Polarization Measurements of the Radio Emission of Cosmic Ray Air Showers with LOPES



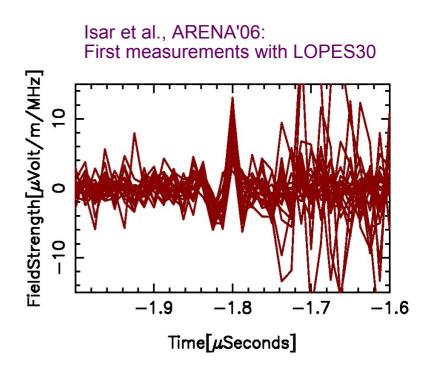
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ARENA 2008, Roma, Italy

Polarization Measurements

Motivation:

 Initially, all 30 antennas of the LOPES experiment were aligned to measure the East-West polarization direction of the air shower radio emission only.



- End of 2006, for recording the full radio signal, LOPES30 has been reconfigured to perform polarization measurements.
- Polarization measurements provide the 'tool' that can verify the geosynchrotron mechanism of the radio emission in air showers.

LOPES at the KASCADE-Grande Array



Polarization Configuration

Location: most antennas inside the original KASCADE array

Configuration:

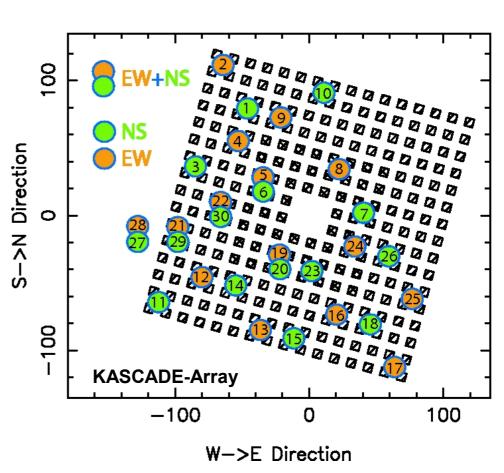
5 EW+NS oriented antennas [28+27,21+29,22+30,5+6,19+20]

10 NS oriented antennas [1,3,7,10,11,14,15,18,23,26]

10 EW oriented antennas [2,4,8,9,12,13,16,17,24,25]

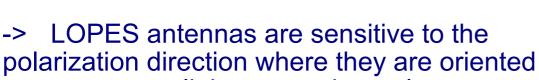
Trigger sources: KASCADE and KASCADE-Grande

LOPES Antenna Layout

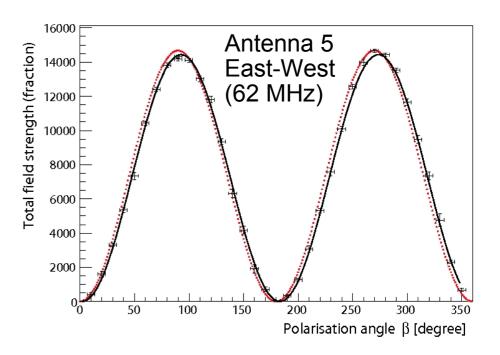


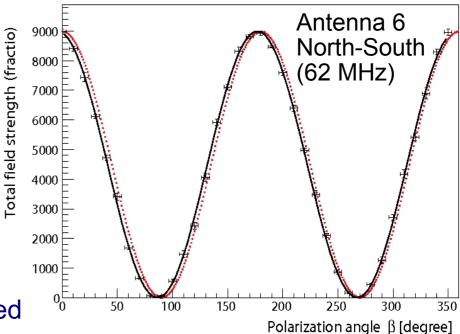
Amplitude Calibration





-> no cross-talk between channels

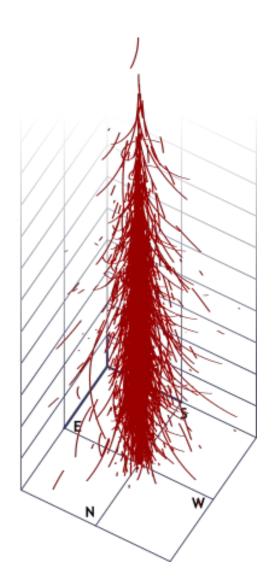




Theoretical Predictions

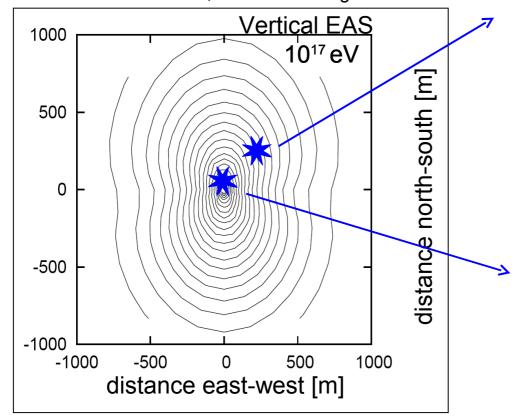
What simulations tell us: [by T. Huege et al]

