

PIC modelling of laser-solid interactions: from ion acceleration to high fields plasmonics

Andrea Sgattoni - INO CNR Pisa **Italy**



Coworkers: our team and collaborators

Pisa INO - CNR and University



INO - CNR



Andrea Macchi
Senior scientist



[Andrea Sgattoni](mailto:andrea.sgattoni@gmail.com)
Post Doc
andrea.sgattoni@gmail.com



Luca Fedeli
PhD student



Italy »
Università di Bologna



Politecnico Milano



Università di Pisa



France »
CEA-Saclay
École polytechnique
Paris



Czech Republic »
Tech. University Prague



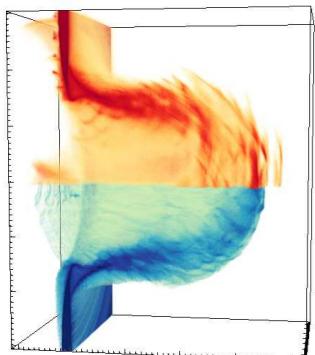
UK »
Queen's Univ. Belfast



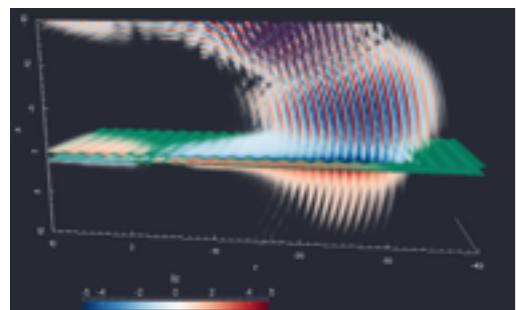
South Korea »
GIST Gwangju

PIC modelling of laser-solid interactions

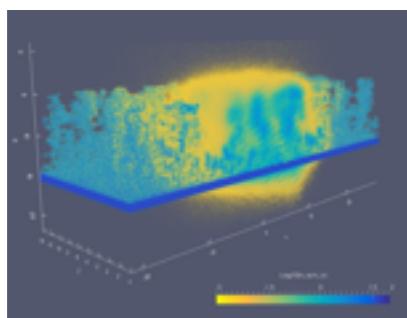
Introduction:



Ion Acceleration using thin foils:
Light-Sail acceleration



Electron Acceleration with gratings:
acceleration by relativistic surface waves



Foam Targets for ion acceleration:
enhanced TNSA

Introduction

Laser-Solid Interaction

Ion Acceleration

Electron Acceleration

High Harmonic generation

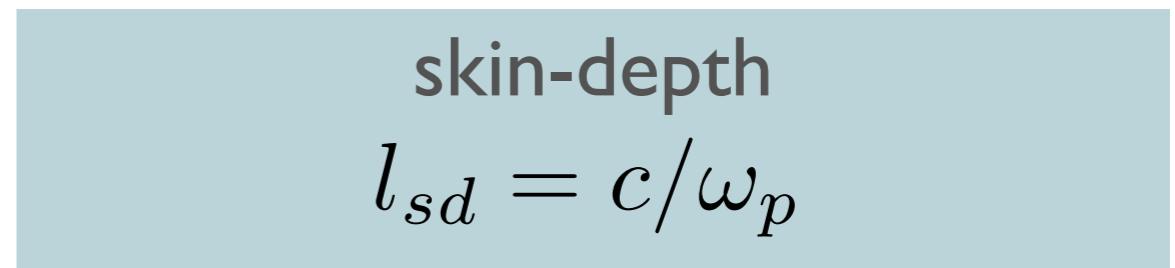
Laser-Solid interaction

ultra-high fields and over-dense plasma

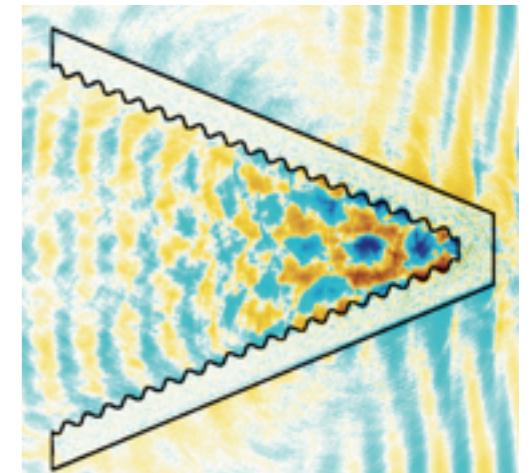
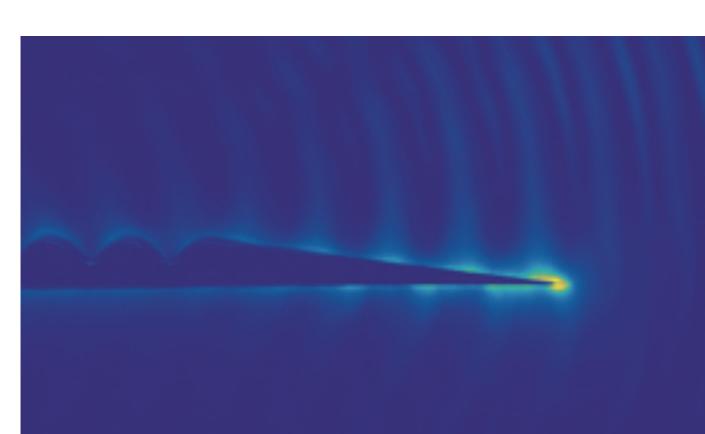
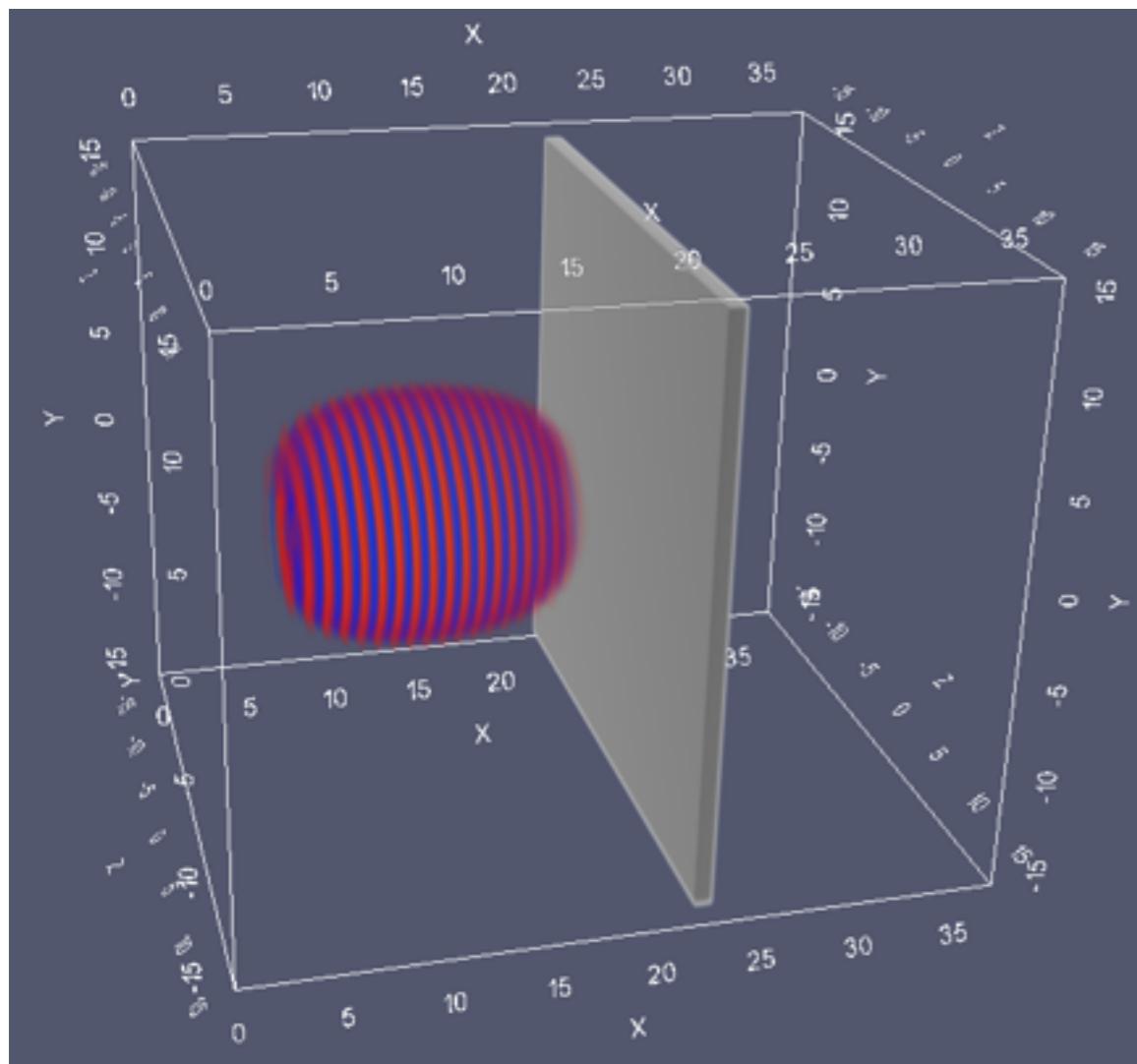
relativistic intensity laser



overdense plasma



- Large box
- many particles in a small volume
- complex target structures



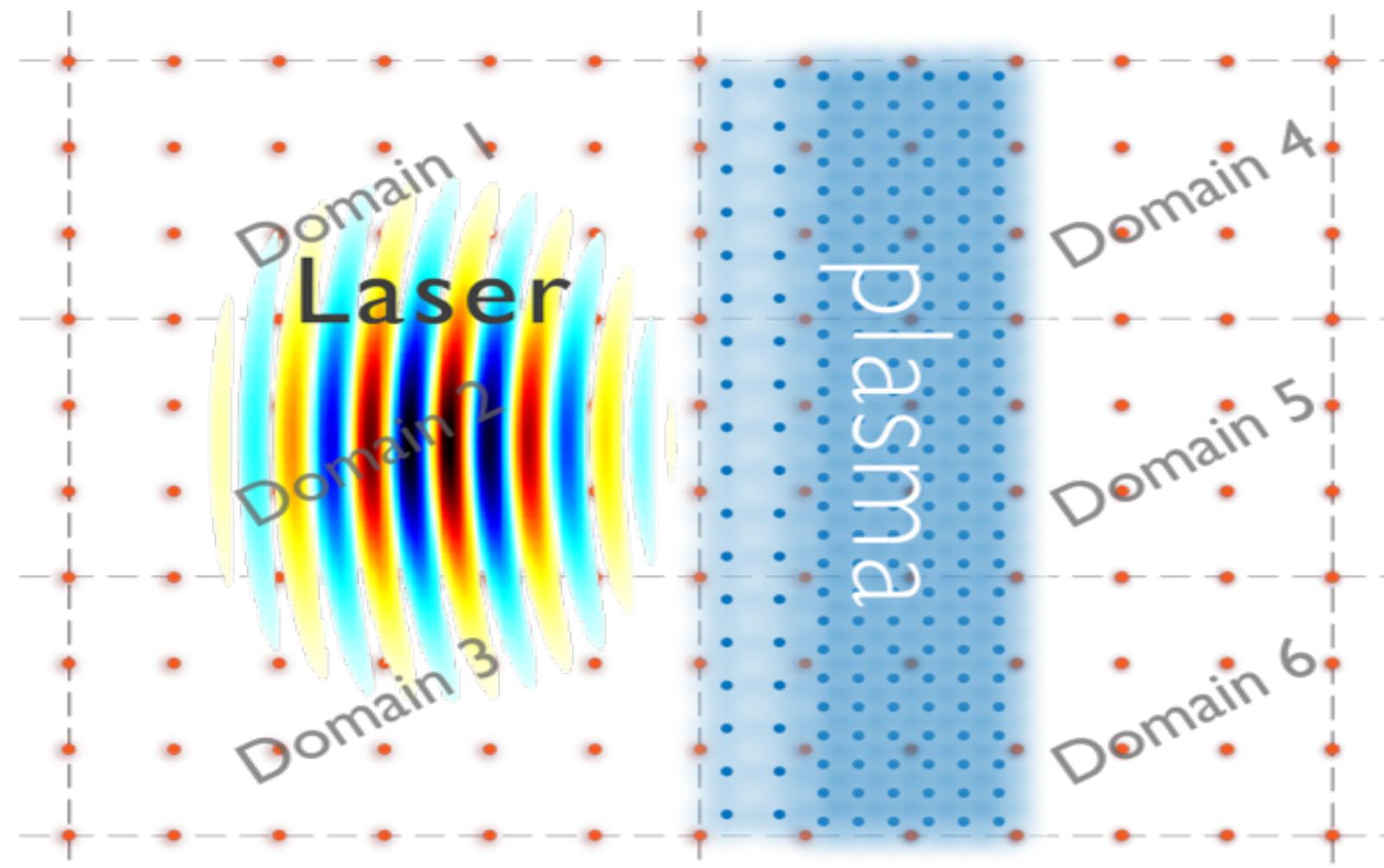
Particle In Cell

From the Maxwell -Vlasov equation to the PIC method

space discretised
on a **GRID**

time steps

plasma density sampled with
numerical particles



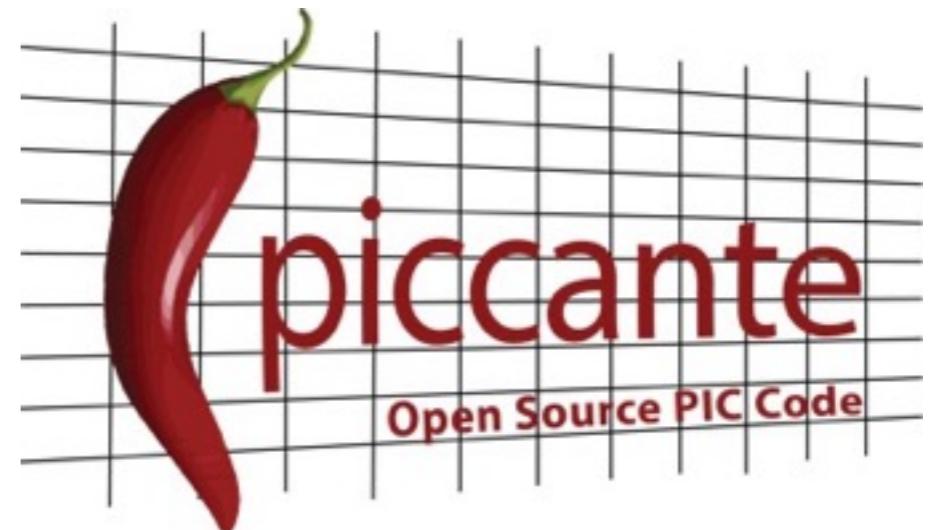
PICCANTE: our open source PIC

Goal:

1D-2D-3D **flexible** in-house

- **PICCANTE** (*italian*): /pik'kante/ = spicy, hot adj.

- 3D fully parallelized C++ MPI
- Open source: hosted on [GitHub.com](#)
- Started from scratch: Nov. 2013
- First production runs: Feb. 2014
- Tested up to 32k cores



GitHub



thanks GitHub for free unlimited account

Ion acceleration using thin foils

Light-Sail acceleration

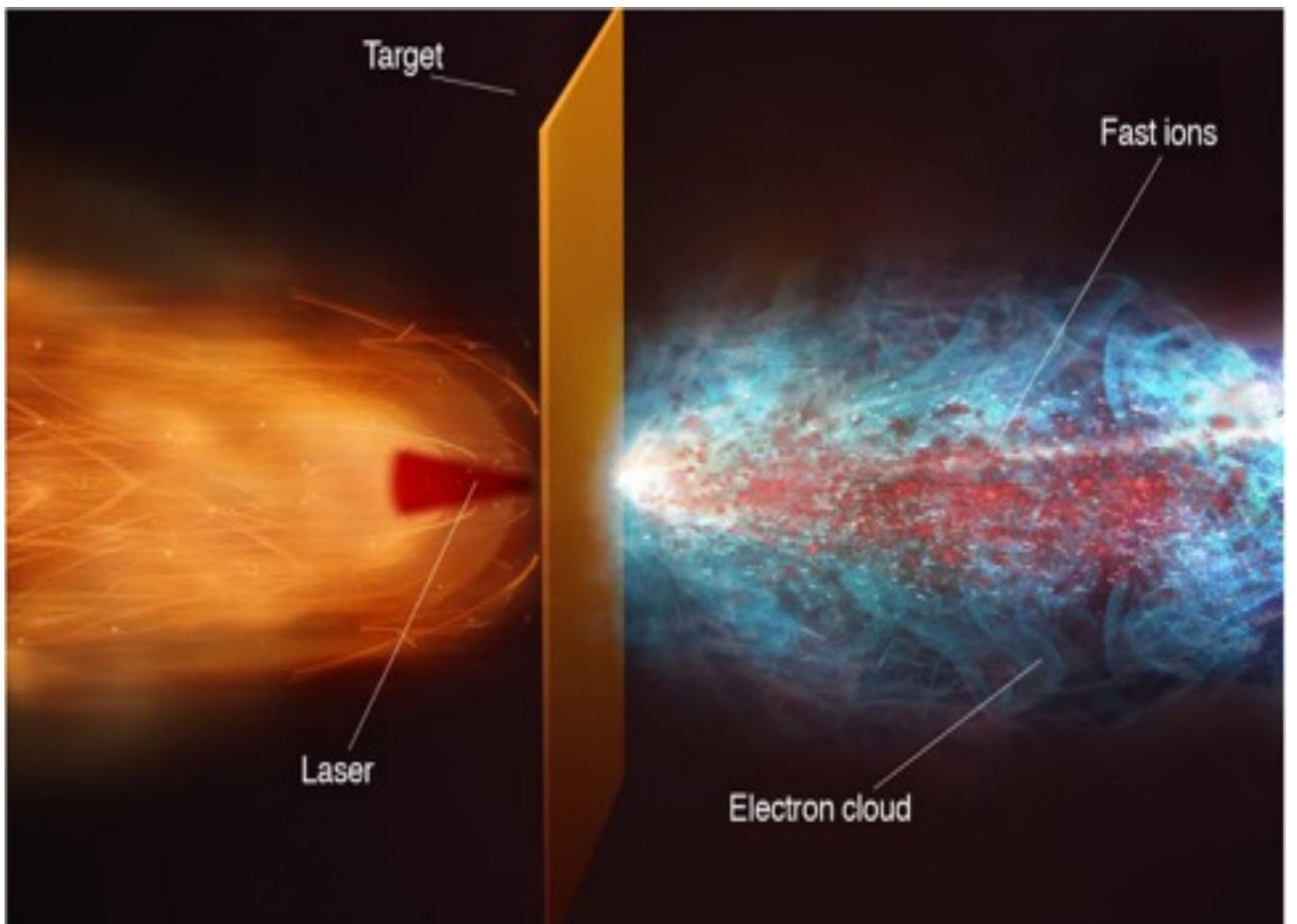
Laser driven ion acceleration

Year 2000

- » Snavely et. al. PRL 85
- » Clark et. al. PRL 84
- » Maksimchuk et. al. PRL 84
- » Wilks et. al. PoP 8 (2001)

Target Normal Sheath Acceleration

- » acceleration of surface contaminants
- » short bunches 10s fs
- » low emittance
- » **exponential energy spectrum**

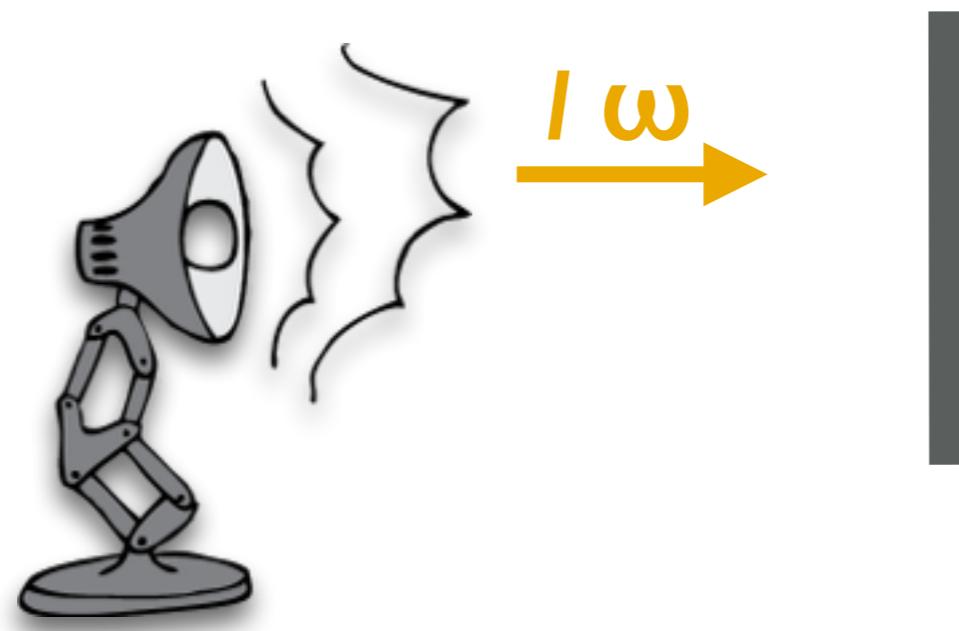


- » Macchi, Borghesi, Passoni Rev. Mod. Phys 85 (2013)
- » Daido, Nishiuchi, Pirozhkov, Rep. Prog. Phys. 75 (2012)

Radiation Pressure Acceleration: Light-Sail sailing with the laser light

- ultrathin foils (10s nm) behaves like mirrors
- accelerated by the light pressure

$$P = \frac{2I}{c}$$

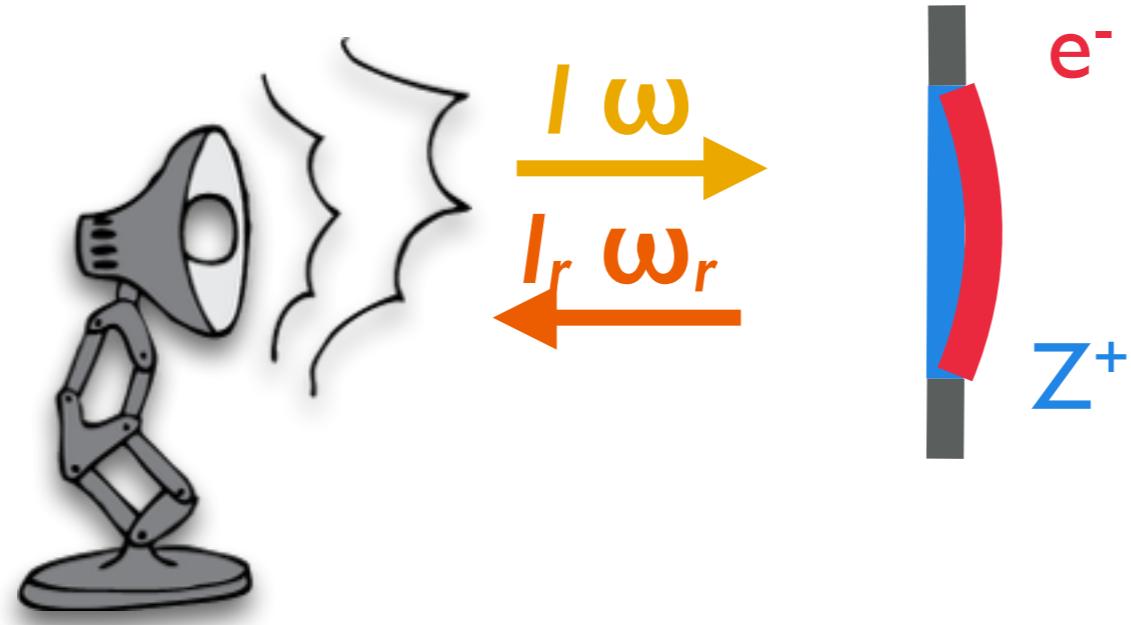


T. Esirkepov, et al. PRL., **92** (2004)
APL Robinson et al, NJP, **10** (2009)

Radiation Pressure Acceleration: Light-Sail sailing with the laser light

- ultrathin foils (10s nm) behaves like mirrors
- accelerated by the light pressure

$$P = \frac{2I}{c}$$

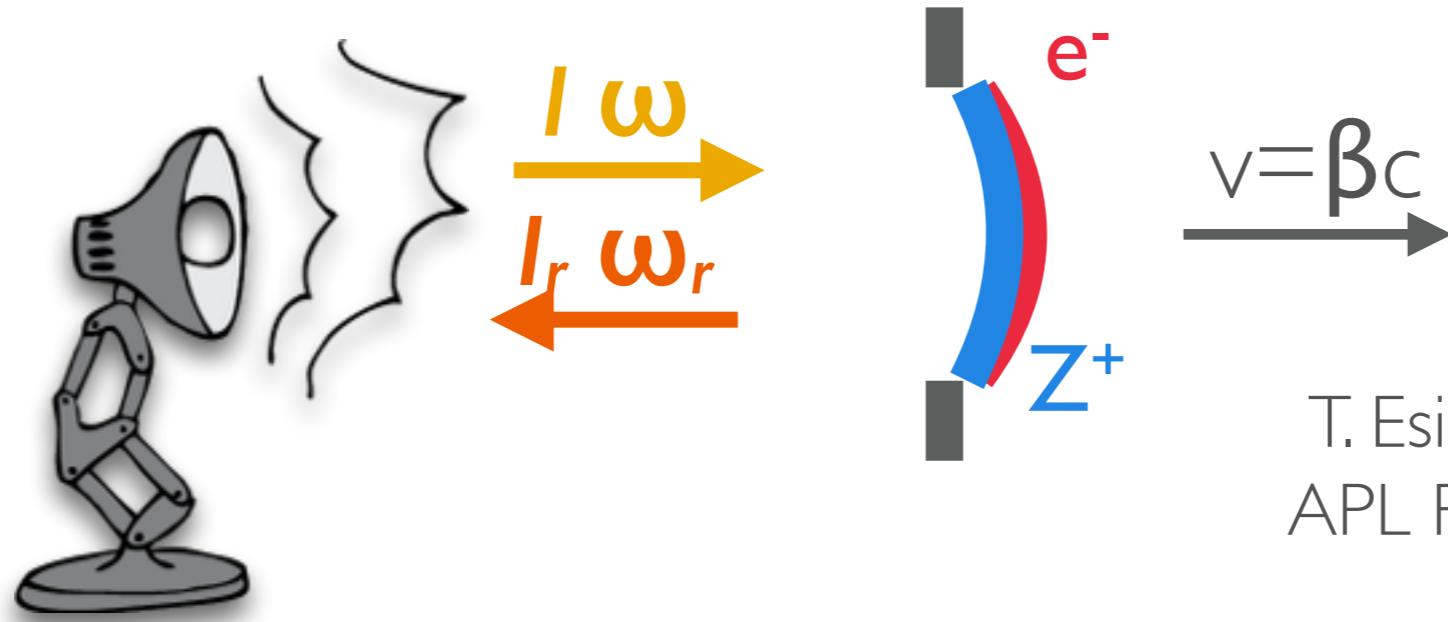


T. Esirkepov, et al. PRL., **92** (2004)
APL Robinson et al, NJP, **10** (2009)

Radiation Pressure Acceleration: Light-Sail sailing with the laser light

- ultrathin foils (10s nm) behaves like mirrors
- accelerated by the light pressure

$$P = \frac{2I}{c}$$



T. Esirkepov, et al. PRL., **92** (2004)
APL Robinson et al, NJP, **10** (2009)

good

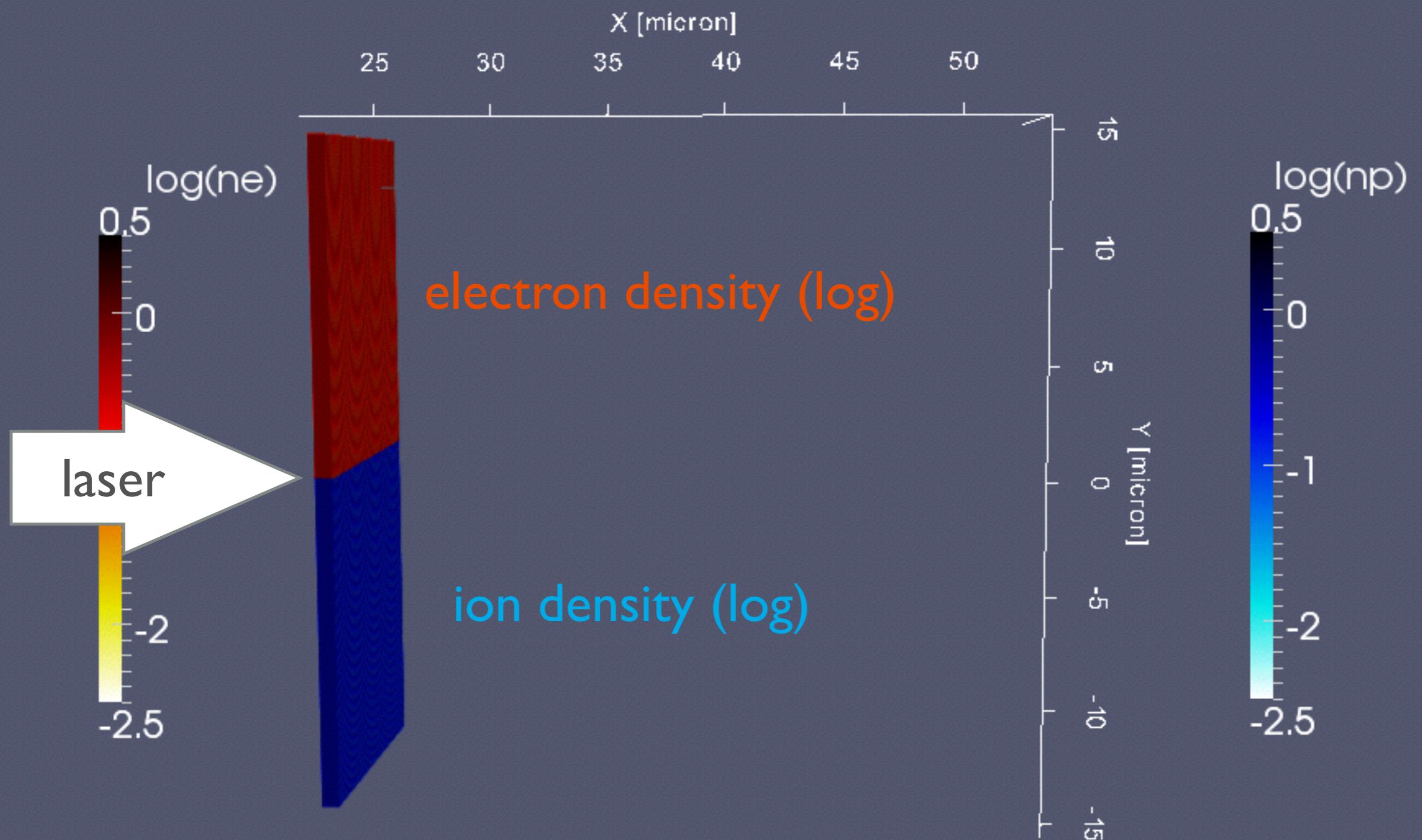
narrower energy spread: acceleration of the bulk of the target
better scaling than TNSA: $E_{\max} \sim I$ vs. $\sim I^{1/2}$

critical

light polarisation: circular is preferred
target integrity: heating, target disassembly, instabilities

