



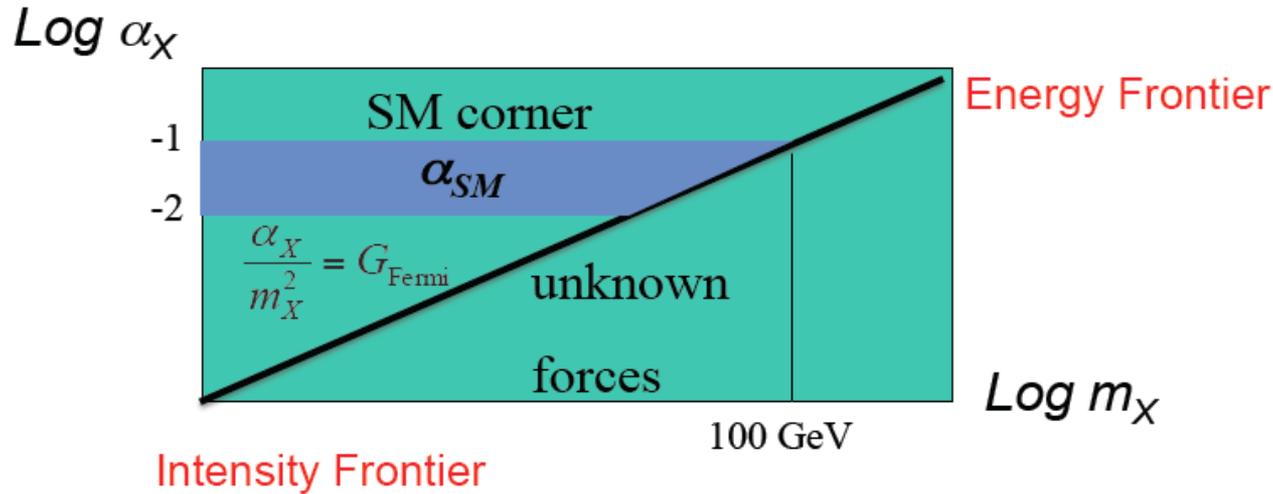
Search for New Physics at Medium Energy

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INFN – Genova

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3° Incontro Nazionale di Fisica Nucleare
Frascati 14-16 Novembre 2016



How to Search for New Physics



LHC regime: $m_X \sim 1 \text{ TeV}$, $\alpha_X \sim \alpha_{SM}$

First results show no hints of new strongly-interacting states or new heavy EW bosons (other than Higgs)

What about if: $m_X \sim 1 \text{ GeV}$, $\alpha_X < 10^{-6}$?

Search for new physics at medium energy

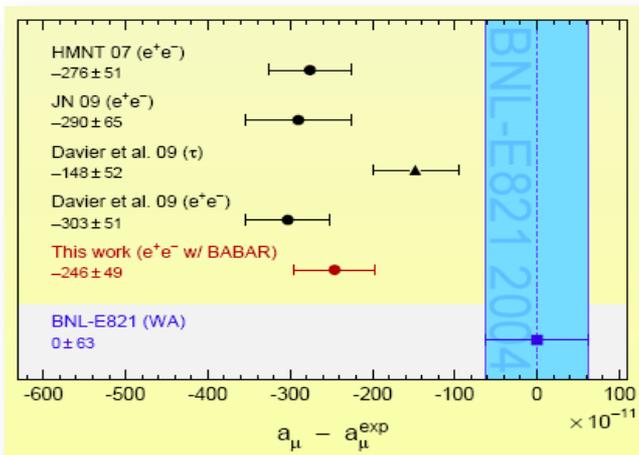
Small effects...high precision and high intensity

Precision
experiments at low
or medium energy!

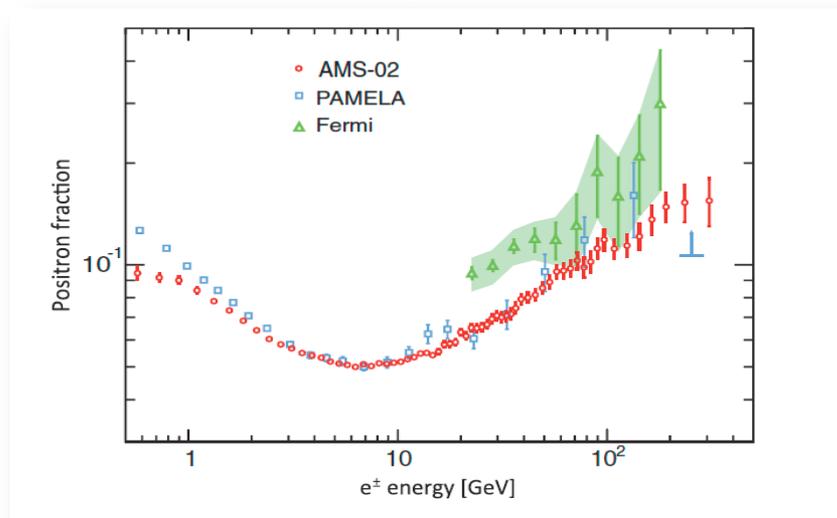
Hints of New Physics...

Indications for (potentially) new physics from anomalies...

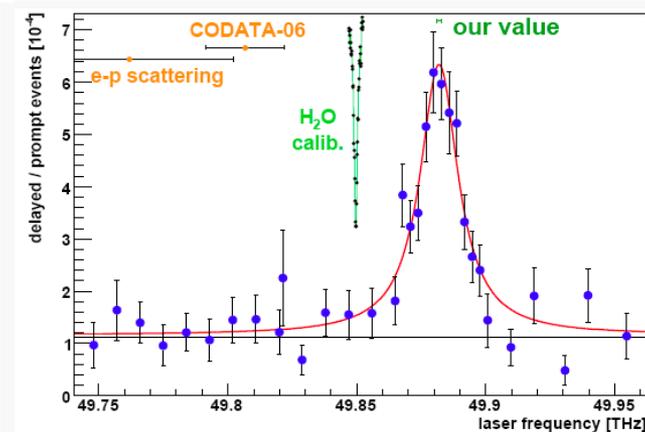
g-2 of muon



Positron and antiproton abundance



muonic hydrogen Lamb shift



$$r_p = 0.84184(67) \text{ fm} \quad u_r^{\text{th}} = 8 \times 10^{-4}$$

CODATA 2006: $r_p = (0.8768 \pm 0.0069) \text{ fm}$, from H
 e-p scattering: $r_p = (0.895 \pm 0.018) \text{ fm}$ (2%)

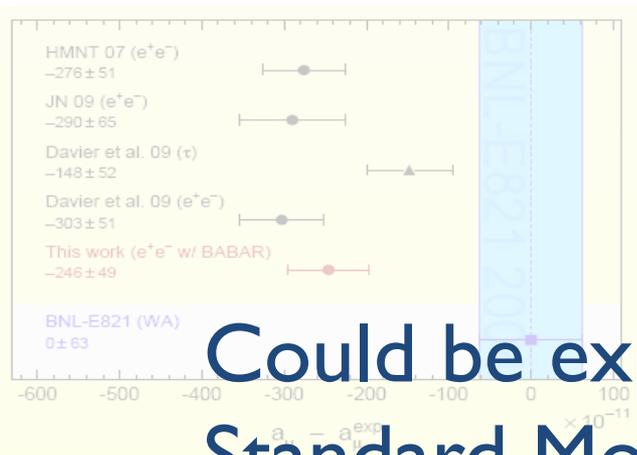
- ★ Muon 200 times closer to p (w.r.t. hydrogen)
- ★ New forces for muon?



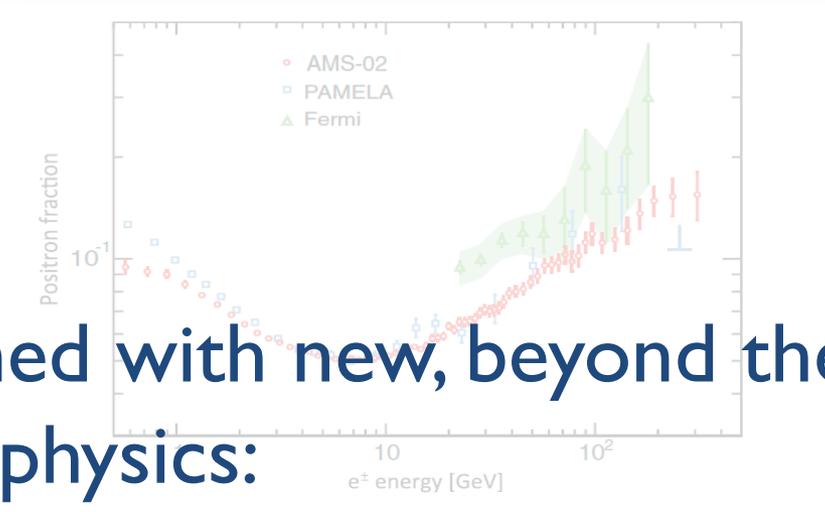
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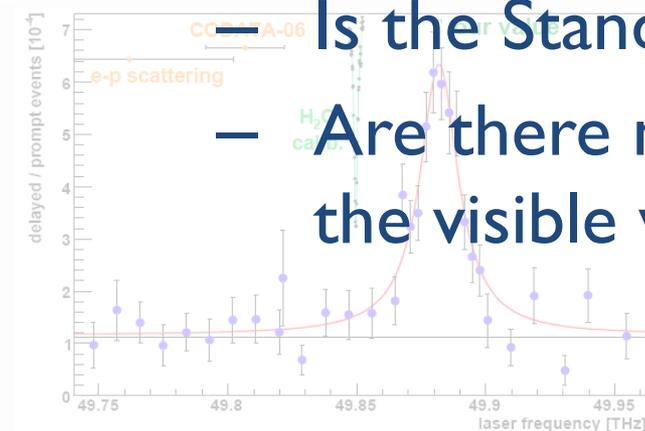


Positron and antiproton abundance



Could be explained with new, beyond the Standard Model, physics:

muonic hydrogen Lamb shift



- Is the Standard Model correct?
- Are there new forces and particles that affect the visible world?

$$r_p = (0.81 \pm 0.11) \text{ fm} \quad \mu\text{-p} \quad 10^{-4}$$

CODATA 2006: $r_p = (0.8768 \pm 0.0069) \text{ fm}$ from H
 e-p scattering: $r_p = (0.895 \pm 0.18) \text{ fm}$ (2%)

- ★ Muon 200 times closer to p (w.r.t. hydrogen)
- ★ New forces for muon?



