Giovanna Benzoni (Milano) Strong, weak and electromagnetic forces at work in atomic nuclei, decay properties

Outline of the talk:

Introduction alpha decay beta decay exotic decay modes examples of measurements





• Fundamental interactions:



Introduction





~3000 known nuclei, both natural (228) and artificial 7000+/- 500 current estimates for bound nuclei

Boundaries of the bound nuclei are not defined yet



Introduction



4 interactions:

-strong: acting btw nucleons, keeps nucleus bound -weak: involving presence of leptons -electromagnetic: btw charged particles/γ rays

→ Gravitational : usually neglected owing to small masses involved

Decay modes

cluster emission



The Nuclear Chart: Radioactive Decay Modes



Introduction

Introduction

Useful definitions for decay:

Law of radioactive decay:

$$N(t) = N_0 e^{-\lambda t}$$





Branching ratio: decay probability of two or more competing processes

Mother/Precursor == Initial/Emitting nucleus Daughter/Successor = Final nucleus **Bateman equations**: mathematical model describing abundances and activities in a decay chain as a function of time, based on the decay rates and initial abundances.

If at a time t, there are $N_i(t)$ atoms of the i-th isotope which decays into the i+1 one with a decay rate λ_i , the amounts of isotopes in the k-th step of the decay chain evolves as:

$$\frac{dN_1(t)}{dt} = -\lambda_1 N_1(t)$$

$$\frac{dN_i(t)}{dt} = -\lambda_i N_i(t) + \lambda_{i-1} N_{i-1}(t)$$

$$\frac{dN_k(t)}{dt} = \lambda_{k-1} N_{k-1}(t)$$

This can be written in an explicit form as:

$$N_n(t) = \prod_{j=1}^{n-1} \lambda_j \sum_{i=1}^n \sum_{j=1}^n \left(\frac{N_i(0)e^{-\lambda_j t}}{\prod_{p=1, p\neq j}^n (\lambda_p - \lambda_j)} \right)$$

This can be extended to cases in which we have many branches

H. Bateman. "Solution of a System of Differential Equations Occurring in the Theory of Radio-active Transformations," Proc. Cambridge Phil. Soc. IS, 423 (1910) <u>https://archive.org/details/cbarchive_122715_solutionofasystemofdifferentia1843</u>

Secular equilibrium:

The rate of production and decay of an element is constant It can be achieved in a decay chain if the daughter B has a much shorter half-life than the mother A

The decay rate of A, which is translated into the production rate of B is constant

The quantity of B accumulates since the number of nuclei which decay in 1s if equal to that produced in 1 s.



This depends on N_A(t=0), λ_A and λ_B

- Each member of the family is a dewar. They are connected and the fill each other in chain
- The rate of emptying (-dN/dt) are function of the level in the dewar N and the decay rate λ
- At equilibrium the rate of evacuation in each dewar is equal
- In nuclear physics rate of decay is the Activity and depend on the decay constant λ
- At equilibrium all activities are equal

Detecting $\alpha \; \beta \; \gamma$ particles: reminder of basic properties

 $\boldsymbol{\alpha}$ particle: massive, charged

- → interested in energy, angular distribution
- → spectrometers, Si, plastic detectors

 β particles: light, charged

- \rightarrow interested in energy and correlations
- \rightarrow e⁺ annihilates into two 511 keV γ rays
- → spectrometers, Si, Plastic detectors

γ particles: massless, neutral

- → interested in energy and angular distribution:
- → HPGe detectors, Organic scintillators





6 x 10

Example of detailed study of a decay:

1) Produce and Identify PARENT nucleus: reactions/ISOL/fragmentation

2) Detect the decay: identify the emitted particle α/β identify competing mechanisms identify decaying state, g.s. or excited measure half-life

3) Study properties of DAUGHTER nucleus: which levels are populated Branching Ratio btw levels subsequent decays



Producing radioactive beams

relativistic fragmentation/fission

of heavy nuclei on thin targets

- > 50 MeV/u → production of cocktail beams of many nuclei
- Use of spectrometers to transport/separate nuclei of interest

