

Chiral three-nuclear forces up to $N^4\text{LO}$

Hermann Krebs
Ruhr-Universität-Bochum

The 8th International Workshop on Chiral Dynamics
Pisa, Italy

July 1, 2015

With V. Bernard, E. Epelbaum, A. Gasparyan, U.-G. Meißner



LENPIC

Low Energy Nuclear Physics International Collaboration



Binder, Calci, Hebeler,
Langhammer, Roth



Furnstahl



Maris, Potter, Vary



Epelbaum, Krebs



Bernard



Golak, Skibinski,
Topolniki, Witala



Meißner



Nogga

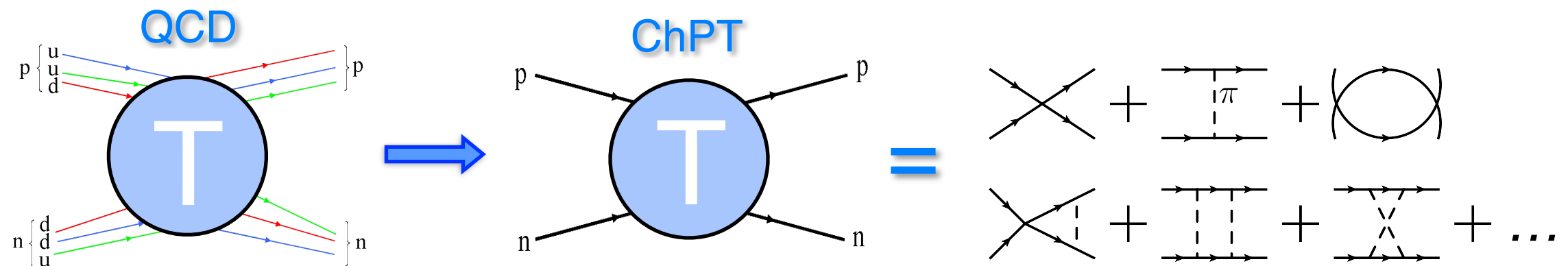


Kamada

Outline

- Nuclear forces in chiral EFT
- 3NF's upto $N^4\text{LO}$
- PWD of the three-nucleon forces
- Summary & Outlook

ChPT pros and cons



ChPT as an effective field theory of QCD

- ✓ is the most general field theory with pions, nucleons (deltas) as dofs in line with the symmetries of QCD
- ✓ is systematically improvable
- ✓ gives a unified description of $\pi\pi$, πN , NN , (axial) vector currents etc.
- ✓ naturally explains the hierarchy $V_{2N} \gg V_{3N} \gg V_{4N}$
- ✓ predicts the long range behavior of nuclear forces
- ✓ allows doing precision physics with/from light nuclei

- ✗ number of free parameters (LEC) increases with increasing order in ChPT
- ✗ does not provide an explanation on the size of a particular LEC
- ✗ is only applicable in the low energy region
- ✗ convergence radius of ChPT is a priori unknown

ChPT nuclear forces

	V_{NN}	V_{3N}	V_{4N}
Worked out up to the order	N ⁴ LO Evgeny's talk	N ³ LO N ⁴ LO in progress	N ³ LO
Regularization used	Dim. Reg In combination with semi-local regularization in Schrödinger eq.	Dim. Reg.	—

