AN APOLOGY FOR ADS3

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- ADS3 IS A TOY MODEL THAT CAPTURES SOME OF THE MOST PUZZLING ASPECTS OF QUANTUM GRAVITY
- THERE ARE NO GRAVITATIONAL WAVES IN PURE ADS3 GRAVITY (AND THIS IS GOOD)
- A "SQUARE ROOT" OF PURE GRAVITY EXISTS
- ADS3 POSSESSES INFINITE-DIMENSIONAL ASYMPTOTICAL ALGEBRAS THAT IN PART REDUCE DYNAMICS TO KINEMATICS
- HIGH-SPIN FIELDS PROPAGATE ZERO DEGREES OF FREEDOM BUT THEY ARE NEITHER FREE NOR EQUIVALENT TO FREE FIELDS

THE EASY PROBLEM WITH QUANTUM GRAVITY: RENORMALIZABILITY

EINSTEIN ACTION HAS A COUPLING CONSTANT WITH DIMENSION [square length] (NEWTON'S CONSTANT)

$$S = \frac{1}{16\pi G} \int d^4x \sqrt{-g} [R(g) + L_{matter}]$$

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$$S_{div} \sim G^{n-1} \log(\Lambda/\mu) \int d^4x \sqrt{-g} R^{n+1}$$

SCALAR MADE OUT OF RIEMANN TENSORS

INFINITE SET OF COUNTERTERMS

WHAT IF THEORY IS FINITE (N=8 SUPERGRAVITY?)

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QUESTIONS SUCH AS UNITARITY IN BLACK-HOLE EVAPORATION OR THE FATE OF SINGULARITIES REMAIN AS UNKNOWN AS IN THE NON-RENORMALIZABLE THEORY

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WE STILL DON'T KNOW WHAT THE THEORY DOES AT VERY HIGH ENERGY

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EXAMPLE: AVERAGE CMB TEMPERATURE AS A CLOCK

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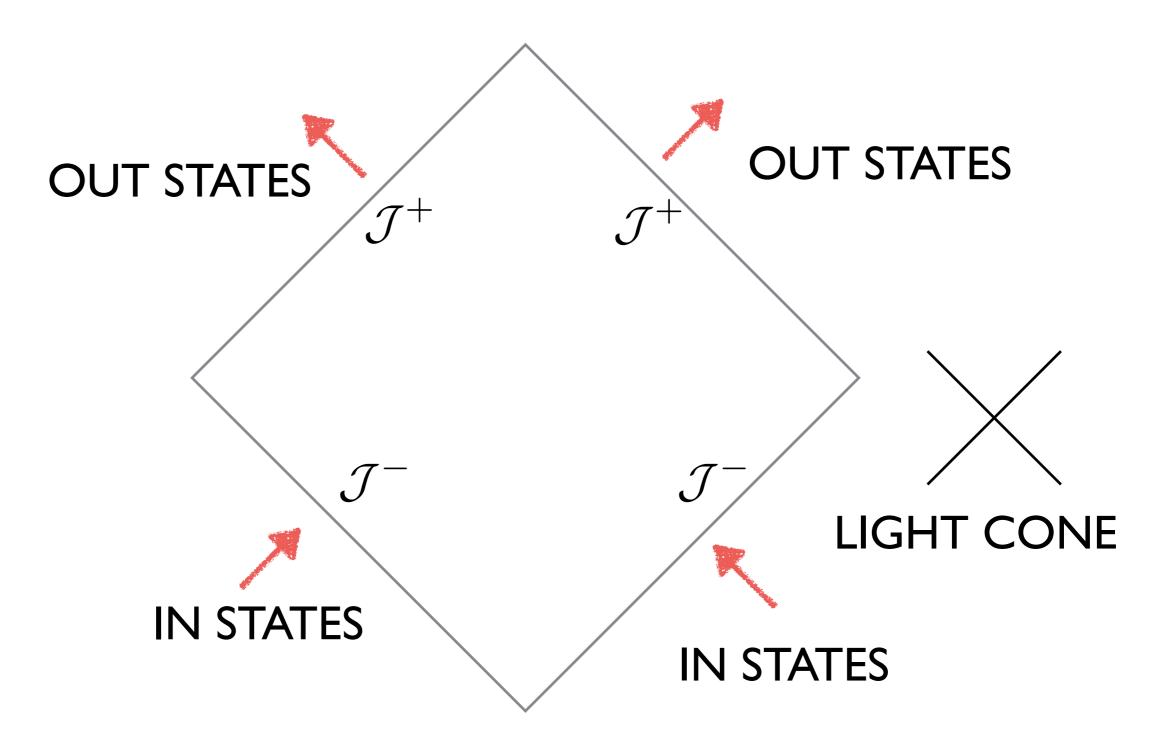
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ONLY NON-PERTURBATIVELY DEFINED OBSERVABLES DEPEND ON THE ASYMPTOTIC BEHAVIOR OF SPACETIME

