Status of GEM Trackers for Super Bigbite Spectrometer at JLab

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On behalf of the SBS Collaboration

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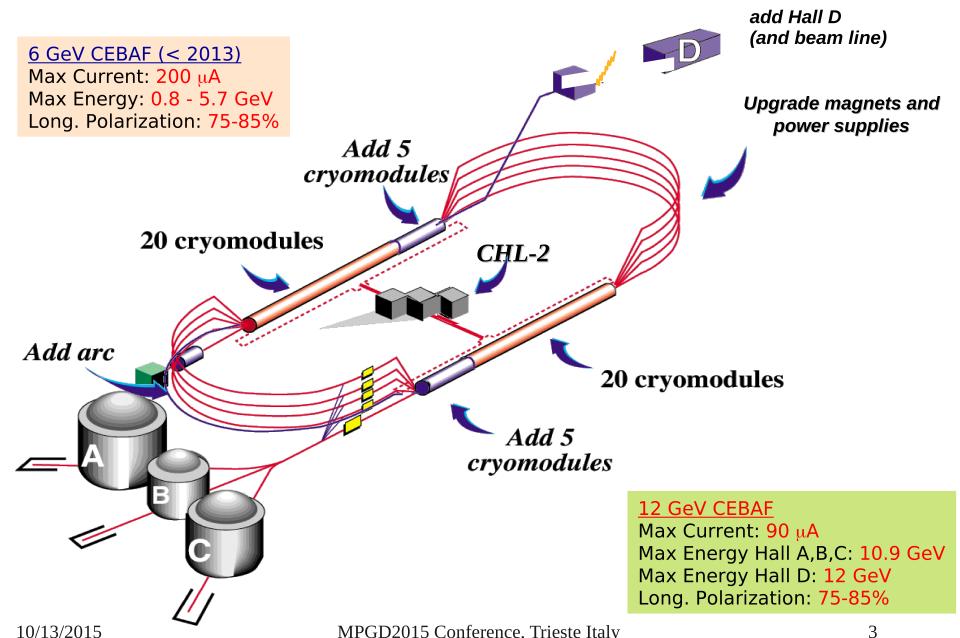
INFN Roma, Catania, Genoa

E. Cisbani, P. Musico

Outline

- GEM Trackers for Super Bigbite Spectrometer (SBS)
- Issues with large-area & light-weight GEM detectors
- APV25 readout electronics

The 12 GeV upgrade of CEBAF accelerator @ JLab



Physics program in Hall A for the CEBAF 12 GeV era @ JLab

SBS physics program

GEP: 12 (GeV/c)²
 GMN: 13.5 (GeV/c)²
 GEN: 10 (GeV/c)²

SSA in nSIDIS: 30,000 gain vs HERMES

➤ A1n/d2n – gain ~ 20-30 compared with HMS/SHMS

> TDIS meson DIS

➤ WACS-ALL, full proposal, 100x gain in productivity

GEnRP, ready for full proposal,10+x gain in productivity

 \triangleright pol H(γ , ϕ p), H(γ , π ° p)

➤ PVDIS – gain 10-15 compared with two HRSs

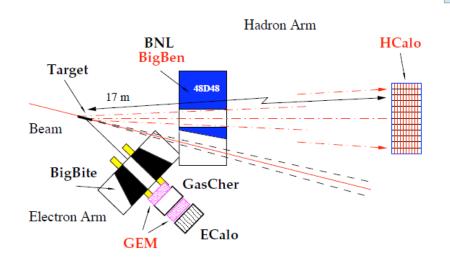
➤ A1p/d2p – gain ~20-30

➤ D(e,e'd) - A,T20

> J/Psi as gluon probe of QCD - well matched to BB/SBS

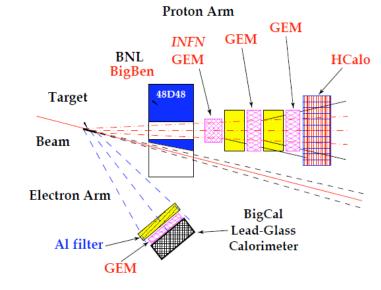
 \triangleright A(e,e'p), A(e,e' $\pi^{+/-}$)

Neutron form factors, E12-09-016 and E12-09-019

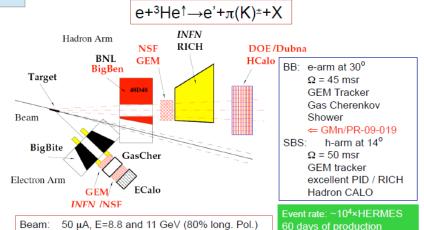


Proton form factors ratio, GEp(5) (E12-07-109)

