



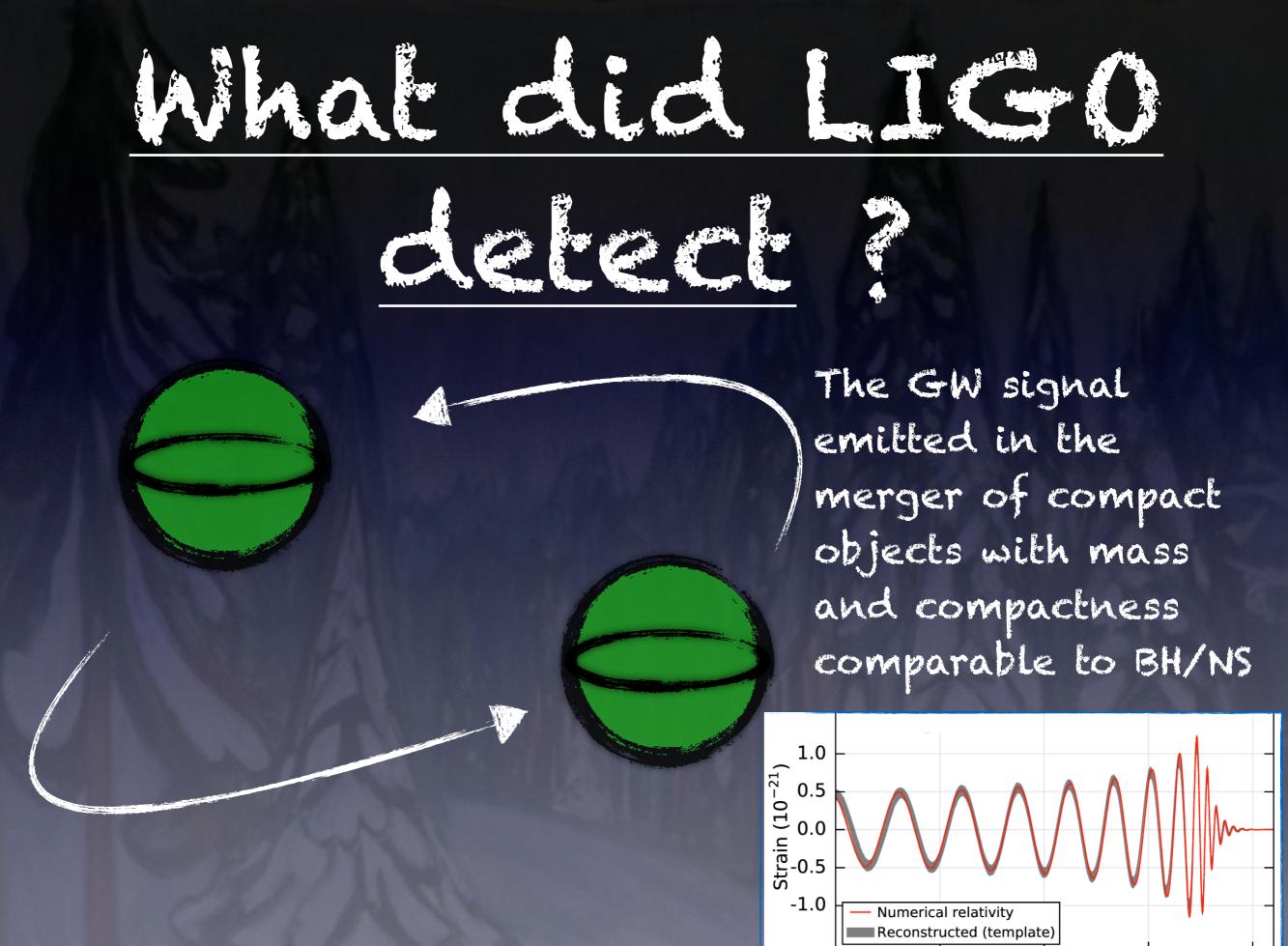
Alfredo Urbano

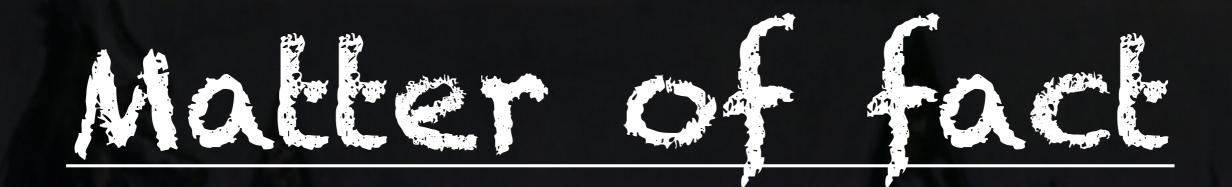
INFN, sez. di Trieste and CERN

Les Rencontres de Physique de la Vallée d'Aoste La Thuile 2018









Everything seems consistent with General Relativity



General Relativity cries for a UV completion

... but I will not discuss GW beyond GR...



