

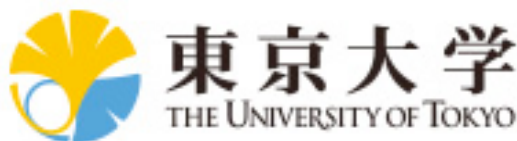
Hyper-Kamiokande

Masato Shiozawa

Kamioka Observatory, Institute for Cosmic Ray Research, U of Tokyo, and
Kamioka Satellite, Kavli Institute for the Physics and Mathematics of the Universe (WPI), U of Tokyo

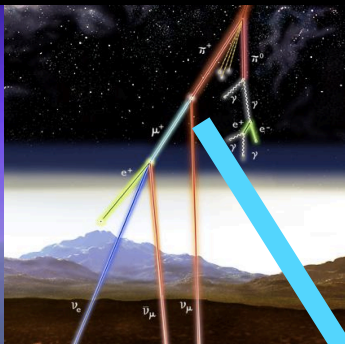
Neutrino Telescope 2015

March 5, 2015

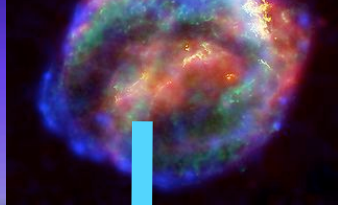


Overview of the Hyper-K

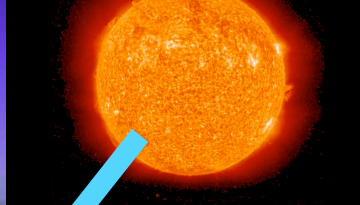
Atmospheric ν



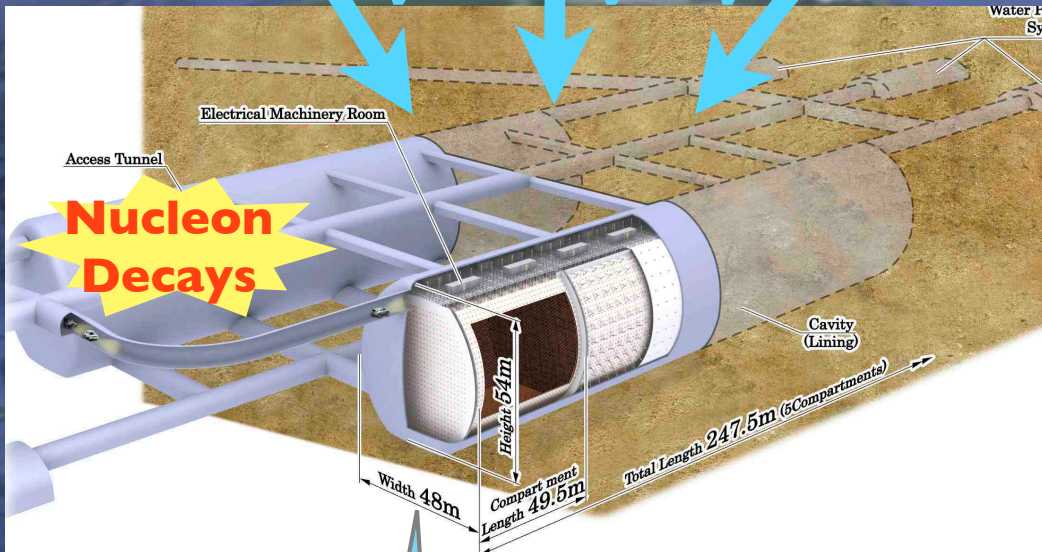
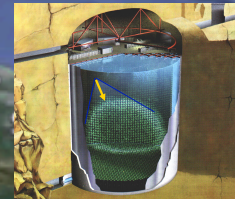
Supernova ν



Solar ν



Super-Kamiokande



Hyper-Kamiokande

25 x Super-K fiducial mass as neutrino target and proton decay source

J-PARC

High intensity neutrino and anti-neutrino beam

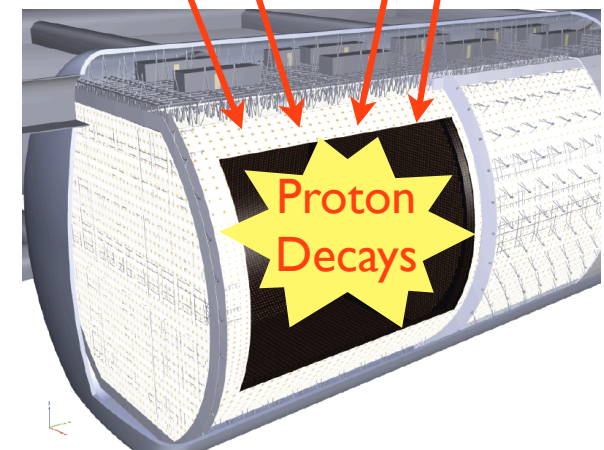
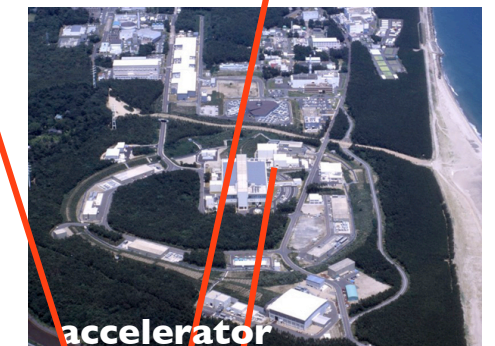
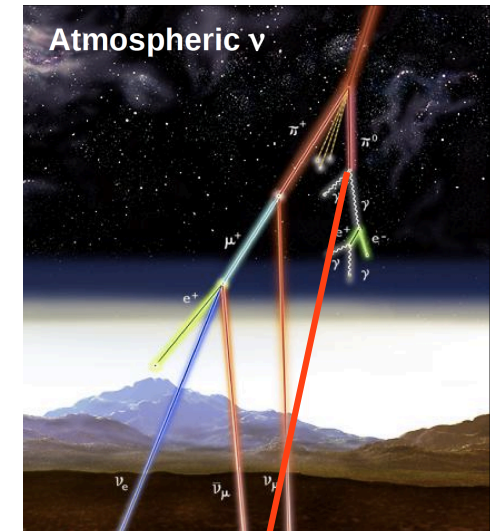
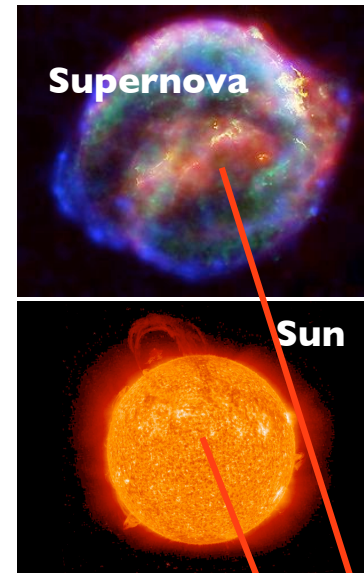


Multi-purpose detector, Hyper-K

Letter of Intent, Hyper-K WG,
arXiv:1109.3262 [hep-ex]

LBL study, Hyper-K WG,
arXiv:1502.05199 and
submitted to PTEP

- **Proton decay 3σ discovery potential**
 - 5×10^{34} years for $p \rightarrow e^+ \pi^0$
 - 1×10^{34} years for $p \rightarrow \nu K^+$
- **Comprehensive study on ν oscillations**
 - CPV (76% of δ space at 3σ), $<20^\circ$ precision
 - MH determination for all δ by J-PARC/Atm ν
 - θ_{23} octant: $\sin^2\theta_{23} < 0.47$ or $\sin^2\theta_{23} > 0.53$
 - $<1\%$ precision of Δm_{32}^2
 - test of exotic scenarios by J-PARC/Atm ν
- **Astrophysical neutrino observatory**
 - Supernova up to 2Mpc distance, ~ 1 SN / 10 years
 - Supernova relic ν signal ($\sim 200\nu$ events/10yrs)
 - Dark matter neutrinos from Sun, Galaxy, and Earth
 - Solar neutrino $\sim 200\nu$ events/day



Hyper-K proto-collaboration w/ cooperation of KEK-IPNS and UTokyo-ICRR

Inaugural Symposium on 1/31, 2015



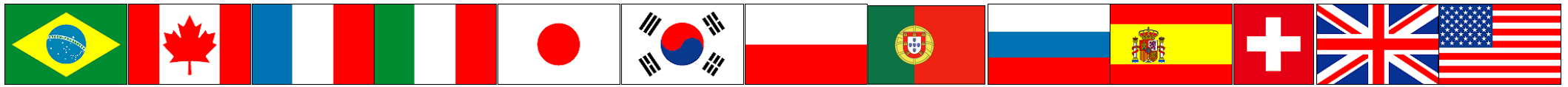
Hyper-K Proto-Collaboration has been formed

MoU signing by KEK/ICRR

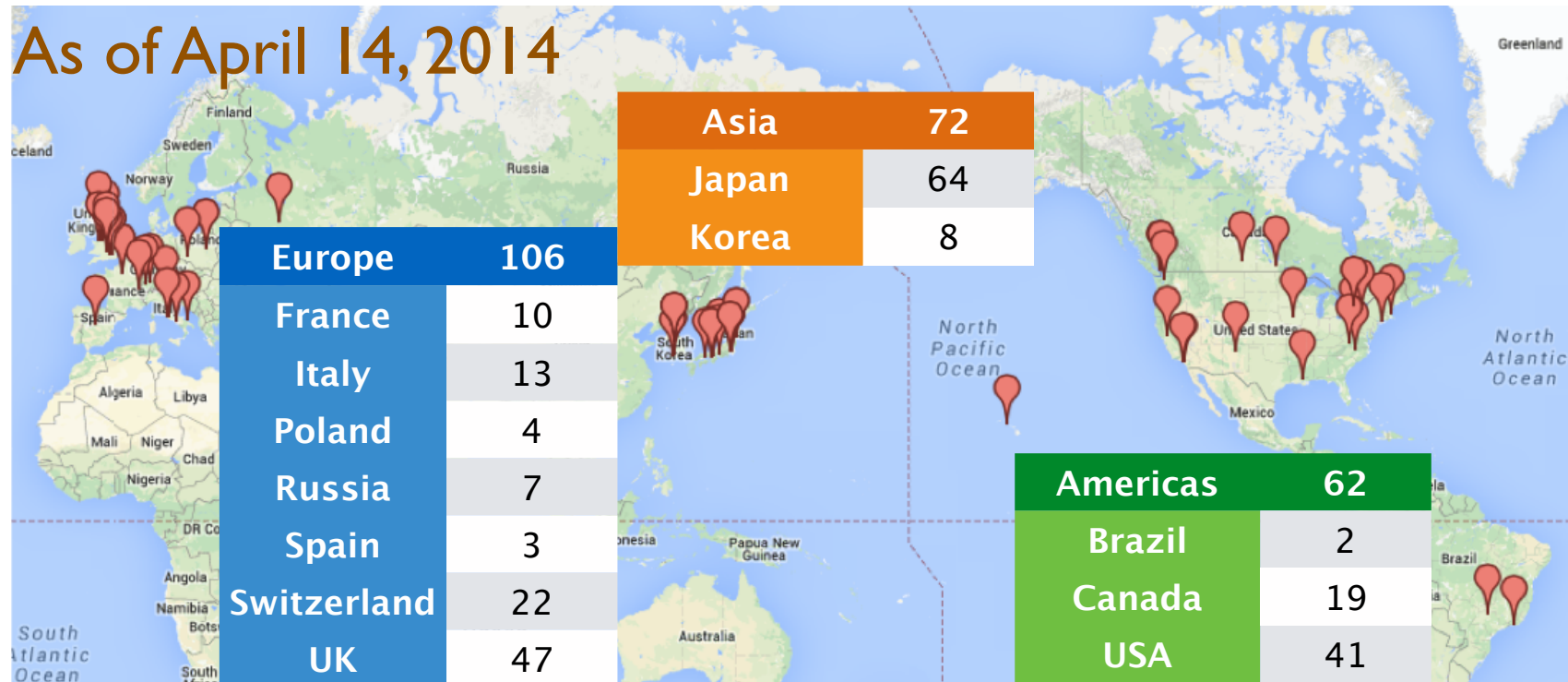
- **KEK-IPNS** and **Tokyo-ICRR** signed the **MOU** of the cooperation in promoting the Hyper-Kamiokande.



Hyper-Kamiokande International Group



As of April 14, 2014



- 240 people and growing!
- Hyper-K Governance Structure has been defined
 - Steering Committee, International Board Representatives, and Convener Board
- R&D fund and travel budget already secured in some countries, and more in securing processes.

