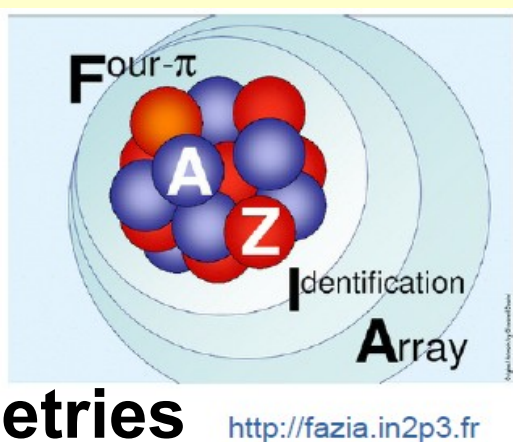


The recent Fazia activity and its programmes

Gaspard Hyde Trace Workshop in Padua, 25/27 march 2015



Symmetries

Summary

The Fazia collaboration

Detectors and Electronics

Very recent results

Programme

Cooperation perspectives (simple hints)

Fazia collaboration



Participating countries: Italy France and Poland
Manpower Total: about 45 people (technologists and technicians included)
Consumed funds: 2006-2014 (around 800ke)

A special and warm thank to **Elio Rosato** for the contribution he gave to FAZIA

Fazia collaboration

now

Az4pi

NucleX

Spiral2pp

MoU

Demonstrator @ LNS

Physics @ Ganil

2003

2006

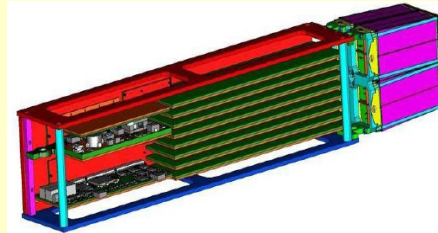
2011

2014

2015

2019

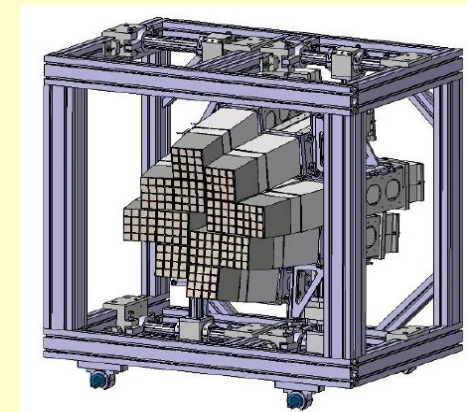
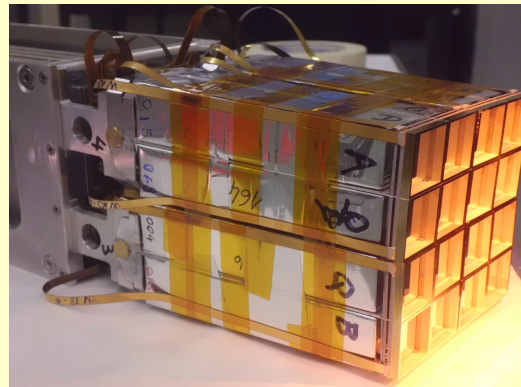
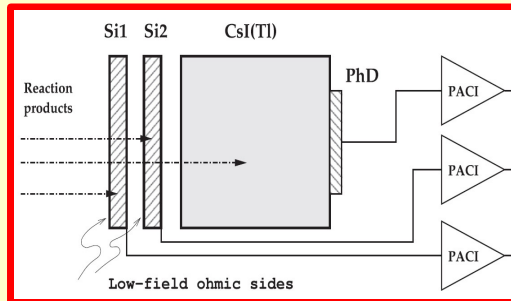
Phase0 **STARTING**:
group assessment,
Si-Si-CsI Telescope
validation



Phase2 **CONSTRUCTION**:
from Teles to Blocks;
commissioning

Phase1 **R&D**: Block
design, FEE and DAQ
prototyping

Phase3 **DEMONSTRATOR**
(Faziatto): experiments from
LNS to GANIL with INDRA



Why more powerful telescopes?

Some major PHYSICS themes

PHASE TRANSITIONS and EVENT CALORIMETRY

Reducing uncertainties from secondary to primary fragments

MULTIFRAGMENTATION

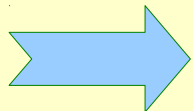
Better reconstruction of spinodal and freeze-out phases

SYMMETRY ENERGY and Nuclear EOS

A more complete view of isospin sensitive variables both in central (main hot source) and in semiperipheral events (QP and neck dynamics)

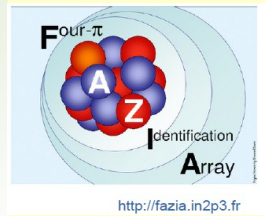
RARE DECAYS OF EXCITED EXOTIC SPECIES

Isospin effects and fragment formation at SPES energies (or higher)



Z and A identification with relatively low thresholds, with wide dynamic range and with high acceptance

New solutions for telescopes and electronics



Silicon recipe:

- **nTD** detectors for very high doping homogeneity (less than 3%)
- **Random wafer cut** to avoid channeling waveform fluctuations
- **Reverse mounting** to stress the signal shape variations for various ions
- Automatic continuous Silicon operating **voltage follower**

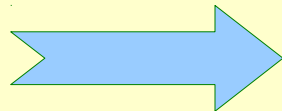
Csl(Tl) recipe:

- **Controlled Tl doping** with 'severe' constraints on the manufacturing
- **Custom Photodiode** sensors for maximum light collection
- New wrapping **reflecting material** to minimize fluorescence absorption

Original performing Electronics :

- **Custom low noise preamp** with charge and current output
- **Fast Digital stages** (100 and 250 MHz, 14 bit ADC)
- **Embedded functions/regulations** on the same multilayer board

Review article



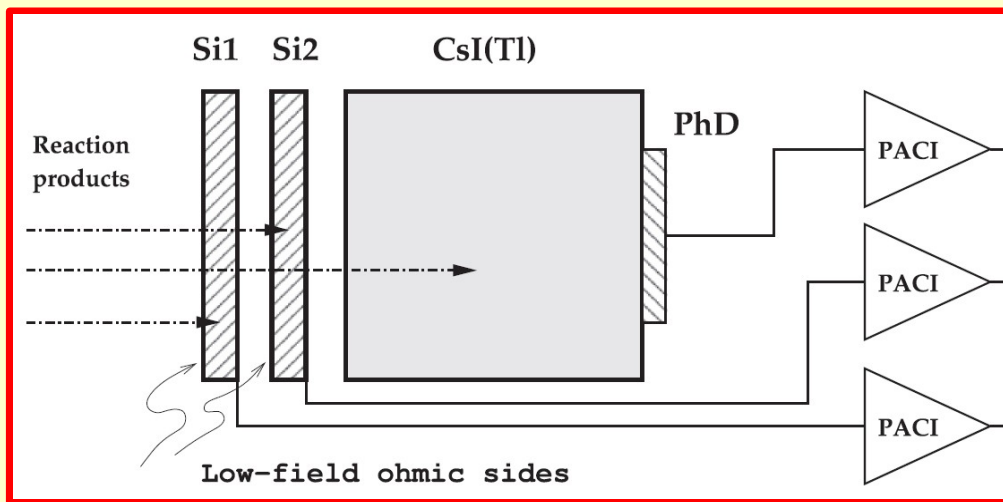
R. Bougault et al. (FAZIA Coll.), Eur. Phys. J. A 50, 47 (2014)

The basic Fazia detectors

Valuable work of
E.Scarlini on Slii-det
(Firenze)

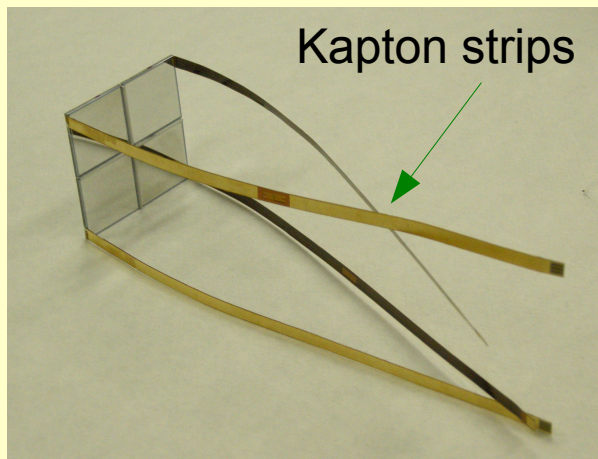
Silicons

20x20mm²
nTD type
 $\rho \sim 3-4000 \text{ ohm*cm}$
300 and 500 μm
7deg cut off <100>

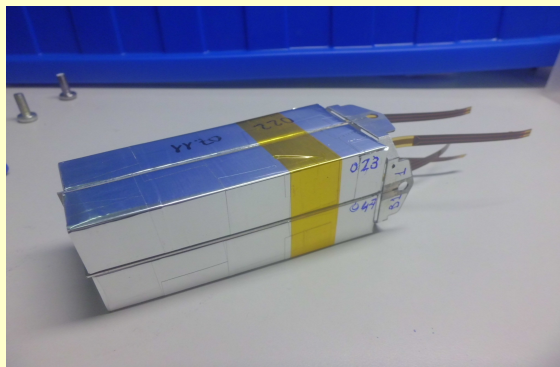


CsI(Tl)

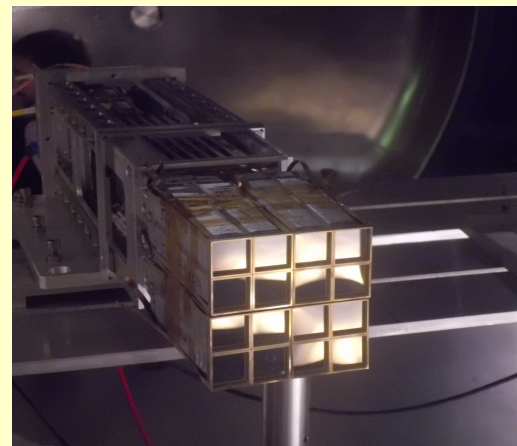
20x20mm² tapered
1500-2000ppm TI-
doping
Uniform doping
10 cm thick



Quartetto of Si mounted on Al (Ergal) support



Quartetto of CsI(Tl)



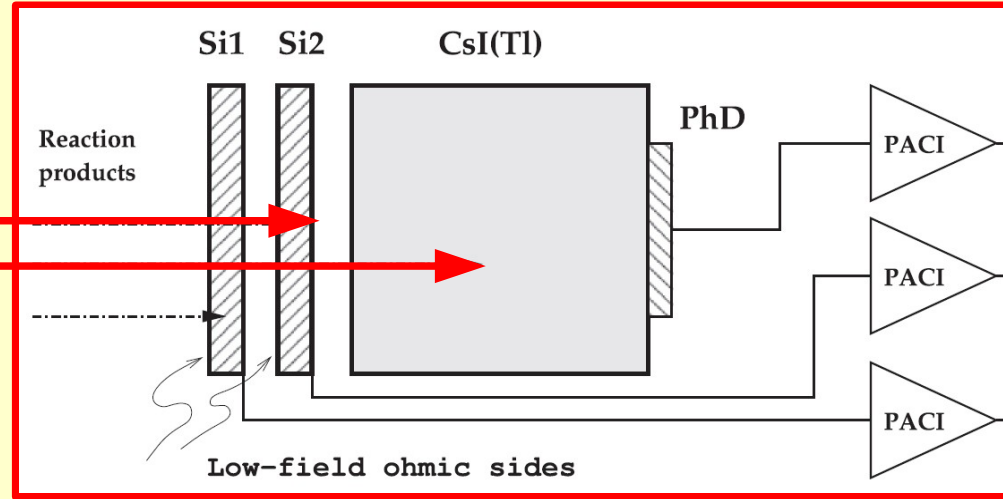
A BLOCK of 16 Telescopes
(brass collimators in front)

