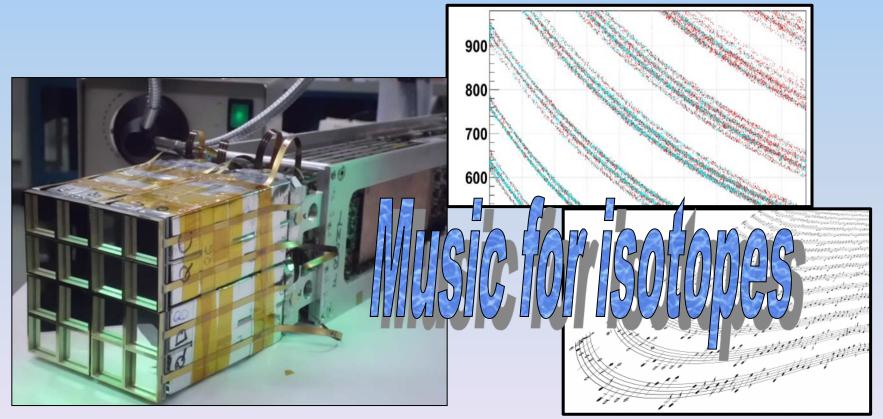
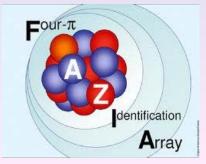
FAZIA: a new detector for charged particles. Developments and first results on isospin sensitive observables.

S. Barlini for the FAZIA collaboration

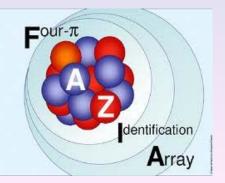




LNS -User Meeting - 2 Dicember 2014

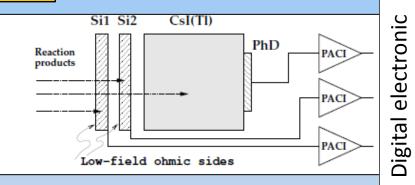
Contents:

- The FAZIA project: a brief introduction
- The R&D phases on detectors (2005-2012)
 - The "FAZIA recipe"
 - The main results
- The Demonstrator (2012-2015)
 - Introduction
 - Some details about the "block"
- The experiments with FAZIA: first results and future prospective
 - Isospin transport
 - Staggering
 - Approved experiments
 - Future prospective



The FAZIA project: a brief introduction.

FAZIA is a new array for charged particles
based on a three stages telescope Si (300 μm)
+ Si (500 μm) + CsI (TI) (10 cm) with a completely digital electronic.



INTERNATIONAL COLLABORATION

Countries: Italy, France, Poland, Romania (+support from Spain) People: about 30 physicists + engineers + technicians

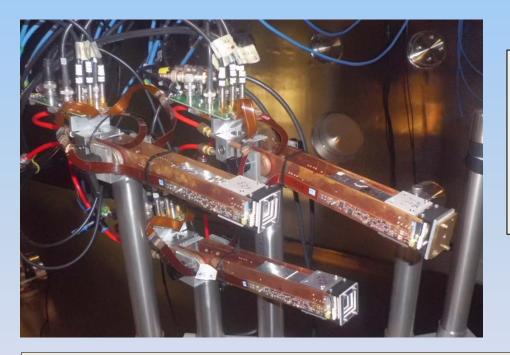
OBJECTIVE

Build-up a new array with unprecedented capabilities of ion identification, with "low" energy thresholds, modular, versatile and transportable (in view of a 'spread' use in various labs)

MAIN PHASES

- R&D on detectors and electronics (2005-2012)
- construction of a Demonstrator (2012-2015)
- experiments with stable and unstable heavy-ion beams (>2014)

