



# Comparing GSI and HIT results

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# Ensure that we are “consistent”

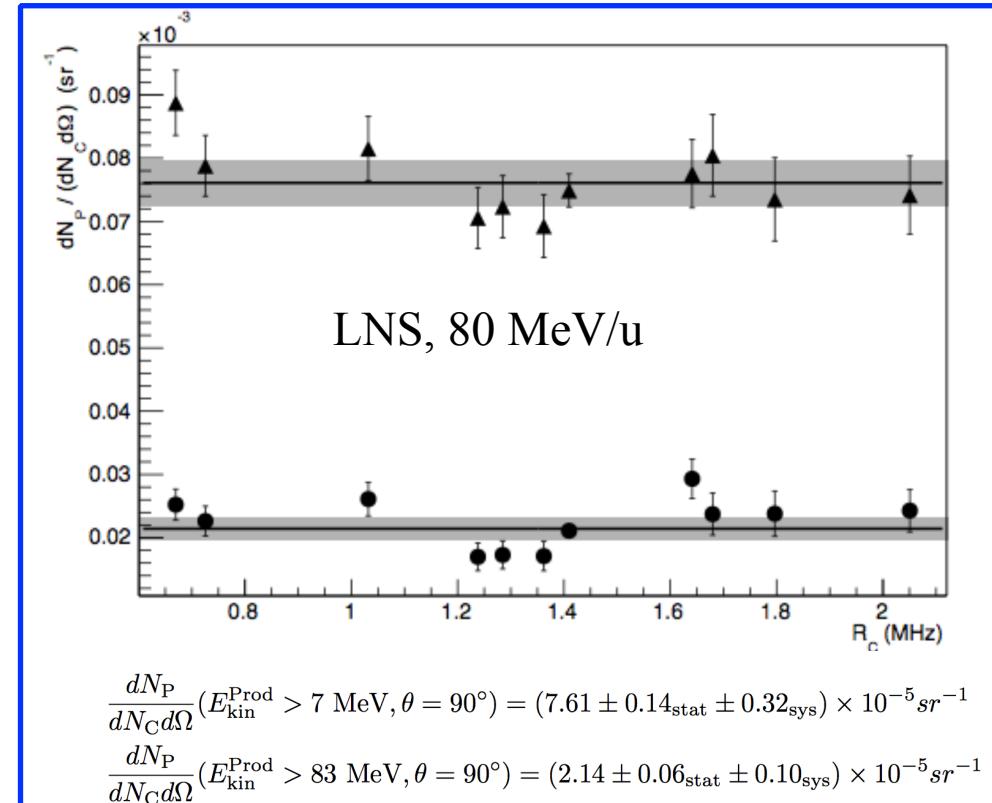
- Both charged and neutral analyses have been redone with  $\sim$  the same setup. A Xchk of the numbers BEFORE publication is crucial.
- For the GSI neutral analysis **we are still in time to catch any mistake:** paper still in preparation.
- The considered inputs are
  - Charged ana:
    - LNS paper (errata)
    - GSI paper
  - Neutrals ana:
    - LNS paper (errata)
    - GSI: preliminary BaF and Lyso papers
    - Ganil latest paper on PG does not provide any flux number...

# Flux of charged particles (OLD)

- What we have published so far:
  - LNS: @ $90^\circ$ , 80MeV/u,  
 $E_{\text{kin}} > 80 \text{ MeV}$ :  $\Phi_p = 2.14 \cdot 10^{-5}$
  - GSI: @ $90^\circ$ ,  $60^\circ$ , 220 MeV/u:
    - $\Phi_p = 8.8 \cdot 10^{-3} (60^\circ)$
    - $\Phi_p = 1.8 \cdot 10^{-3} (90^\circ)$

→ The net effect, going from 80 to 220 MeV/u is to enhance the flux for a factor  $\sim 100$ .

