# Status and perspectives of fragmentation beams at LNS with CHIMERA detector



G.Cardella for the EXOCHIM collaboration



International
Workshop on
Multi facets of
Eos and
Clustering
Clustering

Multi facets of E Fos and E Fos and

6<sup>th</sup> – 9<sup>th</sup> May 2014 Catania, Italy

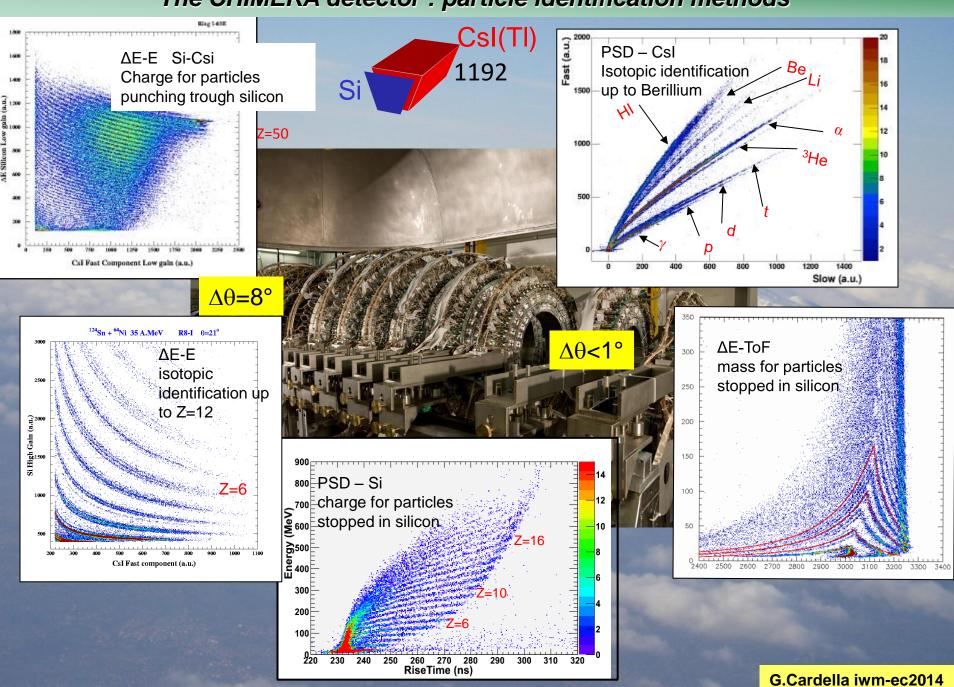




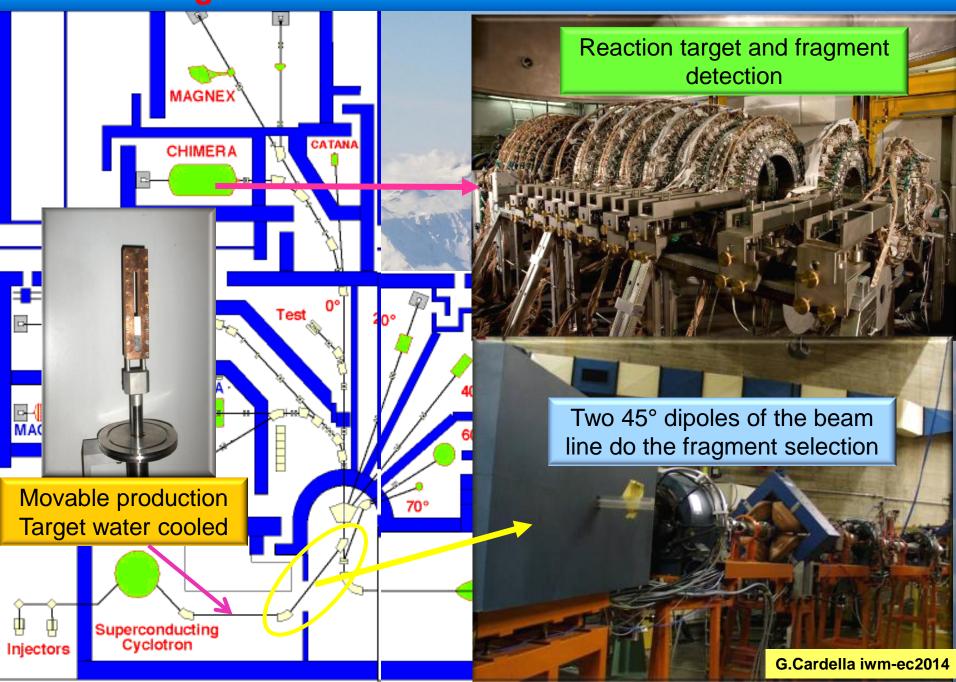
## **Outline**

- >CHIMERA detector and LNS fragmentation beams with selected examples:
  - **≻Elastic inelastic and Transfer reactions:** ¹⁰Be+p→⁰Be+d
    - **≻**<sup>8</sup>He induced reactions
      - ➤ Break-up: ¹6C case
  - $\triangleright \gamma$ -ray detection with CHIMERA & PIGMY resonance study
    - > Conclusions and perspectives

