

LNGS beyond 2020, LNGS, April 28th, 2015

INFN: Post-2020 glimpses

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INFN and Univ. of Padova

much depends on the next 5 years ...

- **LHC14** (high energy: ATLAS, CMS; flavor: LHCb; quark-hadron phase transition: ALICE)
- **Flavor**: NA62; upgraded MEG, Mu-e; BELLEII; EDMs; g-2
- **DM** 1-ton exps. → $10^{-10} - 10^{-11}$ pb
- **Neutrinoless double β** → ν mass degenerate region; enter IH region
- **SBN** → sterile ν ?
- **Gravitational waves** → discovery
- **DE**: BOSS → DESI; DES → LSST
- **CMB**: final PLANCK; B-modes of the polariz.+ black-body spectrum : EU exps. QUBIC, LSPE, QIJOTE + many others on ground and balloons in US, Japan

if

- **Some signal** of physics beyond what the Standard Models of Fundamental Interactions and Cosmology predict → focus and converge on that “area” of New Physics
- **No signal** of BSM physics shows up whatsoever → i) go on and move the frontier of the existing exploration paths; ii) look for alternative, “non-ballistic” paths

- By the end of the 20th century ... **we have a comprehensive, fundamental theory of all observed forces of nature which has been tested and might be valid from the Planck length scale [10^{-33} cm.] to the edge of the universe [10^{+28} cm.]**

D. Gross 2007

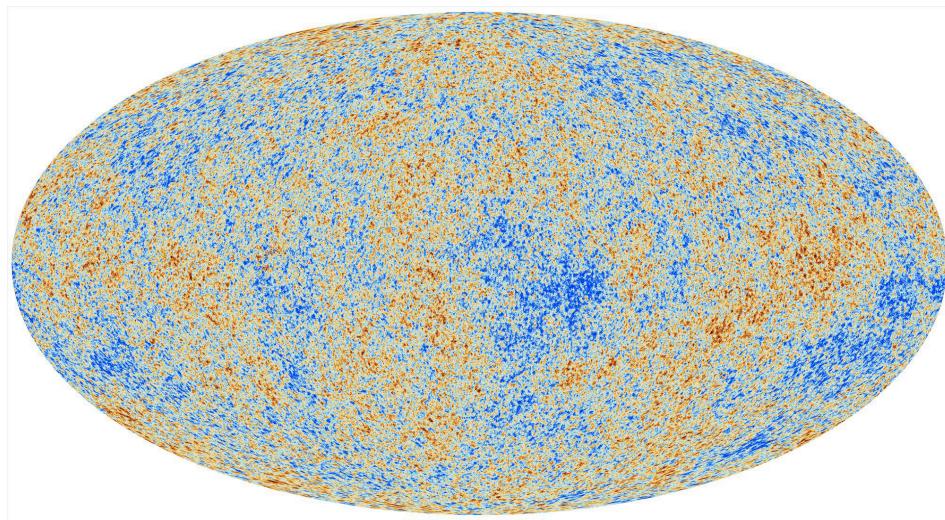
2013: the thiumph of the **STANDARD**

- **PARTICLE STANDARD MODEL**

Three Generations of Matter (Fermions) spin $\frac{1}{2}$					
mass →	I	II	III		
charge →	$\frac{2}{3}$	$\frac{1}{3}$	$\frac{1}{3}$		
name →	u Left up	c Left charm	t Left top		
Quarks	2.4 MeV $\frac{2}{3}$ d Left down	1.27 GeV $\frac{1}{3}$ s Left strange	173.2 GeV $\frac{1}{3}$ b Left bottom		
	ν_e Left electron neutrino	ν_μ Left muon neutrino	ν_τ Left tau neutrino		
Leptons	0.511 MeV e Left electron	105.7 MeV μ Left muon	1.777 GeV τ Left tau		

Bosons (Forces) spin 1					
g gluon	0 0	γ photon	0 0		
Z weak force	93.2 GeV 0 0	Higgs boson	126 GeV 0 0		
W^\pm weak force	80.4 GeV ± 1				

- **COSMOLOGY STANDARD MODEL**



Λ CDM + “SIMPLE” INFLATION

$$\Omega_\Lambda = 0.686 \pm 0.020$$

$$\Omega_m = 0.314 \pm 0.020$$

$$\Omega_b h^2 = 0.02207 \pm 0.00033$$

$$h = 0.674 \pm 0.014$$

WHY BSM

“OBSERVATIONAL” REASONS:

- Dark Matter
- Neutrino Masses
- Cosmic Matter-Antimatter Asymmetry (twofold problem: disappearance of primordial antimatter and extreme reduction of the number of baryons w.r.t the number of photons – initially ~ equal, today $n_{\text{baryons}}/n_{\text{photons}} \sim 10^{-9}$)
- Primordial Inflation
- Dark Energy

WHY BSM

Theoretical reasons (of dissatisfaction towards the SM as a “final” theory rather than actual problems for the SM)

- Lack of the theory of **Flavor** (why three fermion families, why hierarchical mass spectrum, why mixing angles so different)
- **CPV in strong interactions**, i.e. the θ -problem
- **Unification** of the fundamental interactions (running the SM gauge couplings \rightarrow clear trend for unification of the interactions, but “pure SM” fails)
- **Gauge hierarchy** – twofold puzzle: why M_{GUT} or M_{planck} $>>> M_W$; stabilization of the higgs mass at M_W at any order in perturbation theory *

New LHC / HL-LHC Plan

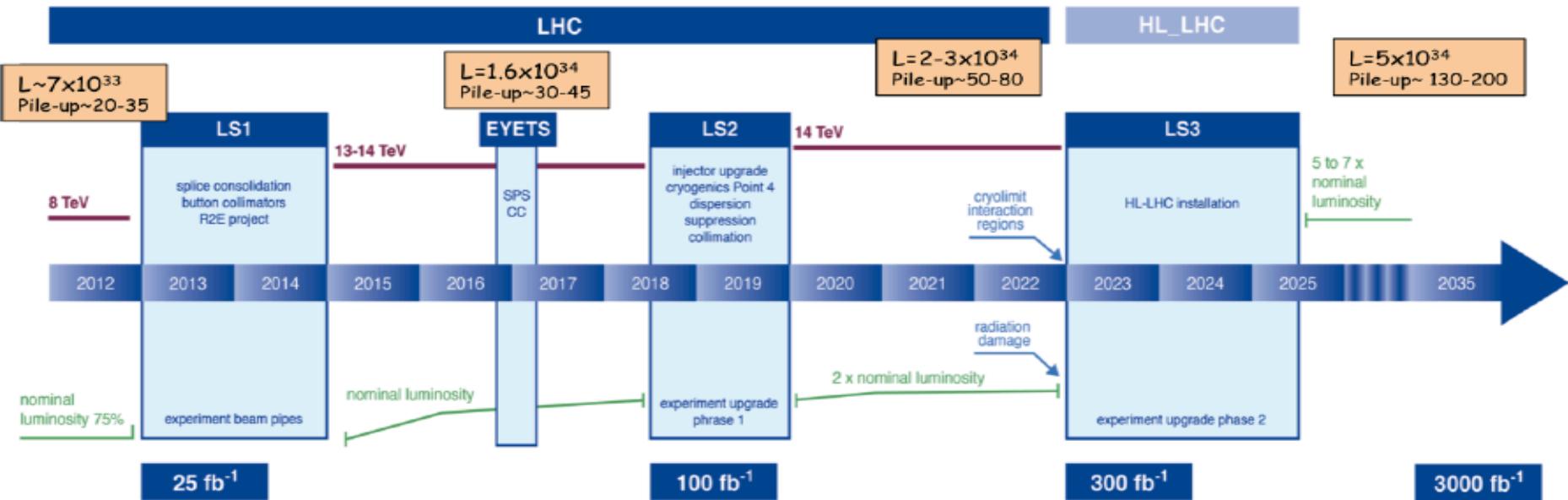


Fig. 1: LHC evolution plan

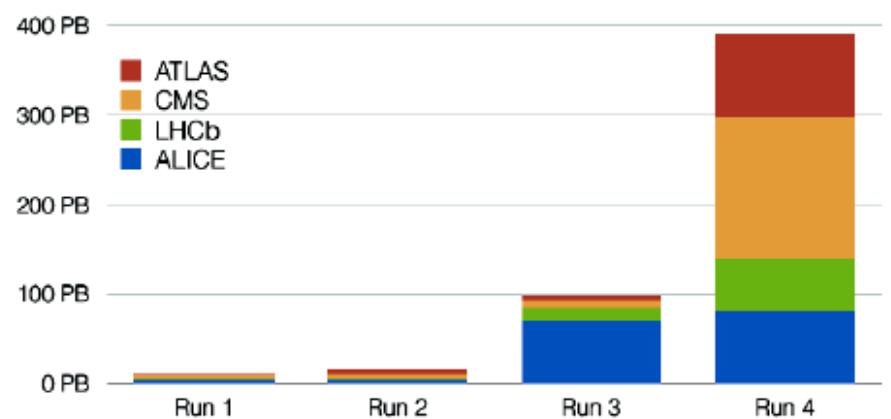
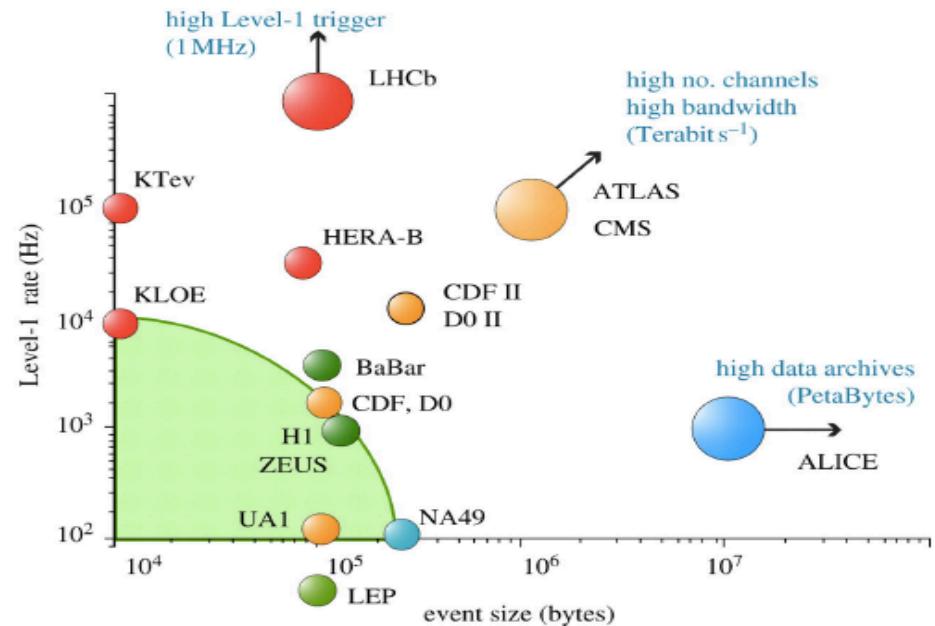
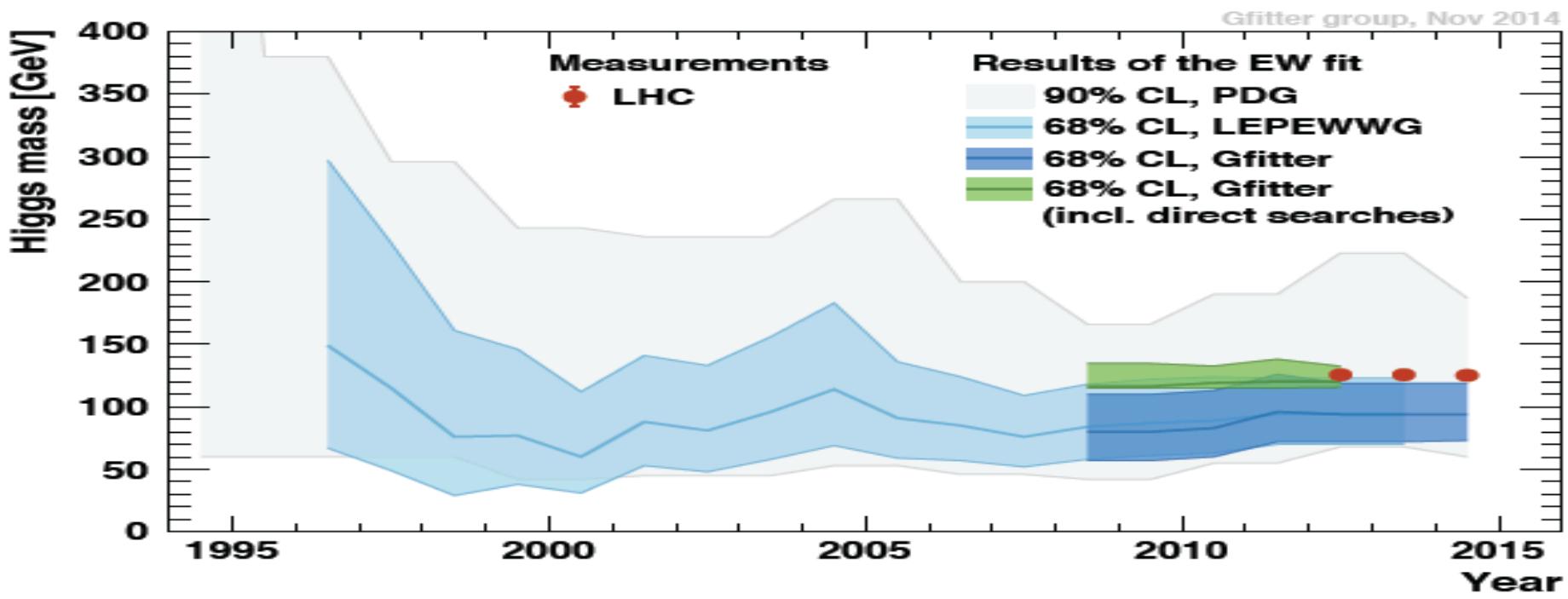
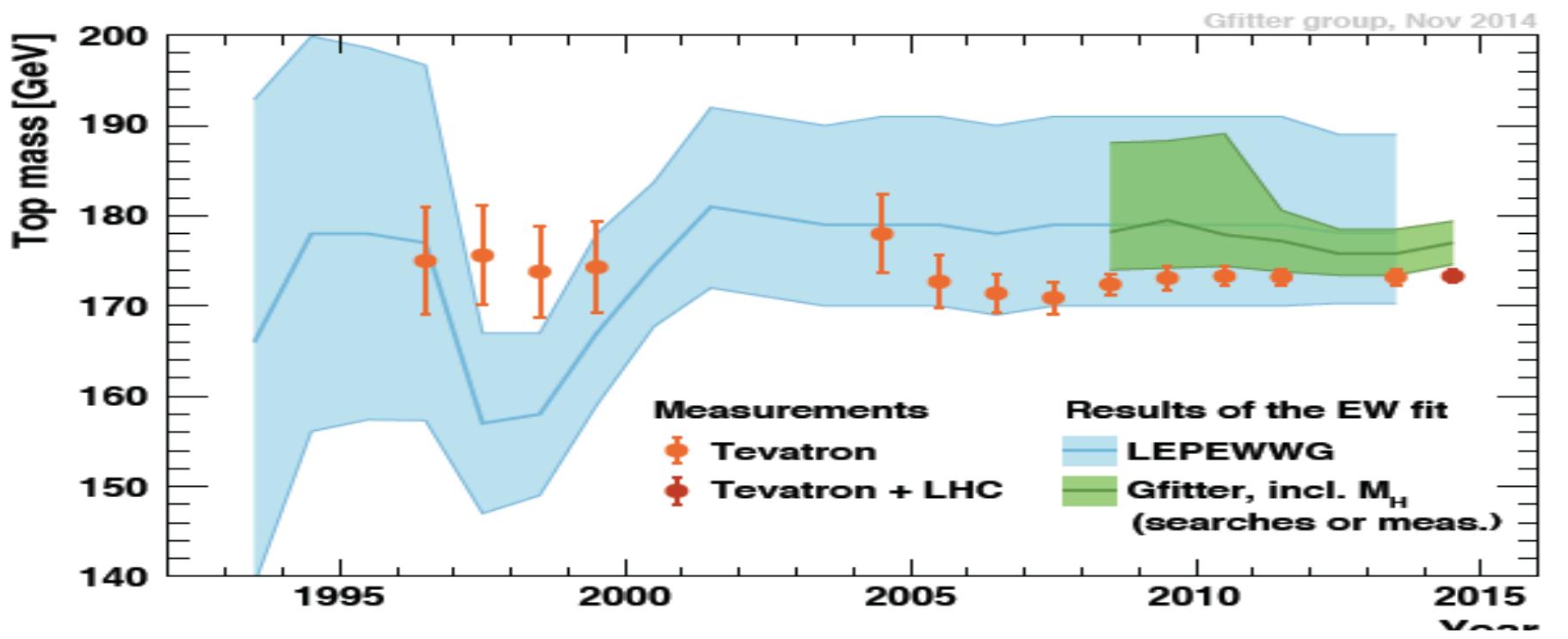