Nucleon Form FactorsSIDIS, TMDsNucleon Structures



SIDIS experiment (conditionally approved)



Target: 65% polarized 3He \Leftarrow GEn(2)/PR-09-016

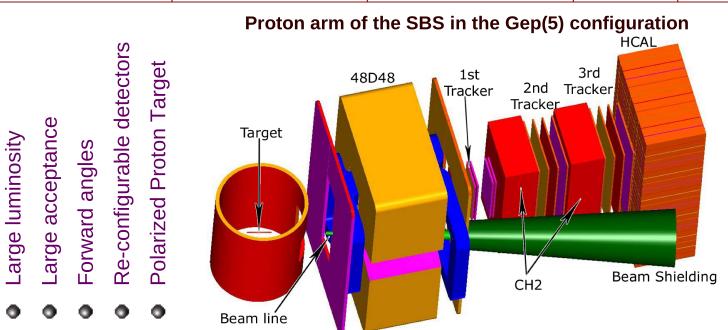
⇒ Luminosity: 1.4×10³⁷ cm⁻²s⁻¹, 0.05 sr

1/10 of proton HERMES

expected stat. accuracy:

Requirements for the Super Bigbite Spectrometer (SBS)

Experiments	Luminosity (s·cm²)⁻¹	Tracking Area (cm²)	Resolution		
			Angular (mrad)	Vertex (mm)	Momentum (%)
GMn - GEn	up to 7·10 ³⁷	40x150 and 60x200	< 1	<2	0.5%
GEp(5)	up to 8·10 ³⁸	40x150, 60x200	<0.7	~ 1	0.5%
Most demanding		and 80x300	~1.5		
SIDIS	up to 2·10 ³⁷	40x150 and 60x200	~ 0.5	~1	<1%
	High rate	Large area	Spatial resolution < 100 microns		

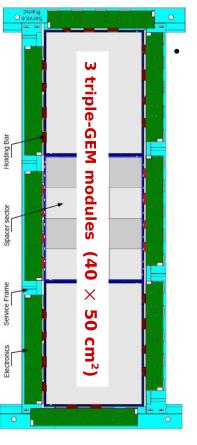


High photon backgound up to 250 MHz/cm² and electron background 160 kHz/cm²

10/13/2015 MPGD2015 Conference, Trieste Italy

SBS GEM Trackers

Front Tracker Layer,,INFN (Rome, Catania)

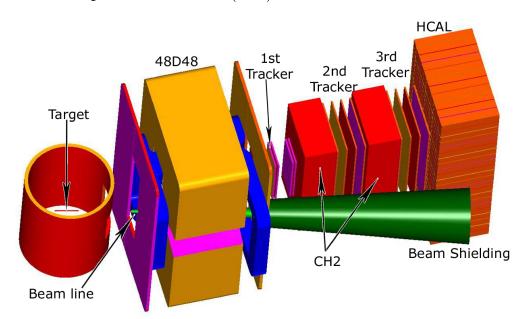


Front Tracker (FT): Track of the recoil protons

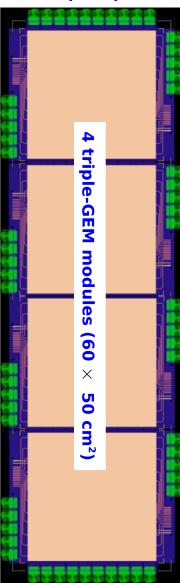
- 1st tracker: 6 GEM layers, active area of $150 \times 40 \text{ cm}^2$
- Each layers: vertical stack of 3 GEM modules $(50 \times 40 \text{ cm}^2)$
- Total production of 18 modules

Back Tracker (BT): Proton Polarimetry

- Polarization of the recoil protons
- 2nd & 3rd Trackers: 10 layers, active area of 200 \times 60 cm²
- Each layer: vertical stack of 4 GEM modules $(60 \times 50 \text{ cm}^2)$
- Total production of 40 (+5) modules



Back Tracker layer (UVa)



Assembly of the SBS Triple-GEM modules

Standard COMPASS triple-GEM

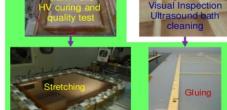
Assembly steps of the Front Tracker GEMs

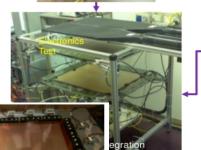
Module production fully established in INFN-Catania
Electronics preliminary QA in Genoa