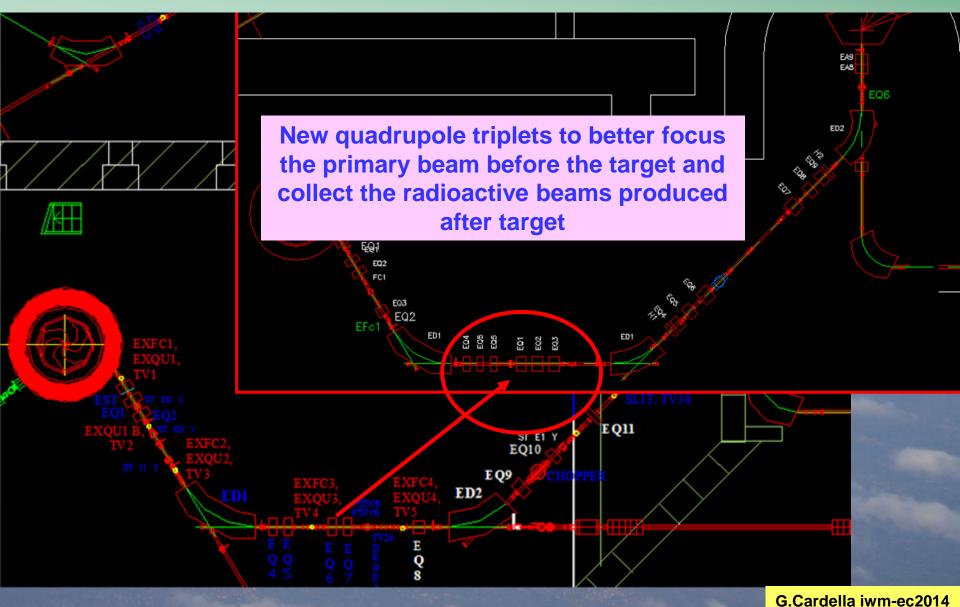
#### The CHIMERA detector : particle identification methods



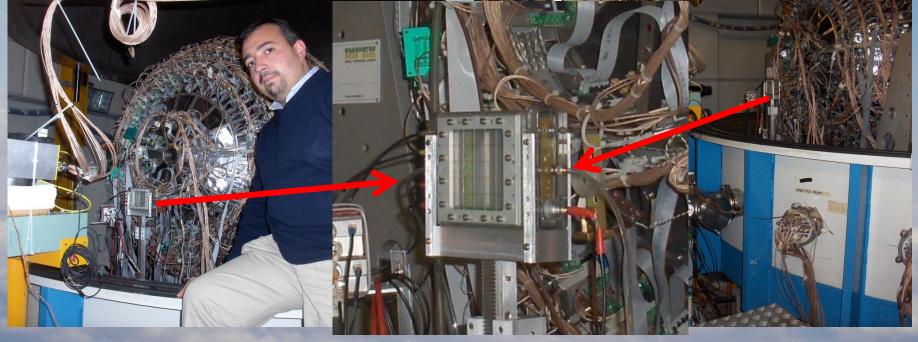
## Fragmentation beams at INFN-LNS - Catania

