

# Amorphous optical coatings for gravitational-wave interferometers

M. Granata, E. Barthelemy, G. Cagnoli, Q. Cassar, J. Degallaix,  
V. Dolique, D. Forest, C. Michel, R. Pedurand, L. Pinard, B. Sassolas  
*Laboratoire des Matériaux Avancés - CNRS*

E. Coillet, C. Martinet, V. Martinez, A. Mermet  
*Institut Lumière Matière – CNRS / UCBL*

A. Amato, M. Canepa  
*Università di Genova*

# outline

motivations

experimental setups

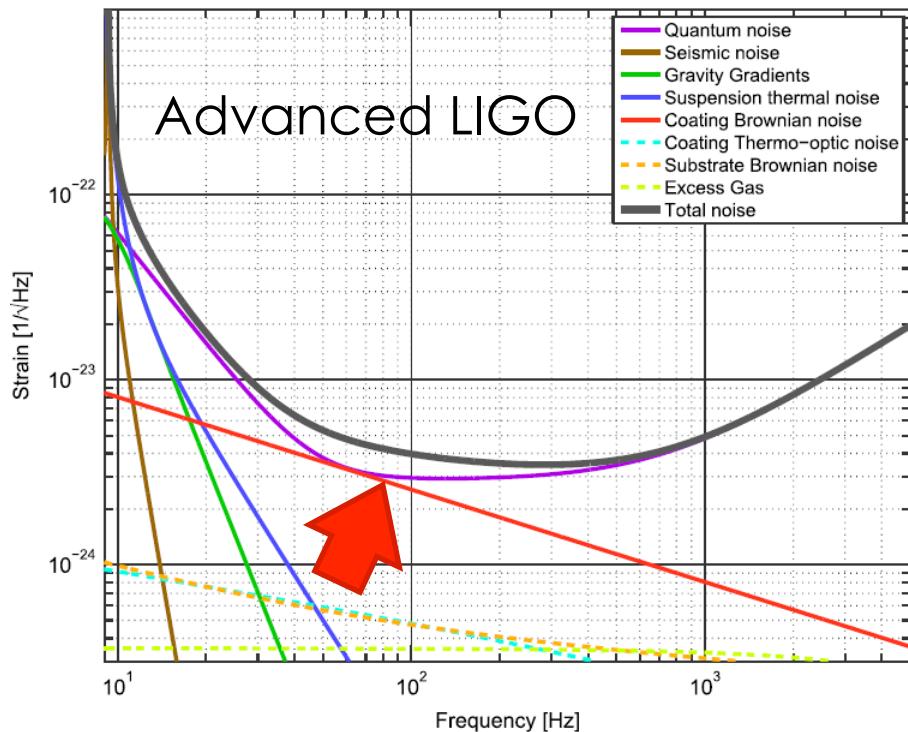
latest results

- measured dilution factor
- deposition parameters and annealing
- advanced detectors' coatings
- some new coating materials

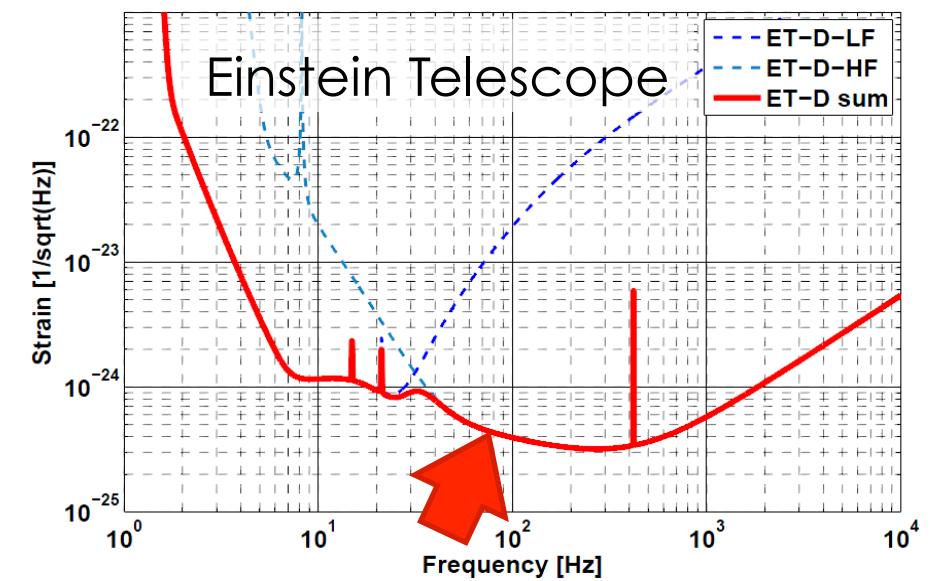
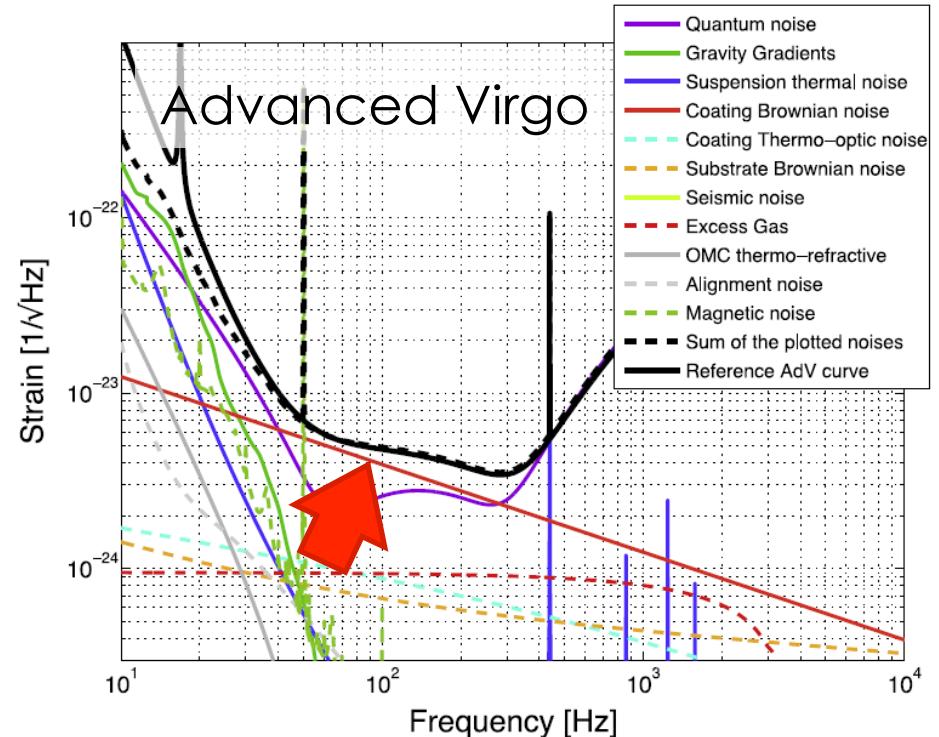
summary and conclusions

# coating thermal noise

issue for present/future detectors



Aasi & al, Class. Quantum Grav. 32 (2015)  
Acernese & al, Class. Quantum Grav. 32 (2015)  
Abernathy & al, ET-0106C-10 (2011)



# requirements

mitigation from:

temperature

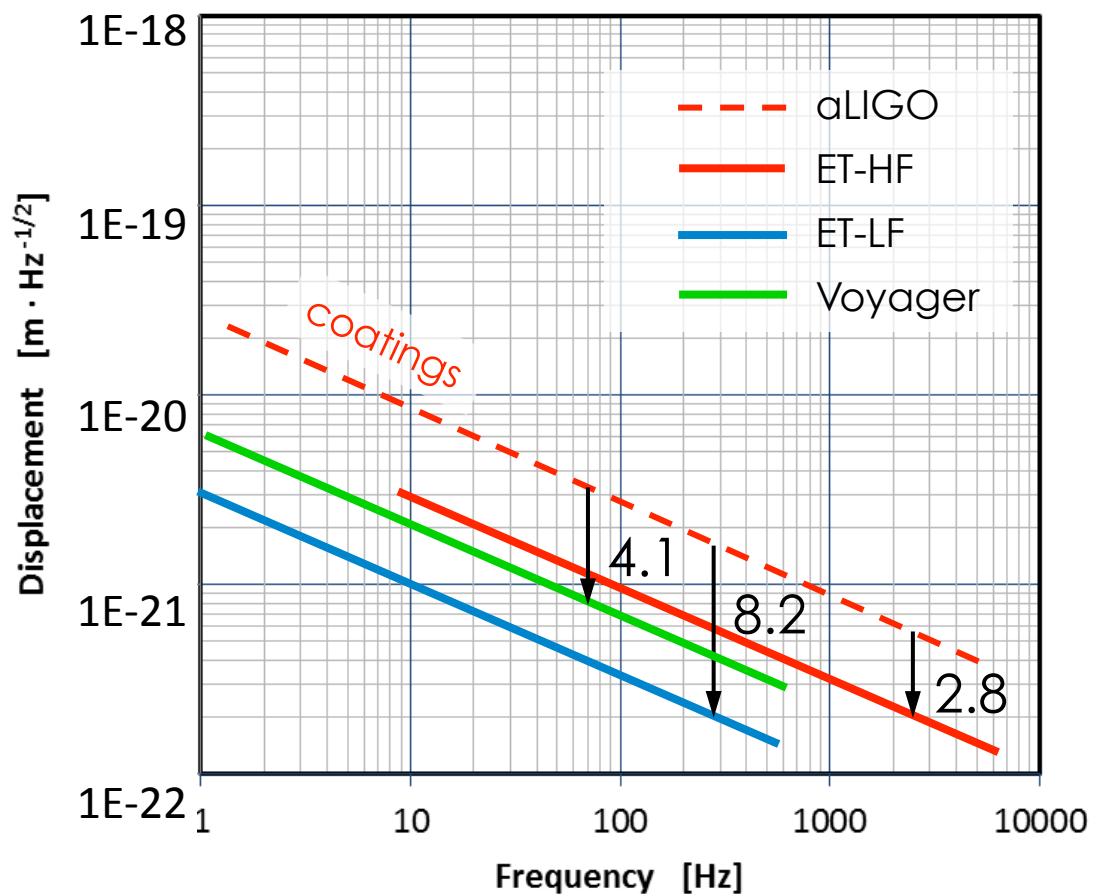
beam size/shape

loss angle

$$\text{ET-HF: } 2.8 = 1.7 \sqrt{2.7}$$

$$\text{ET-LF: } 8.2 = \sqrt{30} 1.7 \sqrt{0.8}$$

$$\text{Voyager: } 4.1 = \sqrt{2.4} \sqrt{7}$$



# requirements

mitigation from:

temperature

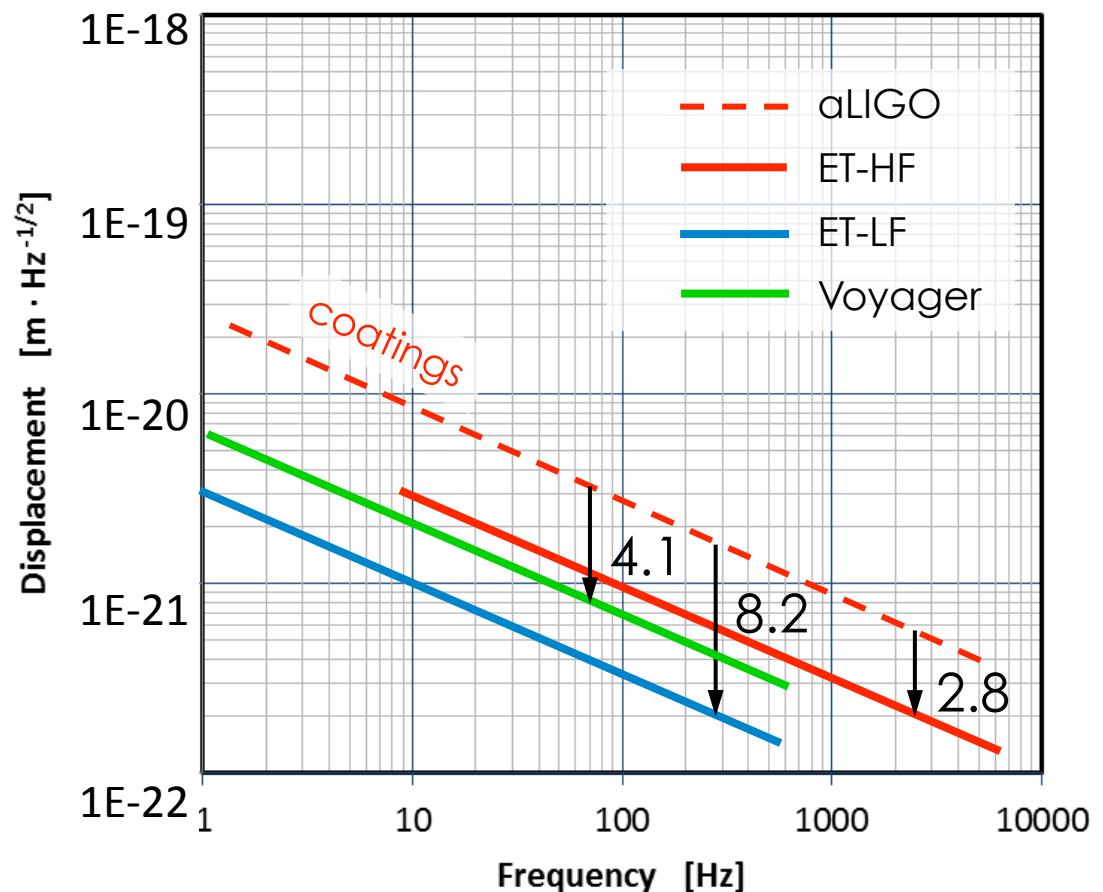
beam size/shape

loss angle

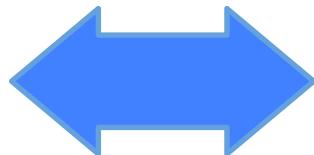
$$\text{ET-HF: } 2.8 = 1.7 \sqrt{2.7}$$

$$\text{ET-LF: } 8.2 = \sqrt{30} 1.7 \sqrt{0.8}$$

$$\text{Voyager: } 4.1 = \sqrt{2.4} \sqrt{7}$$



new materials  
and  
thickness uniformity



upgrades of  
Advanced LIGO/Virgo  
Einstein Telescope  
Cosmic Explorer / Voyager

