



Consiglio di sezione INFN

Presentazione delle attività relative all'iniziativa specifica QFT-HEP a Lecce

Luigi Delle Rose – Lecce, 21/07/2014

Generalities

- ✓ People:
 - C. Corianò, M. Serino, A. Costantini, C.Marzo, L. D. R.
- ✓ Collaborations:
 - E. Mottola (Los Alamos)
 - L. Trentadue (Parma and INFN Milano)
 - E. Gabrielli (CERN and INFN Trieste)
 - P. Colangelo (INFN Bari)

Publications 2013-2014

- ✓ C. Corianò, A. Costantini, L. D. R., M. Serino
 "Superconformal sum rules and the spectral density flow of the composite dilaton (ADD) multiplet in N=1 theories"
 Published in JHEP 1406 (2014) 136, arXiv:1402.6369 [hep-th]
- ✓ C. Corianò, L. D. R., E. Gabrielli, L. Trentadue "Fermion Scattering in a Gravitational Background: Electroweak Corrections and Flavour Transitions" Published in JHEP 1403 (2014), arXiv:1312.7657 [hep-ph]
- ✓ C. Corianò, L. D. R., C. Marzo, M. Serino
 "The dilaton Wess-Zumino action in six dimensions from Weyl gauging: local anomalies and trace relations"
 Published in Class.Quant.Grav. 31 (2014) 105009, arXiv:1311.1804 [hep-th]
- ✓ C. Corianò, L. D. R., C. Marzo, M. Serino "Conformal Trace Relations from the Dilaton Wess-Zumino Action" Published in Phys.Lett. B726 (2013) 4-5. 896-905, arXiv:1306.4248 [hep-th]

Publications 2013-2014

- ✓ C. Corianò, L. D. R., E. Mottola, M. Serino "Solving the Conformal Constraints for Scalar Operator in Momentum Spaces and the Evaluation of Feynman's Master Integrals" Published in JHEP 1307 (2013) 011, arXiv:1304.6944 [hep-th]
- ✓ C. Corianò, L. D. R., E. Gabrielli, L. Trentadue
 "Mass Corrections to Flavor-Changing Fermion-Graviton Vertices in the Standard Model"
 Published in Phys. Rev. D88 (2013) 085008, arXiv:1303.1305 [hep-th]
- C. Corianò, L. D. R., E. Gabrielli, L. Trentadue
 "One loop Standard Model corrections to flavor diagonal fermion-graviton vertices"
 Bublished in Bhys. Boy. D87 (2012) 05 (020, arXiv:1212.5020 [hop.ph]

Published in Phys. Rev. D87 (2013) 054020, arXiv:1212.5029 [hep-ph]

Generalities

✓ Organization of QCD@Work 2014 --- 16-19 June 2014, Giovinazzo

✓ Thesis

- A. Quintavalle "Dilaton Interactions and the Anomalous Breaking of Scale Invariance of the Standard Model" – Dottorato, Luglio 2013
- L.D.R. "The Standard Model in a Weak Gravitational Background: Dilatons, Scale Anomalies and Conformal Methods" – Dottorato, Luglio 2013
- A. Politano "L'equazione di Dirac in campi esterni" Triennale, 2013
- M. Maglio "Scattering di un fotone in un background gravitazionale" Triennale, 2014
- M. Serino "Conformal anomaly actions and dilaton interactions" Dottorato, Luglio 2014

Outline

- \checkmark Superconformal sum rules and the spectral density flow
- ✓ Fermion interactions in a gravitational background
- ✓ Vacuum stability in abelian extensions of the Standard Model
- ✓ Dilaton Wess-Zumino action from the conformal anomaly

Superconformal sum rules and the spectral density flow

Review: The AVV diagram and the pion

The anomalous AVV diagram with an axial-vector current (A) and two vector currents (V) is characterized by a massless singularity

 $\Delta^{\lambda\mu\nu} = a_n \frac{k^{\lambda}}{k^2} \epsilon^{\mu\nu\alpha\beta} k_{1\alpha} k_{2\beta}$ Dolgov, Zakharov. Nucl. Phys. B27, 525 (1971) The axial anomaly pole appears in the AVV diagram

