

# Search for dark matter with the XENON1T detector

*Les Rencontres de Physique de la Vallée d'Aoste*

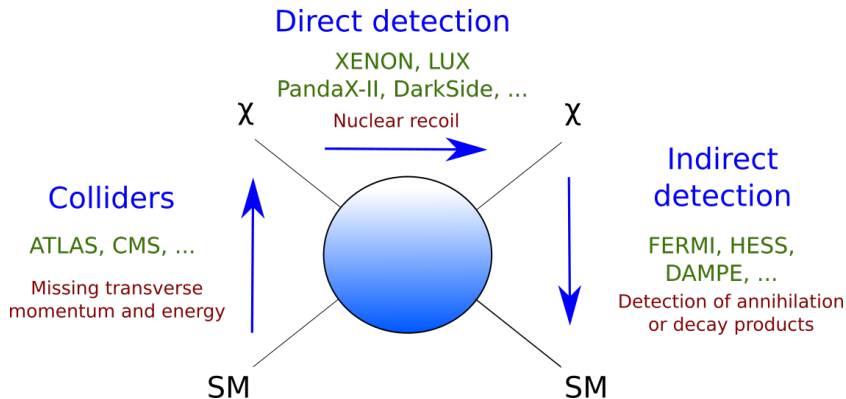
**Guillaume Eurin**  
for the XENON collaboration

Max-Planck-Institut für Kernphysik, Heidelberg

2018/02/27



# Paths to search for particle dark matter

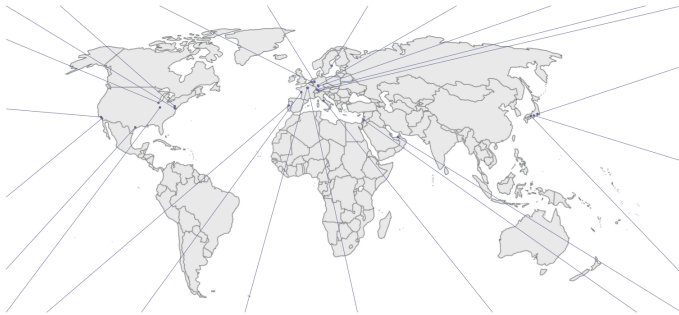


- ▶ Direct: recoil of [1-100] keV for scattering of WIMPs with masses in [1-1000]  $\text{GeV}/c^2$
- ▶ Colliders: collision of SM particles producing DM particles
- ▶ Indirect: self-annihilation thanks to gravitational accumulation or dark matter decay

# The XENON collaboration

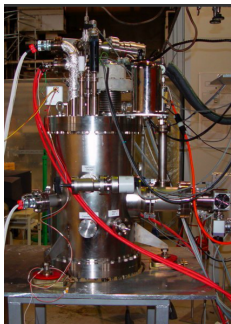


UCSD



# From XENON10 to XENONnT

- ▶ Since 2005, XENON detectors operated at LNGS
- ▶ XENON10: Target (total) mass of 14 (25) kg
- ▶ XENON100: Target (total) mass of 62 (161) kg
- ▶ XENON1T: Target (total) mass of 2000 (3200) kg
- ▶ XENONnT: Planned target (total) mass of 6000 (8000) kg