What's next?: Design Report

- **Design Report** to be prepared in 2015
 - Optimum design, Construction cost/period, Beam&Near detectors, International responsibility
- **International review** under **KEK-IPNS/ICRR** to promote the project
- Start **budget request** in 2015~2016
- **Start construction** in 2018
→ start operation in ~2025

It is critical period to promote the project

Still open for new collaborators

Hyper-Kamiokande EU meeting@CERN

27-28 April 2015

- Meeting to discuss the European effort in Hyper-K
- Open to anyone who has interest in Hyper-K, or is planning to join Hyper-K, or is contributing
- <http://indico.cern.ch/e/ThirdEUHyperK>



Hyper-Kamiokande EU meeting

27-28 April 2015

CERN

Europe/Zurich timezone



Overview

Timetable

Registration

Participant List

Accommodation

- Meeting to discuss the European effort in the [Hyper-Kamiokande experiment](#).
- Open to anyone who has interests in Hyper-K, or is planning to join Hyper-K, or is contributing.
- Detailed information about your Country in Hyper-K can be discussed with your representatives in the [Hyper-Kamiokande International Board Representatives](#), its chair if no representatives are available yet or the [international Steering Committee chair](#).



Starts 27 Apr 2015 11:00

Ends 28 Apr 2015 18:00

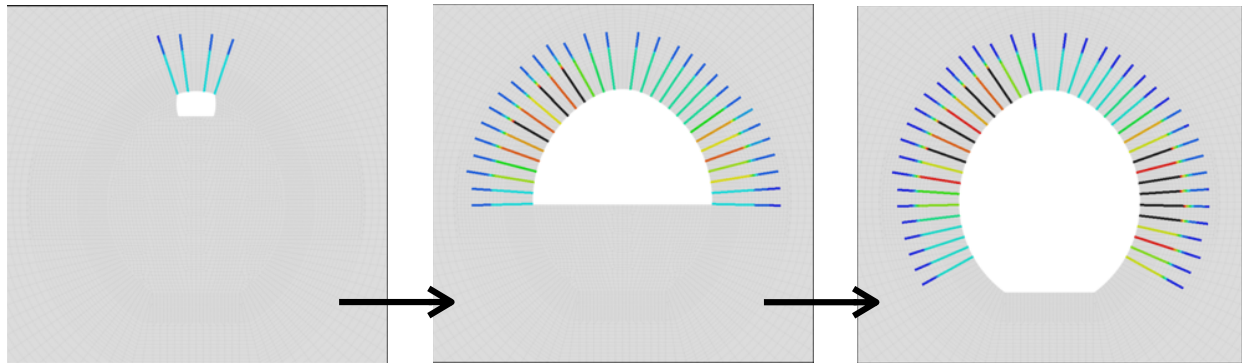
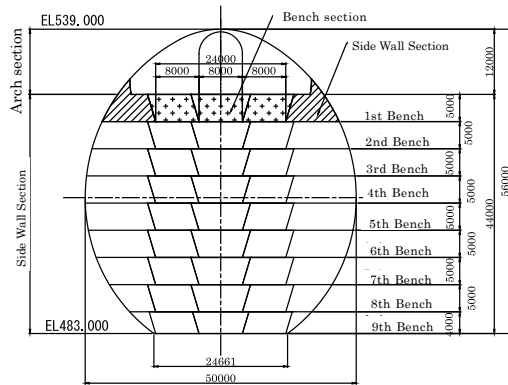
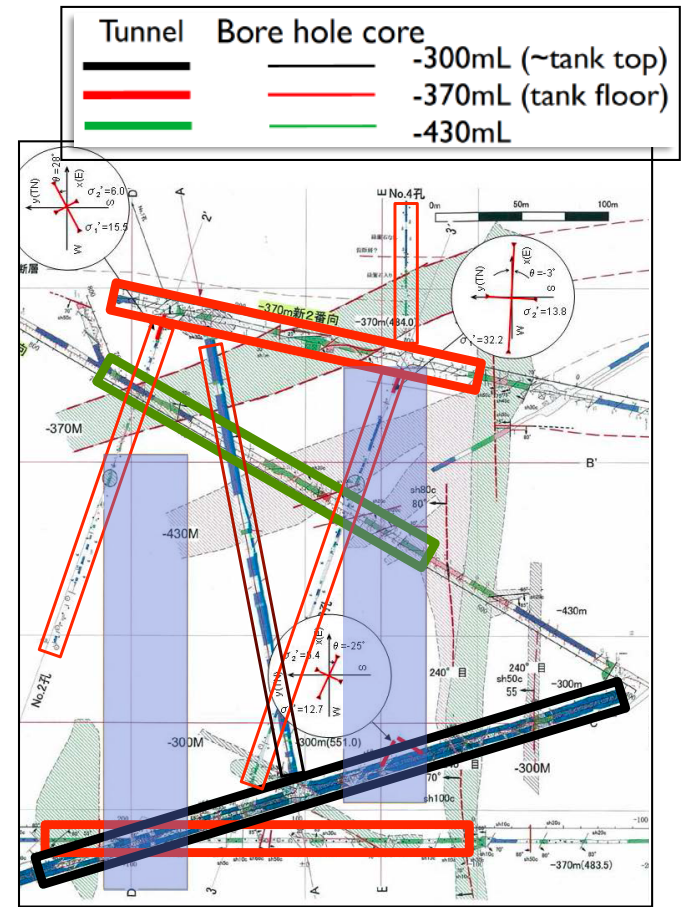
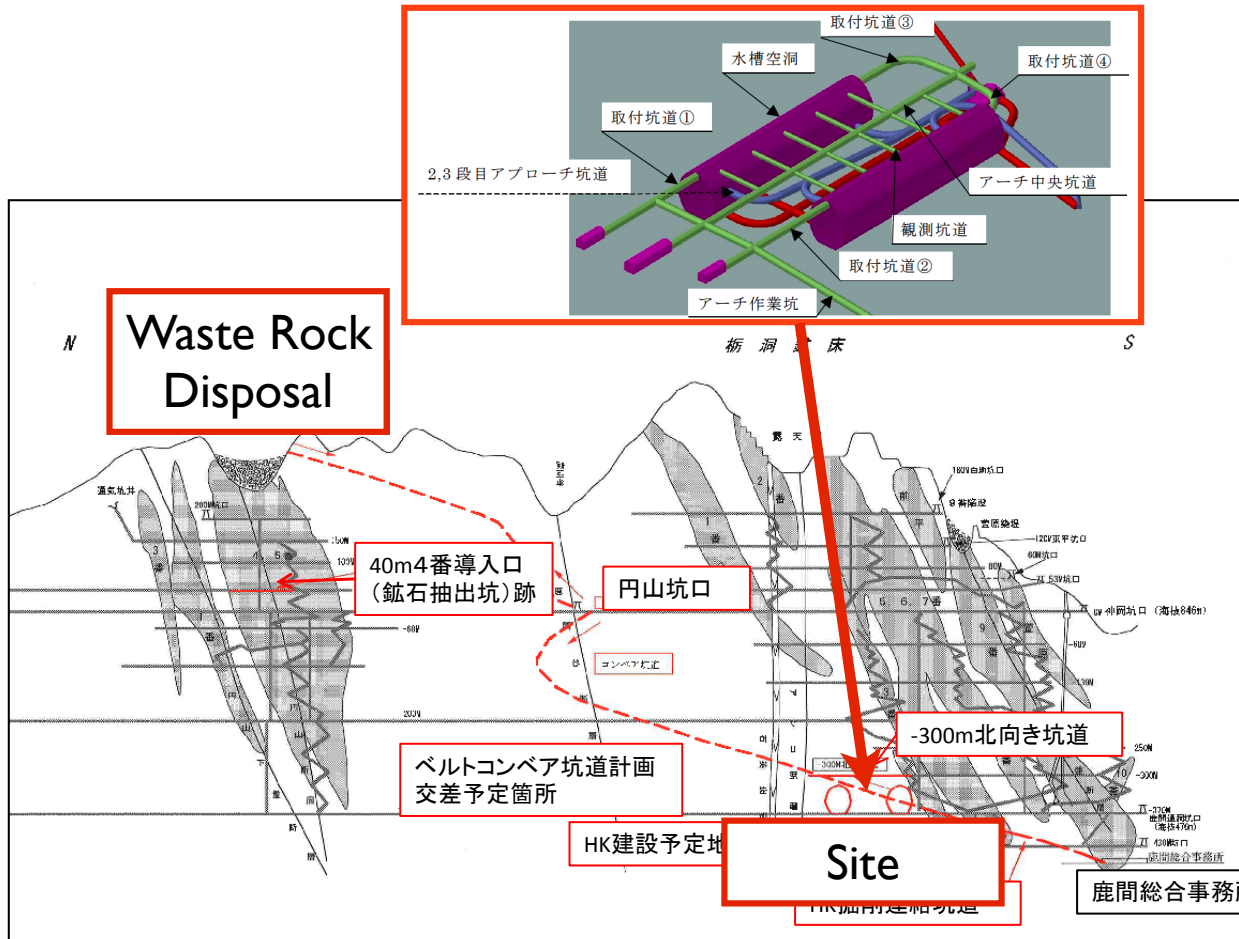
Europe/Zurich



CERN

IT Amphitheatre

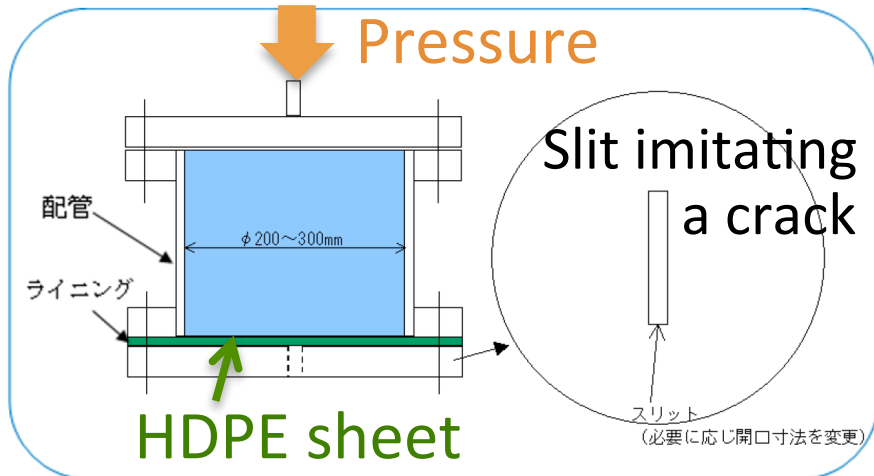
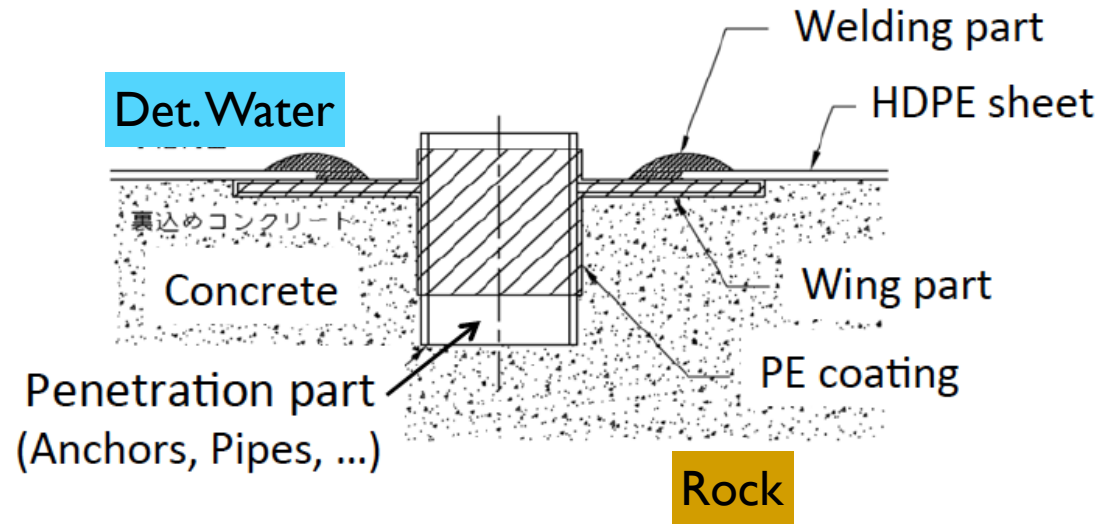
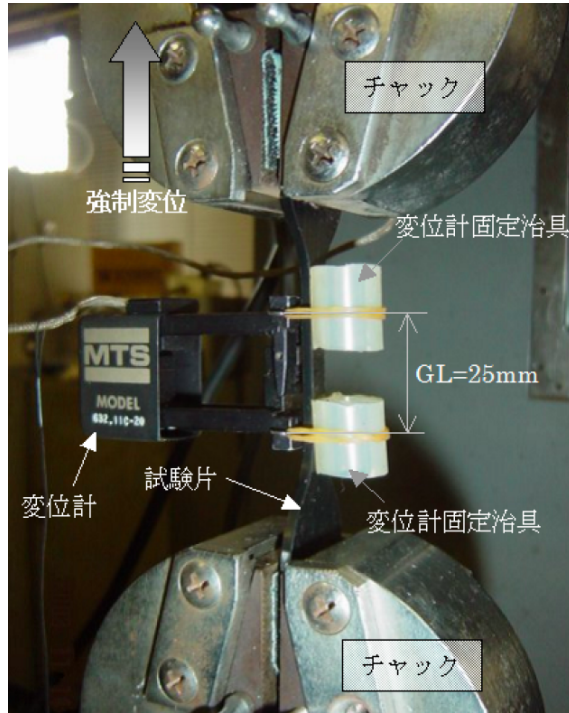
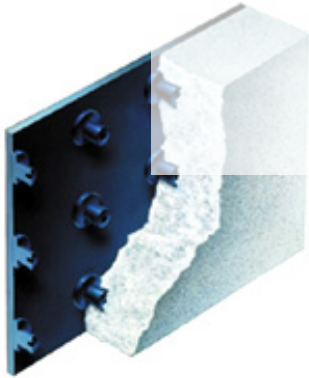
Design and ongoing studies



- Conceptual design of caverns, supports have been made.
- Estimation of excavation period/cost → Optimization

Tank liner material

5mm High Density Polyethylene



- Soak test
 - pure water, 1% $\text{Gd}_2(\text{SO}_4)_3$ loaded
- Tensile creep test
- pressure test
- leak test at the penetrating part

Satisfactory results for Hyper-K

Photo-sensor candidates

Super-K PMT



Venetian blind dynode

highQE/CE PMT



Box&Line dynode

highQE/CE Hybrid Det.



