Re-writing Nuclear Physics textbooks: 30 years of radioactive ion beam physics Basic concepts in nuclear reaction theory

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Bibliography

General scattering theory:

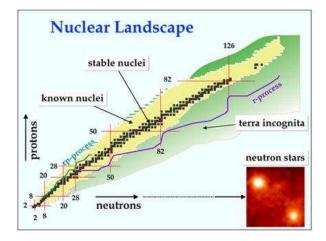
• Quantum collision theory, C.J. Joachain.

Scattering theory applied to nuclear reactions:

- Introduction to nuclear reactions, G.R. Satchler.
- Direct Reactions, G.R. Satchler.
- Direct Nuclear Reactions, N. Glendenning.
- Nuclear reactions for astrophysics, I.J. Thompson and F.M. Nunes.
- Quantum scattering theory and direct nuclear reactions, course notes by A.M.M.

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Unstable nuclei and the limits of stability

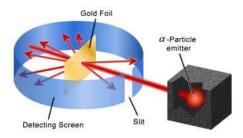


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Motivation of reaction theory

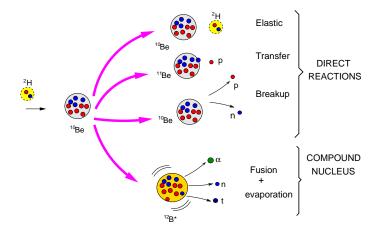
⇒ The aim of reaction theory is to provide a mathematical description of quantum scattering experiments, in order to extract information on the structure of the colliding nuclei and on their mutual interaction dynamics.

 \Rightarrow The first experiment of this kind was the α scattering experiment by Rutherford, who lead to the proposal of his celebrated model of the atom and the subsequent formula for the angular dependence of the scattered α particles.



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Types of reactions: direct vs. compound nucleus processes



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Direct versus compound reactions

DIRECT: elastic, inelastic, transfer,...

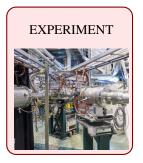
- "fast" collisions (10^{-21} s) .
- only a few modes (degrees of freedom) involved
- small momentum transfer
- angular distribution asymmetric about $\pi/2$ (peaked forward)

COMPOUND: complete, incomplete fusion.

- many degrees of freedom involved
- large amount of momentum transfer
- "loss of memory" \Rightarrow almost symmetric distributions forward/backward

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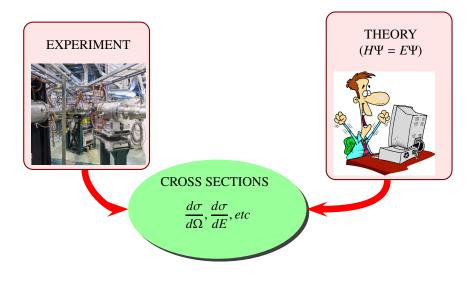
Linking theory with experiments: the cross section



THEORY
$$(H\Psi = E\Psi)$$

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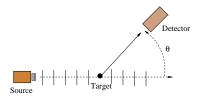
Linking theory with experiments: the cross section



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Experimental cross section

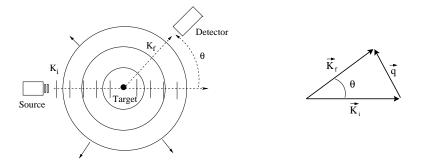


$$\Delta I = I_0 \ n_t \ \frac{d\sigma}{d\Omega} \Delta \Omega$$

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- ΔI : detected particles per unit time in $\Delta \Omega$
- I_0 : incident particles per unit time
- *n_t*: number of target nuclei per unit surface
- $\Delta \Omega$: solid angle of detector
- $d\sigma/d\Omega$: differential cross section

 $\frac{d\sigma}{d\Omega} = \frac{\text{flux of scattered particles through } dA = r^2 d\Omega}{\text{incident flux}}$



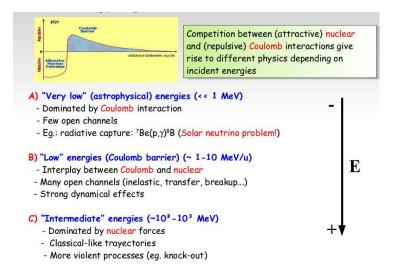
Among the many mathematical solutions of $[H - E]\Psi = 0$ we are interested in those behaving asymptotically as:

$$\Psi_{\mathbf{K}_{\alpha}}^{(+)} \to \Phi_{\alpha}(\xi_{\alpha})e^{i\mathbf{K}_{\alpha}\cdot\mathbf{R}_{\alpha}} + (\text{outgoing spherical waves})$$

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Energy domains



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Elastic scattering

What can we learn by measuring elastic scattering?