Fig. 5: Estimate of the disk space needed by the LHC experiments.



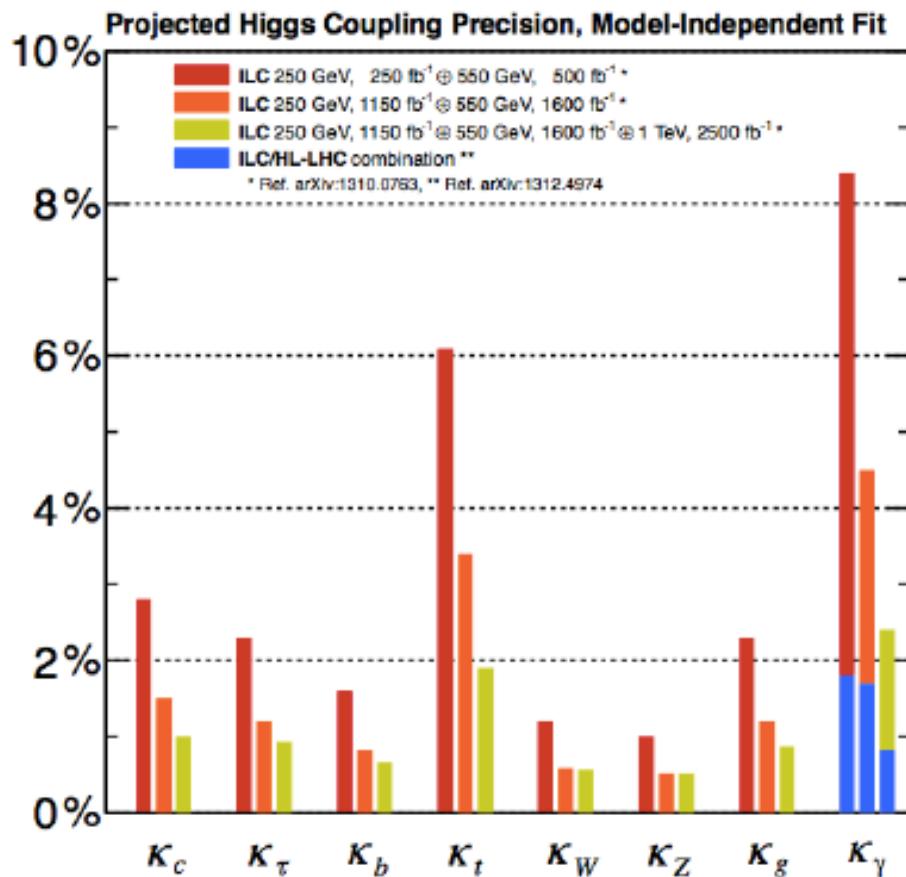
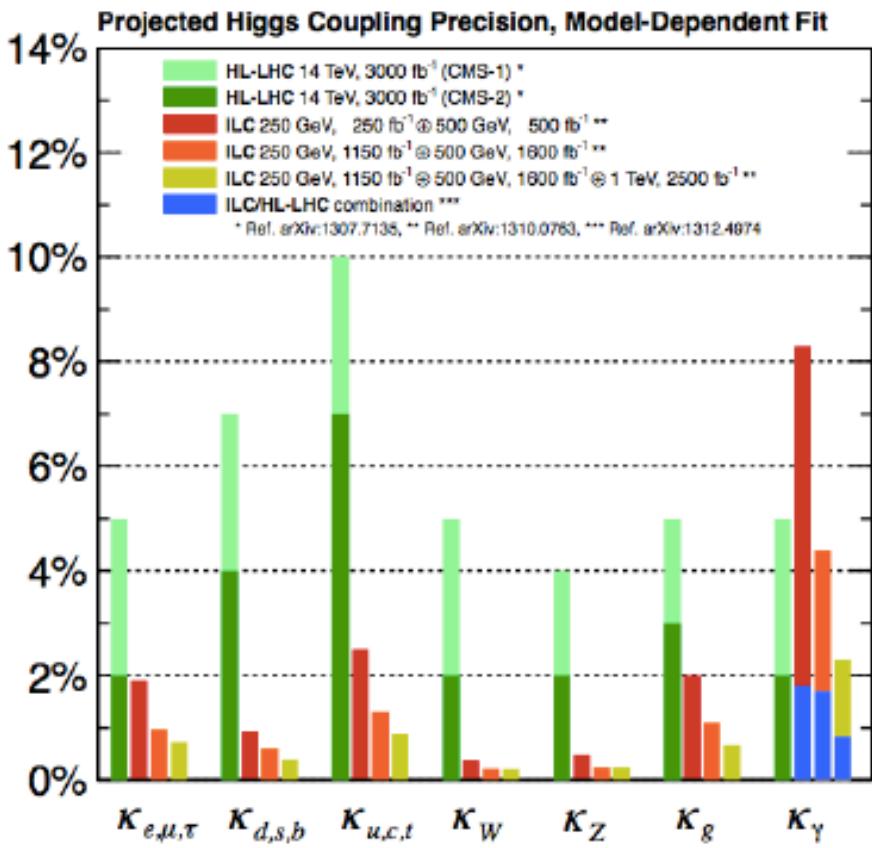
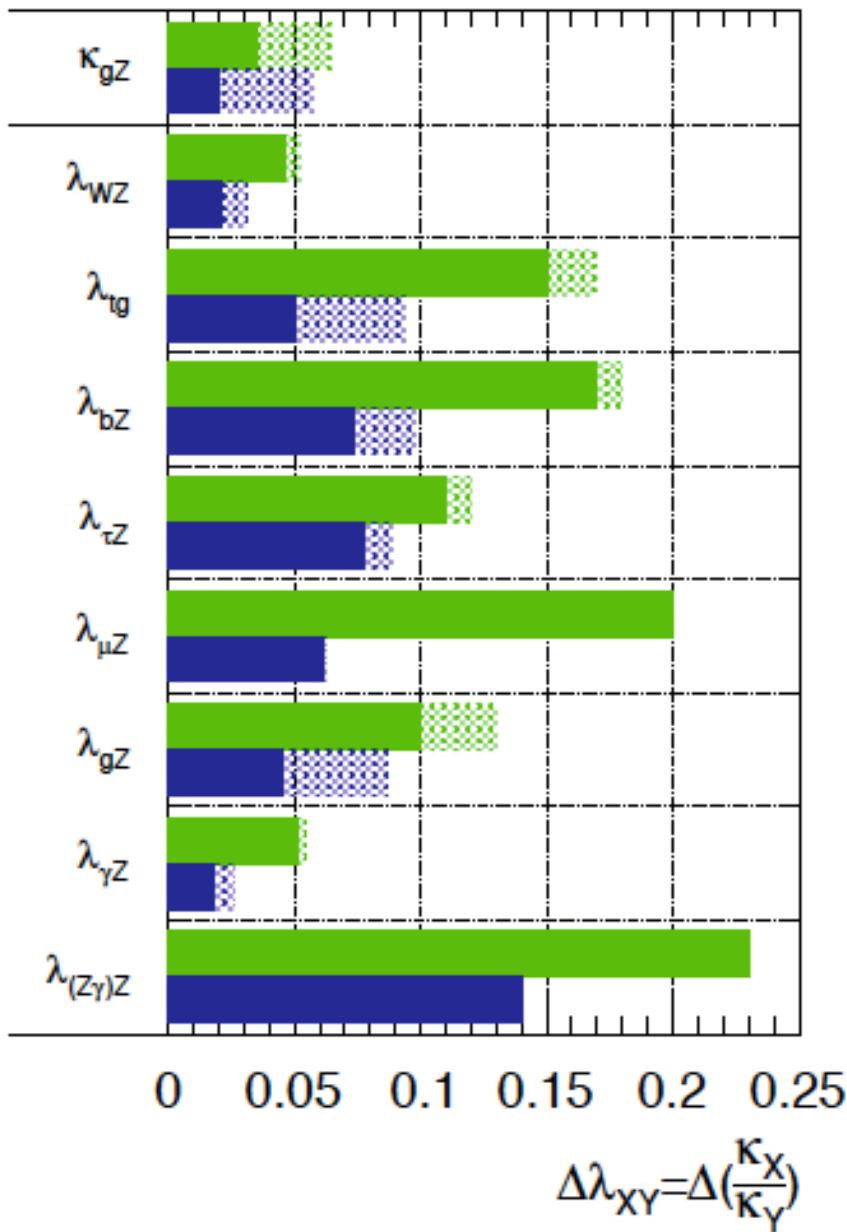


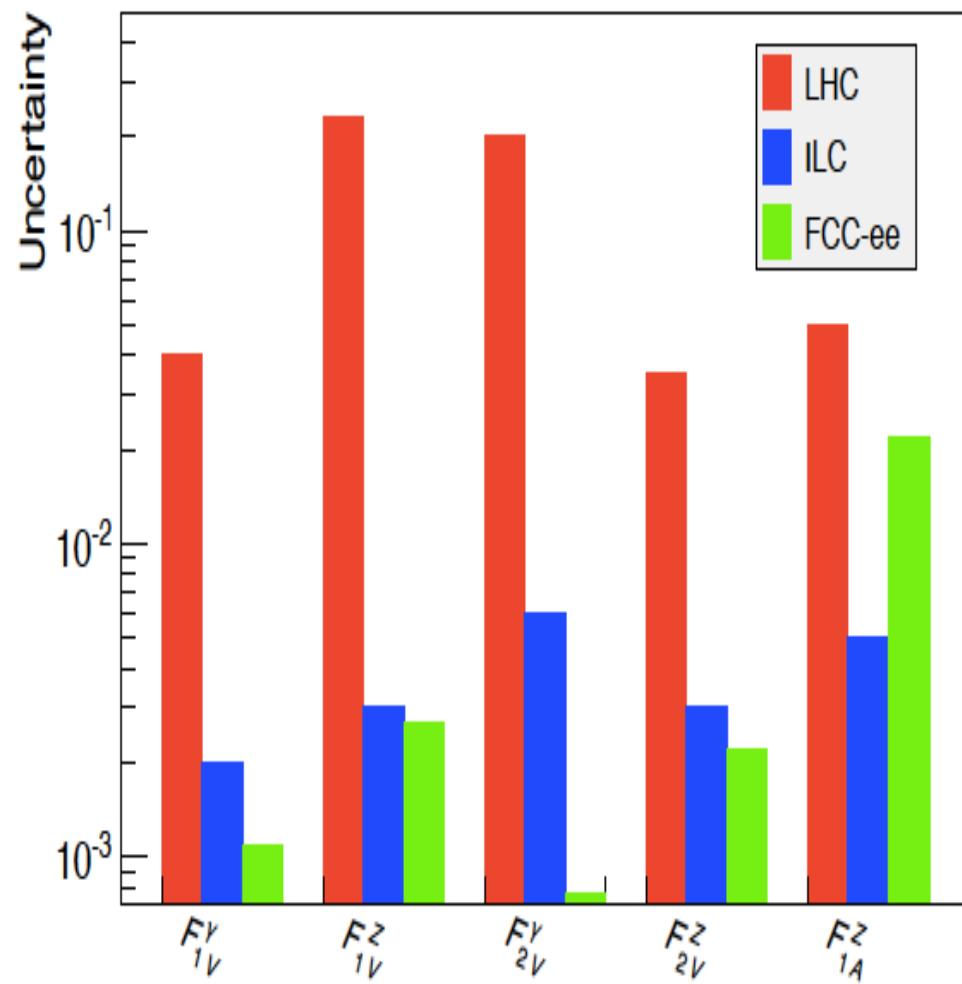
Fig. 11: Model-Dependent (left) and Model-Independent (right) projected precision on Higgs-boson coupling scaling factors. Here $\kappa_i \equiv g_{hi} / g_{hi}^{(SM)}$. Green bands refer to HL-LHC projections; blue bands refer to combinations of ILC and HL-LHC outputs.

