MEG II drift chamber characterisation with the silicon based cosmic ray tracker at INFN Pisa

Alessandro Baldini^a, Elisabetta Baracchini^d, C. Bemporad^{ab}, Fabrizio Cei^{ab}, Adelina D'Onofrio^{ab}, Simeone Dussoni^a, <u>Luca Galli</u>^a, Marco Grassi^a, Donato Nicolò^{ab}, Giovanni Signorelli^a, Marco Venturini^{ac}

a- Istituto Nazionale di Fisica Nucleare sezione di Pisa
b- Università degli studi di Pisa c- Scuola Normale Superiore
d- ICEPP Tokyo University

Frontier Detectors for Frontier Physics 13th Pisa Meeting on Advanced Detectors 24-30 May 2015 - La Biodola Isola d'Elba





The cosmic ray telescope

In order to compare the performance of many possible small-size tracking prototypes, a high-resolution cosmic ray tracker may be used as an external track reference. We have constructed a telescope with an intrinsic tracking resolution of 15–30 μ m by assembling four spare ladders of the external layers of the Silicon Vertex Tracker of the BaBar experiment[1]. The original DAQ was replaced by custom-designed boards coupled with an acquisition front-end personal computer through commercial field-programmable gate array evaluation boards. The hit detection efficiency is >90% for each module. The tracking performance is better than 40 μ m for detector under test thickness of about 1% X₀[2].



MEG II drift chamber

It is a 2π single volume chamber filled with He-Isobutane (85-15%). The 1920 drift cells in are arranged in a ~7 deg stereo geometry. The wires are read out at both ends by a GHz bandwidth DAQ chain. Many of the ~13/cm ionisation clusters will be identifiable in the digitised waveforms and with cluster timing algorithms. The chamber operates at ~1600 V corresponding to a gain of ~5 x 10⁵ electrons.



Alignment

Iterative alignment into the **telescope frame** with ~150μm **precision**. The **height** comes from the **crossing** of the **wire projections** with **tracks** from **opposite directions**. **Wire imaging** is used for **x-z** view.

Simulation studies show that, given a single hit resolution ~120 μ m, the momentum resolution is expected to be ~130 keV and the angular of ~5 mrad[3].



Drift curve

data

Garfield++

The three-cell prototype

The pattern of three 7 mm squared cells is implemented on two FR4 PCB with the anode wire (25 μ m (Au)W) of the central cell staggered by 500 μ m for self tracking capabilities. It stays in a plexiglas gas tight box having two thin windows (1mm thick) looking at the telescope sensitive area. The geometry and working parameters are equivalent to the final chamber. The sense wire signals are amplified by the DC front-end prototype[4] and digitised by a CAEN V1729, 300 MHz bandwidth, 2 GSPS waveform digitiser.



<u>Results</u>

The **r-t relations** are **extracted** form **data** by using the **external track reference**. The results **are in agreement** with the **Garfield++ predictions**.

The difference between the impact parameter measured by the telescope and the estimation from the first cluster is the single hit resolution. It is fitted by a gaussian distribution plus an exponential tail (starting from a distance δ from μ), the tail models the bias expected in this gas admixture[5]. The measured resolution is ~110 μ m in the whole cell. The bias is present in particular for lower impact parameters.



<u>MEG II</u>

160

t (ns)

140

120

The MEG II goal is to reach a sensitivity of 5 10⁻¹⁴ on the $\mu \rightarrow e\gamma$ branching fraction, a factor 10 better than MEG. In addition to the presented tracker the experiment will be equipped with pixelated Timing Counter detectors, consisting in scintillator tiles read by SiPMs to measure the positron time-of-flight. The γ -detector inherits the LXe detector used during the first phase of the experiment, the only difference being related to the inner face read out: 2" PMTs will be replaced by MPPCs to improve detector uniformity (notably for shallow γ -rays), single photo electron response and imaging capability as well.

Bibliography

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Contact: <u>luca.galli@pi.infn.it</u>