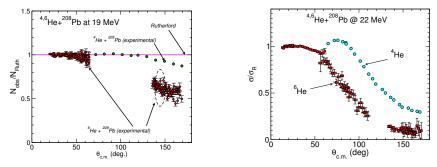
- Studying the angular dependence of elastically scattered particles we can infer information on:
 - The interplay between Coulomb and nuclear forces.
 - The presence of non-elastic channels, that will show up as a reduction of the elastic cross section with respect to the case of inert objects (*absorption*).
- From scattering theory, the angular distribution is calculated from the scattering wavefunction as:

$$\Psi^{(+)}(\mathbf{K},\mathbf{R}) \to e^{i\mathbf{K}\cdot\mathbf{R}} + f(\theta)\frac{e^{iKR}}{R}$$

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2$$

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Rutherford experiment...100 years later



• ⁴He follows Rutherford formula at 19 MeV but not at 22 MeV.Why?

• ⁶He drastically departs from Rutherford formula at both energies. Why?

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Reaction and interaction cross sections

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Interaction cross sections

⇒ Interaction cross section:

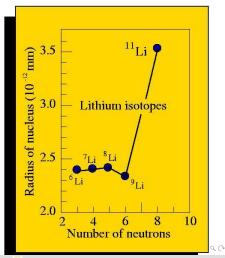
 $\sigma_I = \sigma_{\rm tot} - \sigma_{\rm inel} - \sigma_{\rm el}$

⇒ Interaction radius:

$$\sigma_I = \pi \left(R_I^{proj} + R_I^{targ} \right)^2$$



Tanihata et al, Phys. Rev. Lett. 55 (1985) 2676

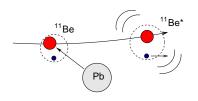


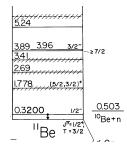
Inelastic scattering

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Inelastic scattering

- Nuclei are not inert or *frozen* objects; they do have an internal structure of protons and neutrons that can be modified (excited) during the collision.
- Quantum systems exhibit, in general, an energy spectrum with bound and unbound levels.





Inelastic scattering

- Direct reactions \rightarrow nuclei make "glancing" contact and separate immediately.
- Energy/momentum transferred from relative motion to internal motion so the projectile and/or target are left in an excited state.
- Involve small number of degrees of freedom.
- The colliding nuclei preserve their identity: $a + A \rightarrow a^* + A^*$
- Typically, they are peripheral (surface) processes.

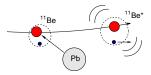
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Models for inelastic excitations

• COLLECTIVE: Involve a collective motion of several nucleons which can be interpreted macroscopically as rotations or surface vibrations of the nucleus.



FEW-BODY/SIGLE-PARTICLE: Involve the excitation of a nucleon or cluster.

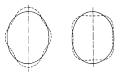


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Types of collective excitations

The nucleons can move inside the nucleus in a coherent (collective) way.

• Vibrations (spherical nuclei): small surface oscillations in shape.



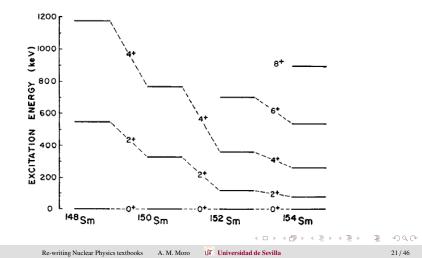
- Rotations (non-spherical nuclei): permanent deformation.
- Monopole (*breathing*) mode: oscillations in the size (radius).
- Isovector excitations (protons and neutrons move out of phase) (eg. giant dipole resonance)

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Types of collective excitations

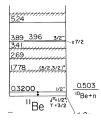
The type of collective motion is closely related to the kind of energy spectrum.

- Rotor: $E_J \propto J(J+1)$
- Vibrator: $E_J \approx n\hbar\omega$

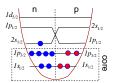


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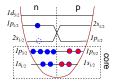
Microscopic description in the IPM: the ¹¹Be case







First excited state $(1/2^{-})$

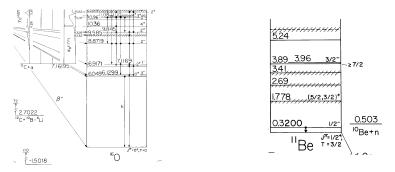


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Models for inelastic excitations

Microscopically, what we describe in both cases are quantum transitions between discrete or continuum states:



©Collective excitations can be regarded as a coherent superposition of many single-particle excitations.

