Dark Matter Indirect Searches, multi-wavelength and multi-messenger data

Cosmic rays

Fiorenza Donato Torino University and INFN

Fermi Open Day Torino, 04.09.2015 Cosmic rays at galactic energies:

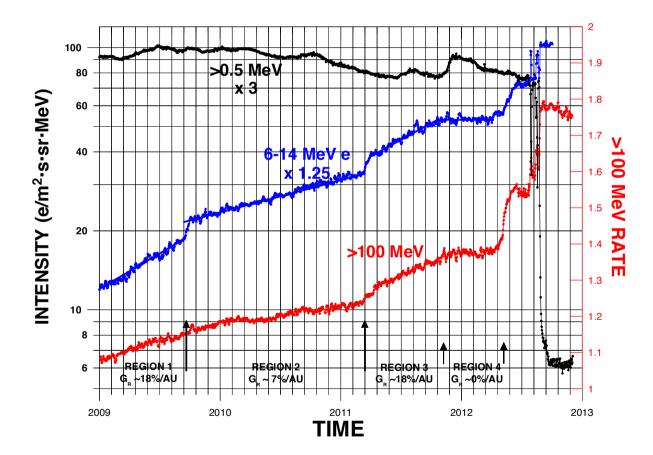
a breakthrough in observations,

a demanding effort for phenomenology

Here, very few remarks out of many open problems, ideas, strategies ...

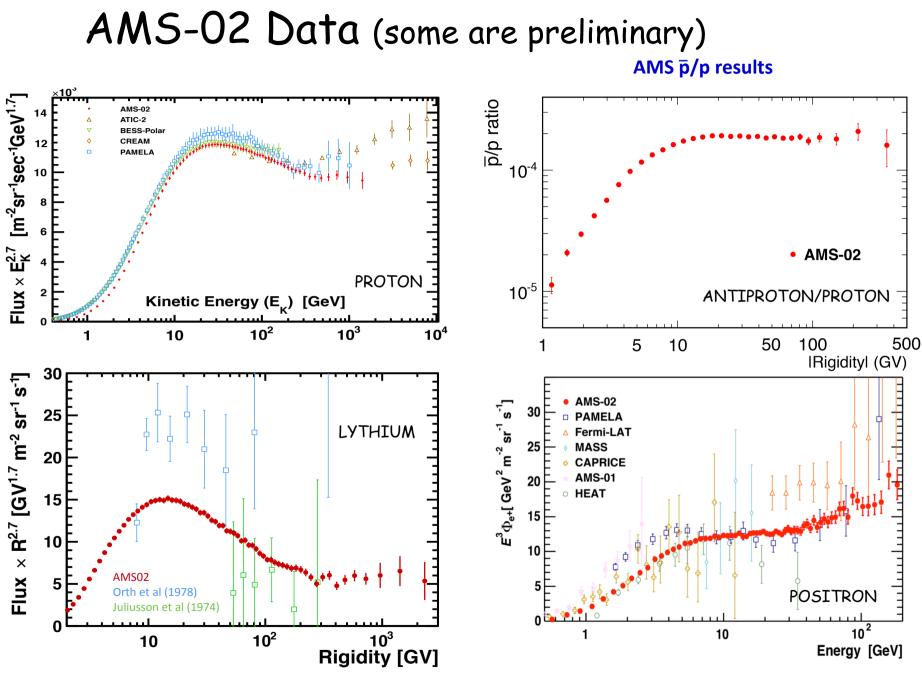
AT VOYAGER 1 STARTING ON ABOUT AUGUST 25, 2012 AT A DISTANCE OF 121.7 AU FROM THE SUN, A SUDDEN SUSTAINED DISAPPEARANCE OF ANOMALOUS COSMIC RAYS AND AN UNUSUALLY LARGE SUDDEN SUSTAINED INCREASE OF GALACTIC COSMIC RAY H AND HE NUCLEI AND ELECTRONS OCCURRED

> W.R. Webber¹, F.B. McDonald²⁺, A.C. Cummings³, E.C. Stone³, B. Heikkila⁵ and N. Lal⁵



The Voyager probe is sending data from the true INTERSTELLAR SPACE!! Now many decades in energy are covered by data

arxiv:1212.0883



From AMS Days at Cern, April 2015

From P. LIPARI, pHe workshop, 07/2015

$$\phi_{j}(E) = K_{j} E^{-\alpha}$$

$$E > 20 \text{ GeV}$$

$$\chi^{2}_{\min} = 1.70 \quad (15 \text{d.o.f.})$$

$$\alpha \simeq 2.82 \pm 0.02$$

$$\overline{p}$$

$$\chi^{2}_{\min} = 18.21 \quad (37 \text{d.o.f.})$$

$$\alpha \simeq 2.80 \pm 0.01$$

$$e^{+}$$

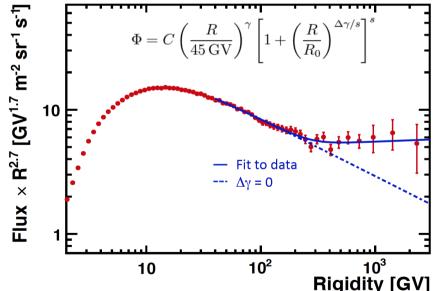
In the "Conventional Model" the similarity of the positron and anti-proton fluxes is simply a *coincidence*.