→ Relatively long decay paths ∆t > 150-300 ns

- Nuclei are brought to rest in final focal plane and let decay
- + cocktail beam: many nuclei at once
- + both short and long-living species
 + get information already with few particles
- Low cross sections
- Limitation on rate to distinguish contribution from each species

ISOL method,

- spallation/fission/fragmentation on
- thick targets, followed by
- chemical/physical processes to extract desired nuclei
 - beams produced at very low energies (60 keV)

 - + high cross section
 - + no need to re-accelerate beams
 - + high rates accepted
 - short-living species might not be accessed easily
 - Refractory elements
- Presence of long-living impurities (isobaric contamination)

Alpha Decay of a Uranium-238 nucleus



 $^{A}_{Z}X \rightarrow ^{A-4}_{Z-2}Y + \alpha$

One of the first discoveries in modern physics: 1908 Rutherford shows that emission of radiation is made of nuclei of 4He. $Q=(Mx-My-M\alpha)c^2$

 $E_{\alpha}(MeV)$

No α emitters with A<146: A< 140 Eα<0 : no emission 140<A<210: Eα<3 MeV:decay possible but not probable A>210 spontaneous decay





⁴He:

2p + 2n

Artificial isotopes >Pb decay by α emission Exceptions: ⁸Be ⁹Be ¹⁶²Gd, ¹⁷⁴Hf N-poor with Z=60-80, very short half-lives (s-ms) Geiger-Nuttall formula:

In λ = A+B In E \rightarrow the higher the energy the lower the T_{1/2} In λ = a+ b In R, R=Eⁿ is the range of α particles in air

E ÷ 4-9 MeV

Egs: ²¹⁸Th Q = 9.85 MeV, $T_{1/2}$ = 10⁻⁷ s ²³²Th Q= 4.08 MeV $T_{1/2}$ = 1.4 10¹⁰ y

Phenomenologic rule, can be explained by quantum mechanics



Geiger-Nuttal formula explains why α decay is not seen for Q<4 MeV

Large Q→ large difference btw X and Y X more unstable X decays faster



Dependence of Q(A,Z) starting from semi-empirical mass formula: Q=28.3-4a_v +8/3a_s $A^{-1/3}$ +4a_c $ZA^{-1/3}(1-Z/3A)$ -4a_{symm}(1-2Z/A)²+3a_p $A^{-7/4}$

Theory of the emission of a particles: α particle pre-exists in the nucleus and has to overcome the Coulomb barrier





We can consider the Coulomb barrier as a series of infinitesimal rectangular barriers

The probability for penetration of each infinitesimal barrier is:

$$dP = \exp(-2dr \sqrt{\frac{2m}{\hbar^2}(V(r) - Q)})$$

Or $P=e^{-2G}$ where G is the GAMOW FACTOR

$$G = \sqrt{\frac{2m}{\hbar^2}} \int_a^b \sqrt{(V(r) - Q} dr$$

Low momentum approximation:

 $Q = \frac{1}{2}mv^2 \ll \frac{zZe^2}{4\pi\epsilon_0 R}$

$$G \cong \sqrt{m} \frac{zZe^2}{4\pi\epsilon_0 R}$$

That is

$$G \propto Z \sqrt{\frac{m}{Q}}$$

Then λ = fP, P=e^{-2G}

And
$$log\lambda = 57 - 1.7Z \sqrt{\frac{m}{Q}}$$

Particle energy lower barrier



Approximations:

- Probability of pre-formation
- Fermi golden rule
- Only decay from and to gs
- Spherical symmetry, no deformation \rightarrow usually not true
- Angular momentum



I=0, n and p coupled into pairs Total ang. momentum is only given by orbital component, $I\alpha$



For I > 0 one needs to account for centrifugal barrier

Conservation of angular momentum Ji=Jf+I α Conservation of parity, depending on $l\alpha$ Conservation of isospin I

Alpha decay is sensitive to specific shell-model orbitals: *Strong Q value dependence favours population of low-lying states * Formation probability is sensitive to overlap of initial and final wave functions

