

**International
Workshop on
Multi facets of
Eos and
Clustering**

IWM-EC 2014

IWM-EC 2014

6th – 9th May 2014 Catania, Italy

**The ASY-EOS experiment at GSI:
investigating symmetry energy at supra-saturation densities**

P. Russotto*

for the ASY-EOS collaboration



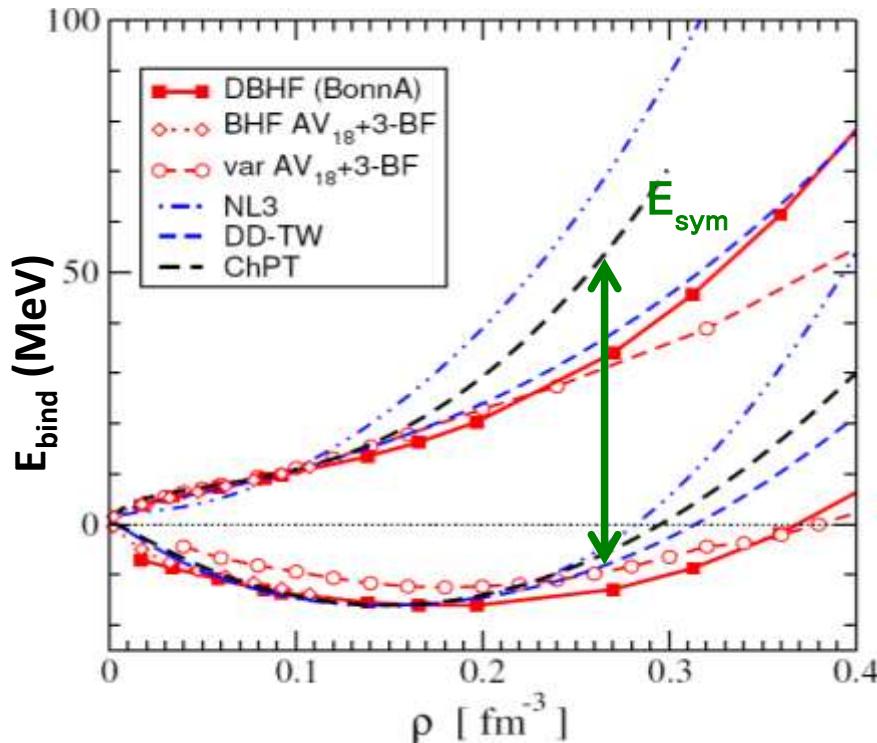
*INFN-Sez. di Catania, Italy

Summary

- Brief intro on High Density investigation of Symmetry Energy with H-I
- Suggestion for a route toward model independency...
- First results of the ASY-EOS experiment
- Future possibilities
- Conclusions

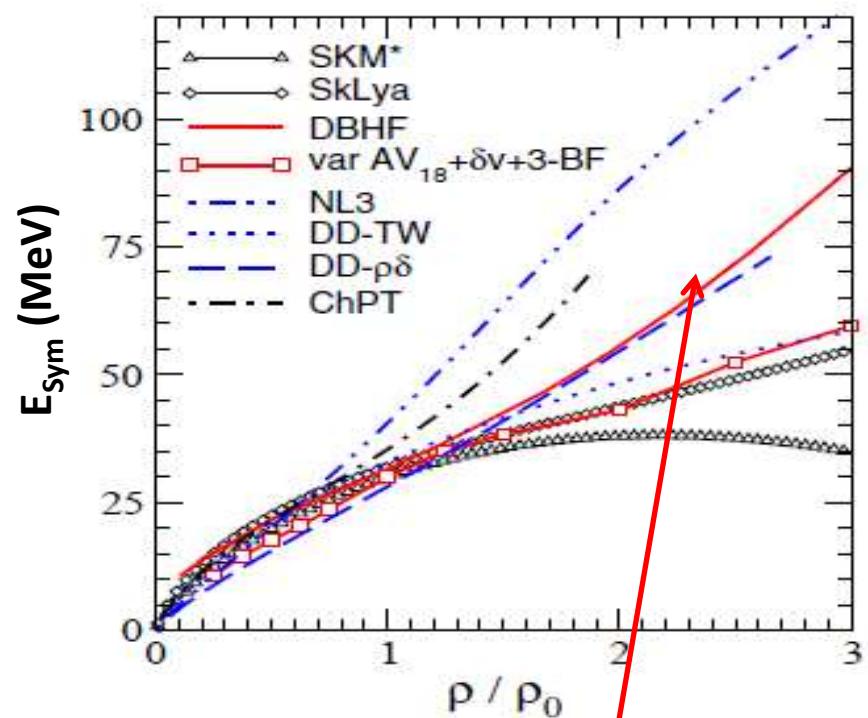
Symmetry Energy

$$E_{sym}(\rho) = E(\rho, I=1) - E(\rho, I=0)$$



EOS of symmetric nuclear
and neutron matter
from
Ab initio calculations (red)
and phenomenological approaches

$$I = \frac{N - Z}{N + Z}$$



High density...so important!

Constraints of the Symmetry Energy

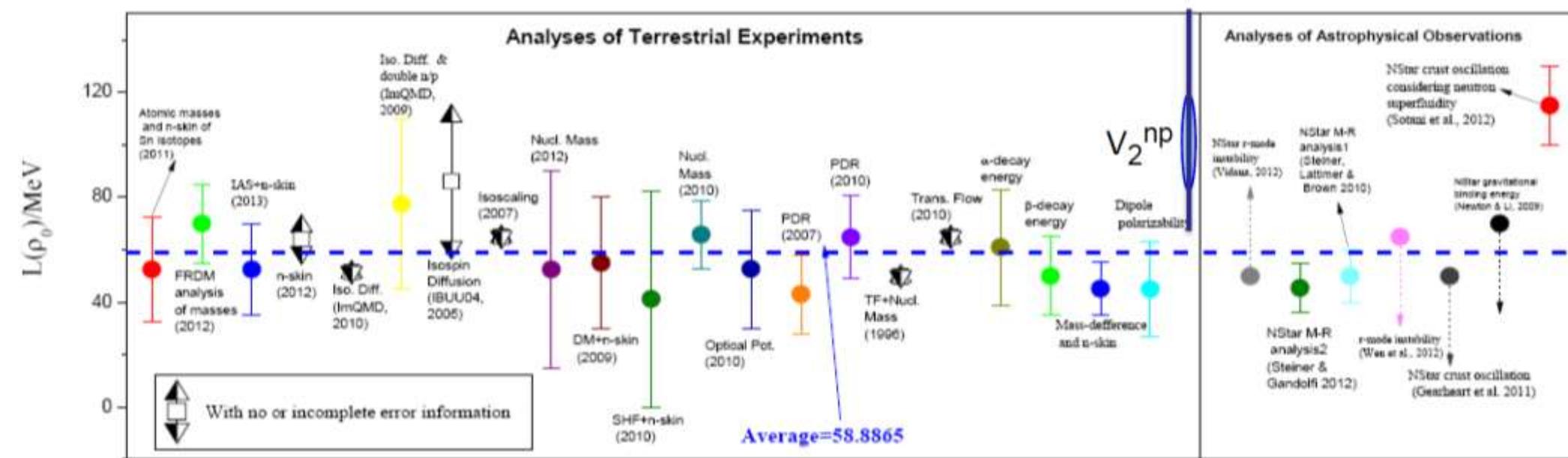
B.A. Li NuSym13
summary talk

$$E_{Sym}(\rho) = S(\rho) = S_0 + \frac{L}{3} \left(\frac{\rho - \rho_o}{\rho_o} \right) + \frac{K_{\text{sym}}}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots,$$

Terrestrial laboratories

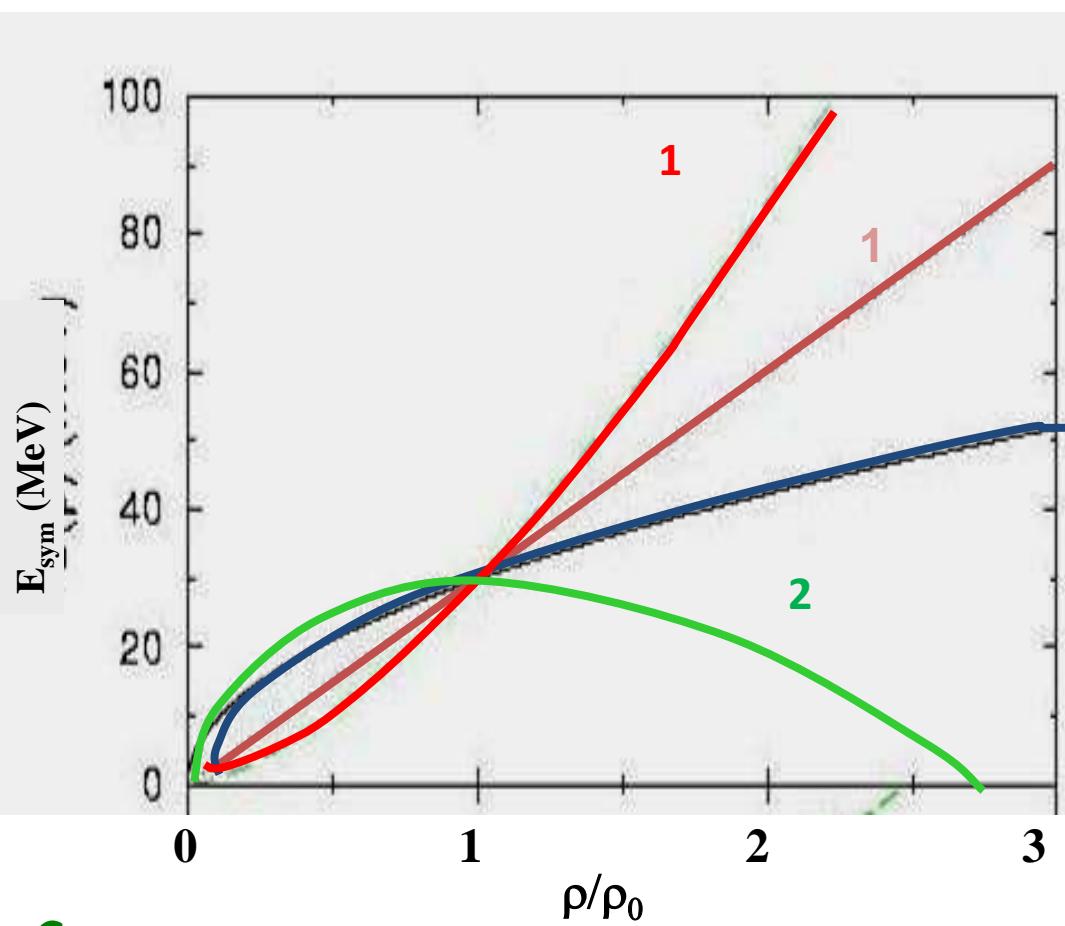
	S_0	L
average of the means	31.55415	58.88646
standard deviation	0.915867	16.52645

- Several constraints (quite consistent among them) around and below ρ_0
- Few constraints above ρ_0



E_{sym} at high density: π^-/π^+ ratio

(measured by FOPI, for different systems and energies, as compared to different models)



See:

Z. Xiao et al., PRL 102 (2009) IBUU04

Z.Q. Feng, PLB 683 (2010) ImIQMD

W.J. Xie , et al., PLB 718 (2013) ImIBL

G. Ferini, et al., NPA 762 (2005) RMF

Ad. from IWM 2011 - Y. Leifels

- Results model dependent
- Density dependence of symmetry energy unambiguously soft or hard
BUT
 - symmetry energy \rightarrow n/p ratio, number of nn, np, pp collisions
 - asystiff $\frac{n}{p} \downarrow \Rightarrow \frac{Y(\Delta^{0,-})}{Y(\Delta^{+,++})} \downarrow \Rightarrow \frac{\pi^-}{\pi^+} \downarrow$
- medium \rightarrow effective masses (N, π, Δ), cross sections \rightarrow thresholds
- asystiff $\Rightarrow \frac{\pi^-}{\pi^+} \uparrow$
- \rightarrow Interpretation of pion data not straight forward

Kaons: more sensitive probes?

- Higher thresholds
- Weakly interacting in medium
- Freeze-out already at 20 fm/c: more reliable than pions as high ρ probes

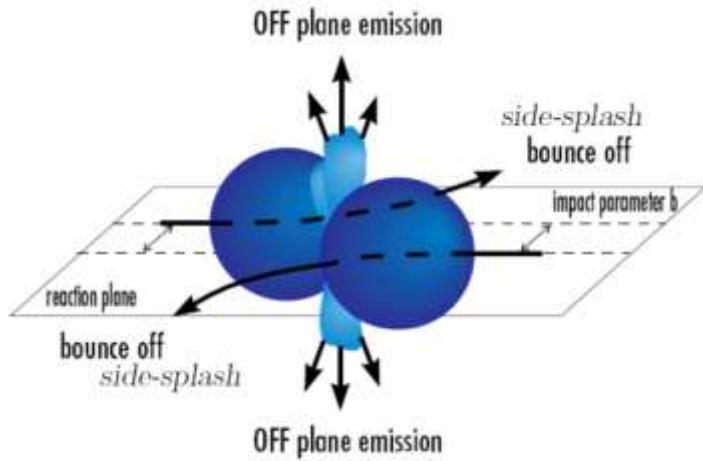
High densities: flows

$$\frac{dN}{d(\phi - \phi_R)}(y, p_t) = \frac{N_0}{2\pi} \left(1 + 2 \sum_{n \geq 1} v_n \cos n(\phi - \phi_R) \right)$$

y = rapidity
 p_t = transverse momentum

$$V_2(y, p_t) = \left\langle \frac{p_x^2 - p_y^2}{p_t^2} \right\rangle$$

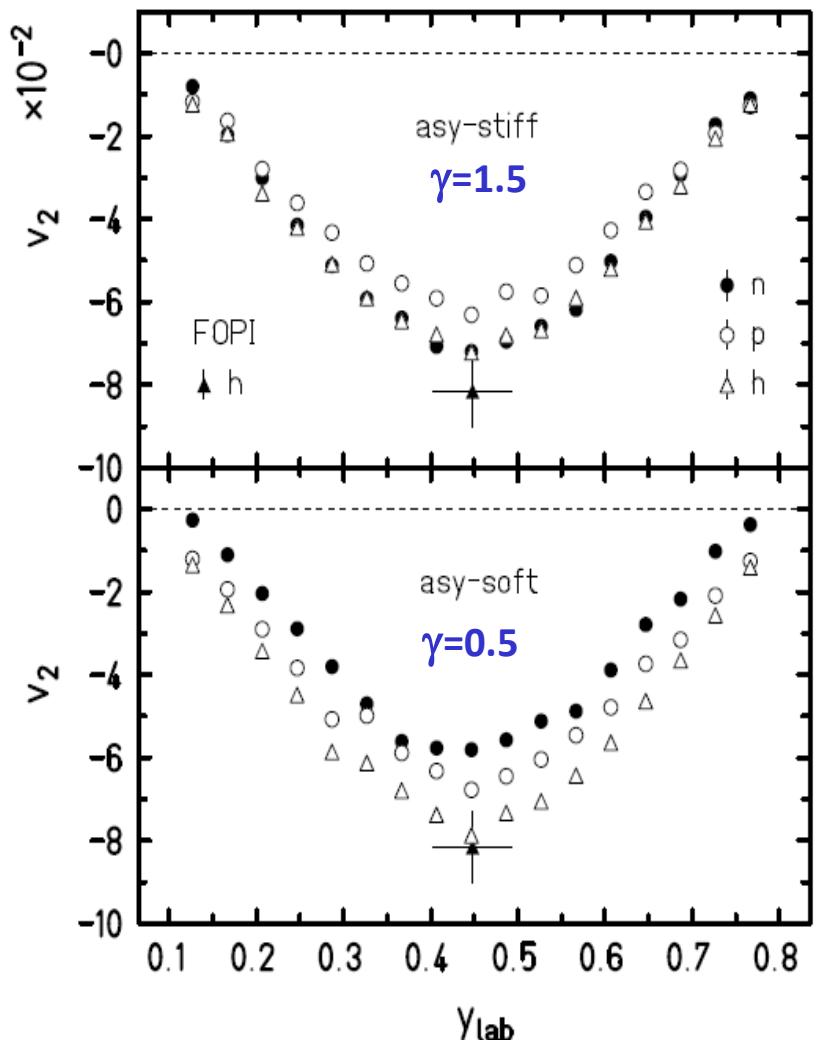
Elliptic flow: competition between in plane ($V_2 > 0$) and out-of-plane ejection ($V_2 < 0$)



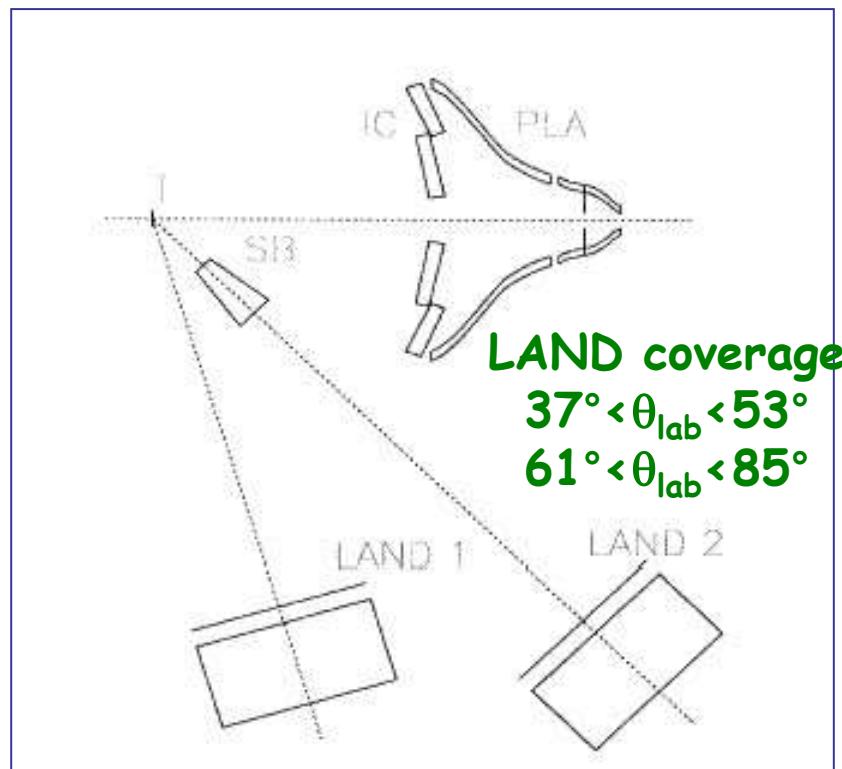
$$E_{\text{sym}} = E_{\text{sym}}^{\text{pot}} + E_{\text{sym}}^{\text{kin}}$$

$$= 22 \text{ MeV} \cdot (\rho/\rho_0)^{\gamma} + 12 \text{ MeV} \cdot (\rho/\rho_0)^{2/3}$$

UrQMD : Au+Au @ 400 AMeV
 $5.5 < b < 7.5 \text{ fm}$



FOPI/LAND experiment on neutron squeeze out (1991)