- highly linearly polarized radio emission
- signal usually present in both polarization components:
 East-West and North-South
- polarization directly related with the shower azimuth for a given zenith angle (see actual analysis)
- dependence of the signal on the position of the observer relative to the incoming shower (see next for illustration)

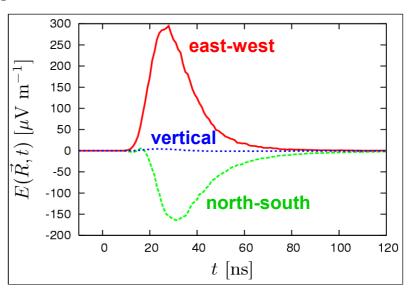


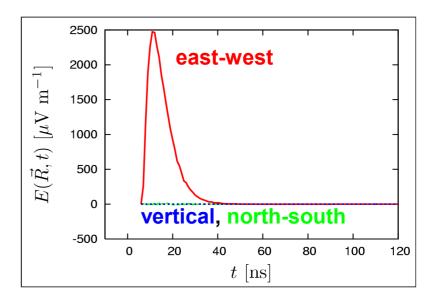
Simulated Pulse Example

45° inclined air shower, total field strength at 10MHZ

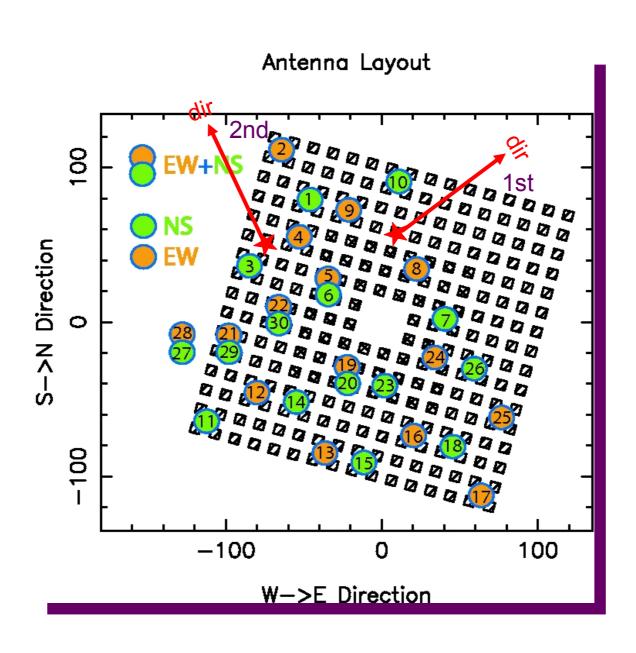


Huege & Falcke, AstropPhys 24 (2005) 116-136





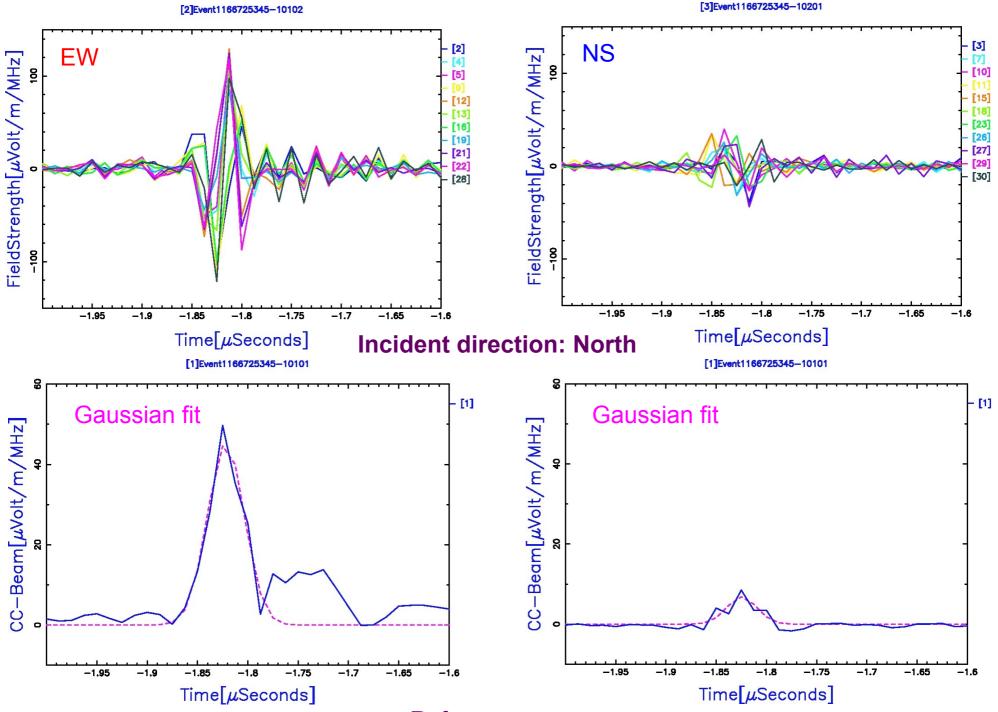
Measured Event Examples