The detectors

Valuable work of
E.Scarlini on Slii-det
(Firenze)

Silicons

20x20mm²
nTD type
 $\rho \sim 3-4000 \text{ ohm} \cdot \text{cm}$
300 and 500 μm
7deg cut off <100>



CsI(Tl)

20x20mm² tapered
1500-2000ppm Tl-
doping
Uniform doping
10 cm thick

DE-E technique

Si1 vs. Si2

Si1+Si2 vs CsI

Deposited Energy

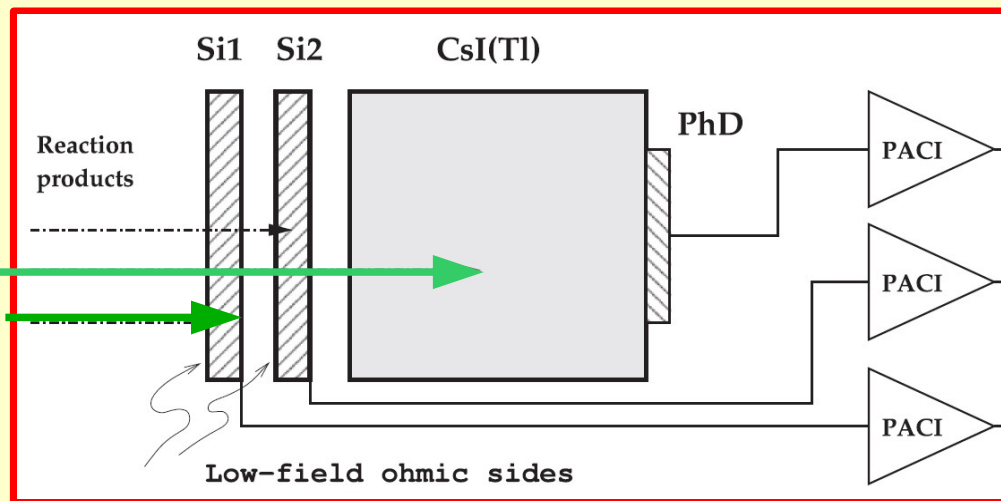
Deposited Energy

The detectors

Valuable work of
E.Scarlini on Slii-det
(Firenze)

Silicons

20x20mm²
nTD type
 $\rho \sim 3-4000 \text{ ohm*cm}$
300 and 500 μm
7deg cut off <100>



CsI(Tl)

20x20mm² tapered
1500-2000ppm Tl-
doping
Uniform doping
10 cm thick

DE-E technique

Si1 vs. Si2

Deposited Energy

Si1+Si2 vs CsI

Deposited Energy

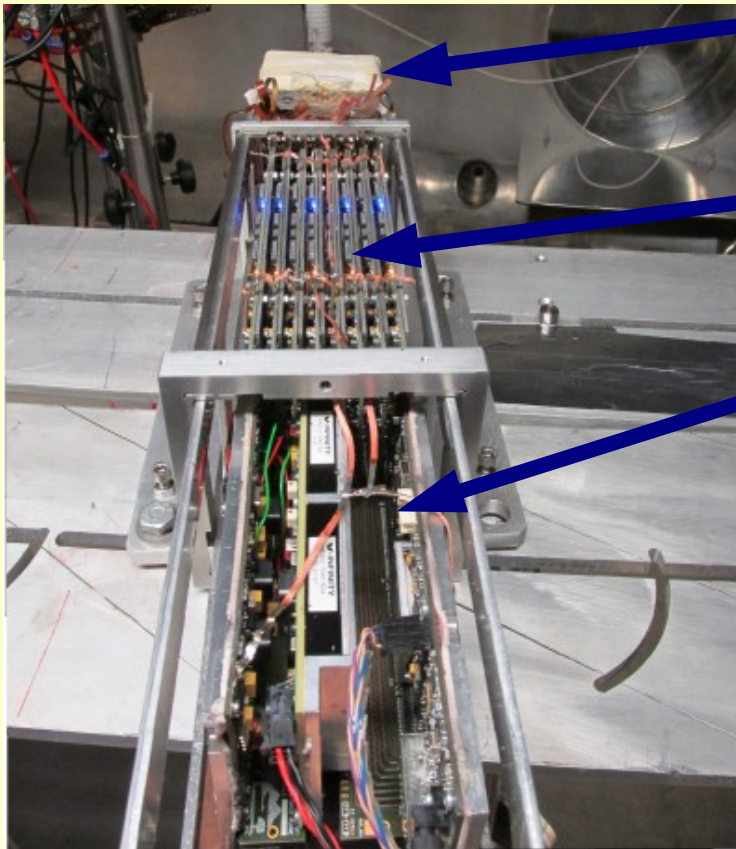
PSA

Si1 (and Si2) IMAX, Qrisetime

CsI fast-slow

The basic FAZIA Block

@ LNS



telescopes

FEE boards, 6x8 channels

HV stage, Block Card,
electro-optical coupling

BLOCK: is the stand-alone FAZIA Module

- 16 Si-Si-Telescopes
- 8 FEE Cards wit 48 lines of preamplifiers, ADC and
- FPGA
- Power Generators and regulators. Pulser generation and control

Powerful electronics

High Power, cooling, large data Bandwidth: all on a few cables !

COOLING

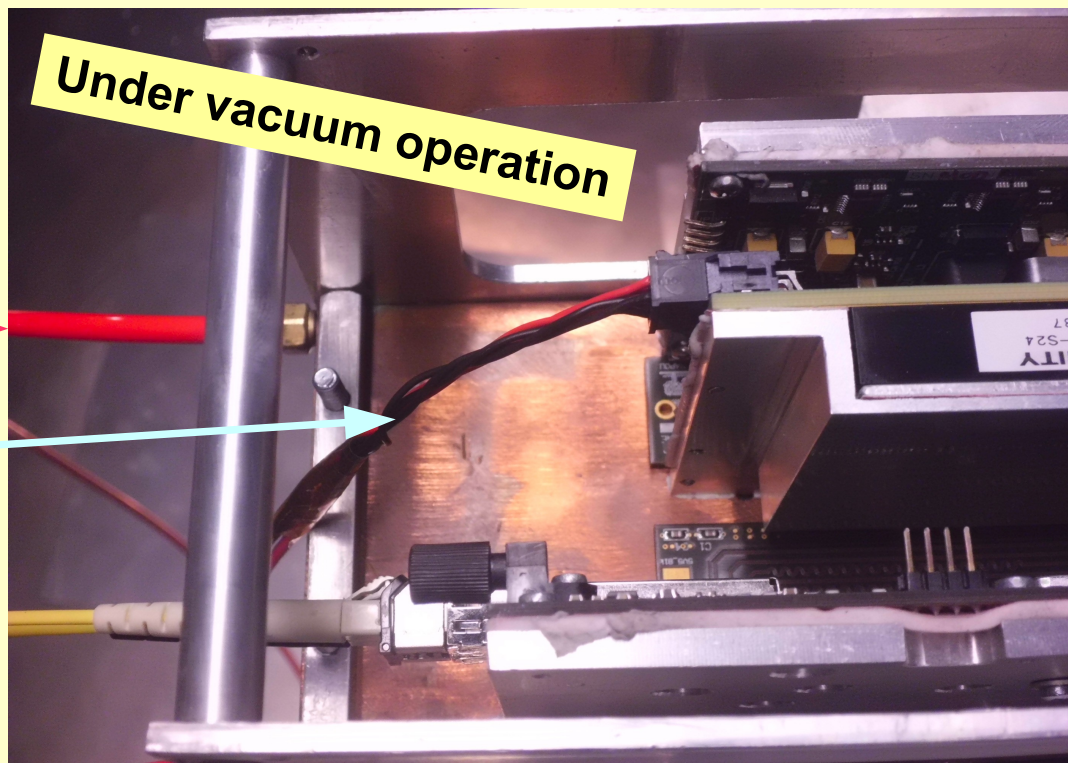
Water cooling
pipe

POWER

48V 'parent'
Power supply

I/O DATA

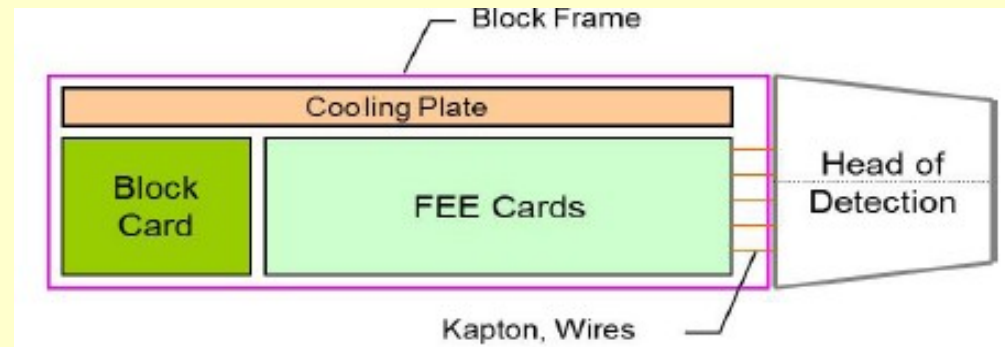
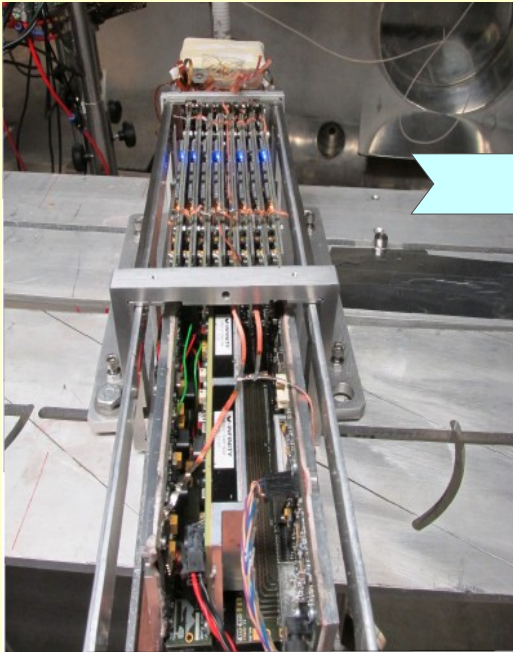
3GB/s optical
fibers



MAJOR CONTRIBUTIONS: IPN (Orsay), INFN Naples and INFN Florence

An original electronics

Charge and current I.n. Preamps; 48 channels on 8 FEE cards. Each FEE contains with 6 channels



Stage 1 (300 μm silicon detector):

Charge: 250 Ms/s 14 bit (250MeV full scale)

Charge: 100 Ms/s 14 bit (4 GeV full scale)

Current: 250 Ms/s 14bit

Stage 2 (500 μm silicon detector):

- Charge: 100 Ms/s 14 bit (4 GeV full scale)

- Current: 250 Ms/s 14bit

Stage 3 (10 cm CsI(Tl) + photo-diode):

- Charge: 100 Ms/s 14 bit (4 GeV silicon-equivalent full scale)

IPN (Orsay)

An original electronics

UP TO PRESENT

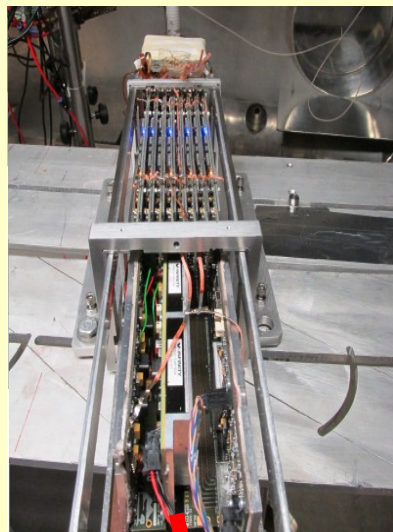
TEST CARD PCB

A.Boiano
G.Tortone
A.Ordine
R. LaTorre (grant)

