



# Synchrotron radiation applications with UV radiation

**Emanuele Pace**

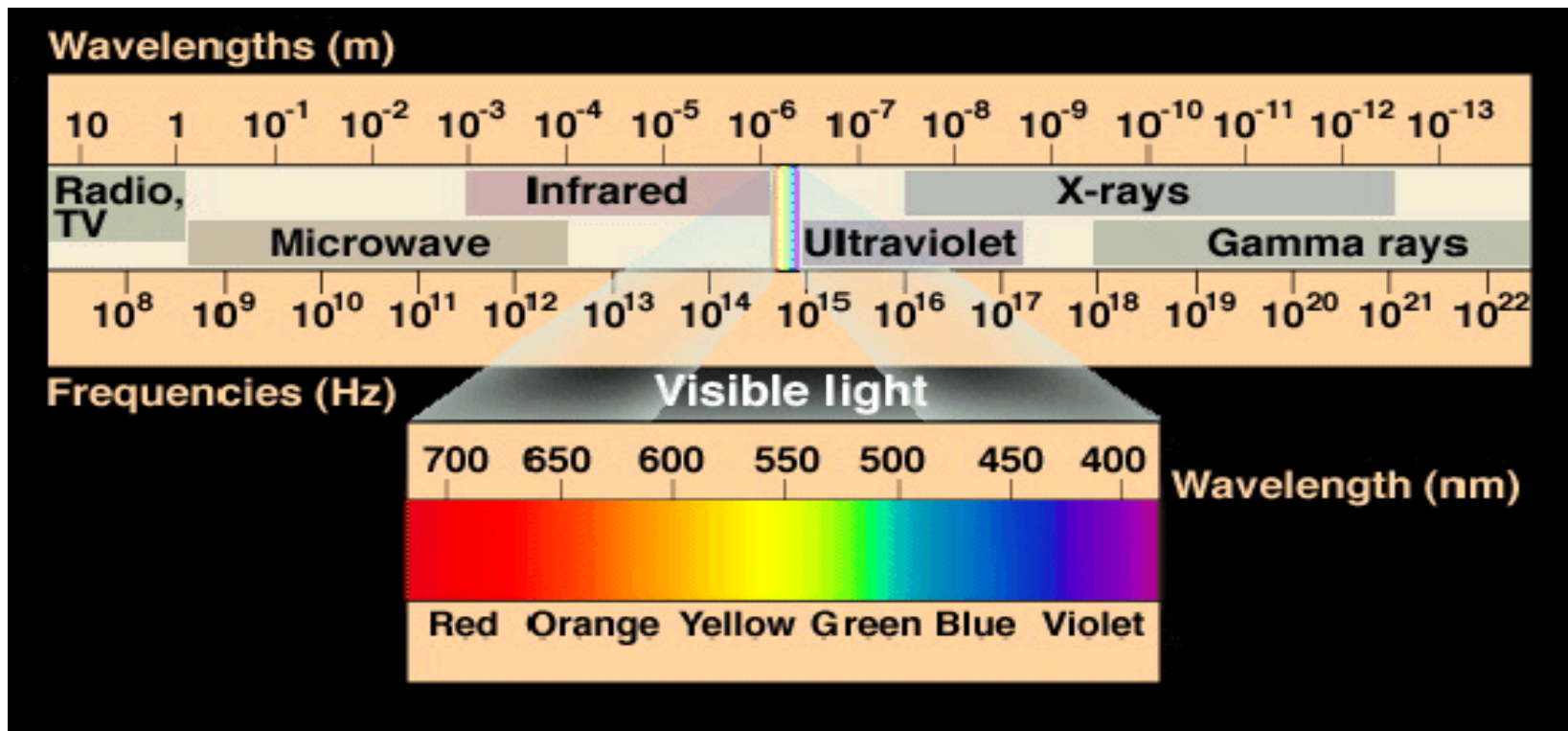
Dipartimento di Fisica e Astronomia

Università degli Studi di Firenze

LNF-INFN

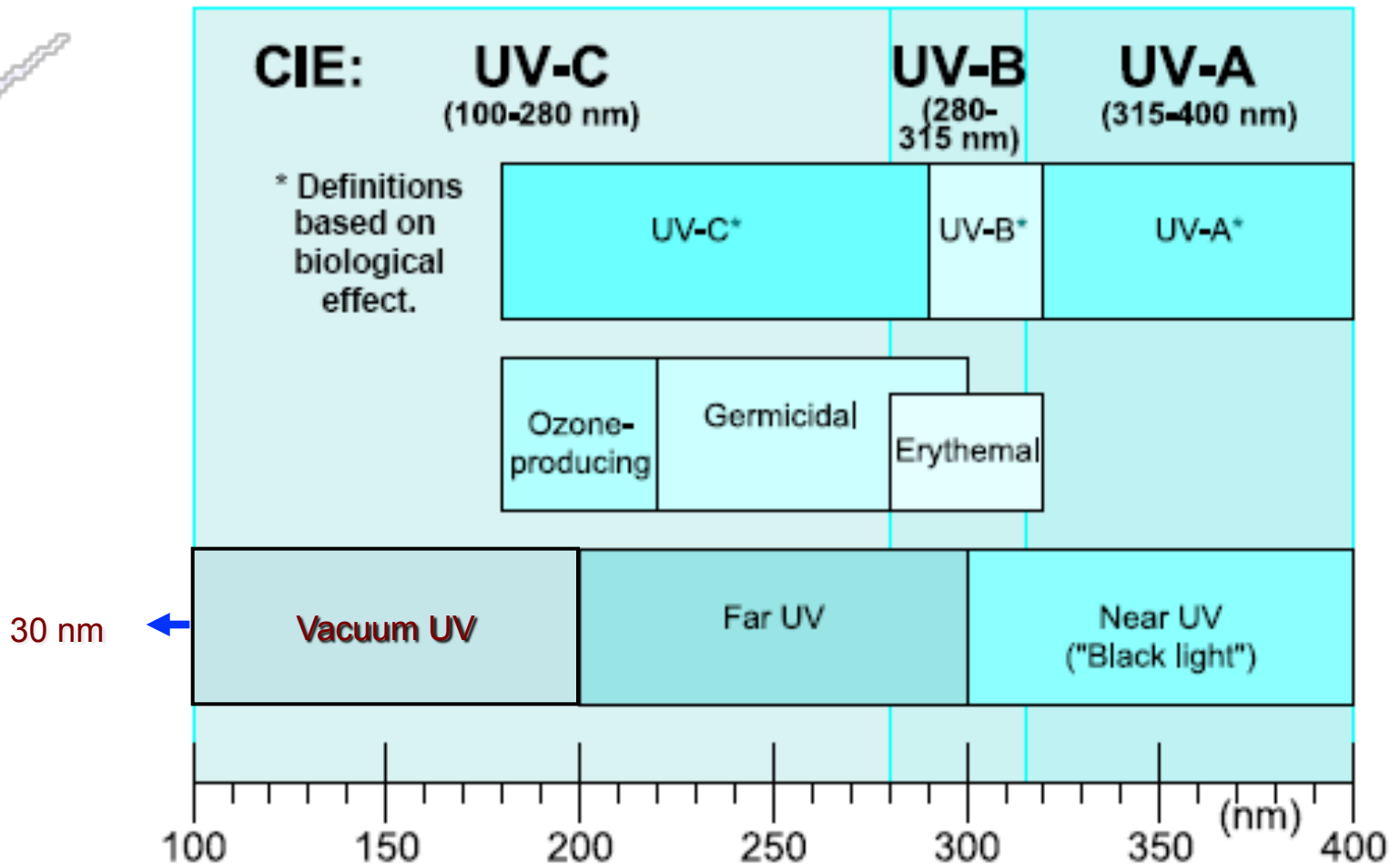
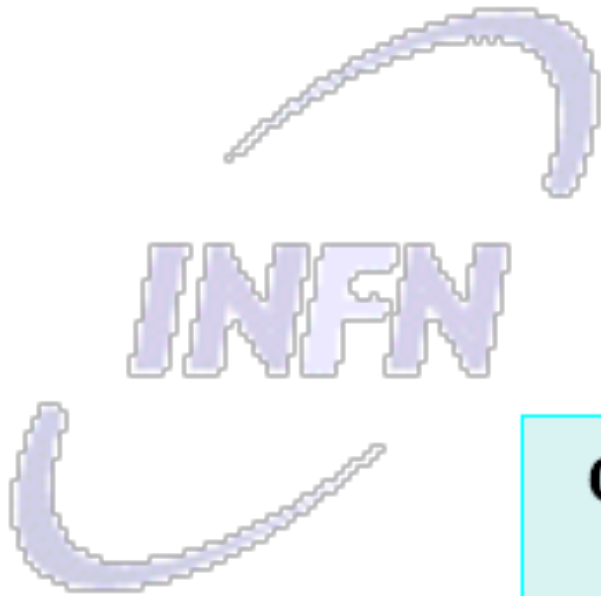


