

# Cryogenic Silicon for LIGO

Nicolas Smith-Lefebvre

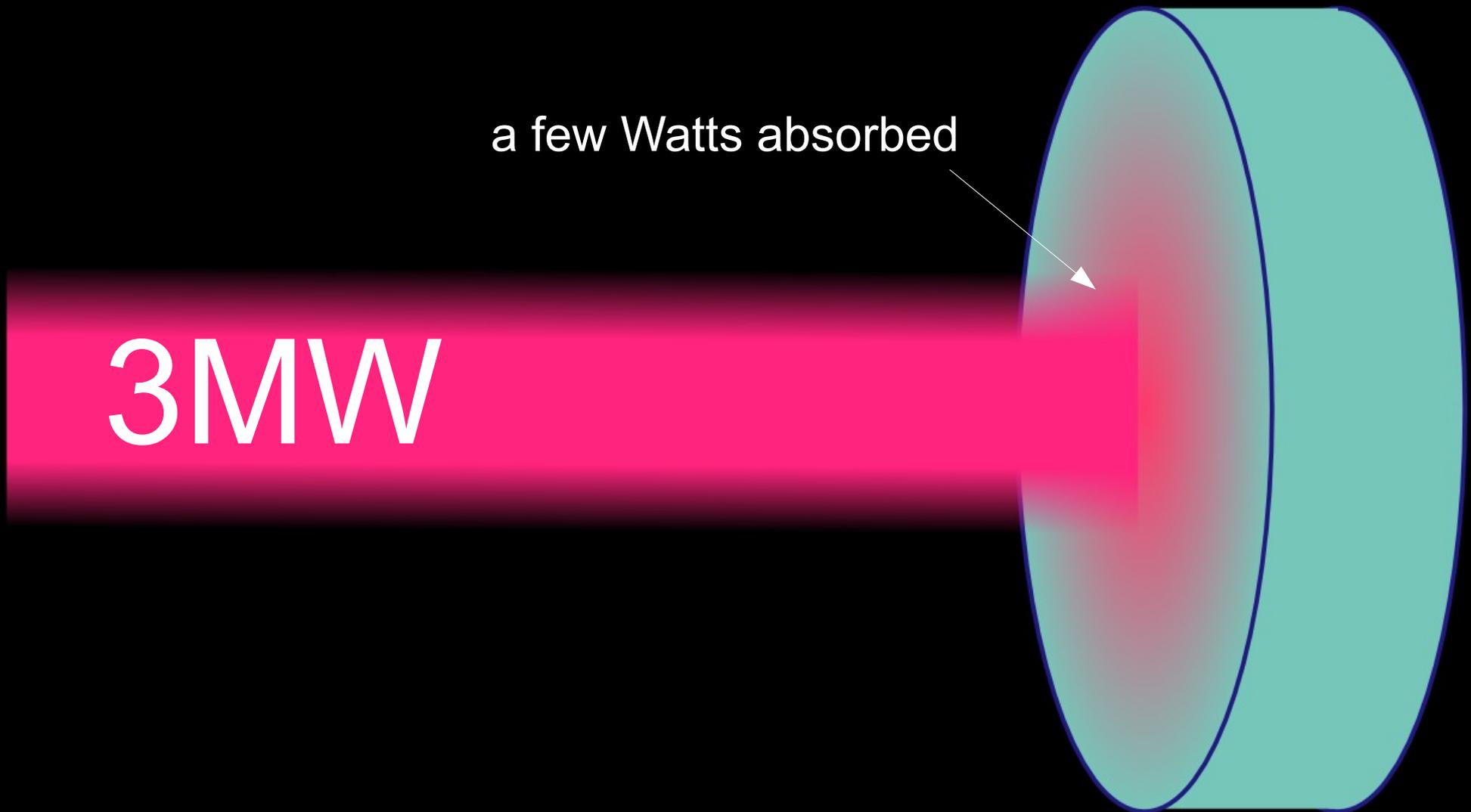
Caltech

GWADW 2013

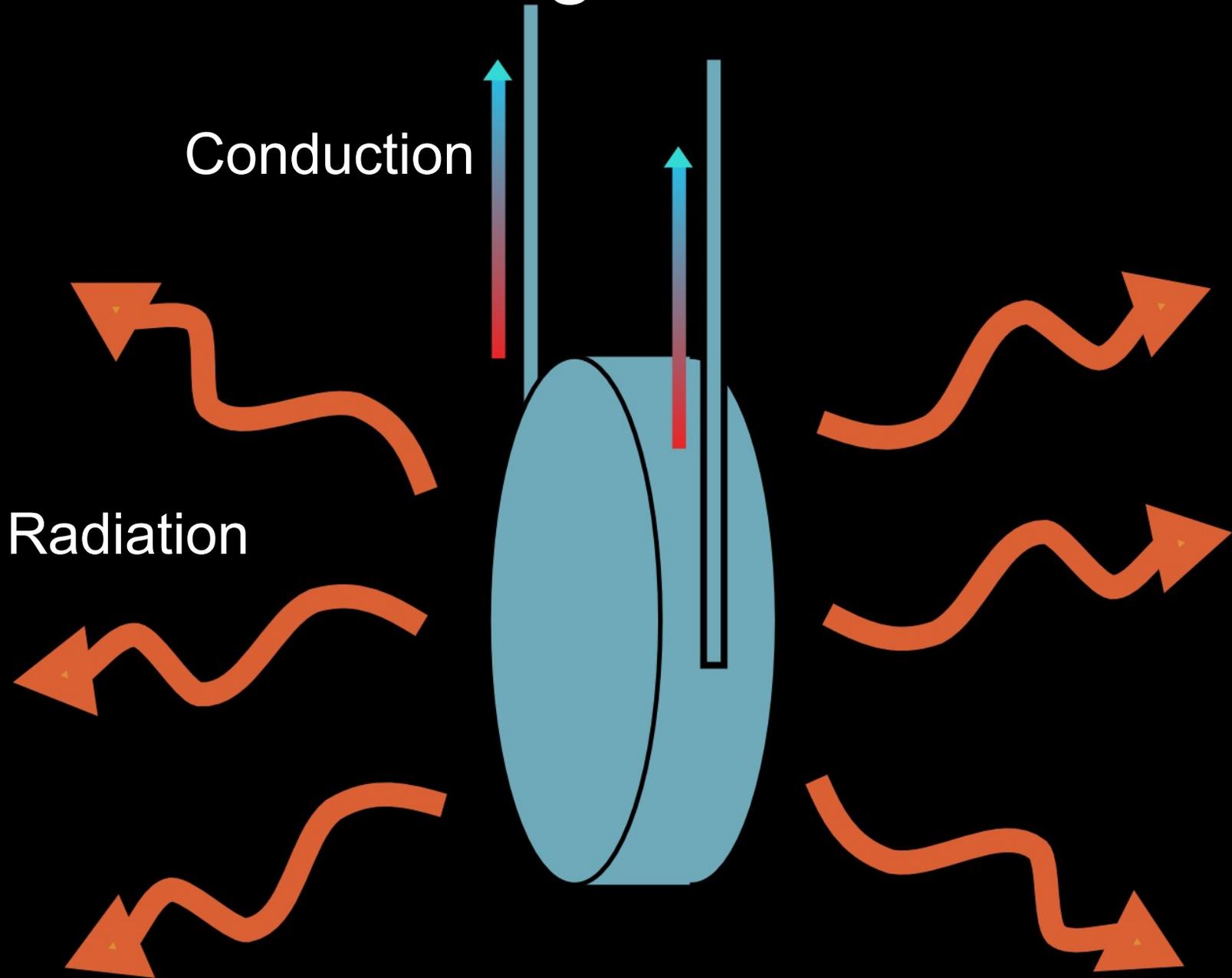
# Outline

- How can one keep the interferometer cold?
- Research that is being done regarding cryogenic silicon suspensions.
- A bid for interest in non-equilibrium thermal noise.

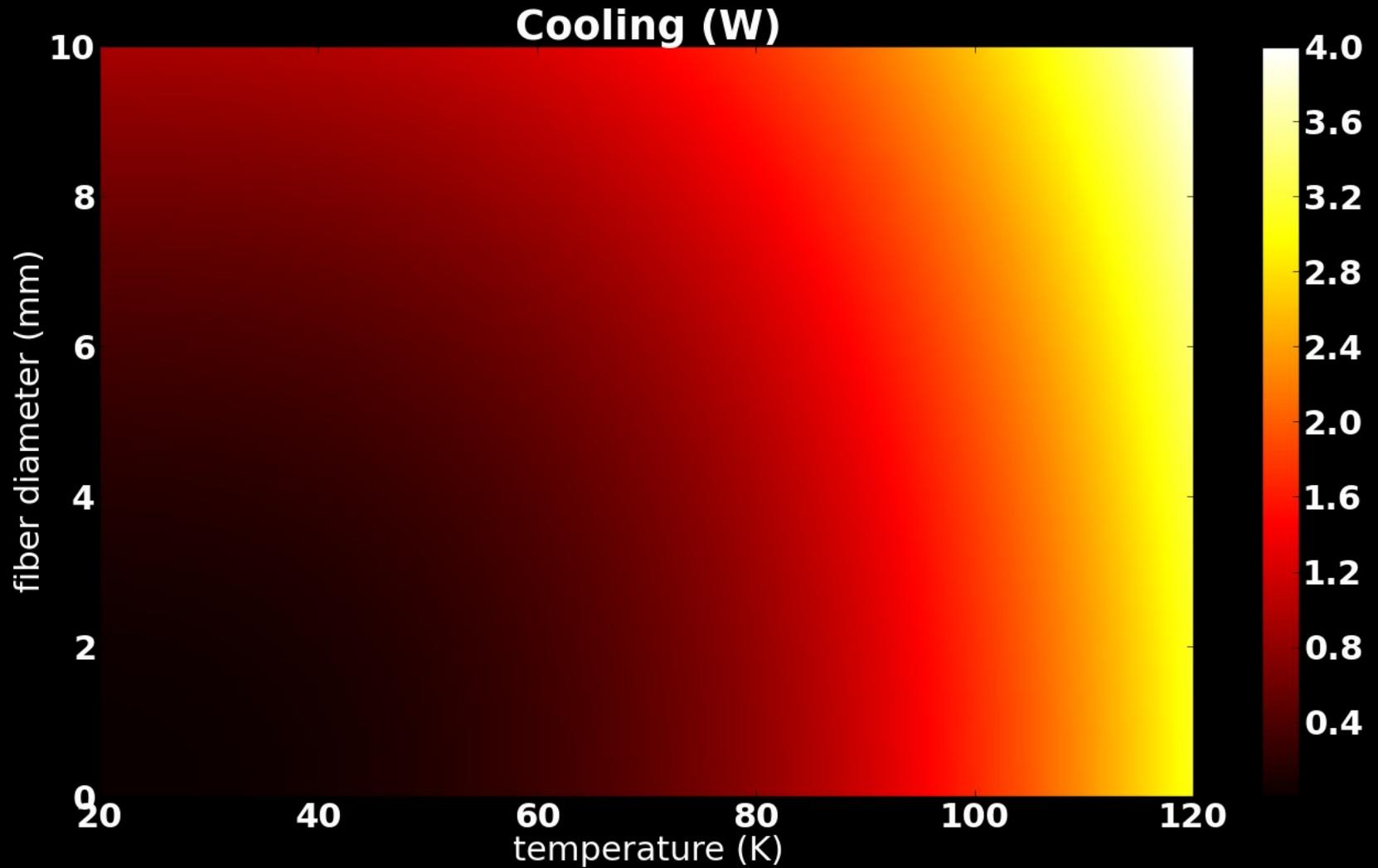
# Cryogenics in Gravitational Wave Interferometers: the cooling problem