Avalanche photo-detector

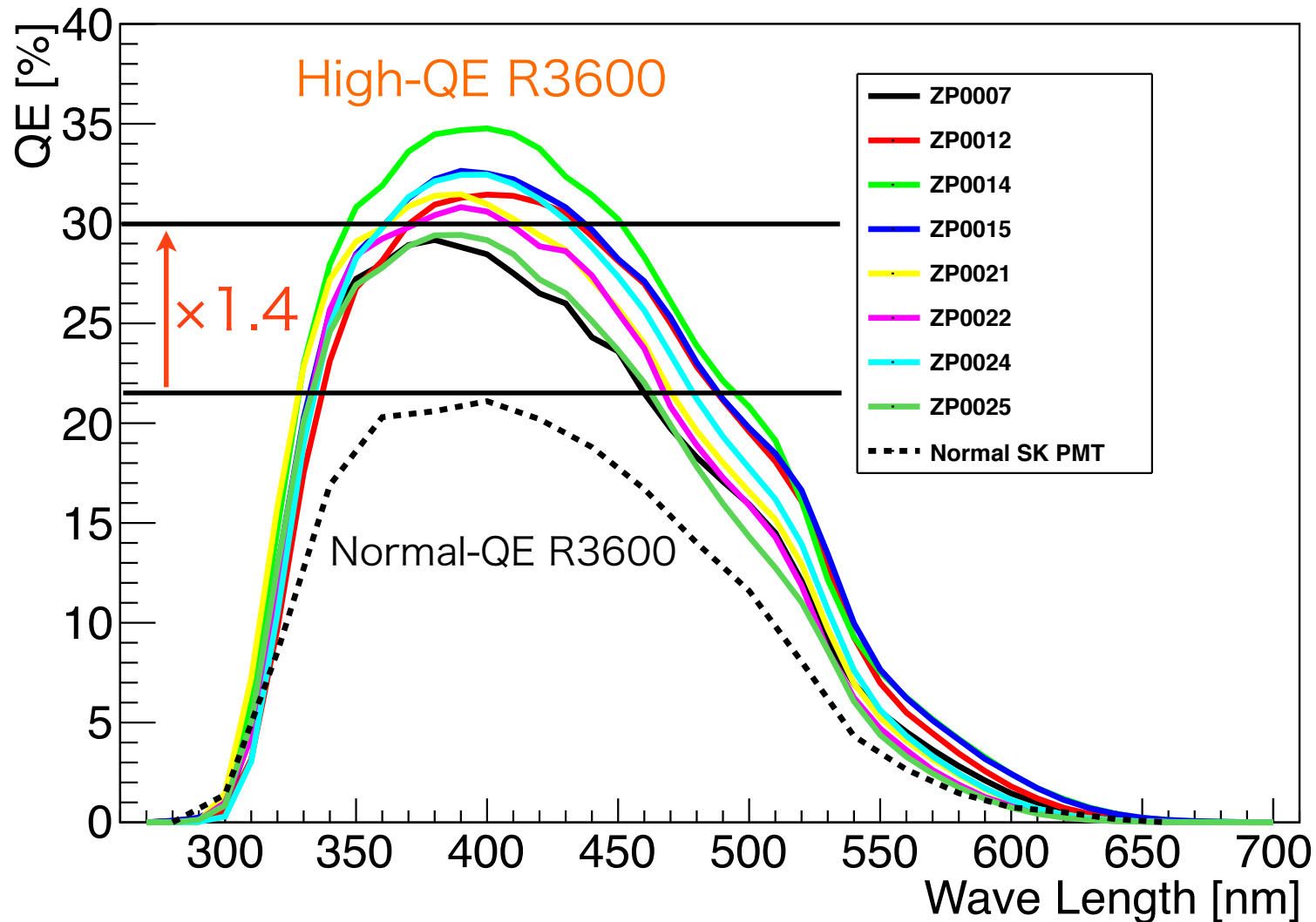
Established by SK
price known

in R&D
lower price expected

in R&D
lower price exp'd

Quantum Efficiency	22%	30%	30%
Collection Efficiency	80%	93%	95%

Higher QE achieved

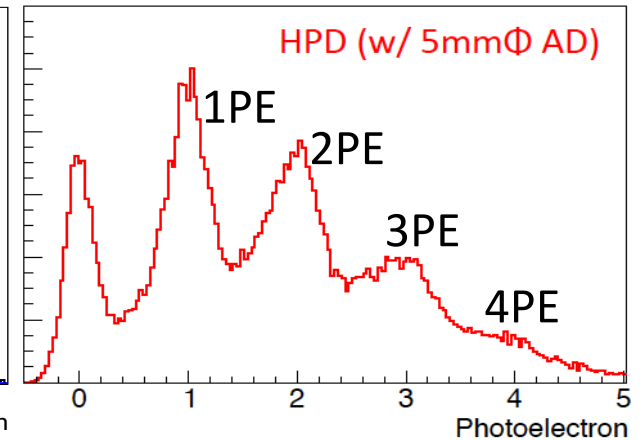
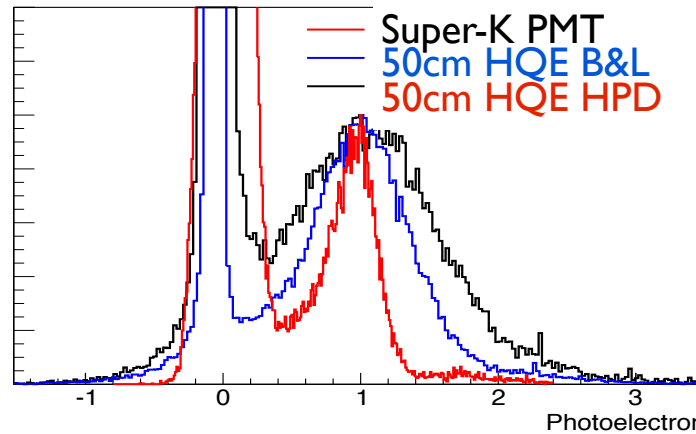
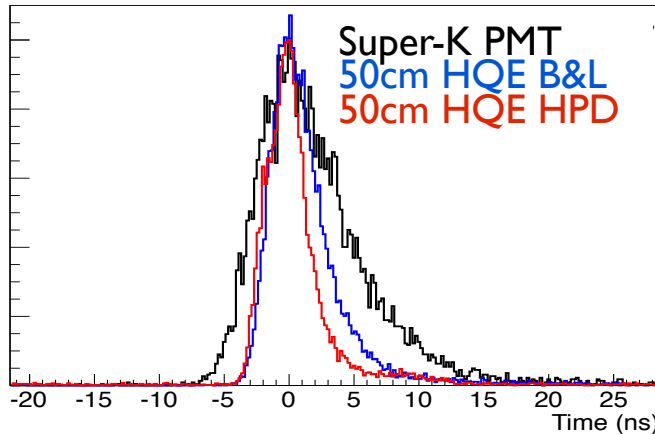


High Quantum Efficiency (QE) of ~30% has been achieved !
for 50cm B&L PMT and HPD

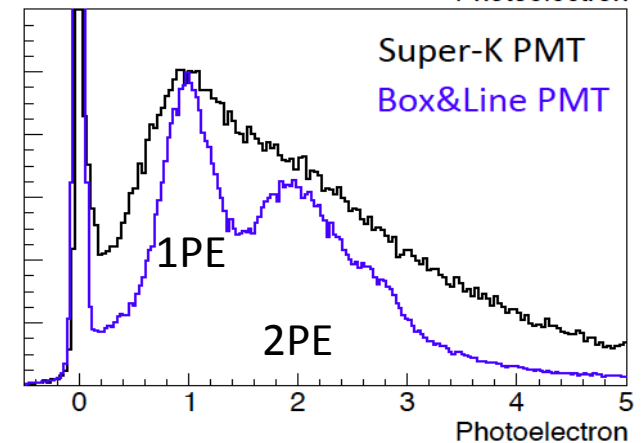
Photo-sensor performance

T distribution @ 1 p.e. 1 p.e. Q distribution

multi-p.e. Q dist.



	SK PMT	B&L PMT	HPD
1 p.e. Δt (ns)	2.1	1.1	1.4
1 p.e. $\Delta Q/Q$ (%)	53	35	16
Peak/Valley ratio	2.2	4.3	3.9



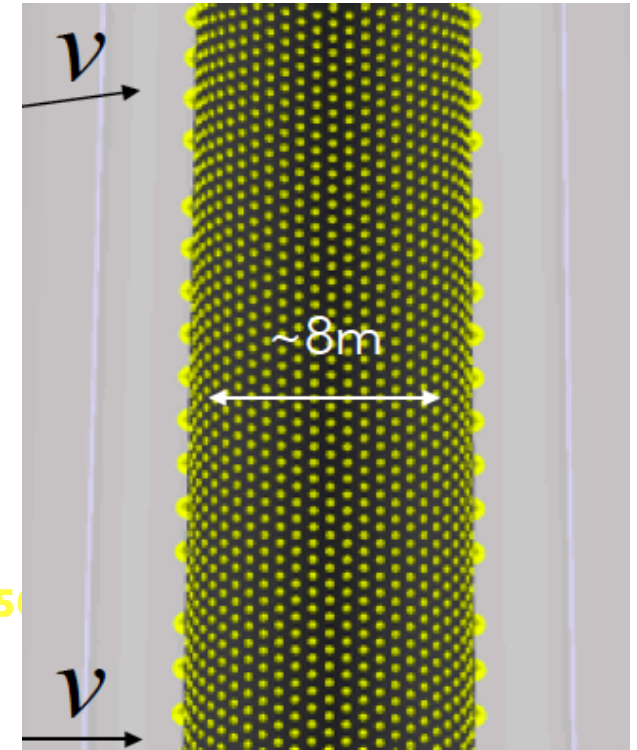
- Achieved better T&Q resolution
- Further tests are planned (test in water, long-term stability etc.)
- to be concluded by 2016

Near Detectors for J-PARC beam

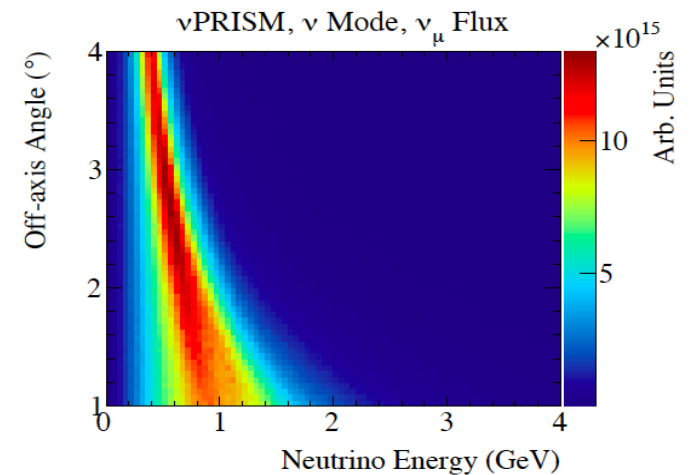
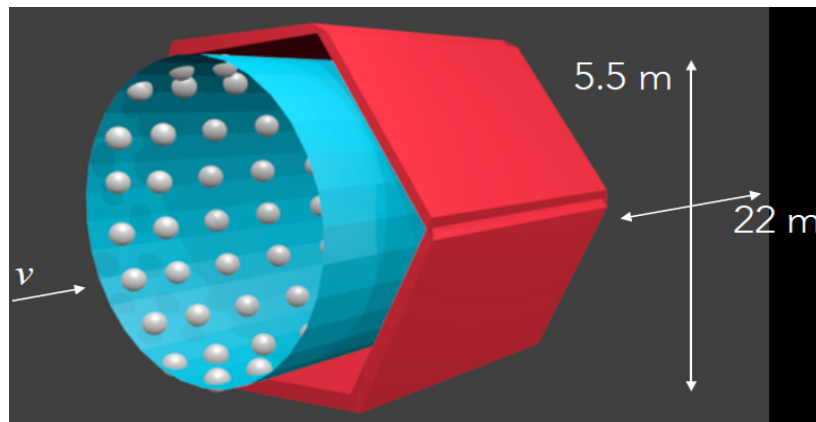
Conceptual design