# Radiazione e.m. Spettro



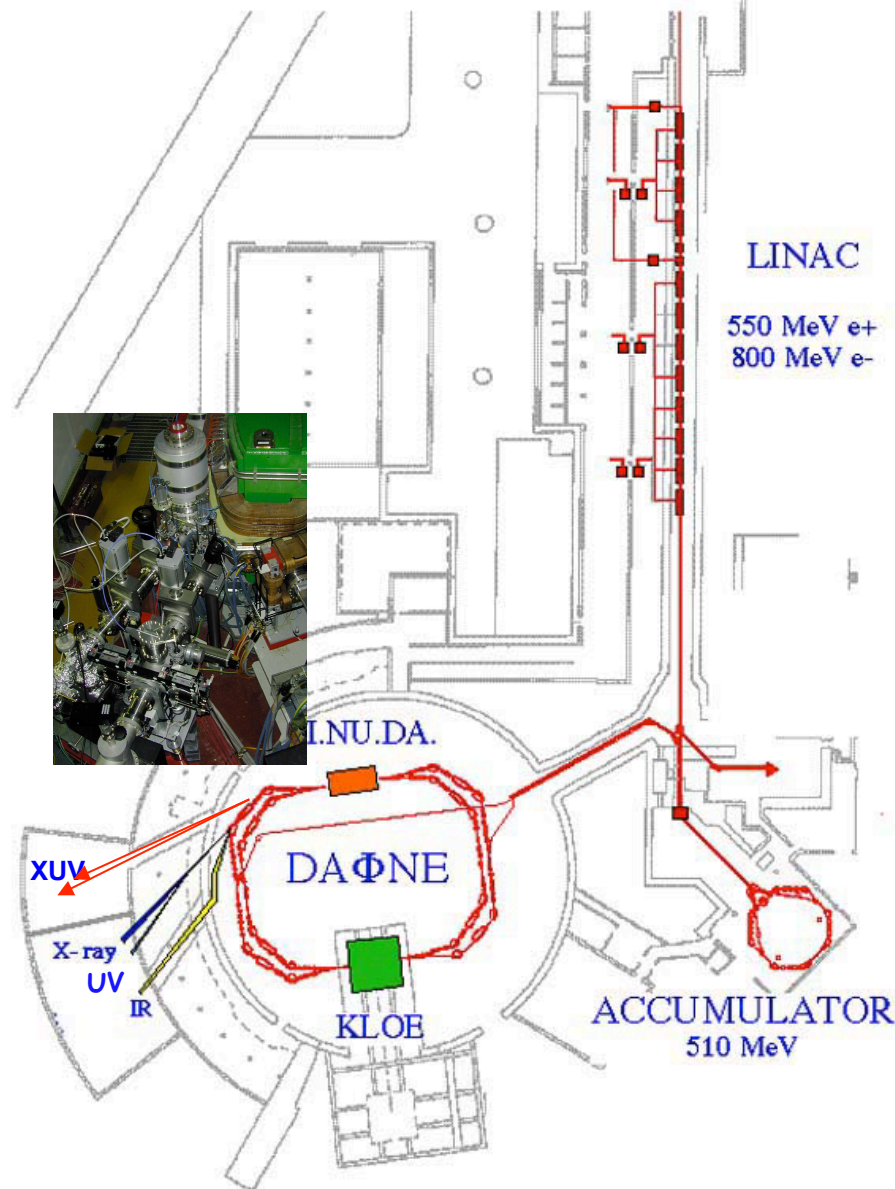


# Radiazione e.m. Spettro UV





# The Frascati $\Phi$ -Factory



LINAC

$550 \text{ MeV } e^+$   
 $800 \text{ MeV } e^-$

I.N.U.D.A.

DAΦNE

XUV

X-ray

UV

IR

KLOE

ACCUMULATOR

$510 \text{ MeV}$



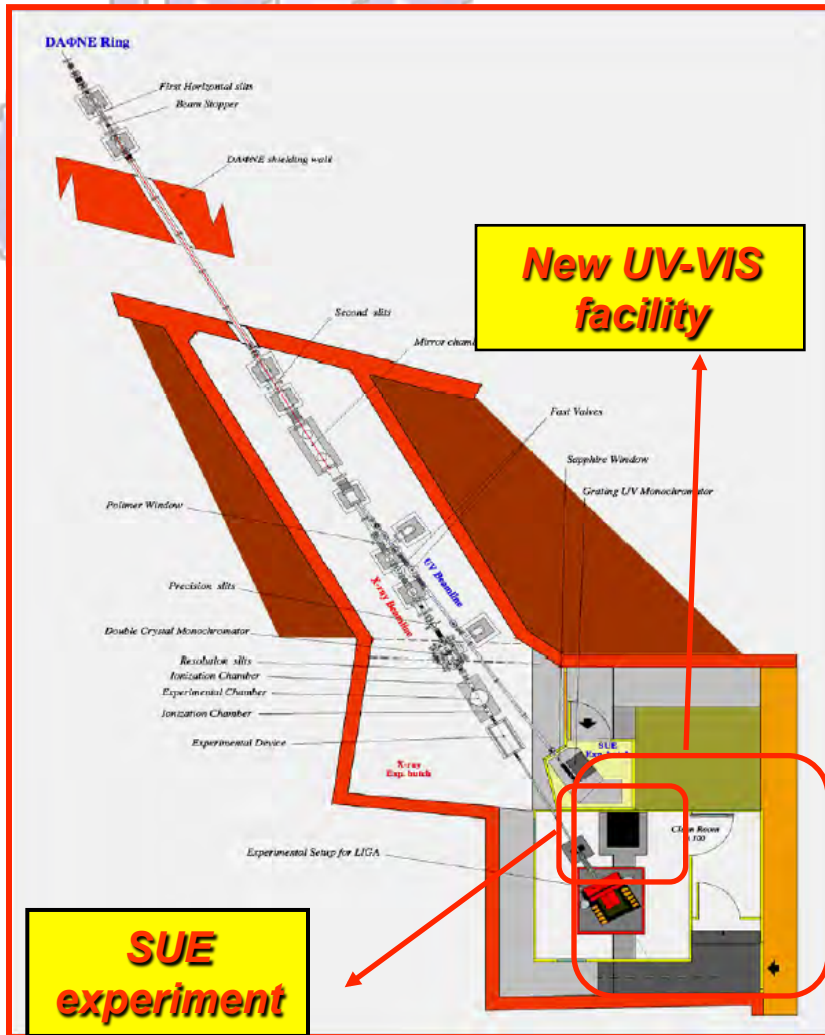
# DXR-2/SOURCE: UV-VIS branchline @ LNF

Test Facility to exploit the synchrotron radiation at DAFNE-L for electro-optical testing, calibration and characterization of optical system and detectors in the VIS-UV domain (2 eV to 10 eV).

