Status and perspectives of India based neutrino observatory (INO)



- Create experimental facility in the country where we can carry out front ranking experiments in the field of particle & astroparticle physics.
- Underground laboratory with ~1 km all-round rock cover accessed through a 2 km long tunnel. A large and several smaller caverns to facilitate many experimental programmes.
- Frontline neutrino issues e.g., mass parameters and other properties, will be explored in a manner complementary to ongoing efforts worldwide.
- The ICAL detector, with its charge identification ability, will be able to address questions about the neutrino mass ordering.
- Will support several other experiments when operational. Neutrino-less Double Beta Decay and Dark Matter Search experiments foreseen in the immediate future.
- Open to International participation

First atmospheric neutrino was reported from Kolar Gold Field (KGF) at a depth of 2.3km way back in 1965 by the TIFR-Osaka-Durham group

Physics Letters 18, (1965) 196, dated 15th Aug 1965

- We had a long tradition of carrying out experiments deep underground.
- KGF laboratory was the deepest underground laboratory during the period 1951-1992.
- KGF by TIFR-Osaka collaboration to look for proton decay.
- KGF mines closed its operation in 1992.

~30 muon /year/m²/sr at KGF, increased by a factor of ~100 at INO





INO Collaboration



Ahmadabad: Physical Research Laboratory **Aligarh:** Aligarh Muslim University **Allahabad: HRI Bhubaneswar:** IoP, Utkal University **Calicut:** University of Calicut **Chandigarh:** Panjab University **Chennai:** IIT-Madras, IMSc **Delhi:** University of Delhi Kalpakkam: IGCAR Kolkata: SINP, VECC, University of Calcutta Lucknow: Lucknow University Madurai: American College Mumbai: BARC, IIT-Bombay, TIFR, **CMEMS Mysore:** University of Mysore **Srinagar:** University of Kashmir Varanasi: Banaras Hindu University

Nearly 100 scientists from 23 research institutes & universities all over India One of the largest basic science projects in India in terms of man power and money

INO site : Bodi West Hills



9° 58' N, 77° 16' E, Pottipuram Village, Theni District Tamil Nadu State

Warm, low rainfall area, low humidity throughout the year

Physics goal for a large mass detector with charge identification capability

- Reconfirm atmospheric neutrino oscillation
- Improved measurement of oscillation parameters
- Search for potential matter effect in neutrino oscillation.
- Determining the sign of Δm^2_{23} using matter effect
- Measuring deviation from maximal mixing and octant of θ_{23}
- Probing Lorentz and CPT violation.
- Ultra high energy muons.

Physics goals to be complementary to other experiments worldwide. There is a growing realization that both atmospheric and accelerator experiments are needed to obtain best values of the parameters.

- •Accelerators : narrow range of L,E; high precision
- •Atmospheric : Large range of L,E, not-so-high precision

INO-ICAL detector



- Total number of RPCs in the ICAL = $3 \times 150 \times 64 = 28,800$
- Total gas volume = $28,800 \times 190$ cm $\times 190$ cm $\times 0.2$ cm ~ 200 m³
- Total surface area = 28,800 \times 1.9m \times 1.9m \sim 10⁵ m²
- Standard gas composition for the avalanche mode:
 - $Freon, R134a(C_2H_2F_4): Isobutane (i-C_4H_{10}): Sulphur \ Hexafluoride (SF_6):: 95.5: 4.3: 0.2$

INO-ICAL Detector

Parameter	ICAL	ICAL-Engineering module	
No. of modules	3	1	
Module dimensions	$16.2m \times 16m \times 14.5m$ $8m \times 8m \times 2m$		
Detector dimensions	$49m \times 16m \times 14.5m$	$8m \times 8m \times 2m$	
No. of layers	150	20	
Iron plate thickness	56mm	56mm	
Gap for RPC trays	40mm	40mm	
Magnetic field	1.3Tesla	1.3Tesla	
RPC dimensions	1950mm × 1910mm × 30mm 1950mm × 1910mm × 3		
Readout strip pitch	30mm 30mm		
No. of RPCs/Road/Layer	8	4	
No. of Roads/Layer/Module	8	4	
No. of RPC units/Layer	192	16	
No. of RPC units	28,800 (107,266m ²)	320 (1,192m ²)	
No. of readout strips	3,686,400	40,960	

A journey through RPC road



Now with 200 cm \times 200 cm Glass RPC in Avalanche mode

Prototype Magnet & RPC setup



500AT : ~1.5 Tesla

- 35 Ton prototype with 12 gaps to house 1m × 1m RPCs
- Long term operational experience
- Operate both glass & bakelite RPCs
- Reconstructed muon track with & without magnetic fields.
- Stability & suitability of LV, HV & electronics.
- Lab environmental condition.