# research & development @ LMA

## technological developments

thickness uniformity over  $\varnothing$  600 mm – **ongoing**

in-situ control system – **ongoing**

improved repeatable ARs

twin mirrors from simple rotation

## new materials

ambient temperature high-index materials – **ongoing**

cryogenic temperature high-index materials – **ongoing**

cryogenic temperature low-index materials – **starting soon**

sapphire payload (substrate+coatings+suspensions) – **in progress**

# experimental setups

# ion-beam sputtering (IBS)



photos: C. Fresillon – photothèque CNRS / E. Le Roux

# metrology – optics

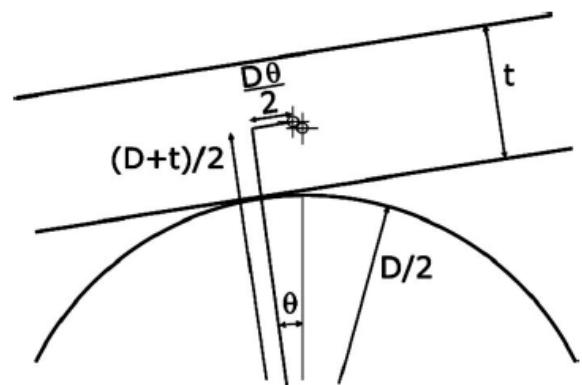
- scattering
- surface defects
- wavefront
- absorption [ambient/cryogenic]



photos: C. Fresillon – photothèque CNRS

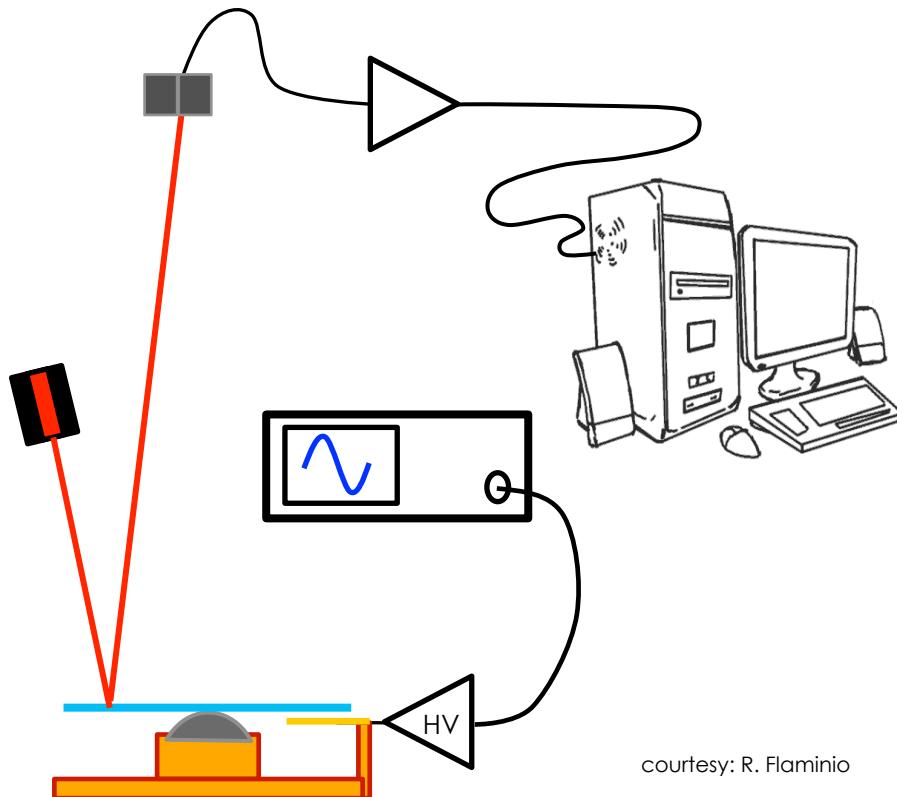
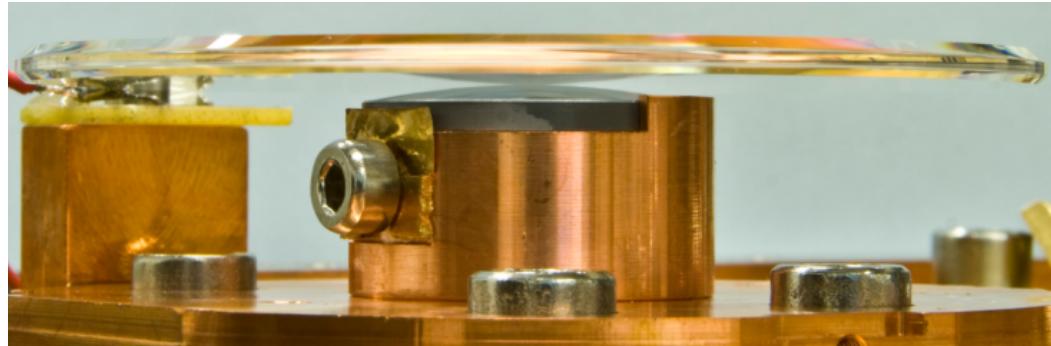
# metrology – mechanical loss

Gentle Nodal Suspension  
(GeNS)



electrostatic excitation  
optical lever + qPD readout

LabVIEW-based acquisition software



courtesy: R. Flaminio

Cesarini & al, Rev. Sci. Instrum. 80 (2009)  
Cesarini & al, Class. Quantum Grav. 27 (2010)

# LMA coating loss measurement protocol

substrates – SiO<sub>2</sub> disks

- optically polished Corning 7980
  - $d = 3"$   $t = 1$  mm [flats]
  - $d = 2"$   $t = 1$  mm [flats]
- micropolished Corning 7980
  - $d = 3"$   $t = 2.5$  mm

coatings

- on both surfaces

annealing

- 10 h @ 900° C before coating
- 10 h @ 500° C after coating

# measured dilution factor<sup>[\*]</sup>

[\*] Li & al, Phys. Rev. D 89, 092004 (2014)

# dilution factor

$$D = \frac{E_c}{E_{cs}} \simeq 1 - \left( \frac{f_0}{f} \right)^2 \frac{m_0}{m}$$

coating  
measured before coating  
coated sample  
measured after coating

The diagram illustrates the components of the dilution factor formula. Arrows point from the words 'coating' and 'measured before coating' to the term  $E_c$  in the numerator. Arrows point from the words 'coated sample' and 'measured after coating' to the term  $m$  in the denominator.

# dilution factor – TiO<sub>2</sub>Ta<sub>2</sub>O<sub>5</sub> GC

$$D = \frac{E_c}{E_{cs}} \simeq 1 - \left( \frac{f_0}{f} \right)^2 \frac{m_0}{m}$$

ANSYS

$$t_s = 1.05 \text{ mm}$$

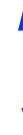
$$Y_s = 72.3 \text{ Gpa}$$

$$\nu_s = 0.164$$

$$\nu_c = 0.230$$

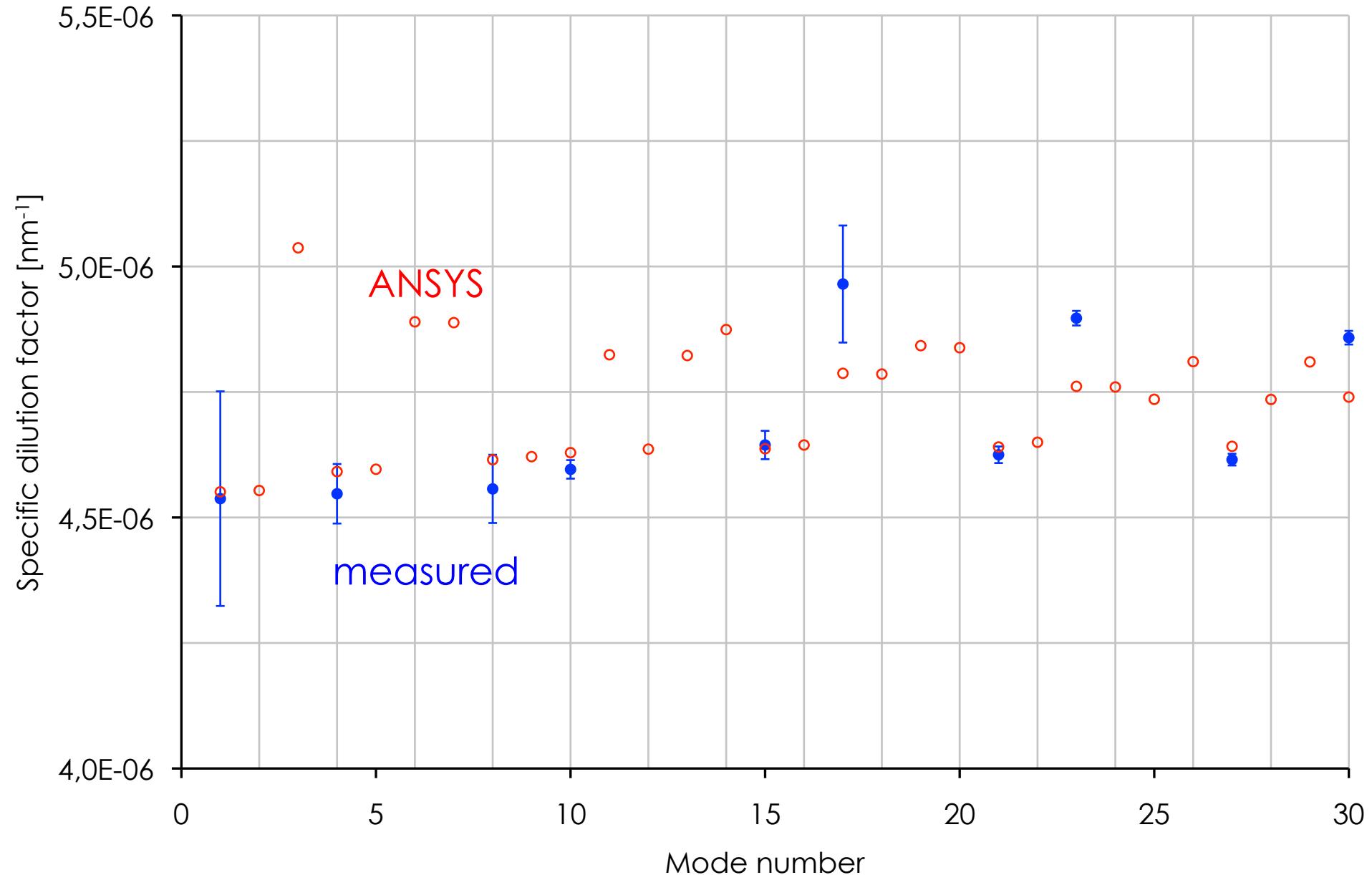
$$\rho_s = 2200 \text{ kg/m}^3$$

$$\rho_c = 6640 \text{ kg/m}^3$$



measured  
[mass / thickness (spectroscopic  
ellipsometry)]

# specific dilution factor – TiO<sub>2</sub>Ta<sub>2</sub>O<sub>5</sub> GC



# dilution factor – TiO<sub>2</sub>Ta<sub>2</sub>O<sub>5</sub> GC

estimation of coating Young's modulus:

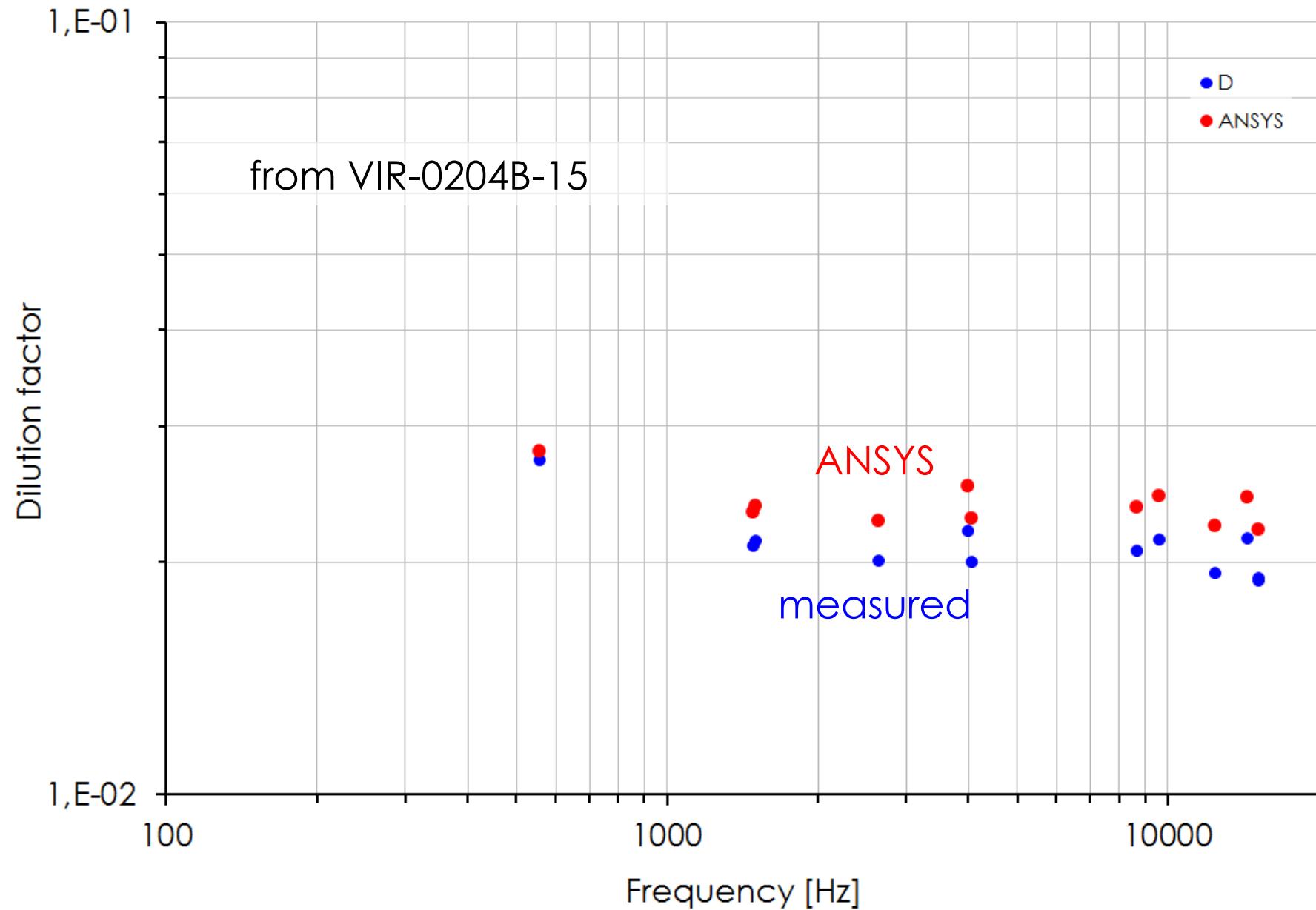
ANSYS

$$D = \frac{E_c}{E_{cs}} \simeq 1 - \left( \frac{f_0}{f} \right)^2 \frac{m_0}{m}$$