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• By doing inelastic scattering experiments we *measure* the *response* of the nucleus to an external field (Coulomb, nuclear). This response is related to some structure property of the nucleus.

Example: for a Coulomb field:

$$B(E\lambda; i \to f) = \frac{1}{2I_i + 1} |\langle \Psi_f | \mathcal{M}(E\lambda) | \Psi_i \rangle|^2$$

where $\mathcal{M}(E\lambda,\mu)$ is the electric multipole operator:

$$\mathcal{M}(E\lambda,\mu) \equiv e \sum_{i}^{Z_p} r_i^{\lambda} Y_{\lambda\mu}^*(\hat{r}_i)$$

• The structure $\Psi_{i,f}$ can be described in a collective, few-body or microscopic model.

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Energy balance for inelastic scattering

• For projectile excitation: $a + A \rightarrow a^* + A$

$$E_{\rm cm}^{i} + M_a c^2 + M_A c^2 = E_{\rm cm}^{f} + M_a^* c^2 + M_A c^2$$

$$M_{a^*} = M_a + E_x$$
 (*E*_x=excitation energy)
Q-value:

$$Q = M_a c^2 + M_A c^2 - M_a^* c^2 - M_A^2 c^2 = -E_x < 0$$

$$E_{\rm cm}^f = E_{\rm cm}^i + Q$$

• So

$$E_x = E^i_{cm} - E^f_{cm}$$

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What do we measure in an inelastic scattering experiment?

In general, one measures the scattering angle and energy of outgoing particles.

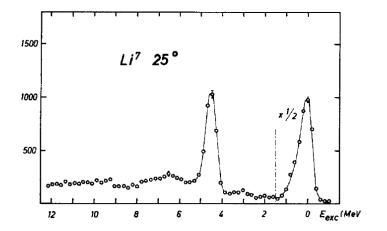
Example: $p+^{7}Li \rightarrow p+^{7}Li^{*}$ proton beam \bullet T_{Li} Target

See Eg. energy and angular distribution of the outgoing protons.

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What do we measure in an inelastic scattering experiment?

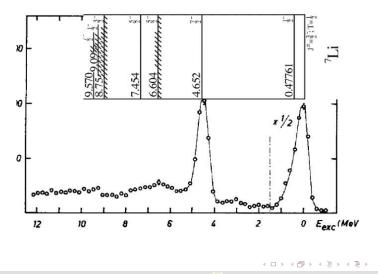
The proton energy carries information on the ⁷Li excitation spectrum.



Data from Nuclear Physics 69 (1965) 81-102

What do we measure in an inelastic scattering experiment?

The proton energy carries information on the ⁷Li excitation spectrum.

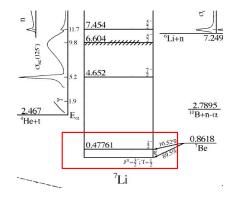


What information do we get from an inelastic scattering experiment?

- The proton energy spectrum shows peaks which correspond to the states of the target (⁷Li)
- The heights of peak (~ cross section) are different for each state ⇒ not all states are populated with the same probability.
- Some peaks are narrow, other are broad. Why?...
- Above a certain excitation energy, the spectrum becomes continuous and structureless.

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What information do we get from an inelastic scattering experiment?



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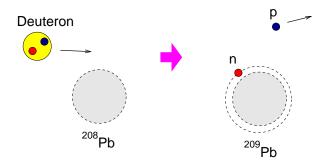
Transfer reactions

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Transfer reactions

Example: $d + {}^{208}Pb \rightarrow p + {}^{209}Pb$



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Transfer reactions: Q-value considerations

Consider: $a + A \rightarrow b + B$

• Energy balance (in CM frame):

$$E_{\rm cm}^i + M_a c^2 + M_A c^2 = E_{\rm cm}^f + M_b c^2 + M_B c^2$$

• Q_0 value:

$$Q_0 = M_a c^2 + M_A c^2 - M_b c^2 - M_B c^2$$

$$E_{\rm cm}^f = E_{\rm cm}^i + Q_0$$

- $Q_0 > 0$: the system gains kinetic energy (exothermic reaction)
- $Q_0 < 0$: the system loses kinetic energy (endothermic reaction)

