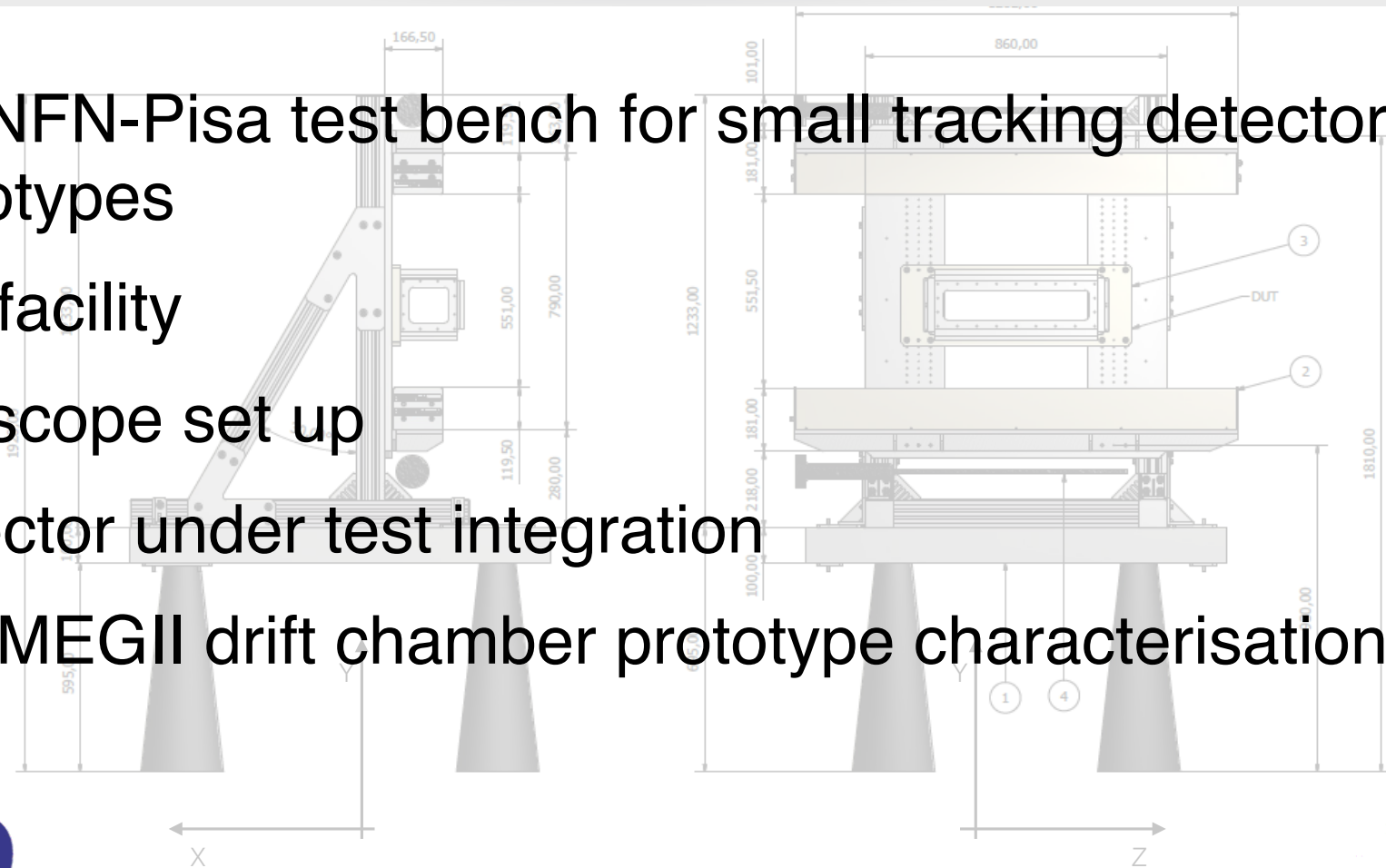


A silicon based cosmic ray telescope as an external tracker to measure detector performance

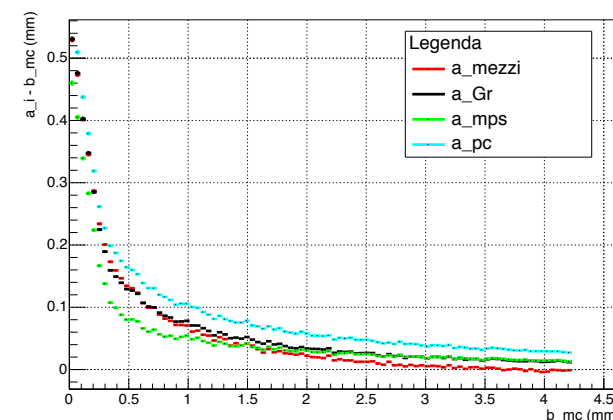
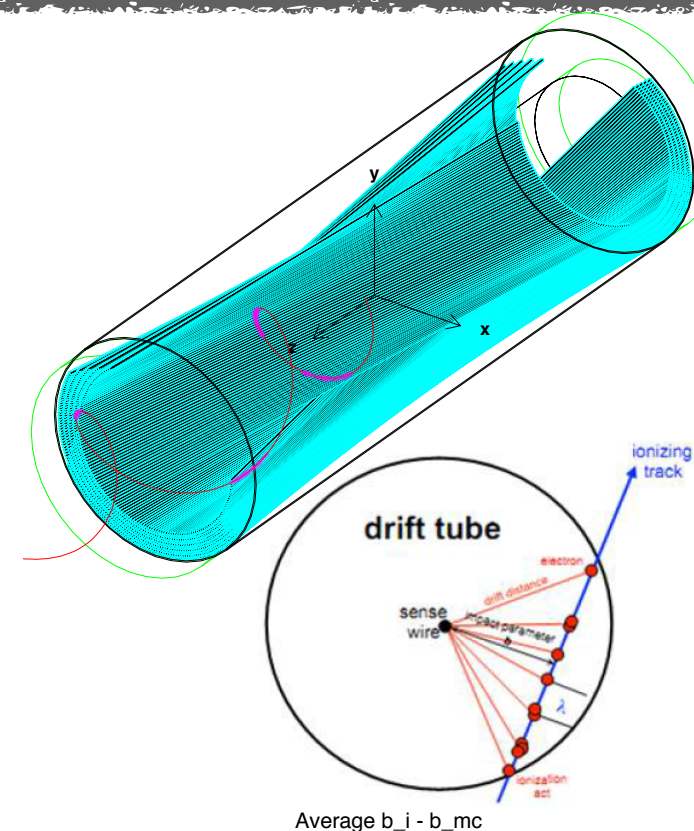
- An INFN-Pisa test bench for small tracking detector prototypes
- The facility
- Telescope set up
- Detector under test integration
- The MEGII drift chamber prototype characterisation



- The INFN telescope facility working group
 - *MEG II: Baldini, Bemporad, Cavallaro*, Cei, D'Onofrio, Dussoni, Galli, Grassi, Nicolò, Signorelli, Tenchini, Venturini*
 - *BaBar: Bettarini, Forti, Lusiani, Walsh*
 - *Electronics group: Avanzini, Morsani*
 - *High technology group: Bosi, Ceccanti, Mammini, Minuti*
- I am speaking on behalf of this group
- *worked on the facility during his master thesis period



- INFN took the responsibility to construct the new drift chamber for the MEG II experiment
 - *single volume wire chamber*
 - *stereo read out*
 - *low mass ionising gas He:IsoButane (85:15)*
 - important to study the position resolution with such a low ionising density: 1.29 cluster/mm
 - the measurement based only on the first ionisation cluster is biased
- Pisa has the leadership, Marco Grassi is the responsible
 - *need for a test bench to study the single hit resolution for small prototypes*



- Despite it is built to serve an experiment, it has to be as general purpose as possible
 - *designed to accommodate small and medium size detector under tests (DUT)*
 - *flexible DAQ scheme to include the DUT read out in the telescope code*
- Effective test bench for tracking detectors
 - *high resolution: resolution at the DUT plane of $\sim 20\text{-}30\ \mu\text{m}$*
- Cheap, “fast” and efficient solution
 - *free beam line: cosmic muons!*
 - *make use of spare modules of the BaBar SVT stored here in Pisa*
 - single hit resolution match the requirements
 - *design a new version of the read out electronics to be more compact and allow a multiplexed read out*
 - DAQ carried out by a CPU

- Operated five years at PEP-II in the BaBar experiment

- *inner tracker: b-tagging*

- z-vertex resolution $250\ \mu\text{m}$, same as B^0 mean path length

- *95-99% single hit efficiency*

- Five layers from the assembly of six wafer types

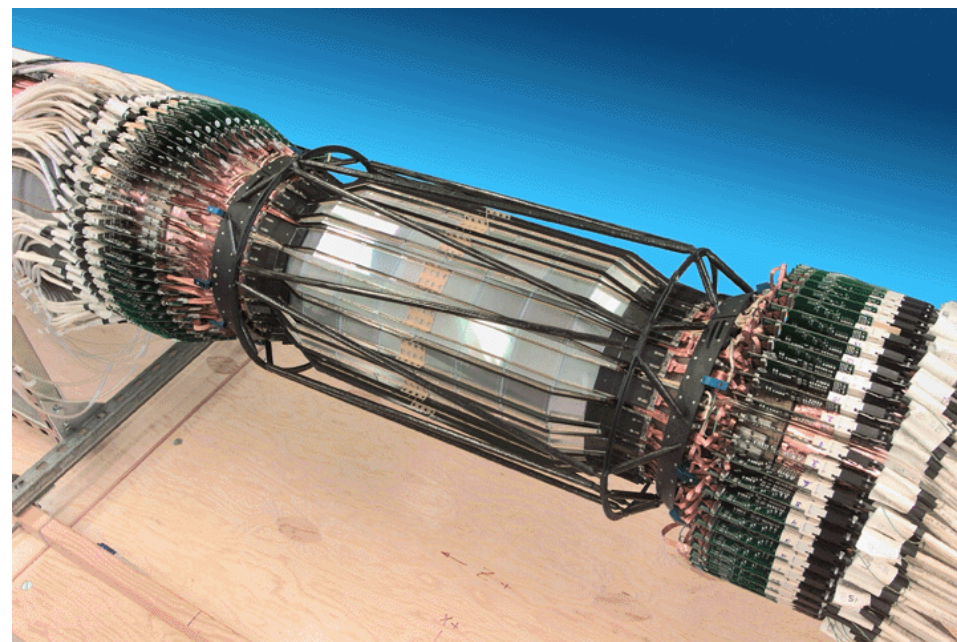
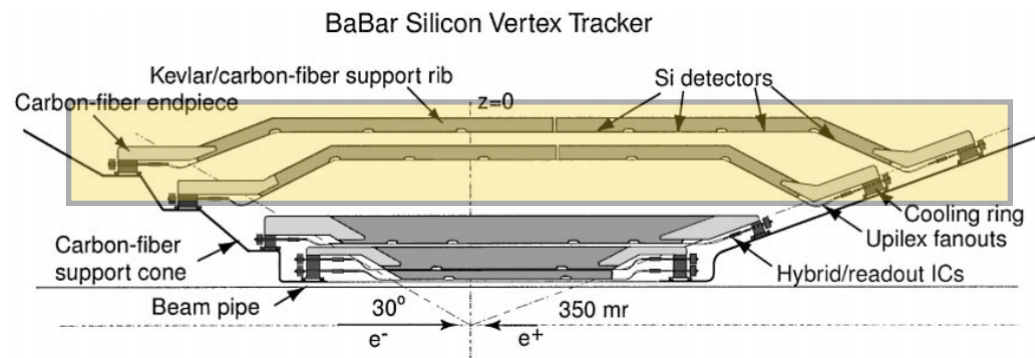
- *thickness $300\ \mu\text{m}$*

- *double sided read out*

- Type IV-V (used in the L4-5) pitch x-z = $100\text{-}210\ \mu\text{m}$
 - Layer 4 active area $\sim 5 \times 30\ \text{cm}^2$
 - Layer 5 active area $\sim 5 \times 40\ \text{cm}^2$

- Layer 4 and 5 spare modules are stored at INFN Pisa and ready to be used

- *what about muon tracking with L4 and L5?*



C. Bozzi, et al., "The BaBar silicon vertex tracker", NIMA, vol 453, pp 78-83, year 2000.

B. Aubert, et al (Babar Collaboration), "The BaBar Detector", NIMA, vol 479, pp 1-116, year 2002.

- Tracking precision as a function of

- single hit resolution*

- 29 μm x-, 61 μm z-view

- tracking precision from single hit resolution (using 4 points) at DUT plane 15 μm x- and 30 μm z-view*

- multiple scattering sources*

- telescope

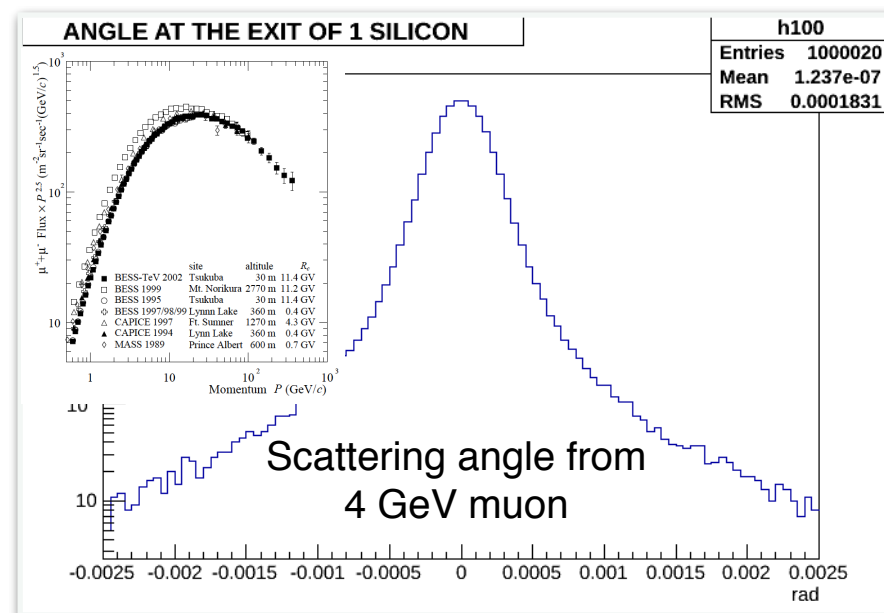
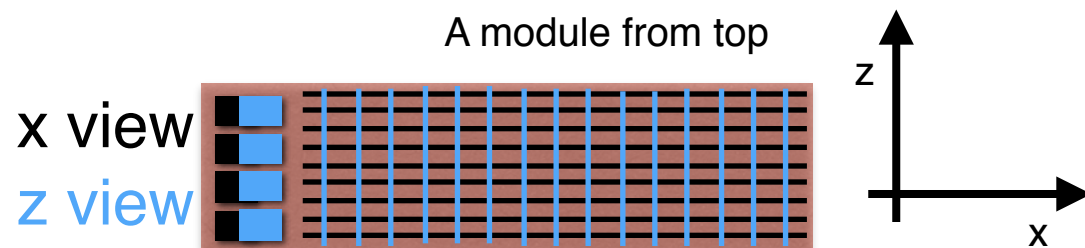
- main contribution by the SVT modules, 300 μm Si*

- DUT

- Resolution and MS is adequate to the goal

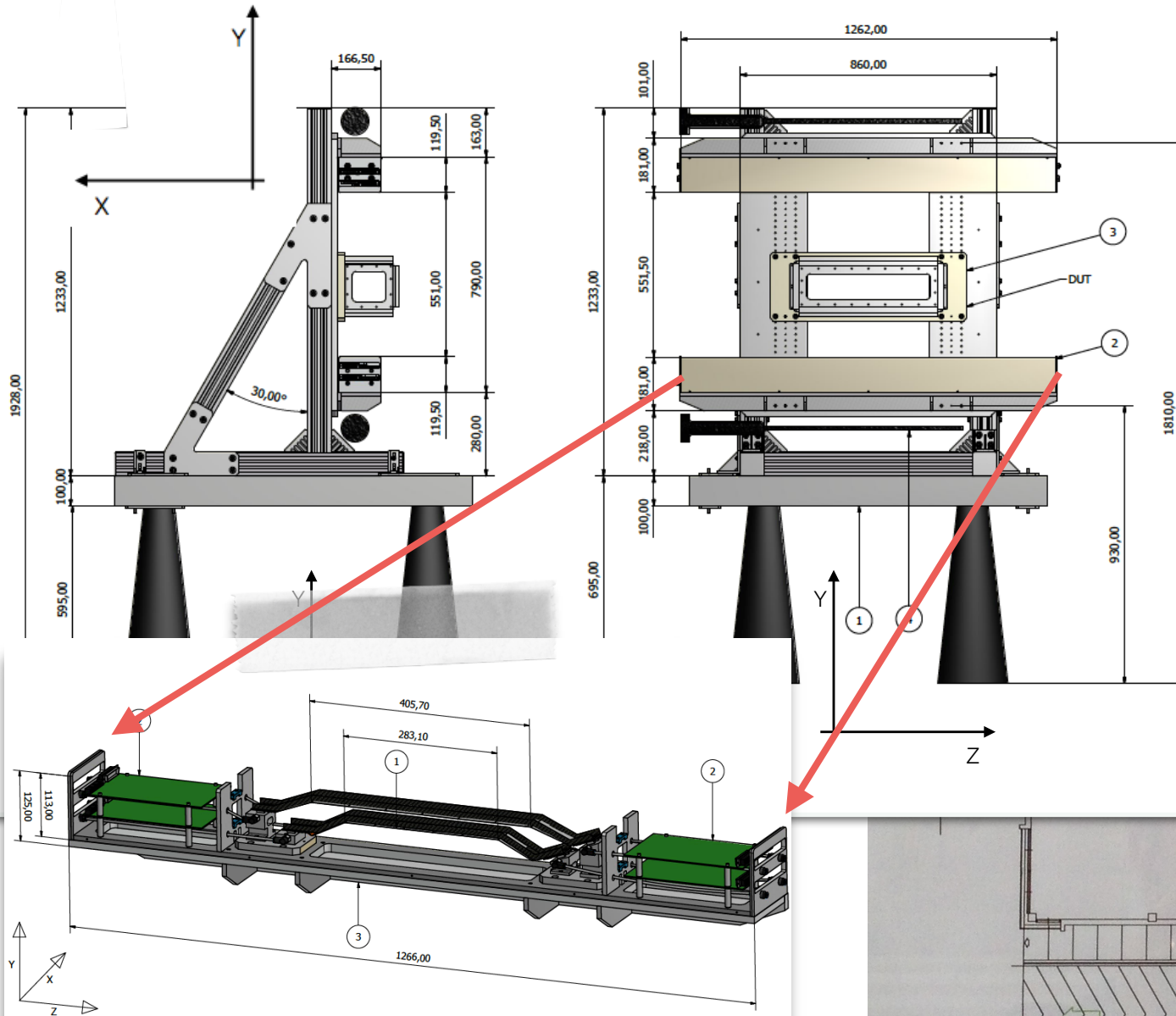
- but the DUT design must follow resolution requirements*

- Let's assemble it!!