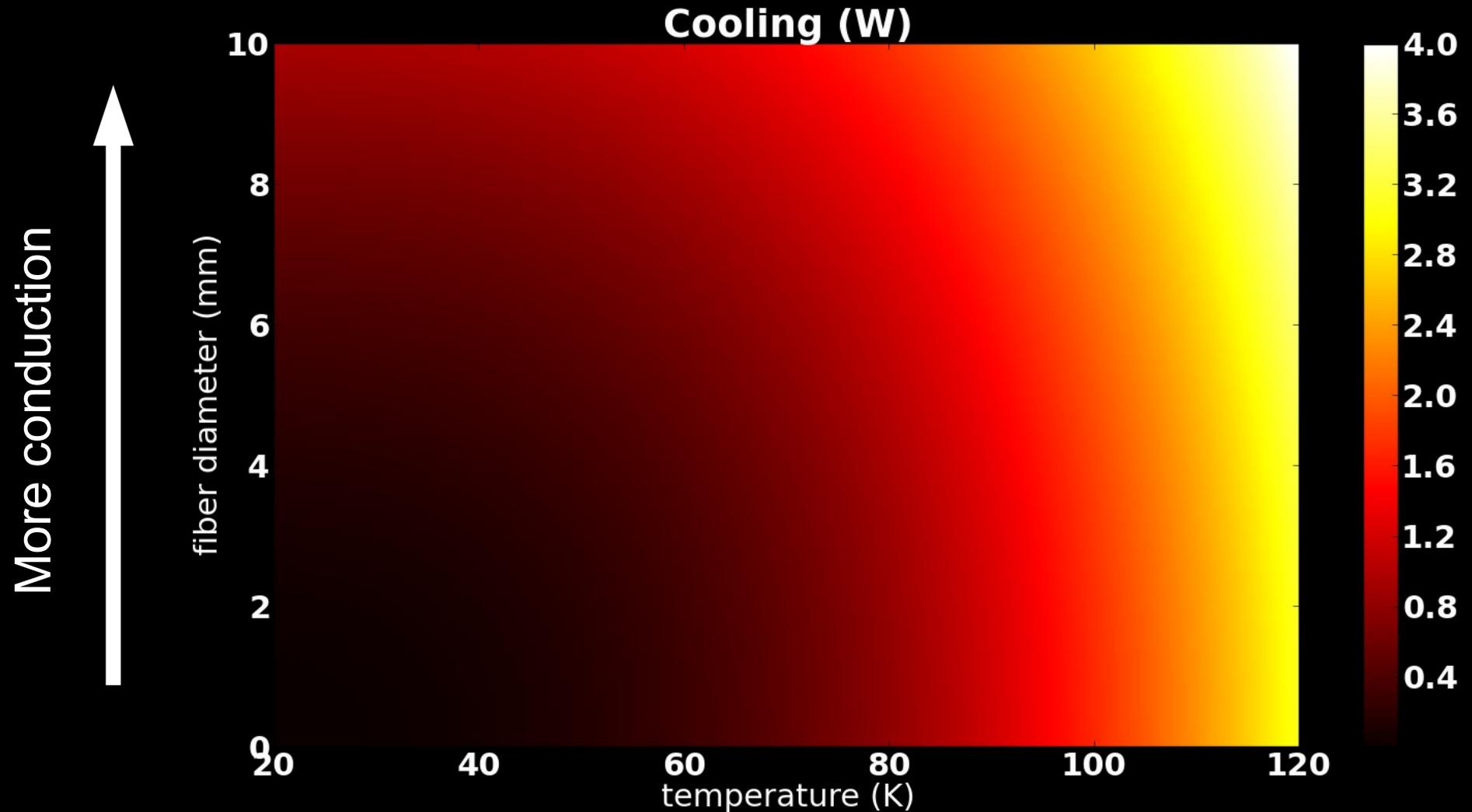
# Cooling methods



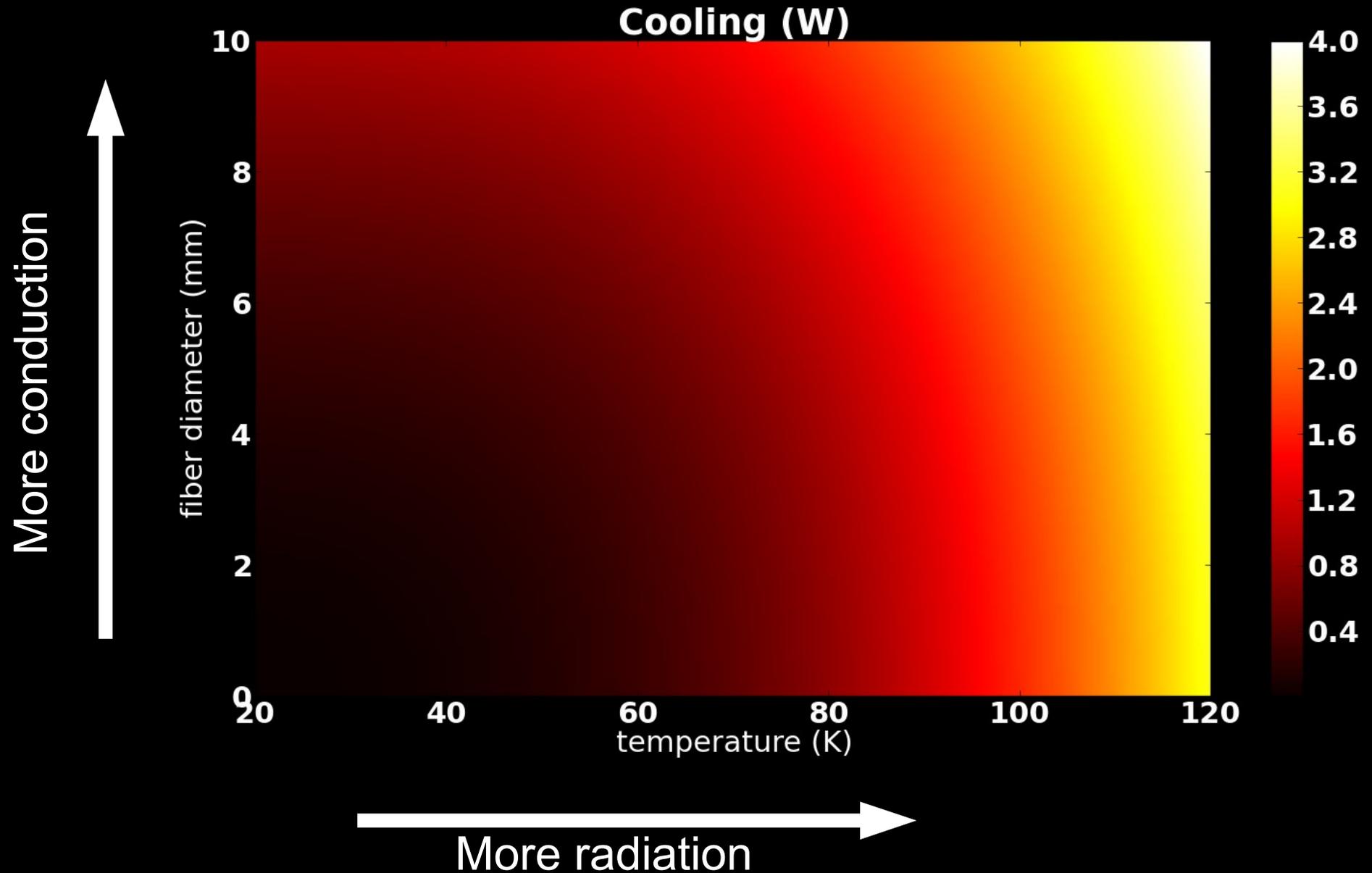
# Cooling power as a function of fiber diameter and temperature



