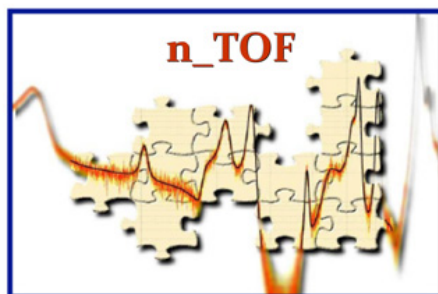
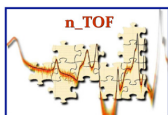


# Nuclear Astrophysics @ n\_TOF, CERN



**Tagliente Giuseppe**

*Istituto Nazionale Fisica Nucleare, Sez. di Bari*



# The n\_TOF Collaboration

*(~100 Researchers from 30 Institutes)*

## **CERN**

**Technische Universitat Wien**

Austria

**IRMM EC-Joint Research Center, Geel**

Belgium

**Charles Univ. (Prague)**

Czech Republic

**IN2P3-Orsay, CEA-Saclay**

France

**KIT – Karlsruhe, Goethe University, Frankfurt**

Germany

**Univ. of Athens, Ioannina, Demokritos**

Greece

**INFN Bari, Bologna, LNL, LNS, Trieste, ENEA – Bologna**

Italy

**Univ. of Tokio**

Japan

**Univ. of Lodz**

Poland

**ITN Lisbon**

Portugal

**IFIN – Bucarest**

Rumania

**CIEMAT, Univ. of Valencia, Santiago de Compostela,  
University of Cataluna, Sevilla**

Spain

**University of Basel, PSI**

Switzerland

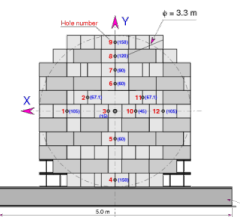
**Univ. of Manchester, Univ. of York**

UK

# n\_TOF Scientific Motivations

- Neutron cross sections relevant for Nuclear Astrophysics
- Measurements of neutron cross sections relevant for Nuclear Waste Transmutation and related Nuclear Technologies (ADS)
- Neutrons as probes for fundamental Nuclear Physics

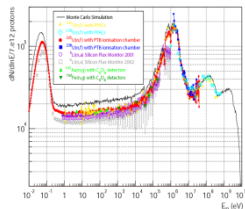
# n\_TOF Time line



**1995-1997**  
TARC  
experiment

**May 1998**  
Feasibility  
CERN/LHC/  
98-02+Add

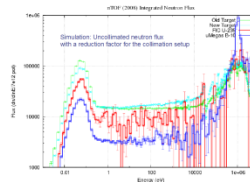
**2000**  
Commissioning



**2008**  
New Target  
construction

**May 2009**

Commissioning



**2009-2012**

Phase II  
Isotopes  
Capture: 14  
Fission: 6  
(n,α): 2

**July 2014**  
Commissioning  
EAR2

**1996**

**2009**

**2014**

**1997**

Concept  
by C.Rubbia  
CERN/ET/Int.  
Note 97-19

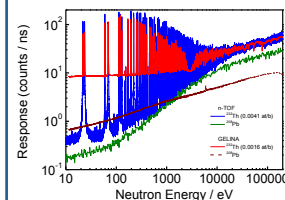
**Aug 1998**  
Proposal  
submitted

**1999**  
Construction  
started



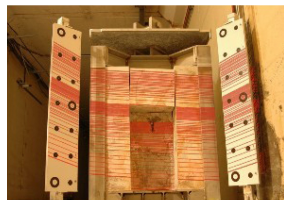
**2001-2004**

Phase I  
Isotopes  
Capture: 25  
Fission: 11



**2004-2007**

Problem  
Investigation



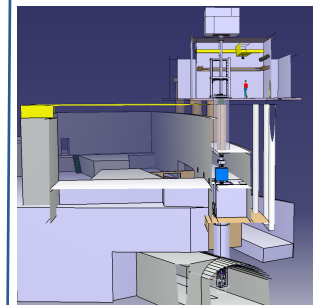
**2009**  
Phase II

**2010**

Upgrades:  
Borated-H<sub>2</sub>O  
Class-A

**2013-2014**

Construction  
second beam  
line  
EAR2



# n\_TOF Facility



## n\_TOF features

broad neutron energy range

n\_TOF  
200m

high instantaneous flux

## Use

proton beam, neutron capture  
s-process studies

intensity (dedicated mode)  
small capture

sample size (enriched samples)  
pulse width

radioactive samples (low intrinsic)

Proton Beam  
20 GeV/c  
7x10<sup>12</sup> ppp

target dimensions 80x80x60 cm<sup>3</sup>

1.4 GeV  
range dominated cross sections  
slowing & moderation H<sub>2</sub>O

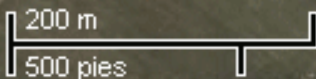
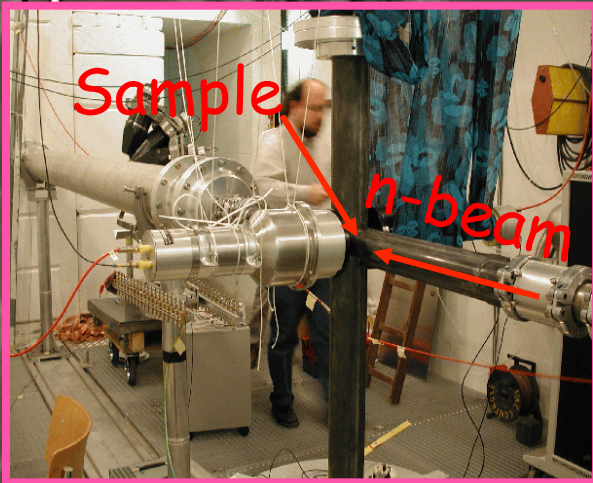
neutron material cross section measurements

factor thickness in 5 cm

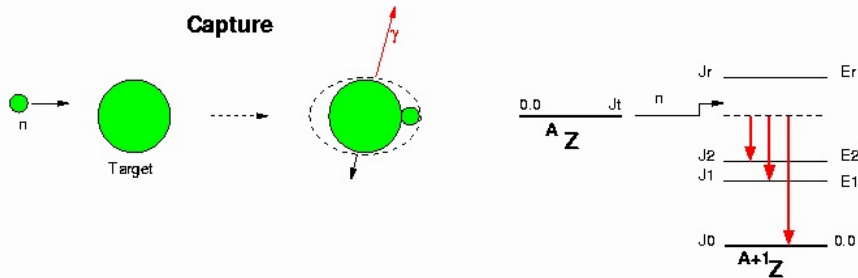
the exit face

50 MeV

PS 20 GeV

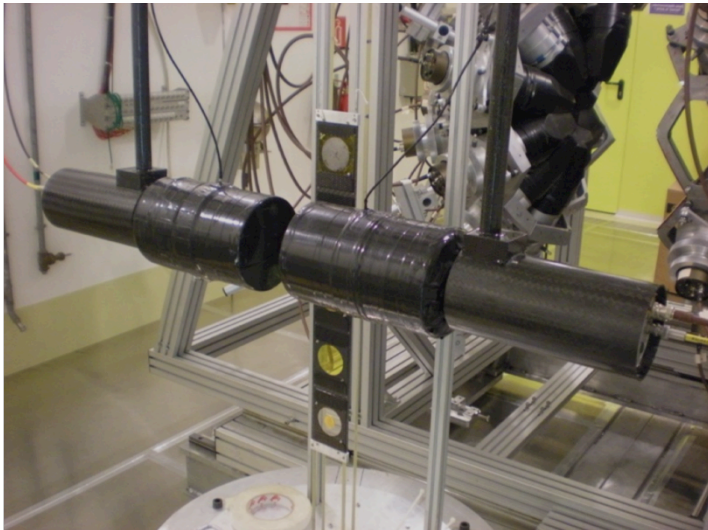


# Detectors for capture reactions



Capture reactions are measured by detecting  $\gamma$ -rays emitted in the de-excitation process.