From P. LIPARI, pHe workshop, 07/2015

The propagation of electrons/positrons and protons/anti-protons/nuclei is considered as very different

Fit of Lithium flux

Same model as the one used for proton and helium (double power law with smooth transition) between 45 GV and 3 TV:

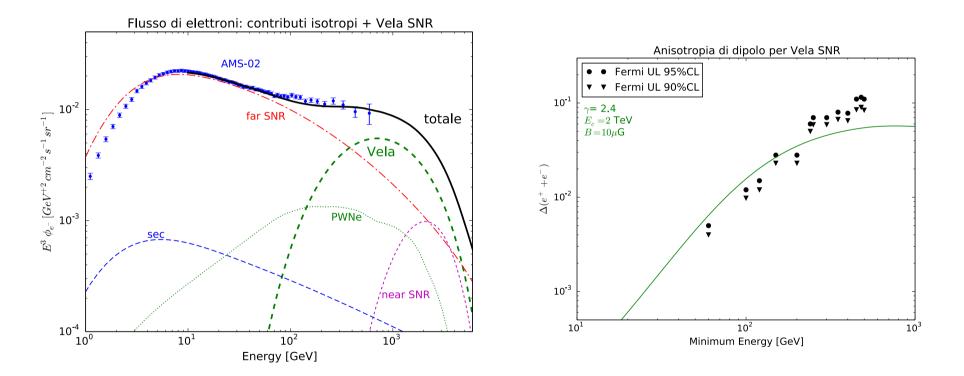


Is this all a coincidence? Probably yes, but then The physics is in the details

 \rightarrow Change of slope at the same range than for the one found for Proton and Helium.

About sources - electrons

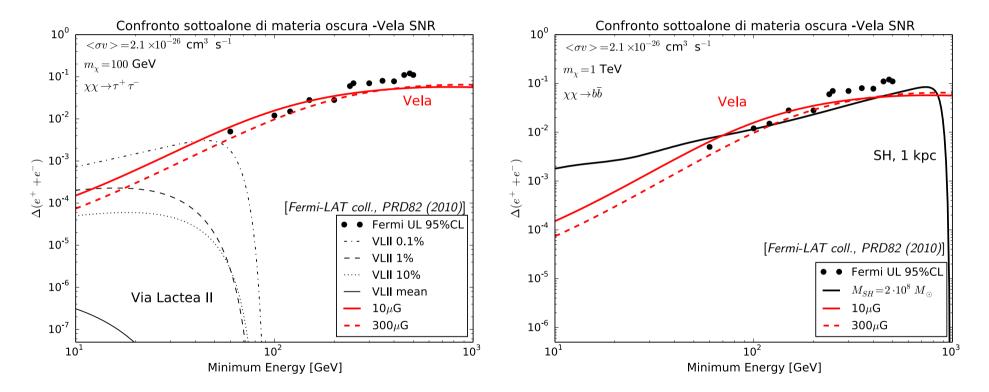
Silvia Manconi, master thesis, 2015



Vela SNR can explain electron high energy data (Di Mauro+2014) The source gives a sizeable anisotorpy in e⁺e⁻

About sources: adding Dark Matter annihilation Anisotropies in the electrons

Silvia Manconi, master thesis, 2015



Vela SNR and 1 Dark Matter sub-halo

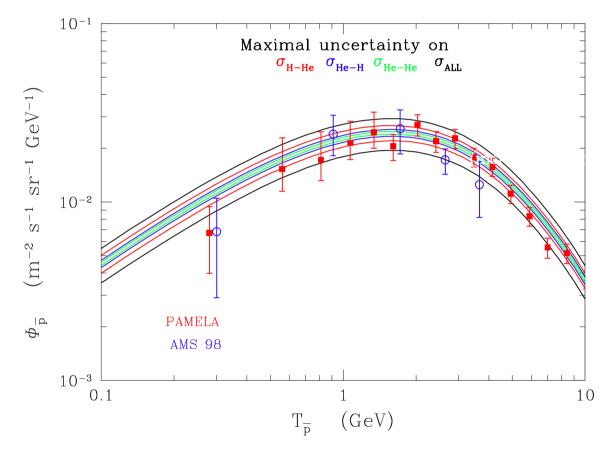
For positrons: we do not have data for e⁺ anisotropy