IN ASYMPTOTICALLY FLAT SPACETIME THE ONLY OBSERVABLE IS THE S-MATRIX



ANTI DE SITTER SPACE

SOLUTION OF EINSTEIN EQUATIONS WITH NEGATIVE COSMOLOGICAL CONSTANT (PROPAGATE SAME DEGREES OF FREEDOM AS PURE GRAVITY)

 $R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G T_{\mu\nu}, \qquad T_{\mu\nu} = g_{\mu\nu}(d-1)(d-2)/16\pi G L^2$ NEGATIVE CONSTANT ENERGY DENSITY

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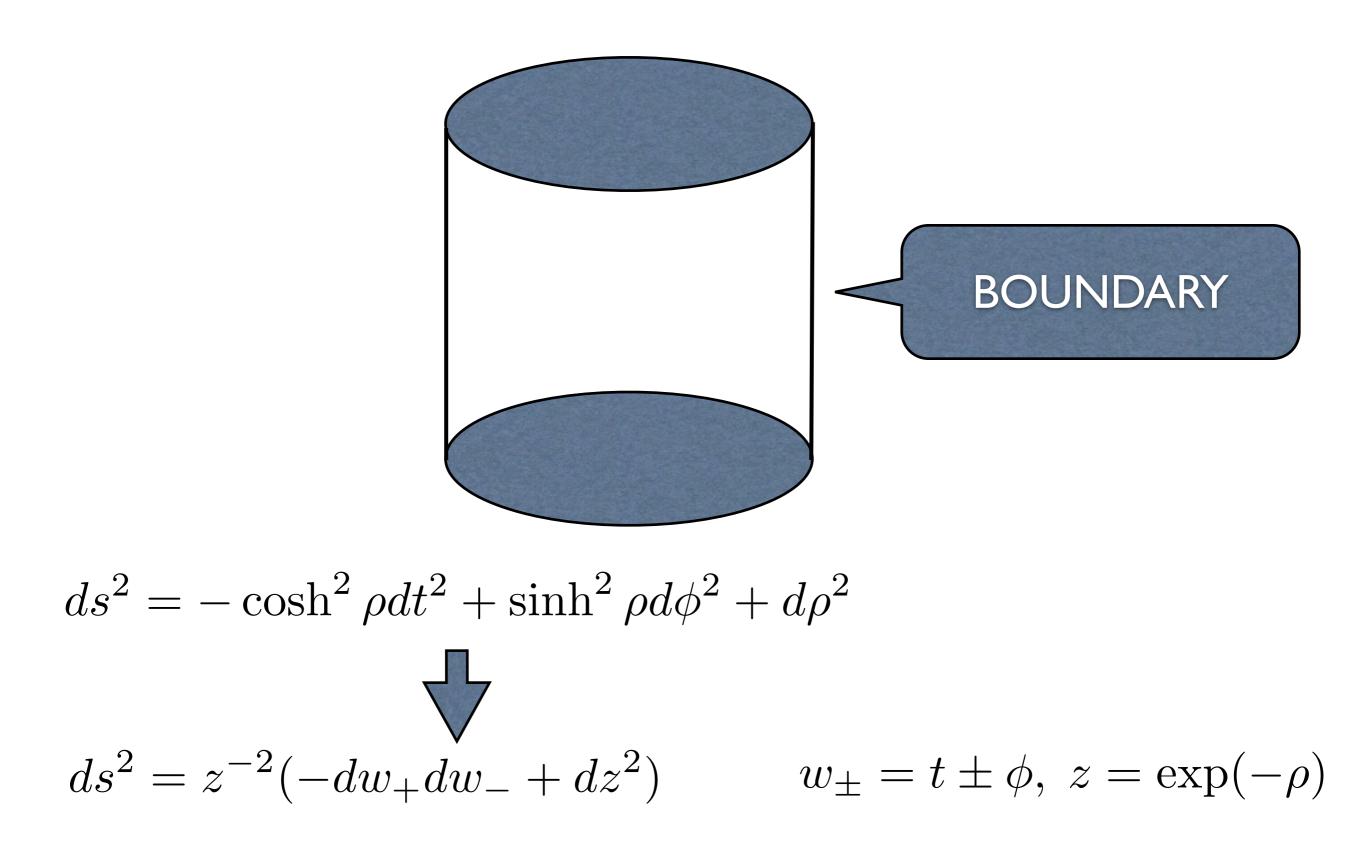
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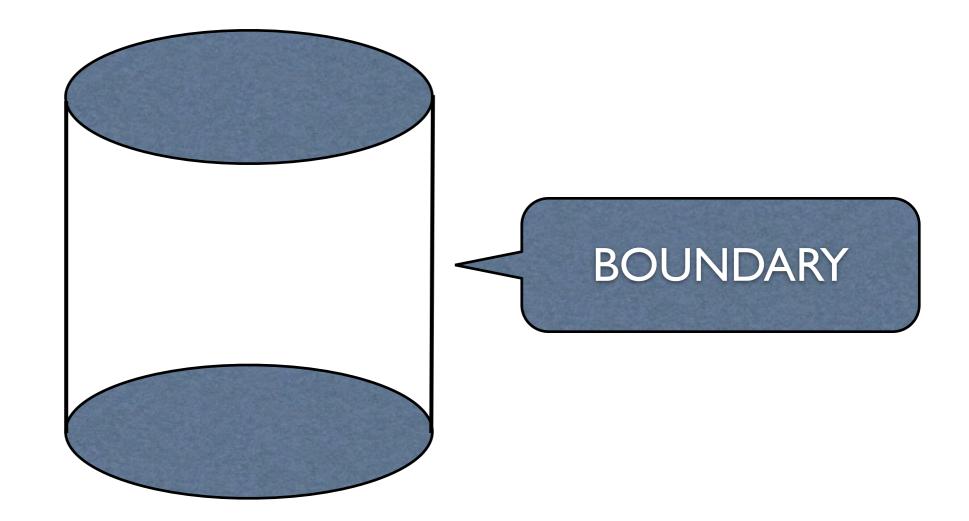
GEOMETRICALLY IT IS A d-DIMENSIONAL HYPERBOLA IN A (2+d)-DIMENSIONAL SPACE OF SIGNATURE (2,d)

$$-(x^0)^2 - (x^{d+1})^2 + \sum_{i=1}^d (x^i)^2 = -L^2$$

SIMPLEST CASE: d=2 ADS3 IS CONFORMAL TO A SOLID CYLINDER



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OBSERVABLES: LOCAL OPERATORS DEFINED ON THE BOUNDARY

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$$w^{\pm} \to w^{\pm} + \epsilon^{\pm}(w^{\pm}) + \frac{1}{2}z^{2}\epsilon''^{\mp}(w^{\mp})$$

$$z \to z - \frac{1}{2}z[\epsilon'^+(w^+) + \epsilon'^-(w^-)]$$

THE ALGEBRA DEFINED BY THE CHARGES ASSOCIATED TO ASYMPTOTIC DIFFEOMORPHISM HAS A CLASSICAL CENTRAL CHARGE

$$[L_n, L_m] = (n - m)L_{n+m} + \frac{c}{12}\delta_{n, -m}m(m^2 - 1)$$
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IN ANY THEORY OF QUANTUM GRAVITY ON ADS3, STATES FIT IN UNITARY REPRESENTATIONS OF THE VIRASORO ALGEBRA