- Observables for α decay are: E α , I α , T_{1/2} ,BR α branching ratio
- A decay can occur either from gs or excited states and decay can be fragmented onto several excited states: FINE STRUCTURE of α decay
- Decay from each nucleus has specific α lines



 α spectrum measured in a Si detector





a tagging to measure new half-lives PHYSICAL REVIEW C 89, 014324 (2014

<u>ם</u> et A.I.Morales

B-decay studies of neutron-rich Tl, Pb, and Bi isotopes

Counts / keV

Counts / keV

200

100

0

50

X-70/15/ rays)

c)



1000 1200 Energy (keV)

1200

1400

1400

²¹⁷Po

339

264

200

485 - 511 (e`e')

(a, a)

400

580

600

841 88

1000

500 800 Energy (keV)





General Properties of α decay

Observables for a decay are: Ea, Ia, $T_{1/2}$,BRa branching ratio

Formation probability, expressed as reduced α width:

$$\delta^2 = \frac{hln(2)b_{\alpha}}{T_{1/2}100P}$$
 P barrier penetration probability

If the decay occurs to many levels in the daughter we can define the HF(hindrance factor) to the excited state α_2 relative to the gs α_1 as:

 $HF = \frac{I_{\alpha 1} P_{\alpha 2}}{I_{\alpha 2} P_{\alpha 1}}$

And HF is sensitive to changes in spin and parity btw initial and final levels

In ΔI =0, 0+-> 0+ transitions HF÷0.5-85



- α decay can be a complementary tool to study low-lying O+ states in even-even nuclei
- → Applied to the region above ²⁰⁸Pb where shape coexistence is predicted

Potential Energy Surfaces in (β . γ) plane show presence of competing minima

letters to nature

A triplet of differently shaped spinzero states in the atomic nucleus ¹⁸⁶Pb

Nature 405 (2000) 430 al., et Andreyev Ż

¹⁸⁶Pb Three competing minima Three 0+ states

¹⁹⁰Po**→**¹⁸⁶Pb+α

¹⁹⁰Pb populated in fusionevaporation reaction → large bg owing to open reaction ch.

SHIP velocity filter separator

Detector array: PSSD Si det. for ion- α correlations 6 Si for conversion electron HPGe for α - γ correlations







(2000) 430 405 Nature al., et Andreyev Ŕ

Measurement of α particles with energies corresponding to different low-lying 0* states, $T_{1/2}$ and HF

$HF \rightarrow$ can be related to single-particle \rightarrow Nature of the levels



Radioactive families: defined by specific energies of a decays

1)4n: Th series 2)4n+1: Np series 3)4n+2: U series 4) 4n+3:At series

> B⁻ stabile 20 8⁻⁻ instabile 15 Energia liberata (MeV) 10 91 92 93 88 89 90 87 u Th Pa Bn Fr Ra Ac Numero atomico Z

Heavy nuclei have neutron excess, but with a decay they lose in percentage more protons: daughter nuclei are beta emitters



One important application is that of a decay tagging of the decay of SHE nuclei

Limits of stability of nuclear systems with very high number of proton is not well defined



To recognize the decay from an unknown nucleus one follows the decay chain down to known cases



Very low production cross sections Very low count rates Long experiments





$$\begin{array}{ll} {}^{A}_{Z}X \rightarrow {}^{A}_{Z+1}Y + e^{-} + v_{e} & \beta^{-} \text{ decay} \\ \\ {}^{A}_{Z}X \rightarrow {}^{A}_{Z-1}Y + e^{+} + v_{e} & \beta^{+} \text{ decay} \\ \\ \\ {}^{A}_{Z}X + e^{-} \rightarrow {}^{A}_{Z-1}Y + v_{e} & \text{EC} \end{array}$$



 β decay is a weak interaction "semi-leptonic" decay

The quark level Feynman diagram for β^{-} decay is shown here:



The Q value in β decay is effectively shared between the electron and antineutrino.