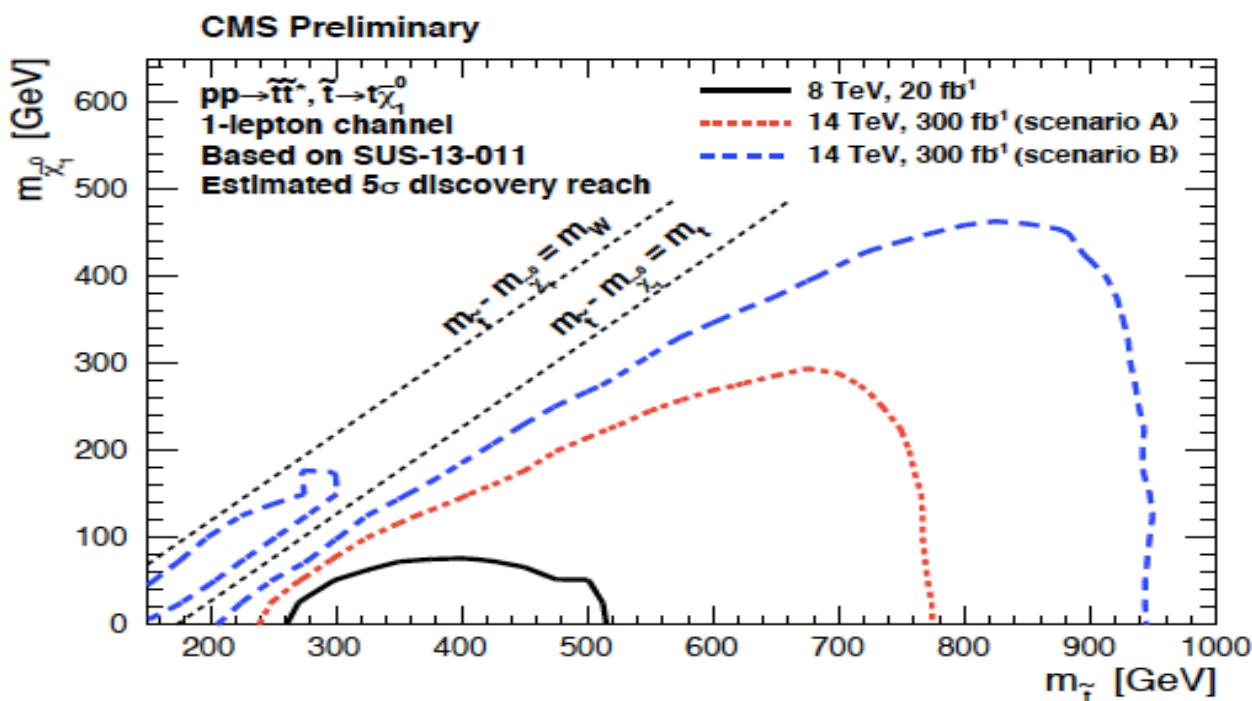
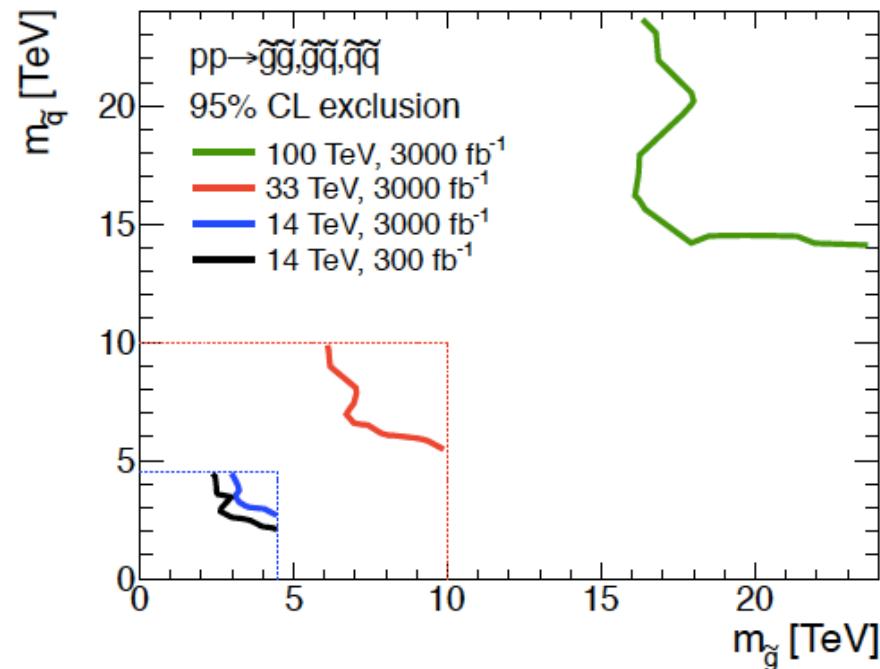
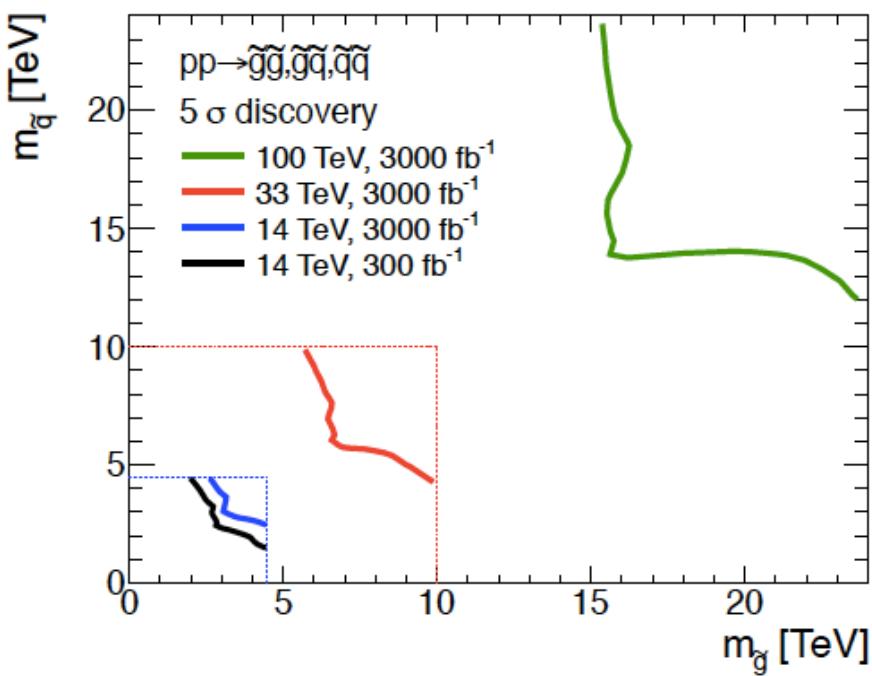
ATLAS Simulation Preliminary

$\sqrt{s} = 14 \text{ TeV}$: $\int L dt = 300 \text{ fb}^{-1}$; $\int L dt = 3000 \text{ fb}^{-1}$



Statistical precisions on vector and axial form factors at LHC with 300 fb^{-1} , at ILC500 with 500 fb^{-1} , at FCC-ee with 2.4 ab^{-1}





THE FLAVOUR PROBLEMS

FERMION MASSES

What is the rationale hiding
behind the spectrum of fermion
masses and mixing angles
(our “**Balmer lines**” problem)

→ LACK OF A FLAVOUR “THEORY”

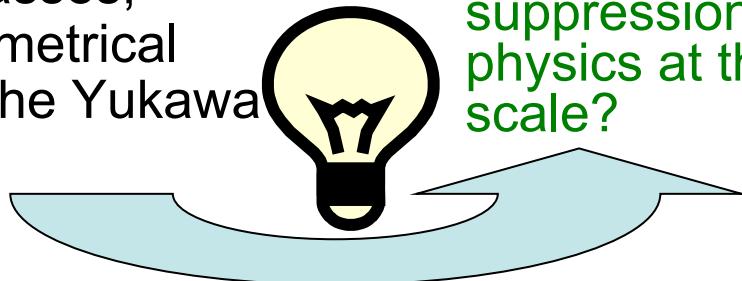
(new flavour – horizontal
symmetry, radiatively induced
lighter fermion masses,
dynamical or geometrical
determination of the Yukawa
couplings, ...?)

FCNC

Flavour changing neutral
current (FCNC) processes are
suppressed.

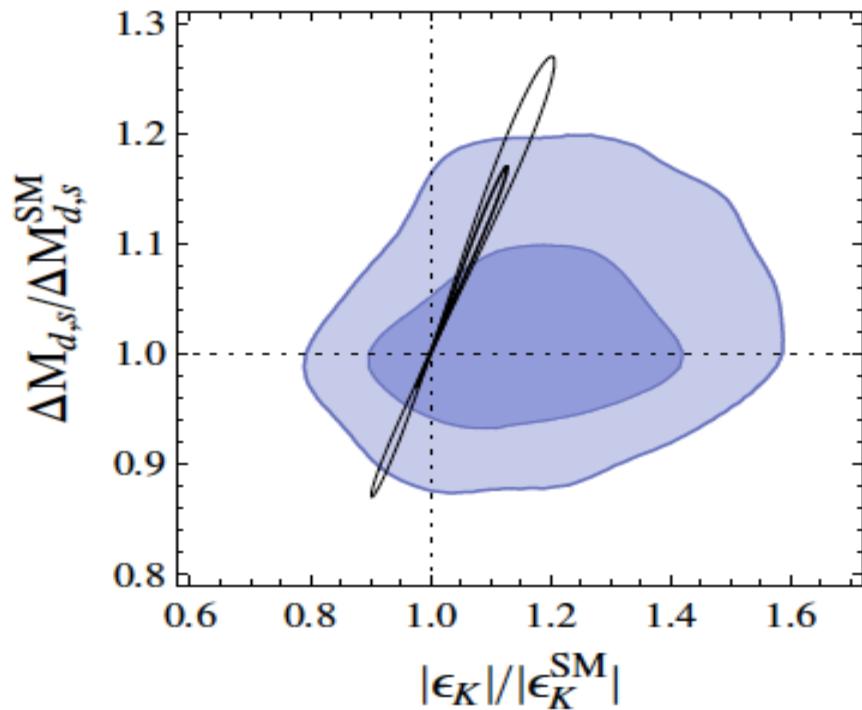
In the SM two nice
mechanisms are at work: the
GIM mechanism and the
structure of the **CKM mixing
matrix**.

How to cope with such delicate
suppression if there is new
physics at the electroweak
scale?

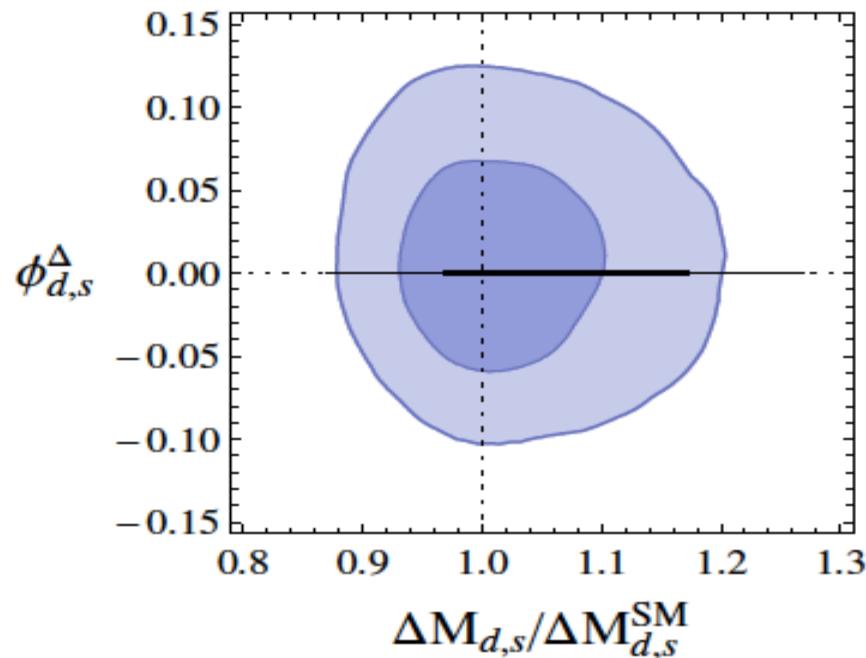


IF SOME SYMMETRY – OR “DYNAMICAL” RATIONALE EXISTS FOR THE FLAVOR PUZZLE → POSSIBILITY OF VERY SELECTIVE PLACES WHERE DEVIATIONS FROM THE CKM PARADIGM APPEAR