Running Prototype RPC Stack at TIFR

Operational since last six years



Now this system is used for developing/testing of ICAL electronics Understood all aspects of RPC Gas gap and its signal, resolution, alignment/offset

Running Prototype RPC Stack at TIFR

0.8

0.6

0.4

0.2



Zenith angle of muon, measurement of cosmic muon flux as well as it angular dependency



Input to detector simulation and digitisation



Close loop gas recirculation and purification system



Magnet design



Soft Iron Plates
Copper Coils
Power Supply
Coil Cooling System

ICAL MAGNETQTY - 3Size :- $16(L) \times 16(W) \times 15(H)$ meterCOIL SIZE :- $15m \times 8m$ Ampere-Turn :- 80,000 ATMax. Design Value:- 100,000 ATMax. Power requirements in each magnet:-<150KW (Coils & Power Supply)</td>Each coil gap = $1300mm \times 80 mm$



20 × 8 × 8 Prototype at Madurai

Mini ICAL to build and commission at ICHEP within 2.5 year Only 20 layers of RPC









Electronics system for ICAL



Major elements

- Front-end board
- RPCDAQ board
- Segment Trigger Module
- Global Trigger Module
- Global Trigger Driver
- Tier1 Network Switch
- Tier2 Network Switch
- DAQ Server
- VLSI, FPGA and ASIC chips; high density connectors

ICAL Trigger Scheme



Detector simulation and event reconstruction



ICAL detector performance : Muon



• Fully contained event has better momentum resolution (upto 6% from pathlenght)

ICAL detector performance : Neutral



- Momentum resolution is worse in neutral energy, which is expected.
- Neutrino physics analysis was done using only information of muon as well as combined information of muon and other neutral particles.

Determination of neutrino mass hierarchy



- Hadronic information is added separately
- Larger value of $\sin^2\theta_{13}$, more confidence level/less year to have same CL

Octant of θ_{23}



- Identify the correct octant for lower value of $sin^2\theta_{23}$
- But certainly add values while combining result of NOvA and T2K

Precision of neutrino mass matrix



• Hierarchy sensitivity of ICAL excludes the wrong-hierarchy minimum for the CPV discovery

• ICAL will play an important role to determine these parameter more precisely

INO site at BodiHills



INO experimental hall



Under ground Lab : a Closer Look



Status of related activities@INO site

- 26.82.5 ha. of revenue land at Pottipuram village
- Civil Work water, electricity, approach road has started
- Building for RPC test setup along with hostel/guest house, workshop, hostel









Summary

- INO a well conceived project with pan-India presence of scientists.
- RPC as a choice for detector well suited and detector R&D matured for mass production.
- Electronics, DAQ, trigger etc, are in final stage of prototyping.
- Magnet : the 1/30 scale prototype work in progress, should be installed within 2 years.
- Simulation and physics studies in various stages of maturity, white paper is ready for submission.
- Expect construction of tunnel to begin by end of this year.
 - Tunnel and experimental hall : 3-4 years,
 - One module per year
- INO to be relevant to the world science in the coming decade.

Physics input in simulation

Parameter	Best-fit values	3σ ranges	Relative 1σ precision
			(in percent)
$\Delta m_{21}^2 \; (eV^2)$	7.5×10^{-5}	$[7.0, 8.1] \times 10^{-5}$	$\sigma(\Delta m_{21}^2) = 2.4\%$
$\Delta m_{31}^2 \; (eV^2)$	2.46×10^{-3} (NH)	$[2.32, 2.61] \times 10^{-3}$ (NH)	$\sigma(\Delta m_{31}^2) = 2.0\%$
$\Delta m_{32}^2 \; (eV^2)$	-2.45×10^{-3} (IH)	$-[2.59, 2.31] \times 10^{-3}$ (IH)	$\sigma(\Delta m_{32}^2) = 1.9\%$
$\sin^2 \theta_{12}$	0.3	[0.27, 0.34]	$\sigma(\sin^2\theta_{12}) = 4.4\%$
$\sin^2 \theta_{23}$	0.45 (NH), 0.58 (IH)	[0.38, 0.64]	$\sigma(\sin^2\theta_{23}) = 8.7\%$
$\sin^2 \theta_{13}$	0.022	[0.018, 0.025]	$\sigma(\sin^2\theta_{13}) = 5.3\%$
$\delta_{\rm CP}(^\circ)$	306	[0, 360]	_
$P_{m} = \sin^2 2\theta_{12} \sin^2 \theta_{02} \frac{\sin^2 \left[(1 - \tilde{A}) \Delta \right]}{2 \sqrt{2C - m}}$			

$$P_{\mu e} = \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2 \left[(1-A) \Delta \right]}{(1-\hat{A})^2}$$

+ $\alpha \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos (\Delta + \delta_{\rm CP}) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1-\hat{A})\Delta]}{(1-\hat{A})}$

$$\vdash \alpha^2 \sin^2 \theta_{12} \cos^2 \theta_{23} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2} + \mathcal{O}(\alpha^3, \alpha^2 s_{13}, \alpha s_{13}^2, s_{13}^3) .$$

$$\widehat{A} = \frac{2\sqrt{2G_F n_e}}{\Delta m_{31}^2}$$
$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E}$$
$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$$

- Error in systematic
 - Flux normalization : 20%
 - Cross section normalization : 20%
 - Energy dependent tilt factor 5%

Dark Matter search with ICAL

- Anomalous events in Kolar Gold Field (KGF) : Is it due to Dark Matter
- Many experiment tried to determine this
- Looked for $\Phi_{DM} \rightarrow l^+ l^- \ (l = e, \mu, \tau)$
- Added extra scintillator/RPC detector in the wall of the cavern to detect second particle (increase the acceptance ranges)



• Either observe it or put an upper limit on it.



Search of Magnetic Monopole

- No experimental evidence, but theoretically expected
- -dE/dx is very much different than charged(electric) particle
- Almost no background (large time gap between signal in consecutive layer, no bending)
- Need special trigger for this study





With larger year of ICAL data, either discover it or put more stringent limit than present one (MACRO)

In absence of signal in 10 years, upper limit of flux