This is the case of gs-> gs decay

Q value for
$$\beta^-$$
 decay is
 $Q_{\beta^-} = (M(A, Z) - M(A, Z + 1))c^2$



$$\begin{split} &Q_{\beta} \quad \text{Sum of the energy of the electron (positron) and antineutrino (neutrino)} \\ &Q_{\beta^{-}} = T_{M(A,Z+1)^{+}} + T_{e} + T_{v} \approx T_{e} + T_{v} \quad \left(\text{since } T_{M(A,Z+1)^{+}} < keV\right) \\ &Q_{\beta^{-}} = \left(T_{e}\right)_{\max} = \left(T_{v}\right)_{\max} \quad \text{If the other term is null} \\ &Q_{\beta^{+}} = M(A,Z) - M(A,Z-1) - 2m_{e}c^{2} \end{split}$$

For electron capture: $Q_{EC} = (M(A,Z) - M'(A,Z-1))c^2 - B_n$ B_n binding energy of n - shell electron

NB: EC and $\beta\text{+}$ are not always competing: if $\beta\text{+}$, EC is possible, the contrary is not guaranteed

Note these are atomic masses



when $L_{\beta} = n = 0$

and $\pi_i \pi_f = +1$

Electron and neutrino do not carry angular momentum when the angular momentum conservation requires that $L_{\beta} = n > 0$ and/or $\pi_{i}\pi_{f} = -1$

$$\Delta I = \left| I_i - I_f \right| \equiv 0,1$$

I.

Selection rules	Type of transition	Order of forbiddenness	ΔΙ	$\pi_i \pi_f$
decays	Allowed		0,+1	+1
		1	∓2	-1
	Forbidden unique	2	∓ 3	+1
		3	∓ 4	-1
		4	Ŧ 5	+1
		•	•	•
		1	0, ∓1	-1
	Forbidden	2	∓ 2	+1
		3	Ŧ 3	-1
		4	∓ 4	+1

Classification of allowed decays







Useful empirical rules

The fifth power beta decay rule: the speed of a β transition increases approximately in proportion to the fifth power of the total transition energy



depends on spin and parity changes between the initial and final state

additional hindrance due to nuclear structure effects :

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isospin, "I-forbidden", "K-forbidden", etc.
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Fermi Golden Rule

Treat beta decay as a transition that depends upon the strength of coupling between the initial and final states Decay constant is given by Fermi's Golden Rule

$$\lambda_{\beta} = \frac{2\pi}{\hbar} |M|^2 \rho(E_f); M = \int \psi_f V \psi_i dv$$

→Electron and neutrino do not pre-exist in atom but are formed at the time of decay

→The decay is the result of the interaction btw the nucleon and the field produced by the electron-antineutrino couple \rightarrow weak interaction

- Perturbation theory can be applied since the interaction is "weak" M matrix element which couples the initial and final states •
- ٠
- Rate proportional to the strength of the coupling between the initial and final states factored by the density of final states available to the system ٠ electron-antineutrino

 $V=g \, \delta(\mathbf{r}_e - \mathbf{r}) \delta(\mathbf{r}_v - \mathbf{r})$

g is the "intensity" of the interaction There are two values: $q_F e q_{GT}$

 $t \equiv T_{1/2}^{\beta_i} = \frac{T_{1/2}^{\exp}}{P_{\beta_i}}$ partial half-life of a given $\beta^{-}(\beta^{+}, EC)$ decay branch (*i*) $\frac{\ln 2}{T_{1/2}^n} = \frac{g^2}{2\pi^3} \int_1^W p_e W_e (W_0 - W_e)^2 F(Z, W_e) C_n dW_e$ g - weak interaction coupling constant p_e - momentum of the β particle W_e - total energy of the β particle W_{ρ} - maximum energy of the β particle $F(Z, W_{\rho})$ - Fermi function - distortion of the β particle wave function by the nuclear charge C_n - shape factor Z - atomic number 1

$$K = 64\pi^4 m_o^5 c^4 g^2 / h^7 \qquad \lambda = \frac{\ln 2}{t_{1/2}} = K \left| M_{if} \right|^2 f_o$$

$$f_o = \int_{1}^{W_o} F(Z, W) W (W^2 - 1)^{1/2} (W_o - W)^2 dW$$

Comparative Half Lives

Based on probability of electron energy emission coupled with spectrum and the Coulomb correction $f_0 t_{1/2}$ is called the comparative half life of a transition

$$f\tau = \frac{2\pi^3\hbar^7 c^3}{g^2 |M_{if}|^2}$$

Assumes matrix element is independent of energy (true for allowed transitions) Yields f_{τ} (or $f_o t_{1/2}$), comparative half-life may be thought of as the half life corrected for differences in Z and energy

