

Higgs and flavour physics near the Fermi scale

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“Rewarding science”, LNF Frascati, 24 June 2015



The Standard Model



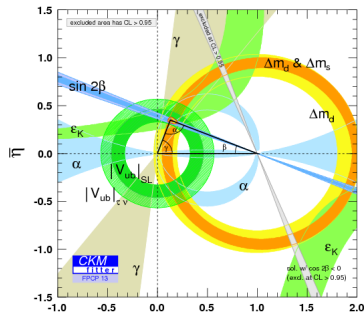
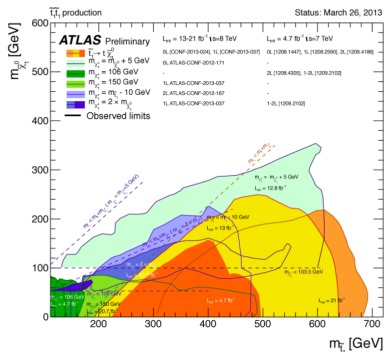
? The Standard Model ?

? Supersymmetry ? ??? ? Composite models ?

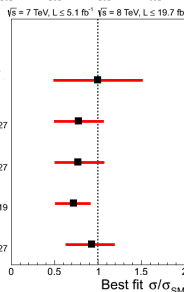
Where is New Physics? Experiments

Neat indications of NP:

Gravity, Dark Matter, ν , ... but:



- CMS Preliminary
 Individual Results
- $V H \rightarrow b\bar{b}$ arXiv:1310.3667
 $\mu(m_H=125.0 \text{ GeV}) = 1.0 \pm 0.5$
 - $H \rightarrow \tau\tau$ arXiv:1401.5041
 $\mu(m_H=125.0 \text{ GeV}) = 0.78 \pm 0.27$
 - $H \rightarrow \gamma\gamma$ HIG-13-001
 $\mu(m_H=125.0 \text{ GeV}) = 0.78 \pm 0.27$
 - $H \rightarrow WW$ arXiv:1312.1129
 $\mu(m_H=125.6 \text{ GeV}) = 0.72 \pm 0.19$
 - $H \rightarrow ZZ$ arXiv:1312.5353
 $\mu(m_H=125.6 \text{ GeV}) = 0.93 \pm 0.27$



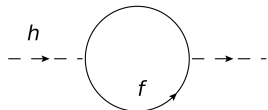
$$m_h \approx \Lambda$$

[Λ = highest scale h couples to, e.g. M_{Planck}]

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Λ = "cutoff"? Misleading: SM is renormalizable and divergences do not appear.

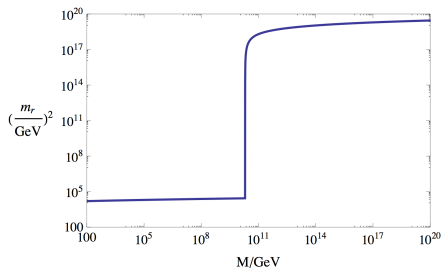


$$\mathcal{L} \supset y h \bar{f} f \Rightarrow \frac{dm_h^2}{d \log \mu} = -\frac{3y^2}{4\pi^2} M_f^2$$

$$m_h^2(m_h^2) \simeq m_h^2(M_{NP}) - a M_{NP}^2 \log \frac{M_{NP}^2}{m_h^2}$$

$$\text{Fine tuning } \Delta \simeq a M_{NP}^2 / m_h^2$$

[e.g. $a \propto y^2$]



Hierarchy Problem: initial condition $m_h^2(M_{NP})$ to be chosen with precision $\sim 1/\Delta$

Physics at the Fermi scale depends on details of way shorter distances!

Where is New Physics? Theory

Answer \simeq attitude towards the hierarchy problem

- 1 **Protect the mass of the scalars** from any NP [’t Hooft 1979, ...]

The Fermi scale is “natural” small $\Delta \Rightarrow M_{NP} \lesssim \text{TeV}$

Examples: Supersymmetry composite Higgs models

No evidence for NP puts pressure on this “standard” attitude

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2 Short distance assumptions

M_{NP} can be $\gg \text{TeV}$

[W. Bardeen, Foot, Shaposhnikov, Dubovsky, Strumia,...]

$\Delta \simeq a M_{NP}^2 / m_h^2$ with $a \ll 1$ (delicate, e.g. gravity?)

3 Accept a large tuning

M_{NP} can be $\gg \text{TeV}$

e.g. anthropic selection of parameters [Weinberg 1987, Agrawal et al 1997, ...]

General motivation

Look for “natural” New Physics at current and future experiments

Where?

- Higgs (SM and new ones) Barbieri Buttazzo Kannike S Tesi 1304.3670
“ “ 1307.4937
- Flavour Barbieri Campli Isidori S Straub 1108.5125
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This seminar

Out: a lot, e.g. composite Higgs models, vacuum stability

In: more recent developments of above topics

Can new “Higgses” be the lightest new particles around?

Yes, in various natural models. For example in Supersymmetry

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MSSM

$$m_h^2 \leq m_Z^2 \cos^2 2\beta + \Delta_t^2 \Rightarrow \Delta_t \gtrsim 85 \text{ GeV} \Rightarrow \text{stops heavier than } \sim 1.5 \text{ TeV}$$

Fine tuning worse than 1%!

$$\frac{dv^2}{dm_{H_u}^2} \simeq \frac{4}{g^2}$$

$$\delta m_{H_u}^2 \sim -\frac{3y_t^2}{4\pi^2} m_{\tilde{t}}^2 \log \frac{\Lambda}{m_{\tilde{t}}}$$

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NMSSM

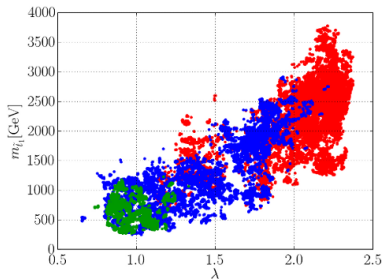
Extra singlet S $\Delta W = \lambda S H_u H_d + f(S)$

$$m_h^2 \leq m_Z^2 c_{2\beta}^2 + \Delta_t^2 + \lambda^2 v^2 s_{2\beta}^2$$

Fine tuning better than 5%!

