Magnetic oscillations in holographic liquids (and monopoles)

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Based on arXiv: 1501.06459 with S. Nowling, L. Thorlacius, and T. Zingg

and on work in progress with T. Alho, and R. Pourhasan

Gauge/Gravity Duality 2015, Firenze, April 15, 2015

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Outline

Part I

Overview and Motivations from CMT-side (I) and AdS-side (I)

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- Prelude: Magnetic oscillations in cond-matt
- Holographic liquids: our model
- Results
- Discussions and Open problems

Part II

- Motivations from CMT-side (II) and AdS-side (II)
- Monopole operators: our model
- Preliminary results
- Discussion and outlook

Motivation: study of compressible quantum states of matter

- Quantum: T = 0 or very low T ($T \ll \mu$)
- Compressible [Huijse-Sachdev'11][Sachdev'12]: continuum, translationally invariant, with a global charge Q; Ground state of $H \mu Q$ is characterised by $\langle Q \rangle$ a smooth non-constant function of μ

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- Known examples:
 - Solids (translational symmetry is broken);
 - 2 Superfluids (U(1) is broken);
 - 3 Fermi liquids
 - 4 Non-Fermi liquids

Zoom in: Non-Fermi liquids

Fermi liquids vs non-Fermi liquids: intro

Fermi liquids (FL)

- \blacksquare resistivity $\rho \sim {\rm T}^2$
- Fermi surfaces
- quasi-particles
- perturbative field theory

Non-Fermi liquids (NFL)

- resistivity $ho \sim T$
- Fermi surfaces
- NO quasi-particles
- strongly interacting

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- Left: Doping-temperature phase diagram of the hole-doped cuprate superconductors [Broun 108]
- Right: Magnetic field-Temperature in heavy-fermions metals [Gegenwart et al '08]
 - Quantum critical point behind NFL

A more specific perspective from CMT

Magnetic oscillations in finite density systems

- quantum oscillations in the magnetization as a function of 1/B, present in metals at low temperature T and strong magnetic field B
- standard tool to diagnostic and analyse Fermi surfaces
- in ordinary metals (Fermi liquids): Landau-Kosevich-Lifshitz formula
- in "exotic phases": some surprises:



Figure: phase diagram from under-doped to over-doped high-T_c SC [Sebastian-Harrison-Lozarich'12]

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In NFL: Magnetic oscillations and ARPES suggest gapless excitations and no quasi-particle descriptions [Sebastian-Harrison-Lozarich'12, Sebastian-Harrison-Lozarich'11]

Prelude: Magnetic oscillations in Fermi liquids

Apply a magnetic field to a gas of electrons in (3 + 1)-dimensions:

- Landau levels + Zeeman splitting:

$$\epsilon_{\ell}^2 = \mathbf{k}^2 + \mathbf{m}^2 + (2\ell + 1)\gamma \mathbf{B} \pm \gamma \mathbf{B}, \qquad \ell = 0, 1, \dots$$

- B is parallel along z-axis and $k \equiv k_z$
- γ is proportional to the gyromagnetic ratio
- What happens?
 - closed quantized orbits in the plane perpendicular to k
 - separation between two orbits: $\Delta_{\ell,\ell+1} \sim \mathbf{B}$
 - degeneracy: $\delta \sim \mathbf{B}$
- Recall: Free energy

$$\Omega \sim \textit{TB}\sum_{\ell} \int \textit{dk} \sum_{\sigma=\pm} \log\left(1 + \exp\frac{\mu - \epsilon_{\ell,\sigma}(\textit{k})}{\textit{T}}\right)$$

Increasing B: what happens? (2+1)-d example



 \blacksquare Increasing B: then $\Delta_{\ell,\ell+1}$ increases but also the degeneracy (area enclosed) δ increases

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Increasing B: what happens? (3+1)-d example



extra degeneracy in k: the crossing of Fermi surface is maximized when this degeneracy is minimal: extremal cross-sectional Fermi surface!

period is controlled by the (extremal) Fermi surface: A_F

Prelude: Magnetic oscillations in Fermi liquids: Summary

Increasing B: what happens?

- continuous jump of free energy (Ω) \iff oscillations of magnetization $M = -\frac{1}{V} \frac{\partial \Omega}{\partial \mu}$ (crossing of FS by highest occupied Landau level)
- conditions

Low T (thermal effects suppressed by oscillations) and strong B: T \lesssim B $<<\mu$ 2 (pure metal)

■ For FL: Landau-Kosevich-Lifshitz formula in 3 + 1-d

$$\mathcal{M}_{\rm osc} \sim \sqrt{B} \sum_{\ell} \frac{m^{\star} T/B}{\sqrt{\ell} \sinh\left(\#\,\ell m^{\star} T/B\right)} \cos\left(A_F \ell/B\right) \label{eq:Mosc}$$

For FL: Landau-Kosevich-Lifshitz formula in 2 + 1-d

$$\mathcal{M}_{\rm osc} \sim \frac{1}{B} \sum_{\ell} \frac{m^{\star} T/B}{\ell \sinh\left(\# m^{\star} T/B\right)} \sin\left(A_F \ell/B\right)$$

with m^{\star} the effective mass, $m^{\star} \sim \frac{\partial A_{\rm F}}{\partial \epsilon}$.

