

Dark Matter: Collider vs. Direct Searches M33 rotation curve

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La Thuile

15q.cm



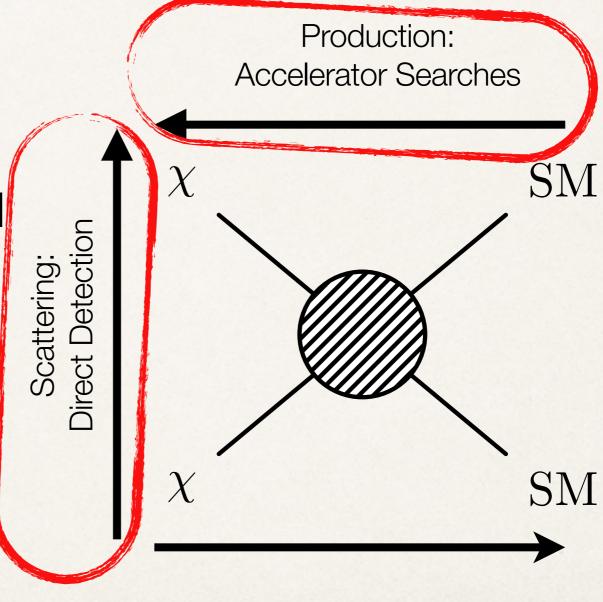
Center for Astroparticle Physics GENEVA

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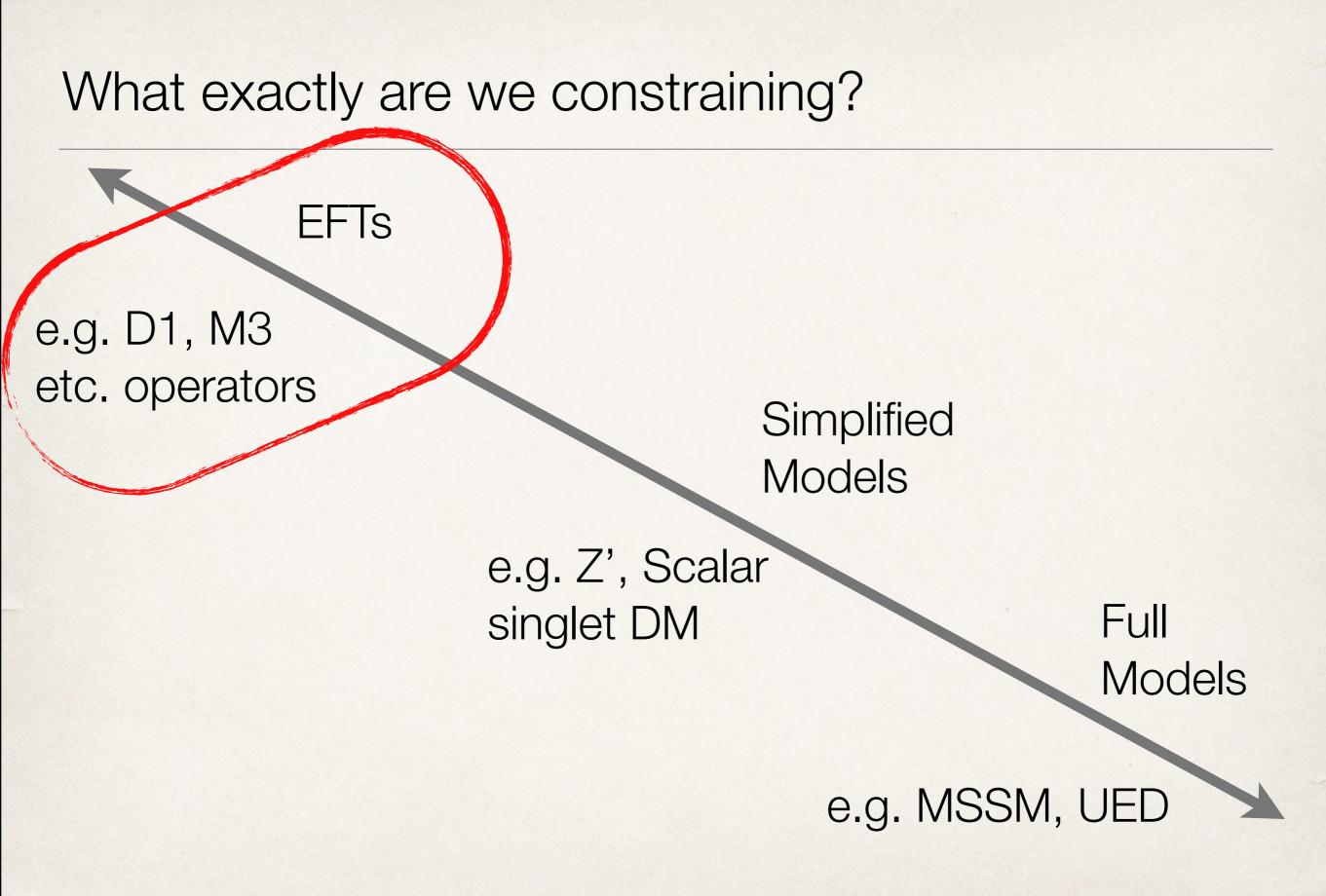
R (kpc)