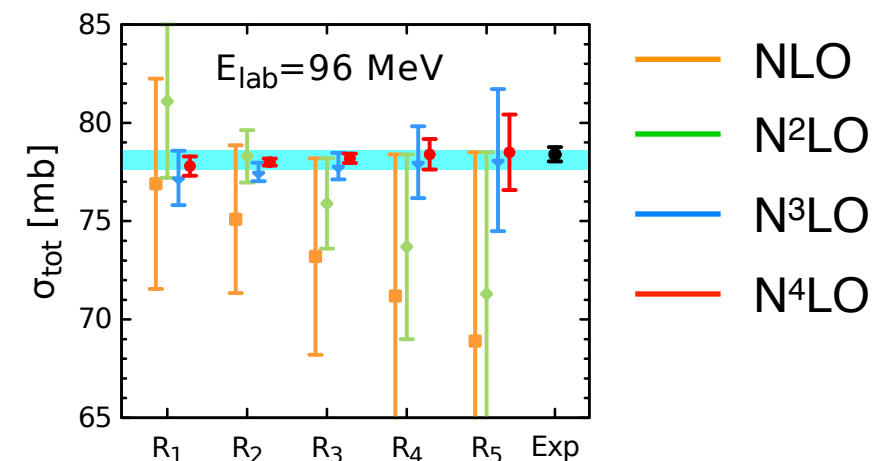
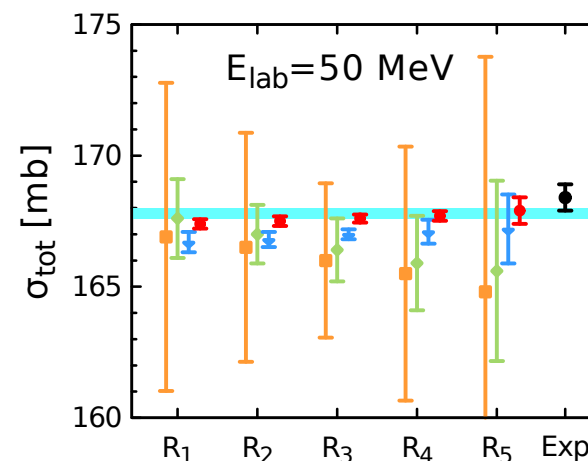
Partial N⁵LO calculation → Ruprecht's talk

Novelties in NN sector (beside the construction of N⁴LO NN)

- Local regularization in coordinate space: $V_{\text{long-range}}(\vec{r}) \rightarrow V_{\text{long-range}}(\vec{r}) \left[1 - \exp\left(-\frac{r^2}{R^2}\right) \right]^n$