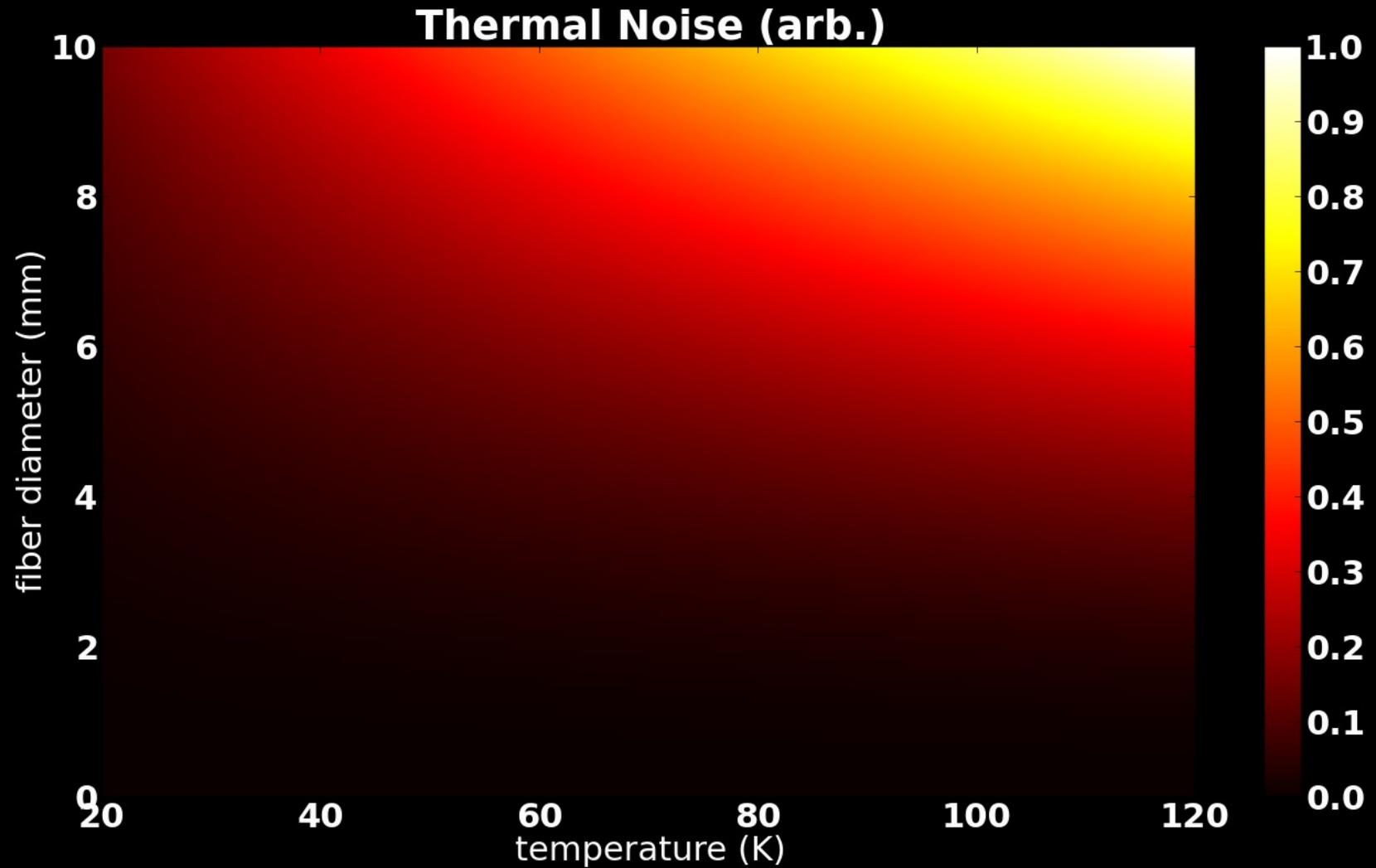
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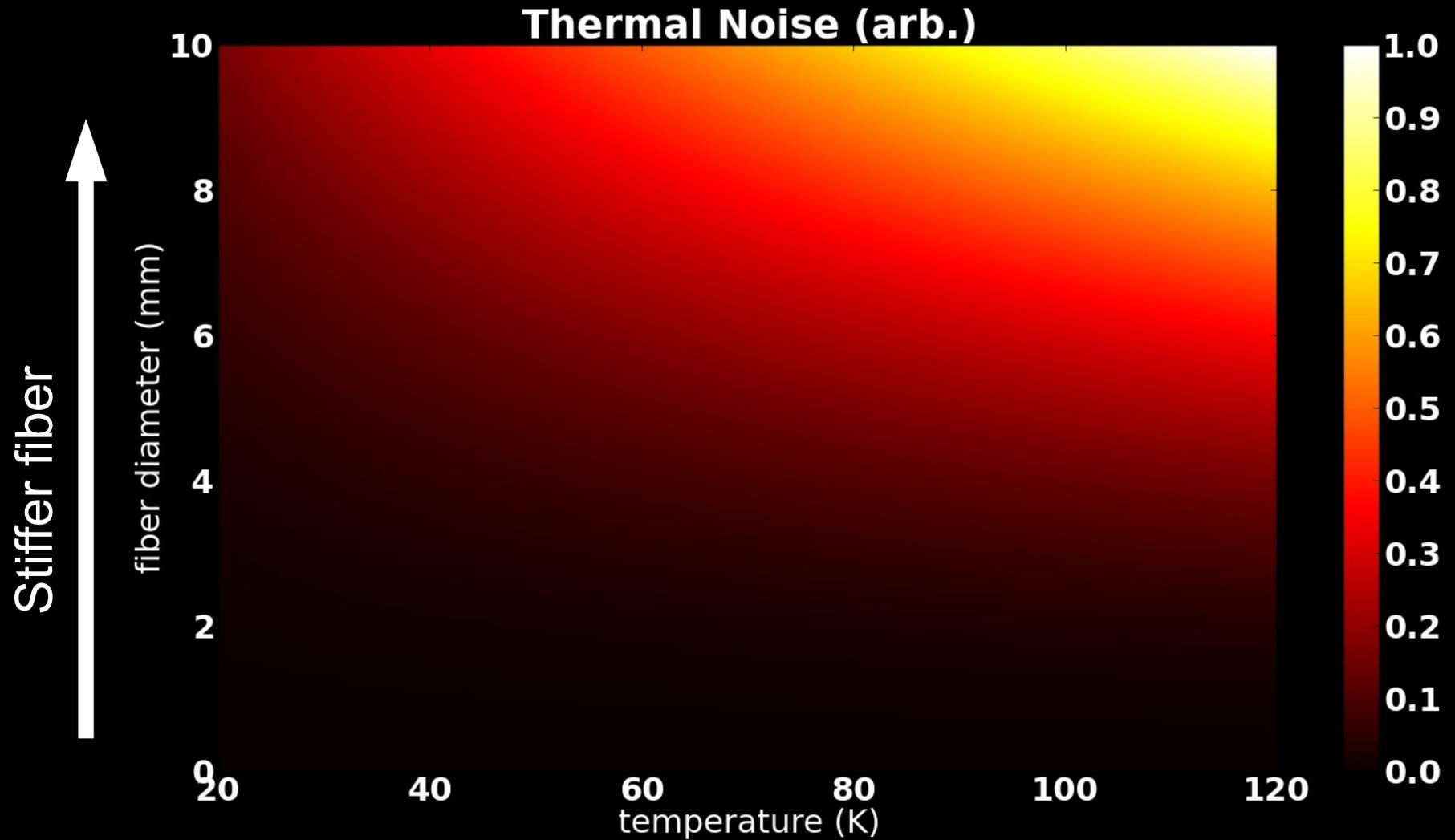
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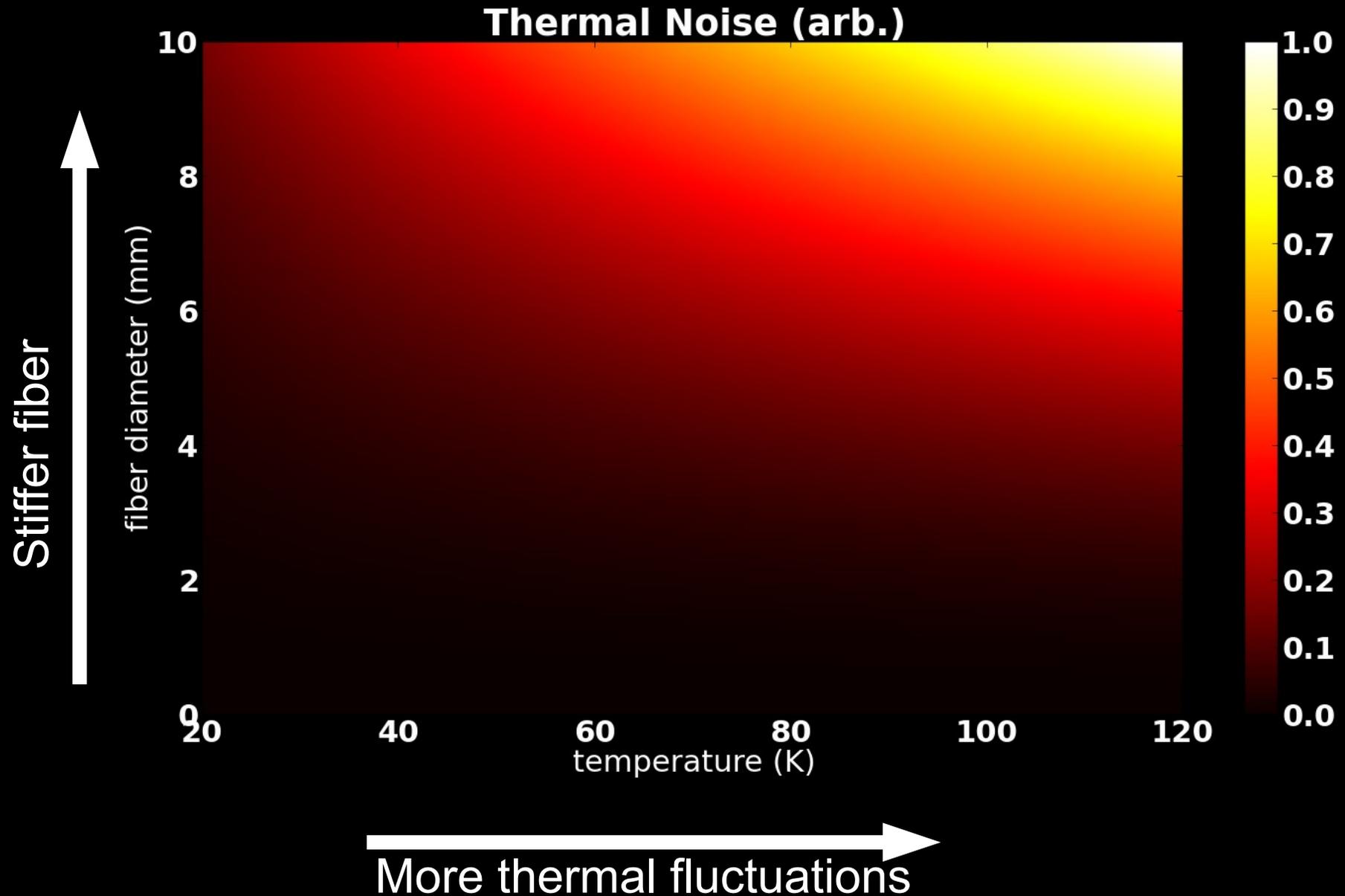
# Thermal Noise as a function of fiber diameter and temperature