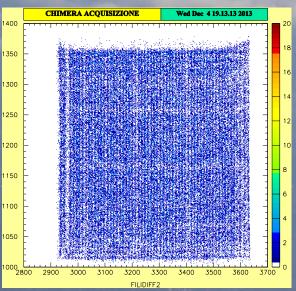


# The 2010 upgrading of LNS Fragmentation beam

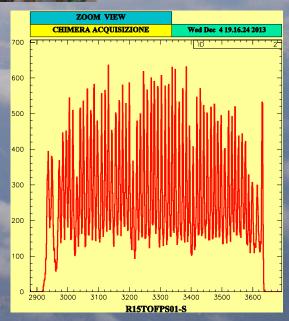


## **Upgrading of tagging system: test PPAC at LNS**





A PPAC was built to better measure beam trajectory



## Intensities available from the most recent beams produced

| the state of the s | The part of the part of the part |                         |
|--|----------------------------------|-------------------------|
| primary beam   | beam                             | intensity<br>(kHz/100W) |
| 18O 55MeV/A  | 16C                              | 120                     |
| setting 11Be   | 17C                              | 12                      |
|  | 13B                              | 80                      |
|  | 11Be                             | 20                      |
|  | 10Be                             | 60                      |
|  | 8Li                              | 20                      |
| 18O 55MeV/A  | 14B                              | 3                       |
| setting 12Be   | 12Be                             | 5                       |
|  | 9Li                              | 6                       |
|  | 6He                              | 12                      |
| 13C 55 MeV   | 11be                             | 50                      |
| setting 11Be   | 12B                              | 100                     |
| 36Ar 42 MeV  | 37K                              | 100                     |
| setting 34Ar   | 35Ar                             | 70                      |
|  | 36Ar                             | 100                     |
|  | 37Ar                             | 25                      |
|  | 33CI                             | 10                      |
|  | 34CI                             | 50                      |
|  | 35CI                             | 50                      |
| 20Ne 35 MeV  | 18Ne                             | 50                      |
| setting ne18   | 17F                              | 20                      |
|  | 21Na                             | 100                     |
| 70 <b>Z</b> n 42MeV  |                                  |                         |
| setting 68Ni   | 68Ni                             | 20                      |

## New beams to be used during 2014

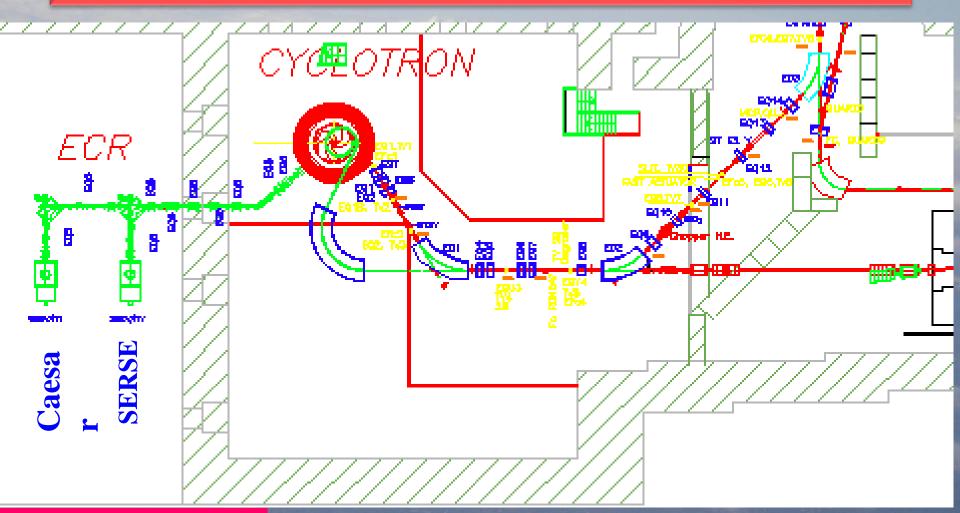
<sup>8</sup>He (CHIMERA)

<sup>14</sup>Be (test experiment) collaboration with Leuven

38S (Magnex)

#### **Future**

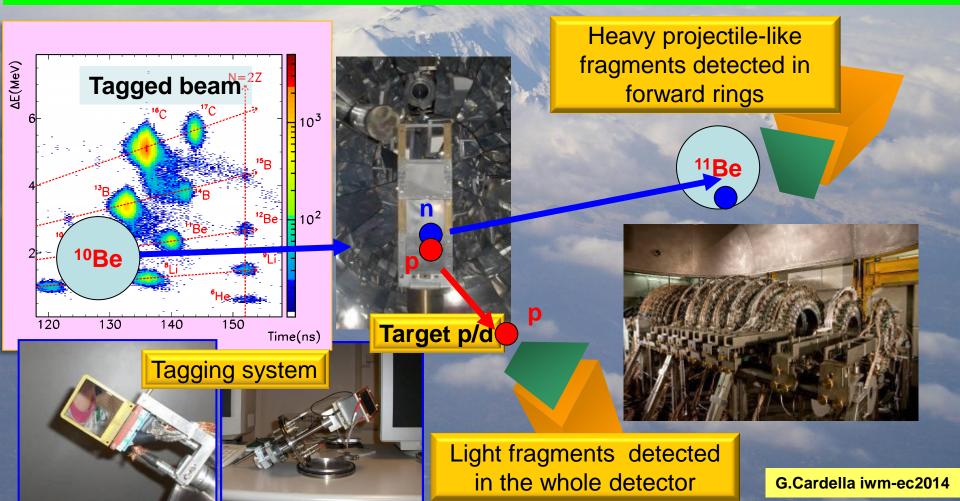
New cryostat and coils for the CS – possibility to extract beam trough stripping – the acceleration of beams with lower charge status will allow an increase of intensity from 20 to 100 – This will be a fantastic perspective for the fragmentation beam



## -Neutron transfer reactions near halo nuclei -

We study direct reactions using light exotic nuclei impinging on p, d targets useful to investigate on various structure effects

EVENT SELECTION performed with kinematic coincidences – measuring in binary/ternary reactions all reaction partners we clean the events