Questions, goals, and methods

Questions

- Can gauge/gravity duality give an alternative prediction and description for magnetic oscillations in strongly correlated systems?
- Can we attack the problem from a different point of view?
- Our goals
 - Magnetic oscillations in systems at finite density and strongly correlated via gauge/gravity duality
 - bottom-up approach without introducing probe fermions
- Our approach
 - Extension of electron star/cloud model [Hartnoll-Silverstein-Polchinski-Tong '10][Hartnoll-Tavanfar '10][VGMP-Nowling-Thorlacius-Zingg'10][Hartnoll-Petrov'10] See also previous related works: [deBoer-Papadodimas-Verlinde'10][Arsiwalla-deBoer-Papadodimas-Verlinde'10]
 - Previous work on magnetic effects in holographic metals within probe approx [Denef-Hartnoll-Sachdev'09][Hartnoll-Hofman'09][Hartnoll-Hofman-Tavanfar'10][Gubankova-Brill-Cubrovic-Schalm-Schijven-Zaanen'11] [Blake-Bolognesi-Tong-Wong'12] [Albash-Johnson-MacDonald'12][Gubankova-Brill-Cubrovic-Schalm-Schijven-Zaanen'13] [Hartnoll-Hofman-Vegh'11]

The bulk model: Electron cloud at B = 0

Action

$$S = S_{EH} + S_M + S_{ff}$$

= $\frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left(R - 2\Lambda\right) - \frac{1}{4e^2} \int d^4x \sqrt{-g} F_{\mu\nu} F^{\mu\nu} - \int d^4x \sqrt{-g} \mathcal{L}_{ff}$

where on-shell $\mathcal{L}_{fl} = p$ [Schutz' 70][Brown '93][Bombelli-Torrence '90][de Ritis-Lavorgna-Platania-Stornaiolo '85]

- Ingredients
 - degenerate charged perfect fluid of non-interacting fermions of mass m in 4-dimensions coupled to
 - Maxwell-Einstein theory with $\Lambda = -\frac{3}{l^2}$ (asymptotically AdS₄, with L the AdS radius)

- we search finite T configurations
- $\kappa^2 = 8\pi G_N$ is the gravitational coupling, e Maxwell coupling constant, here: $\frac{\kappa}{L} \ll 1$ (classical gravity regime)
- Dual: strongly correlated fermions at finite density $\mu = \lim_{r \to 0} A_t$ and at finite temperature $T = T_H$

The electron cloud at B = 0



 fermions are treated in a Thomas-Fermi approximation (or Tolman-Oppenheimer-Volkoff) [Hartnoll-Tavanfar'10]:

$$mL >> 1$$
, and $e^2 \sim \frac{\kappa}{L} \ll 1$ (1)

- actually it works beyond (1) [Allais-McGreevy'13][Gubankova-Brill-Cubrovic-Schalm-Schijven-Zaanen'13]
- fermions are characterised by a local chemical potential

$$\mu_{\rm loc}(\mathbf{r}) = \mathsf{A}_{\mu}(\mathbf{r})\mathbf{u}^{\mu}(\mathbf{r}) = \mathsf{A}_{t}(\mathbf{r})\mathbf{e}_{0}^{t}(\mathbf{r})$$

(static fluid: 4-velocity $u^{\mu} = (e_0^t, 0, 0, 0)$, with e_a^{μ} the tetrad)

The electron cloud at B = 0

 \blacksquare Eq. of state for the fermions in the rest frame: $\varepsilon^2=k^2+m^2$

density of states

$$n(arepsilon)=etaarepsilon\sqrt{arepsilon^2-m^2}$$
 electrons: $eta=rac{1}{\pi^2}$

If use the fluid is described by pressure p(r), energy density $\rho(r)$, and charge density $\sigma(r)$

$$\mathbf{p} = \mathbf{p} \left(\mu_{\text{loc}}(\mathbf{r}) \right), \quad \rho = \rho \left(\mu_{\text{loc}}(\mathbf{r}) \right), \quad \sigma = \sigma \left(\mu_{\text{loc}}(\mathbf{r}) \right)$$

for example

$$\sigma = \beta \int_{m}^{\mu_{\rm loc}(\mathbf{r})} \mathrm{d}\varepsilon \,\varepsilon \sqrt{\varepsilon^2 - m^2} = \frac{\beta}{3} \left(\mu_{\rm loc}^2(\mathbf{r}) - m^2 \right)^{3/2}$$

fluid is supported if

 $\mu_{\mathrm{loc}}(\mathbf{r}) \geq \mathbf{m}$

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The electron cloud at B = 0: Geometry