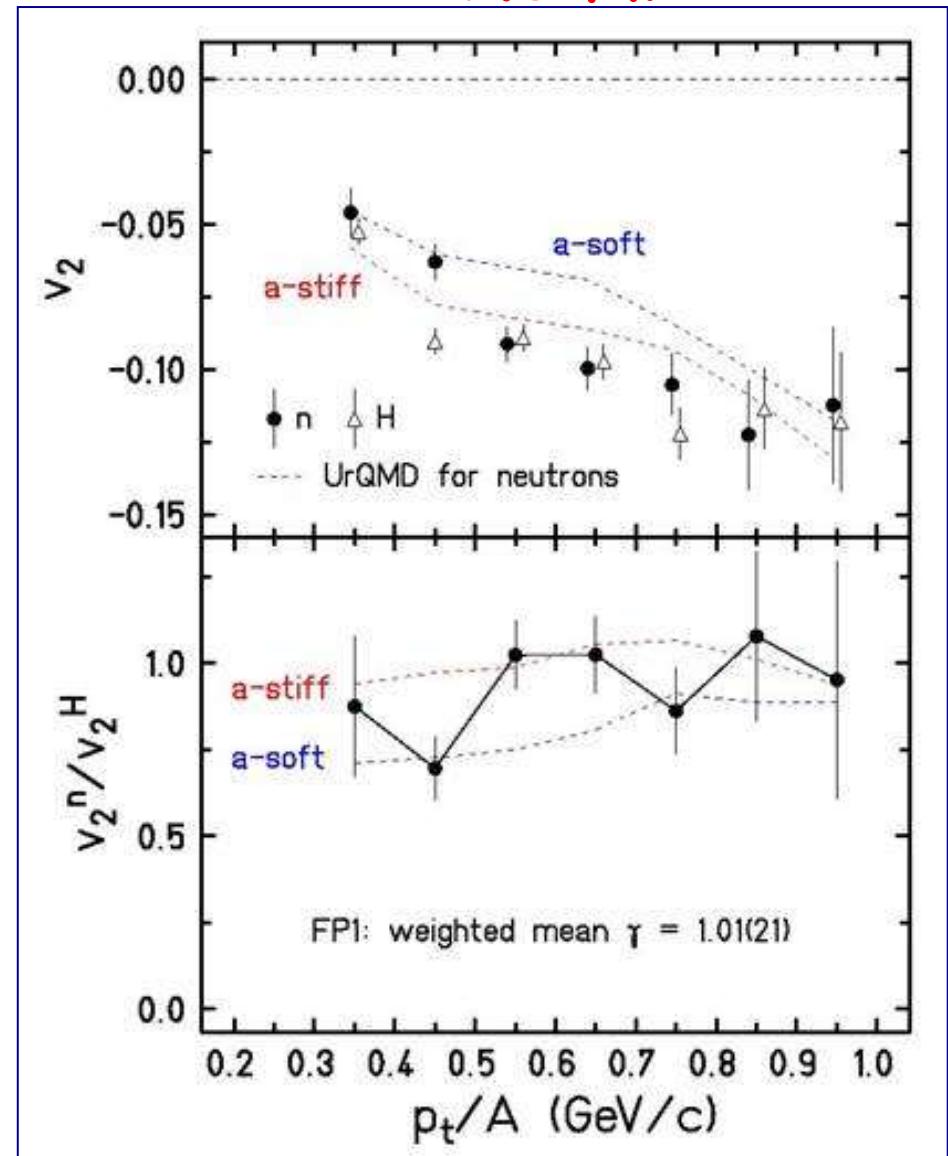
$$\gamma = 0.9 \pm 0.4$$

$$L = 83 \pm 26$$

$$E_{\text{sym}} = E_{\text{sym}}^{\text{pot}} + E_{\text{sym}}^{\text{kin}}$$

$$= 22 \text{ MeV} \cdot (\rho/\rho_0)^\gamma + 12 \text{ MeV} \cdot (\rho/\rho_0)^{2/3}$$

Au+Au 400 A MeV
 $b < 7.5 \text{ fm}$



Y. Leifels et al., PRL 71, 963 (1993)
P. Russotto et al., PLB 697 (2011)

Results with Tübingen QMD

UrQMD:

momentum dep. of isoscalar field

momentum dep. of NNECS

momentum independent power-law

parameterization of the symmetry energy

Tübingen-QMD:

density dep. of NNECS

asymmetry dep. of NNECS

soft vs. hard EoS

width of wave packets

momentum dependent (Gogny inspired)
parameterization of the symmetry energy

M.D. Cozma, PLB 700, 139 (2011);
[arXiv:1102.2728](https://arxiv.org/abs/1102.2728)

$$x = -1.35 \pm 1.25$$

M.D. Cozma et al., Towards a model-independent constraint of the high-density dependence of the symmetry energy.

[arXiv:1305.5417](https://arxiv.org/abs/1305.5417) [nucl-th] PRC88 044912 (2013)

Au+Au 400 A MeV $b < 7.5$ fm

stiffness

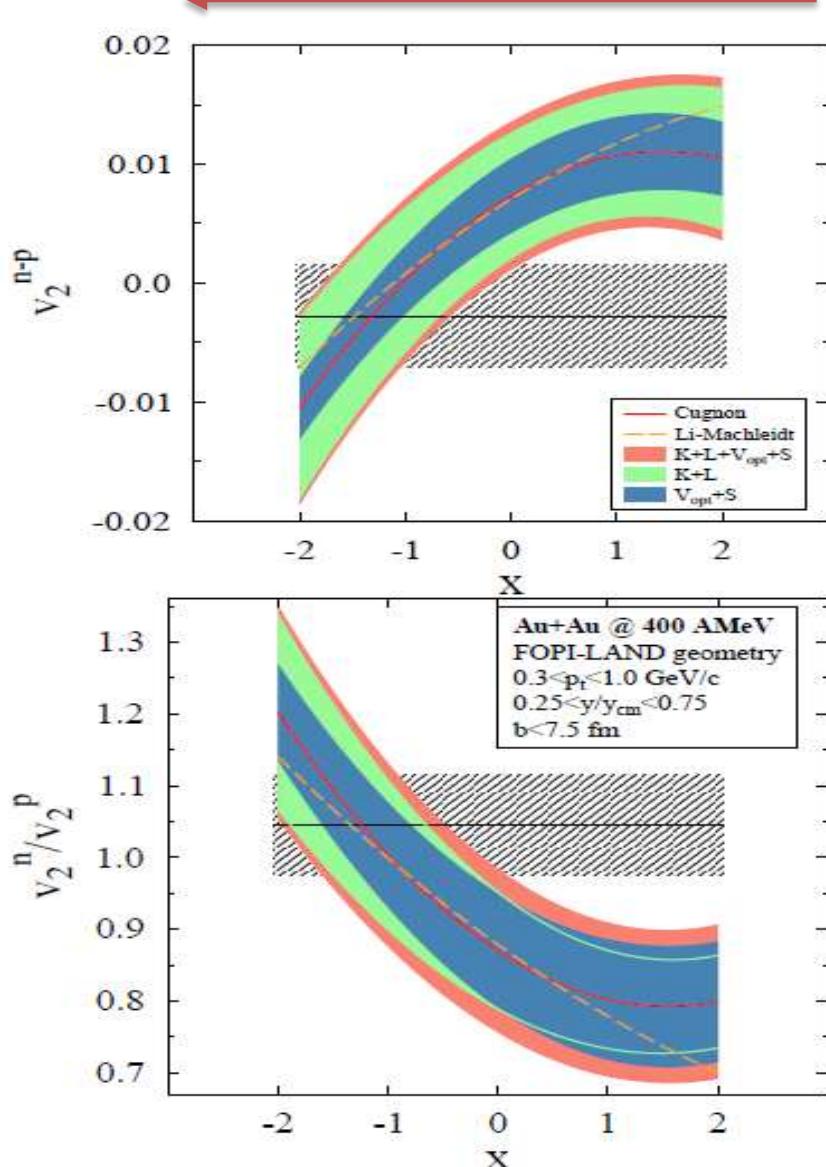
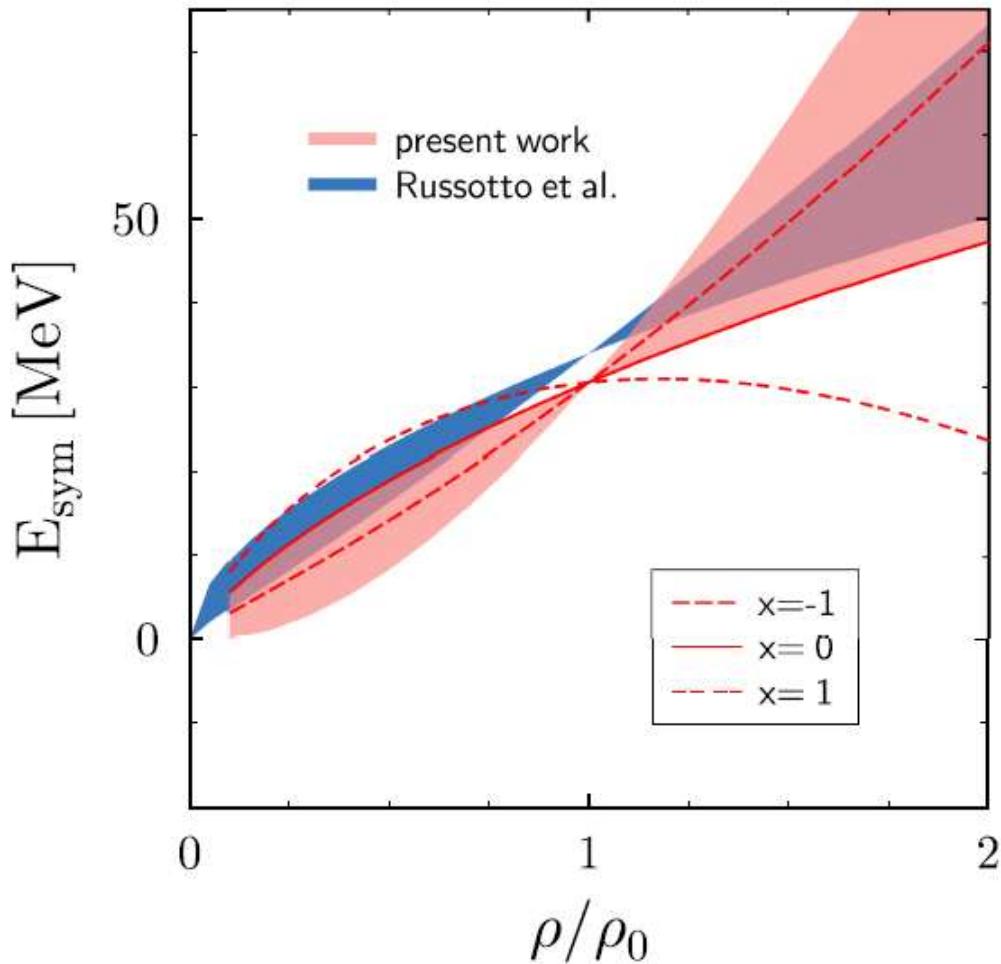


FIG. 2: Model dependence of npEFD and npEFR and comparison with FOPI-LAND experimental data, integrated over impact parameter $b \leq 7.5$ fm. Sensitivity to the different model parameters, compressibility modulus (K), width of nucleon wave function (L), optical potential (V_{opt}) and parametrization of the symmetry energy (S) are displayed. The total model dependence is obtained by adding, in quadrature, individual sensitivities.

Results with Tübingen QMD and UrQMD



$$x = -1.0 \pm 1.0$$

M.D. Cozma et al., Towards
a model-independent
constraint of the high-
density dependence of the
symmetry energy,

[arXiv:1305.5417 \[nucl-th\]](https://arxiv.org/abs/1305.5417)

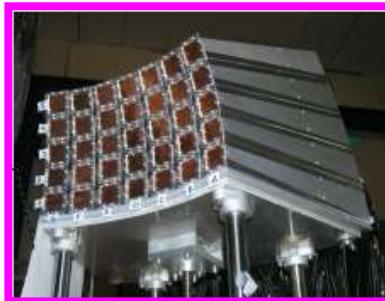
PRC88 044912 (2013)

ASY-EOS S394 experiment @ GSI Darmstadt (May 2011)

Au+Au, $^{96}\text{Zr}+^{96}\text{Zr}$, $^{96}\text{Ru}+^{96}\text{Ru}$ @ 400 AMev



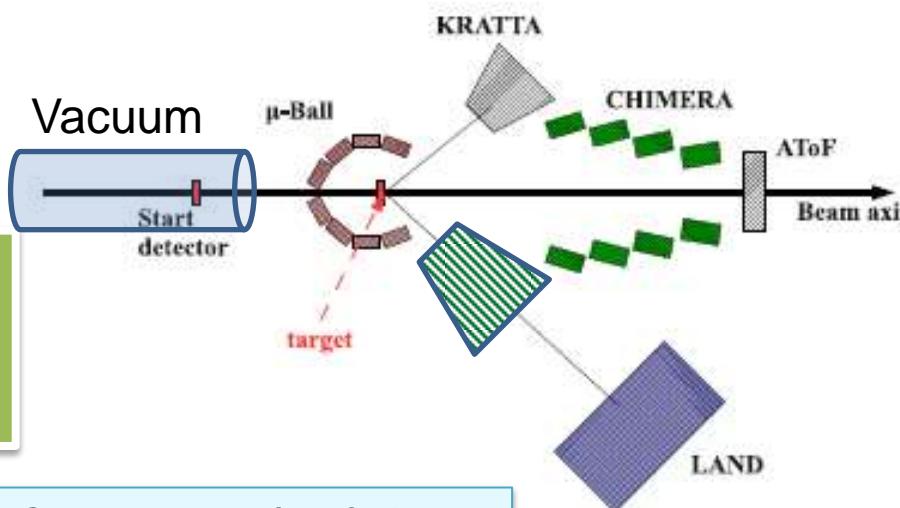
muBall: 4 rings 50 CsI(Tl), $\Theta > 60^\circ$. Discriminate target vs. reactions with air. Multiplicity and reaction plane measurements.



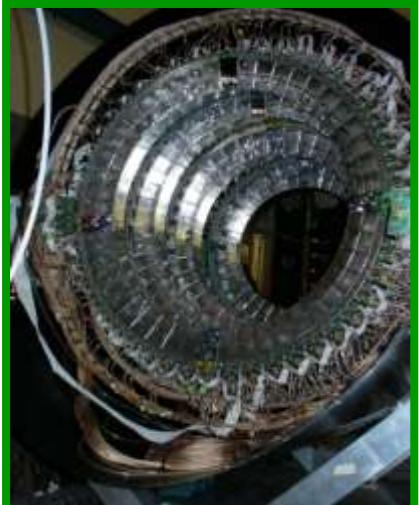
KraTTA: 35 (5x7) triple telescopes (Si-CsI-CsI) placed at $21^\circ < \Theta < 60^\circ$ with digital readout . Light particles and IMFs emitted at midrapidity



Shadow bar:
evaluation of
background neutrons
in LAND



TOFWALL:
96 plastic bars;
ToF, ΔE , X-Y
position.
Trigger, impact
parameter and
reaction plane
determination

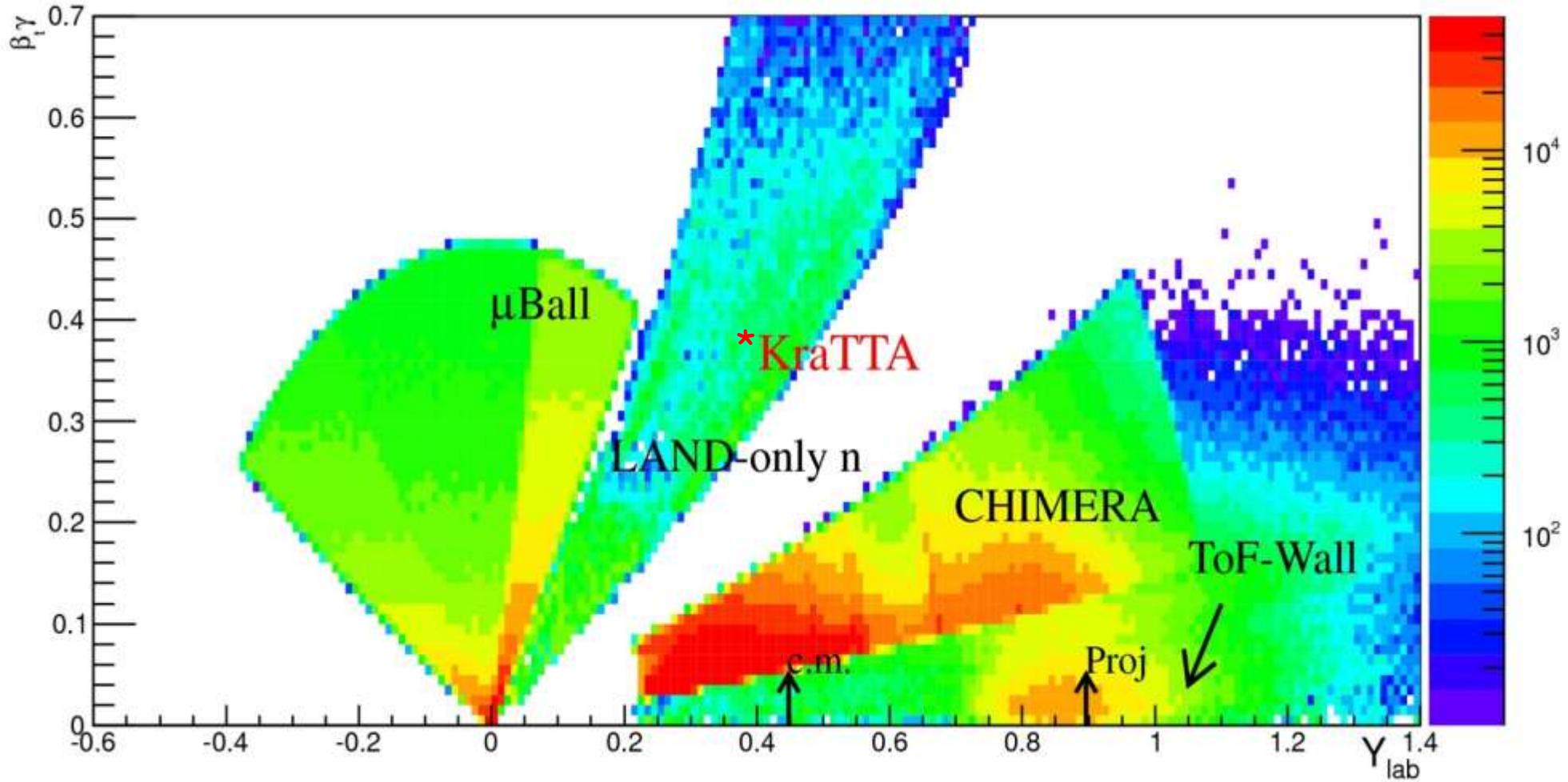


CHIMERA: 8 (2x4) rings, high granularity CsI(Tl), 352 detectors $7^\circ < \theta < 20^\circ$ + 16x2 pads silicon detectors. Light charged particle identification by PSD. Multiplicity, Z, A, Energy: impact parameter and reaction plane determination

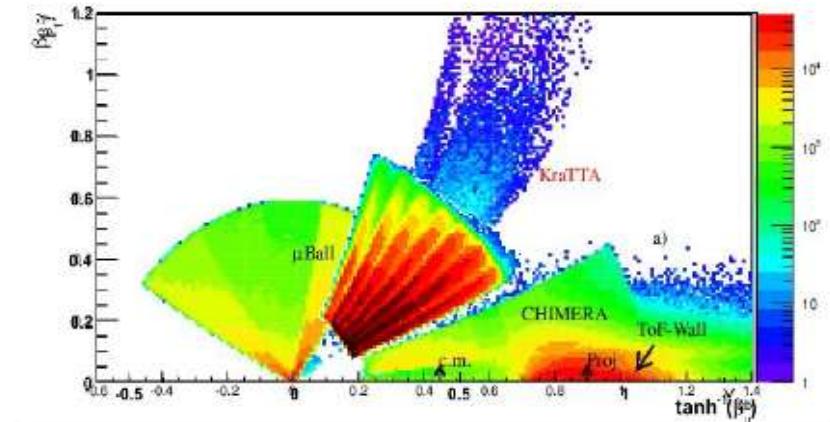


LAND: Large Area Neutron Detector . Plastic scintillators sandwiched with Fe $2 \times 2 \times 1 \text{ m}^3$ plus plastic veto wall. New Taquila front-end electronics. Neutrons and Hydrogen detection. Flow measurements