[green points, $\tan \beta \lesssim 5$]

$$\frac{dv^2}{dm_{H_u}^2} \simeq \frac{\kappa}{\lambda^3} \frac{1}{t_{2\beta}}$$



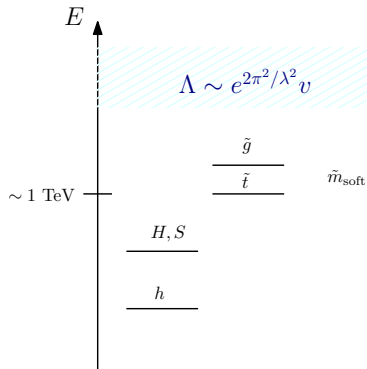
Gherghetta et al. 1212.5243 [$\Lambda = 20 \text{ TeV}$]

Model, and plan of the next 6 slides

NMSSM with $\lambda \sim 1$ and heavy stops & gluinos

[$\lambda \gtrsim 0.7$ needs a completion before the GUT scale]

Goal = strategy to look for the extra Higgses



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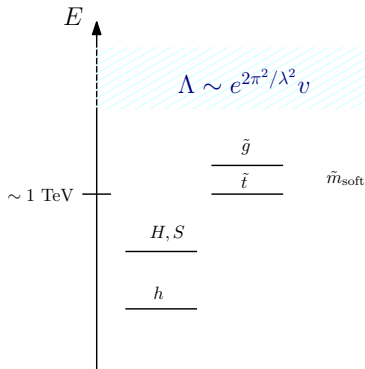
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They are:

CP-even h_1, h_2, h_3 (from h, S, H)

CP-odd A, A_5

H^\pm



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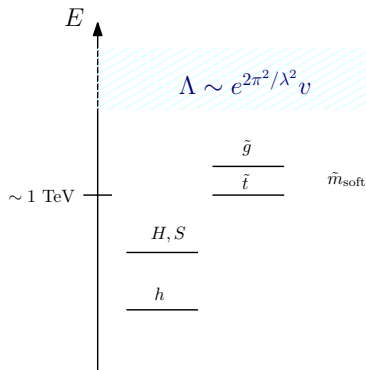
They are:

CP-pari h_1, h_2, h_3 (da h, S, H)

Assumptions

Only loop contribution = top-stop Δ_t

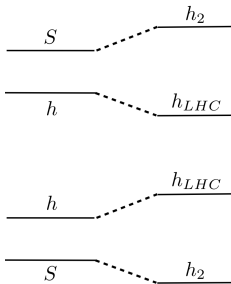
No invisible decays $h_1, h_2 \rightarrow \chi\chi, \dots$



$$\mathcal{H}_{\text{ph}} \equiv \begin{pmatrix} h_3 \\ h_1 \\ h_2 \end{pmatrix} = R^T \begin{pmatrix} H \\ h \\ S \end{pmatrix}, \quad R = R_\delta^{12} R_\gamma^{23} R_\sigma^{13}$$

$$m_{h_1} = 125 \text{ GeV}$$

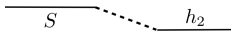
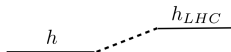
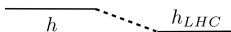
$$M^2 = \begin{pmatrix} \widetilde{M}^2(m_{H^\pm}^2, \lambda, t_\beta, \Delta_t) & & \\ vM_1 & vM_2 & M_3^2 \end{pmatrix} = R^T \text{diag}(m_{h_3}^2, m_{h_1}^2, m_{h_2}^2) R$$



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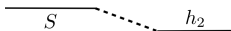
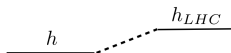
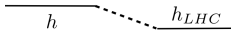
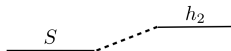
Analytic relations!

$$\delta, \gamma, \sigma(m_{h_2}, m_{h_3}, m_{H^\pm}, \lambda, t_\beta, \Delta_t)$$

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A motivated limiting case

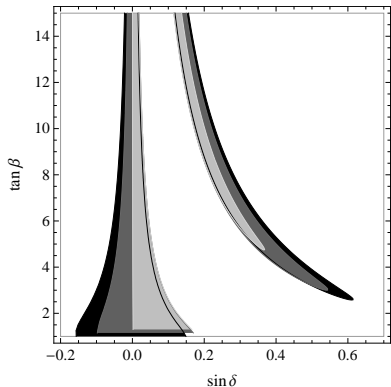
$$m_{h_3} \gg m_{h_{1,2}} \text{ and } \sigma, \delta \rightarrow 0 \quad [\text{free pars: } m_{h_2}, t_\beta, \Delta_t, \lambda]$$

Higgs couplings and fit

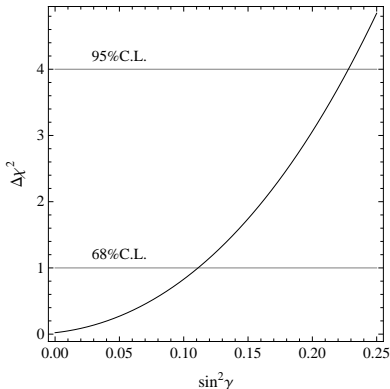
$$[h_{\text{LHC}} = h_1 = c_\gamma(c_\delta h - s_\delta H) + s_\gamma S]$$

$$\frac{g_{h_1 tt}}{g_{htt}^{\text{SM}}} = c_\gamma \left(c_\delta + \frac{s_\delta}{\tan \beta} \right), \quad \frac{g_{h_1 bb}}{g_{hbb}^{\text{SM}}} = c_\gamma (c_\delta - s_\delta \tan \beta), \quad \frac{g_{h_1 VV}}{g_{hVV}^{\text{SM}}} = c_\gamma c_\delta$$