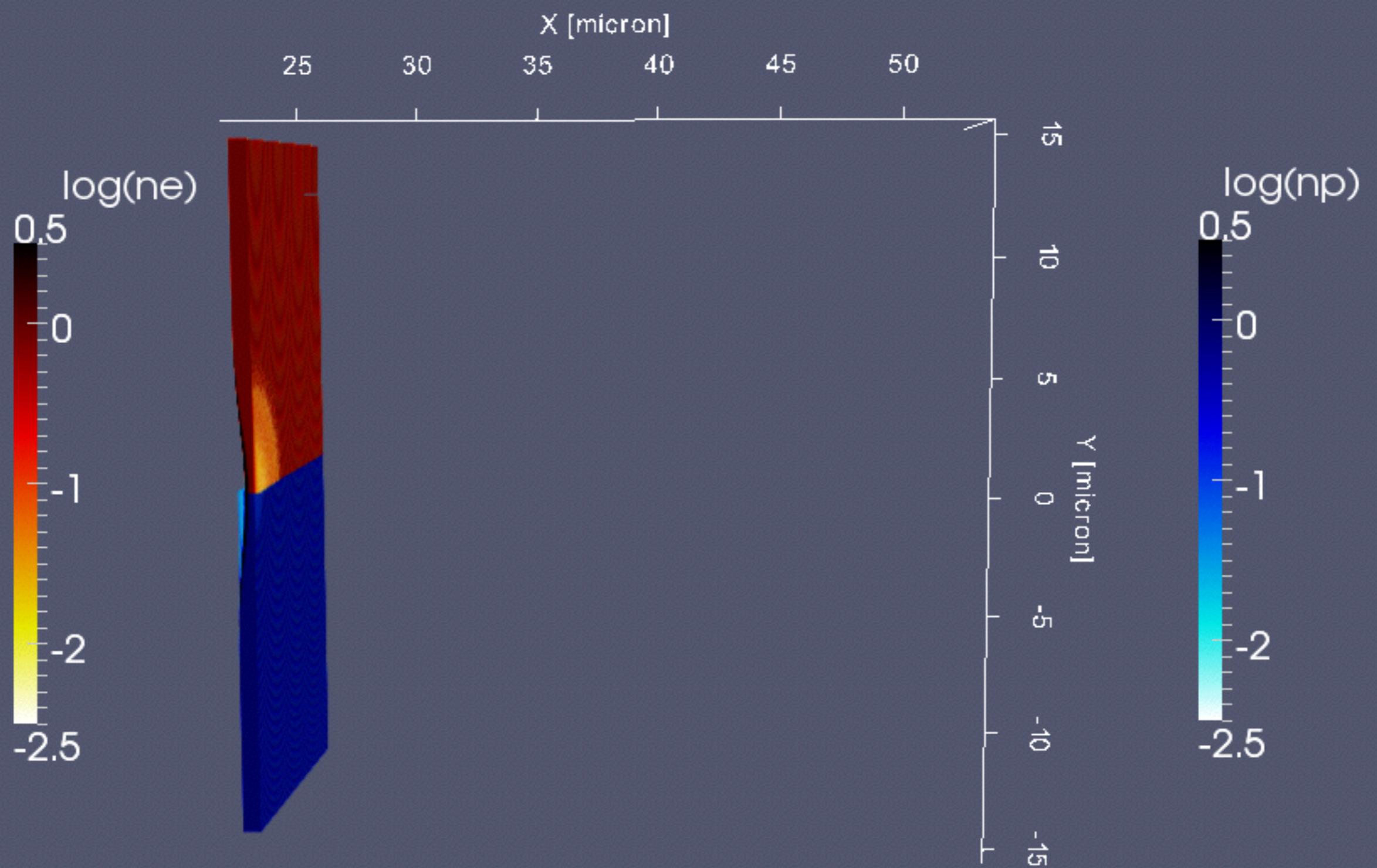
4096x1720x1720 grid points 10 billion particles

$t = 0 \text{ fs}$



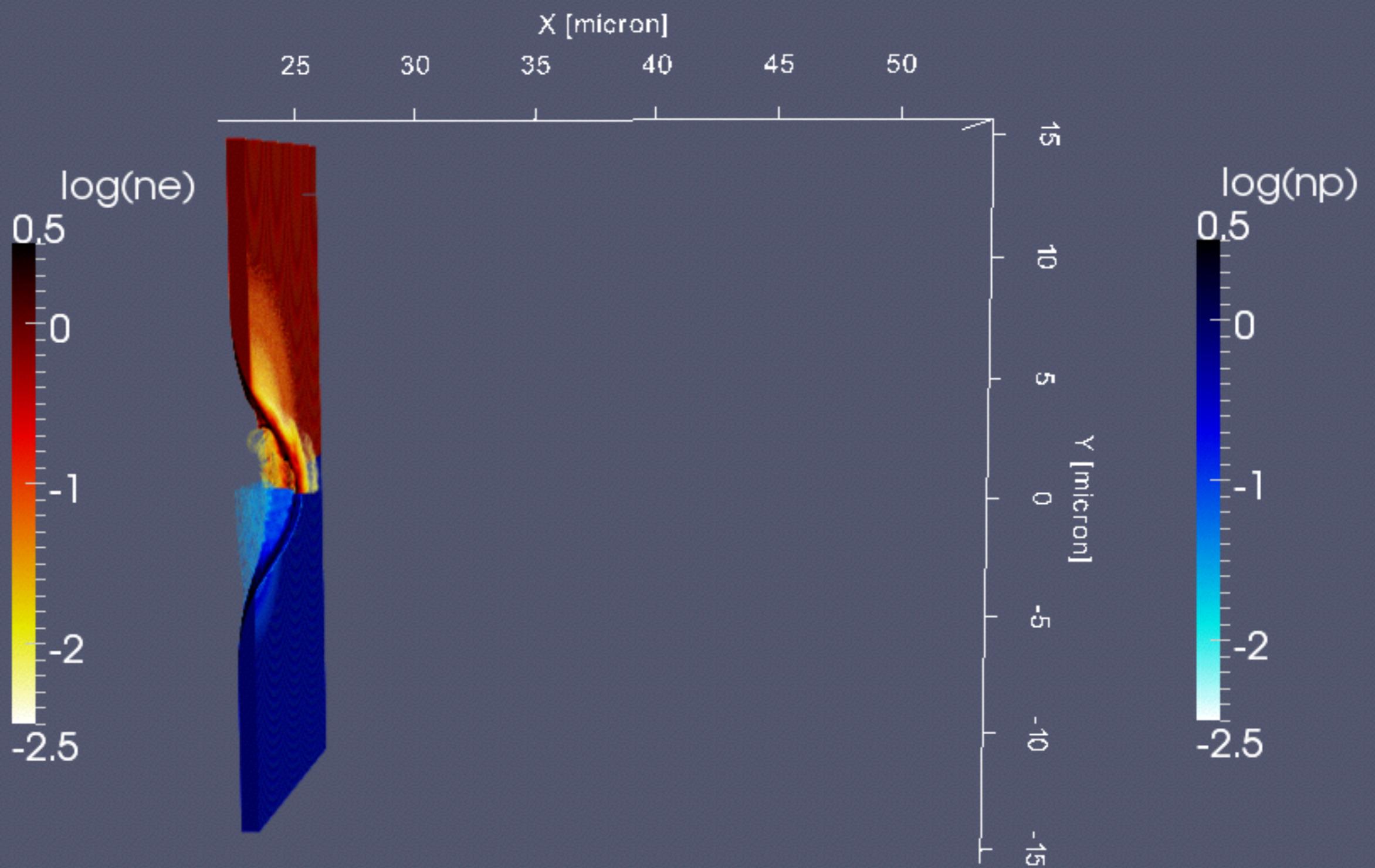
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3.5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0.8 \mu\text{m}$
e-p plasma: $d = 0.8 \mu\text{m}$ $n_e = 64 n_c$

$t = 16,6 \text{ fs}$



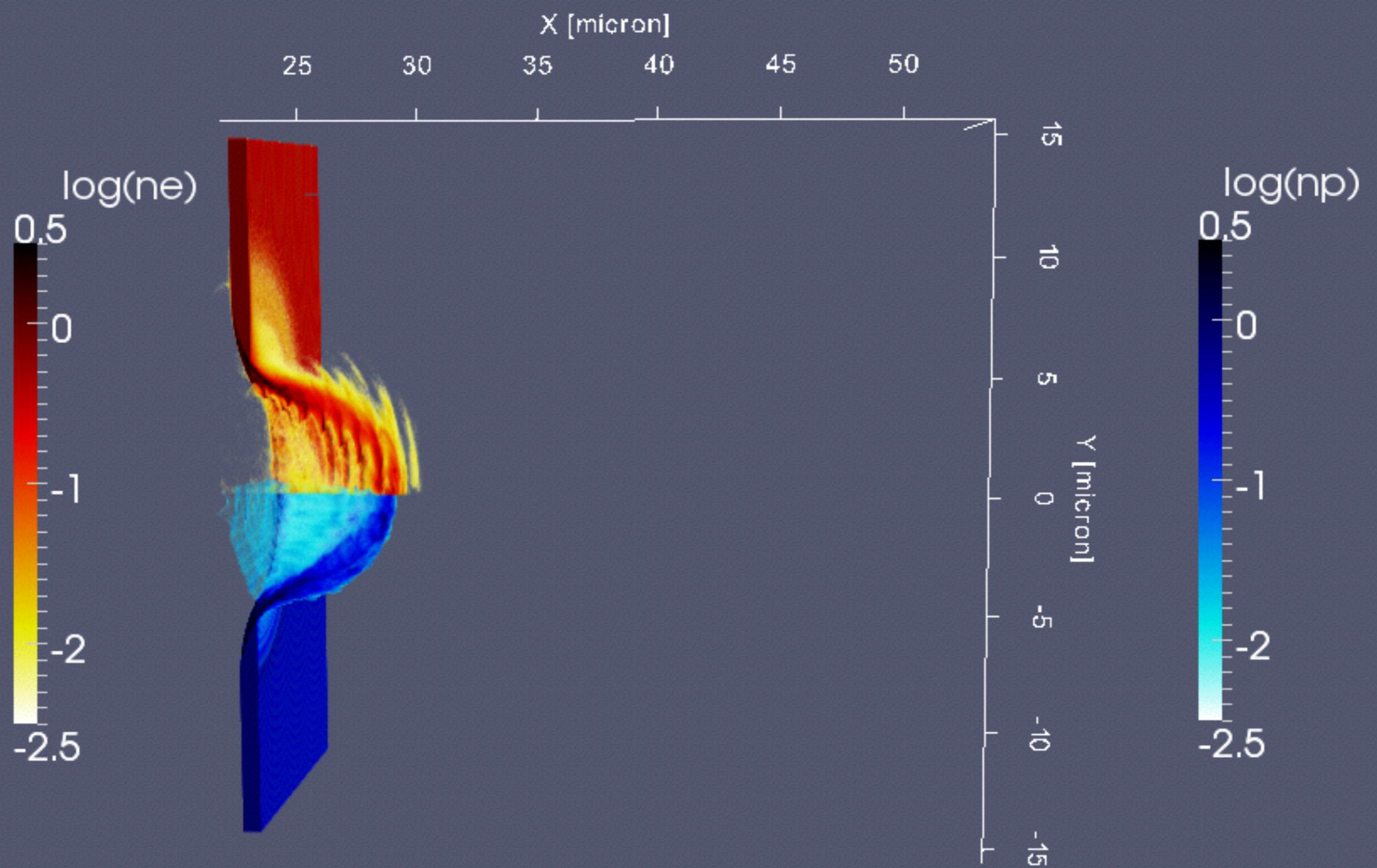
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3,5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0,8 \mu\text{m}$
e-p plasma: $d = 0,8 \mu\text{m}$ $n_e = 64 n_c$

$t = 33,3 \text{ fs}$



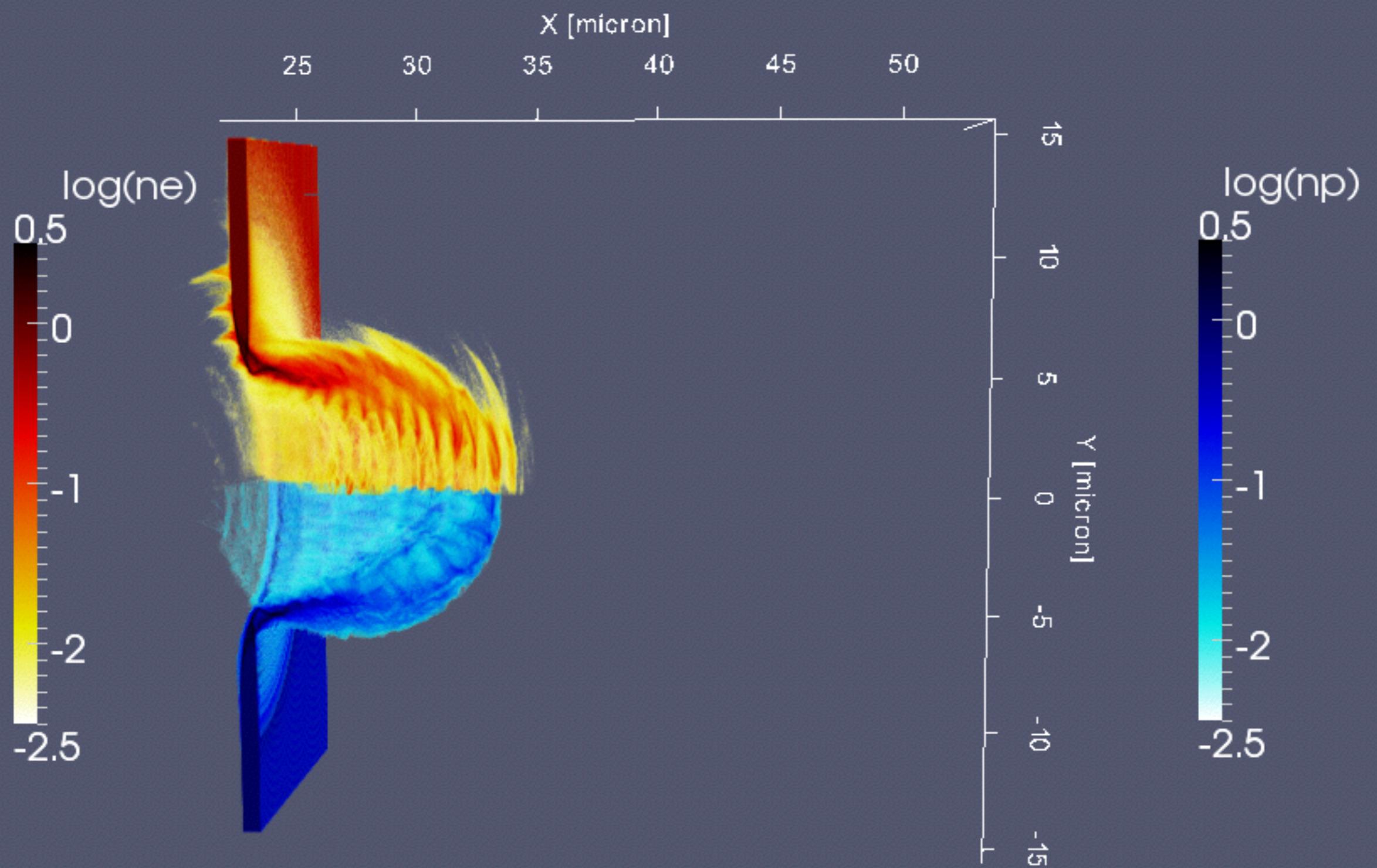
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3,5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0,8 \mu\text{m}$
e-p plasma: $d = 0,8 \mu\text{m}$ $n_e = 64 n_c$

$t = 50 \text{ fs}$



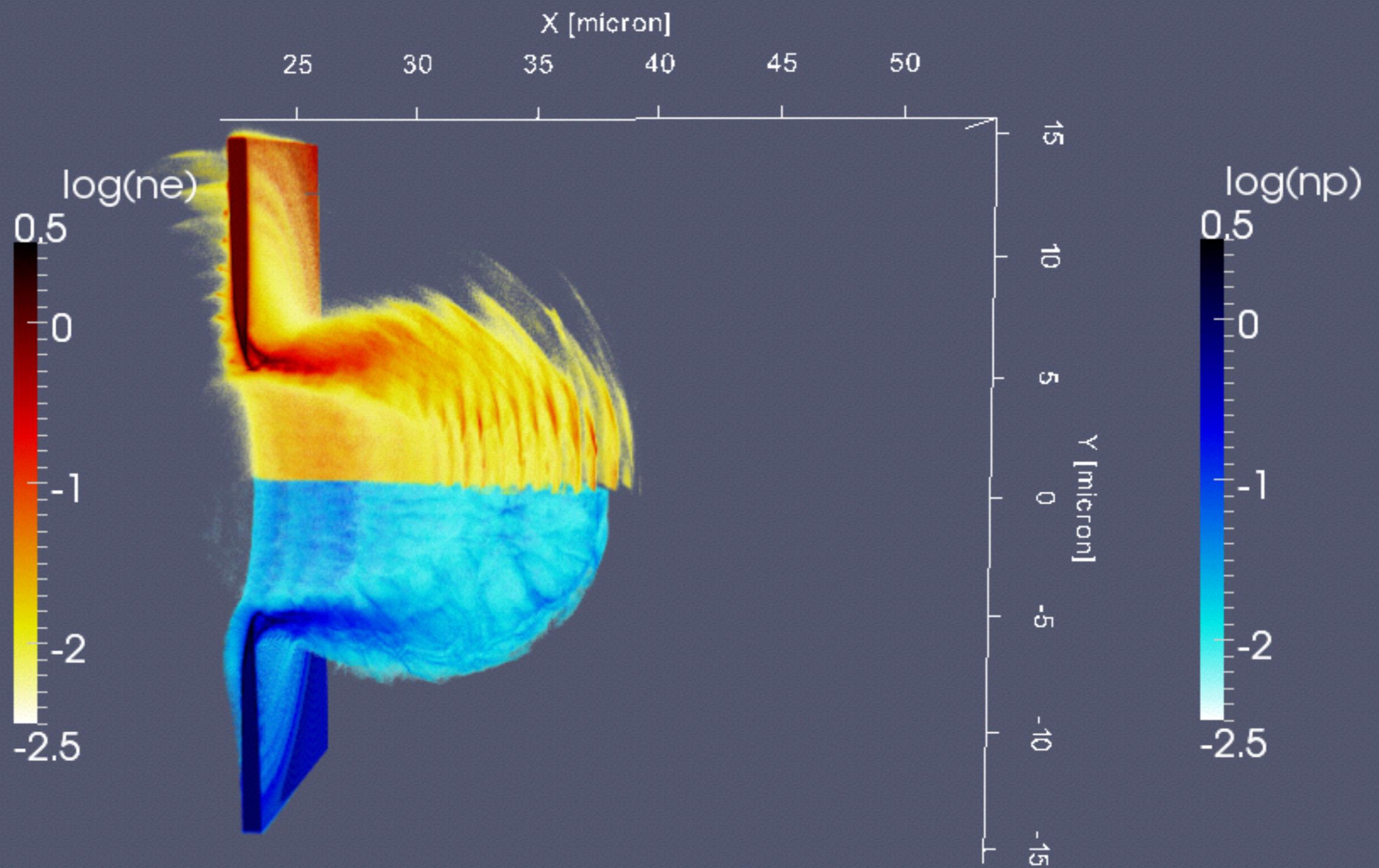
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3.5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0.8 \mu\text{m}$
e-p plasma: $d = 0.8 \mu\text{m}$ $n_e = 64 n_c$

$t = 66,6 \text{ fs}$



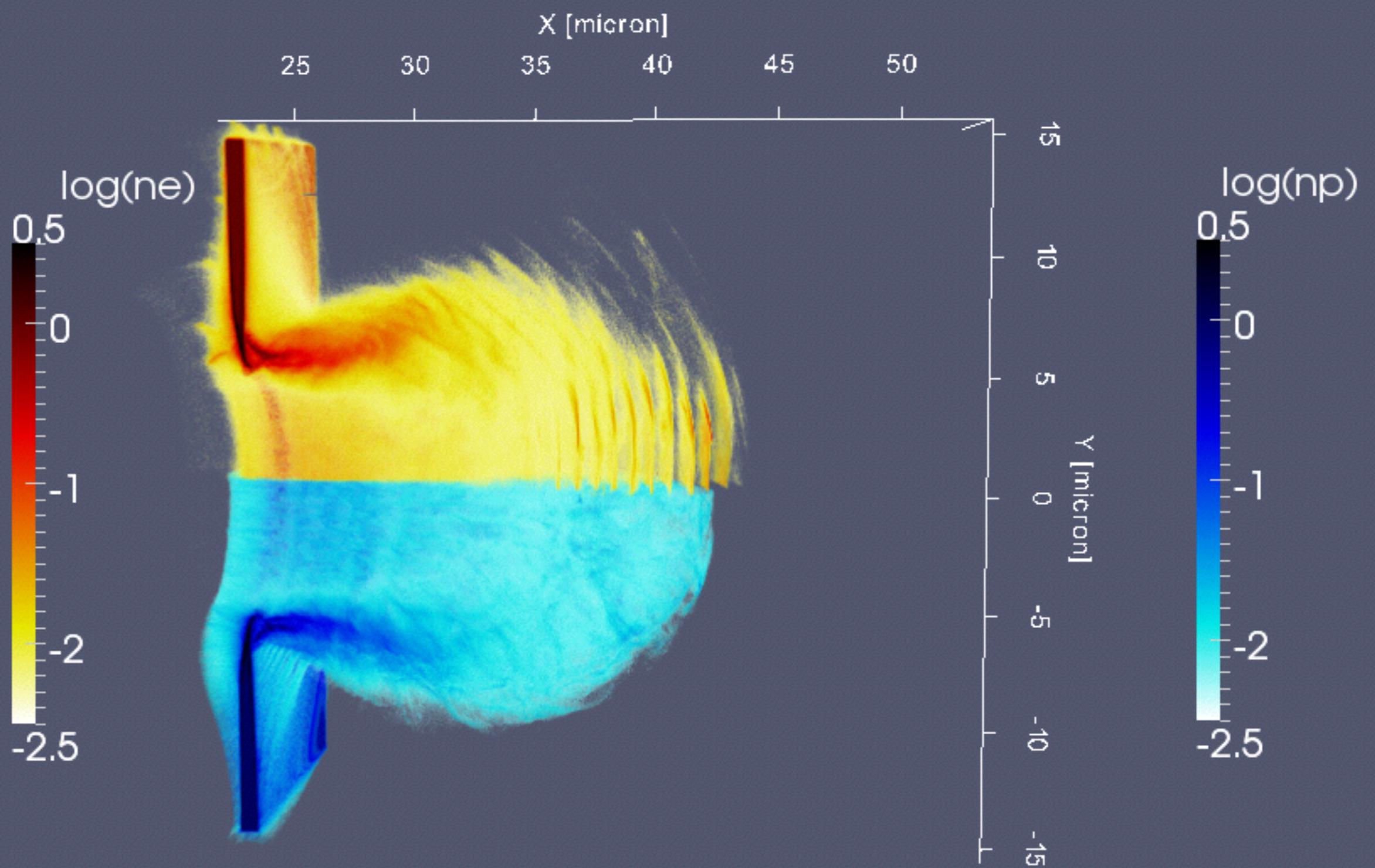
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3,5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0,8 \mu\text{m}$
e-p plasma: $d = 0,8 \mu\text{m}$ $n_e = 64 n_c$

$t = 83,3 \text{ fs}$



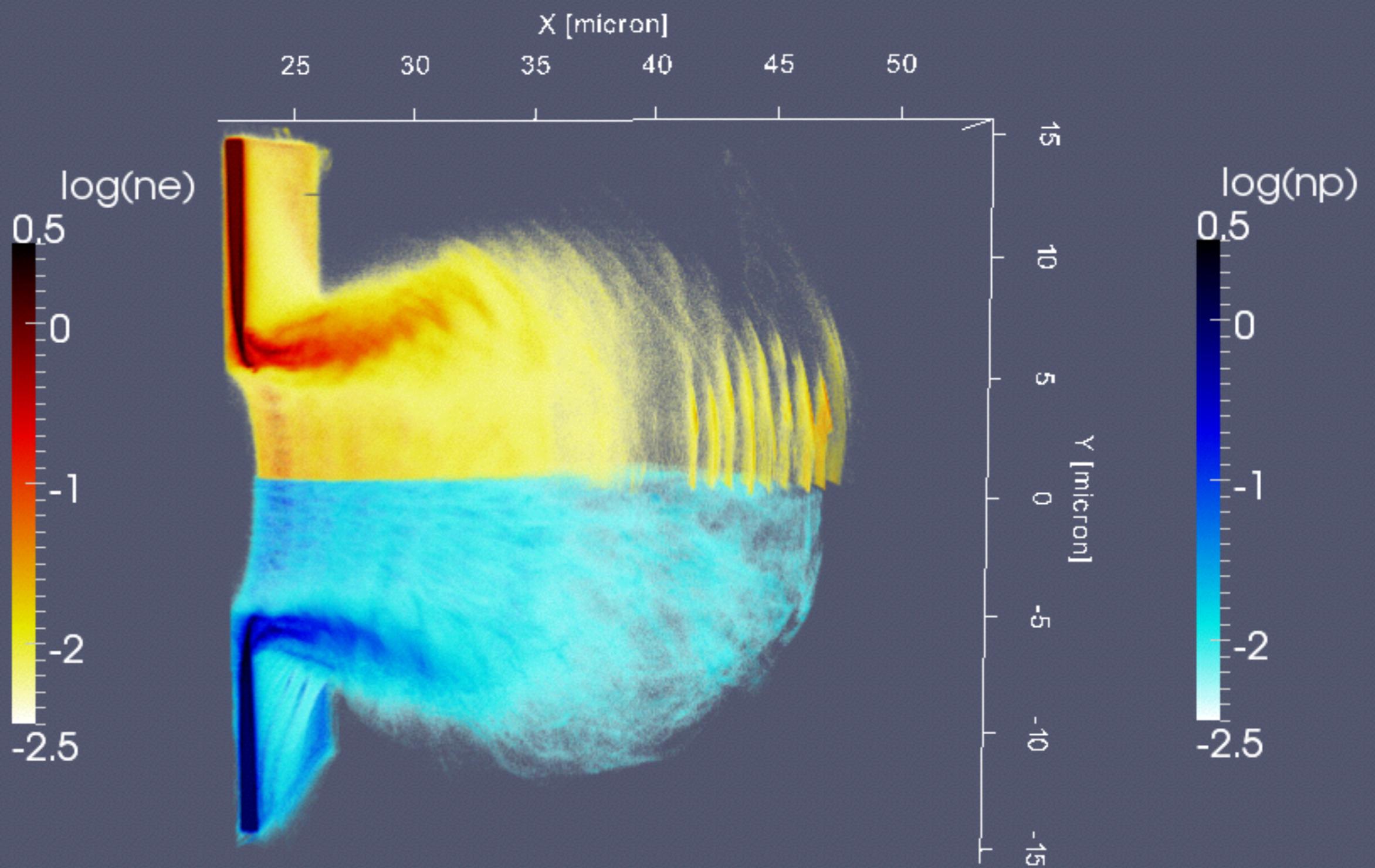
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3,5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0,8 \mu\text{m}$
e-p plasma: $d = 0,8 \mu\text{m}$ $n_e = 64 n_c$