Au+Au @ 400 A.MeV: Some kinematics



* Random uniform distribution $E_{\text{Kin}} < 100$ Mev

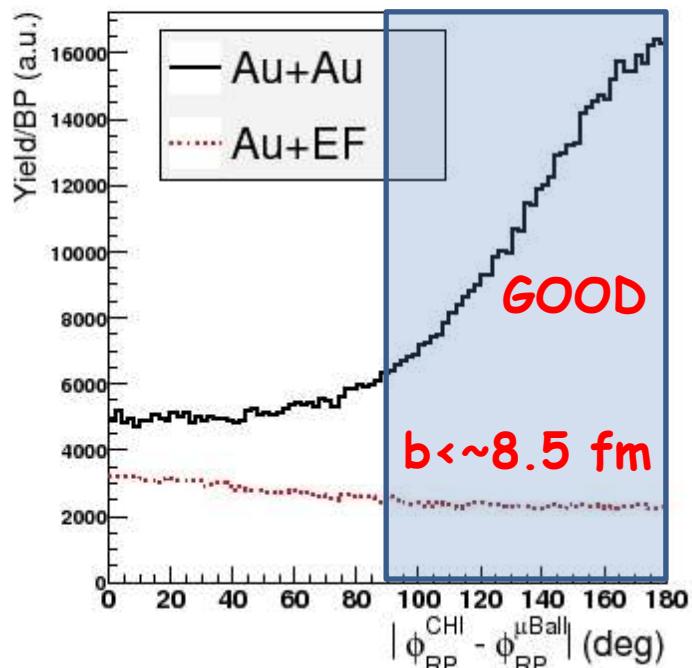
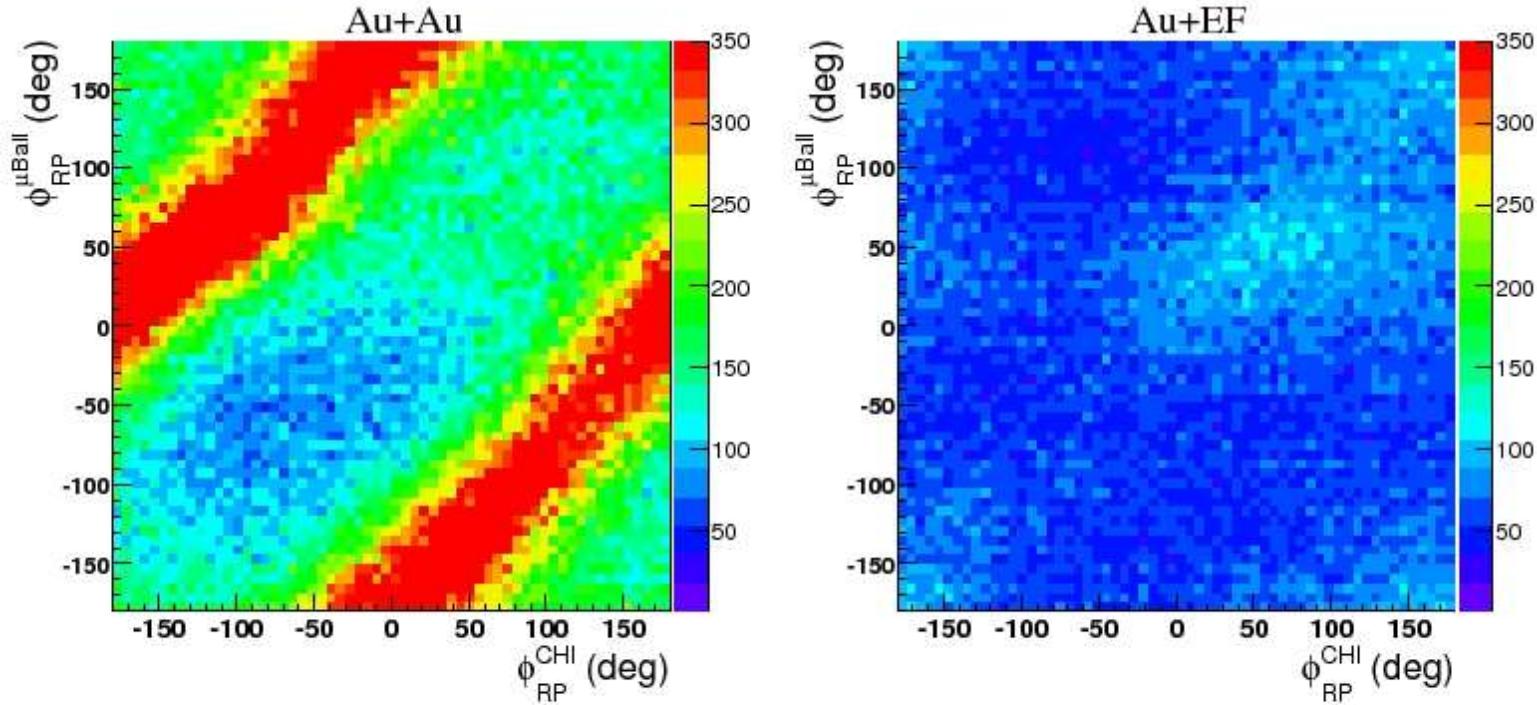


P. Russotto et al., EPJA 50, 38 2014.

P. Russotto et al., Procs. of INPC2013, EPJ Web of Conf.

P. Russotto et al., Journal of Phys. Conf. Series 420, 012092, (2013)

Au+Au @ 400 A.MeV: Background rejection



CHIMERA $M(Y_{c.m}>0.1) \geq 4$

$$\vec{Q} = \sum_{i=1}^M w_i Z_i \gamma \vec{\beta}_t^i \quad w_i = \begin{cases} 1 & \text{for } Y_{cm} > 0.1 \\ 0 & \text{for } Y_{cm} < 0.1 \end{cases}$$

$\mu Ball$ $M \geq 2$

$$\vec{Q} = \sum_{i=1}^M \hat{\beta}_t^i$$

ad. from P. Danielewicz et al., PLB 1985

Au+Au @ 400 A.MeV: Centrality selection

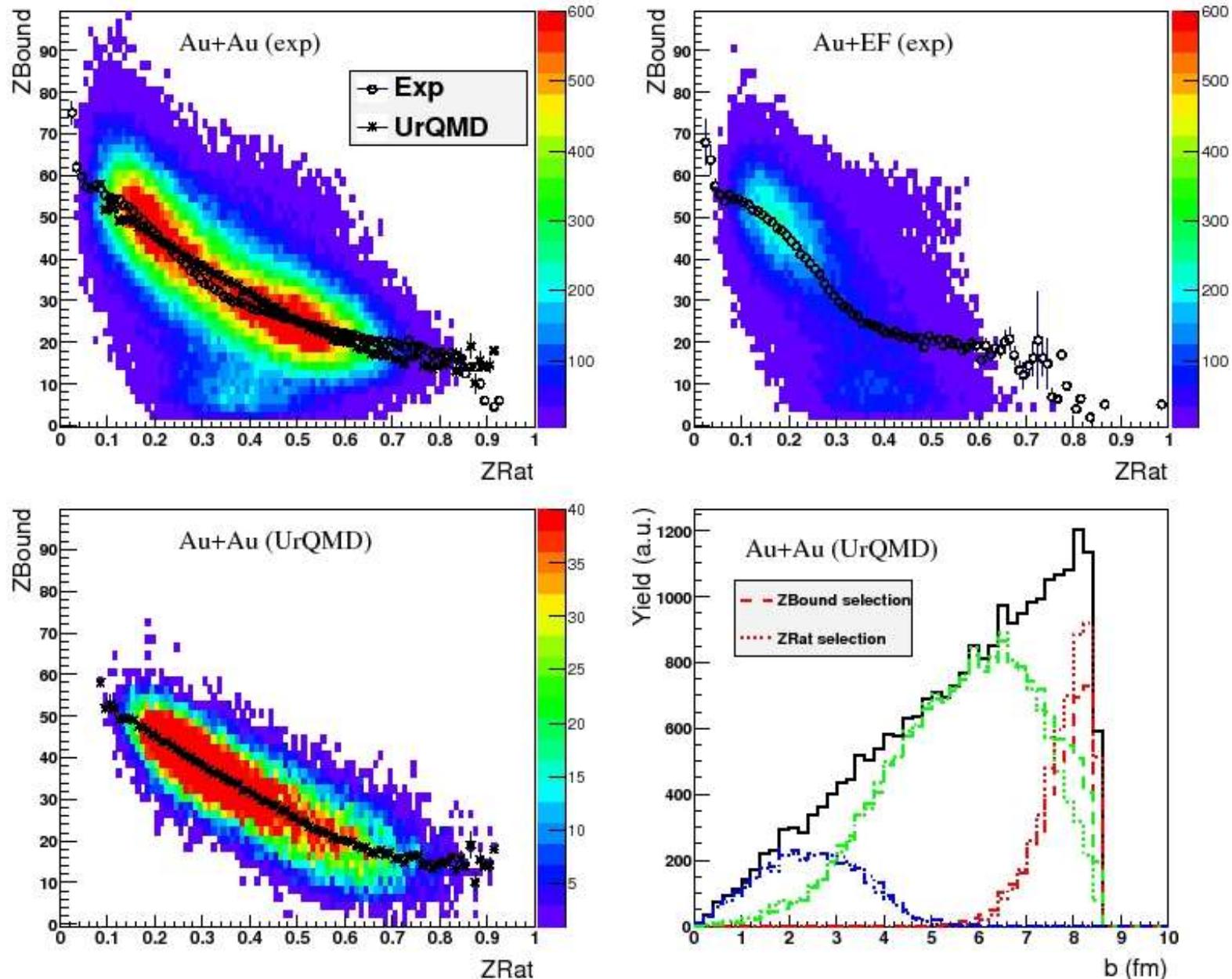
$$Z\text{Bound} = \sum_i Z_i$$

with $Z_i \geq 2$

From
AToF-Wall
and
CHIMERA

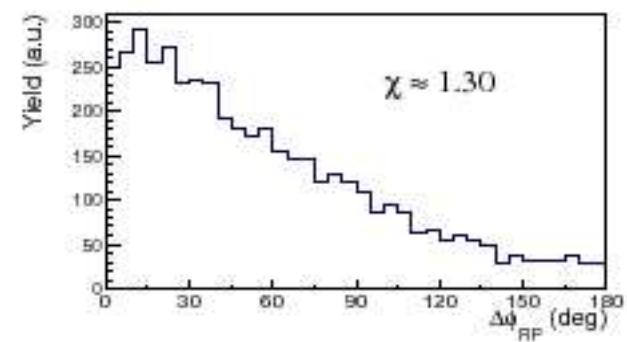
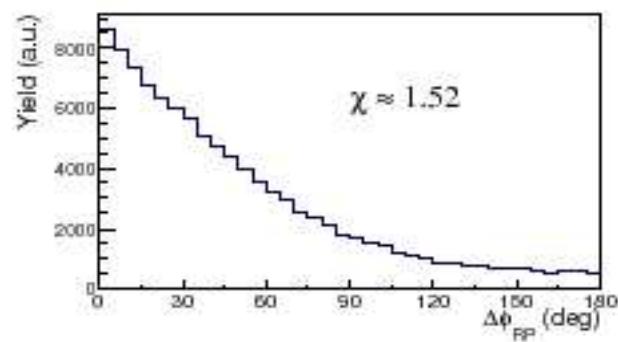
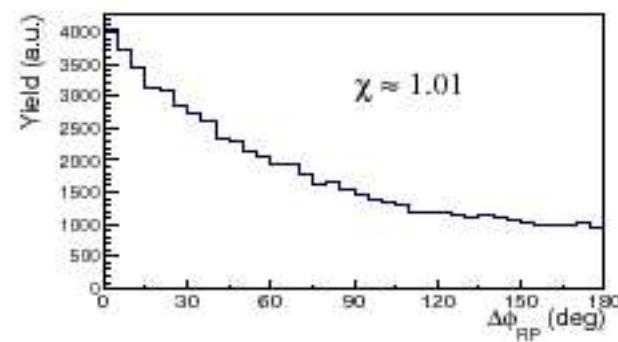
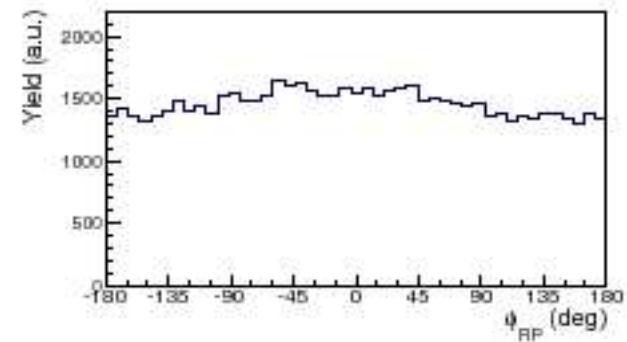
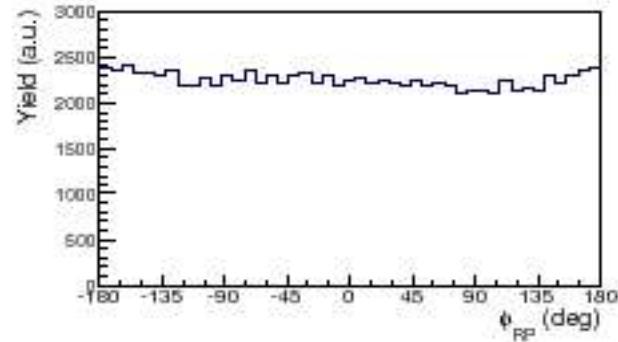
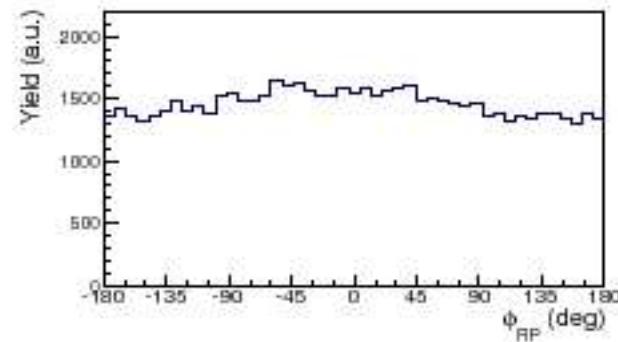
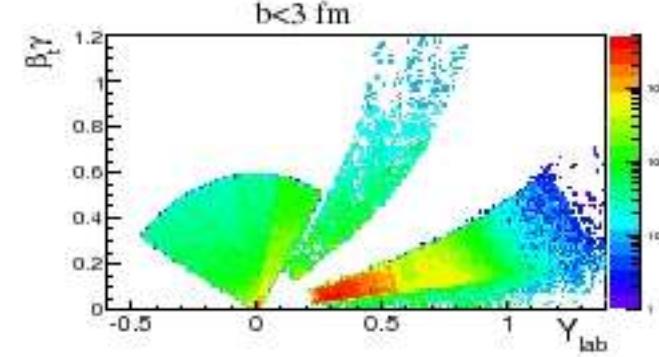
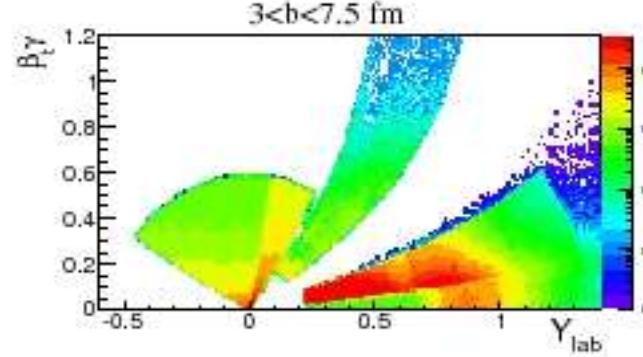
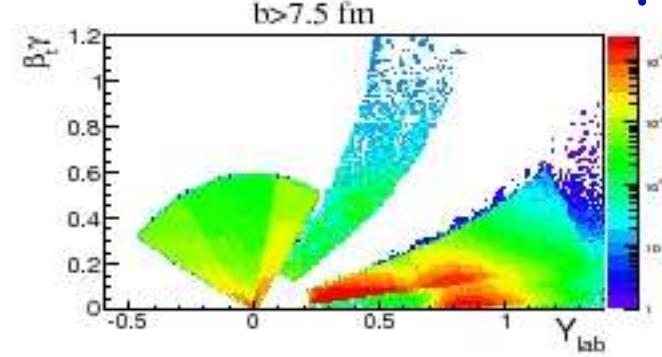
$$Z\text{Rat} = \frac{Z_{\text{trans}}}{Z_{\text{long}}} =$$

$$= \frac{\sum_i Z_i \cdot \sin^2(\vartheta_i^{\text{lab}})}{\sum_i Z_i \cdot \cos^2(\vartheta_i^{\text{lab}})}$$



Reaction plane orientation

Au+Au @ 400 AMeV



CHIMERA

$M(Y_{c.m.} > 0.1) \geq 4$

$$\vec{Q} = \sum_{i=1}^M w_i Z_i \gamma \vec{\beta}_t^i \quad w_i = \begin{cases} 1 & \text{for } Y_{cm} > 0.1 \\ 0 & \text{for } Y_{cm} < 0.1 \end{cases}$$