LHC8 status



$$s_\gamma^2 = 0, 0.15, 0.3$$



Higgs couplings and fit

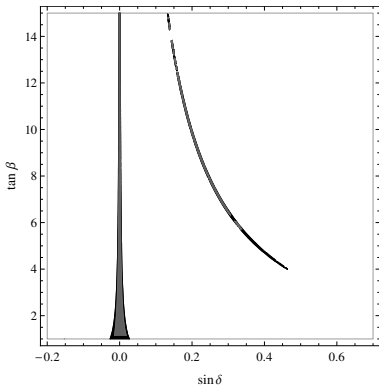
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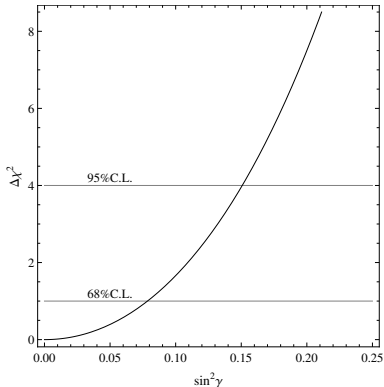
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LHC14 (300 fb⁻¹) projections



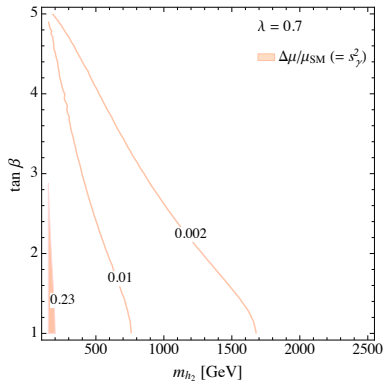
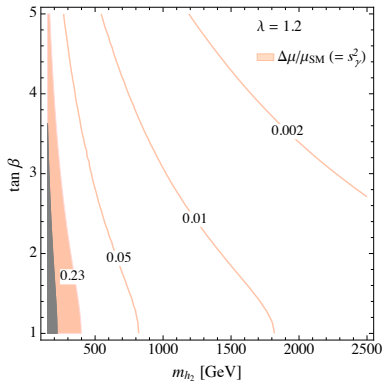
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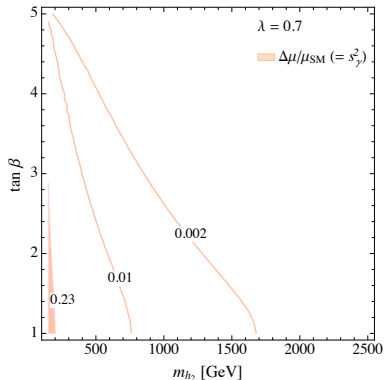
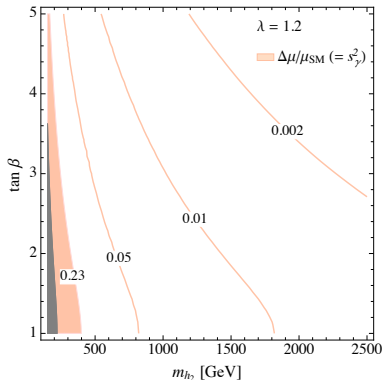
$[\Delta_t = 70 \div 80 \text{ GeV, influent}]$



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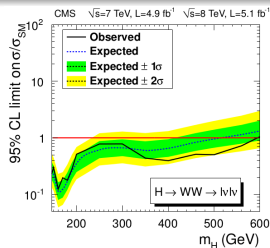
Direct searches of h_2 : $m_{h_2} > 250 \text{ GeV}$: VV dominates [second channel hh]

$m_{h_2} < 250 \text{ GeV}$: signal = $s_\gamma^2 \times \text{SM}$

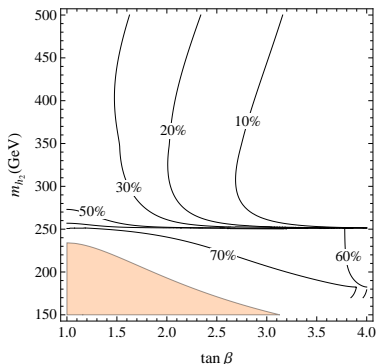
h_2 at the LHC and beyond

In 2013 everybody was looking for top partners!

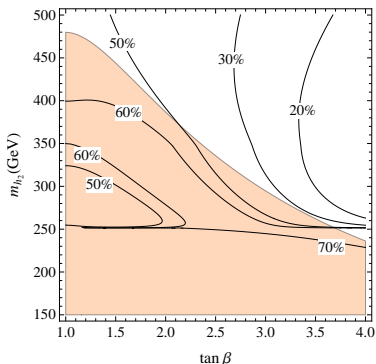
Only available searches in VV :



$\lambda = 0.8$



$\lambda = 1.4$

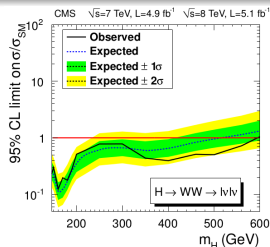


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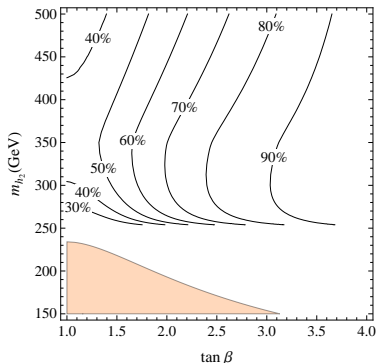
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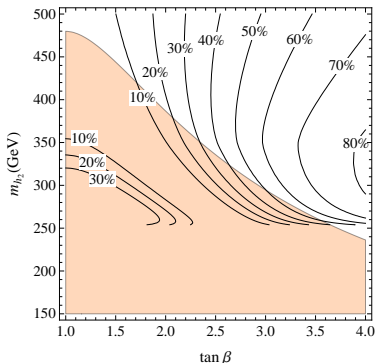
Also $h_2 \rightarrow hh$ seemed promising:



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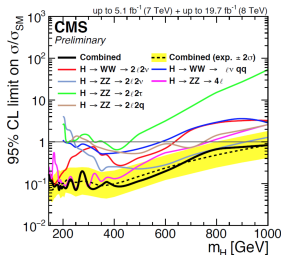


Now plenty of scalar searches!