Dark Matter Searches

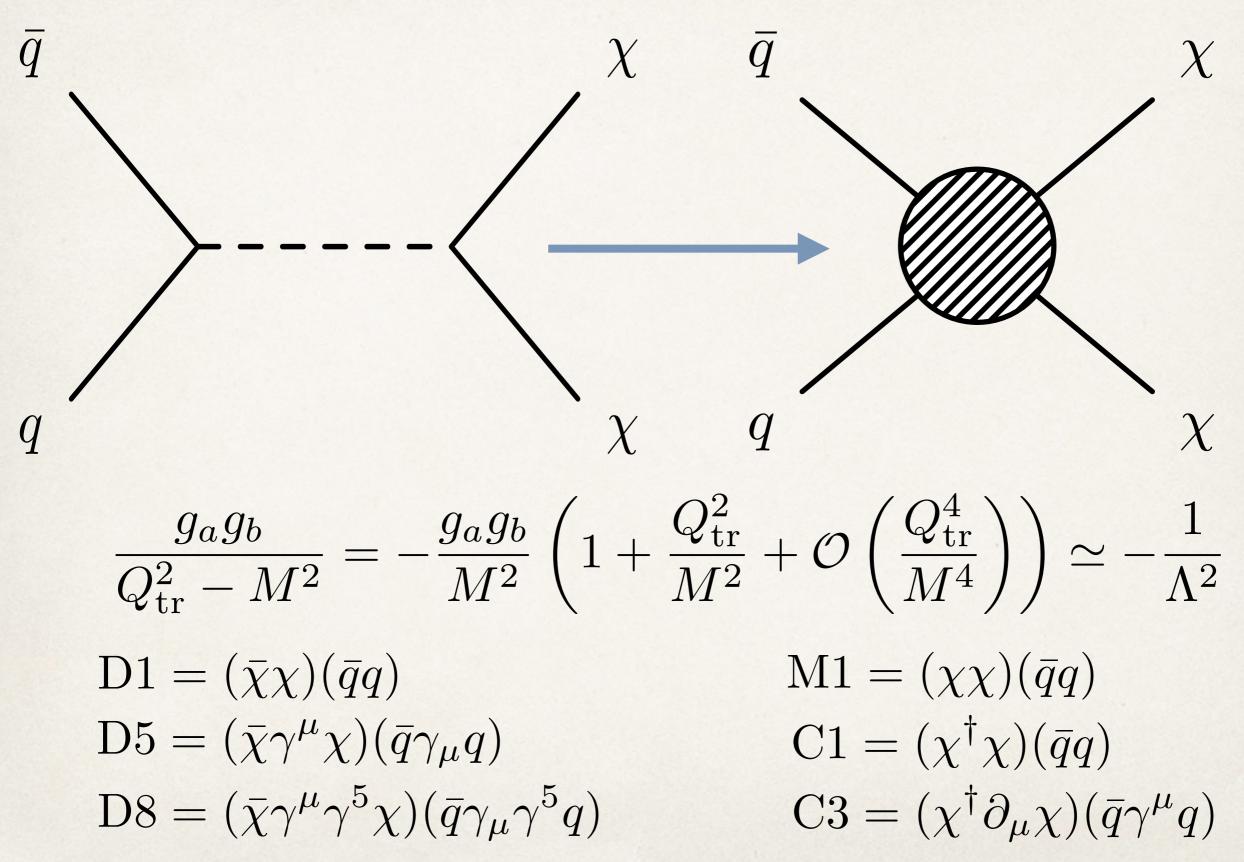
- Regardless of the underlying particle physics, there are several complementary ways to search for DM
- Each technique has its own strengths and challenges