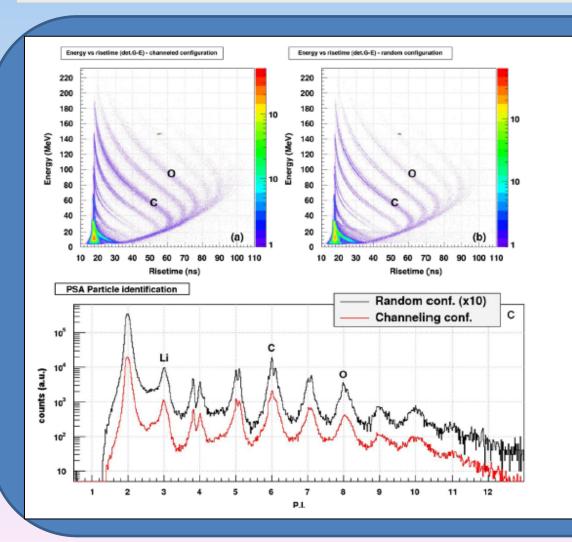
The R&D phases

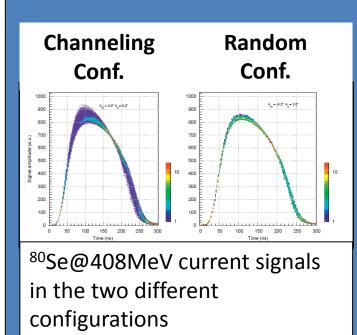


During this phase, we focus in the detector building in order to improve the particle identification. The electronic was always digital, but not in the final configuration. Also the mechanical mounting is not the final one.

- -Many test in different laboratories (LNS, GANIL. LNL)
- -PACI: preamplifier with charge and current outputs
- -Digital electronic cards outside scattering chamber with:
 - Fast sampling ADC's (100MS/s@14bit for Si,125MS/s@12bit for CsI)
 - Powerful computation unit, FPGA
 - accurate Base Line (BL) estimation (pre-trigger 2-4μs)
 - large memory depth (typical 16 μ s, BL + signal)

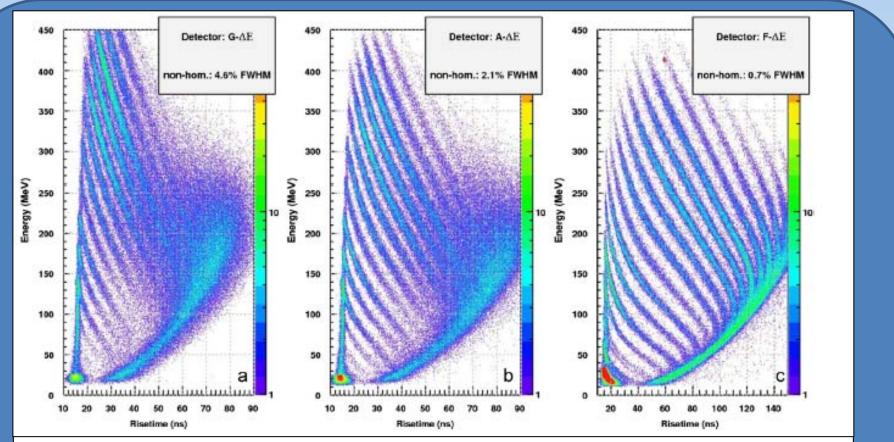
• "Random" cut of the Silicon wafers tilted with respect to the major crystal direction