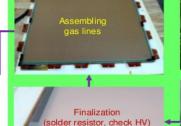
Module integration and characterization in INFN-Sanità



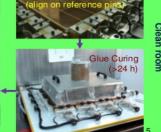
ation by rad. osmics



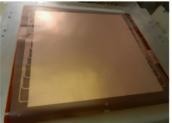




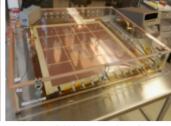
SBS Coll. Meeting - Front Tracker GEM



Parts of the Back Trackers
Polarimeter GEM



GEM foil with the visible contact of the HV sectors

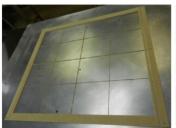


GEM foil in the N₂ box for leakage current test



GEM foil on the mechanical stretcher

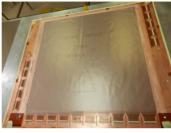
Standard COMPASS 2D readout board



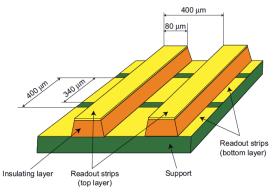
Support frame for GEM with 300 µm spacers inside the active area



Frames on a custom holder for cleaning in Ultrasonic bath

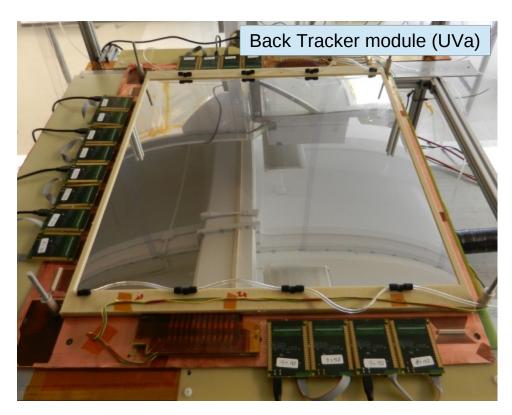


Two dimensional flexible readout board

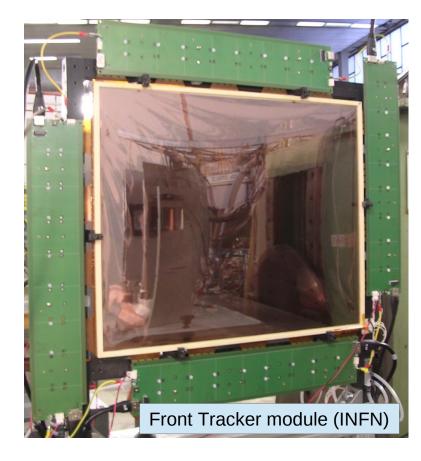


SBS GEM modules

- Spatial resolution < 0.1 mm; high radiation tolerance
- Lightweight triple-GEM detectors (0.7% radiation length)
- Readout layer: 2D x/y strip ala COMPASS (0.4 mm pitch)
- APV25-based electronics with VME64x modules (total channels > 120K channels)







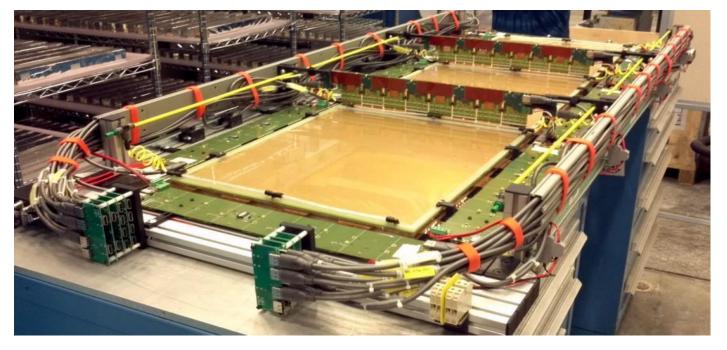
Production Status

Front Tracker GEMs

- 18 modules to be completed by mid 2017
- 8 modules already assembled with 4 tested
- One full layer integrated with APV25 cards @ JLab
- 4 layers expected by end 2016



Carbon fiber Holding frame
More compact and more rigid option
minimize thermal deformation



Production Status

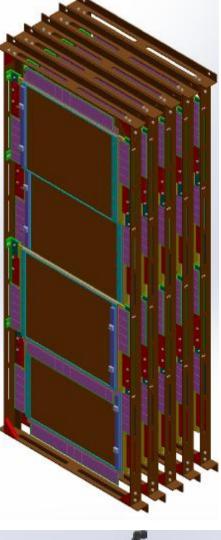
Back Tracker GEMs

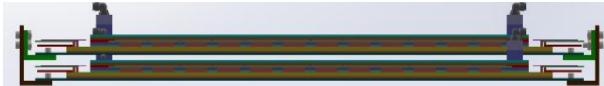
- 45 modules to be completed by mid 2017
- Production rate of 2 modules / month.
- 19 modules successfully tested as of Oct. 2015