About sources - secondary antiprotons

A case for cross sections

Uncertainties on the antiproton flux from nuclear cross sections

(Donato+ ApJ 2001, PRL 2009)

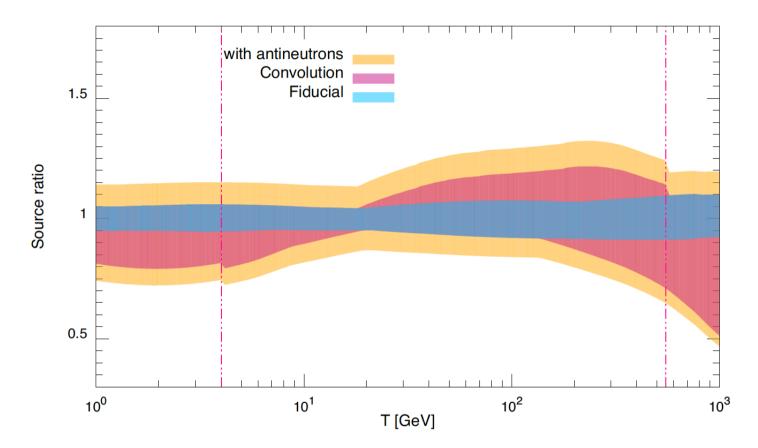


- pp: Tan& Ng
- H-He, He-H, He-He: DTUNUC MC
- Functional form for the cross section derived from other reactions, given NO DATA!!

Maximal uncertainty from p-He cross sections: 20-25%!

Data from AMS-02 on cosmic antiprotons are at ~ 10% accuracy

Uncertainties due p-p scattering

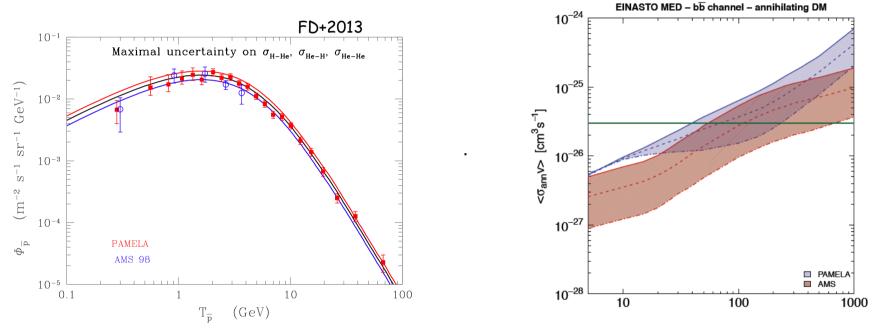


Uncertainties in the pbar production spectrum from p-p scattering are at least 10%. Conservative: 20% at low energies (GeV) up to 50% (TeV) (data expected at least up to ~ 500 GeV)

Reactions involving helium & higher energies

Uncertainties due to helium reactions range 40-50% on Secondary CR flux

Effect of cross section uncertainty on DARK MATTER interpretation Fornengo, Maccione, Vittino JCAP2014



m_{DM} [GeV]

AMS-02 is providing data with much higher precision up to hundreds of GeV!!! Their interpretation risks to be seriously limited by nuclear physics

Final (or starting!) observations - I

- Voyager is taking data in the interstellar space!! Now, models should include them and work out effects down to O(10)MeV
- AMS is providing high precision data, not easy to arrange them in one simple model
- High precision CR data imply **refined models** for: the sources, the propagation, the evolution of the Galaxy with time
- Physics and astrophysics of the sources will take a prominent role in the models: astrophysical sources (primary & secondary CRs); production through particle scatterings (secondaries in the sources, in the ISM)
 →astronomy with charged CRs?1

Final (or starting!) observations - II

- The data are such that charged CRs and photons can be studied in the same physics context. Multi-species and multi-wavelength approach
- Does it still make to speak of *standard* CR models and *standard* codes?
- The <u>nuclear processes</u> from the highest nuclei to light antimatter (total inelastic, production, branching ratios, inelastic nonannihilating, nuclear fusion, ...) are modeled according to LAB experiments and many cross sections are NOT measured at all, or data are largely insufficient
- The lack of data on several lab cross sections puts serious limits in the interpretation of forthcoming cosmic ray data:
 i.e.: Cosmic antiproton data are expected with few% errors, while nuclear physics may bring uncertainties ~ 50%!

→ A HUGE LAB EXPERIMENTAL PROGRAM IS NEEDED