The pole structure clearly describes the pseudoscalar pion π

the anomalous spectral density in the AVV possesses a sum rule

 $\frac{1}{\pi} \int_{0}^{\infty} \rho(s, m^2) ds = f$ Horejsi. Phys. Rev. D32, 1029 (1985)

this is a general feature of anomalies (chiral, conformal, superconformal)

Sum rules

The existence of a sum rule implies a particular behaviour of the spectral density

we consider a sum rule with a mass deformation parameter m to control the spectral density away from the chiral/conformal point

 $\frac{1}{\pi}\int_0^\infty \rho(s,m^2)ds = f$

with f ≠ 0 and mass independent

- the spectral density is integrable
- its scale dimension is fixed
- ρ (s, m²) flows towards a δ (s) as m goes to zero
- a pole-like behaviour appears also in the UV

anomalies are IR/UV phenomena $\lim_{k^2 \to \infty} k^2 F(k^2, m^2) = f$

this corresponds to a massless state propagating in the theory

The conformal anomaly case

- Conformally anomalous correlators possess spectral densities with a convergent sum rule
- As the AVV case, it has been shown that the TVV correlator, (T is the energy-momentum tensor EMT) shares a similar behaviour dictated by the conformal anomaly this has been proved in QED, QCD and in EW sector of the SM

this can be interpreted as an effective scalar, the **dilaton**, coupling to the trace of the EMT, in full analogy with the pion

exchange of a conformal anomaly pole

In the following the superconformal anomaly case is considered, in which chiral, conformal and superconformal symmetries are treated in a unified way

Anomalies in N=1 Super Yang-Mills

Consider a N=1 Super Yang-Mills (for instance SuperQCD)

vector (gauge) supermultiplet chiral (matter) supermultiplet $V = (A^a_{\mu}, \lambda^a, D^a)$ $\Phi = (\phi_i, \chi_i, F_i)$

we can introduce the Ferrara-Zumino (FZ) hypercurrent

$$\mathcal{J}_{A\dot{A}} = \operatorname{Tr}\left[\bar{W}_{\dot{A}}e^{V}W_{A}e^{-V}\right] - \frac{1}{3}\bar{\Phi}\left[\stackrel{\leftarrow}{\bar{\nabla}}_{\dot{A}}e^{V}\nabla_{A} - e^{V}\bar{D}_{\dot{A}}\nabla_{A} + \stackrel{\leftarrow}{\bar{\nabla}}_{\dot{A}}\stackrel{\leftarrow}{D}_{A}e^{V}\right]\Phi$$
$$W_{A} = 2gW_{A}^{a}T^{a} = -\frac{1}{4}\bar{D}^{2}e^{-V}D_{A}e^{V}$$

which describes a chiral supermultiplet containing the **EMT**, the **R**-current and the supersymmetric current $\mathcal{J} = (R^{\mu}, T^{\mu\nu}, S^{\mu})$

- in a classical superconformal theory the FZ hypercurrent is conserved $ar{D}^{\dot{A}}\mathcal{J}_{A\dot{A}}=0$

the chiral superpotential ${\mathcal W}$ must be cubic or vanishing



Anomalies in N=1 Super Yang-Mills

in the quantum mechanical framework, the FZ develops a superconformal anomaly

$$\bar{D}^{\dot{A}}\mathcal{J}_{A\dot{A}} = a_n D_A W^2$$
 $a_n = -\frac{2}{3}g^2 \frac{3T_G - \sum_f T(R_f)}{16\pi^2}$