ALLOWED transitions second term is independent on nucleus,
 → ft has the same value for all allowed transitions
 For forbidden decays ft increases with degree of forbiddenness



<u>Measuring logft</u>

Superallowed Gamow–Teller decay of th doubly magic nucleus ¹⁰⁰Sn







Relativistic Fragmentation reaction ¹²⁴Xe beam on ⁹Be target

~260 100 Sn nuclei produced (0.75/h) ~ 126 fully reconstructed decay chains



Figure 2 | Time distribution of first decay events. The histogram shows the observed time distribution of all first decay events in the nearest-neighbouring pixels after implantation of 100Sn nuclei. Decay curves resulting from the MLH analysis are shown individually for 100Sn (dashed) and its daughter nucleus 100 In (dash-dot). The solid line shows the sum of these decay curves and takes into account a small amount of random background.



Figure 5 | Distribution of the positron energies emitted in the β -decay of ¹⁰⁰Sn. The spectrum contains only decay events that can be assigned to ¹⁰⁰Sn decays with a probability of at least 75%. The MLH fit was applied to the region between 400 and 2,600 keV, which is indicated with markers. The solid curve illustrates the shape of the best-fitting single-component β-decay phase space function determined by MLH analysis.

Standard model description of the mixing btw quark flavours: CKM (Cabibbo Kobayashi Maskawa) mixing matrix

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} \approx \begin{bmatrix} 0.974 & 0.225 & 0.003 \\ 0.225 & 0.973 & 0.041 \\ 0.009 & 0.040 & 0.999 \end{bmatrix}$$
Quark
Weak
States
States
States

CKM is a unitary matrix. If not there is physics beyond SM To test this we concentrate to the first line, which is known with greater precision



Test of fundamental symmetries of the weak interaction

- 1) Verify the Conserved Vector Current hypothesis:
 - $G_{\rm F}$ is independent on nuclear interaction
- 1) Determine the Vud element



Vud extracted with different decay studies: most precise value is given by superallowed Fermi decays



Superallowed $0^+ \rightarrow 0^+$ decay between T=1 analog states depends uniquely on the vector part of the weak interaction and, according to the Conserved Vector Current (CVC) hypothesis, its experimental ft value should be directly related to the vector coupling constant, which is a fundamental constant, equal to all the transitions.

The expression for ft includes small correction terms, which account to 1% of its value. It is therefore convenient to define a "corrected" Ft value:



Transition dependent

Transition independent in CVC hypothesis

Being: f statistical rate, t half-life, G=G_V vector coupling constant for semileptonic weak int. δ'_{R}, δ_{NS} and Δ_{R} are radiative corrections,

 $\delta_{\mathcal{C}}$ is an isospin-symmetry breaking term.

Quantities to extract JYFLŤRAP **ISOLTHAP** Canadian PT experimentally are: Ð. 25 LEBIT. TITAN. Half-life Branching ratio with different 30 ⁵⁴Co transitions ⁵⁰lln 25 Z^{-20} 15'CI 3090 (a) ° AI T 10 3080 Ŧ 3070 Ð (s) # 3060 3050 O, łŦ 5 $\mathbf{20}$ Π 10 15 25 3040 Ν 3030 ²²Mg ³⁸Ca ⁴⁶V ¹⁰C (b) 14O 26mAI 34CI 38mK ⁵⁰Mn ⁶²Ga ⁷⁴Rb 14 transitions 8 of which with 34Ar 42Sc 54Co 3090 *7t*(s) uncertainties lower than 1% 3080 3070 3060 30 10 20 40