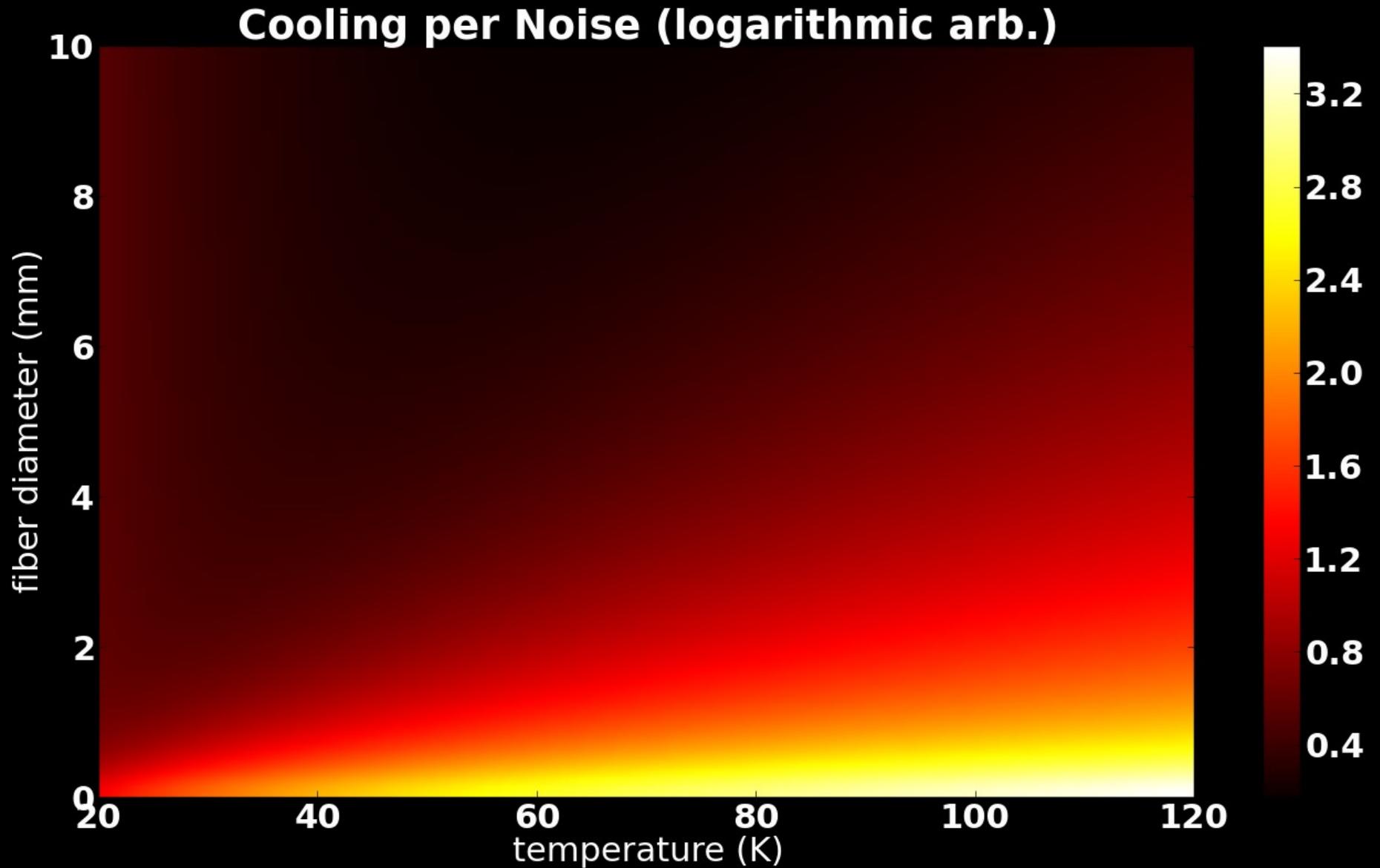
# Thermal Noise as a function of fiber diameter and temperature



# Thermal Noise as a function of fiber diameter and temperature



# Cooling vs Noise trade-off



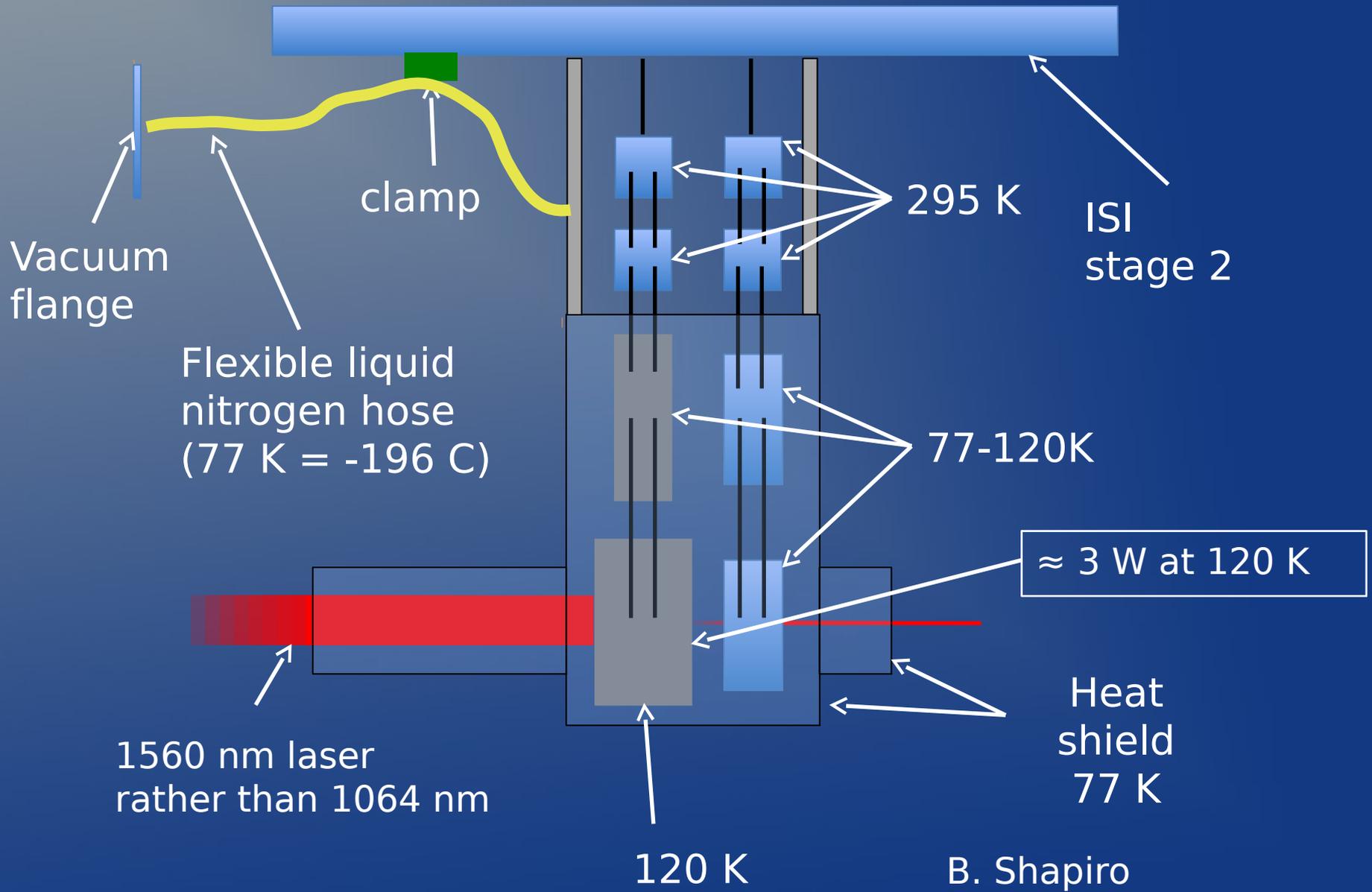
# Temperature and cooling trade-off

- Differing approaches
- KAGRA - Sapphire 20K Thick fibers (higher thermal noise than thin)
- ET - Silicon 20K Low power (requires two independent interferometers)
- LIGO3 Blue - Silicon 120K radiative cooling (not as cold)

# 120K Silicon

- The only way for one detector to achieve significant broadband improvement over advanced detectors (high power AND low thermal noise)
- Low frequency improvement due to thermal noise (zero thermo-elastic noise)
- High frequency improvement due to high stored arm power
  - 120K Silicon has higher thermal conductivity than copper, thermal distortion greatly reduced