The components of the FZ supermultiplet are

$$\begin{split} R^{\mu} &= \bar{\lambda}^{a} \bar{\sigma}^{\mu} \lambda^{a} + \frac{1}{3} \left(-\bar{\chi}_{i} \bar{\sigma}^{\mu} \chi_{i} + 2i \phi_{i}^{\dagger} \mathcal{D}_{ij}^{\mu} \phi_{j} - 2i (\mathcal{D}_{ij}^{\mu} \phi_{j})^{\dagger} \phi_{i} \right), \\ S^{\mu}_{A} &= i (\sigma^{\nu \rho} \sigma^{\mu} \bar{\lambda}^{a})_{A} F^{a}_{\nu \rho} - \sqrt{2} (\sigma_{\nu} \bar{\sigma}^{\mu} \chi_{i})_{A} (\mathcal{D}_{ij}^{\nu} \phi_{j})^{\dagger} - i \sqrt{2} (\sigma^{\mu} \bar{\chi}_{i}) \mathcal{W}_{i}^{\dagger} (\phi^{\dagger}) \\ &- i g (\phi_{i}^{\dagger} T_{ij}^{a} \phi_{j}) (\sigma^{\mu} \bar{\lambda}^{a})_{A} + S^{\mu}_{IA}, \\ T^{\mu\nu} &= -F^{a\,\mu\rho} F^{a\,\nu}{}_{\rho} + \frac{i}{4} \left[\bar{\lambda}^{a} \bar{\sigma}^{\mu} (\delta^{ac} \vec{\partial}^{\nu} - g t^{abc} A^{b\,\nu}) \lambda^{c} + \bar{\lambda}^{a} \bar{\sigma}^{\mu} (-\delta^{ac} \vec{\partial}^{\nu} - g t^{abc} A^{b\,\nu}) \lambda^{c} + (\mu \leftrightarrow \nu) \right] \\ &+ (\mathcal{D}^{\mu}_{ij} \phi_{j})^{\dagger} (\mathcal{D}^{\nu}_{ik} \phi_{k}) + (\mathcal{D}^{\nu}_{ij} \phi_{j})^{\dagger} (\mathcal{D}^{\mu}_{ik} \phi_{k}) + \frac{i}{4} \left[\bar{\chi}_{i} \bar{\sigma}^{\mu} (\delta_{ij} \vec{\partial}^{\nu} + i g T^{a}_{ij} A^{a\,\nu}) \chi_{j} \right. \\ &+ \left. \bar{\chi}_{i} \bar{\sigma}^{\mu} (-\delta_{ij} \vec{\partial^{\nu}} + i g T^{a}_{ij} A^{a\,\nu}) \chi_{j} + (\mu \leftrightarrow \nu) \right] - \eta^{\mu\nu} \mathcal{L} + T^{\mu\nu}_{I}, \\ &\text{ and satisfy the anomaly equations} \\ \left\{ \begin{array}{c} \partial_{\mu} R^{\mu} &= -\frac{a_{n}}{2} F^{a}_{\mu\nu} \tilde{F}^{a\,\mu\nu} \\ T^{\mu}_{\mu} &= \frac{3}{4} a_{n} F^{a}_{\mu\nu} F^{a\,\mu\nu} \\ \bar{\sigma}_{\mu} S^{\mu} &= 3i a_{n} \bar{\lambda}^{a} \bar{\sigma}^{\mu\nu} F^{a}_{\mu\nu} \end{array} \right. \right.$$



Perturbative computation

We look at the three correlation functions responsible for the appearance of the superconformal anomaly

$\delta^{ab} \Gamma^{\mulphaeta}_{(R)}(p,q)$	\equiv	$\langle R^{\mu}(k) A^{a\alpha}(p) A^{b\beta}(q) \rangle$	$\langle RVV \rangle$,
$\delta^{ab}\Gamma^{\mu\alpha}_{(S)A\dot{B}}(p,q)$	≡	$\langle S^{\mu}_{A}(k) A^{a\alpha}(p) \bar{\lambda}^{b}_{\dot{B}}(q) \rangle$	$\langle SVF \rangle$,
$\delta^{ab} \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q)$	≡	$\langle T^{\mu\nu}(k) A^{a\alpha}(p) A^{b\beta}(q) \rangle$	$\langle TVV \rangle$

they are constrained by the vector current conservation

 $p_{\alpha} \Gamma^{\mu\alpha\beta}_{(R)}(p,q) = 0, \quad q_{\beta} \Gamma^{\mu\alpha\beta}_{(R)}(p,q) = 0,$ $p_{\alpha} \Gamma^{\mu\alpha}_{(S)}(p,q) = 0,$ $p_{\alpha} \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q) = 0, \quad q_{\beta} \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q) = 0$

and by the conservation of the supercurrent and of the EMT

$$i k_{\mu} \Gamma^{\mu\alpha}_{(S)}(p,q) = -2p_{\mu} \sigma^{\mu\alpha} \hat{\Gamma}_{(\lambda\bar{\lambda})}(q) - i\sigma_{\mu} \hat{\Gamma}^{\mu\alpha}_{(AA)}(p) ,$$

$$i k_{\mu} \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q) = q_{\mu} \hat{\Gamma}^{\alpha\mu}_{(AA)}(p) \eta^{\beta\nu} + p_{\mu} \hat{\Gamma}^{\beta\mu}_{(AA)}(q) \eta^{\alpha\nu} - q^{\nu} \hat{\Gamma}^{\alpha\beta}_{(AA)}(p) - p^{\nu} \hat{\Gamma}^{\alpha\beta}_{(AA)}(q)$$