 - ✓ By construction long - range physics is unaffected by this regulator
 - ✓ No additional SFR is needed
- Theoretical uncertainty estimation due to chiral expansion for every fixed cutoff R

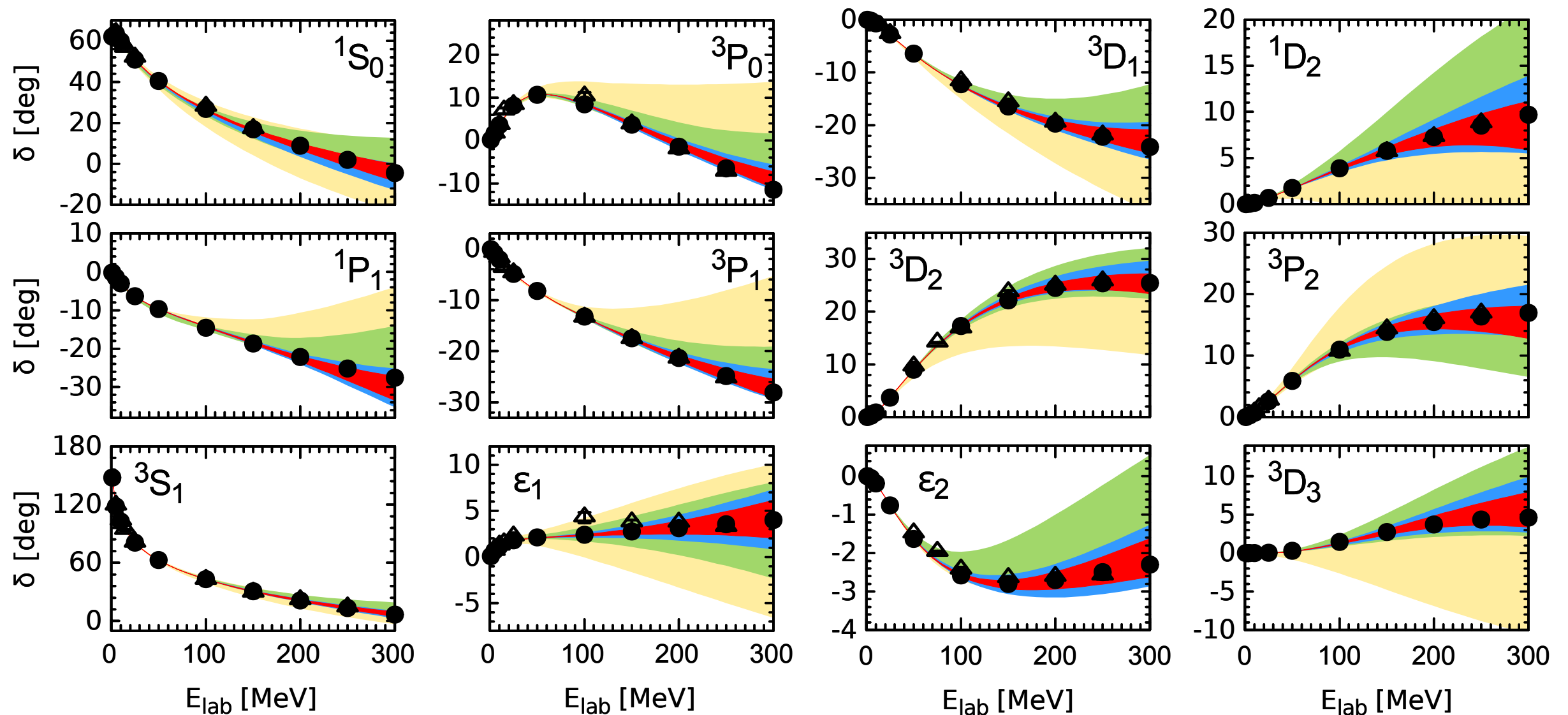
np total cross section
at several cutoffs
 $R_i = (0.7 + i \times 0.1) \text{ fm}$