- Oscillation study
 - Water target (same w/ the far detector, minimize nuclear uncertainty)
 - $NCT\pi^0$ BG measurement
 - beam ν_e BG
- Other physics
 - $\nu\mu$, νe interaction studies
 - Sterile ν searches

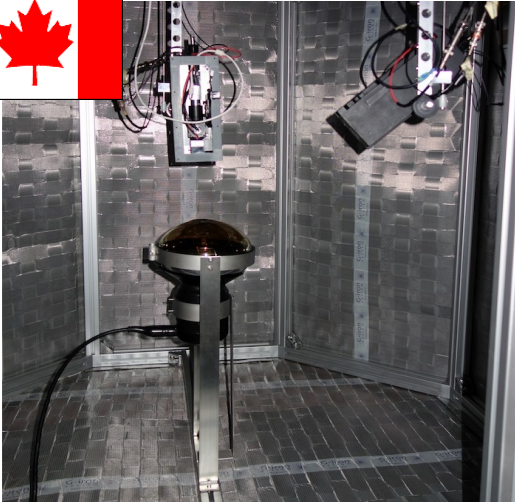
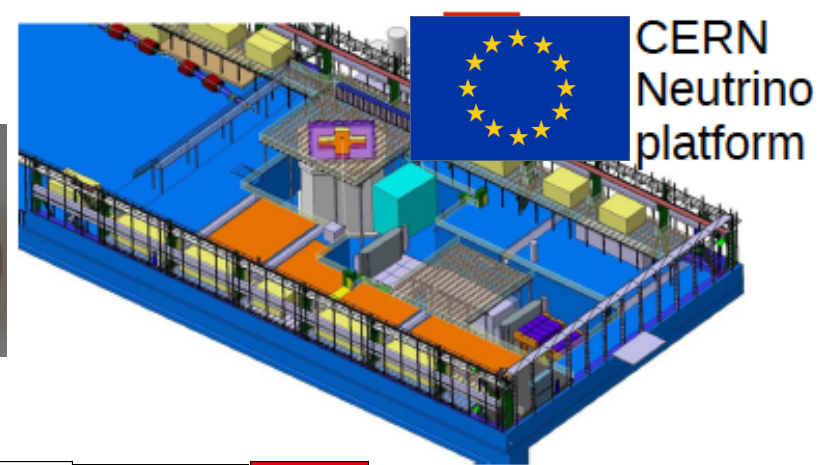
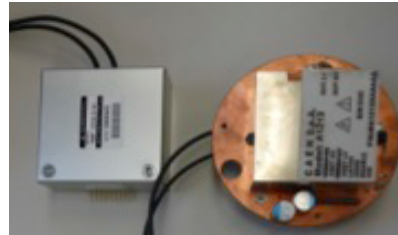
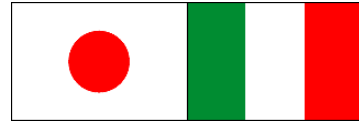
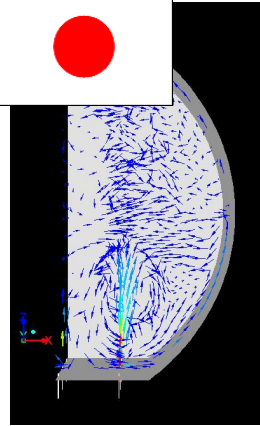
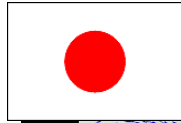
ν PRISM
50m tall WČ



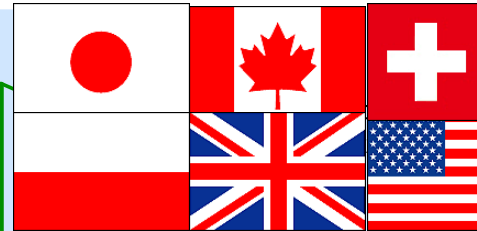
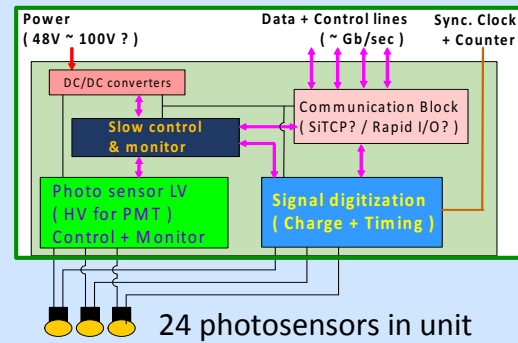
TITUS
WČ+MRD



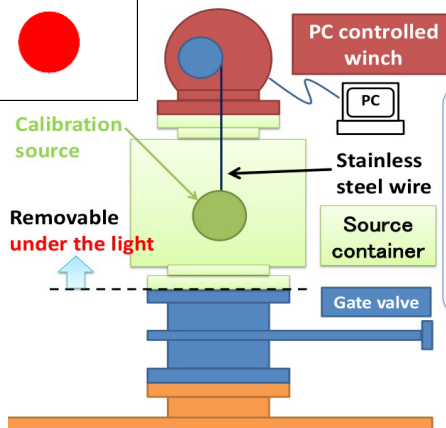
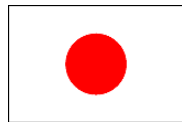
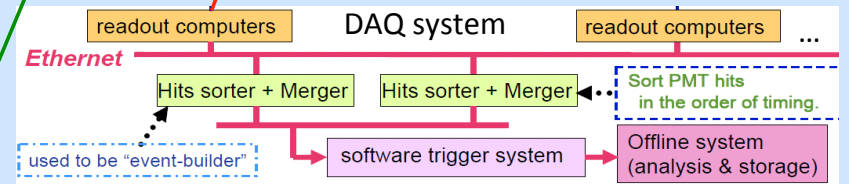
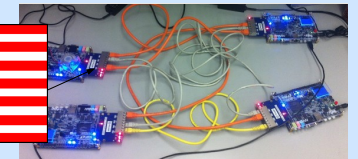
Worldwide R&D



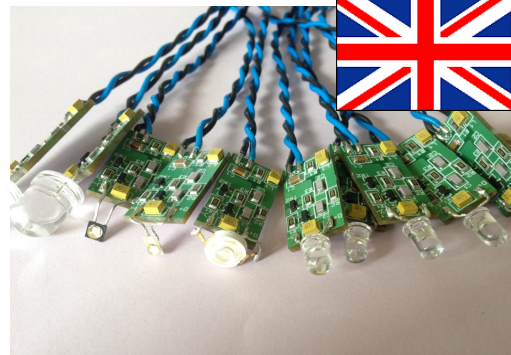
Elec. + HV modules in water



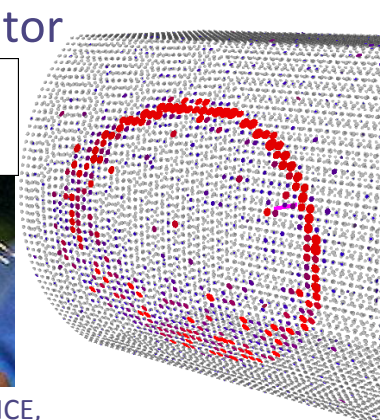
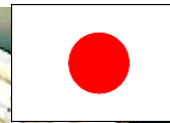
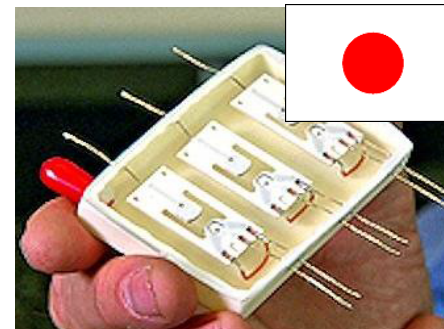
Trial for communication (RapidIO in FPGA boards)



LED



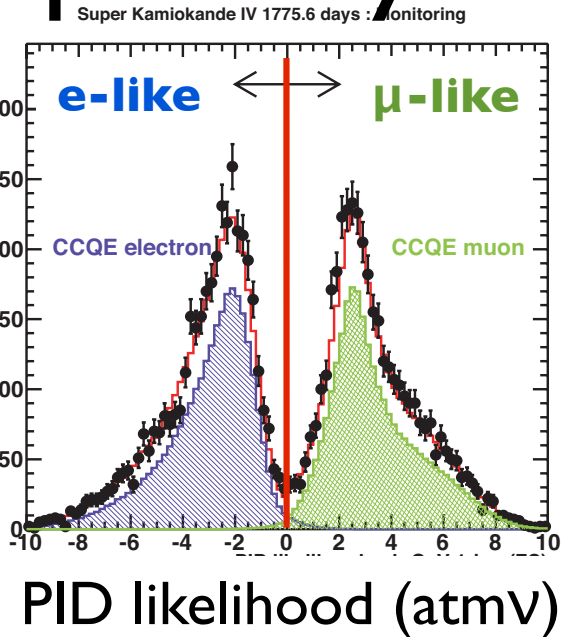
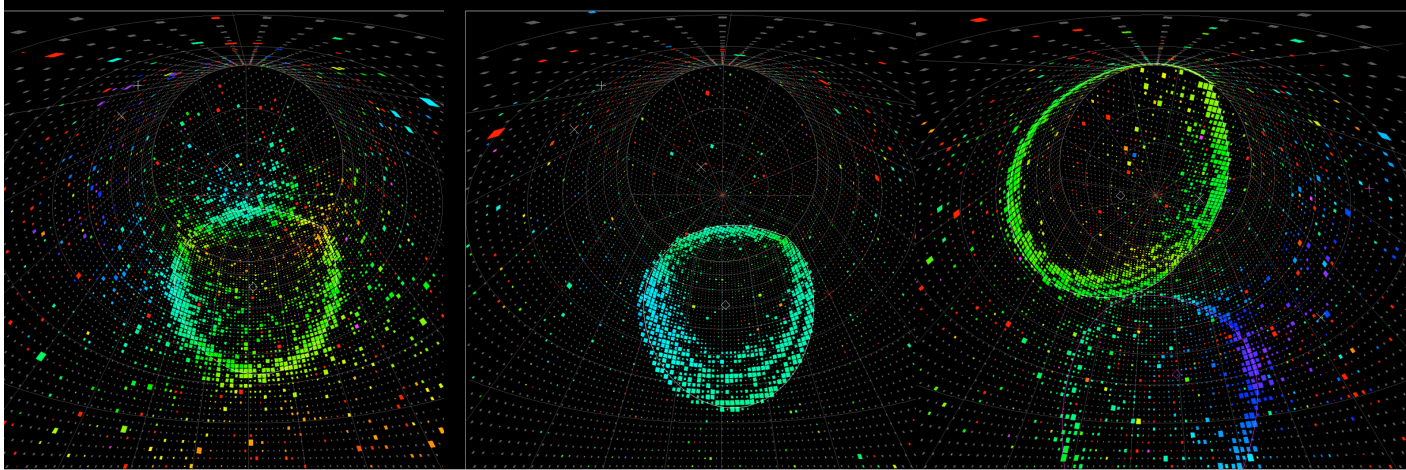
Compact neutron generator



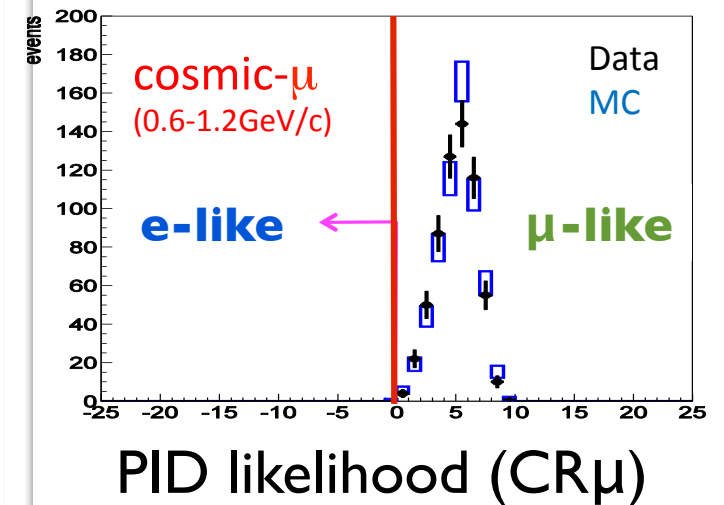
IEEE TRANSACTIONS ON PLASMA SCIENCE,
VOL. 40, NO. 9, SEPTEMBER 2012

Hyper-K Physics Potentials

Detector performance for p -decays



- High mass (1Mton scale, 20×Super-K)
- Good ring-imaging capability at ~ 1 GeV
 - atmospheric ν , proton decays, accelerator ν
- Excellent particle ID (e or μ) capability $> 99\%$
- Energy resolution for e and $\mu \sim 3\%$
- opportunity to improve more
- for proton decay search via $p \rightarrow e^+ \pi^0$
 - good $\sim 5\%$ invariant proton mass resolution
 - high 40% signal efficiency
 - 99.998% atmospheric ν BG rejection