Perturbative computation

Samples of the one-loop perturbative expansion of the three anomalous correlators

 $A^{a\,\alpha}(p)$

 $A^{b\,\beta}(q)$

 $A^{a \alpha}(p)$

 $\bar{\lambda}^b_{\dot{B}}(q)$

 $A^{a\,\alpha}(p)$

 $A^{b\,\beta}(q)$





Explicit results

- We show the chiral matter contribution in the on-shell gauge currents kinetimatic region
- A massive chiral supermultiplet is employed to stay away from the conformal point and to study the flow of the spectral densities

$$\begin{split} \Gamma^{\mu\alpha\beta}_{(R)}(p,q) &= i \frac{g^2 T(R)}{12\pi^2} \Phi_1(k^2,m^2) \frac{k^{\mu}}{k^2} \varepsilon[p,q,\alpha,\beta] \,, \\ \Gamma^{\mu\alpha}_{(S)}(p,q) &= i \frac{g^2 T(R)}{6\pi^2 k^2} \Phi_1(k^2,m^2) \, s_1^{\mu\alpha} + i \frac{g^2 T(R)}{64\pi^2} \Phi_2(k^2,m^2) \, s_2^{\mu\alpha} \,, \\ \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q) &= \frac{g^2 T(R)}{24\pi^2 k^2} \Phi_1(k^2,m^2) \, t_{1S}^{\mu\nu\alpha\beta}(p,q) + \frac{g^2 T(R)}{16\pi^2} \Phi_2(k^2,m^2) \, t_{2S}^{\mu\nu\alpha\beta}(p,q) \end{split}$$



Explicit results

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$$\begin{split} \Gamma^{\mu\alpha\beta}_{(R)}(p,q) &= i \frac{g^2 T(R)}{12\pi^2} \Phi_1(k^2,m^2) \frac{k^{\mu}}{k^2} \varepsilon[p,q,\alpha,\beta] & \text{superconformal anomaly poles} \\ \Gamma^{\mu\alpha}_{(S)}(p,q) &= i \frac{g^2 T(R)}{6\pi^2 k^2} \Phi_1(k^2,m^2) \frac{k^{\mu}}{k^2} + i \frac{g^2 T(R)}{64\pi^2} \Phi_2(k^2,m^2) \frac{k^{\mu\alpha}}{k^2}, \\ \Gamma^{\mu\nu\alpha\beta}_{(T)}(p,q) &= \frac{g^2 T(R)}{24\pi^2 k^2} \Phi_1(k^2,m^2) \frac{k^{\mu\nu\alpha\beta}}{k^3}(p,q) + \frac{g^2 T(R)}{16\pi^2} \Phi_2(k^2,m^2) \frac{k^{\mu\nu\alpha\beta}}{k^3}(p,q) \end{split}$$

• The first form factor is the anomaly contribution

$$\begin{split} \Phi_1(k^2, m^2) &= -1 - 2 \, m^2 \, \mathcal{C}_0(k^2, m^2) \,, \\ \Phi_2(k^2, m^2) &= 1 - \mathcal{B}_0(0, m^2) + \mathcal{B}_0(k^2, m^2) + 2m^2 \mathcal{C}_0(k^2, m^2) \end{split}$$
 this is interpreted as the residue of the pole

Anomalous spectral density

Spectral densities can be computed using cutting rules or exploiting directly the analytic continuations of the two- and three-point scalar integrals

The spectral density of the anomalous form factor

$$\rho_{\chi}(s,m^2) = \frac{1}{2i} \text{Disc } \chi(s,m^2) = \frac{2\pi m^2}{s^2} \log \left(\frac{1+\sqrt{\tau(s,m^2)}}{1-\sqrt{\tau(s,m^2)}}\right) \theta(s-4m^2) \\ \tau(k^2,m^2) = \sqrt{1-4m^2/k^2}$$
Representatives of the family of spectral densities plotted versus s in units of m².
The family flows towards the s=0 region becoming a $\delta(s)$ function as m² goes to zero.