Z of daughter

^{A4}Bb

20

35

Superallowed decays

Superallowed decays



Weights of the different contributions to the uncertainties in the measurements

Evolution of the precision in the determination of the Vud values in past 30 years



- * Improved productions and separation of beams
 - → lower bg conditions
 - →higher statistics allow for exclusive measurements
- * Simulations help to define properties of detectors
- * Definition of goal-standard in data analysis

Exotic nuclei with excess of neutrons have different Fermi energies for n and p Allowed decays might not be possible owing to large mismatch in wave functions

First forbidden transitions play a role



²⁰⁸Pb





Impact on r-process abundances

During "Freeze-out": detour of β -decay chains \Rightarrow *r-abundance changes*

During "Freeze-out": enhancement of neutron flux ⇒ *r-abundance changes*

Neutron number N



Quantities that can be extracted in a β decay experiment



"pandemonium effect" - neutron rich nuclei - log ft is a just upper limit

Pandemonium effect





• HPGe detectors are conventionally used to construct the level scheme populated in the decay

→ Higher Qvalue higher possibility of missing feeding

 $\boldsymbol{\cdot} \textbf{From}$ the γ intensity balance we deduce the B-feeding

Pandemonium effect implies

ightarrow Wrong definition of gamma feeding and branching ratios

 ${\tt I}_{\beta}$ and logft

Measuring β decay



Beam

HPGe

PHYSICAL REVIEW C 87, 045502 (2013)

Improved half-life determination and eta-delayed γ -ray spectroscopy for $^{18}\mathrm{Ne}$ decay

Evolution in the precision of half-life measurements, example of ¹⁸Ne



Setup SCEPTAR+8 π gamma spectr. @TRIUMF (CANADA) + Tape system





* 8π array of 20 high-purity germanium (HPGe) gamma-ray detectors that provides high-resolution measurements of the gamma rays emitted from excited states of the atomic nucleus.



Part III: Examples

PHYSICAL REVIEW C 87, 045502 (2013)

Improved half-life determination and eta-delayed γ -ray spectroscopy for ¹⁸Ne decay



Best fit - - Background ¹⁸Ne 18 F 20 40 100 120 140 60 80 160 Time (s) (b) ٠ Gate on 1042 keV Best fit - - - Backer 20 40 60 80 100 120 140 160 Time (s)

Gate on 511 keV

٠

Gating on a specific γ line of the daughter nucleus cleans the spectrum \rightarrow only contributions from wanted decay

Evolution of precision of $T_{1/2}$ meas.:

- Better reduction of contaminants
- Higher efficiency and selectivity
- Faster electronics

PHYSICAL REVIEW LETTERS

(2014)

PRL 112,

 β Decay

week ending 14 MARCH 20

from Mirror -Breaking Isospin of test Sensitiv of ³⁸Ca:

Superallowed $0^+ \rightarrow 0^+$ Transitions

 $^1\text{H}(^{39}\text{K},\,\text{2n})^{38}\text{Ca}$ reaction on a LN_2-cooled hydrogen gas target with 30A-MeV ^{39}K .

Ejectiles mass separated

Implanted on aluminized Mylar tape of a fast tape transport system.

1-mm-thick BC-404 scintillator to detect β particles, and an expecially calibrated 70% HPGe



Careful determination of β and γ efficences

Branching ratio is defined as:

$$R_i = \frac{N_{\beta\gamma_i}}{N_{\beta}\varepsilon_{\gamma_i}} \frac{\varepsilon_{\beta}}{\varepsilon_{\beta_i}},\tag{3}$$

where $N_{\beta\gamma_i}$ is the total number of β - γ coincidences in the γ_i peak; N_{β} is the total number of beta singles corresponding



Exotic decay modes: Rare decay modes

Proton emission: * emission from both $gs \rightarrow$ direct emission * emission from excited states $\rightarrow \beta$ delayed proton emission

Direct emission * usually in competition with β decay *: Ep small: nucleus decays via β decay Ep big: T_{1/2}(p) very short and p cannot be detected

Typically seen in Z~50, Z~80, con 10⁻¹² s<T_{1/2}(p) < 0.1 s

Is again described as a decay as a tunneling through barrier Here centrifugal barrier is very important Delayed emission if p is emitted after β decay \rightarrow competition with γ decay



¹⁵¹Lu can decay via

- Beta decay
- Proton decay
- β -delayed proton emission





2 proton emission

Unbound even-Z nuclei Described by eg. Di-proton model: The 2p leave the nucleus correlated in a s wave

Recent predictions show that there are numerous examples but soon start deviating and become not accessible exp.