Electro-Weak Couplings and SM Tests

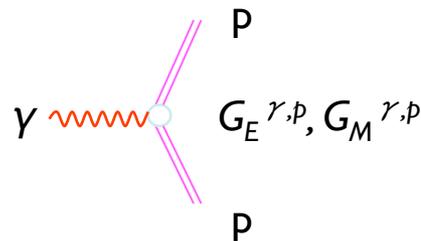
Electroweak charged fermion couplings

	q^γ	q^Z	a^Z
e	-1	$-(1 - 4 \sin^2 \theta_W)$	+1
u	$\frac{2}{3}$	$1 - \frac{8}{3} \sin^2 \theta_W$	-1
d	$-\frac{1}{3}$	$-1 + \frac{4}{3} \sin^2 \theta_W$	+1
s	$-\frac{1}{3}$	$-1 + \frac{4}{3} \sin^2 \theta_W$	+1

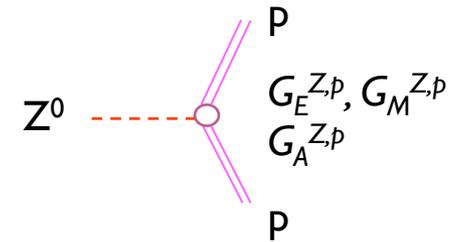
★ $Q_W^p = | -4 \sin^2 q_W \sim 0.071 |$ (SM)

★ $Q_W^n = -1$

Electromagnetic interaction



Neutral weak interaction



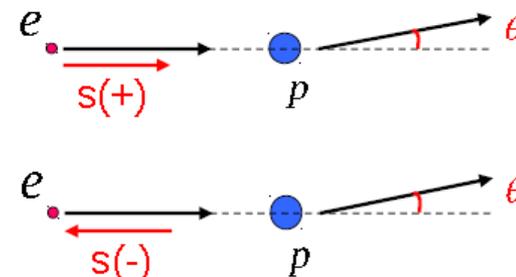
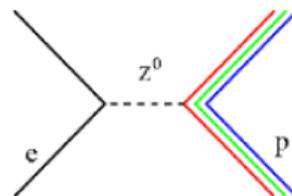
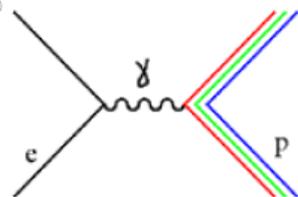
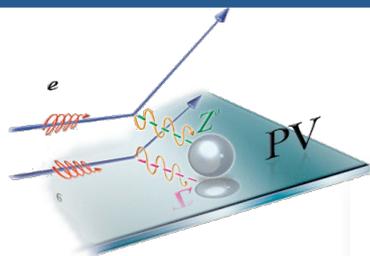
$$G_E^\gamma = \frac{2}{3}G_E^u - \frac{1}{3}G_E^d - \frac{1}{3}G_E^s \quad G_{E,M}^Z = (1 - \frac{8}{3}\sin^2\theta_W)G_{E,M}^u + (-1 + \frac{4}{3}\sin^2\theta_W)G_{E,M}^d + (-1 + \frac{4}{3}\sin^2\theta_W)G_{E,M}^s$$

$$G_M^\gamma = \frac{2}{3}G_M^u - \frac{1}{3}G_M^d - \frac{1}{3}G_M^s$$

Access to rich physics:

- Novel Probes of Nucleon Structure
- Strange Quark Form Factors
- Neutron skin of a heavy nucleus
- Indirect Searches for New Interactions

Parity Violating Electron Scattering



Electromagnetic (PC) + Neutral-weak (PV)

Parity-Violating Asymmetry: $A_{PV} \equiv (\sigma_+ - \sigma_-)/(\sigma_+ + \sigma_-)$

$$A_{PV}(p) = \frac{-G_F Q^2}{4\pi\alpha\sqrt{2}} \left[\frac{\epsilon G_E G_E^Z + \tau G_M G_M^Z - (1 - 4\sin^2\theta_W)\epsilon' G_M G_A^Z}{\epsilon (G_E)^2 + \tau (G_M)^2} \right]$$

As $Q^2 \rightarrow 0, \theta \rightarrow 0$: Forward angles Backward angles

$$A_{PV}(p) \Rightarrow \frac{-G_F Q^2}{4\pi\alpha\sqrt{2}} [Q_W^p + Q^2 B(Q^2)]$$

~230 ppb

Leading order term: $Q_W^p(E)$

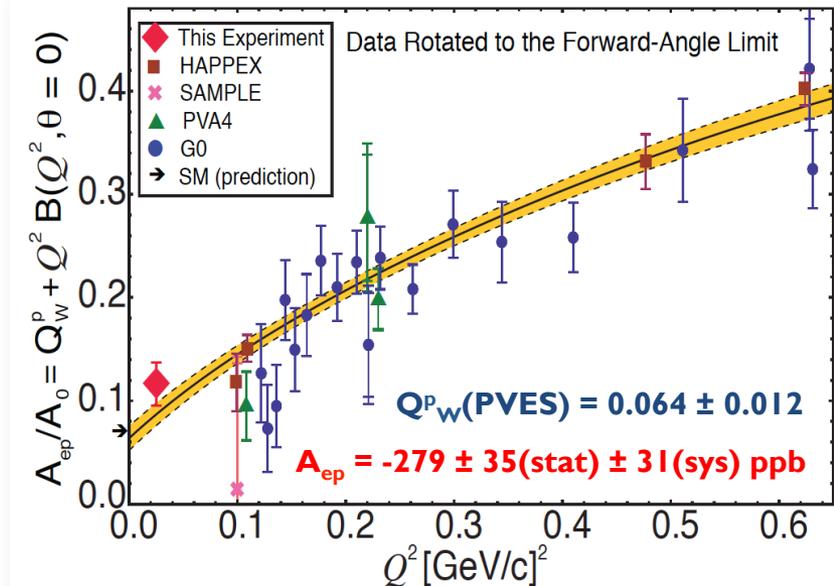
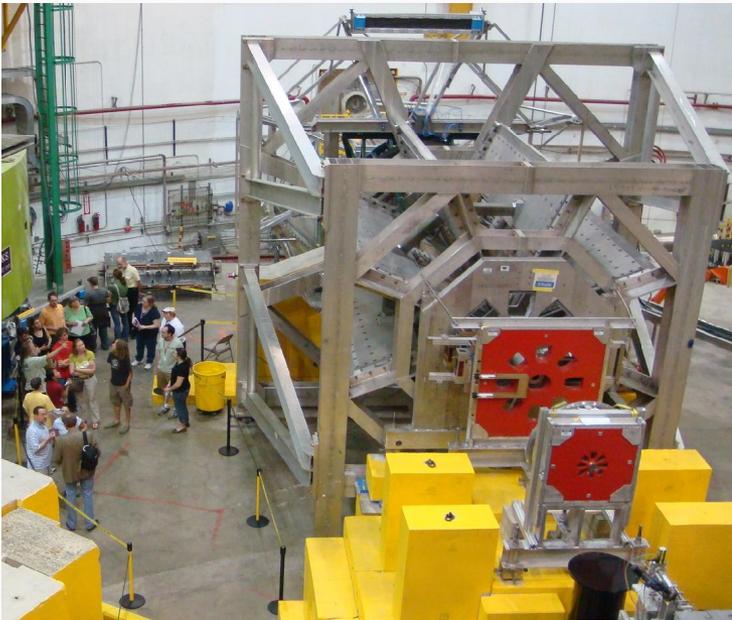
$B(Q^2)$: contains $G_{E,M}^\gamma, G_{E,M}^Z$ which are constrained by previous PC and PV form factor measurements

Using Isospin symmetry ($p \leftrightarrow n$) A_{PV} is expressed as a combination of strange-quark G_M^s and G_E^s only

Qweak at Jefferson Lab

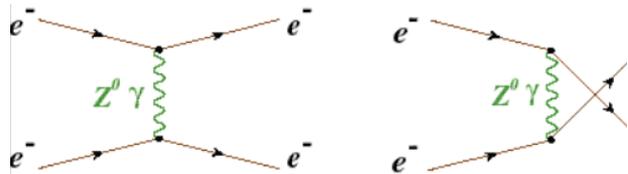
“A Search for New Physics at the TeV scale via a Measurement of the Proton’s Weak Charge”

- Parity violating elastic electron-proton scattering at Jefferson Lab
- First direct measurement of the proton’s weak charge Q_W^p and determination of $\sin^2\theta_W$ at low Q^2 ($\delta\sim 5\%$ goal)
- Analysis of commissioning run (4% of exp. statistics)