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 $R_{0\mu} - g_{0\mu}R + g_{0\mu}\Lambda = 0 \rightarrow -3$ non-dynamical equations = constraints

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FOR ADS3, BOUNDARY CFT IS 2d. BEST KNOWN BECAUSE OF INFINITE-DIMENSIONAL VIRASORO ALGEBRA

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CLASSICAL GRAVITY AS CHERN-SIMONS $A = e/l - \omega, \qquad \tilde{A} = e/l + \omega$ $S_E = S(A) - S(\tilde{A})$ $S = \frac{k}{4\pi} \int \operatorname{tr} (AdA + \frac{2}{3}AAA), \qquad k = \frac{l}{4G}$

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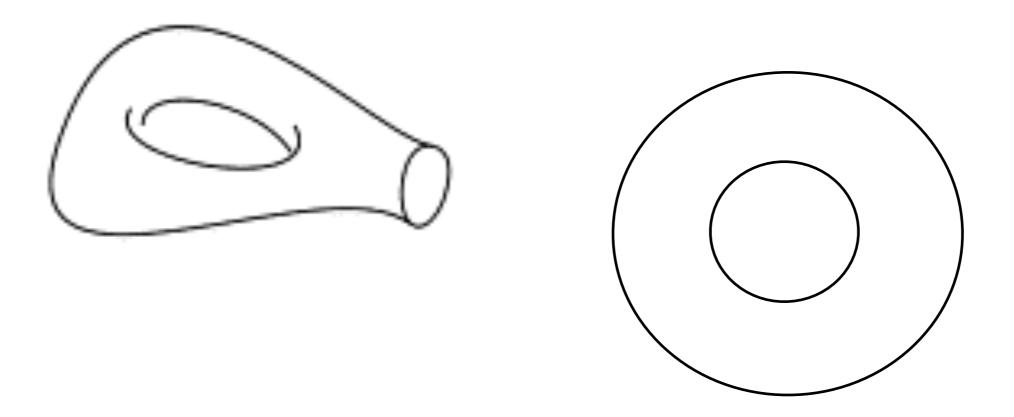
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CONSTRAINT EQUATION (GAUSS LAW)

$$F|_{\Sigma} = 0 \to A = dUU^{-1}$$
 (locally)

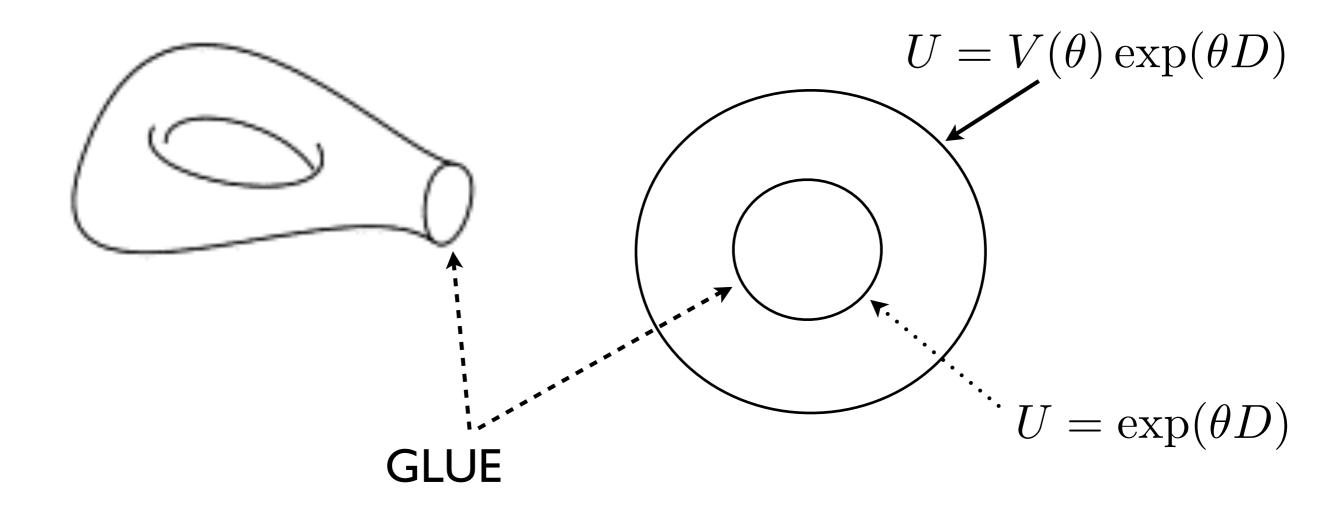
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(IN)FINITE DIMENSIONAL SPACE WITH SEVERAL CONNECTED COMPONENTS.