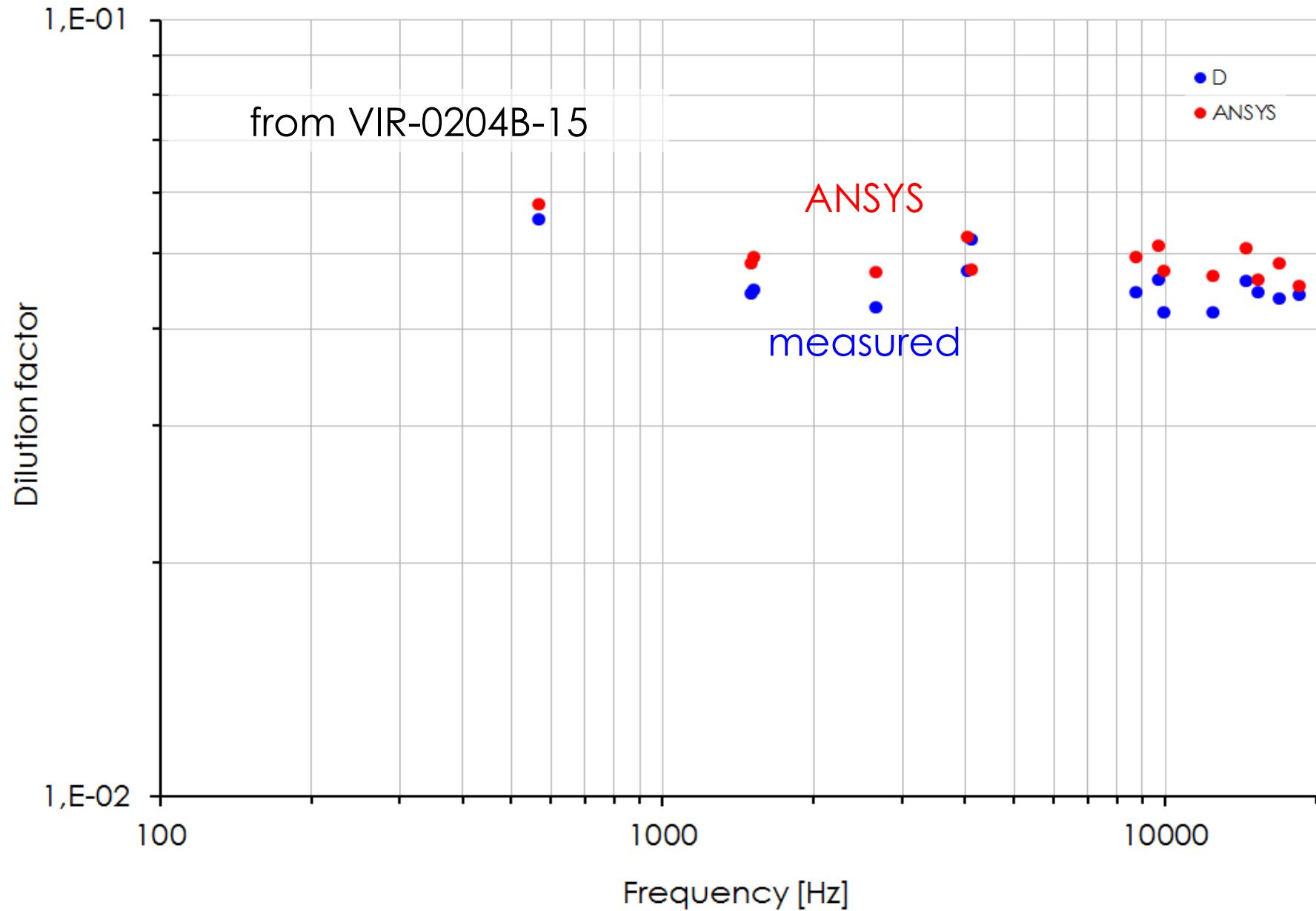
$$\begin{aligned} t_s &= 1.05 \text{ mm} \\ Y_s &= 72.3 \text{ Gpa} \\ v_s &= 0.164 \\ v_c &= 0.230 \\ \rho_s &= 2200 \text{ kg/m}^3 \\ \rho_c &= 6640 \text{ kg/m}^3 \end{aligned}$$