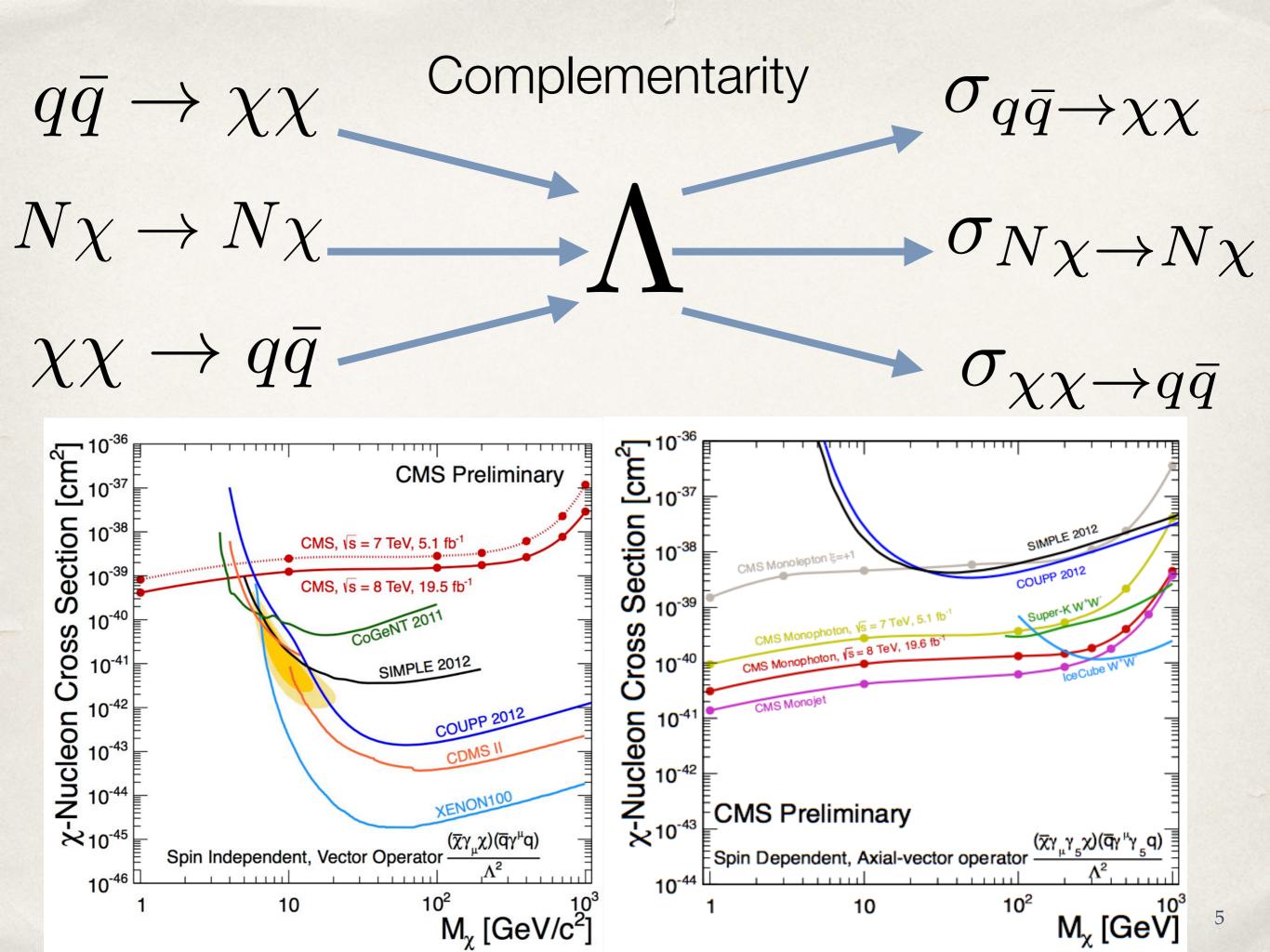


Annihilation: Indirect Detection



Effective Field Theories





Spin Independent scattering rate:

$$\frac{dR}{dE_R} = \frac{\sigma_{\rm SI}}{2m_{\chi}\mu_{\chi N}} (Z + \frac{f_n}{f_p}(A - Z))^2 F^2(E_R) \int_{v_{\rm min}}^{\infty} \rho_0 \frac{f(\vec{v})}{|\vec{v}|} d^3 v$$

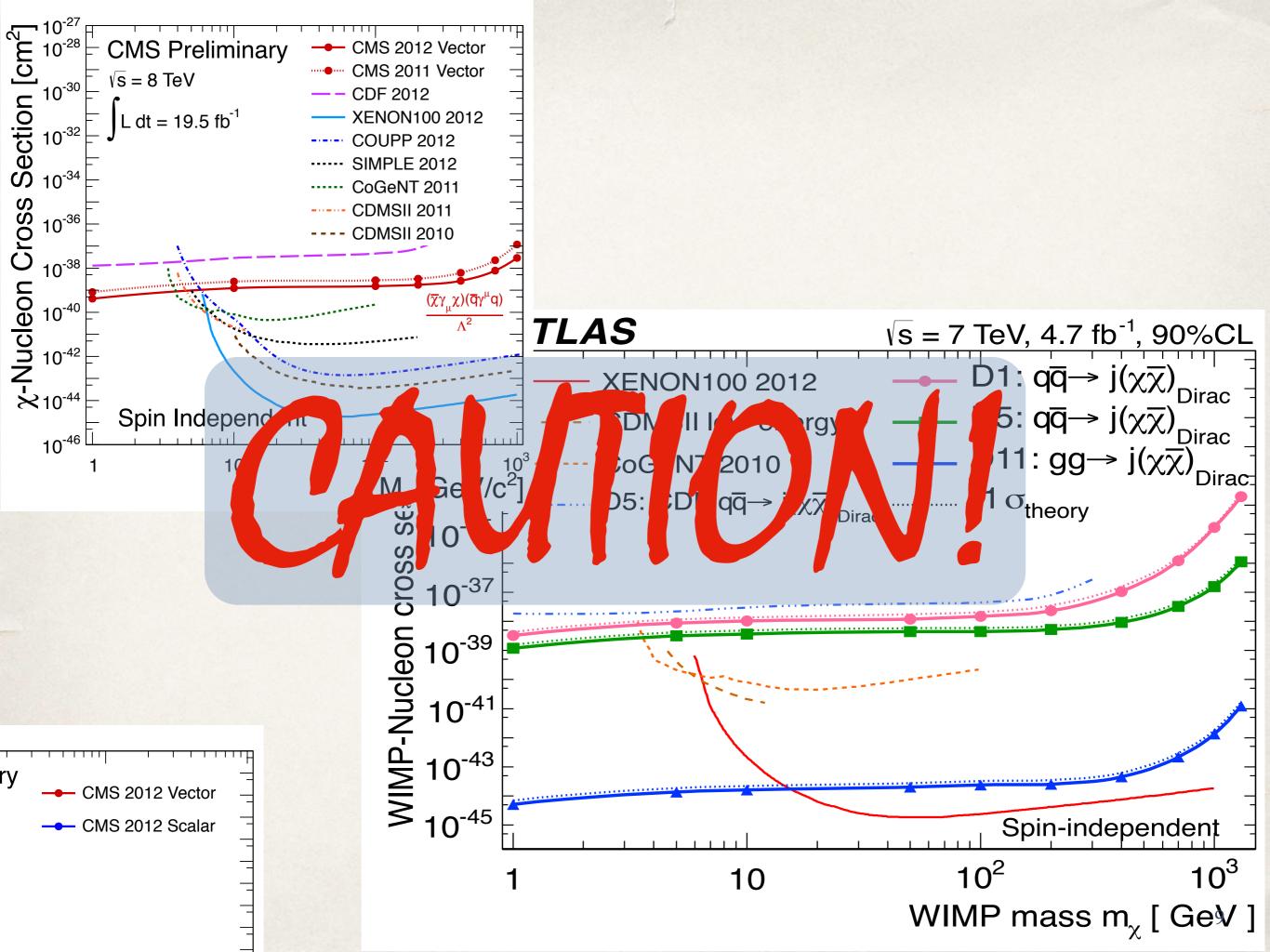
applies to the scalar DM-nucleon coupling; But other DM-nucleon operators lead to the same non-relativistic operator, $(\bar{\chi}\chi)(\bar{N}N), (\bar{\chi}\gamma^{\mu}\chi)(\bar{N}\gamma_{\mu}N), (\phi^{*}\phi)(\bar{N}N), i(\phi^{*}\overleftrightarrow{\partial_{\mu}}\phi)(\bar{N}\gamma^{\mu}N)$

so constraints on $\sigma_{\rm SI}$ still apply after renormalisation. Similarly, spin dependent scattering comes from operators:

$$(\bar{\chi}\gamma^{\mu}\gamma^{5}\chi)(\bar{N}\gamma_{\mu}\gamma^{5}N),(\bar{\chi}\sigma^{\mu\nu}\chi)(\bar{N}\sigma_{\mu\nu}N)$$