$t = 100 \text{ fs}$



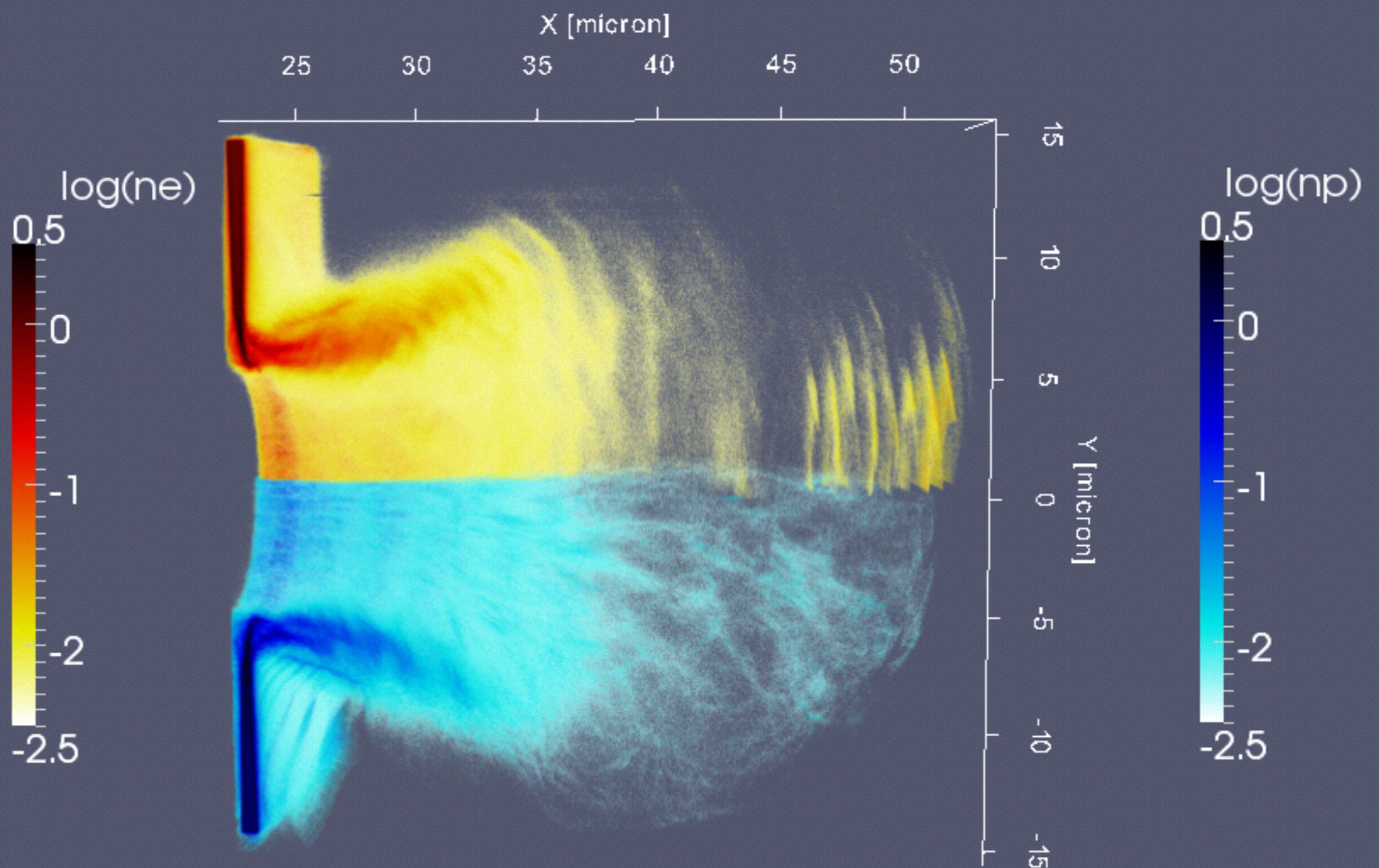
CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3.5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0.8 \mu\text{m}$
e-p plasma: $d = 0.8 \mu\text{m}$ $n_e = 64 n_c$

$t = 116 \text{ fs}$



CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3.5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0.8 \mu\text{m}$
e-p plasma: $d = 0.8 \mu\text{m}$ $n_e = 64 n_c$

$t = 133 \text{ fs}$

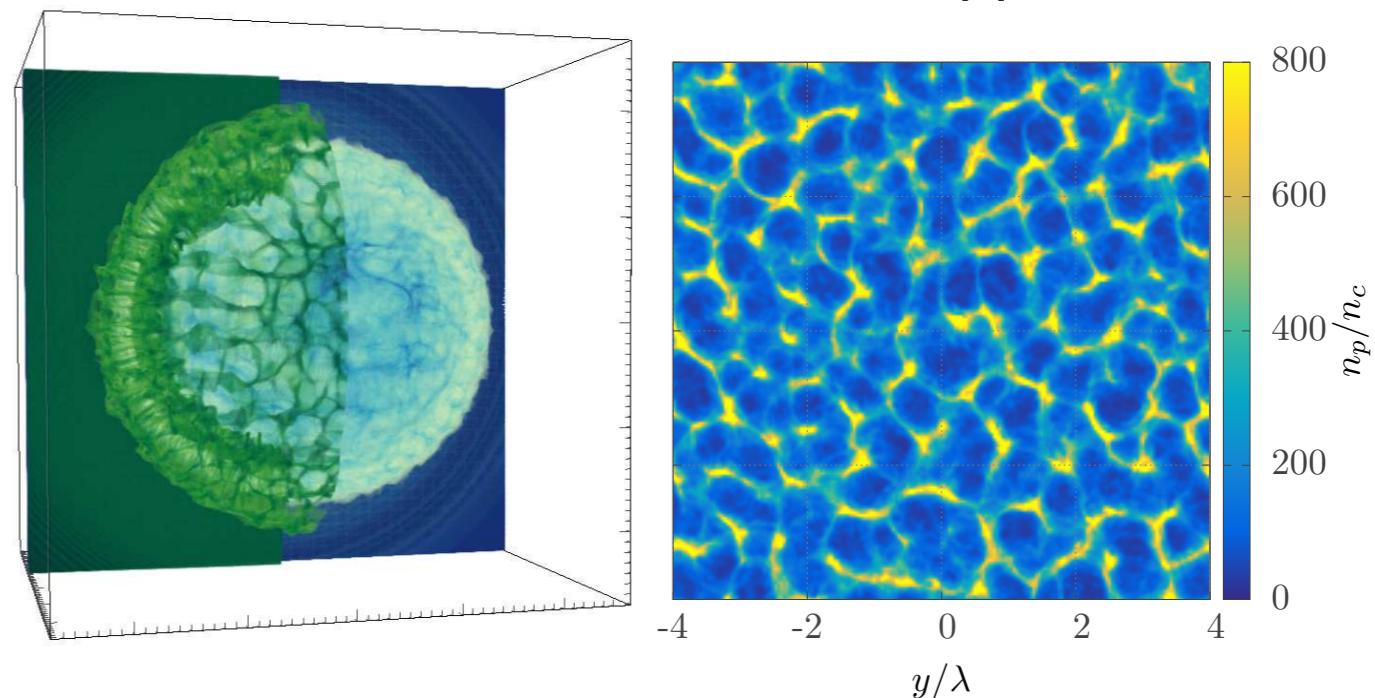
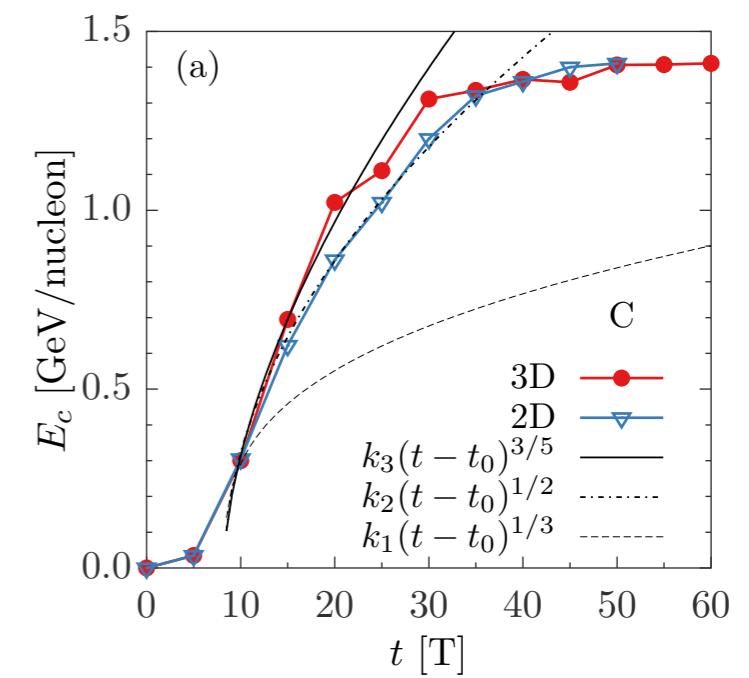


CP laser: $I = 1.7 \cdot 10^{23} \text{ W/cm}^2$ $w_0 = 3.5 \mu\text{m}$ $\tau = 24 \text{ fs}$ $\lambda = 0.8 \mu\text{m}$
e-p plasma: $d = 0.8 \mu\text{m}$ $n_e = 64 n_c$

3D PIC simulations of Light Sail acceleration

- » interaction followed until transparency
- » **acceleration in 3D is faster than 2D**
in agreement with analytical model

Sgattoni et. al. APL 105 (2014)



Rayleigh Taylor Instability:
net-like structures

S. I. Abarzhi, PRE 59 (1999)

Sgattoni et. al. PRE 91(2015)

B. Eliasson NJP 17 (2015) Pegoraro, Bulanov, PRL(2007)

Light-Sail @ GEMINI (RAL) UK

group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

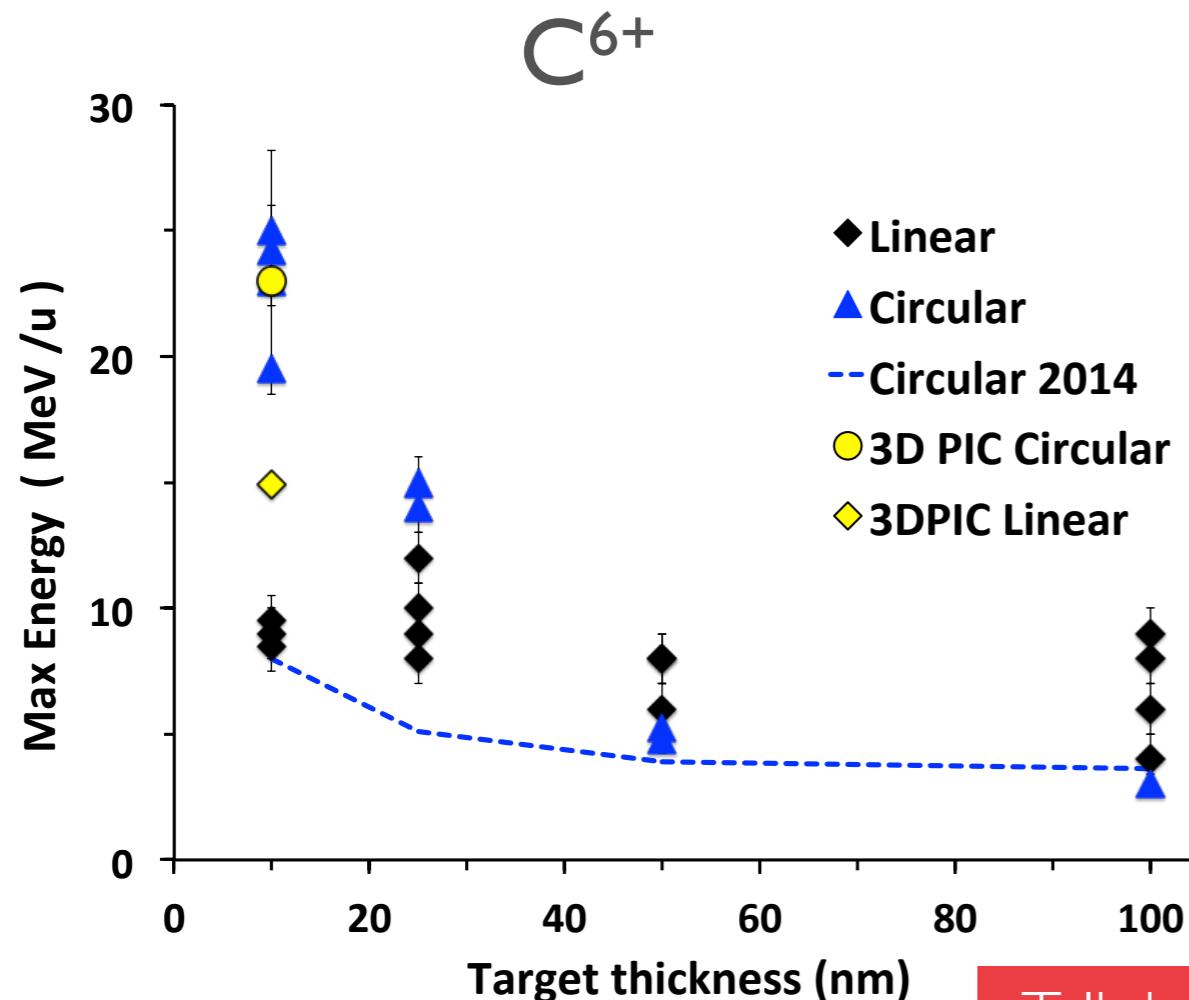
$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, $\text{FWHM}\sim4\mu\text{m}$, $I\sim3\cdot10^{20}\text{W/cm}^2$

Light-Sail @ GEMINI (RAL) UK

group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, $\text{FWHM}\sim4\mu\text{m}$, $I\sim3\cdot10^{20}\text{W/cm}^2$

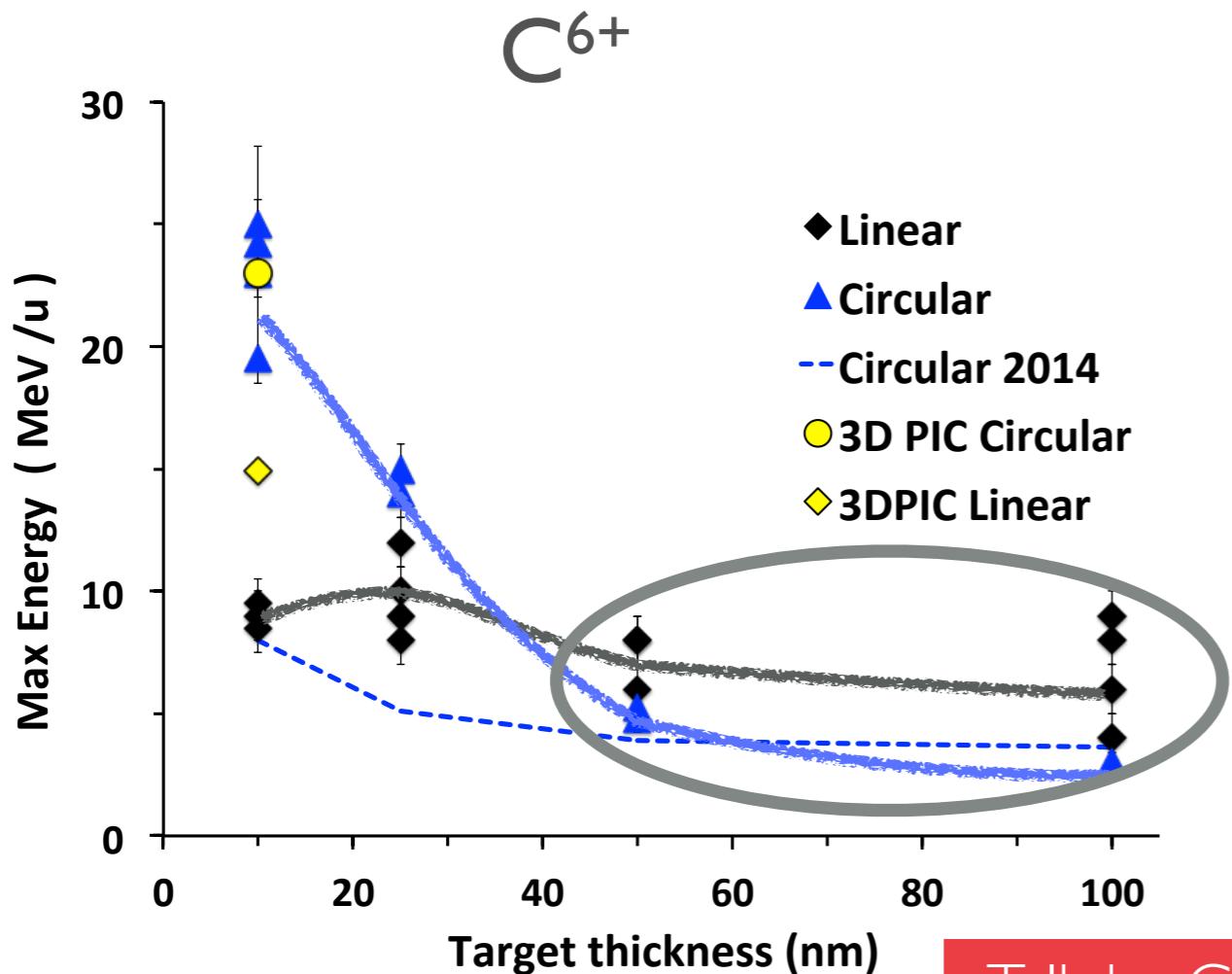


Light-Sail @ GEMINI (RAL) UK

group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, FWHM $\sim 4\mu\text{m}$, $I\sim 3\cdot 10^{20}\text{W/cm}^2$



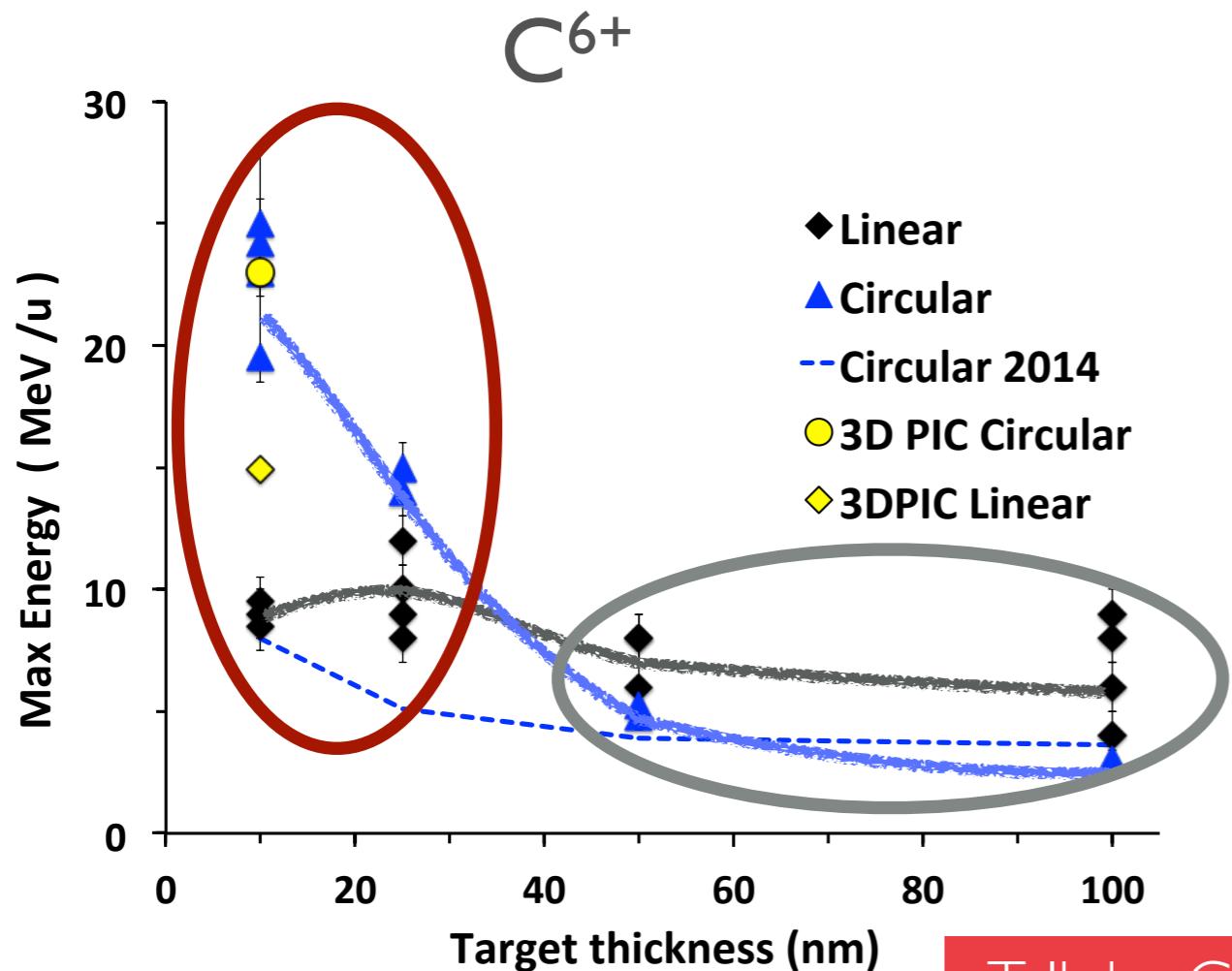
- for “thick target” linear pol. is better than circular

Light-Sail @ GEMINI (RAL) UK

group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, FWHM $\sim 4\mu\text{m}$, $I\sim 3\cdot 10^{20}\text{W/cm}^2$



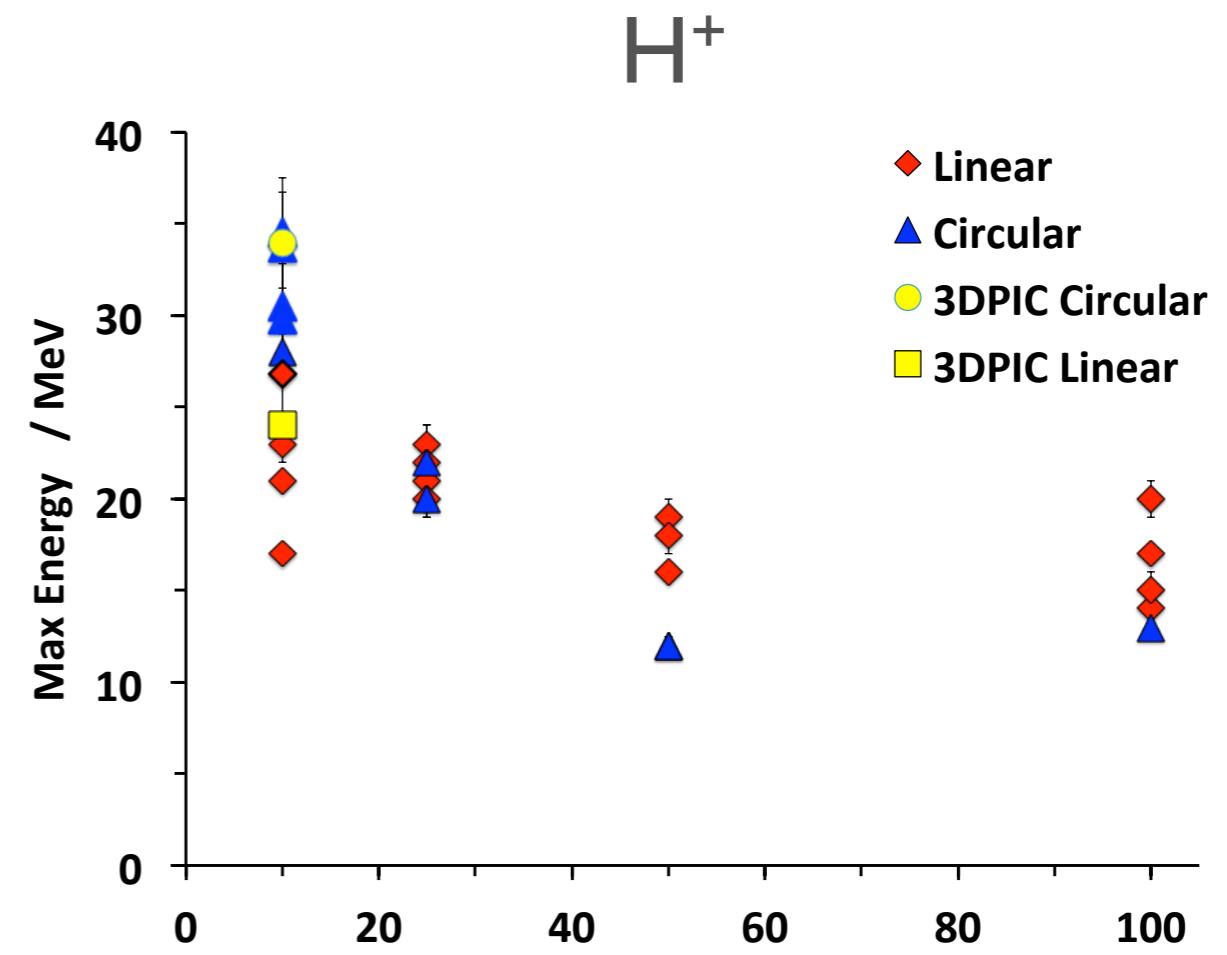
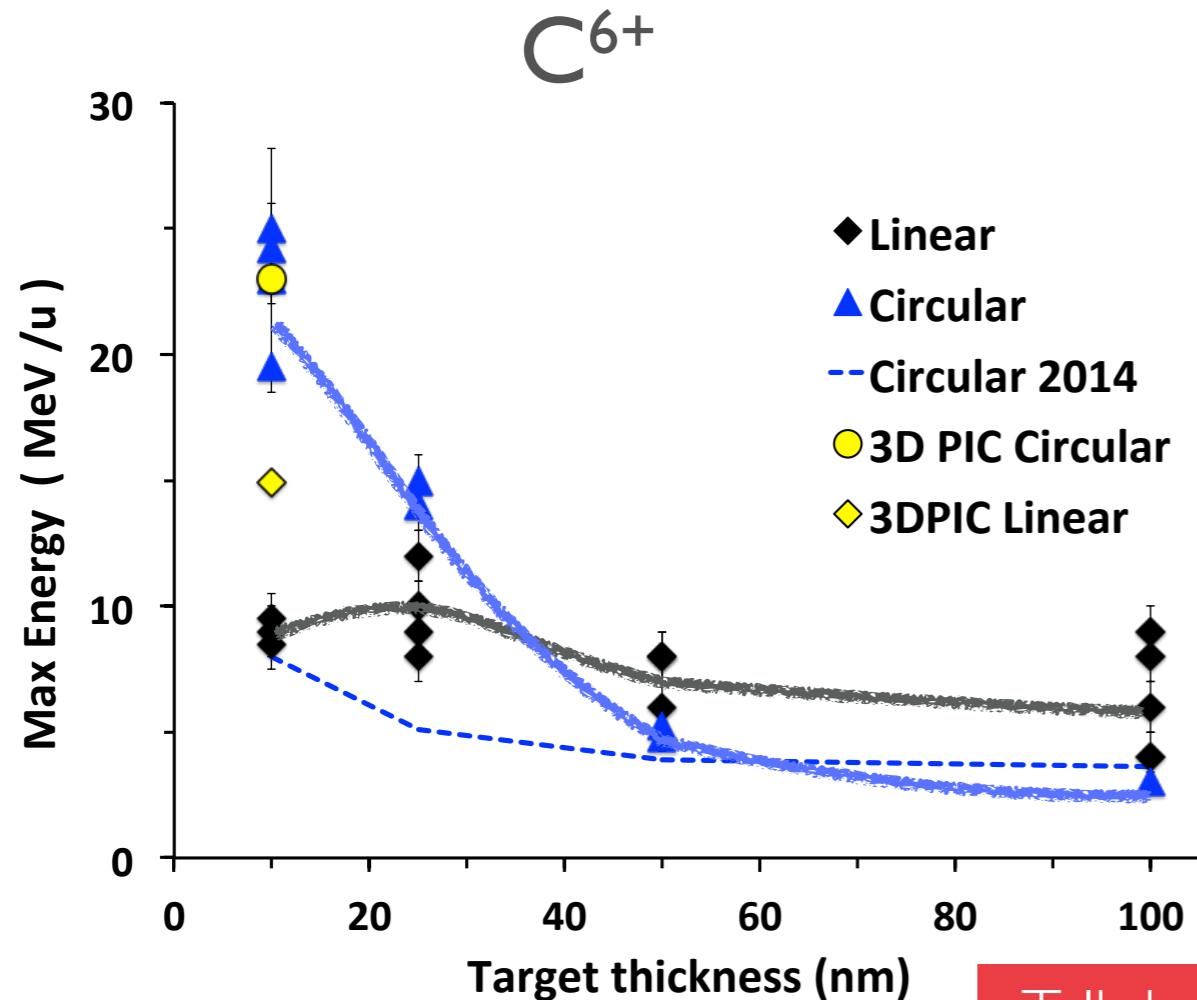
- for “thick target”
linear pol. is better than circular
- for thin target
circular pol is best
- Light-Sail dominates for the thinnest targets

Light-Sail @ GEMINI (RAL) UK

group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, FWHM $\sim 4\mu\text{m}$, $I\sim 3\cdot 10^{20}\text{W/cm}^2$