2005 - 2007

Guillaume Eurin MPIK

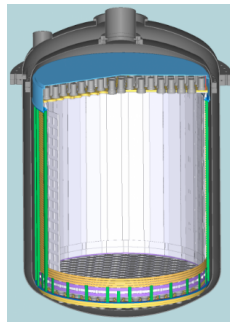


2008 - 2016

Search for dark matter with XENON1T



2013 - 2018



From 2019

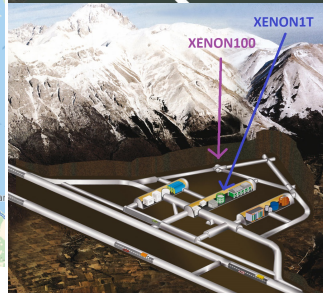
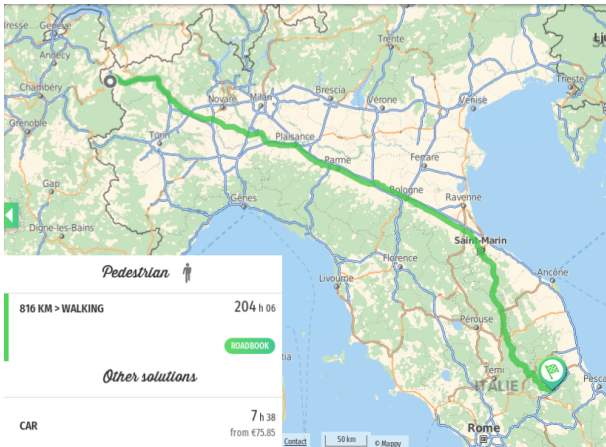
2018/02/27

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# XENON1T @ LNGS

- ▶ *Laboratori Nazionali del Gran Sasso* in central Italy
- ▶ Located inside the 10 km long *Traforo del Gran Sasso*
- ▶ XENON1T in Hall B with an overburden of 3600 m.w.e.





EPJ C (2017), 77, 881

Purification systems  
Cryogenic systems

DAQ and Slow Control

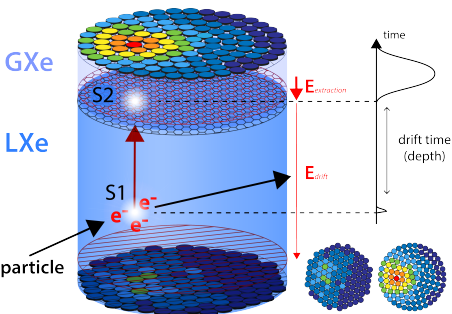
Krypton distillation column

ReStoX (Xenon storage and  
recuperation)  
Bottle rack (Xenon filling)

Water tank: Time Projection Chamber (TPC)  
and Cherenkov muon veto

# XENON1T: Dual phase Xenon TPC

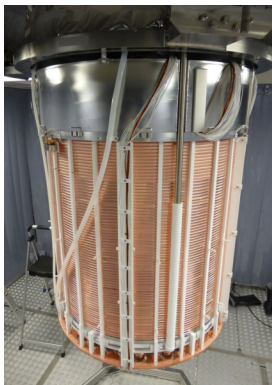
- ▶ Energy deposits from interaction  $\Rightarrow$  excitation and ionization of LXe
- ▶ Light signal (S1) from scintillation after deexcitation



- ▶ Ionization  $e^-$ 's drift upwards in E field
- ▶  $e^-$ 's extracted at LXe/GXe interface to excite and ionize GXe atoms
- ▶ Secondary scintillation  $\propto N_{e^-}$  extracted (S2)
- ▶ S2 observed by both PMT arrays, S1 mostly by bottom array
- ▶ Drift time and S2 pattern provide 3D position of the initial interaction

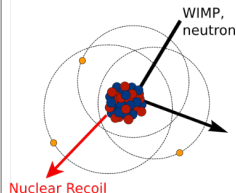
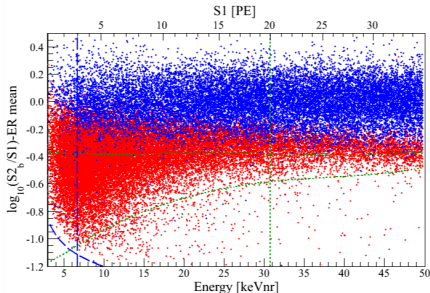
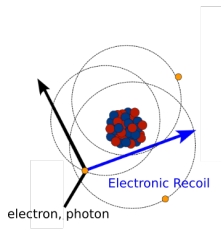
# XENON1T @ LNGS

- ▶ Dual-phase xenon TPC with 2t of LXe in active volume  
EPJ C (2017), 77, 881
- ▶ 2 PMT arrays (127, top & 121, bottom) with average  
QE of 34.5% @ 178 nm EPJC 75 (2015) 11, 546
- ▶  $\sim 1$  m drift length