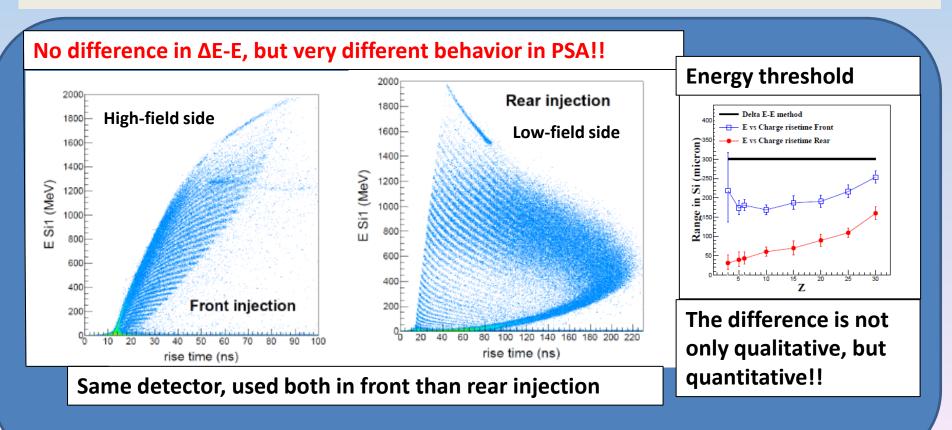
LNS User Meeting, 2 December 2014

- "Random" cut of the Silicon wafers tilted with respect to the major crystal direction
- Usage of nTD Silicon detectors with good dopant homogeneity (1-3%)



Effect of different doping homogeneity on Energy vs Charge rise-time PSA

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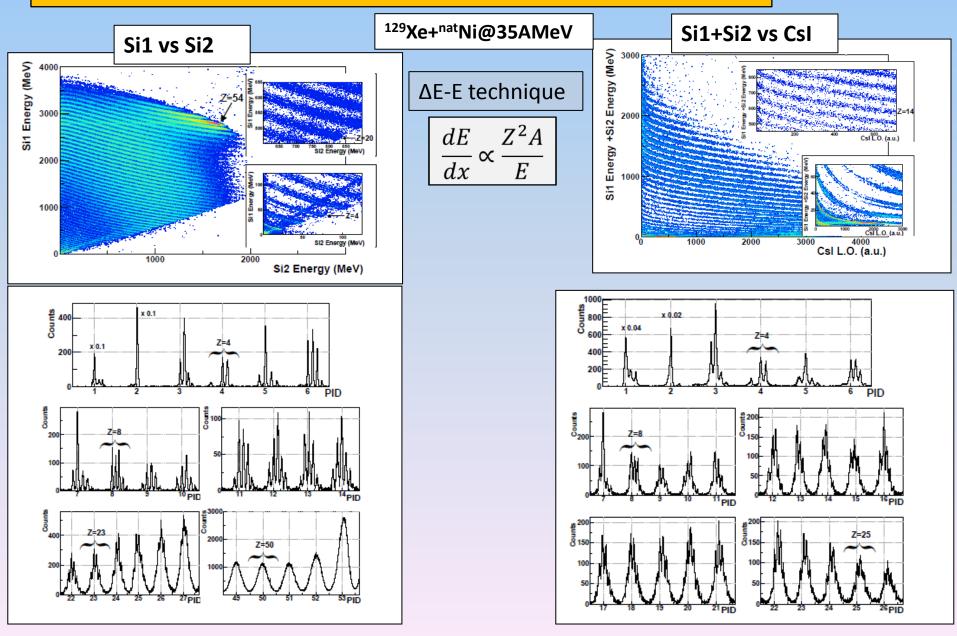
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Main results published on

Attention to the radiation damage!

S.Barlini at al., Nucl.Instr.and Meth. A600 (2009), 644 L.Bardelli et al., Nucl.Instr.and Meth. A605 (2009), 353 L.Bardelli et al., Nucl.Instr.and Meth. A654 (2011), 272 S.Carboni et al., Nucl.Instr.and Meth. A664 (2012), 251 G.Pasquali et al., Europ. Phys. J. A48, (2012), 158 N.Le Neindre et al., Nucl.Instr.and Meth. A701(2013), 145 S.Barlini et al., Nucl.Instr.and Meth. A707(2013), 89