EX.with U(3) or SU(2)³ flavor symmetries



Barbieri et al



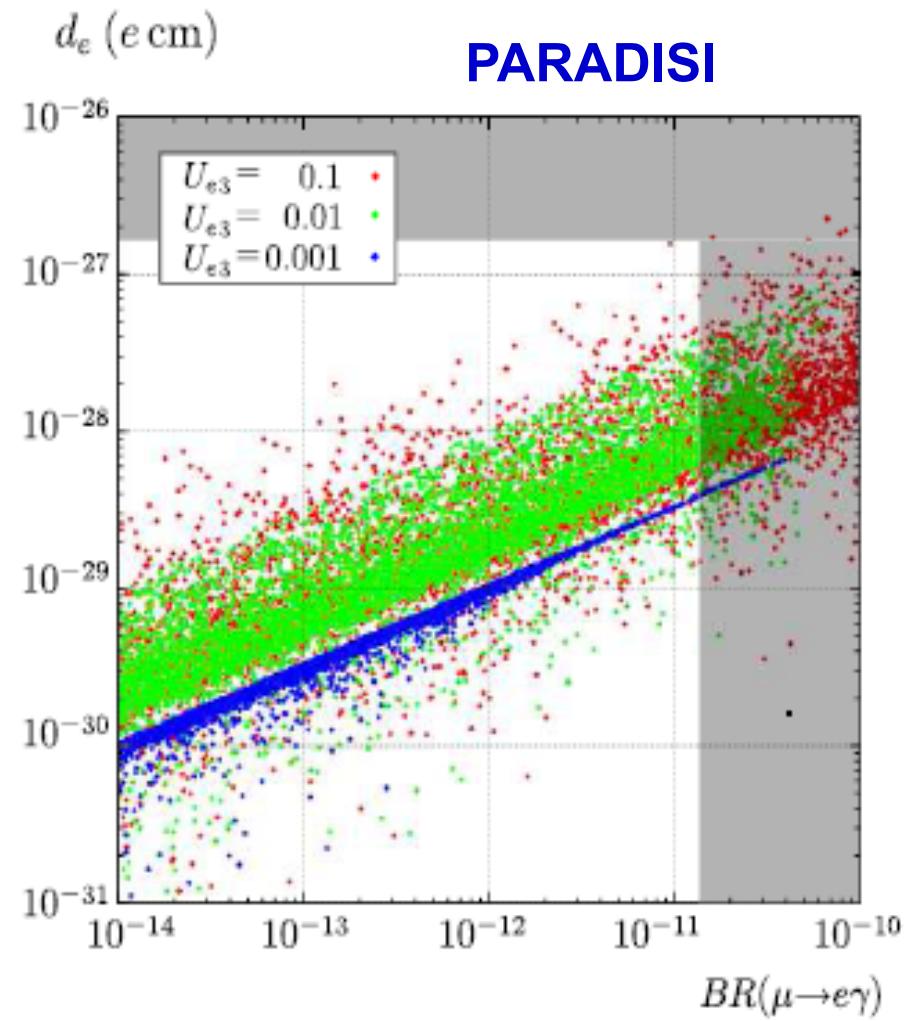
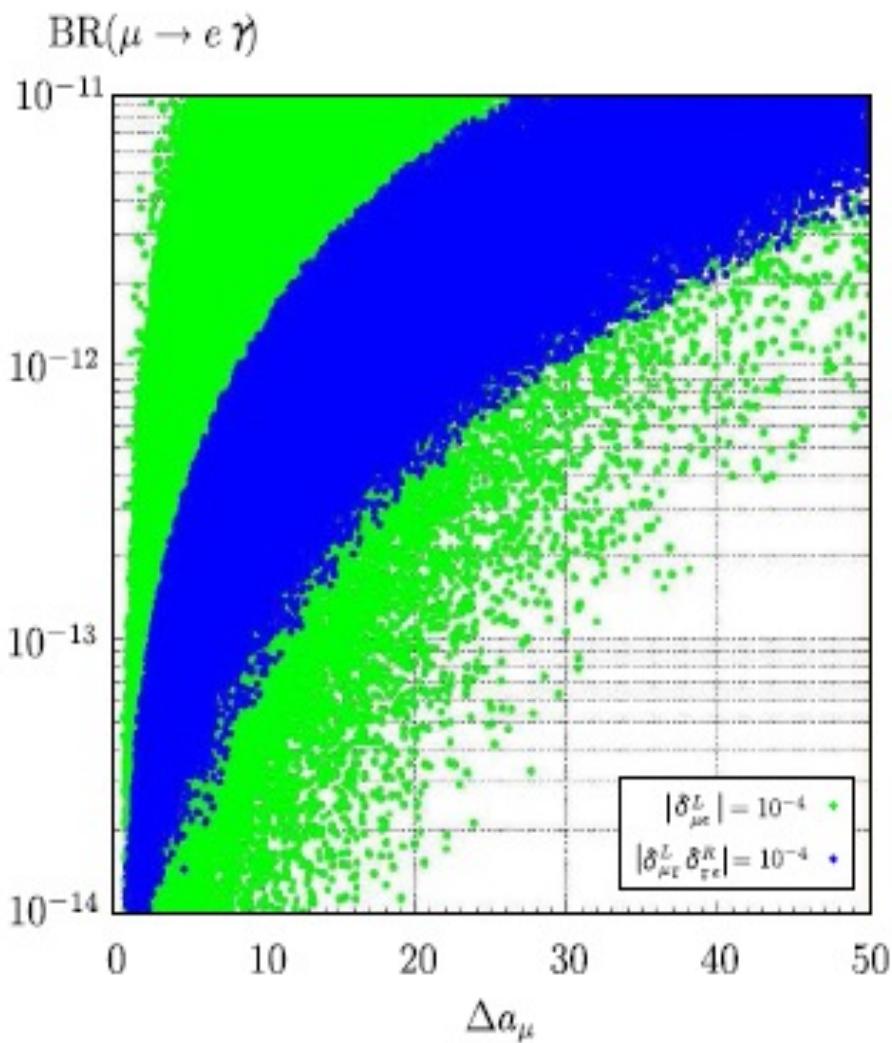
Theoretical-experimental **CHALLENGE**: can we reach the **1%** accuracy in the determination of the CKM unitarity triangle? → the “extreme flavor” path: a dream? Punzi et al

Hadronic Parameter	2002 [1115]	2013 [1116]
$f_+^{K\pi}(0)$	-	0.4%
B_K	17%	1.3%
f_{B_s}	13%	2%
f_{B_s}/f_B	6%	1.7%
B_{B_s}	9%	7%
B_{B_s}/B_B	3%	10%
$F_{D^*}(1)$	3%	2%
$B \rightarrow \pi$	20%	10%

C. Tarantino et al

Hadronic Parameter	What Next Era (2025)
$f_+^{K\pi}(0)$	0.1%
B_K	0.1 – 0.5%
f_{B_s}	0.5%
f_{B_s}/f_B	0.5%
B_{B_s}	0.5 – 1%
B_{B_s}/B_B	0.5 – 1%
$F_{D^*}(1)$	0.5%
$B \rightarrow \pi$	$\geq 1\%$

LFV, g – 2, EDM: a promising correlation in SUSY SEESAW



- A) Multimessenger astronomy,
- B) neutrino properties,
- C) dark side of the Universe and CMB

- A) **Photon, cosmic ray, neutrino , gravitational** astronomies (some in their maturity, some in their youth, some just baby or even still to be born)
- B) **neutrino mass** and its relation to the global symmetry of the SM, **Lepton number** (Dirac vs. Majorana nature of the neutrinos); measuring the full neutrino mass parameters (neutrino mass hierarchy, CP violation)
- C) **Dark Matter; Dark Energy** and **their role in the evolution of the Universe** (primordial inflation, elw. phase transition, quark-hadron phase transition, nucleosynthesis, matter-antimatter cosmic asymmetry)

What the combined BICEP2 +PLANCK analysis is telling us

- Only upper limit on $r \rightarrow$ **no hint for the presence of a large energy scale (related to primordial inflation)**
- Most **inflationary scalar potentials** still allowed
- Strong constraints on the **number of neutrino-like particles** present at the photon decoupling and on the **sum of the neutrino masses**

Info from Planck: Neutrino # and mass

$\sum m_\nu < 0.23 \text{ eV (95% CL)}$

$N_{\text{eff}} = 3.15 \pm 0.23$

Planck + Lyman alpha

$\sum m_\nu < 0.14 \text{ eV (C.L.)}$

Prospects for PLANCK + EUCLID

$\Delta m_\nu \approx 0.03 \text{ eV} \text{ & } \Delta N_\nu \approx 0.08$

Going beyond the SM: the NEUTRINO MASS

A. GIULIANI, SAC APPEC 2013

Cosmology, single and double β decay measure different combinations of the neutrino mass eigenvalues, constraining the **neutrino mass scale**

In a standard three active neutrino scenario:

$$\Sigma \equiv \sum_{i=1}^3 M_i$$

cosmology
simple sum
pure kinematical effect

$$\langle M_\beta \rangle \equiv \left(\sum_{i=1}^3 M_i^2 |U_{ei}|^2 \right)^{1/2}$$

β decay
incoherent sum
real neutrino

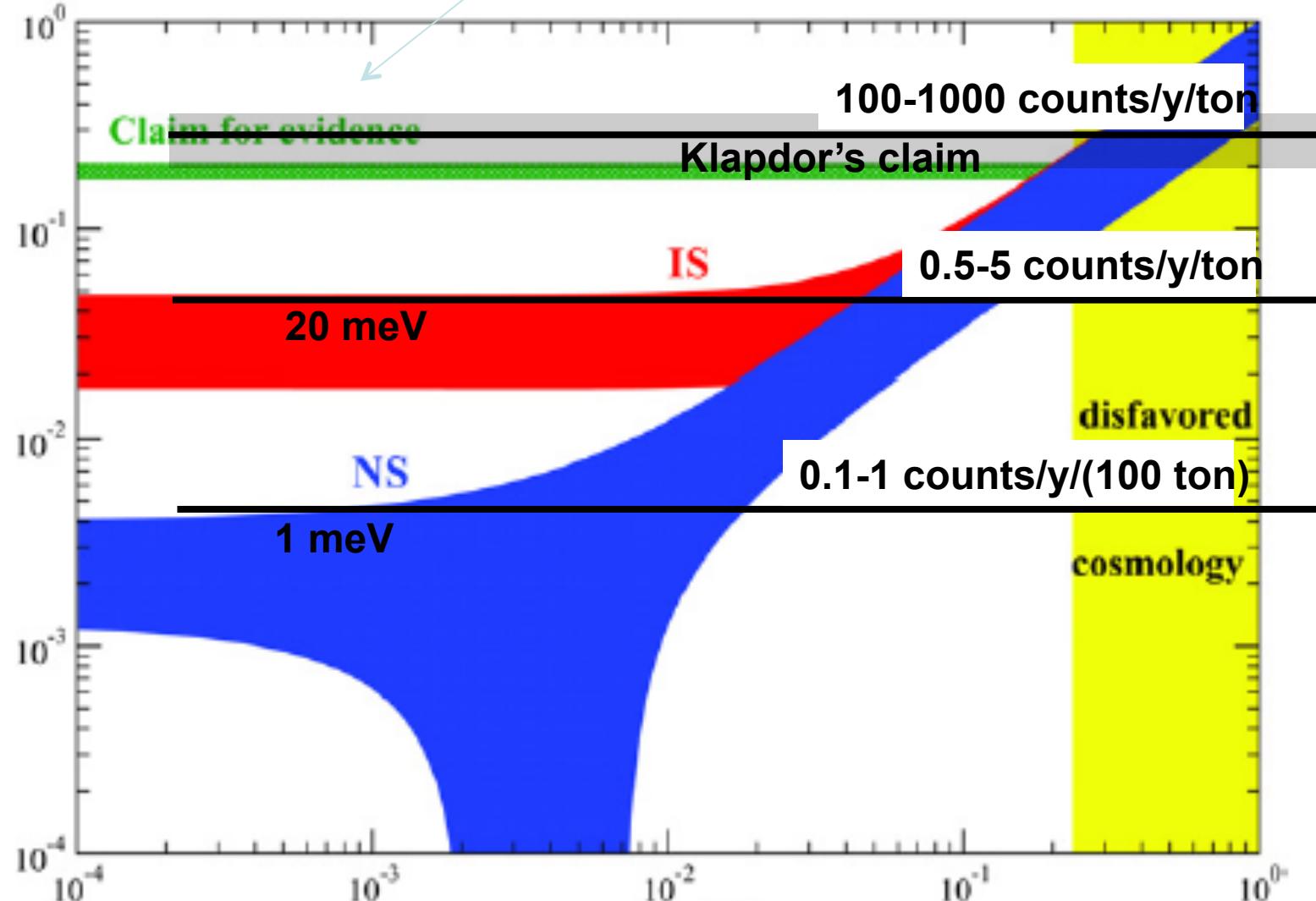
$$\langle M_{\beta\beta} \rangle \equiv \left| \sum_{i=1}^3 M_i |U_{ei}|^2 e^{i\alpha_i} \right|$$

double β decay
coherent sum
virtual neutrino
Majorana phases

Three challenges for 0ν -DBD search

$\langle M_{\beta\beta} \rangle$ [eV]

Klapdor Krivosheina Modern Physics Letters A 21, No. 20 (2006) 1547



CONNECTION DM – ELW. SCALE

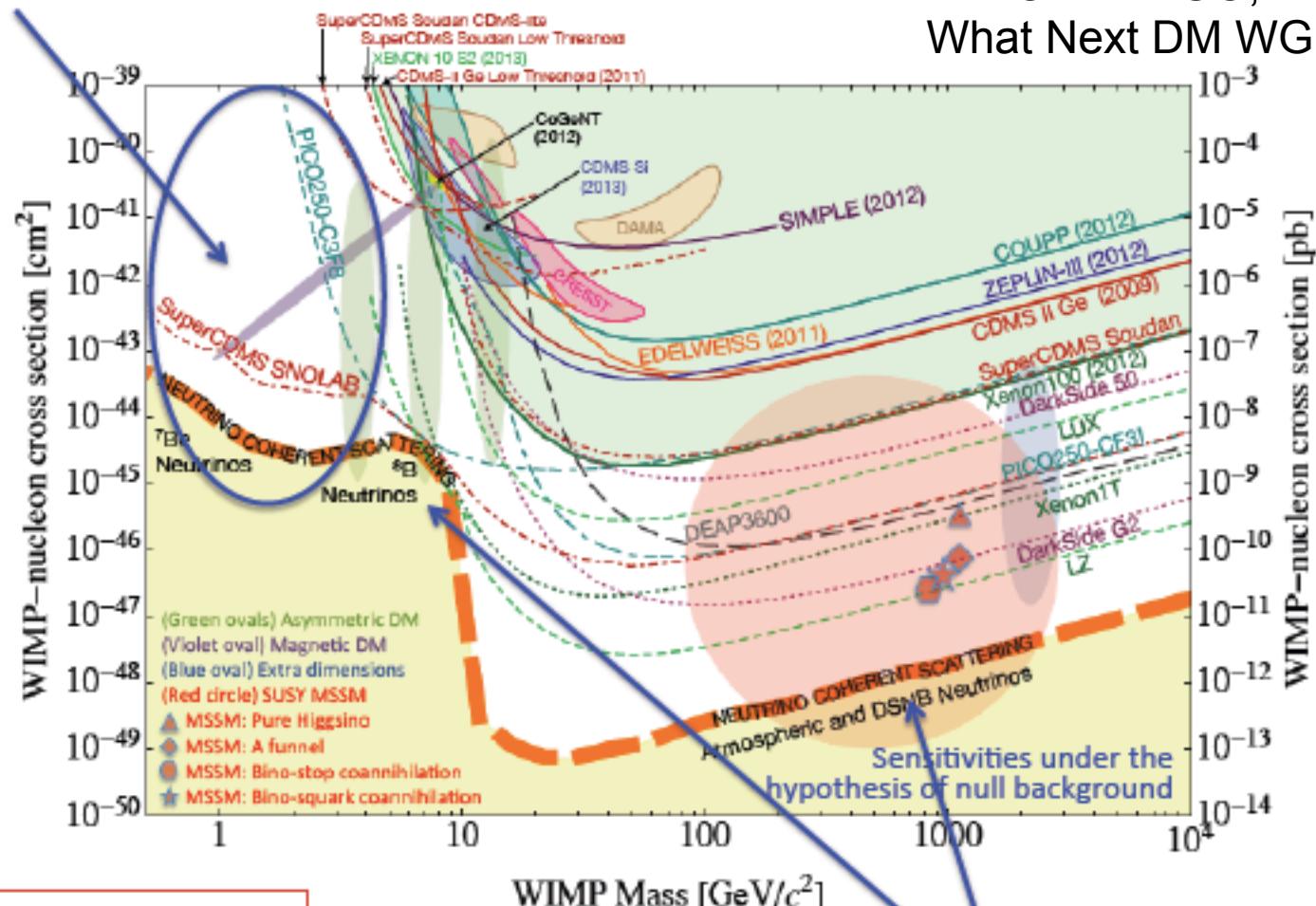
THE WIMP MIRACLE: STABLE ELW. SCALE WIMPs