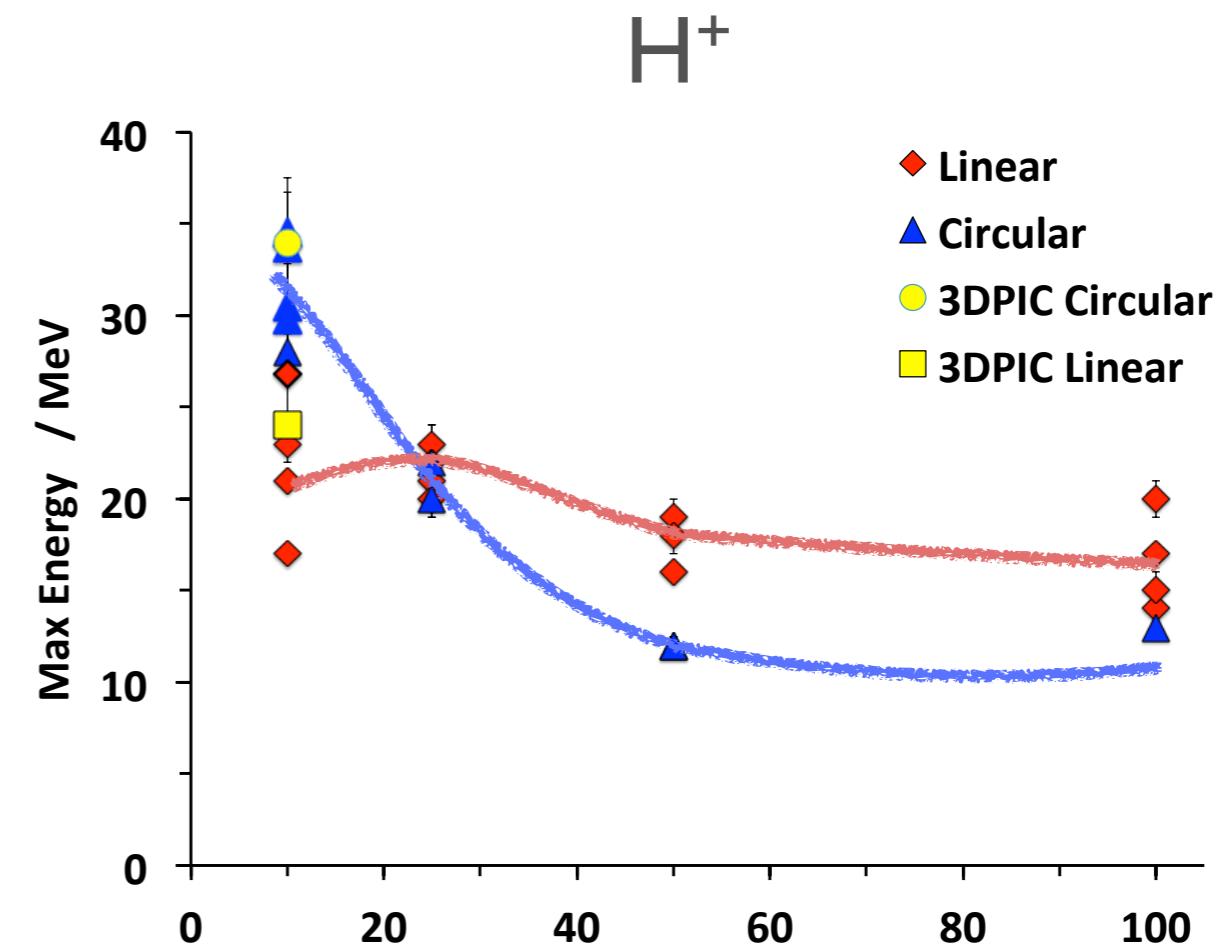
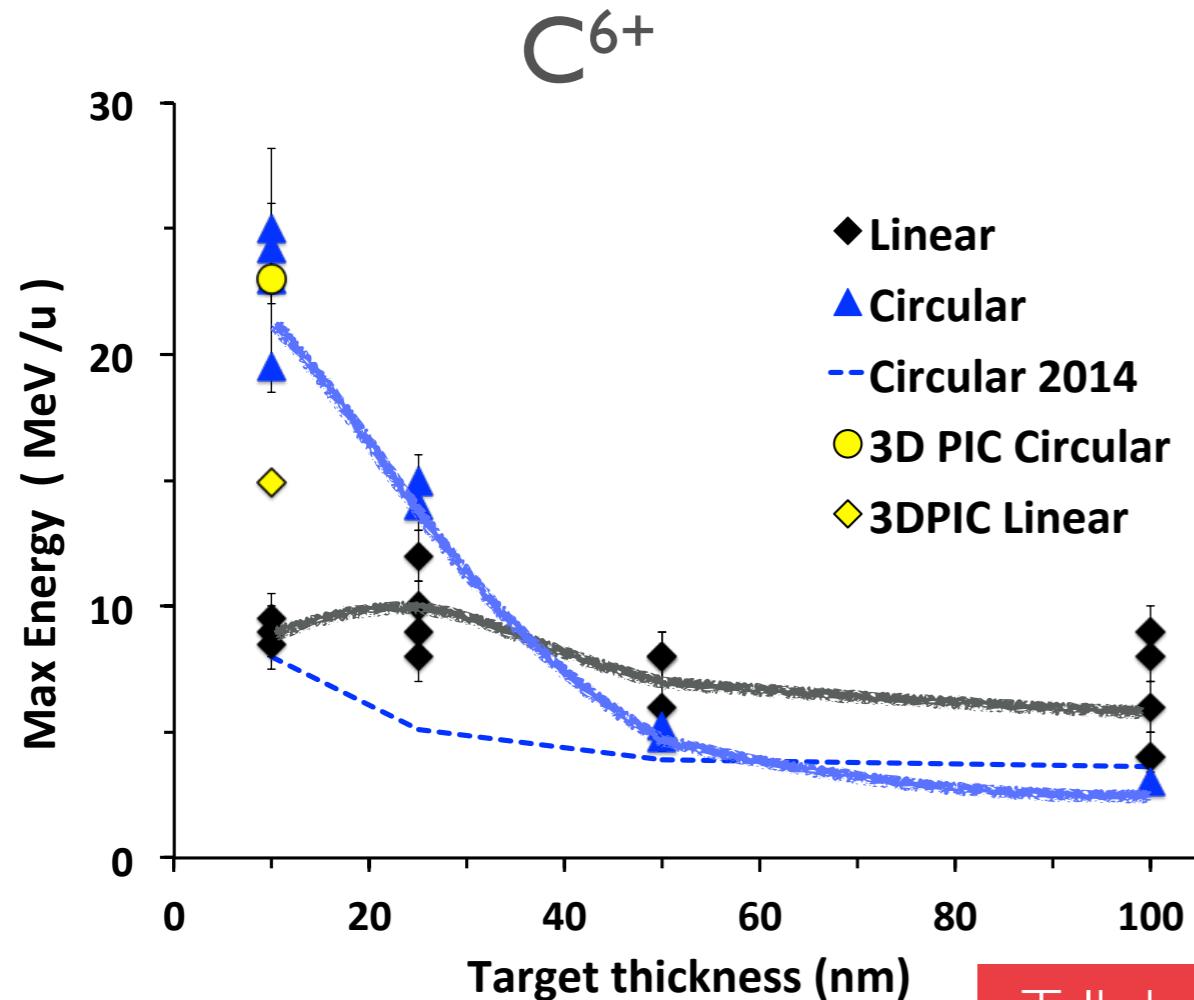


Light-Sail @ GEMINI (RAL) UK

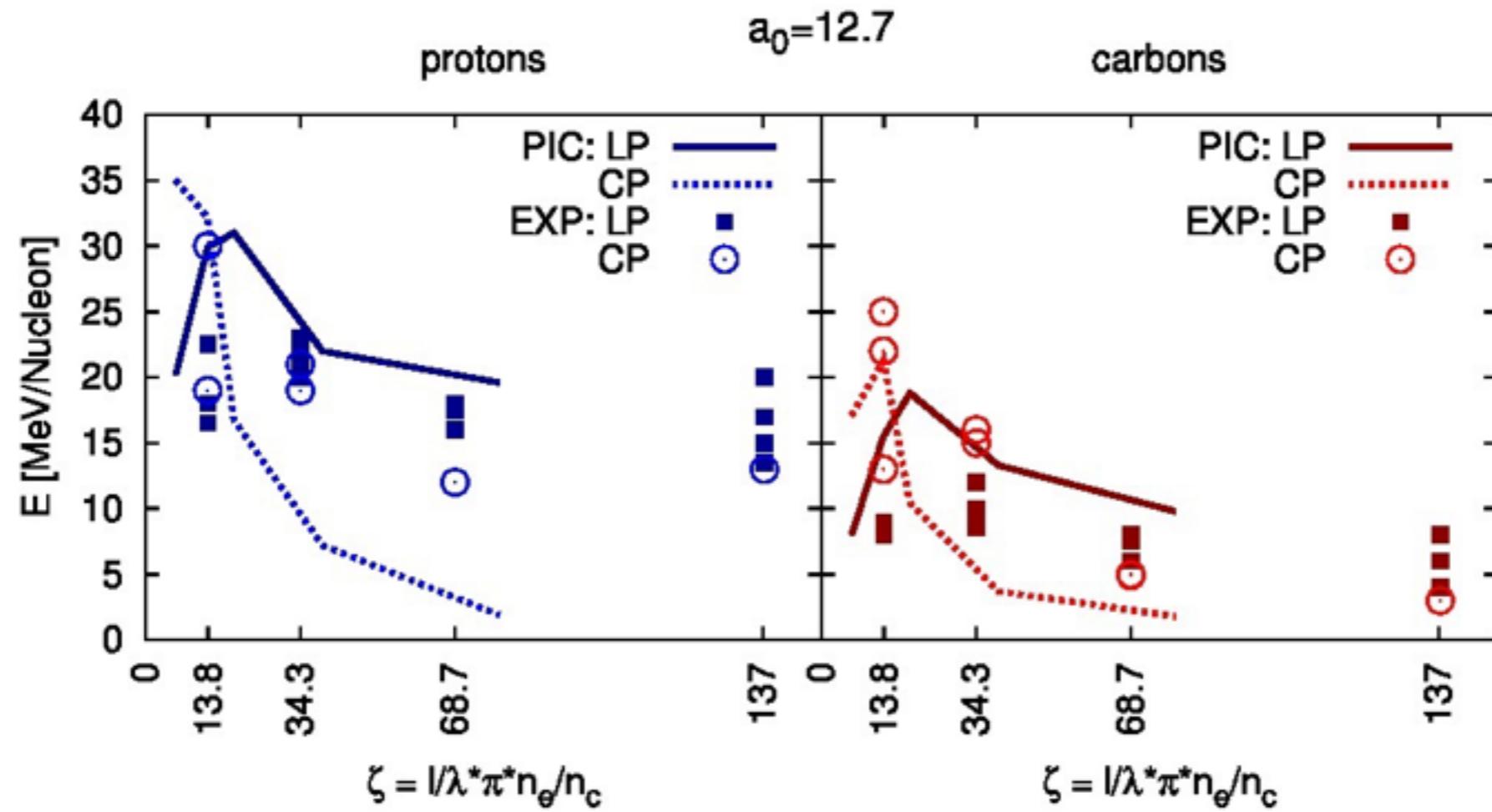
group of Borghesi at Queen's University Belfast

- Very high contrast (10^{12})
- control polarisation
- ultrathin carbon targets 10-100nm
- normal incidence

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, FWHM $\sim 4\mu\text{m}$, $I\sim 3\cdot 10^{20}\text{W/cm}^2$



2D simulations



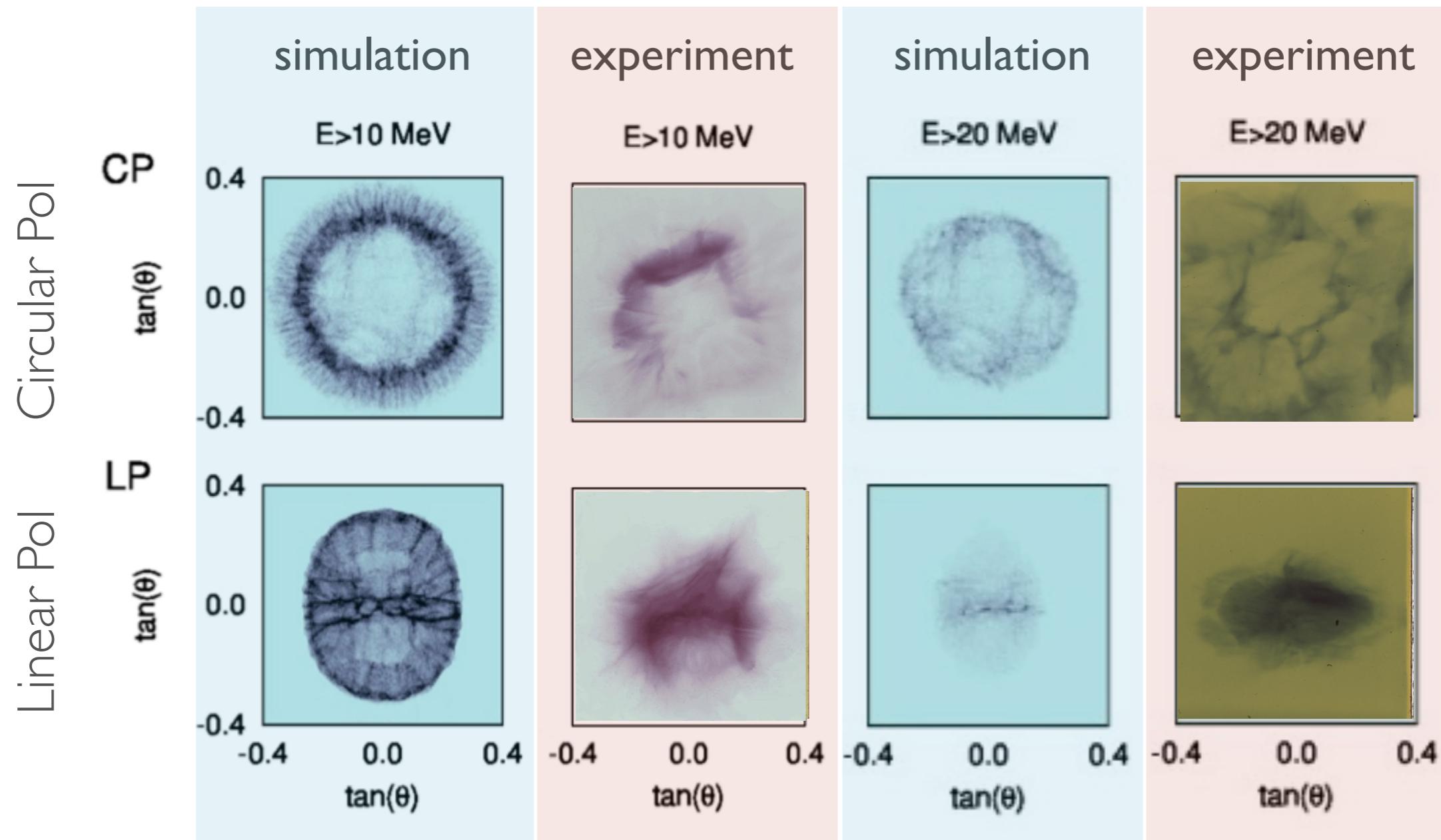
For the thinnest targets

linear-pol: electron heating leads to **transparency**

circular-pol: target stays opaque and **accelerated by radiation pressure**

Light Sail experiment (Gemini) vs. 3D PIC

$\tau=40\text{fs}$, $\lambda=0.8\mu\text{m}$, $E=6.5\text{J}$, FWHM $\sim 4\mu\text{m}$, $I\sim 3 \cdot 10^{20}\text{W/cm}^2$



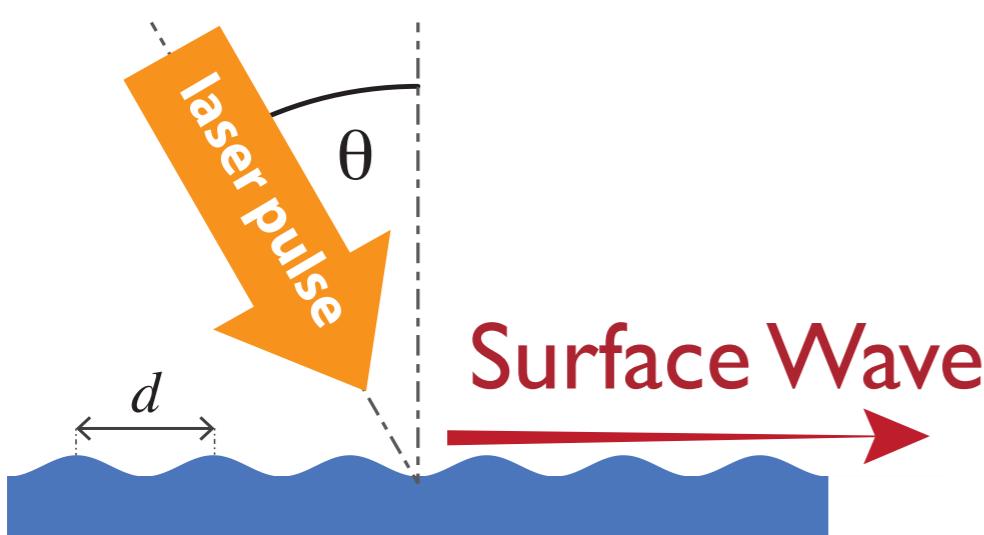
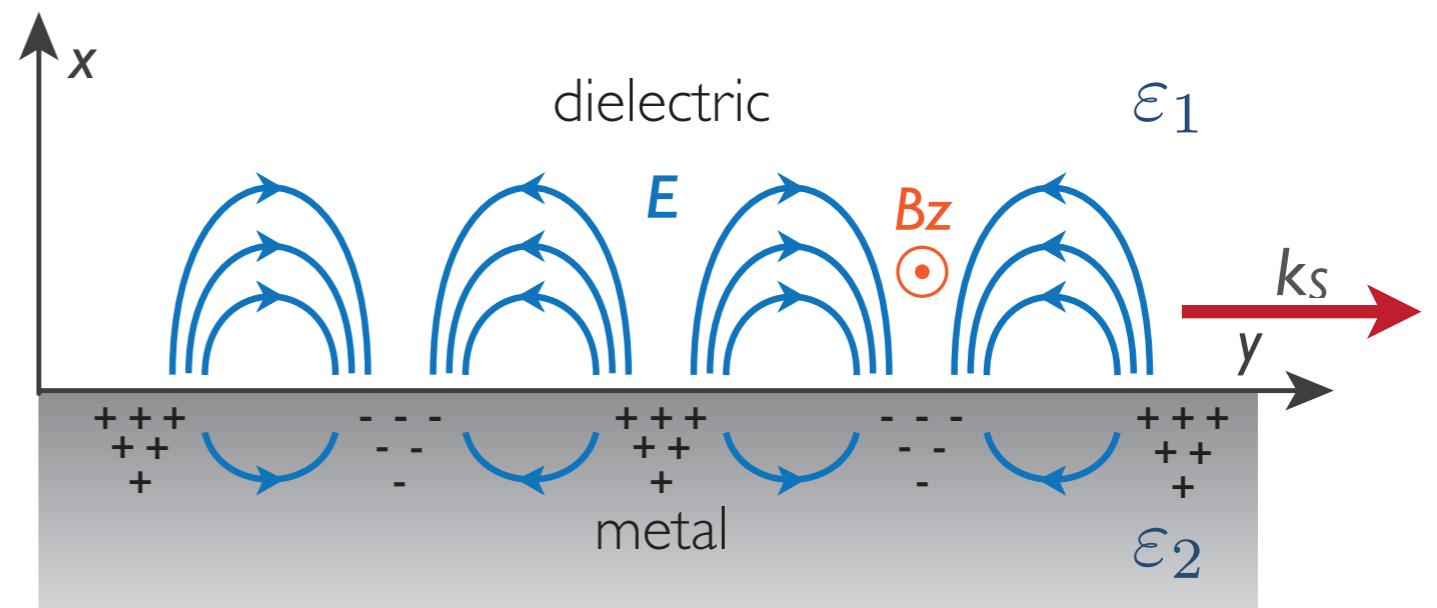
Electron acceleration using grating targets

acceleration by relativistic surface waves

Surface waves excitation using grating

collective e- oscillation
at the surface of a metal
or a plasma

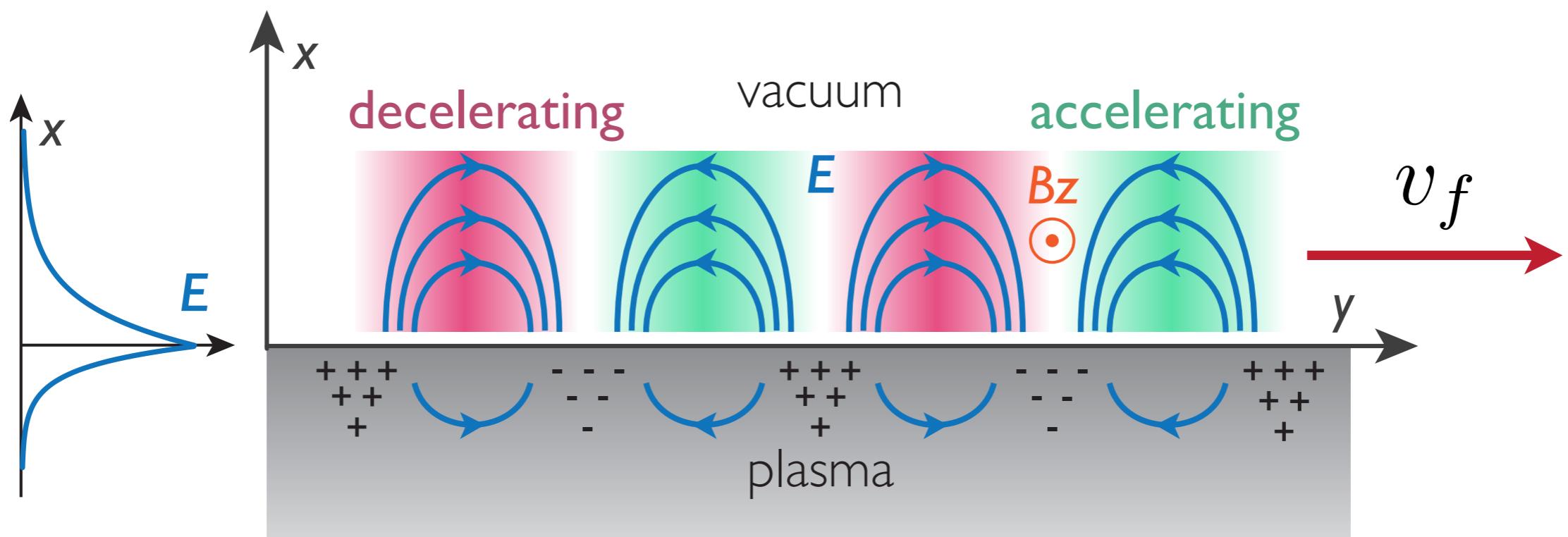
$$k_{SW}(\omega) = \frac{\omega}{c} \sqrt{\frac{\omega_p^2 - \omega^2}{\omega_p^2 - 2\omega^2}}$$



can be excited by a laser on a grating

$$\omega_L \sin(\theta) = k_{SW}(\omega_L) \pm nq$$

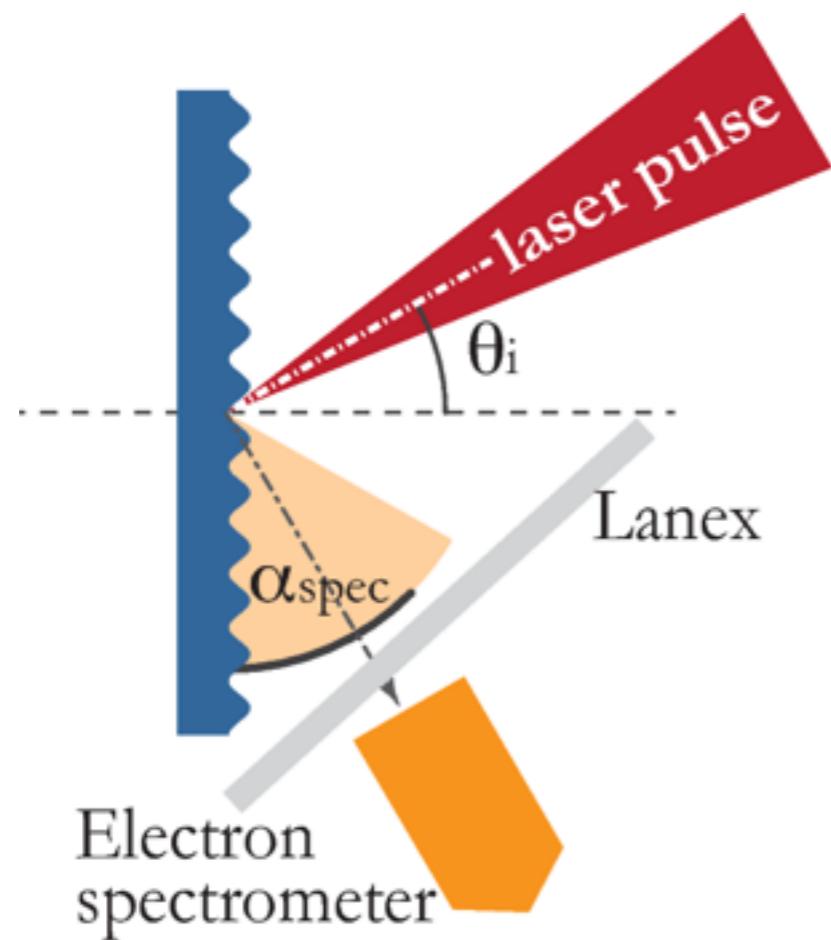
Electron acceleration by relativistic SW



electrons can phase-lock with SW
and be accelerated

Experiment

@ SLIC CEA-Saclay France (group of T. Ceccotti)



Laser:
UHII00 laser at CEA Saclay - France
P-polarised, contrast 10^{12}

$$I \sim 5 \cdot 10^{19} \text{W/cm}^2$$

$$\tau \sim 25 \text{fs}$$

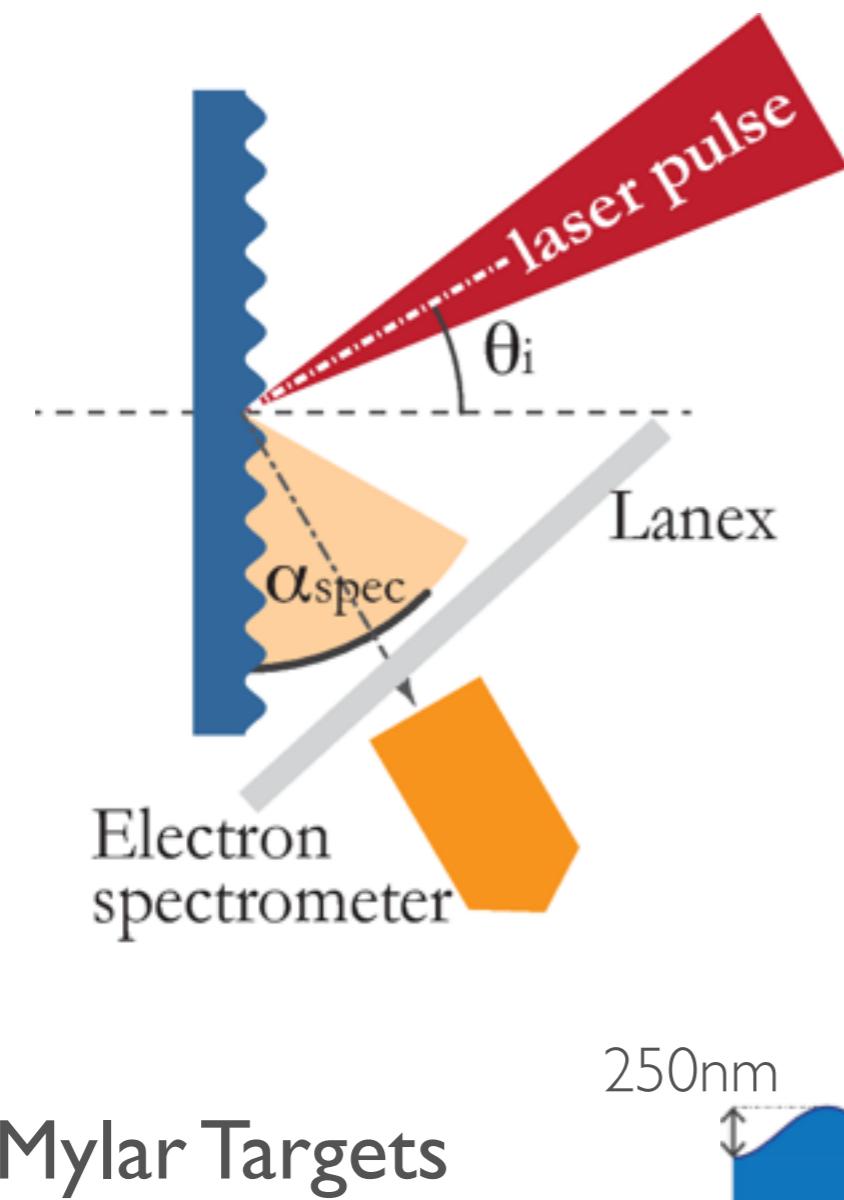
$$\lambda \sim 0,8 \mu\text{m}$$

$$w_0 \sim 4 \mu\text{m}$$



Experiment

@ SLIC CEA-Saclay France (group of T. Ceccotti)