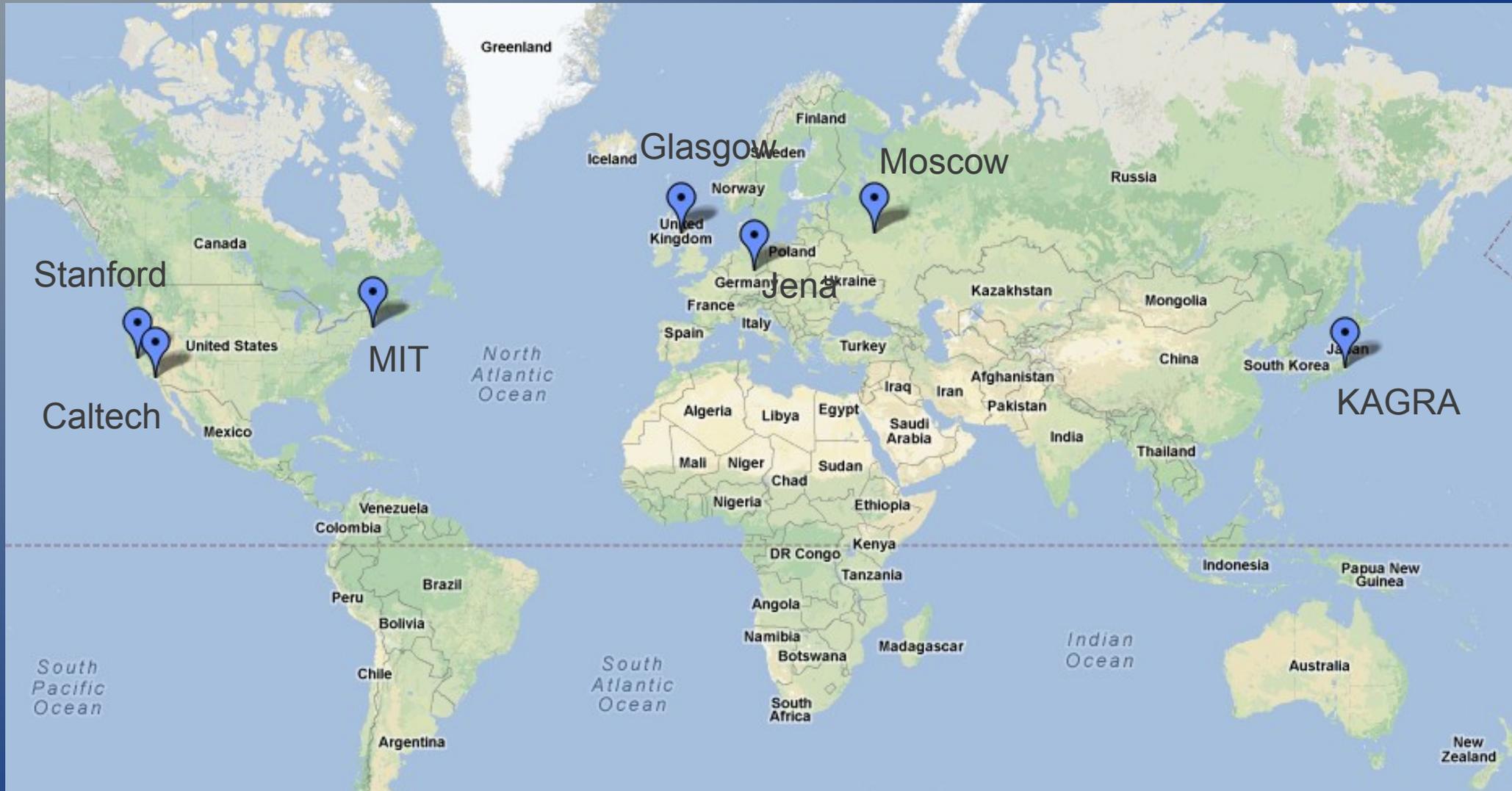
# Silicon Quad Suspension



# Some unknowns

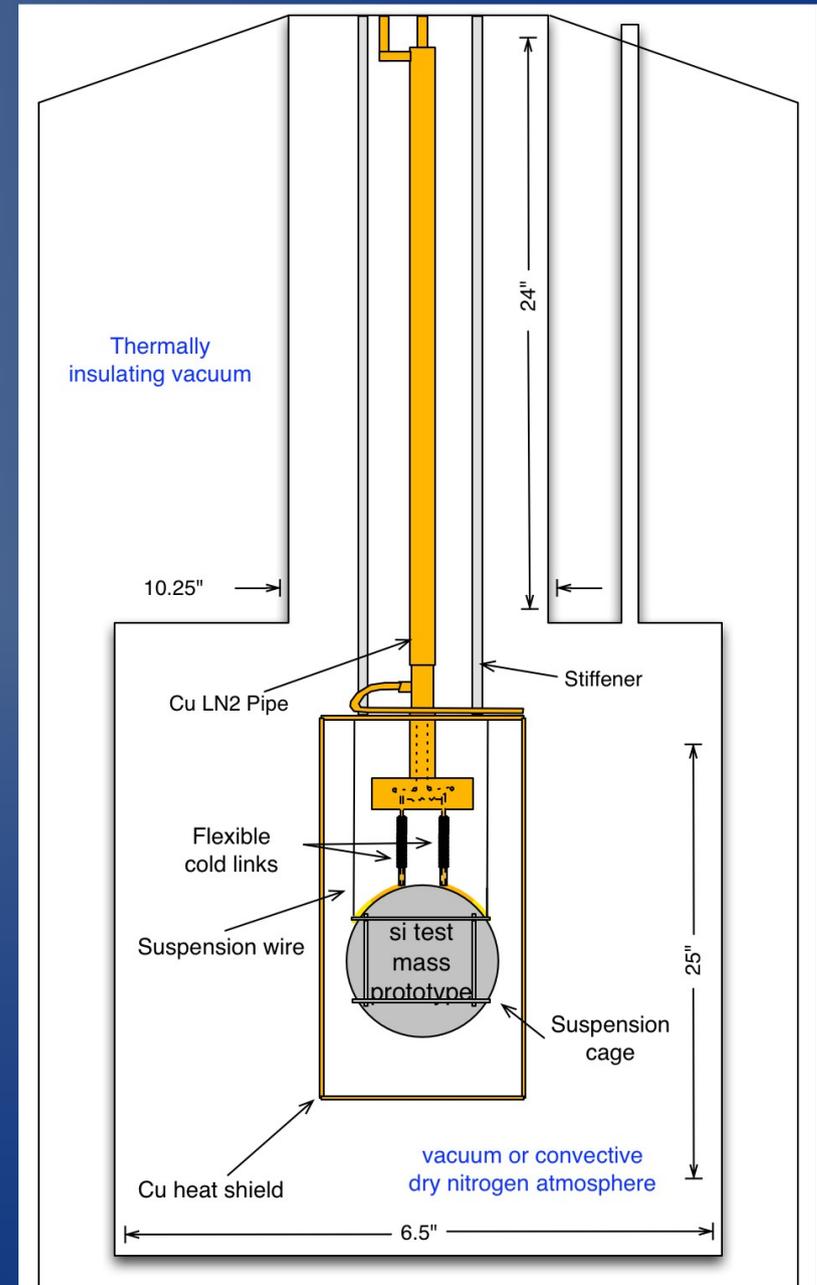
- Initial cooling method
- Laser techniques at 1550nm
- High emissivity coating
- How to mitigate surface losses
- Bonding of silicon to silicon
- How to create high purity masses at LIGO3 scale
- Unknown unknowns