1) ENLARGEMENT OF THE SM	SUSY (x^μ, θ)	EXTRA DIM. (x^μ, j^i)	LITTLE HIGGS. SM part + new part
	Anticomm. Coord.	New bosonic Coord.	to cancel Λ^2 at 1-Loop
2) SELECTION RULE	R-PARITY LSP	KK-PARITY LKP	T-PARITY LTP
→ DISCRETE SYMM.	Neutralino spin 1/2	spin1	spin0
→ STABLE NEW PART.	↓ m_{LSP} ~100 - 200 GeV	↓ m_{LKP} ~600 - 800 GeV	↓ m_{LTP} ~400 - 800 GeV
3) FIND REGION (S) PARAM. SPACE WHERE THE “L” NEW PART. IS NEUTRAL + $\Omega_L h^2$ OK			

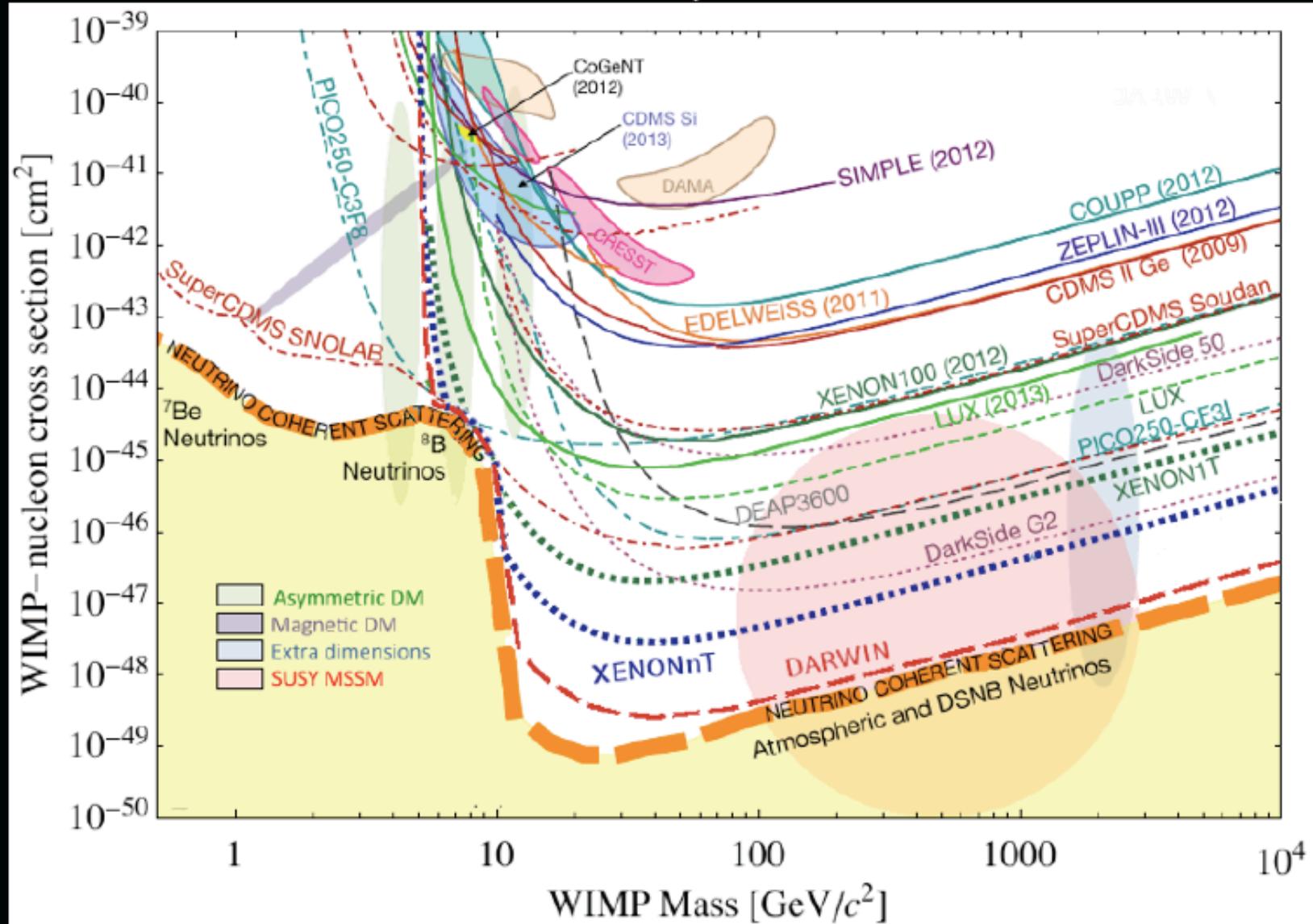
Direct detection

Light WIMPs window

N. FORNENGO,
What Next DM WG 2015

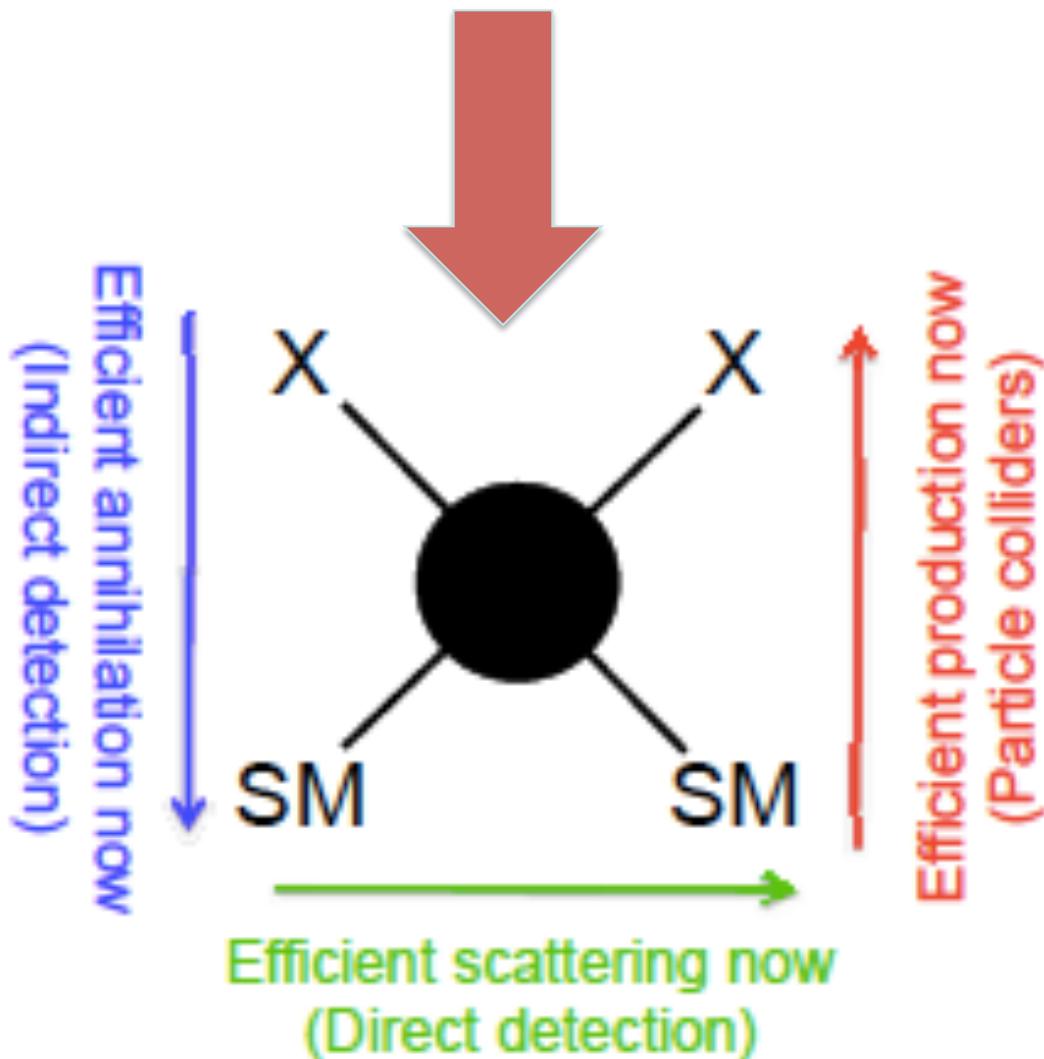


1) Science Goals: Dark Matter Projected Sensitivities



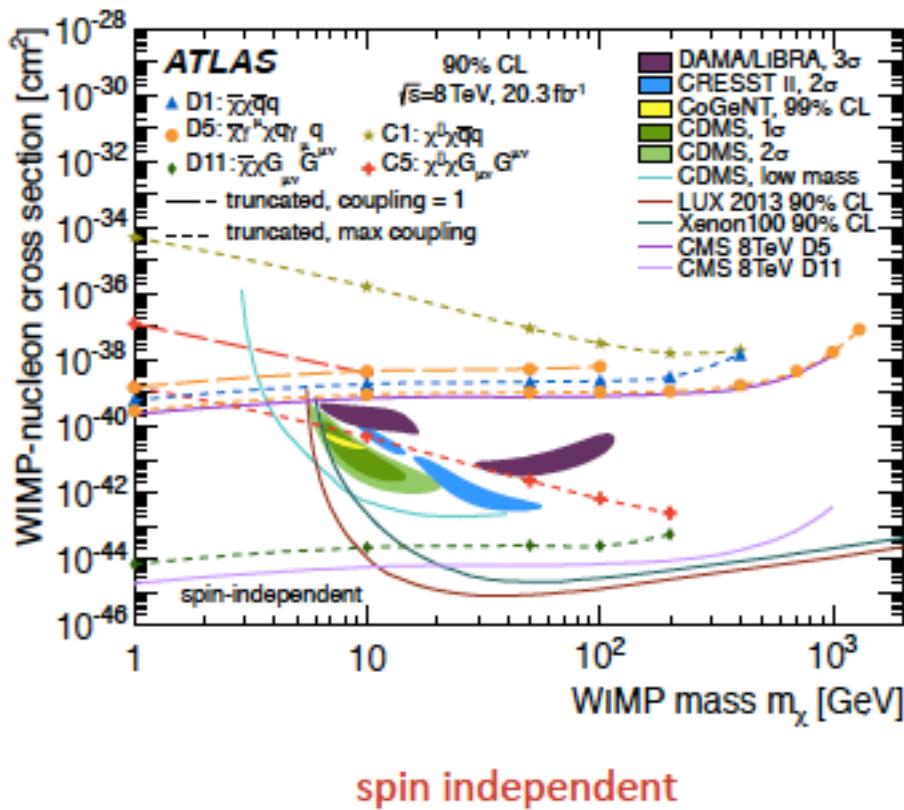
What if 2+ of these experiments observe strong candidate dark matter signals?
Build a directional detector to establish astrophysical origin.