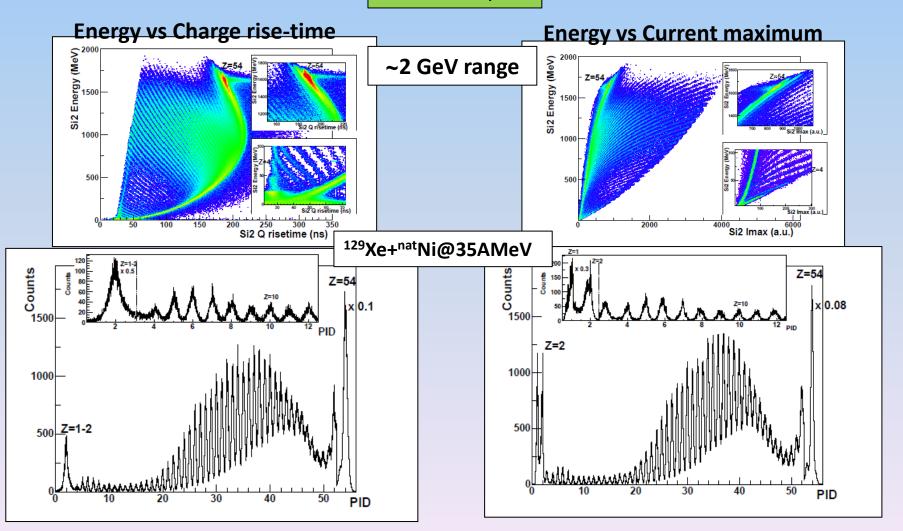
The R&D phases main results: the Isotopic resolution



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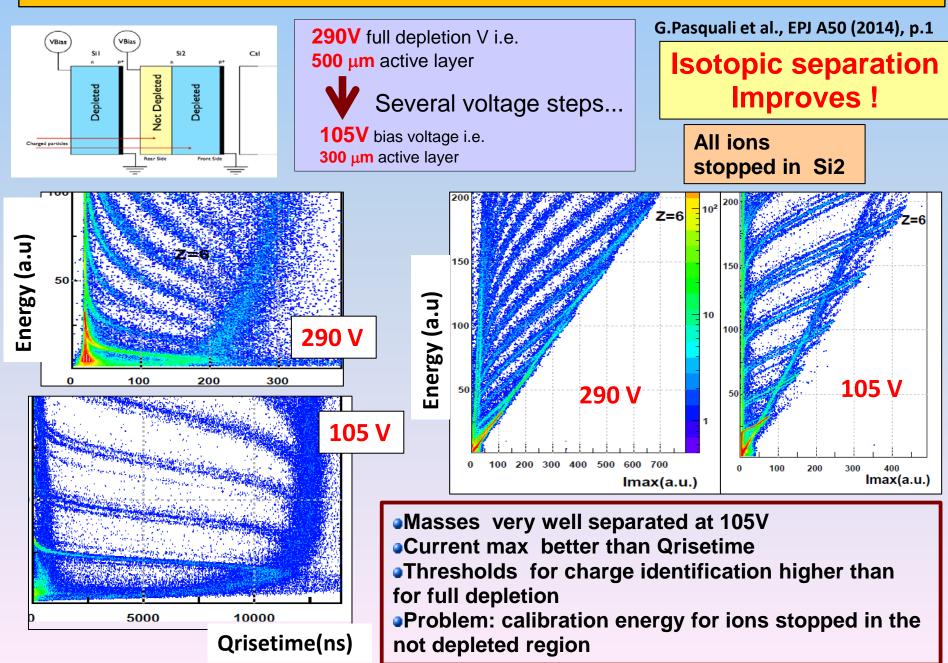
The R&D phases main results: PSA identification