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THIS SPACE ADMITS A KAHLER STRUCTURE AND A KAHLER FORM: WE CAN QUANTIZE USING COHERENT STATES

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EXISTENCE OF THIS THEORY STILL DEBATED

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FOR GENERIC μ bad theory with ghosts at $\mu L=1$ ghosts disappear for brown-henneaux boundary conditions

ASYMPTOTIC ALGEBRA IS ONLY ONE COPY OF VIRASORO

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BOUNDARY CONDITIONS CAN BE RELAXED TO GIVE

 $Q(\epsilon^{-}) \neq 0$

BUT THEORY HAS GHOST

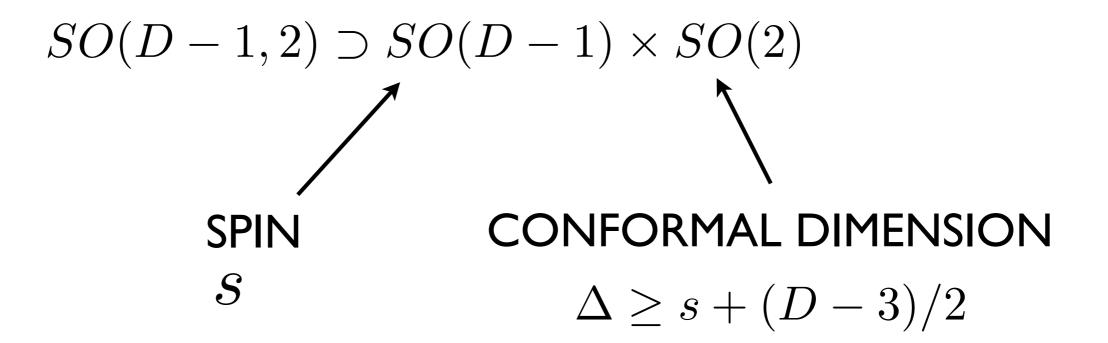
IN ANY DIMENSION D>2 CONFORMAL GROUP IS

 $SO(D-1,2) \supset SO(D-1) \times SO(2)$