DM COMPLEMENTARITY: efficient annihilation in the early Universe implies today

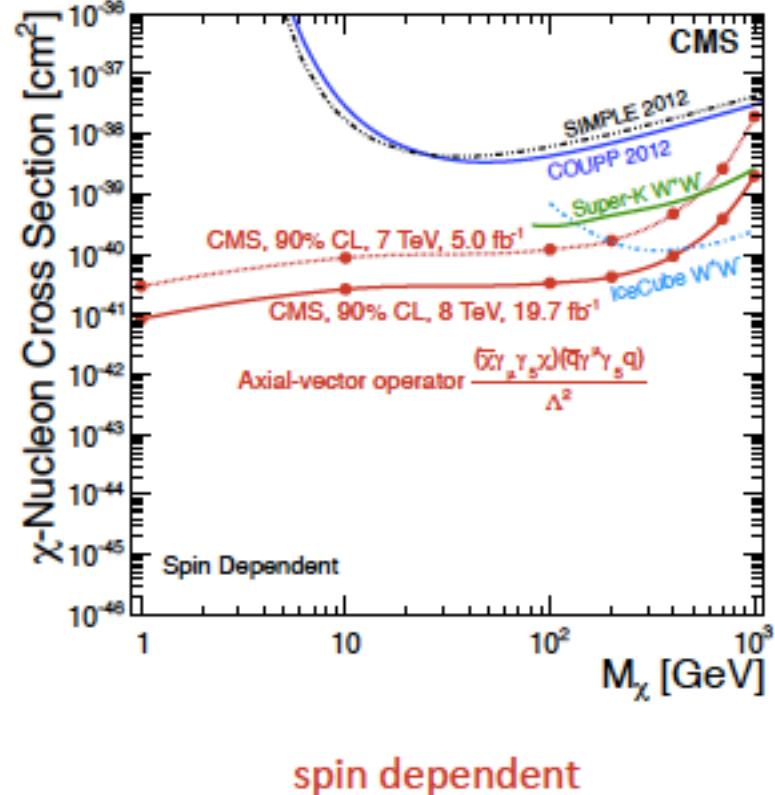


Mono-X searches

N. Fornengo, What Next DM WG 2015

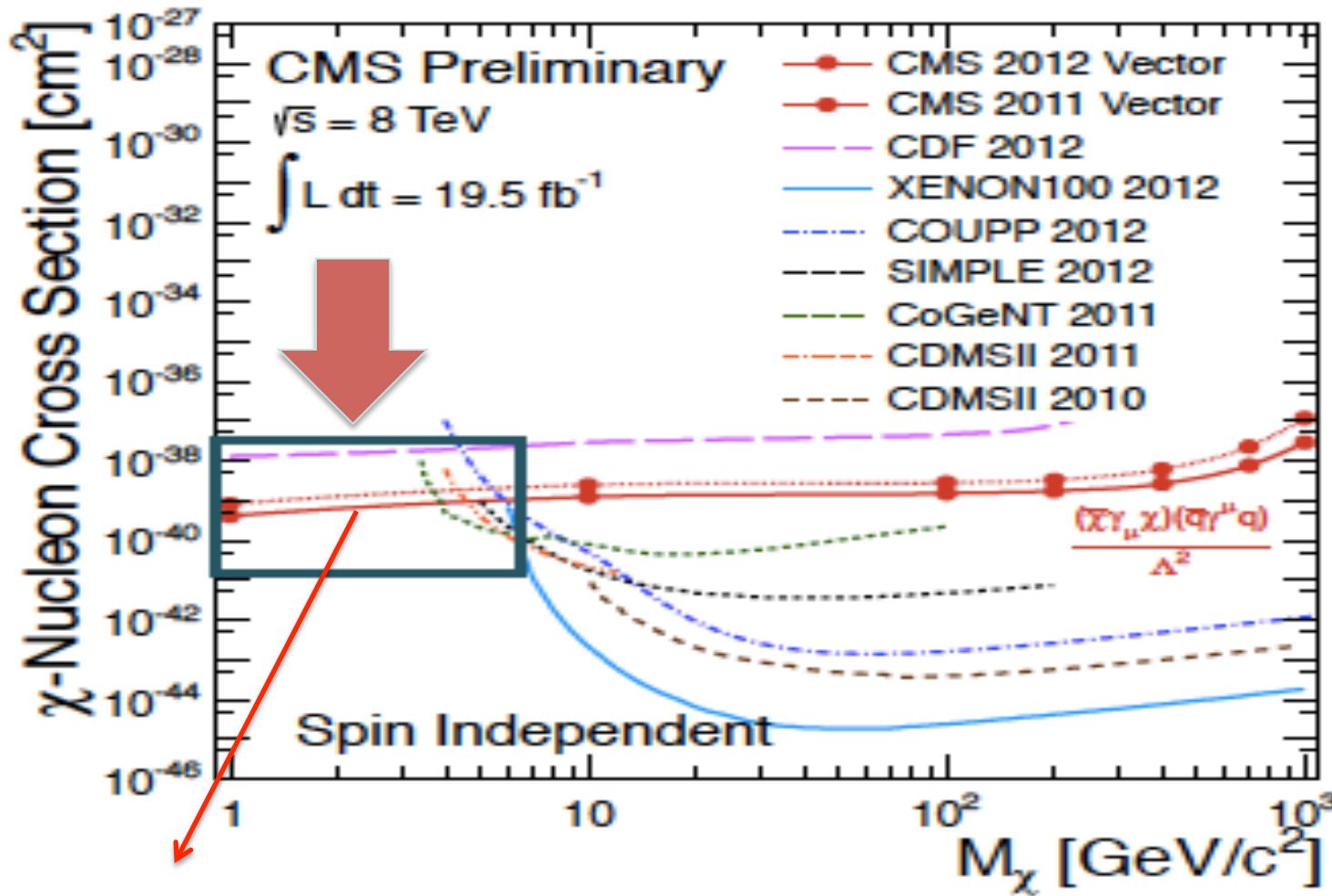


spin independent

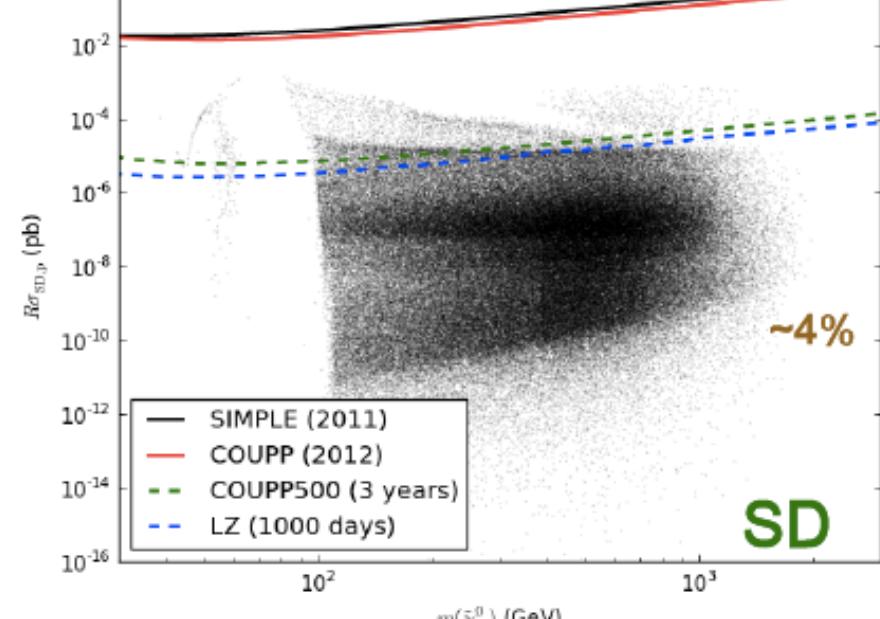
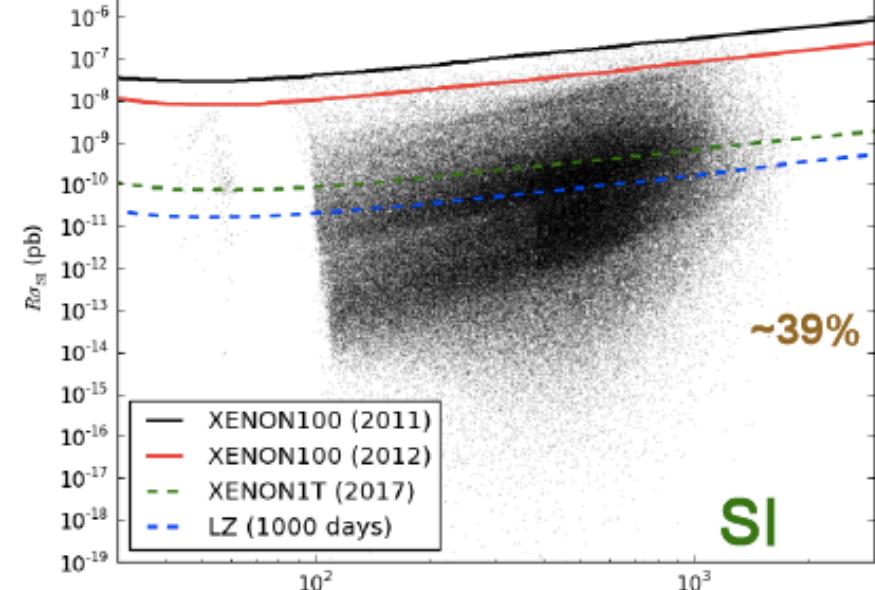


spin dependent

Inferred bounds on the direct detection scattering cross section
(subject to the validity range of EFT)

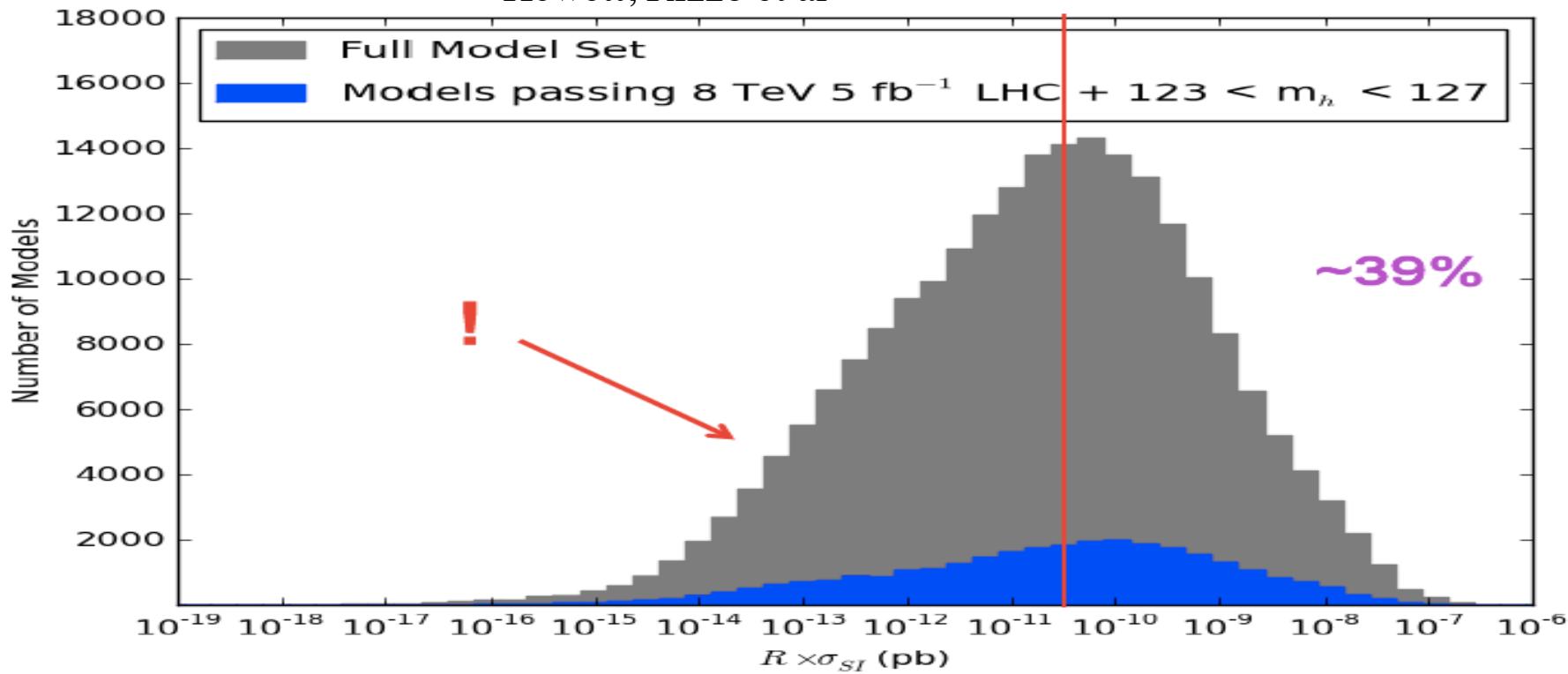


Relevant to intensify the efforts here: ex.
asymmetric DM with **DM particles of**
mass~ baryon mass given that ρ_{DM} not
much different from ρ_B

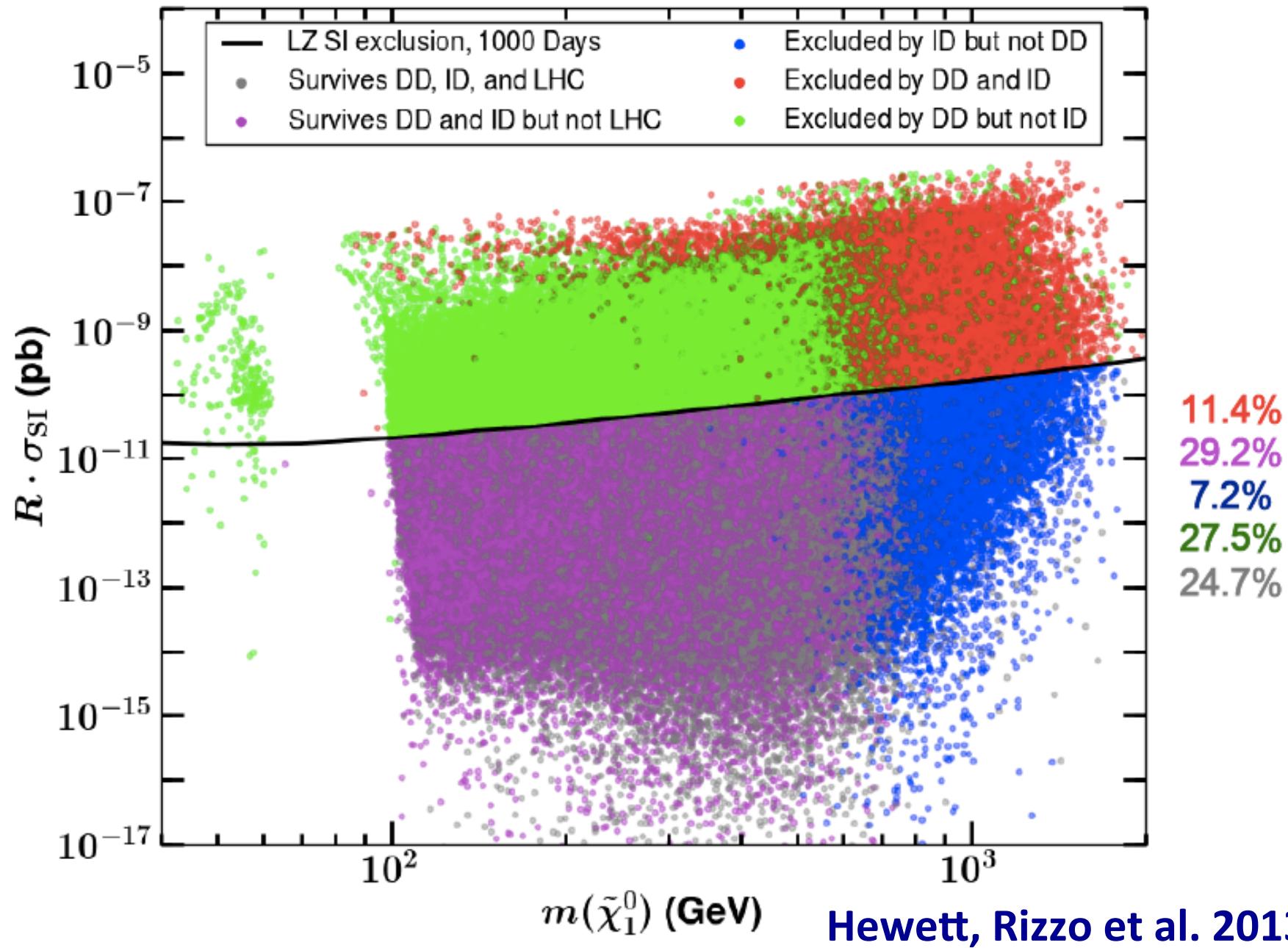


125k models pMSSM under scrutiny

Hewett, Rizzo et al



pMSSM models DD = LZ both SI + SD ID = FERMI + CTA



Challenges for next DM, $\beta\beta$ frontiers; Challenges for LNGS

- Attack and cover the IH region → 1-ton neutrinoless $\beta\beta$
- WIMPS DM : Reach the neutrino background → n-ton exps. $n > 5$

LNGS → largest ultra low-background facility ...
(see Terranova's talk)

LNGS at the next frontier: national lab or ERIC?

- To cope with the impressive future challenges
→ need for a truly international lab
- Possibility: European Research Infrastructure Consortium (ERIC)
- Possible partners: Netherland, Germany, Spain, Poland, ... (UK?, Sweden?; later on, also France should join)

Overall, in the next three years the APPEC agencies will need to take a decision on

- a) the construction of the phase 1.5 of KM3Net,
- b) a major investment as a contribution to a neutrino long baseline program in US or Japan,
- c) a European-led dark matter multi-ton experiment
- d) a ton-scale neutrino mass detector (double beta decay technique)
- e) a major contribution on ground and/or space to the cosmology program probing the param. of inflation.

The SAC recommends that APPEC prepares these decisions, in **coordination with its global partners, as well as with nearby fields**, e.g. particle physics national and international laboratories (CERN, FNAL, KEK and JPARC).

3 crucial ingredients for the success of the recipe:

- **Intense theoretical activity** with large integration of it with the experimental activities (PACT ...)
- **INNOVATIVE R&D**
- **Modern Computing**

The importance of being **SMALL**

My recommendation: beware the temptation of going ONLY for LARGE enterprises

The protective shield of large, Big Science: too big to fail!

Richness of small, “unorthodox” projects based more on clever ideas than on muscular, managerial strength!