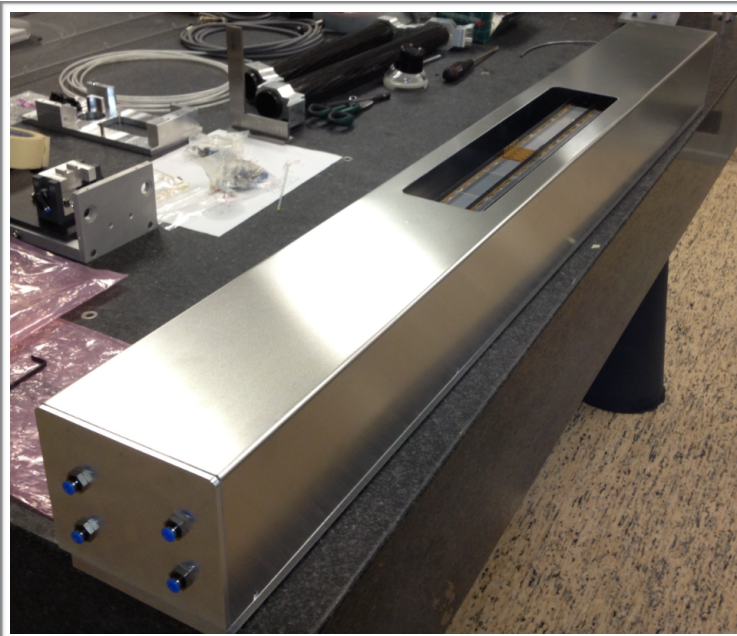
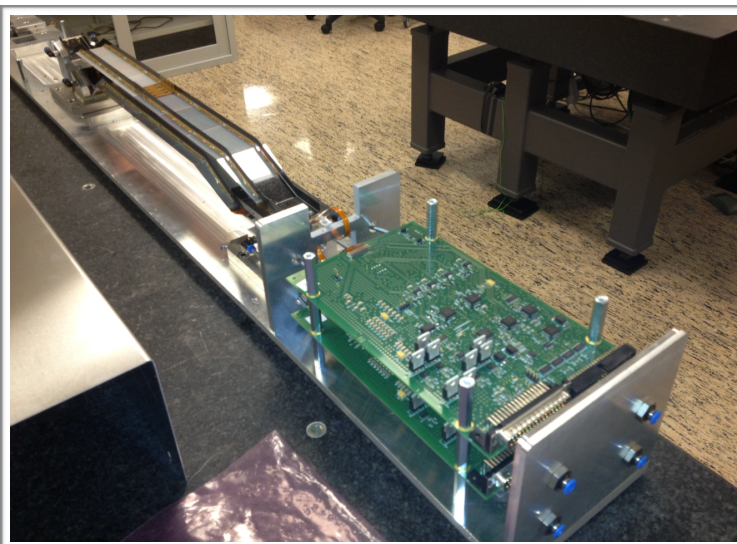
DUT thick (# SVT)	MS-uncert. (μm)	σ (μm , x — z views)
0	2	15 — 30
1	13	20 — 32
2	18	24 — 35
4	26	30 — 40

**GROUND
FLOOR**



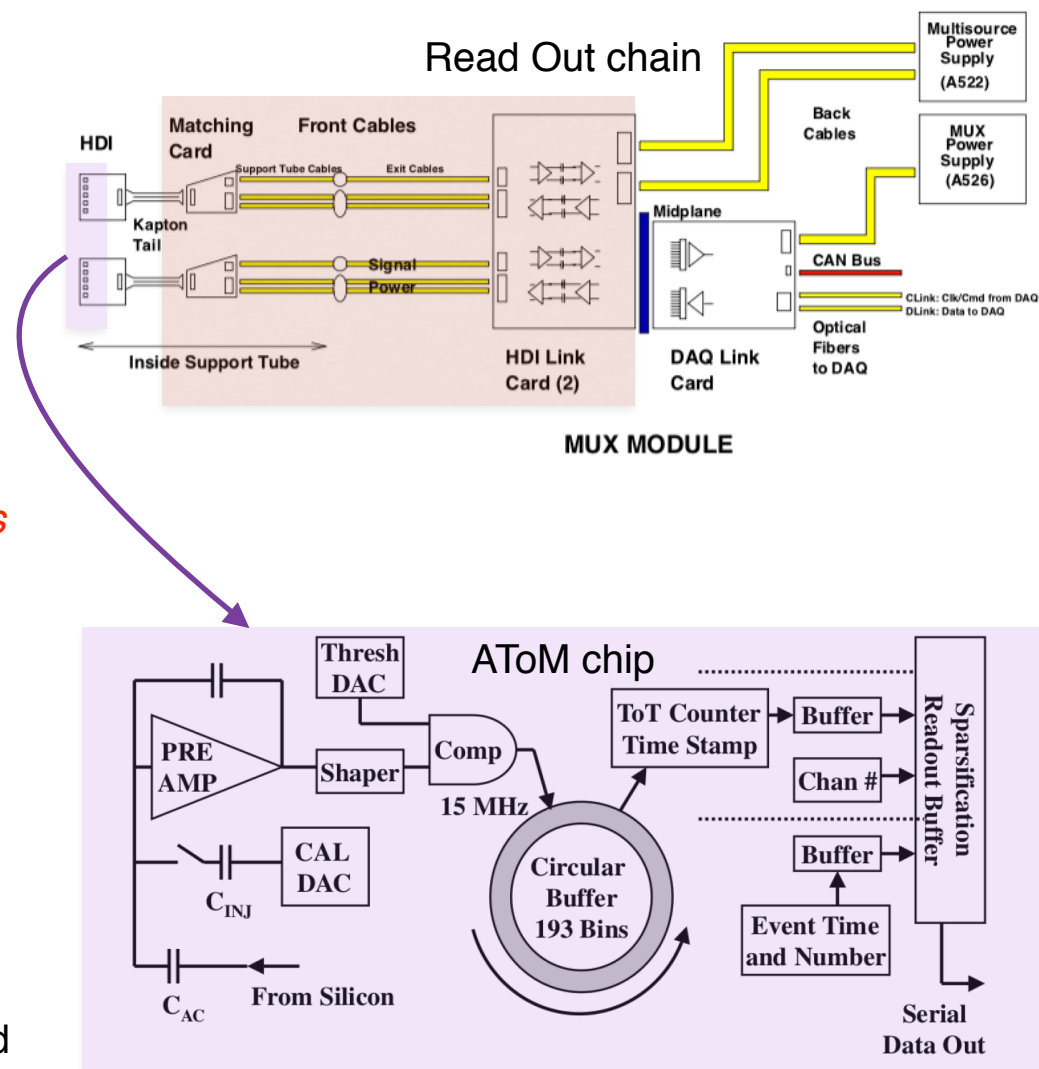
OFFICINE

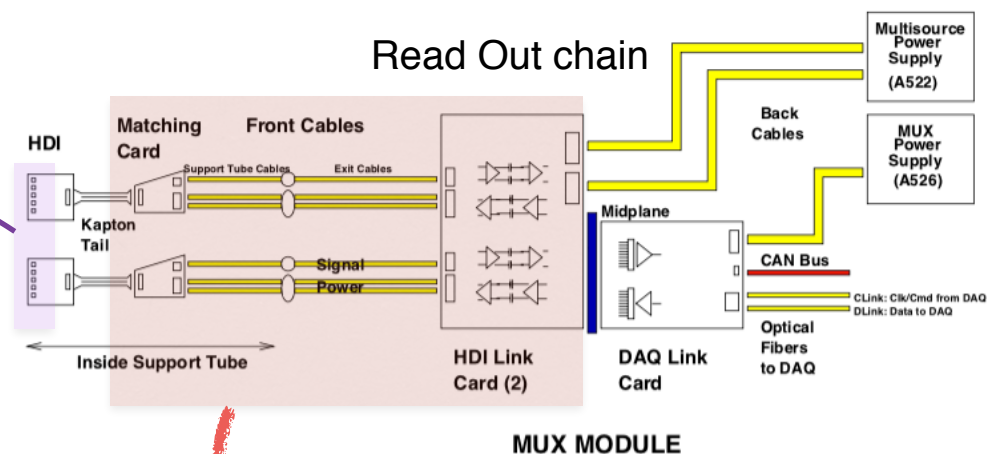
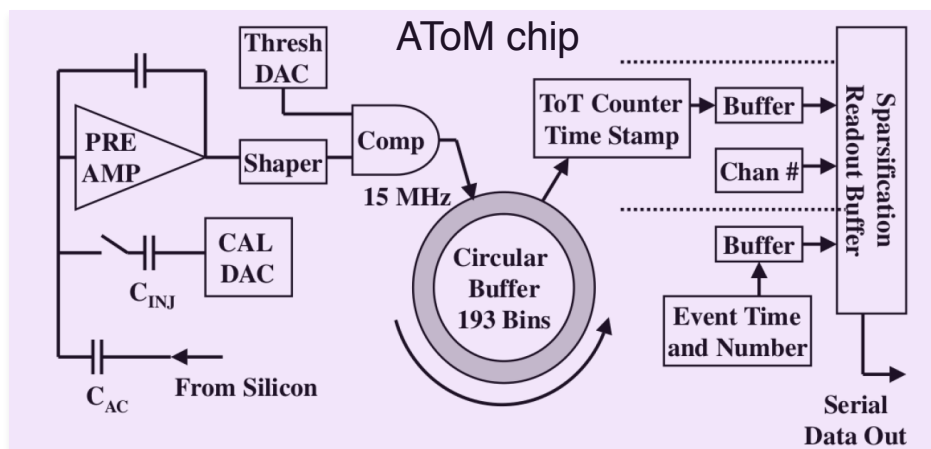
**Telescope
lab**



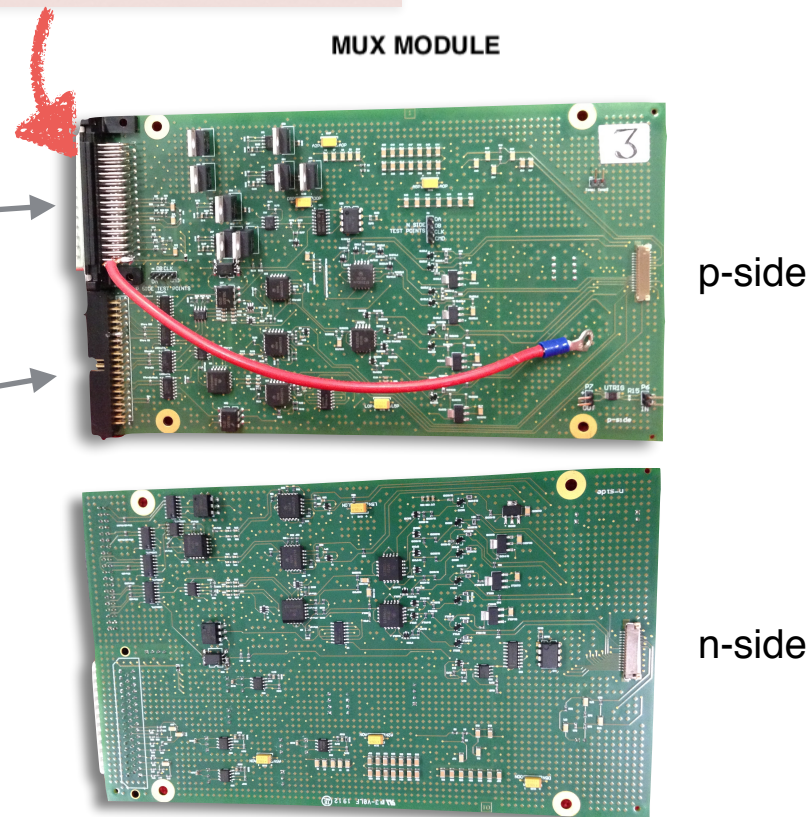
- Designed on the model of the spare module holders to host
 - *a layer4 (below) and a layer5*
 - *custom boards*
 - one each detector halves
- the box provides
 - *cooling with water pipes*
 - *dry air*
 - *light tightness*
 - *thin window in the telescope acceptance*
 - 20 μm aluminated-Mylar
- position relative to marker on the box plate with optical survey
 - *$\sim 100 \mu\text{m}$ accuracy*

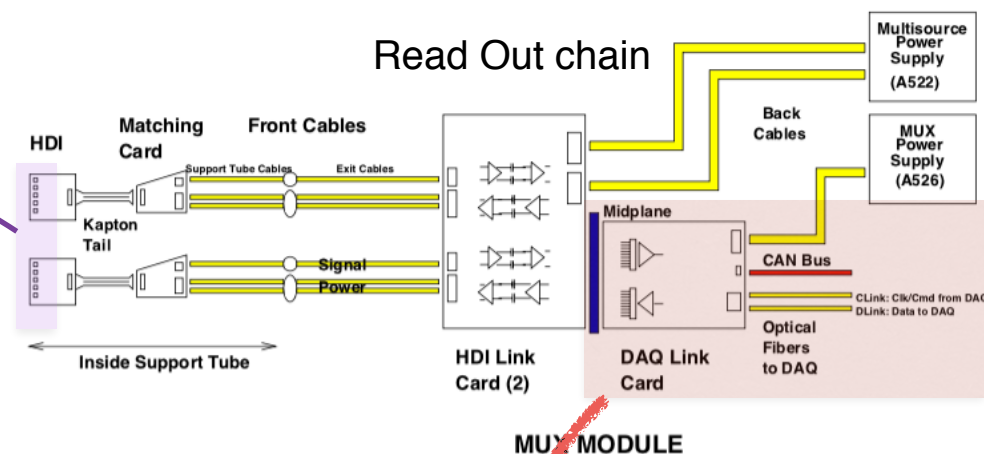
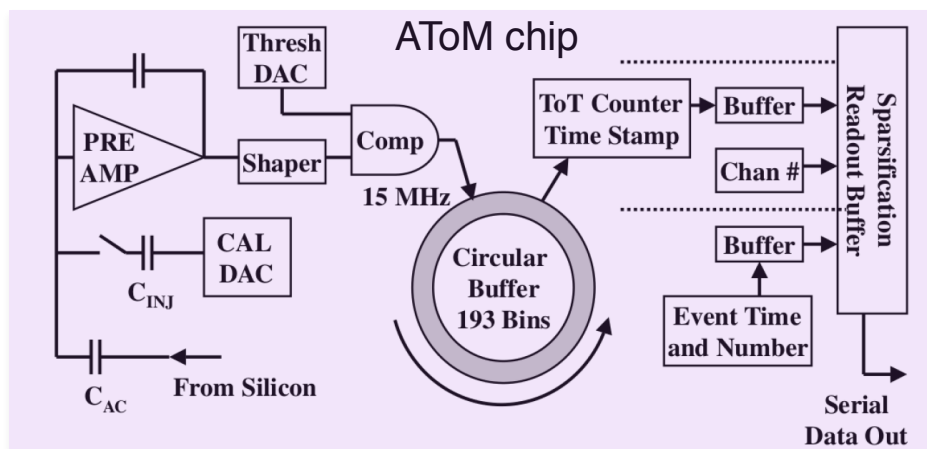
- DAQ card - **no usable anymore**
 - *generate clk and command signals and interpret the data back from AToM*
 - *connected to a SUN workstation*
- HDI link card + Matching card - **no usable anymore**
 - *connected to the detectors with kapton tails*
 - *handles signal distribution*
 - *supply to AToM*
- AToM read out chip - **attached to Si modules**
 - *A Time over threshold Machine*
 - connected to 128 channels
 - given an external trigger the list of the fired channels in output
 - internal calibration circuit