PSA technique

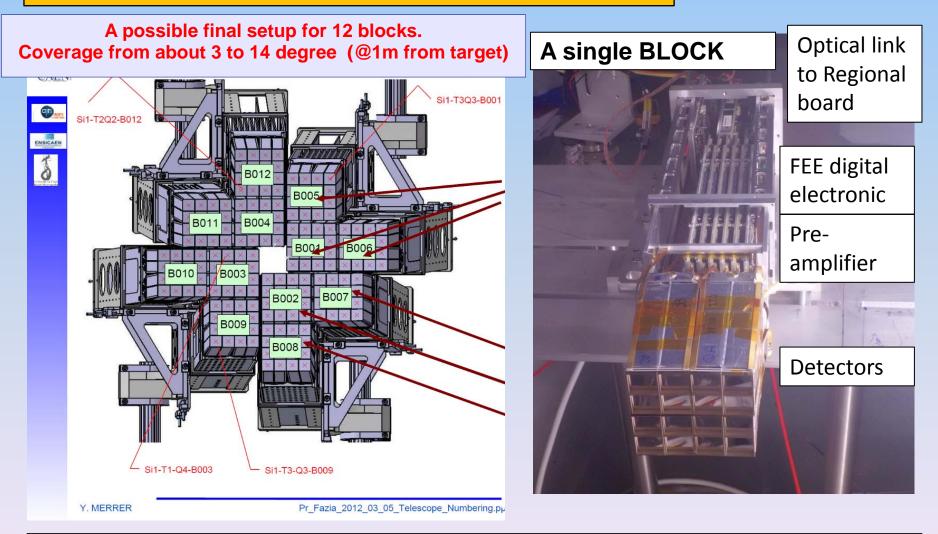


With a higher gain (as in LNL test showed in previous slide), we can see also some masses!

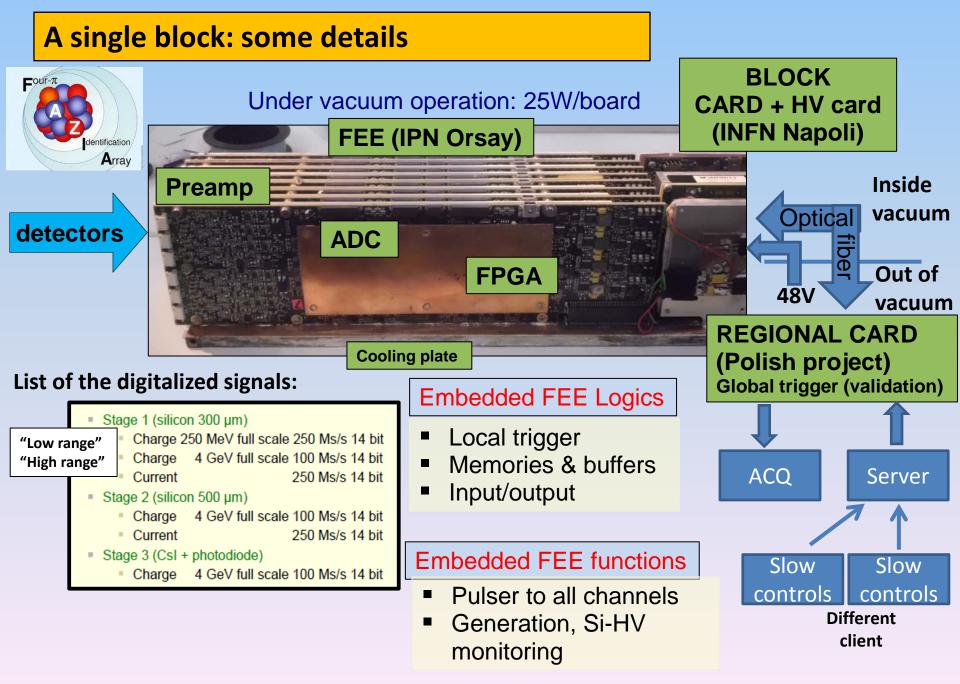
The R&D phases main results: Behavior of partially depleted silicons



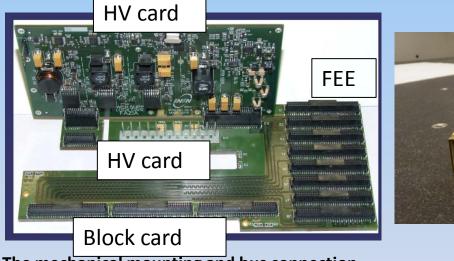
The Demonstrator (2012-2015): introduction



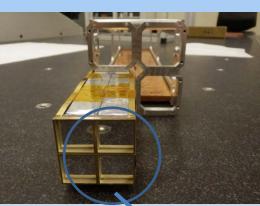
Block: the elementary structure formed by 4x4 telescopes (Si-300um + Si-500um +Csl). The FEE is mounted behind the detectors under vacuum!