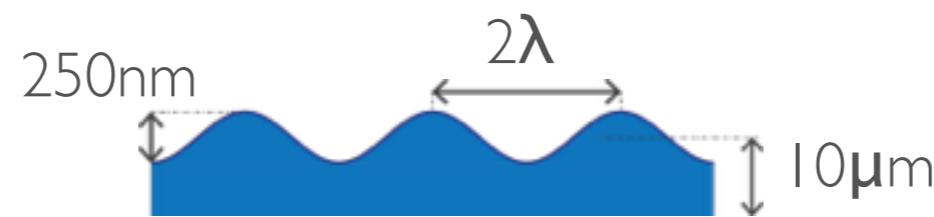
Laser:
UHII00 laser at CEA Saclay - France
P-polarised, contrast 10^{12}

$$I \sim 5 \cdot 10^{19} \text{ W/cm}^2$$

$$\tau \sim 25 \text{ fs}$$

$$\lambda \sim 0,8 \mu\text{m}$$

$$w_0 \sim 4 \mu\text{m}$$

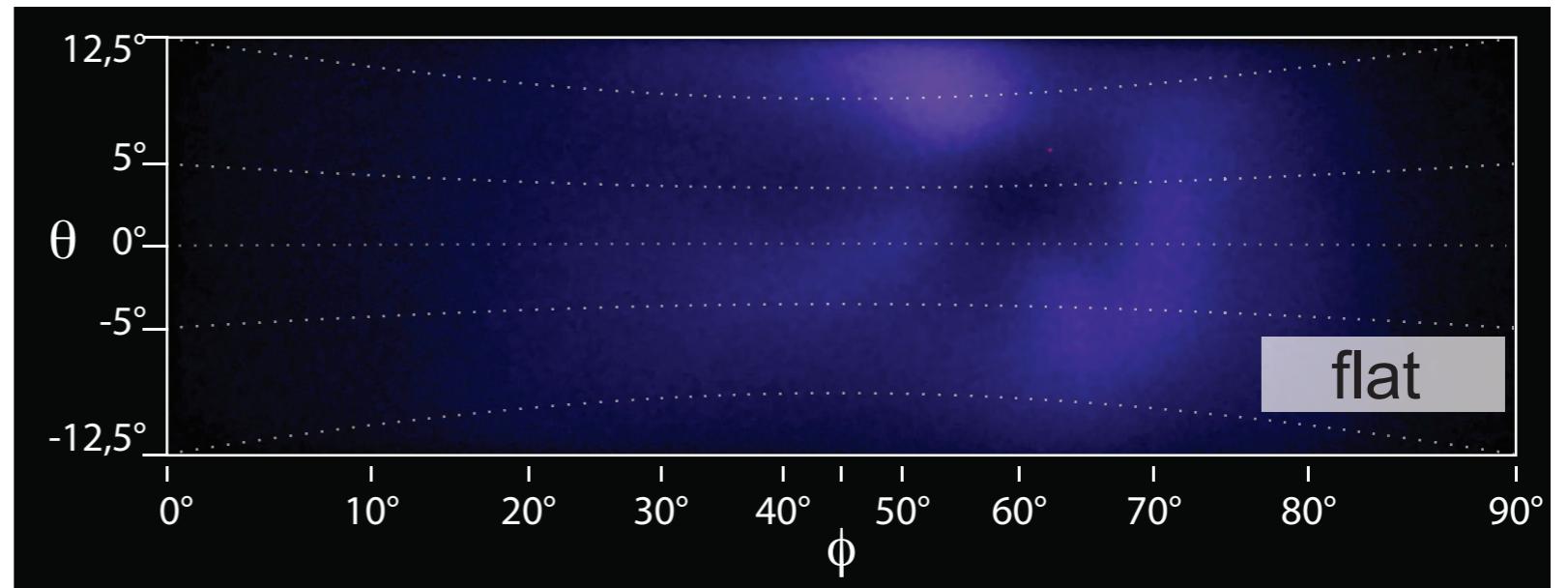
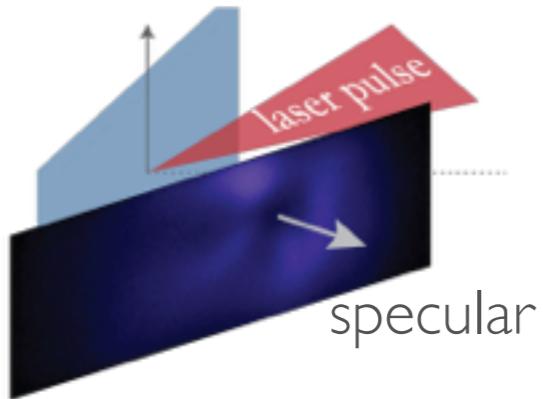


$$\theta_{\text{res.}} = 30^\circ$$

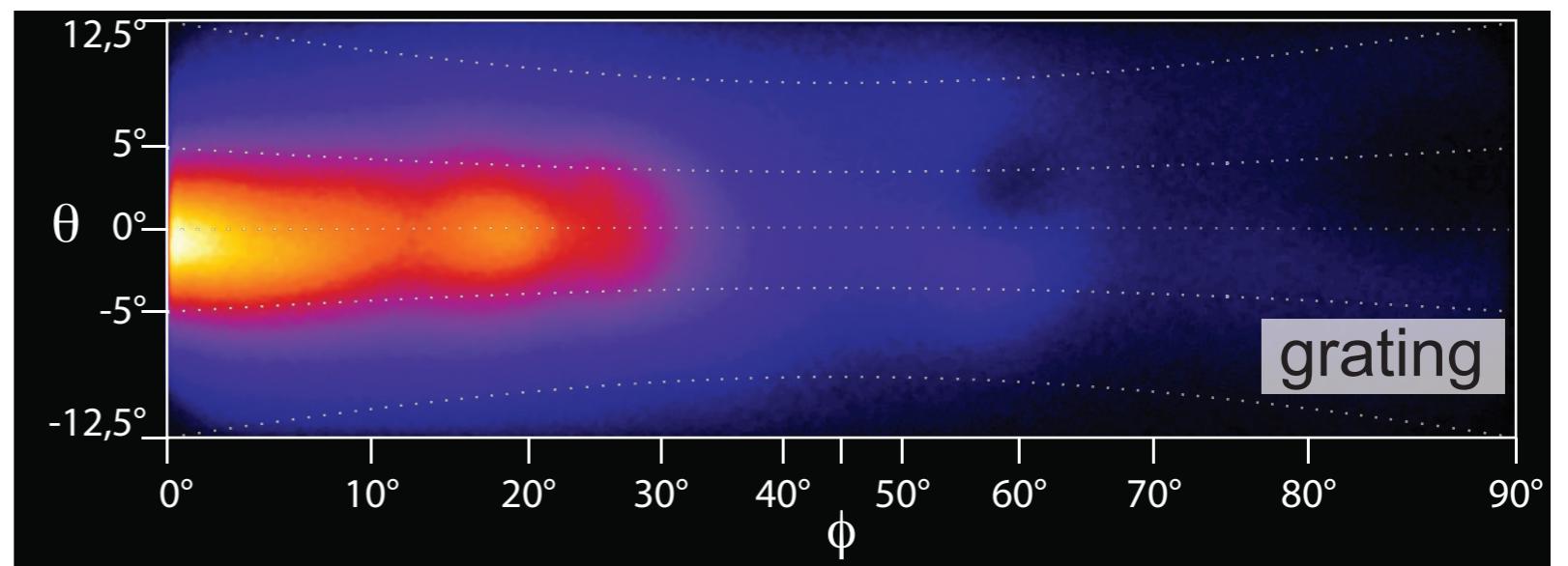
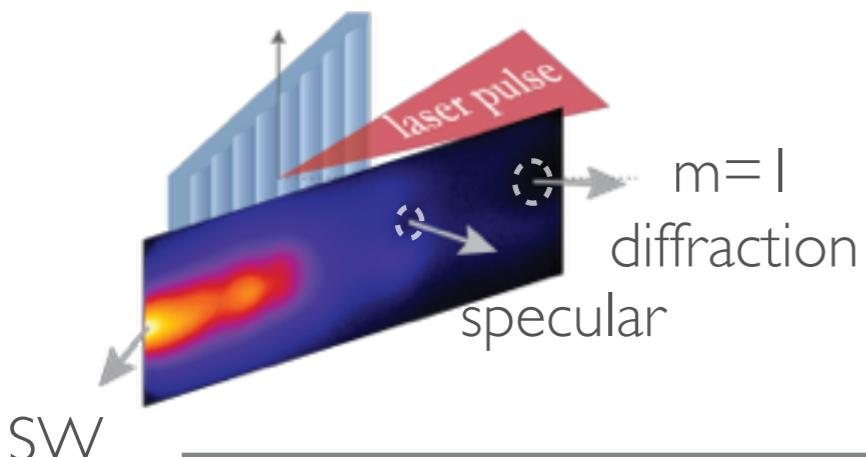
Electron emission from the front side

$E > 1.5 \text{ MeV}$

Flat target:



Grating @ resonance

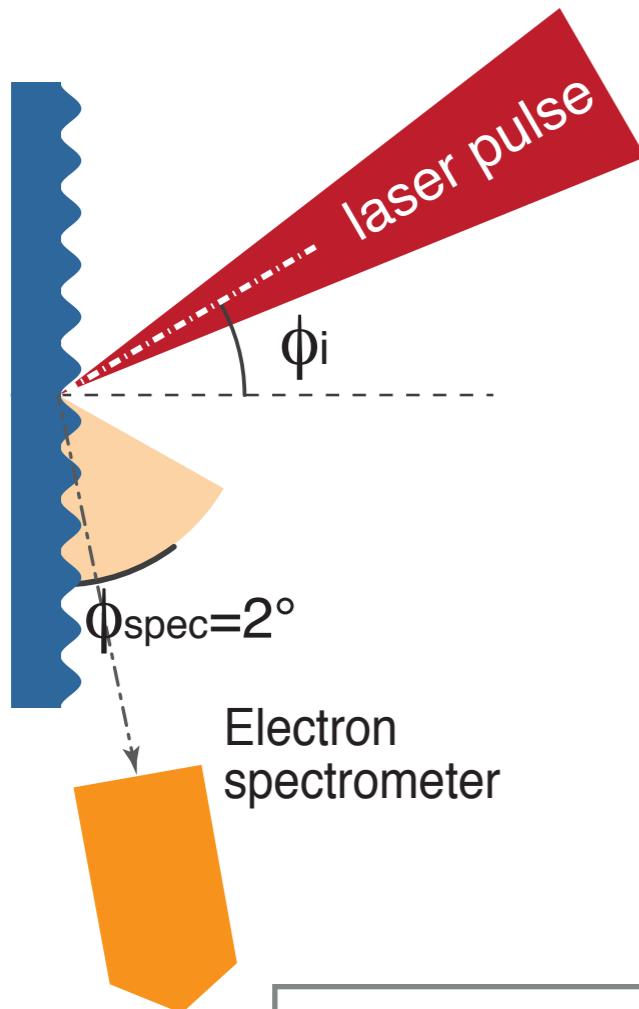


SW

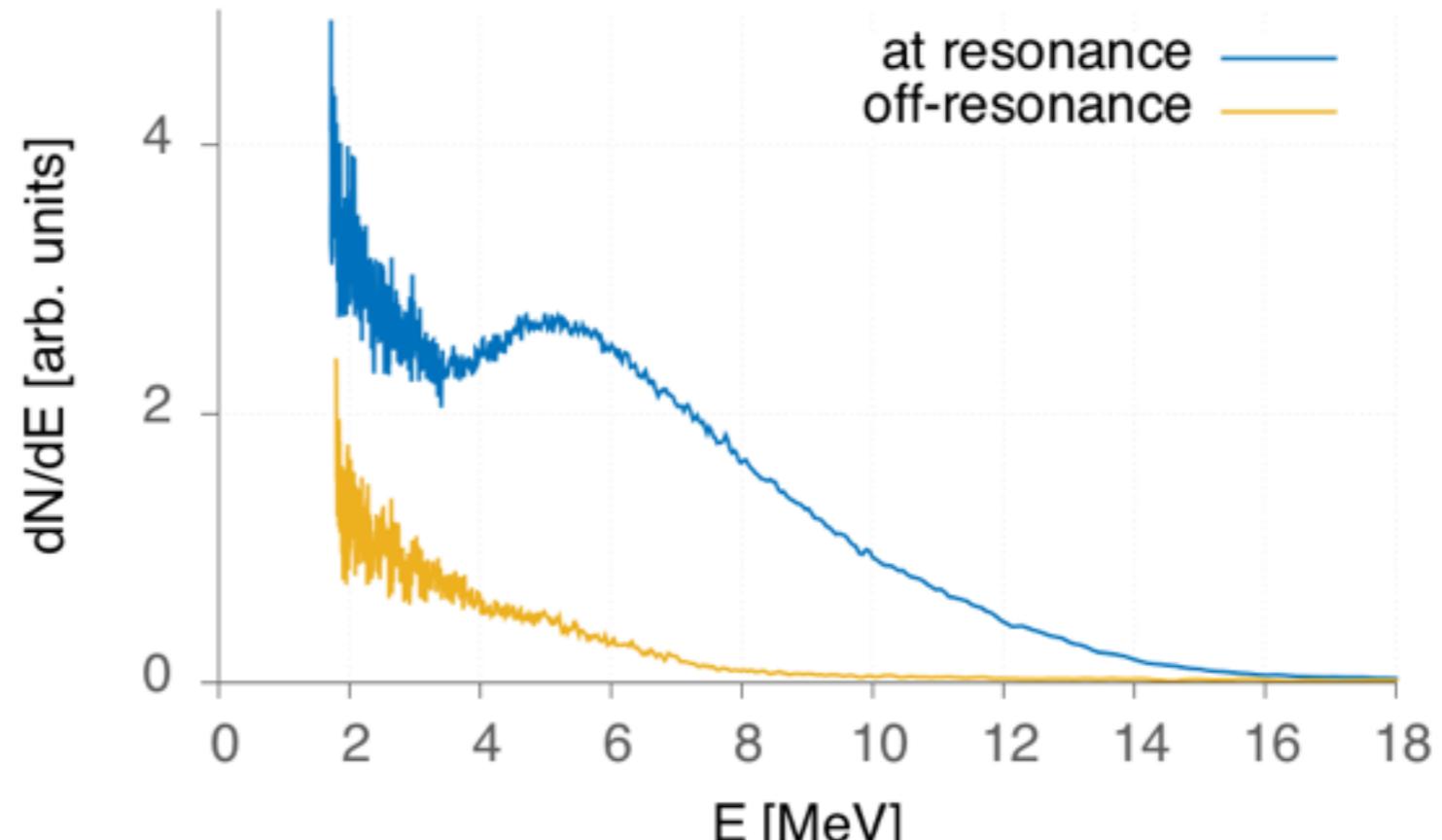
- » **narrow** emission along the **surface**
- » with grating 10x higher flux: **$> 100 \text{ pC in } \sim 100 \text{ mrad}$**

Electron spectrometer: grating targets

$\phi_{\text{res.}} = 30^\circ$



spectrometer kept at $\Phi_{\text{spec}} \sim 2^\circ$

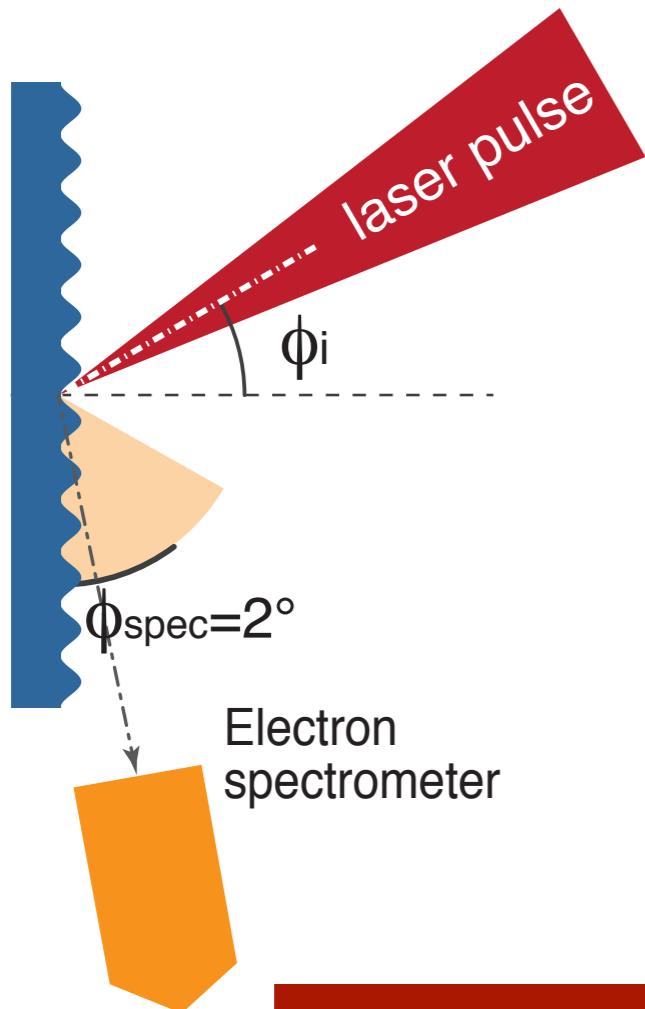


» non-thermal spectrum
» peak at 6-10 MeV

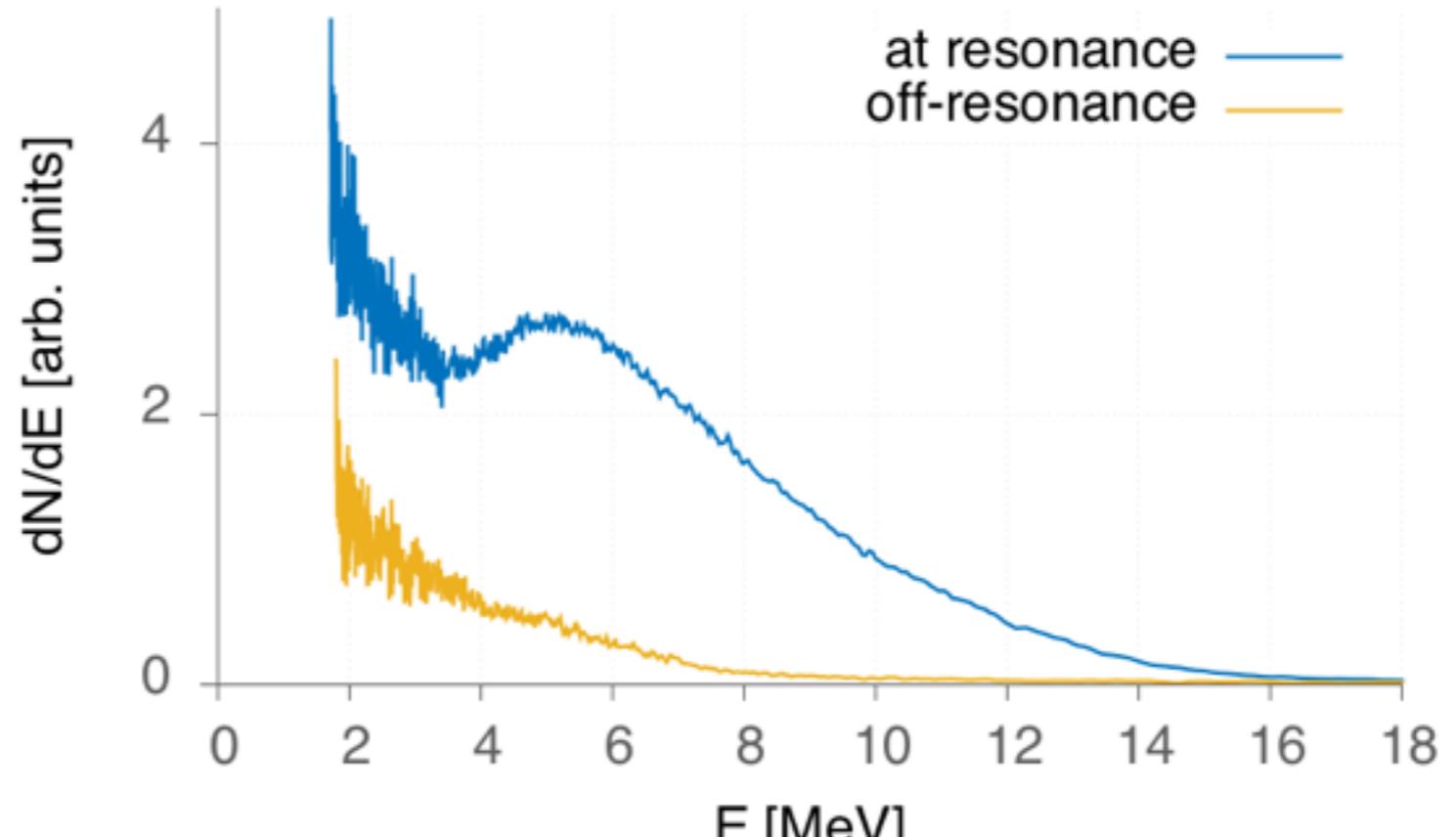
» high energy
» small signal out res.

Electron spectrometer: grating targets

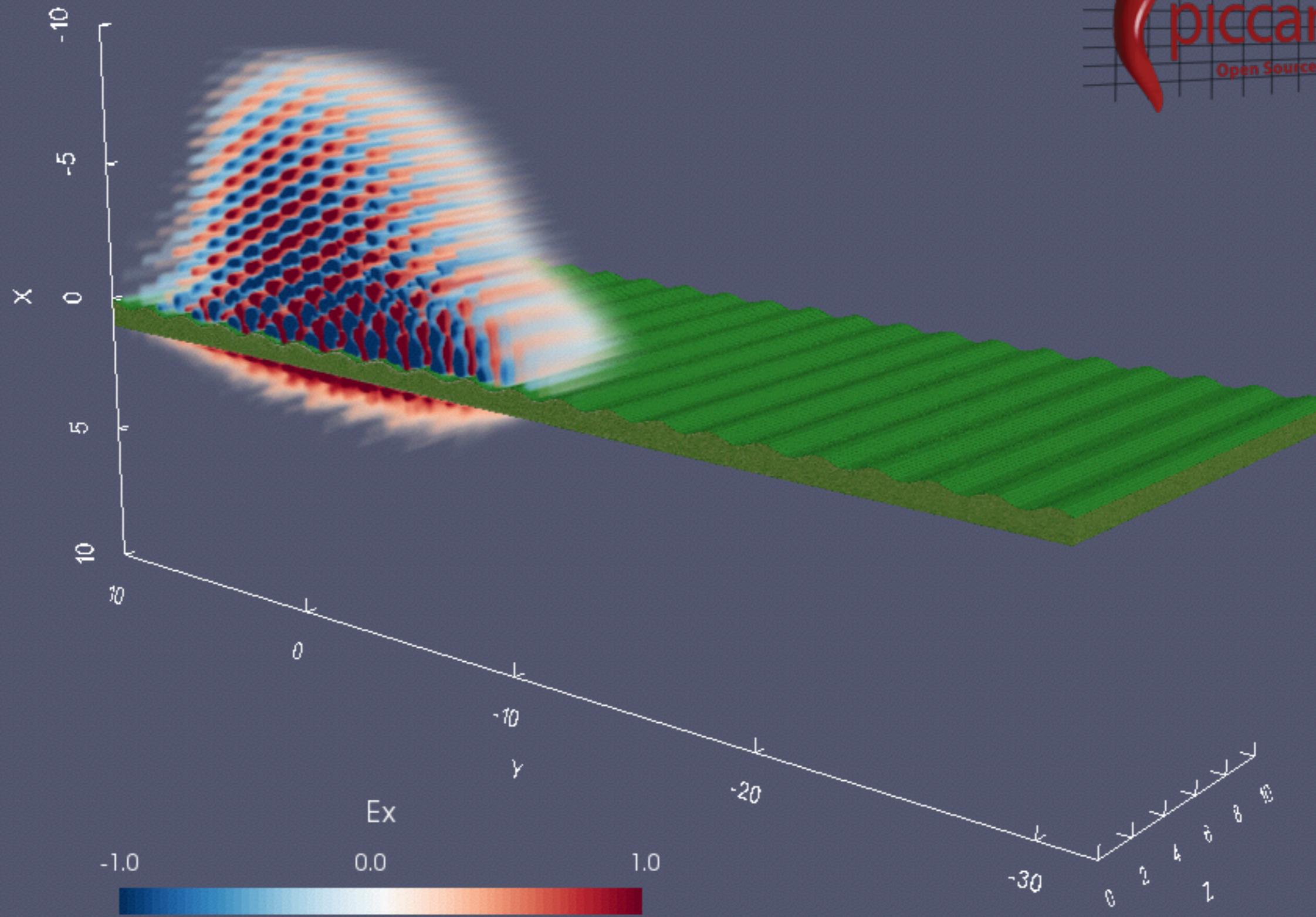
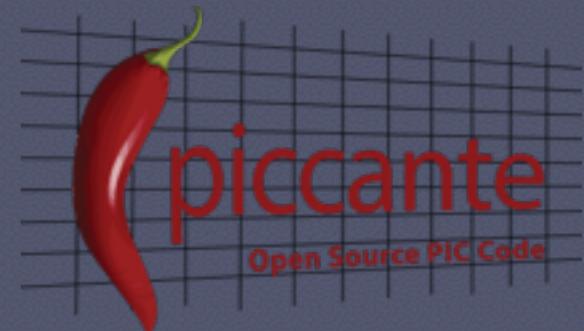
$\phi_{\text{res.}} = 30^\circ$

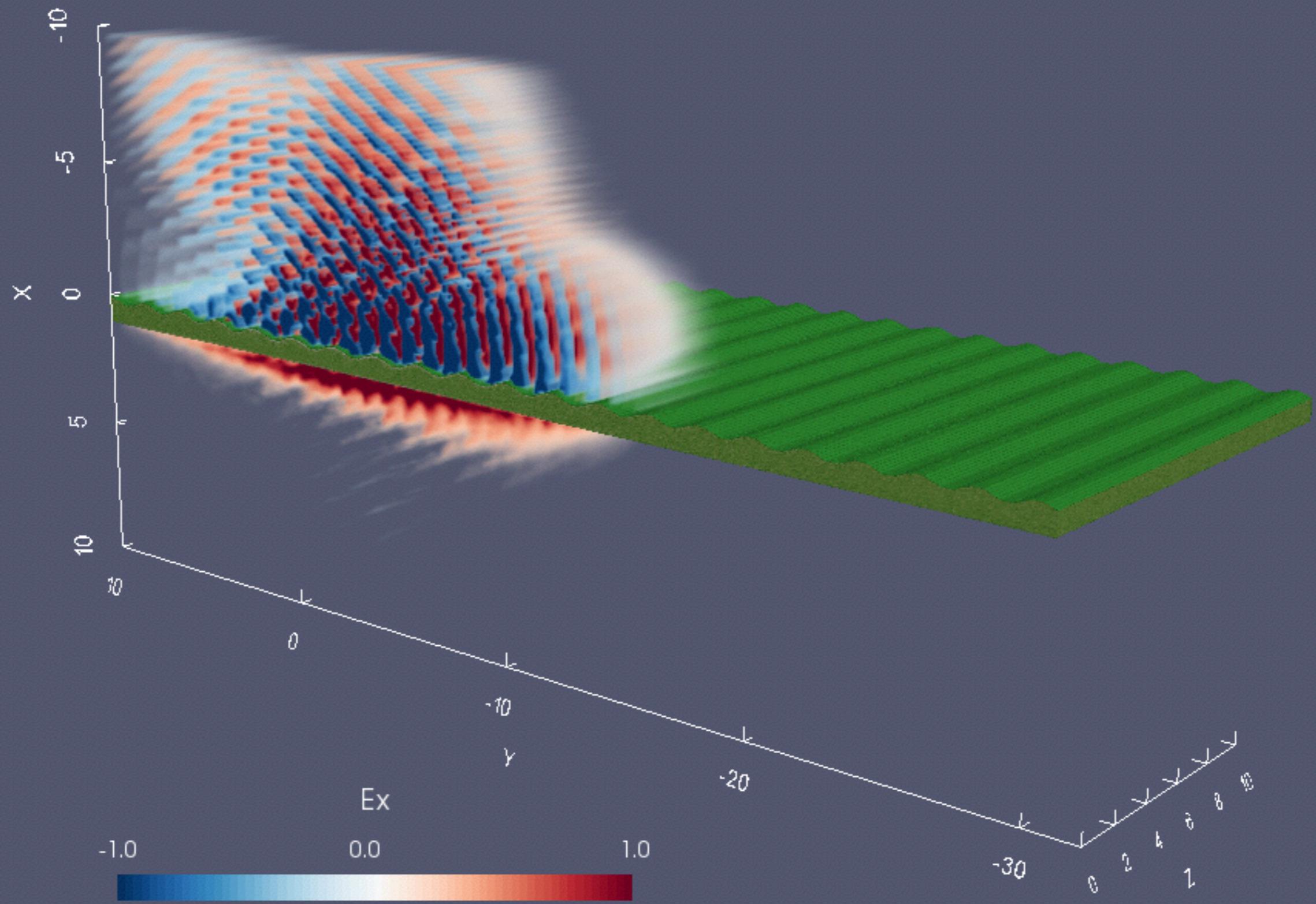


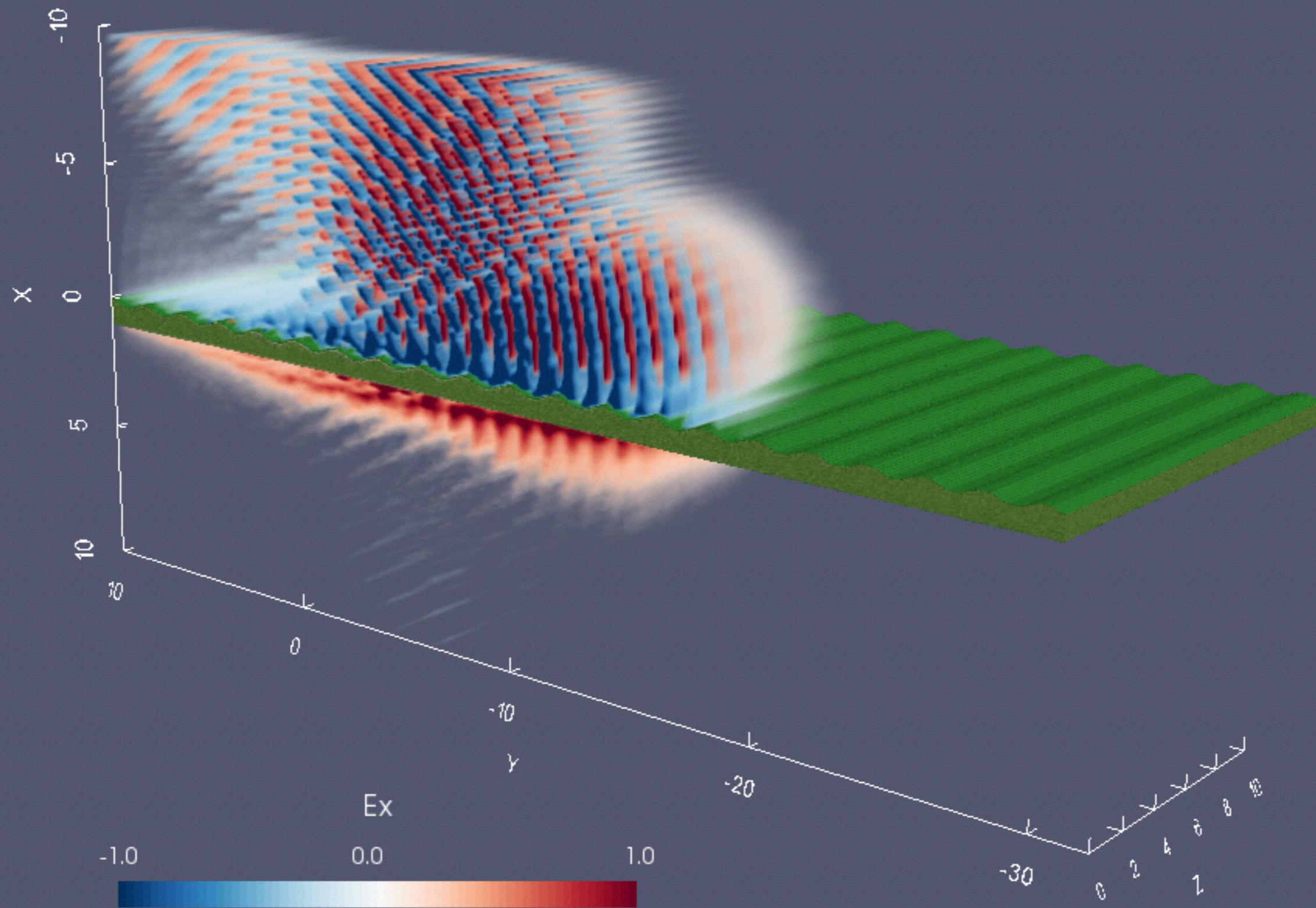
spectrometer kept at $\Phi_{\text{spec}} \sim 2^\circ$