And we need new particles beyond the standard Model of particle physics... (isn't it?)



Could they change standard GW signals?



I Correction to observable quantities aka "direct detection"

II Statistical effects aka "indirect detection"

III Additional material I will not cover: BH superradiance and PBHs

I.a) Change in the waveform during the inspiral phase

I.b) Change in the waveform during the ringdown phase

I.c) Change in the frequency evolution during the inspiral phase (but I do not have time to discuss)

I.a) Change in the waveform during the inspiral phase

Goal: Highlight the presence of "new physics corrections" compared to what we expect in standard General Relativity

I.a) Change in the waveform during the inspiral phase

In this example I will consider the merger of two black holes.

And what new physics you have in mind? Simple U(1) dark QED $\mathcal{L}_V = -\frac{1}{4}V_{\mu\nu}V^{\mu\nu} + \frac{1}{2}m_V^2 V$

$$\mathcal{L}_{\chi} = ar{\chi} \left[\gamma^{\mu} (i \partial_{\mu} - g' V_{\mu}) - g' V_{\mu} \right]$$

I.a) Change in the waveform during the inspiral phase

In this example I will consider the merger of two black holes.

ASSUMPTION

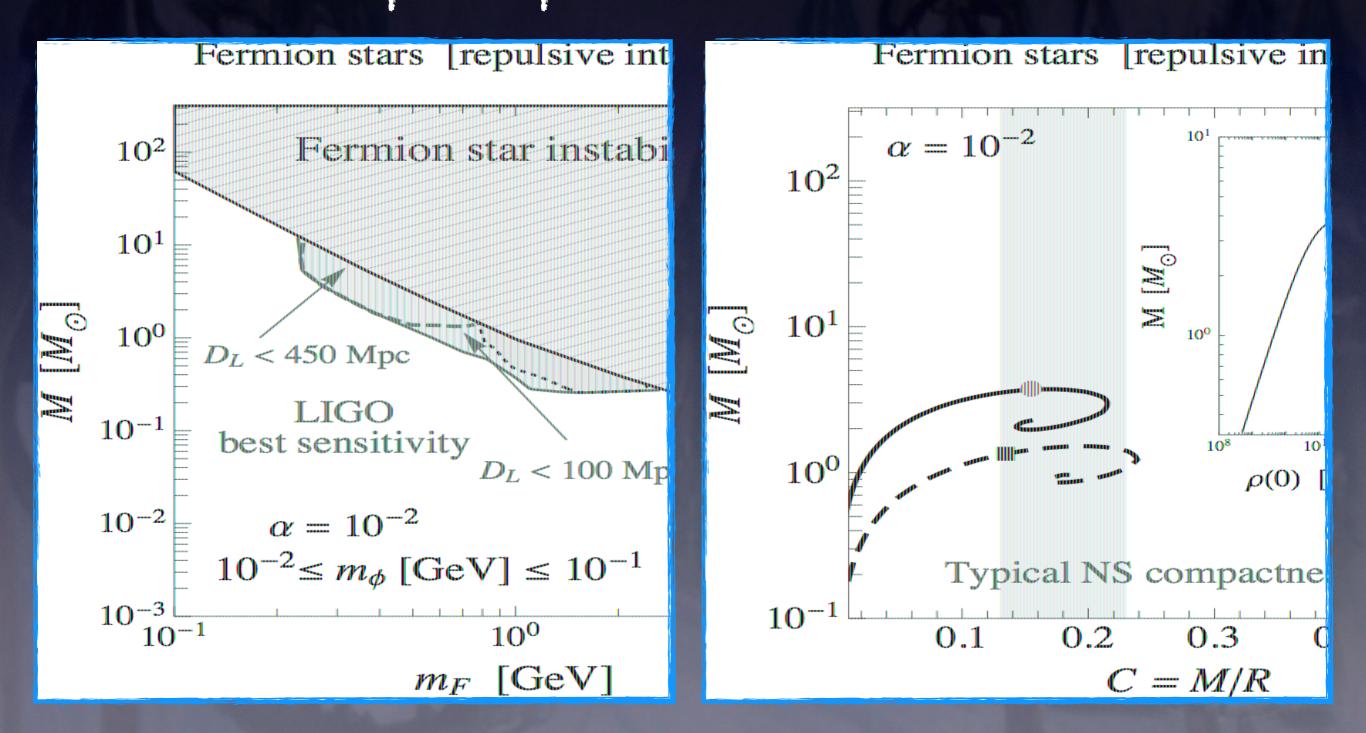
The new fermions condensate

to form

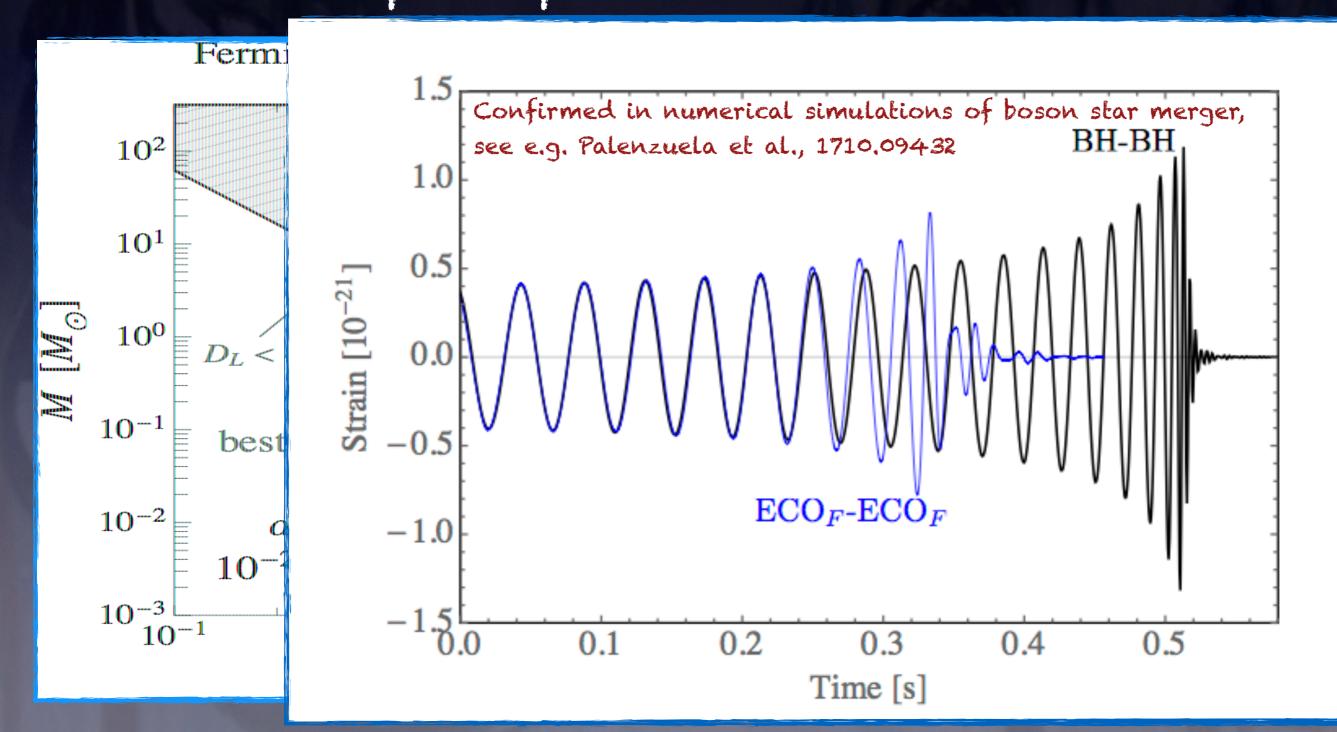
a "dark neutron star"

 $egin{array}{rcl} \mathcal{L}_V &=& -rac{1}{4}V_{\mu
u}V^{\mu
u}+rac{1}{2}m_V^2V \ \mathcal{L}_\chi &=& ar{\chi}\left[\gamma^\mu\left(i\partial_\mu-g'V_\mu
ight)ight. \end{array}$