Big Bang

Quark-Gluon

Protoni e
neutroni

Protoni e
Nuclei leggeri

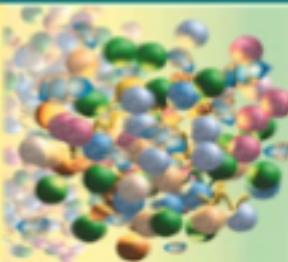
Atomi

Gravità

Nucleare forte

Nucleare debole

→Galassie
→Molecole→DNA



10^{-43} sec

10^{-32} sec

10^{-10} sec

10^{-35} m

10^{-32} m

10^{-18} m

10^{19} GeV

10^{16} GeV

10^2 GeV

10^{-4} sec

10^{-16} m

1 GeV

100 sec

10^{-15} m

1 Mev

300KY → 15GY

10^{-10} m

10 eV

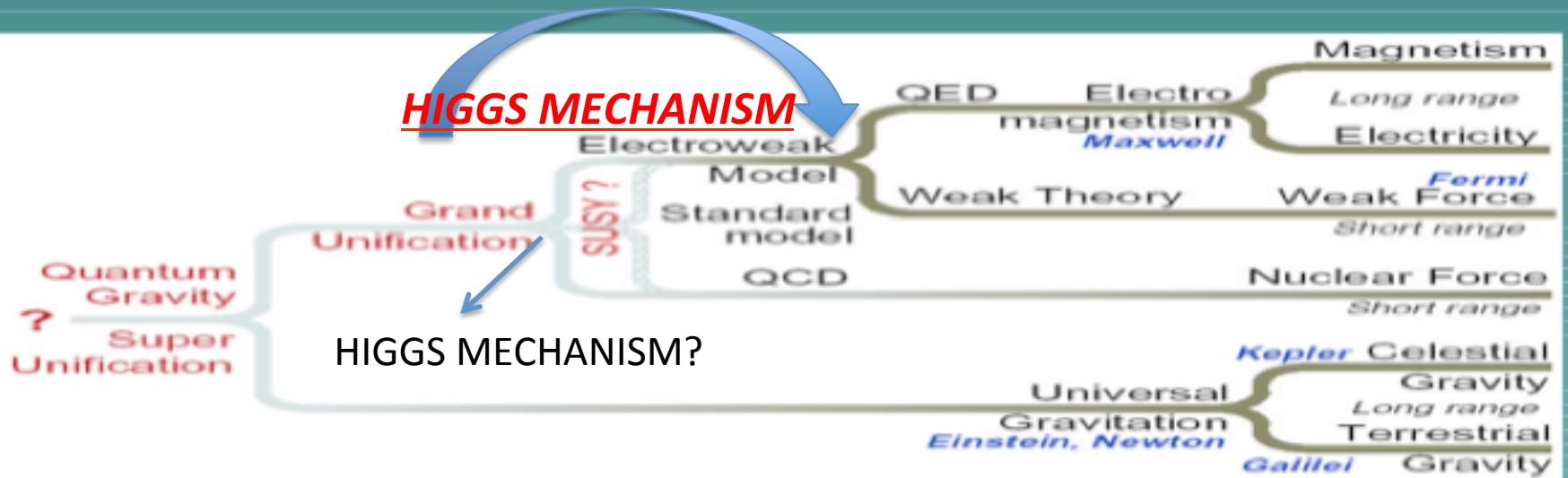
???

LHC

LEP

As tronomia→

HIGGS MECHANISM



Theories:

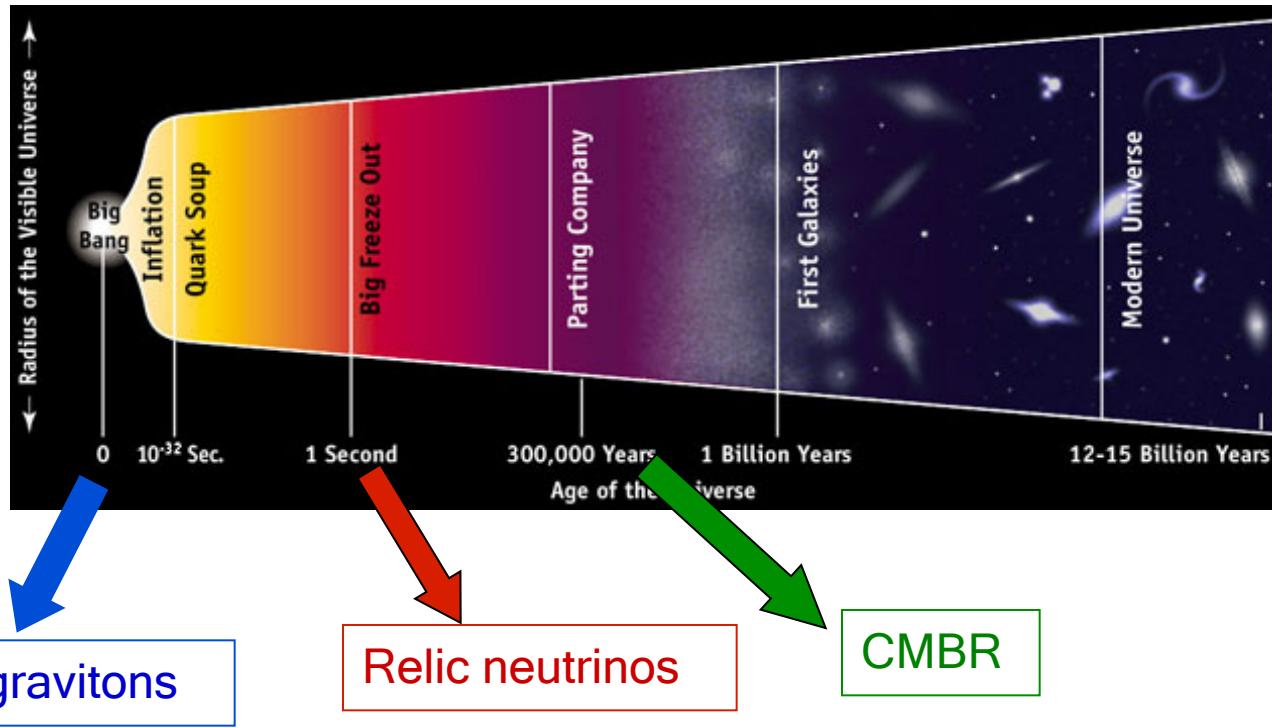
STRINGS?

RELATIVISTIC/QUANTUM

CLASSICAL

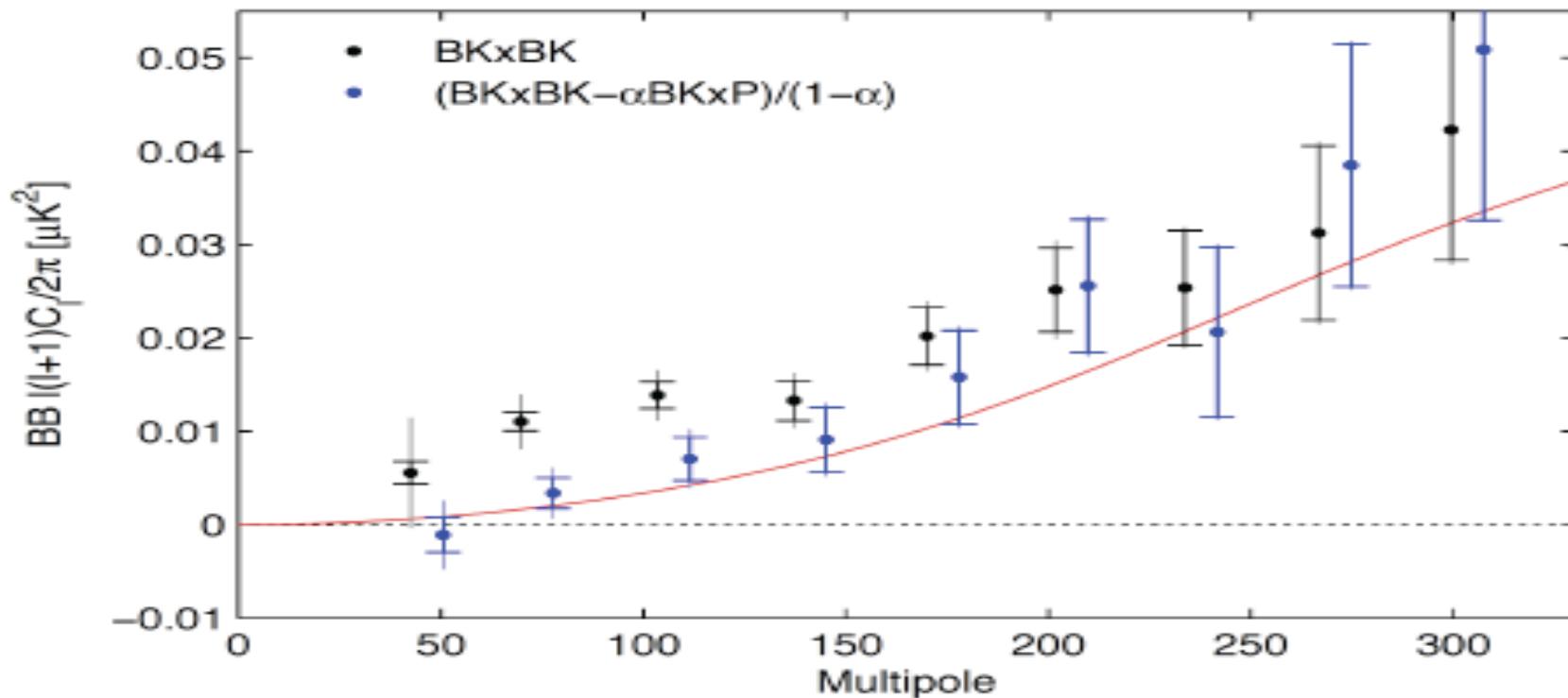
BACK-UP SLIDES

Relic Stochastic Background



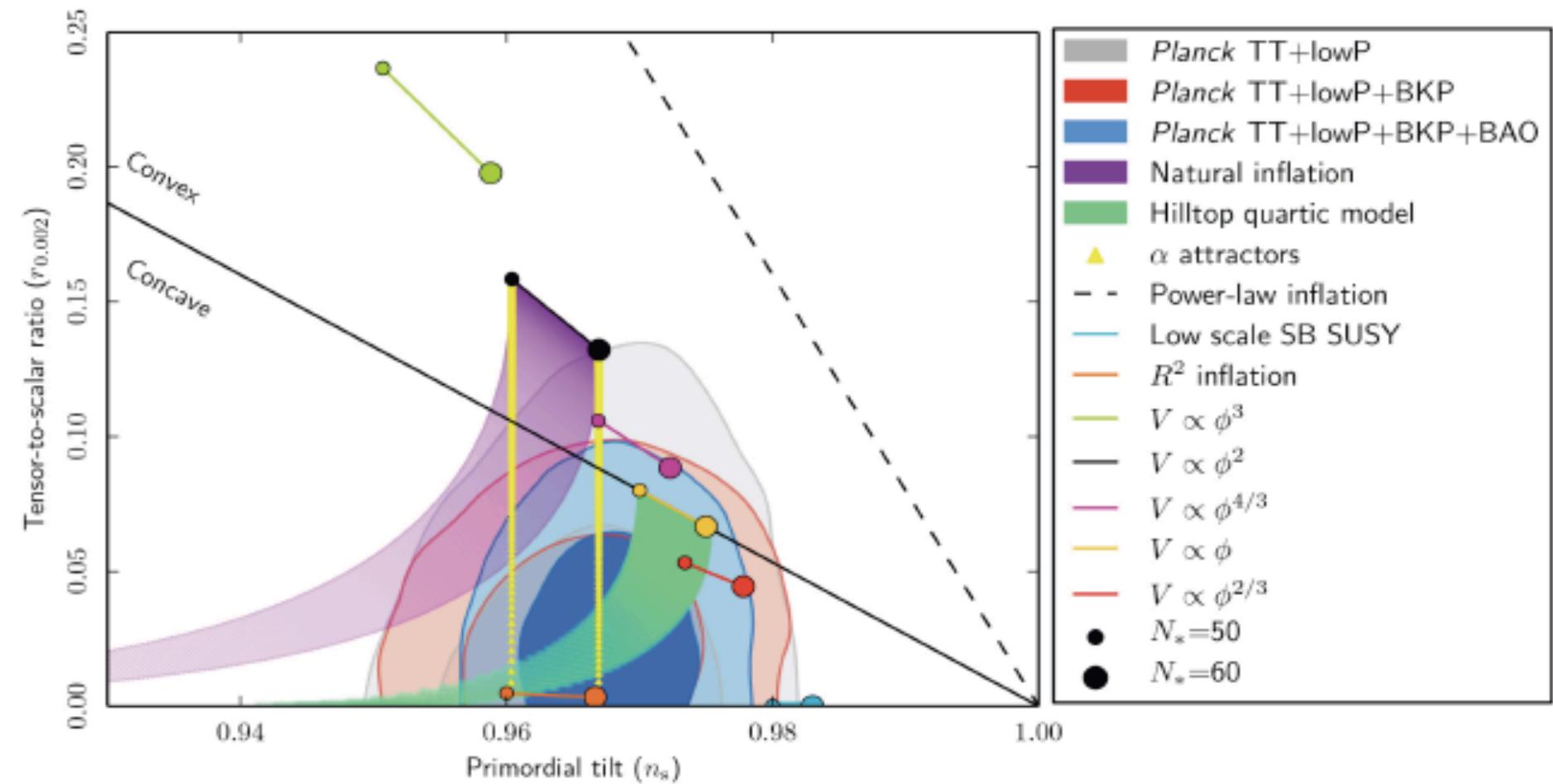
- Imprinting of the early expansion of the universe
- Correlation of at least two detectors needed

Subtracting PLANCK “dust” from BICEP2 results



r<0.12 @95% CL

THE INFLATON SCALAR POTENTIAL



$$V^{1/4} = 1.94 \times 10^{16} \left(\frac{r}{0.12} \right)^{1/4} \text{ GeV}$$

Is gauge hierarchy a “fictitious” problem?

- * Notice: gauge hierarchy is not a problem of the SM in itself; indeed, it arises when the **SM particles come in touch with much more massive particles** probably related to a new, large energy scale in physics
- But, how about having the SM as the final theory (“forever”) of all the interactions, but for gravity whose quantization has its own story independently from the SM of electroweak+ strong interactions? – **softened gravity** (Giudice, Isidori, Salvio, Strumia 2014)

2012: the conquest of a new energy scale in physics

- ~1900 ATOMIC SCALE 10^{-8} cm. $1/(\alpha m_e)$
- ~1970 STRONG SCALE 10^{-13} cm. $M_e^{-2\pi/\alpha_s^b}$
- ~2010 WEAK SCALE 10^{-17} cm. TeV^{-1}

FUNDAMENTAL OR DERIVED SCALE?