### – Advantages of binary kinematics: the ¹ºBe+p→ºBe+d case -

Nuclear Instruments and Methods in Physics Research A 715 (2013) 56-61



Contents lists available at SciVerse ScienceDirect

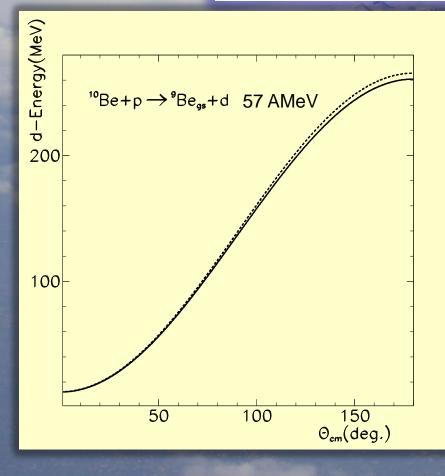
#### Nuclear Instruments and Methods in Physics Research A

journal homepage: www.elsevier.com/locate/nima



#### Kinematical coincidence method in transfer reactions

L. Acosta <sup>b</sup>, F. Amorini <sup>b</sup>, L. Auditore <sup>d</sup>, I. Berceanu <sup>h</sup>, G. Cardella <sup>a,\*</sup>, M.B. Chatterjiee <sup>i</sup>, E. De Filippo <sup>a</sup>, L. Francalanza <sup>b,c</sup>, R. Giani <sup>b,c</sup>, L. Grassi <sup>a,k</sup>, A. Grzeszczuk <sup>j</sup>, E. La Guidara <sup>a,g</sup>, G. Lanzalone <sup>b,e</sup>, I. Lombardo <sup>b,f</sup>, D. Loria <sup>d</sup>, T. Minniti <sup>d</sup>, E.V. Pagano <sup>b,c</sup>, M. Papa <sup>a</sup>, S. Pirrone <sup>a</sup>, G. Politi <sup>a,c</sup>, A. Pop <sup>h</sup>, F. Porto <sup>b,c</sup>, F. Rizzo <sup>b,c</sup>, E. Rosato <sup>f</sup>, P. Russotto <sup>b,c</sup>, S. Santoro <sup>d</sup>, A. Trifirò <sup>d</sup>, M. Trimarchi <sup>d</sup>, G. Verde <sup>a</sup>, M. Vigilante <sup>f</sup>



The lab energy of the detected particle determines the CM emission angle

Due to the relatively good energy resolution we can obtain an angular distribution with much better resolution than the one determined by the size of the detectors

### STEPS of the analysis ¹⁰Be+p→⁰Be+d

We select only complete events with two detected particles and with total detected charge Ztot=Zbeam+1

We can plot the  $\Delta \phi$  angle between the two coincidence detectors - due to momentum conservation Δφ must be 180°

 $\Delta \phi$  width due to the finite opening of the detectors

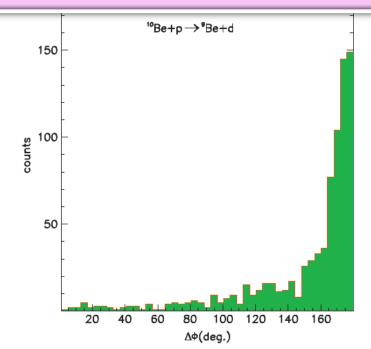


Fig. 5. Relative angle  $\Delta \phi$  between the telescopes selected in coincidence in the reaction <sup>10</sup>Be+p→ <sup>9</sup>Be+d. The peak at 180° is due to kinematical coincidences.

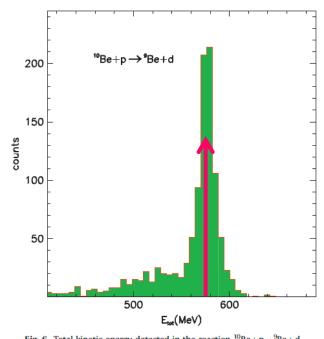


Fig. 6. Total kinetic energy detected in the reaction <sup>10</sup>Be+p→ <sup>9</sup>Be+d.