At n\_TOF, two detection systems are used, for different purposes.

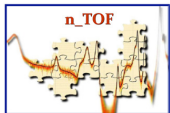
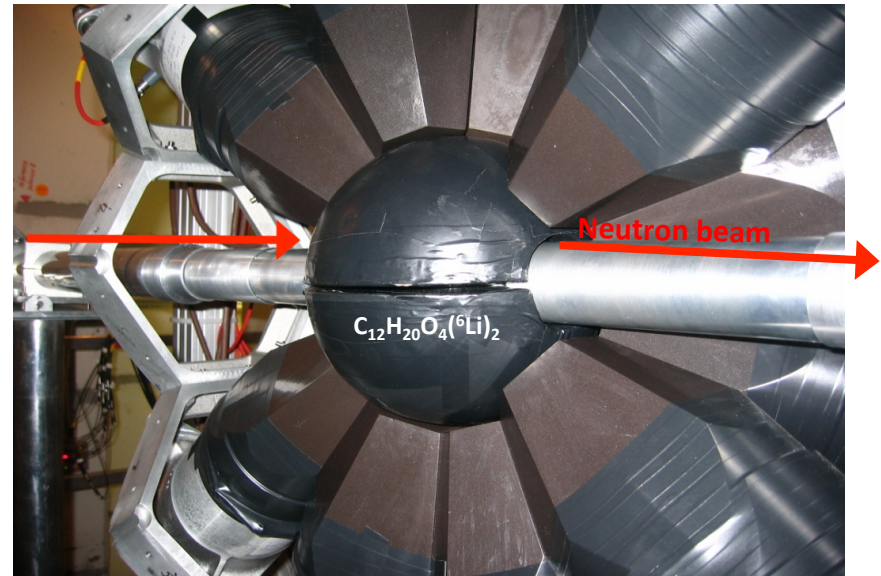


## $C_6D_6$ (deuterated liquid scintillators)

- low neutron sensitivity device
- used for low cross-section samples

## Total Absorption Calorimeter (TAC)

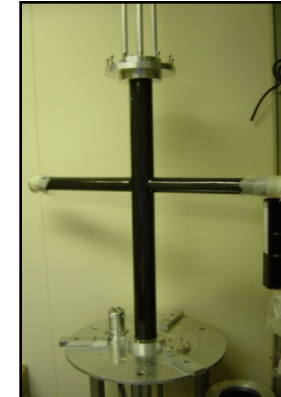
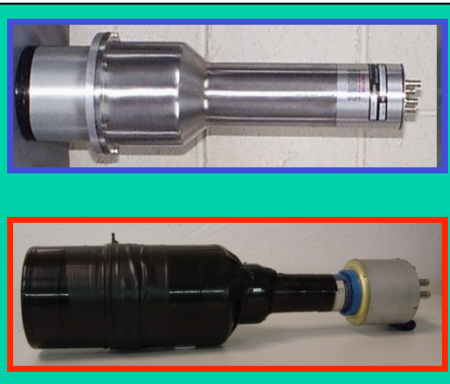
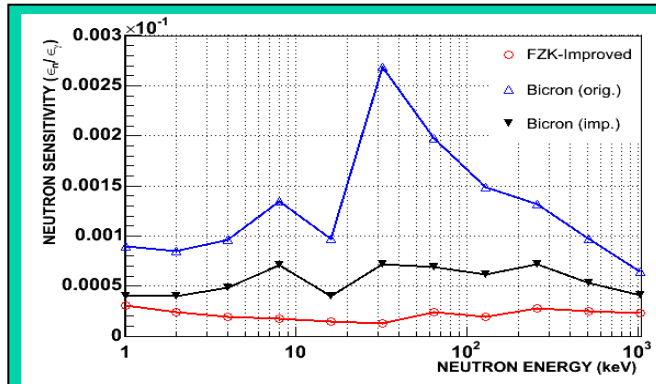
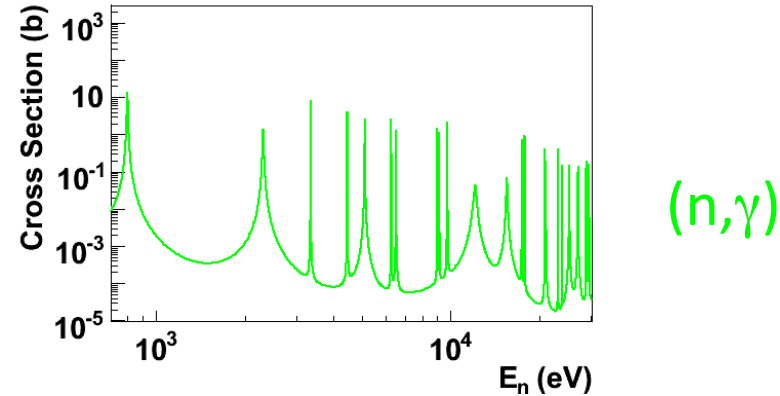
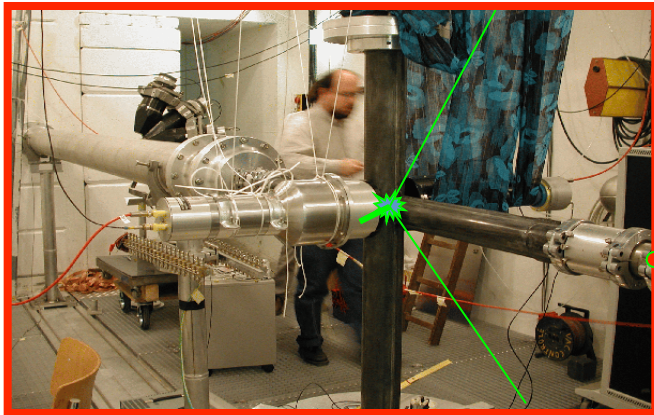
- High-efficiency  $4\pi$  detector (40  $BaF_2$  scintillators with neutron shielding)
- mostly used for fissile isotopes (actinides)