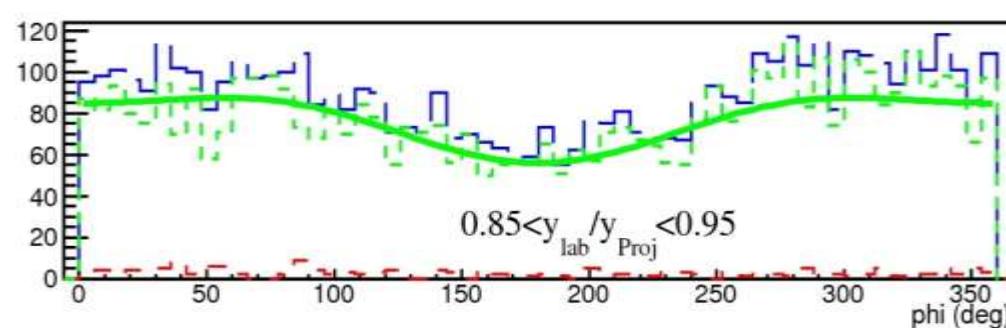
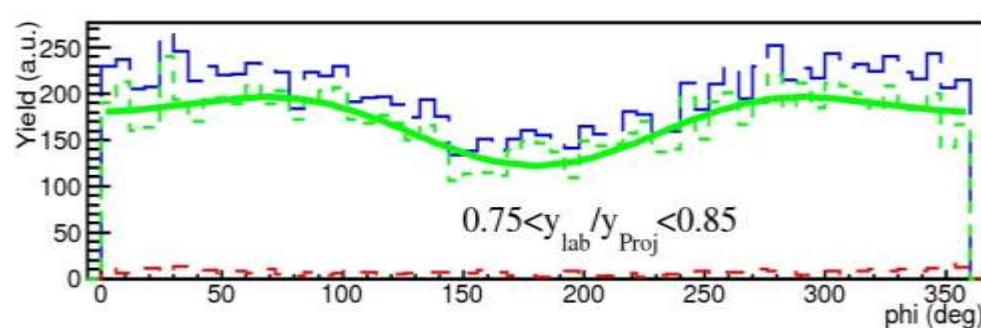
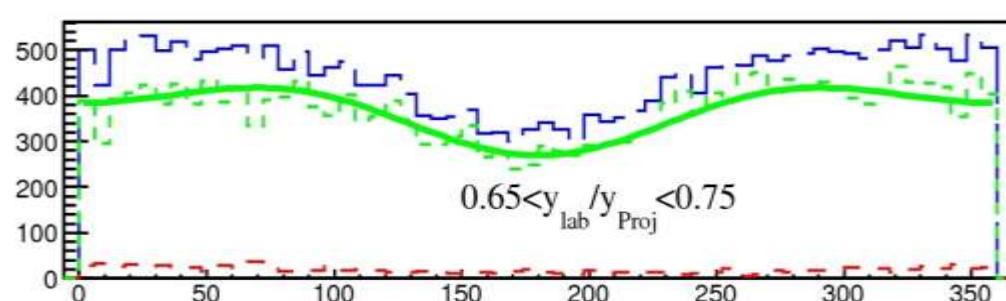
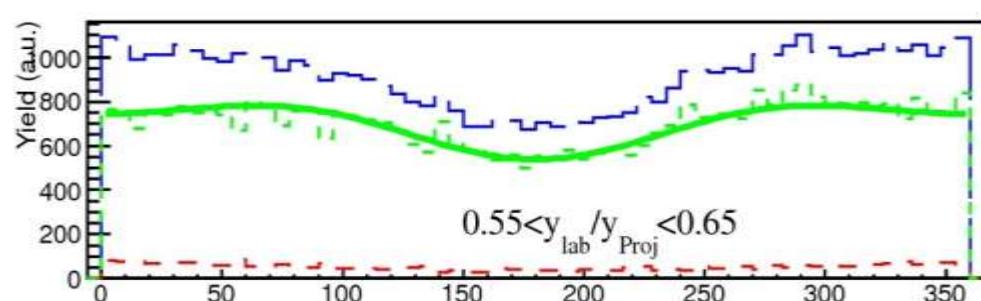
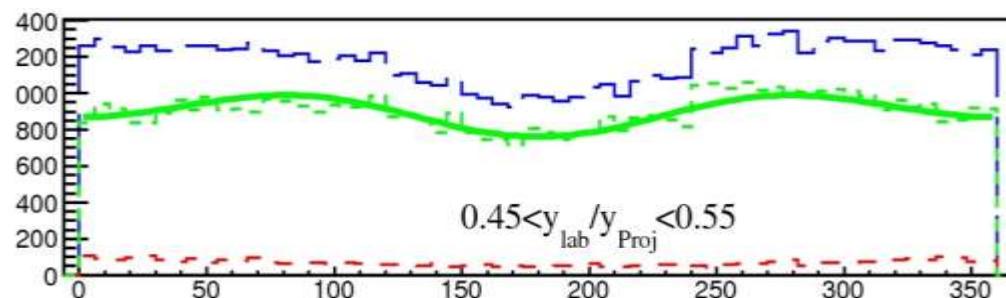
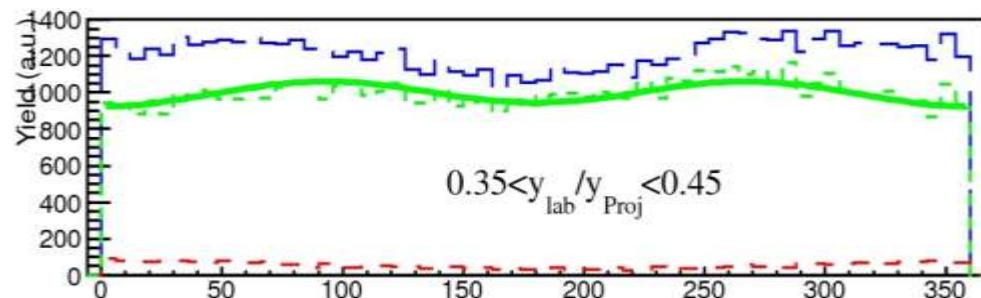
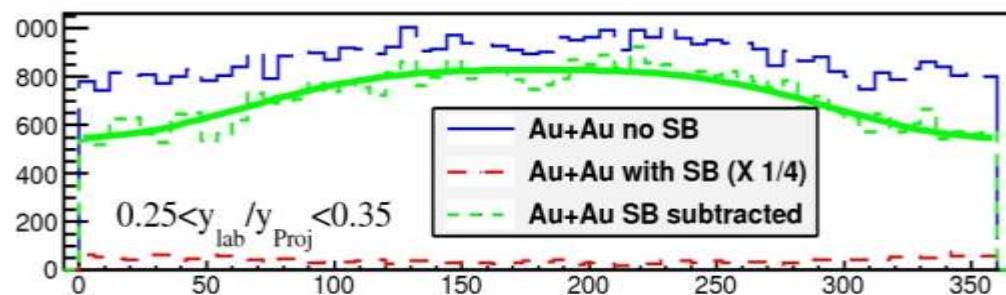
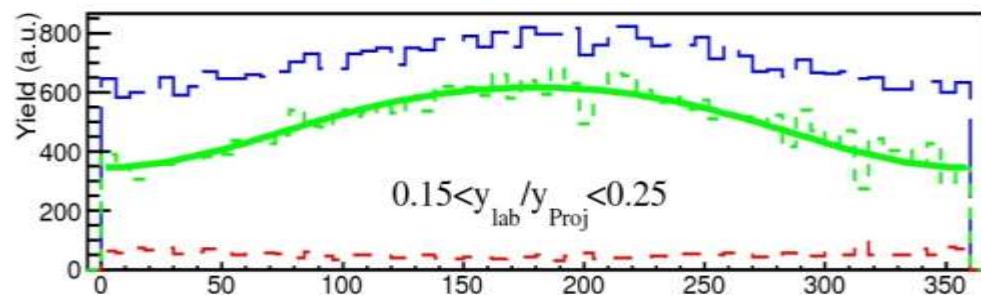
b selection	$x \sim$	$\langle \cos(2\Delta\Phi) \rangle \sim$	$\Delta\Phi \text{ (deg)} \sim$
$b > 7.5 \text{ fm}$	1.01	0.36	34
$3 < b < 7.5 \text{ fm}$	1.52	0.60	26
$b < 3 \text{ fm}$	1.30	0.52	29

ad. from P. Danielewicz et al., PLB 1985

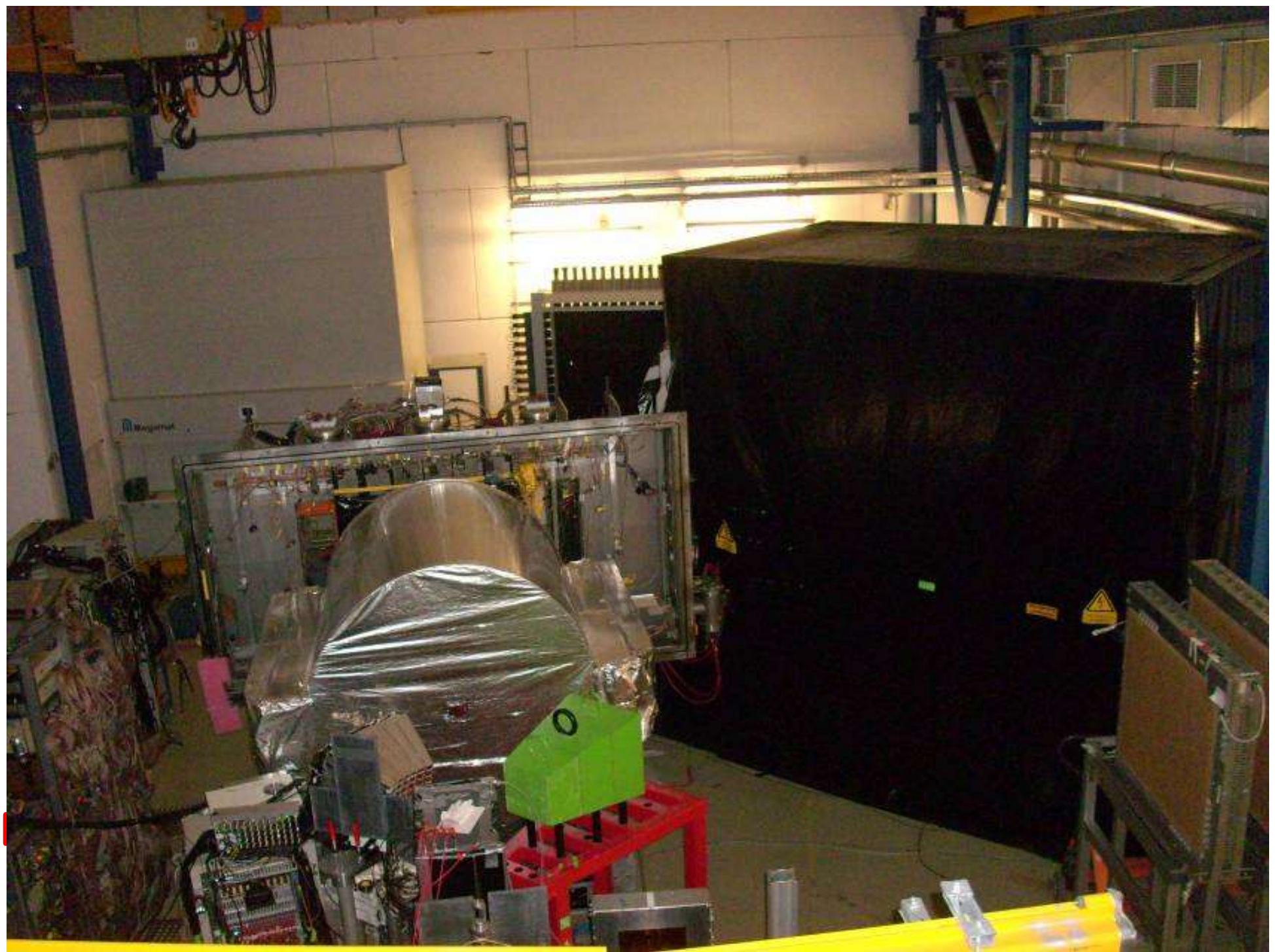
J-Y Ollitrault arXiv:nucl-ex/9711003v2

Neutron azimuthal distributions from LAND

Au+Au @ 400 AMeV
b < 7.5 fm



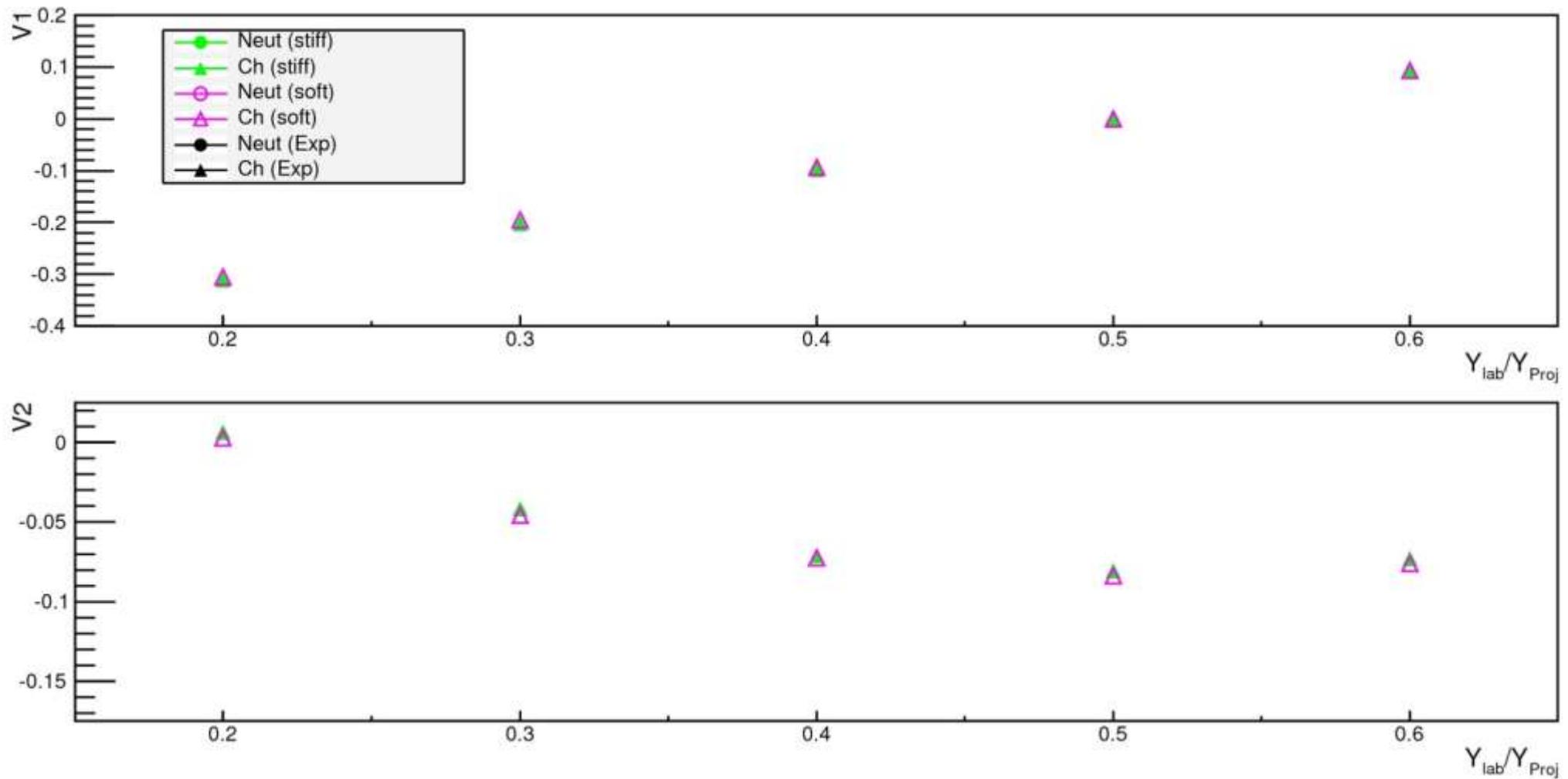
How much background to subtract ?



Comparison with UrQMD

Au+Au @ 400 AMeV $b < 7.5$ fm

Wbg= 110 %



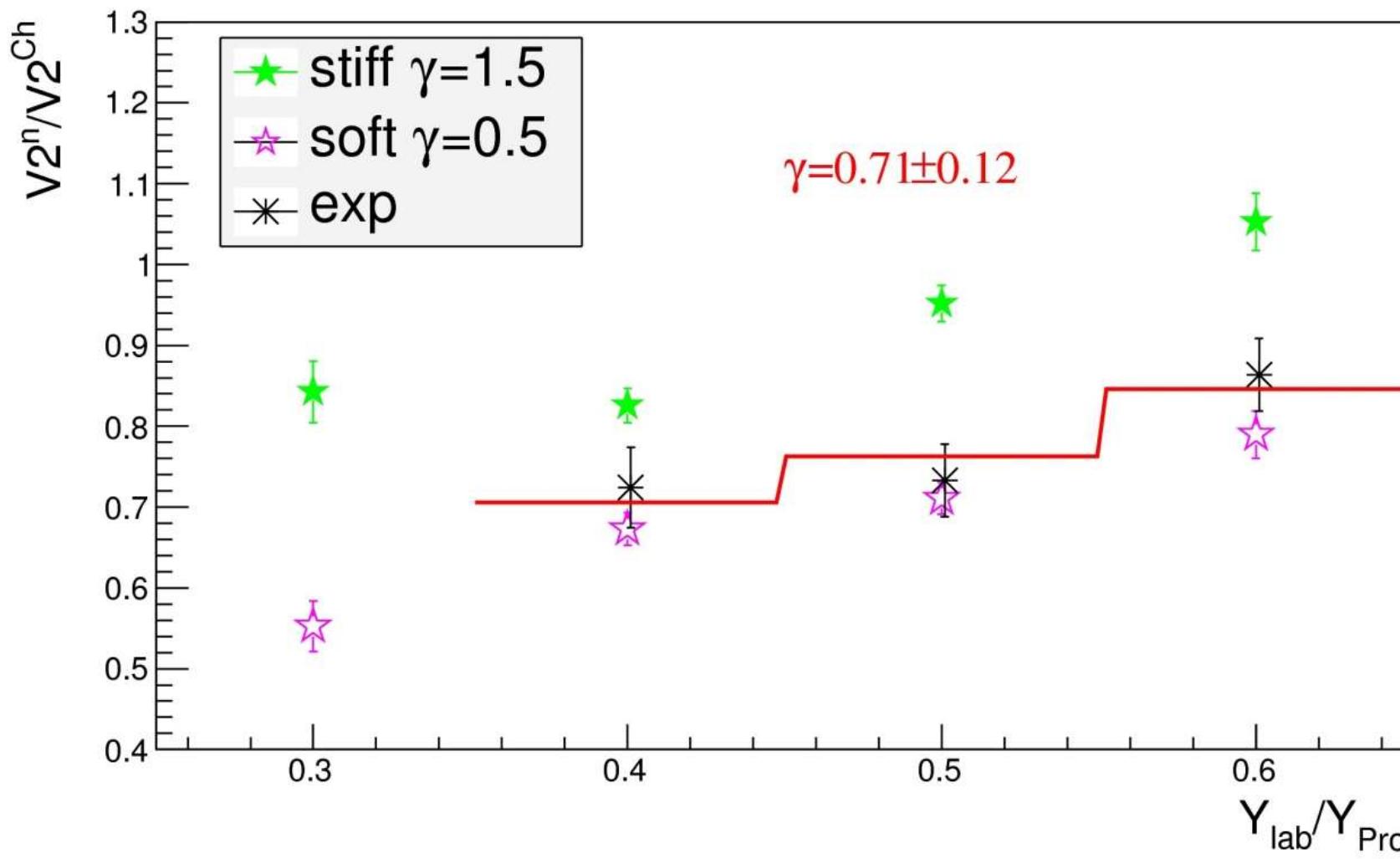
Only charged particles!!!