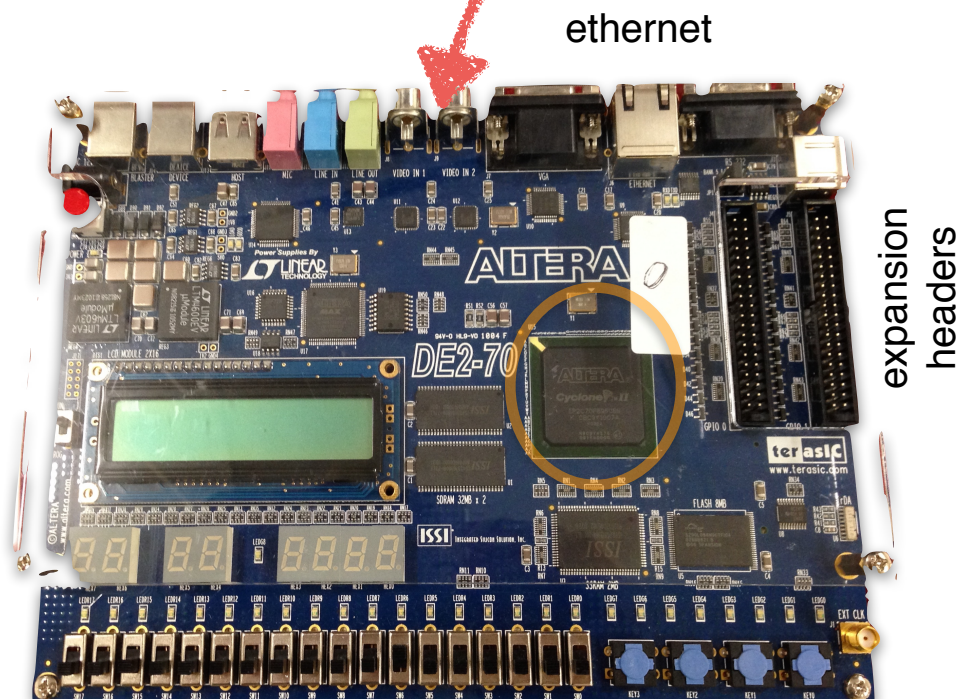


- Analog, digital and bias voltage supplier
 - *connected to original HV module*
- clk and serial lines translator
 - *LVDS signals from lower connector converted to PECL ATOM standard and vice versa*
 - electric scheme partially taken from original project
- connected to ATOM chips with kapton tail
 - *one custom per half a detector*
 - *a tail per side*
- placed inside the box close to the SVT module

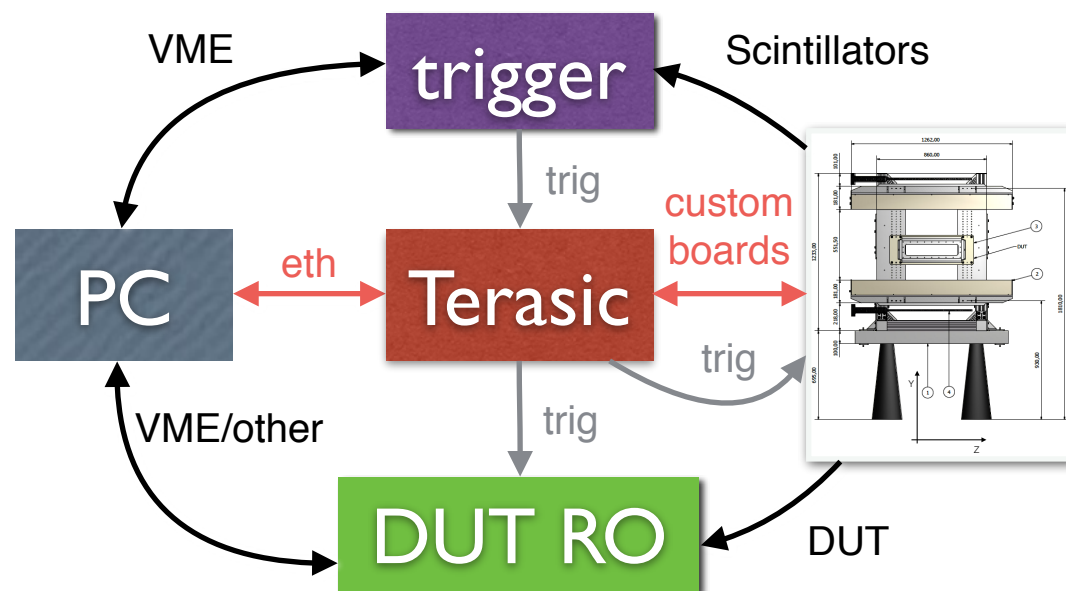


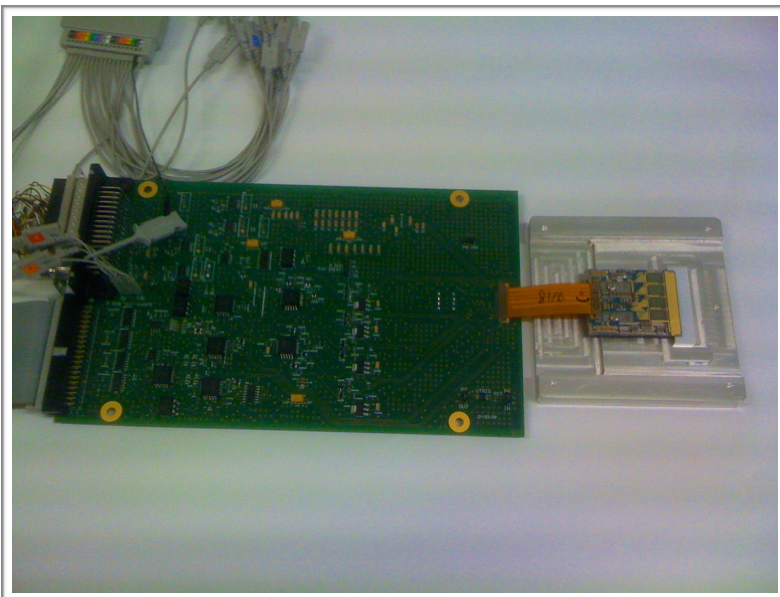


- Educational board with CycloneII FPGA to handle communication from DAQ PC to modules
 - *expansion header for the connection to a custom board*
 - *ethernet to an internal network*
 - network the master DAQ PC
- ATOM communication protocol programmed in the FPGA
- Ethernet link with the NIOSII processor

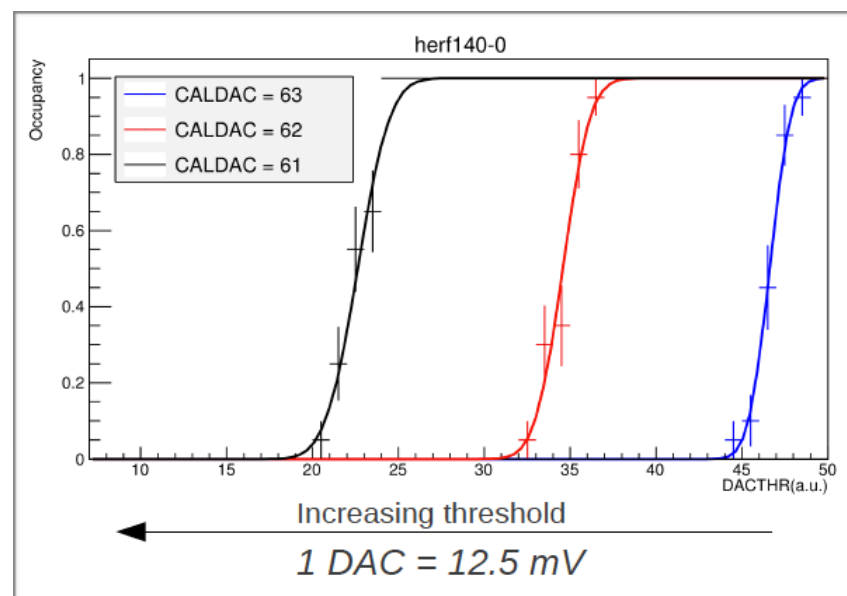
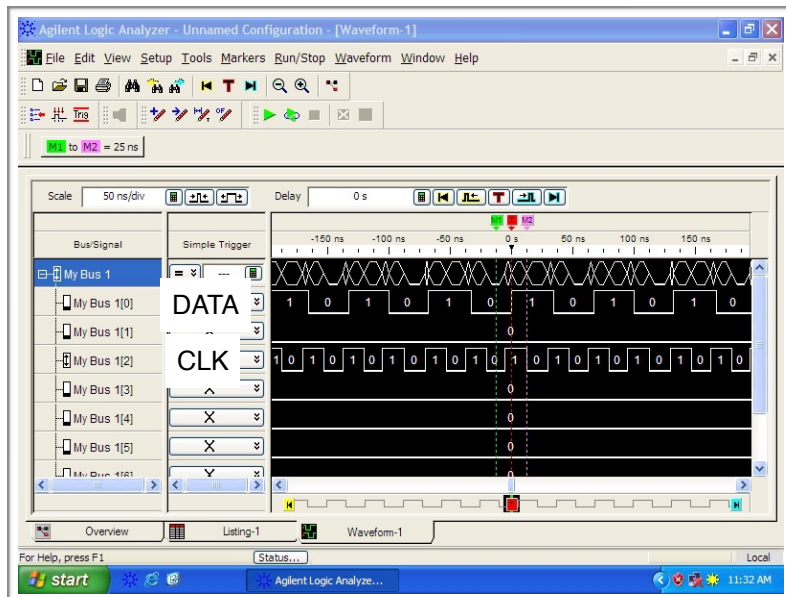


- PC - master
 - *unique DAQ supervisor*
 - no busy and event building issues
 - *C++ code*
 - server based ethernet connection
 - VME link
 - *trigger management*
 - high flexibility to accept many DUT read out
- Terasic boards
 - *client, send out data only upon request*
 - *linked to the detector halves/custom boards*
- Trigger
 - *MEG trigger VME board*
 - search for coincidence between the scintillators
 - programmable selection in case of other applications
 - *LVDS signal to stop AToM and DUT electronics*

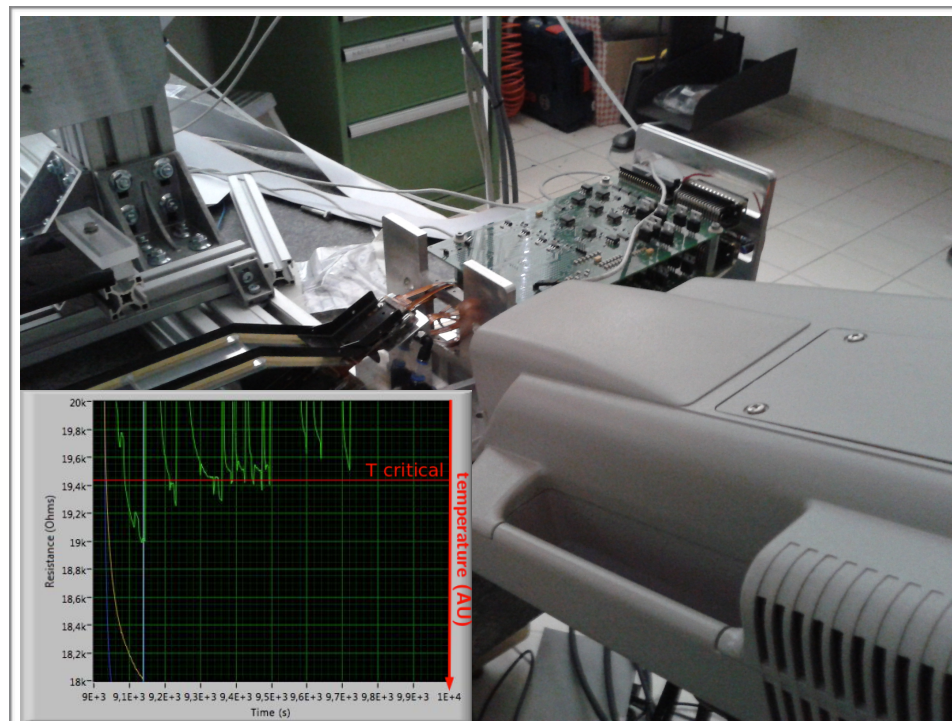
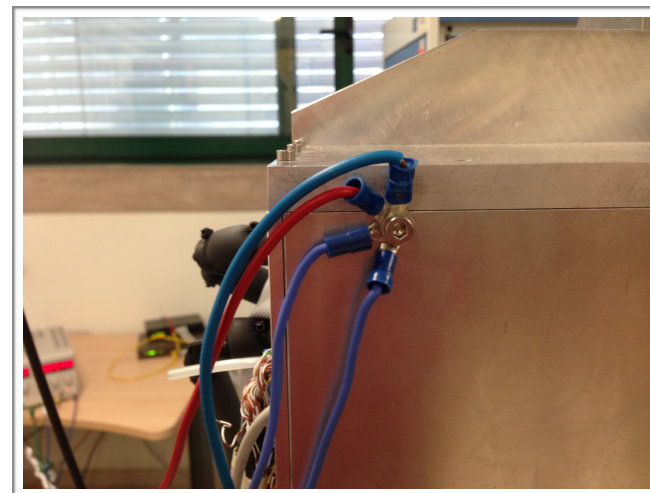




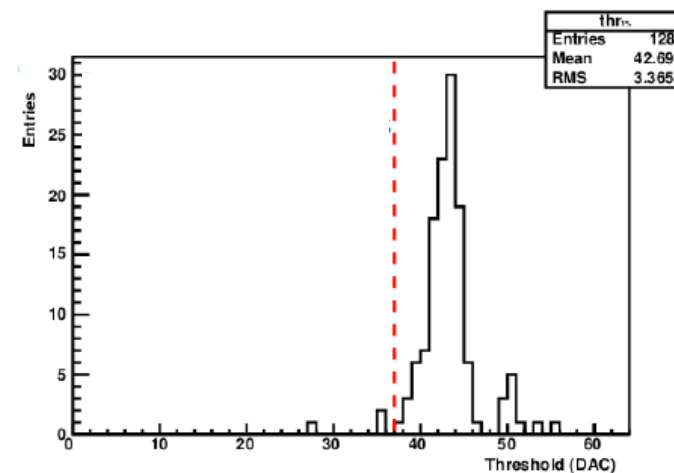
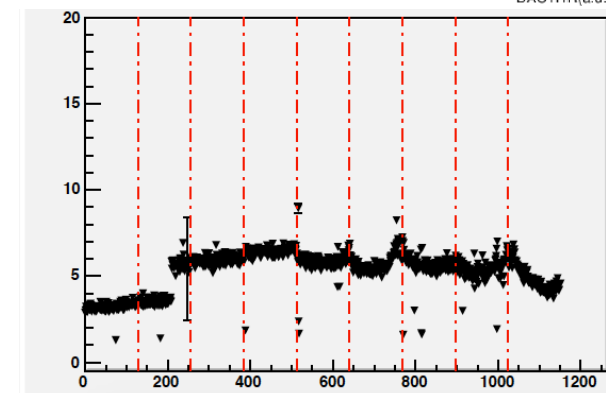
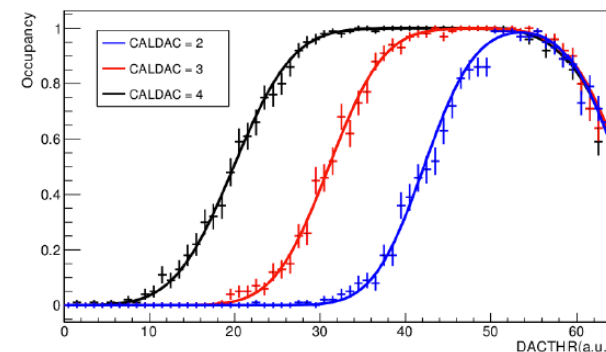
- First implementation of PC-Terasic-custom-AToM communication chain
 - *AToM answered to CLK THROUGH command*
 - debugging with logic state analyser
 - great success!!!
- Firmware development
 - *full calibration of the naked hybrid*
 - ECN measurement
 - discriminator linearity