# (n,γ) Total energy detection @ n\_TOF

## Improvements in the Experimental Setup & Data Analysis

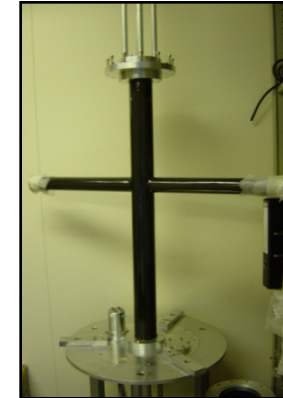
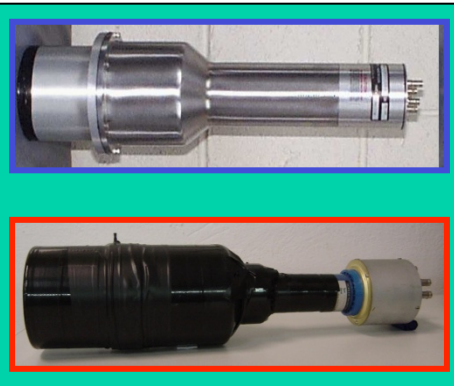
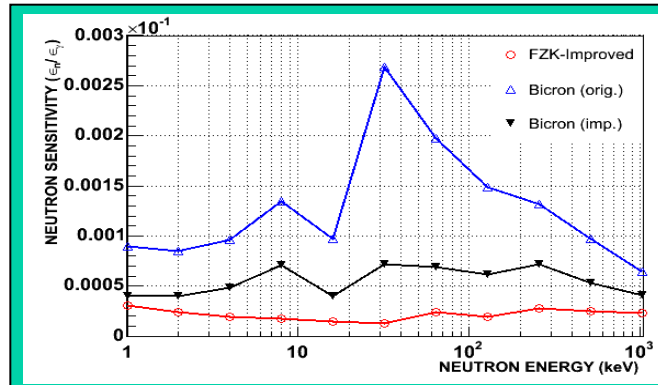
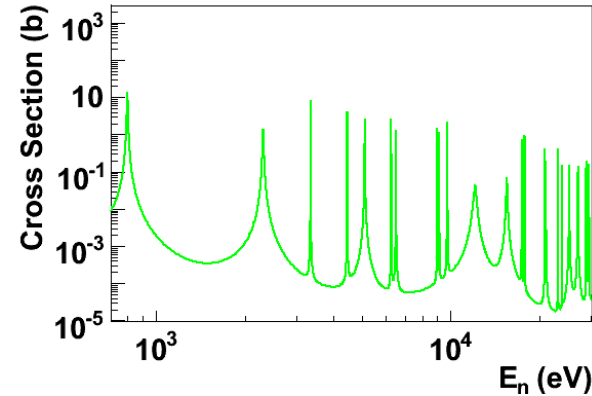
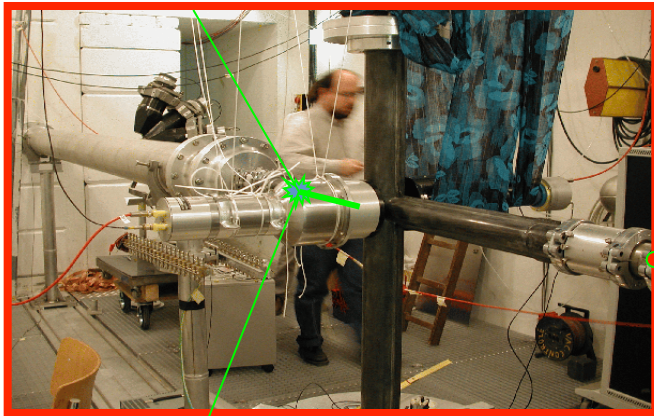
- Lowest neutron sensitivity → No neutron background corrections !



# (n, $\gamma$ ) Total energy detection @ n\_TOF

## Improvements in the Experimental Setup & Data Analysis

- Lowest neutron sensitivity  $\Rightarrow$  No neutron background corrections !



- n\_TOF: first facility with a neutron sensitivity optimized below measurable levels.
- All the (n, $\gamma$ ) measurements with  $C_6D_6$  (since start in 2002) were made with this improved setup.

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# The experimental activity at n\_TOF: Ph I

- **C**ross sections relevant in Nuclear Astrophysics
  - s-process: branchings
  - abundancies in presolar grains
  - Magic nuclei
  - Isotopes of particular interest

<sup>151</sup>Sm

204,206,207,208Pb, <sup>209</sup>Bi

24,25,26Mg

90,91,92,94,96Zr, <sup>93</sup>Zr

<sup>139</sup>La

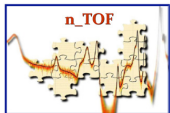
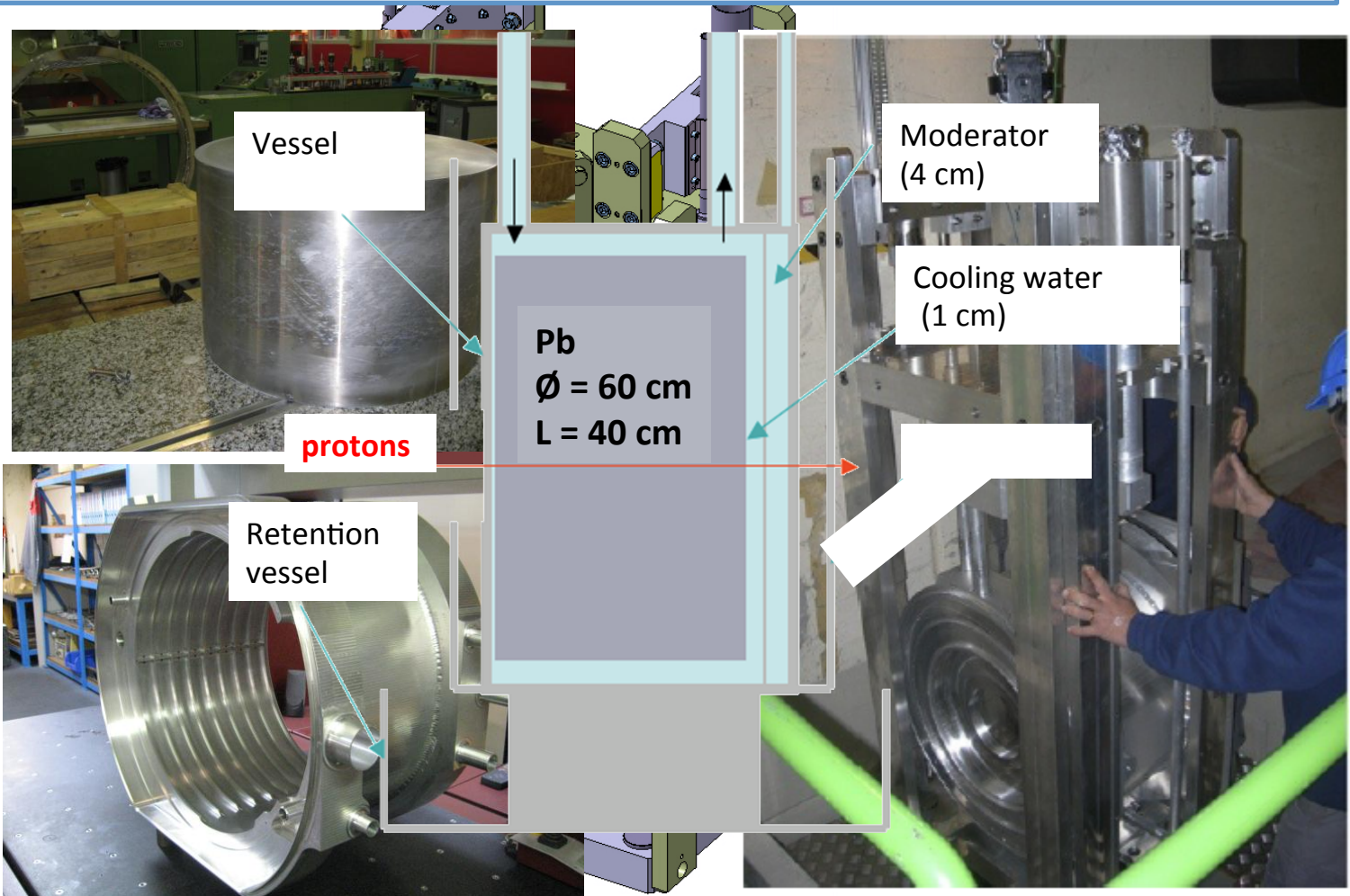
186,187,188Os