We also clean the events putting constraints on the total detected energy must be equal to the beam energy 580  $MeV + Q_{value} (-4.58 MeV)$ 

Notwithstanding the scarce total energy resolution we see only GS events

$$^{9}$$
Be\* $\rightarrow$  n+α+α S<sub>n</sub>=1.66 MeV

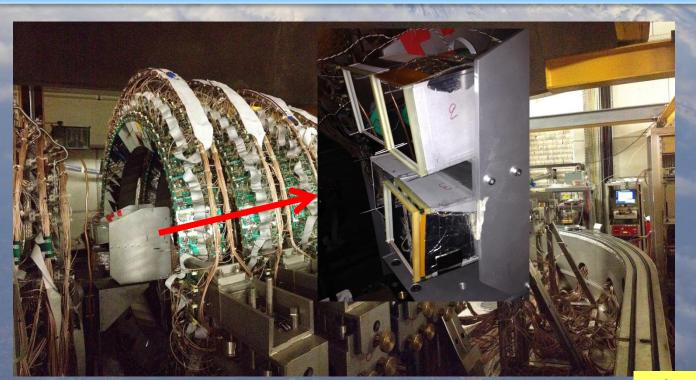
- The ¹ºBe+p→9Be<sub>q,s,</sub>+ d angular distribution -From the analysis we get the deuteron energy spectrum Using kinematics we can convert it from  $dN/dE \rightarrow dN/d\theta$  $^{10}Be+p \rightarrow ^{9}Be_{m}+d$ 20 200 dσ/dΩ(mb/sr) counts **DSSSD-tag** DSSSD-traj 100  $^{10}$ Be+p $\rightarrow$  $^{9}$ Be+d (58AMeV) 10  $^{10}$ Be+d $\rightarrow$  $^{9}$ Be+t  $\theta_{\text{beam}}$ 50 100 150 50 d-Energy(MeV) 0.8 0.6 Fig. 7. Deuteron energy spectrum from the reaction  ${}^{10}\text{Be} + \text{p} \rightarrow {}^{9}\text{Be}_{gs} + \text{d}$ . 0.4 0.2 10 After solid angle and efficiency correction we get the angular strip 20 60 80 40 distributions  $\Theta_{cm}(\text{deg.})$ 

Note that angular distributions are automatically corrected for the fragmentation beam angular spread

## Improvement of the detection system

#### **FARCOS 0°**

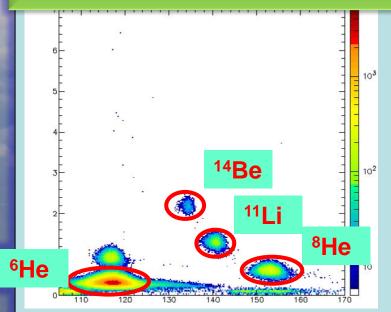
For next experiments we will use FARCOS (see Emanuele talk) in configuration around 0° in order to have a kind of spectrometer to measure the quasi-projectile – light particles will be detected with CHIMERA using kinematical coincidence and beam trajectory measurement we will clean from background and extract more accurate excitation energies and CM angular distributions



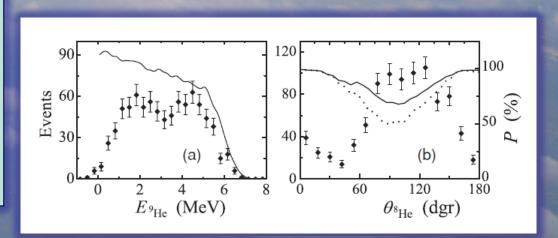
## Coming Experiments: Neutron transfer on <sup>8</sup>He

One programmed experiment is the <sup>8</sup>He+d reaction performed to study the <sup>9</sup>He resonance - production will by done using a <sup>11</sup>B primary beam – We expect around 2000 particles/sec of beam intensity

Preliminary production test performed with <sup>18</sup>O primary beam



Reaction studied at Dubna at 25 AMeV
We will increase the beam energy
searching for higher excitation energy
structures



M. S. Golovkov et al PHYSICAL REVIEW C 76, 021605(R) (2007)

## Coming Experiments: Break-up study - CLIR

#### (Clusters in Light Ion Reactions)

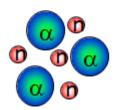
Break-up of RIBs: coincidences

Self-conjugated nuclei → cluster effect in GS

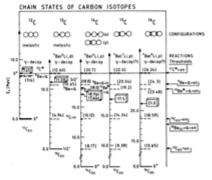
→ role of quartetting in nuclei → unveiled from the existence of rotational bands

Neutron-rich light isotopes → cluster effect leading to nuclear molecules -> covalence bonding due to neutrons

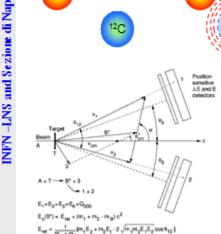
> Light isotopes approaching ndrip line > exotic clustering effects -> also influence of pairing in covalent neutrons

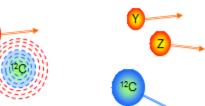


16C

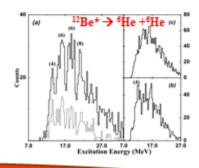






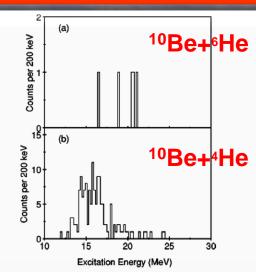


ith RIBs



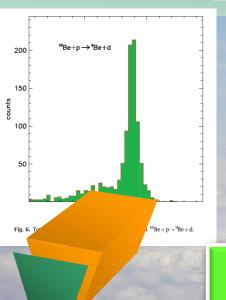
Low statistics & high resolution

<sup>16</sup>C+<sup>12</sup>C Search for exotic decay of <sup>16</sup>C ( and the other available beams <sup>10,11</sup>Be <sup>13</sup>B .....)