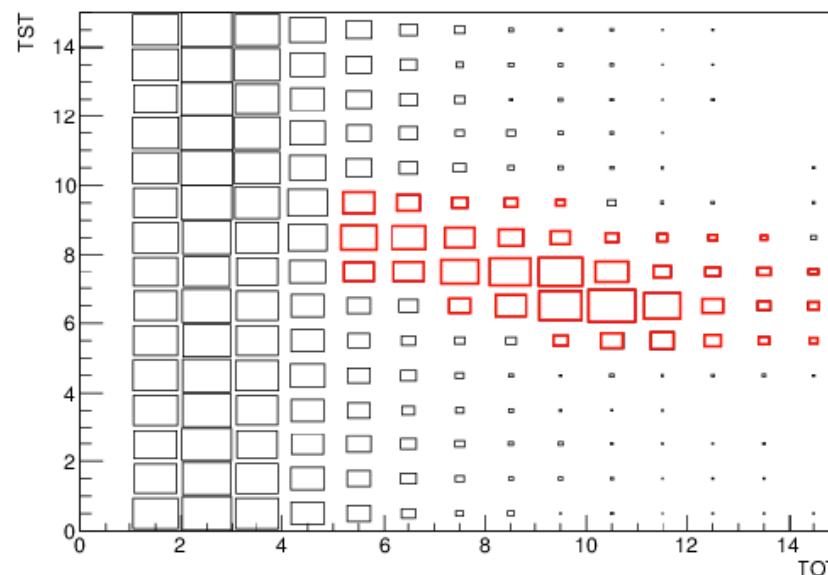
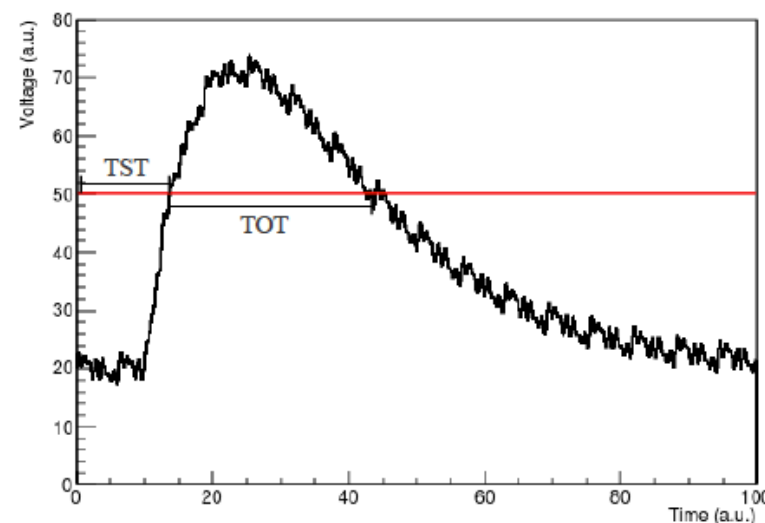
- Once each single module has been tested on bench we passed to detector assembly
- two major issues
 - *grounding*
 - suffering from ground loops
 - *in the test bench configuration the ground was the HV module*
 - *modified the HV module ground connection configuration and defined as aluminum support structure*
 - *chip temperature control*
 - the original chilling system not sufficient
 - *measured the hot point with thermal camera*
 - *at the T critical the communication with the module interrupted with severe risk of damage*
 - the solution was to install a dedicated air conditioning system to maintain the laboratory temperature around 24 degrees
 - *without up to 34 deg in summer...*
 - the temperature is monitored continuously during DAQ
 - *if it exceed the maximum value the HV system is shut down*



- Objectives
 - *dead channel list*
 - *set the chip discrimination threshold*
- Noise and gain from the excitation curves
 - *calculated for any of the 128 chip channels*
 - *threshold associated to a channel associated to the probability of noise induced hit $< 1\%$*
 - *distribution threshold and definition of noisy channels*



- Hit parameters
 - *Time Over Threshold*
 - proportional to the energy deposit
 - *Time STamp*
 - hit time
 - *used to find coincidence with the trigger*
 - TST peak used to calibrate the trigger latency
- Safe hit selection as a starting point for track reconstruction
 - *high charge and in time with the trigger*
 - red region

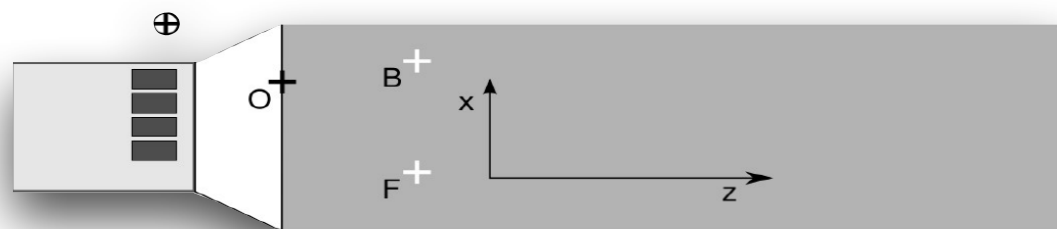


```
root [7] rec->Scan("", "tot>4&&tst<13&&tst>3&&nhit<100")
```

* Row *	* Instance *	* run *	* eve *	* typ *	* nhit *	* boa *	* chi *	* cha *	* tot .tot *
* 0 *	20 *	3027 *	0 *	0 *	63 *	0 *	4 *	61 *	9 *
* 0 *	21 *	3027 *	0 *	0 *	63 *	0 *	4 *	61 *	9 *
* 0 *	22 *	3027 *	0 *	0 *	63 *	0 *	6 *	88 *	10 *
* 0 *	23 *	3027 *	0 *	0 *	63 *	1 *	3 *	52 *	8 *
* 0 *	24 *	3027 *	0 *	0 *	63 *	1 *	3 *	52 *	8 *
* 0 *	25 *	3027 *	0 *	0 *	63 *	1 *	3 *	53 *	5 *
* 0 *	26 *	3027 *	0 *	0 *	63 *	1 *	3 *	53 *	5 *
* 0 *	28 *	3027 *	0 *	0 *	63 *	1 *	6 *	80 *	7 *
* 0 *	53 *	3027 *	0 *	0 *	63 *	6 *	1 *	122 *	11 *
* 0 *	59 *	3027 *	0 *	0 *	63 *	6 *	7 *	4 *	9 *
* 0 *	60 *	3027 *	0 *	0 *	63 *	7 *	4 *	60 *	9 *
* 0 *	61 *	3027 *	0 *	0 *	63 *	7 *	4 *	60 *	9 *
* 0 *	62 *	3027 *	0 *	0 *	63 *	7 *	7 *	12 *	11 *
* 2 *	2 *	3027 *	2 *	0 *	57 *	0 *	3 *	5 *	11 *
* 2 *	3 *	3027 *	2 *	0 *	57 *	0 *	3 *	5 *	11 *
* 2 *	14 *	3027 *	2 *	0 *	57 *	0 *	8 *	75 *	7 *
* 2 *	15 *	3027 *	2 *	0 *	57 *	0 *	8 *	76 *	6 *

- The first tracks...

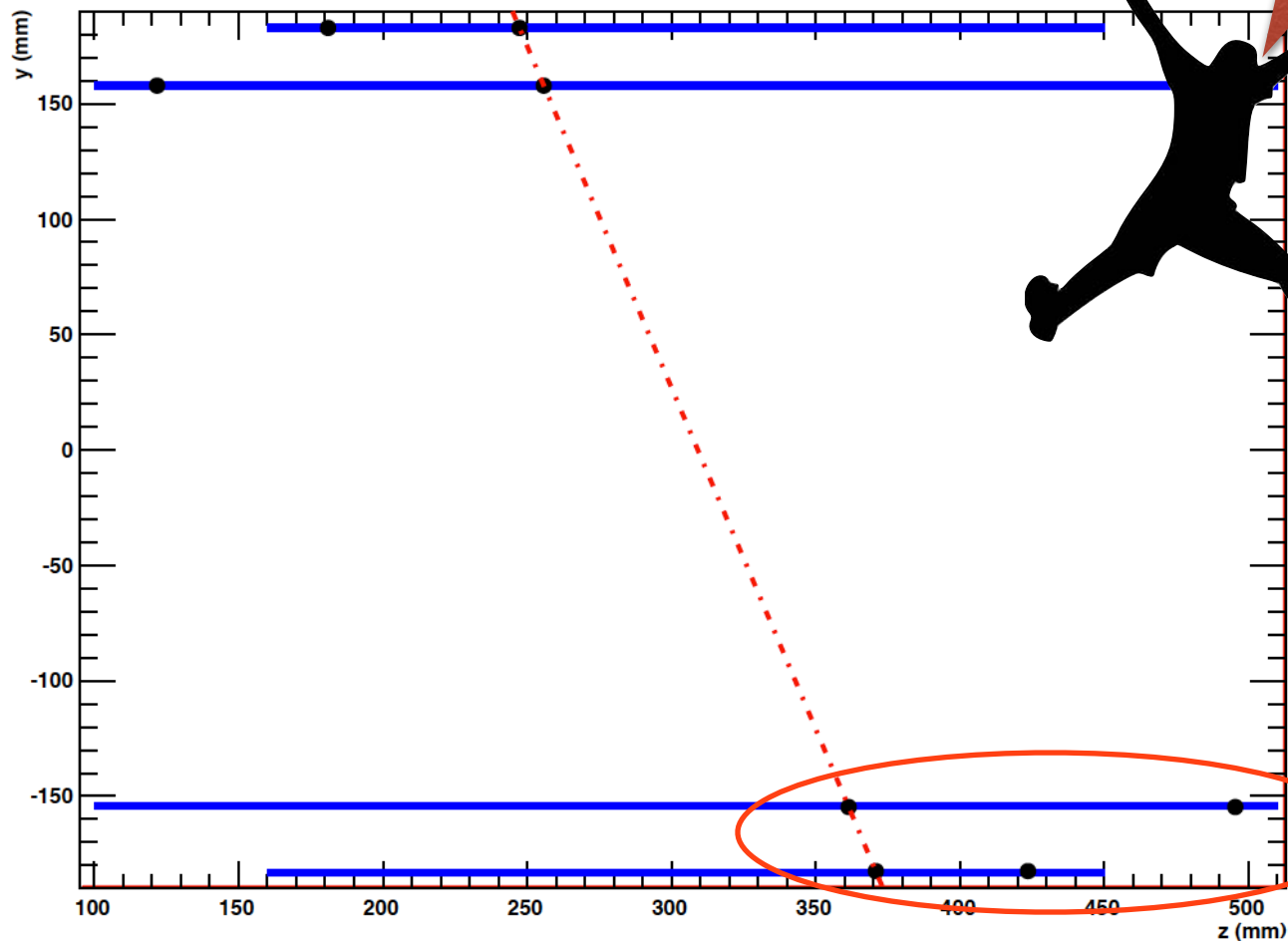
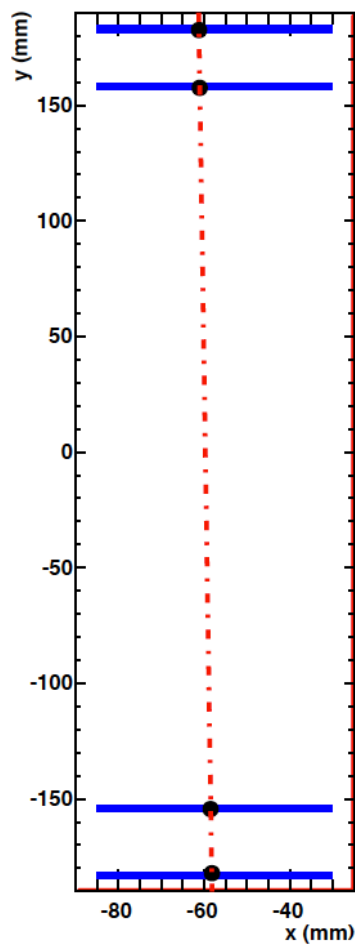
- *hit on 4 different boards*
 - 4 SVT modules
- *both p- and n-side*
 - by looking at chip number



- We always require for a 8-hit track to push at the best the tracking performance
- Local coordinates to telescope with rotations and translations

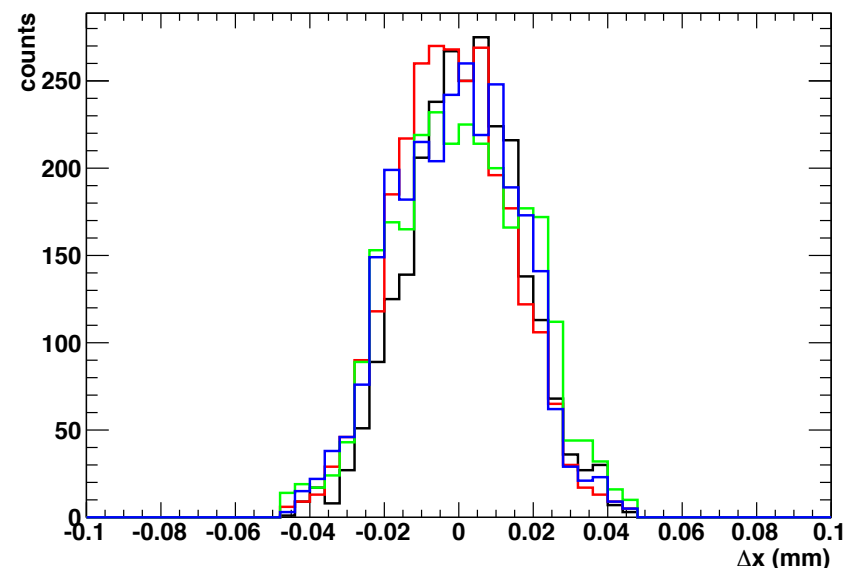
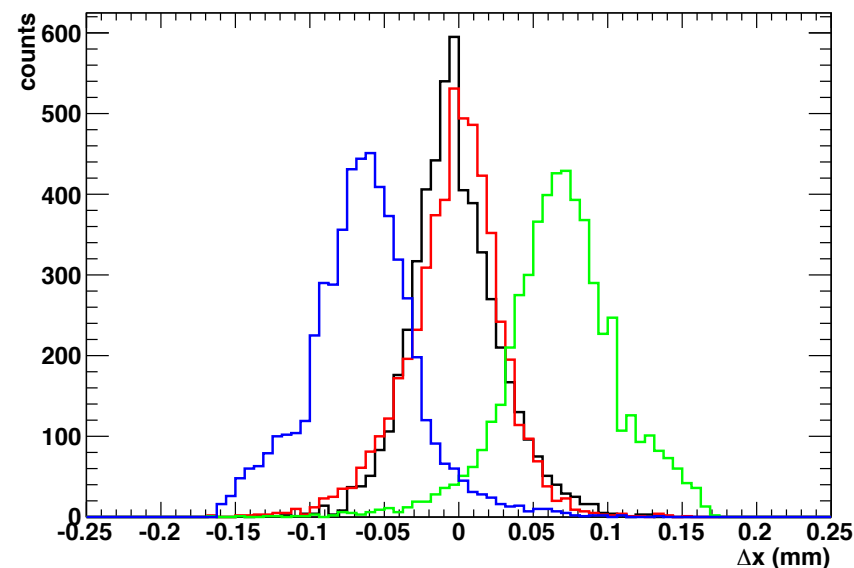
A track!