Future experiments at JLab: MOLLER

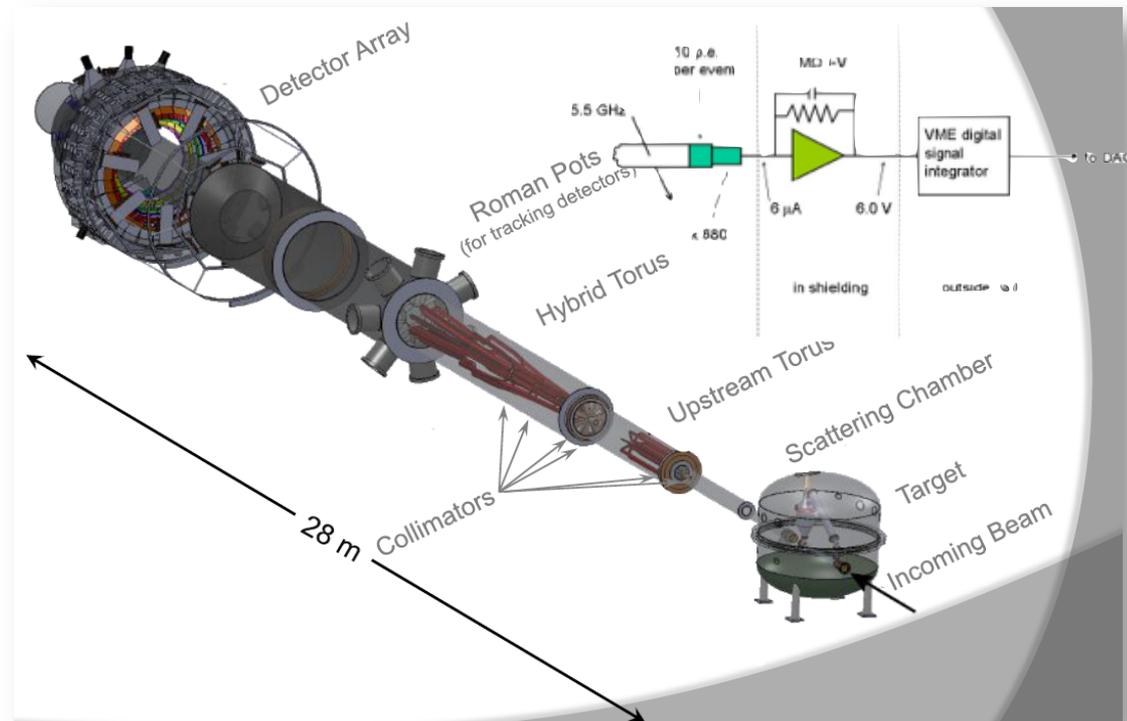
Measurement Of Lepton Lepton Elastic Reaction (Q_w^e measurement)



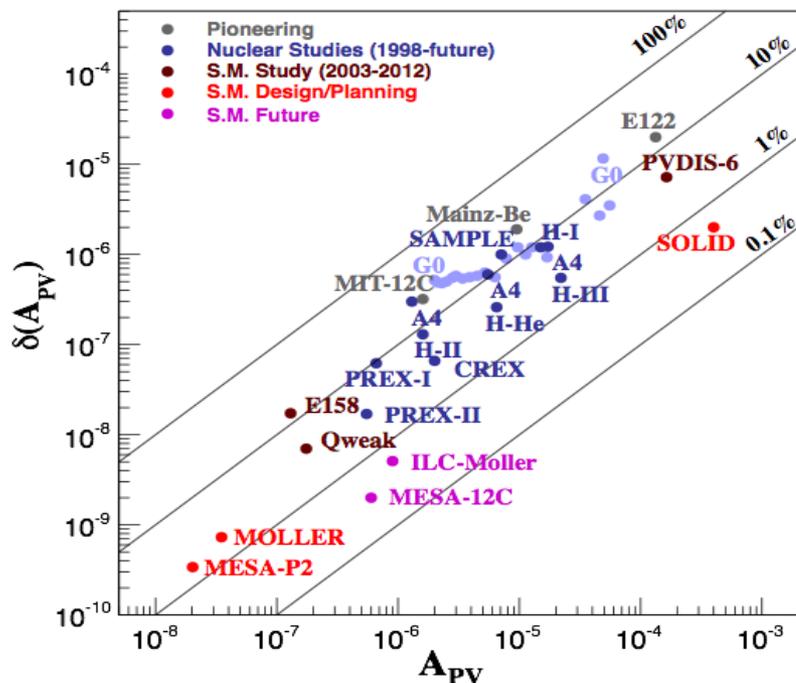
$$Q_w^e = (1 - 4\sin^2 \theta_w) + \text{1-loop corrections}$$

$$A_{PV} = mE \frac{G_F}{\sqrt{2}\pi\alpha} \frac{4 \sin^2 \theta}{(3 + \cos^2 \theta)^2} Q_w^e \sim 30 \text{ ppb}$$

- Purely leptonic probe
- Unprecedented polarized luminosity and beam stability
- 11 GeV, 75 μA , $P > 85\%$
- Expected rate: $\sim 150 \text{ GHz}$, $< 10\% \text{ bg}$
- 5kW dissipated power on H2 target
- e-p elastic scattering bg suppression
- Goal: $\delta A/A \sim 2\%$

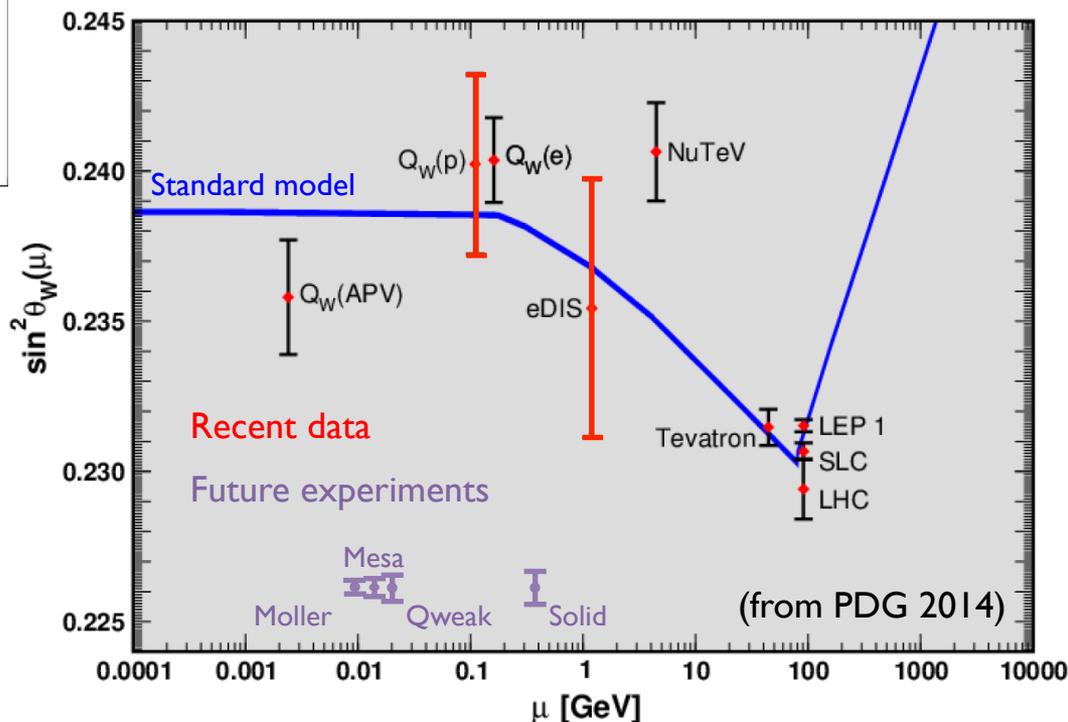


Testing the Standard Model



New experiments planned in near future will reach a precision of 1ppb in PV asymmetries

Measurements at GeV scale can be sensitive to new physics at the TeV scale



Lead (^{208}Pb) Radius Experiment: PREX

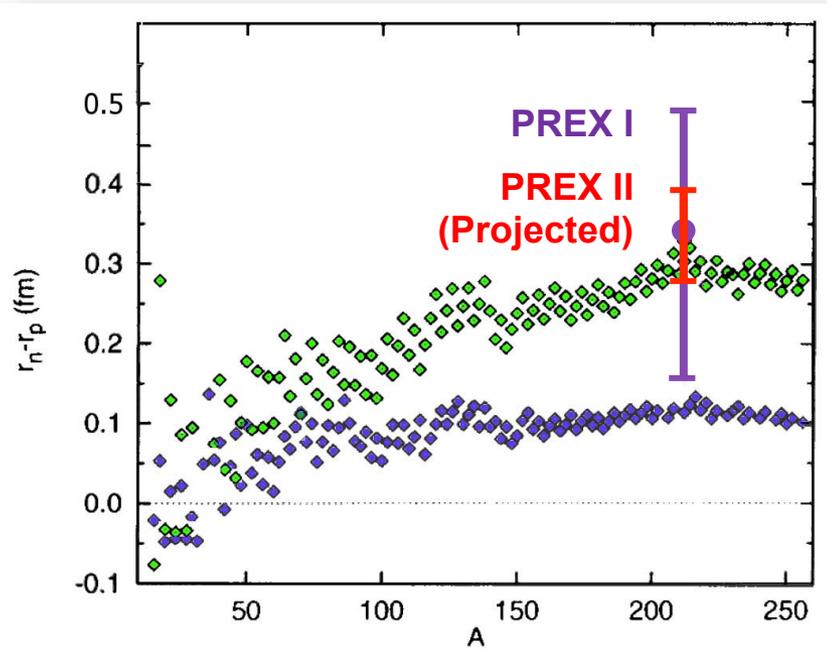
Parity Violating Electron Scattering from Nuclei
Exploit weak coupling of the neutron