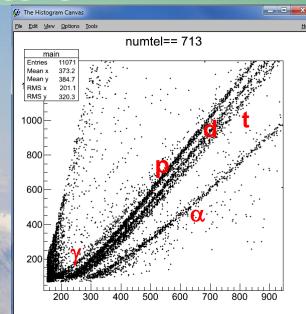
N. I. Ashwood et al Phys.Rew.C 70, 064607 (2004)

- Another improvement -  $\gamma$ -ray tagging? -

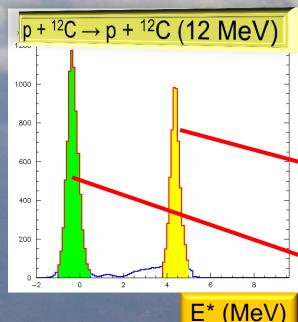


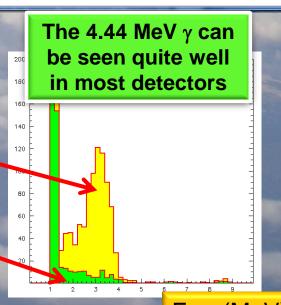
γ-ray tagging could be a solution to improve the kinematical coincidence method in case of excited levels - How to combine efficient γ-ray detectors and CHIMERA?

Csl(TI) have a large efficiency for gamma ray detection



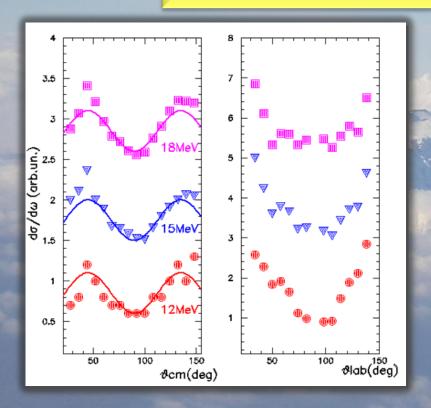
We have measured the γ-ray signals from Csl(tl) using proton beams on carbon target and looking at excitation and decay of the 4.44 MeV <sup>12</sup>C first excited state





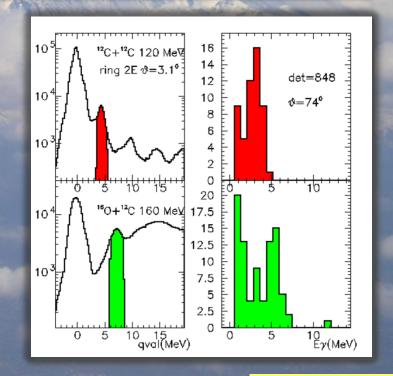
– Using the  $\gamma$ -ray tagging: can we measure  $\gamma$ -rays angular distributions?-

We can also measure γ-ray angular distribution from 4.44 <sup>12</sup>C excited state

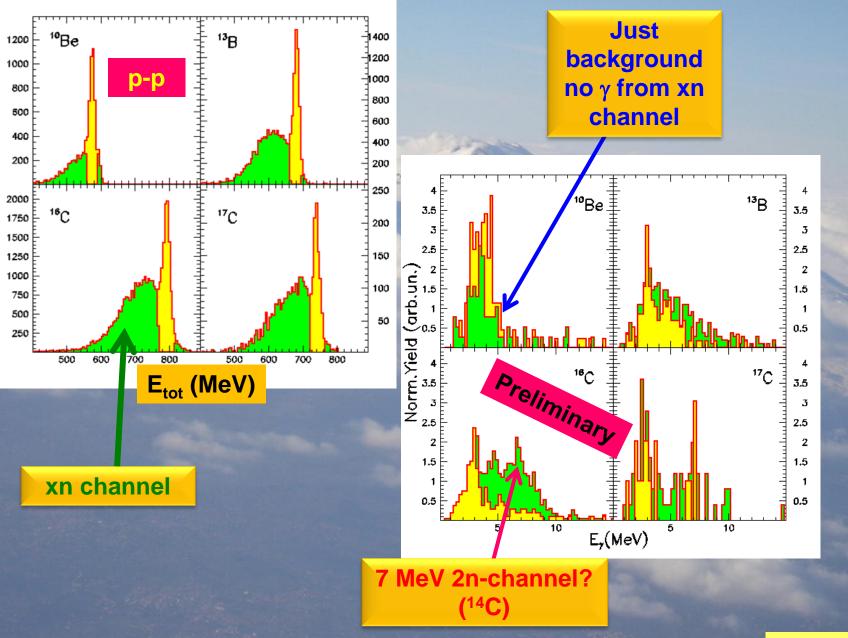


CM angular distribution was evaluated in the system of recoiling <sup>12</sup>C after inelastic scattering of protons at various energies