 $\bar{\chi}\chi\,\bar{N}N = f\left(\bar{\chi}\chi\,\bar{q}q,\bar{\chi}\chi\,G^a_{\mu\nu}G^a_{\mu\nu}\right) \qquad \bar{\chi}\gamma^\mu\chi\,\bar{N}\gamma_\mu N = f\left(\bar{\chi}\gamma^\mu\chi\,\bar{q}\gamma_\mu q\right)$ χ χ χ χ \boldsymbol{Q} $\phi^* \phi \, \bar{N}N = f(\phi^* \phi \, \bar{q}q, \phi^* \phi \, G^a_{\mu\nu} G^a_{\mu\nu}) \quad \phi^* \stackrel{\leftrightarrow}{\partial_\mu} \phi \, \bar{N}\gamma^\mu N = f(\phi^* \stackrel{\leftrightarrow}{\partial_\mu} \phi \, \bar{q}\gamma^\mu q)$

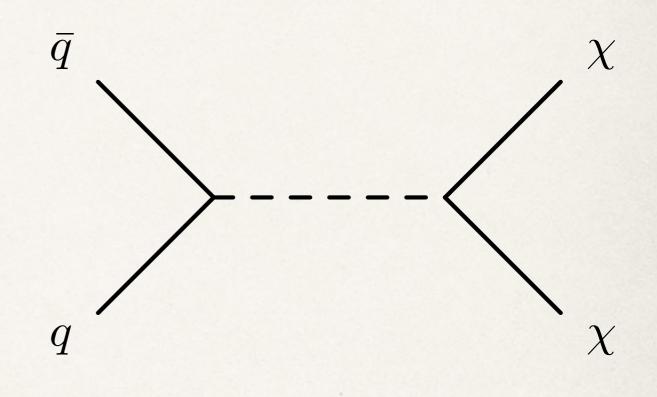
 $(\bar{\chi}\chi)(\bar{q}q)$ $(\bar{\chi}\gamma^{\mu}\chi)(\bar{q}\gamma_{\mu}q)$ $(\bar{\chi}\chi)(G^a_{\mu\nu}G^a_{\mu\nu})$ σ_{SI} $(\phi^*\phi)(\bar{q}q)$ $\left(\phi^* \stackrel{\leftrightarrow}{\partial_{\mu}} \phi\right) \left(\bar{q}\gamma^{\mu}q\right)$ $(\phi^*\phi)(G^a_{\mu\nu}G^a_{\mu\nu})$



Fundamental Limit to Validity

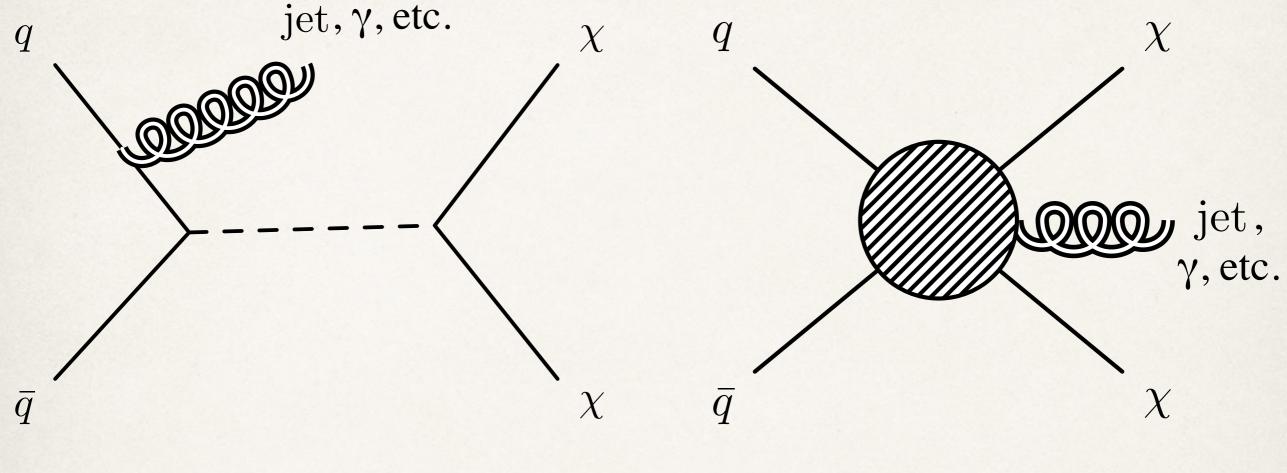
• In s-channel:

 $Q_{\rm tr} > 2m_{\rm DM}$ $M > 2m_{\rm DM}$ $\Lambda = \frac{M}{\sqrt{g_a g_b}} \ge \frac{M}{4\pi}$

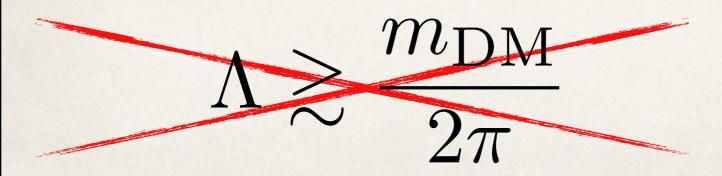


 $m_{\rm DM}$ 1 9-

Effective Field Theories



 $\frac{g_a g_b}{Q_{\rm tr}^2 - M^2} = -\frac{g_a g_b}{M^2} \left(1 + \frac{Q_{\rm tr}^2}{M^2} + \mathcal{O}\left(\frac{Q_{\rm tr}^4}{M^4}\right) \right) \simeq -\frac{1}{\Lambda^2}$