➤ Important features are

- ☺ Photon flux known and reproducible
- ☺ Photon flux independent from the energy
- ☺ Ultra-fast and high frequency pulses (0.3 ns, 1GHz)
- ☺ Continuous spectrum radiation covering a wide domain
- ☺ Large spot size (~ mm)

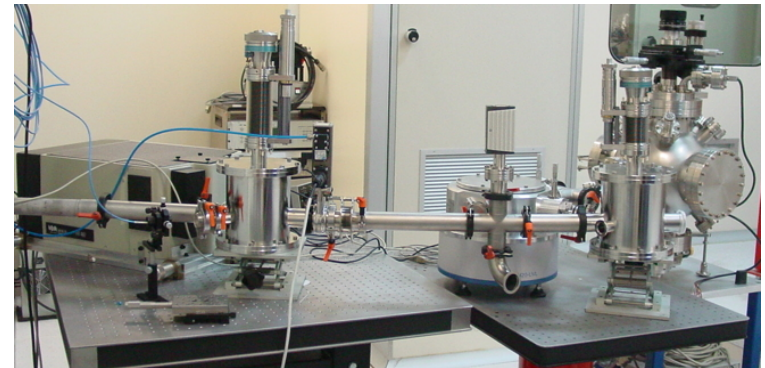
# DAΦNE-L UV branch line



Wiggler UV branch line-deflection by a grazing incidence gold coated mirror (about  $2^\circ$ )

UV-VIS beamline new setup  $2-10\text{ eV}$  (120-650 nm)

*Branch line in a 1000-class cleanroom*



- Space applications
- Astrobiology and photo-biology
- Optical technology
- Detector technology
- Instrumentation testing and calibration
- Optical properties of materials



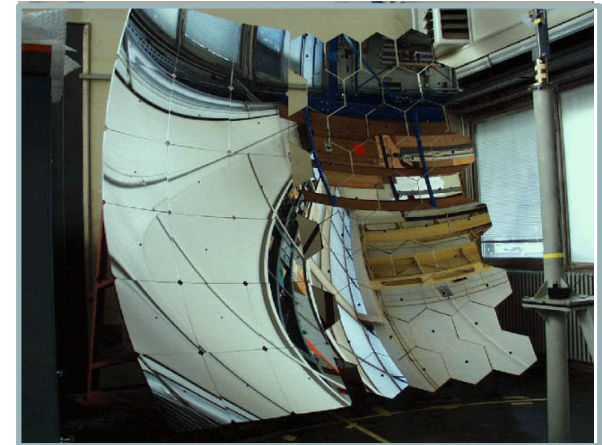
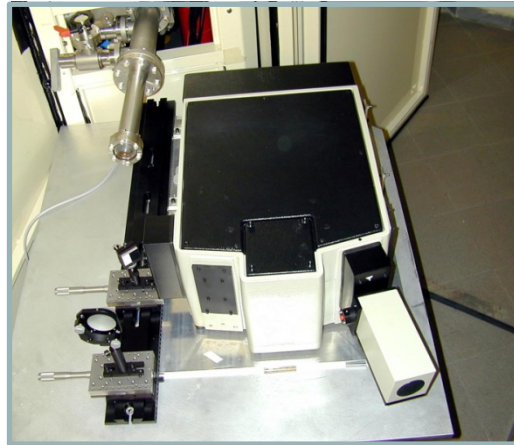
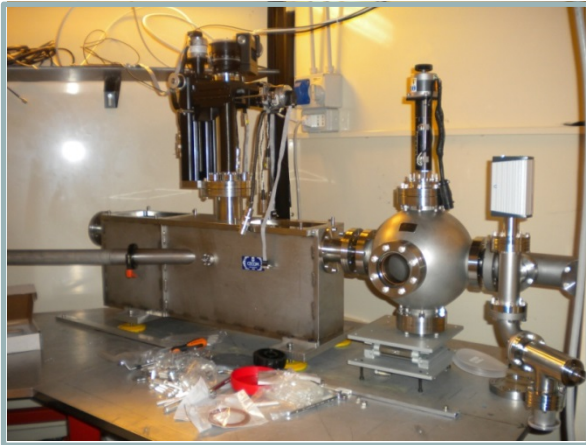
INFN

# DAΦNE-L UV branch line

UV-VIS monochromatic radiation source (180-650 nm)