Using calibration beams we see also higher energy  $\gamma$ -rays as those emitted by  $^{16}\text{O}$  around 6 MeV with overall efficiency around 30% of  $4\pi$  (GEANT-IV)



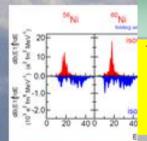
## – Using the $\gamma$ -ray tagging: Preliminary $\gamma$ spectra with radioactive beams-



## **Coming Experiments: PIGMY**

## Search for iso-scalar excitation of the PIGMY resonance in <sup>68</sup>Ni nuclei

Spokes: G.Cardella, E.G.Lanza for the EXOCHIM collaboration

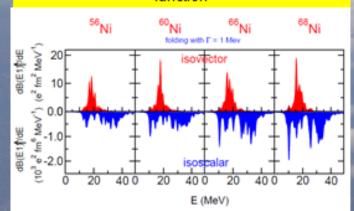


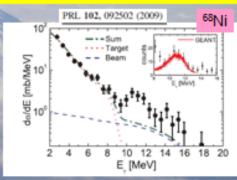


## The Pigmy resonance

The search for population and decay of the Pigmy resonance was particular stressed in the last years especially due to the results obtained with neutron rich nuclei at GSI. The interest was high also because its sensitivity to the symmetry term of the nuclear equation of state - A recent revue can be found in Progress in Particle and Nuclear Physics 70 (2013) 210 by D. Savran, T. Aumann, A. Zilges

Experiments at GSI were performed using <sup>132</sup>Sn and <sup>68</sup>Ni – The resonance was excited by virtual photons generated by the Coulomb field of heavy target nuclei, so probing its isovector response function

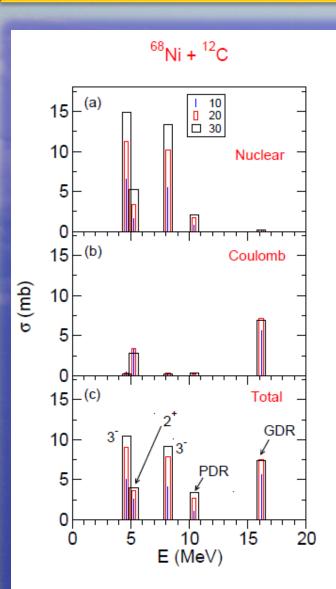




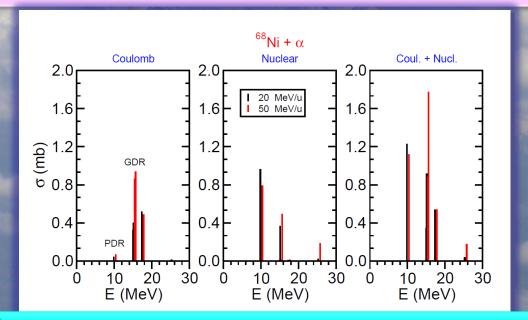
However various calculations show that this resonance can be excited also using isoscalar probes

## The Pigmy resonance: cross section predictions

A search for isoscalar population of the pigmy resonance on unstable nuclei is very interesting in order to fully understand the characteristics of such resonance



Semi-classical calculations (based on RPA strengths) predict that the best beam energy for such a search is around 30 A MeV - A sizeable cross section is predicted for an isoscalar probe as <sup>12</sup>C (around 3mb)



<sup>12</sup>C seems the best target – its cross section is expected to be at least 2 times larger than  $\alpha$ -particles G.Cardella iwm-ec2014

## <sup>68</sup>Ni Beam for the Pigmy experiment

68Ni beam was recently produced at LNS in the framework of the TIMESCALEZN experiment

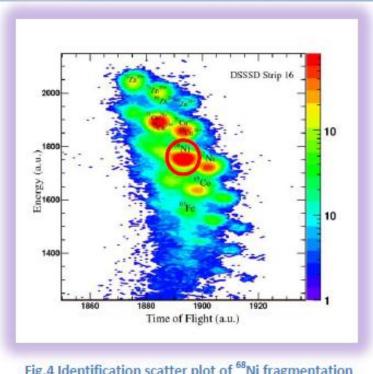
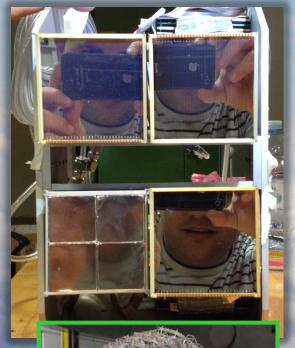