Silicon detector Hits

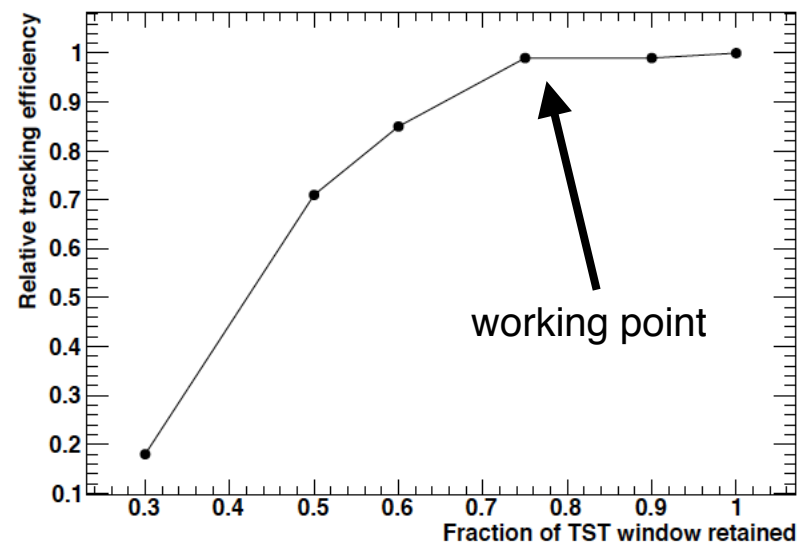
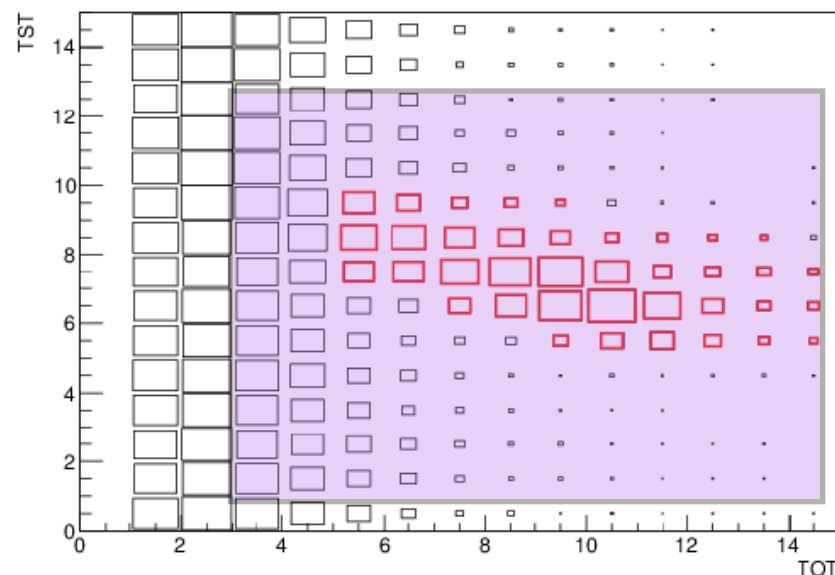


HW ambiguity, 1 AToM channel associated to 2 different strips

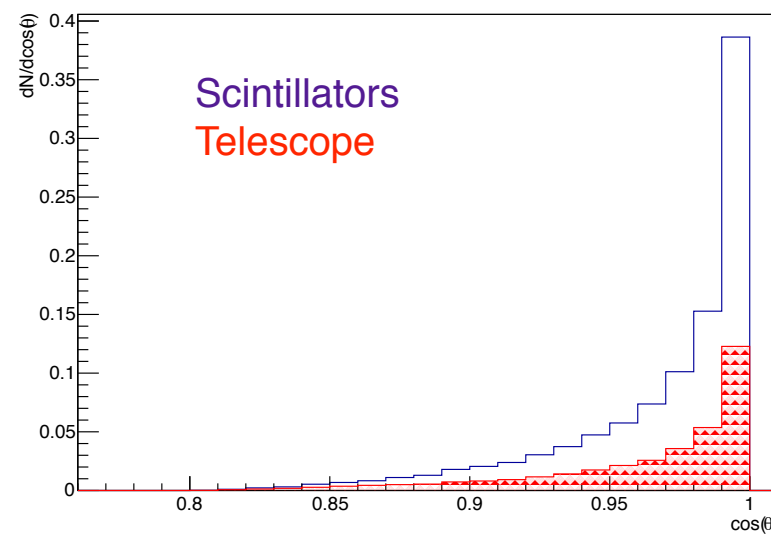
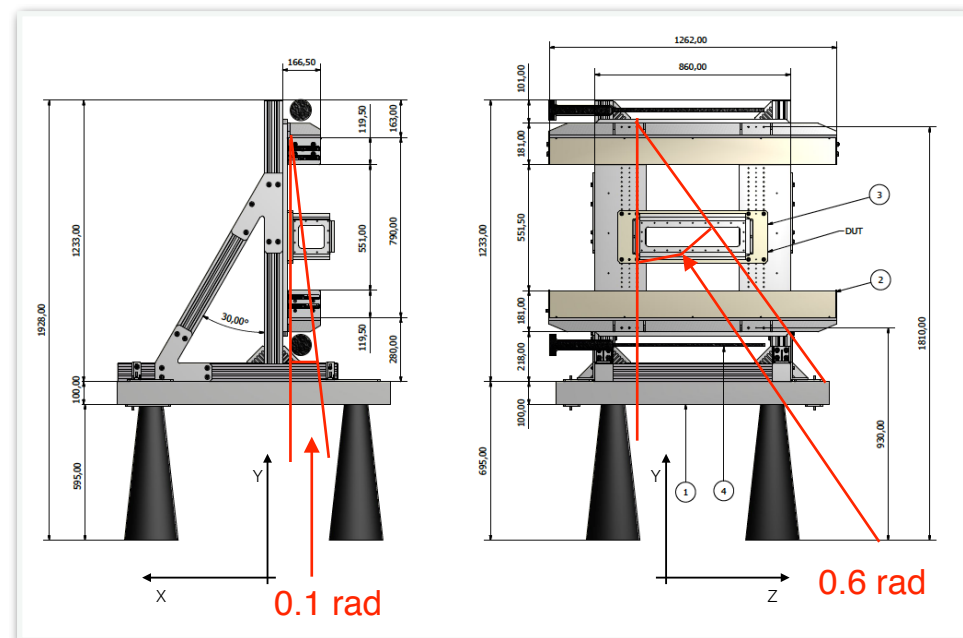
- Optical survey accuracy $100\ \mu\text{m}$
 - *need of software alignment*
- Track selection with rather relaxed chi2 cut
 - *we expect “large” residuals*
 - we do see up to $\sim 100\mu\text{m}$ shifts
- alignment by
 - *shift in x-z directions*
- After SW alignment
 - *all residual centered at 0*
 - *gaussian distribution and sigma*
 \sim hit resolution: $\sim 20\mu\text{m}$



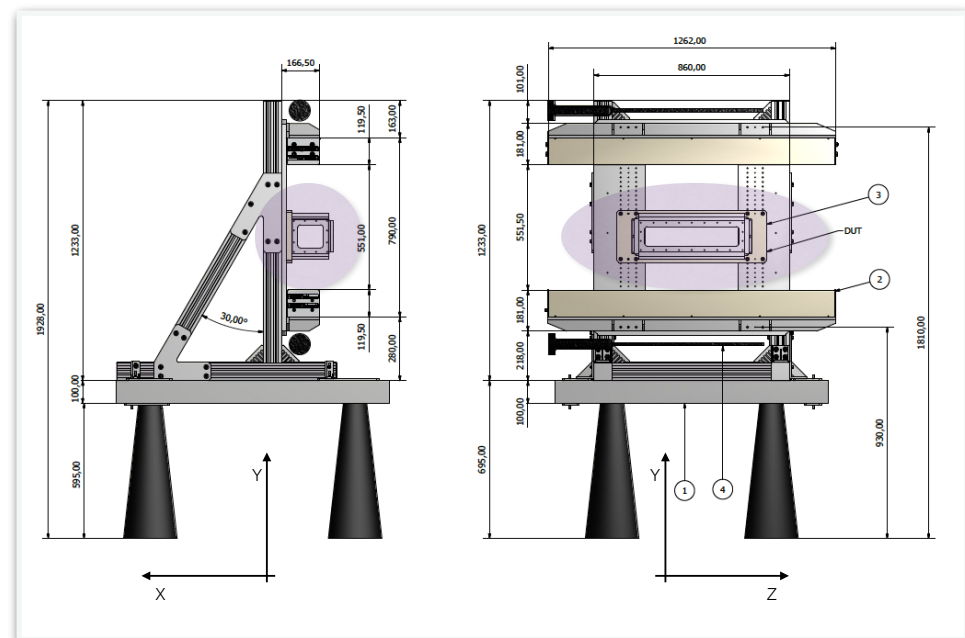
- Reconstruction code training with clean events
 - *very low noise-hit contamination*
 - *looks for the best chi-square combination*
 - not sophisticated but adequate to this system
- aim at highest efficiency
 - *relax hit selection condition*
 - *looking for saturation of the relative tracking efficiency*
 - no performance degradation looking at residuals and chi-square



- Numbers computed with 50cm gap between the box
- Small telescope acceptance
 - *max angle = 0.1 rad in the x-y view*
 - *0.6 rad in the z-y*
- expected (and measured) rate $\sim 0.3\text{Hz}$
- 30% relative scintillator - telescope acceptance
 - *order of 130 reconstructed tracks/run*
 - we can extract each plane efficiency to be larger than 90%
- At the DUT plane
 - *$\sim 3\text{ tracks/hour/cm}^2$*
 - *$\sim 500\text{ tracks/week/cm}^2$*

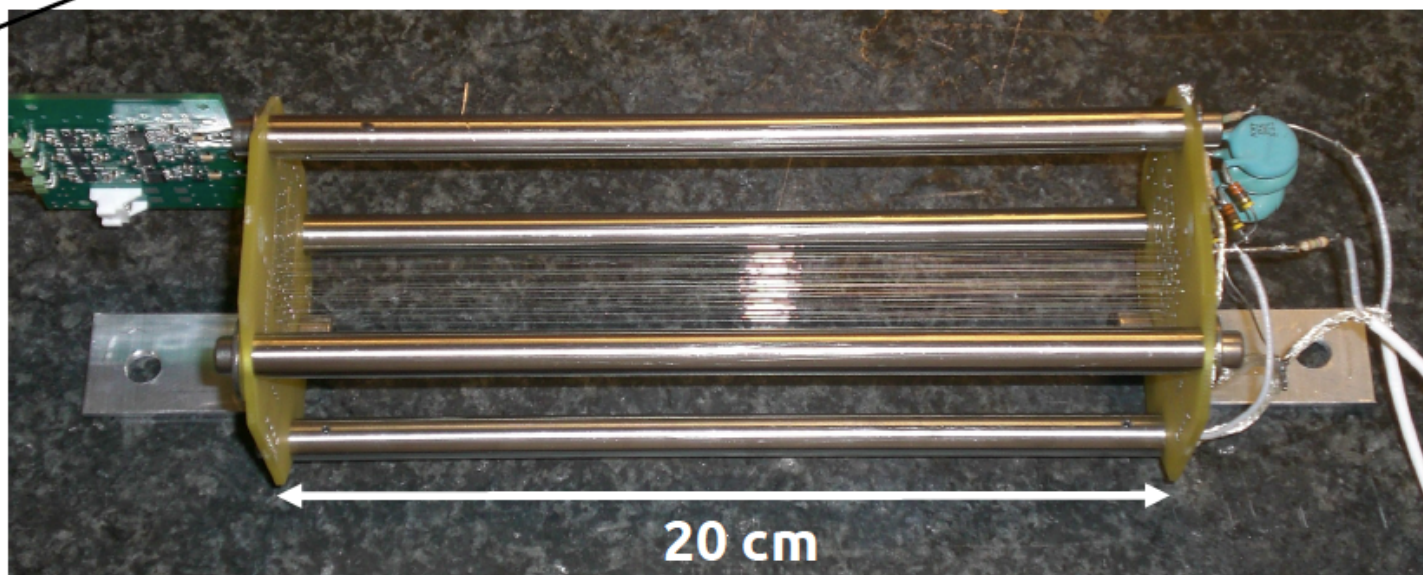
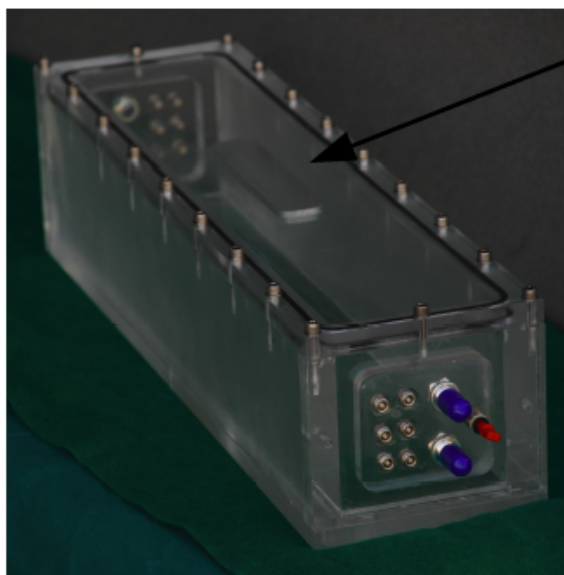
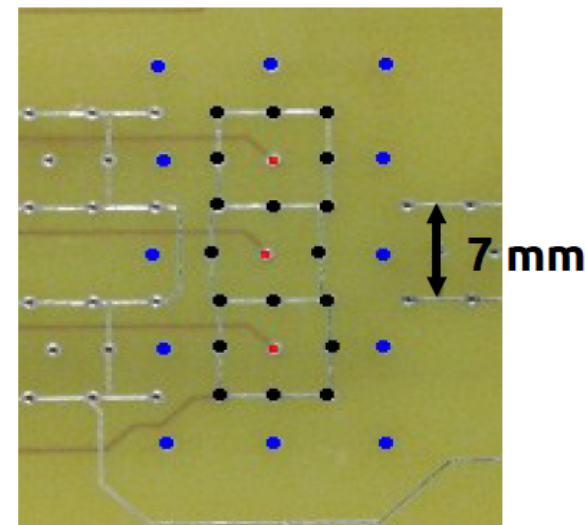


- Small size detector
 - *the telescope active area $5 \times 20 \text{ cm}^2$*
 - *take care about thickness*
 - it can degrade the performance
 - *largest box distance allowed: $\sim 50 \text{ cm}$*
- Read out
 - *ideally linkable to master PC*
 - busy and event building not an issue...
 - VME
 - ethernet
 - USB
 - ...
 - *if not possible, solution tailored on needs*
 - hardware and software modifications

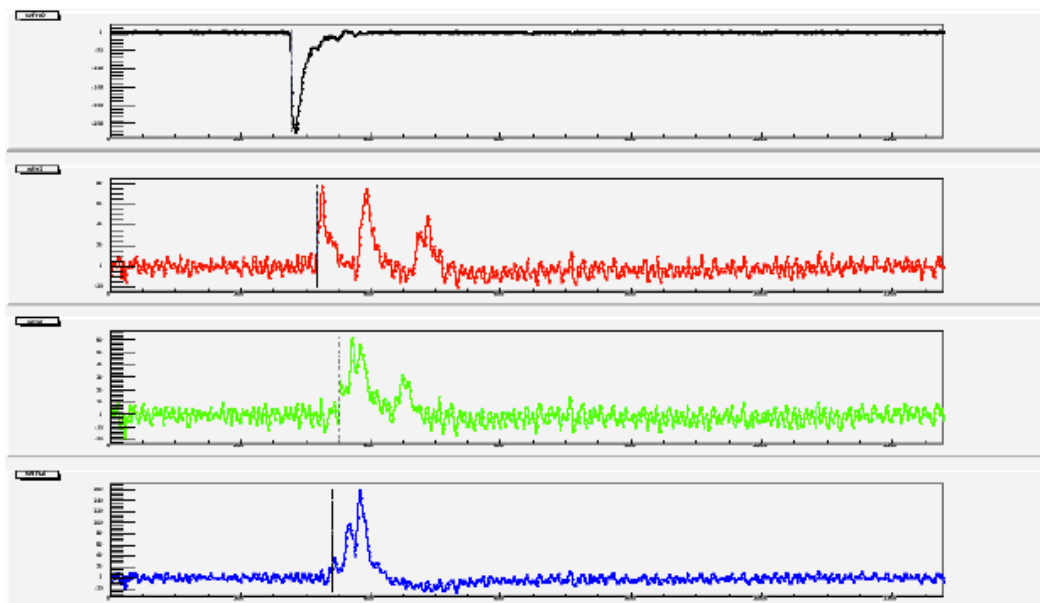


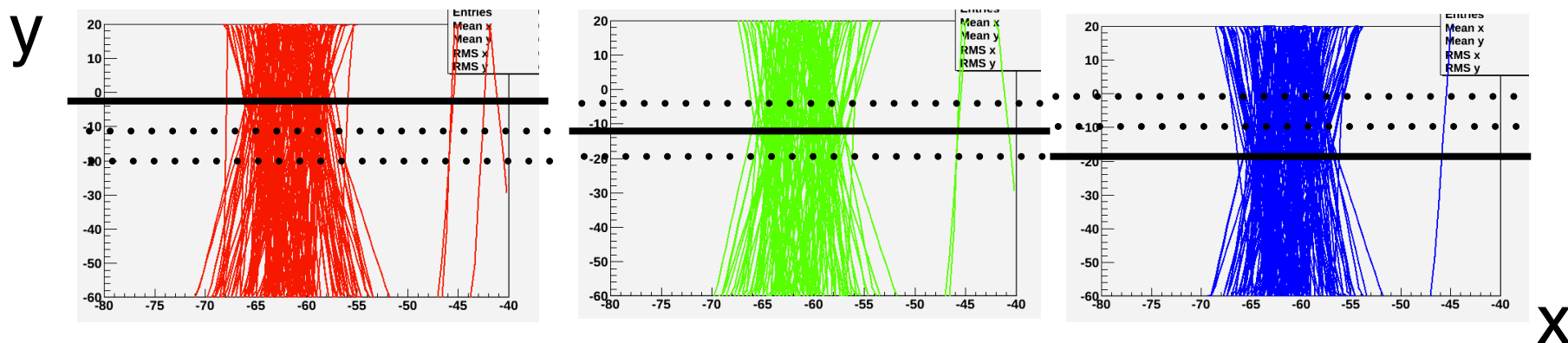
DUT thick (# SVT)	MS-uncert. (μm)	σ (μm , x — z views)
0	2	15 — 30
1	13	20 — 32
2	18	24 — 35
4	26	30 — 40