Epelbaum, HK, Meißner
EPJA 51 (2015) 5



Phase shifts and mixing angles

Epelbaum, HK, Meißner, arXiv: 1412.4623

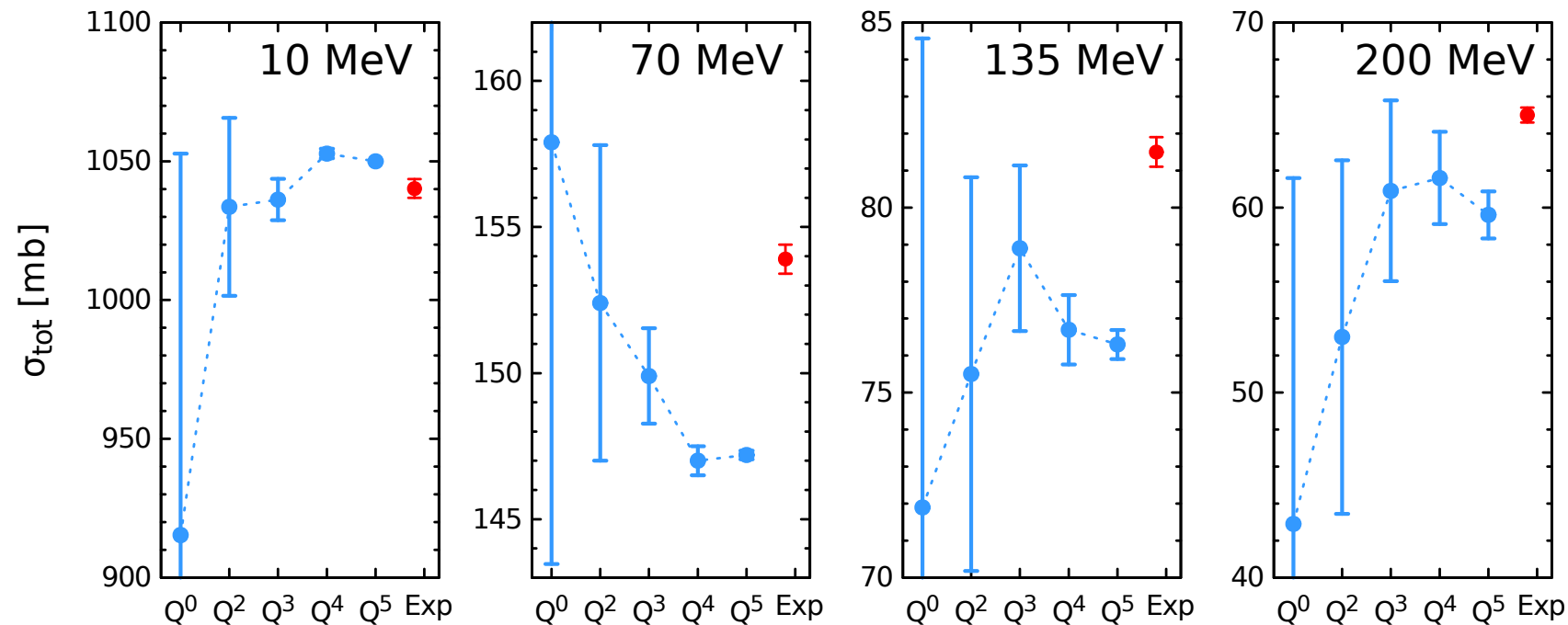


$R = 0.9$ fm — NLO — N²LO — N³LO — N⁴LO

- Good convergence of chiral expansion
- Error bands are consistent with each other → strong support of chiral uncertainty estimation
- Excellent agreement with NPWA data

Role of the 3NFs

LENPIC collaboration: Binder et al. arXiv:1505.07218



Total cross section
for Nd scattering

chiral predictions
without 3NF at $R = 0.9$ fm

- The discrepancy at 10 MeV is much lower than at other energies
- Significant discrepancy between experiment and theory

Cross section at low energy is governed by S-wave spin-doublet and spin-quartet Nd scattering lengths:

$^4a \gg ^2a$ (one order of magnitude)

4a is much less sensitive to 3NF (Pauli principle)

Clear evidence of missing 3NFs at higher energy

3NF up to N⁴LO

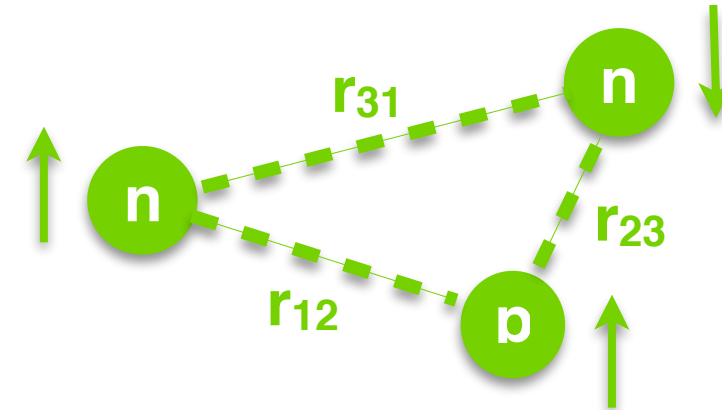
	Long - range			Short - range		
NLO						
N ² LO	van Kolck '94, Epelbaum et al. '02					
N ³ LO						
N ⁴ LO	HK, Gasparyan, Epelbaum PRC85 (12); PRC87 (13)			Work in progress		

Most general structure of a local 3NF

Up to N⁴LO, the computed contributions are local → it is natural to switch to r-space.