Ansatz

$$ds^2 = \frac{L^2}{r^2} \left(-\frac{\hat{c}(r)^2}{\hat{g}(r)^2} dt^2 + dx^2 + dy^2 + \hat{g}(r)^2 dr^2 \right) \,, \qquad A_t = \frac{eL}{\kappa} \frac{\hat{c}(r)\hat{a}(r)}{r\hat{g}(r)}$$

Then

$$\mu_{\rm loc}(\mathbf{r}) = \frac{\mathsf{e}}{\kappa} \hat{a}(\mathbf{r})$$

where ^ denotes dimensionless quantities

Solutions at finite temperature



Figure: Left: fluid profile at finite T [VGMP-Nowling-Thorlacius-Zingg'10]. Right: Cartoon of the geometry

Free energy

$$\mathcal{F} = \mathcal{E} - \mu \mathcal{Q} - s \mathsf{T}$$

The electron cloud geometry is the preferred solution for $T \le T_c$ compared to AdS charged black brane [VGMP-Nowling-Thorlacius-Zing'10].



3rd order phase transition [VGMP-Nowling-Thorlacius-Zingg'10][Hartnoll-Petrov'10]

The anisotropic electron cloud: $B \neq 0$

- Now: we add a magnetic field B supported by the black brane, B is constant and pointing along the radial direction
- Action

$$S = S_{EH} + S_M + S_{fl}$$

= $\frac{1}{2\kappa^2} \int d^4x \sqrt{-g} (R - 2\Lambda) - \frac{1}{4e^2} \int d^4x \sqrt{-g} F_{\mu\nu} F^{\mu\nu} - \int d^4x \sqrt{-g} \mathcal{L}_{fl}$

where on-shell $\mathcal{L}_{\mathbf{f}} = p$

- we search for finite T and B configurations
- charged perfect fluid of non-interacting fermions of mass m in 4-dimensions coupled to Maxwell-Einstein theory with $\Lambda=-\frac{3}{l^2}$
- Spin fluid models known since 70's [Schutz'70][Ray'72][Bailey-Israel'75][deOliveira-Salim'91][Brown'93][deOliveira-Salim'95]
- Dual: strongly correlated fermions at finite density $\mu = \lim_{r \to 0} A_t$, at finite temperature $T = T_H$ and at finite magnetic field $\mathcal{B} = \lim_{r \to 0} F_{xy}$.

The anisotropic electron cloud: $B \neq 0$

degenerate gas of electrons experience

$$\mu_{\rm loc}(\mathbf{r}) = \mathbf{A}_{\mu} \mathbf{u}^{\mu} , \qquad \mathbf{H}_{\rm loc}(\mathbf{r}) = \mathbf{e}_{1}^{[\mu} \mathbf{e}_{2}^{\nu]} \mathbf{F}_{\mu\nu}$$

Fluid equation of state: Landau levels and Zeeman splitting

$$\varepsilon_{\ell}^2 = \mathbf{m}^2 + \mathbf{k}^2 + (2\ell + 1)\gamma \mathbf{H}_{\rm loc} \pm \gamma \mathbf{H}_{\rm loc}$$

Density of states

$$\mathbf{n}(\varepsilon) = \beta \gamma \mathbf{H}_{\rm loc} \sum_{\ell \ge 0}^{\prime} \theta(\varepsilon^2 - \mathbf{m}^2 - 2\ell\gamma \mathbf{H}_{\rm loc}) \frac{\varepsilon}{\sqrt{\varepsilon^2 - \mathbf{m}^2 - 2\ell\gamma \mathbf{H}_{\rm loc}}} \,,$$

fluid is described by thermodynamic variables

 $\mathbf{p} = \mathbf{p}\left(\mu_{\rm loc}(\mathbf{r}), \mathbf{H}_{\rm loc}(\mathbf{r})\right),$

same for $\rho(\mathbf{r})$, $\sigma(\mathbf{r})$, and the magnetization $\eta(\mathbf{r})$. For example

$$\begin{aligned} \sigma(\mathbf{r}) &= \int_{\sqrt{m^2 + 2\ell\gamma H_{\rm loc}}}^{\mu_{\rm loc}} \mathbf{n}(\varepsilon) d\varepsilon \\ &= \beta\gamma \mathbf{H}_{\rm loc} \sum_{\ell \ge 0}' \theta(\mu_{\rm loc}^2 - \mathbf{m}^2 - 2\ell\gamma \mathbf{H}_{\rm loc}) \sqrt{\mu_{\rm loc}^2 - \mathbf{m}^2 - 2\ell\gamma \mathbf{H}_{\rm loc}} \,. \end{aligned}$$