Holding frame:

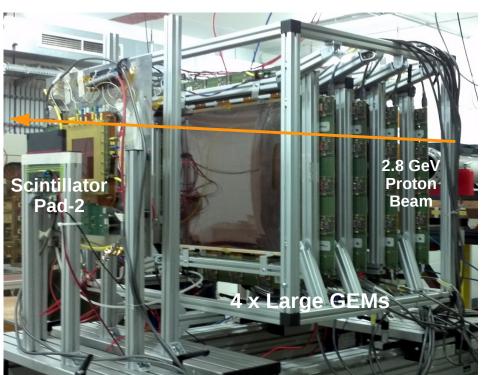
- on the frame (bottom plane), other 2 modules on L-shape (top plane)
 - This minimizes dead area
 - And allow easy replacement of the modules and of the FE cards
- The holding frames are under production @ JLab

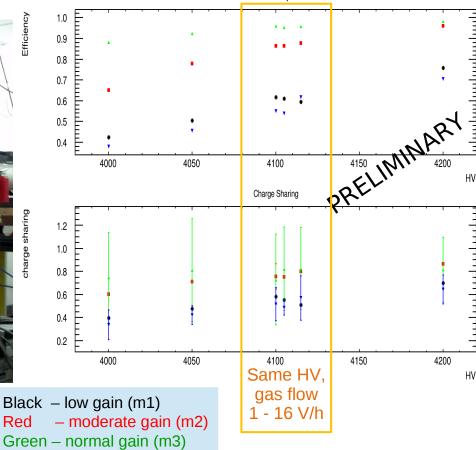




Performance in Test Beam

FT GEM modules high Intensity Proton beam in Julich COSY Test Beam (Oct. 2014)





- Study GEM response in high intensity proton beam (small spot ~ few cm²)
- Different dividers on different module
- Investigate HV and gas flow

- Efficiency slightly affected by the high beam intensity
- No noticeable effects from gas flow rate

- very low gain (m4)

Performance in Test Beam

SRS + SRU Readout using DATE @ FTBF

BT GEM modules in Test Beam @ FNAL (Oct. 2013)

- Two SBS BT GEM prototypes tested at FTBF
- APV25-SRS electronic tested at trigger rate 400 Hz
- Data analysis for spatial resolution, gain efficiency, gain uniformity, timing of the APV25 signal ...
- FNAL test beam data reveals big issues (Gas flow, Quality of X/Y readout board etc)

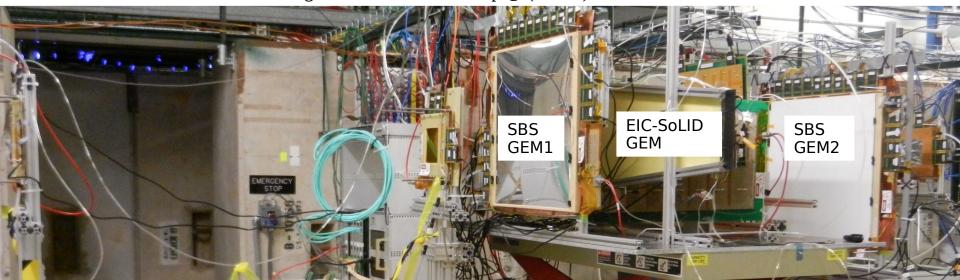




- · 64 APV's read out by SRS
- · Acquiring data from FECs with an SRU
- Current DAQ rate is ~150 Hz
- · Using 6-9 25ns time slices for digitization
- Beam structure: 4s spills, 1min rep. time, 10 20k particles/spill
- · Trigger: coincidence of 3 scintillators