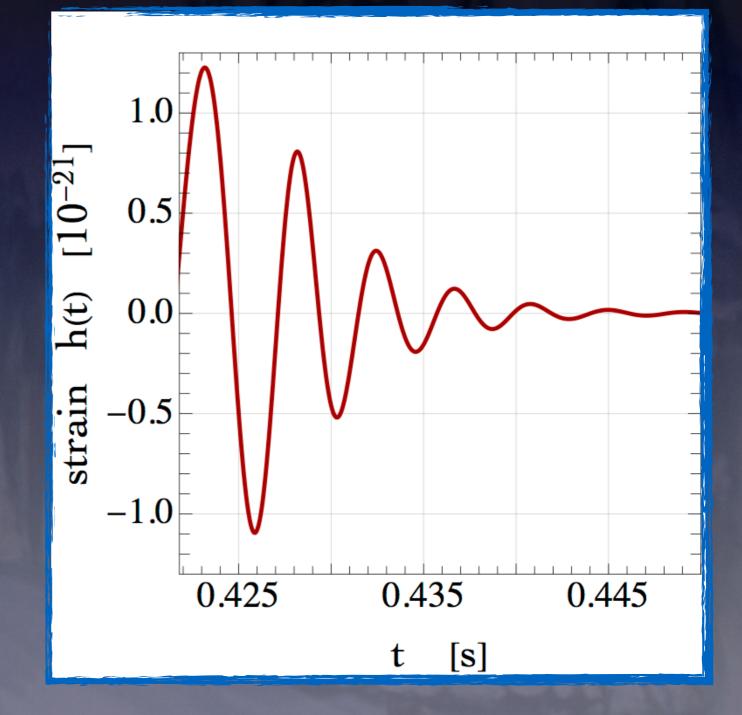
I.a) Change in the waveform during the inspiral phase



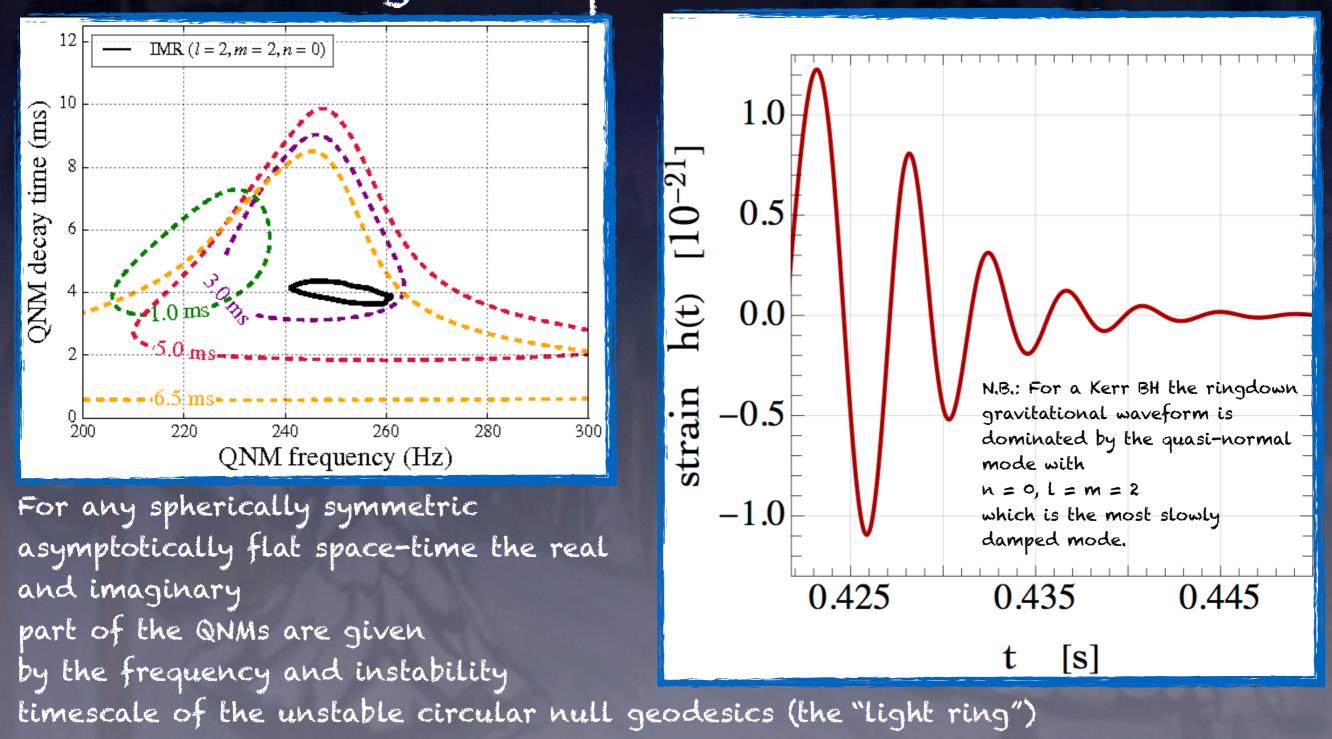
I.a) Change in the waveform during the inspiral phase