First measurement:

clear indication of neutron skin

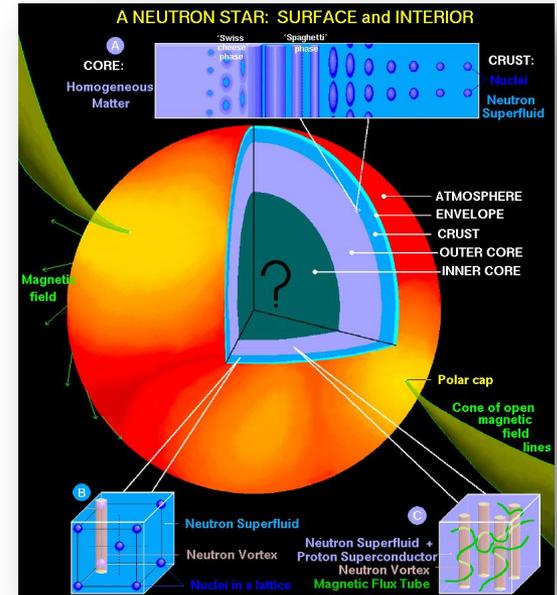
Future measurement:

constrain symmetry equation of state



Relativistic
mean field

Non-relativistic
skyrme

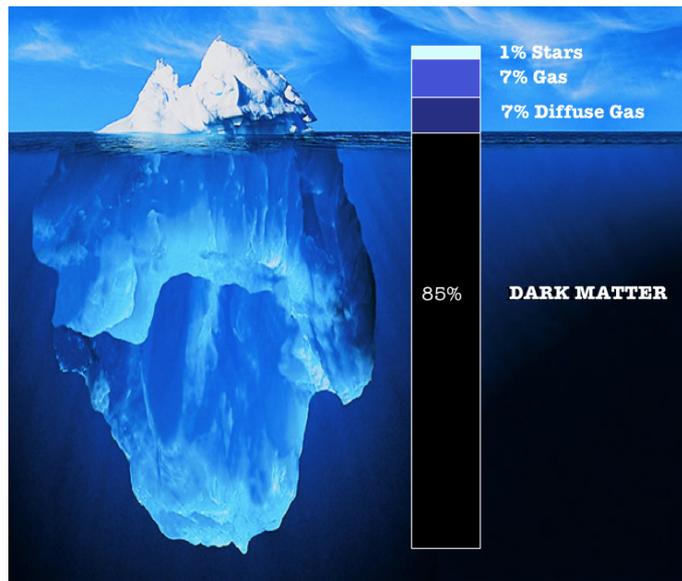


A neutron skin of 0.2 fm or more has implications for our understanding of neutron stars and their ultimate fate.

^{48}Ca experiment also approved at Jlab
Parallel program at Mainz MESA facility

Dark Matter vs. Baryonic Matter

Compelling astrophysical indications about DM existence



★ Does DM participate to non-gravitational interactions?

★ Is DM a new particle?

★ Constraint on DM mass and interactions:

- should be “dark” (no EM interaction)
- should weakly interact with SM particles
- should provide the correct relic abundance
- should be compatible with CMB power spectrum
... assuming that the gravity is not modified and DM undergoes to other interactions

★ We can use what we know about standard model particles to build a DM theory

Use the SM as a template: $SM = U(1)_{EM} \times SU(2)_{Weak} \times SU(3)_{Strong}$

Particles, interactions and symmetries

known particles
&
force-carriers

particles:
quarks, leptons

force-carriers:
gluons, γ , W, Z, graviton (?), Higgs, ...

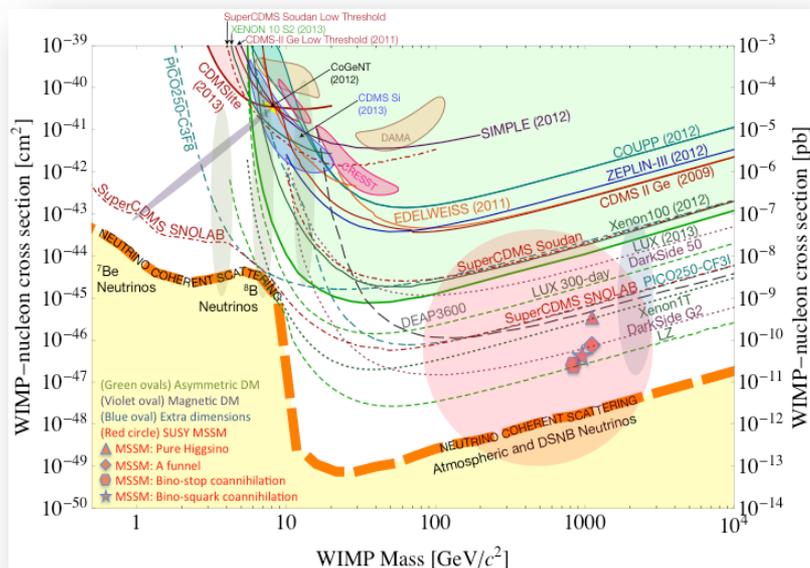
Two options:

★ **New matter** interacting through the **same forces**

★ **New matter** interacting through **new forces**

Weakly Interacting Massive Particles

Limits on DM_{Xsec} vs. DM_{Mass} from Direct Detection



The WIMPs paradigm:

- slowly-moving cosmological Weakly Interacting Massive Particles interacting with the SM-Weak force
- DM detection by nuclear recoil
- Experiments optimized for heavy DM ($M > 10$ GeV)

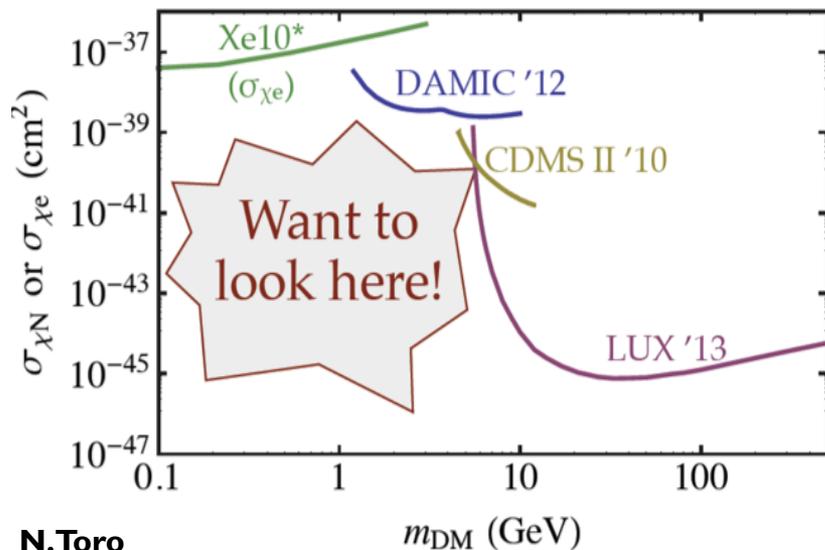
No (firm) evidence (so far):

- weak-interaction at tree level ruled-out
- exp. limits close to irreducible neutrino background



WIMPs paradigm is not the only theoretically well-motivated option
 Light DM: extending the search to unconventional (and unexplored!) territory

Hunting for Dark Matter: light DM

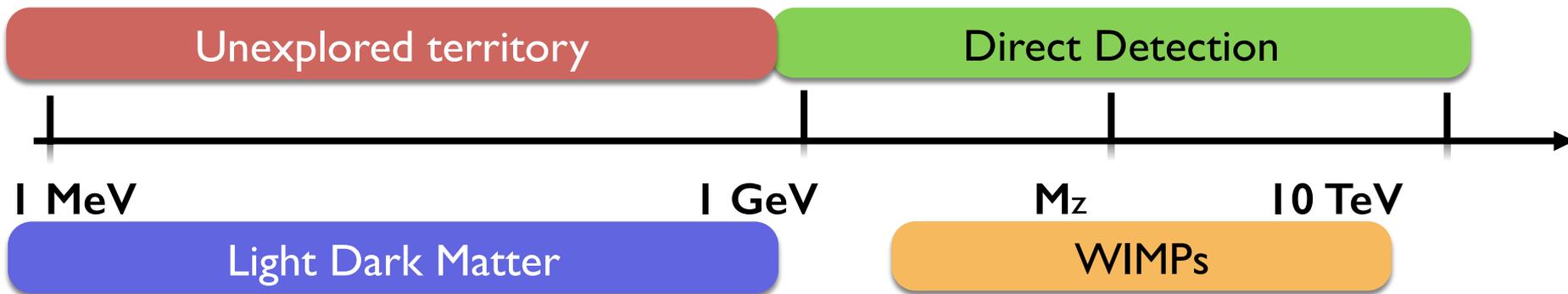


Dark/Hidden Sector Light Dark Matter couples to SM with a new force

- Light Dark Matter (X) in 1-1000 MeV mass range
- New vector boson (A') responsible for DM-DM and DM-SM interaction
- (Traditional) Direct Detection is (almost) impossible

Accelerator-based DM search

- ★ High intensity
- ★ Medium energy

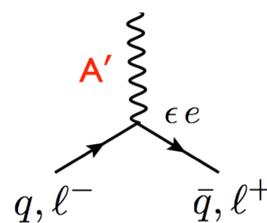
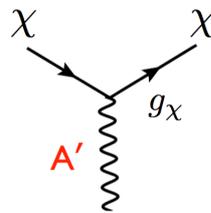
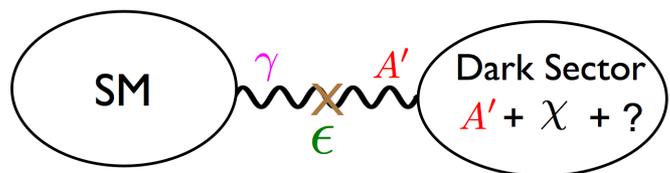


Dark Sector or Hidden Sector
DM not directly charged under SM interactions

Can be explored at accelerators!