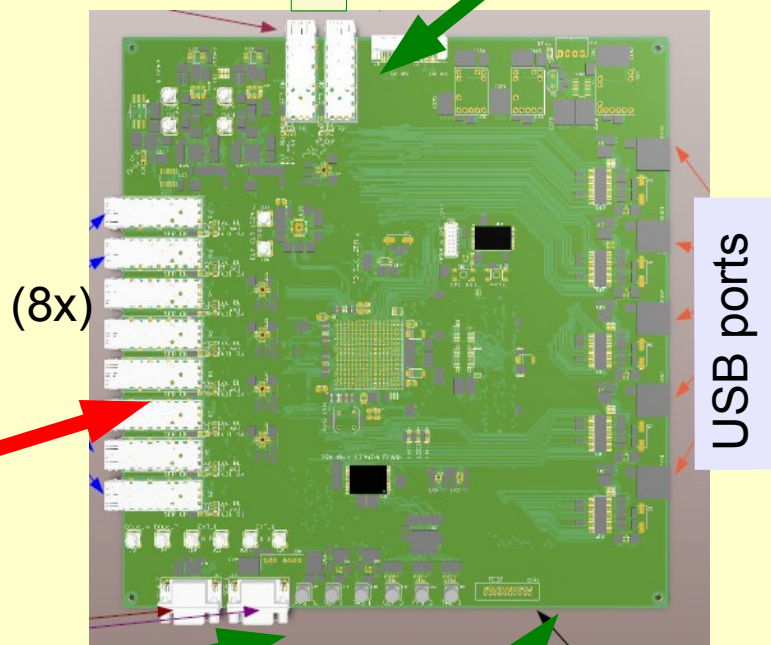
INFN Na

SW recent developments
M.Bini
S.Valdrè (PhD)

INFN & Uni Fi



To PC
Ethernet
1000BaseSX
(50MBps)



(8x)

USB ports

Fiber from BLOCK
CARD

RS232 ports

Zone for possible CENTRUM
connector (GANIL DAQ dialogue)

An original electronics

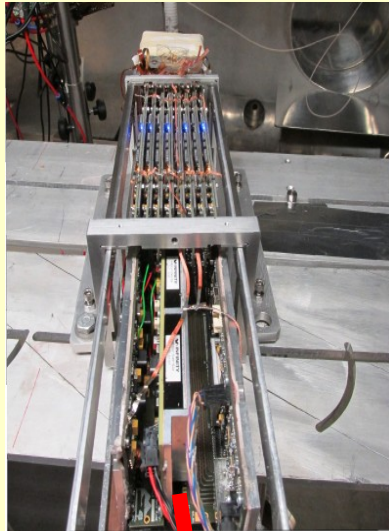
NEXT FUTURE

REGIONAL BOARD

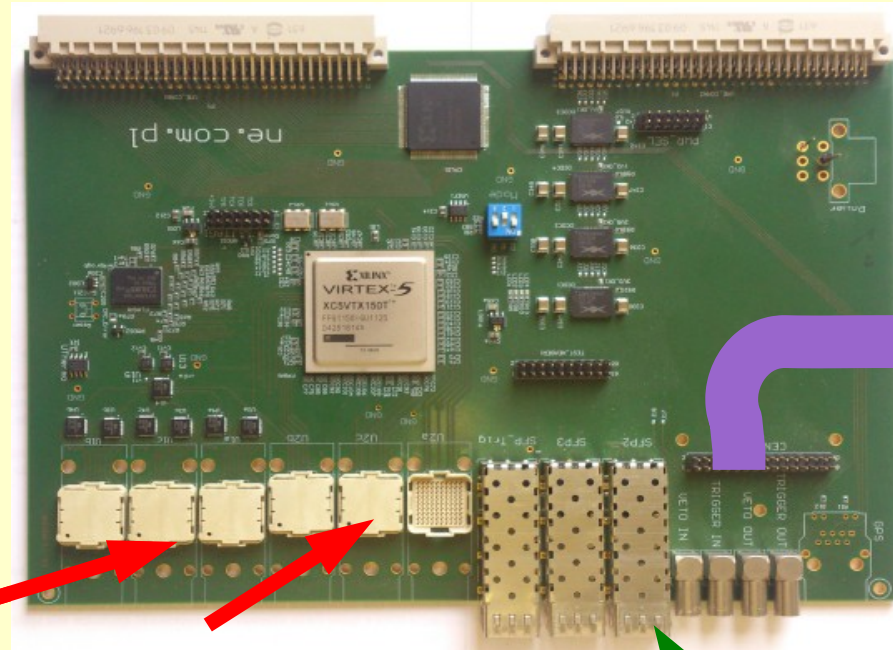
**M. Palka and
UNI Krakow**

+ Na and Fi contributions

The **Test Card** is for 2015 experiments; then it will be replaced with this **Regional Board**

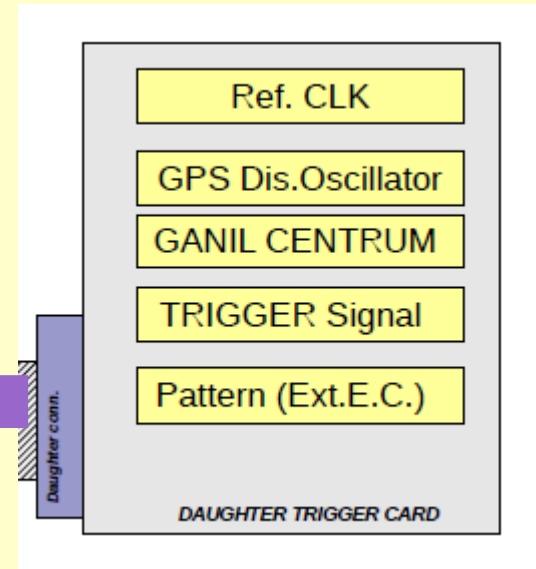


Fiber from BLOCK CARD



Snap12 Minipods
(transmitters and receivers)

SFP 3Gbit

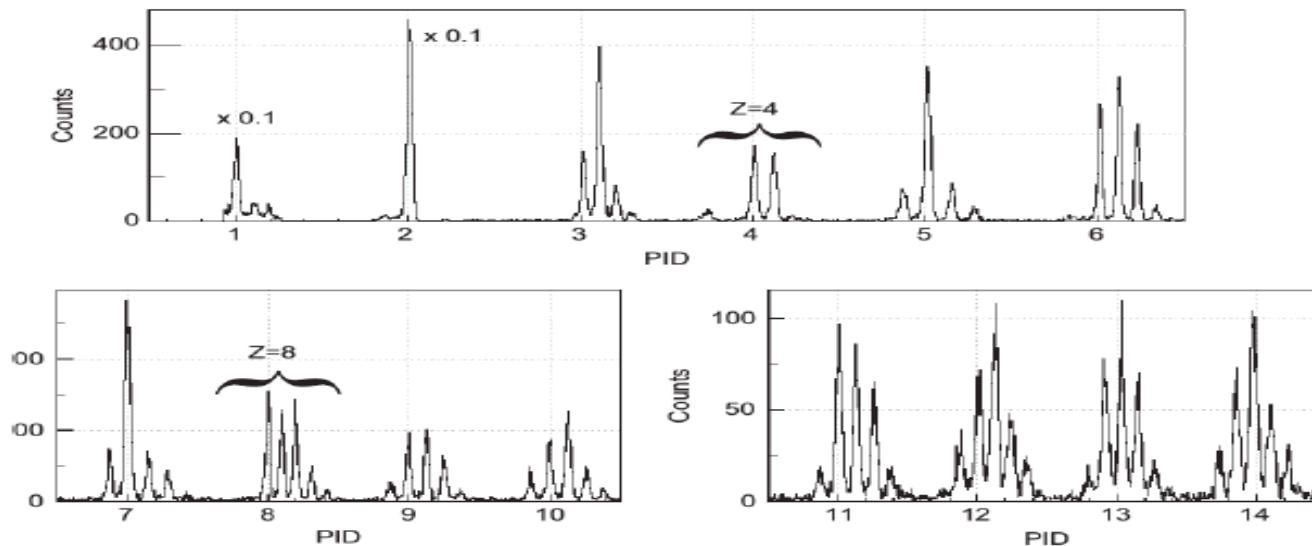


**Capable of
handling 36 blocks
(3x12 sets)**

Results from Phase 1

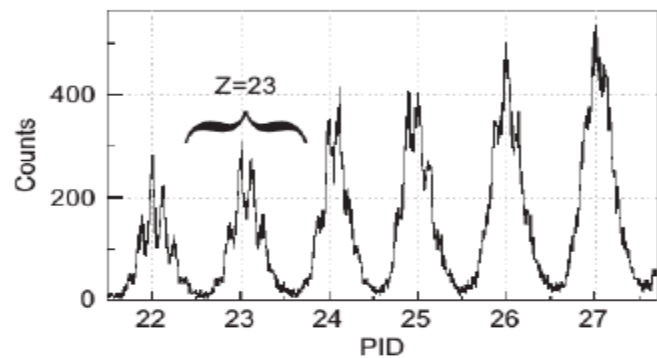
Improved ΔE -E technique

Test with
single
telescopes

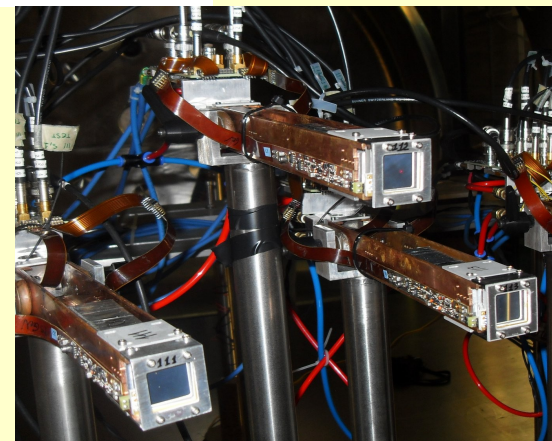
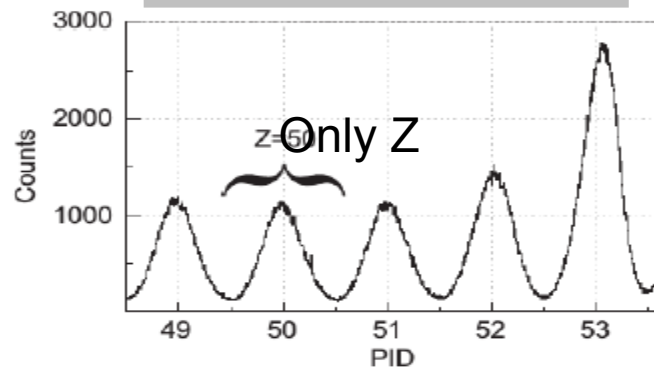