General validity, NMSSM particular application:

$$\sin^2 \gamma = \frac{M_{hh}^2 - m_h^2}{m_{h_2}^2 - m_h^2}, \quad \text{only 2 free pars.!}$$

$$M_{hh}^2 = m_Z^2 c_{2\beta}^2 + \frac{\lambda^2 v^2}{2} s_{2\beta}^2 + \Delta_t^2, \quad \Delta_t \text{ little impact}$$

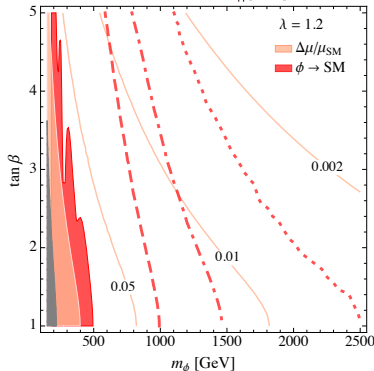
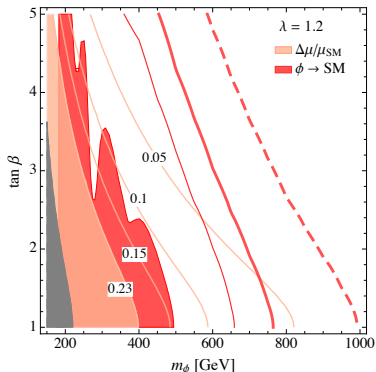
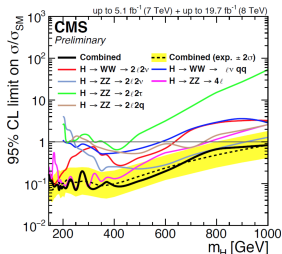


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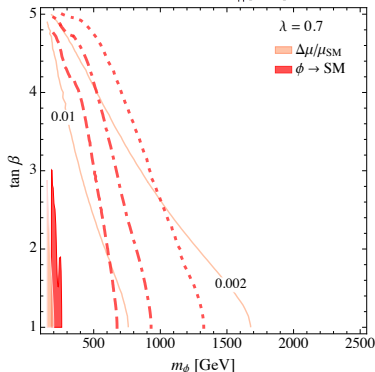
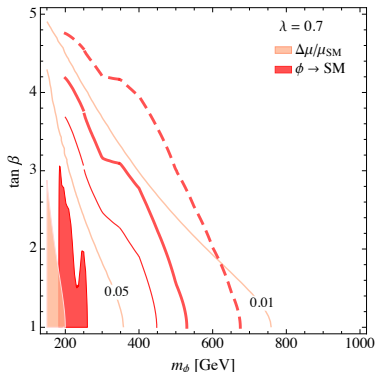
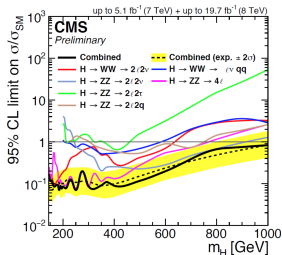


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Higgses self couplings

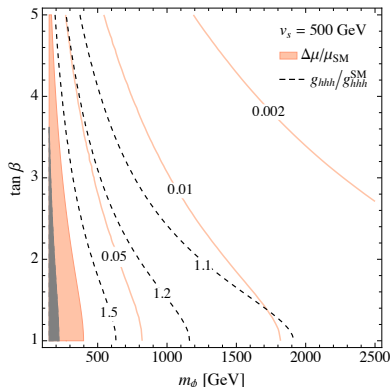
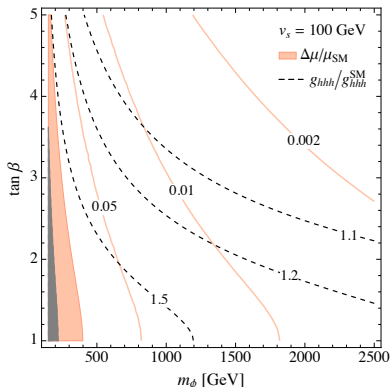
Actually all V pars. matter for triple Higgs couplings...but dominant one is $v_s!$
[Buttazzo S Tesi 1505.05488]

- ◇ $h_2 \rightarrow hh$ typically less important than $h_2 \rightarrow VV$, especially for $m_{h_2} \gg m_W$

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- ◇ $h_2 \rightarrow hh$ typically less important than $h_2 \rightarrow VV$, especially for $m_{h_2} \gg m_W$
- ◇ g_{hhh} : deviations expectable at the HL-LHC!



HL-LHC sensitivity: $\sim 50\%$ [from Snowmass report Dawson et al 1310.8361]

Flavour and $U(2)^3$

Flavour: motivation and $U(2)^3$

Flavour and the SM $|V_{\text{CKM}}| \sim \begin{pmatrix} 1 & 0.2 & 4 \cdot 10^{-3} \\ 0.2 & 1 & 4 \cdot 10^{-2} \\ 9 \cdot 10^{-3} & 4 \cdot 10^{-2} & 1 \end{pmatrix}$

$(y_u, y_c, y_t) \sim (10^{-6}, 10^{-2}, 1)$ $(y_d, y_s, y_b) \sim (10^{-5}, 10^{-3}, 10^{-2})$

Flavour and NP $\mathcal{L}_{\text{NF}} = \sum_i \frac{1}{\Lambda_i^2} \mathcal{O}_i \Rightarrow \Lambda_i \gtrsim 10^4 \div 10^5 \text{ TeV}$