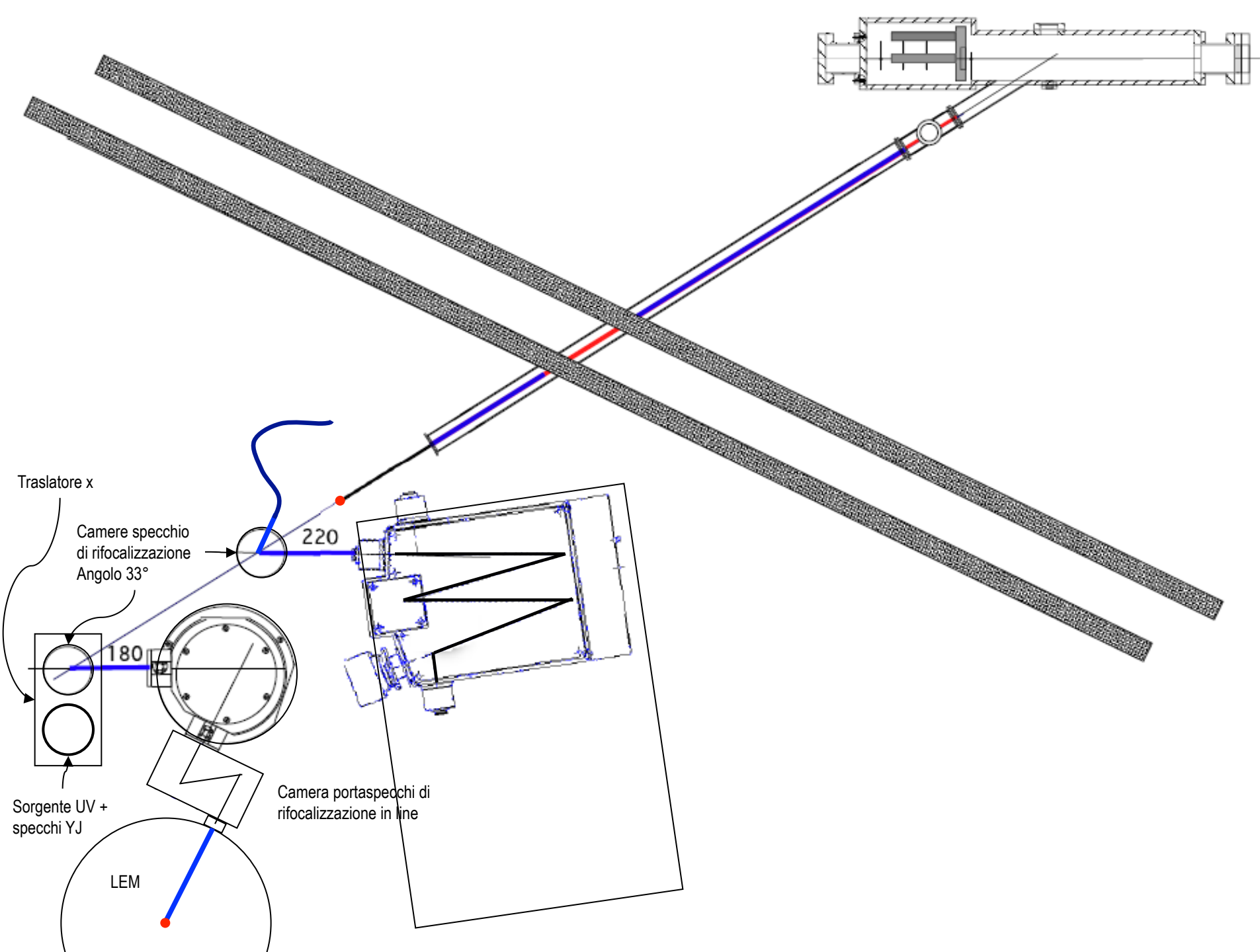
VUV monochromatic radiation source (120-250 nm)

UV-VIS radiation source (200-650 nm)



- Large optical systems (up to 4 mt) surface characterization
- UV photoageing of optical components and materials
- Detector calibration
- Photobiology and exobiology experiments