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Transfer reactions: Q-value considerations

Example: $d + {}^{208}Pb \rightarrow p + {}^{209}Pb$

$$\frac{d + ^{208}Pb}{Q_0 = +1.7 \text{ MeV}}$$

$$Q_0 = M_d c^2 + M(^{208}\text{Pb})c^2 - M_p c^2 - M(^{209}\text{Pb})c^2 = +1.7 \text{ MeV}$$

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Transfer reactions: Q-value considerations

Example: $d + {}^{208}Pb \rightarrow p + {}^{209}Pb$

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IS $Q_0 > 0$: the outgoing proton will gain energy with respect to the incident deuteron.

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Transfer reactions: Q-value considerations

Example: $d + {}^{208}Pb \rightarrow p + {}^{209}Pb$

$$\frac{d + 208 Pb}{Q_0 = +1.7 MeV}$$

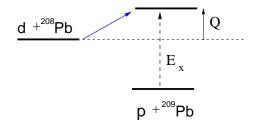
$$Q_0 = M_d c^2 + M(^{208}\text{Pb})c^2 - M_p c^2 - M(^{209}\text{Pb})c^2 = +1.7 \text{ MeV}$$

 $\mathbb{I} \otimes Q_0 > 0$: the outgoing proton will gain energy with respect to the incident deuteron.

For a transfer reaction, the Q value is just the difference in binding energies of the transferred particle/cluster in the initial and final nuclei:

Transfer reactions: Q-value considerations

If the transfer leads to an excited state, the *Q*-value will change, and hence the kinetic energy of the outgoing nuclei.



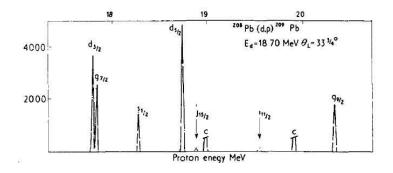
Energy balance:

$$E_{\rm cm}^f = E_{\rm cm}^i + Q = E_{\rm cm}^i + Q_0 - E_x$$

If we know Q_0 we can infer the excitation energies (E_x) measuring the final kinetic energy of outgoing fragments.

What we do observe in a transfer experiment?

Example: $d + {}^{208}Pb \rightarrow p + {}^{209}Pb$

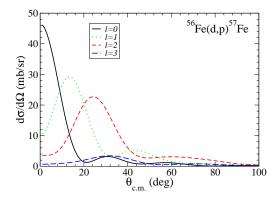


- The proton energy spectrum shows some peaks which reflect the excitation energy spectrum of the residual nucleus (²⁰⁹Pb).
- The population probability will depend on the reaction dynamics and on the structure properties of these states.

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Transfer example: ⁵⁶Fe(d,p)⁵⁷Fe

Selectivity of ℓ :

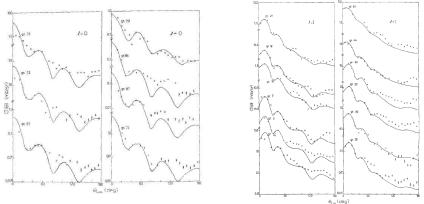


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Transfer example: ⁵⁶Fe(d,p)⁵⁷Fe

Selectivity of ℓ :



H.M. Sen Gupta et al, Nucl. Phys. A160, 529 (1971)

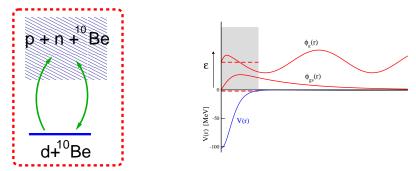
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Breakup reactions

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Breakup reactions

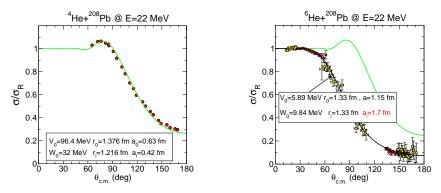
- Direct processes of the from: $a + A \rightarrow b + x + A$
- Can be interpreted (and modelled) as an inelastic excitation to the continuum spectrum.