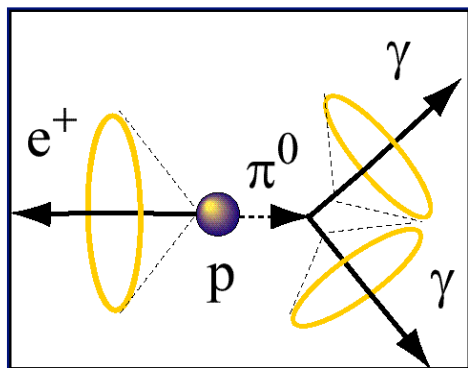
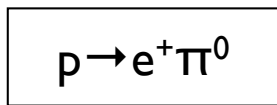


mis-PID:

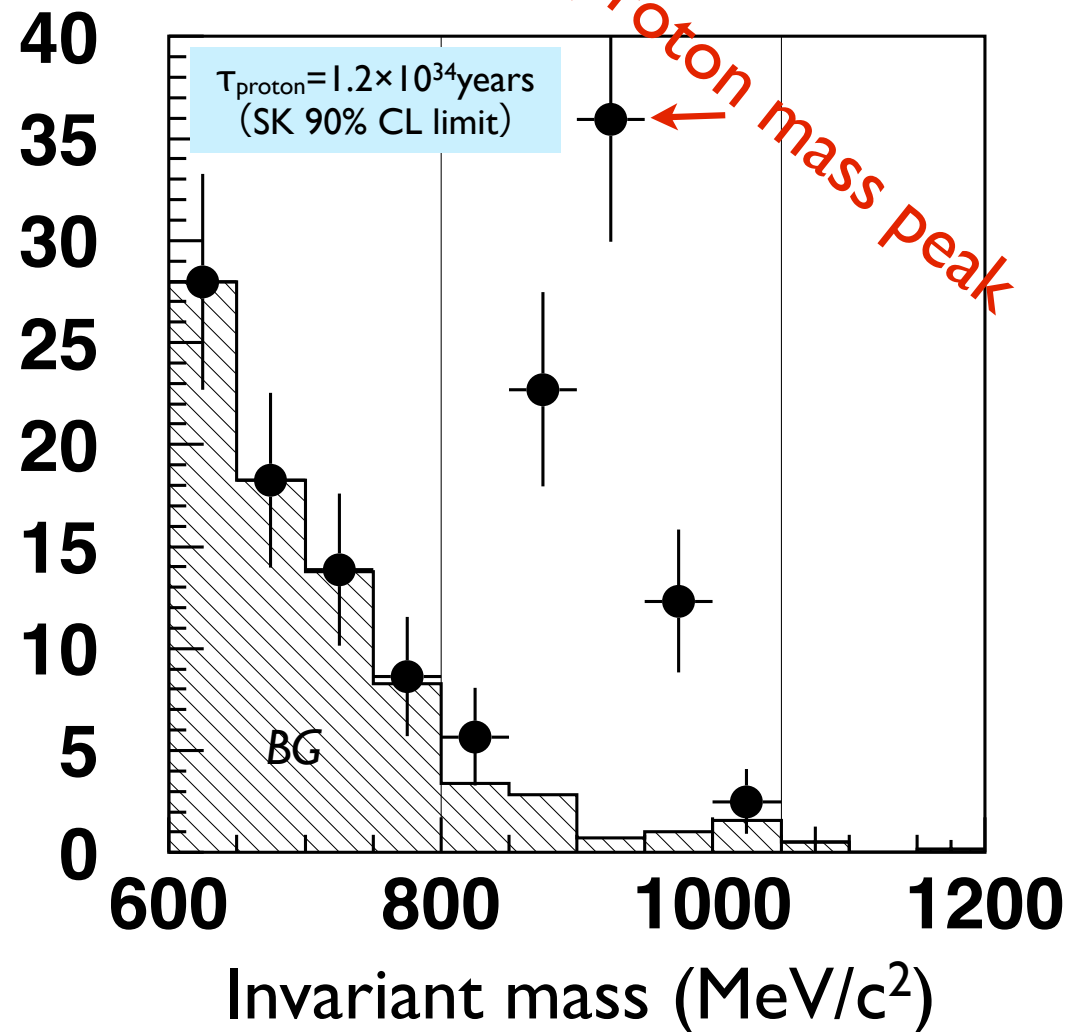
Data: 0.00 ± 0.16 (stat.)%

MC : 0.10 ± 0.10 (stat.)%

Discovery potential in Hyper-K

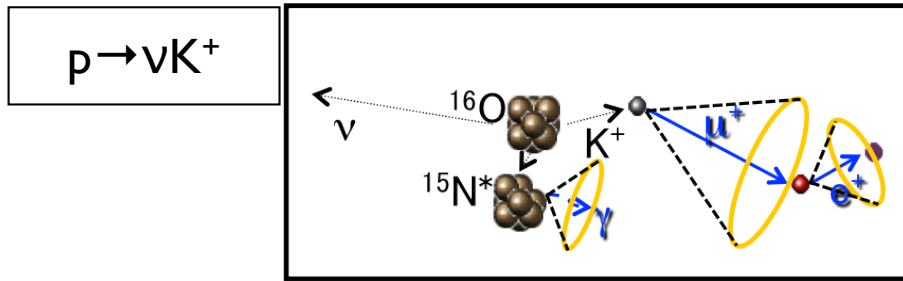


- ▶ Discovery reach (3σ)
 - ▶ $\tau(p \rightarrow e^+ \pi^0) \sim 5 \times 10^{34}$ years (HK 10yrs)
- ▶ Limit (90%CL)
 - ▶ $\tau(p \rightarrow e^+ \pi^0) > 1 \times 10^{35}$ years (HK 10yrs)

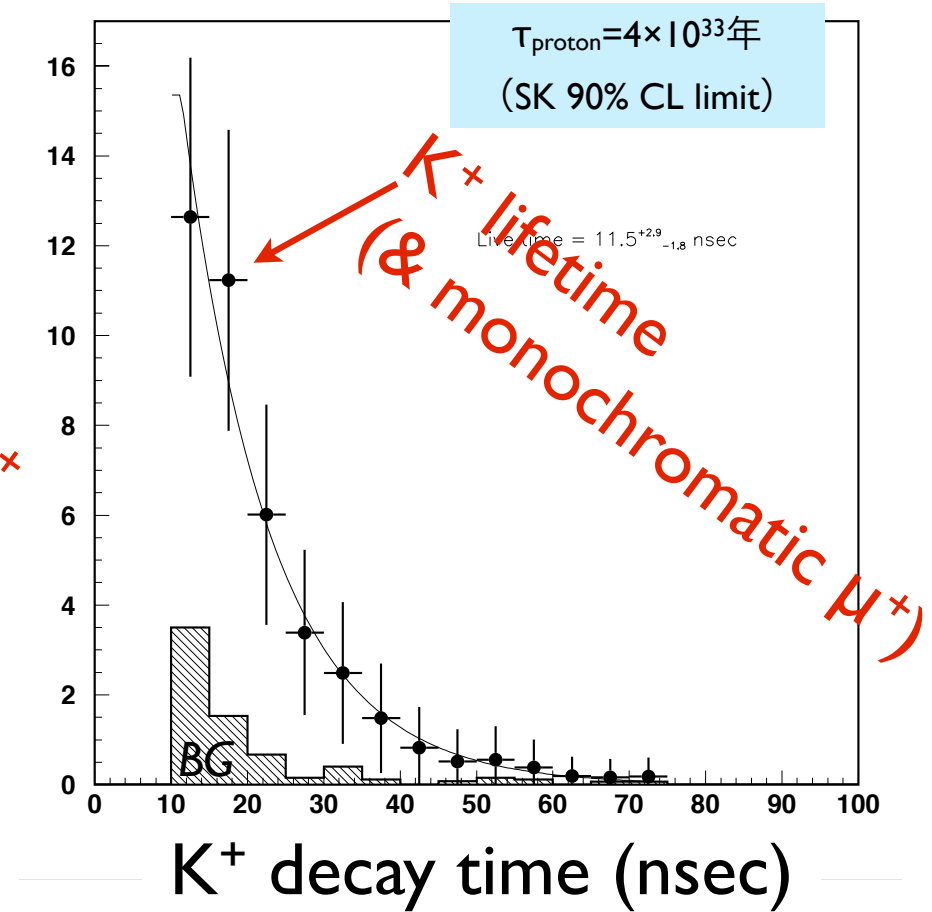
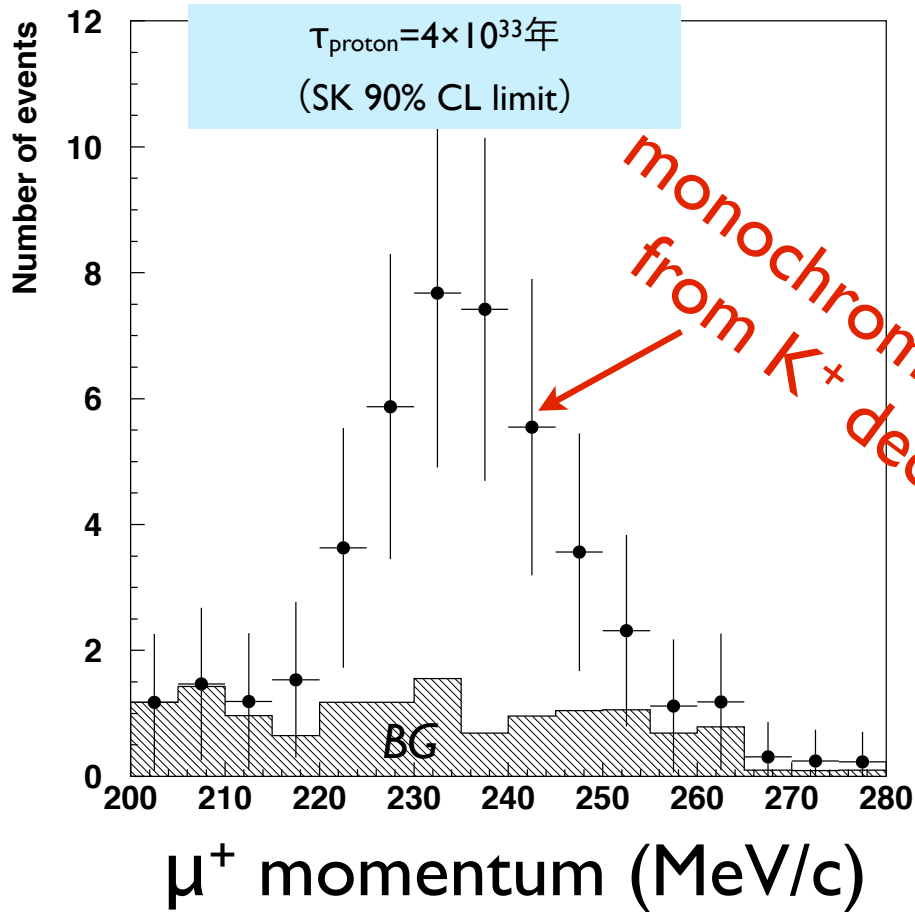


Only realistic proposal to reach the lifetime of 10^{35} years
for $p \rightarrow e^+ \pi^0$

Discovery potential (2)

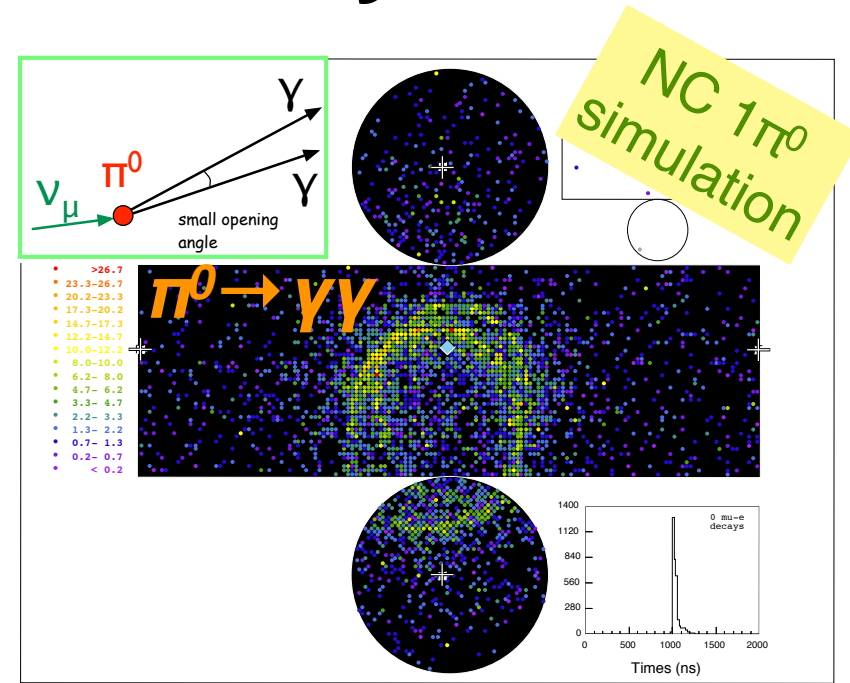
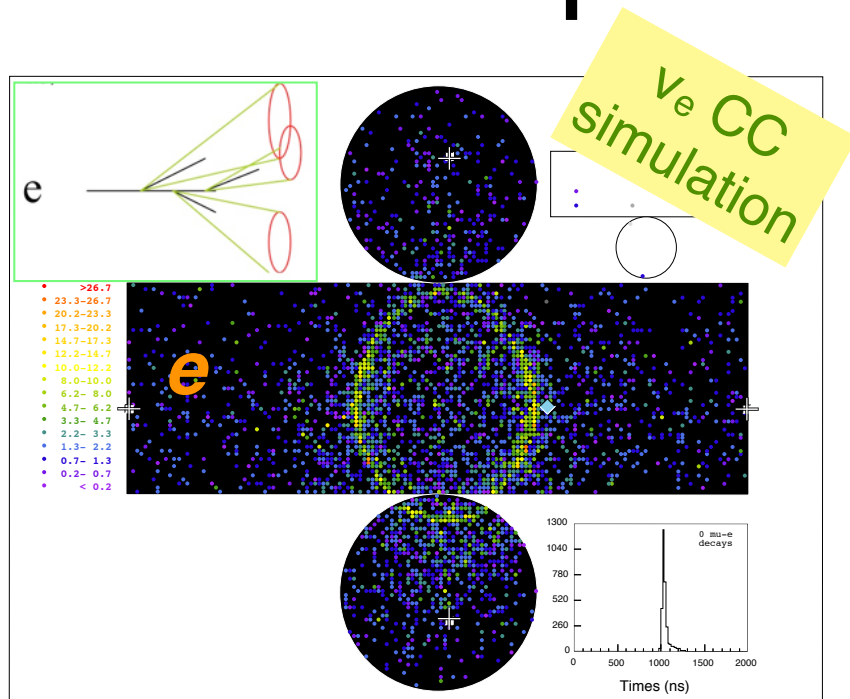


