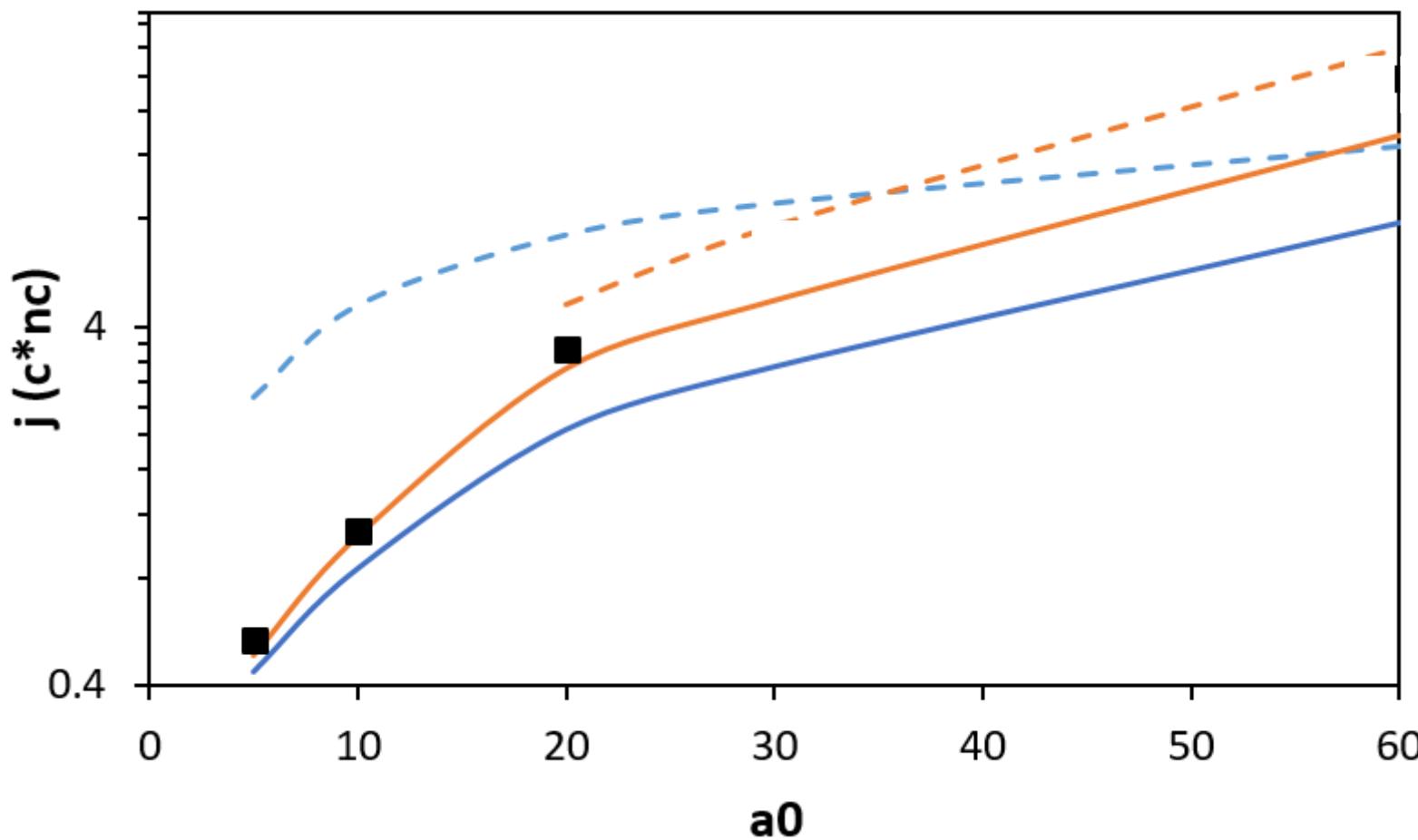


$$j_h(t) = \chi \frac{a_0^2}{2} \frac{1}{\left(\frac{(\gamma - 1)\sqrt{\gamma^2 - 1}}{\gamma^2} \right)} \frac{\sqrt{\gamma(t) - 1}}{\gamma(t)^2}$$