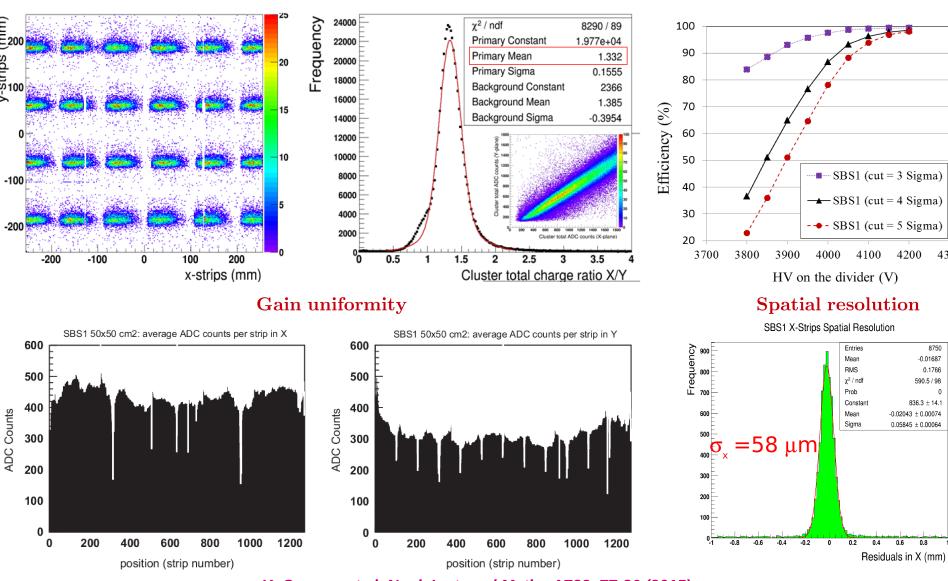


Large GEM Test Beam Setup @ (FNAL) UVa & FIT



Performances in test Beam

Charge sharing



K. Gnanvo et al. Nucl. Inst. and Meth., A782, 77-86 (2015)

Efficiency curve vs. HV

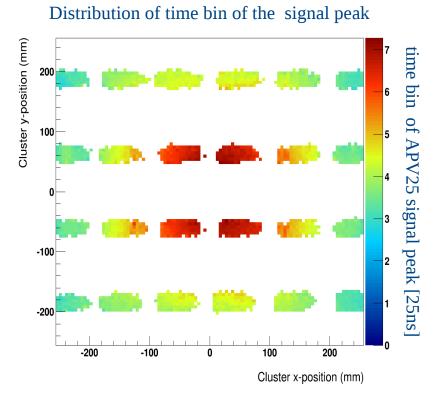
Hadron beam reconstruction

Outline

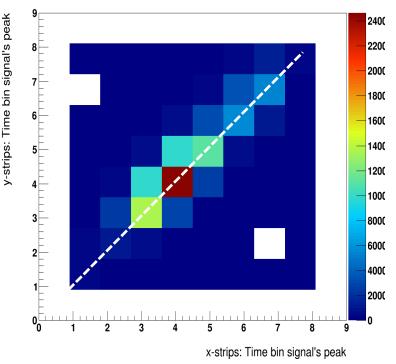
- GEM Trackers for Super Bigbite Spectrometer SBS
- Issues with large-area & light-weight GEM detectors
- APV25 readout electronics

Analysis of the APV25 signal timing from the FNAL Test Beam data

- We looked at the spatial distribution of the the APV25 signal peak w.r.t. the trigger delay (arbitrary reference)
- Strong spatial non uniformity of the signal timing → Induced charge signal collected later by the readout strips
 in the center of the detector than at the edges.
- Difference as high as 4 time bins (100 ns) between center and edges
- Excellent timing correlation of the signal in x-strips and y-strips → the readout electronics not the source



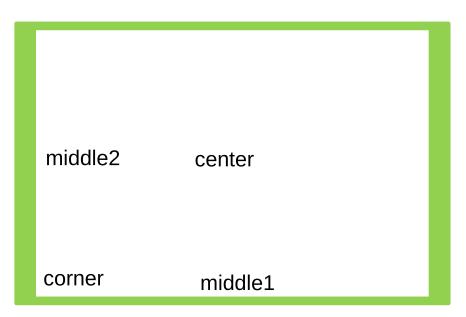


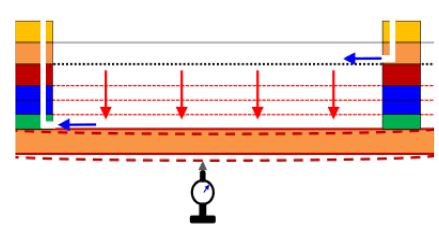


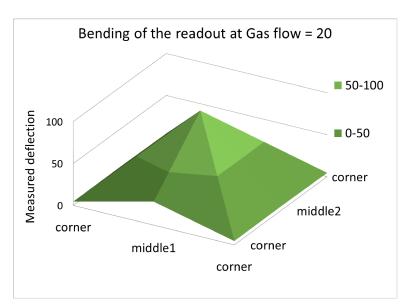
Cause: Deformation of the readout board due to over pressure caused by the gas flowing inside the detector