preliminary

γ extrapolation from rapidity dependence

Au+Au @ 400 AMeV $b < 7.5$ fm

Wbg= 110 %

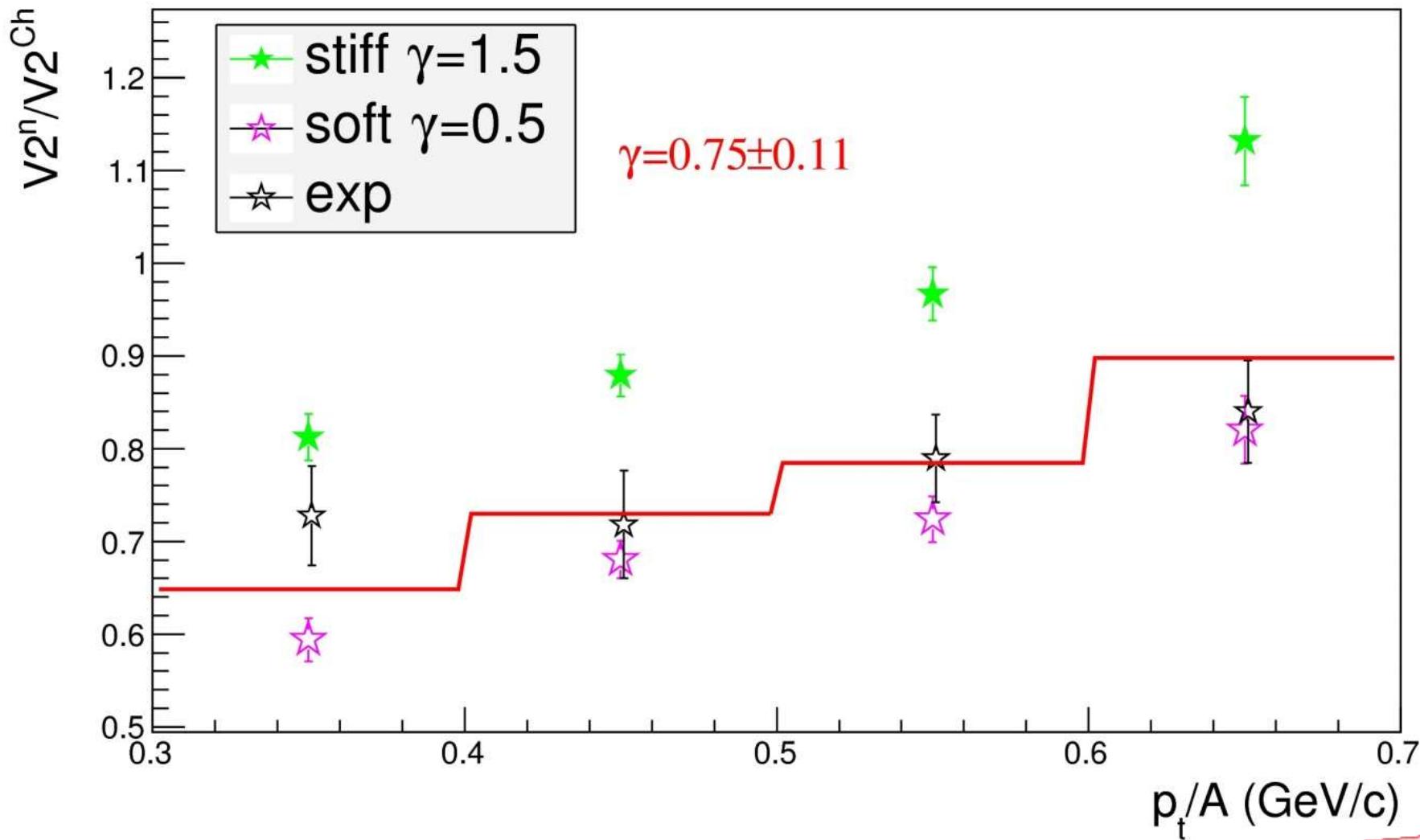


preliminary

γ extrapolation from transverse momentum dependence

Au+Au @ 400 AMeV $b < 7.5$ fm

Wbg= 110 %

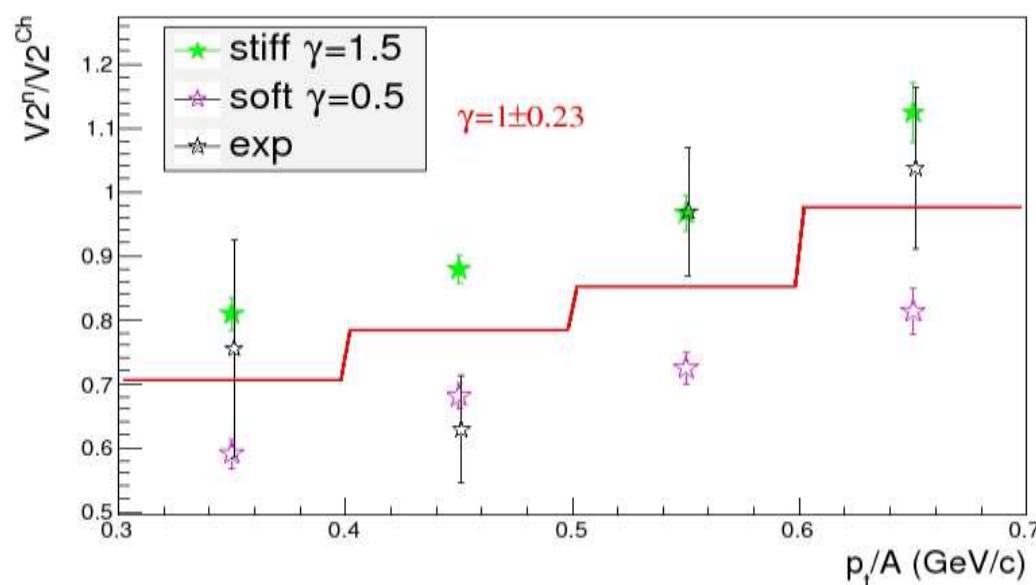
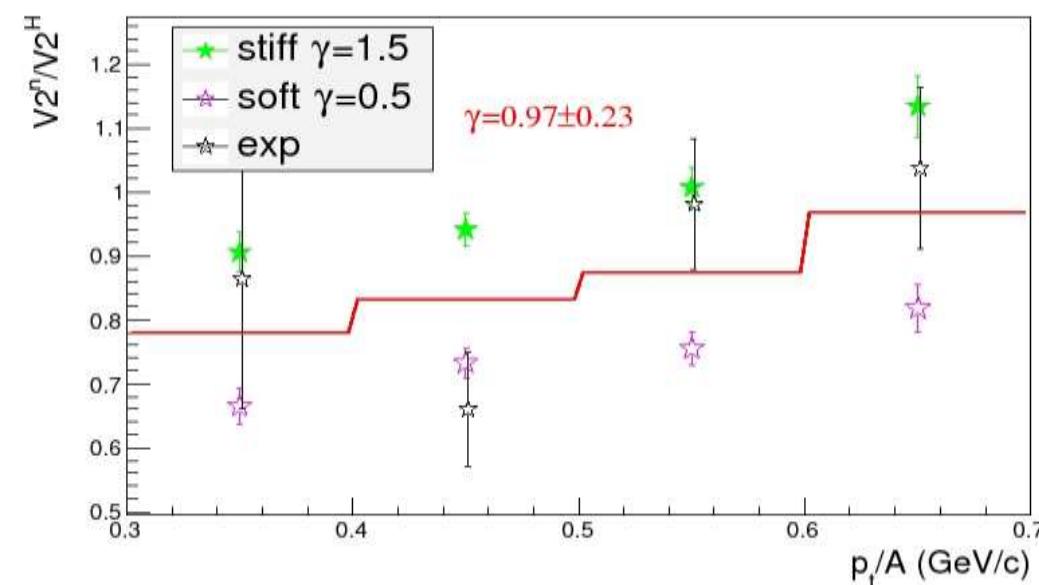
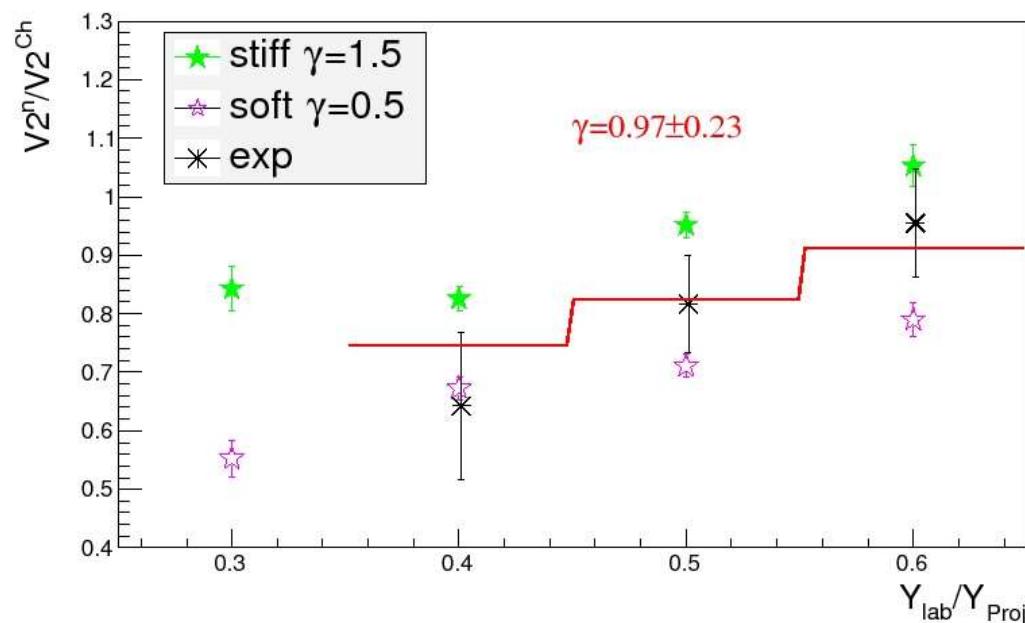
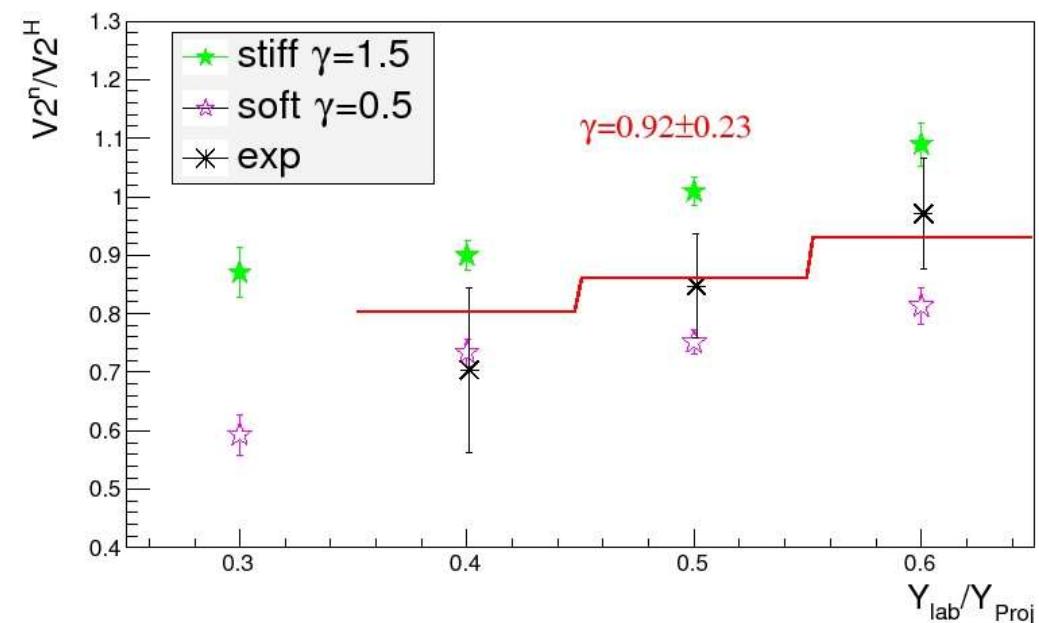


preliminary

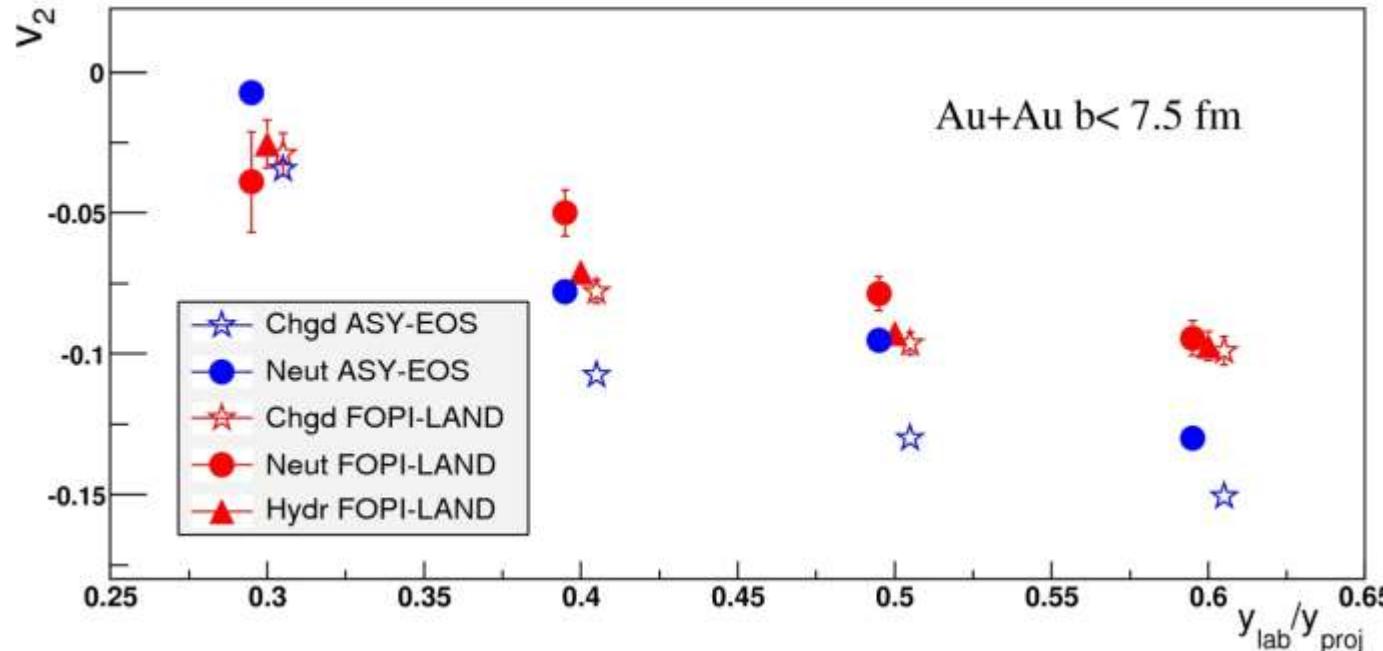
γ extrapolation : $v2n/v2ch$ vs $v2n/v2H$ (FOPI-LAND data)

Au+Au @ 400 AMeV
b<7.5 fm

$0.25 < y_{\text{lab}}/y_{\text{proj}} < 0.75$
 θ_{lab} cut as in ASY-EOS set-up



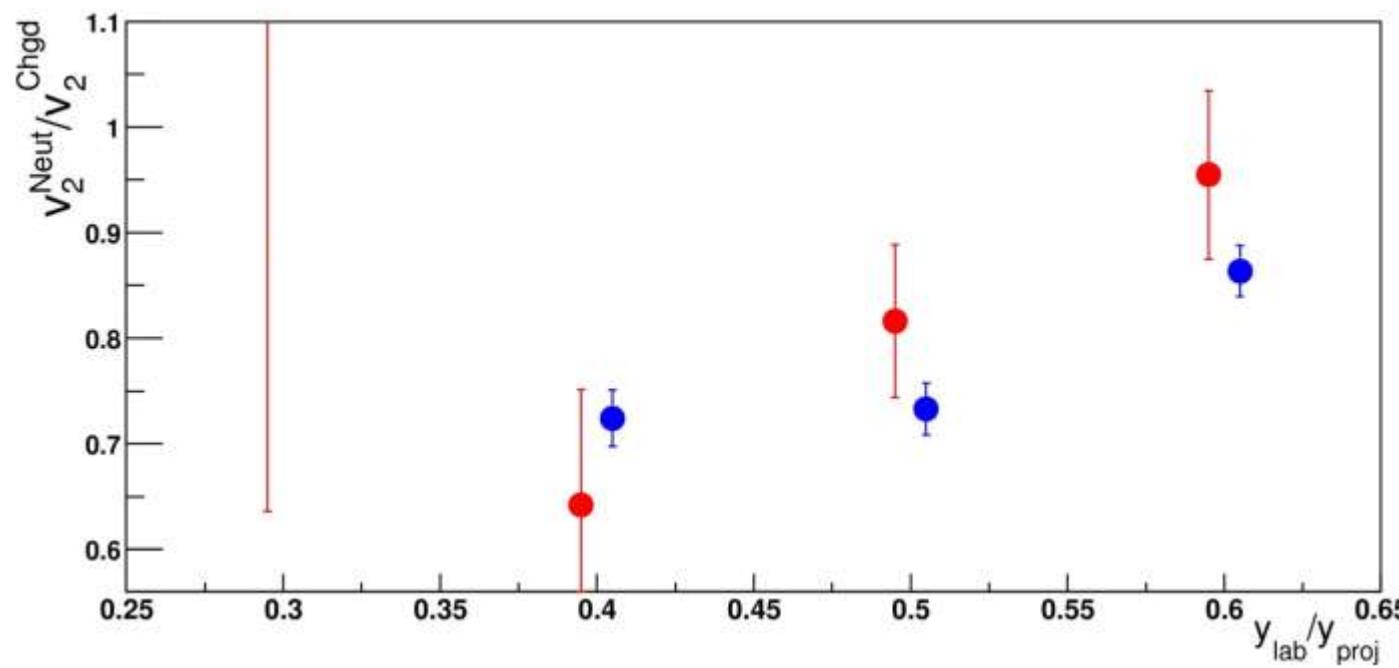
Comparing ASY-EOS with FOPI-LAND exp: rapidity dependence



Au+Au @ 400 AMeV
 $b < 7.5 \text{ fm}$
 $0.25 < y_{\text{lab}}/y_{\text{proj}} < 0.75$

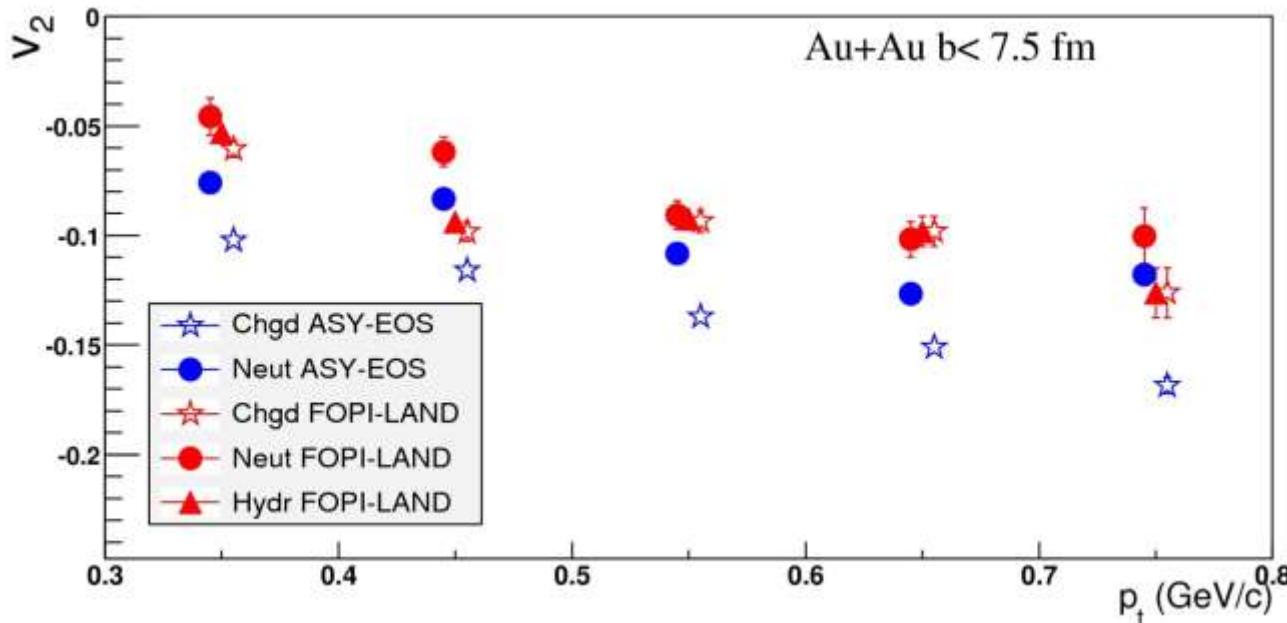
θ_{lab} cut as in ASY-EOS set-up

Wbg= 110 %



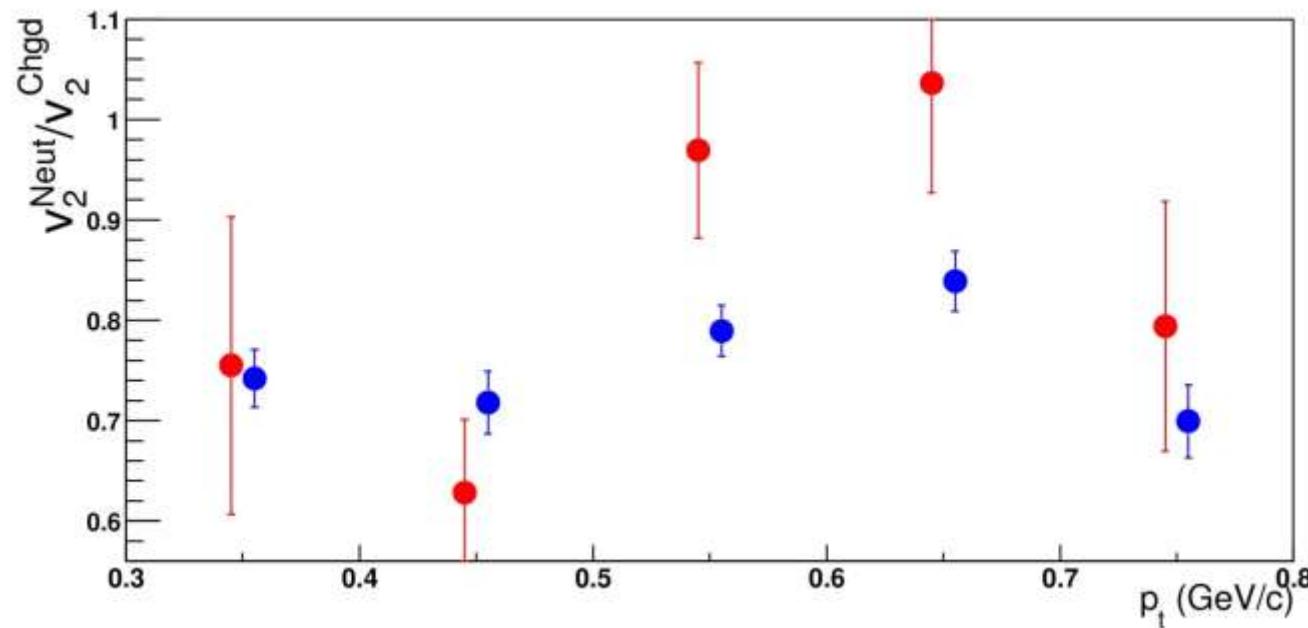
preliminary

Comparing ASY-EOS with FOPI-LAND exp: transverse momentum dep.