- In the period 2002-2004 measured long-needed **capture and fission** cross-sections for **36 isotopes**, 18 of which radioactive.
- The unprecedented combination of **excellent resolution, unique brightness and low background** has allowed to collect **high-accuracy data**, in some cases for the **first time ever**.

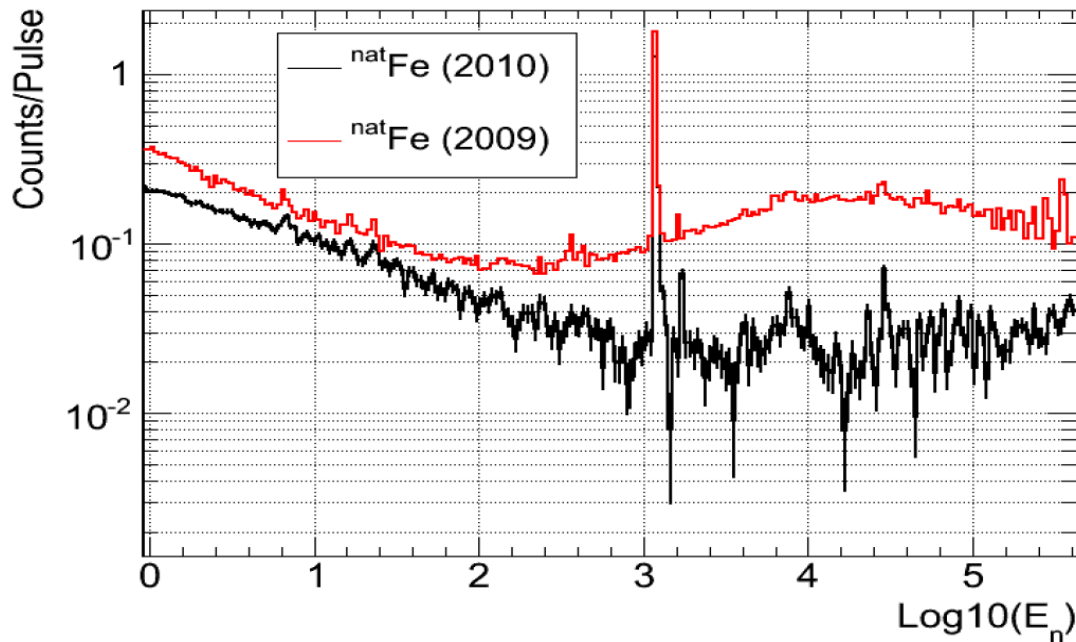
# **n\_TOF Phase II**

# The new spallation Target

The cooling and the moderator systems in the target are separated, so to optimize neutron spectrum or minimize background



# The new spallation Target



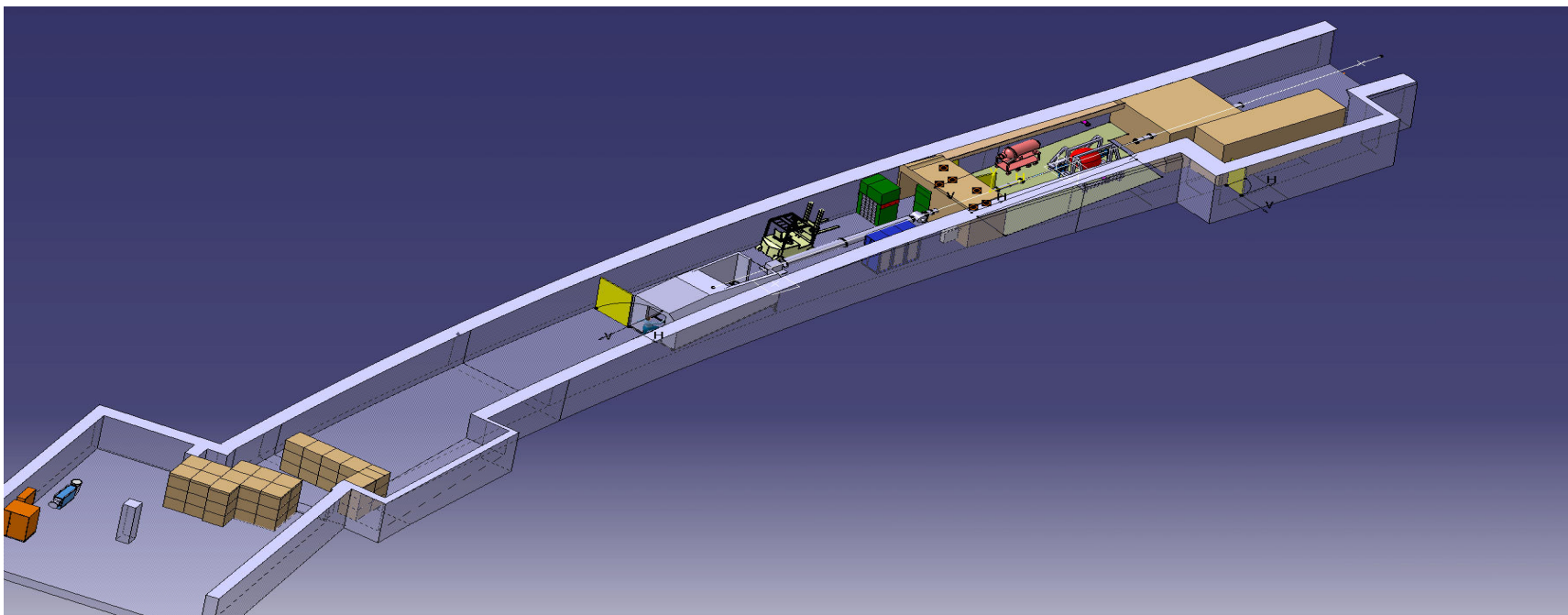
**Moderator**

2009  $\text{H}_2\text{O}$

2010  $\text{H}_2\text{O} + \text{H}_3\text{BO}_3$  (borated water)

**The borated water as moderator reduces the background of a factor 10!!  
In the energy region 1-100 keV !**

# Work Sector of Type A



Since 2010 the n\_TOF experimental area was transformed in work sector type A. It allows to measure sample with very high activity.