# Spectral emission



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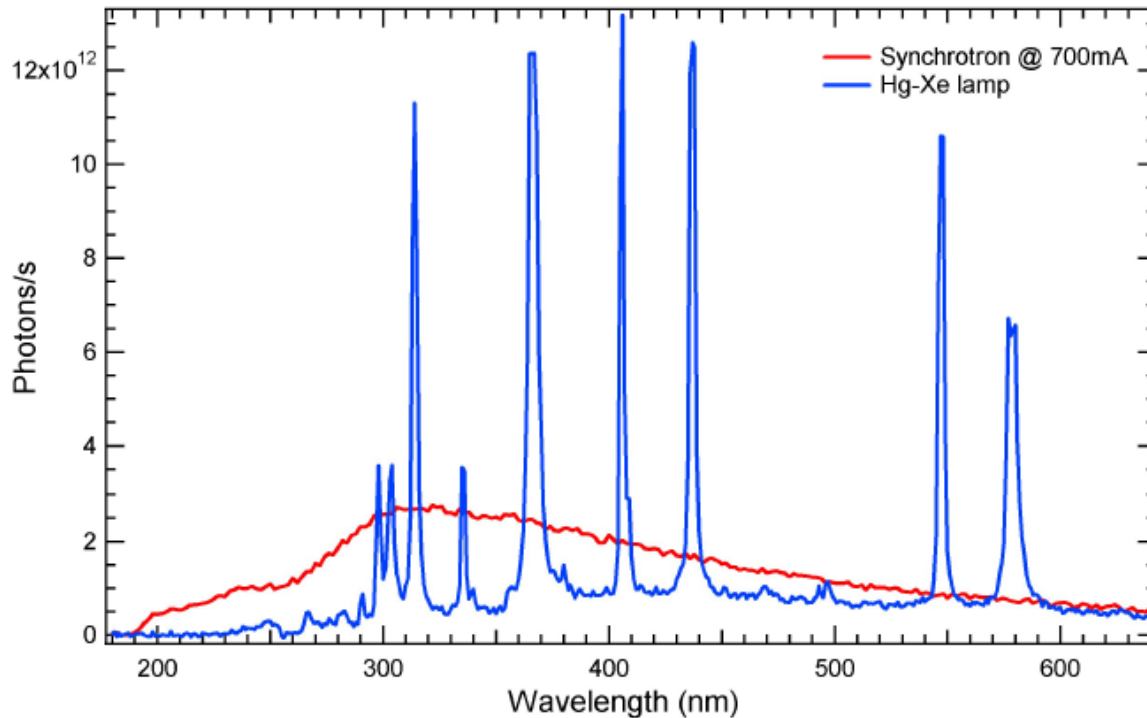
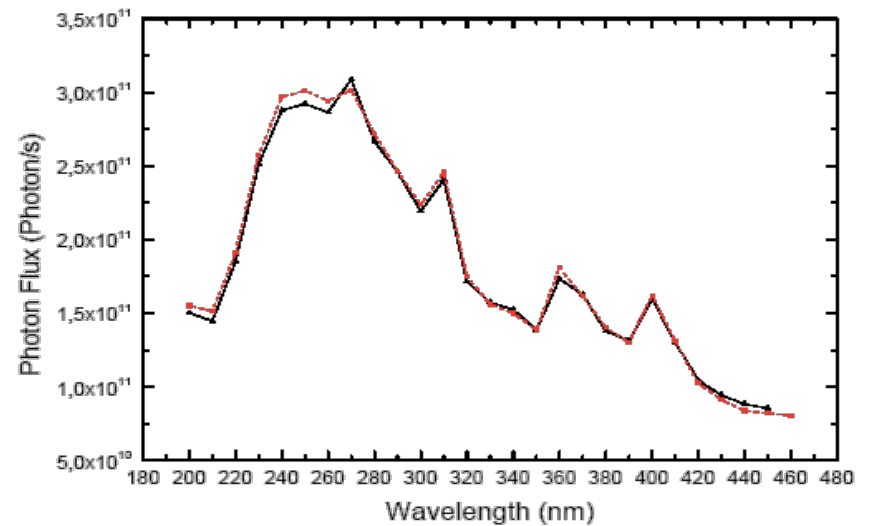
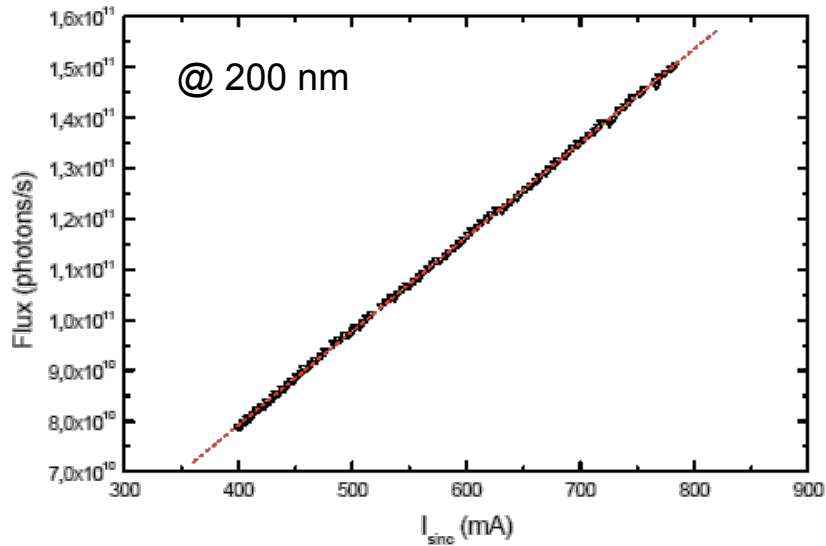
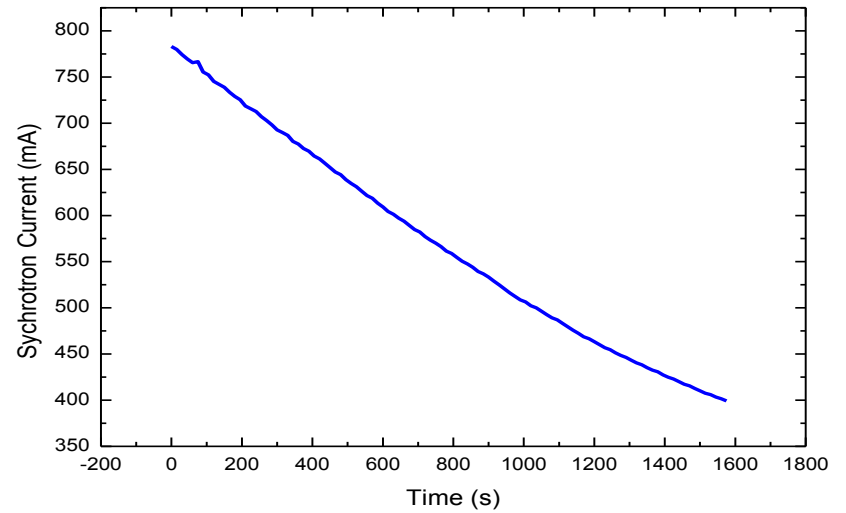


Figure 2. Emission spectrum of the Hamamatsu 500 W Hg-Xe lamp compared to the SR spectrum emitted in the UV-VIS range by an electron current of 0.7 A circulating in the storage ring.

# The synchrotron source

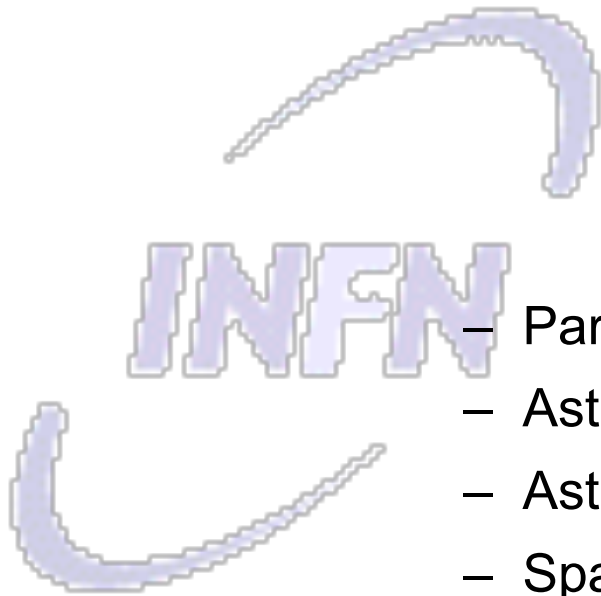


$I_{\text{norm}} = 750 \text{ mA}$





# Science



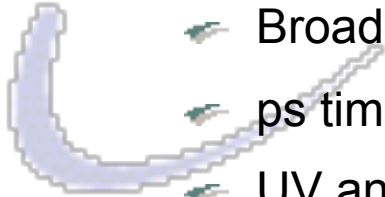
- Particle experiments (using Cherenkov light)
- Astro-particle experiments (using fluorescence light)
- Astronomy experiments
- Space experiments
  - Cosmic rays
  - Astronomy
  - Earth Observation
- Astrobiology and photo-biology
- Photochemistry
- Optical technology
- Detector technology
- Instrumentation testing and calibration
- Optical properties of materials