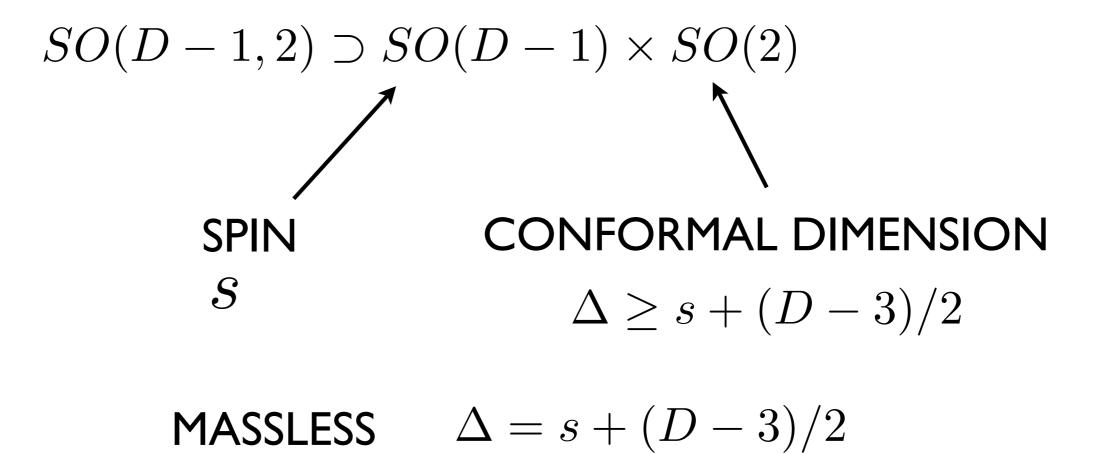
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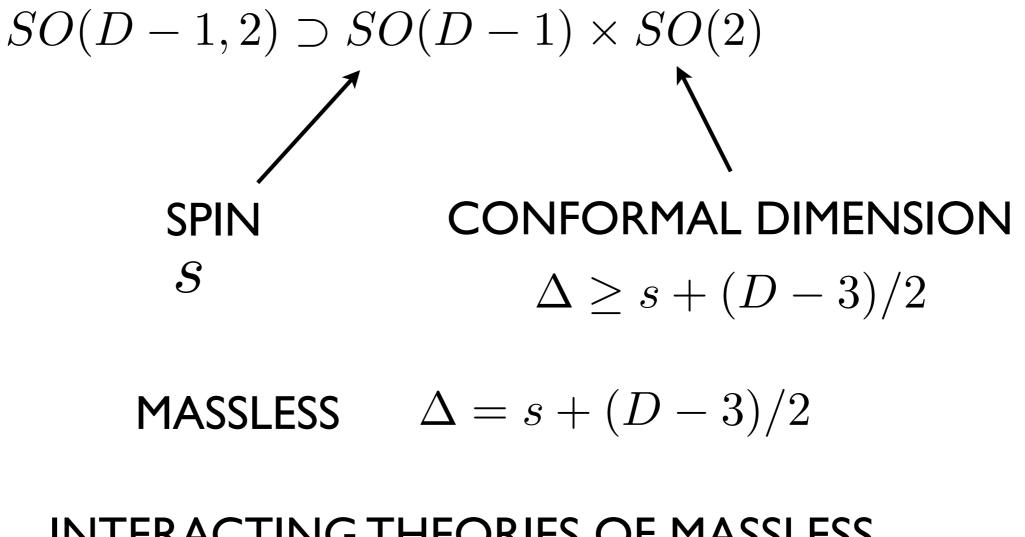
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INTERACTING THEORIES OF MASSLESS HIGH SPIN PARTICLES EXIST IN ADS (VASILIEV)

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MALDACENA-ZHIBOEDOV: IN D>3 SAME AS IN FREE THEORY

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BUT NOT IN D=3!

THEY ARE ALSO ASYMPTOTIC SYMMETRY ALGEBRAS OF CHERN-SIMONS

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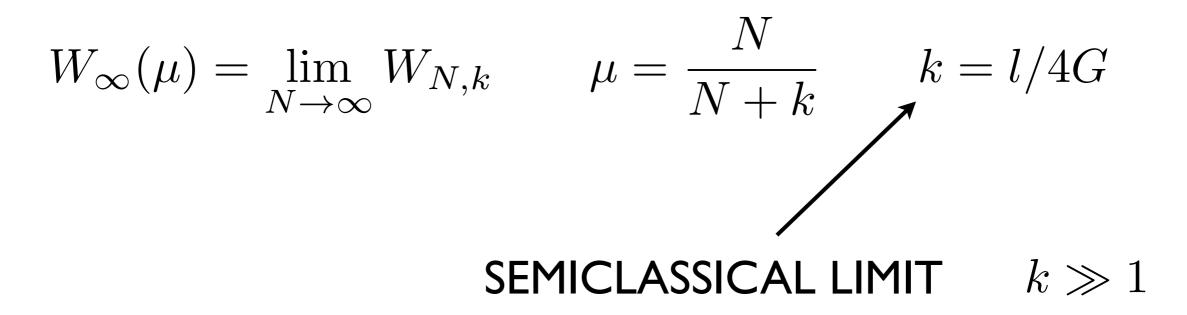
"ANALYTIC CONTINUATION" OF SL(N,R)

$$hs[\mu] = UEA[J^a] / \{J_0^2 - \frac{1}{2}(J_-J_+ + J_+J_-) - (\mu^2 - 1)/4\}$$