 ϕ =51.18°, θ =66.44°, Geo-Angle:83°, $\log(N_e)$ = 5.3, $\log(N_u)$ = 5.4, E_p (estimate)= 1.8 x 10¹⁸ eV [4]Event1166245311-10202 EW [4] NS FieldStrength[*\m*Volt/m/MHz] FieldStrength[**Wolt/m/MHz] [8] [12] [21] [22] -1.85-1.8-1.65-1.6 -1.55 -1.75-1.85-1.8-1.6-1.55Time[µSeconds] Time[µSeconds] **Incident direction: North-East** [1]Event1166245311-10101 [1]Event1166245311-10101 - [1] **–** [1] Gaussian fit Gaussian fit CC—Beam[*w*Volt/m/MHz] CC-Beam[µVolt/m/MHz] -1.85-1.8 -1.75 -1.7 -1.65-1.55 -1.6-1.85-1.8-1.55-1.75-1.6Time[µSeconds] $Time[\mu Seconds]$

Ref: Isar et al, ICRC'07

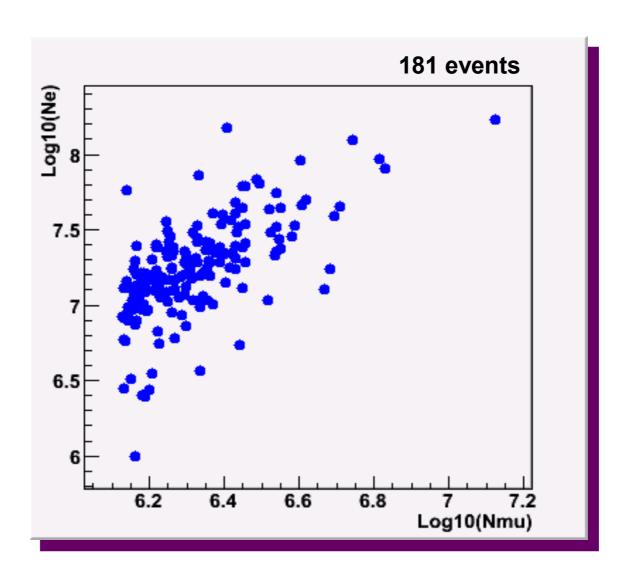
 ϕ =332.7°, θ =54.2°, Geo-Angle:77°, $\log(N_e)$ = 5.8, $\log(N_u)$ = 5.4, E_p (estimate)= 2.9 x 10¹⁷ eV