Au+Au @ 400 AMeV
b < 7.5 fm
 $0.25 < y_{lab}/y_{proj} < 0.75$

θ_{lab} cut as in ASY-EOS set-up



Wbg= 110 %

preliminary

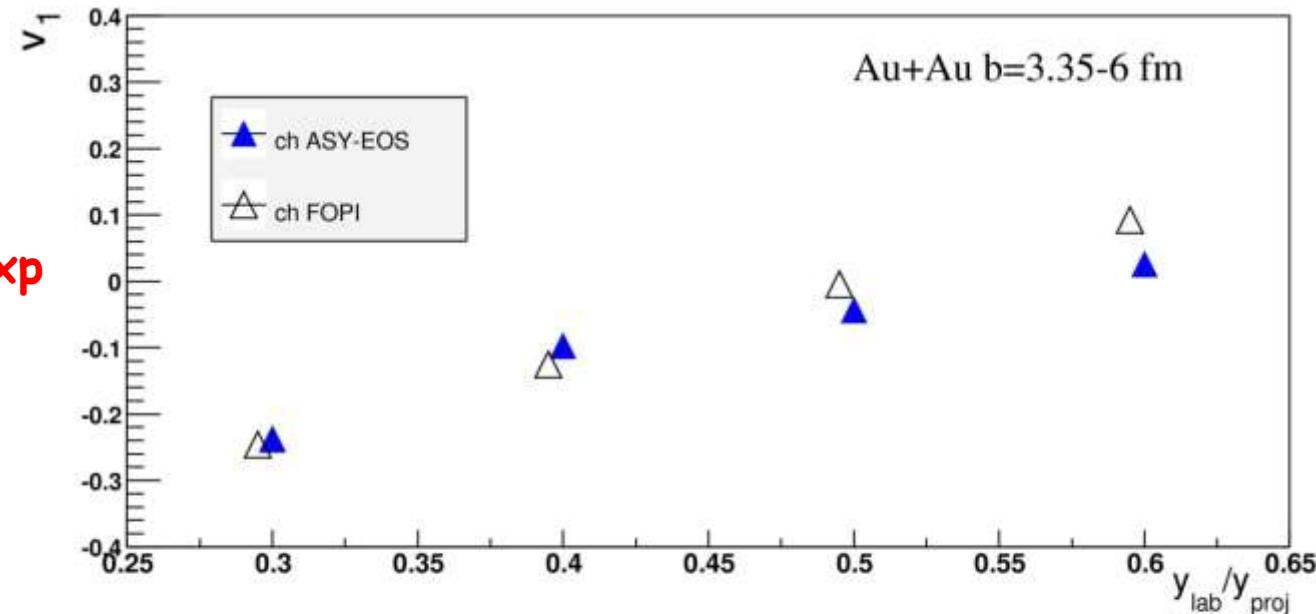
Comparing ASY-EOS with FOPI: rapidity dep. of charged particles

Au+Au @ 400 A.MeV

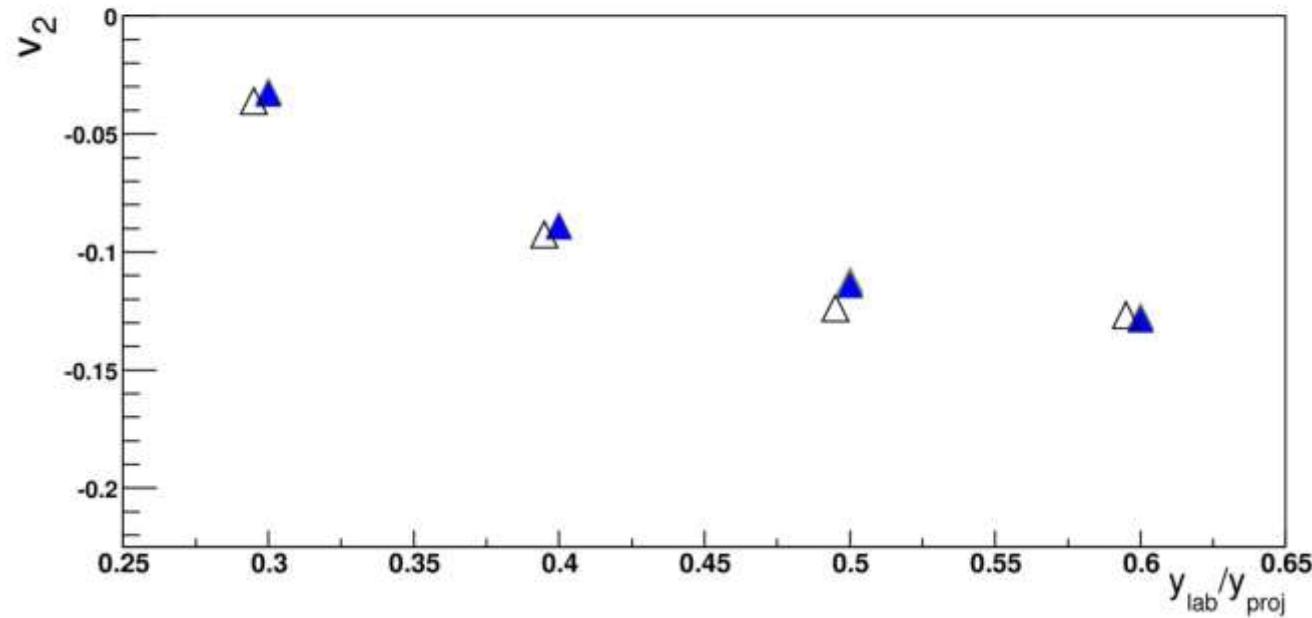
$3.35 < b < 6 \text{ fm}$ (c2)

θ_{lab} cut as LAND in ASY-EOS exp

Courtesy of W. Reisdorf



Good agreement between v_2 of charged particles of FOPI and ASY-EOS exps ...good benchmark for us



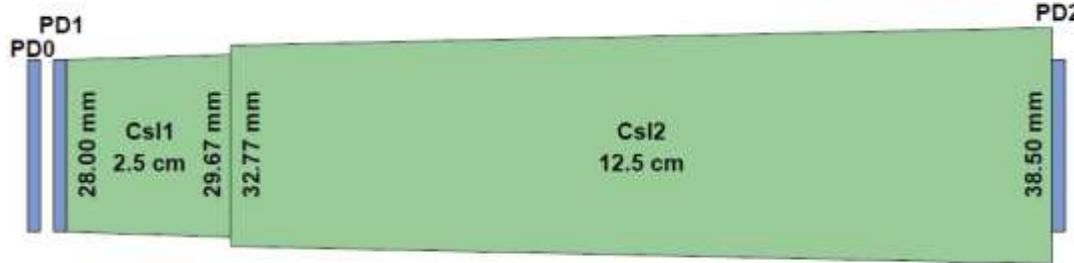
preliminary

Comparing KraTTA* with FOPI: rapidity dep. of isotopes

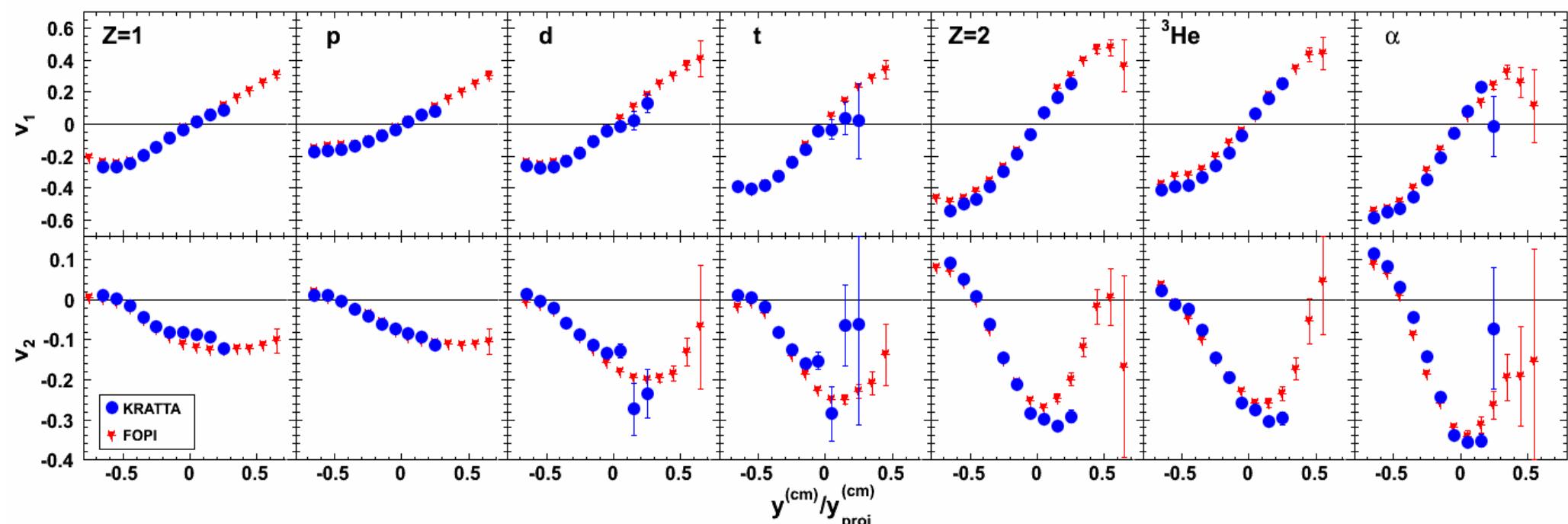
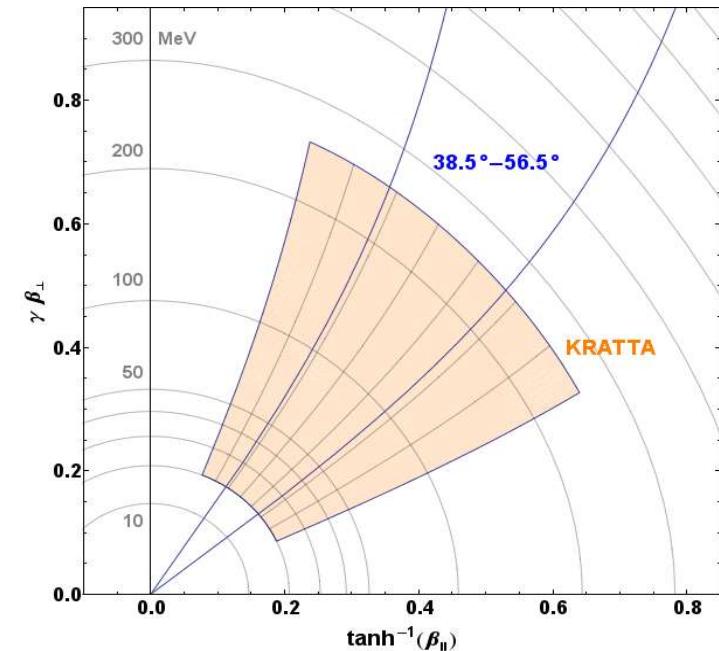
35 modules (5×7), $20.7^\circ < \theta < 63.5^\circ$

40 cm from target.

Digitized with 100 MHz, 14 bits Flash ADCs



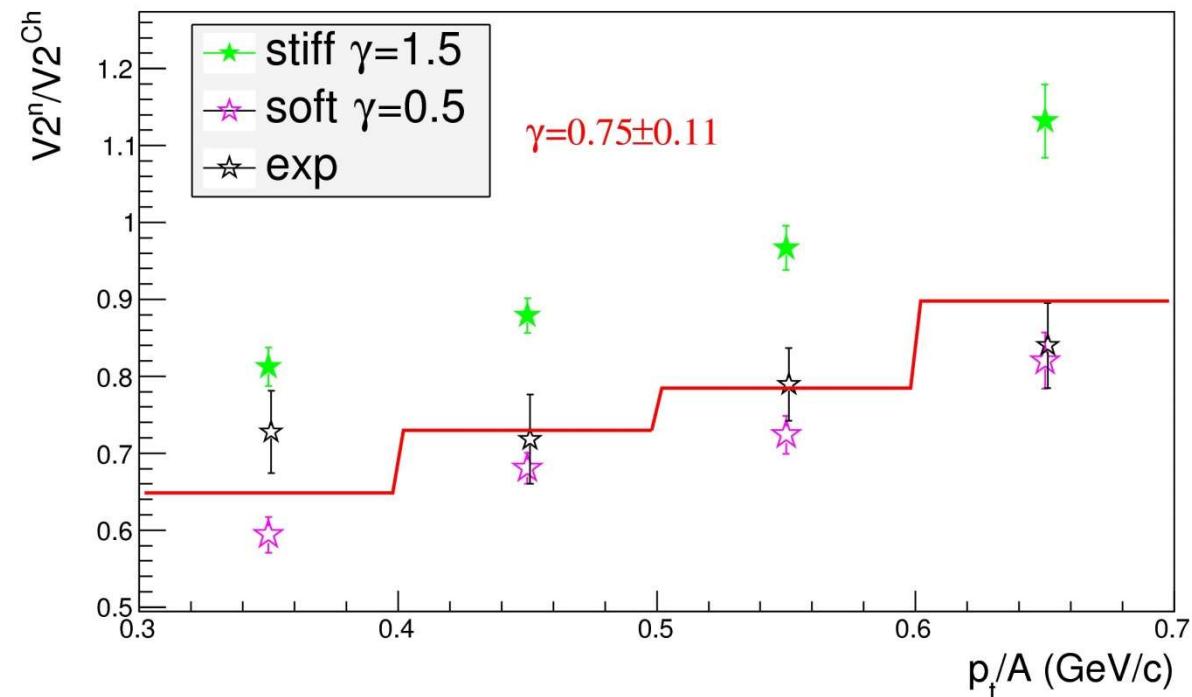
Au+Au @ 400 A.MeV
 $3.35 < b < 6$ fm (c2)
 Θ_{lab} cut as LAND



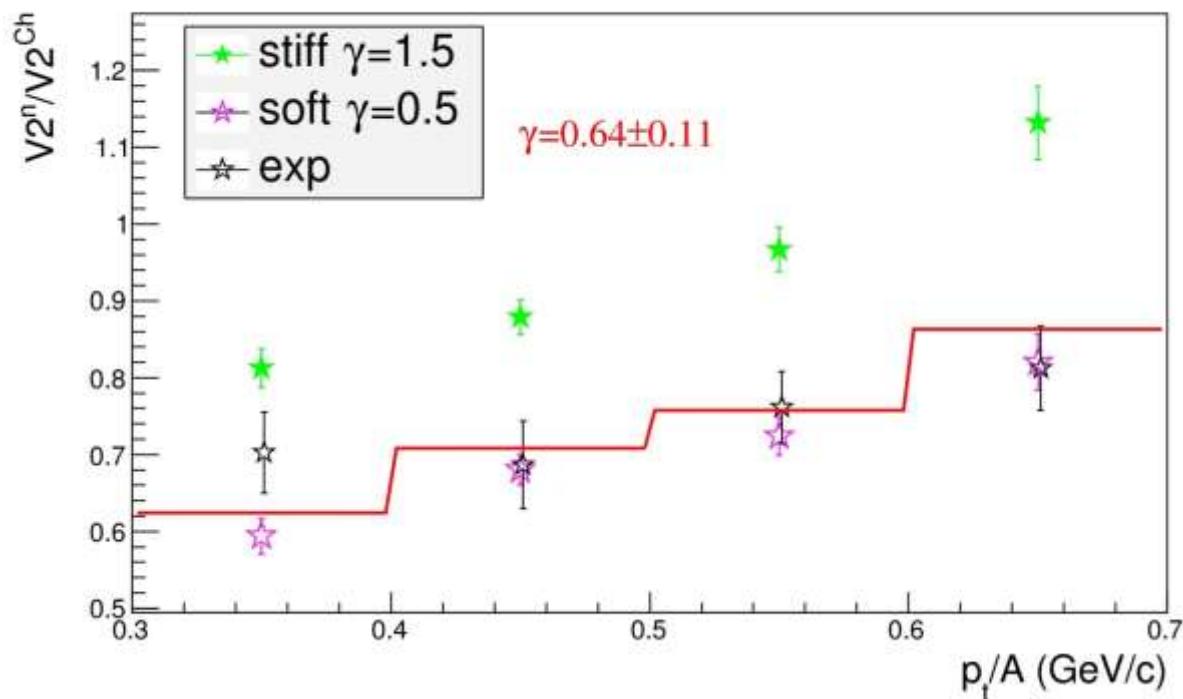
γ extrapolation from transverse momentum dependence : Wbg influence