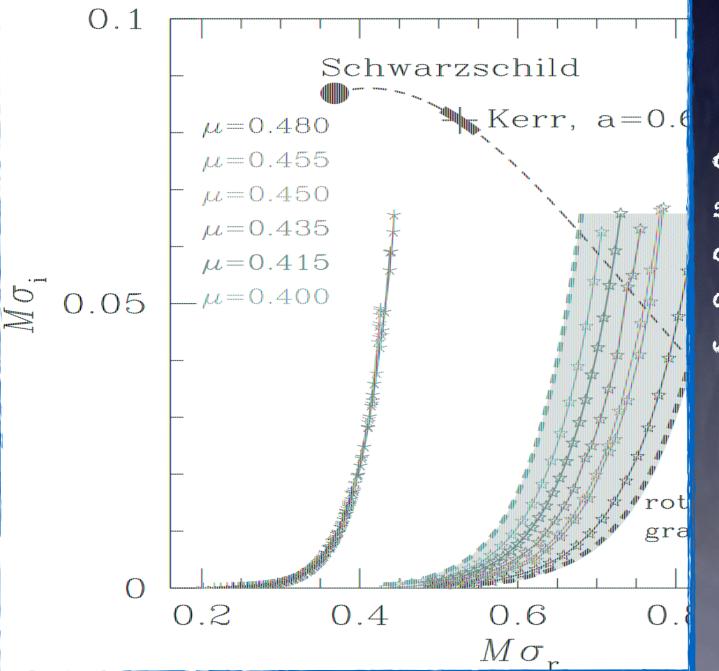
I.b) Change in the waveform during the ringdown phase



I.b) Change in the waveform during the ringdown phase

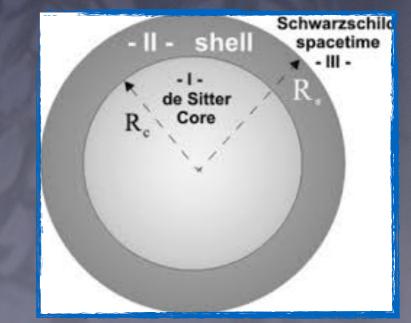


I.b) Change in the waveform during the ringdown phase chirentia

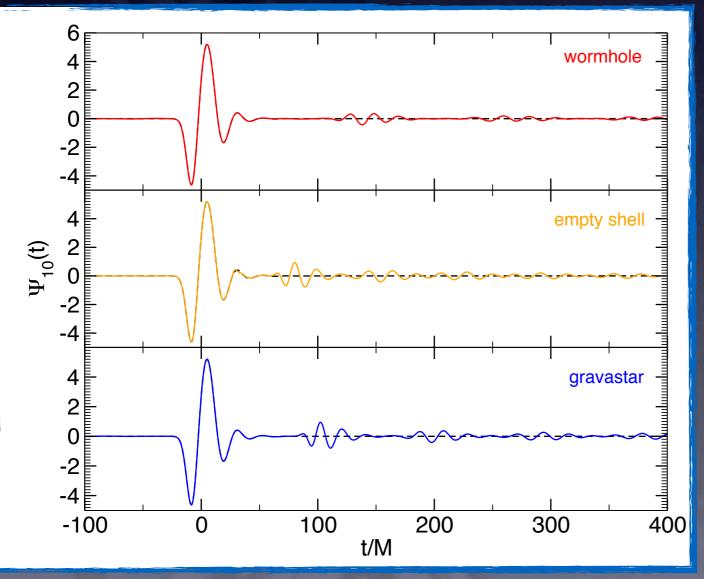


Chirenti and Rezzolla, 1602.08759

Gravastars are a toy model for the so-called "black-hole mimickers", objects that can be almost as compact as a black hole but without an event horizon.