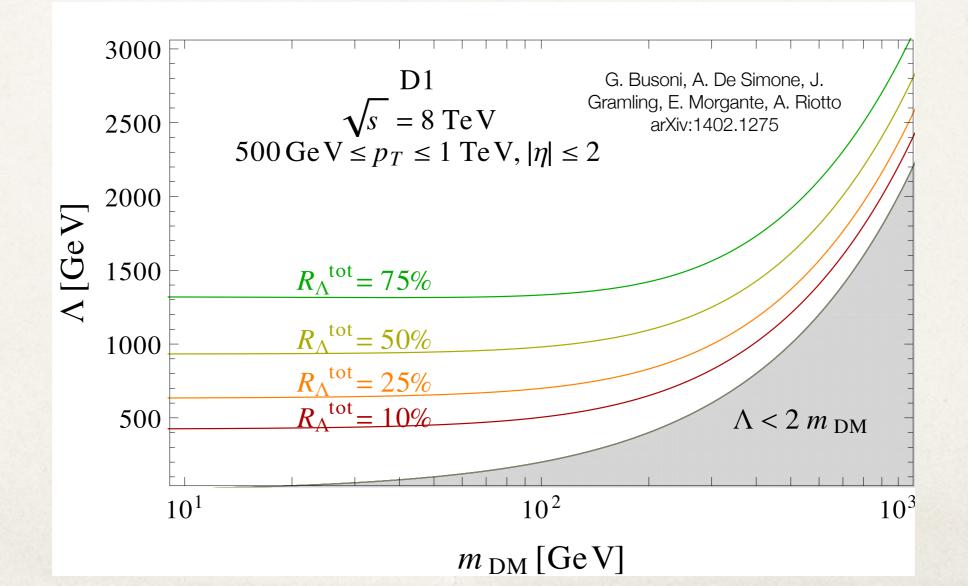
< M

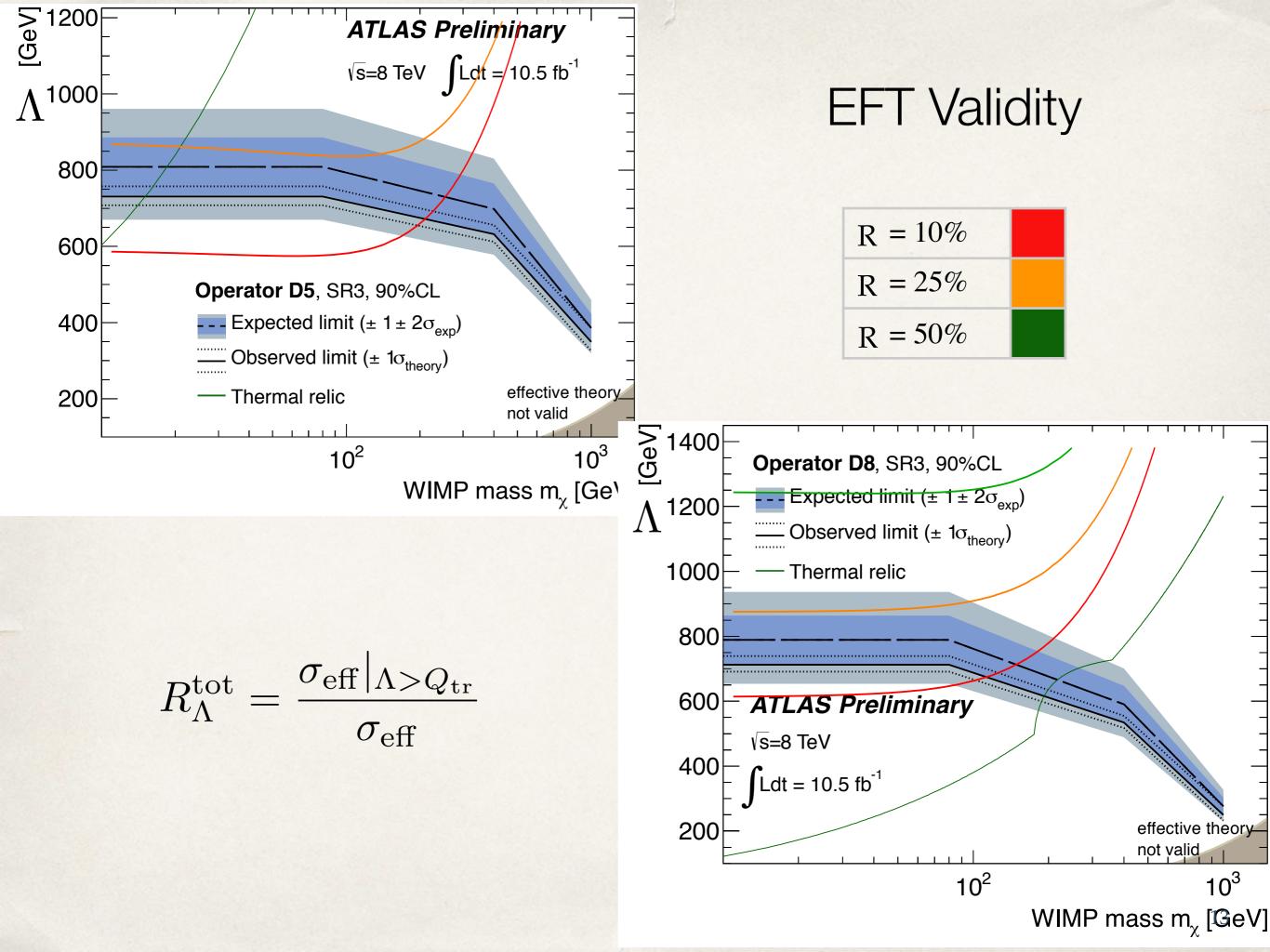
Measuring the Validity

Calculate or measure the fraction of events that pass the condition $Q_{tr} < \Lambda$, for a given choice of Λ and m_{DM} , and assuming $g \ge 1$.