Measurement of the deformation of the readout board

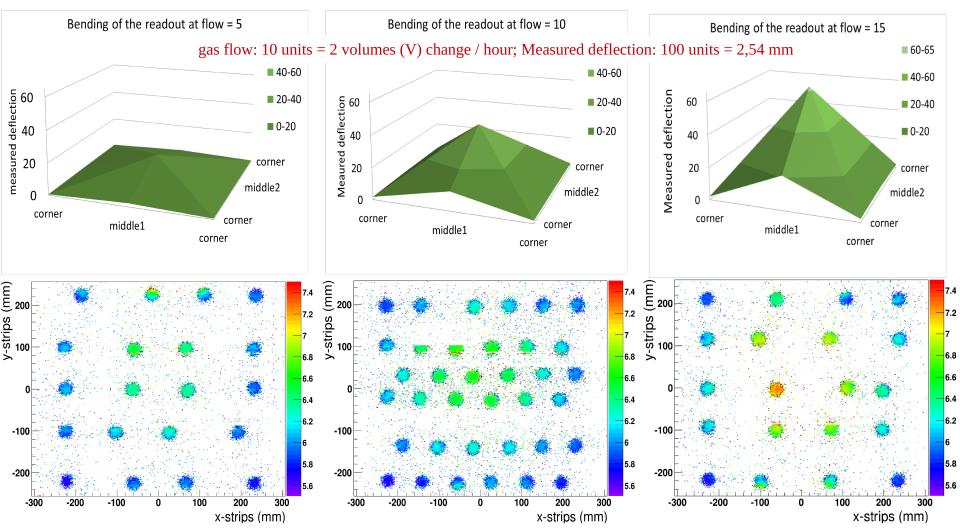
- Setup of a test (see cartoon on the left) to measure the bending of the readout board (honeycomb support) with the Ar/CO_2 flow rate inside the chamber
- Measurement were taken at 4 location on the bottom side of the honeycomb support
- The measured deflection of 100 units is equivalent to 2.54 mm
- A gas flow = 10 units represent about 2 volumes (V) change / hour in the GEM chamber (V = 3.6 L)







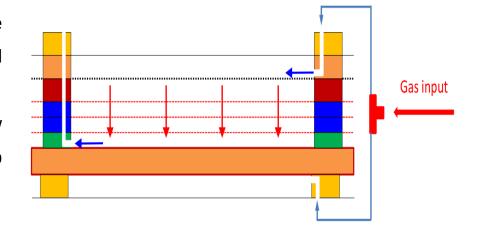
Measurement of the deformation of the readout board

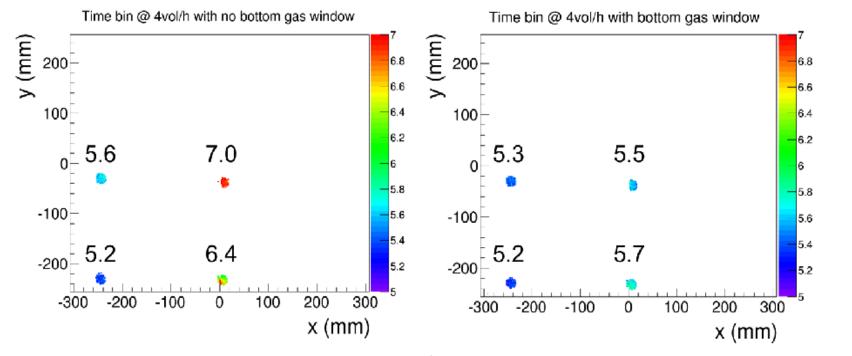


- APV25 signal peak time bin measured with Sr90 source at different flow rates.
- Amplitude of the non uniformity depend on the gas flow (more precisely built-up pressure in the chamber)
- Clear correlation between the time bin of signal peak and the deformation of the readout board

Solution: Compensate the deformation of the readout board with a bottom gas volume

- Adding the bottom gas window significantly reduce considerably the spatial non uniformity of the signal speak time bin at high gas flow rate
- In addition, we also slightly change the gas flow design of the chamber to reduce the pressure built-up inside the chamber