 $\chi(k^2, m^2) \equiv \Phi_1(k^2, m^2)/k^2$

Anomalous spectral density

• The spectral density is integrable and satisfies a convergent sum rule

$$\frac{1}{\pi} \int_{4m^2}^{\infty} ds \rho_{\chi}(s, m^2) = 1$$

the anomaly coefficient has been factorized for convenience

• The discontinuity of the anomalous form factor is characterized by a cut for $k^2 > 4 m^2$.

There is any resonant state of the pole at $k^2=0$ in the massive case (decoupling)

• The spectral density flows towards a $\delta(s)$ as m goes to zero $\lim_{m \to 0} \frac{1}{\pi} \rho_{\chi}(s, m^2) = \delta(s)$

anomaly accounts for the appearance of massless states in the spectrum, one for each component of the superconformal hypercurrent



Non-anomalous spectral density

The non-anomalous form factor Φ_2

- needs a subtraction for its integrability (is affected by renormalization)
- does not possess a sum rule
- tends to a uniform distribution as m goes to zero

$$\frac{1}{\pi} \lim_{m \to 0} \rho_{\Phi_2}(k^2, m^2) = 1$$

there is a sort of *duality* between the two spectral densities for m->0





Conformal anomaly in QCD

The anomalous correlators (TVV) expands onto three form factors

$$\Gamma^{\mu\nu\alpha\beta}(p,q) = \Gamma^{\mu\nu\alpha\beta}_{q}(p,q) + \Gamma^{\mu\nu\alpha\beta}_{g}(p,q) \qquad \Gamma^{\mu\nu\alpha\beta}_{q/g}(p,q) = \sum_{i=1}^{3} \Phi_{i\,q/g}(k^2,m^2)\,\phi^{\mu\nu\alpha\beta}_{i}(p,q)$$

The explicit results in quark sector are

$$\Phi_{1q}(k^2, m^2) = \frac{g^2}{6\pi^2 k^2} \left\{ -\frac{1}{6} + \frac{m^2}{k^2} - m^2 \mathcal{C}_0(k^2, m^2) \left[\frac{1}{2} - \frac{2m^2}{k^2} \right] \right\}, \quad \text{anomalous}$$

$$\Phi_{2q}(k^2, m^2) = -\frac{g^2}{4\pi^2 k^2} \left\{ \frac{1}{72} + \frac{m^2}{6k^2} + \frac{m^2}{2k^2} \mathcal{D}(k^2, m^2) + \frac{m^2}{3} \mathcal{C}_0(k^2, m^2) \left[\frac{1}{2} + \frac{m^2}{k^2} \right] \right\}, \quad \text{anomalous}$$

$$\Phi_{3q}(k^2, m^2) = \frac{g^2}{4\pi^2} \left\{ \frac{11}{72} + \frac{m^2}{2k^2} + m^2 \mathcal{C}_0(k^2, m^2) \left[\frac{1}{2} + \frac{m^2}{k^2} \right] + \frac{5m^2}{6k^2} \mathcal{D}(k^2, m^2) + \frac{1}{6} \mathcal{B}_0^{\overline{MS}}(k^2, m^2) \right\} \right] \quad \text{anomalous}$$

$$\rho_{2q}(s,m^2) = -\frac{g^2}{24\pi} \frac{m^2}{s^2} \left[3\sqrt{\tau(s,m^2)} - \left(1 + \frac{2m^2}{s}\right) \log \frac{1 + \sqrt{\tau(s,m^2)}}{1 - \sqrt{\tau(s,m^2)}} \right] \theta(s - 4m^2)$$

This spectral density is not anomalous, but is integrable and with a convergent sum rule

another massless state in the spectrum?