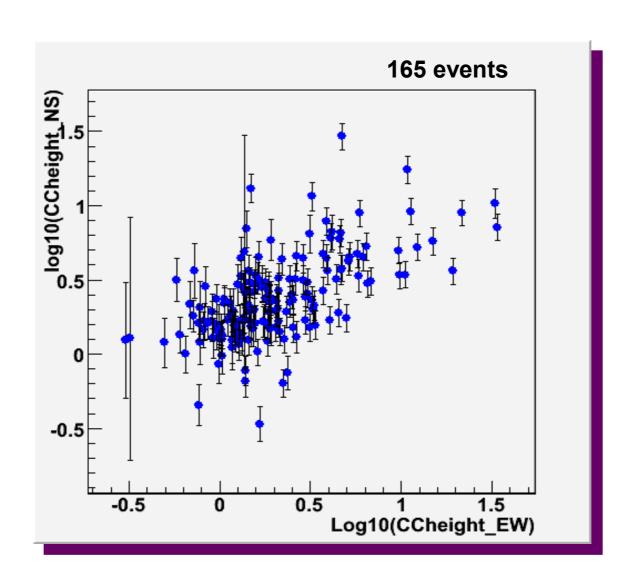
Ref: Isar et al, ICRC'07

Analysis

- 1) analysed events so far: Dec. 2006 Aug. 2007
- 2) selection: well-reconstructed by KASCADE Grande: θ <50°, log(N_{μ})>6.1

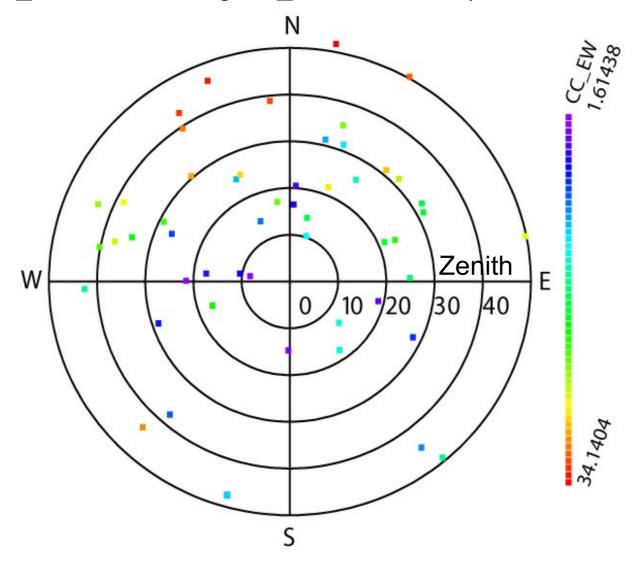


3) complete pipeline applied for each EW and NS channel, to calculate the CC-beam in EW and NS polarization components independently