$$R_{\Lambda}^{\text{tot}} = \frac{\sigma_{\text{eff}} |\Lambda > Q_{\text{tr}}}{\sigma_{\text{eff}}}$$

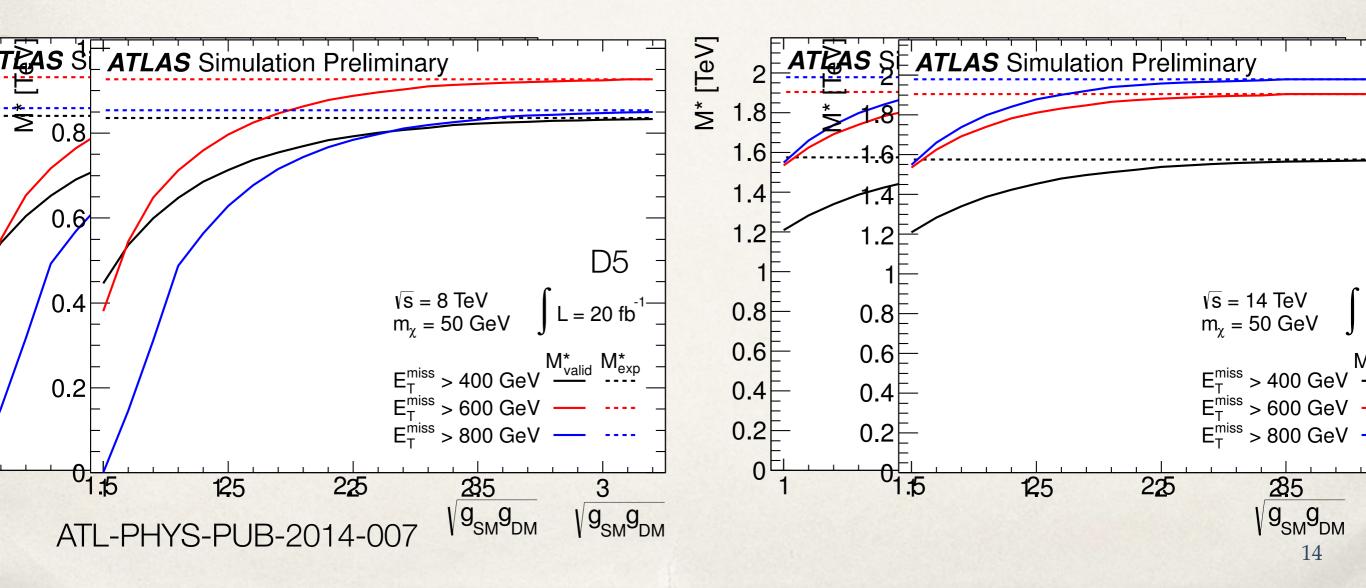
D1 = $(\bar{\chi}\chi)(\bar{q}q)$ $q\bar{q} \to \chi\chi + \text{jet}$ $Q_{\text{tr}} = p_q + p_{\bar{q}} - p_{\text{jet}}$