Extra pole cancellation in susy theories

Supersymmetric theories force the extra pole to cancel

In a general Yang-Mills theory the contribution to the non-anomalous form factor with the pole-like behaviour is

$$f_{2}(k^{2}) = \frac{N_{f}}{2} f_{2}^{(f)}(k^{2}) + N_{s} f_{2}^{(s)}(k^{2}) + N_{A} f_{2}^{(A)}(k^{2})$$
 Bosonic and fermionic contributions have

$$= \frac{g^{2}}{144\pi^{2}k^{2}} \left[-\frac{N_{f}}{2}T(R_{f}) + N_{s} \frac{T(R_{s})}{2} + N_{A} \frac{T(A)}{2} \right]$$
 opposite signs

 $f_2^{(f)}$ for a Weyl fermion, $f_2^{(s)}$ for a complex scalar, $f_2^{(A)}$ for a gauge vector

- chiral supermultiplet $N_f = 1, N_s = 1, N_A = 0$ $T(R_f) = T(R_s)$
- vector supermultiplet $N_f = 1, N_s = 0, N_A = 1$ $T(R_f) = T(A)$

there is no ambiguity in a susy theory





Remarks

- ✓ Analysis of chiral, conformal and superconformal anomalies and their sum rules in perturbation theory
- ✓ Anomalous spectral densities flow towards massless states
- ✓ Axion-dilatino supermultiplet naturally emerges from the quantum breaking of the superconformal symmetry
- Unambiguous interpretation in a supersymmetric context:
 one-to-one correspondece between anomalies and anomaly poles

Fermion interactions in a gravitational background

Fermion interactions in a gravitational background

- One-loop Standard Model (QCD & EW) corrections to <Tf_if_j> in the flavor diagonal and off-diagonal sector
 - classification of the operators
 - computation of the form factors (including the full mass dependece)
 - computation of the deflection angle of a fermion in a gravitational background (gravitational lensing of neutrino)



Fermion interactions in a gravitational background

- Compute the SM effective action at one-loop order in a dilaton
 and a gravitational background
- Inspection of the electroweak corrections to the Newton's potential
- Scattering of a fermion in an external gravitational field generated by heavy sources (star, black hole, galaxy)
- Analysis of the gravitational interactions of neutrinos (differences between Dirac and Majorana neutrinos?)

Vacuum stability in abelian extensions of the Standard Model (in preparation)



Effective potential:



Vacuum stability in abelian extensions of SM

- The requirement of vacuum stability of a SM extension, up to a given scale (GUT or Planck), can be used to constrain its parameter space
- We are analyzing the stability problem in one of the simplest extensions of the SM $SU(3)_c imes SU(2)_w imes U(1)_Y imes U(1)'$
- The spectrum of the SM is augmented by new states:
 - ✓ A gauge boson Z'
 - A gauge boson 2
 A complex scalar φ to achieve U(1)' SSB (vev ~ TeV)

explains small neutrino masses through seesaw

✓ SM singlet fermions to cancel U(1)' anomalies

Vacuum stability in abelian extensions of SM

• Consider for instance the B-L (Baryon – Lepton numbers) extension

 \checkmark The type-I seesaw is implemented in the Yukawa sector

$$\mathcal{L}_{yuk} = \mathcal{L}_{SM\,yuk} - Y_{\nu} L \cdot H\nu_R^c - Y_N \phi^{\dagger} \nu_R^c \nu_R^c$$

 \checkmark The scalar potential is given by

$$V(H,\phi) = m_1^2 H^{\dagger} H + m_2^2 \phi^{\dagger} \phi + \lambda_1 (H^{\dagger} H)^2 + \lambda_2 (\phi^{\dagger} \phi)^2 + \lambda_3 (H^{\dagger} H) (\phi^{\dagger} \phi)$$

 $\checkmark\,$ The vacuum stability conditions are

$$\lambda_1 > 0, \quad \lambda_2 > 0, \quad 4\lambda_1\lambda_2 - \lambda_3^2 > 0$$

RGEs are employed to study vacuum stability up to the Planck scale

RGE@NLO

Mathematica package for the automatic computation and the numerical solution of the two-loop RGE in a general QFT

- Input: Gauge group (SU, SO, Sp, G2, F4, E6, E7, E8),
 Field (Superfield) content, (Irreps may be specified by Dynkin labels)
 Potential (Superpotential)
- Output: β functions and anomalous dimensions at two-loop order results are saved in Mathematica and LaTeX files a code for the numerical analysis is generated



(GeV)