→ The LNS and GSI analyses **did not account for the SC discrimination time correction (Traini's corr.)**: fluxes overestimated



$$\frac{dN_p}{dN_C d\Omega}(E_{\text{kin}}^{\text{Prod}} > 7 \text{ MeV}, \theta = 90^\circ) = (7.61 \pm 0.14_{\text{stat}} \pm 0.32_{\text{sys}}) \times 10^{-5} \text{ sr}^{-1}$$

$$\frac{dN_p}{dN_C d\Omega}(E_{\text{kin}}^{\text{Prod}} > 83 \text{ MeV}, \theta = 90^\circ) = (2.14 \pm 0.06_{\text{stat}} \pm 0.10_{\text{sys}}) \times 10^{-5} \text{ sr}^{-1}$$

$$\frac{dN_p}{N_C d\Omega}(\theta = 60^\circ) = (8.78 \pm 0.07_{\text{stat}} \pm 0.64_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\frac{dN_d}{N_C d\Omega}(\theta = 60^\circ) = (3.71 \pm 0.04_{\text{stat}} \pm 0.37_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\frac{dN_t}{N_C d\Omega}(\theta = 60^\circ) = (0.91 \pm 0.01_{\text{stat}} \pm 0.21_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

GSI

$$\frac{dN_p}{N_C d\Omega}(\theta = 90^\circ) = (1.83 \pm 0.02_{\text{stat}} \pm 0.14_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\frac{dN_d}{N_C d\Omega}(\theta = 90^\circ) = (0.78 \pm 0.01_{\text{stat}} \pm 0.09_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\frac{dN_t}{N_C d\Omega}(\theta = 90^\circ) = (0.128 \pm 0.005_{\text{stat}} \pm 0.028_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

# Flux of charged particles (NEW)

→ Preliminary results from HIT analysis are already available

fluxes at 60deg setup are calculated using efficiency simulations for 90deg setup

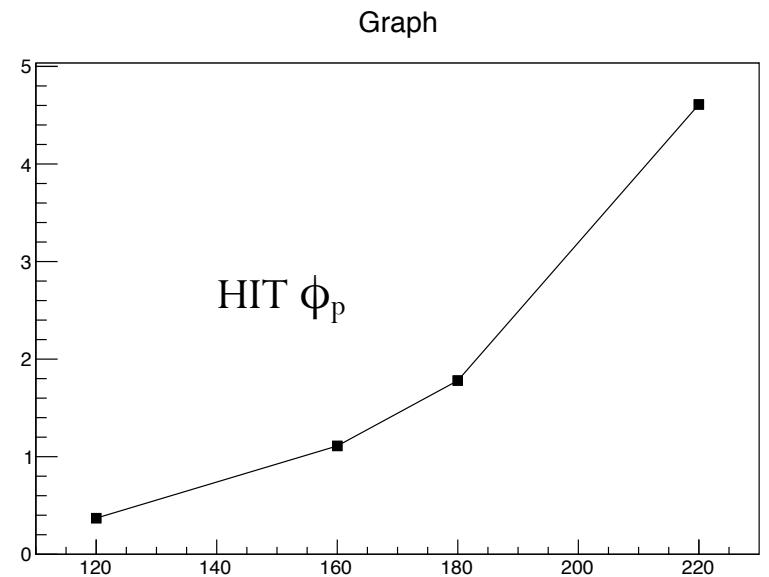
	HIT 90deg	HIT 60 deg	60deg/ 90deg	Sim incl. primaries
Oxy300	$10.66 \times 10^{-3}$	$65.00 \times 10^{-3}$	6.1	
Oxy260	$4.90 \times 10^{-3}$	$36.74 \times 10^{-3}$	7.5	
Oxy210	$2.84 \times 10^{-3}$	$19.88 \times 10^{-3}$	9.9	
C220 (*GSI)	$2.73 \times 10^{-3} <all>$ $1.83 \times 10^{-3} <p>$ $0.78 \times 10^{-3} <d>$ $0.12 \times 10^{-3} <t>$	$13.34 \times 10^{-3} <all>$ $8.78 \times 10^{-3} <p>$ $3.71 \times 10^{-3} <d>$ $0.91 \times 10^{-3} <t>$	4.9	
C220 $<r_{12-14}>$	$5.02 \times 10^{-3}$			
C181	$1.95 \times 10^{-3}$			
C160	$1.24 \times 10^{-3}$			
C120	$0.39 \times 10^{-3}$			
C80 (*Catania)	$0.27 \times 10^{-3}$			
He145	$1.21 \times 10^{-3}$	$14.02 \times 10^{-3}$	11.5	$0.38 \times 10^{-3} (0.02)$
He125	$0.65 \times 10^{-3}$	$9.57 \times 10^{-3}$	14.7	$0.40 \times 10^{-3} (0.02)$
He102	$0.61 \times 10^{-3}$	$4.00 \times 10^{-3}$	6.6	$0.31 \times 10^{-3} (0.02)$