Fig. 4 Identification scatter plot of <sup>68</sup>Ni fragmentation beam The beam was produced via fragmentation of <sup>70</sup>Zn on a 0.25 mm <sup>9</sup>Be target at 40 MeV/A

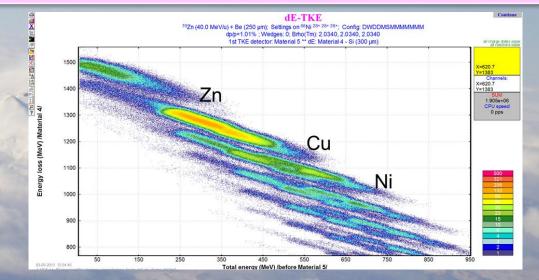
A beam intensity of about 2x10<sup>4</sup> part/sec/ 100 W primary beam was obtained

We also demonstrated that we can clean our beam from not fully stripped ions - The mylar foil of the tagging MCP is a stripper foil cleaning most of such contaminants – thank to this <sup>68</sup>Ni is the most intense beam transported

## **Detection system for the Pigmy experiment**



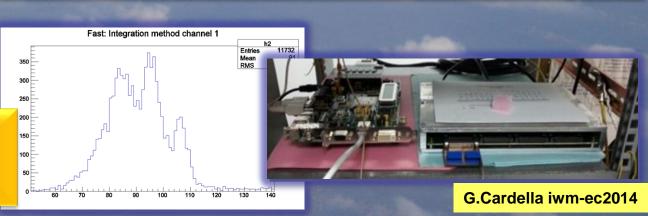
FARCOS will detect and identify <sup>68</sup>Ni with good energy resolution (stopped in the two silicon stages of the telescopes)



The Sphere CsI with the new digital GET electronics will provide  $\gamma$  detection

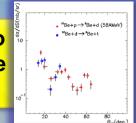


The 4.44 γ-ray seen by a farcos CsI with GET electronics

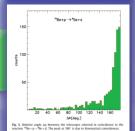


## **Conclusions and perspectives**

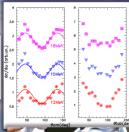
Using cocktail of neutron rich beams with the CHIMERA detector we are able to extract angular distributions for many reaction channels searching for structure effects on cross sections



The  $4\pi$  detection efficiency is very useful and allows extensive use of the kinematical coincidence technique



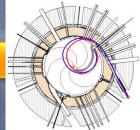
We can also detect  $\gamma$ -rays with our CsI(TI) detectors in order to tag excited levels



For the future experiments we are working to improve our detection capabilities and resolutions also coupling CHIMERA to a new high resolution strip telescope array FARCOS



Very exciting long term perspectives by the intensity upgrade of CS



## I wish to thank all my collaborators

L.Acosta<sup>1</sup>, L.Auditore<sup>4</sup>, I.Berceanu<sup>9</sup>, G.C.<sup>2</sup>, M.B.Chatterjee<sup>10</sup>, E.De Filippo<sup>2</sup>, L.Francalanza<sup>1,3</sup>, G.Lanzalone<sup>1,6</sup>, I.Lombardo<sup>8</sup>, E.Morgana<sup>2,3</sup>, T.Minniti<sup>4</sup>, S.Norella<sup>4</sup>, A.Pagano<sup>2</sup>, E.V.Pagano<sup>1,3</sup>, M.Papa<sup>2</sup>, S.Pirrone<sup>2</sup>, G.Politi<sup>2,3</sup>, A.Pop<sup>9</sup>, L.Quattrocchi<sup>4</sup>, F.Rizzo<sup>1,3</sup>, E.Rosato<sup>8</sup>, P.Russotto<sup>2,3</sup>, A.Trifirò<sup>4</sup>, M.Trimarchì<sup>4</sup>, G.Verde<sup>2</sup>, M.Vigilante<sup>8</sup>

- INFN Laboratori Nazionali del Sud, Catania, Italy
   INFN, Sezione di Catania, Catania, Italy
- 3) Dipartimento di Fisica e Astronomia Università di Catania, Italy
- 4) INFN gruppo collegato di Messina & Dipartimento di Fisica e Astronomia Università di Messina, Italy
  - 5) Centro Siciliano di Fisica Nucleare e Struttura della Materia, Italy 6) Università Kore di Enna, Enna, Italy
- 7) INFN Sezione di Milano & Dipartimento di Fisica e Astronomia Università di Milano, Italy
- 8) INFN Sezione di Napoli & Dipartimento di Fisica e Astronomia Università di Napoli, Italy
  - 9) Institute for Physics and Nuclear Engineering, Bucharest, Romania 10) Saha Institute of Nuclear Physics, Kolkata, India