Dark Forces & Dark Matter

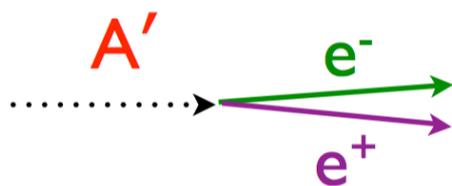
(Light DM - Light Mediators)



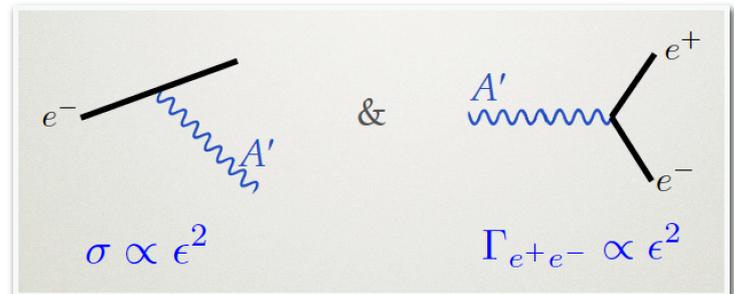
4 parameters:

- $m_\chi, m_{A'}, \epsilon, \alpha_D$
- $m_\chi, m_{A'} : \text{MeV} - \text{GeV}$

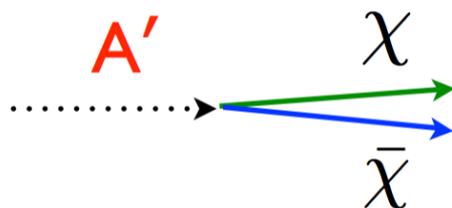
Visible



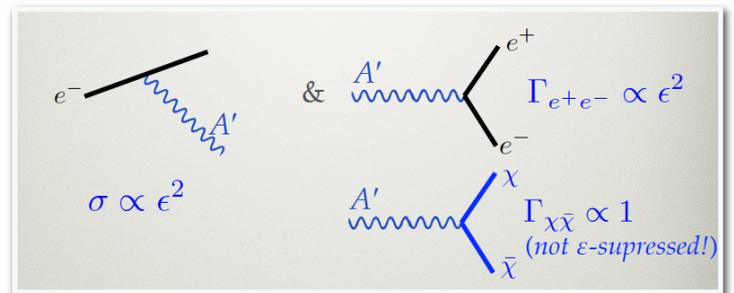
- Minimal decay
- Decay regulated by ϵ^2
- Independent of m_χ
- Requires $m_{A'} < 2m_\chi$ (on-shell)



Invisible



- Depends on 4 parameters
- $m_{A'} > 2m_\chi$ (on-shell)
- $\alpha_D = g_\chi^2 / 4\pi \gg \epsilon^2 \alpha_{EM}$



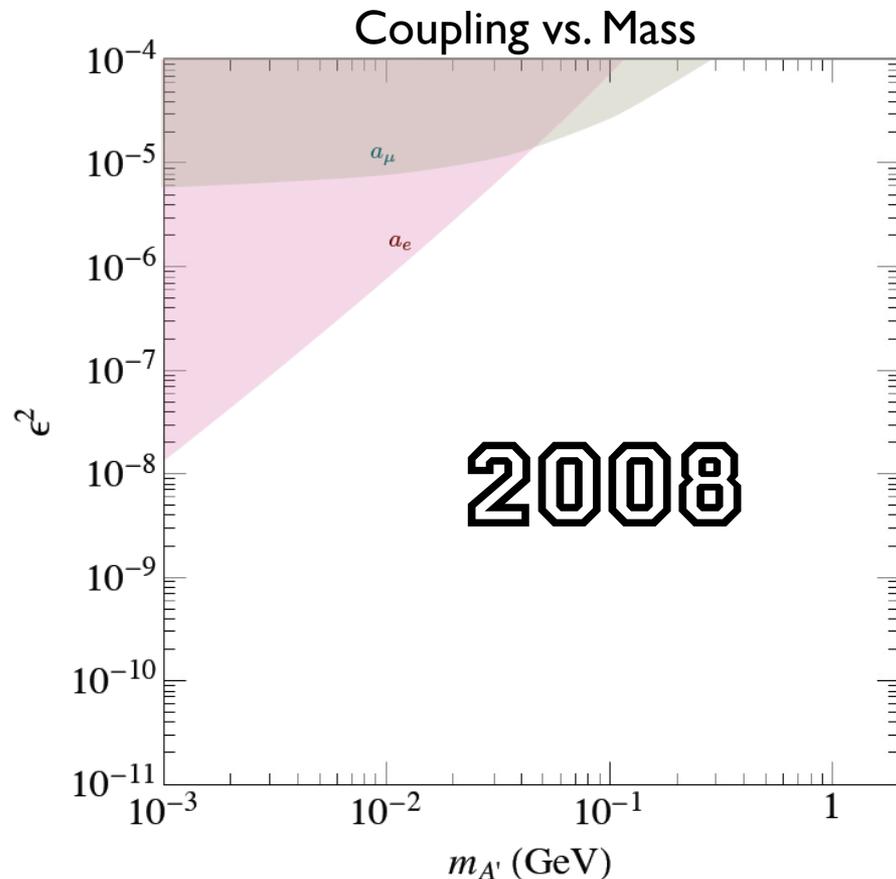
Hunting for the Dark Photon: visible decays

Fixed target: $e N \rightarrow e N A' \rightarrow e N l^+ l^-$
→ **JLAB, MAINZ**

Fixed target: $p N \rightarrow N A' \rightarrow p l^+ l^-$
→ **FERMILAB, SERPUKHOV**

Annihilation: $e^+ e^- \rightarrow A' \gamma \rightarrow l^+ l^- \gamma$
→ **BABAR, BELLE, KLOE, CLEO**

Meson decays: $\pi^0, \eta, \eta', \omega, \phi \rightarrow A' \gamma (M)$
→ $l^+ l^- \gamma (M)$
→ **KLOE, BES3, WASA-COSY,
PHENIX, NA48, LHCb**



Hunting for the Dark Photon: visible decays

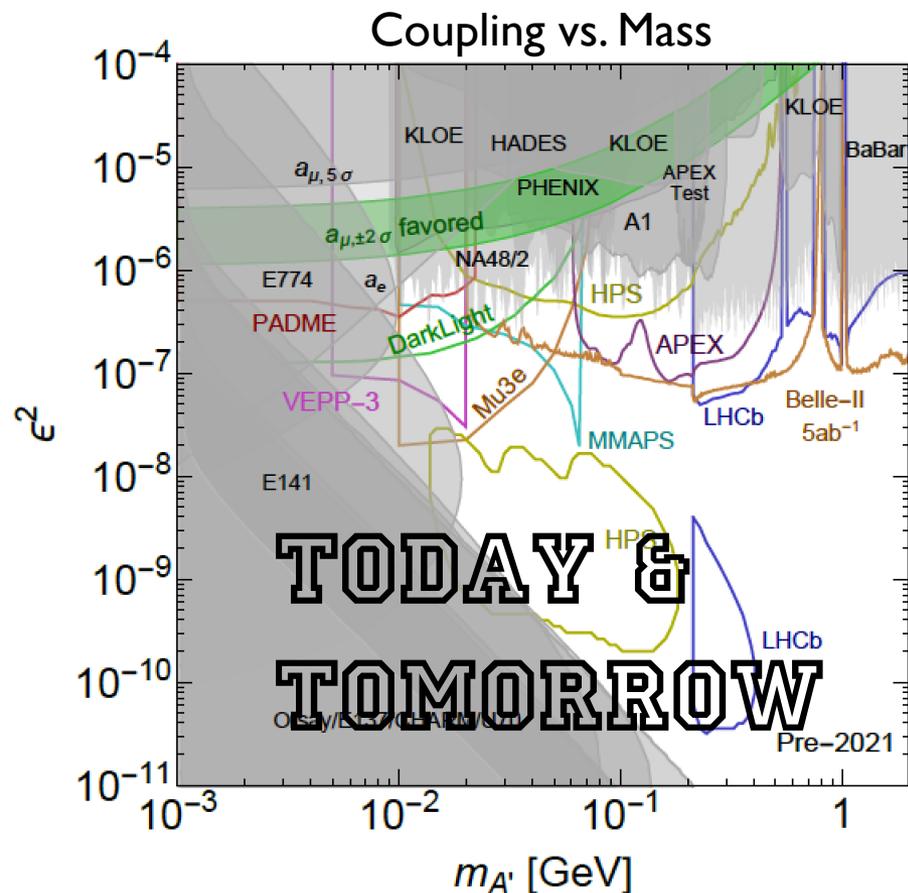
Fixed target: $e N \rightarrow e N A' \rightarrow e N l^+ l^-$
→ **JLAB, MAINZ**

Fixed target: $p N \rightarrow N A' \rightarrow p l^+ l^-$
→ **FERMILAB, SERPUKHOV**