Sys error is of about one order of magnitude lower than obtained flux value.  
Eps\_LYSO(Ekin) is not implemented

# A more detailed comparison

- To check the order of magnitude:
  - Total Np, N12C, Solid angle, efficiencies, Dead Time.

exp. data	$\phi_p @ 90^\circ$	
LNS (80MeV)	$2.15 \cdot 10^{-5}$	no traini's corr
GSI (220MeV)	$1.8 \cdot 10^{-3}$	no traini's corr
HIT C220	$4.61 \cdot 10^{-3}$	with traini's corr
HIT C181	$1.78 \cdot 10^{-3}$	
HIT C160	$1.11 \cdot 10^{-3}$	
HIT C120	$0.37 \cdot 10^{-3}$	



# Neutral analysis (I)

- The prompt photons analysis has already a lot of inputs...
  - LNS (80 MeV/u, 90°), LYSO
  - GSI (220 MeV/u, 90°, 60° and 120°), LYSO and BaF
  - HIT
- Only published result is LNS one and did not account for the SC discrimination time correction (Traini's corr.): flux is likely to be overestimated.

$$\frac{dN_\gamma}{dN_C d\Omega} (E > 2 \text{ MeV}, \theta = 90^\circ) = (2.32 \pm 0.01_{\text{stat}} \pm 0.15_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}.$$

- The GSI LYSO analysis is ∼ completed (traini's Corr not yet applied)

$$\Phi_\gamma(E > 2 \text{ MeV} @ 60^\circ) = (6.59 \pm 0.22_{\text{stat}} \pm 1.07_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\Phi_\gamma(E > 2 \text{ MeV} @ 90^\circ) = (7.39 \pm 0.38_{\text{stat}} \pm 1.27_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

$$\Phi_\gamma(E > 2 \text{ MeV} @ 120^\circ) = (5.02 \pm 0.24_{\text{stat}} \pm 1.34_{\text{sys}}) \times 10^{-3} \text{ sr}^{-1}$$

Preliminary

# Neutral analysis (II)

→ From LNS to GSI..

- The flux increases  $\sim$  a factor 3
- The huge increase in the uncertainties is due to the unfolding procedure!  
[might change now that we are recomputing the unfolding matrix and maybe the unfolding procedure ;) ...]

→ Also GSI BaF results have to be considered.... (traini's Corr not yet applied)

Preliminary BaF, GSI, 220 MeV/u

	60°	90°
Data	$(1.15 \pm 0.11) \times 10^{-2}$	$(1.29 \pm 0.22) \times 10^{-2}$
BIC	$(2.15 \pm 0.02) \times 10^{-2}$	$(1.83 \pm 0.02) \times 10^{-2}$
QMD	$(2.33 \pm 0.03) \times 10^{-2}$	$(1.88 \pm 0.03) \times 10^{-2}$
INCL	$(1.31 \pm 0.02) \times 10^{-2}$	$(1.09 \pm 0.02) \times 10^{-2}$

# From HIT, raw yields

From I. Mattei

CARBON RUNS	eff_DT	N primaries	correction	Nprimaries_corr	Ngamma(2-9MeV)	Ngamma(2-9MeV)/(Nprim_corr*eff_DT)
ProductionRun_230214_Test_Carbon120_19.dat	0,9083	131726755	1,16	152803035,8	5130,260	3,69639E-05
ProductionRun_230214_Test_Carbon160_17.dat	0,8681	577842727	1,10	635626999,7	3591,820	6,50942E-06
ProductionRun_230214_Test_Carbon160_18.dat	0,8655	311660686	1,11	345943361,5	1870,050	6,2457E-06
ProductionRun_230214_Test_Carbon181_15.dat	0,8328	414377928	1,11	459959500,1	3025,240	7,89768E-06
ProductionRun_230214_Test_Carbon181_16.dat	0,8271	162961117	1,11	180886839,9	1226,590	8,1985E-06
ProductionRun_230214_Test_Carbon220_12.dat	0,9022	145602650	1,08	157250862	1407,110	9,91818E-06
ProductionRun_230214_Test_Carbon220_13.dat	0,5439	231759061	1,24	287381235,6	1046,710	6,69652E-06
ProductionRun_230214_Test_Carbon220_14.dat	0,7394	341284877	1,08	368587667,2	2768,890	1,01598E-05