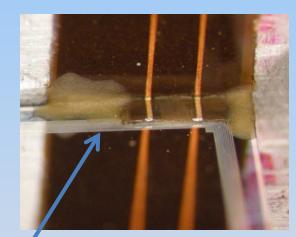


A single block: some details









Many detectors ... many Kapton to be connected!!

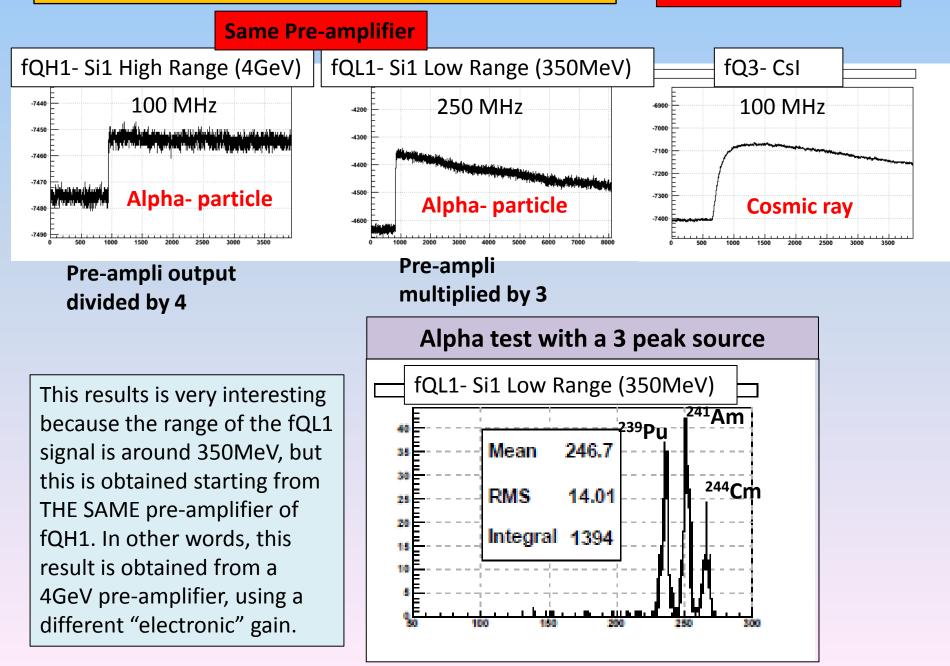


Kapton bounding! Kapton to transmit the signal



The Block: first "real" test (with 2 blocks)

LAST WEEK RESULTS

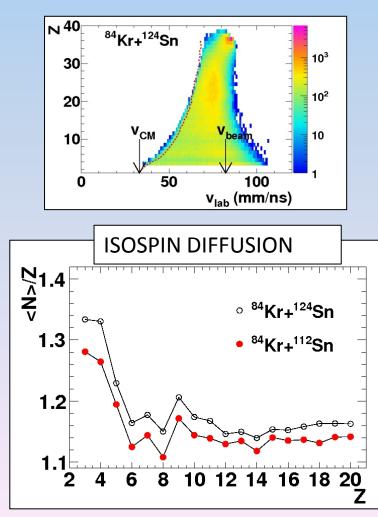


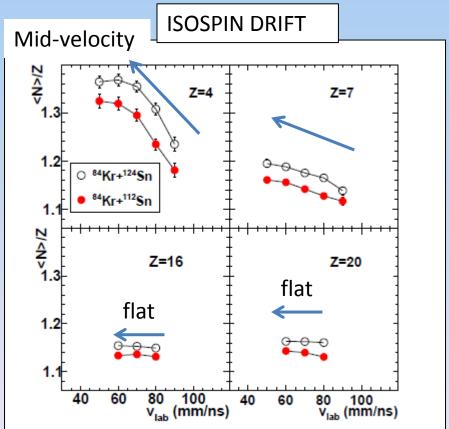
The experiments with FAZIA: ISOSPIN TRANSPORT

⁸⁴Kr + ^{112,124}Sn@35AMeV

Only one telescope of FAZIA with R&D configuration between θ =4.8° and 6° (grazing 4.1° and 4.0°)