# Expected signal and backgrounds

- ▶ Nuclear and electronic recoils through interactions with nucleus (WIMPs, neutrons) or atomic electrons ( $\beta$ s,  $\gamma$ s)  
 $\Rightarrow$  excitation and ionization of LXe
- ▶ Discrimination between ER and NR possible due to different contributions to ionization and excitation ( $S2/S1$  ratio)



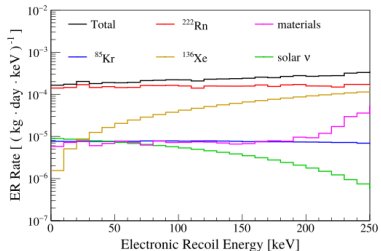
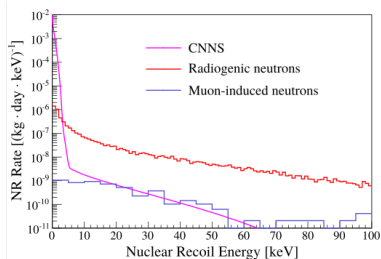
# Background sources for XENON detectors

## ► Nuclear recoils:

- Neutrons from muon spallation  
⇒ underground laboratory + Cherenkov muon veto
- $(\alpha, n)$  reactions and spontaneous fission  
⇒ material selection
- Coherent Neutrino-Nucleus Scattering (CNNS)

## ► Electronic recoils:

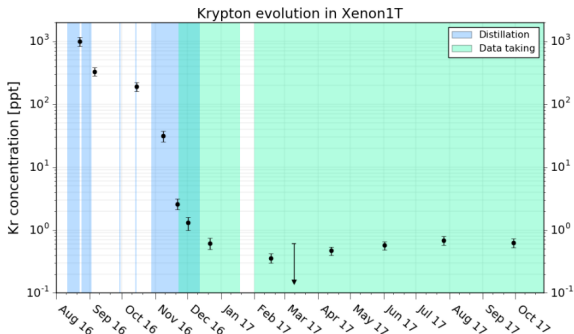
- $\gamma$ s from natural radioactivity  
⇒ xenon self-shielding and material selection
- Internal contamination:
  - $^{136}\text{Xe}$  Two-neutrino double- $\beta$  decay
  - $^{85}\text{Kr}$  from nuclear power plant operation  
⇒ cryogenic distillation
  - Radon and daughters  
⇒ material selection and cryogenic distillation



JCAP 1604 (2016), 04, 027

# Background from krypton

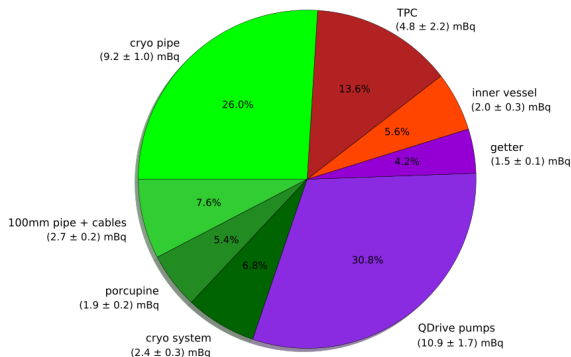
- ▶ Rare Gas Mass Spectrometer (RGMS) for Kr concentration measurements EPJ C (2014) 74, 2746
- ▶ Regular samples taken from liquid phase



- ▶ Successful krypton distillation allowing sub-ppt level EPJ C 77 (2017) 5, 275
- ▶  $(1.7 \pm 0.3) \cdot 10^{-4}$  events / (kg.day.keV) with 1300 kg FV and 5-40 keV (NR): Lowest ER background in a DM detector
- ▶ Background dominated by  $^{222}\text{Rn}$  progenies

# Background from radon

- ▶  $^{222}\text{Rn}$ : current most critical background source
- ▶ Radon budget in XENON1T:  $\sim 10\mu\text{Bq/kg}$