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Flavour and NP $\mathcal{L}_{\text{NF}} = \sum_i \frac{1}{\Lambda_i^2} \mathcal{O}_i \Rightarrow \Lambda_i \gtrsim 10^4 \div 10^5 \text{ TeV}$

Possible solution: $\mathcal{L}_{\text{NP}} = \sum_i \xi_i \frac{c_i}{\Lambda_i^2} \mathcal{O}_i \quad c_i \sim \mathcal{O}(1)$

ξ_i small thanks to some “propriety” (e.g. a **flavour symmetry**)

Flavour: motivation and $U(2)^3$

Flavour and the SM $|V_{CKM}| \sim \begin{pmatrix} 1 & 0.2 & 4 \cdot 10^{-3} \\ 0.2 & 1 & 4 \cdot 10^{-2} \\ 9 \cdot 10^{-3} & 4 \cdot 10^{-2} & 1 \end{pmatrix}$

$$(y_u, y_c, y_t) \sim (10^{-6}, 10^{-2}, 1) \quad (y_d, y_s, y_b) \sim (10^{-5}, 10^{-3}, 10^{-2})$$

Flavour and NP $\mathcal{L}_{\text{NF}} = \sum_i \frac{1}{\Lambda_i^2} \mathcal{O}_i \Rightarrow \Lambda_i \gtrsim 10^4 \div 10^5 \text{ TeV}$

Possible solution: $\mathcal{L}_{\text{NP}} = \sum_i \xi_i \frac{c_i}{\Lambda_i^2} \mathcal{O}_i \quad c_i \sim \mathcal{O}(1)$

ξ_i small thanks to some “propriety” (e.g. a **flavour symmetry**)

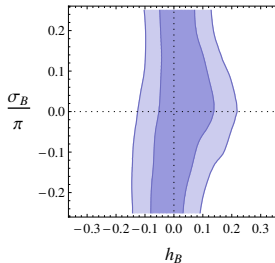
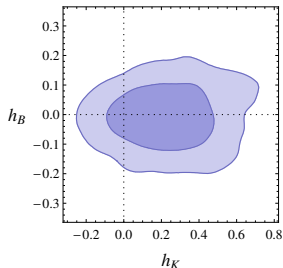
$$U(2)^3 = U(2)_{Q_L} \times U(2)_{U_R} \times U(2)_{D_R}$$

Barbieri Isidori et al 1105.2296

- ✓ $\xi \sim V_{CKM}^{2 \div 4} \Rightarrow \Lambda \sim \text{a few TeV}$ is OK with flavour constraints
- ✓ interesting phenomenology behind the corner
- ✓ separate 3rd generation from the first two \rightarrow ideal for natural theories!

$$\Delta F = 2$$

$$h_{K,B} \simeq c_{K,B} \left(\frac{3\text{TeV}}{\Lambda} \right)^2$$



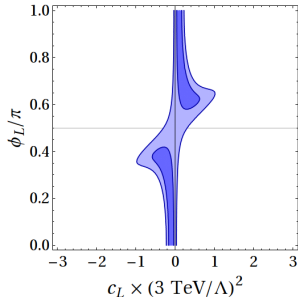
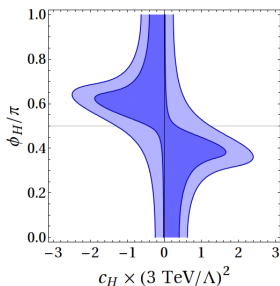
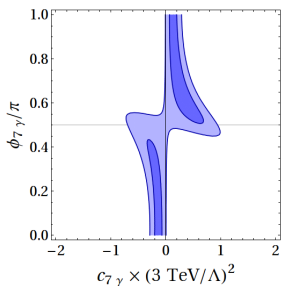
Take-home messages

- Data consistent with $\mathcal{L}_{\text{NP}} = \sum_i \xi_i \frac{c_i}{\Lambda_i^2} \mathcal{O}_i$ and $|c_i| = 0.2 \div 1$

- Larger effects allowed than in other models (e.g. $U(3)^3$)
see also [Buras, Girschbach 1206.3878](#)

$$\Delta F = 1$$

$$b \rightarrow s(d)l\bar{l}, b \rightarrow s(d)\nu\bar{\nu}$$



Take-home messages

- Data consistent with $\mathcal{L}_{\text{NP}} = \sum_i \xi_i \frac{c_i}{\Lambda_i^2} \mathcal{O}_i$ and $|c_i| = 0.2 \div 1$

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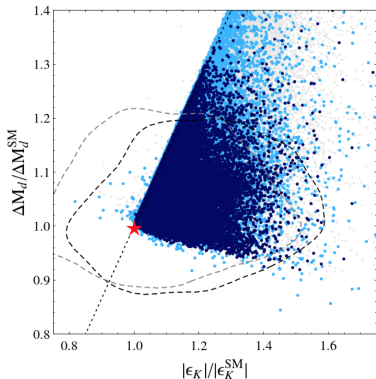
see also [Buras, Girschbach 1206.3878](#)

$U(2)^3 + \text{SUSY}: \Delta F = 2$ vs the LHC

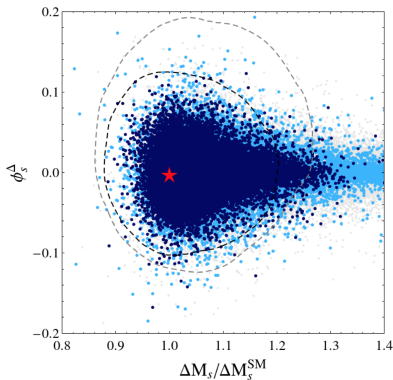
All points: allowed by LHC8 searches

Darker: conservative exclusions
Lighter: compressed spectra, ...