A meaningful comparison requires a **complete set of independent operators**

$$\begin{aligned}
 \tilde{\mathcal{G}}_1 &= 1 \\
 \tilde{\mathcal{G}}_2 &= \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \\
 \tilde{\mathcal{G}}_3 &= \vec{\sigma}_1 \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_4 &= \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \vec{\sigma}_1 \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_5 &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \vec{\sigma}_1 \cdot \vec{\sigma}_2 \\
 \tilde{\mathcal{G}}_6 &= \boldsymbol{\tau}_1 \cdot (\boldsymbol{\tau}_2 \times \boldsymbol{\tau}_3) \vec{\sigma}_1 \cdot (\vec{\sigma}_2 \times \vec{\sigma}_3) \\
 \tilde{\mathcal{G}}_7 &= \boldsymbol{\tau}_1 \cdot (\boldsymbol{\tau}_2 \times \boldsymbol{\tau}_3) \vec{\sigma}_2 \cdot (\hat{r}_{12} \times \hat{r}_{23}) \\
 \tilde{\mathcal{G}}_8 &= \hat{r}_{23} \cdot \vec{\sigma}_1 \hat{r}_{23} \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_9 &= \hat{r}_{23} \cdot \vec{\sigma}_3 \hat{r}_{12} \cdot \vec{\sigma}_1 \\
 \tilde{\mathcal{G}}_{10} &= \hat{r}_{23} \cdot \vec{\sigma}_1 \hat{r}_{12} \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_{11} &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \hat{r}_{23} \cdot \vec{\sigma}_1 \hat{r}_{23} \cdot \vec{\sigma}_2 \\
 \tilde{\mathcal{G}}_{12} &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \hat{r}_{23} \cdot \vec{\sigma}_1 \hat{r}_{12} \cdot \vec{\sigma}_2 \\
 \tilde{\mathcal{G}}_{13} &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \hat{r}_{12} \cdot \vec{\sigma}_1 \hat{r}_{23} \cdot \vec{\sigma}_2 \\
 \tilde{\mathcal{G}}_{14} &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \hat{r}_{12} \cdot \vec{\sigma}_1 \hat{r}_{12} \cdot \vec{\sigma}_2 \\
 \tilde{\mathcal{G}}_{15} &= \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \hat{r}_{13} \cdot \vec{\sigma}_1 \hat{r}_{13} \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_{16} &= \boldsymbol{\tau}_2 \cdot \boldsymbol{\tau}_3 \hat{r}_{12} \cdot \vec{\sigma}_2 \hat{r}_{12} \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_{17} &= \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3 \hat{r}_{23} \cdot \vec{\sigma}_1 \hat{r}_{12} \cdot \vec{\sigma}_3 \\
 \tilde{\mathcal{G}}_{18} &= \boldsymbol{\tau}_1 \cdot (\boldsymbol{\tau}_2 \times \boldsymbol{\tau}_3) \vec{\sigma}_1 \cdot \vec{\sigma}_3 \vec{\sigma}_2 \cdot (\hat{r}_{12} \times \hat{r}_{23}) \\
 \tilde{\mathcal{G}}_{19} &= \boldsymbol{\tau}_1 \cdot (\boldsymbol{\tau}_2 \times \boldsymbol{\tau}_3) \vec{\sigma}_3 \cdot \hat{r}_{23} \hat{r}_{23} \cdot (\vec{\sigma}_1 \times \vec{\sigma}_2) \\
 \tilde{\mathcal{G}}_{20} &= \boldsymbol{\tau}_1 \cdot (\boldsymbol{\tau}_2 \times \boldsymbol{\tau}_3) \vec{\sigma}_1 \cdot \hat{r}_{23} \vec{\sigma}_3 \cdot \hat{r}_{12} \vec{\sigma}_2 \cdot (\hat{r}_{12} \times \hat{r}_{23})
 \end{aligned}$$



Building blocks:

$$\boldsymbol{\tau}_1, \boldsymbol{\tau}_2, \boldsymbol{\tau}_3, \vec{\sigma}_1, \vec{\sigma}_2, \vec{\sigma}_3, \vec{r}_{12}, \vec{r}_{23}$$

Constraints:

- Locality
- Isospin symmetry
- Parity and time-reversal invariance

$$\rightarrow V_{3N} = \sum_{i=1}^{20} \tilde{\mathcal{G}}_i F_i(r_{12}, r_{23}, r_{31}) + 5 \text{ perm.}$$