- squared cells, 7×7 mm², drift tube length ~ 20 cm
- staggered cells for self tracking capability
- light gas admixture: He-Isobutane 85:15 : 12.9 cluster/cm
- HV 1600-1700 V, max drift time ~ 150 ns
- MEG II fast pre-amplifier, BW ~ 1 GHz for cluster identification
- Gas tight plexiglas box with 1 mm window in the active region



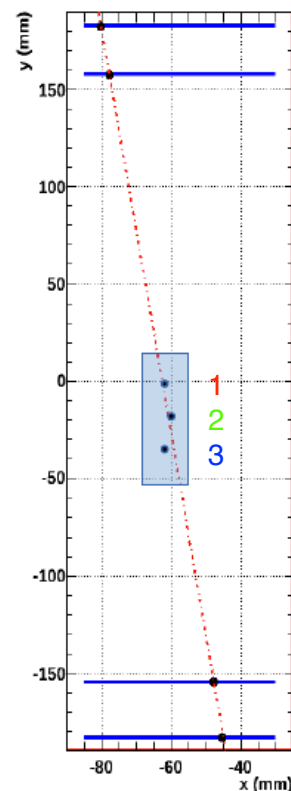
- Read out with VME CAEN V1729
 - *12 bit*
 - *2 GSPS*
 - *300 MHz BW*
 - *differential input*
- configuration and data transfer code into telescope C++ DAQ
 - *upper scintillator PMT digitised for track T0*



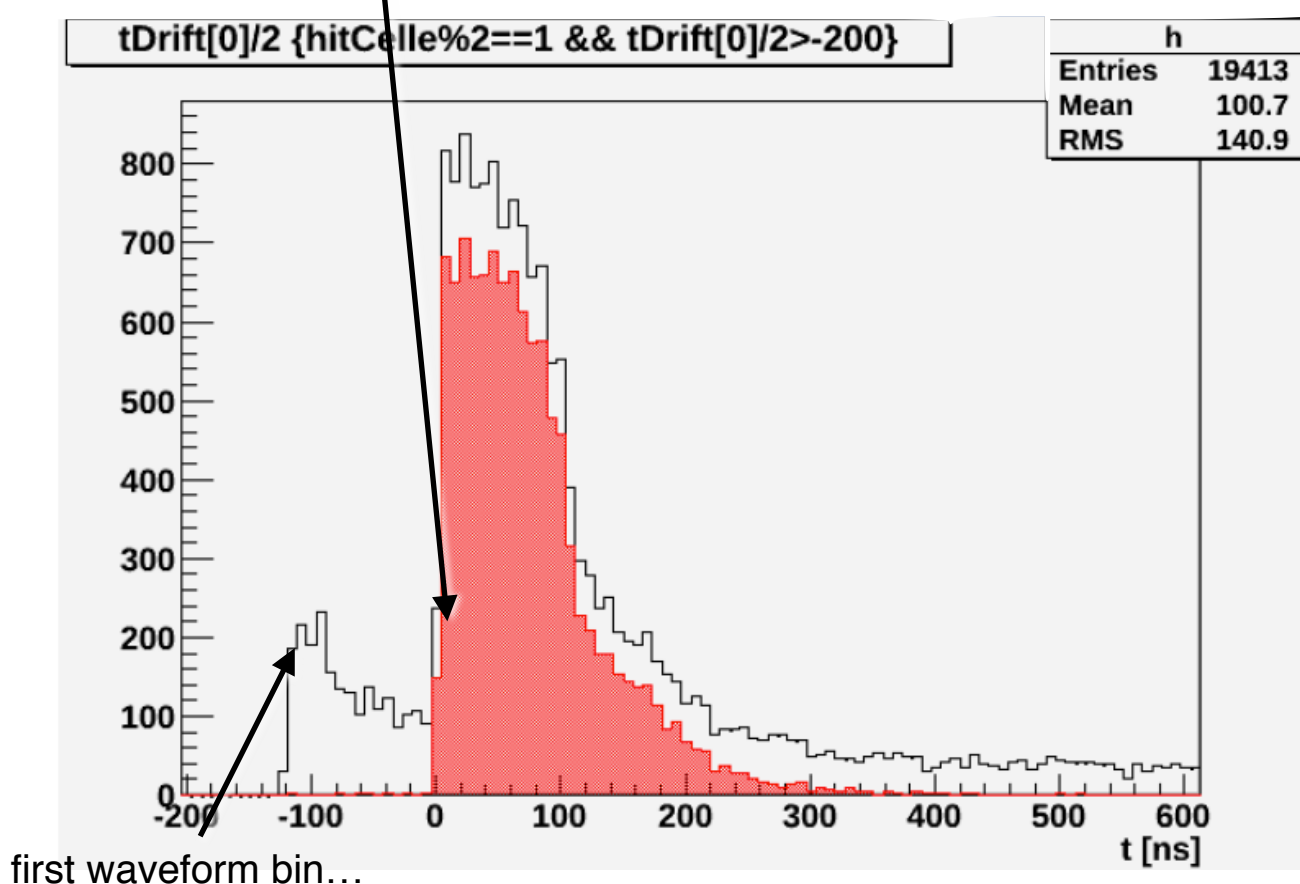
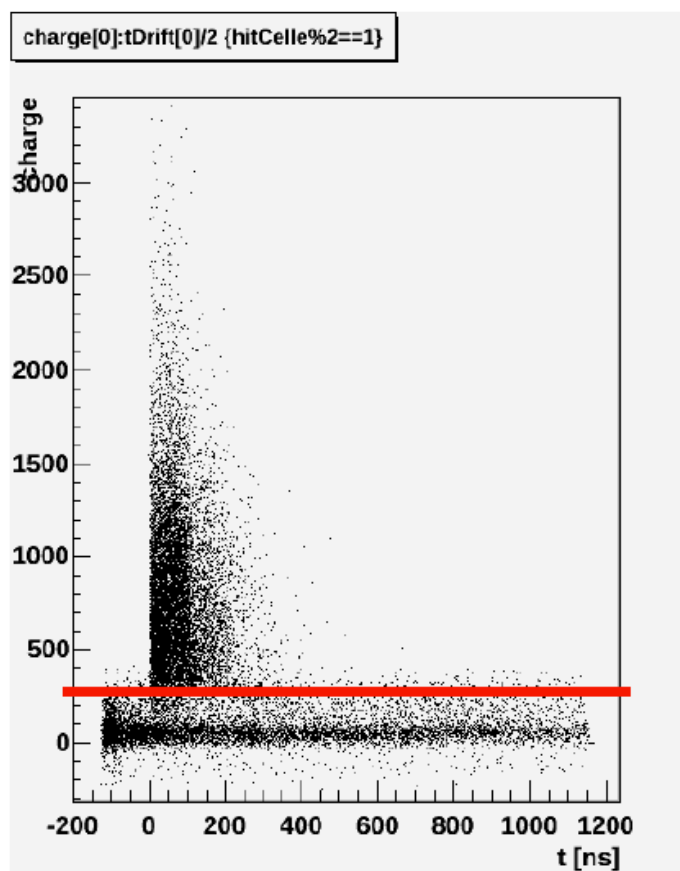


- The cell position in the y coordinate can be identified by looking at the distribution of the tracks having an hit on the cells (1 2 3)

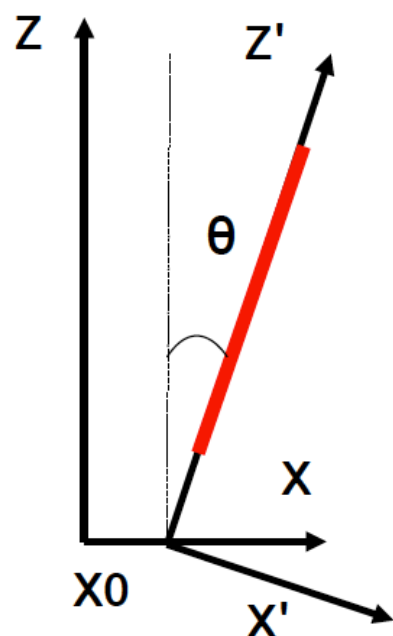
- *shift in y visible by eye*
 - it has to be with a properly working telescope!
- *x-stagger not visible (only 500 μm ...)*
 - this is not the right way to look at it



- Time offset calibration w.r.t. PMT reference
 - *time from SW double threshold discriminator*
 - *noise rejection looking at charge*
 - we present a set of run with high noise (grounding lost), it is usually much better

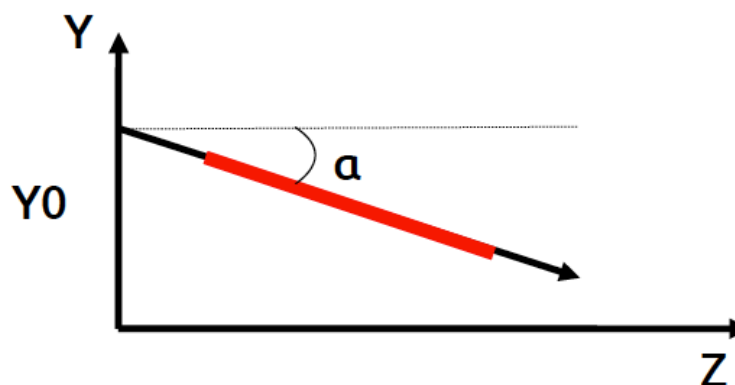


- wire equation in the telescope reference system to compute the impact parameter b
 - *top view: wire imaging*
 - *front view: strabismus effect*



"Top view"

4 parameters: $(\theta, X_0, \alpha, Y_0)$



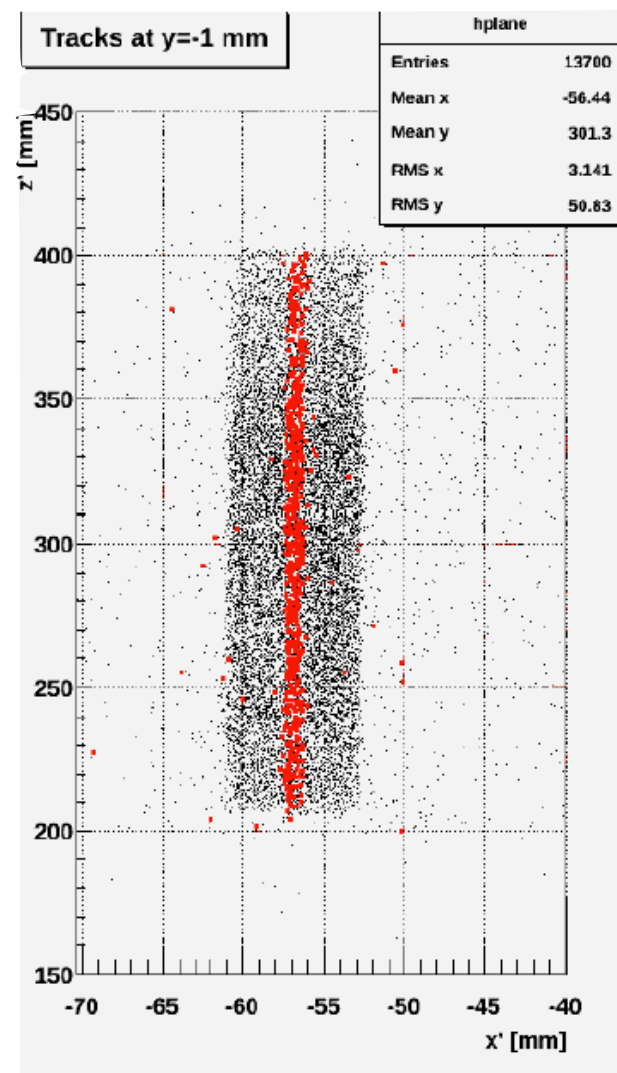
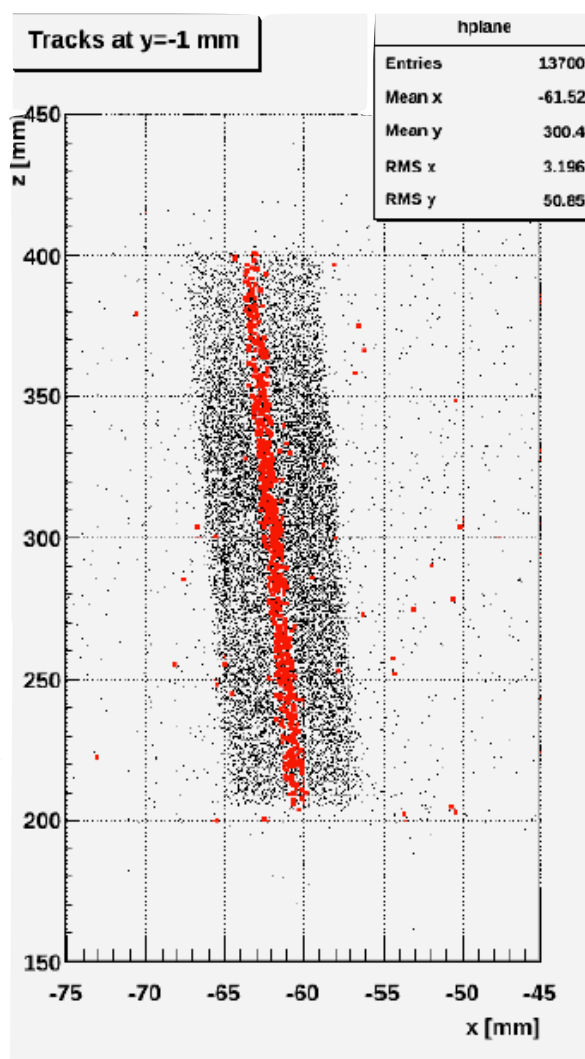
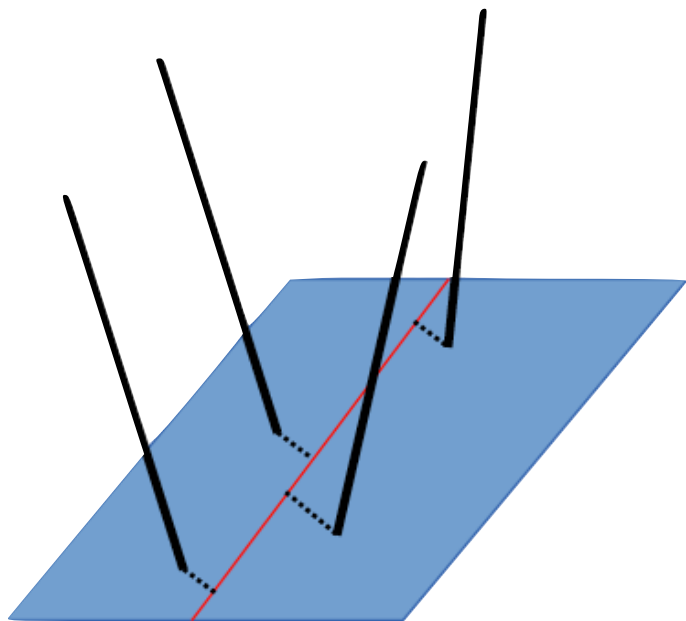
"Front view"