# The experimental activity @ n\_TOF: Ph II

- Cross sections relevant in Nuclear Astrophysics
  - s-process: seeds isotopes

$^{54,56,57}\text{Fe}$

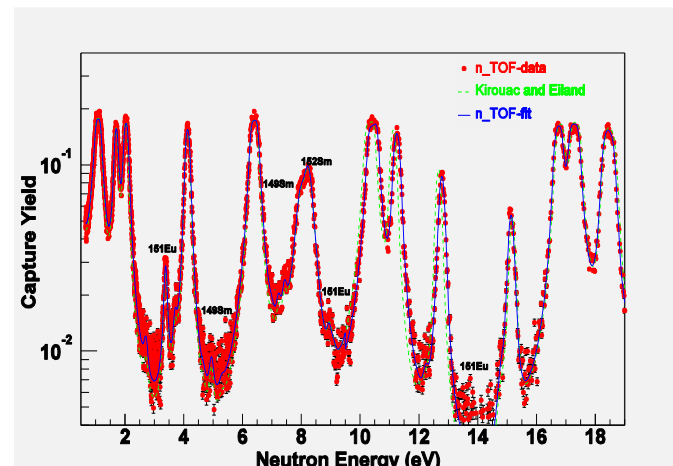
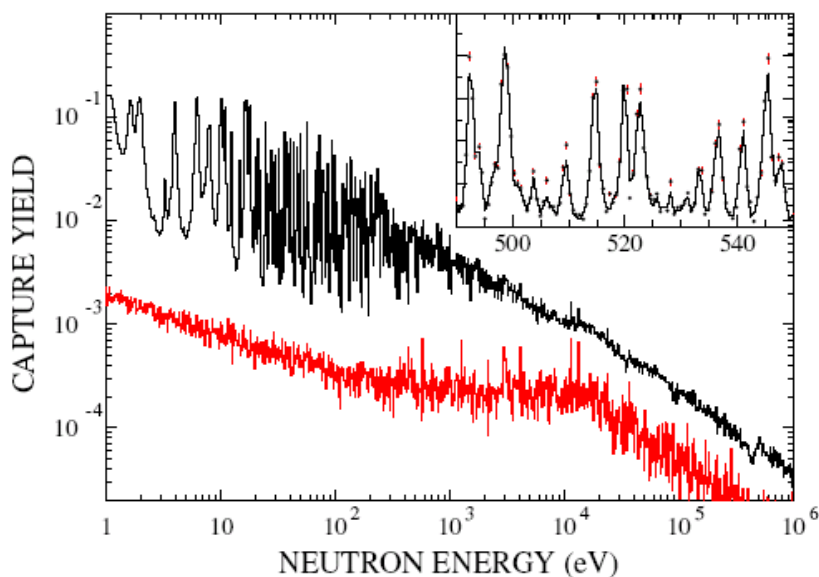
$^{58,60,62}\text{Ni}$ ,  $^{63}\text{Ni}$

$^{25}\text{Mg}$

$^{93}\text{Zr}$

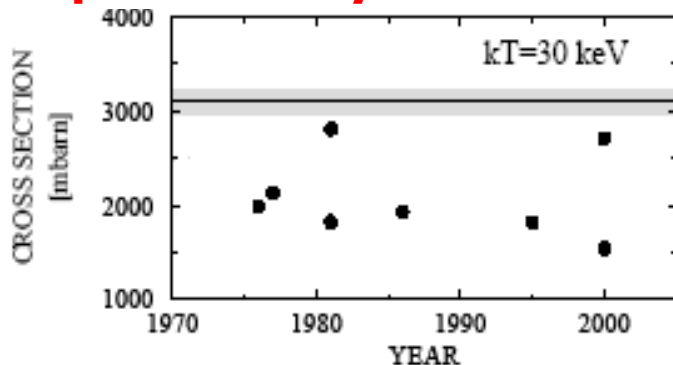
In the period 2009-2012 measured long-needed **capture and fission** cross-sections for **22 isotopes**, 14 of which radioactive.

# The experimental results: $^{151}\text{Sm}$



Measured for the first time at a time-of-flight facility  
Resonance analysis with SAMMY code.

**Maxwellian averaged cross-section  
experimentally determined for the first time**

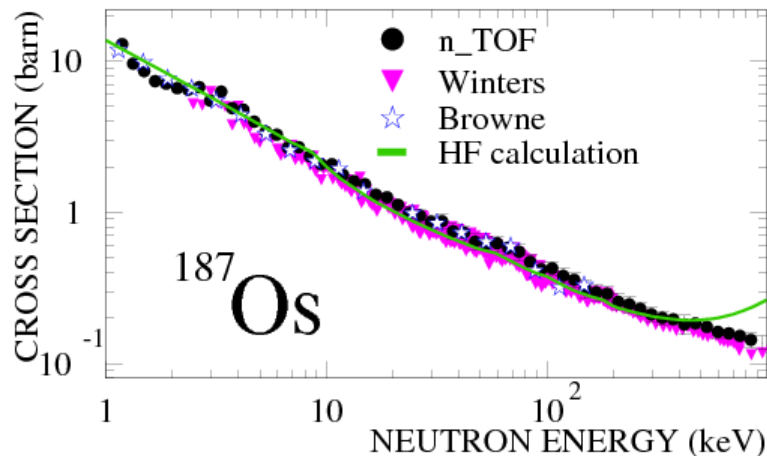
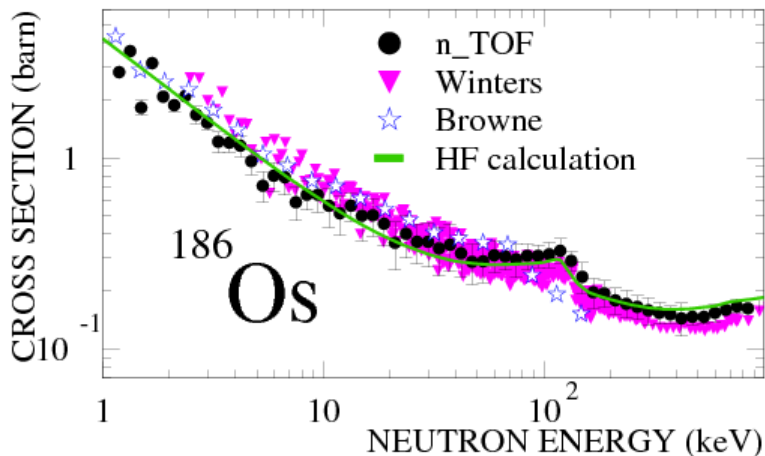


s-process in AGB stars  
produces 77% of  $^{152}\text{Gd}$ ,  
23% from p process