- ▶ Discovery reach (3σ)
 - ▶ $\tau(p \rightarrow \nu K^+) \sim 1 \times 10^{34}$ years (HK 10yrs)
- ▶ Limit (90%CL)
 - ▶ $\tau(p \rightarrow \nu K^+) > 3 \times 10^{34}$ years (HK 10yrs)



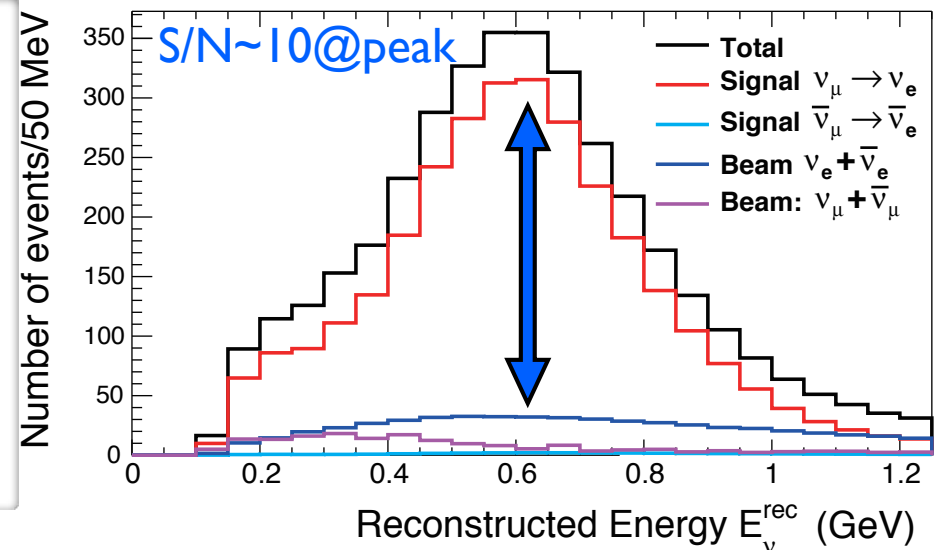
Experimental test on Supersymmetry

Detector performance for J-PAR Cv



Appearance ν mode

- For ν_e appearance in J-PARC ν_μ beam
 - high 60% ν_e signal efficiency
 - >99.9% ν_μ CC rejection, 99% NC π^0 rejection
- opportunity to improve more



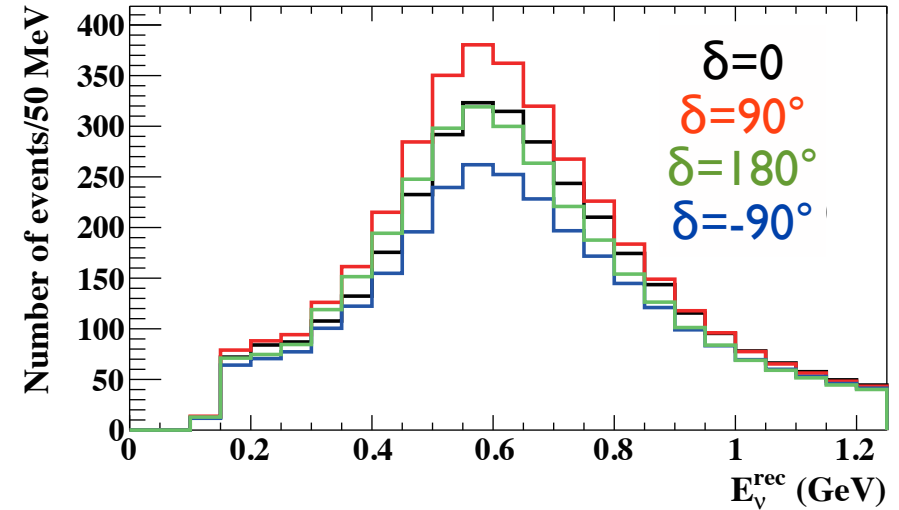
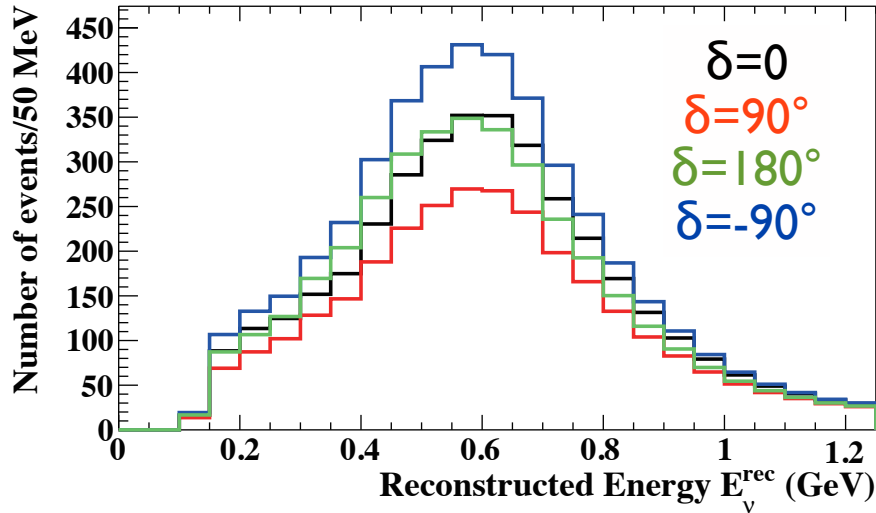
δ_{CP} dependence of observables

7.5MW $\times 10^7$ s (1.56×10^{22} POT)

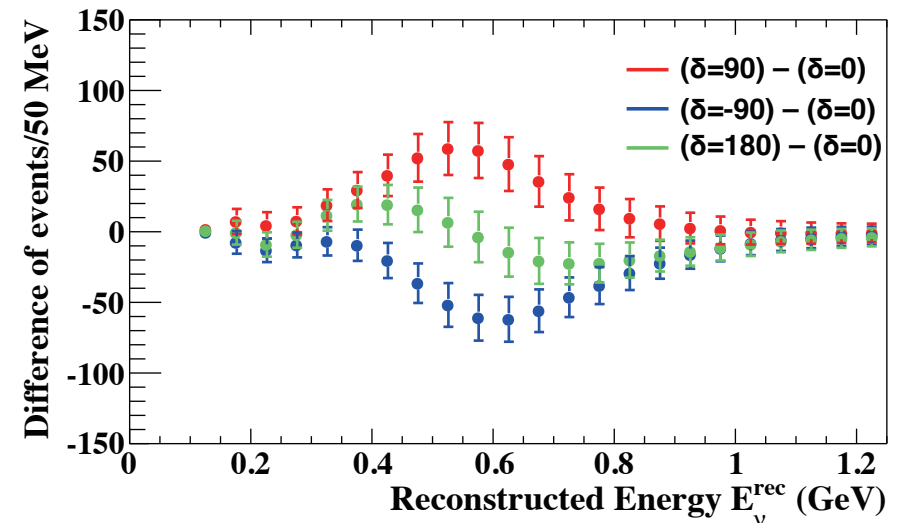
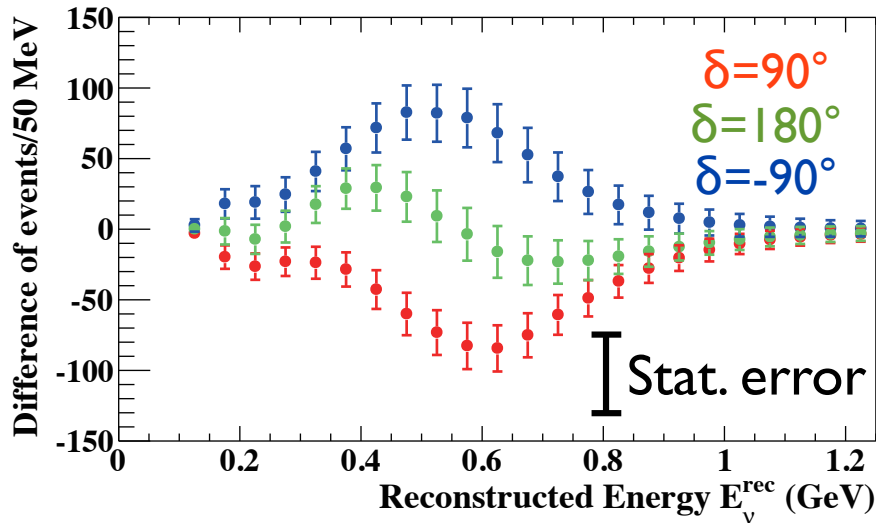
Neutrino mode: Appearance

Antineutrino mode: Appearance

ν_e candidates



Difference from $\delta=0$



Sensitive to all values of δ with numbers + shape

Assumed systematic uncertainties

Realistic estimation based on SK/T2K

- Beam flux + near detector constraint
 - Conservatively assumed to be the same
- Cross section uncertainties not constrained by ND
 - Nuclear difference removed assuming water measurements
- Far detector
 - Reduced by increased statistics of atmospheric ν control sample

Uncertainty on the expected number of events at Hyper-K (%)