S. Carboni, NIM A 664 (2012)

A up to $Z=24$



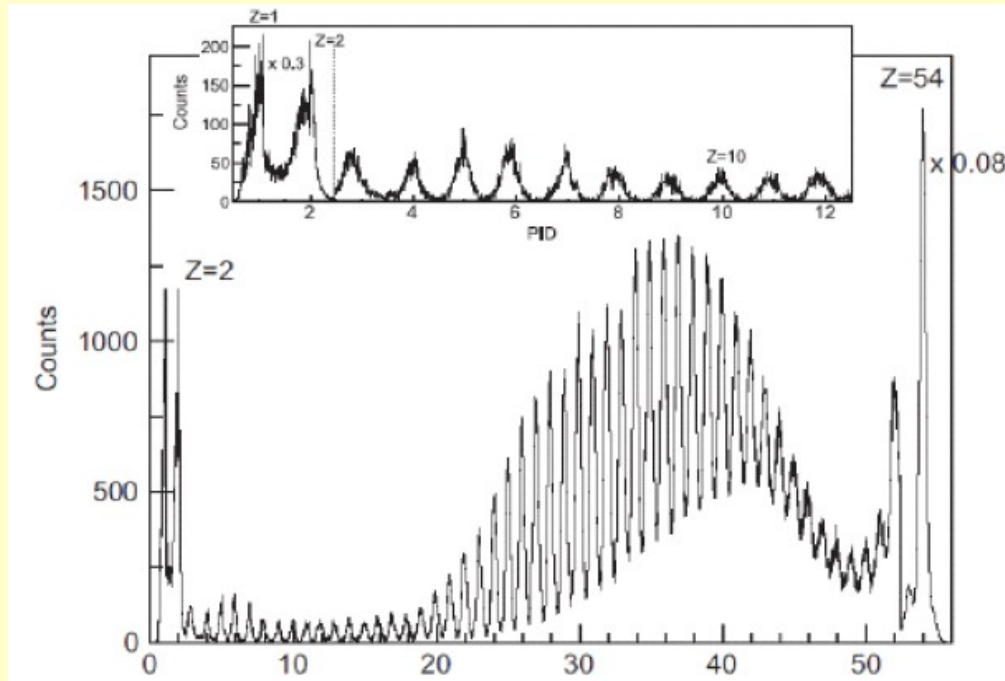
Charge up to $Z=54$



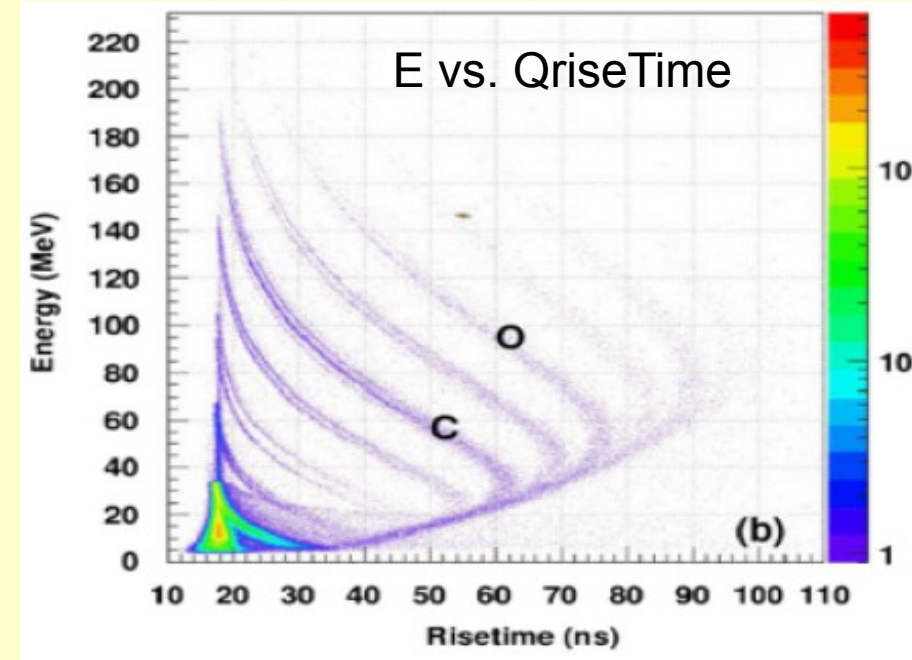
Results from Phase 1

Strong improvement of PSA

IONS STOPPED IN A SINGLE LAYER



For path in the Silicon detector lower than 30 μm (for a C) to 100 μm (for a Xe), no discrimination is possible using PSA.



- ◆ Charge separation is obtained up to 54
- ◆ Mass separation is more critical; up to $Z=14$ in best detector-electronics conditions
- ◆ Reverse mounting needed
- ◆ IMAX is better than QriseTime as a shape parameter

LNS commissioning test

december 2014

OBJECTIVES:

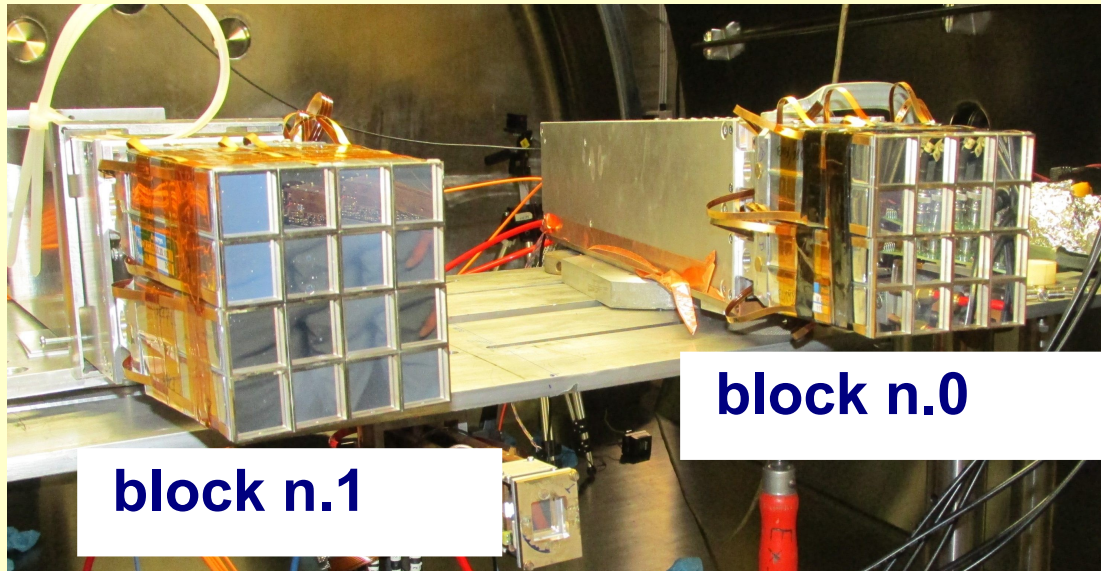
- ★ performance scalability, from single telescopes to telescope array (now 32 pieces)
- ★ complete electronics chains
- ★ data transport and DAQ
- ★ triggering and coincidence issues
- ★ Infrastructure equipments (cooling)

BEAM TEST with two blocks

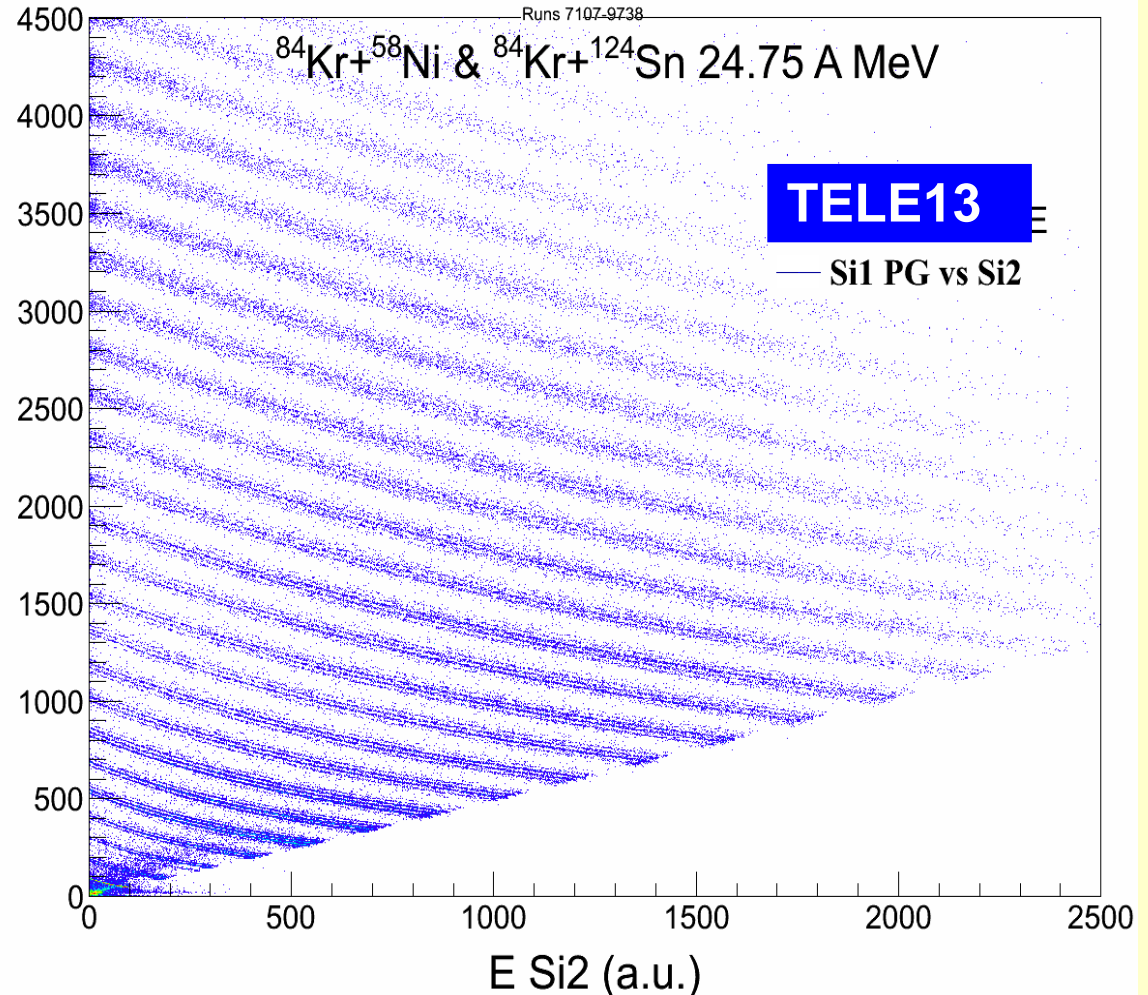
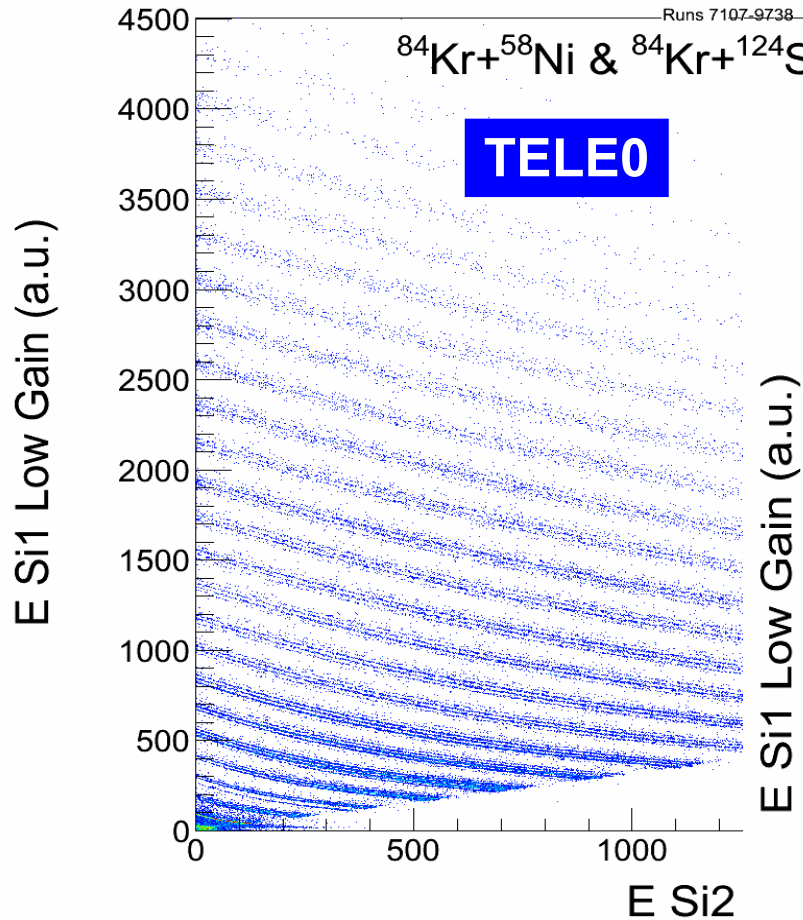
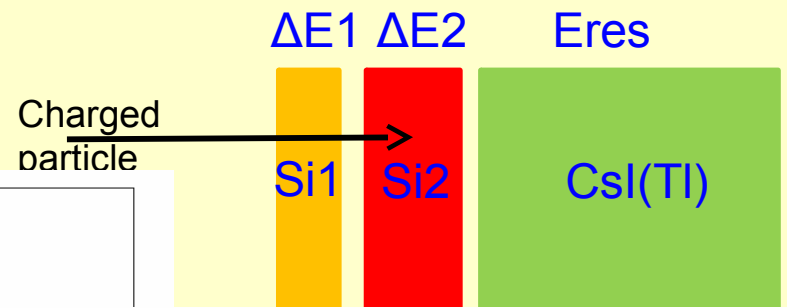
**Ciclope Scattering
chamber**