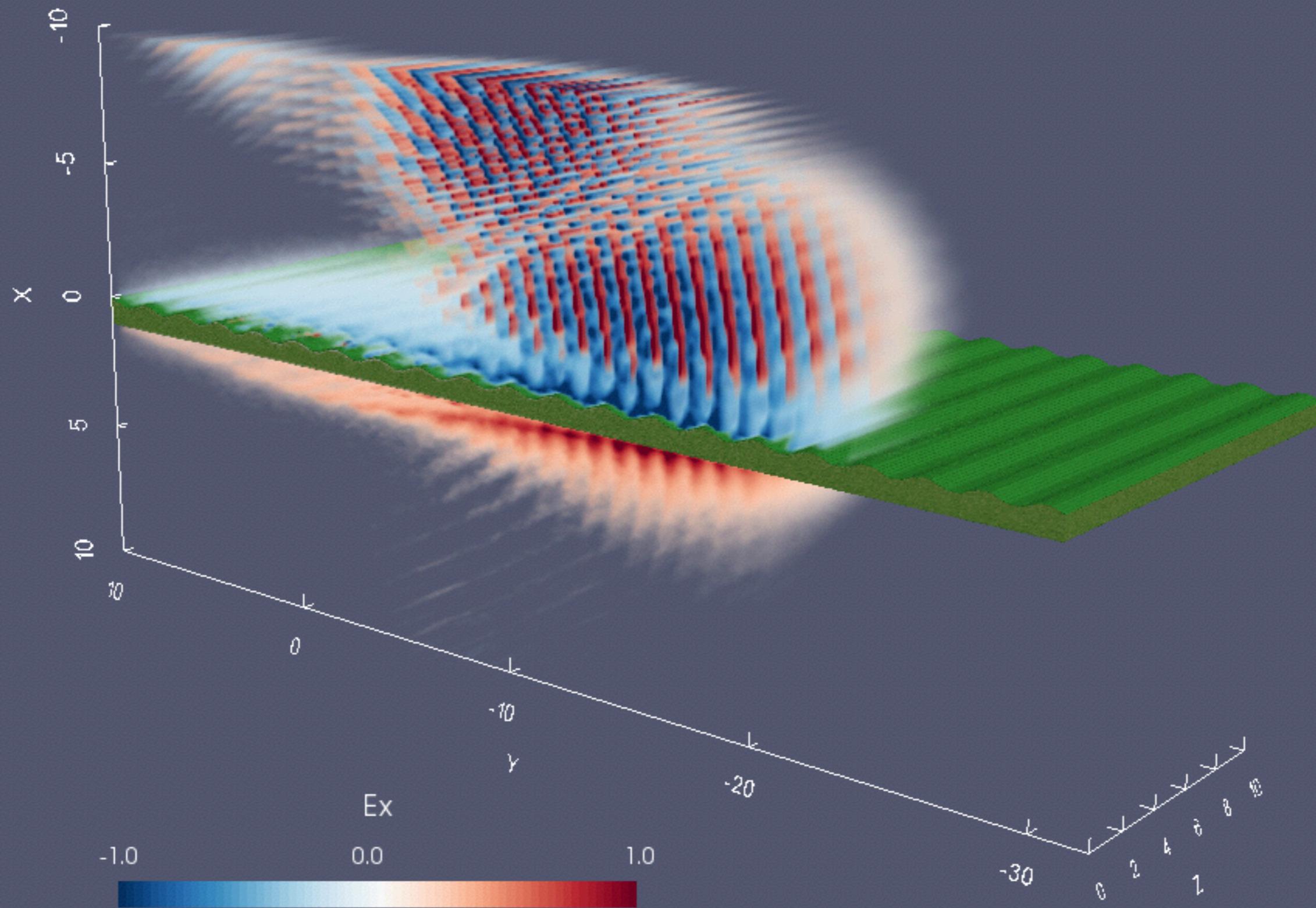


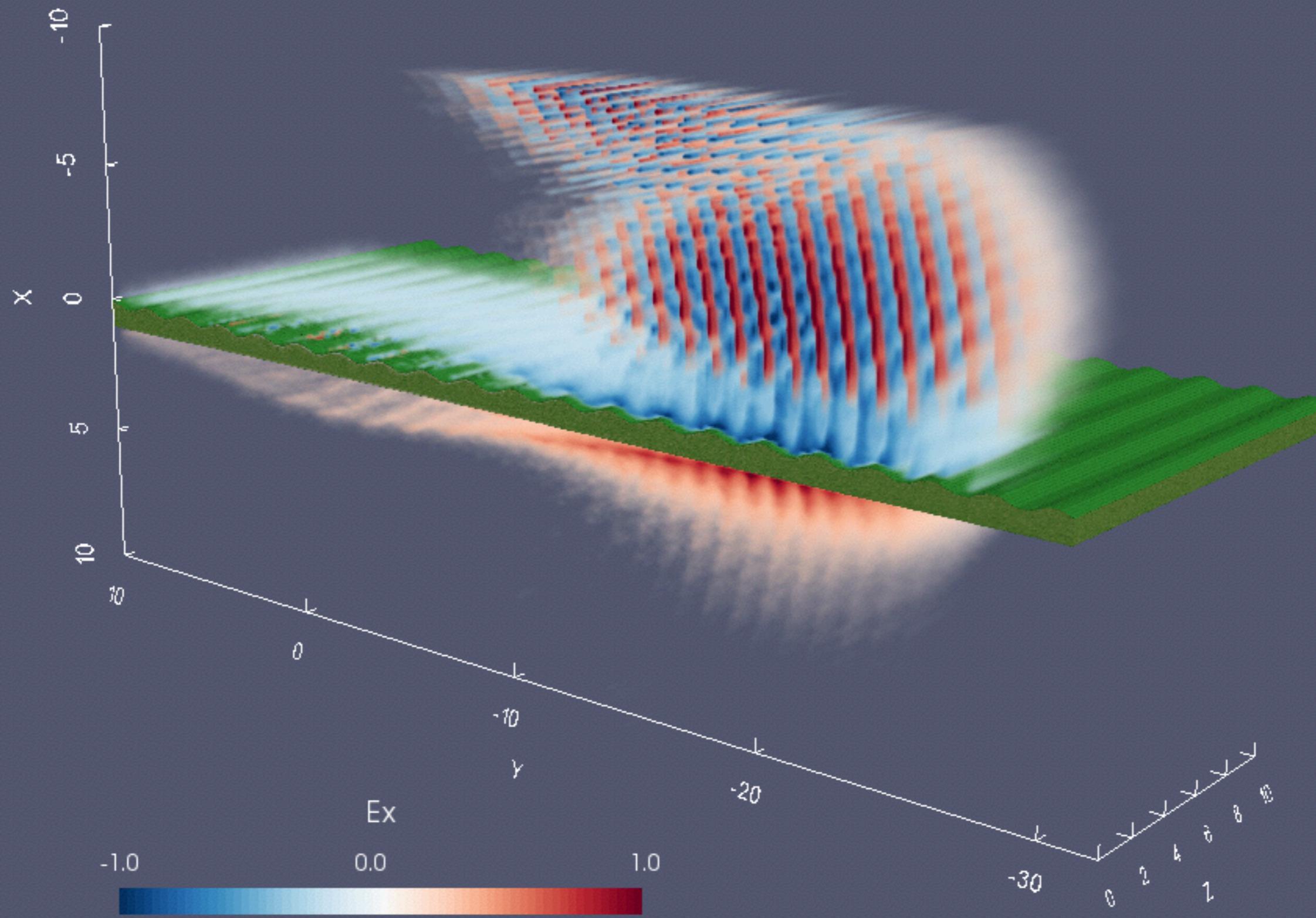
Similar results obtained with grating for different angle of resonance!!! $\phi_{\text{res.}} = 15^\circ$ or 45°

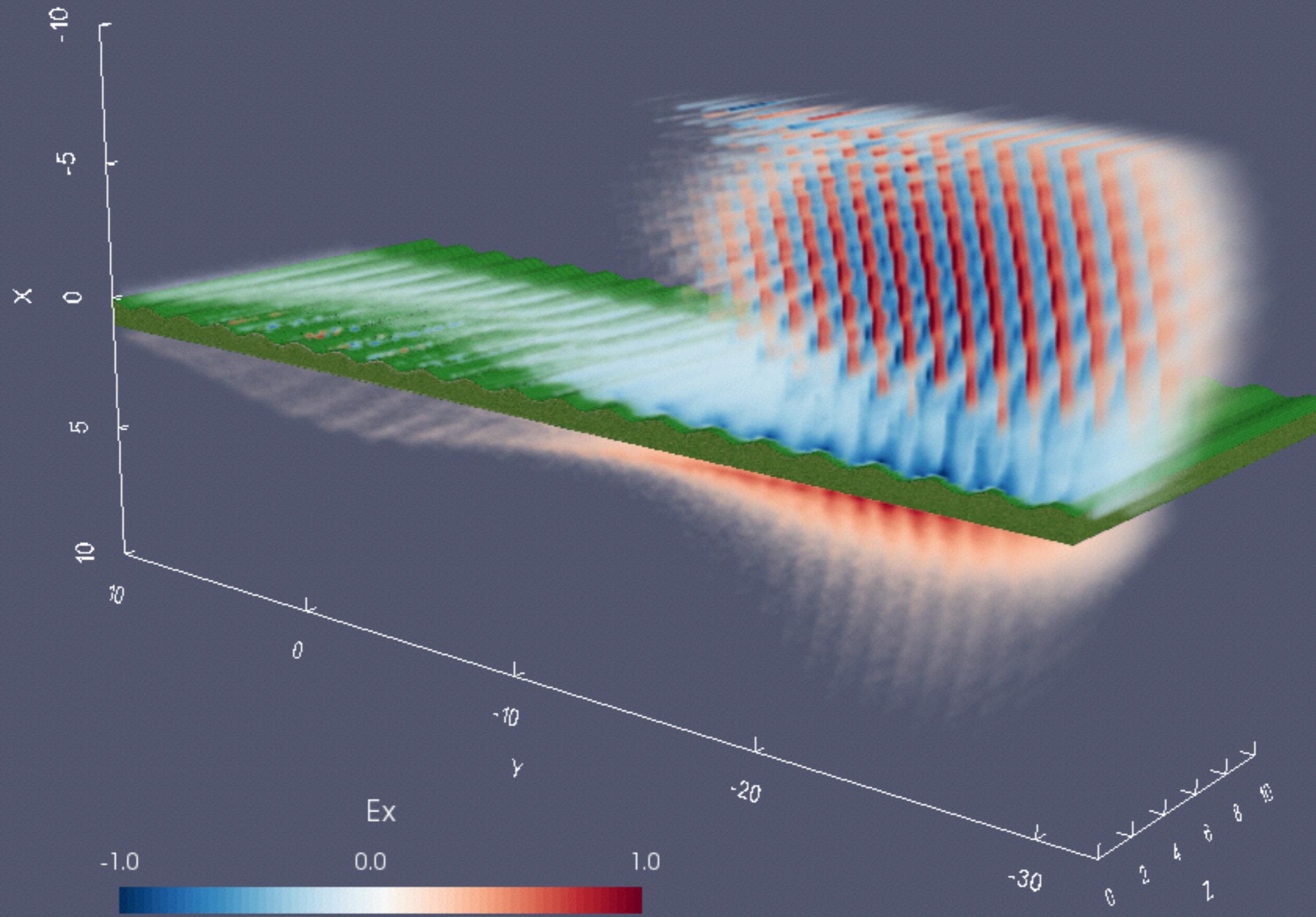




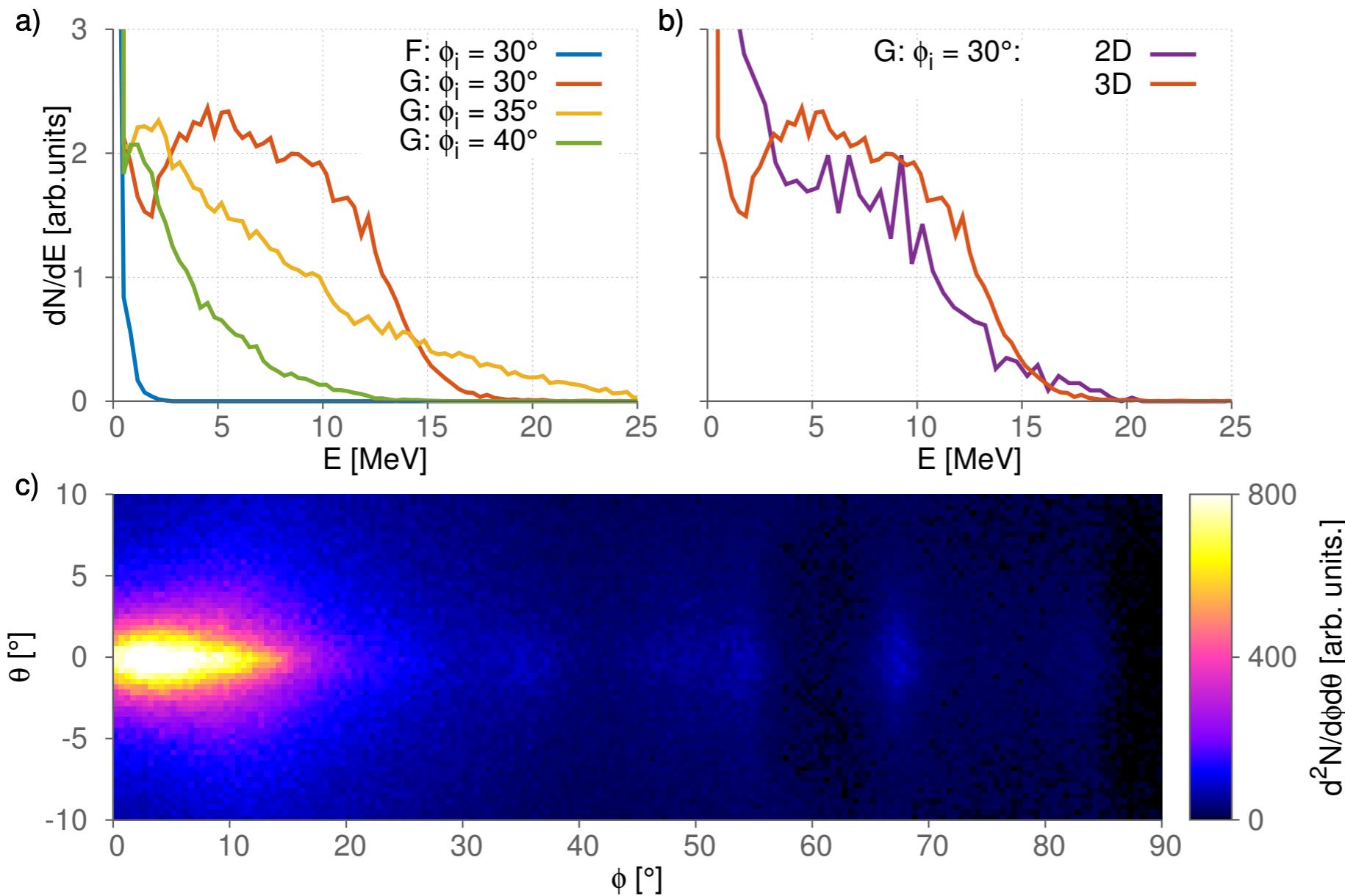




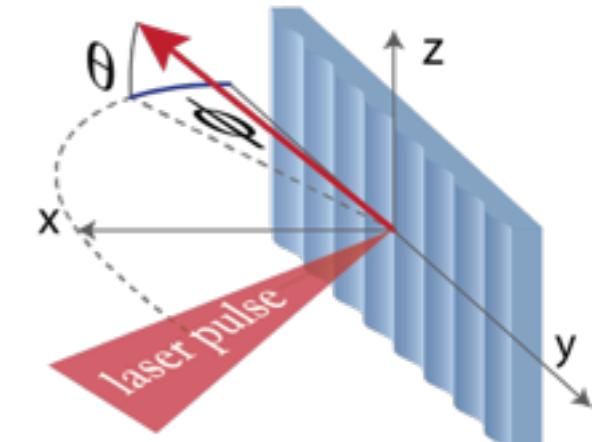




3D simulations



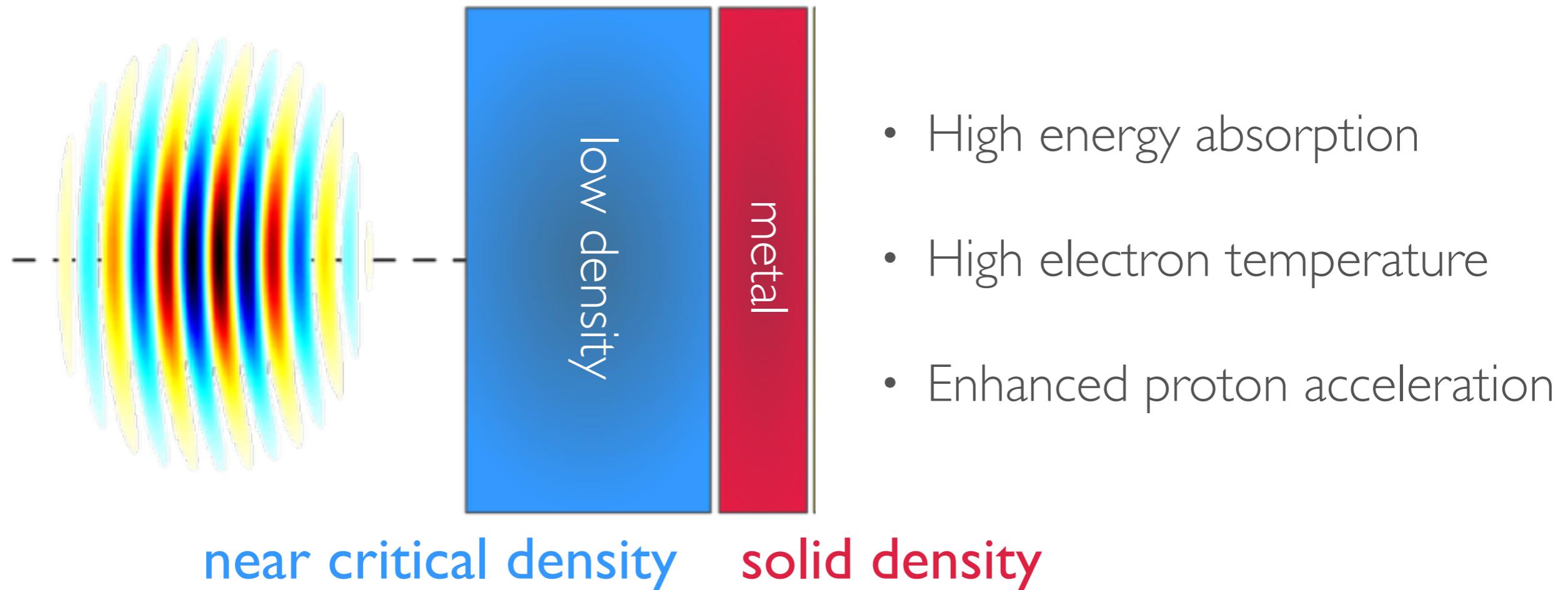
at $\phi=2^\circ \Theta=0^\circ$



Only the 3D simulation reproduces correctly the **e-spectra**

Targets for enhanced TNSA
thin solid foils with a thick low density foam layer

2 Layer target for enhanced TNSA



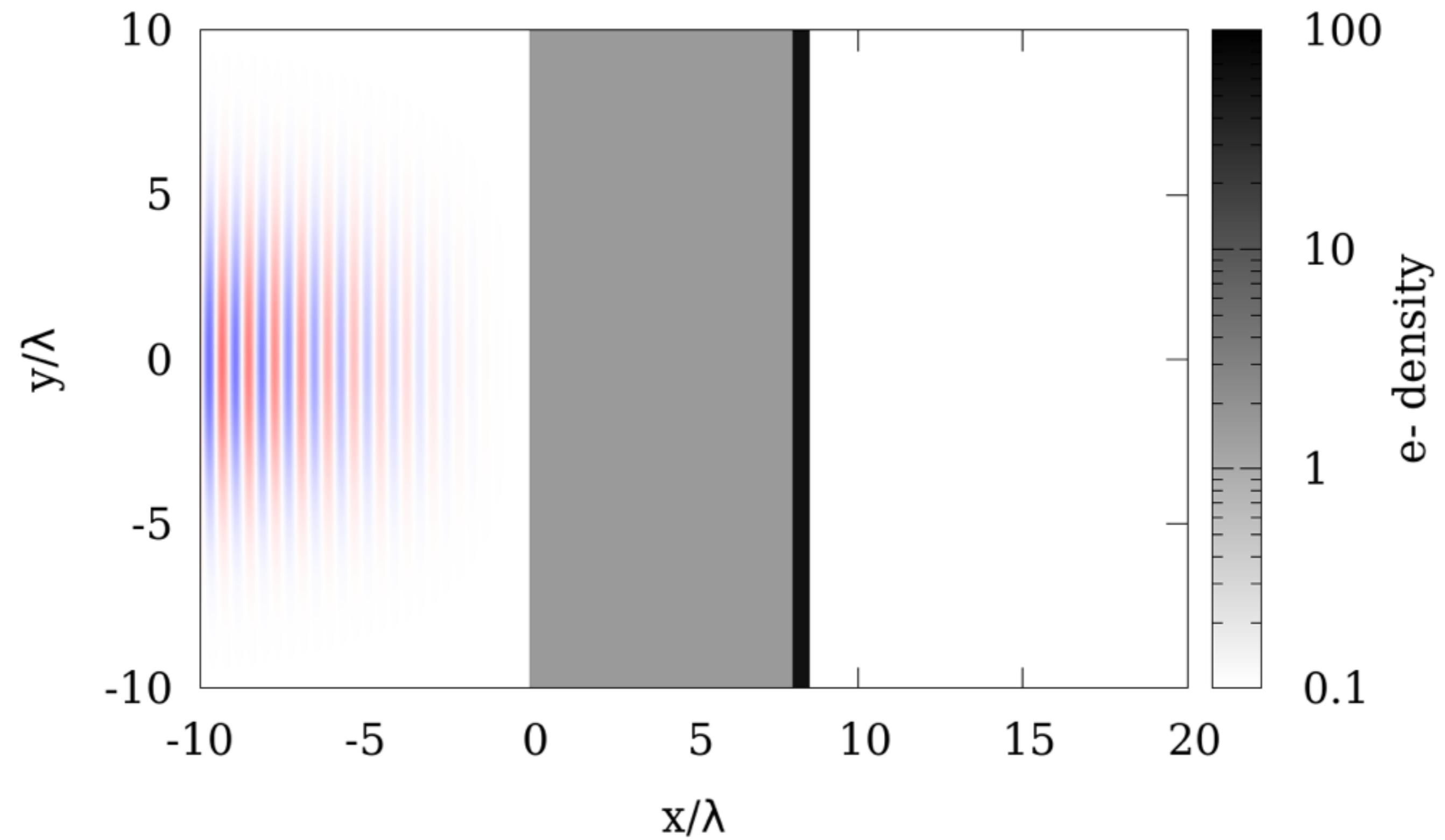
Nakamura et. al. PoP **17**, (2010)

Passoni et. al. PPCF 56(2014)

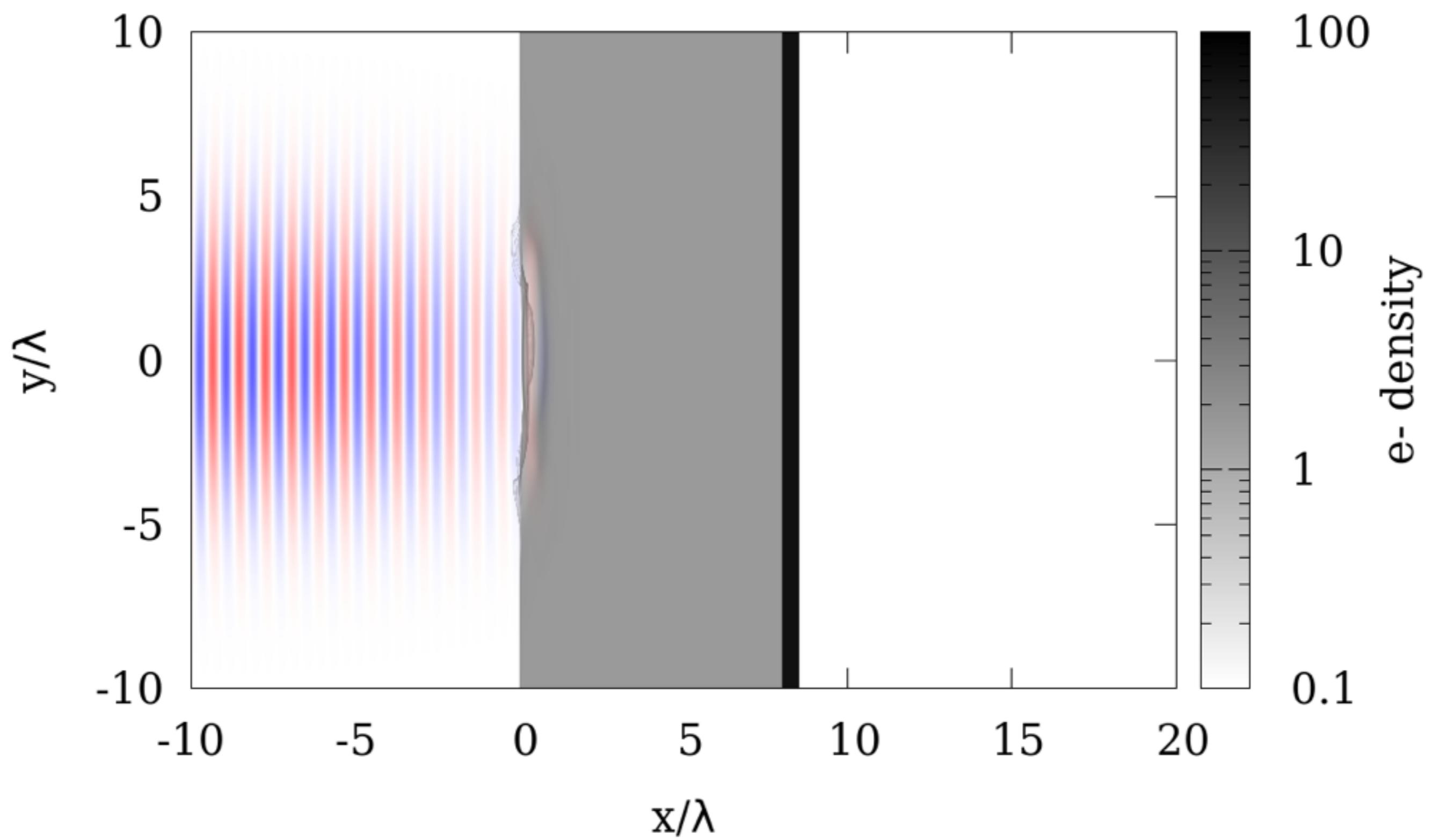
Sgattoni et. al. PRE **85** (2012)