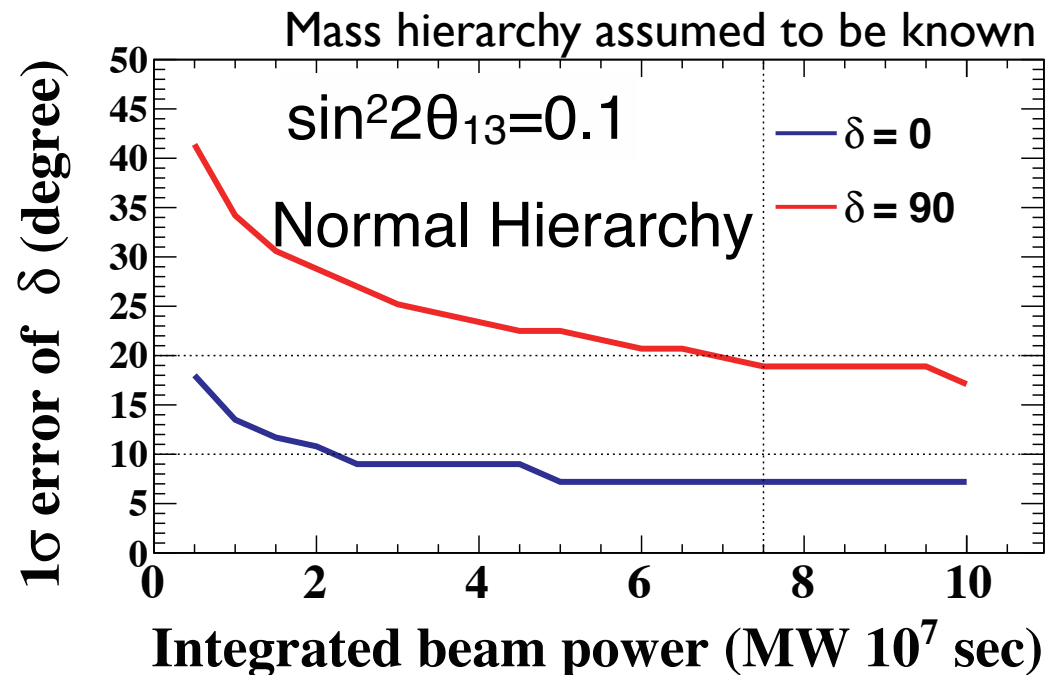
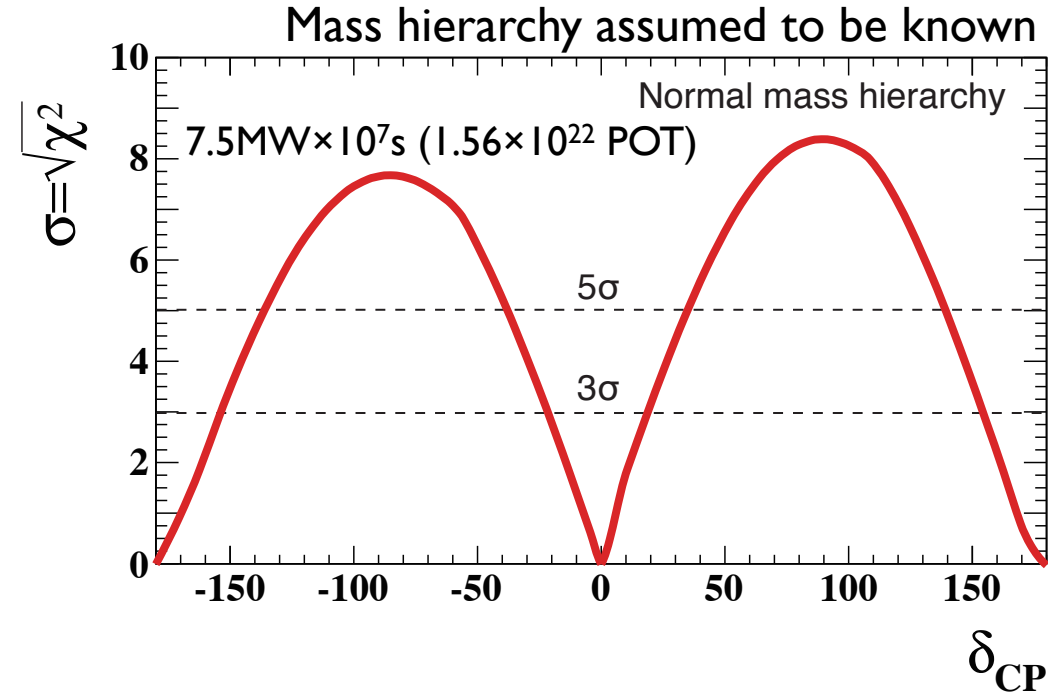
	ν mode		anti- ν mode		(T2K 2014)	
	νe	$\nu\mu$	$\bar{\nu}e$	$\bar{\nu}\mu$	νe	$\nu\mu$
Flux&ND	3.0	2.8	5.6	4.2	3.1	2.7
XSEC model	1.2	1.5	2.0	1.4	4.7	5.0
Far Det. +FSI	0.7	1.0	1.7	1.1	3.7	5.0
Total	3.3	3.3	6.2	4.5	6.8	7.6

- Further reduction by new near detectors under study

Sensitivity to CP violation

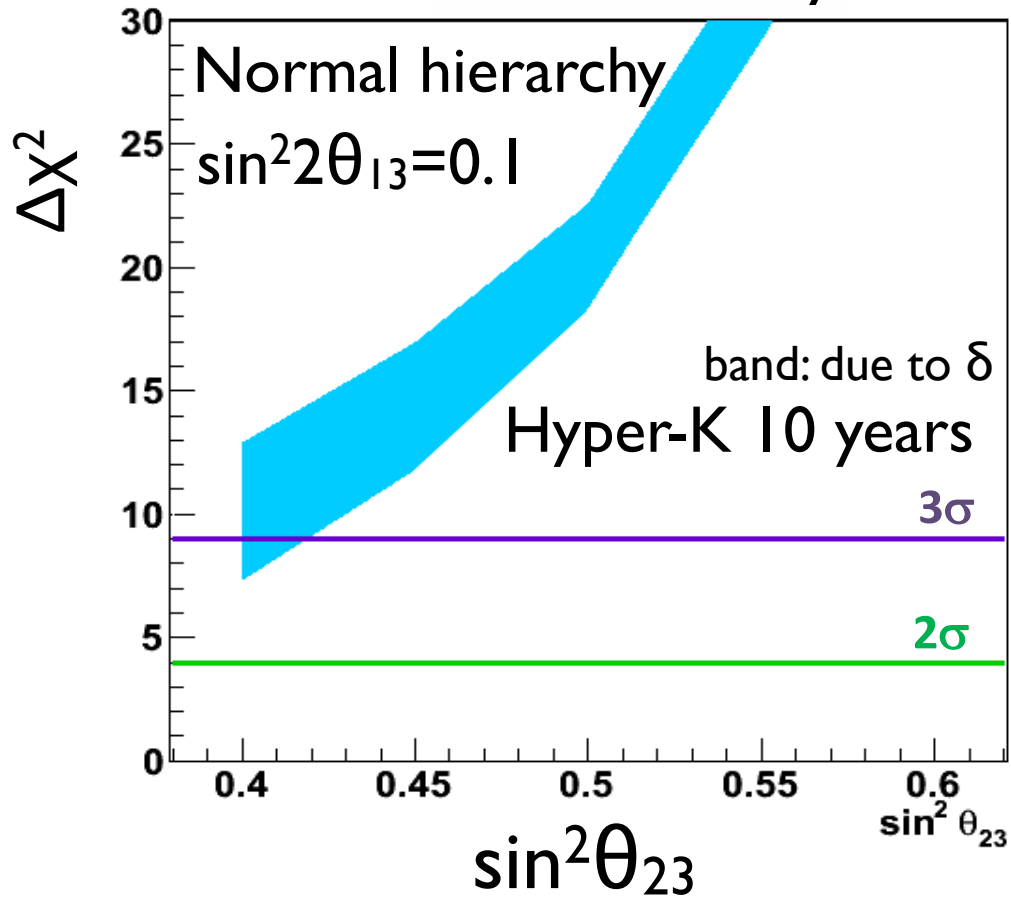
arXiv:1502.05199 and
submitted to PTEP

- Exclusion of $\sin\delta=0$
 - $>3\sigma$ for 76% of δ
 - $>5\sigma$ for 58% of δ
- 8° - 19° precision depending on the true value of δ

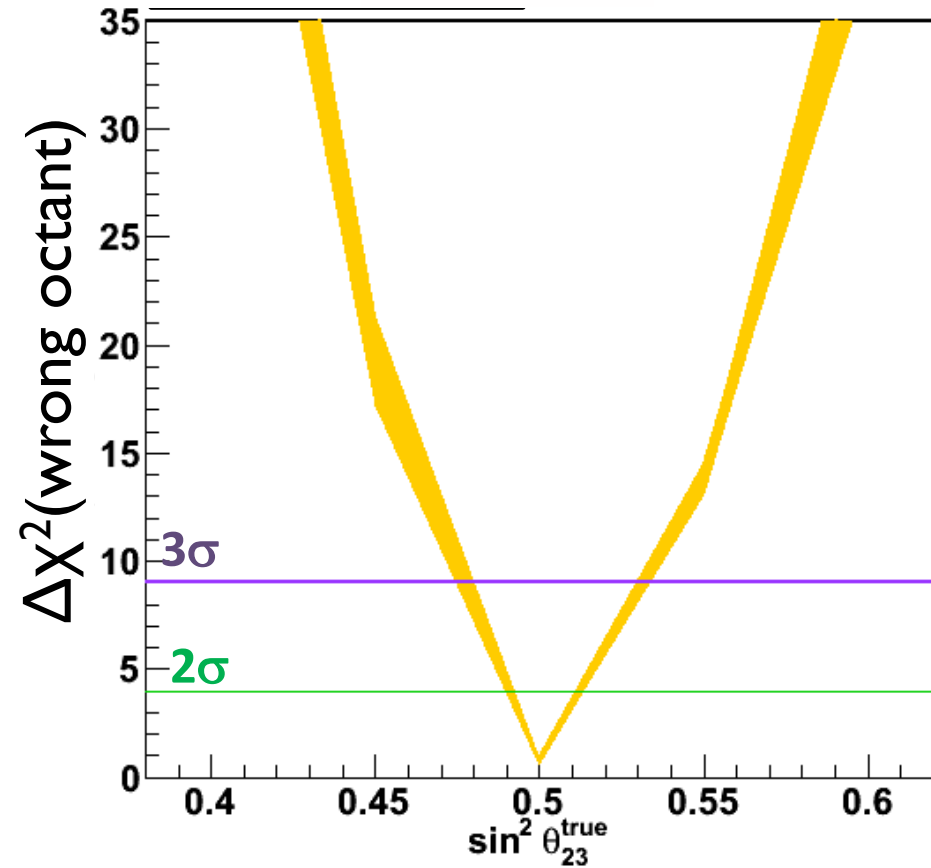


Atmospheric ν

Mass hierarchy



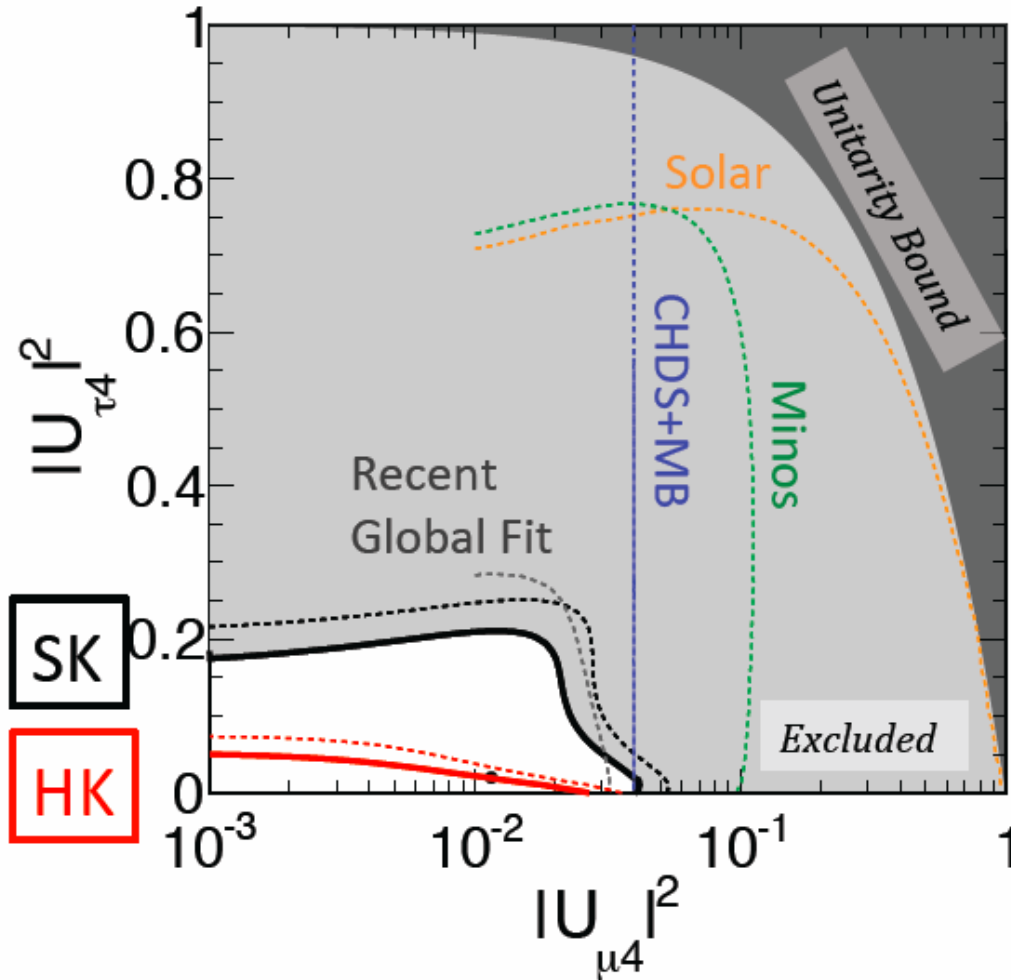
θ_{23} octant



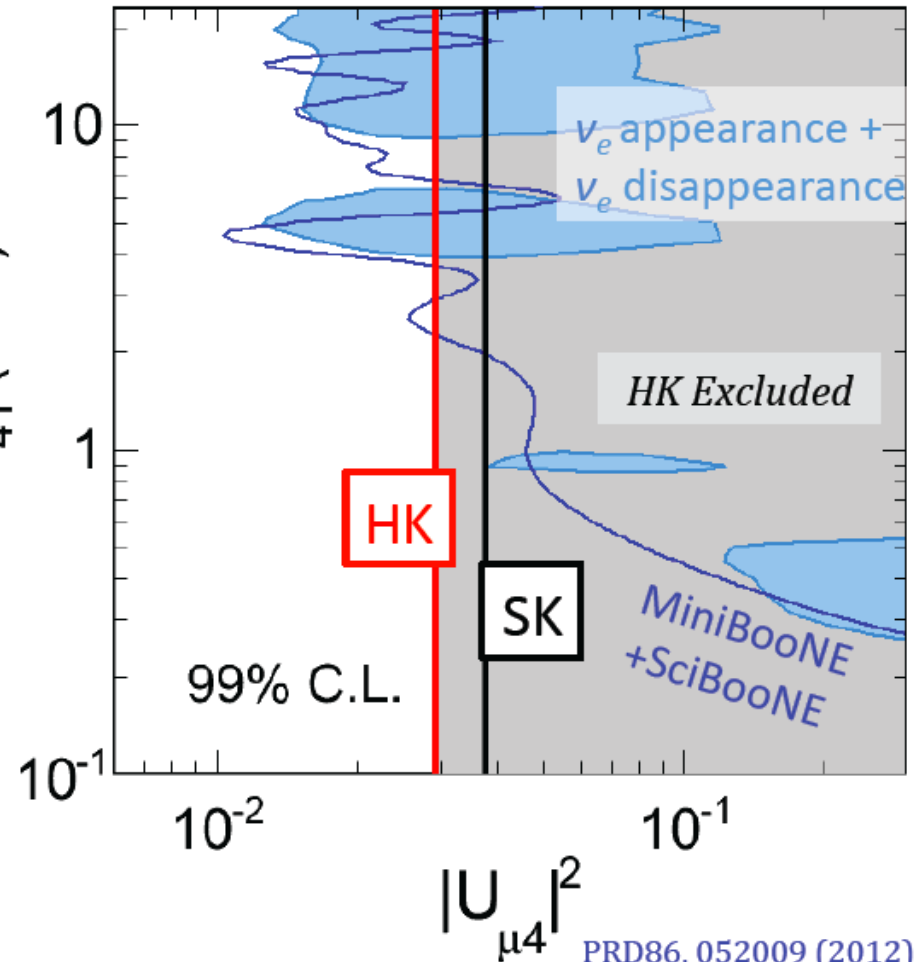
Complementary measurements to accelerator ν
Combined analysis of acc + atm ν will enhance capability