Valuable support from
CICLOPE Staff LNS

Data Analysis mainly done by
N.LeNeindre, O.Lopez, S.Barlina



LNS commissioning



ONE BLOCK had an electrical shock
maybe related to a vacuum break.

LNS commissioning

4000MeV

ZOOM

Charged
particle

$\Delta E1$

$\Delta E2$

Eres

Si1

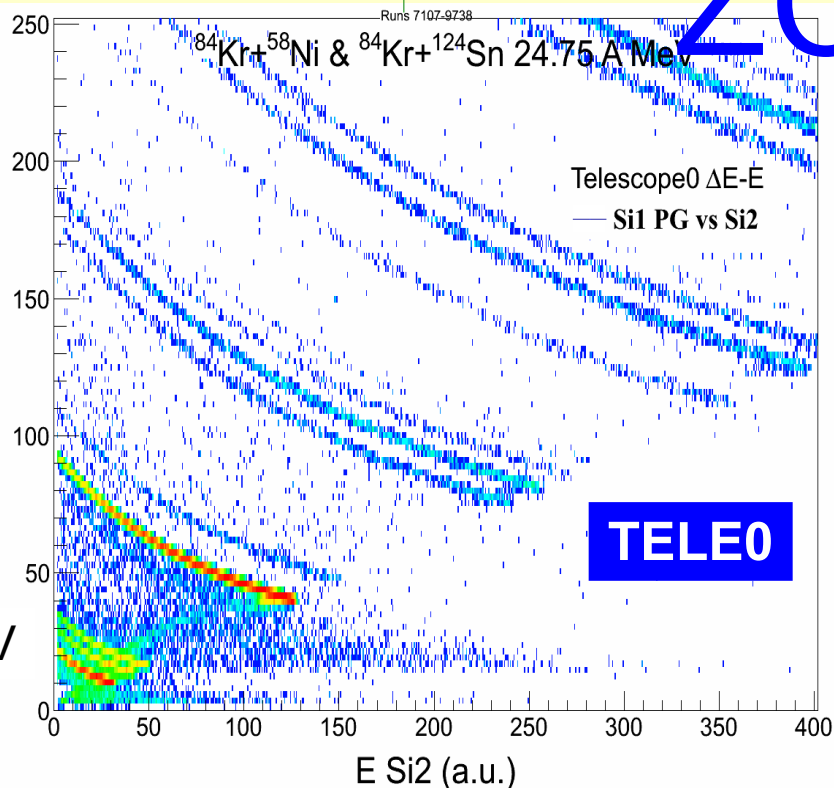
Si2

CsI(Tl)

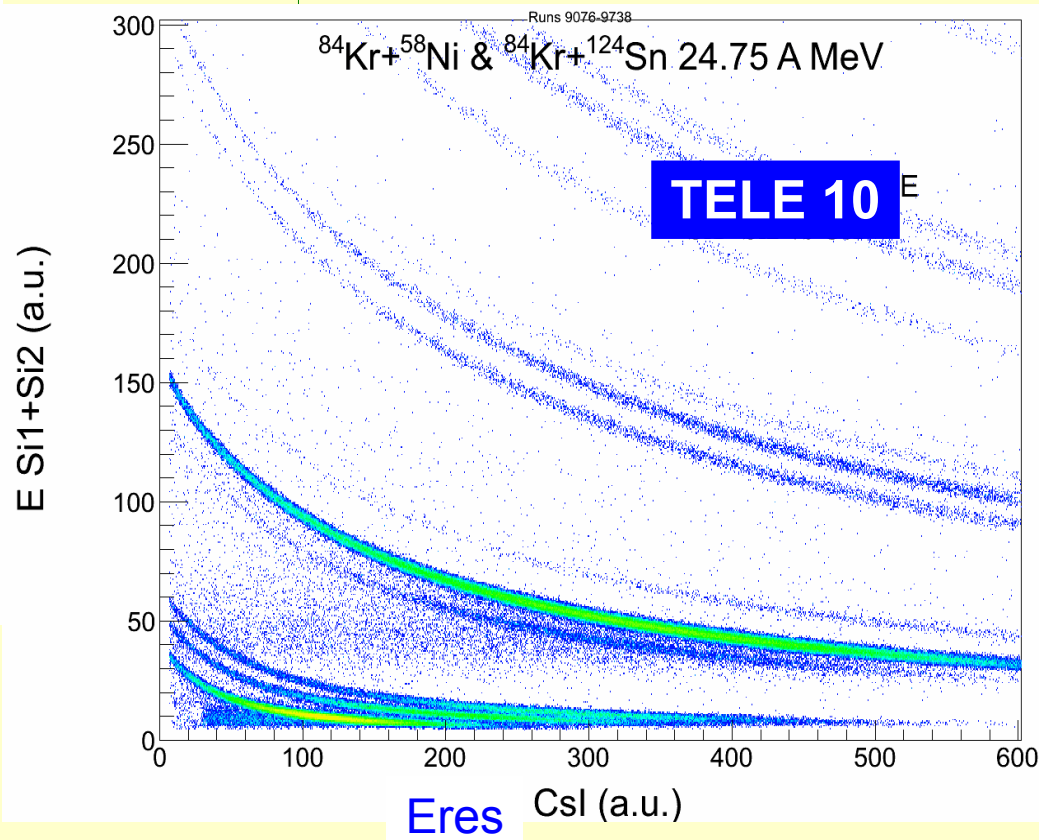
asini - G.H.T. Workshop Pd 2015

E Si1 Low Gain (a.u.)

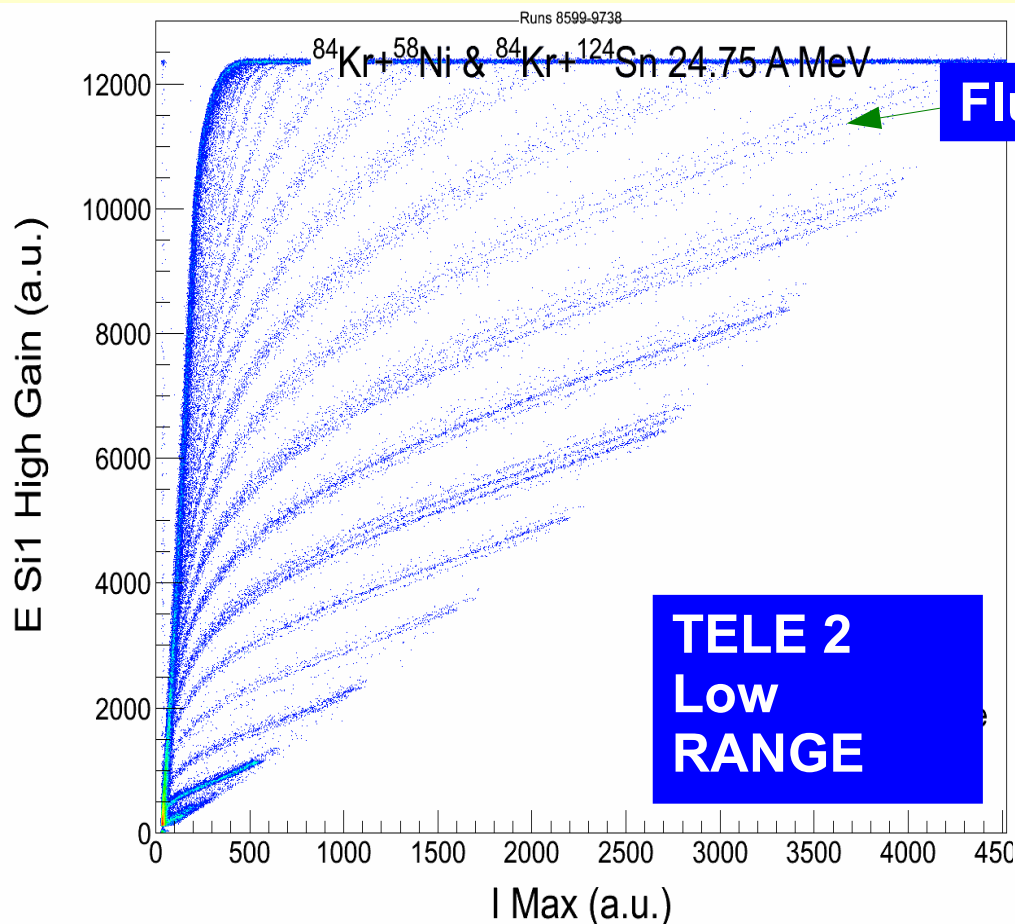
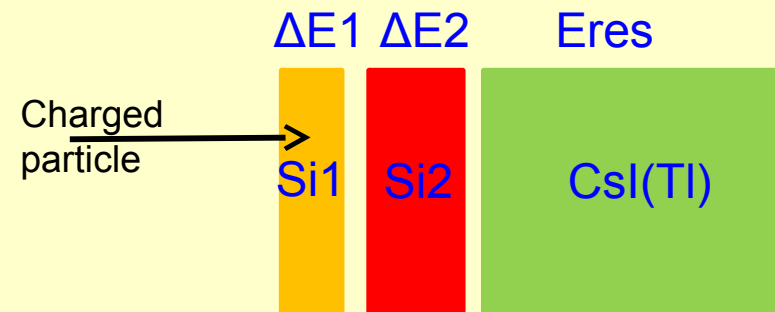
6MeV