Au+Au @ 400 AMeV
 $b < 7.5$ fm

Wbg= 110 %



Wbg= 100 %



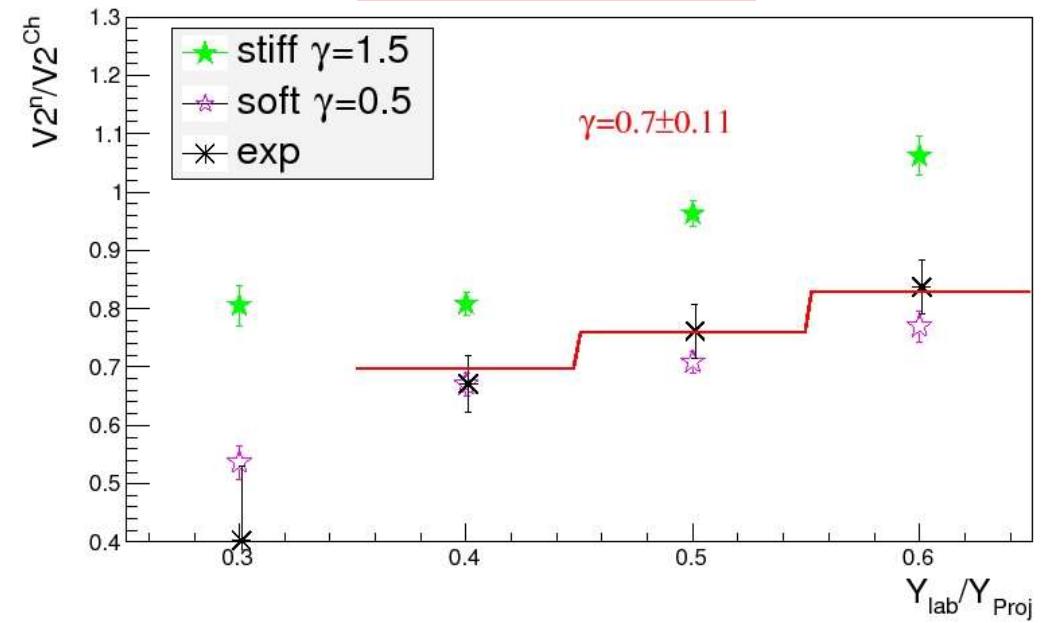
preliminary

γ extrapolation: impact parameter selection influence

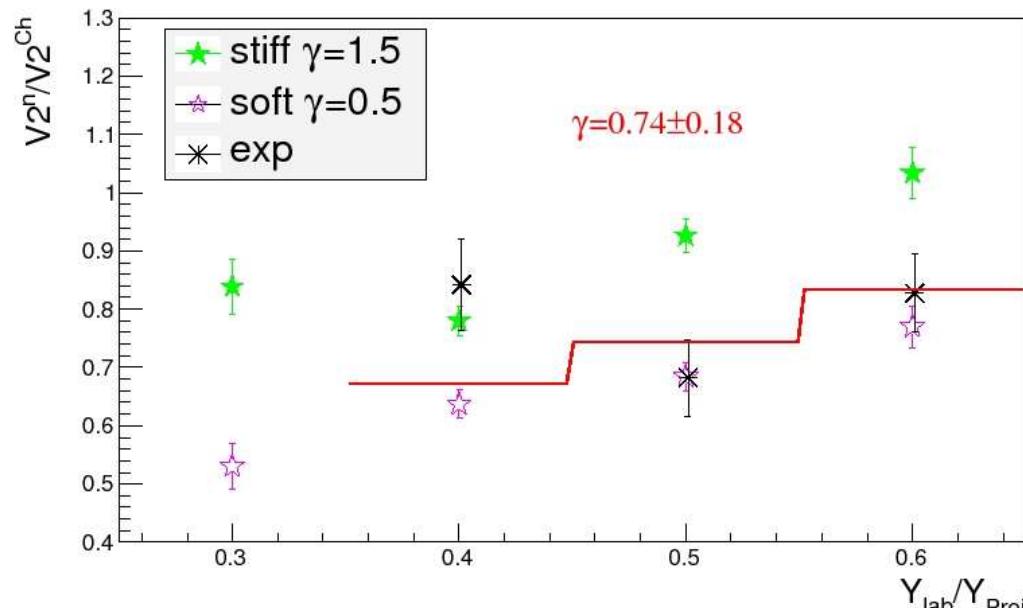
Au+Au $3 < b < 7.5$ fm

Au+Au $3.35 < b < 6$ fm

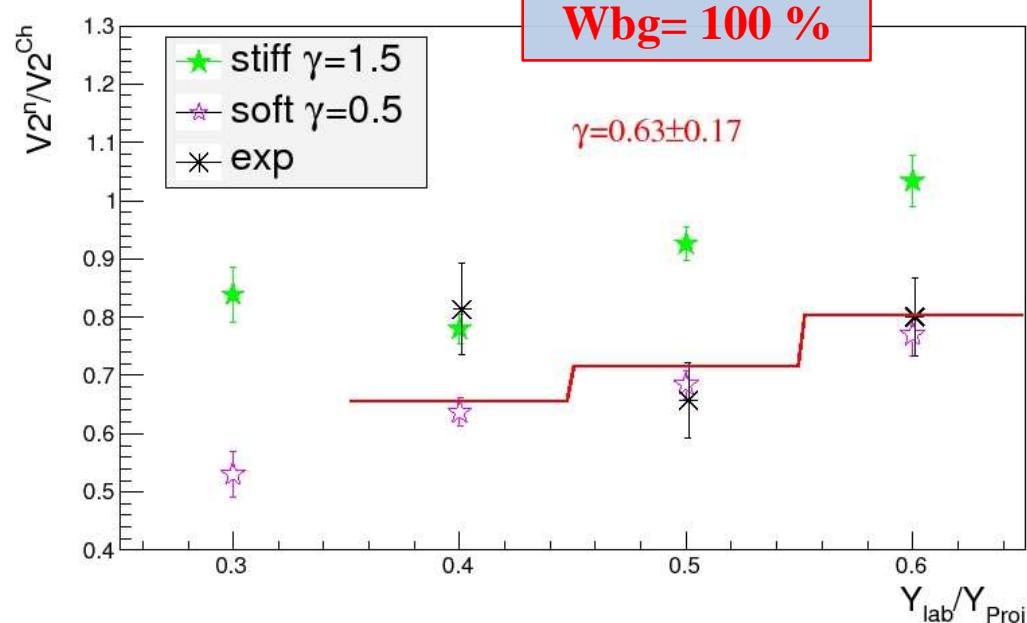
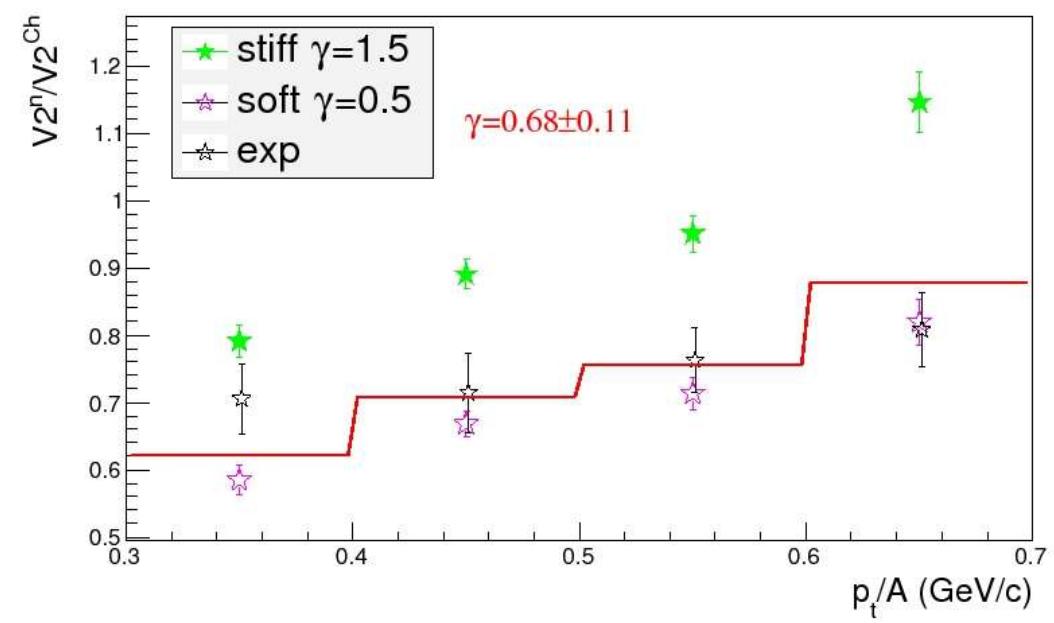
Wbg= 110 %



Wbg= 110 %



Wbg= 100 %

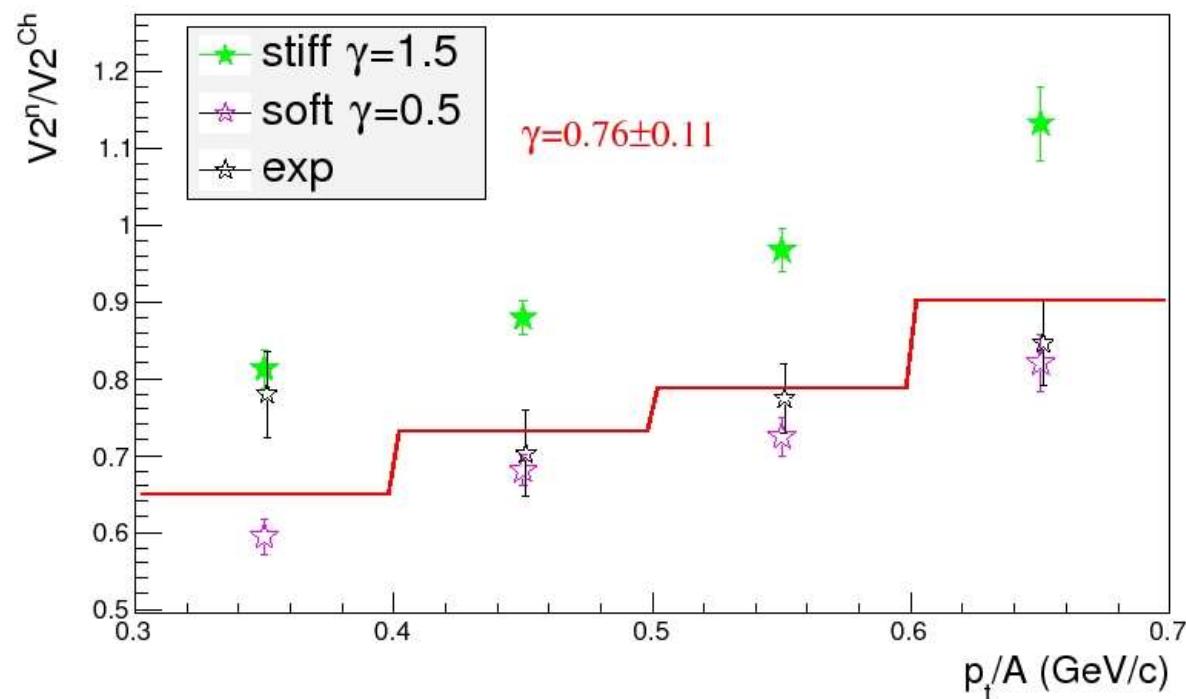
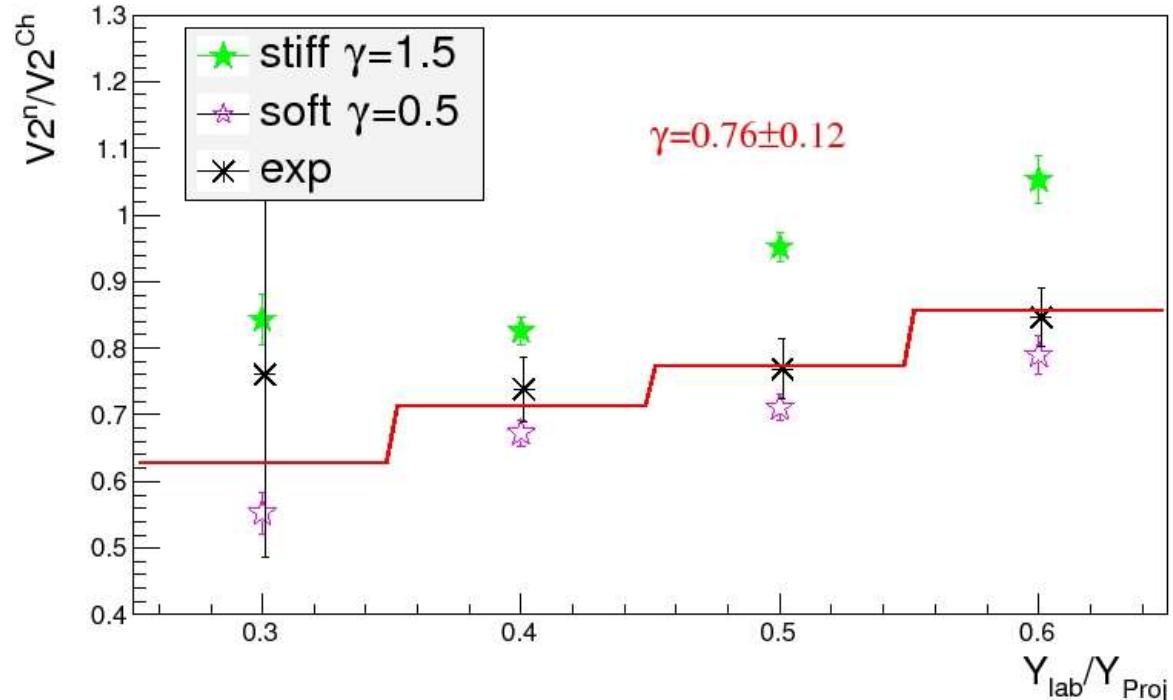


Last result: improved background evaluation

Au+Au @ 400 AMeV
 $b < 7.5$ fm

Neutrons:
 $(\text{Au+Au}) - (\text{Au+Au with SB}) +$
 $- (\text{Au+EF}) + (\text{Au+EF with SB})$

Charged Particles:
 $(\text{Au+Au}) - (\text{Au+EF})$



The analysis is in progress...

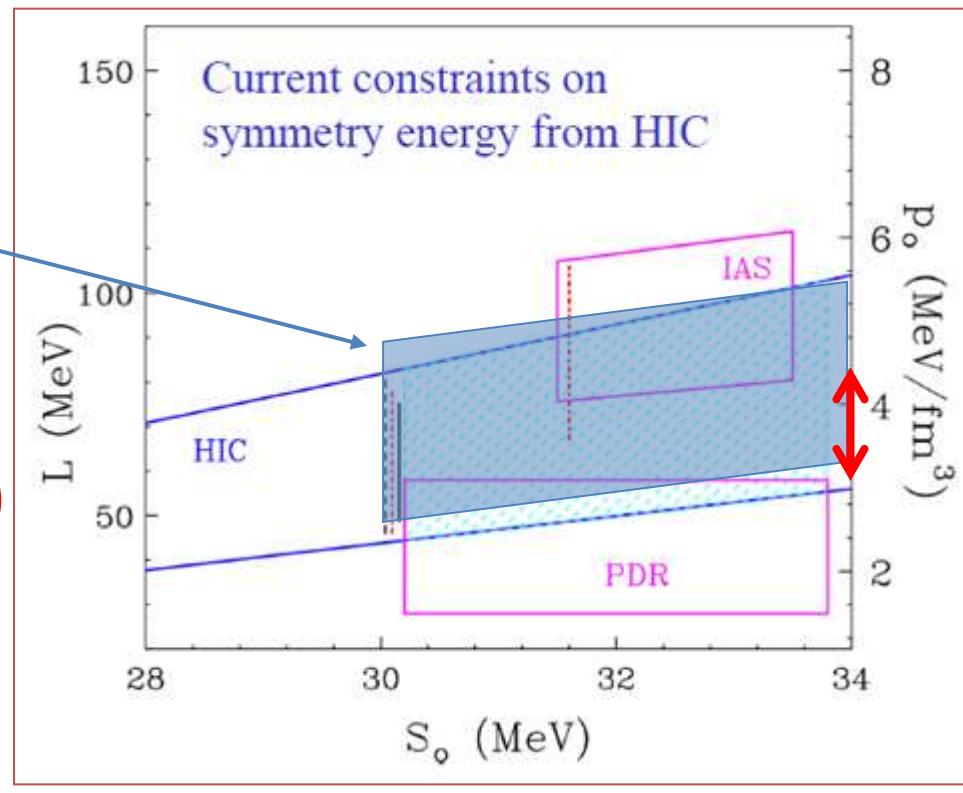
$$S(\rho) = S_0 + \frac{L}{3} \left(\frac{\rho - \rho_o}{\rho_o} \right) + \frac{K_{\text{sym}}}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots,$$

$$L = 3\rho_0 \cdot \partial E_{\text{sym}} / \partial \rho$$

From “old”
elliptic
n/p flow*

$$\gamma = 0.7 \pm 0.2$$

$$L = 70 \pm 13 \text{ (FP1)}$$



from M.B. Tsang et al., PRL 102, 122701 (2009)
 vertical lines: analyses with ImQMD (Zhang et al.)
 and IBUU04 (Li and Chen)

IAS

isobaric analog states
 Danielewicz/Lee 2008

HIC

heavy-ion collisions
 isospin diffusion, n/p ratios
 Tsang et al., 2009

PDR

pygmy dipole resonance
 Klimkiewicz et al. 2007

see also “Complete Electric Dipole Response in ^{208}Pb ”
 Tamii et al.,
 PRL 107, 062502 (2011)

symmetry
 pressure
 $P_0 = (L/3)\rho_0$

$$S_0 = E_{\text{sym}}(\rho_0)$$

NeuLAND @ FAIR/GSI

TDR finalized in Oct 2011 and submitted
total volume $2.5 \times 2.5 \times 3 \text{ m}^3$

each bar readout by two PMT

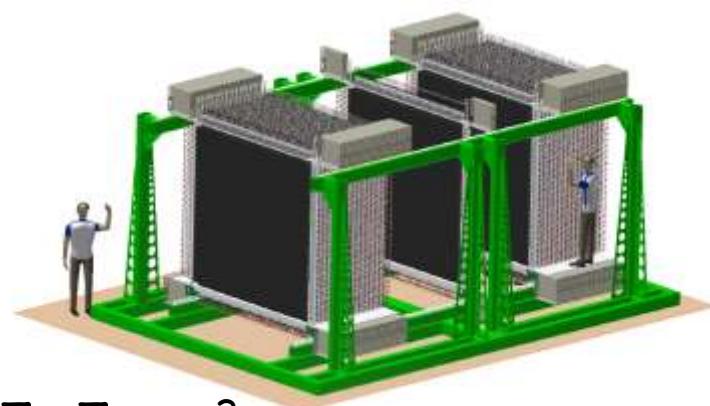
3000 modules (plastic scintillator bars) $250 \times 5 \times 5 \text{ cm}^3$

30 double planes with 100 bars each, bars in neighboring planes
mutually perpendicular