## 12C (\*10<sup>-6</sup>)

$$120\_19 = 36.96$$

$$160\_17, 160\_18 = 6.51, 6.25$$

$$181\_15, 181\_16 = 7.90, 8.20$$

$$220\_12, 220\_13, 220\_14 = 9.92, 6.70, 10.16$$

# HIT raw yields

From I. Mattei

FLUKA parametrization:  $Z_{BP} = -0.8272 + 0.01629 * E_{beam} + 0.0001202 * E^2_{beam}$

	BP position in PMMA $Z_{BP}$ (cm)
$^{12}\text{C}$ 120 MeV/u	2.86
$^{12}\text{C}$ 160 MeV/u	4.86
$^{12}\text{C}$ 181 MeV/u	6.06
$^{12}\text{C}$ 220 MeV/u	8.57

In carbon runs (just 90°)  
the PMMA was always  
10 cm long.

	HIT ( $10^{-6}$ Gamma(2-9MeV)/ $(^{12}\text{C}_{corr} * \text{epsDT})$ )	FLUKA ( $10^{-6}$ Gamma/primary)
$^{12}\text{C}$ 120 MeV/u	36.96	 Sim processed with AnaFluka new
$^{12}\text{C}$ 160 MeV/u	6.25 - 6.51	
$^{12}\text{C}$ 181 MeV/u	7.90 - 8.20	
$^{12}\text{C}$ 220 MeV/u	6.70 - 10.16	11.9(@90°; 13.0@60°)

# The prompt $\gamma$ overview

	(a)	(b)
$N\gamma$		
$N^{12}C$		
$\Omega$		
$\epsilon$		
DT		

exp. data	$\phi @ 90^\circ$	
LNS LYSO (80MeV)	$0.23 \cdot 10^{-2}$	no traini's corr
GSI LYSO (220MeV)	$0.74 \cdot 10^{-2}$	no traini's corr
GSI BAF (220MeV)	$1.15 \cdot 10^{-2}$	no traini's corr

# Comparison with Ganil

→ Phys. Med. Biol. 60 (2015) 565–594

- At the entrance of a PMMA target, where the contribution of secondary nuclear reactions is negligible, prompt-gamma counts per incident ion, per millimetre and per steradian equal to  $(124 \pm 0.7\text{stat} \pm 3\text{sys}) \times 10^{-6}$  for 95 MeV/u carbon ions,  $(79 \pm 2\text{stat} \pm 23\text{sys}) \times 10^{-6}$  for 310 MeV/u carbon ions, and  $(16 \pm 0.07\text{stat} \pm 1\text{sys}) \times 10^{-6}$  for 160 MeV protons were found for prompt gammas with energies higher than 1 MeV
- As such, if an average value between the entrance and the maximum yield immediately upstream of the prompt-gamma falloff is used, conditioned by a 2 MeV energy threshold, an absolute yield of  $(174 \pm 0.9\text{stat} \pm 50\text{sys}) \times 10^{-6}$  counts/ion\*mm\*sr is obtained. The corresponding value published by Agodi and colleagues after considering the projected ion range of 80 MeV/u carbon ions in PMMA (15.42 mm, estimated with SRIM 2013) is  $(150 \pm 0.6\text{stat} \pm 10\text{sys}) \times 10^{-6}$  counts/ion\*mm\*sr. The agreement with previously published results from independent collaborations is remarkably good.
- Problem for GSI?
  - Our friends, going from 95 to 310 are reducing the yield per mm of a factor 0.64. **For us, going from 80 to 220** means a  $\sim *3$  flux and a  $\sim *5.36$  range [assuming 8.25 cm] for a 0.56 net scaling factor.... If we use the BaF result, the reduction is smaller (0.83).. compatible with the lower energy range?