I.b) Change in the waveform during the ringdown phase

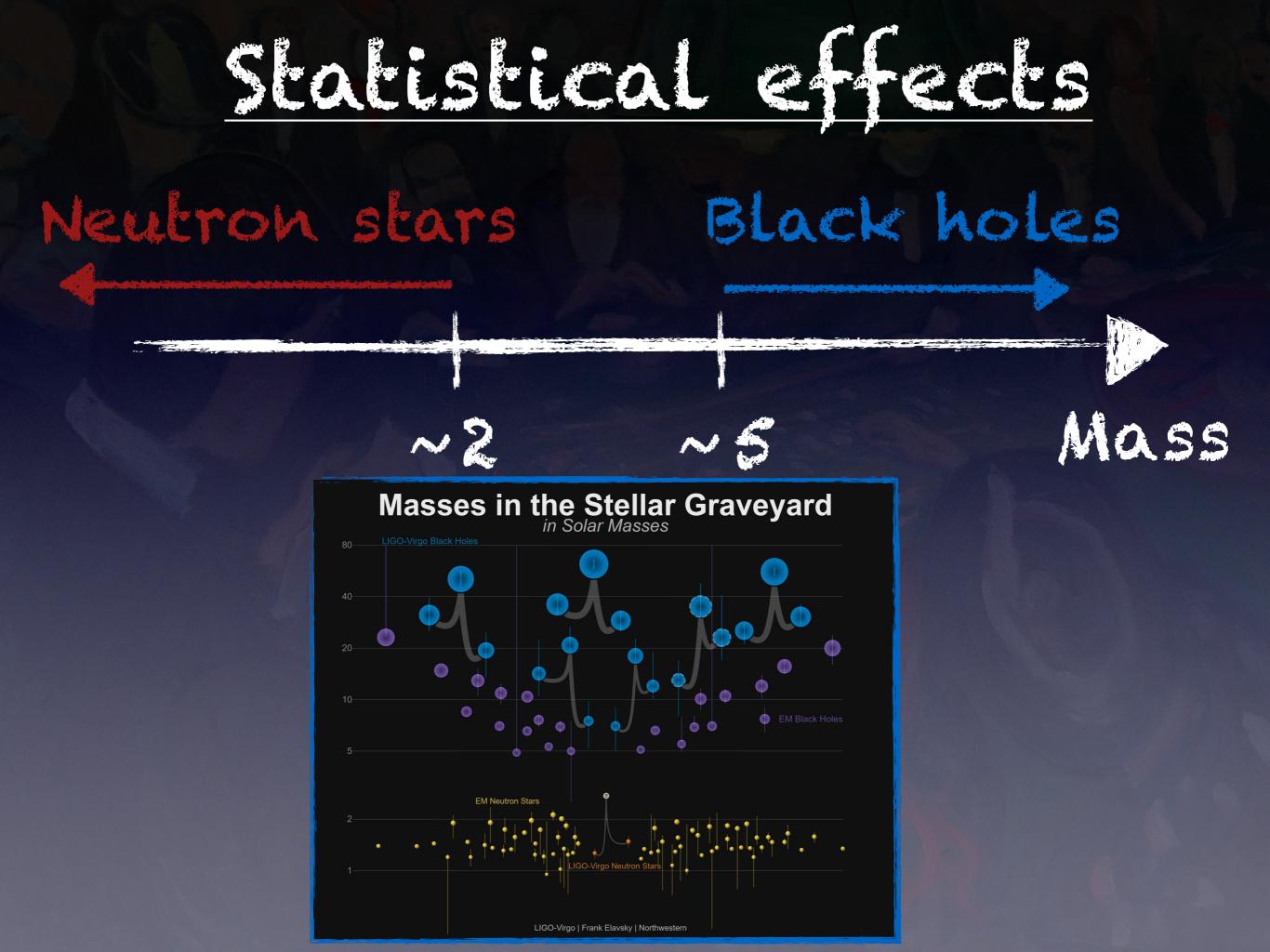


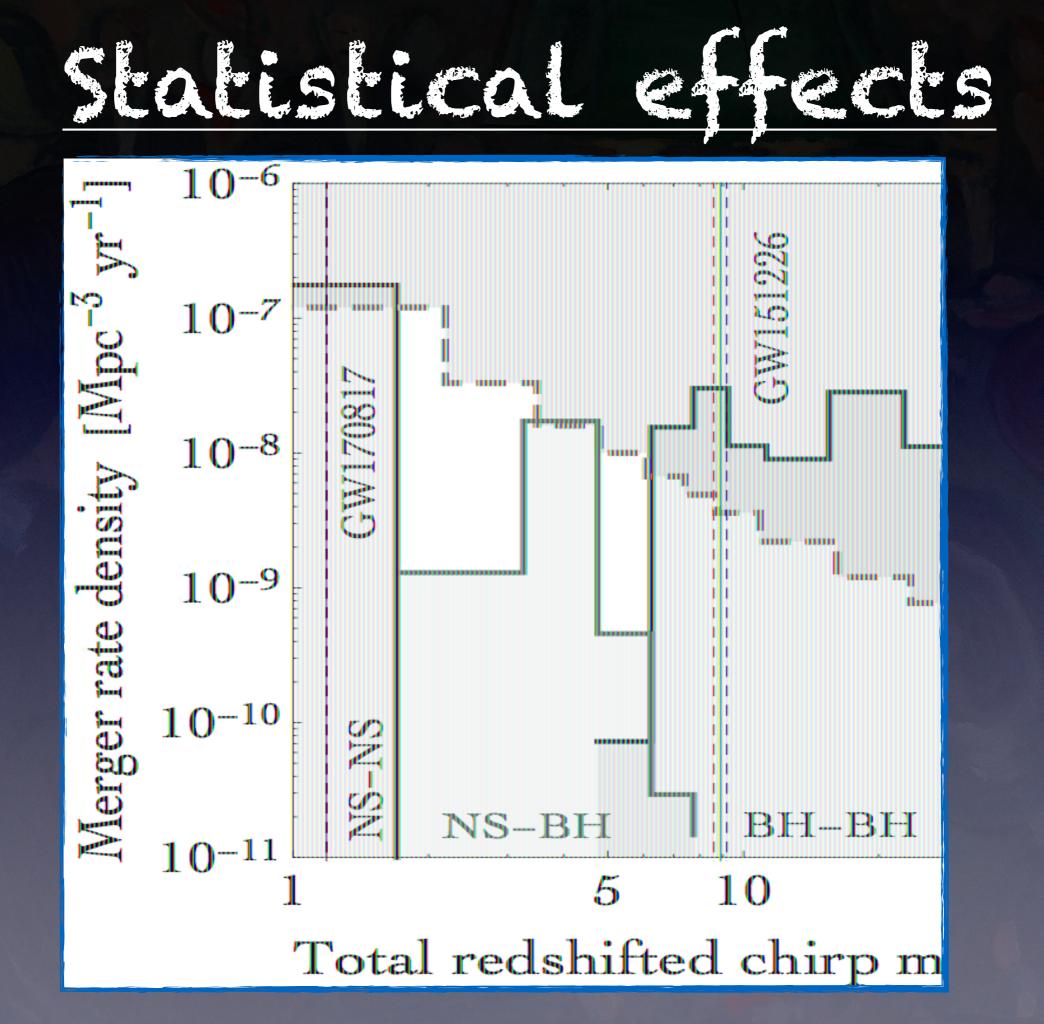
V. Cardoso et al., 1608.08637

Even in the limit in which the ECO is as compact as a black hole, we can still distinguish the former from the latter because of the presence of additional pulsations ("echoes" of ECOs) in the ringdown phase.

Statistical effects

Simulations from "The Synthetic Universe project", http://www.syntheticuniverse.org/





Conclusion/Open questions

Conclusion/Open questions

i) The hunking season is now open

ii) ECOs: formation and stability + numerical simulations

iii) Many things I didn't have time to discuss... (BH superradiance, PBHs, et cetera...)



I.c) Change in the frequency evolution during the inspiral phase

$$f_{\rm GW}(t) \sim rac{1}{\pi} \left(rac{G_N M_C}{c^3}
ight)^{-5/8} \left(rac{G_N M_C}{c^3}
ight)^{-5/8}$$

$$m_V^{-1} = \mathcal{O}(10 - 1000)$$

 $m_V = \mathcal{O}(10^{-11} - 10)$

ASSUMPTION

The charged DM particles are trapped by a NS during its lifetime The NS carries a net "dark charge"

