JLab12 status and activities

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JLab from 6 to 12 GeV



Nucleon structure functions

How can we build a multidimensional picture of the nucleon ? Non-perturbative QCD is complicated

- Lattice QCD
- Phenomenological description → structure functions



- Elastic Form Factors, Parton Distributions Functions
- Nucleon charges
- Parton Angular momentum

GPD measurements





 $IM(H) \rightarrow$ unpol. \rightarrow electric charge

- growing as $t \rightarrow 0$
- FT → spread in the nucleon volume

 $IM(\widetilde{H}) \rightarrow$ helicity \rightarrow axial charge

- almost flat
- FT → center of the nucleon

CLAS Coll., S. Pisano et al. PR D91 (2015) 5, 052014 PRL 114 (2015) 3, 032001

The CLAS12 spectrometer

- Luminosity up to 10³⁵ cm⁻² s⁻¹
- High polarization electron beam
- H and D polarized targets
- Wide acceptance

CENTRAL DETECTOR

- solenoidal field
- vertex tracker
- time-of-flight

FORWARD DETECTOR

- toroidal field
- 6 sector geometry
- vertex tracker
- three regions of drift chambers
- time-of-flight
- two threshold Cherenkov counters
- preshower
- EM calorimeter



The CLAS12 spectrometer

- Luminosity up to 10³⁵ cm⁻² s⁻¹ Region 3 High polarization electron beam H and D polarized targets Region 2 • Wide acceptance Region 1 **CENTRAL DE** Good PID for electrons, pions and hadrons with almost 4π coverage solenoidal f Kaon ID not sufficient • vertex track time-of-flig **PAC30 report (2006):** Measuring the kaon asymmetries is likely to be as important as pions The present capabilities of the present **FORWARD D** CLAS12 design are weak in this respect and should be strengthened. toroidal fiel 6 sector geometry Solenoid vertex tracker three regions of drift chambers HTCC time-of-flight two threshold Cherenkov counters
- preshower
- EM calorimeter



The CLAS12 RICH project

- Initially proposed by INFN LNF and Fe (2010)
- Construction of the first sector approved by JLab & DOE (Sep. 2013)
- Special funding by MIUR for the second sector (Sep. 2013)

The goal:

4σ π/K/p separation for 3-8 GeV/c momenta

First RICH sector:

 to be ready for the beginning of CLAS12 operation with unpolarized and longitudinally polarized targets

Second RICH sector:

 necessary to extend the kinematic coverage at higher p_T and to operate with transversely polarized targets The RICH Collaboration INFN Ba, Fe, Ge, LNF, ISS/Roma1 JLab, Argonne, Duquesne U., Connecticut U. (USA), Glasgow U. (UK), Mainz (Germany), Kyungpook U. (Korea), UTFSM (Chile)



The RICH detector

Components

- aerogel radiator: imposed by momentum range
- MultiAnode PMT: to match the aerogel photon spectrum
- mirrors: to reduce the photodetector area Main constraints:
- geometry fixed by CLAS12
- minimal material budget in the acceptance

LNF activities

- R&D and prototyping
- selection of the photon detectors
- design of the module
- test of the mirrors



The RICH concept

- Hybrid solution: proximity gap plus mirror focusing
 - Small polar angle particles
 - 1m gap
 - direct imaging of the Cherenkov photons
 - thin radiator
 - Large polar angle particles
 - about 3m path length after mirror reflections
 - double passage of Cherenkov photons in aerogel
 - thicker radiator to compensate photon loss

Simulations show that the required π/k/p separation can be achieved in the whole momentum range http://www.lnf.infn.it/~mirazita/RICH/RICH_TDR.pdf



Photodetectors

- Multi-Anode Photomultiplier tubes Hamamatsu H8500 & H12700
- 8x8 matrix of 5.8x58 mm² pixels
- packing fraction ~90%
- gain ~10⁶
- mature technology
- H8500 extensively studied at LNF
- validation of the SPE capabilities
- ageing tests (n and γ)
- New H12700 released
- better SPE detection



R. Montgomery et al., arXiv:1409.3622, NIM in print



The mirror system



Ten spherical mirrors in Carbon Fiber sandwich with a honeycomb core

- minimal material budget
- radius of curvature 2.7 m

Four planar mirror made by a sandwich of thin glass layers with a honeycomb core

- small material budget
- lower cost than carbon fiber

Four planar mirrors on the lateral sides and one on the floor

- out of acceptance, no material budget issues

The CFRP spherical mirror is the more critical one in terms of performances

A careful R&D phase has been carried on to identify the producer

CFRP mirror demos produced by different companies →CMA (Tucson USA): produced mirrors for HERMES, LHCb



Spherical mirror tests - 1

- The surface quality of the mirrors must be tested at various scales
- surface accuracy, μm : related to the resolution on the reflection angle
- A spot size measurement provide information on the average surface accuracy
- the image of a pointlike source at the CoC is a point
- $D_0 \sim mm$ $\sigma_{\theta} \sim mrad \rightarrow$ **mirro** light source camera