# Who's working on it?



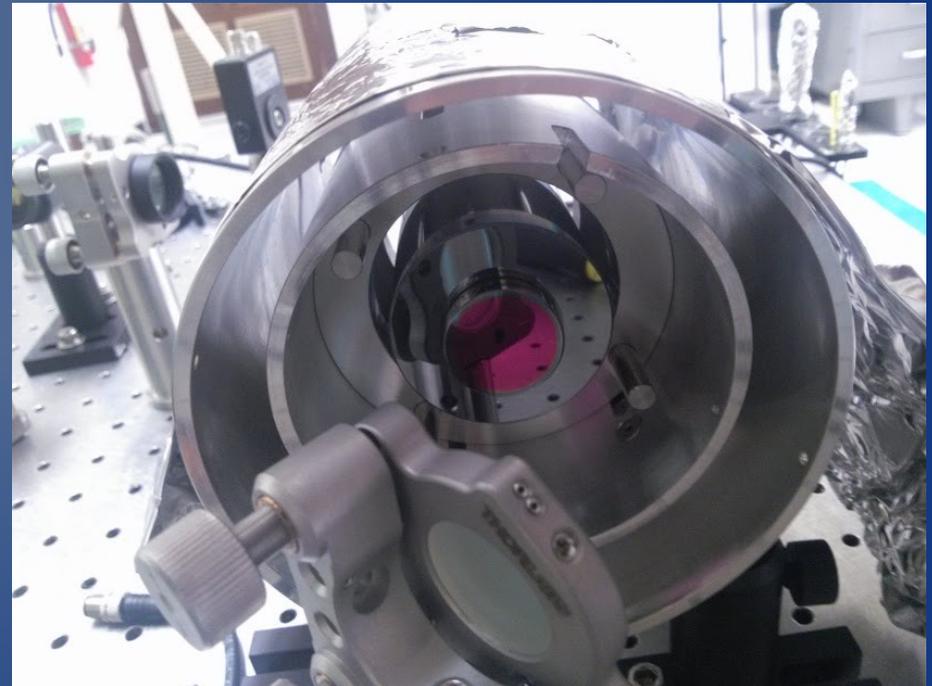
# Initial test mass cooling @ Stanford

B. Shapiro

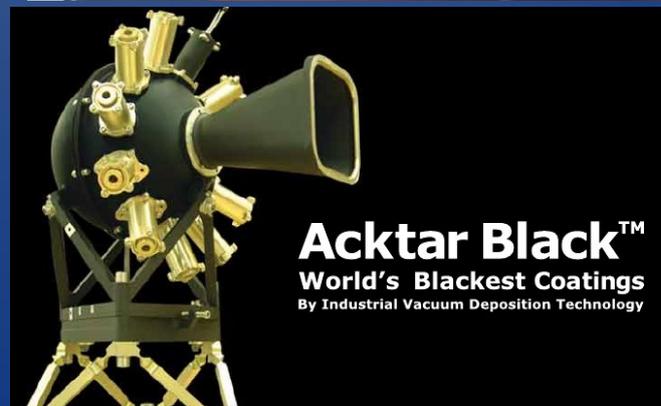
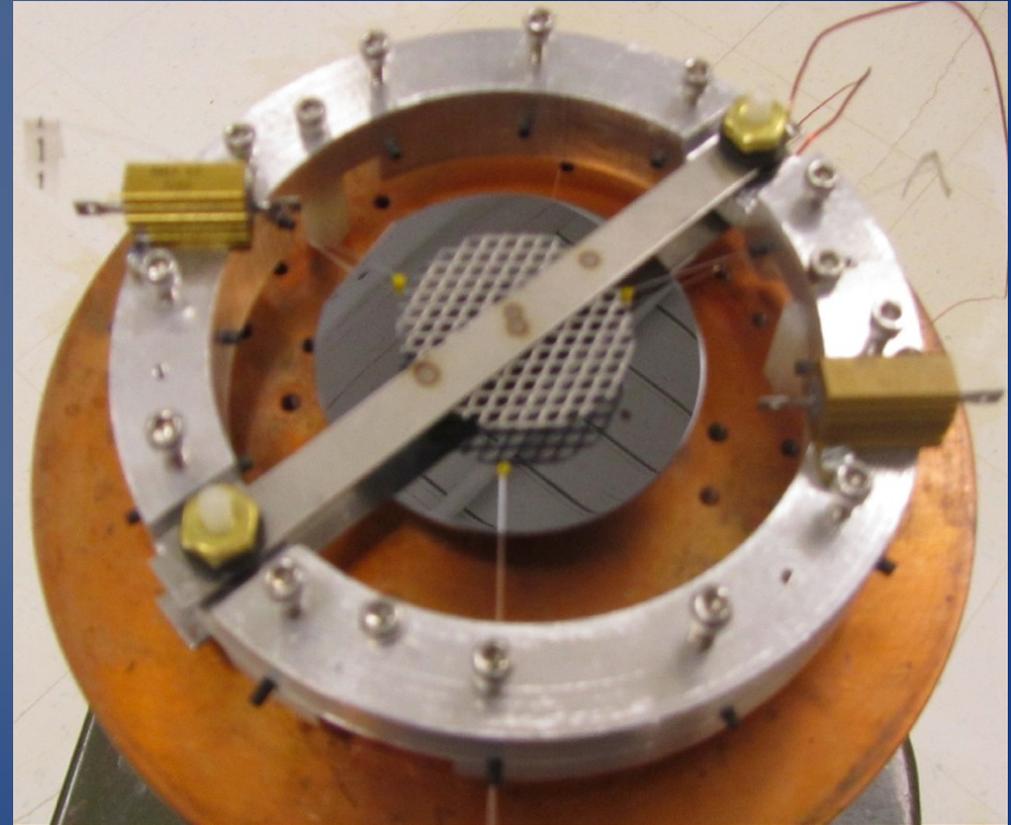
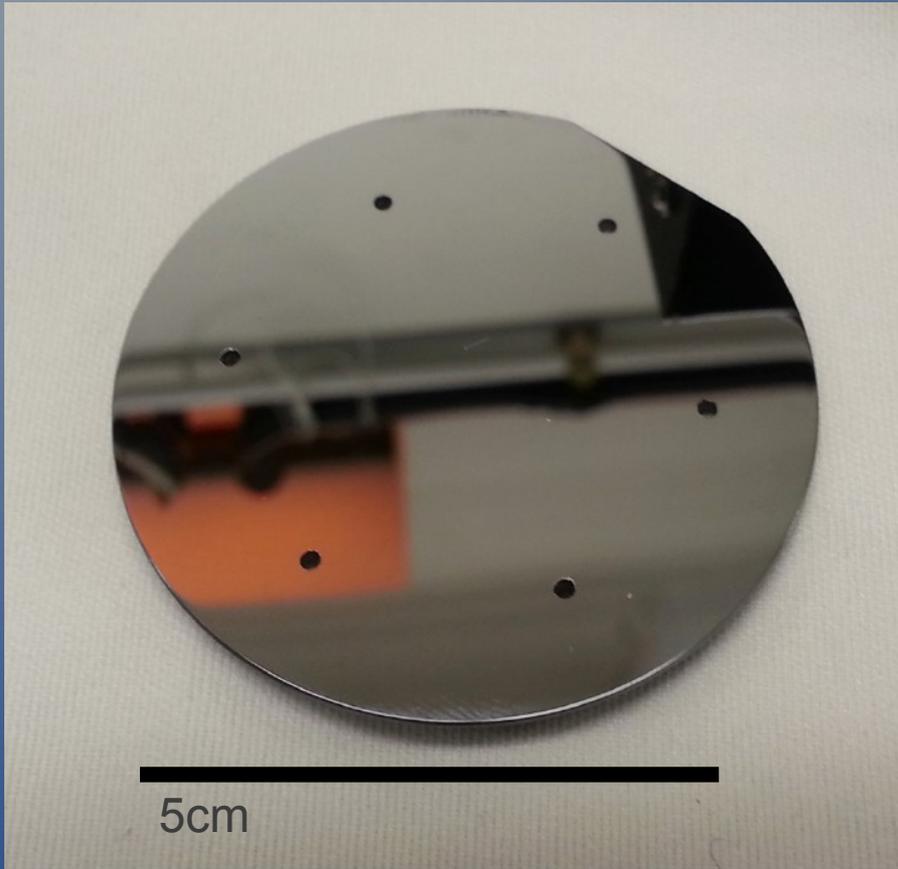


# Cryogenic Reference Cavities @Caltech

- Provides experience for many relevant technologies
- Ultra-stable DC frequency reference
- Potentially interesting system for studying macroscopic quantum mechanics

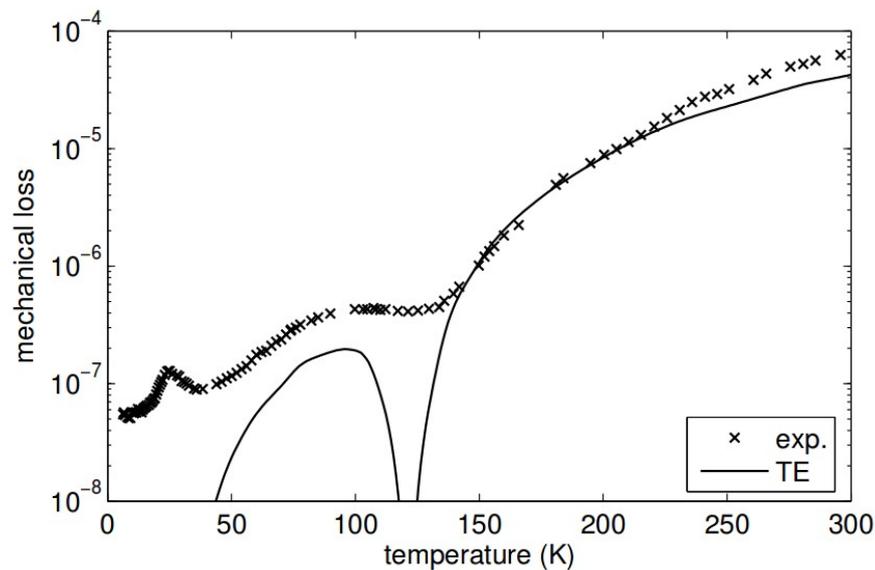


# High emissivity coating experiment @MIT



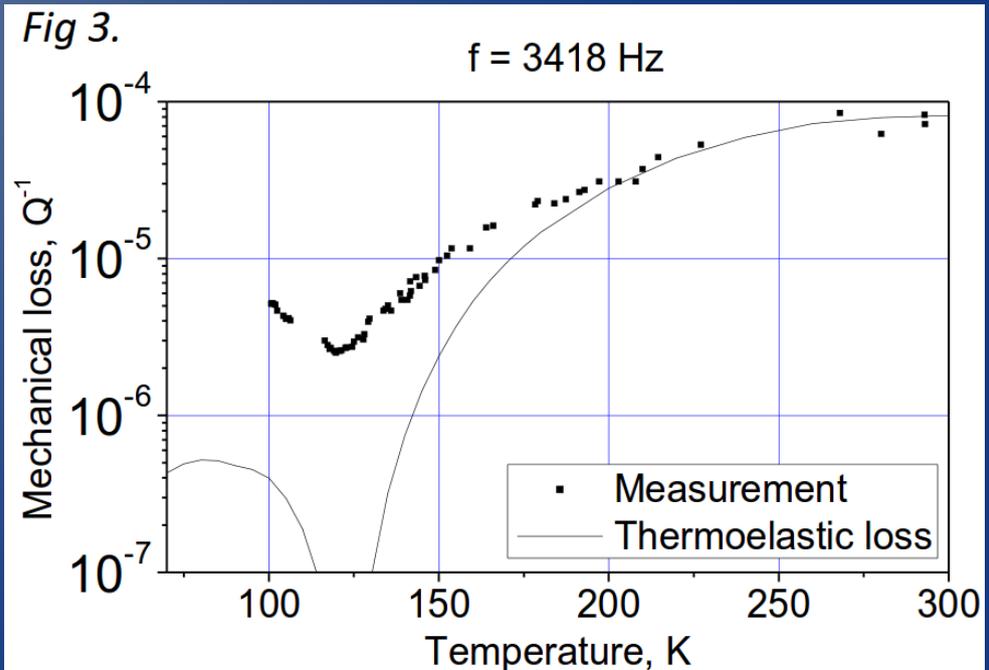
# Losses in Silicon Samples @Jena/Glasgow/Moscow

- Losses in silicon samples still limited by surface quality or other dirty physics
- More tests required to hit the true loss limit



(c) 19980 Hz

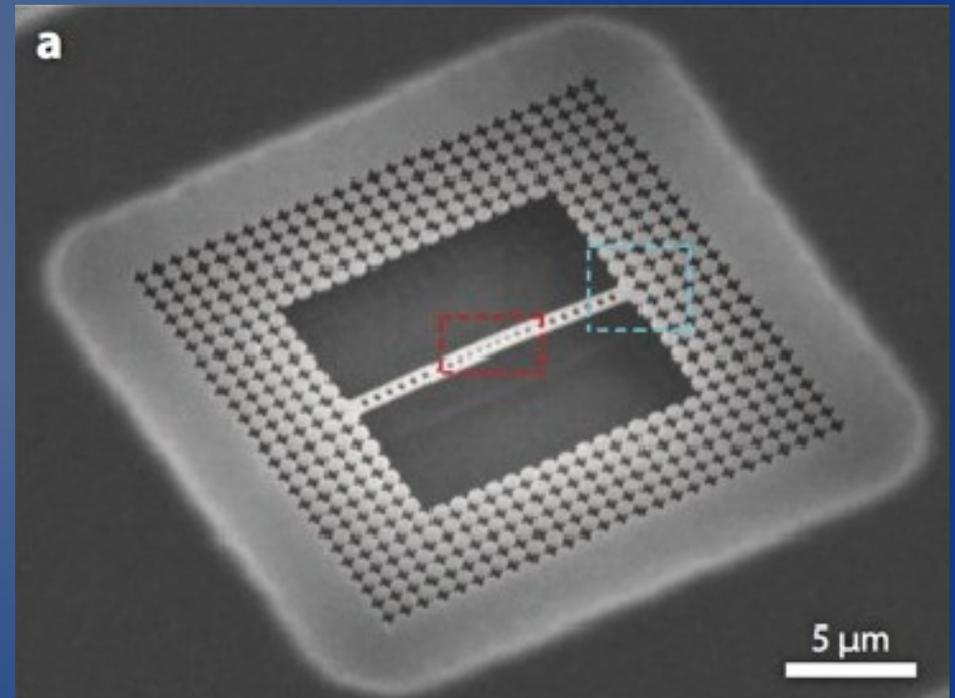
Silicon Cantilever @ Jena/Glasgow  
Nawrodt et al. arXiv:1003.2893