$$Y_c = 122 \text{ Gpa}$$

# ITM coating [on Si]

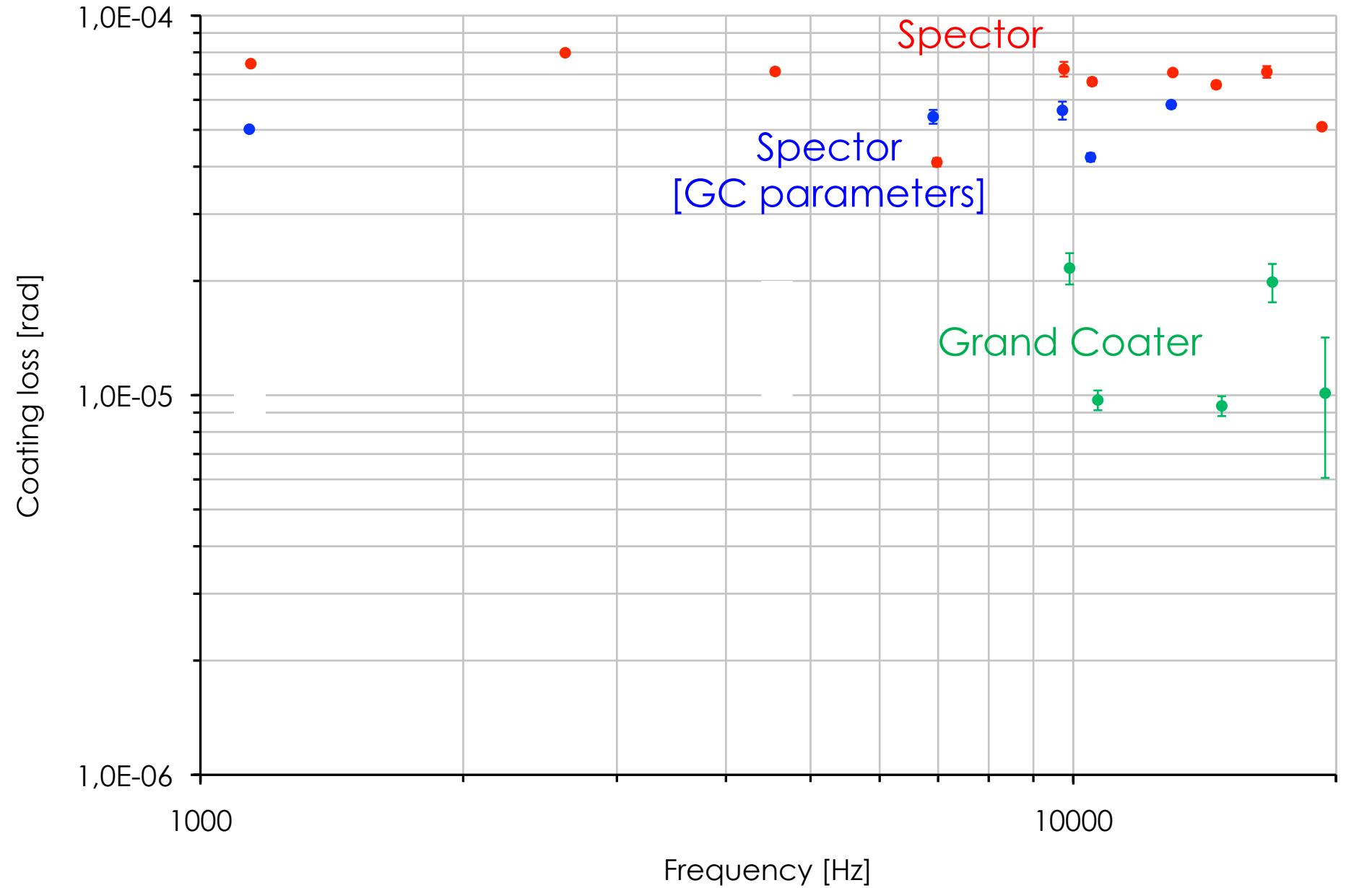


# ETM coating [on Si]

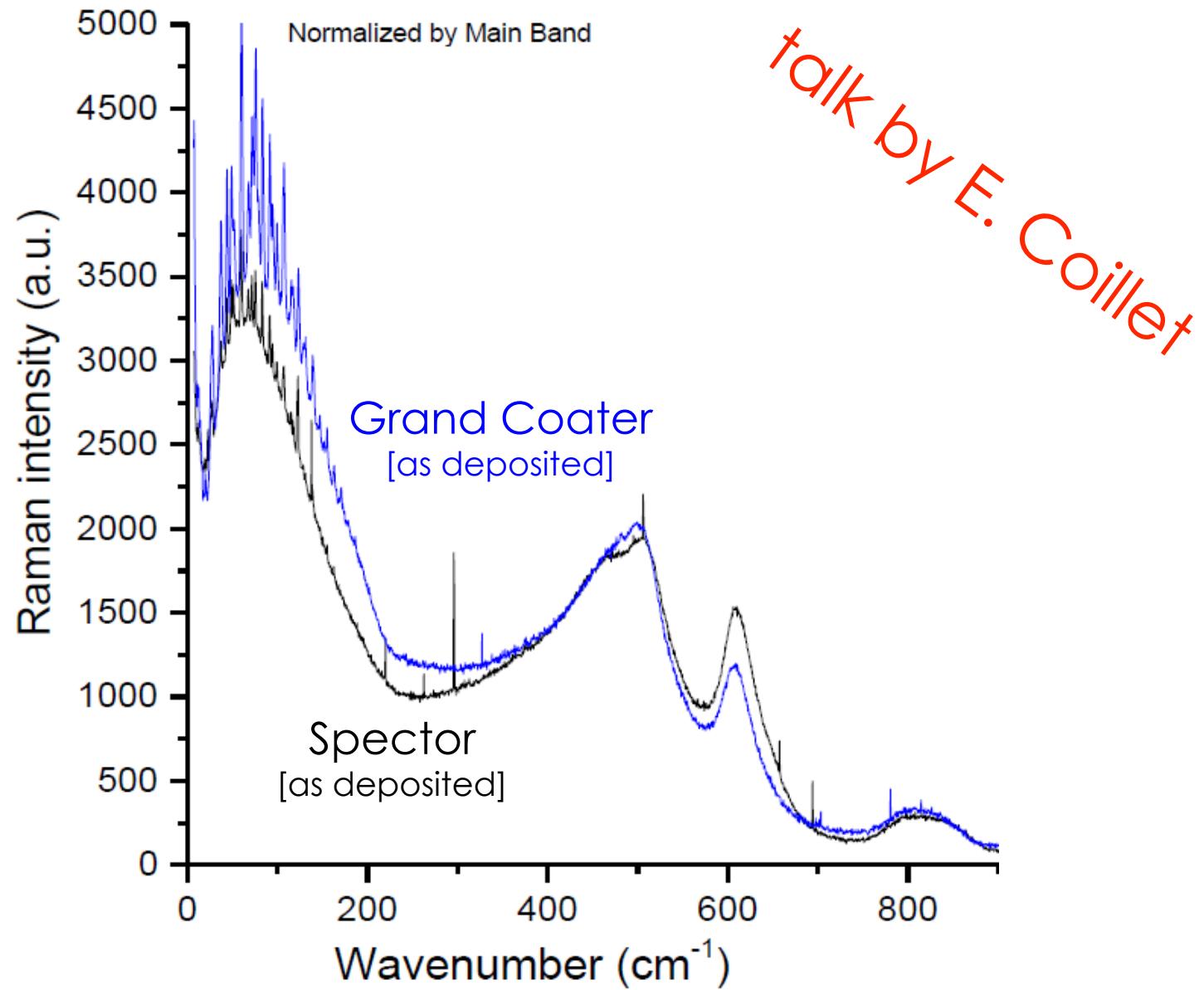


# deposition parameters and annealing - SiO<sub>2</sub> -

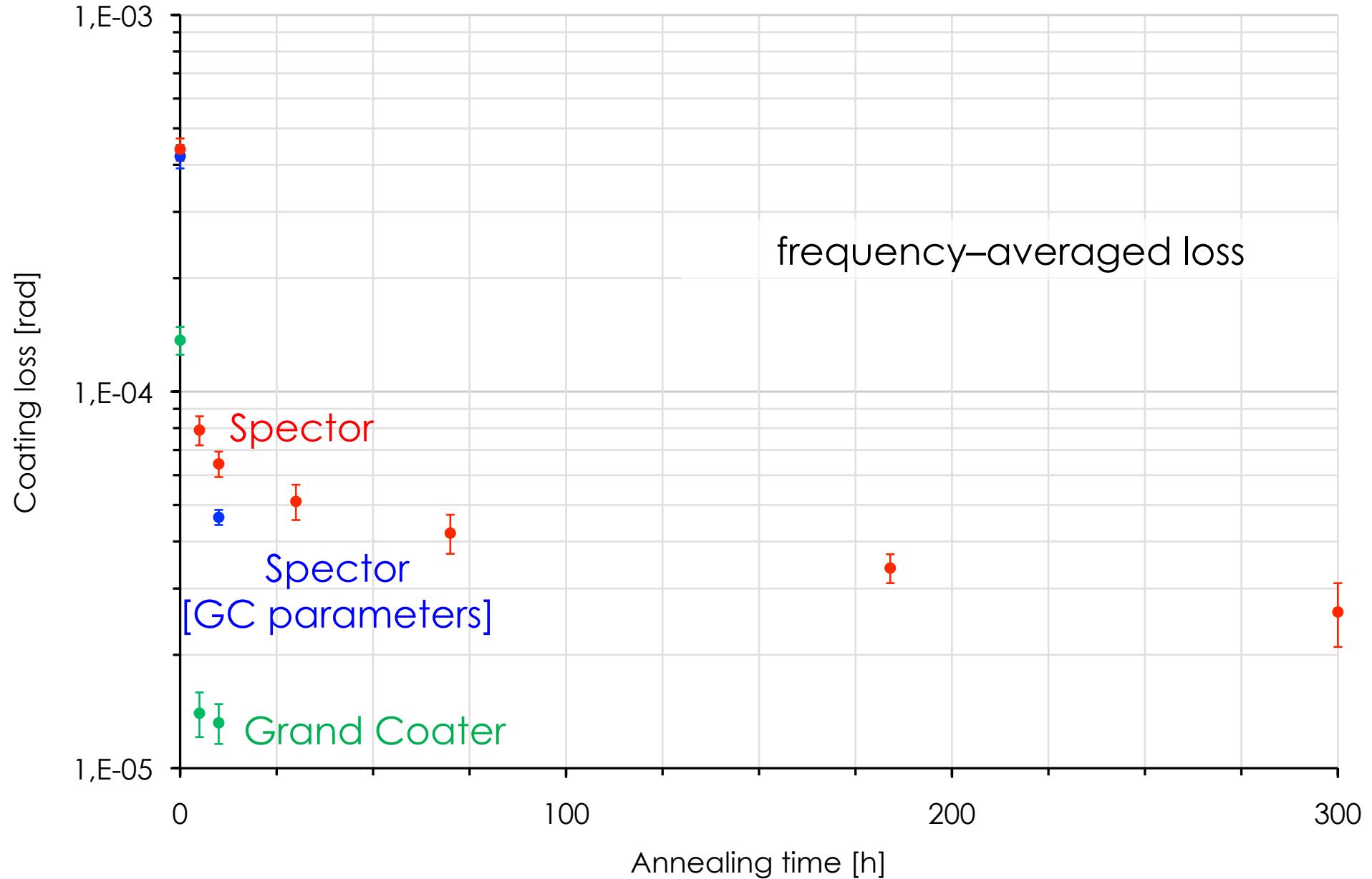
# deposition parameters



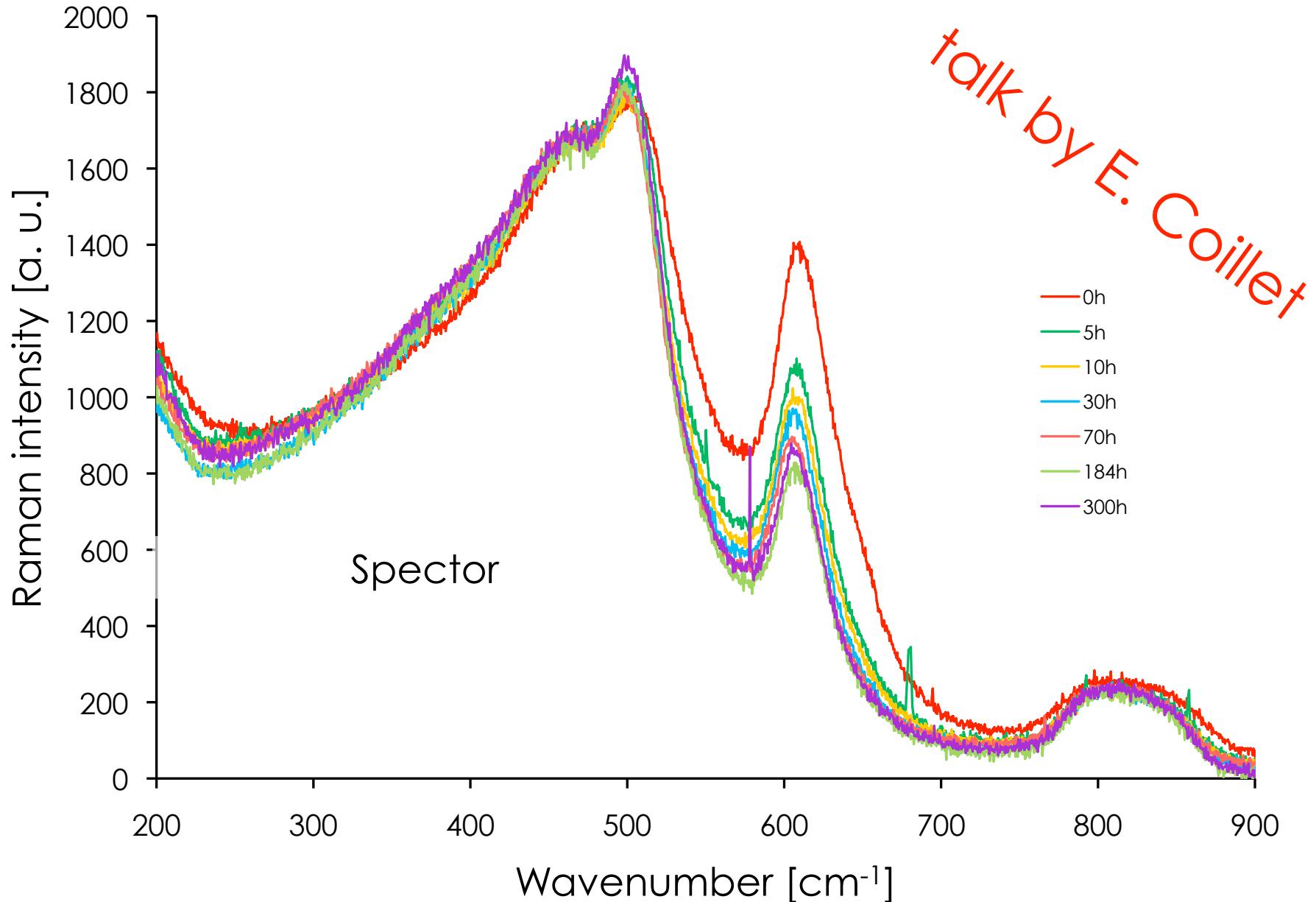
# structure



# annealing

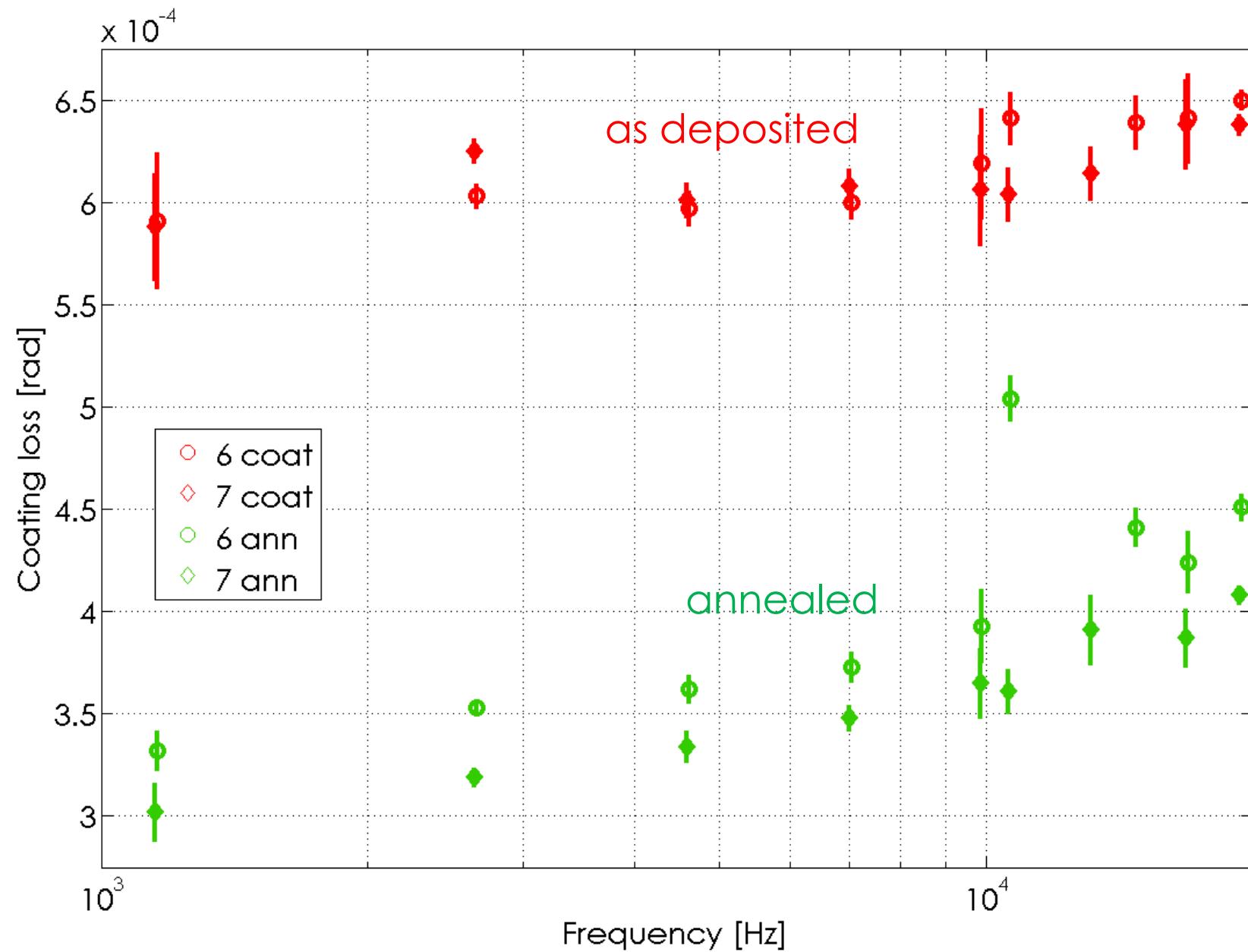


# structure

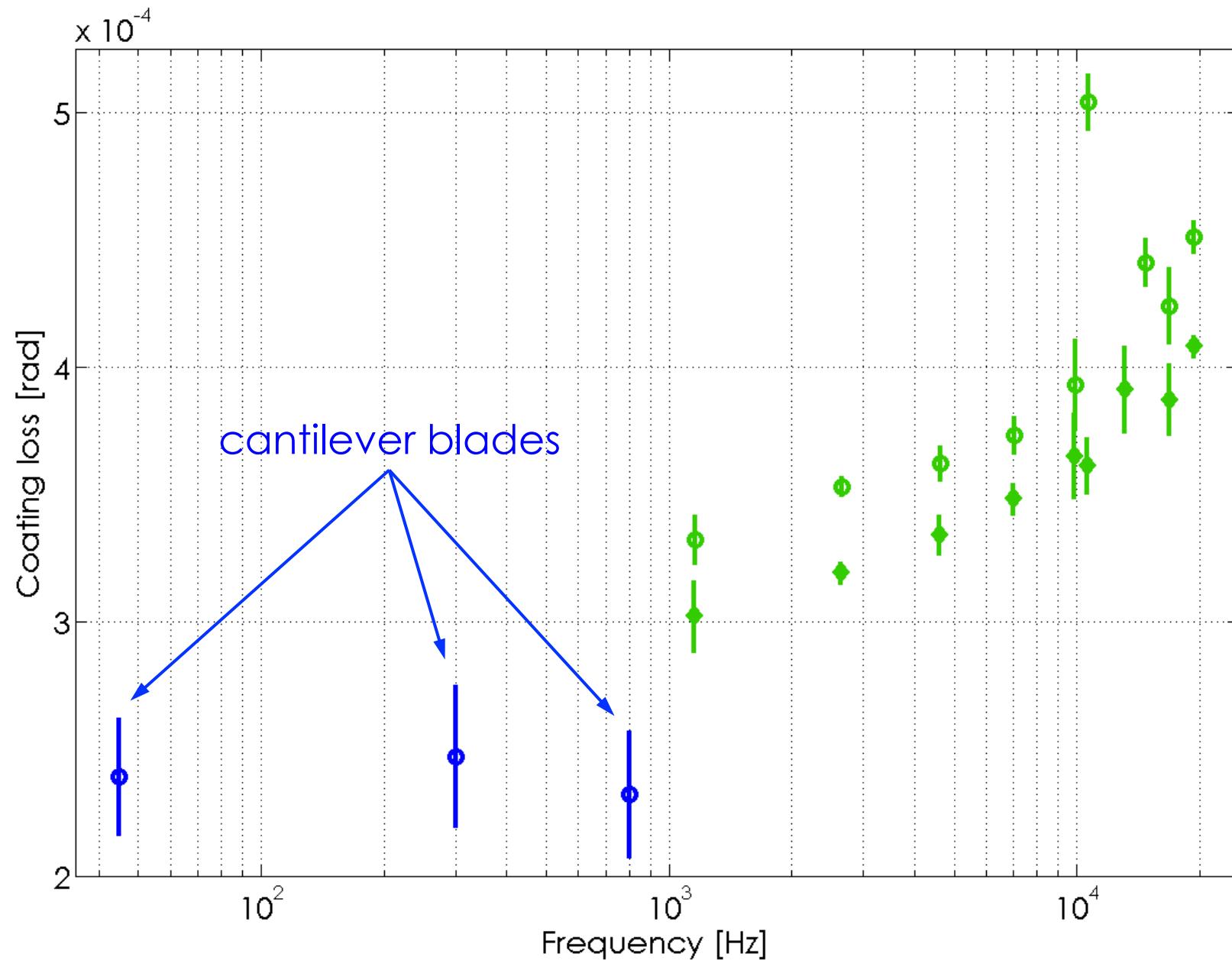


advanced detectors'  $\text{TiO}_2\text{Ta}_2\text{O}_5$

# samples 6 & 7



# previous results – comparison



# coating loss [ $\times 10^4$ ]

neglecting frequency dependence:

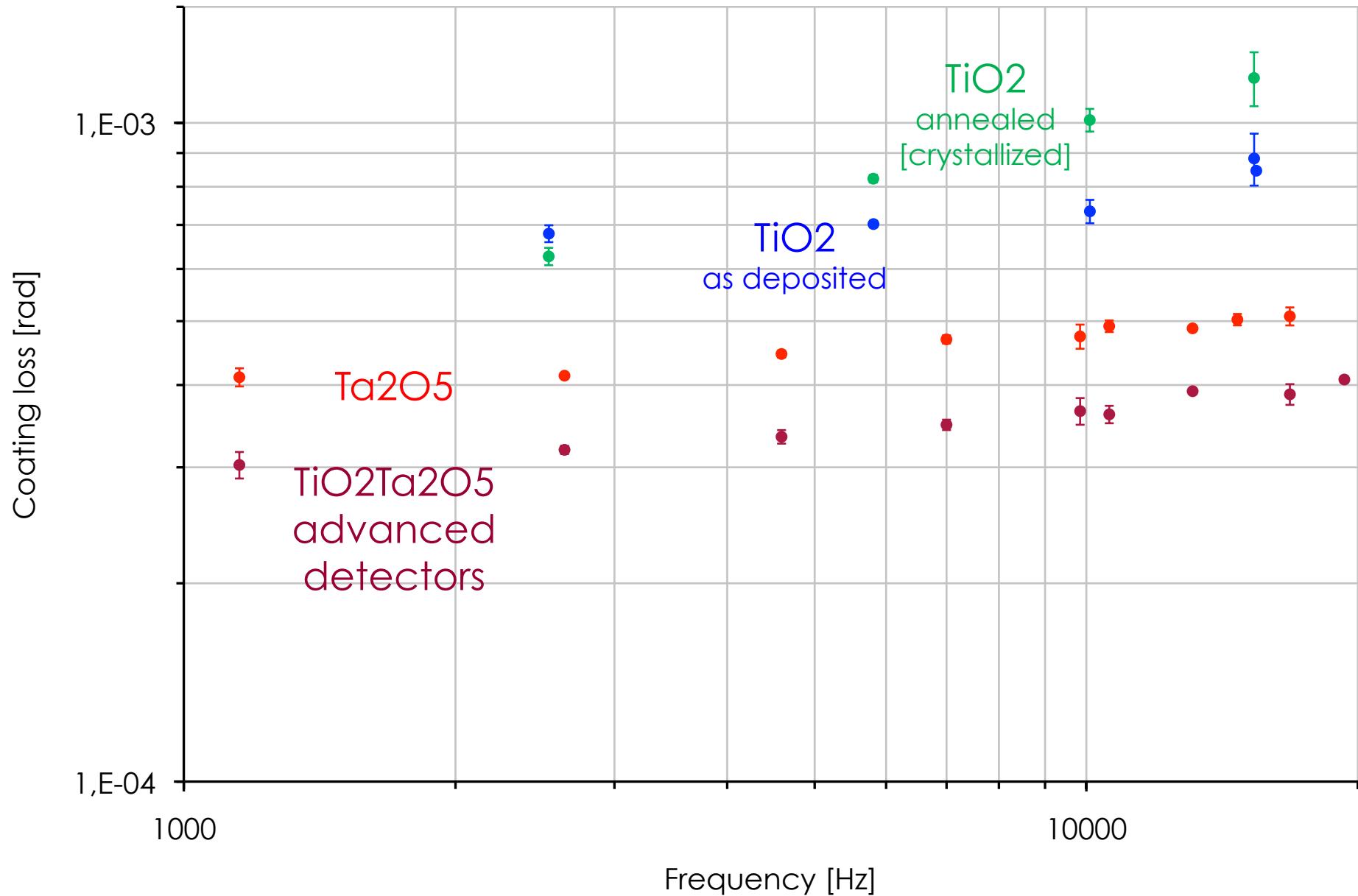
	sample 6	sample 7
$\Phi_{\text{asDep}}$	$6.2 \pm 0.2$	$6.1 \pm 0.2$
$\Phi_{\text{ann}}$	$4.0 \pm 0.6$	$3.6 \pm 0.4$
Thermal-Noise Interferometer <sup>[1]</sup> cantilever blades <sup>[2]</sup>		$3.66 \pm 0.26$ $2.4 \pm 0.3$

[1] Principe & al, Phys. Rev. D 91, 022005 (2015)

[2] Granata & al, Phys. Rev. D 93, 012007 (2016)