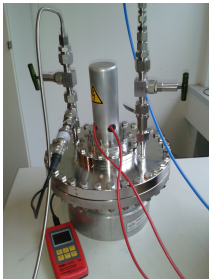


- ▶ Individual radon sources identified by emanation measurement
- ▶ Fighting strategies: material selection (HPGe, ICPMS, Rn emanation), surface cleaning, cryogenic distillation

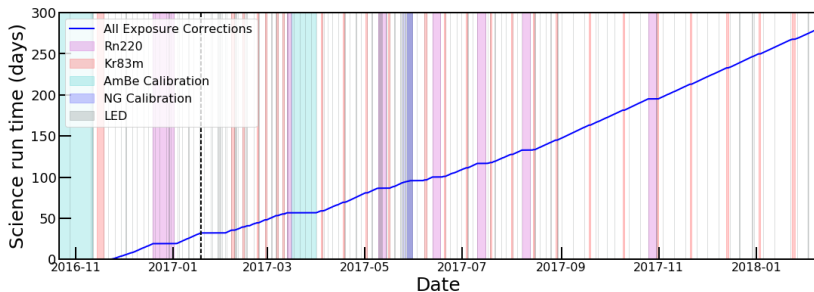


# Radon emanation measurement strategies

- ▶ Proportional counters for sensitive radon emanation measurement
- ▶ Electrostatic radon monitors
- ▶ Parallel measurements available for high sample throughput
- ▶ Automatized emanation measurements with Auto-Ema setup for reproducibility

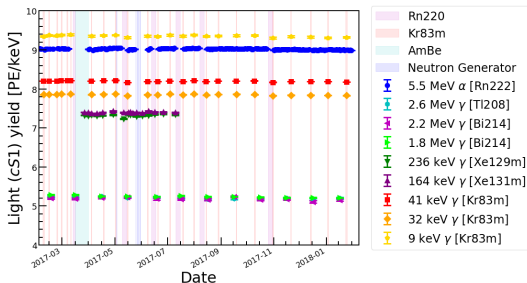


# Exposure in XENON1T science runs

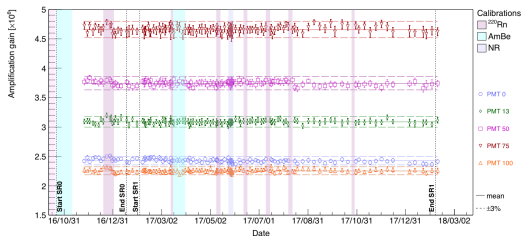


- ▶ Total exposure in XENON1T science runs (SR0+SR1):  
32.1+246.7 live days and 92% of detector uptime
- ▶ Regular ER ( $^{220}\text{Rn}$ ) and NR calibrations (AmBe and Neutron Generator)
- ▶ PMT response monitoring with LED
- ▶  $^{83\text{m}}\text{Kr}$  calibration for study of xenon purity with  $e^-$ -lifetime evolution