J. H. Bin, et. al. PRL **115**(2015)

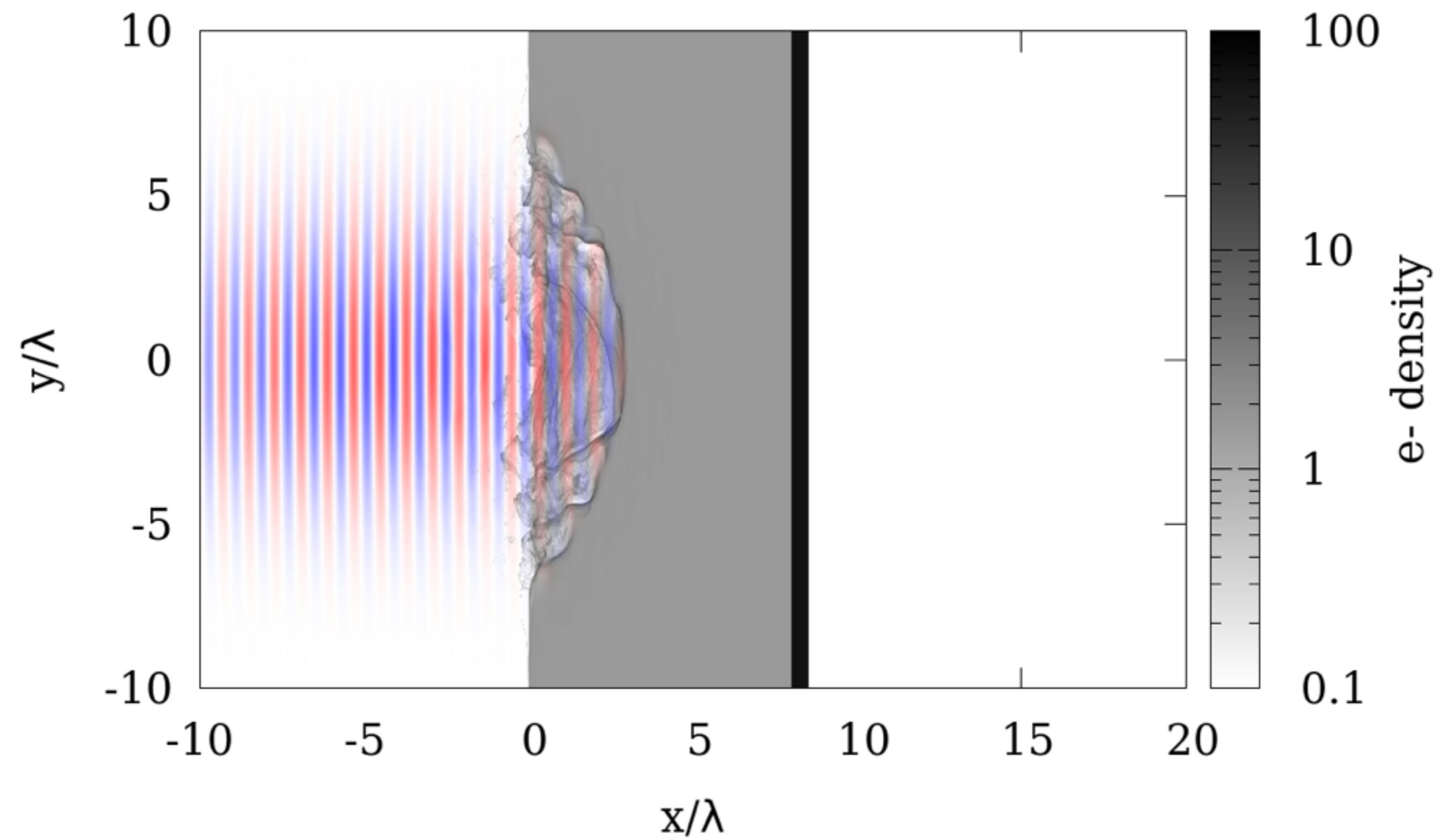
$T = 00$



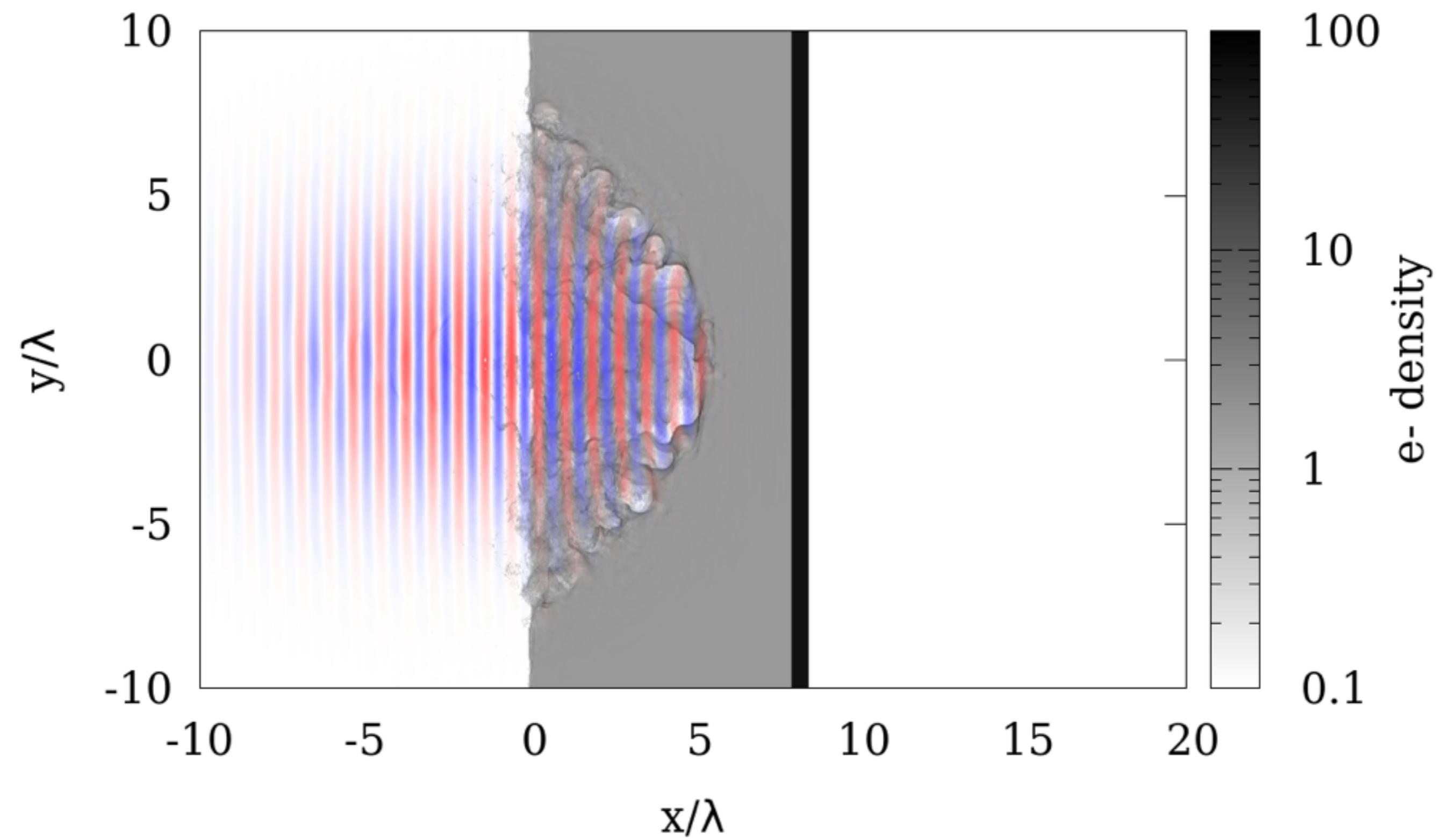
$T = 04$



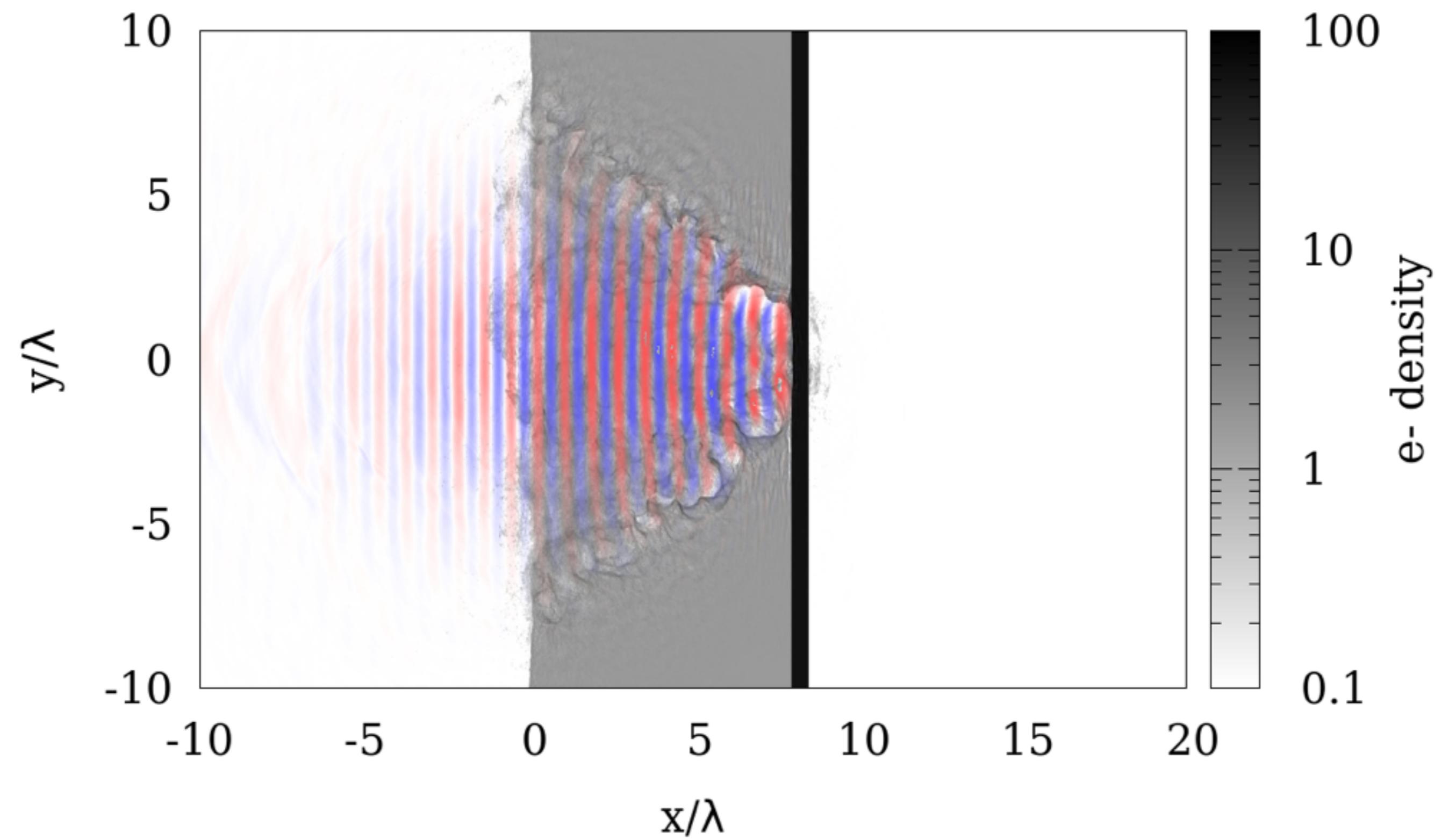
$T = 08$



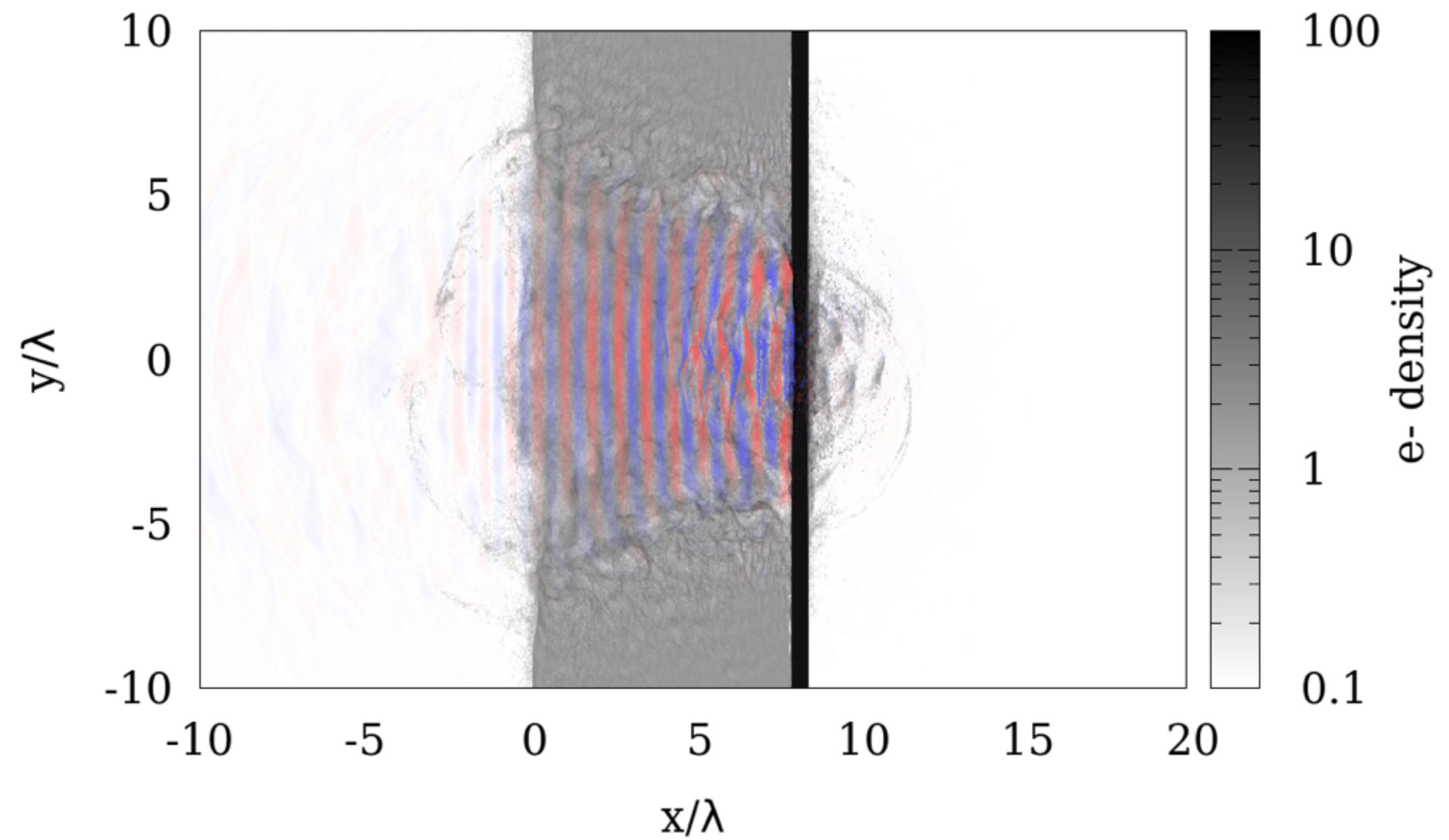
$T = 12$



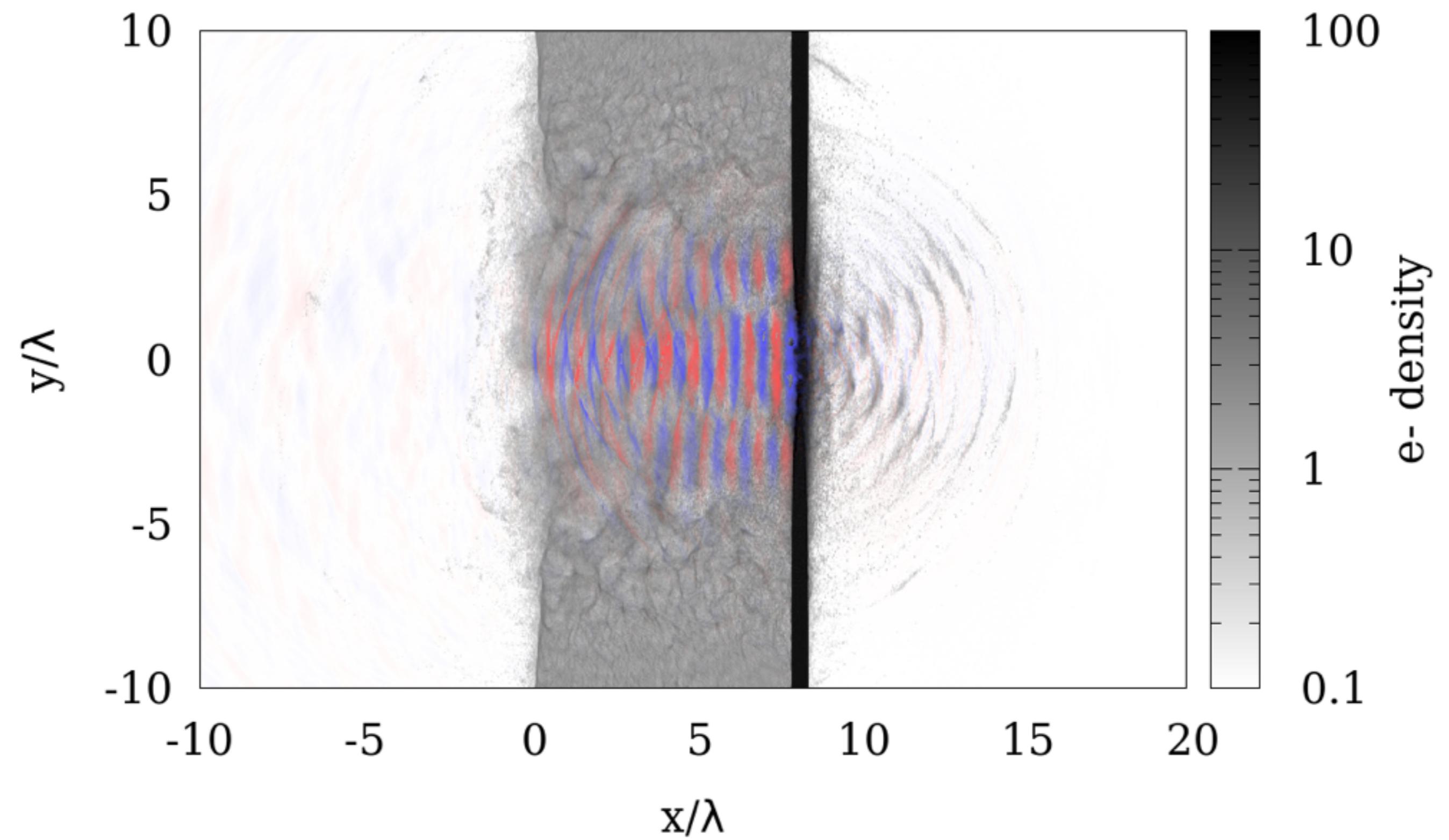
$T = 16$



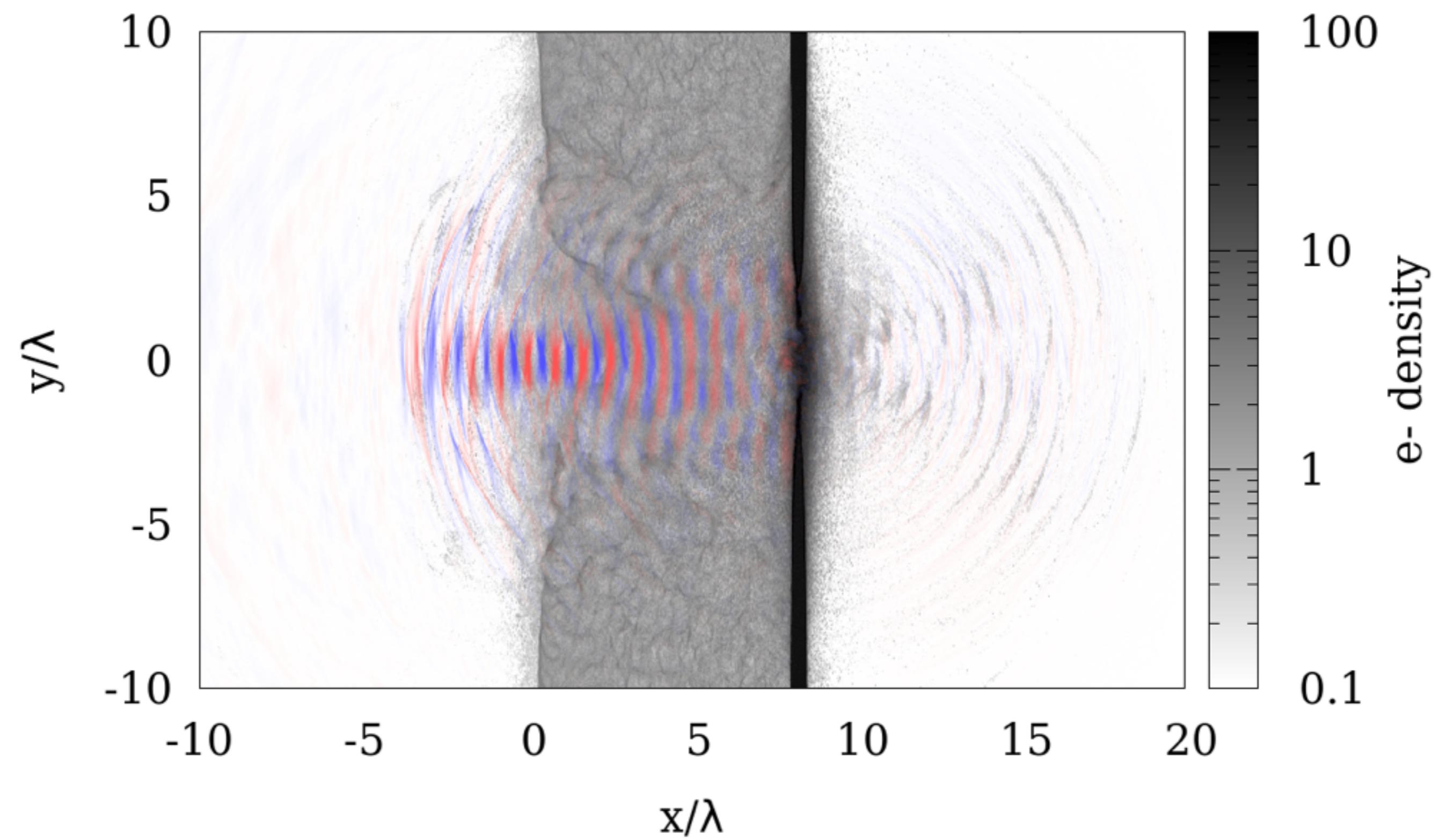
$T = 20$

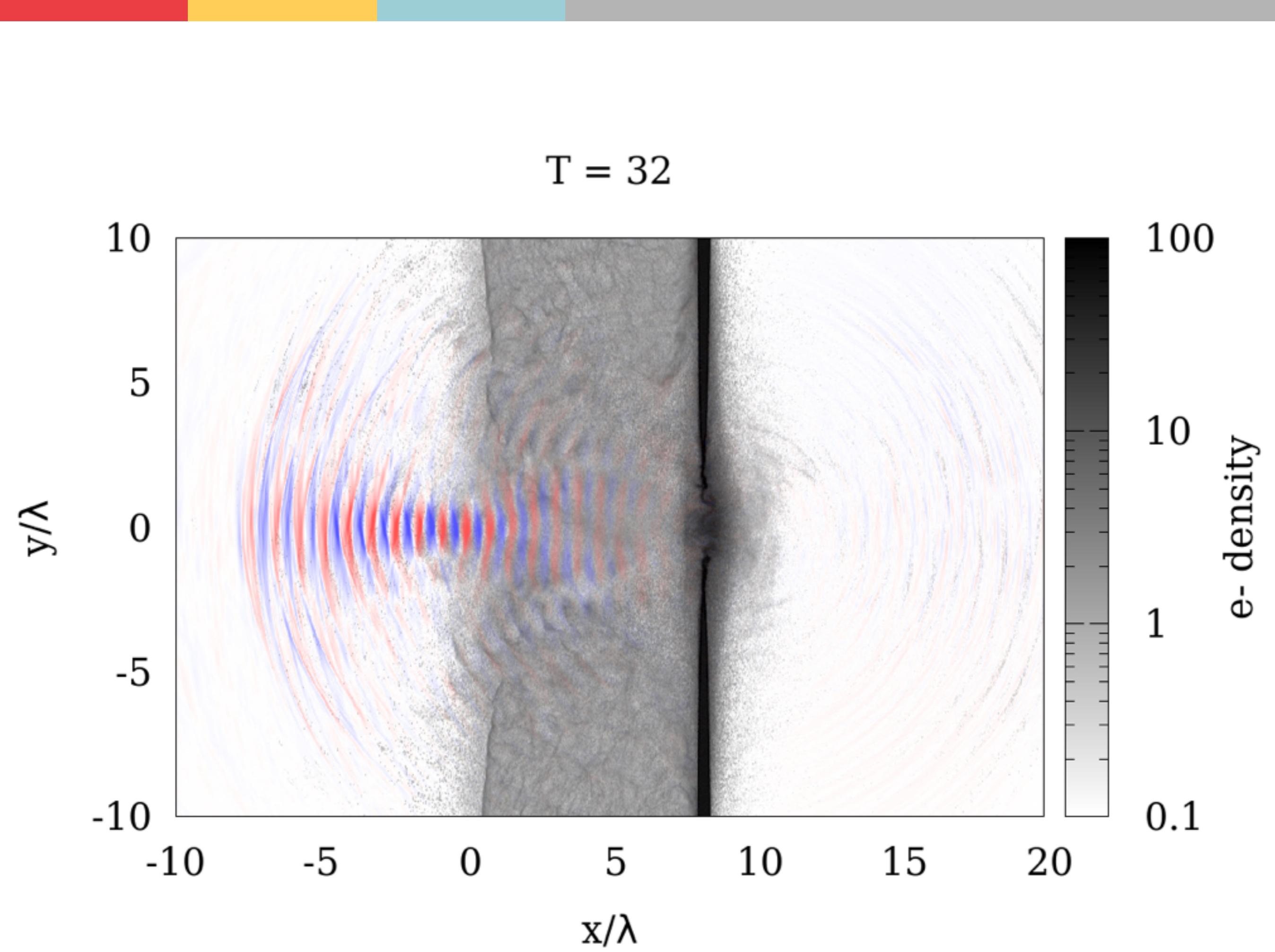


$T = 24$



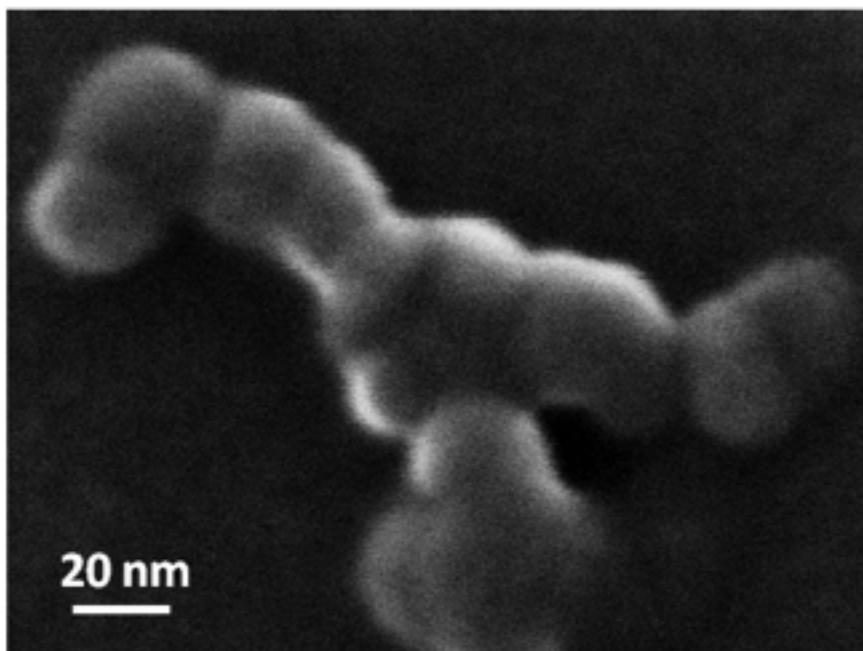
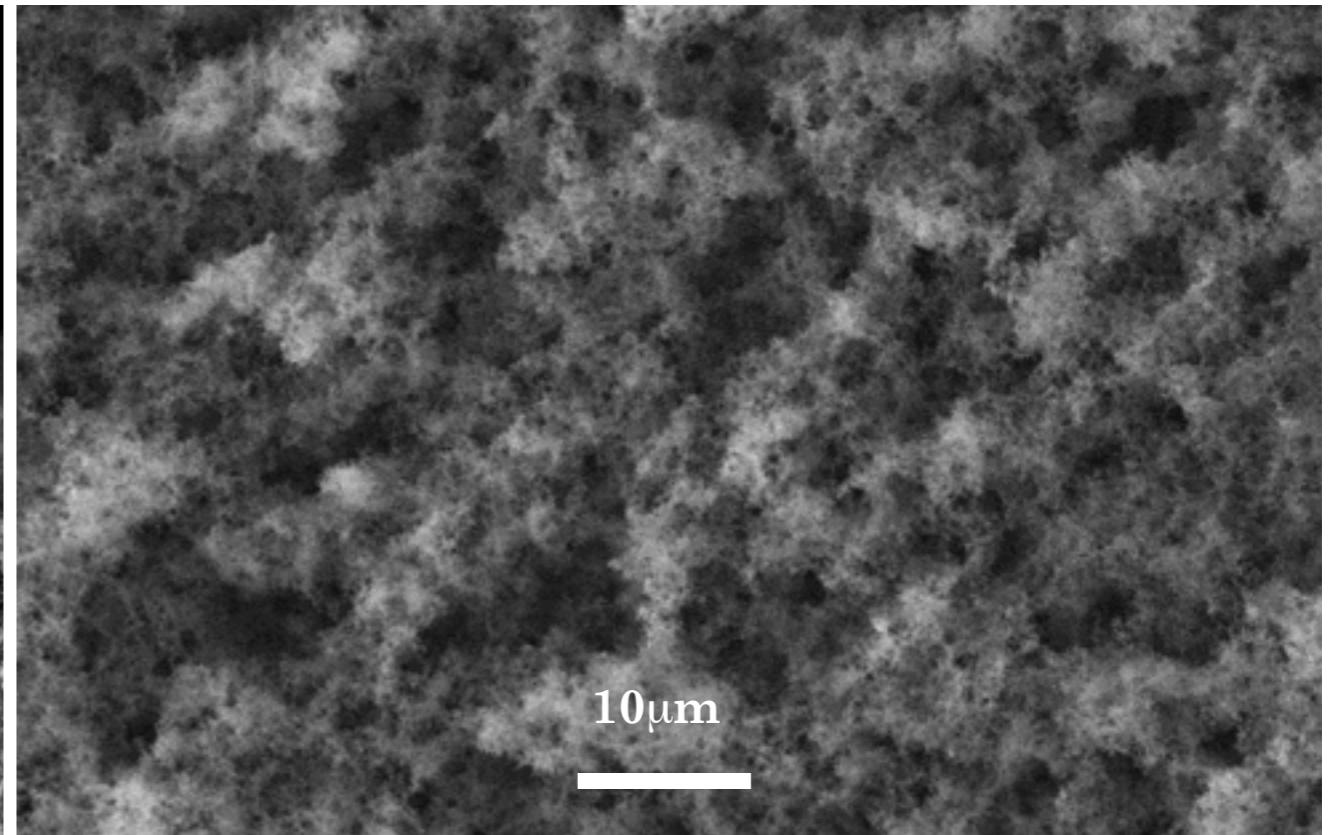
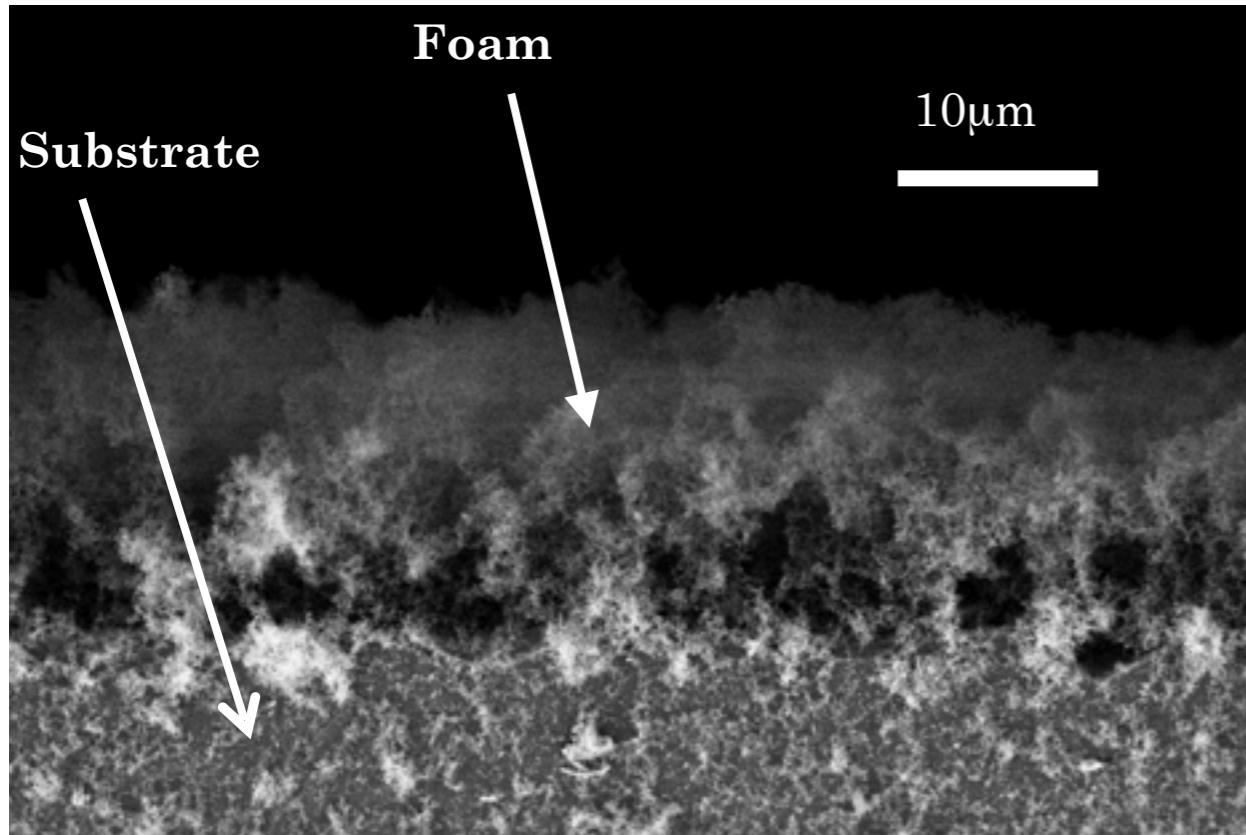
$T = 28$