Silicon Wafer @ Moscow  
Prokhorov, Mitrofanov

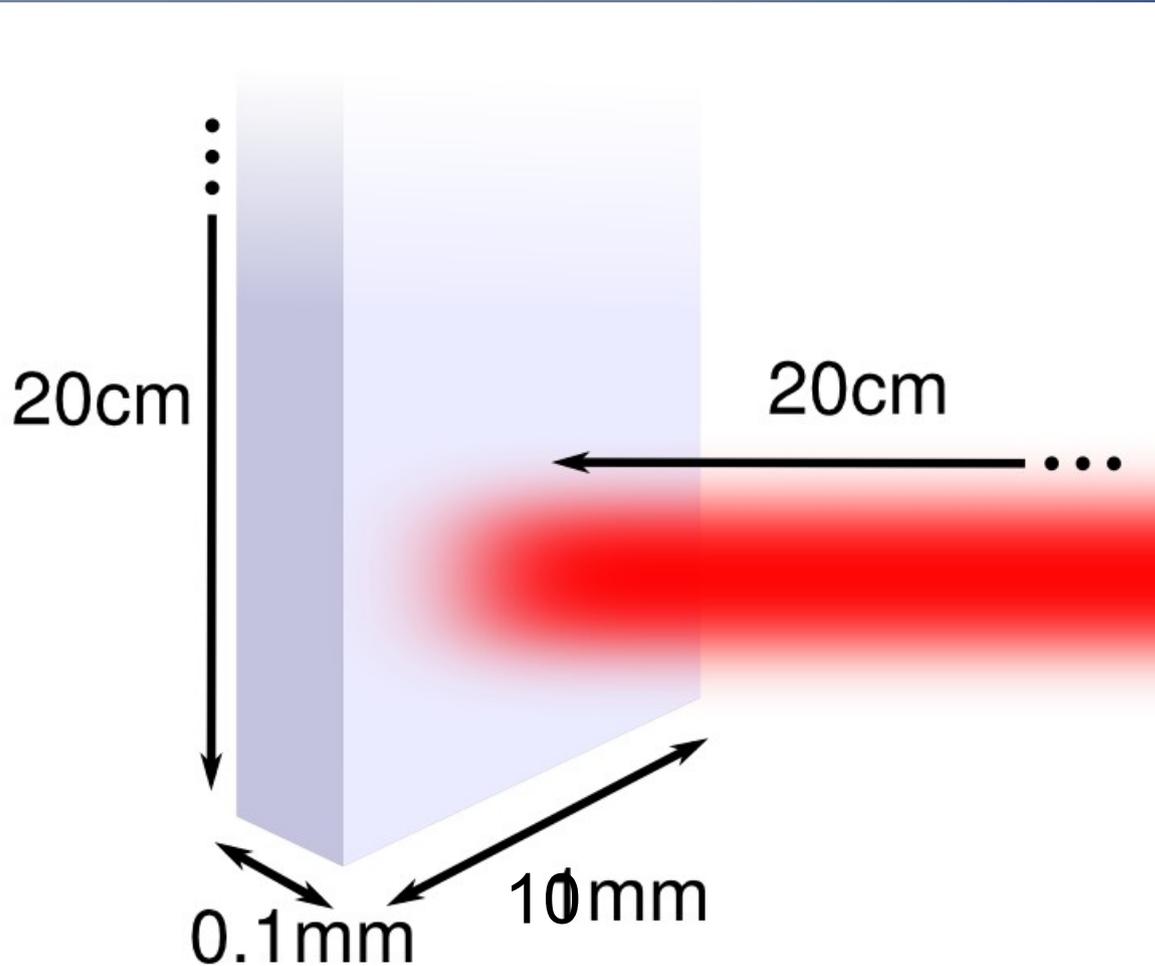
# Oskar Painter Silicon Etching @Caltech

- Micro-mechanical silicon structures
- Demonstrated laser cooling of mechanical oscillator to quantum ground state
- Willing to collaborate with us on Silicon etching techniques for low surface losses



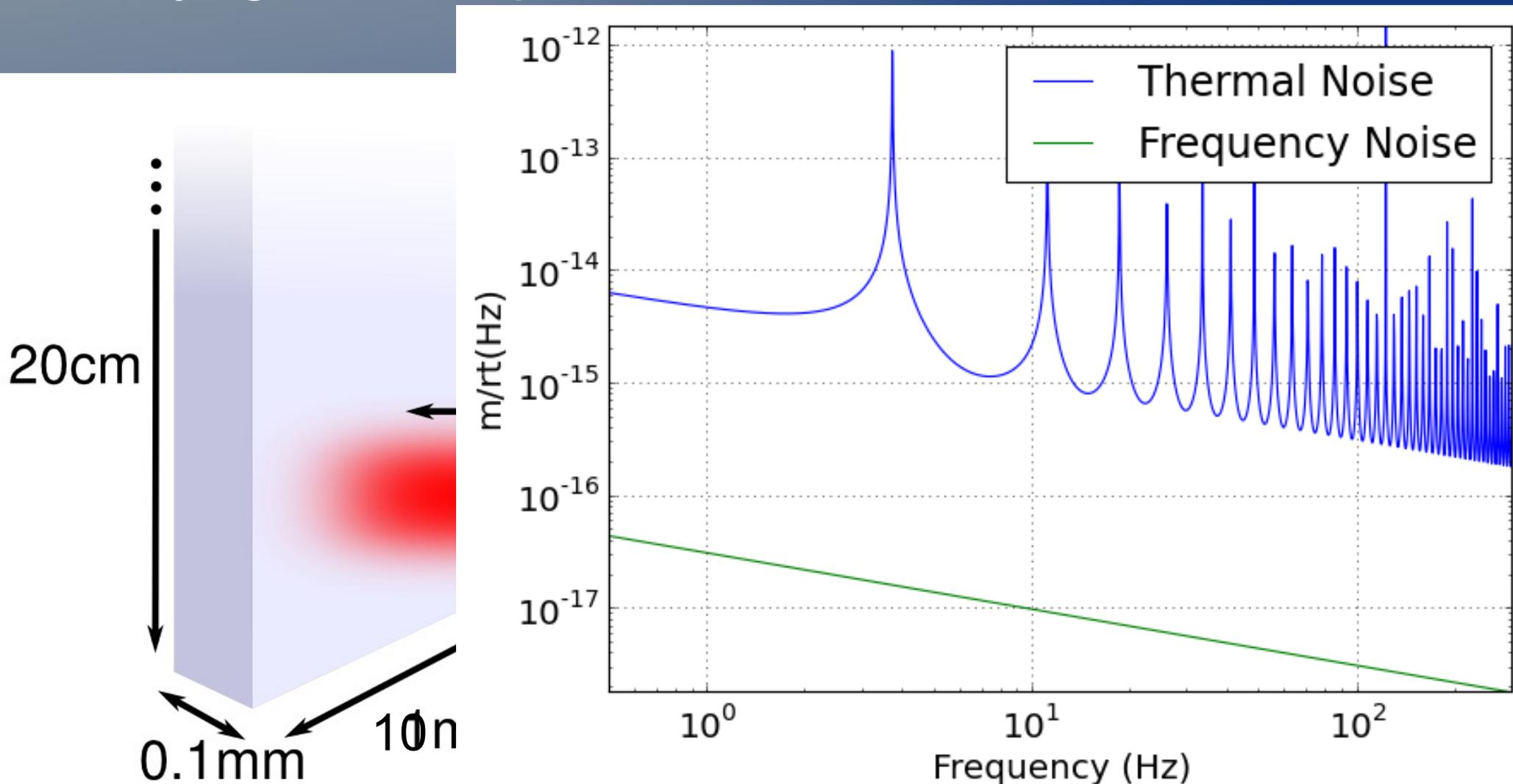
# Silicon macro-mechanical cantilever @Caltech

- Allows direct measurement of thermal noise at cryogenic temperatures



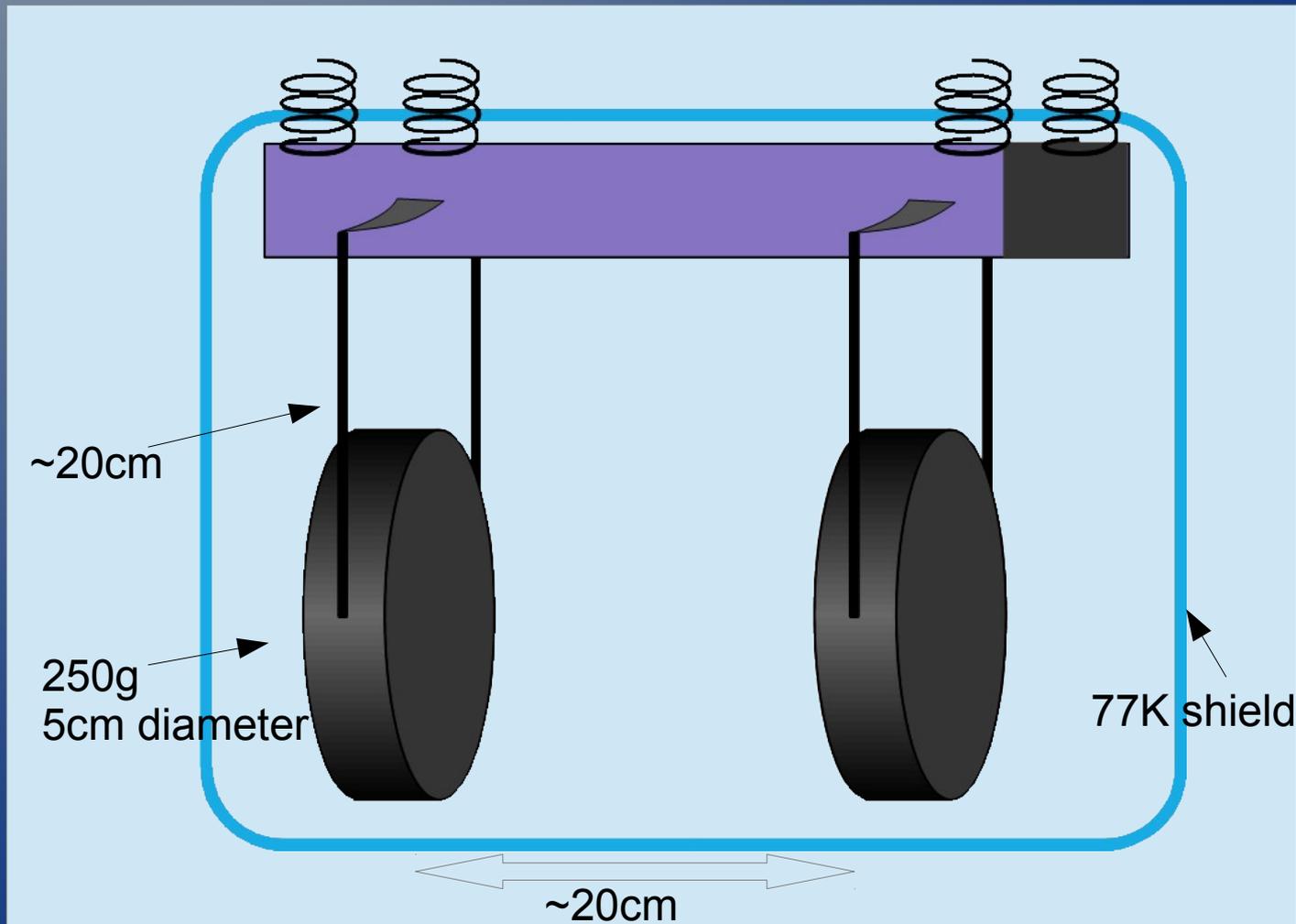
# Silicon macro-mechanical cantilever @Caltech