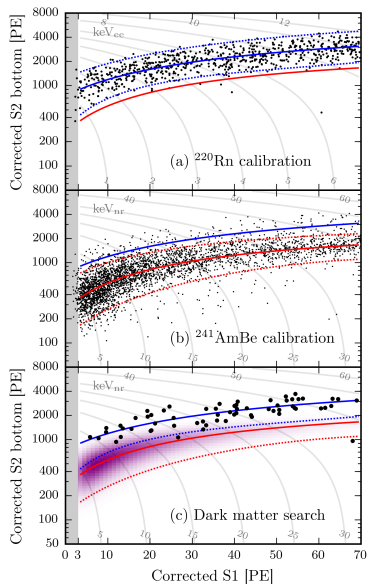
# Detector stability



- Stable charge yield, light yield and PMT gains for science data



# First result from XENON1T science runs

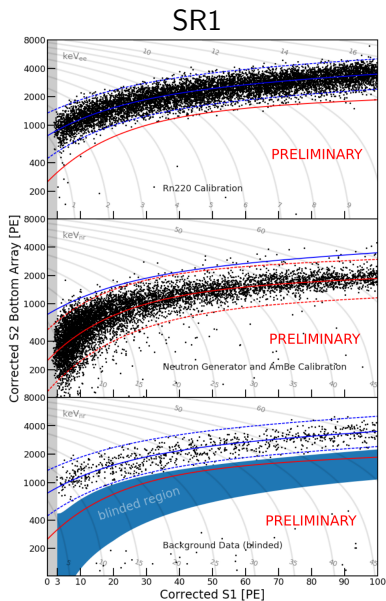
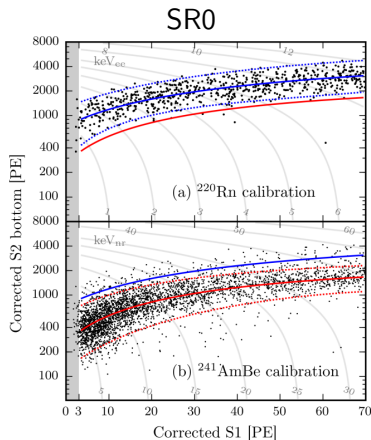


- ▶ ER band from  $^{220}\text{Rn}$  calibration data
- ▶ NR band from  $\text{AmBe}$  calibration data
- ▶ No excess above expected background in ROI  
 $\Rightarrow$  exclusion limit
- ▶ Spin-independent WIMP-nucleon scattering cross-section reached for the first time:

$$7.7 \times 10^{-47} \text{ cm}^2$$

$$@ m_\chi = 35 \text{ GeV}/c^2 \quad \text{PRL 119 (2017)}$$

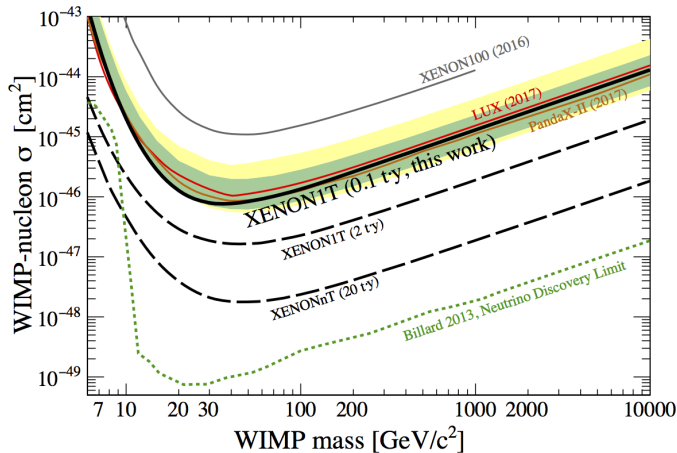
# Latest result from XENON1T science runs



- Calibration statistics improved for SR1
- Extra NR calibration with NG

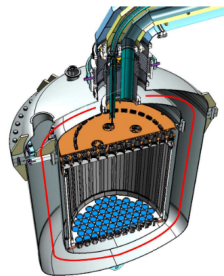
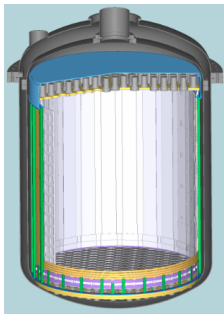
# Latest result from XENON1T science runs

- Spin-independent WIMP-nucleon scattering cross-section reached for the first time:  $7.7 \times 10^{-47} \text{ cm}^2$  @  $m_\chi = 35 \text{ GeV}/c^2$   
PRL 119 (2017)



- Sensitivity significantly improved for XENONnT

# XENONnT: direct upgrade for XENON1T

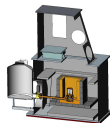
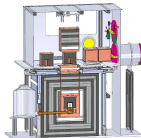


- ▶ Most subsystems of XENON1T designed for quick upgrade
- ▶ Minor modifications to cryogenic and purification system
- ▶ New TPC with an extra 246 PMTs
- ▶ Second emergency recovery system for 10t of xenon
- ▶ Should access the CNNS floor at low WIMP mass
- ▶ Preparation ongoing during XENON1T operations