0.02

0.00

- 0.04

- 0.06 L_____ 100

105

Some preliminary results

- m_{v_h} = 100 GeV

---- m_{v_h} = 1000 GeV

 $.... m_{v_h} = 1500 \, \text{GeV}$

1017

10¹⁴



1011

Q(GeV)



Regions in which the stability conditions are preserved up to 10⁵ GeV (blue) 10⁹ GeV (green) 10¹⁵ GeV (yellow) 1019 GeV (red)

mh2 - mass of the heaviest scalar $\boldsymbol{\theta}$ - mixing angle in the scalar sector mvh - mass of the heavy neutrinos

Remarks

- The requirement of vacuum stability may constrain the parameter space of a model
- ✓ Different seesaw mechanisms can be analyzed in the context of vacuum stability and RG evolution
- ✓ The complete two-loop analysis requires the computation of effective potential, matching conditions at the EW scale and possible threshold corrections



Backup slides

The N=1 supersymmetric Lagrangian in the component formalism is

$$= -\frac{1}{4}F^{a}_{\mu\nu}F^{a\,\mu\nu} + i\lambda^{a}\sigma^{\mu}\mathcal{D}^{ab}_{\mu}\bar{\lambda}^{b} + (\mathcal{D}^{\mu}_{ij}\phi_{j})^{\dagger}(\mathcal{D}_{ik\,\mu}\phi_{k}) + i\chi_{j}\sigma_{\mu}\mathcal{D}^{\mu\dagger}_{ij}\bar{\chi}_{i}$$
$$-\sqrt{2}g\left(\bar{\lambda}^{a}\bar{\chi}_{i}T^{a}_{ij}\phi_{j} + \phi^{\dagger}_{i}T^{a}_{ij}\lambda^{a}\chi_{j}\right) - V(\phi,\phi^{\dagger}) - \frac{1}{2}\left(\chi_{i}\chi_{j}\mathcal{W}_{ij}(\phi) + h.c.\right)$$

The component expansion of the superfields are

 \mathcal{L}

$$\Phi_{i} = \phi_{i} + \sqrt{2}\theta\chi_{i} + \theta^{2}F_{i}$$

$$W_{A}^{a} = \lambda_{A}^{a} + \theta_{A}D^{a} - (\sigma^{\mu\nu}\theta)_{A}F_{\mu\nu}^{a} + i\theta^{2}\sigma_{A\dot{B}}^{\mu}\mathcal{D}_{\mu}\bar{\lambda}^{a\dot{B}},$$

$$V^{a} = \theta\sigma^{\mu}\bar{\theta}A_{\mu}^{a} + \theta^{2}\bar{\theta}\bar{\lambda}^{a} + \bar{\theta}^{2}\theta\lambda^{a} + \frac{1}{2}\theta^{2}\bar{\theta}^{2}(D^{a} + i\partial_{\mu}A^{a\mu})$$

The terms of improvement, necessary only for a scalar field, are

$$\begin{split} S^{\mu}_{IA} &= \frac{4\sqrt{2}}{3}i\left[\sigma^{\mu\nu}\partial_{\nu}(\chi_{i}\phi^{\dagger}_{i})\right]_{A} & \text{These ensure the vanishing of the classical trace of the EMT and of the gamma-trace of the supercurrent for a classical superconformal theory} \end{split}$$

The two-point functions appearing in the Ward identities are

$$\begin{split} \Gamma^{(AA)}_{\mu\nu}(p) &= -i\delta^{ab}\left(\eta_{\mu\nu}-\frac{p_{\mu}p_{\nu}}{p^2}\right)\Sigma^{(AA)}(p^2) \\ \Gamma^{(\lambda\bar{\lambda})}_{A\dot{B}}(p) &= i\delta^{ab}\,p_{\mu}\sigma^{\mu}_{A\dot{B}}\,\Sigma^{(\lambda\bar{\lambda})}(p^2)\,, \end{split}$$

with

$$\Sigma^{(AA)}(p^2) = \frac{g^2}{16\pi^2} p^2 \left\{ T(R) \mathcal{B}_0(p^2, m^2) - T(A) \mathcal{B}_0(p^2, 0) \right\}$$

$$\Sigma^{(\lambda\bar{\lambda})}(p^2) = \frac{g^2}{16\pi^2} \left\{ T(R) \mathcal{B}_0(p^2, m^2) + T(A) \mathcal{B}_0(p^2, 0) \right\}.$$