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# Prototype suspension @Caltech

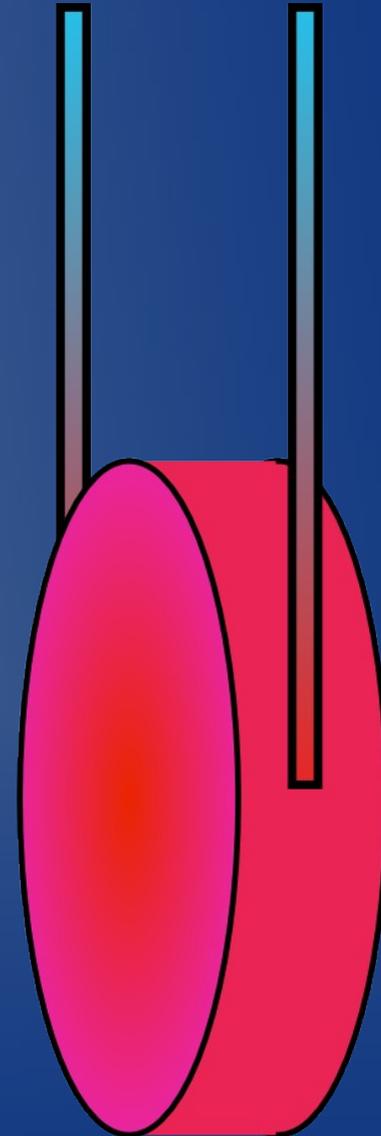
- Good for developing fabrication techniques, though likely difficult to see thermal noise





# An Open Question in Thermal Noise Modeling: Non thermal equilibrium

- Cryogenic interferometers will have thermal gradients along the suspension wire
- It is not yet settled how to calculate the noise correctly



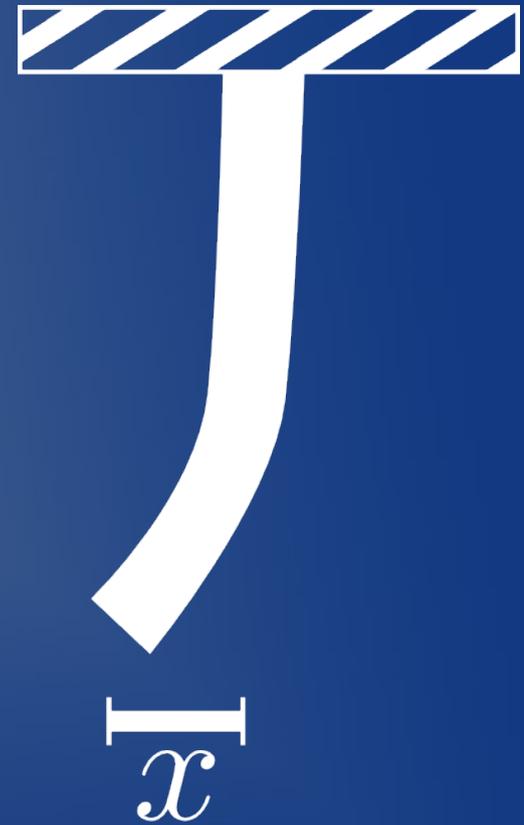
# The Thermal Noise Cookbook: Single degree of freedom oscillator

- Use the expression from Saulson (1990)

$$S_x(f) = \frac{4k_B T \omega_0^2 \phi}{m\omega([\omega_0^2 - \omega^2]^2 + \omega_0^4 \phi^2)}$$

# Multiple degree of freedom system

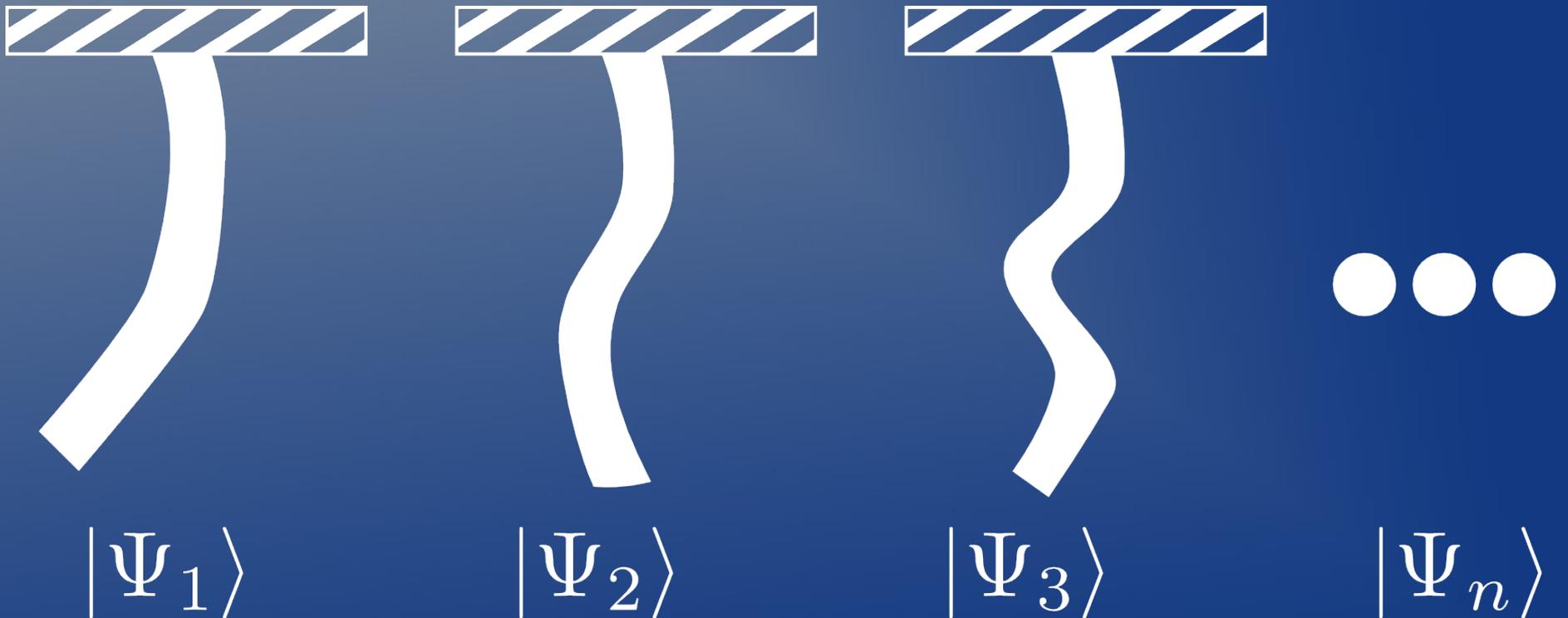
- System may have many 'modes' but you are only interested in the thermal noise measured by some degree of freedom ( $x$ )
- There are multiple recipes in the Cookbook to approach this
- My example: the position of the end of a beam



# Multi-DOF system: the modal approach

Step 1:

Break the system into the normal mode DOFs



# Muti-DOF system: the modal approach

Step 2:

Find the thermal amplitude of each mode (from Saulson)

$$S_{\Psi_n}(f)$$

# Muti-DOF system: the modal approach

Step 3:

Find overlap of each mode with desired DOF,  
and combine them.

$$S_x(f) = \sum_n \langle x | \Psi_n \rangle^2 S_{\Psi_n}(f)$$

- Major drawback: not straightforward to use when loss angle is function of position

# The Levin approach

- Uses the fluctuation dissipation theorem while treating the system holistically

Step 1.

Convert your desired DOF into a conjugate force and apply that force to the system



# The Levin approach

Step 2.  
Calculate the dissipation  
'experienced' by the force



$$W_{\text{diss}} = 2\pi f \int_V \rho(\vec{r}) \phi(\vec{r}) d^3 \vec{r}$$

# The Levin approach

Step 3.

Plug into formula

$$S_x(f) = \frac{4k_B T W_{\text{diss}}}{\pi^2 f^2 F_0^2}$$

$$W_{\text{diss}} = 2\pi f \int_V \rho(\vec{r}) \phi(\vec{r}) d^3 \vec{r}$$

Sum over volume elements instead of sum over modes

# A new recipe

- Use a Levin-like approach to find the coupling of each volume element to your degree of freedom
- Allow each volume element to have its own temperature (and loss angle)

# The Gradient Effect

- Does the thermal gradient contribute an additional effect to the noise?

$$S_x(f) = \frac{4k_B}{\pi f F_0^2} \times$$

$$\int_V \rho(\vec{r}) \left[ T(\vec{r}) + H(\vec{\nabla} T(\vec{r})) \right] \phi(\vec{r}) d^3 \vec{r}$$

# The Need to Complete the Cookbook

- The problem of thermal noise in non-thermal equilibrium is very relevant for cryogenic interferometers
- Very interesting results presented here suggest some (large) effects due to thermal gradients
- Talk by Claudia Lazzaro and poster by Rossana De Gregorio

# Fin

Thanks to:

Caltech: Rana Adhikari, Alastair Heptonstal, Norna Robinson, Eric Gustufson

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Stanford: Brett Shapiro

MIT: Rai Weiss, Lamiya Mowla