**Maxwellian averaged (n, $\gamma$ ) cross section of the  
 $^{151}\text{Sm}$  and previous calculation (symbol)**

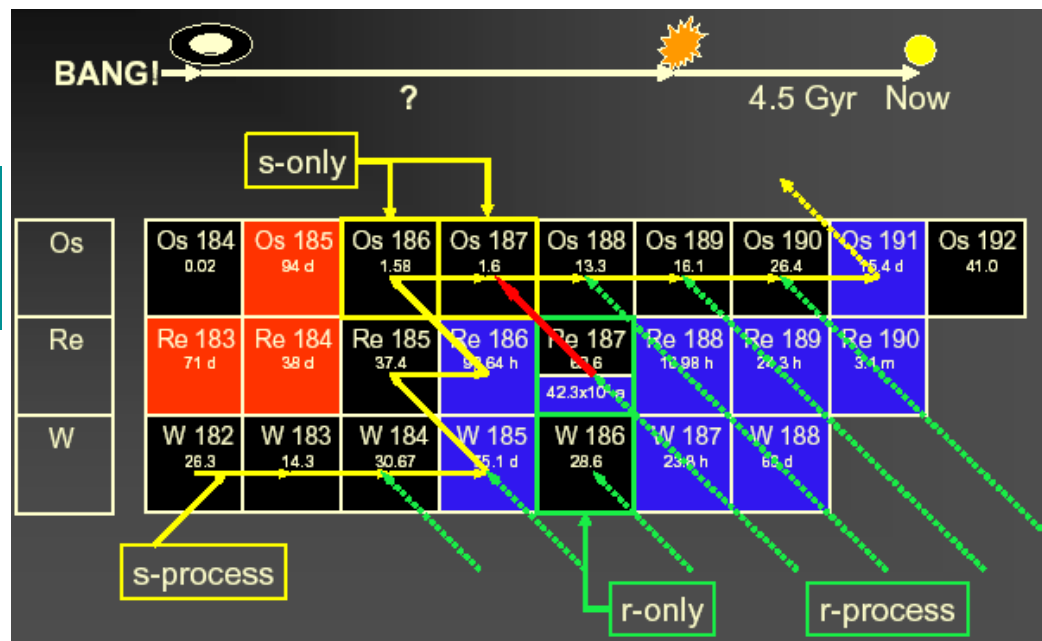
**NO PREVIOUS MEASUREMENTS!**

# The experimental results: $^{186,187}\text{Os}$



$$^{187}\text{Os}_c = ^{187}\text{Os} - \frac{\sigma(186)}{\sigma(187)} ^{186}\text{Os}$$