# Background reduction strategies for XENONnT

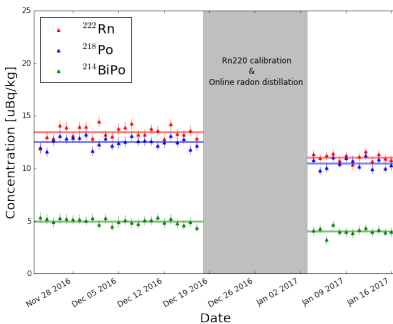
- ▶ Radon is the dominating background in XENON1T
- ▶ Material screening and selection with  $\gamma$ -spectrometry

EPJ C (2017), 77, 890



- ▶ Online reduction using cryogenic distillation,
- ▶ Proofs of principle EPJ. C (2017), 77, 143 and XENON100

EPJ C (2017), 77, 358

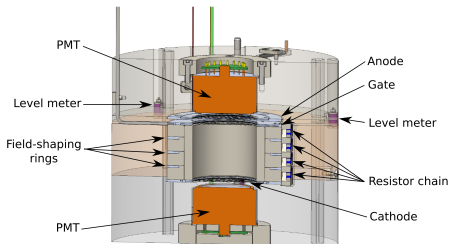
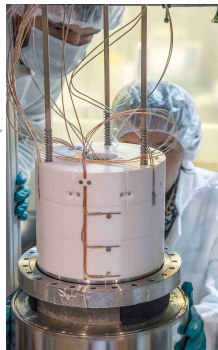
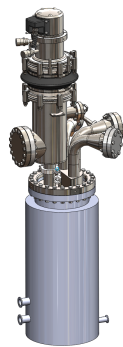


- ▶ Expected reduction factor for the column: 100



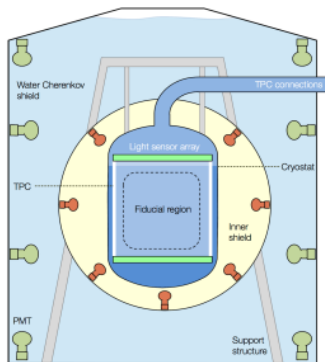
# Background reduction strategies for XENONnT

- ▶ Clean room to prevent deposition of dust containing  $^{238}\text{U}$
- ▶ Surface cleaning for  $^{222}\text{Rn}$  daughters reduction
- ▶ Validation of surface cleaning techniques:
  - 5-20 cm drift length LXe TPC
  - $e^-$ -lifetime measurement for purity monitoring after surface cleaning
  - Inner core exchangeable
  - External ports allowing larger samples to be introduced in GXe flow



HeXe  
HeidelbergXenon

# The DARWIN consortium



- ▶ 50t of xenon and 40t in the active volume
- ▶ Sensitivity to spin-independent WIMP-nucleon scattering cross-section of  $\text{few } 10^{-49} \text{ cm}^2 @ m_\chi = 50 \text{ GeV}/c^2$   
JCAP (2016), 11, 017
- ▶ Science run could start in 2023 lasting at least 7 years

- ▶ Other potential physics goals:
  - Detection of events from CNNS
  - $pp\text{-}\nu$  detection from the Sun
  - Search for neutrinoless double- $\beta$  decay
  - Observation of supernova neutrinos

- ▶ XENON1T: first tonne-scale LXe TPC for dark matter searches
- ▶ Outstanding first results on SI WIMP-nucleon scattering cross-section
- ▶ Lowest background level for DM searches with  $0.2 \text{ events} \times \text{tonne}^{-1} \times \text{keV}^{-1} \times \text{d}^{-1}$
- ▶ 1 tonne  $\times$  year exposure will yield world-leading result
- ▶ Data from second science run still blinded  
Results to be announced in March 2018
- ▶ XENON1T continues data taking until installation of XENONnT, end of 2018

<http://www.xenon1t.org/>  
Twitter: @Xenon1T

# Thank you for your attention!