Entrance gas window foil collapse

Problem

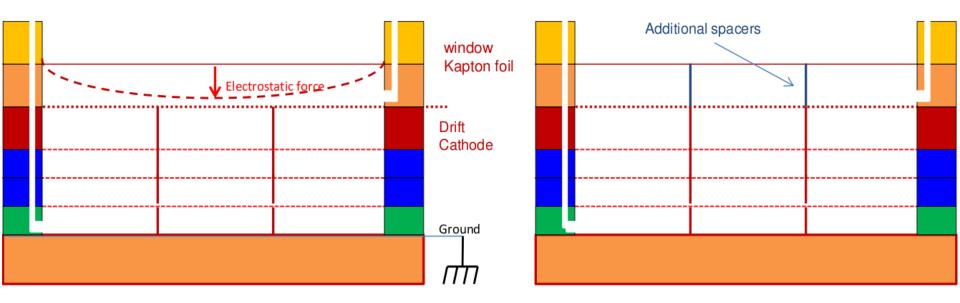
- High particle rate over a large area of the detector

 ⇒ charging up of the Kapton foil

 ⇒ Strong
 electrostatic attraction between gas window & drift cathode
- Strong distortion of the APV25 signal (timing and shape)

Initial proposed solution

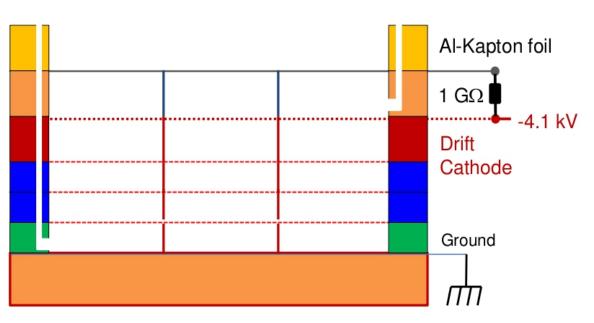
- A simple initial fix was to add some spacers in the gas window region of the chamber
- We saw a improvement but not sure about long term stability of the fix in high rate condition

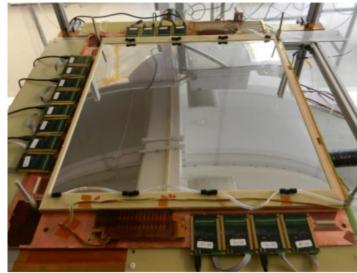


Entrance gas window foil collapse

Final proposed solution

- Use aluminized gas window foil and set it to the same potential as the drift cathode → Faraday
 cage like to prevent charges accumulation on the gas window as well as the top layer of the drift
- Tested with SBS-BT-GEM with x-ray source at high rate > 1 MHz /cm2 equivalent MIP.
 - Without the HV on the gas window \Rightarrow foil collapse after a few hours of x-ray exposure
 - With the HV on, we did not observe any collapse after 5 days of almost continuous exposure

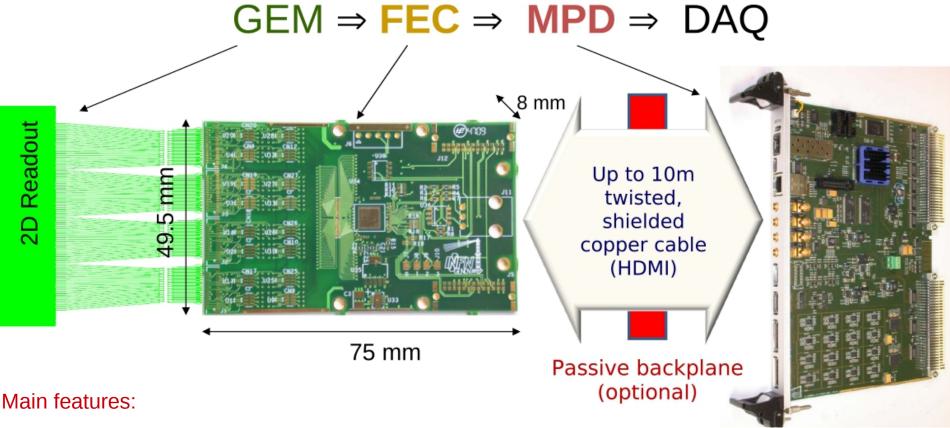




Outline

- GEM Trackers for Super Bigbite Spectrometer SBS
- Issues with large-area & light-weight GEM detectors
- APV25 readout electronics

Readout electronics for SBS GEM Trackers



- Use analog readout APV25 chips (> 100 k Channels)
- 2 actives components: APV25 Front end cards & VME64x module: Multi Purpose Digitizer (MPD)
- HDMI cables to transfer data between these two components

APV25 FE cards and Back planes

Different versions of the APV25 FE card produced

- v 4.10 with ZIF connectors for the FT Modules
- v 4.11 with Panasonic connectors for BT Modules

Different versions of the back planes

- 2 types for the Back Tracker GEMs: 5 and 12 slots
- 1 type for the Front Tracker GEM: long 5-slots