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Ansatz

$$ds^{2} = \frac{L^{2}}{r^{2}} \left(-\frac{\hat{c}(r)^{2}}{\hat{g}(r)^{2}} dt^{2} + dx^{2} + dy^{2} + \hat{g}(r)^{2} dr^{2} \right)$$
$$A_{t} = \frac{eL}{\kappa} \frac{\hat{c}(r)\hat{a}(r)}{r\hat{g}(r)}, \qquad A_{y} = \frac{eL}{\kappa} \hat{B} x$$

Then

$$\mu_{\rm loc}(\mathbf{r}) = \frac{\mathbf{e}}{\kappa} \hat{a}(\mathbf{r}), \qquad \mathbf{H}_{\rm loc}(\mathbf{r}) = \frac{\mathbf{e}}{\kappa \mathbf{L}} \hat{\mathbf{B}} \mathbf{r}^2$$

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Solve for $\hat{c}(r)$, $\hat{g}(r)$, $\hat{a}(r)$.

Solution



fluid is supported when

$$\ell_{\rm filled} = \left\lfloor \frac{\mu_{\rm loc}^2 - \mathbf{m}^2}{2\gamma \mathcal{H}_{\rm loc}} \right\rfloor \ge 0$$

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Results: boundary magnetization \mathcal{M}



Figure: The labels denote temperatures $\hat{\mathcal{T}}/\hat{\mathcal{T}}_{c}=0.9~(a),~0.3~(b),~3\cdot10^{-3}~(c)$

- The magnetization of the electron cloud (solid lines) is lower than that of a dyonic black brane (dashed lines) with the same parameters
- overall amplitude of the magnetization \mathcal{M} is linear in \mathcal{B} : \neq from Landau-Fermi theory and \neq [Hartnoll-Hofman-Tavanfar'10]][Blake-Bolognesi-Tong-Wong'12]: we have back-reaction now!
- ≠ Friedel oscillations, which were not observed in the electron cloud [VGMP-Nowling-Thorlacius-Zingg'11]: continuum of bulk Fermi surfaces





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independent of the temperature T as expected at low T

The anisotropic electron cloud: Thermodynamics

Free energy and thermodynamics relation

$$\mathcal{F} = \mathcal{E} - \mathbf{s}\mathcal{T} - \mu \mathcal{Q}, \qquad \frac{3}{2}\mathcal{E} = \mathbf{s}\mathcal{T} + \mu \mathcal{Q} - \mathcal{M}\mathcal{B}$$

 free energy: 2nd order phase transition between anisotropic electron cloud and dyonic black hole



Compare with B = 0: third order phase transition [VGMP-Nowling-Thorlacius-Zingg'10][Hartnoll-Petrov'10] Expected to be first order taking into account quantum corrections as in B = 0 [Medvedyeva-Gubankova-Cubrovic-Schalm-Zaanen'13] Phase transitions can be studied numerically and analytically close to the critical point [Hartnoll-Petrov'10]

$$\Delta \mathcal{F} \sim \int \delta \mathbf{p} \sim \delta \mu^{3/2} \,\Delta \mathbf{r} \sim (\Delta \mathbf{r})^4 \sim \Delta \mathcal{T}^2$$

- Analytically: They match with numerics up to the third digit
 - vs \mathcal{T} (at fixed \mathcal{B})

$$\frac{\Delta \mathcal{F}}{\mu^3} = -\hat{f}\left(\mathbf{m}, \frac{\mathbf{B}}{\mu^2}, \star\right) \Delta \mathcal{T}^2$$

- vs \mathcal{B} (fixed \mathcal{T})

$$\frac{\Delta \mathcal{F}}{\mu^3} = -\hat{\mathbf{g}}\left(\mathbf{m}, \frac{\mathbf{T}}{\mu}, \star\right) \Delta \mathcal{B}^2$$

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Summary

- a holographic model for a 2+1 dimensional system of strongly correlated electrons in a magnetic field
- The system shows magnetic oscillations dominated by a single sharp Fermi surface
- The oscillation amplitude has a non-Fermi liquid character and it is different from earlier probe fermion computations
- The model: 3+1 dimensional bulk fermions treated in a Thomas Fermi approx in an asymptotically AdS dyonic black brane background with gravitational and electromagnetic back-reaction
- our results confirmed later also by [Carney-Edalati'15]
- Oulook
 - beyond Thomas-Fermi approx: WKB along the lines of [Medvedyeva-Gubankova-Cubrovic-Schalm-Zaanen'13][Carney-Edalati'15] Or approach of [Allais-McGreevy'13]

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- Other systems
- Thermalisation effects