



# Sensing & Control of the Advanced Virgo experiment

#### Diego Bersanetti, Gianluca Gemme, Bas Swinkels

101st SIF National Congress, Sapienza - University of Rome

September 25th, 2015

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GWs ●		Advanced Virgo 000		
Grav	vitational W	Vaves		

• Perturbative wave-like solutions to Einstein's equations:

$$\mathbf{g} = \eta + \mathbf{h}$$
,  $|h_{\mu\nu}| \ll 1 \implies \left(\nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2}\right) h_{\mu\nu} = 0$ 

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#### Gravitational Waves

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• Using the quadrupole approximation, the amplitude is:

$$h_{\mu\nu} = \frac{2G}{c^4} \cdot \frac{1}{r} \cdot \frac{d^2 Q_{\mu\nu}}{dt^2}$$

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• Coalescing NS/NS in the Virgo cluster ( $r \sim 10 \,{\rm Mpc}$ ):  $h \simeq 10^{-21}$ 

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The amplitude we are looking for is **extremely** small!

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### GW Detection through Interferometry (1)



Effect of a GW on an interferometer

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# GW Detection through Interferometry (2)



Michelson Interferometer

- Detection based on  $\Delta L$  between arms
- $\Delta L \approx \frac{1}{2}hL$
- From quadrupole approximation,  $h \simeq 10^{-21}$
- If  $L \simeq 10^3$  m we have to measure  $\Delta L \approx 10^{-18}$  m !

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# GW Detection through Interferometry (2)



Interferometer with Fabry-Pérot cavities

- ✓ Length-to-phase transduction is amplified:
  - Effective length  $L' = L \cdot \frac{2\mathcal{F}}{\pi}$
  - Finesse:  $\tilde{\mathcal{F}} = \frac{\pi R}{1-R}$
  - Maximised phase response
- Drawback: requires
   a resonant condition
   to work

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# GW Detection through Interferometry (2)



Power-Recycled configuration

The idea is to *recycle* the wasted light:

- $\checkmark P_{\rm eff} \gg P_{\rm input}$  $(factor \sim 50)$
- ✓ Shot Noise reduced (factor  $\sim 7$ )
- $\otimes$  Another resonant cavity to be controlled

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# The Virgo Experiment



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### The 2nd generation: Advanced Virgo (1)



- 200 W laser
- Larger beam
- New optical layout
- Heavier mirrors
- Increased Finesse
- Upgraded Superattenuator
- Signal Recycling cavity

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#### Developing a new Lock Acquisition scheme is mandatory!

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## The 2nd generation: Advanced Virgo (2)



Sensitivity: evolution in time



 $Sensitivity:\ different\ configurations$ 

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### **Operation Conditions**



- Uncontrolled mirrors move at the *micrometer* scale
- Error signals only valid at the *nanometer* scale
- Working accuracy is at the picometer scale

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### Operation Conditions



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#### This requires a complicated procedure known as Lock Accurisition

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# Lock Acquisition

- Error signals are available only *around resonance*
- Error signals are extracted at the output ports
- Error signals are used to compute correction signals
- Correction signals are sent to the optics



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# Lock Acquisition

- Once the ITF is locked on its operating point, the control scheme is optimized in order to reduce the control noise
- Use of less noisy error signals
- Use of more aggressive filters



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## Simulations: requirements on actuation force

Study of the sensing noise and impact on actuators:



Electronic noise's impact vs. Force

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#### Simulations: requirements on actuation force

Study of the sensing noise and impact on actuators:



Electronic noise's impact vs. Displacement

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# Simulations: the *Guided Lock* algorithm (1)

Evaluation of residual cavity velocity via optical signals:



Simulation of freely swinging cavity

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Evaluation of residual cavity velocity via optical signals:



Error Signal vs. Cavity Velocity

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# Simulations: the *Guided Lock* algorithm (2)

Single impulse to reduce velocity & lock with transmission signal



Transmission lock of arm cavity (1)

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Conclusions							

- The experimental apparatus is undergoing a significant update
- New challenges due to stricter operating conditions
- Efforts ongoing into characterization, simulations and development:
  - study of the new actuators
  - simulations of the new arm cavities
  - development of a new lock acquisition strategy

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### 2016 is the year of Advanced Virgo

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