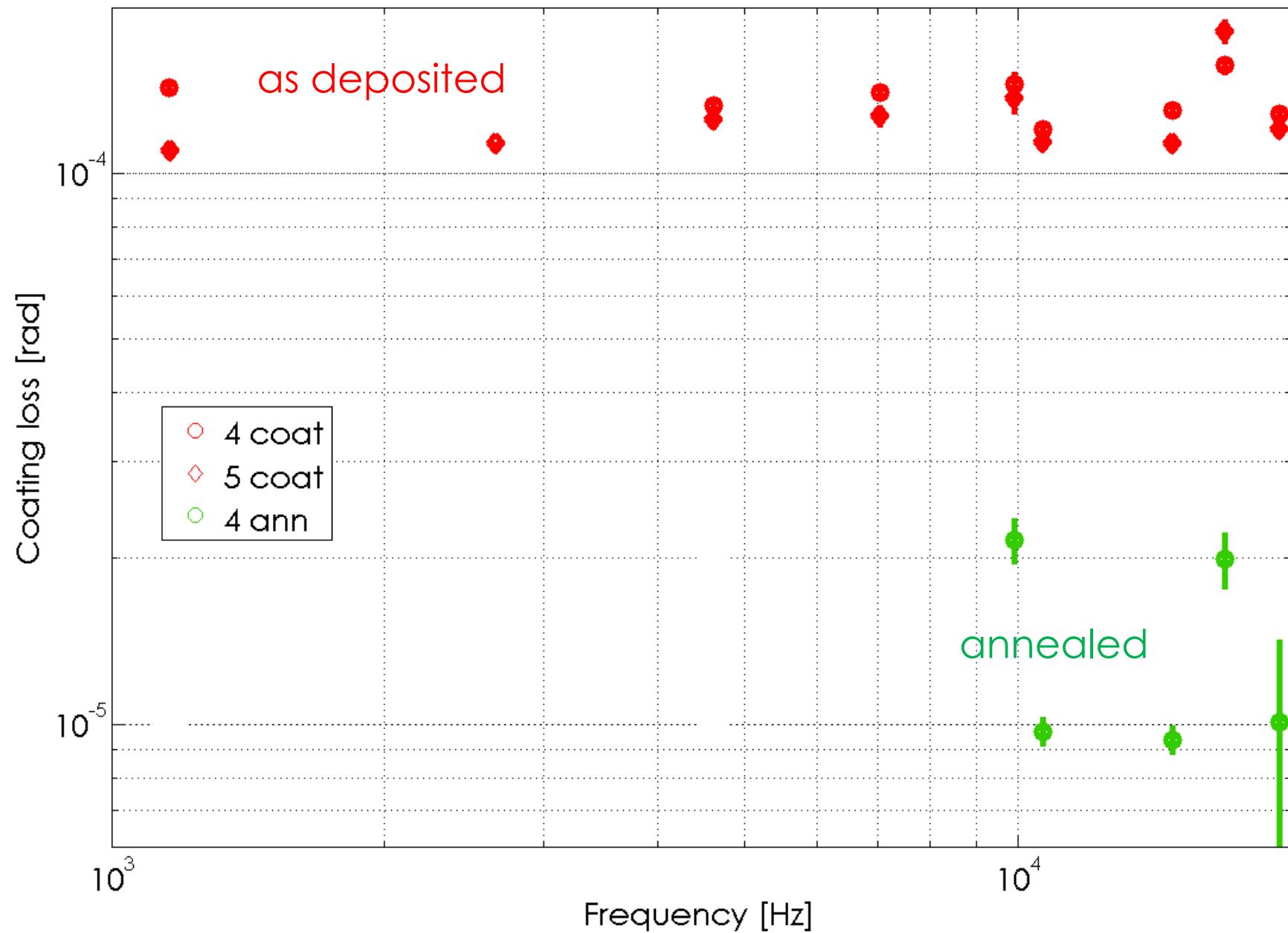
*actual impact of TiO<sub>2</sub> doping*

# TiO<sub>2</sub> doping

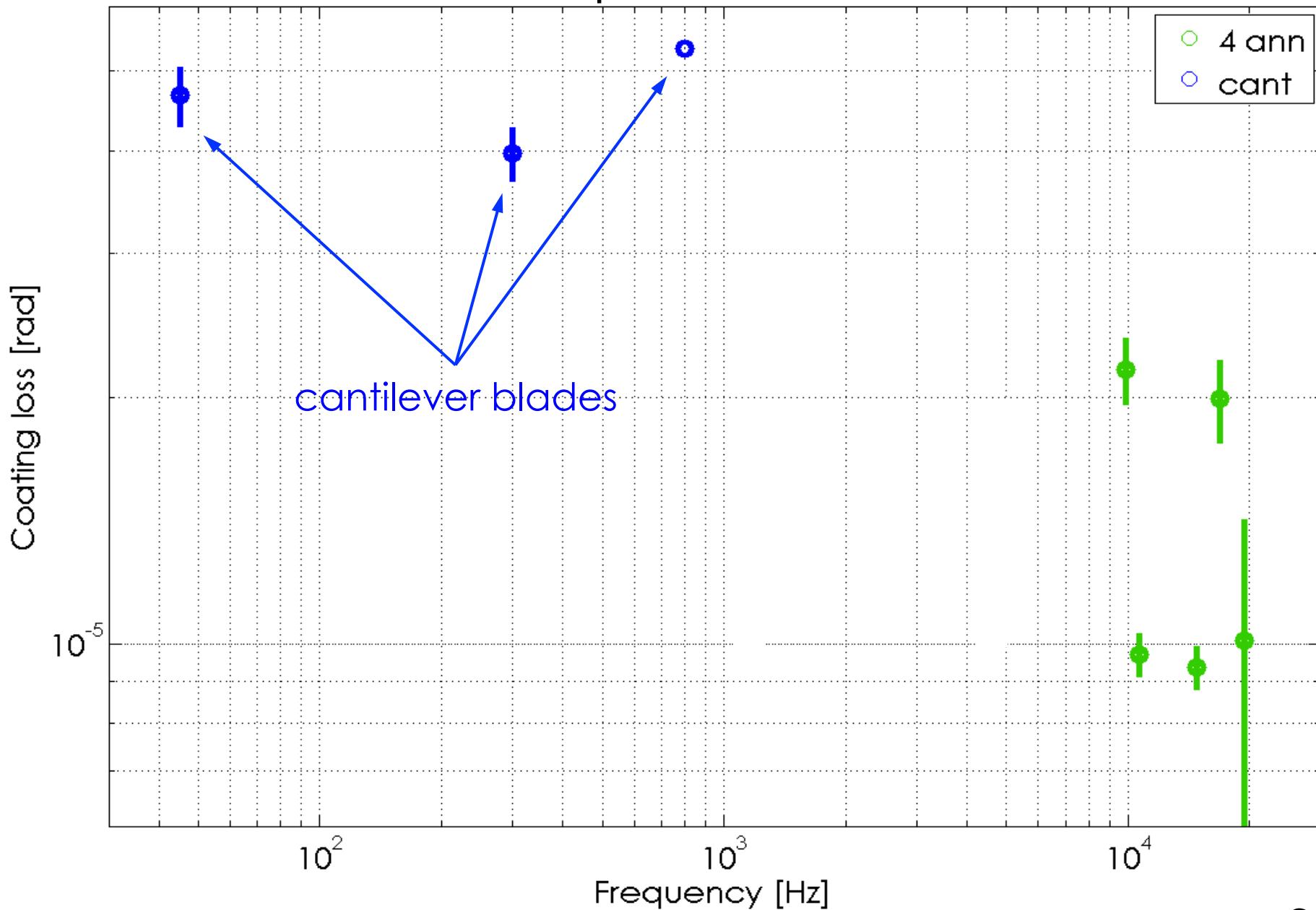


# advanced detectors' SiO<sub>2</sub>

# samples 4 & 5



# previous results – comparison



# coating loss [ $\times 10^4$ ]

neglecting shape dependence:

	sample 4	sample 5
$\Phi_{\text{asDep}}$	$1.4 \pm 0.1$	$1.3 \pm 0.2$
$\Phi_{\text{ann}}$	$0.14 \pm 0.06$	
fiber-suspended disks <sup>[1]</sup>		$0.5 \pm 0.3$
Thermal-Noise Interferometer <sup>[2]</sup>		$0.51 \pm 0.07$
cantilever blades <sup>[3]</sup>		$0.45 \pm 0.03$
q-pITF <sup>[4]</sup>		$0.62 \pm 0.48$

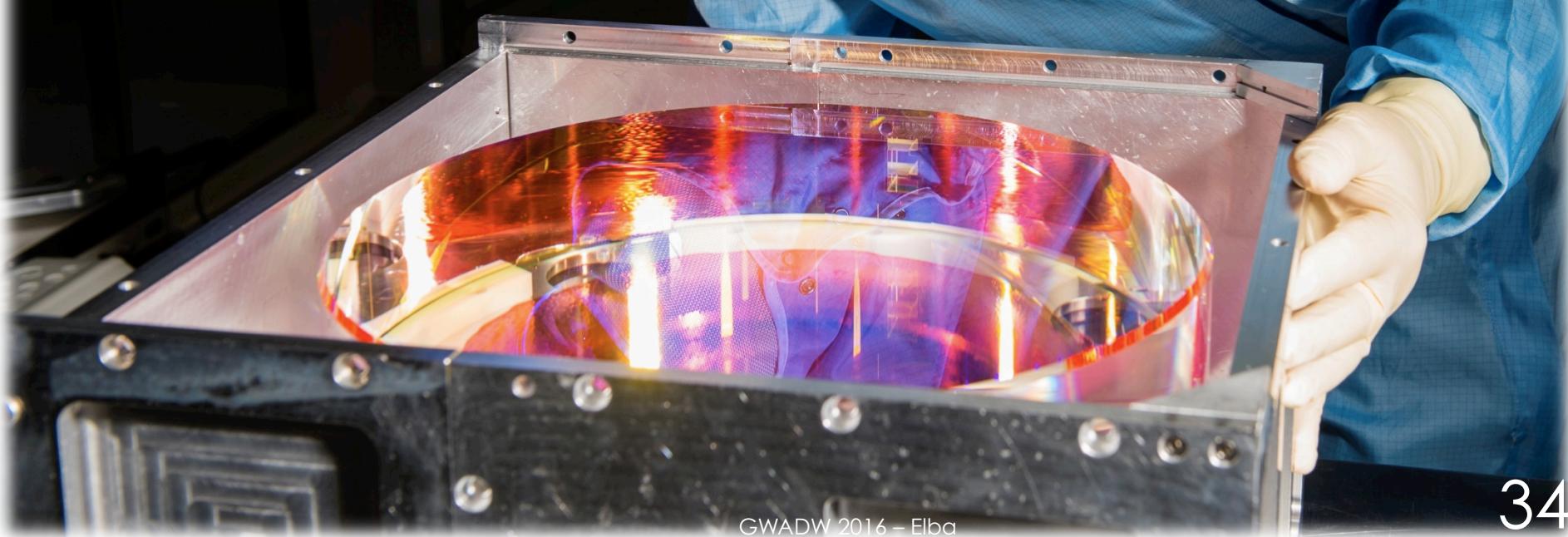
[1] Penn & al, Class. Quantum Grav. 20 (2003)

[2] Principe & al, Phys. Rev. D 91, 022005 (2015)

[3] Granata & al, Phys. Rev. D 93, 012007 (2016)

[4] Li & al, Phys. Rev. D 89, 092004 (2014)

## HR coatings of advanced detectors



# model

optimized  $\text{TiO}_2\text{Ta}_2\text{O}_5/\text{SiO}_2$  stacks

$$N_{ITM} = 8$$

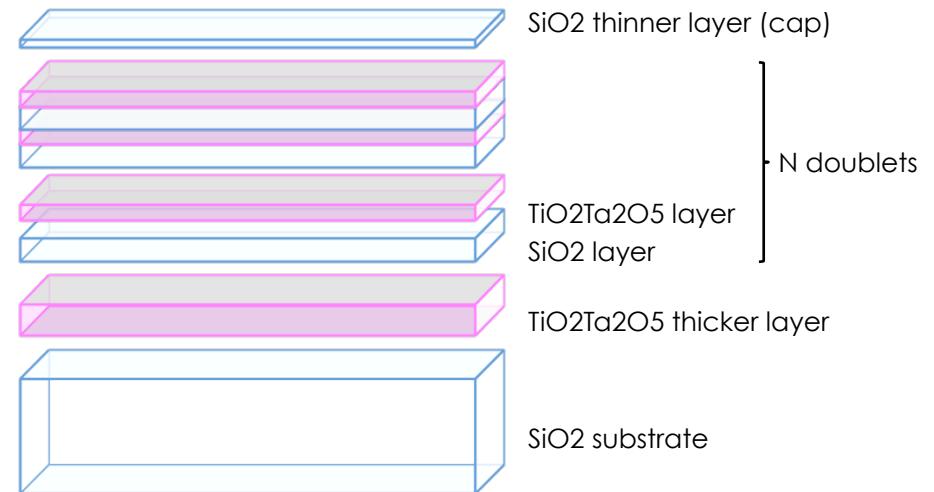
$$N_{ETM} = 18$$

expected loss<sup>[1]</sup>:

$$\phi_{HR} = \frac{\sum t_i Y_i \phi_i}{\sum t_i Y_i}$$

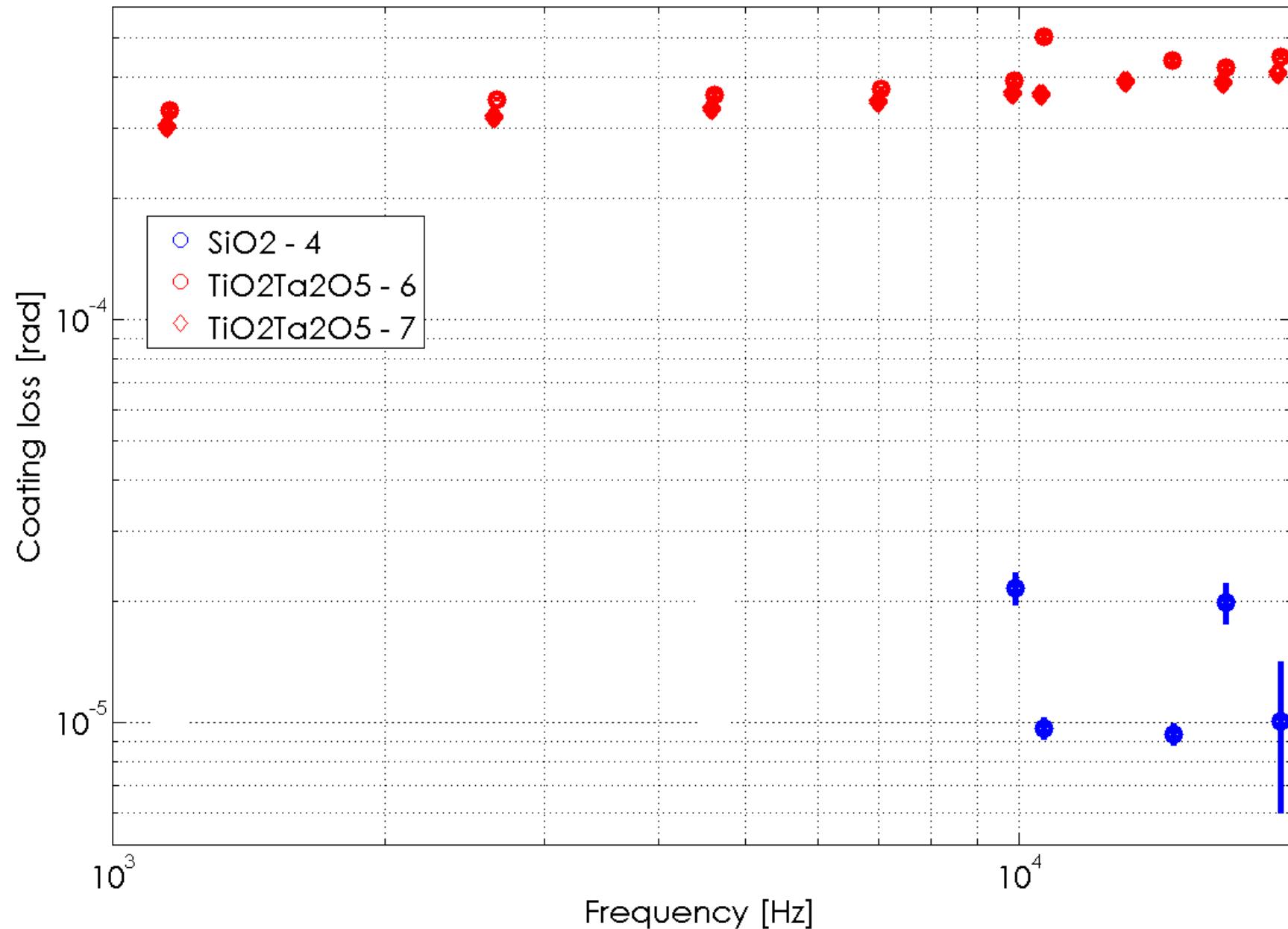
$i = \text{TiO}_2\text{Ta}_2\text{O}_5, \text{SiO}_2$

$$Y_{\text{TiO}_2\text{Ta}_2\text{O}_5} = 140 \text{ GPa}$$
$$Y_{\text{SiO}_2} = 72.3 \text{ GPa}$$