$\sigma_t \leq 150 \text{ ps}$ and $\sigma_{x,y,z} \leq 1.5 \text{ cm}$

one-neutron efficiency $\sim 95\%$ for energies 200-1000 MeV

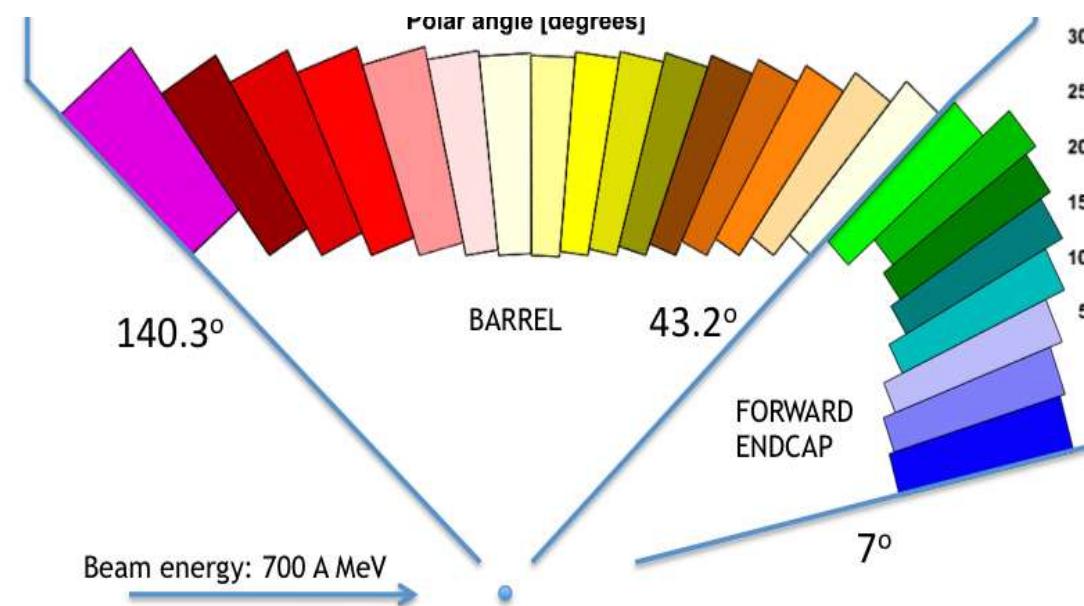
multi-neutron detection capability



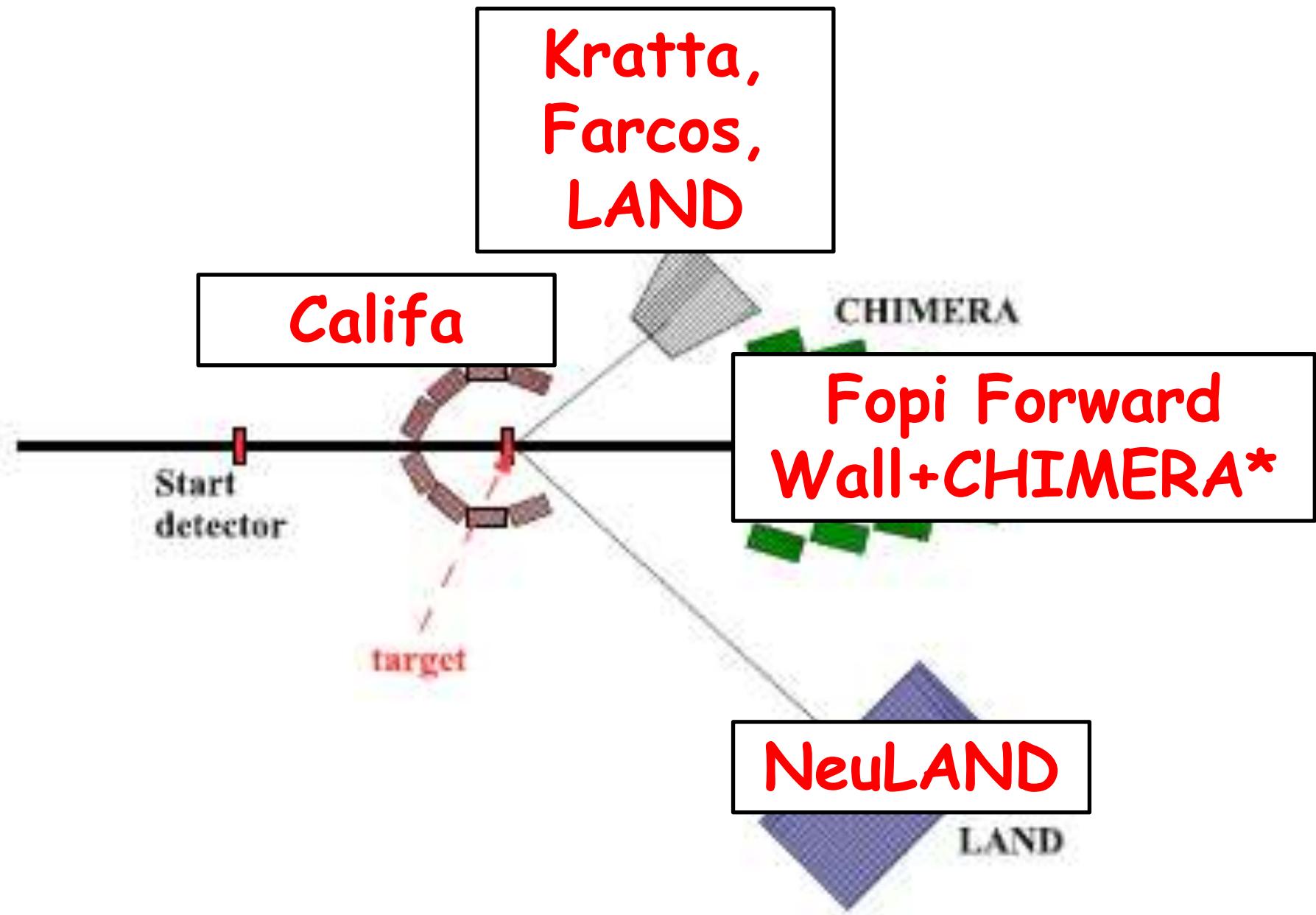
I. Gasparic AsyEOS2012 workshop,
6.9.2012, Siracusa, Italy

Califa
CALorimeter for the In Flight
detection of γ
rays and light charged pArticles

CsI(Tl) read by APD
with digital read-out

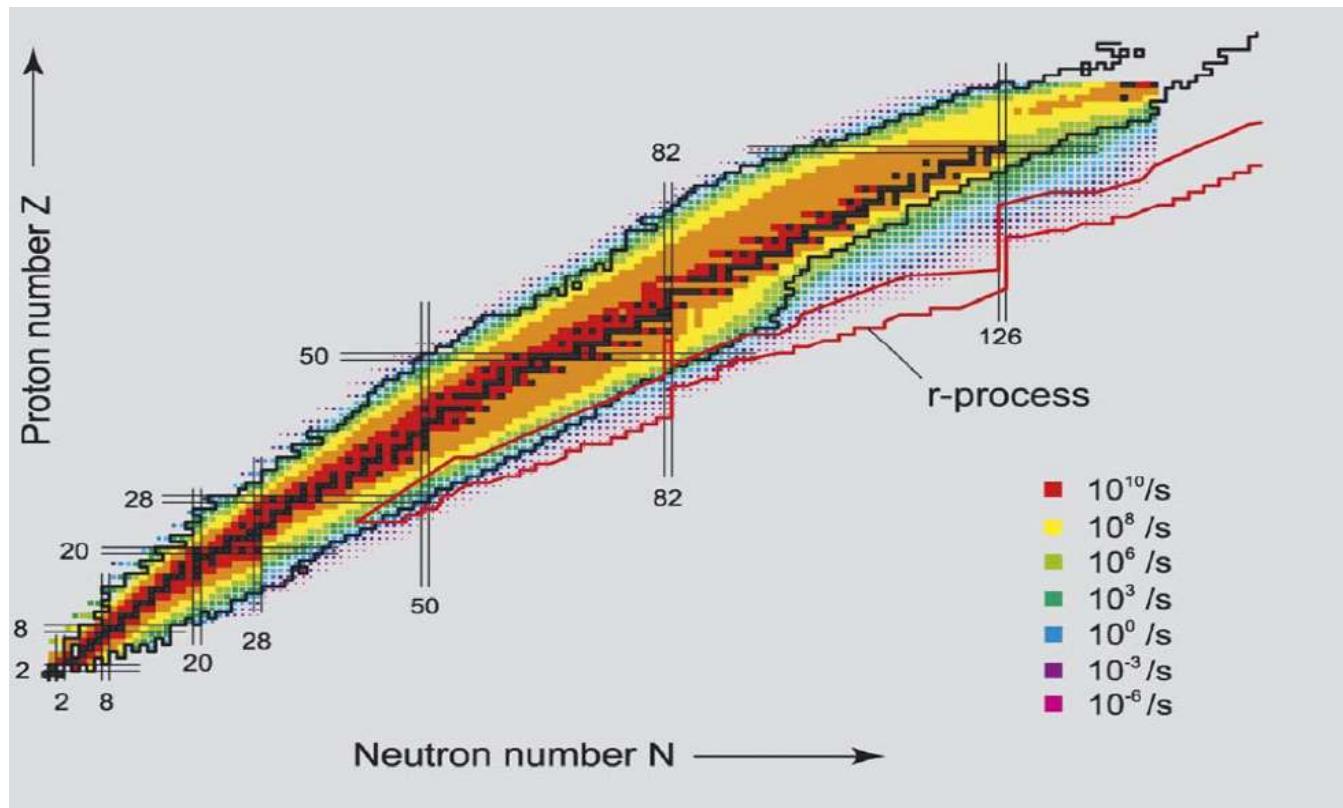


ASY-EOS future set-up?



* Ring 1-2-3
 $(\theta < 7^\circ)$

FAIR rates



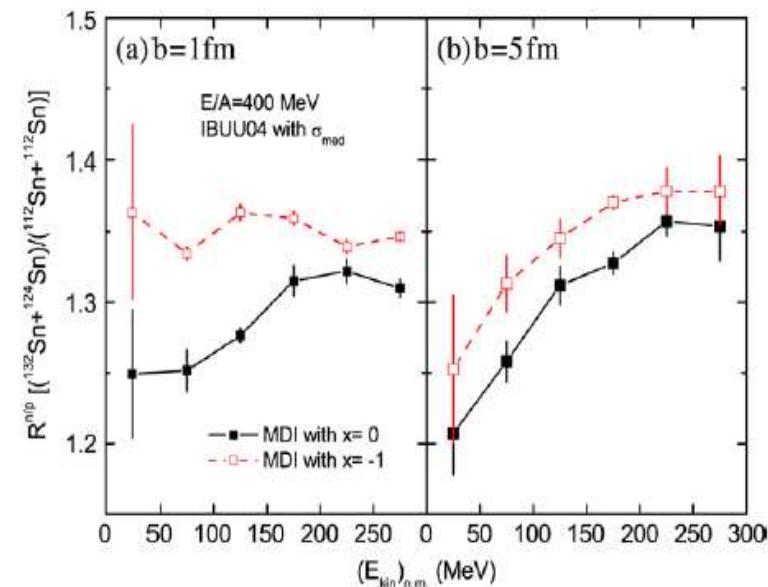
Some interesting beams (and I^2)

$^{197}\text{Au}+^{197}\text{Au}$ @ 600, 800, 1000 AMeV (0.039+0.039)

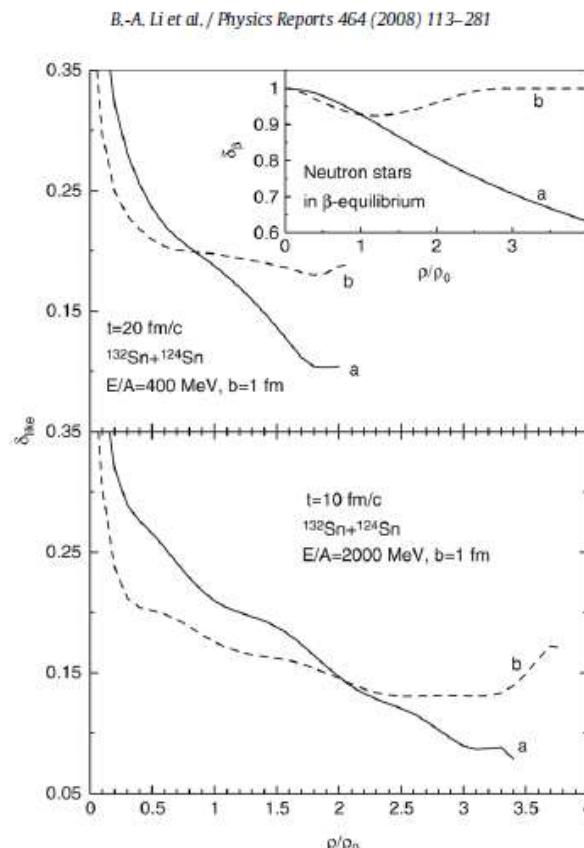
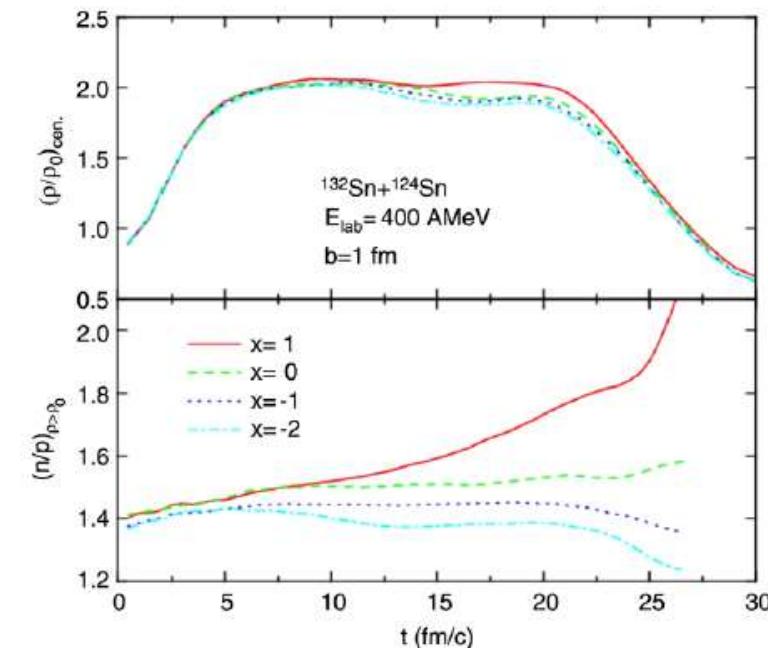
$^{132}\text{Sn}+^{124}\text{Sn}$ @ 400, 800, 1000 AMeV (0.059+0.037)

$^{106}\text{Sn}+^{112}\text{Sn}$ @ 400, 800, 1000 AMeV (0.003+0.011)

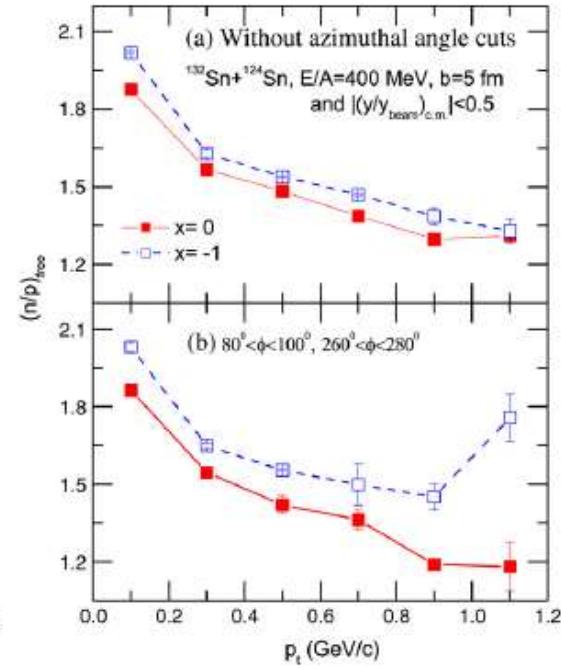
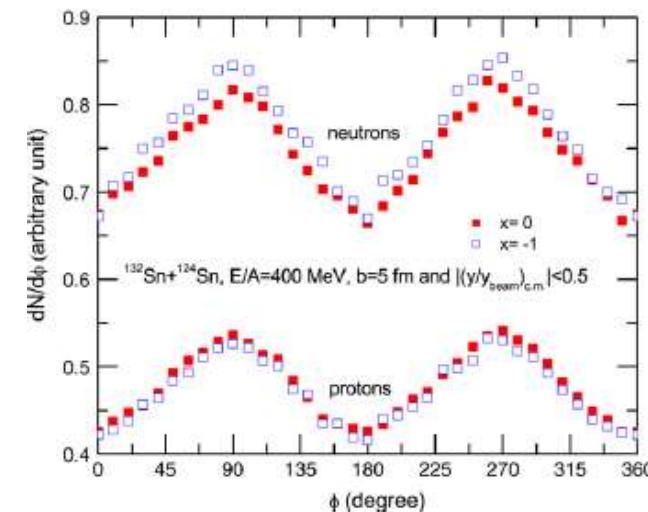
Why ^{132}Sn ?



ratio of free nucleons taken from the reactions of $^{132}\text{Sn} + ^{124}\text{Sn}$ and $^{112}\text{Sn} + ^{112}\text{Sn}$ (right panel). Taken from Ref. [67].



1.5 A GeV ($b = 6 \text{ fm}$) from the three different models f ill circles and solid line: $NL\rho\delta$. Open circles and dash ext and the previous caption.



Conclusions

Symmetry Energy:

- Low densities: several constraints quite consistent
- High density:
 - pion constraints not consistent
 - n/p flows suggests...a route "Towards a model-independent constraint of the high-density dependence of the symmetry energy"
 - **Finalizing ASY-EOS data analysis is in progress**
- Work on code consistency needed...everywhere
- New and better experiments on n,p flows and ratio, pions and kaons, also with high asymmetric beams (e.g. ^{132}Sn) and new detectors (Riken TPC, NeuLand@R3B)
- International collaborations and efforts

On the road.....



The Asy-Eos Collaboration

P. Russotto^{1,a}, M. Chartier², M.D. Cozma³, E. De Filippo¹, A. Le Fèvre⁴, S. Gannon², I. Gašparić^{5,6}, M. Kiš^{4,5}, S. Kupny⁷, Y. Leifels⁴, R.C. Lemmon⁸, Q. Li⁹, J. Łukasik¹⁰, P. Marini^{11,12}, P. Pawłowski¹⁰, S. Santoro^{13,14}, W. Trautmann⁴, M. Veselsky¹⁵, L. Acosta¹⁶, M. Adamczyk⁷, A. Al-Ajlan¹⁷, M. Al-Garawi¹⁸, S. Al-Homaidhi¹⁷, F. Amorini¹⁶, L. Auditore^{13,14}, T. Aumann⁶, Y. Ayyad¹⁹, V. Baran^{16,20}, Z. Basrak⁵, R. Bassini²¹, J. Benlliure¹⁹, C. Boiano²¹, M. Boisjoli¹², K. Boretzky⁴, J. Brzychczyk⁷, A. Budzanowski¹⁰, G. Cardella¹, P. Cammarata¹¹, Z. Chajecki²², A. Chbihi¹², M. Colonna¹⁶, B. Czech¹⁰, M. Di Toro^{16,23}, M. Famiano²⁴, V. Greco^{16,23}, L. Grassi⁵, C. Guazzoni^{21,25}, P. Guazzoni^{21,26}, M. Heil⁴, L. Heilborn¹¹, R. Introzzi²⁷, T. Isobe²⁸, K. Kezzar¹⁸, A. Krasznahorkay²⁹, N. Kurz⁴, E. La Guidara¹, G. Lanzalone^{16,30}, P. Lasko⁷, I. Lombardo^{31,32}, W.G. Lynch²², Z. Matthews³, L. May¹¹, T. Minniti^{13,14}, M. Mostazo¹⁹, A. Pagano¹, M. Papa¹, S. Pirrone¹, R. Pleskac⁴, G. Politi^{1,23}, F. Porto^{16,23}, R. Reifarth⁴, W. Reisdorf⁴, F. Riccio^{21,25}, F. Rizzo^{16,23}, E. Rosato^{31,32}, D. Rossi^{4,22}, H. Simon⁴, I. Skwirczynska¹⁰, Z. Sosin⁷, L. Stuhl²⁹, A. Trifirò^{13,14}, M. Trimarchi^{13,14}, M.B. Tsang²², G. Verde¹, M. Vigilante^{31,32}, A. Wieloch⁷, P. Wigg², H.H. Wolter³³, P. Wu², S. Yennello¹¹, P. Zambon^{21,25}, L. Zetta^{21,26}, and M. Zoric⁵

¹INFN-Sezione di Catania, Catania, Italy

²University of Liverpool, Liverpool, UK

³IFIN-HH, Magurele-Bucharest, Romania

⁴GSI Helmholtzzentrum, Darmstadt, Germany

⁵Ruder Bošković Institute, Zagreb, Croatia

⁶Technische Universität, Darmstadt, Germany

⁷Jagiellonian University, Kraków, Poland

⁸STFC Laboratory, Daresbury, UK

⁹Huzhou Teachers College, China

¹⁰IFJ-PAN, Krakow, Poland

¹¹Texas A&M University, College Station, USA

¹²GANIL, Caen, France

¹³INFN-Gruppo Collegato di Messina, Messina, Italy

¹⁴Università di Messina, Messina, Italy

¹⁵Institute of Physics, Slovak Academy of Sciences, Bratislava, Slovakia

¹⁶INFN-Laboratori Nazionali del Sud, Catania, Italy

¹⁷KACST Riyadh, Riyadh, Saudi Arabia

¹⁸King Saud University, Riyadh, Saudi Arabia

¹⁹University of Santiago de Compostela, Santiago de Compostela, Spain

²⁰University of Bucharest, Bucharest, Romania

²¹INFN-Sezione di Milano, Milano, Italy

²²NSCL Michigan State University, East Lansing, USA

²³Università di Catania, Catania, Italy

²⁴Western Michigan University, USA

²⁵Politecnico di Milano, Milano, Italy

²⁶Università degli Studi di Milano, Milano, Italy

²⁷INFN, Politecnico di Torino, Torino, Italy

²⁸RIKEN, Wako, Japan

²⁹Institute of Nuclear Research, Debrecen, Hungary

³⁰Università Kore, Enna, Italy

³¹INFN-Sezione di Napoli, Napoli, Italy

³²Università di Napoli, Napoli, Italy

³³LMU, München, Germany

THANKS!