Spherical mirror tests - 2

- The surface quality of the mirrors must be tested at various scales
- surface roughness, nm: related to the reflectivity of the mirror

$$R \propto \exp\left[-\left(\frac{4\pi\sigma_r\cos\vartheta_i}{\lambda}\right)^2\right]$$

To have R~1 at λ =400 nm (aerogel) σ_r ~ few nm



Roughness of several sample of CFRP measured at LNF with an Atomic Force Microscopy Thanks to F. Micciulla

RMS of the height profile $\sigma_r = 2.5 \div 5.0 \text{ nm}$

G. Angelini, Thesis Univ. of Rome1 (2014)



Mechanical structure

The construction of the mechanical structure of the RICH is responsibility of LNF

D. Orecchini, A, Orlandi, S. Tomassini, A. Viticchie'

- It includes:
- external envelope in Al alloy and Al honeycomb panels with carbon fiber stiffening elements
- supporting panel for the Front-End electronics
- housing for the aerogel tiles
- mounting systems for the spherical and flat mirrors





Electronic panel

Small prototype of the box made by LNF workshop Tests of the assembly procedure and of the light-tightness of the panel underway Supporting panel Al honeycomb sandwich

MAPMT side

Cooling system

Expected heat production of the Front-End electronics is about 400W

- 100W from the MAROC3 ASICs
- 300W from the readout FPGA

Test of the cooling system underway with a real scale prototype of half panel - resistive wires used to simulate the electronics heating power

RICH schedule

Mechanics

- tender awarded, construction will start soon
- delivery to JLab in march 2016

Aerogel

- production of the first layer of the large angle sector underway
- order for the second layer by this year

Mirrors

- production of flat mirrors started
- production of spherical mirrors must begin by this year, expected completion in 14 months

MAPMTs (JLab)

- production underway, expected completion in may 2016

Electronics

- design of ASIC and FPGA boards completed, validation tests in progress
- production completion by beginning of next year

RICH installation

- beginning of module assembly at JLab in 2016
- installation in the Hall B by June 2017

additional slides

Dihadron SIDIS production

$$\vec{e}p \rightarrow e'\pi^+\pi^- X$$

Beam Spin Asymmetry

$$A_{LU} = \frac{\sigma^+ - \sigma^-}{\sigma^+ + \sigma^-}$$

Under well defined approximations

$$A_{LU} \propto \frac{e(x)H_1^{\angle}(z)}{f_1(x)D_1(z)}$$
 known from Belle

e(x) is related to the scalar charge and to the πN sigma-term

$$\int e_q(x,Q^2)dx = \sigma_q(Q^2)$$

$$\sigma_u + \sigma_d = \frac{\sigma_{\pi N}}{\left(\frac{m_u + m_d}{2}\right)}$$

Prototype tests at CERN

Prototype tests at CERN

reflected light

Aerogel characterization

Refraction index

- The aerogel produced by BINP (Novosibirsk) has been selected for the RICH
- suitable refractive index: n=1.05
- large tiles: 20x20 cm²
- high transmittance: > 0.6 at λ =400 nm
- low chromatic dispersion
- Chromatic dispersion measurements
- Laboratory: prism method with a spectrophotometer
- Test beam at CERN with 8 GeV/c pions and optical filters

Monte Carlo calculation

 dispersion model combining air and quartz

The front panel

The front panel holds:

- the aerogel tiles
- the flat mirrors

Must be a stiff but light structure - CFRP skins with honeycomb core

RICH assembly structure

- The assembly structure must
- provide the same attaching point as in CLAS12
- allow easy access for the installation of the inner components
- allow the smooth handling of the module for transportation in the experimental hall

RICH installation

The assembly of the RICH will be performed at JLab in a clean area (aerogel, mirrors) **RICH module** apr 2016 pre-assembly of some components nm 90mm Y=∼0mm Z=∼0mm oct 2016 RICH module mar 2017 assembly completed (mirrors) apr 2017 start installation Dark room – mirror tests **Electronics Aerogel and** Mirror storage 26

Electronics

Complex system

- About 400 MAPMTs, 25000 channels
- **Modular structure**
- Groups of 2 or 3 MAPMT form a tile
- Each MAPMT is served by a MAROC3 chip
- Each tile is served by LV, HV and an FPGA with optical connection
- The tiles are assembled on a supporting panel made by a sandwich of Al and honeycomb

MAPMT & MAROC3

RICH performances

Simulations tuned to CERN data

Direct detection: higher momenta

- rings are closer
- larger Npe allows clean ID

Detection after reflection: lower momenta

- rings are well separated
- lower Npe doesn't affect the ID

P = 6.3 GeV/c $\theta = 6 \text{ degrees}$ $R_{QP} = 0.59$

$$P = 3.7 \text{ GeV/c}$$
 $\theta = 22 \text{ degrees}$ $R_{QP} = 0.98$

pion kaon proton

pion kaon proton