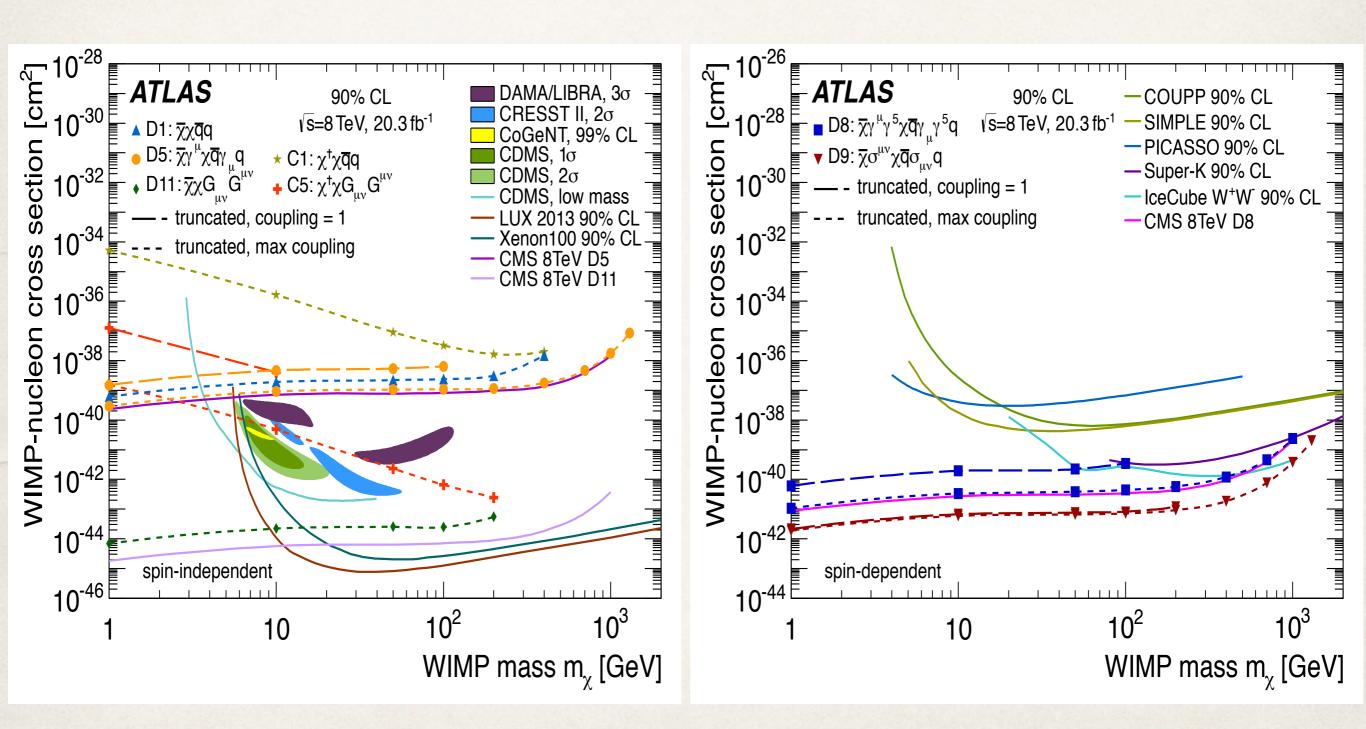


Rescaling the Limits

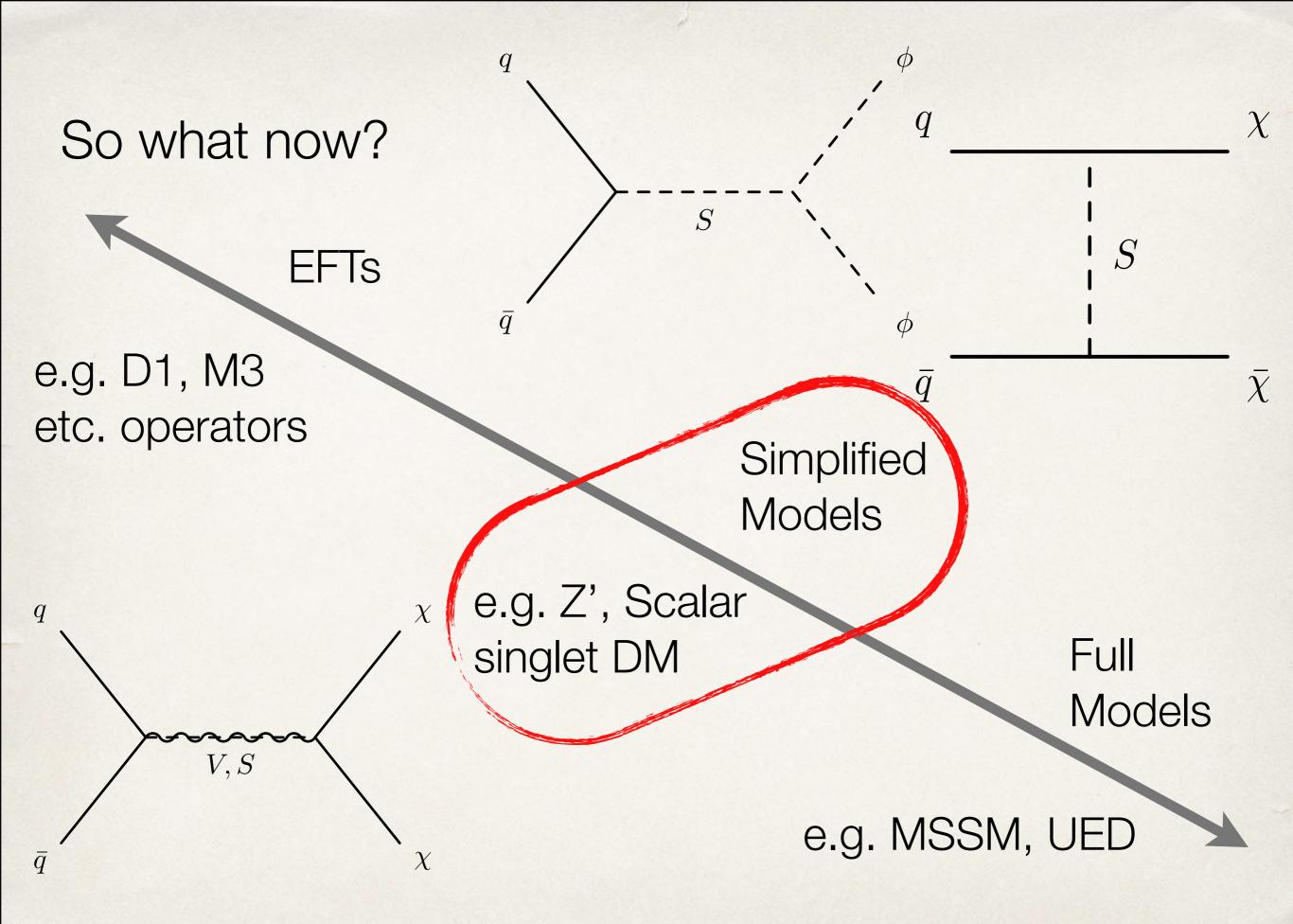
• For a given $\sqrt{g_{\rm SM}g_{\rm DM}}$, cut all events that don't pass $M\equiv\sqrt{g_{\rm SM}g_{\rm DM}}\Lambda\geq Q_{\rm tr}$



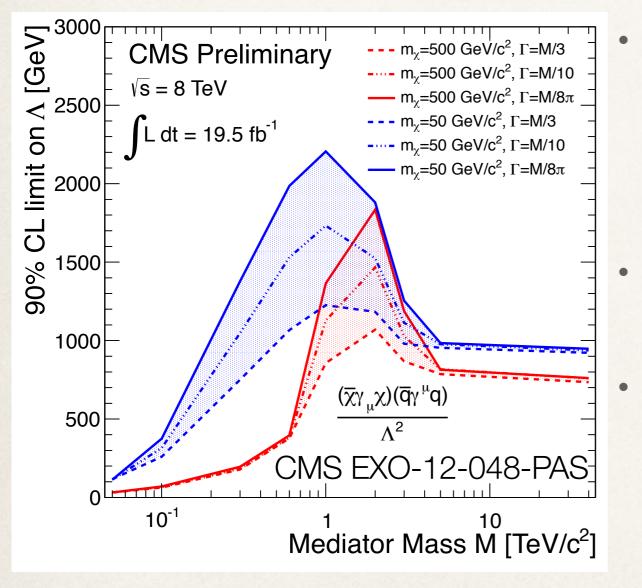
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Resonances & Widths