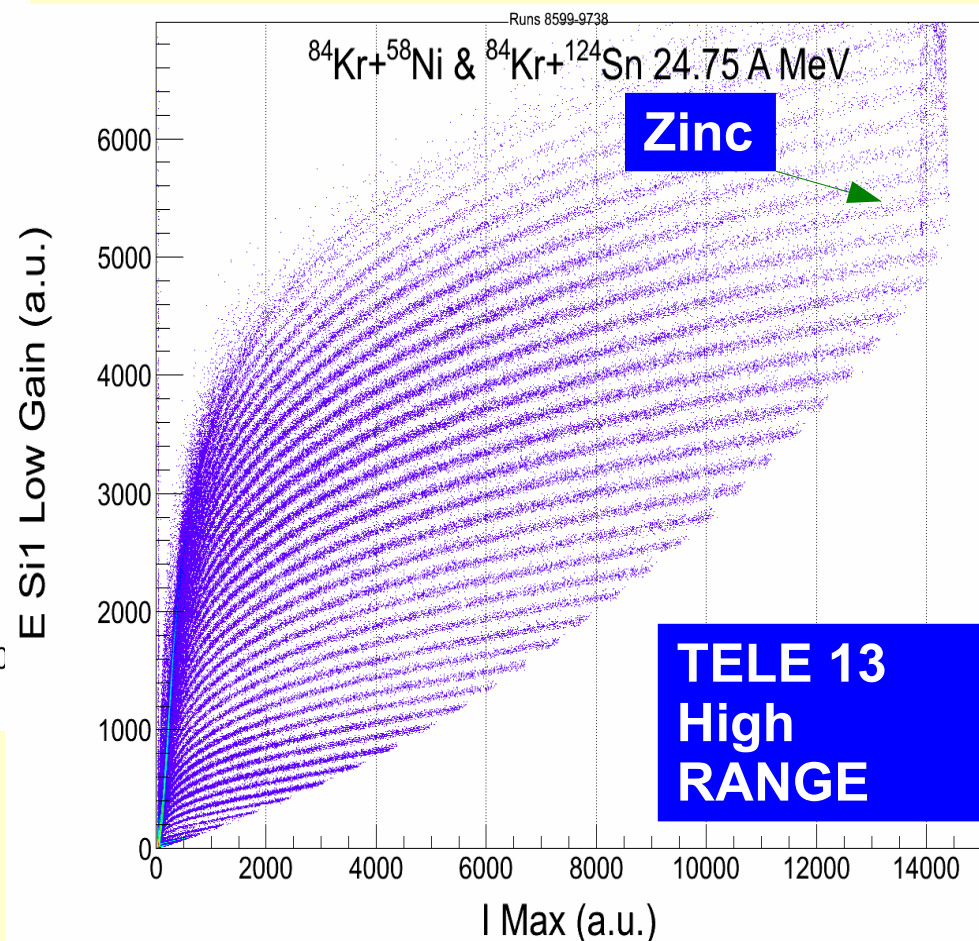
Very good isotopic separation also for light particles even with HighRange channel (ie. 4GeV dynamics!!)



LNS commissioning

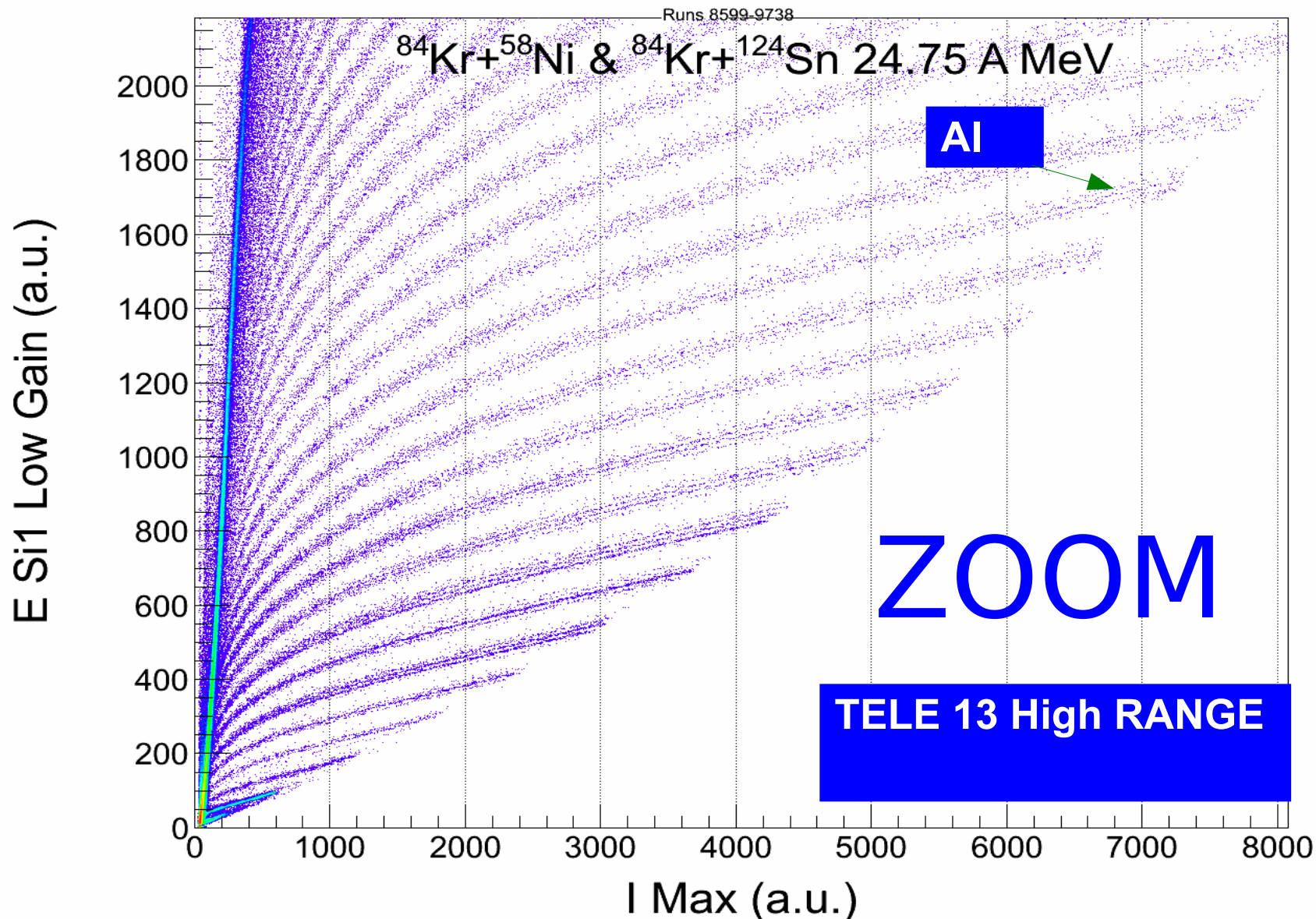


Isotopic separation from PSA up to $Z=9-13$
Full Charge identification



LNS commissioning

Isotopes for Z at least 14 from PSA



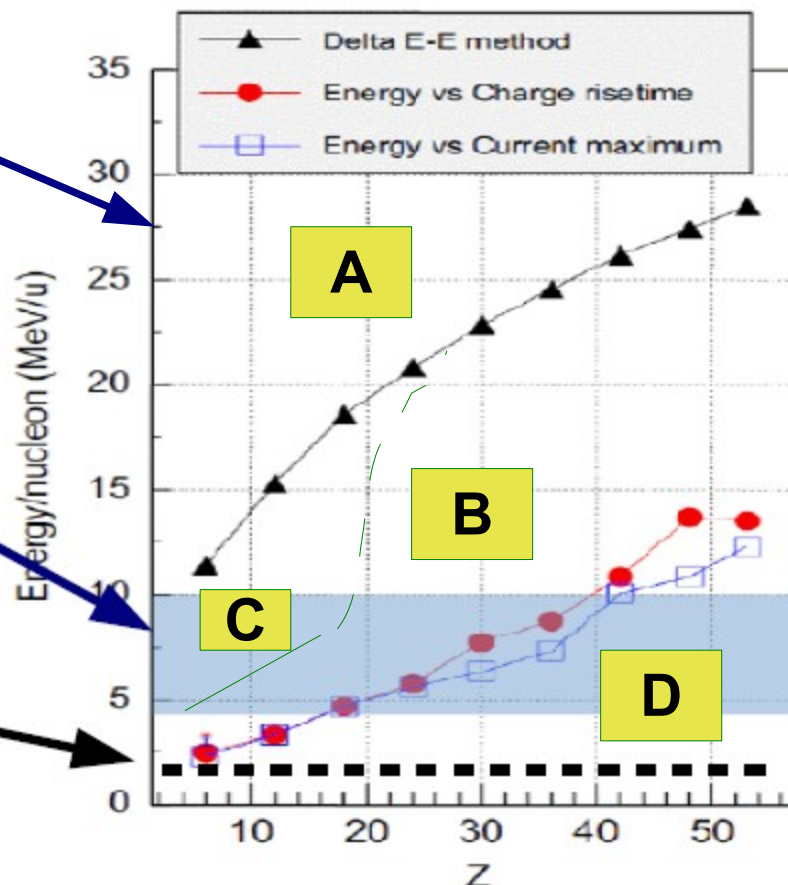
What thresholds needed?

Fazia telescopes are now well suited for Fermi energy regime

Fermi energy
RANGE

SPES beam
energy RANGE

SPES lowest
ER energies



A

Unlimited Z
A up to Z=23-25

B

Unlimited Z

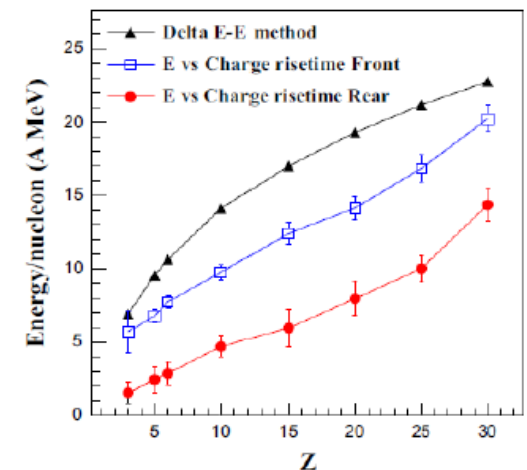
C

A up to Z=13-14

D

No identification

Reverse mounting mandatory



Further developments

Improve particle identification AND/OR reduce energy thresholds

Don't forget lower energy experiments with ISOL machines **SPES and **Spiral2****

DE-E method: thinner stages

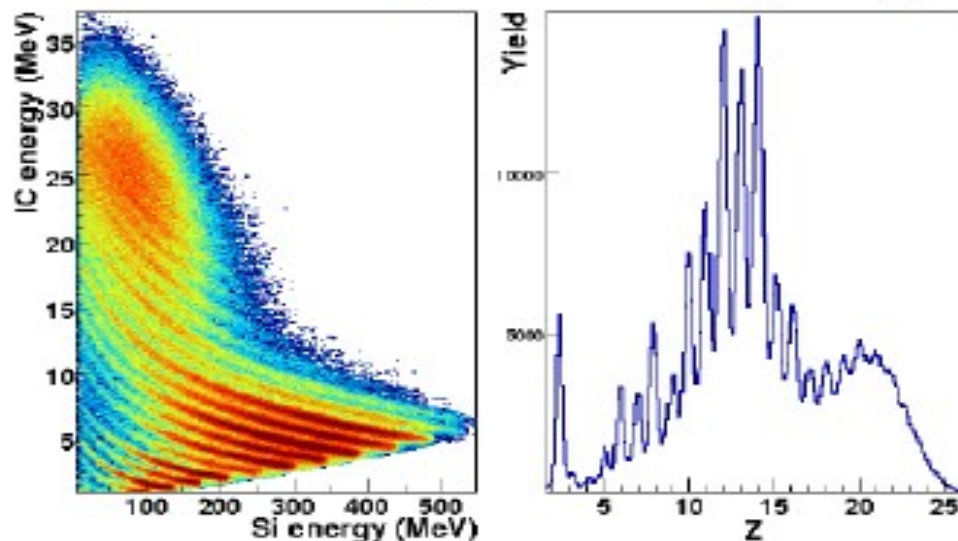
- GAS: lower thresholds, ok for Z (up to.....) but no isotopes**
- Solid state: very thin Silicon, lower thresholds, ok for Z.... some mass resolution?**