Epelbaum, Gasparyan, HK, PRC87 (2013) 054007

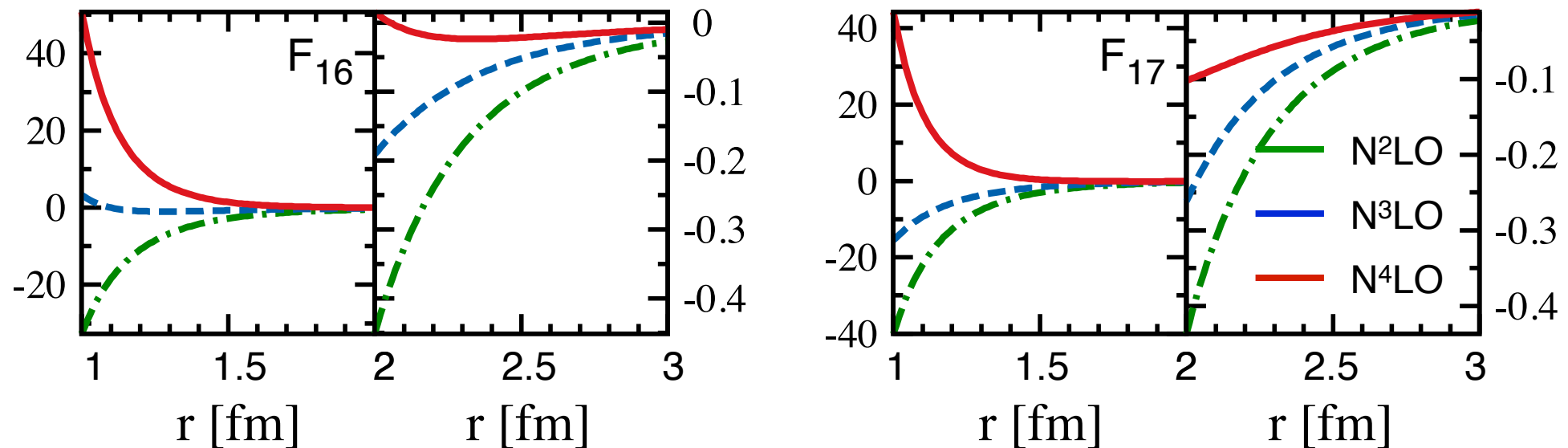
Schat, Phillips, PRC88 (2013) 034002

Epelbaum, Gasparyan, HK, Schat, EPJA51 (2015) 3

Long-range 3NF up to N⁴LO

Representative dominant contributions to profile functions

Epelbaum, Gasparyan, HK, Schat, EPJA51 (2015) 3



- All 22 profile functions start to contribute at N⁴LO
- Large N⁴LO contributions due to sizable c_i 's (hidden Δ dofs)
- No statement about convergence possible
 - explicit Δ treatment needed to clarify convergence issue

Quantitative statements are only possible once observables are calculated

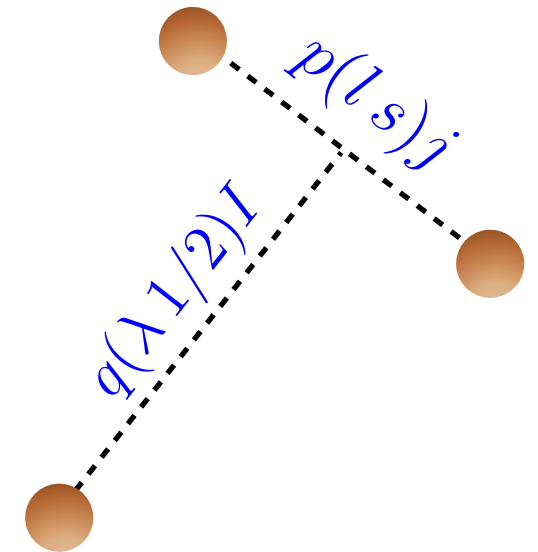
Partial wave decomposition

Golak et al. Eur. Phys. J. A 43 (2010) 241

- Faddeev equation is solved in the partial wave basis

$$|p, q, \alpha\rangle \equiv |pq(ls)j(\lambda\frac{1}{2})I(jI)JM_J\rangle |(t\frac{1}{2})TM_T\rangle$$