[Dashed: $\Delta F = 2$ fit]



Barbieri Buttazzo S Straub 1402.6677

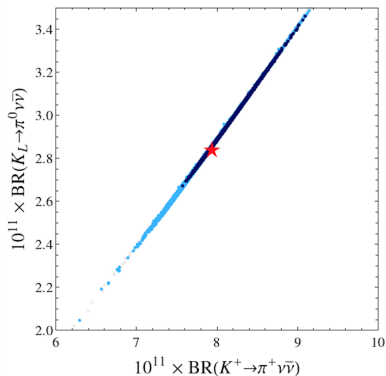
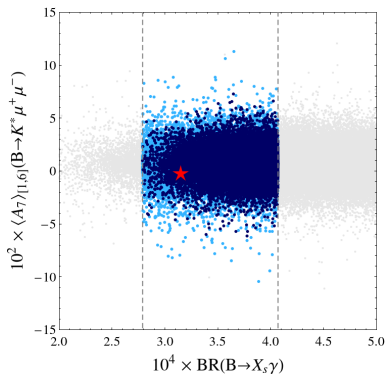


What if no particles observed at the LHC14?

Then only promising observables: ϕ_s $\Delta M_{d,s}$

$U(2)^3 + \text{SUSY}: \Delta F = 1$ vs the LHC

Barbieri Buttazzo S Straub 1402.6677



Target for LHCb: $A_7 \sim 2 - 3\%$

NA62 will measure $\text{BR}_{K^+ \rightarrow \pi^+ \nu \bar{\nu}}$ at $\sim 10\%$

Specific correlation could allow to distinguish $U(2)^3$ from other models

Higgs and flavour measurements: what do they imply for natural New Physics?

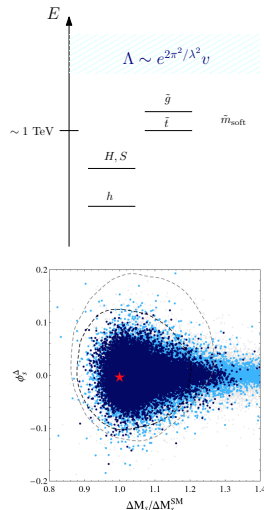
Higgs and flavour measurements: what do they imply for natural New Physics?

→ The NMSSM Higgs sector

$$[h_2 \rightarrow VV, hh, h \text{ couplings}, g_{hhh}]$$

→ A $U(2)^3$ symmetry & the flavour of NP

$$[\phi_s, |V_{ub}|, b \rightarrow s(d)\ell\bar{\ell}, \nu\bar{\nu}, \dots]$$



Higgs and flavour measurements: what do they imply for natural New Physics?

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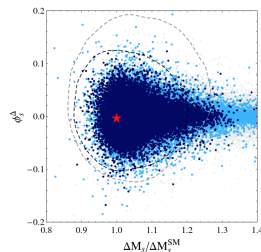
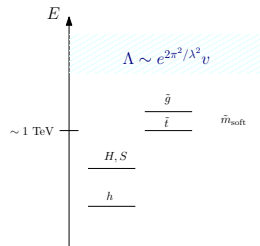
$$[\phi_s, |V_{ub}|, b \rightarrow s(d)\ell\bar{\ell}, \nu\bar{\nu}, \dots]$$

What next? A lot!

Flavour: improve on SM predictions ($\epsilon_K, \epsilon'_K, \dots$)

Higgses: phenomenology of CP-odd ones, H, \dots

NP unrelated to Hierarchy Problem: Dark Matter, ...



Many thanks to...



Riccardo Barbieri

Dario Buttazzo



Andrea Tesi



David M. Straub



Kristjan Kannike

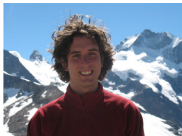


Many thanks to...



Riccardo Barbieri

Dario Buttazzo



Andrea Tesi



David M. Straub



Kristjan Kannike



And to the INFN for the Fubini prize!

Back up Higgses

Extrapolation of direct searches

We started from (and improved)

i) Collider Reach (β) Salam Weiler 2014 ii) Thamm Torre Wulzer 1502.01701

m_0 excluded at LHC8, obtain m_1 at future collider via $B(s_1, L_1, m_1) = B(s_0, L_0, m_0)$

$$B(s, L, m) \propto L \times \int d\hat{s} \frac{1}{\hat{s}} \hat{s} \hat{\sigma}(\hat{s}) \frac{d\mathcal{L}}{d\hat{s}}(s)$$

Assumptions/limitations

→ Not valid if systematics dominate and change significantly from 0 to 1

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$$B(s, L, m) \propto L \times \frac{\Delta \hat{s}}{m^2} c \frac{d\mathcal{L}}{d\hat{s}}(s) \Big|_{\hat{s}=m^2} \quad \hat{s}\hat{\sigma}(\hat{s}) = c \Rightarrow \frac{d\mathcal{L}}{d\hat{s}} \text{ drives the reach}$$

$$L_1 c_{ij} \frac{d\mathcal{L}_{ij}}{d\hat{s}}(s_1) \Big|_{\hat{s}=m_1^2} = L_0 c_{ij} \frac{d\mathcal{L}_{ij}}{d\hat{s}}(s_0) \Big|_{\hat{s}=m_0^2}$$

Assumptions/limitations

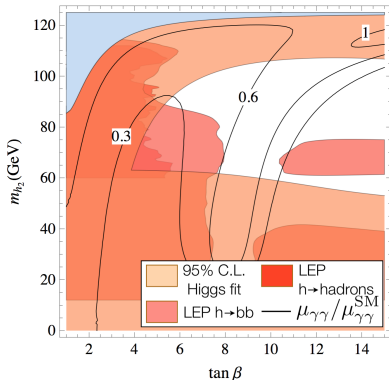
- Not valid if systematics dominate and change significantly from 0 to 1
- $\hat{s} \gg m_{\text{bkg}}$ [i.e. not valid at $\hat{s} \sim 2m_t$ for $h_2 \rightarrow hh(4b)$]
- $\frac{\Delta \hat{s}}{m^2} \ll 1$ i.e. not valid if analysis depends a lot on shape far from peak