EX. EXTRA-DIMENSIONS
or
TeV STRING THEORY

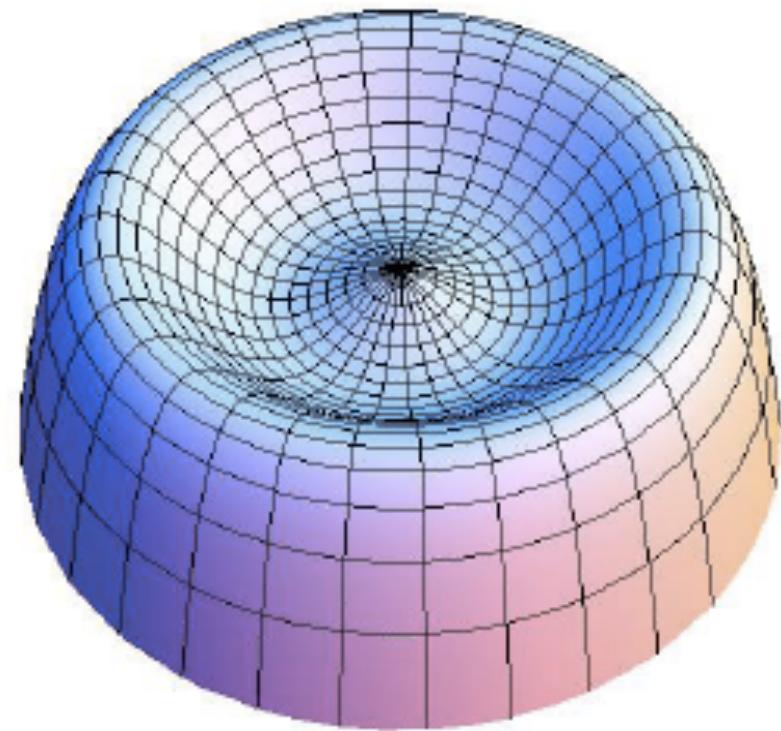
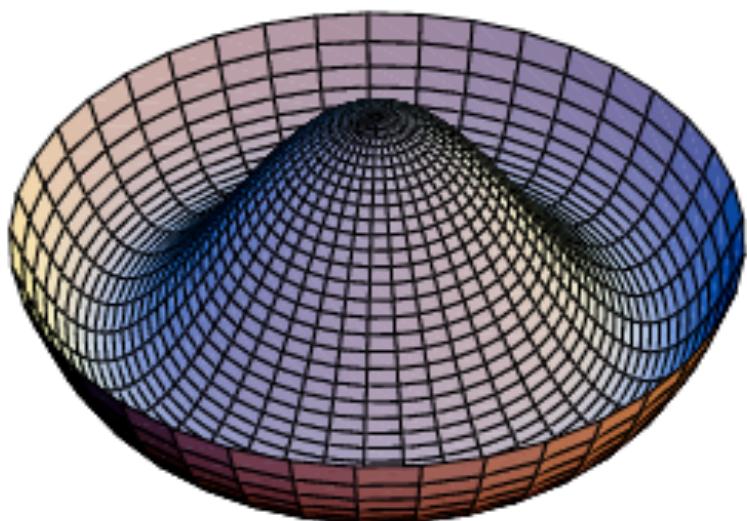
EX.: TECHNICOLOR or
SUSY with ELW RAD. BREAKING

NEW PARTICLES AT THE TEV SCALE?

STABILITY



INSTABILITY



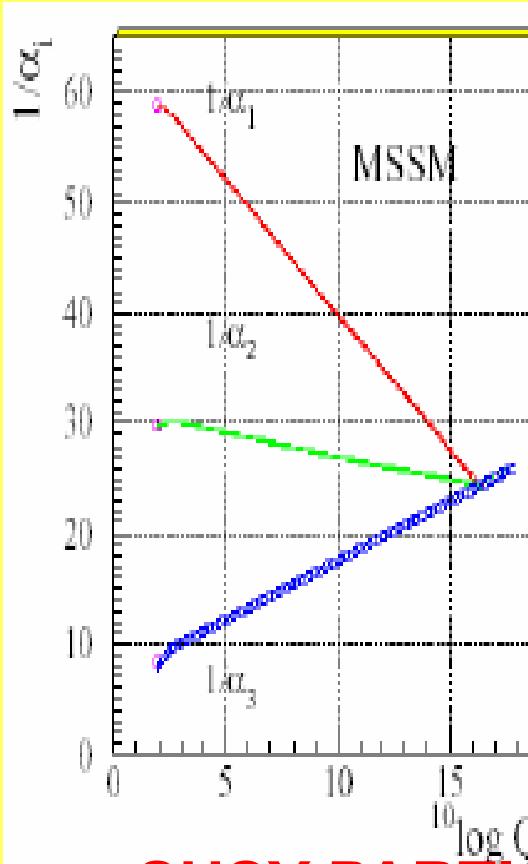
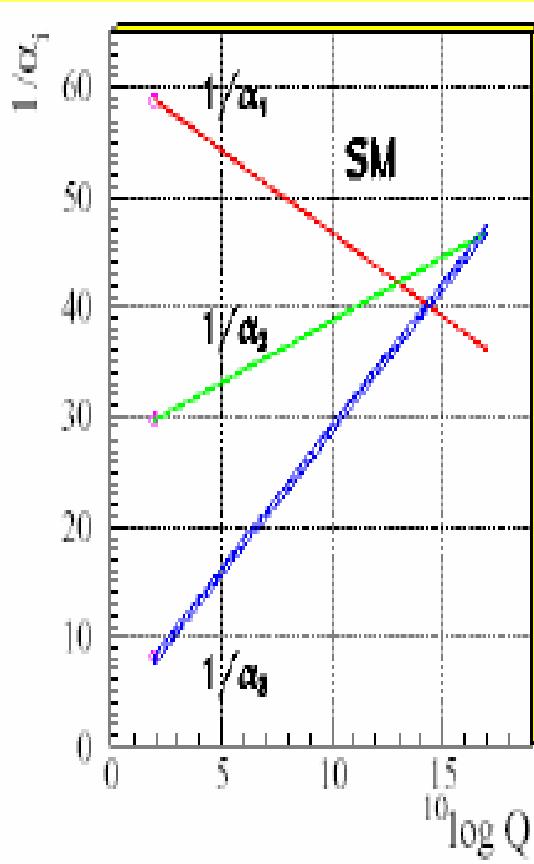
**ON THE IMPORTANCE OF PRECISELY
MEASURING HIGGS and TOP MASSES**

The harvest of the recently opened high energy gamma ray astronomy and the opening of the new astronomies: gravitational waves, neutrinos and the high energy cosmic rays

The understanding of the neutrino sector and its cosmological role

Large theoretical and experimental progress in the dark matter quest, reaching close to the parameter limits of current theories; the precise study of the parameters of the equation of state of dark energy and, eventually, of those of the inflation potential.

LOW-ENERGY SUSY AND UNIFICATION



Input

$$\alpha^{-1}(M_Z) = 128.978 \pm 0.027$$

$$\sin^2 \theta_{\overline{\text{MS}}} = 0.23146 \pm 0.00017$$

$$\alpha_s(M_Z) = 0.1184 \pm 0.0031$$

Output

$$M_{\text{SUSY}} = 10^{3.4 \pm 0.9 \pm 0.4} \text{ GeV}$$

$$M_{\text{GUT}} = 10^{15.8 \pm 0.3 \pm 0.1} \text{ GeV}$$

$$\alpha_{\text{GUT}}^{-1} = 26.3 \pm 1.9 \pm 1.0$$

**SUSY PARTICLES AT
THE TEV SCALE !**

The BIG and the SMALL- $\dim[m] \neq 0$

- $V = \mu^2 |H|^2 + \lambda |H|^4$ what is the value of the energy of its vacuum, i.e. the SM **vacuum energy**?
 $\rightarrow V_0 = \mu^2 \langle H \rangle^2 + \lambda \langle H \rangle^4 \sim (100 \text{ GeV})^2$
observed vacuum energy, i.e. dark energy
accelerating the expansion of the Universe $O(10^{-3} \text{ eV})$
- V defined up to a constant \rightarrow choose such constant to **cancel** the $O(100 \text{ GeV})^2$ contribution
- **10^{-3} eV 10^2 GeV 10^{16} GeV 10^{19} GeV**
- **Why** so different mass scales ?
- **How** to guarantee their separation \rightarrow symmetry vs. multiverse

The BIG and the SMALL – dim[m]=0

- $h_t - h_e$ **flavour** issue
- L_{SM} no symmetry prevents to add a term violating **CP in the strong interactions** whose size depends on a **dimensionless** parameter $\theta \rightarrow$ the bound on the neutron EDM $\rightarrow \theta < 10^{-10}$
- **The θ – problem** : the symmetry solution

Axion from breaking of global chiral symmetry; axion field acts as dynamical theta para-meter,

$$\mathcal{L} \supset -\frac{\alpha_s}{8\pi} \underbrace{\frac{A}{f_A}}_{\bar{\theta}} G_{\mu\nu}^a \tilde{G}^{a,\mu\nu}$$

[Peccei,Quinn 77; Weinberg 78; Wilczek 78]

spontaneously relaxing to zero, $\langle A \rangle = 0$ (thus CP conserved)

- mass due to chiral symmetry breaking $m_A \sim m_\pi f_\pi / f_A$
- has universal coupling to photons, $\mathcal{L} \supset -\frac{\alpha}{8\pi} C_0 \frac{A}{f_A} F_{\mu\nu} \tilde{F}^{\mu\nu}$

Ringwald

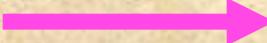
LOW-ENERGY SIGNATURES OF UNIFICATION AT 10^{16} GeV

- PROTON DECAY mediated by new particles (scalars or gauge bosons) related to the unified physics at 10^{16} GeV which DOES NOT respect the BARYON and LEPTON NUMBER SYMMETRIES → for a mediator of mass $\sim 10^{16}$ GeV we expect a proton lifetime in the ballpark of $\sim 10^{34}$ years → exp. accessible
- NEUTRON-ANTINEUTRON OSCILLATION if the unified symmetry (ex. SO(10)) breaks down to an intermediate symmetry subsequently spontaneously broken at $\sim 10^6$ GeV with the breaking of Baryon number of two units (ex. SO(10) → $SU(4)_{PS} \times SU(2)_L \times SU(2)_R \rightarrow SU(3) \times SU(2)_L \times U(1)_Y$) → exp. accessible (for instance , at the ESS)

MATTER-ANTIMATTER ASYMMETRY **NEUTRINO MASSES CONNECTION: BARYOGENESIS THROUGH LEPTOGENESIS**

- Key-ingredient of the SEE-SAW mechanism for neutrino masses: **large Majorana mass for RIGHT-HANDED neutrino**
- In the early Universe the heavy RH neutrino decays with Lepton Number violation; if these decays are accompanied by a new source of CP violation in the leptonic sector, then

VANILLA LEPTOGENESIS !

 it is possible to create a lepton-antilepton asymmetry at the moment RH neutrinos decay. Since SM interactions preserve Baryon and Lepton numbers at all orders in perturbation theory, but violate them at the quantum level, such **LEPTON ASYMMETRY** can be converted by these purely quantum effects into a **BARYON-ANTIBARYON ASYMMETRY** (**Fukugita-Yanagida mechanism for leptogenesis**)

LFV IN SUSY SEE-SAW

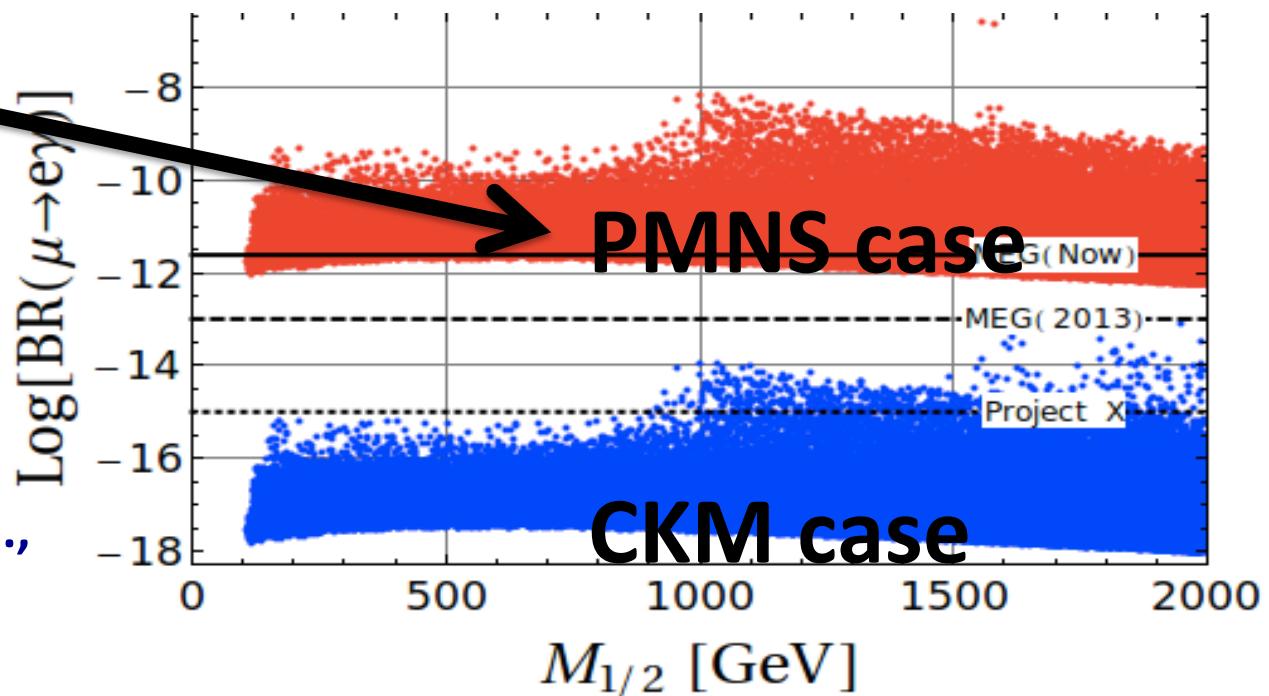
SEE- SAW (type 1) LOW-ENERGY SUSY

New source of
(leptonic) flavor:

YUKAWA COUPLINGS OF THE
NEUTRINO DIRAC MASS
CONTRIBUTIONS, i.e. **THE**
YUKAWAs of the
HIGGS couplings to
the **LETF-** and **RIGHT –**
HANDED NEUTRINOS

The scalar lepton
masses through their
running bring memory of
those new sources of
leptonic flavor at the TeV
scale, i.e. at energies
much below the
(Majorana) mass of the
RH neutrinos

PMNS case in
mSUGRA with
 $\tan\beta = 10$



Calibbi, Chowdhuri, A. M.,
Patel, Vempati 2012