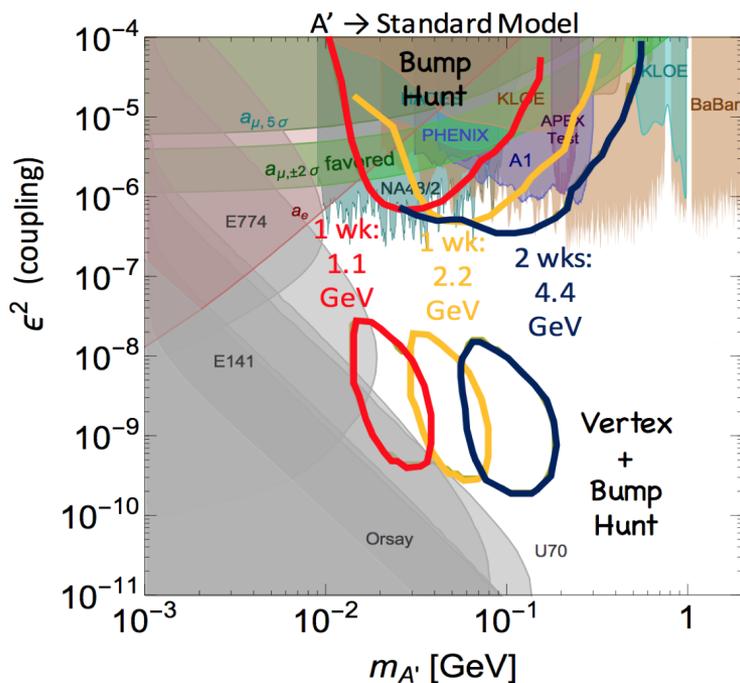
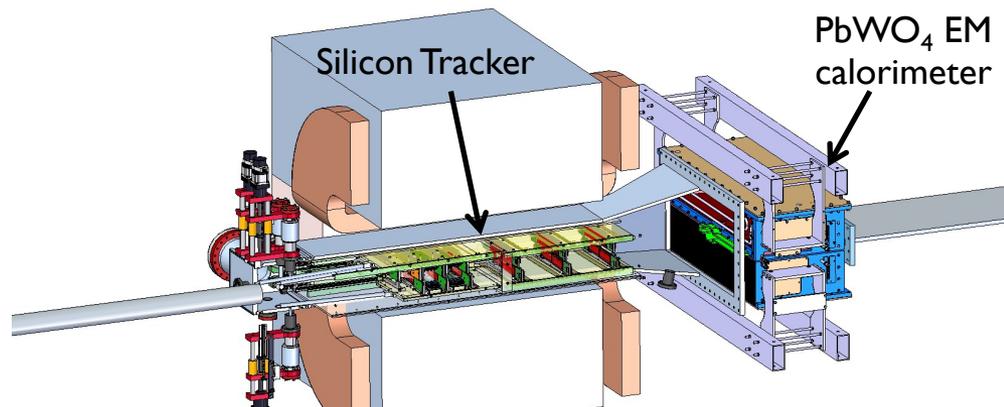
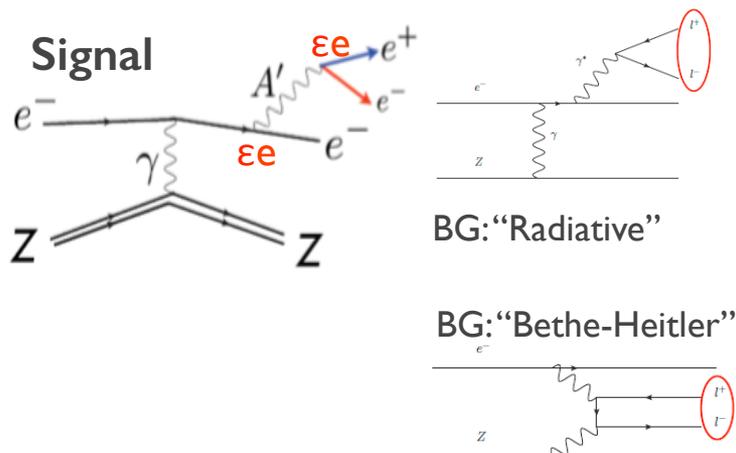
Annihilation: $e^+ e^- \rightarrow A' \gamma \rightarrow l^+ l^- \gamma$
→ **BABAR, BELLE, KLOE, CLEO**

Meson decays: $\pi^0, \eta, \eta', \omega, \phi \rightarrow A' \gamma (M)$
→ $l^+ l^- \gamma (M)$
→ **KLOE, BES3, WASA-COSY, PHENIX, NA48, LHCb**

No positive signal (so far)
but limits in coupling vs.
mass parameter space



The Heavy Photon Search Experiment



HPS search strategy:

- *Bump hunt*: peak in the e^+e^- invariant mass spectrum (mass resolution 1 MeV)
- *Displaced vertex*: e^+e^- detached from primary vertex (target) for long-lived A'

HPS Status:

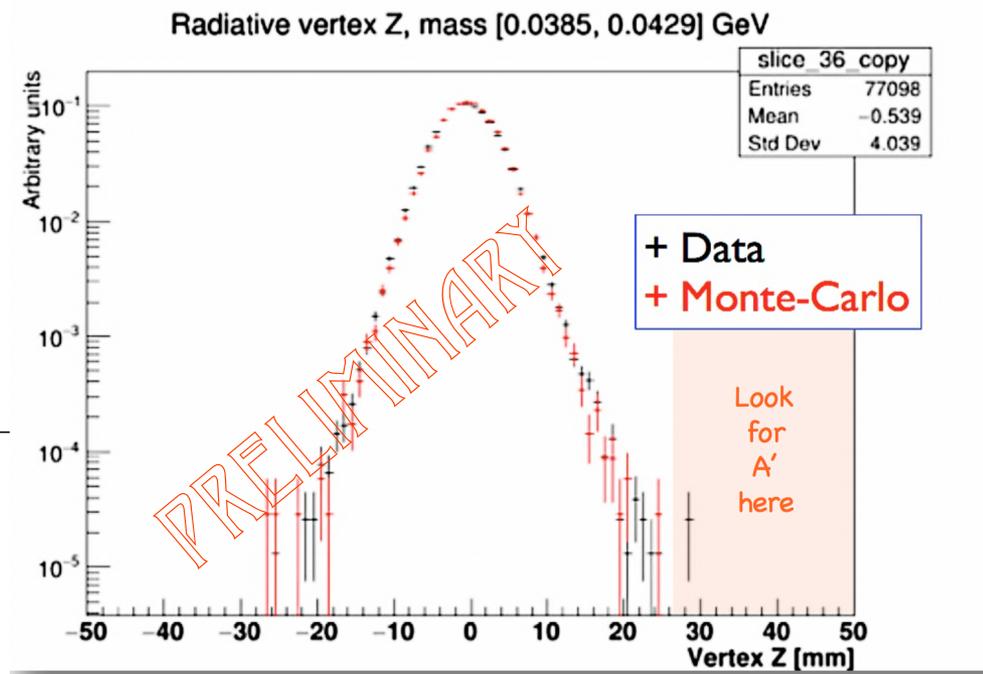
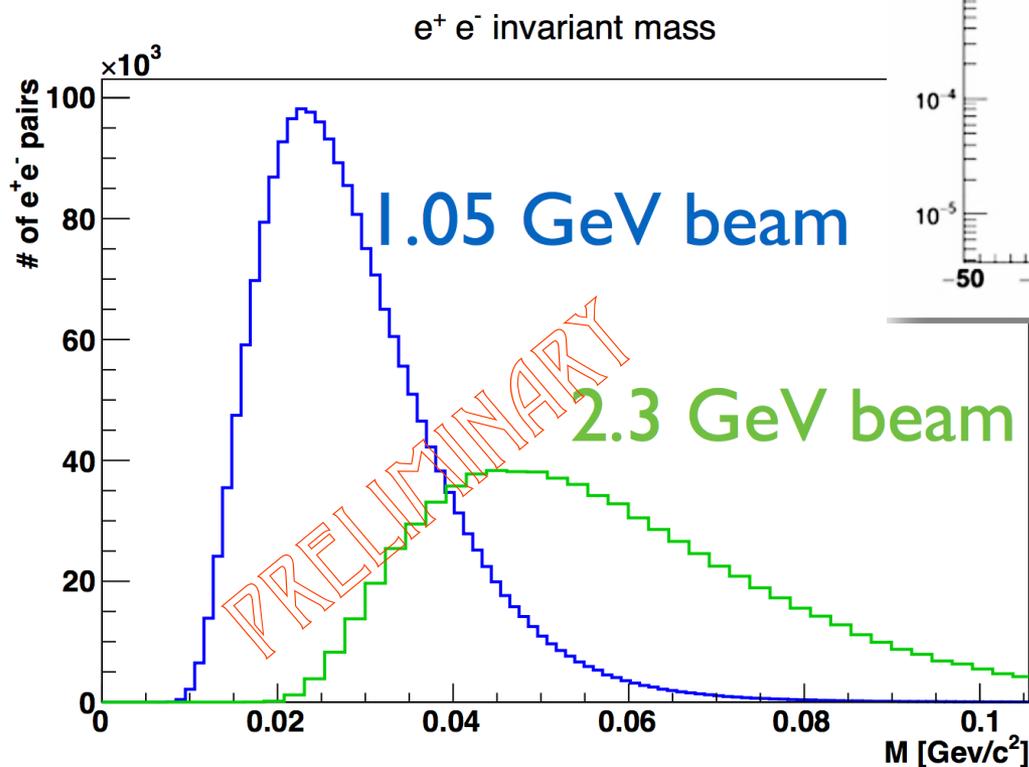
- 180 approved days at 1.1, 2.2 and 4.4 GeV in Hall B at JLab
- Commissioning run successfully completed in Spring 2015: SVT at 0.5 mm from the beam
- Opportunistic running in 2015-2016 at 1.05 and 2.3 GeV
- Data analysis in progress

The Heavy Photon Search Experiment



Searching for the Heavy Photon using a blinded analysis (10% of the data) for both bump hunting and vertexing