- Too many terms for doing PWD by hand \Rightarrow Automatization



$$\underbrace{\langle p'q'\alpha'|V|pq\alpha\rangle}_{\text{matrix} \sim 10^5 \times 10^5} = \int \underbrace{d\hat{p}' d\hat{q}' d\hat{p} d\hat{q}}_{\text{can be reduced to 5 dim. integral}} \sum_{m_l, \dots} \left(\text{CG coeffs.} \right) \left(Y_{l, m_l}(\hat{p}) Y_{l', m_{l'}}(\hat{p}') \dots \right) \underbrace{\langle m'_{s_1} m'_{s_2} m'_{s_3} | V | m_{s_1} m_{s_2} m_{s_3} \rangle}_{\text{depends on spin \& isospin}}$$

- Numerically expensive due to many channels and 5-dim. integration
- PWD matrix-elements can be used to produce matrix-elements in harmonic oscillator basis

Straightforward implementation of high order 3nf's in many-body calc.
within No-Core Shell Model

PWD for local forces

$$\langle m'_s | \vec{\sigma} \cdot \vec{p} | m_s \rangle = \sum_{\mu=-1}^1 p Y_{1\mu}^*(\hat{p}) \sqrt{\frac{4\pi}{3}} \langle m'_s | \vec{\sigma} \cdot \vec{e}_\mu | m_s \rangle \quad \leftarrow \text{momentum-independent part}$$

$$\begin{aligned} \langle m'_{s_1} m'_{s_2} m'_{s_3} | V | m_{s_1} m_{s_2} m_{s_3} \rangle &= \sum_{\mu' s} \langle m'_{s_1} m'_{s_2} m'_{s_3} | \text{Spin matrices \& } \vec{e}_\mu \text{'s} | m_{s_1} m_{s_2} m_{s_3} \rangle (Y_{1\mu}^* s) \\ &\times V((\vec{p}' - \vec{p})^2, (\vec{q}' - \vec{q})^2, (\vec{p}' - \vec{p}) \cdot (\vec{q}' - \vec{q})) \end{aligned}$$

can be reduced to 3 dim. integral

$$\begin{aligned} \langle p' q' \alpha' | V | p q \alpha \rangle &= \sum_{m_l \dots} (\text{CG coeffs.}) \int \overbrace{d\hat{p}' d\hat{q}' d\hat{p} d\hat{q}}^{\text{can be reduced to 3 dim. integral}} Y_{l'_1 m'_1}^*(\hat{p}') Y_{l'_2 m'_2}^*(\hat{q}') Y_{l_1 m_1}^*(\hat{p}) Y_{l_2 m_2}^*(\hat{q}) \\ &\times V((\vec{p}' - \vec{p})^2, (\vec{q}' - \vec{q})^2, (\vec{p}' - \vec{p}) \cdot (\vec{q}' - \vec{q})) \rightarrow \text{Speed up factors} > 1000 \end{aligned}$$

- Unregularized 3NF matrix elements can be used to generate locally regularized 3NFs

$$\langle p' q' \alpha' | V | p q \alpha \rangle \rightarrow \sum_n \langle p' q' \alpha' | V | n \rangle \langle n | R | p q \alpha \rangle \quad \text{with } \langle p' q' \alpha' | R | p q \alpha \rangle \text{ matrix element of local regulator}$$

Hebeler, HK, Epelbaum, Golak, Skibinski PRC91 (2015) 4

Summary

- Chiral 3NF's are studied up to $N^3\text{LO}$ / partly up to $N^4\text{LO}$
- Optimized version of PWD for local 3NF's
- Stored matrix elements can be used within local regularization

Outlook

- $N^4\text{LO}$ Δ -less/ $N^3\text{LO}$ - Δ calc. of shorter range part of 3NF
 - Generation of matrix-elements for 3NF's up to $N^4\text{LO}$ Δ -less/ $N^3\text{LO}$ - Δ
Due to optimized PWD should not cost much