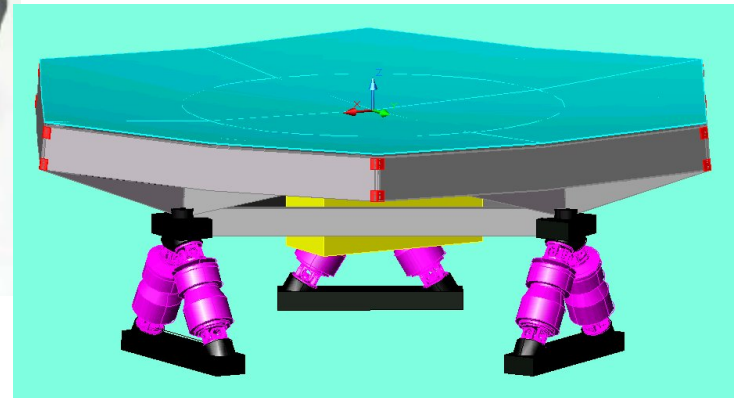
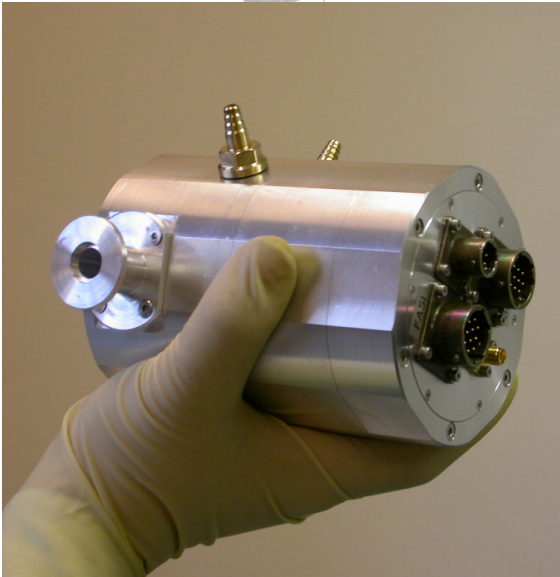
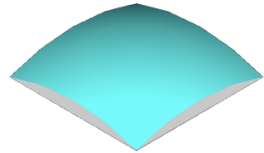
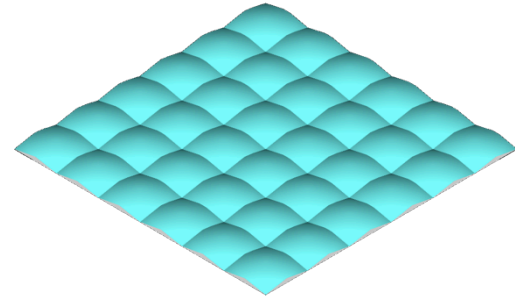
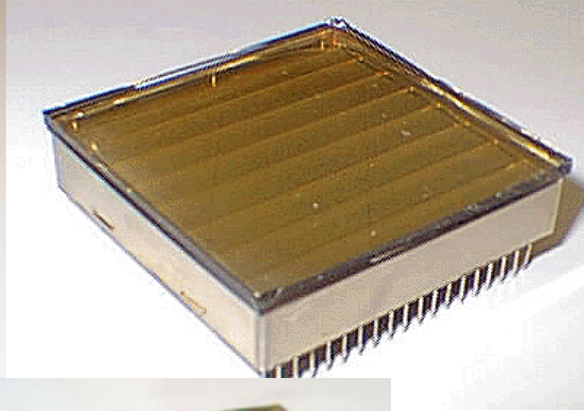
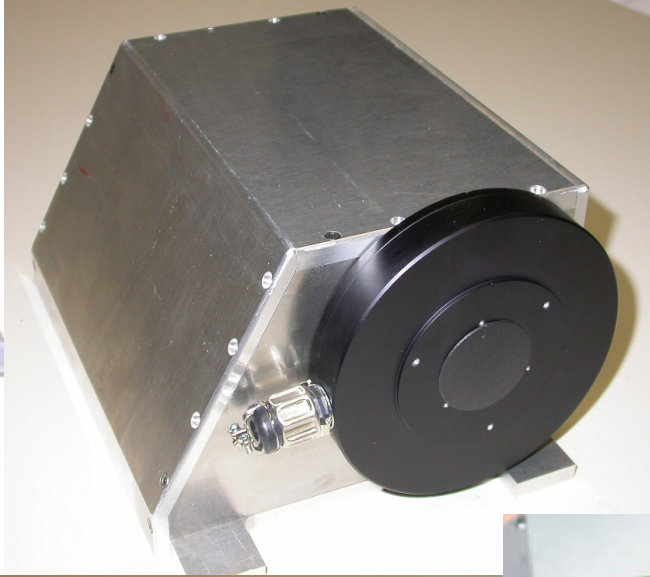
# Which science or technology?