$$\frac{\sigma(186)}{\sigma(187)} = 0.42 \pm 0.02$$

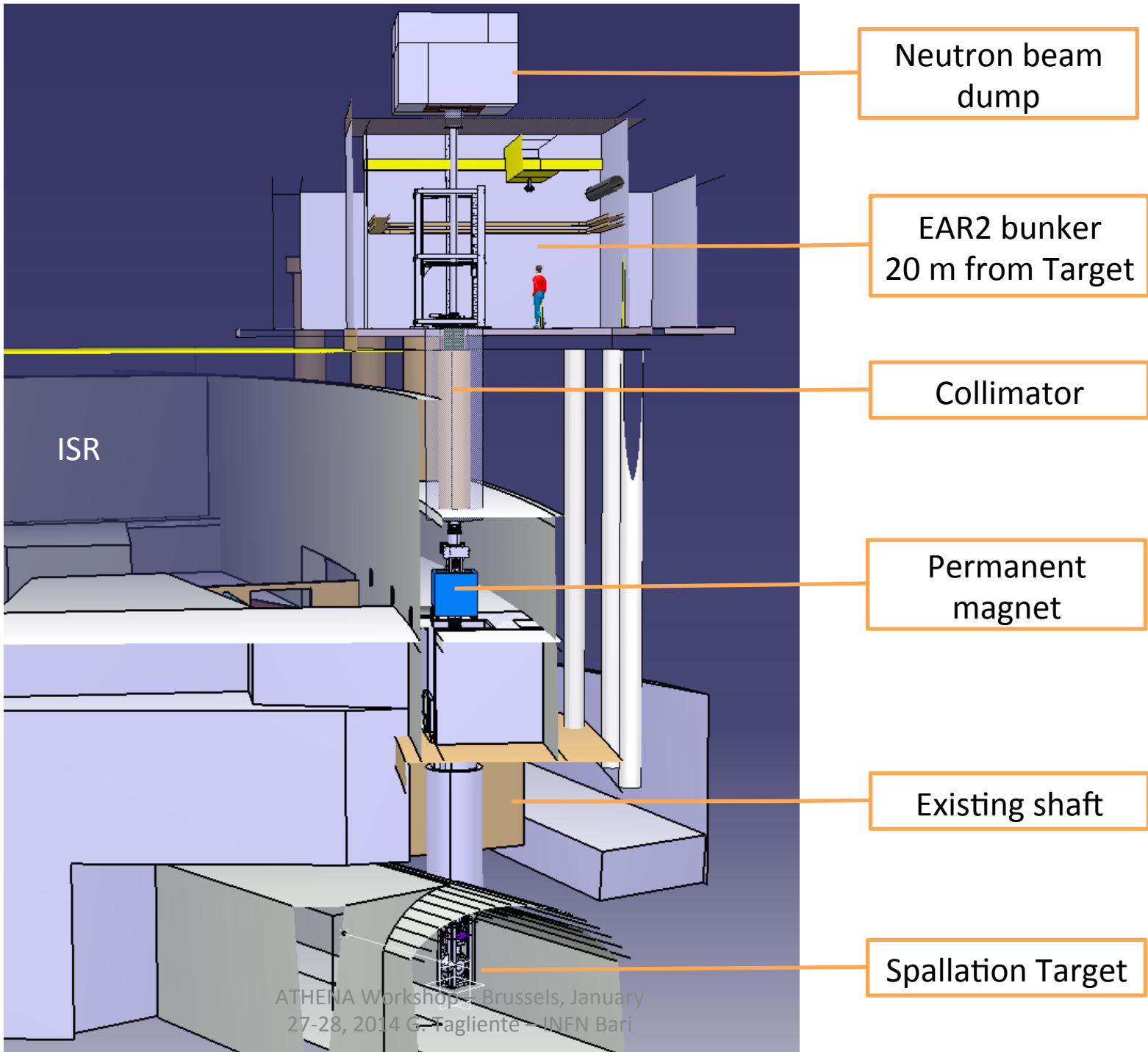




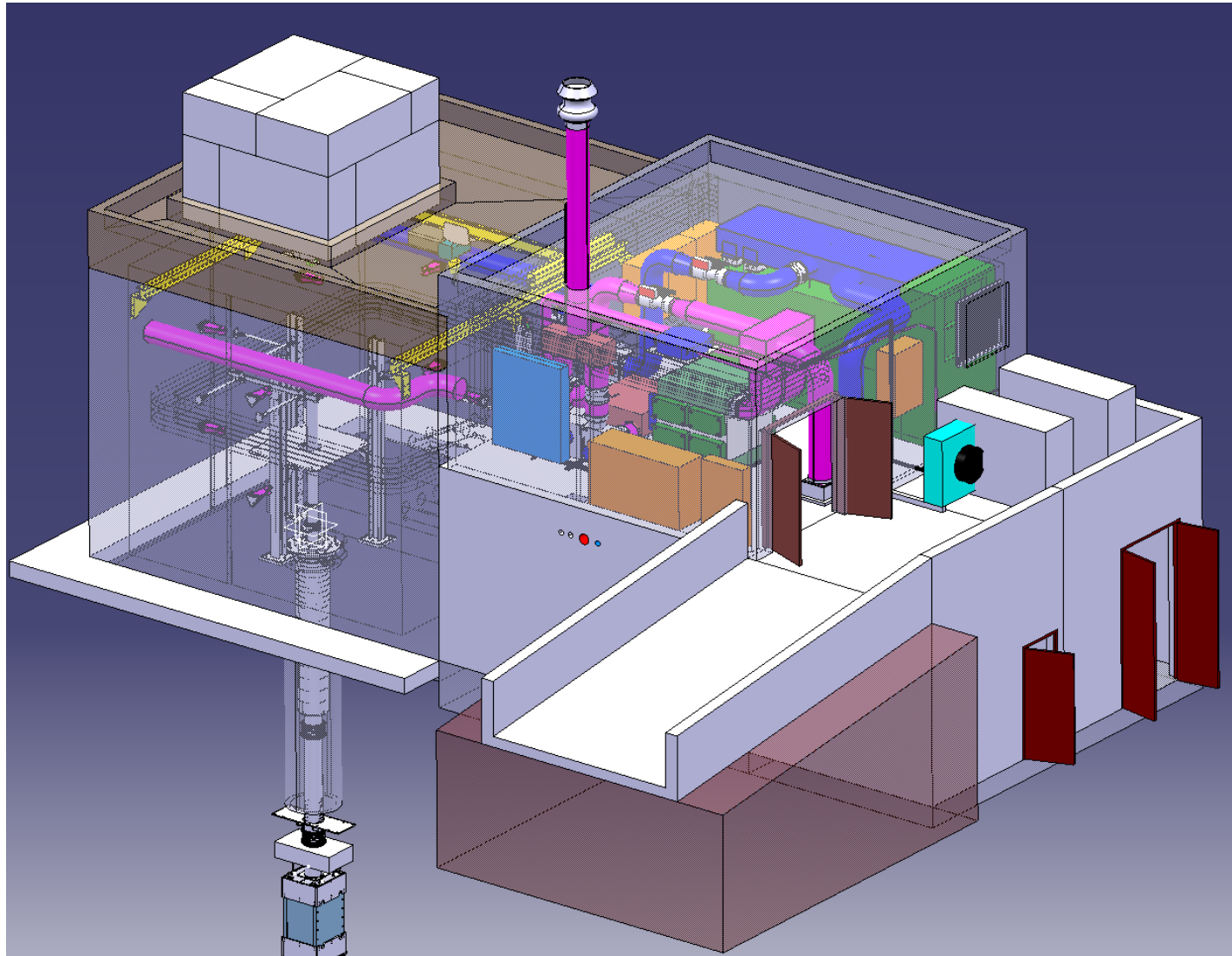
# Publications

| Isotope                   | Reference                                     |
|---------------------------|---|
| $^{24,25,26}\text{Mg}$    | PRC 85 (2012) 044615                          |
| $^{58}\text{Ni}$          | PRC 89 (2014) 014605                          |
| $^{62}\text{Ni}$          | <i>Submitted to PRC</i>                       |
| $^{63}\text{Ni}$          | <i>PRL 110 (2013) 022501</i>                  |
| $^{90}\text{Zr}$          | PRC 77 (2008) 035802                          |
| $^{91}\text{Zr}$          | PRC 78 (2008) 045804                          |
| $^{92}\text{Zr}$          | PRC 81 (2010) 055801 <b>APJ 780 (2014) 95</b> |
| $^{93}\text{Zr}$          | <i>PRC 87 (2013) 014622</i>                   |
| $^{94}\text{Zr}$          | PRC 84 (2011) 015801                          |
| $^{96}\text{Zr}$          | PRC 84 (2011) 055802                          |
| $^{139}\text{La}$         | PRC 75 (2007) 035807                          |
| $^{151}\text{Sm}$         | PRL 93 (2004) 161103 – PRC 73 (2006) 034604   |
| $^{186,187,188}\text{Os}$ | PRC 82 (2010) 015802 – PRC 82 (2010) 015804   |
| $^{204}\text{Pb}$         | PRC 75 (2007) 015806                          |
| $^{206}\text{Pb}$         | PRC 76 (2007) 045805                          |
| $^{207}\text{Pb}$         | PRC 74 (2006) 055802                          |
| $^{209}\text{Bi}$         | PRC 74 (2006) 025807                          |