PSA

Exploring partially depleted Silicon diodes

To lower thresholds: gas

M.Bruno et al. EPJ A49,, 128 2013
32S+40Ca@17MeV/A (LNL)



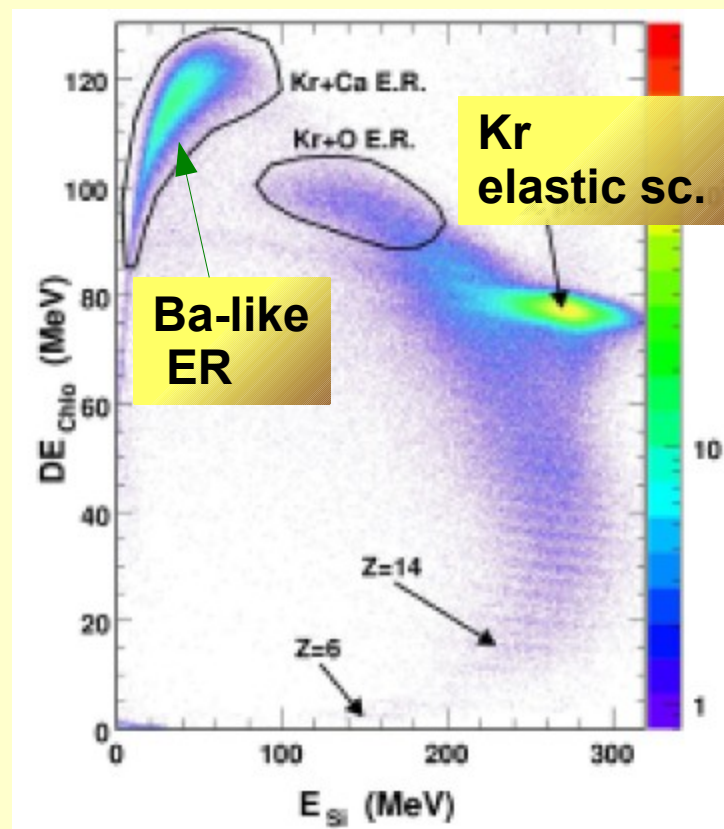
GARFIELD

ADVANTAGE: Z identification is safe up to a certain limit; down to 1-2MeV/u for massive $A > 100$ nuclei

DRAWBACKS: not-easy operation, no isotopes

INDRA

G.Ademard PhD Thesis 2011
Kr+Ca 5.5MeV/u

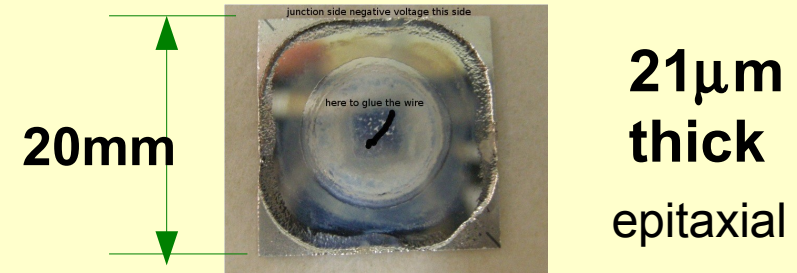
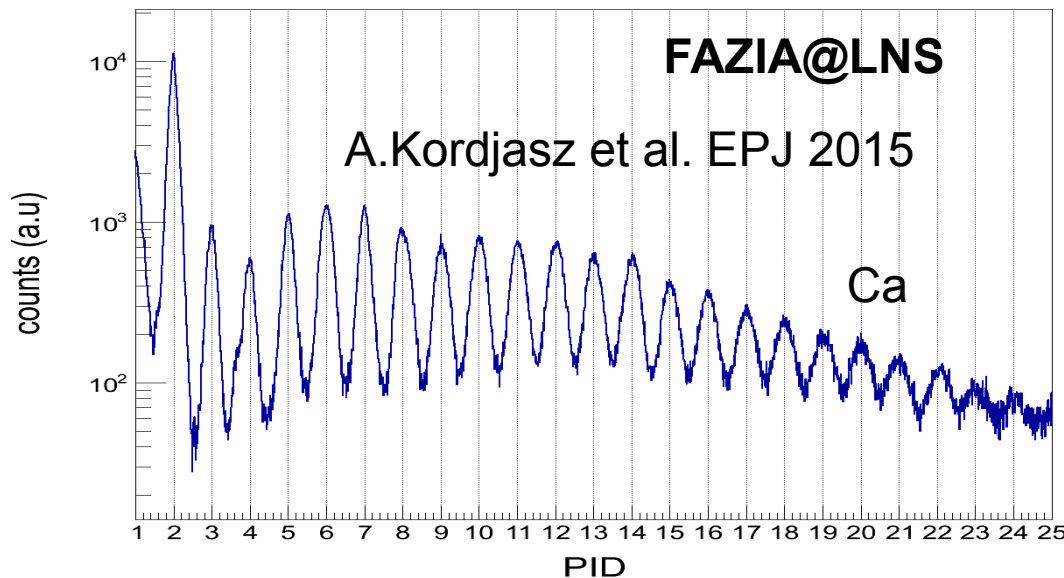


MUCH expertise within FAZIA collaboration

To lower thresholds: thin Si

Warsaw developments of thin 'large' silicon pads

Si(20 μ m)-Si(500 μ m)



20x20 mm² thin Silicon,
mounted on a FAZIA Telescope

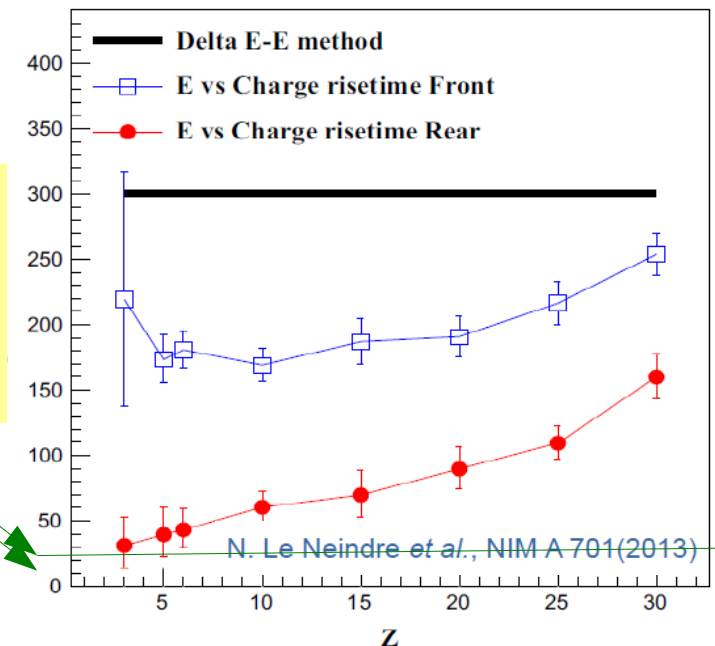
Threshold
20micron

Promising results

Z-identification with thresholds as low as
1.1MeV for protons and 2MeV/u for Mn ions.

NO MASS SEPARATION, YET

thickness



Underbiasing Si-detectors

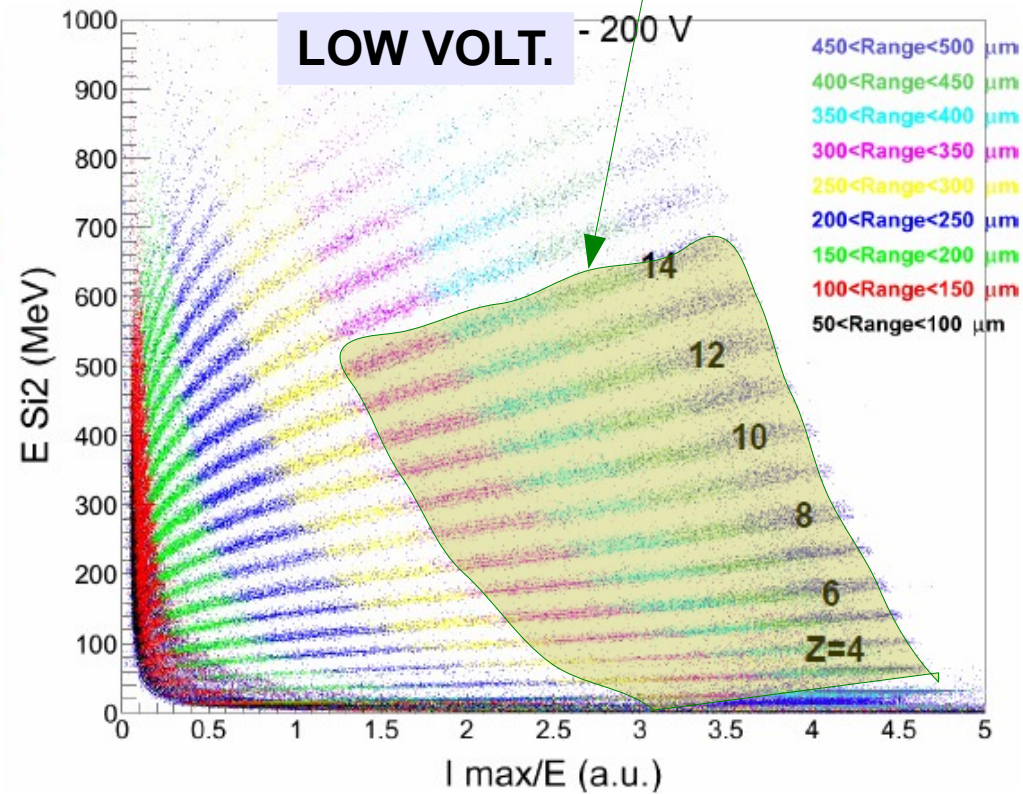
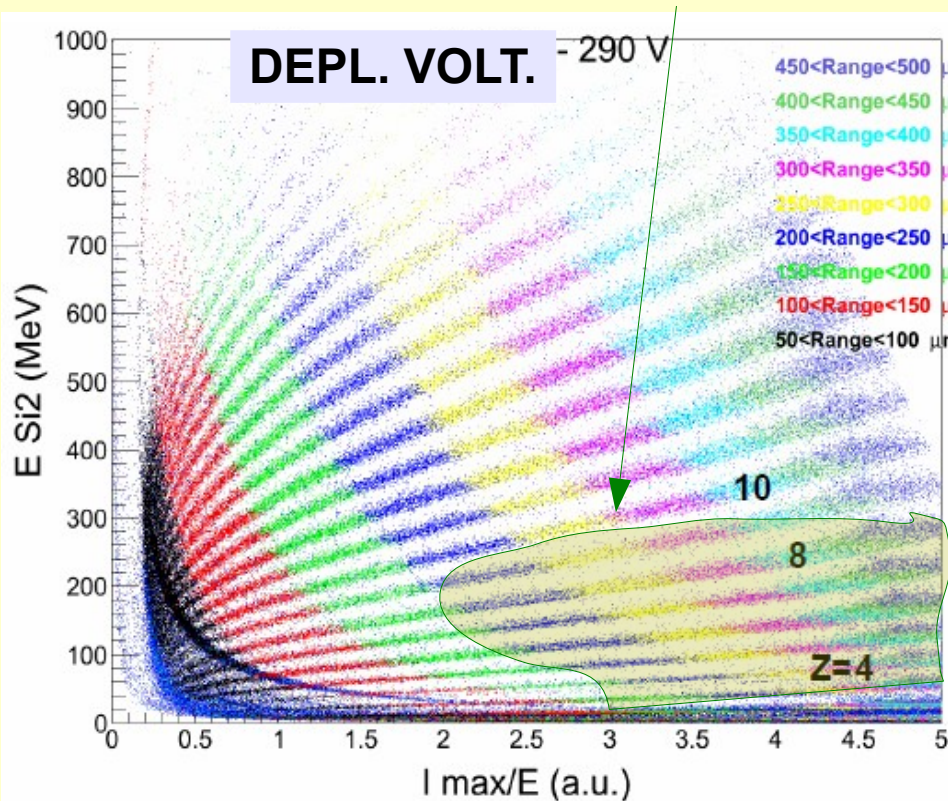
G.Pasquali et al