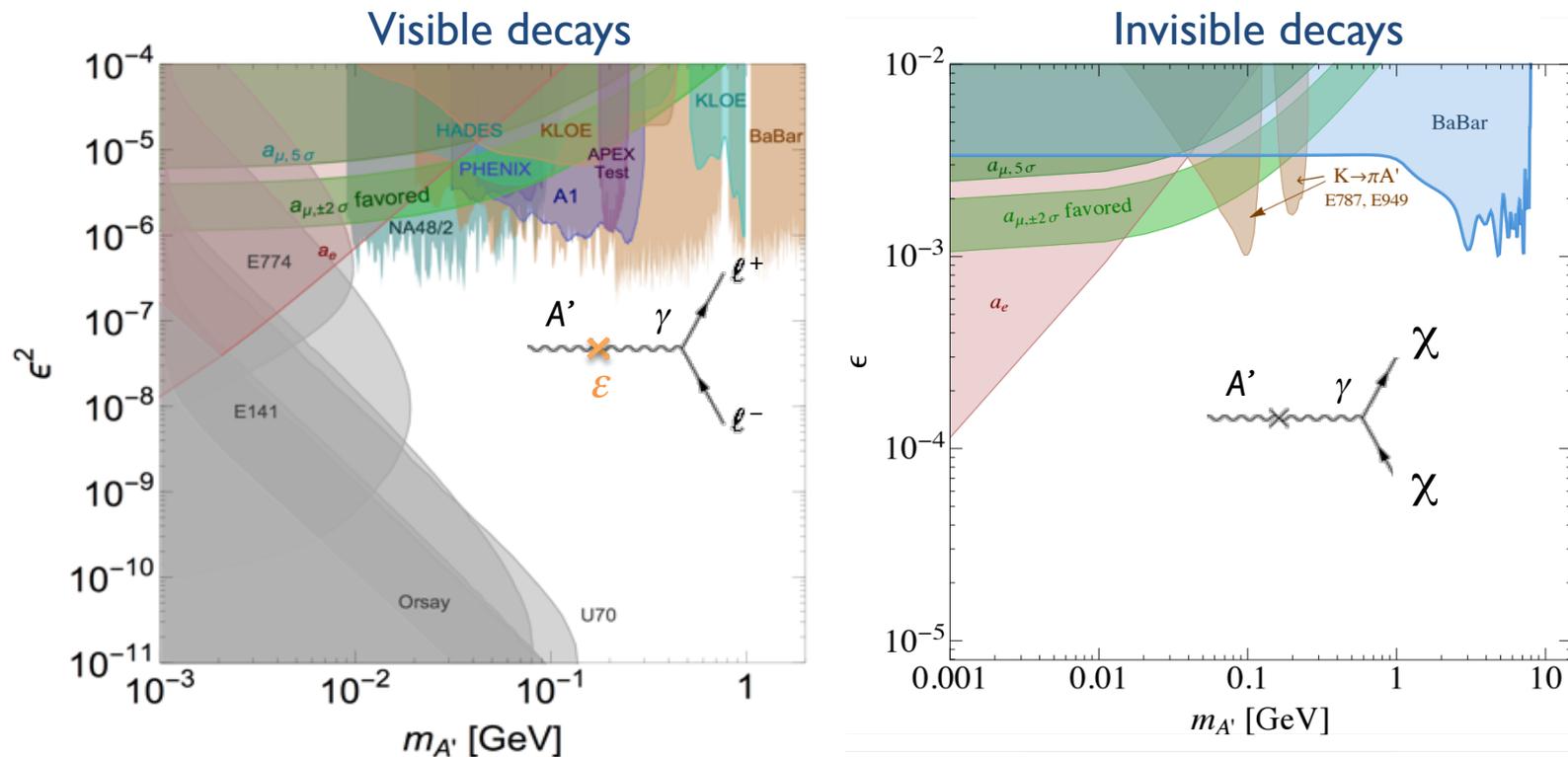
Unblind analysis (full statistics) after finalizing procedure and freezing cuts



- Mass and vertex resolution in agreement with MC
- Results from 2015-2016 data in early 2017
- Detector upgrades to increase reach in progress
- Major data taking period in 2018

Dark Photon Invisible Decays

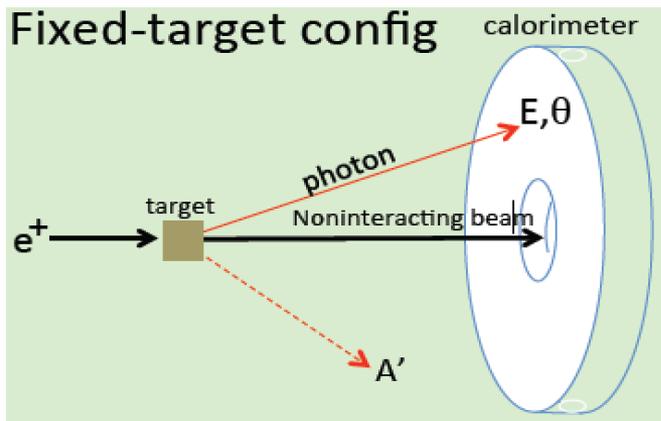
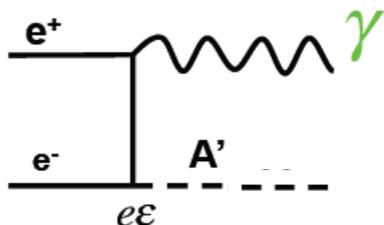
If any χ ($M_\chi < M_{A'}/2$) exists, A' will decay into dark sector particles: very few model independent measurement are available in this scenario



Two experimental methods:

- missing mass A' search
- dark matter χ scattering searches

Dark Photon Invisible Decays



Missing mass search:

- Independent of A' decay mechanism
- Bump hunt
- Need a positron beam
- Limited $M_{A'}$ accessible
 - 1 GeV beam: $M_{A'} < 31$ MeV
 - 5 GeV beam: $M_{A'} < 71$ MeV

VEPP3 @ Novosibirsk

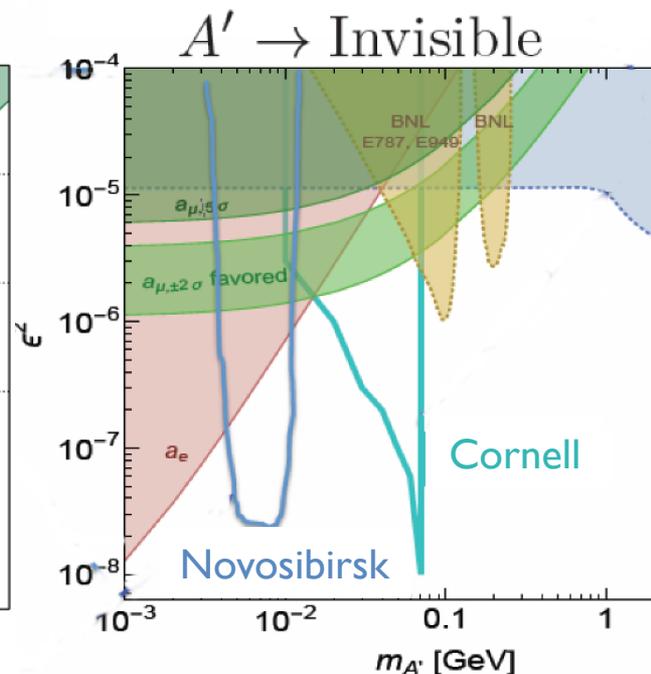
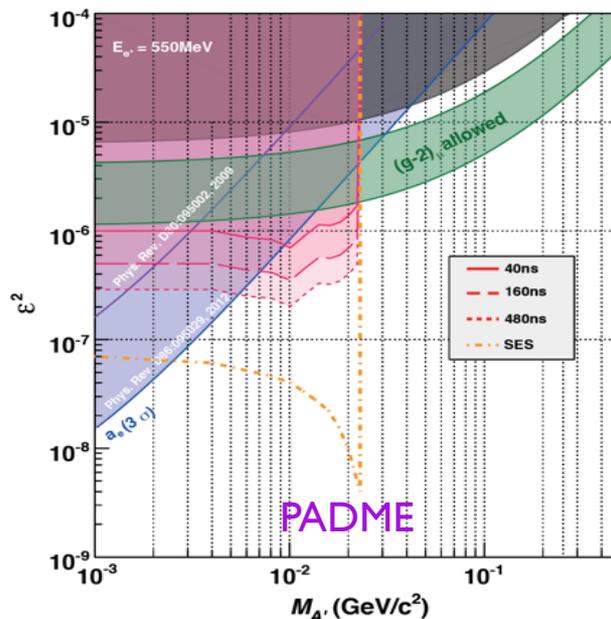
$E_{e^+} = 500$ MeV
EOT $\sim 10^{15} - 10^{16}$ year $^{-1}$

PADME @ LNF

$E_{e^+} = 550$ MeV
EOT $\sim 10^{13} - 10^{14}$ year $^{-1}$

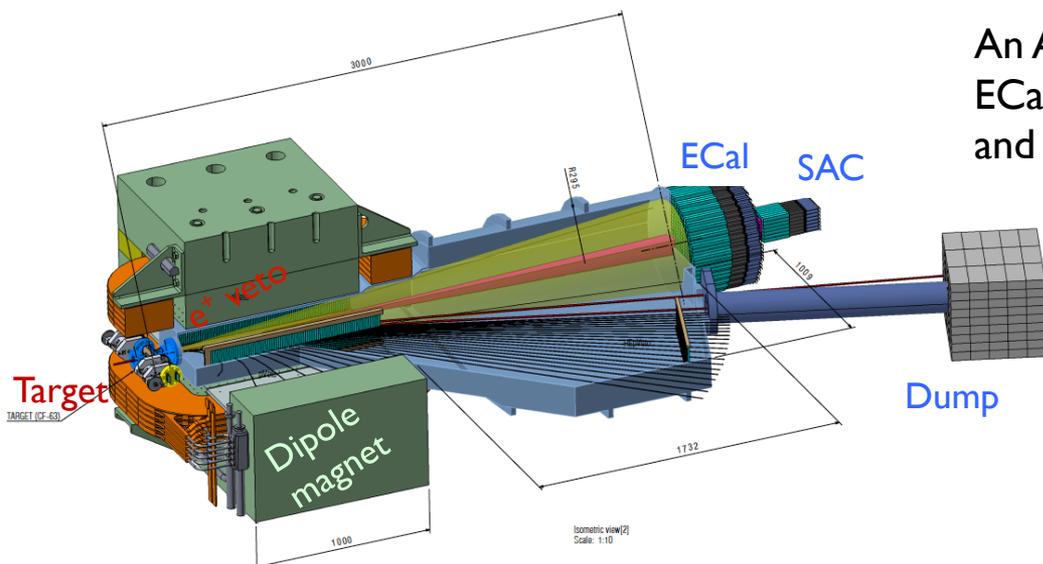
CORNELL

$E_{e^+} = 5.3$ GeV
EOT $\sim 10^{17} - 10^{18}$ year $^{-1}$

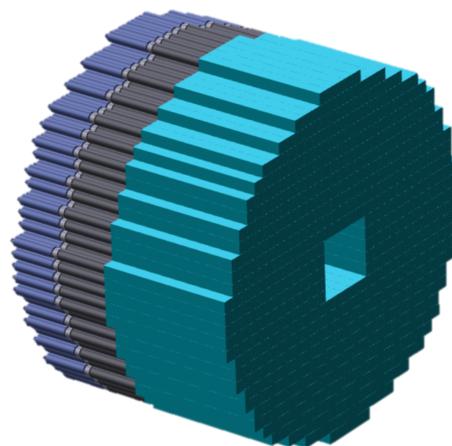


An A' candidate in PADME is a single photon in the ECal having no charged particles in time in the veto and no photons in time in the SAC:

- Beam: $\sim 5000 e^+$ on target in 40 ns bunch, at 50 bunch/s ($\sim 10^{13} e^+$ /year)
- Measure: Energy, time and direction of photons
- Compute the $M_{\text{miss}}^2 = (P_{e^-}^4 + P_{\text{beam}}^4 - P_{\gamma}^4)^2$
- Veto any other particle outside the beam region



- The PADME experiment has been approved from INFN at the end of 2015
- The PADME experiment is financed by the “What Next” INFN program with 1.35M€ (2016-2018)
- Goal: completing the detector assembly by the end of 2017 and to accumulate $10^{13} e^+$ on target by the end of 2018

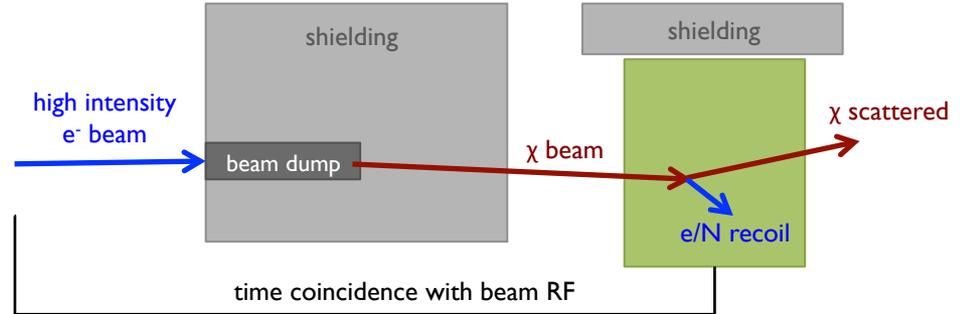
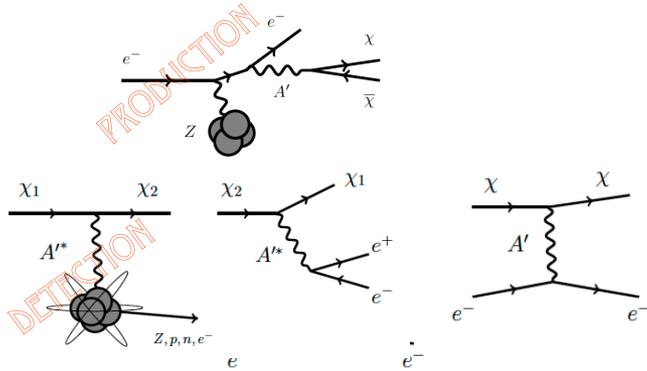


ECal:

- BGO calorimeter
- Cylindrical shape: radius 290 mm, depth of 230 mm, 616 crystals
- Expected performance:
 - $\sigma(E)/E = 1.57\%/\sqrt{E} \oplus 0.35\%$
 - 1.5 mrad angular resolution

Beam Dump e χ periment

A beam of DM particle can be produced by the interaction of a high intensity/high energy (electron) beam in a dump

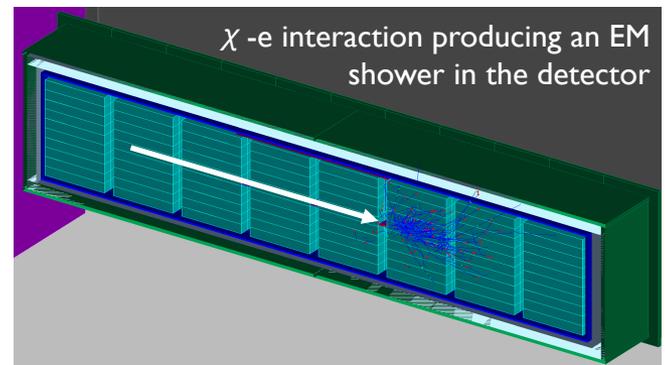
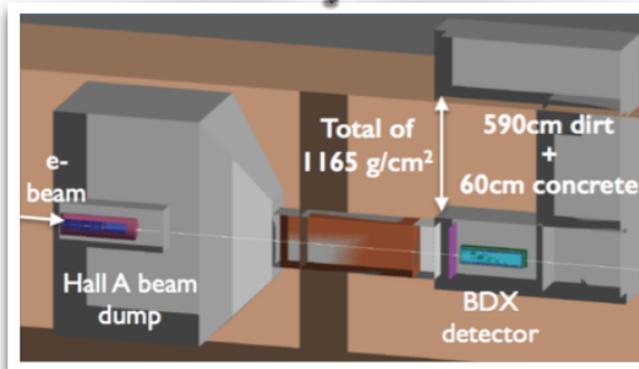


BDX @ JLab:

- New underground (~8m) hall downstream a high intensity dump
- Detector capable of measuring EM shower from electron recoil

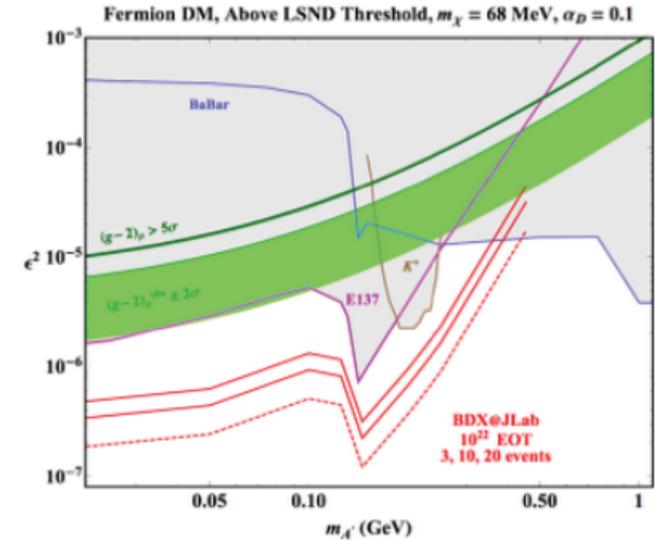
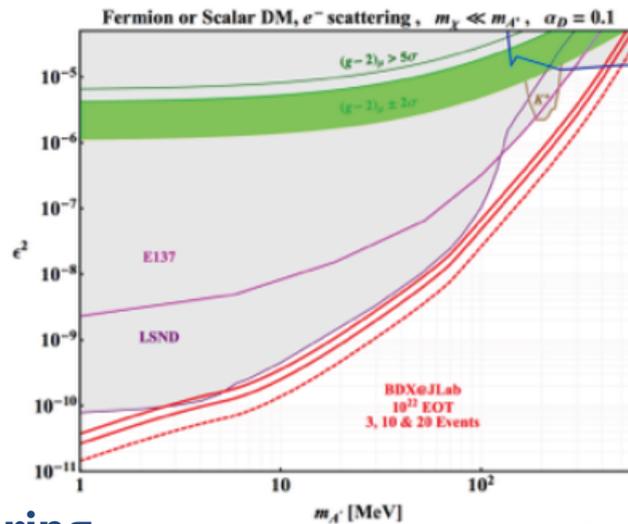
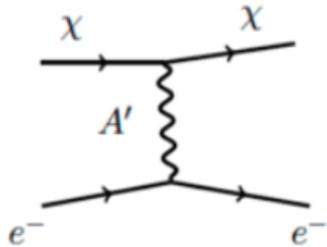
BDX status:

- Proposal approved by JLab PAC in July 2016 (arXiv:1607.01390)
 - 11 GeV electron beam, 10^{22} EOT (65 uA for 285 days)
 - ~800 CsI(Tl) crystals (ex BaBar EC) (~1 m³ active volume) + plastic vetos

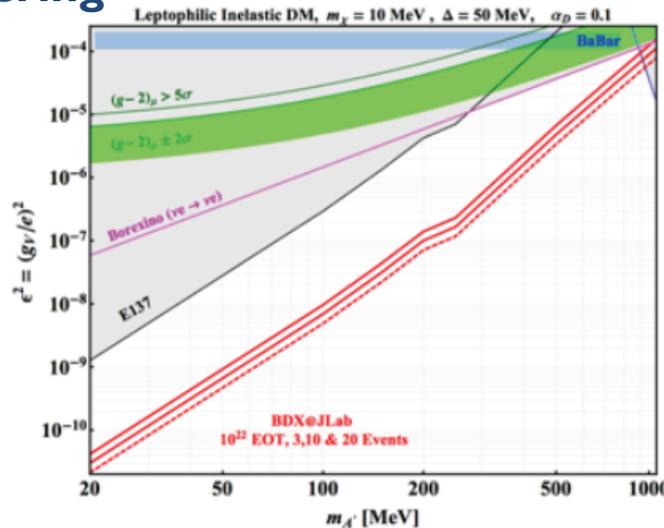
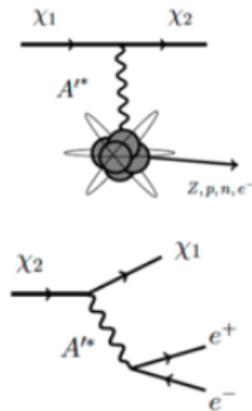


Beam Dump e χ periment

χ -e elastic scattering



χ -e inelastic scattering



10-100 times more sensitive
than previous experiments
excluding a significant area of
the parameter space

Summary

- Medium energy, high intensity experiments offers important opportunities for search for new physics at mass scale of $\sim \text{GeV}$
- Hints for new physics at this scale arise for various anomalous observations
- Ongoing searches explore different avenues:
 - Precision tests of the Standard Model
 - Search for light dark matter at accelerators
 - ...
- A broad experimental program that involves several experiments and facilities that can reach unprecedented precision challenging the potentials of the energy frontier