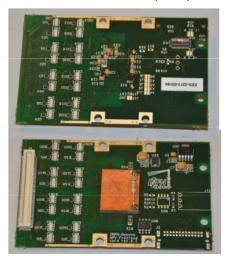
v 4.0

Front Tracker GEM



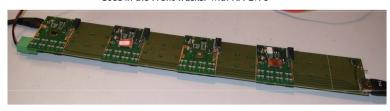


v 4.11 Back Tracker GEM (UVa)

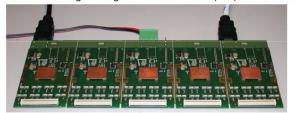


Backplanes

Used in the Front Tracker with RH LVPS



Original design for the Back Tracker (UVa) with RH LVPS





Multi Purpose Digitizer (MPD) card

- VME64x board perform the digitization of analog signals from the FE cards and handle the slow control signals
- DDR2 (128 MB), 110 MHz system clock
- Compliant with JLab VME64x VITA 41 (VXS) standard
- 6 HDMI-A connectors for data and sow control signals

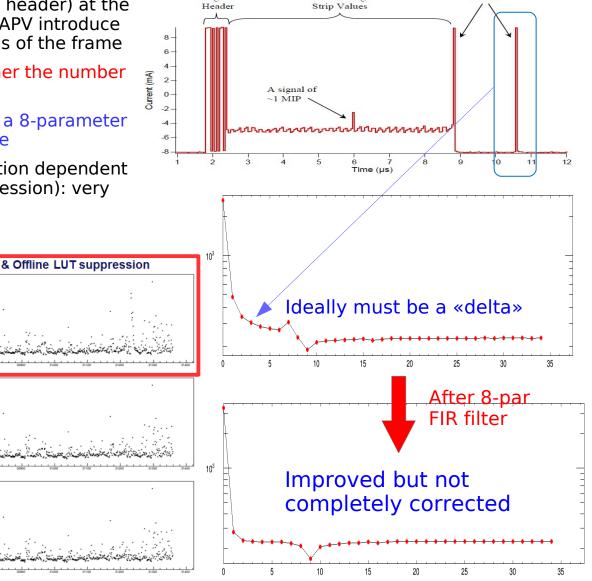
MPD block diagram MPD Board Power Supply Clock Generator VME64x compliant Live Insertion with VXS extension Remote Configuration 4 x HDMI-A ALTERA SN74VMEH22501A ARRIA GX FPGA DAQ EP1AGX60DF780C6N 2 x HDMI-A rigger & Syster Clock & Fast cmd LVDS Distributor **CTRL** I²C Distributor 128 MB SDRAM 1250 Mb/s (FTLF1318P2BCL) 10-100 PHY **PMC** SD-Card interface MT47H128M8 DP83848J 2125 Mb/s (FTLF8524P2BNL) Expansion SFP Optical Transceiver



Long (23 m) HDMI cable effects on APV25 analog signal

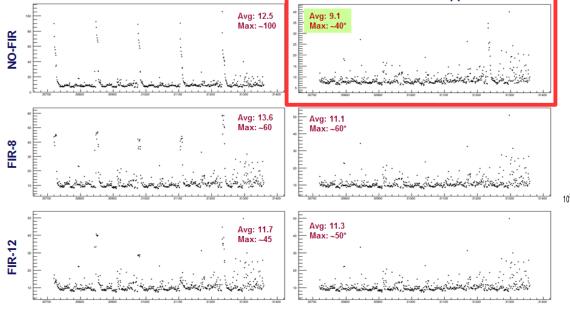
Digital

- The large «binary» information (digital header) at the beginning of the analog signals of the APV introduce a large noise on the first (~20) channels of the frame
- Longer the cable larger the noise, higher the number of channel involved
- Belle (2012 JINST 7 C01082) proposed a 8-parameter FIR filter (12 m long cables) in firmware
- We added an off-line pedestal subtraction dependent on the digital header value (LUT suppression): very noisy channels largely recovered



128 Analog

Ticks



Summary

- The Hall A equipment for the 12 GeV Upgrade of the CEBAF at JLab is the Super Bigbite
 Spectrometer (SBS).
- The 3 tracking stations of the SBS are equipped with large area and light weight GEM detectors.
- The construction and commissioning of all 60 GEM chambers is ongoing at University of Virginia and at INFN Catania & Roma.
- The MPD system, an APV25-based readout electronics is developed at the University of Genoa to read out the GEM trackers.
- The MPD is compliant with the JLab VME64x VITA 41 (VXS) standard