• Important for weakly-bound nuclei (eg. halo nuclei)

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Influence of breakup on elastic scattering

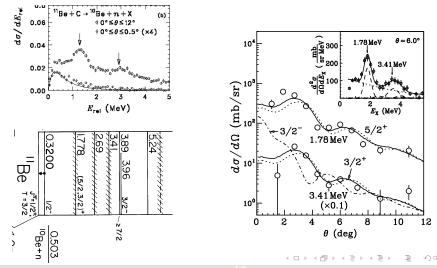


• ${}^{4}\text{He} + {}^{208}\text{Pb}$ shows typical Fresnel pattern \rightarrow *strong absorption*

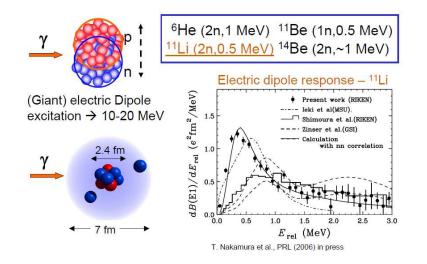
- ⁶He+²⁰⁸Pb shows a prominent reduction in the elastic cross section due to the flux going to other channels (mainly break-up)
- ${}^{6}\text{He}+{}^{208}\text{Pb}$ requires a large imaginary diffuseness $\rightarrow long$ -range absorption

Extracting information from the continuum with breakup reactions

Example: Populating resonances by "inelastic scattering" in ¹¹Be+¹²C Fukuda et al. Phys. Rev. C70 (2004) 054606)



Coulomb response of halo nuclei



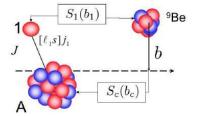
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Knockout reactions

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Knockout reactions

- Fast-moving projectile on a (typically) light target.
- One nucleon suddently removed (absorbed) due to its interaction with the target.
- The remaining nucleons remain unchanged and is detected.
- The momentum of the core is traced back to that of the removed nucleon because in the rest frame of the projectile $\vec{P} = 0$

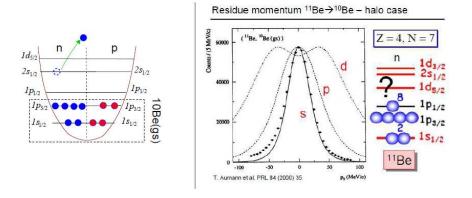


$$\vec{P} = \vec{p}_c + \vec{p}_1 = 0$$

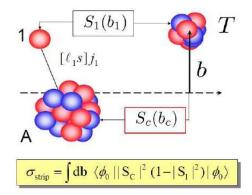
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Knockout reactions

- The shape is determined by the orbital angular momentum ℓ .
- The magnitude is determined by the amount of $s_{1/2}$ (spectroscopic factor)



Knockout reactions



- $|S_c(b_c)|^2$ =probability of survival of the core.
- $1 |S_1(b_1)|^2$ =probability of absorption of the neutron.

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Radiative capture

Radiative capture

Radiative capture: $b + c \rightarrow a + \gamma$

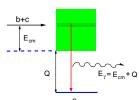
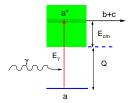


Photo-absorption: $a + \gamma \rightarrow b + c$



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⇒ Related by detailed balance:

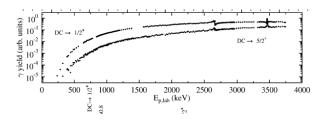
$$\sigma_{E\lambda}^{(rc)} = \frac{2(2J_a + 1)}{(2J_b + 1)(2J_c + 1)} \frac{k_{\gamma}^2}{k^2} \sigma_{E\lambda}^{(phot)} \qquad (\hbar k_{\gamma} = E_{\gamma}/c)$$

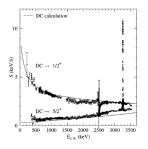
⇒ Astrophysical S-factor:

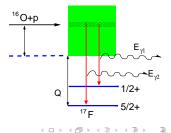
$$S(E_{\rm c.m.}) = E_{\rm c.m.}\sigma_{E\lambda}^{(rc)} \exp[2\pi\eta(E_{\rm c.m.})]$$

Example: $p + {}^{16}O \rightarrow {}^{17}F + \gamma$

Morlock, PRL79, 3837 (1997)







Implications in astrophysics: the r-process

 \Rightarrow Most neutron-rich isotopes of elements heavier than nickel are produced, by the beta decay of very radioactive matter synthesized during the so-called r process.

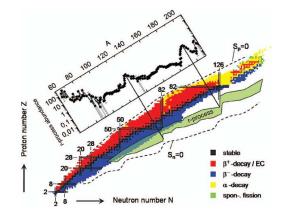


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