S.Barlini et al., PRC 87 (2013), 054607





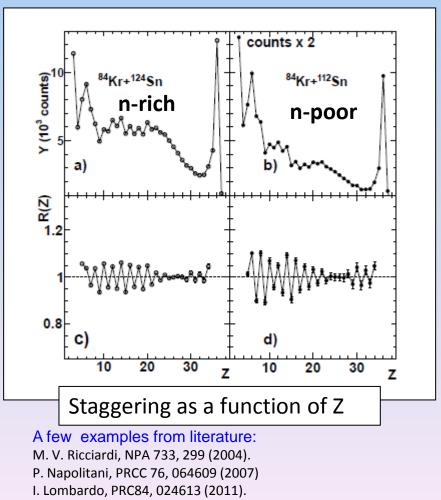
A few examples at comparable energies: Yang PRC60(1999) Lombardo PRC82(2010), JmodPE 20(2011), PRC84(2011); DeFilippo PRC86(2012); Galichet PRC80 (2009); Brown PRC87 (2013)

The experiments with FAZIA: STAGGERING

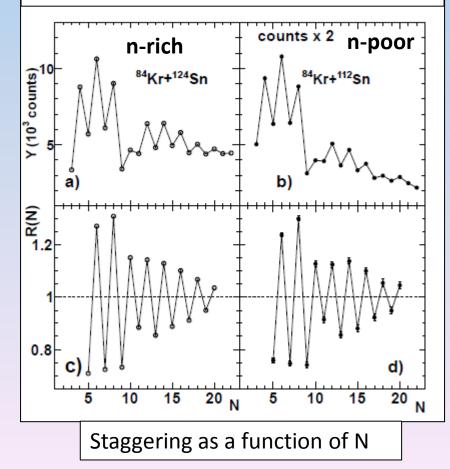
⁸⁴Kr + ^{112,124}Sn@35AMeV

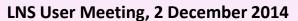
G. Casini, PRC 86, 011602 (2012). M.D'Agostino, NPA875 (2012) 139 Only one telescope of FAZIA with R&D configuration between θ =4.8° and 6° (grazing 4.1° and 4.0°)

S.Piantelli et al., PRC 88, 064607 (2013)



Thanks to the very good mass resolution, we can explore also this!

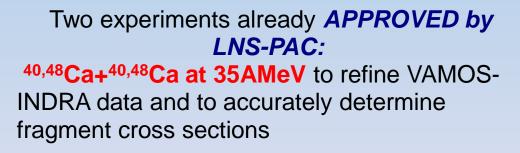




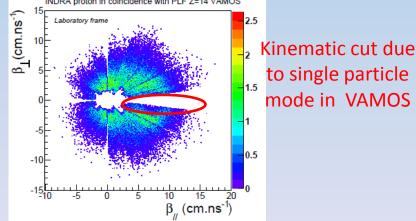
The experiments with FAZIA: approved experiments

2015: 4 (to 6) FAZIA blocks in a stand-alone configuration. Partial covering of forward angles up to about 15 degrees

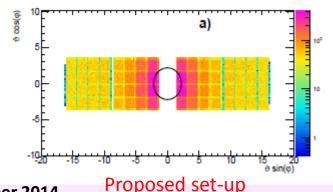
MAIN PHYSICS: QP features, QP fragmentation cross section



Signers (47) and affiliations (16)



⁷⁸Kr+^{46,50}Ti at 35AMeV in reverse kinematics, to investigate the role of isospin diffusion on QP sequential fission