- event with small drift time:
 $t < 5\text{ ns}$

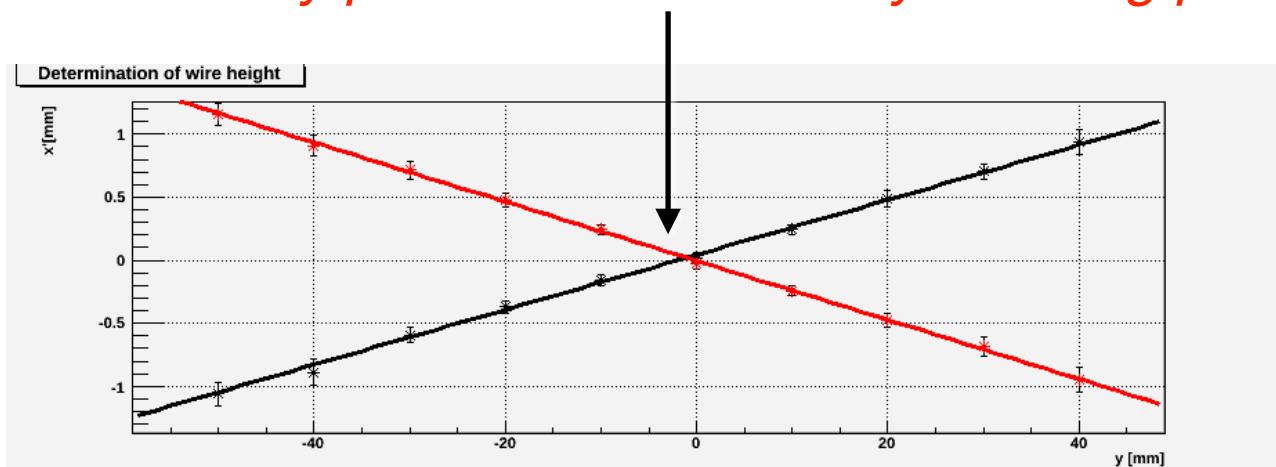
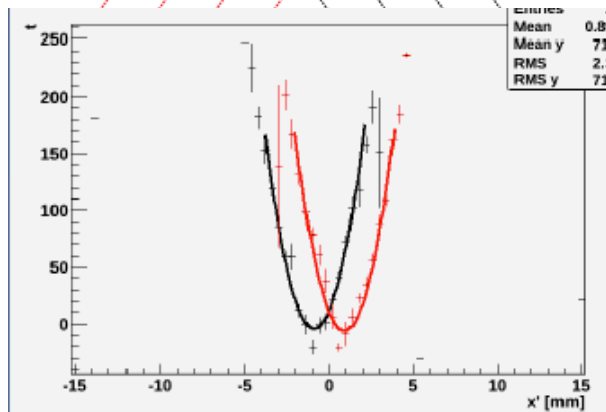
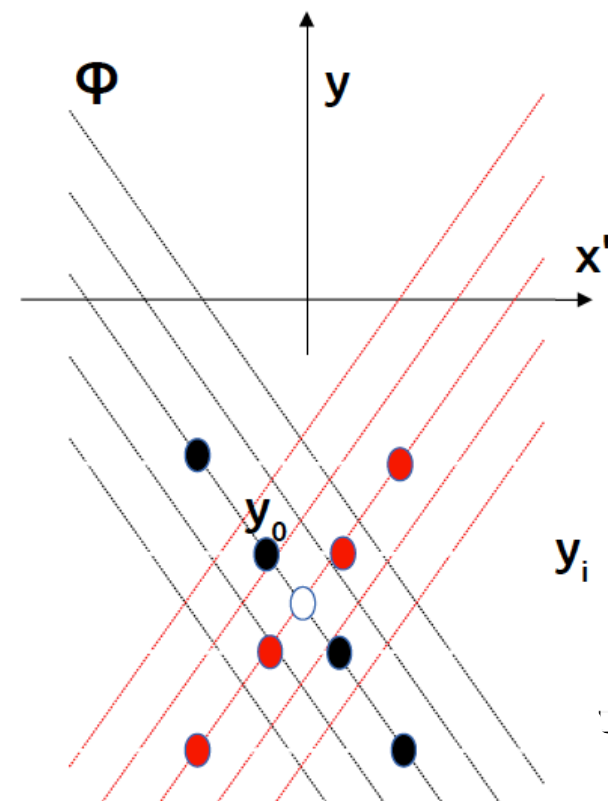
- *wire imaging*

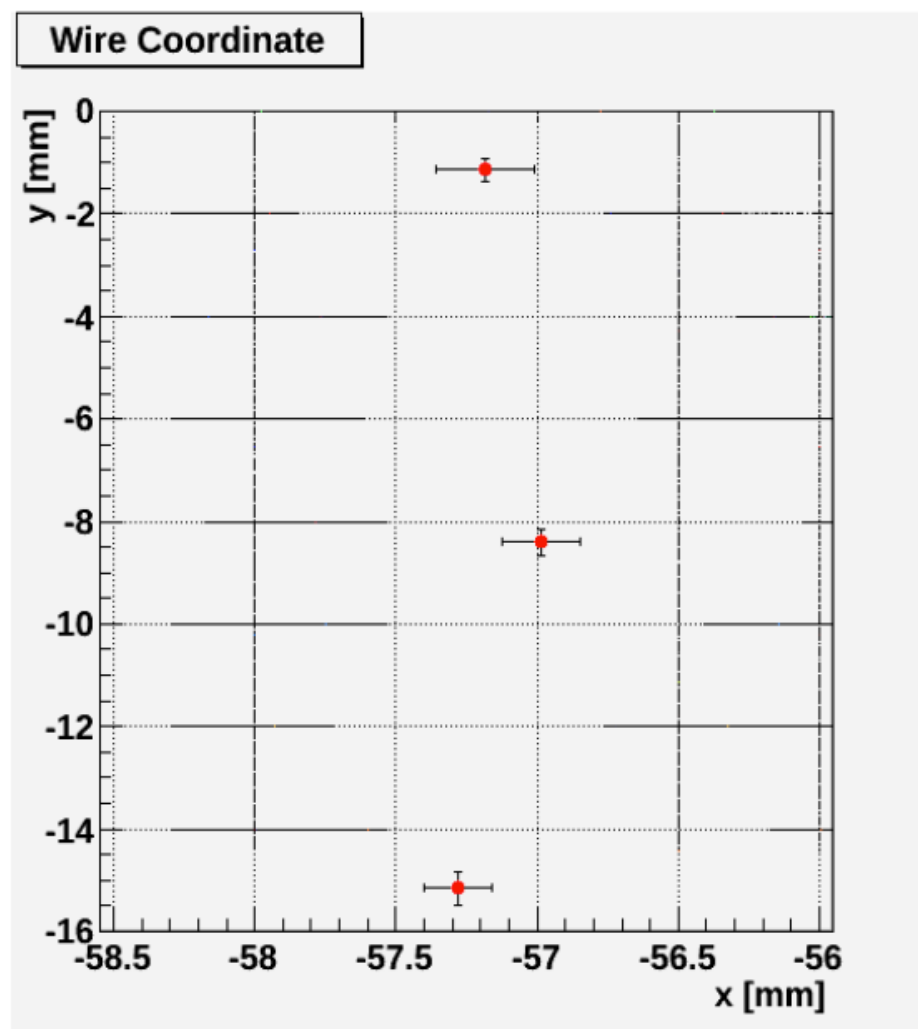
- outliers from noise in the DUT

- *less than 1%*

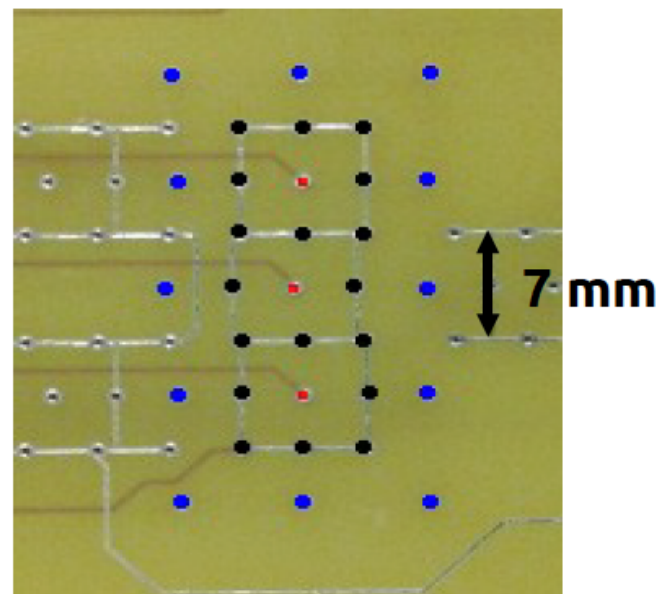


- use of tracks with $\phi = \pm 100$ mrad
- *use of different nominal y planes and fit the $x-t$ relation*
 - the parabola vertex is the wire position
 - the predictions do not agree if the y plane used in the projection is not the real wire position
- *wire y position estimation by crossing point*



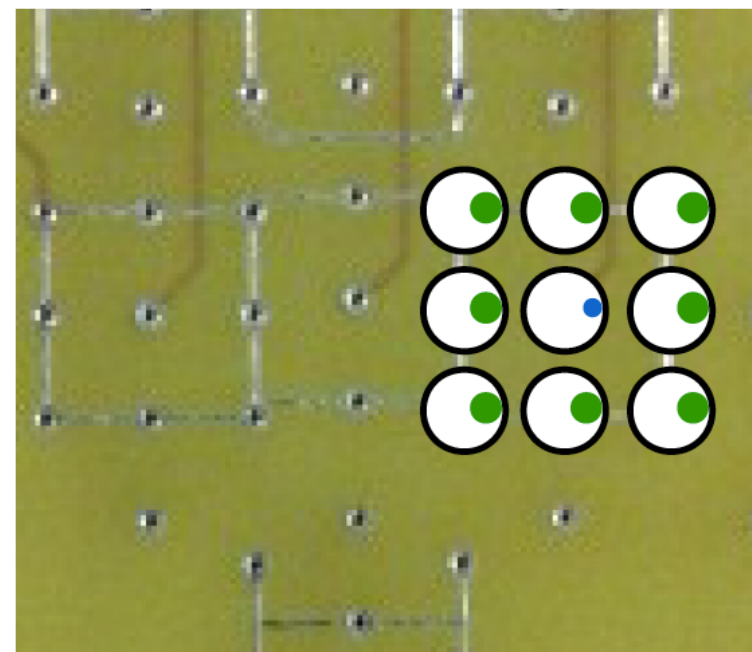
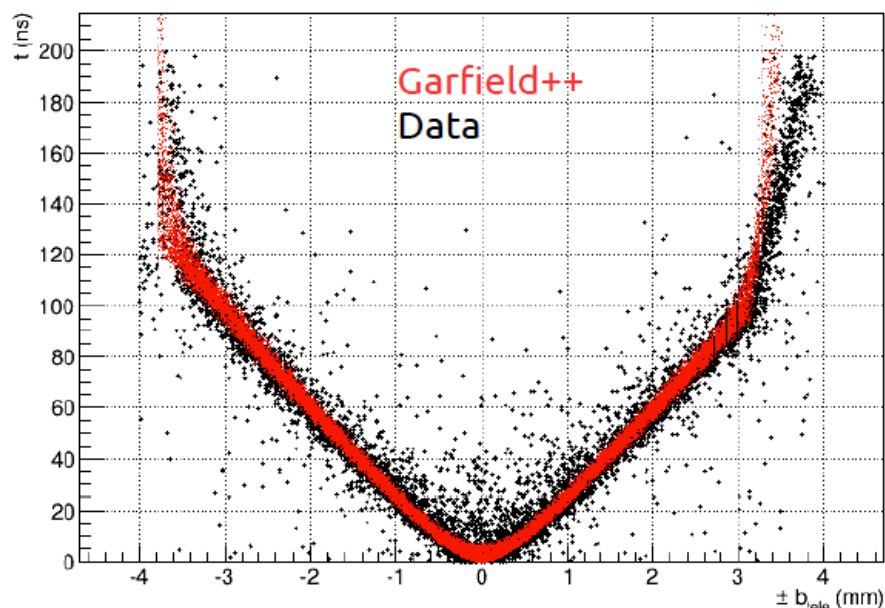


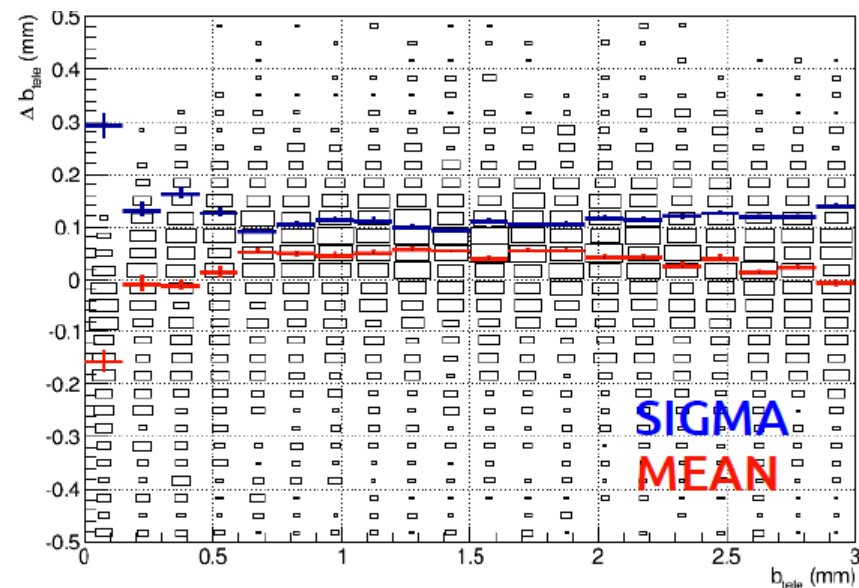
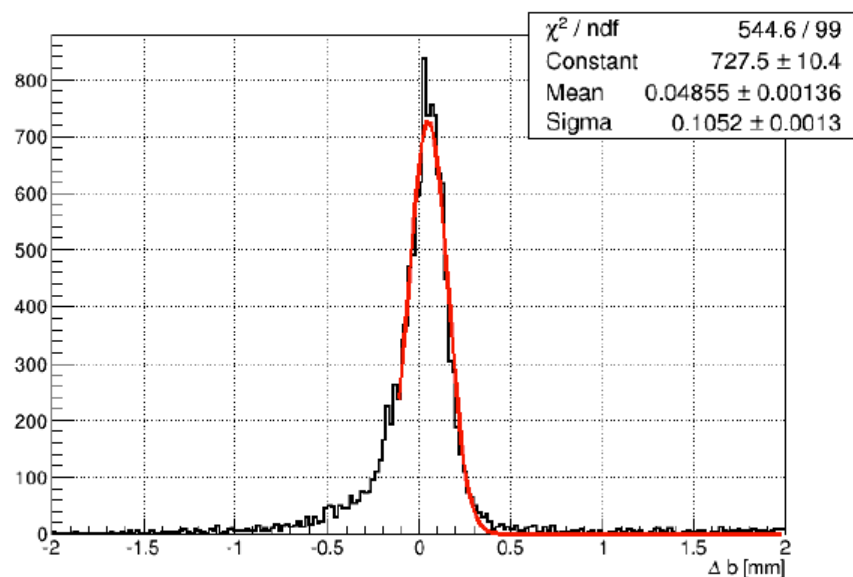
- 7 mm wire distance
 - *and 500 μm stagger*
- now we are confident to measure
 - *x-t relations*
 - *position resolution*



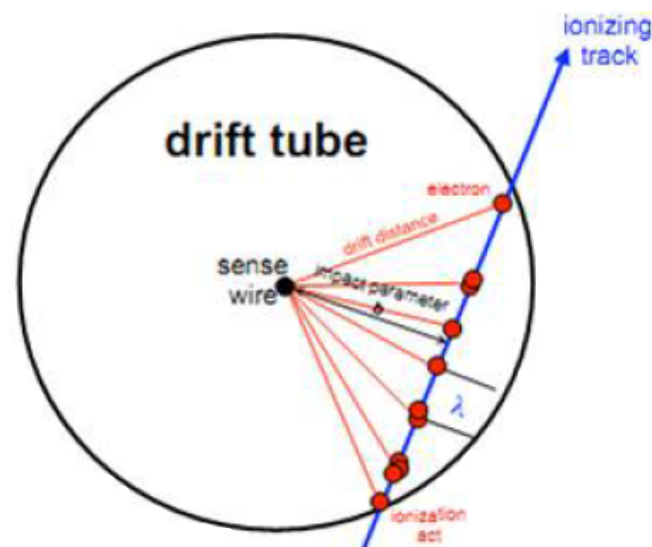
- used to convert time information in position
 - *agreement with garfield++ simulations*
 - *evidence of small cell asymmetries*
- related to wiring procedure