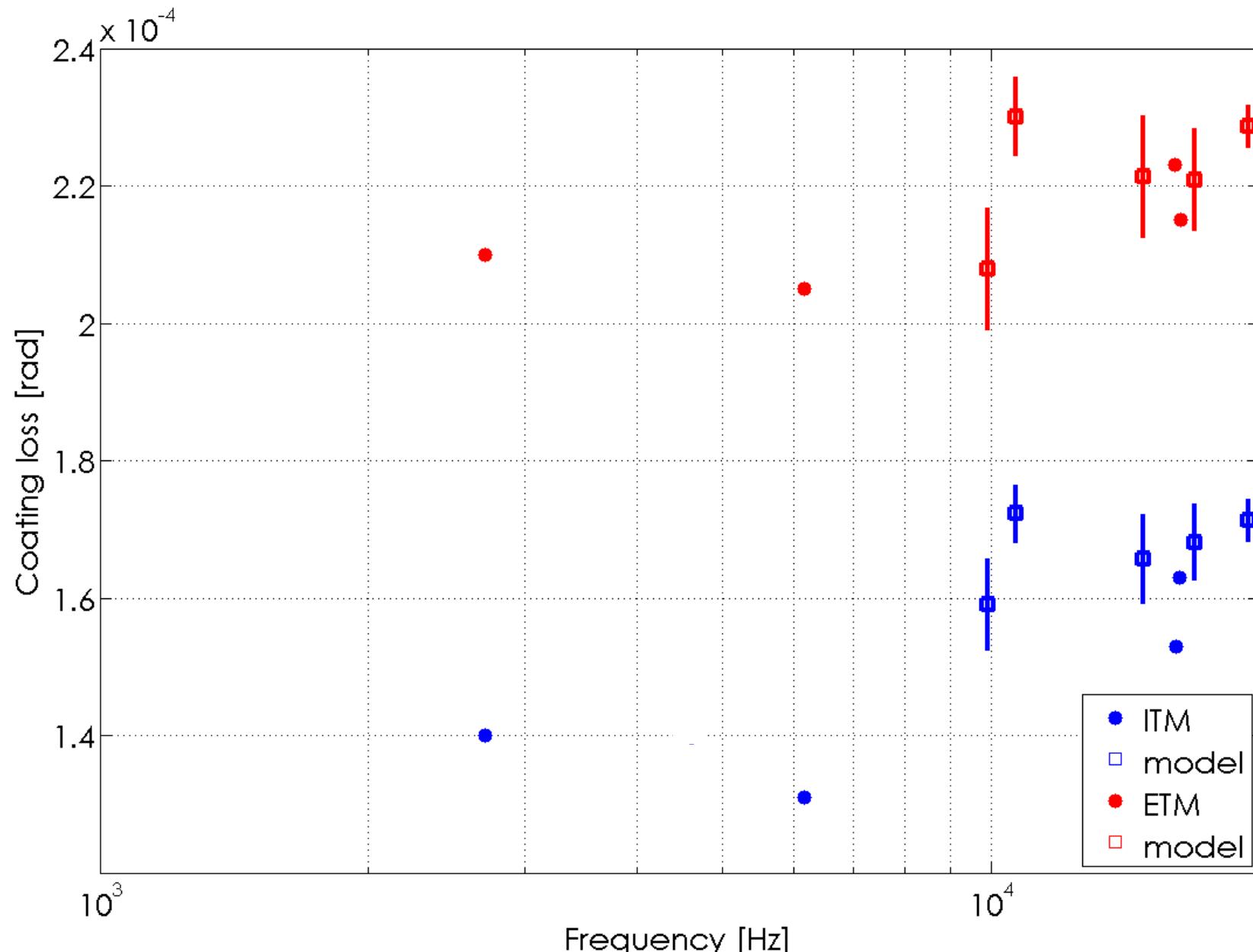


[1] Penn & al, Class. Quantum Grav. 20 (2003)

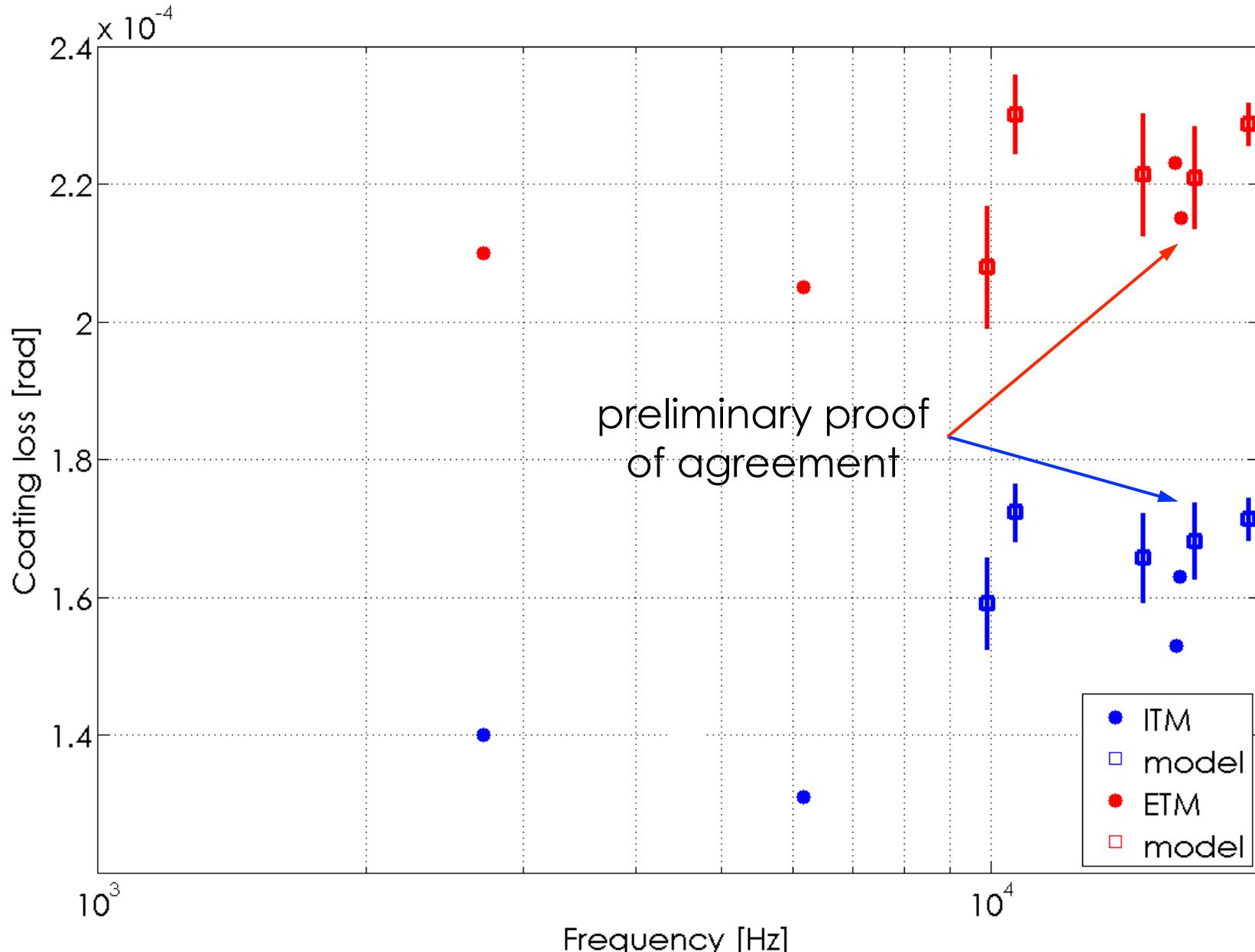
# monolayers



# stack loss – theory & measurements



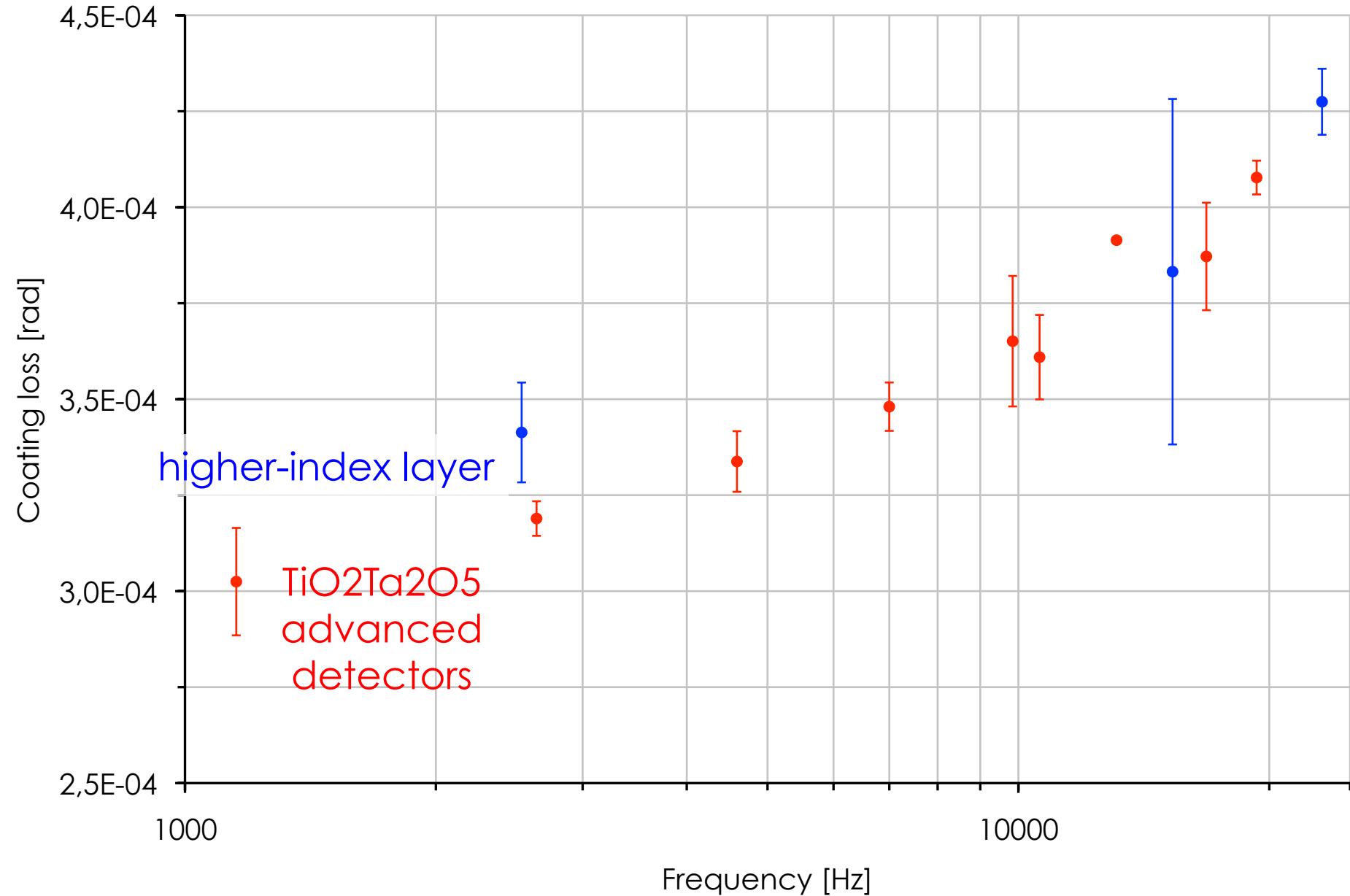
# stack loss – theory & measurements



PRELIMINARY

new higher-index layer

# comparison



# improvement

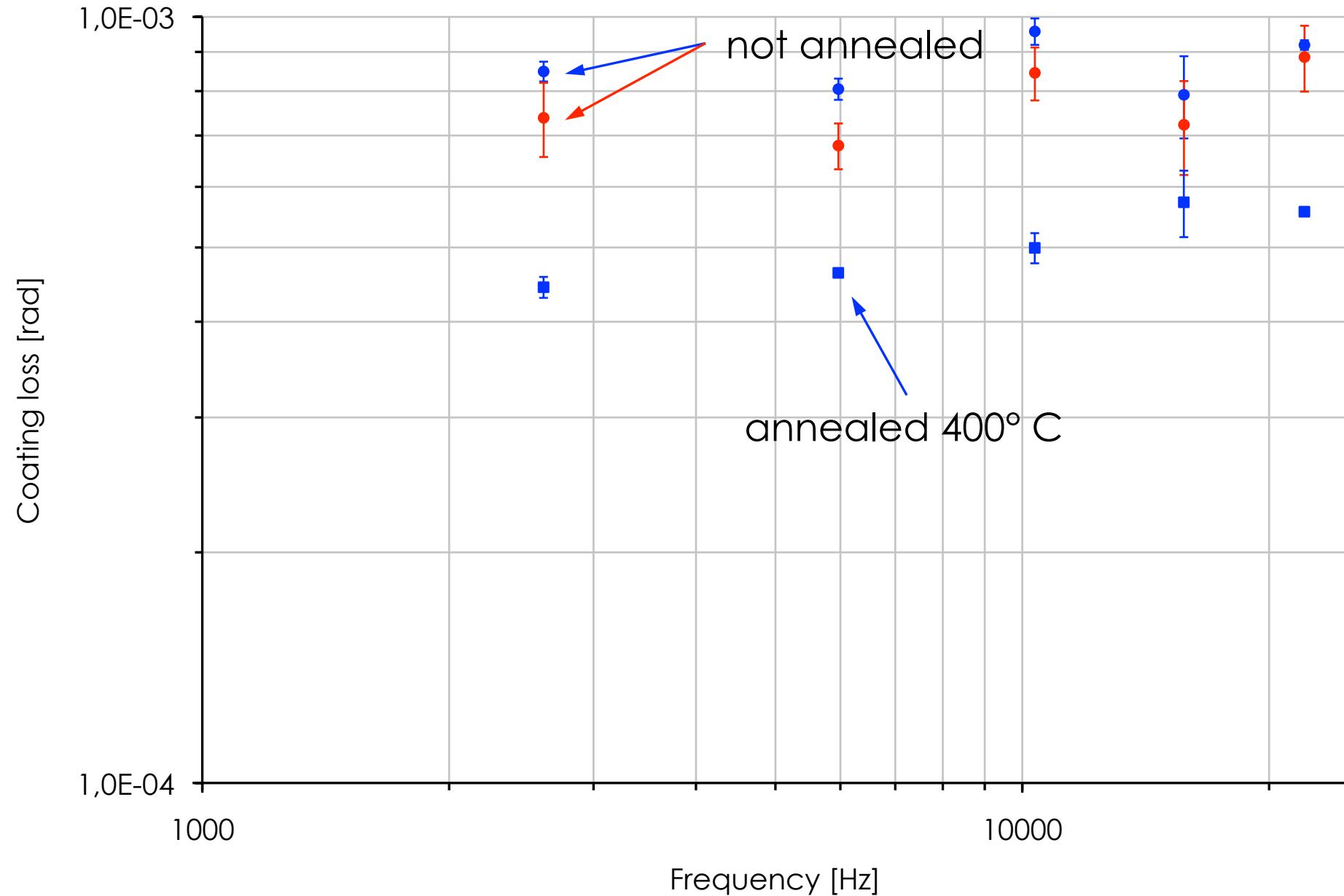
	TiO <sub>2</sub> Ta <sub>2</sub> O <sub>5</sub>	new high-index layer
$\Phi_{\sim 2.5 \text{ kHz}}$	$(3.20 \pm 0.04) \cdot 10^{-4}$	$(3.41 \pm 0.13) \cdot 10^{-4}$
$\Phi_{\sim 14 \text{ kHz}}$	$3.9 \cdot 10^{-4}$	$(3.8 \pm 0.5) \cdot 10^{-4}$
$n_{1064 \text{ nm}}$	$2.050 \pm 0.005$	2.095

→ similar loss  
higher index → thinner layers → 6% thinner ETM stack  
[-1 doublet]  
lower stack loss  
expected