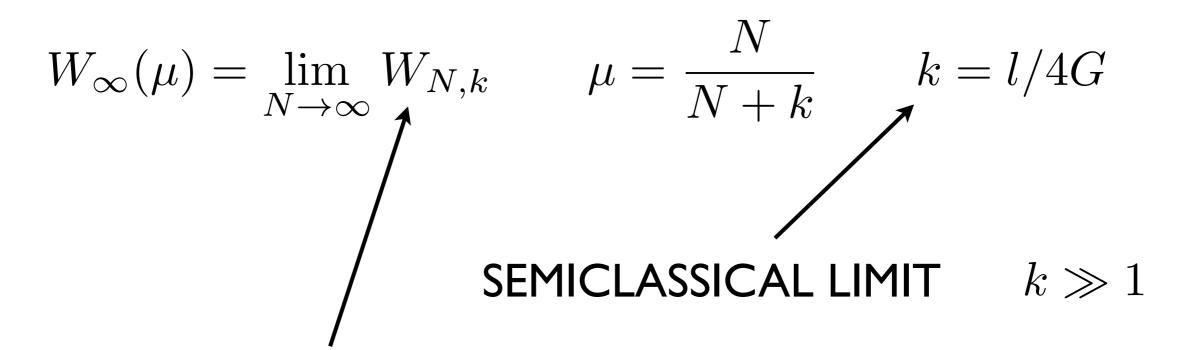
CHERN-SIMONS POSSESS A B-H ASYMPTOTIC ALGEBRA

$$W_{\infty}(\mu) = \lim_{N \to \infty} W_{N,k} \qquad \mu = \frac{N}{N+k} \qquad k = l/4G$$

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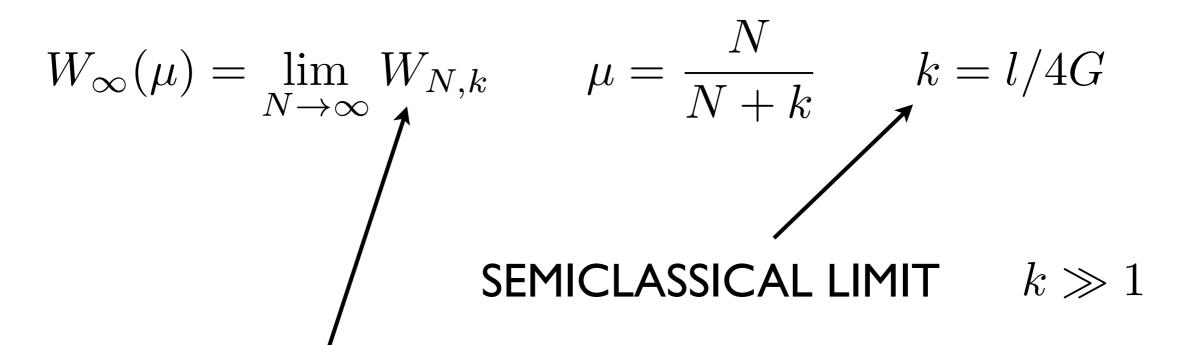


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GABERDIEL-GOPAKUMAR CONJECTURE: THIS IS CFT DUAL TO ADS3 VASILIEV HIGH SPIN THEORY

> TESTS: SPECTRUM OF LIGHT STATES, SYMMETRIES, DEFORMATIONS OK

GRAVITATIONAL WAVES DO NOT PROPAGATE IN ADS3, BUT IT IS NEVERTHELESS A PRECIOUS THEORETICAL LABORATORY FOR QUANTUM GRAVITY BECAUSE IT POSSESSES KEY FEATURES OF 4D GRAVITY, LIKE BLACK HOLES, TOGETHER WITH ESPECIALLY LARGE ASYMPTOTIC ALGEBRAS THAT PARTIALLY REDUCE DYNAMICS TO KINEMATICS