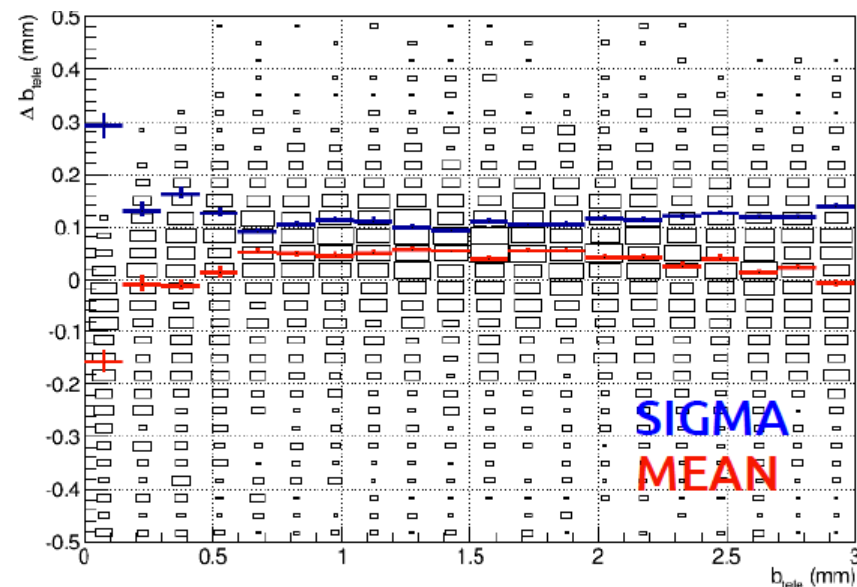
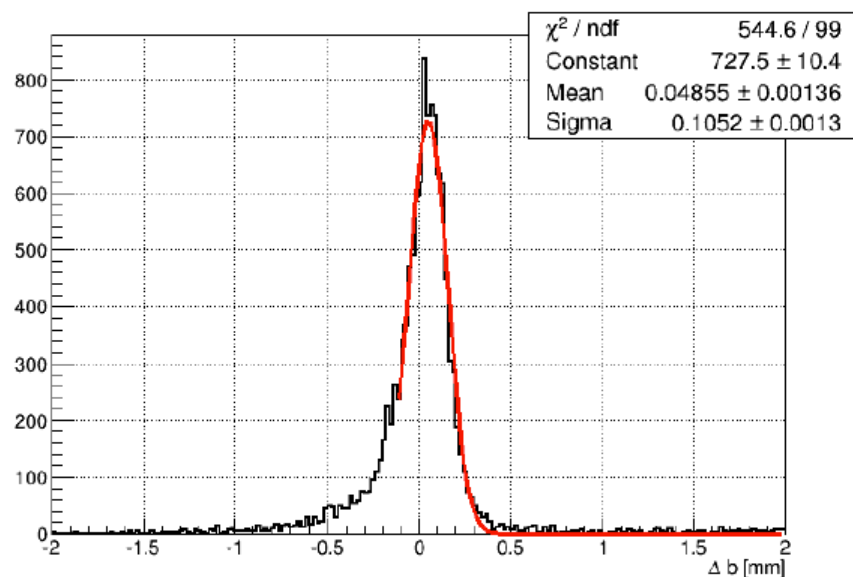
PCB hole: 500 μm
Sense wire: 20 μm
Field wires: 80 μm
(not in scale)



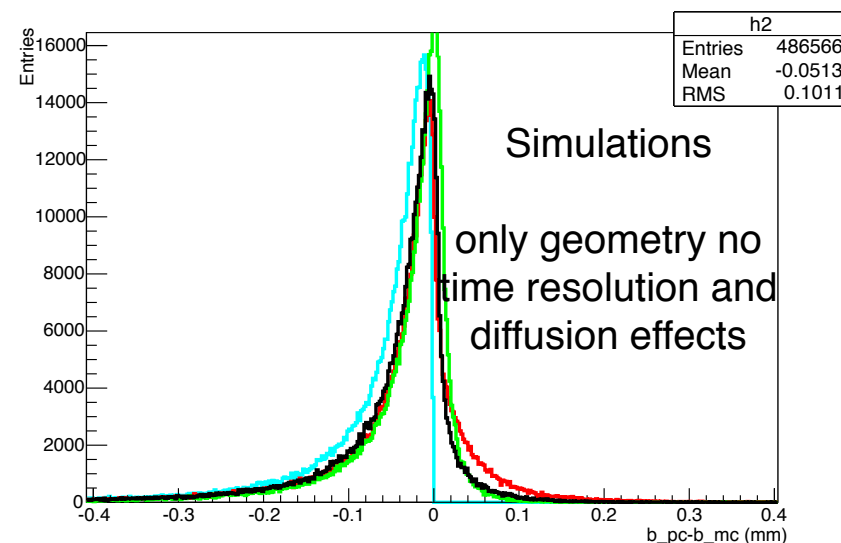


- Use of only first ionisation cluster
 - *evidence of bias for small impact parameters*
 - *resolution of $\sim 110 \mu\text{m}$ as expected*
 - master thesis by A. D'Onofrio to study the resolution improvement by using the information of the other clusters ongoing

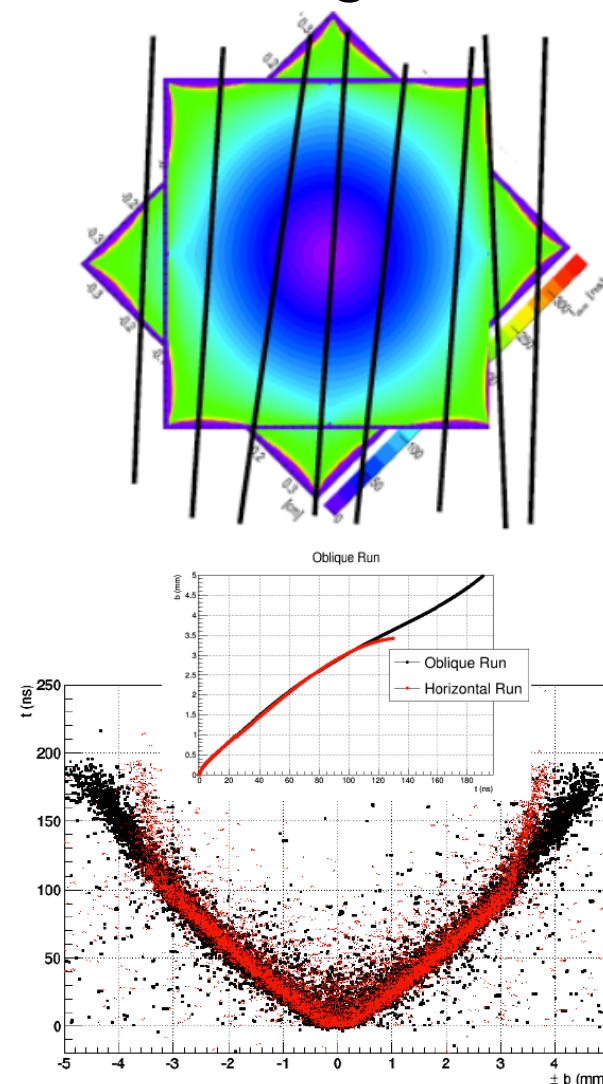
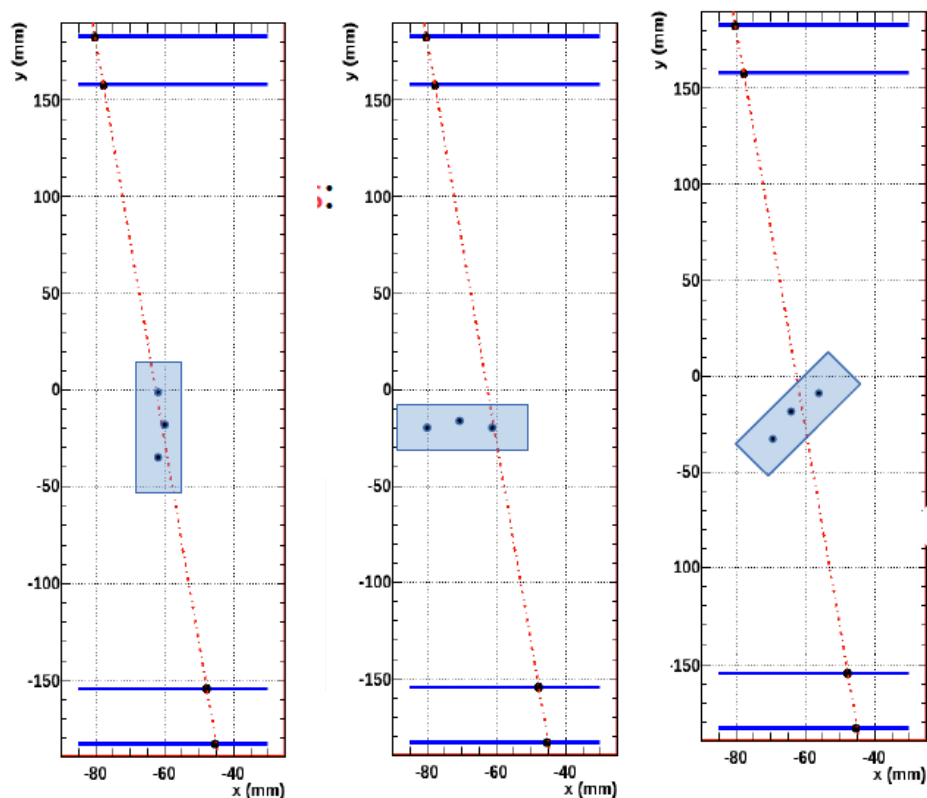


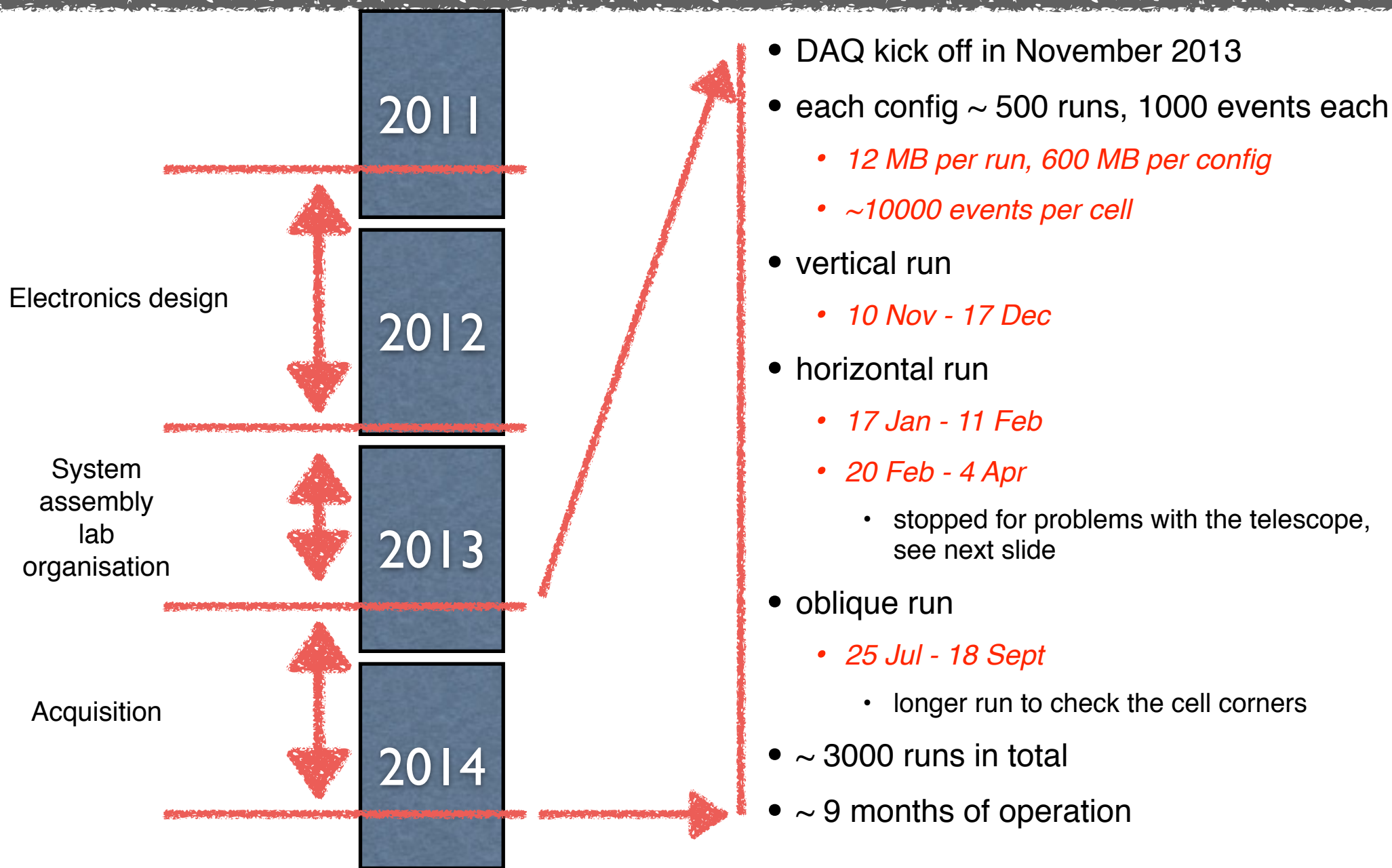


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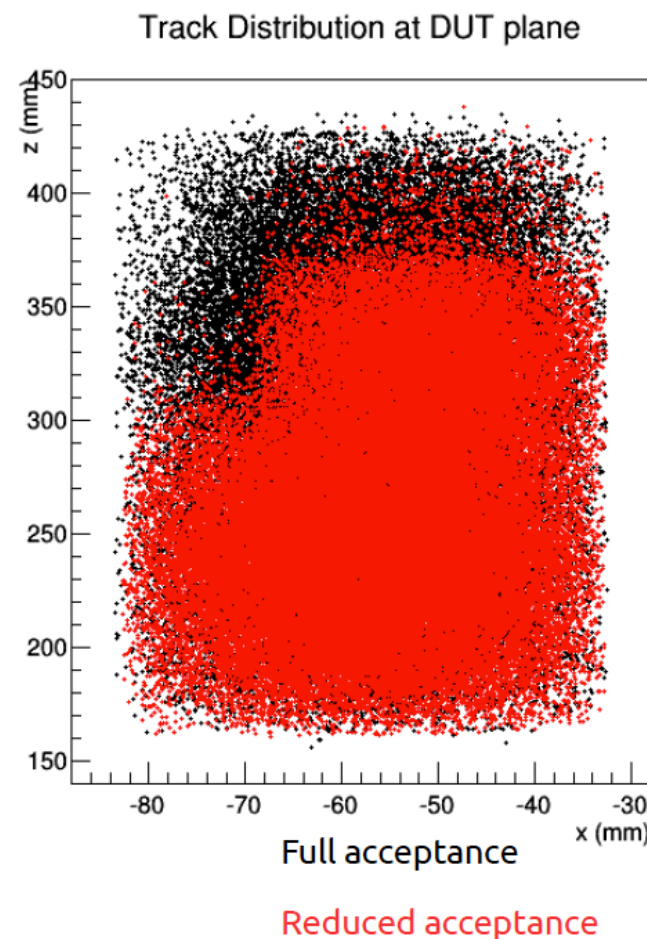


- The measurement was repeated in three configurations
 - *resolution cross check*
 - *x-t relation scan*





- The facility is operative
 - *in one year of operation no sign of performance degradation, but*
 - Suddenly lost of few chips in April
 - 4 chips on a L4 p-side (the whole side) and 2 chips on a L5 n-side (2/5 a side)
 - reason unknown
 - *no temperature variation*
 - *no loss of dry air*
 - *voltage spike?*
 - L4 substituted, L5 still there
- We have few spares for L5 and L4 (~ 4 modules per type)
 - *can be used if needed*
 - *less chips lower acceptance*
 - but telescope usable
 - *we could use 3 hits instead of four with a moderate performance loss*
- New run with MEG II prototype
 - *use of GHz electronics for refined cluster timing studies*
- Lifetime ~ 10 years
 - *responsible: Luca Galli*



- A Si-based cosmic ray telescope as an INFN sezione di Pisa facility
- New TDAQ and electronics to substitute the original from BaBar experiment
 - *no longer usable*
- A MEG II prototype characterised in many configurations
 - *both telescope and DUT resolutions as expected*
- New runs with fast electronics foreseen with MEG II prototypes
 - *the facility is available for any requirements from the groups of this section*
 - Avanzini, Spinella: micro-megas detector?