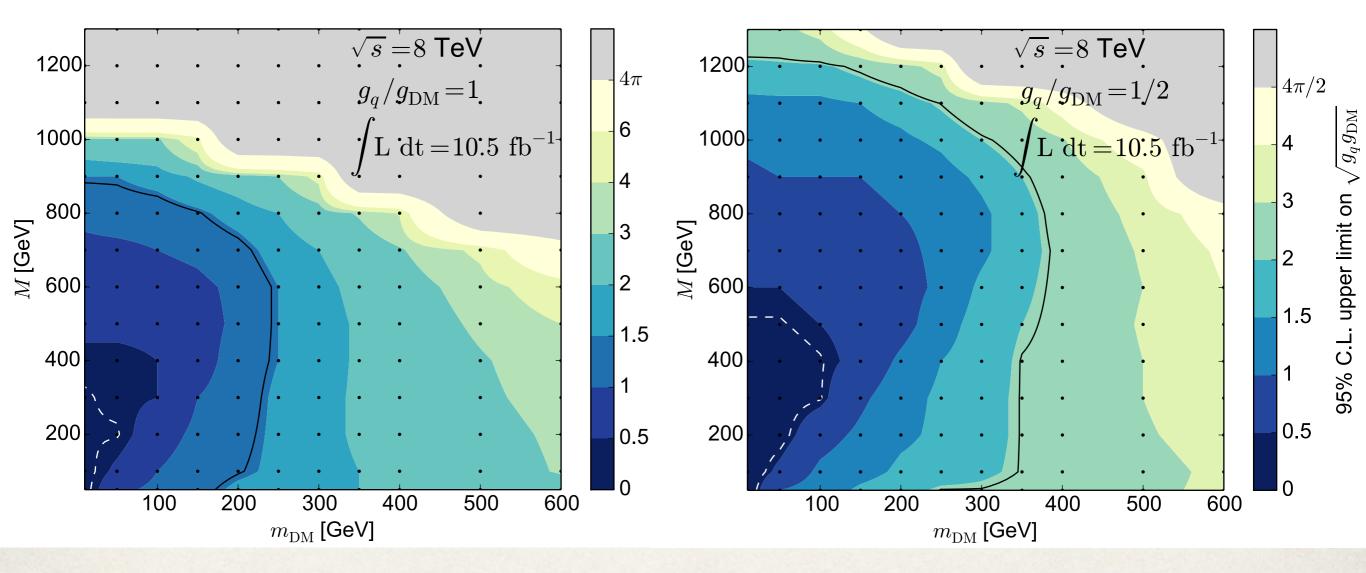
Resonance strengthens constraints relative to EFT, but width adds more parameters

Opens mediator searches

Min width fixed by the model: Beware arbitrary widths

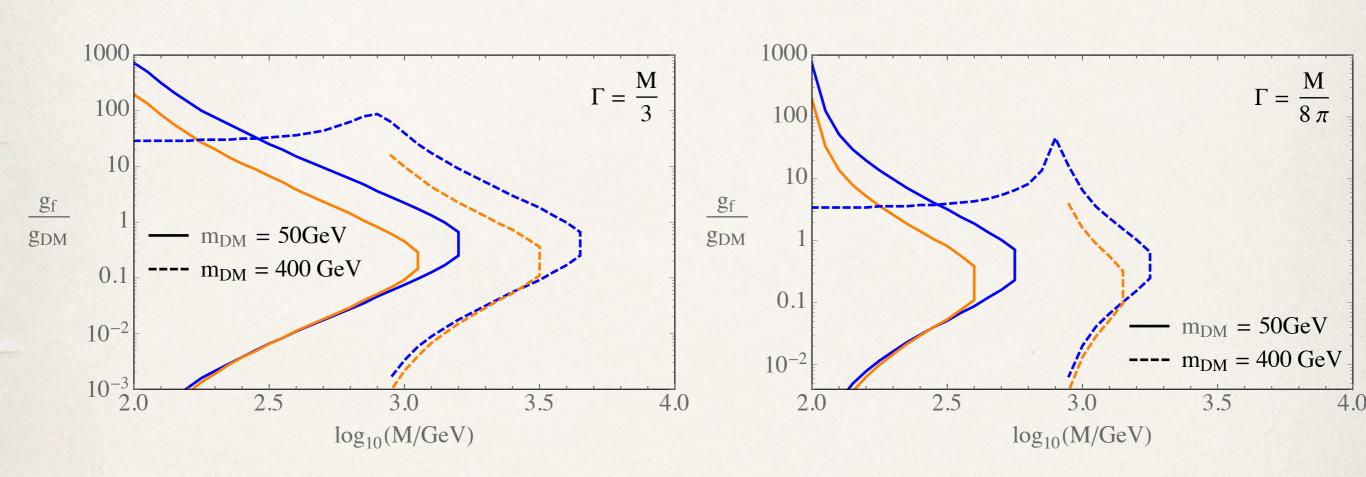
Constraining Simplified Models

$m_{\rm DM}, \Lambda \to m_{\rm DM}, M_{\rm med}, g_{\rm DM}, g_{q_i}$



TDJ, Karl Nordstrom, arXiv:1502.05721

Treatment of the width



d.

To summarise

- Effective Field Theories are a powerful tool allowing comparisons between different classes of experiment
- Facilitates complementary and powerful comparisons between Direct Detection and LHC constraints
- At LHC energies, the approximation begins to break down, and remain fully valid only for large couplings
- Truncation can make EFT constraints robust, but weaker; Simplified Models will give stronger constraints at LHC energies