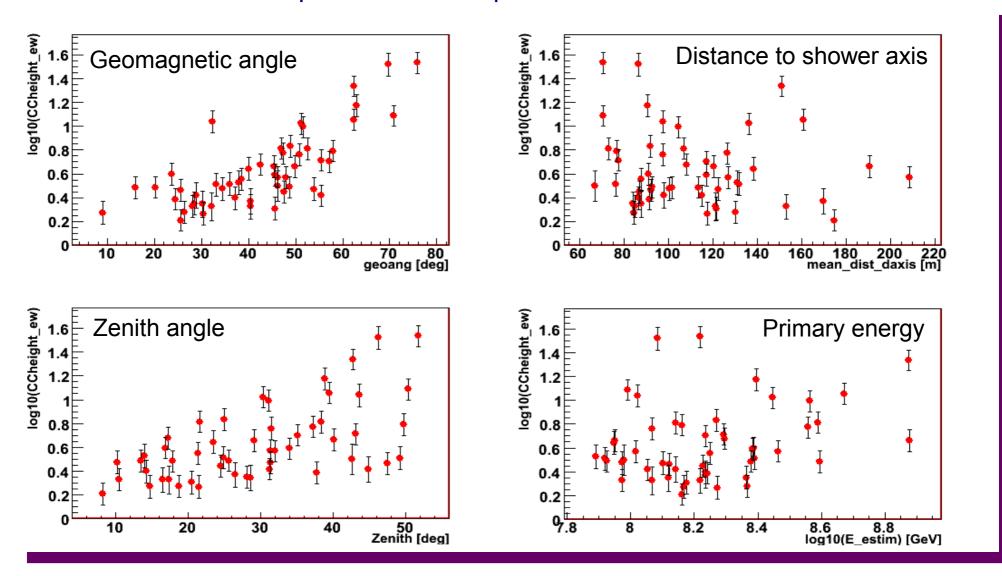
4) Detection

I. East-West polarization: 52 events (CC_EW-beam>1.5, good_EW coherence)

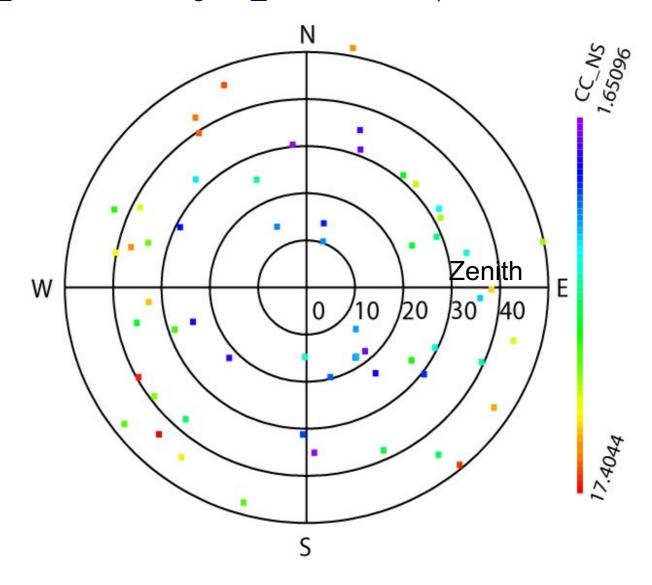


Largest E-W polarized signals detected for showers coming from North

- 5) (unmodified) pulse height correlations with air shower parameters
 - I. East-West polarization component

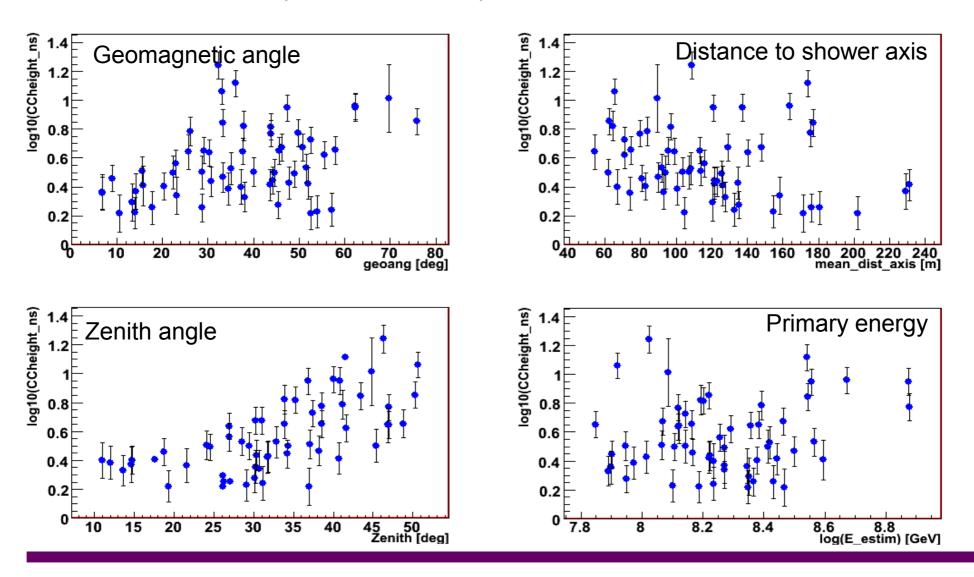


 North-South polarization: 58 events (CC_NS-beam>1.5, good_NS coherence)



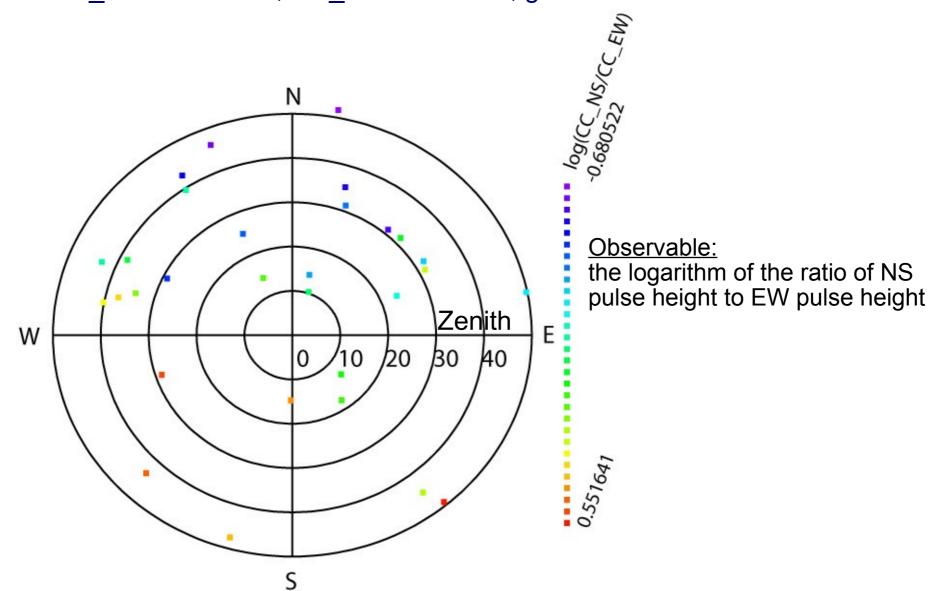
Largest N-S polarized signals detected for showers coming from West or East