Eur. Phys. J. A (2014) 50: 86

Test of a 500mm FAZIA Si, below depletion

Z OK; A for Z=3 to 9 R>200 μ m

Z OK; A for Z<14 R>350 μ m



ADVANTAGE: Improvement of the identification performance. Higher Resolution

DRAWBACK: loss of sensitivity, higher thresholds, long sampling

Experimental programme

Two experiments planned in 2015 at INFN-LNS

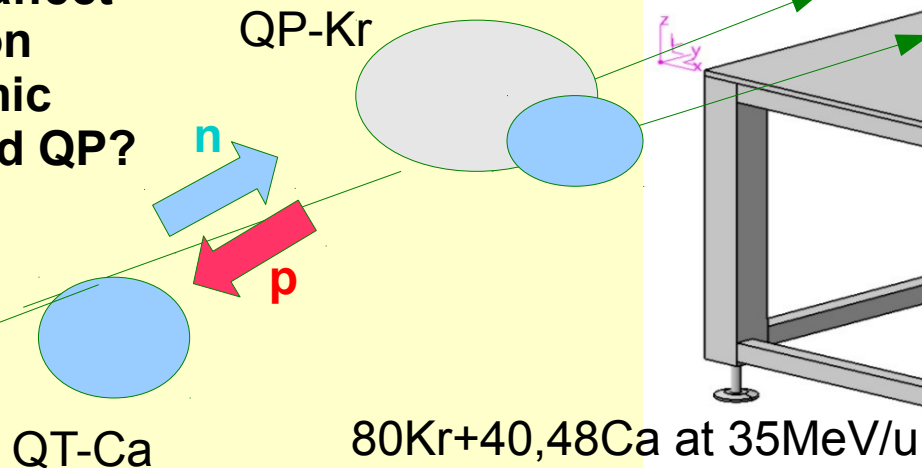
Spokep.
S.Piantelli INFN Fi

ISOFAZIA:

influence of n-p diffusion on the further decay of QP via fission (fast or statistical).

Isospin transport in semiperipheral collisions

How do p-n fluxes affect
the statistical fission
(slow) or the dynamic
split (fast) of excited QP?



Barlini et al., Physical Review C **87**, 054607 (2013)
Piantelli et al., Physical Review C **88**, 064607 (2013)
DeFilippo et al EXOCHIM at AsyEOS Workshop 2015
Russotto et al Phys Rev C **91** 2015

Experimental programme

40,48Ca+40,48Ca at 35MeV/U

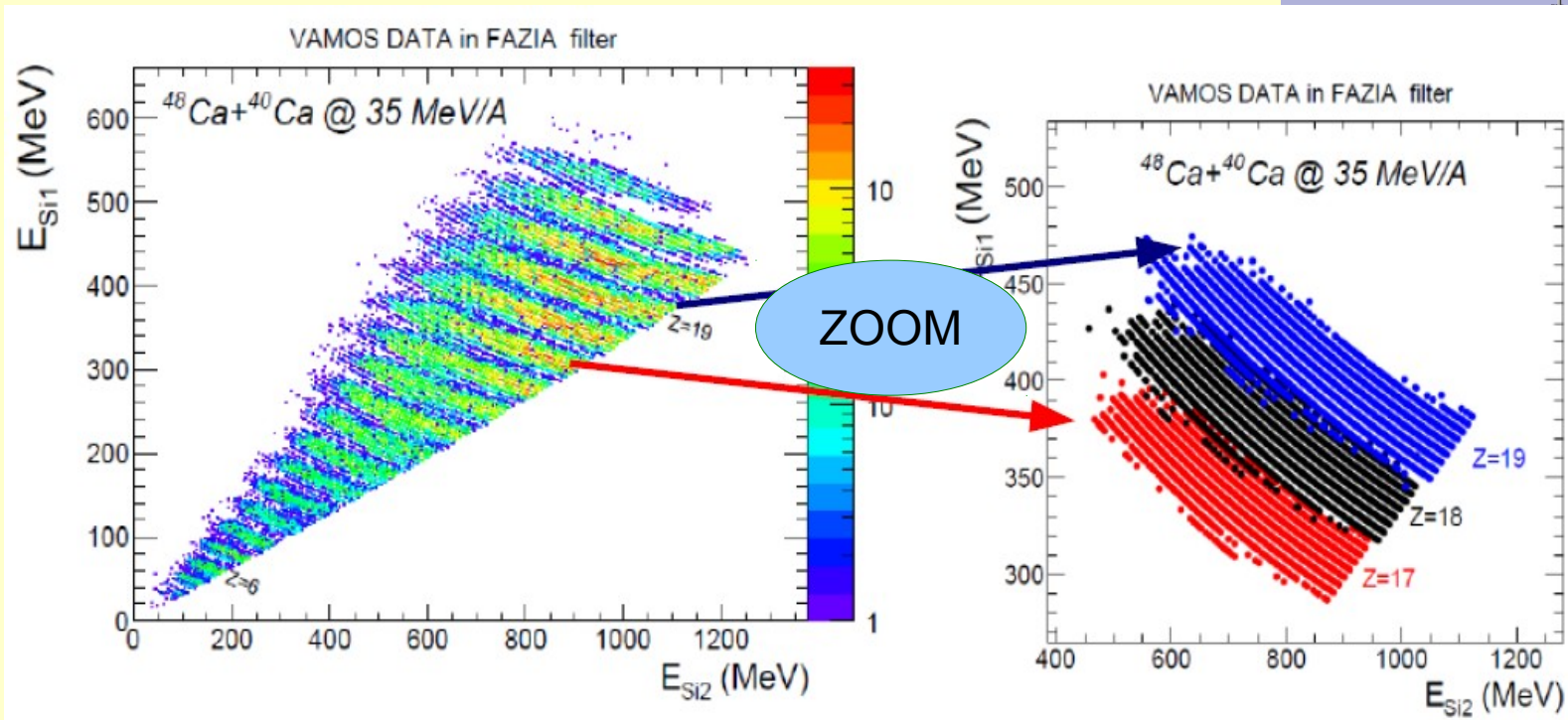
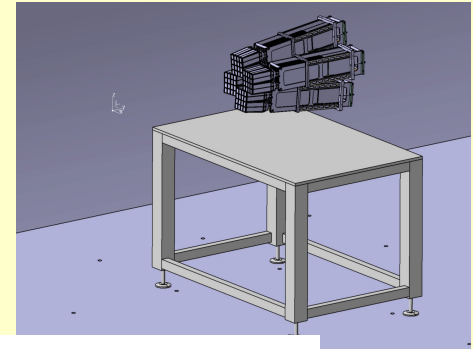
Spokep.
E.Bonnet GANIL

FAZIASYM:

Complete the INDRA-VAMOS campaign

Verify the wide isotopic distributions for n-rich system

Measure absolute cross sections



Data set measured with VAMOS filtered through the FAZIA apparatus

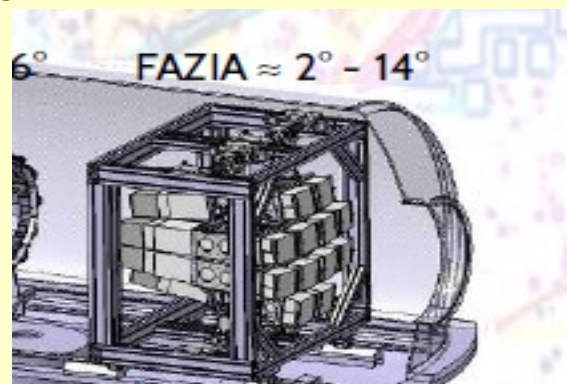
Experimental programme

from 2017 : GANIL experimental phase

CSS2 will return to operation;
stable beams between Ca and Xe
($E=40\text{-}80\text{MeV/u}$)

FAZIA A, Z for $Z < 25$ at forward
angles ($< 14^\circ$); Z for all others

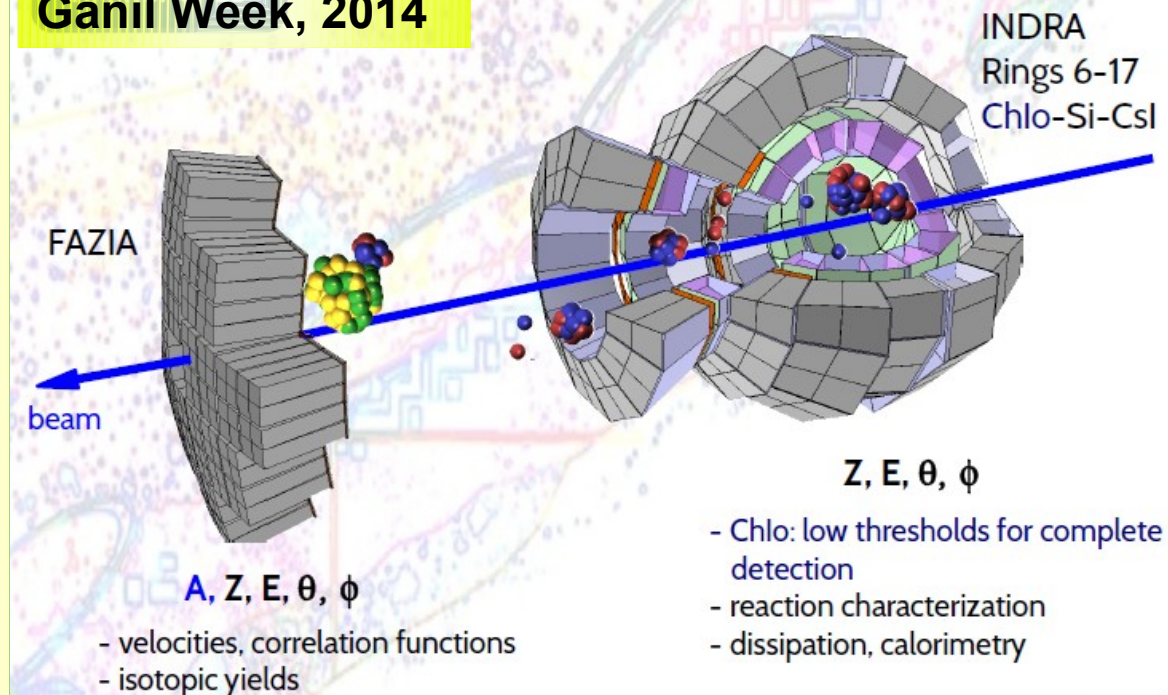
INDRA gives high acceptance to
get reliable event reconstruction



FAZIA on its support in
the INDRA chamber (D5)

From J. Frankland
Ganil Week, 2014

FAZIA@INDRA



Hints for cooperation

On a physical ground

- **Multiparticle correlations at high-energies:** FAZIA vs. Strip detectors for momentum resolution
- **Collective excitation in nuclei.** From Temperature evolution of GDR to shape transition at high spin values to DDR as an isospin equilibration mode

On a detector ground:

- **Optimizing Silicon production and operation (doping, thickness unif., channeling)**
- **Strip detectors**
- **Timing with Silicon detectors**

On an electronics ground:

- **Test our FEE boards coupled with specific 'non-Fazia' detectors**
- **R&D: towards a more integrated electronics, ASIC**

If manpower and money will permit....

Hints for cooperation

Possible experiment with the LISE exotic beams (after 2018)