Two- and three-point scalar integrals

$$\mathcal{B}_{0}(p_{1}^{2}, m_{0}^{2}, m_{1}^{2}) = \frac{1}{i\pi^{2}} \int d^{n}l \frac{1}{(l^{2} - m_{0}^{2})((l + p_{1})^{2} - m_{1}^{2})},$$

$$\mathcal{C}_{0}((p + q)^{2}, p^{2}, q^{2}, m_{0}^{2}, m_{1}^{2}, m_{2}^{2}) = \frac{1}{i\pi^{2}} \int d^{n}l \frac{1}{(l^{2} - m_{0}^{2})((l - p)^{2} - m_{1}^{2})((l - p - q)^{2} - m_{2}^{2})}$$

$$\mathcal{B}_{0}(p_{1}^{2}, m^{2}) \equiv \mathcal{B}_{0}(p_{1}^{2}, m^{2}, m^{2}) \qquad \qquad \mathcal{C}_{0}((p + q)^{2}, m^{2}) \equiv \mathcal{C}_{0}((p + q)^{2}, 0, 0, m^{2}, m^{2}, m^{2})$$



$$\begin{split} s_1^{\mu\alpha} &= \sigma^{\mu\nu}k_\nu \,\sigma^\rho k_\rho \,\bar{\sigma}^{\alpha\beta}p_\beta \\ s_2^{\mu\alpha} &= 2p_\beta \,\sigma^{\alpha\beta}\sigma^\mu \,. \\ t_{1S}^{\mu\nu\alpha\beta}(p,q) &\equiv \phi_1^{\mu\nu\alpha\beta}(p,q) = (\eta^{\mu\nu}k^2 - k^\mu k^\nu) u^{\alpha\beta}(p,q) \,, \\ t_{2S}^{\mu\nu\alpha\beta}(p,q) &\equiv \phi_3^{\mu\nu\alpha\beta}(p,q) = (p^\mu q^\nu + p^\nu q^\mu)\eta^{\alpha\beta} + p \cdot q(\eta^{\alpha\nu}\eta^{\beta\mu} + \eta^{\alpha\mu}\eta^{\beta\nu}) - \eta^{\mu\nu}u^{\alpha\beta}(p,q) \\ &- (\eta^{\beta\nu}p^\mu + \eta^{\beta\mu}p^\nu)q^\alpha - (\eta^{\alpha\nu}q^\mu + \eta^{\alpha\mu}q^\nu)p^\beta \,, \end{split}$$

tensor structures in the qcd computation

$$\begin{split} \phi_1^{\mu\nu\alpha\beta}(p,q) &\equiv t_1^{\mu\nu\alpha\beta}(p,q) = (k^2\eta^{\mu\nu} - k^{\mu}k^{\nu})u^{\alpha\beta}(p,q) \,, \\ \phi_2^{\mu\nu\alpha\beta}(p,q) &\equiv t_3^{\mu\nu\alpha\beta}(p,q) + t_5^{\mu\nu\alpha\beta}(p,q) - 4t_7^{\mu\nu\alpha\beta}(p,q) = -2u^{\alpha\beta}(p,q)[k^2\eta^{\mu\nu} + 2(p^{\mu}p^{\nu} + q^{\mu}q^{\nu}) \\ &- 4(p^{\mu}q^{\nu} + q^{\mu}p^{\nu})] \,, \\ \phi_3^{\mu\nu\alpha\beta}(p,q) &\equiv t_{13}^{\mu\nu\alpha\beta}(p,q) = (p^{\mu}q^{\nu} + p^{\nu}q^{\mu})\eta^{\alpha\beta} + p \cdot q(\eta^{\alpha\nu}\eta^{\beta\mu} + \eta^{\alpha\mu}\eta^{\beta\nu}) - \eta^{\mu\nu}u^{\alpha\beta}(p,q) \\ &- (\eta^{\beta\nu}p^{\mu} + \eta^{\beta\mu}p^{\nu})q^{\alpha} - (\eta^{\alpha\nu}q^{\mu} + \eta^{\alpha\mu}q^{\nu})p^{\beta} \,, \end{split}$$