Fully mixed case and a $\gamma\gamma$ signal

Singlet-like state lighter than 125 GeV

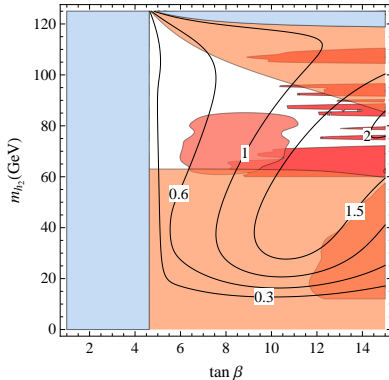
hard to see, but exploration already started

$\lambda = 0.1, \Delta_t = 85 \text{ GeV}$



$[m_{h_3} = 500 \text{ GeV}, s_\sigma^2 = 10^{-3}, v_s = v]$

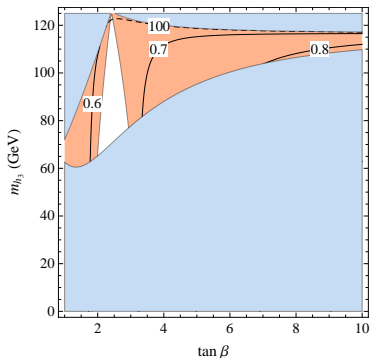
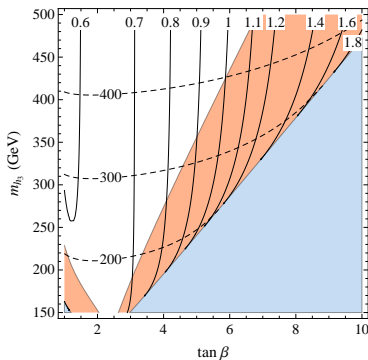
$\lambda = 0.8, \Delta_t = 75 \text{ GeV}$



see e.g. [Badziak et al. 1304.5437,...](#)

$$\frac{g_{h_3 tt}}{g_{htt}^{\text{SM}}} = s_\delta - \frac{c_\delta}{t_\beta} \quad \frac{g_{h_3 bb}}{g_{hbb}^{\text{SM}}} = s_\delta + t_\beta c_\delta \quad \frac{g_{h_3 VV}}{g_{hVV}^{\text{SM}}} = s_\delta \quad [\Delta_t = 75 \text{ GeV}]$$

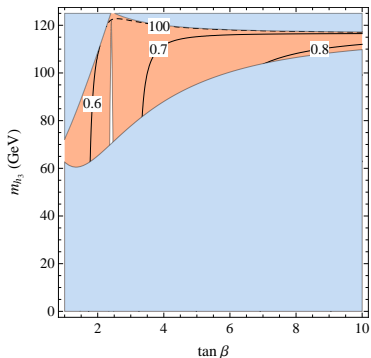
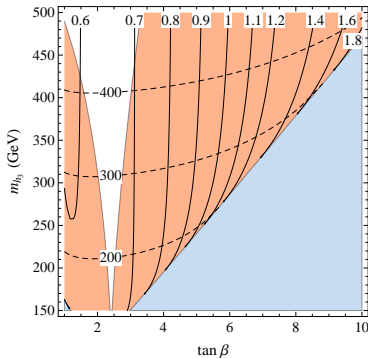
Status fit LHC8:

 $m_{H^\pm} > 480 \text{ GeV}$ from $B \rightarrow X_s \gamma$!dashed: m_{H^\pm} cont: λ 

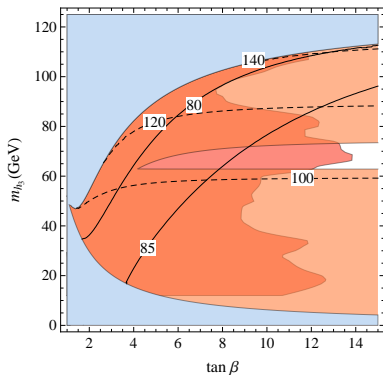
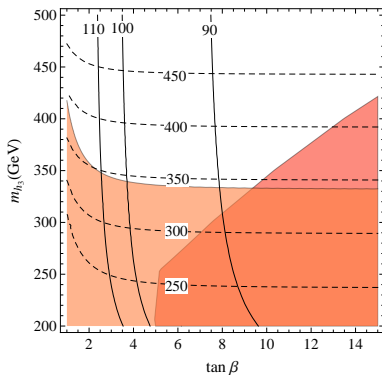
$$[\widetilde{\mathcal{M}}_{12}^2(t_\beta, \dots) = 0 \rightarrow \delta = 0]$$

 h_3 phenomenology: more similar to MSSM

$$\frac{g_{h_3 tt}}{g_{htt}^{\text{SM}}} = s_\delta - \frac{c_\delta}{t_\beta} \quad \frac{g_{h_3 bb}}{g_{hbb}^{\text{SM}}} = s_\delta + t_\beta c_\delta \quad \frac{g_{h_3 VV}}{g_{hVV}^{\text{SM}}} = s_\delta \quad [\Delta_t = 75 \text{ GeV}]$$

Projections fit LHC14 (300 fb^{-1}):dashed: m_{H^\pm} cont: λ  $m_{H^\pm} > 480 \text{ GeV}$ from $B \rightarrow X_s \gamma$! $[\widetilde{\mathcal{M}}_{12}^2(t_\beta, \dots) = 0 \rightarrow \delta = 0]$ h_3 phenomenology: more similar to MSSM

Status fit LHC8:

[dashed: m_{H^\pm} cont: Δt]

Red regions excluded by direct searches at LEP and CMS

Projections fit LHC14: above regions completely excluded

[if $\frac{\mu A_t}{m_{\tilde{t}}^2}$ very large, conclusions could change...]