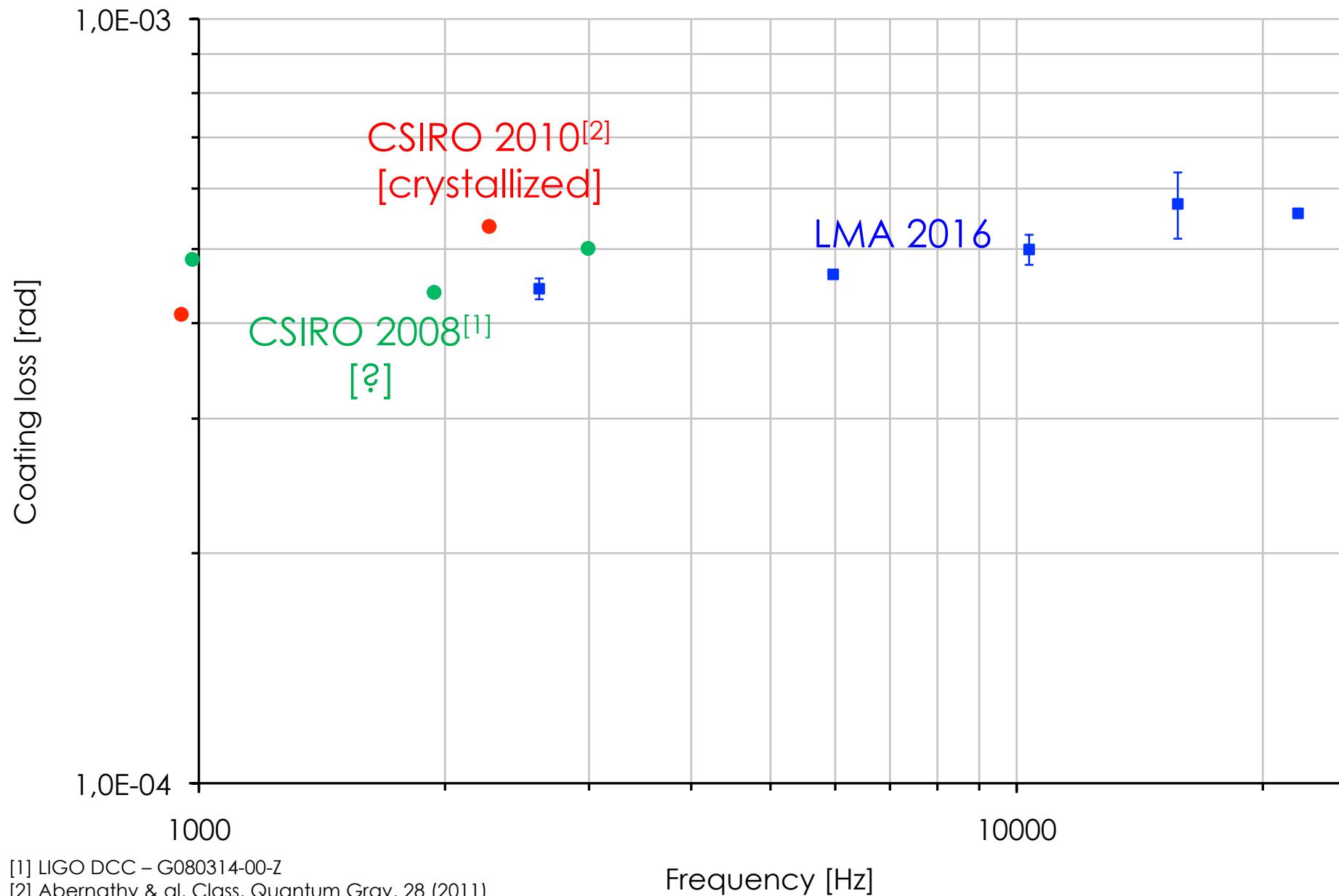
PRELIMINARY

## HfO<sub>2</sub> for cryogenic detectors

# HfO<sub>2</sub> – Spector



# comparison



[1] LIGO DCC – G080314-00-Z

[2] Abernathy & al, Class. Quantum Grav. 28 (2011)

Frequency [Hz]

# characterization

optical absorption / scattering measurements ongoing  
preliminary results → amorphous material

X-ray diffraction measurements ongoing  
as deposited / 300° C- / 400° C-annealed monolayers