II. North-South polarization component



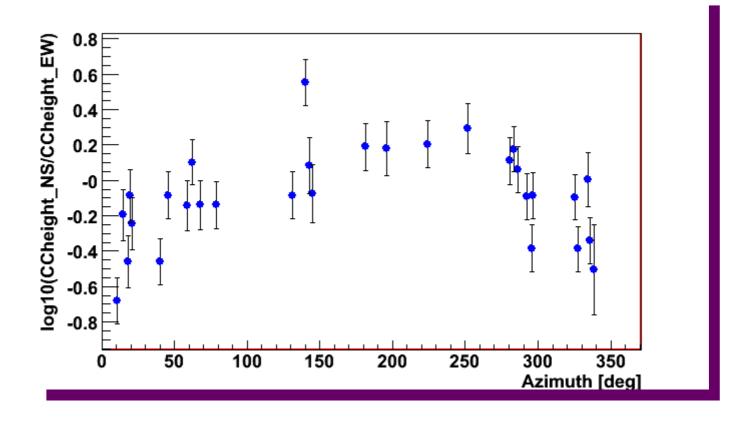
Normal correlations with air shower parameters

III. Dual-Polarization: 30 events CC_EW-beam>1.5, CC_NS-beam>1.5, good EW and NS coherence



red: N-S component dominant; blue: E-W component dominant

III. Dual-Polarization



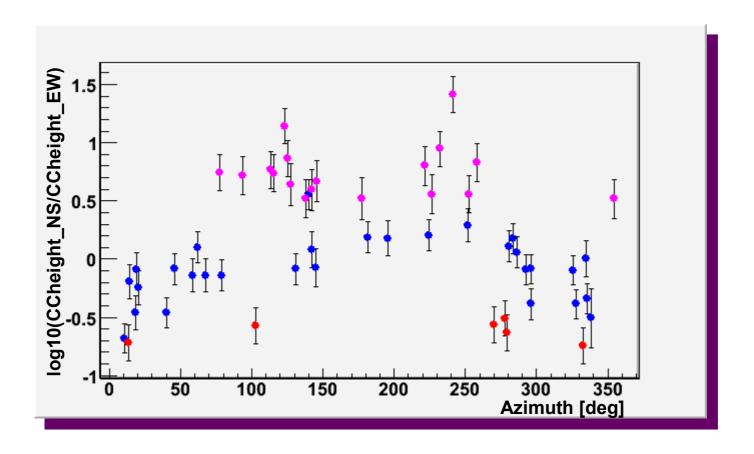
- -> Azimuth dependence of the pulse height ratio!
- > 0: N-S component dominant
- < 0: E-W component dominant

Pink: CC_NS>1.5; good_NS coherence; CC_EW<1.5;

value of the CC_EW set to 0.5 as upper limit

Red: CC_EW>1.5; good EW coherence; CC_NS<1.5;

value of the CC_NS set to 0.5 as upper limit



- -> N-S polarization component dominant for showers coming from East and West
- -> E-W polarization component dominant for showers coming from North

Outlook:

- LOPES is performing polarization measurements and allows a much more detailed analysis of the radio events than with measuring the East-West polarization component only.
- The array is absolute amplitude calibrated in order to estimate the electric field strength of the short radio pulse generated in air showers.
- It follows:
 - -> analysis of the signal for individual antennas
 - -> normalization of the pulse heights to primary energy, distance to shower core, etc.
 - -> estimate of the polarization vector of the radio emission in air showers
 - -> detailed comparisons with simulations

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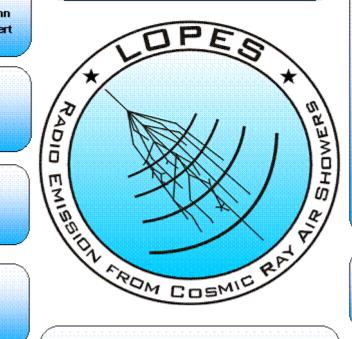
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