Back up Flavour

Modelling $U(2)^3 = U(2)_{Q_L} \times U(2)_{U_R} \times U(2)_{D_R}$

$$U(2)^3 \text{ exact} \quad \longrightarrow \quad m_u = m_d = m_s = m_c = 0, \quad V_{CKM} = 1$$

$$Y_u = y_t \left(\begin{array}{c|c} 0 & 0 \\ \hline 0 & 1 \end{array} \right) \quad Y_d = y_b \left(\begin{array}{c|c} 0 & 0 \\ \hline 0 & 1 \end{array} \right)$$

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- $\Delta Y_u \sim (2, \bar{2}, 1)$, $\Delta Y_d \sim (2, 1, \bar{2})$ to obtain quark masses

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- $\Delta Y_u \sim (2, \bar{2}, 1)$, $\Delta Y_d \sim (2, 1, \bar{2})$ to obtain quark masses
- Minimal $U(2)^3$: only 1 doublet $V \sim (2, 1, 1)$ to explain CKM

$V_{CKM}(s_u, s_d, \delta, \epsilon_L) \rightarrow$ all **4** physical pars. from tree-level observables!

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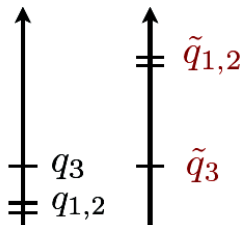
$V_{CKM}(s_u, s_d, \delta, \epsilon_L) \rightarrow$ all 4 physical pars. from tree-level observables!

Assumption: all flavour violation controlled by $\Delta Y_{u,d}, V$

i.e. \mathcal{L}_{NP} built with bilinears like $\bar{q}_L V \gamma_\mu q_{3L}$, $\bar{q}_L \Delta Y_d d_R$

Example: $\mathcal{L}_{NP} \supset \frac{c_L^B e^{i\phi_B}}{\Lambda^2} (\mathbf{V}_{tb} \mathbf{V}_{ti}^*)^2 (\bar{d}_L^i \gamma_\mu b_L)^2, \quad i = d, s \quad [B_{d,s}^0 - \bar{B}_{d,s}^0]$

Supersymmetric realisation of $U(2)^3$

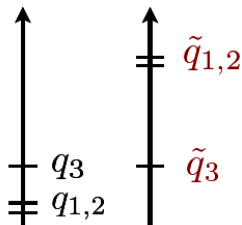


SUSY with heavy 1, 2 generations

- ✓ flavour-blind CP violation (EDMs)
- ✓ ok with collider bounds

Impossible in $U(3)^3$!

Supersymmetric realisation of $U(2)^3$



SUSY with heavy 1, 2 generations

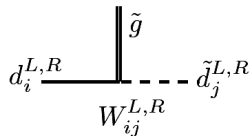
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- ✓ ok with collider bounds

Impossible in $U(3)^3$!

$$\mathcal{L}_{F\text{-breaking}} \sim \tilde{q}^\dagger \tilde{m}^2(\Delta Y, V) \tilde{q}$$

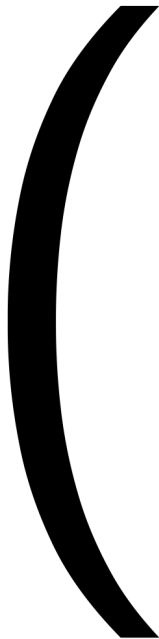
$$W^L \simeq \begin{pmatrix} c_d & \kappa^* & -\kappa^* s_L e^{i\gamma} \\ -\kappa & c_d & -c_d s_L e^{i\gamma} \\ 0 & s_L e^{-i\gamma} & 1 \end{pmatrix}$$

$$W^R \simeq 1 + O(y_s/y_b) \quad \kappa = s_d e^{i\beta}$$



$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} \left(1 + F_{\tilde{g}} + F_{H^\pm} + F_{\tilde{H}^\pm} + F_{\tilde{W}} + F_{\tilde{B}} \right)$$

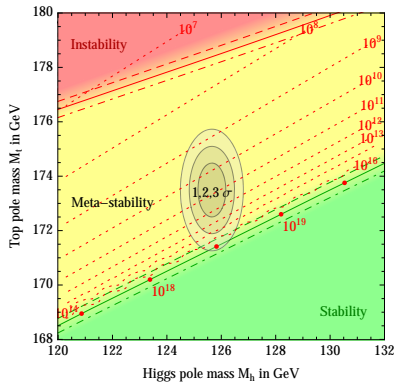
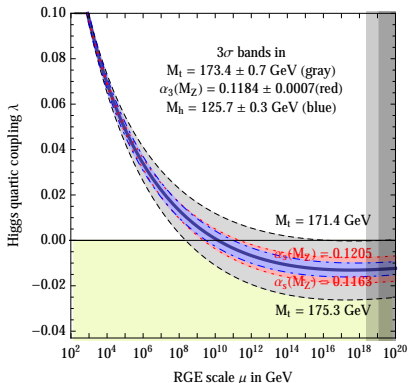
Back up Vacuum stability



Other *neat* indications of a NP scale?

SM vacuum is metastable \Rightarrow does not require NP!

[DM and neutrinos can be included without further destabilising]

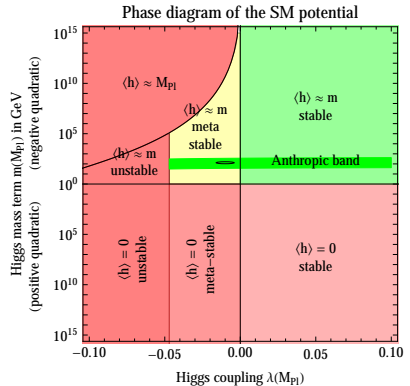
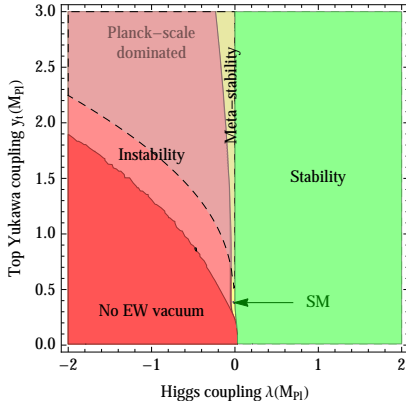


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But: λ and y_t are “critical”, accident or deep meaning?