# The second **E**xperimental **A**Rea @ n\_TOF

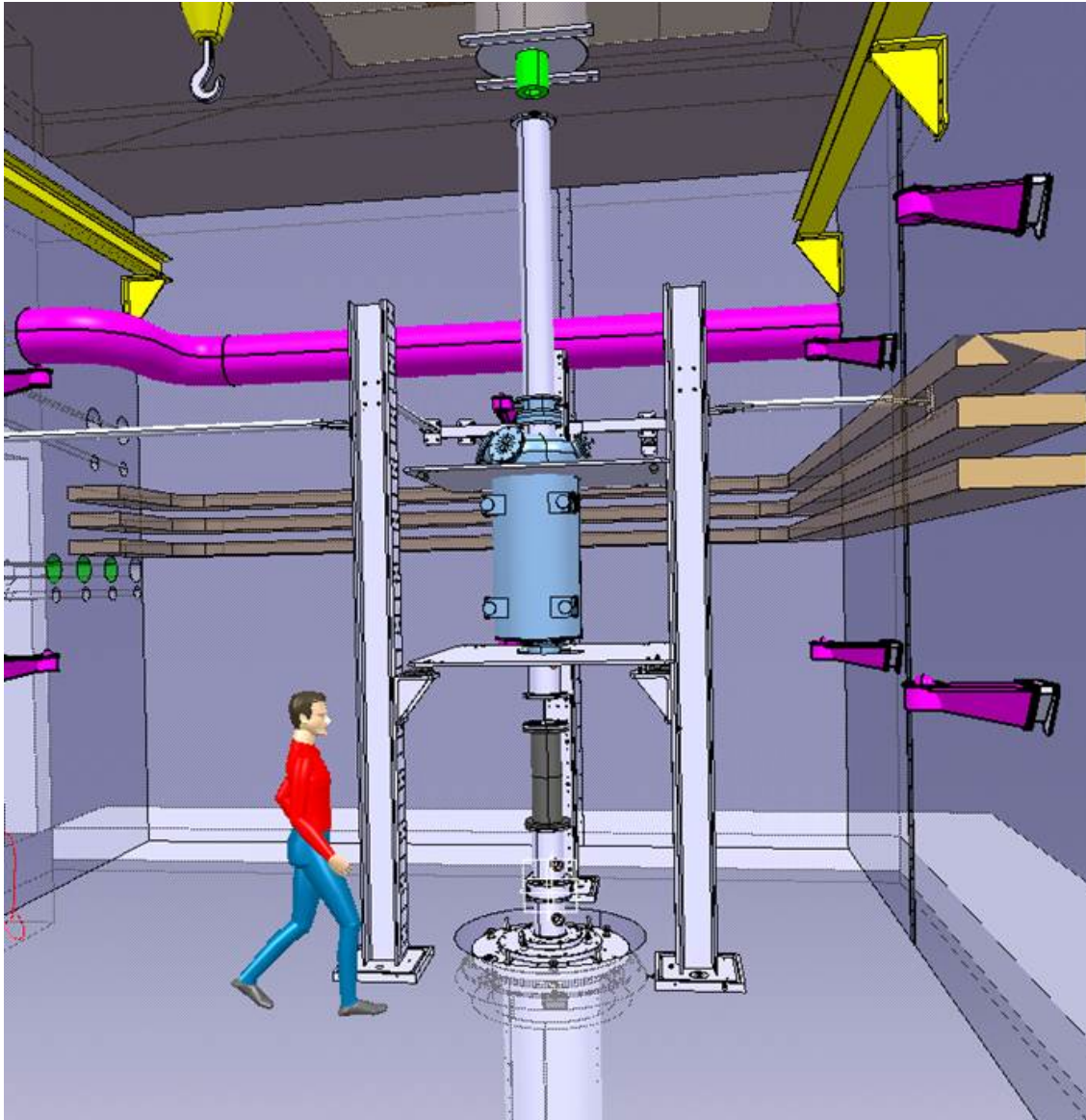


# EAR 2 general view



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# Sketch of possible installation of Detectors



Sketch of possible  
installation of  
Detectors

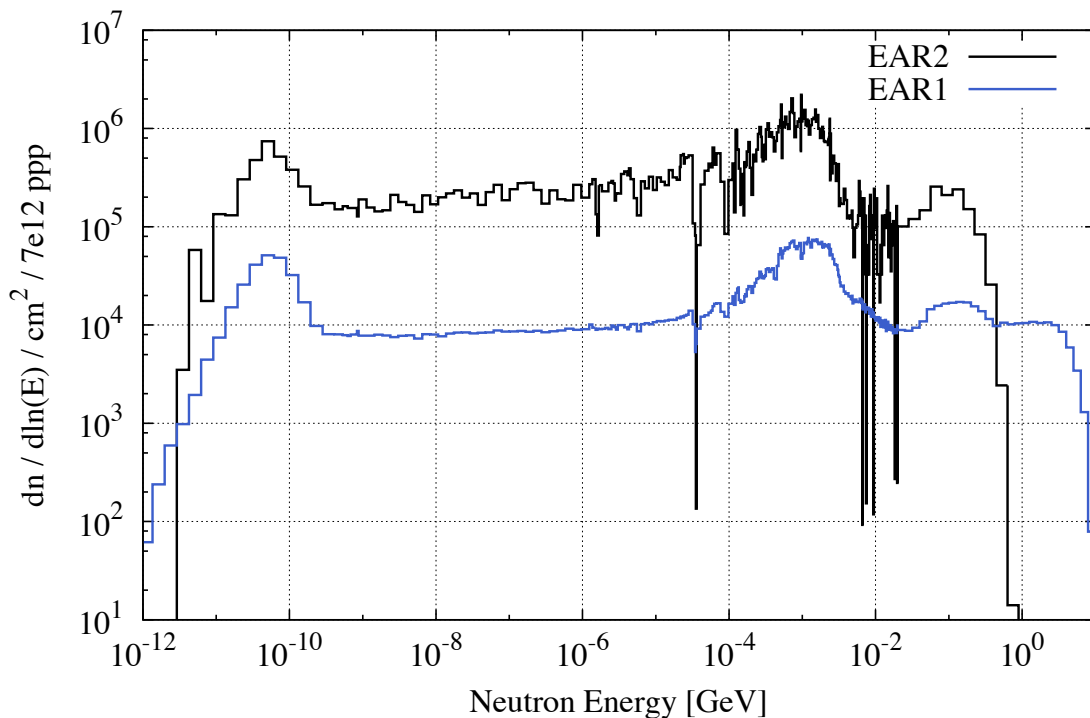
# Status of EAR 2



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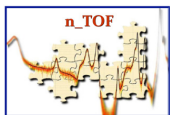
# Main features of EAR 2

Comparison of the Neutron Fluence in EAR1 and EAR2



**Higher fluence**, by a factor of 25, relative to EAR1.  
The **shorter flight path** implies a factor of 10 smaller time-of-flight.  
Global gain by a factor of **250 in the signal/background ratio** for radioactive isotopes!

The huge gain in signal-to-background ratio in EAR2 allows to measure radioactive isotopes with **half lives as low as a few years.**



# The experimental program EAR 2

## The EAR2 will allow to:

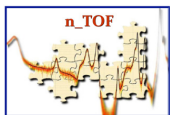
- measure samples of **very small mass (<1 mg)**
- measure **short-lived radioisotopes** (down to a few years)
- collect data on a much **shorter time scale**
- **measure (n,charged particle) reactions with thin samples**

## Letter of intent for measurements in EAR2:

- **(n,p)** and **(n, $\alpha$ )** cross sections on  $^7\text{Be}$ ,  $^{25}\text{Mg}$ ,  $^{26}\text{Al}$
- **Fission** cross sections of the **short lived actinides**  $^{232}\text{U}$ ,  $^{238,241}\text{Pu}$  and  $^{244}\text{Cm}$
- **Capture** cross section of  $^{79}\text{Se}$ ,  $^{245}\text{Cm}$
- Cross section and angular distribution of fragments from  $^{232}\text{U}(n,f)$

## Status of the EAR2:

- **Approved by CERN**, final design phase
- **Start construction** in May 2013
- **Beam ready** in mid-2014
- **Physics start** in 2015





# AstroPhysics program EAR I & EAR II

| Isot.                              | R              | Comments         |
|------------------------------------|----------------|------------------|
| $^{70,72,73}\text{Ge}$             | (n, $\gamma$ ) | s-process flow   |
| $^{171}\text{Tm}, ^{204}\text{Tl}$ | (n, $\gamma$ ) | Branching points |

| Isot.                                    | R                | Comments                             |
|--|------------------|--------------------------------------|
| $^{147}\text{Pm}$                        | (n, $\gamma$ )   | Branching point                      |
| $^{26}\text{Al}$                         | (n,p/ $\alpha$ ) | $^{26}\text{Al}$ galactic abundance  |
| $^{53}\text{Mn}$                         | (n, $\gamma$ )   | Explosive stage of stellar evolution |
| B,C, $^{14}\text{N}$ ,O, $^{19}\text{F}$ | (n, $\alpha$ )   | n capture in light nuclei            |
| $^{79}\text{Se}$                         | (n, $\gamma$ )   | Branching point                      |

# Conclusions



- There is need of **accurate new data** on neutron cross-section both for **astrophysics and advanced nuclear technology**.
- Since 2001, **n\_TOF@CERN** has provided an important contribution to the field, with an intense activity on **capture and fission measurements**.
- Several results of interest for **stellar nucleosynthesis** (Sm, Os, Zr, Ni, Fe, etc...).
- Important data on actinides, of interest for **nuclear waste transmutation**.
- To date, high resolution measurements performed in **EAR1** in optimal conditions (borated water moderator, Class-A experimental area, etc...).
- A second **experimental area at 20 m** for high flux measurements is in construction.
- The EAR2 (starting in 2015) will open **new perspectives** for frontier measurements on short-lived radionuclides.