Test of Sterile ν by atmospheric ν

Look for **extra overall muon deficit** or **shape distortion**



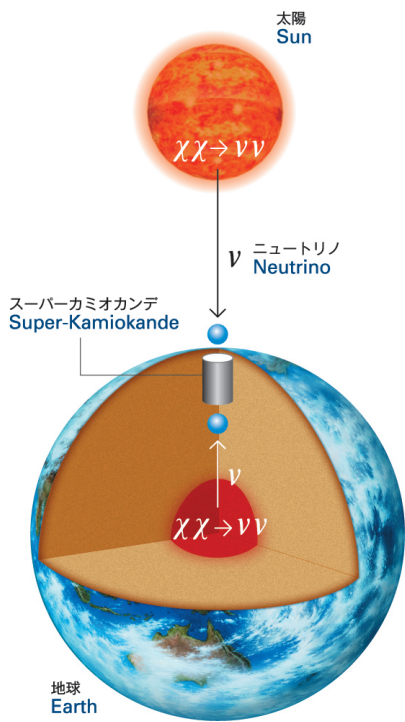
$|U_{\tau 4}|^2 < 0.066$ @99%CL
(0.164 in Super-K)



$|U_{\mu 4}|^2 < 0.029$ @99%CL
(0.038 in Super-K)

PRD86, 052009 (2012)
JHEP1305(2013)050

Complementary to other experiments

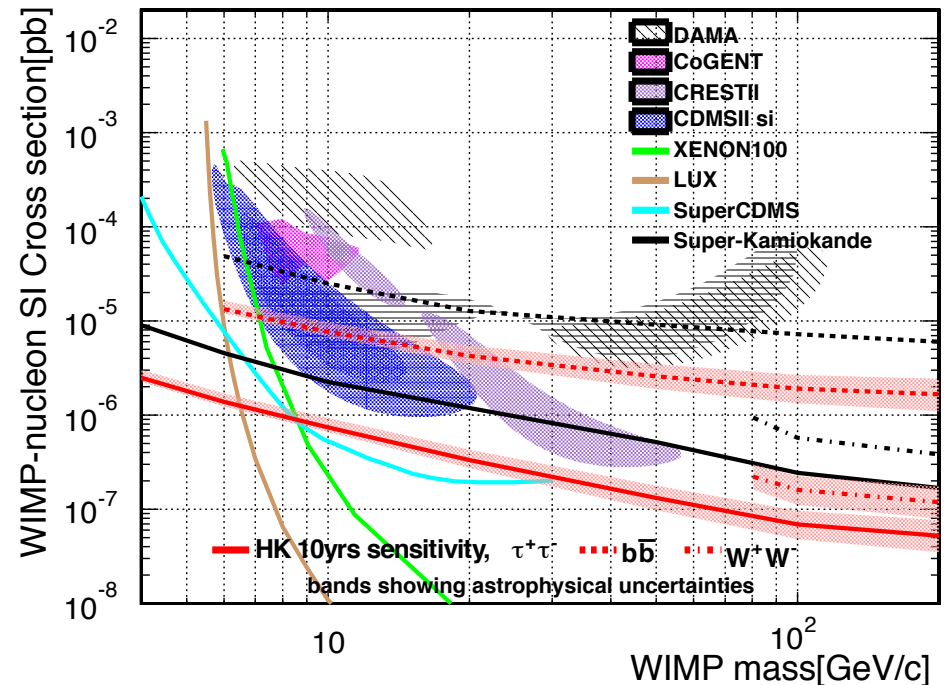
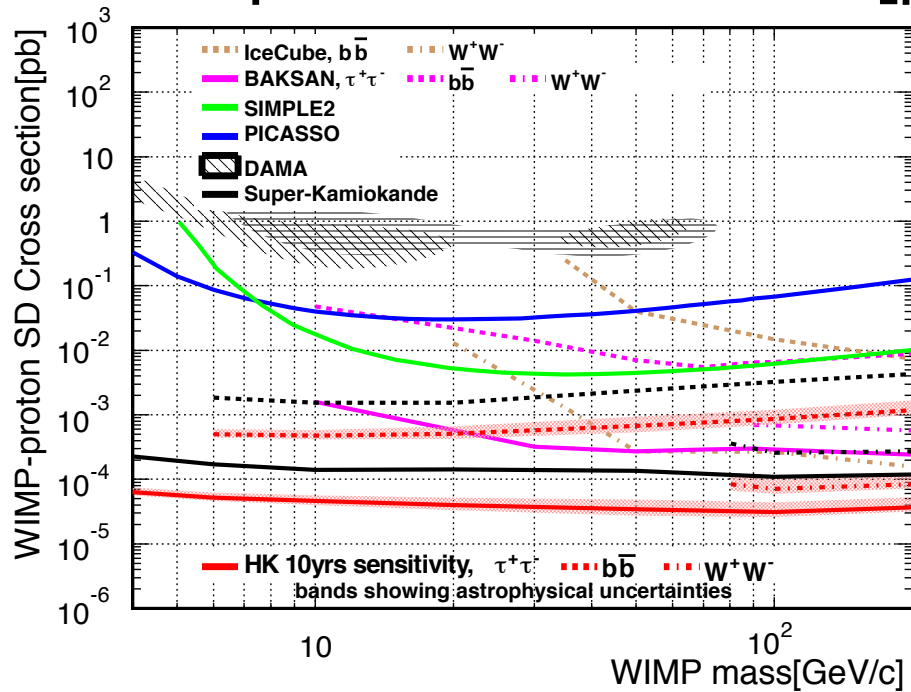


Search for ν 's induced by dark matters

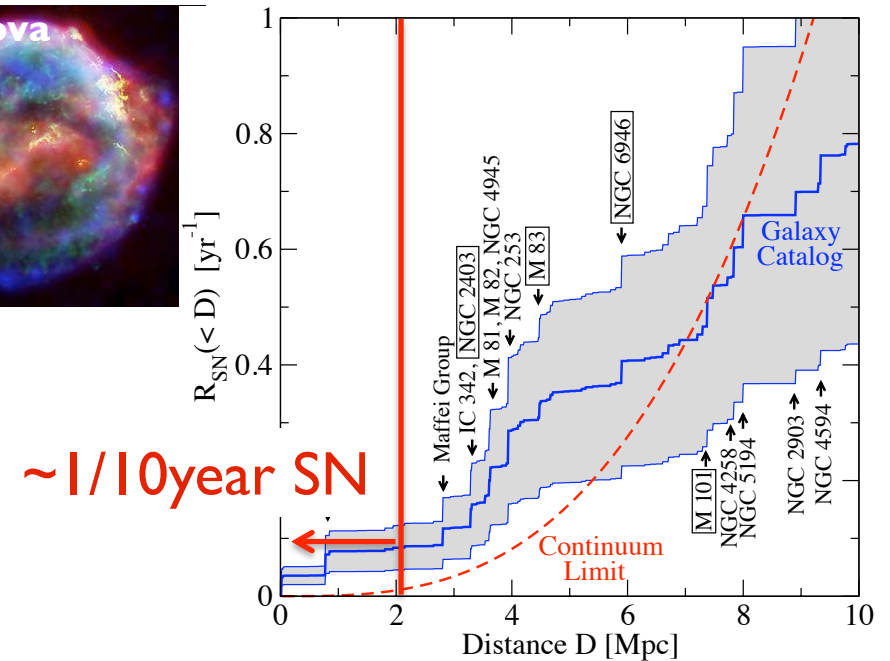
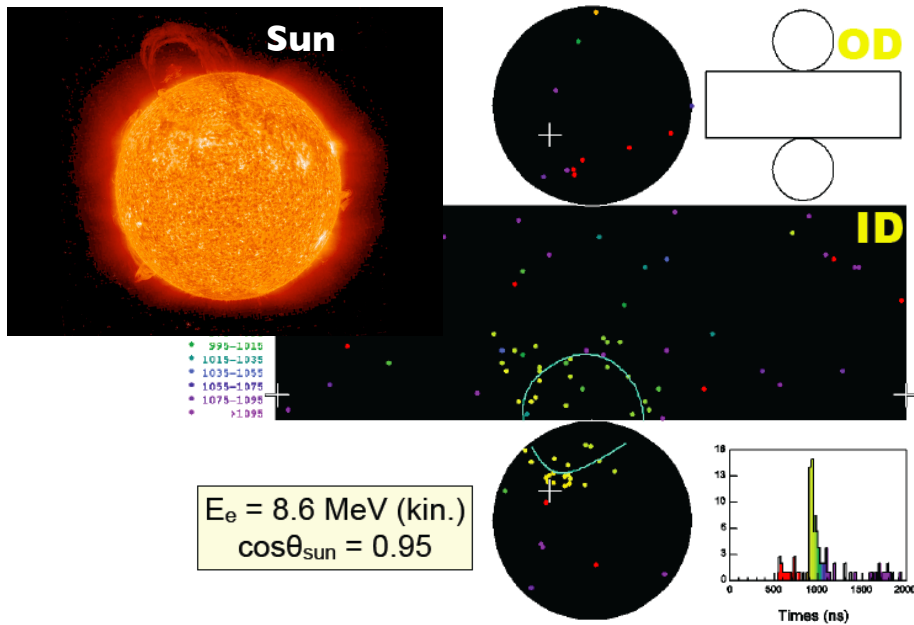
- provide complementary information w/ direct detection experiments
- Sensitive to low mass (GeV/c^2) WIMPs

Expected sensitivity for Solar WIMPs

WIMP-proton cross section [pb] WIMP-nucleon cross section [pb]



Det. performance for astrophysical ν



- Astrophysical neutrinos such as **galactic supernova, supernova in nearby galaxies, relic SN vs**, solar ν s
 - Energy threshold $\sim 5 \text{ MeV}$ by established techniques of water purification, triggering system, analysis algorithms
 - tagging capability of $\mu \rightarrow e \nu \nu$ and nuclear de-excitation γ in $p \rightarrow \nu K^+$
- energy scale stability $\sim 1\%$
- stable operation (small $< 1\%$ deadtime for Supernova observation)

Summary

- **Wide physics topics, many discovery potentials**
 - Proton decay discovery
 - CPV (76% of δ space at 3σ), δ precision of $<20^\circ$
 - SN bursts, relic SN ν , WIMP annihilation ν ...
- **Many good results in development works**
 - Cavity and support design
 - Plastic liner
 - 50cm high sensitivity photo-sensors
 - Many rooms to contribute
- **Boost promoting the project**
 - International proto-collaboration has been formed
 - Cooperation with KEK-IPNS/ICRR to develop the project
 - Design Report to be prepared in 2015
 - Open for new collaborators

Hyper-Kamiokande EU meeting@CERN

27-28 April 2015

- Meeting to discuss the European effort in Hyper-K
- Open to anyone who has interest in Hyper-K, or is planning to join Hyper-K, or is contributing
- <http://indico.cern.ch/e/ThirdEUHyperK>



Hyper-Kamiokande EU meeting

27-28 April 2015

CERN

Europe/Zurich timezone



Overview

Timetable

Registration

Participant List

Accommodation

- Meeting to discuss the European effort in the [Hyper-Kamiokande experiment](#).
- Open to anyone who has interests in Hyper-K, or is planning to join Hyper-K, or is contributing.
- Detailed information about your Country in Hyper-K can be discussed with your representatives in the [Hyper-Kamiokande International Board Representatives](#), its chair if no representatives are available yet or the [international Steering Committee chair](#).



Starts 27 Apr 2015 11:00
Ends 28 Apr 2015 18:00
Europe/Zurich



CERN
IT Amphitheatre