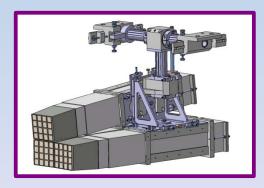


The experiments with FAZIA: future perspective

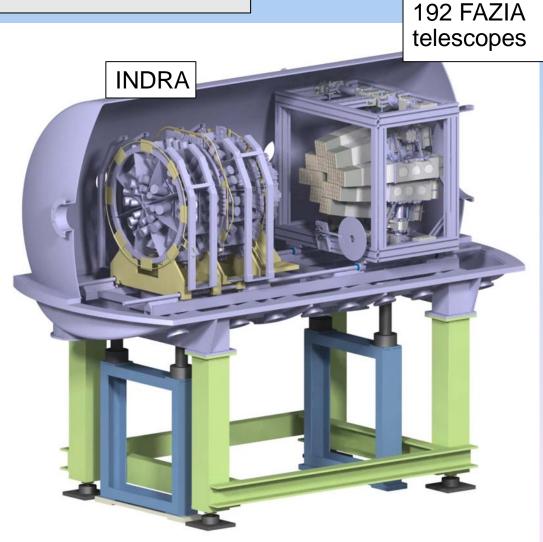
>2015:

FAZIA and the multidetector INDRA at GANIL

3-block supporting arm



...but also to LNS, LNL with stable and exotic beams!



CONCLUSIONS:

Many fragments are produced in heavy ions collisions, with wide range of energy and sizes and coming from several sources
The identification of these fragments in A,Z with low thresholds and over wide solid angles is needed to next exploration of EOS and of role of Symmetry Energy in excited systems

 FAZIA is building a new medium-size array of telescopes with unprecedented capabilities in terms of ion identification

 Specific recipes and solutions for production and use of Si detectors has been proposed and tested with success

Pulse shape analysis and fast sampling electronics are employed

 RIBs (besides stable energetic beams) will be useful to explore reaction dynamics far from stability

The FAZIA project in Europe: R&D phase

R. Bougault¹, G. Poggi^{2,3}, S. Barlini^{2,3}, B. Borderie⁴, G. Casini³, A. Chbihi⁵, N. Le Neindre¹, M. Pârlog^{1,6}, G. Pasquali^{2,3}, S. Piantelli³, Z. Sosin⁷, G. Ademard⁴, R. Alba⁸, L. Bardelli^{2,3}, M. Bini^{2,3}, A. Boiano¹², M. Boisjoli⁵, E. Bonnet⁵, R. Borcea⁶, B. Bougard¹, G. Brulin⁴, M. Bruno¹³, S. Carboni^{2,3}, M. Cinausero¹⁰, L. Ciolacu⁶, I. Cruceru⁶, M. Cruceru⁶, M. Degerlier¹¹, P. Desrues¹, J.A. Dueñas⁹, P. Edelbruck⁴, M. Falorsi², J.D. Frankland⁵, E. Galichet^{4,18}, K. Gasior¹⁵, F. Gramegna¹⁰, D. Gruyer⁵, A. Grzeszczuk¹⁵, M. Guerzoni¹⁴, H. Hamrita⁴, C. Huss⁴, M. Kajetanowicz⁷, K. Korcyl¹⁷, A. Kordyasz¹⁶, T. Kozik⁷, P. Kulig⁷, L. Lavergne⁴, E. Legouée¹, O. Lopez¹, J. Łukasik¹⁷, C. Maiolino⁸, T. Marchi¹⁰, P. Marini⁵, I. Martel⁹, Y. Merrer¹, L. Morelli¹³, F. Negoita⁶, A. Olmi³, A. Ordine¹², C. Pain¹, M. Pałka⁷, P. Pawłowski¹⁷, M. Petcu⁶, H. Petrascu⁶, E. Piasecki¹⁶, E. Rauly⁴, M.F. Rivet⁴, E. Rosato¹², E. Scarlini^{2,3}, G. Tortone¹², T. Twaróg⁷, S. Valdré^{2,3}, E. Vient¹, M. Vigilante¹², E. Wanlin⁴, A. Wieloch⁷, and W. Zipper¹⁵

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- ⁷ Jagiellonian University, Institute of Physics, ul. Reymonta 4, 30-059 Krakow, Poland.
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- ¹² Dipartimento di Fisica, Università di Napoli "Federico II" and INFN, Sezione di Napoli, Compl. Un. Monte S.Angelo ed. 6, 80126 Napoli, Italy.
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