# summary & conclusions

# coating loss

dilution factor can be measured

# coating loss

dilution factor can be measured

loss depends on

coating deposition parameters

post-deposition history

→ shared protocol needed to validate/compare measurements

# coating loss

dilution factor can be measured

loss depends on

coating deposition parameters

post-deposition history

→ shared protocol needed to validate/compare measurements

advanced detectors' materials

new characterization completed

doping impact revised

preliminary proof of HR stacks' loss model

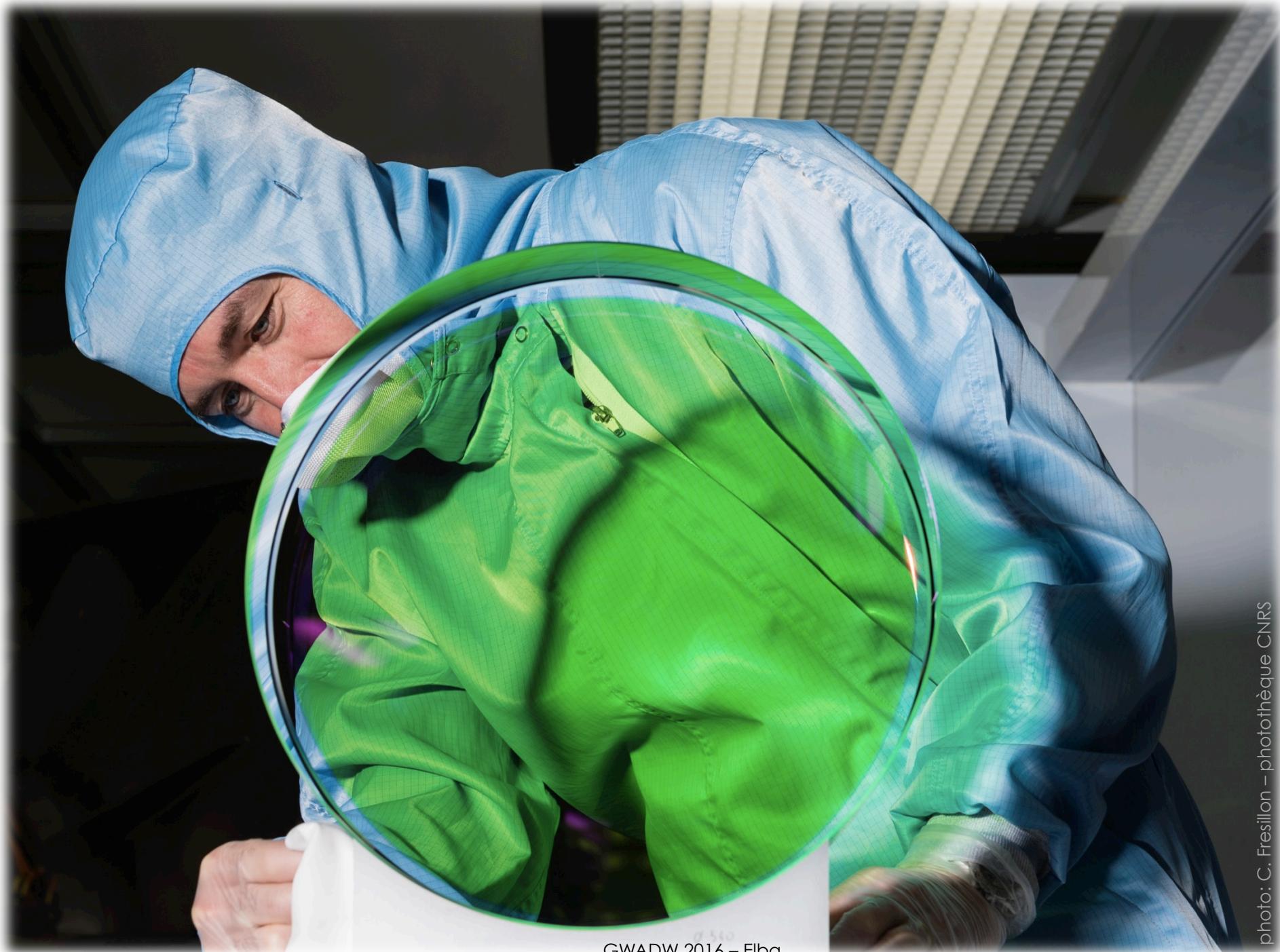
coating loss

new amorphous materials

new higher-index layer

HfO<sub>2</sub> – reconsidered for cryogenic detectors

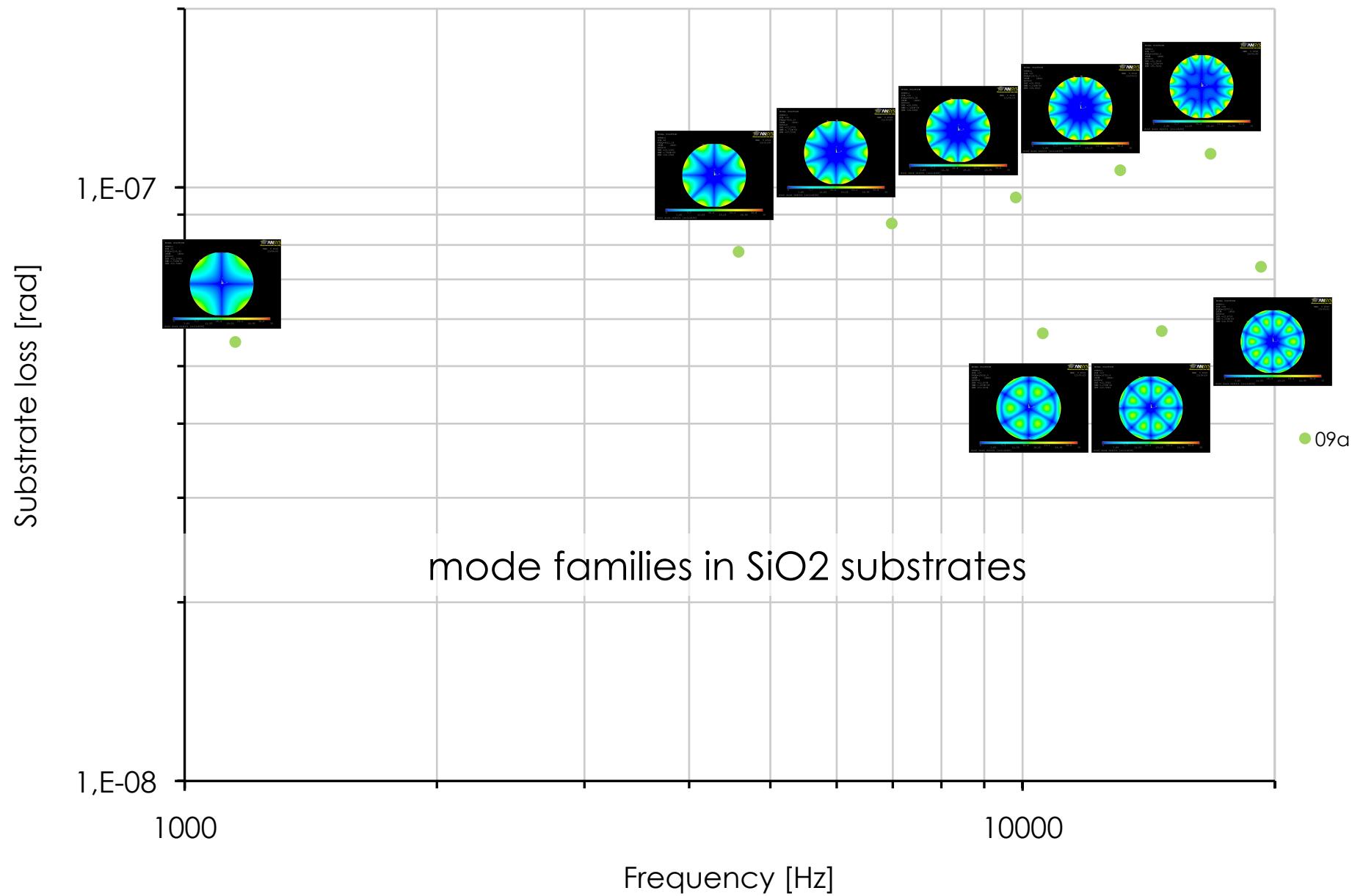
work ongoing



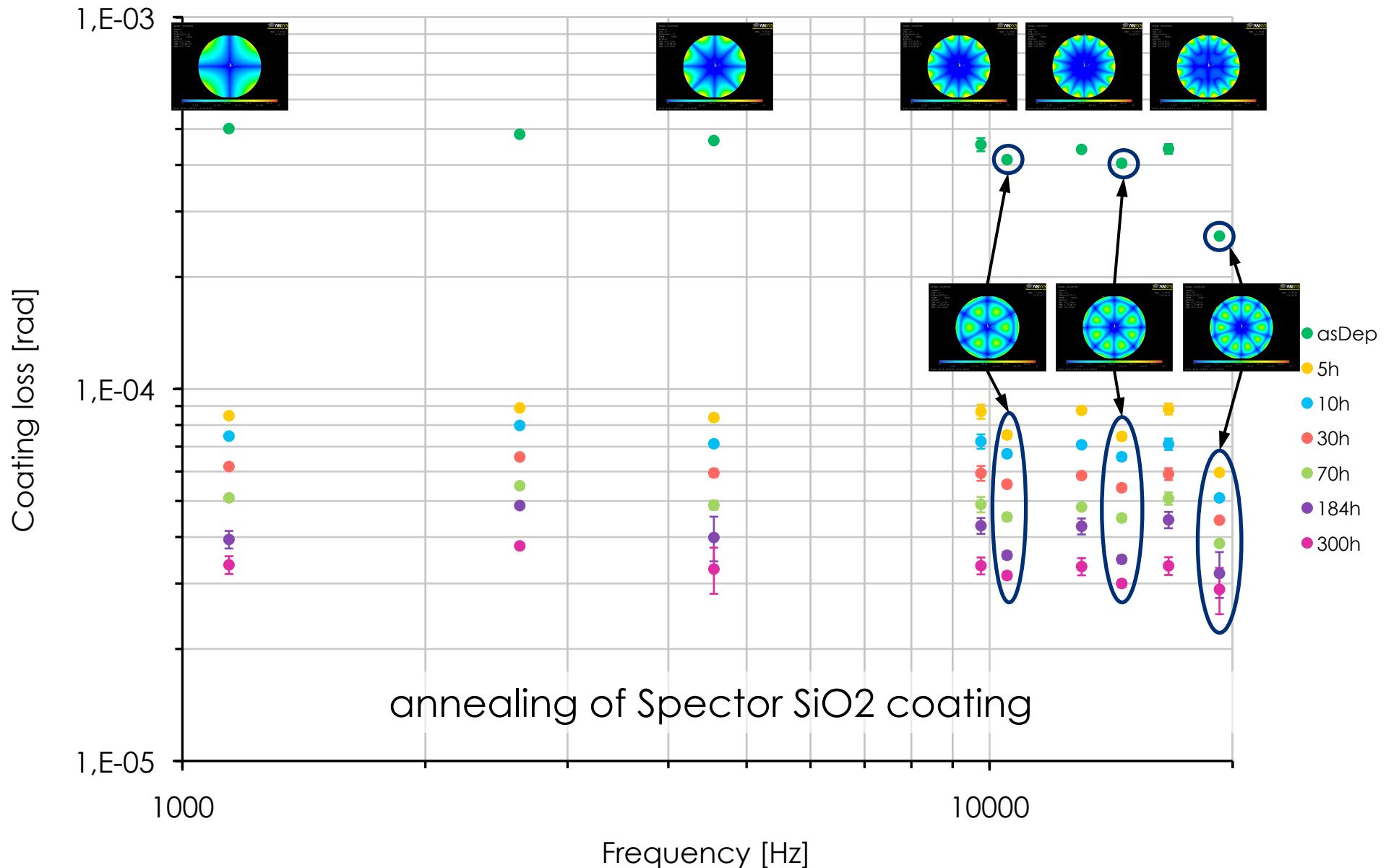
GWADW 2016 – Elba

photo: C. Fresillon – photothèque CNRS

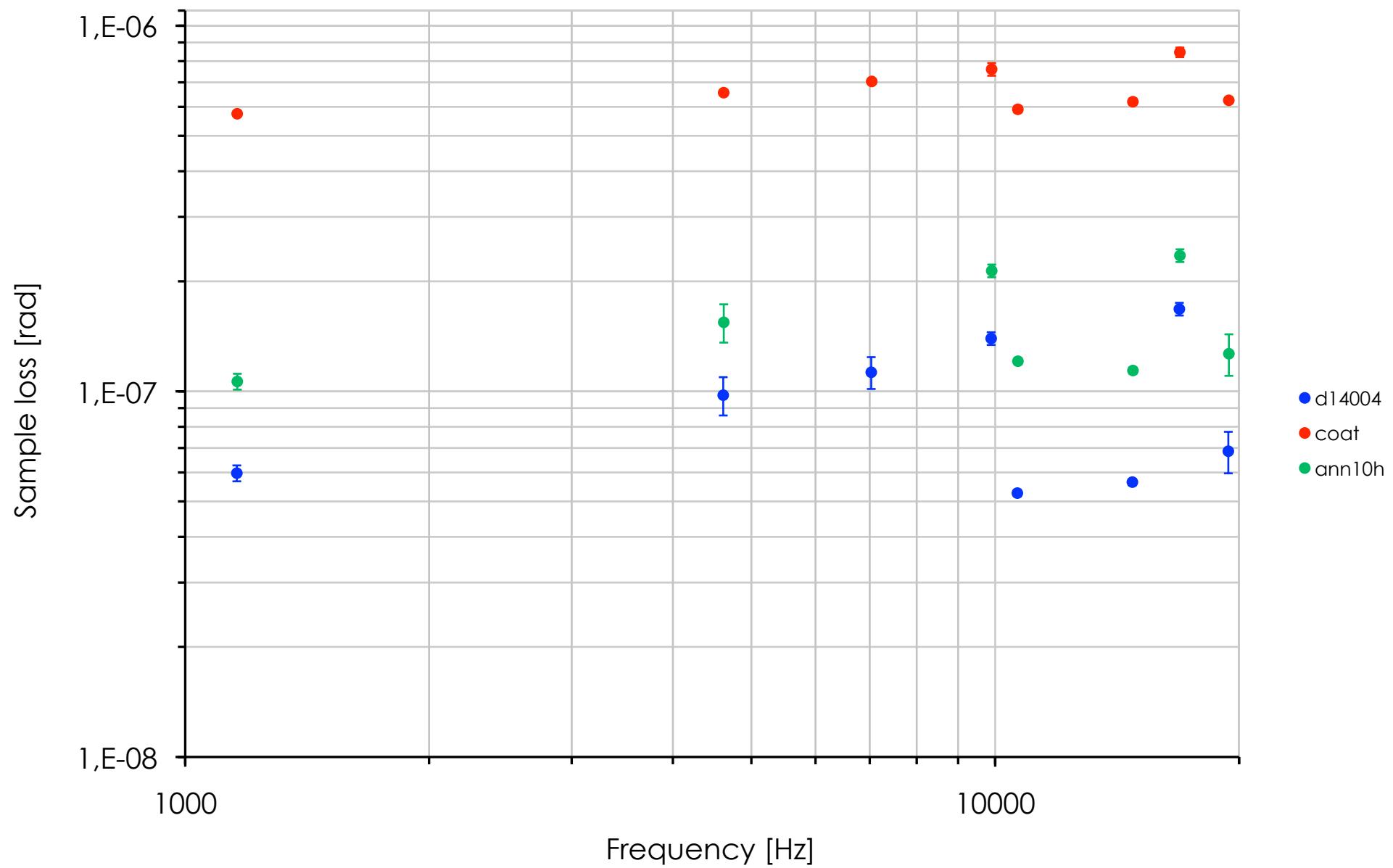
# sample 9



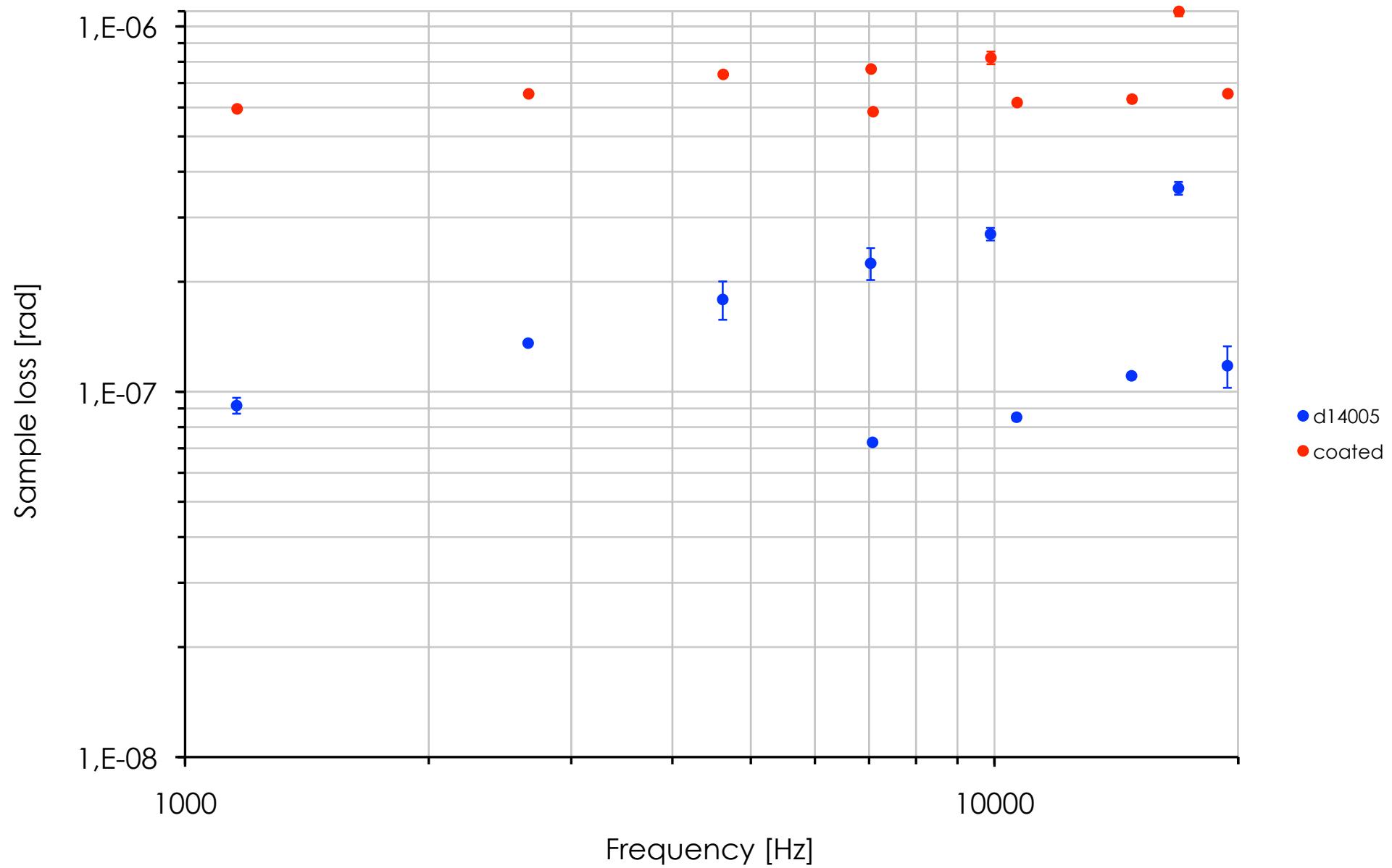
# sample 2



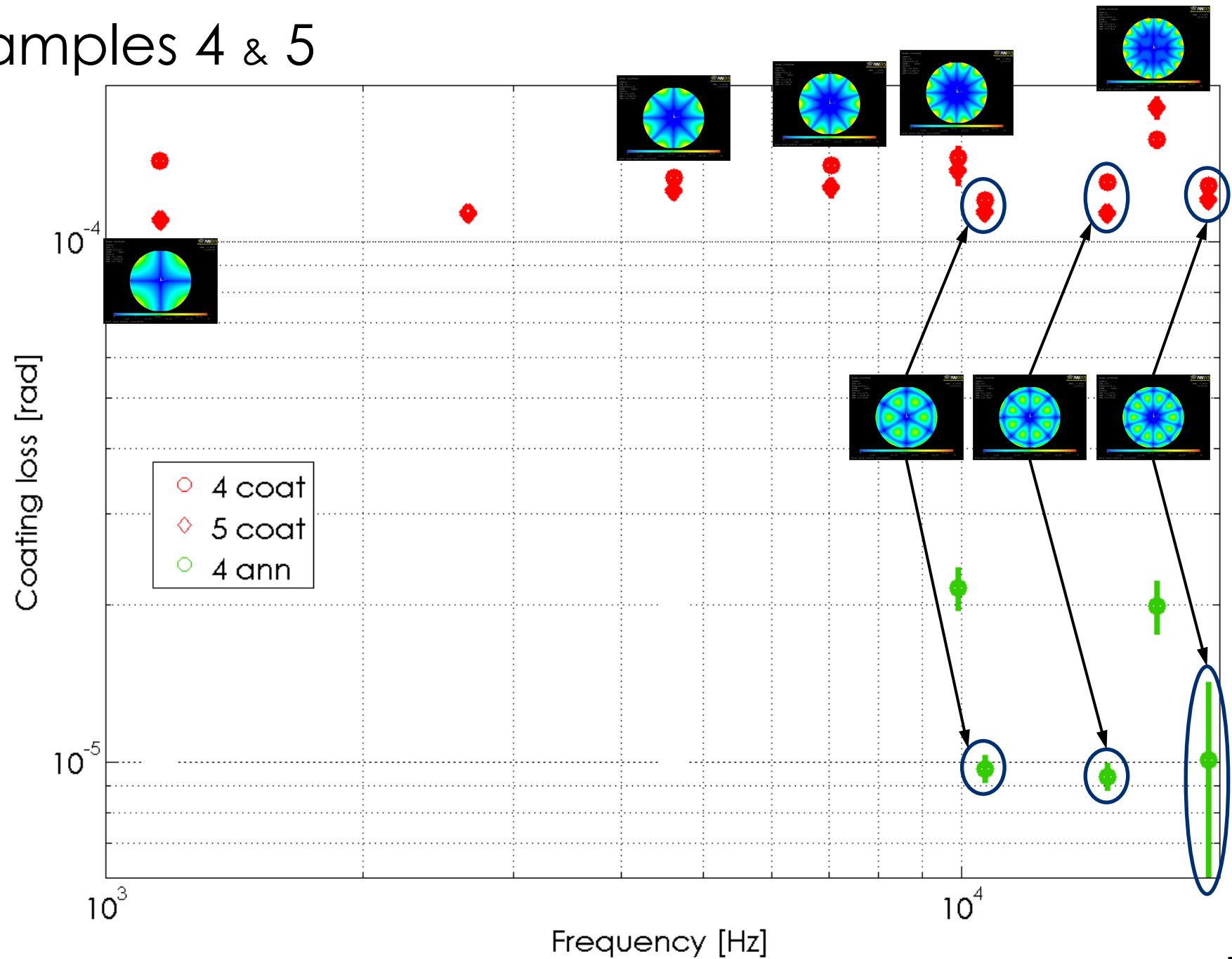
# sample 4



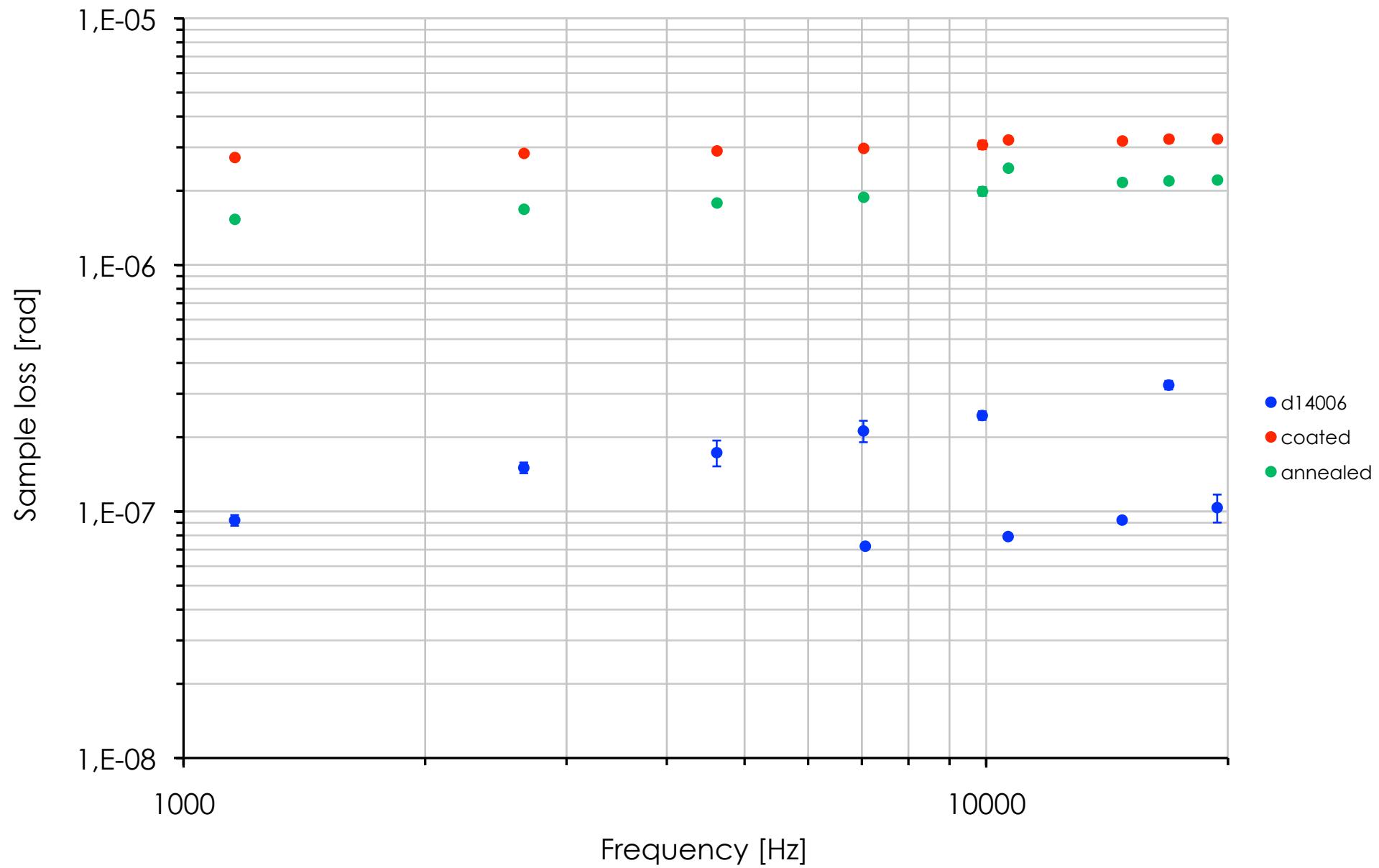
# sample 5



# samples 4 & 5



# sample 6



# sample 7