Pulsed Laser Deposition

group of Passoni @ Politecnico di Milano



- » Pulsed Laser Deposition (PLD)
- » **Carbon** “foam” grown solid foils
- » $5 \sim 10 \text{ mg/cm}^3$

Zani et al, Carbon **56** (2013)

experiment @ GIST Gwangju South Korea



Proton maximum Energy @ GIST

30 fs, spot 4 μm $5 \cdot 10^{20} \text{W/cm}^2$, incidence 30°

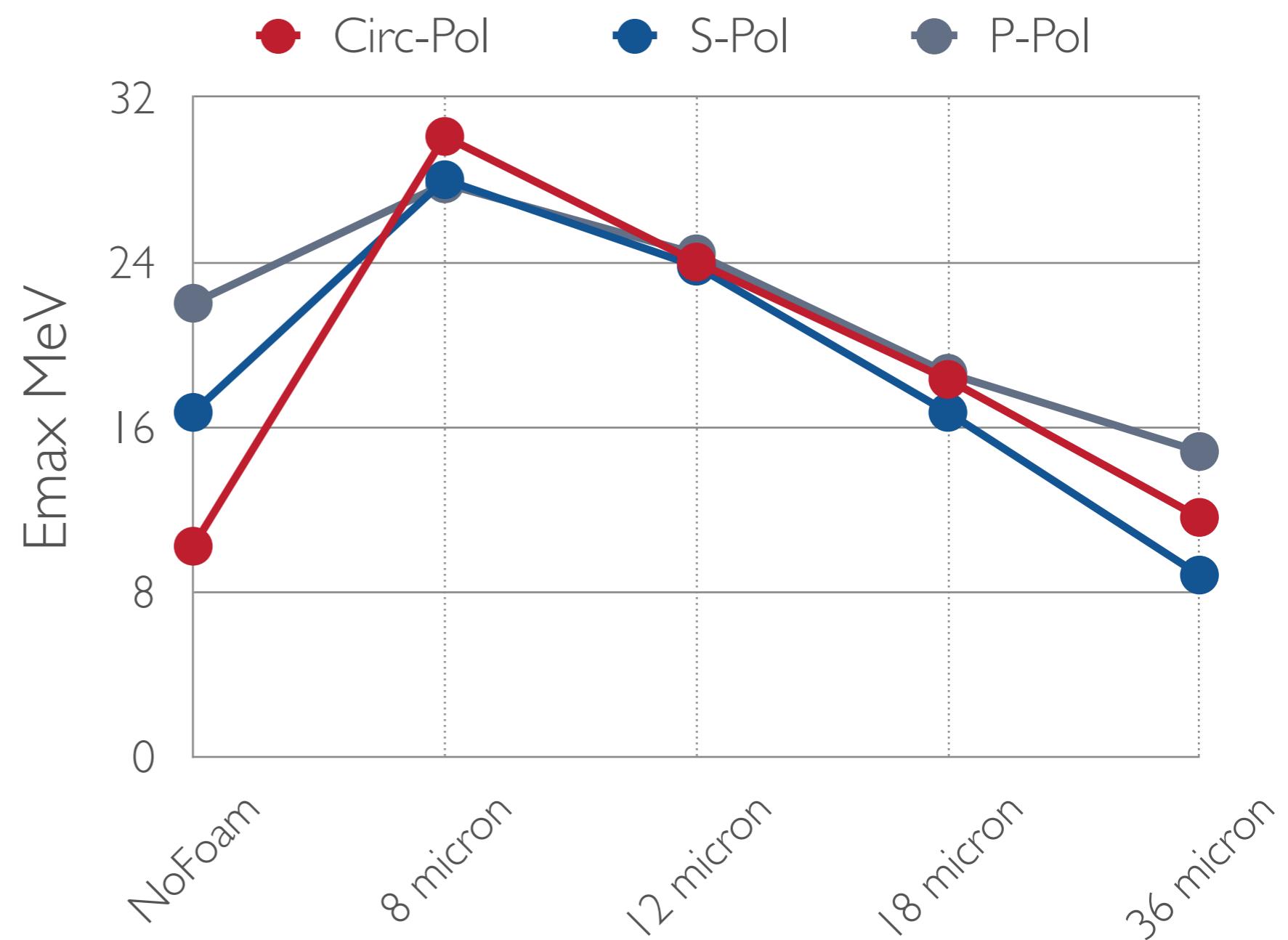
targets:

Al 0,75 μm

+

C foam 6,8 mg/cm³
» 8 μm
» 12 μm
» 18 μm
» 36 μm

(Nov 2014)

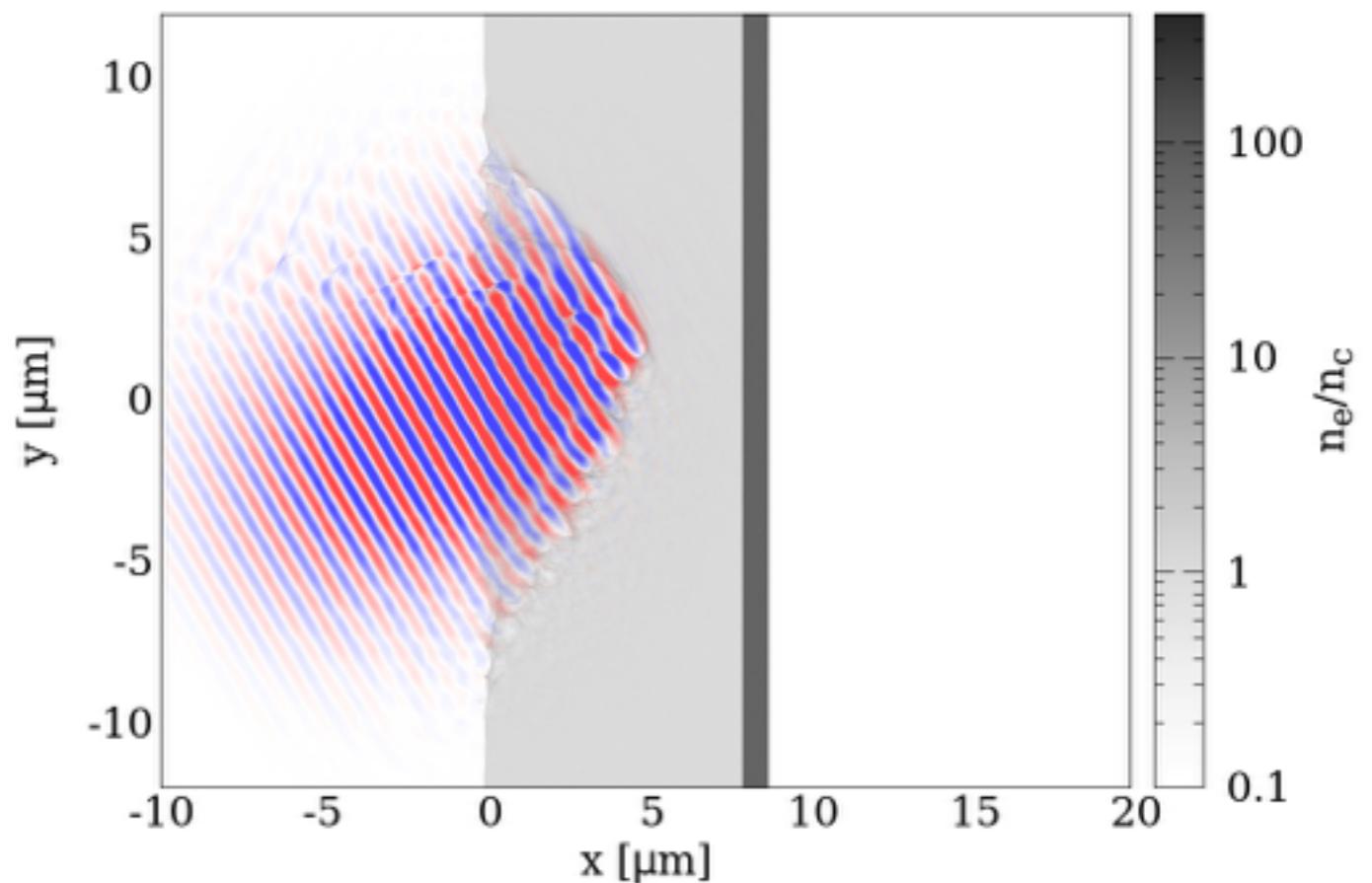


2D simulations

Energy scaling is not comparable with experiments

Huge difference between polarisations

Not predictive!



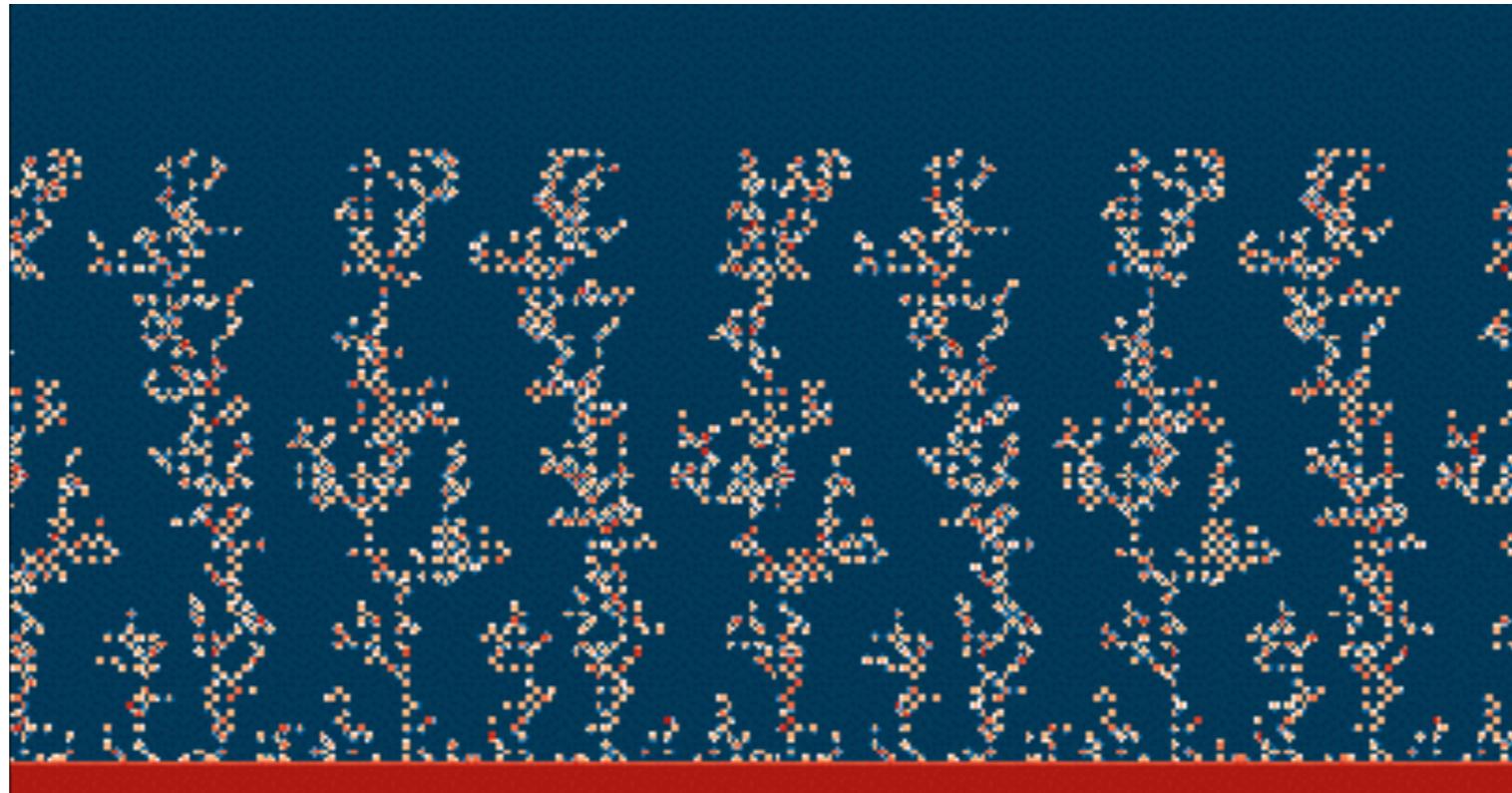
Foam layer as aggregate of cluster work in progress

DLA

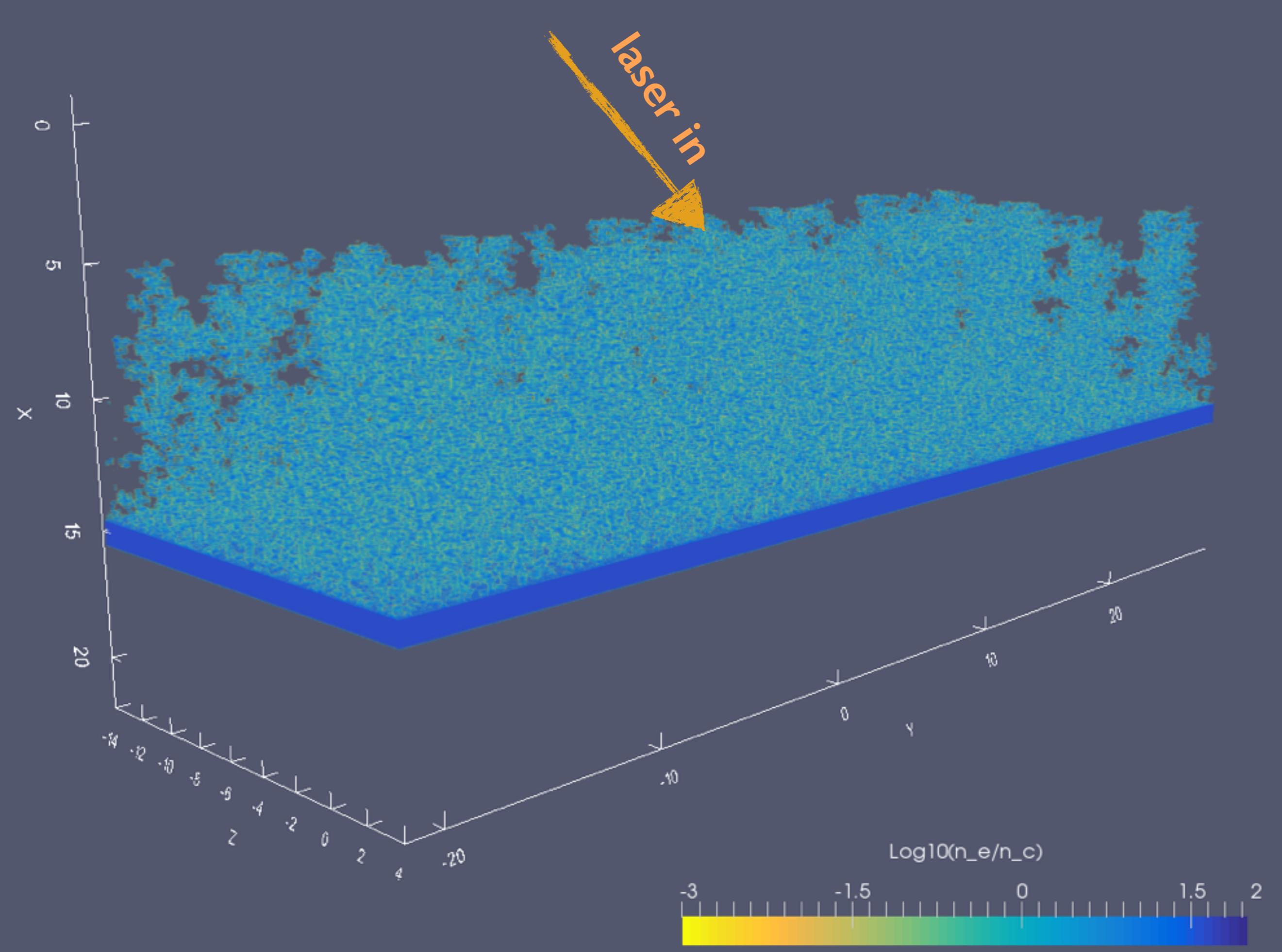
Diffusion Limited Aggregation ♦
algorithm for “growing” fractal structures *Brownian trees*

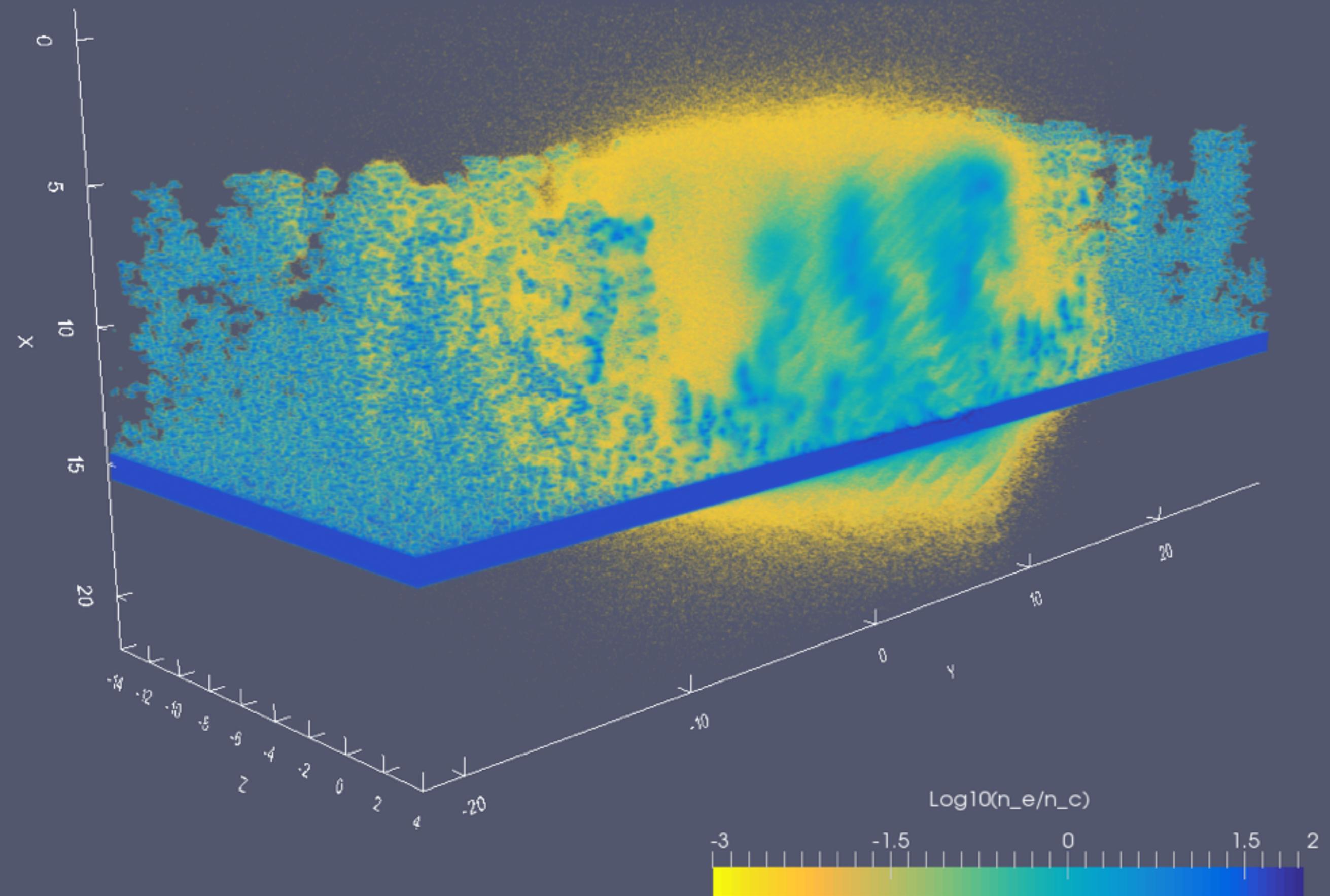
foam

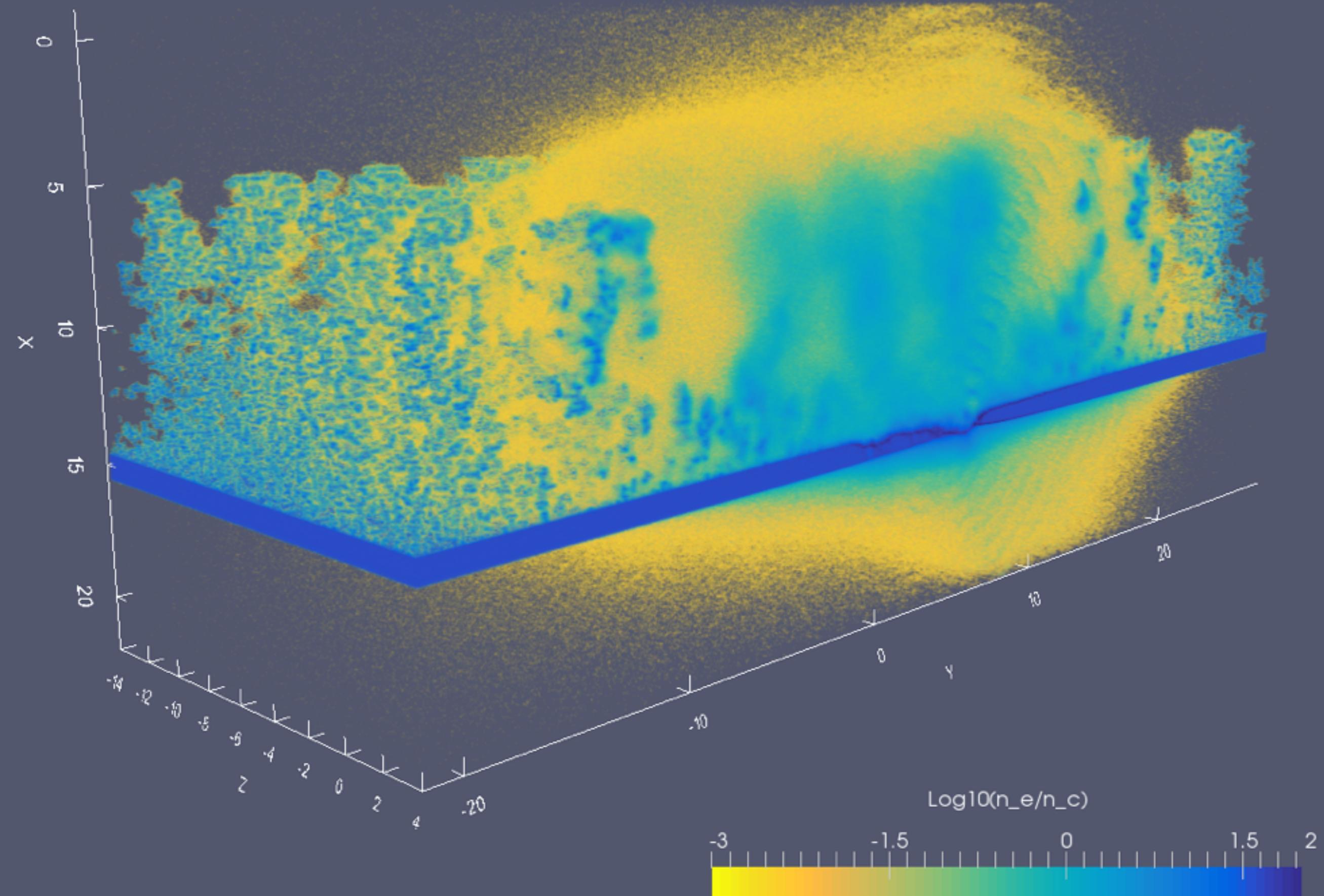
dense clusters $n \sim 50n_c$
average density $n \sim 1n_c$



♦T.A. Witten Jr, L. M. Sander, Phys. Rev. Lett. 47, 1400 (1981)

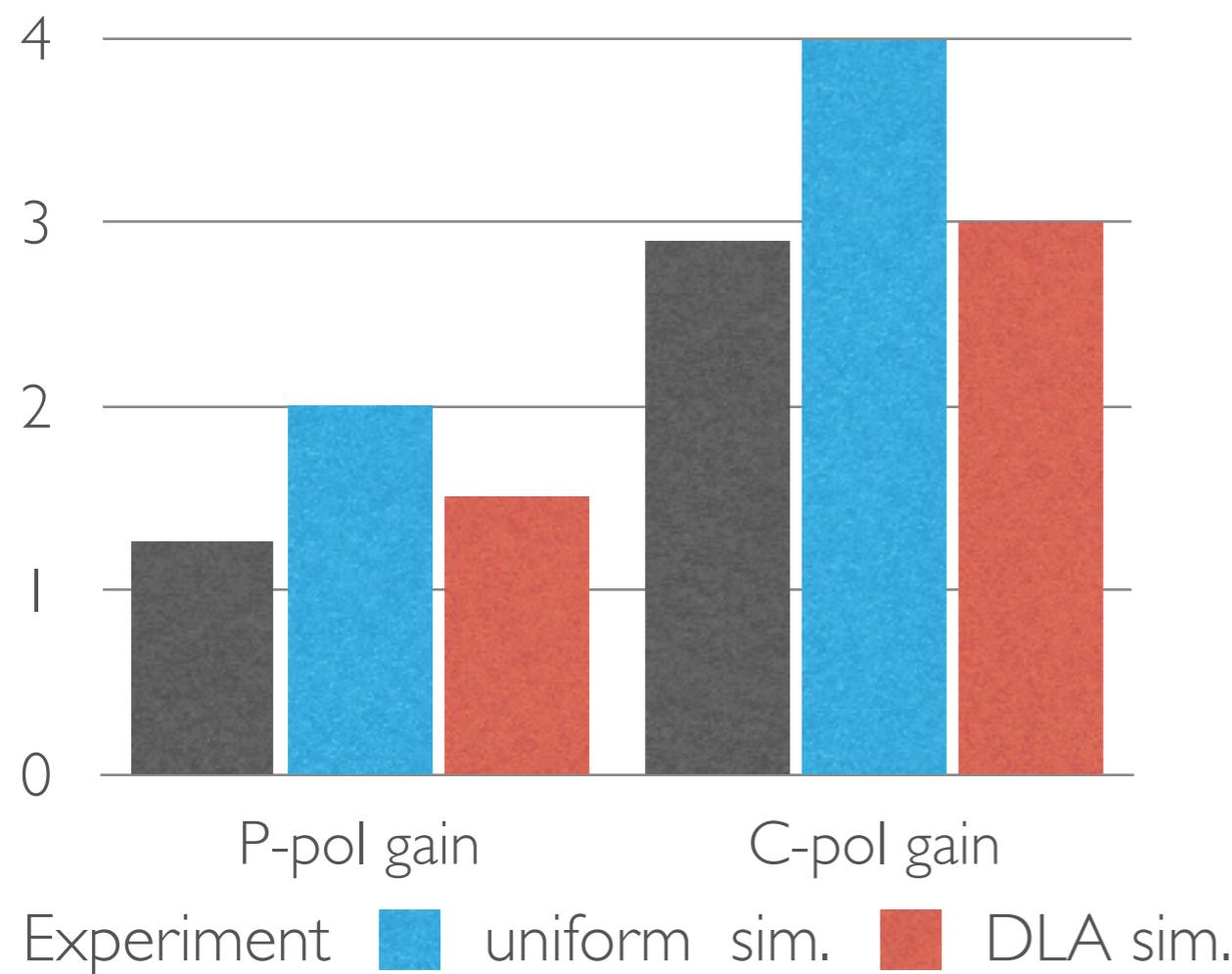






E_{\max} gain with vs. without foam

$E_{\max}^{\text{foam}}/E_{\max}^{\text{flat}}$



20-30% underestimation of the max. ion energy but...

DLA Foam seems to catch the features of the interaction better than uniform foam

... more work to come

Conclusions

PIC

Light-Sail acceleration under same interaction conditions
Circular Pol better than Linear Pol

PIC

Electron acceleration by surface waves:
multi MeV non-thermal e-bunches



3D PIC simulations can be quantitatively predictive
also for laser-solid interaction... if used wisely



2D simulations rarely give quantitative estimations
and should be used with the support of few 3D runs

Funding acknowledgements



LASERLAB-EUROPE



PRACE HighPerformanceComputing consortium
for access to FERMI BlueGene/Q™ at CINECA



Ministero dell'Istruzione, dell'Università e della Ricerca

MIUR (Italy): FIRB and PRIN projects

thank you