Comparison average currents



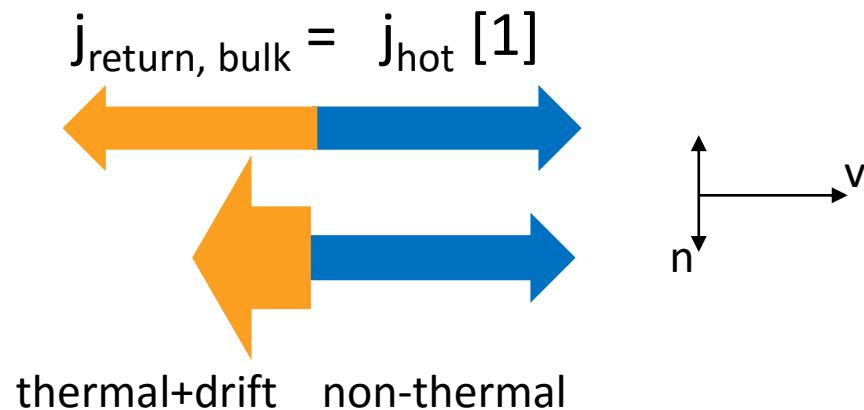
Average current



energy flux conservation solved with

- time-dependent variables, assuming quiver motion
- average, **constant** variables, assuming quiver motion
- - - time-dep. variables assuming $\gamma_h(t)$ modified for co-moving e- in surface fields
(PRL 107 205003 (2011))

Impact on Ohmic heating



- return current is **slow** ($n_{e,\text{hot}}/n_{\text{bulk}}$)

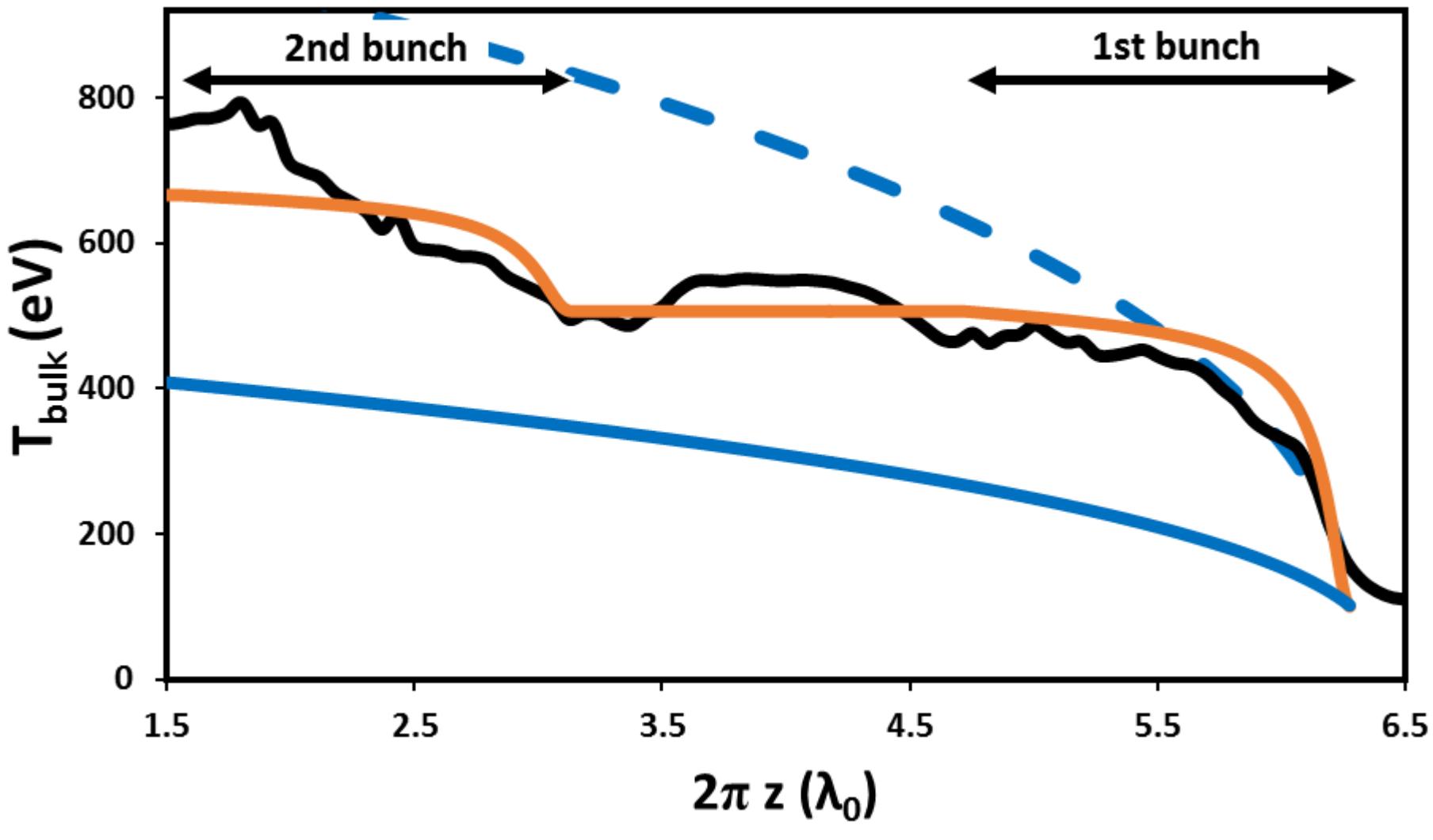
- **Ohmic heating effective** [2]:

$$\frac{3}{2} n_{e,\text{bulk}} \frac{\partial T_{e,\text{bulk}}}{\partial t} = \eta j_{\text{hot}}^2 \text{ with } \eta \propto Z\Lambda/T_{e,\text{bulk}}^{3/2} \text{ from Spitzer model}$$

[1] Plasma Phys. Control. Fusion 39, 653 (1997)

[2] Phys. Plasmas 2, 2796 (1995)

Impact on Ohmic heating

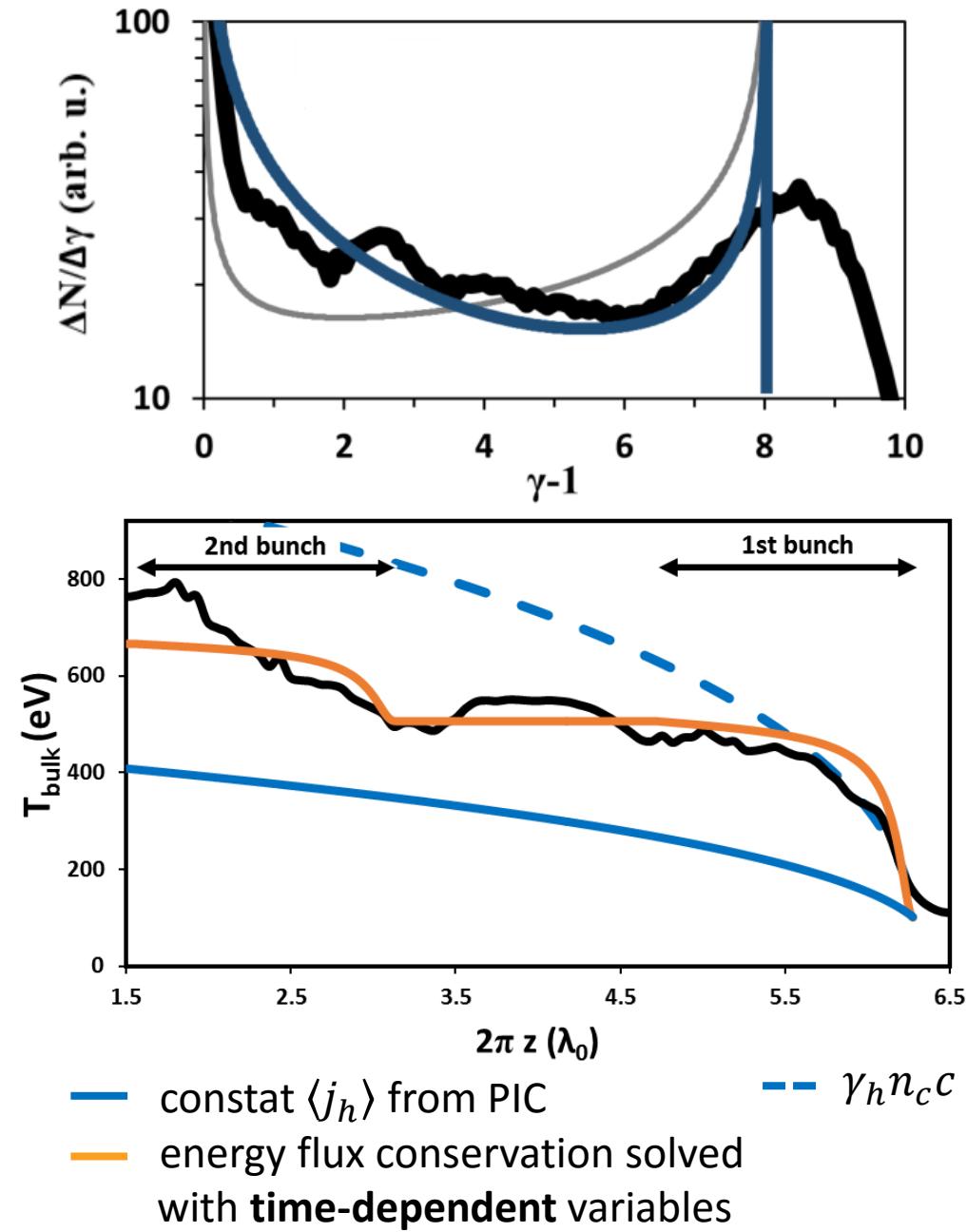


— $\langle j_h \rangle$ from PIC - - - $\gamma_h n_c c$
energy flux conservation solved with
— time-dependent variables, assuming quiver motion

Conclusions and outlook

- generally, at laser-solid interactions of ultrashort lasers, the accelerated electrons are ***not thermal***
- this can have ***experimentally relevant*** effects:
 - spectrum
 - current
 - Ohmic heating
 - magnetic field generation and collimation
 - Others?
 - K α yield
 - K α line broadening and shift
 - (which with temperature models need superposition of several temperatures to explain experiments)
 - For ion acceleration, behind foil quick randomization occurs due to static fields.
Does one see the effect of temporal and spectral structure in ions?
- for **realistic lasers** (prepulses, few cycles, mobile ions + expansion of foil during laser, heating) and thin foils (recirculation) one might still get some ***randomization*** of electron phases = exponential-like spectra

Conclusions and outlook



$$j_h(t) = \chi \frac{a_0^2}{2} \frac{1}{\left(\frac{(\gamma - 1)\sqrt{\gamma^2 - 1}}{\gamma^2} \right)} \frac{\sqrt{\gamma(t) - 1}}{\gamma(t)^2}$$

$$\langle j_h \rangle = \chi \frac{a_0^2}{2} \left(\frac{\tan^{-1}(\hat{a}_0) \sqrt{\hat{a}_0^2 + 1}}{\tanh^{-1}\left(\hat{a}_0 / \sqrt{\hat{a}_0^2 + 1}\right)} - 1 \right)^{-1}$$

$$\approx \chi \sqrt{3} a_0 - 1$$

- ultrashort laser-solid interaction:
non-thermal energy distribution in energetic bunches
- current $\neq \gamma n_c c$