Figure 1 Two-proton decay from ⁴⁵Fe. **a**, One-proton decay to ⁴⁴Mn requires energy input and is not a possible outcome. In two-proton decay to ⁴³Cr, two protons share a kinetic energy of about 1 MeV. **b**, A ⁴⁵Fe ion enters the sensitive region of the chamber from the left, stops in the gas and emits two protons — seen as a V shape.

b

PRL 110, 222501 (2013) PHI SICAL REVIEW LETTERS 31 MAY 2	PRL 110, 222501 (2013)	PHYSICAL	REVIEW	LETTERS	week ending 31 MAY 201
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Landscape of Two-Proton Radioactivity

E. Olsen,^{1,2} M. Pfützner,^{3,4} N. Birge,^{1,2} M. Brown,^{1,5} W. Nazarewicz,^{1,2,3} and A. Perhac^{1,2}

β delayed fission



¹⁸⁰TI: T_{1/2} = 1.1 s, 150 atoms/second, β-delayed fission probability is 10⁻⁶
 ¹⁸⁰Hg (Z=80, N=100): expected to split into two ⁹⁰Zr(Z=40, N=50):
 → symmetric mass split

A. Andreyev et al., PRL 105 (2010) 252502

Cluster decay→ very Asymmetric fission

Spontaneous emission of particles heavier than a: ¹⁴C, ²⁰O,²⁴Ne,²⁴Mg,²⁸Si...³⁴Si Cluster are in general n-rich and 14<A<34

Can be described either as a generalized a decay or as a very asymmetric fission



 $\begin{array}{l} \textbf{Probable Clusters: } \ ^{3}\textbf{Ile}, \ ^{3}\!^{16}\textbf{Be}, \ ^{12+6}\textbf{C}, \ ^{16}\textbf{N}, \ ^{16}\!^{16}\!^{12+22}\textbf{O}, \ ^{22}\!^{23}\textbf{F}, \ ^{22}\!^{24}\!^{12}\!^{16}\textbf{Ne}, \ ^{24}\!^{24}\!^{28}\textbf{Ma}, \ ^{23}\!^{24}\!^{12}\!^{12}\!^{16}\textbf{Mg}, \ ^{31}\!^{32}\textbf{Al}, \ ^{23}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{-36}\!\textbf{Mg}, \ ^{31}\!^{32}\!^{4}\textbf{Al}, \ ^{23}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{-36}\!\textbf{Ng}, \ ^{31}\!^{32}\!^{4}\textbf{Al}, \ ^{23}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{-36}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!^{26}\!^{26}\!\textbf{Si}, \ ^{3}\!^{24}\!^{24}\!^{24}\!^{24}\!^{24}\!^{24}\!^{24}\!^{24}\!^{24}\!^{26}\!$

3 out of 4 fundamental interactions are at play in the nucleus: strong/em/weak

They determine the stability/instability of the system They define the means for a nucleus to decay

More frequent decays are:

- α decay \rightarrow tunneling through barrier
 - → heavy systems, 2 body process, well defined a energies
- β decay \rightarrow most frequent
 - → weak interaction "semi-leptonic" decay
 - → 3 body system, continuous emission spectrum
 - \rightarrow governed by selection rules

Spontaneuos fission > heavy systems, tunneling thr. barrier

Going further away of stability line more exotic decays are registered:

- β delayed emission of n, β delayed fission
- Proton decay and 2p decay
- Cluster decays.....





7000+/- 500 current estimates for bound nuclei

Boundaries of the bound nuclei are not defined yet