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- Electro-optical characterization of optical materials and detectors
- Broad-band and precise spectro-photometry
- ps time-resolved measurement
- UV and VIS polarimetry and spectro-polarimetry
- UV and VIS detector calibration (CCD, MCP, WBM, etc.)
- Experiments on meteorites, micrometeorites and astrobiology
- Photobiology
- Development and tests of innovative detectors and materials (diamond, nanodiamond, carbon nanotubes, nitrides, etc.)
- Test and development of innovative optics and coatings
- UV tracks and background for UHECR
- Test of large optics (ALSO project by ESA)
- AIV and OGSE for space project (EUSO, SCORE/HERSCHEL, WSO/UV)



# Technologies

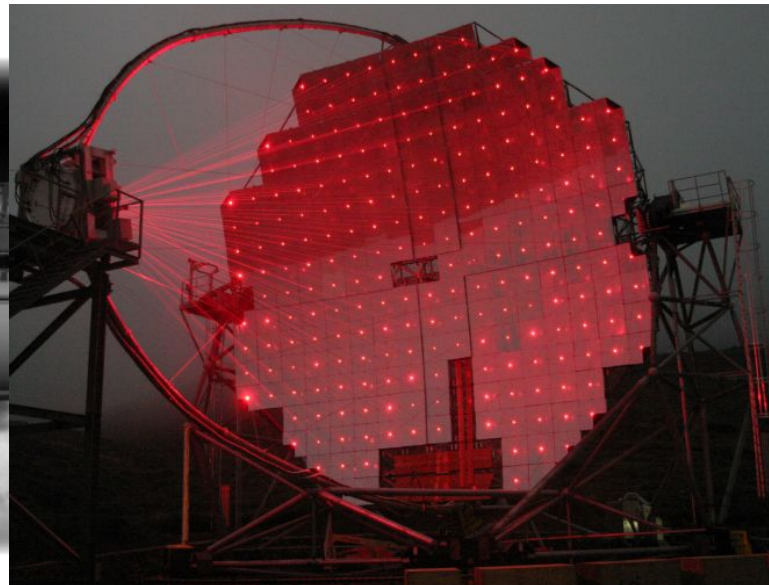
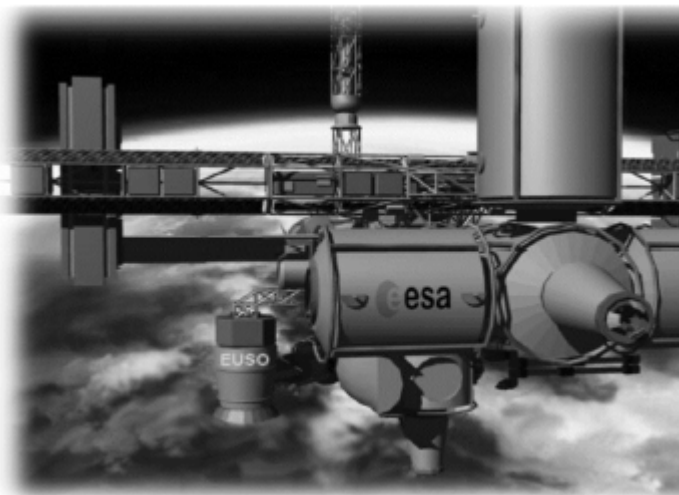
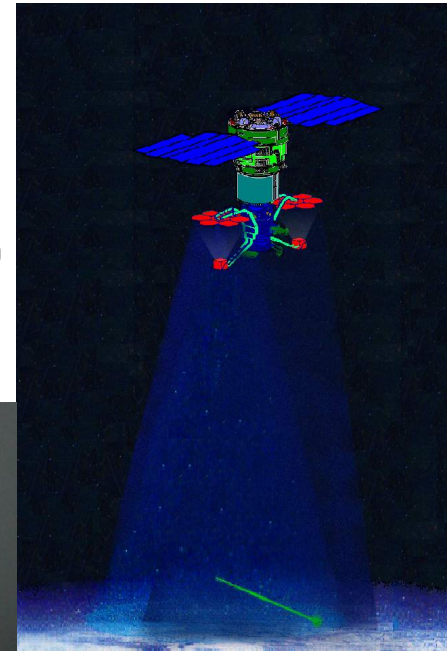
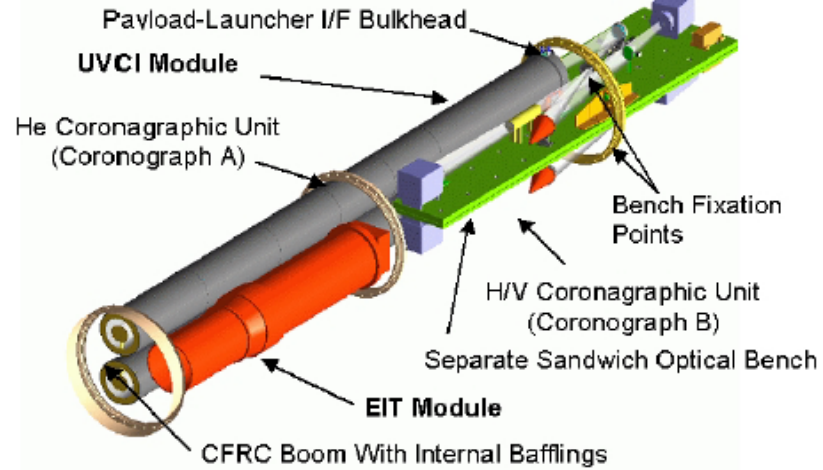
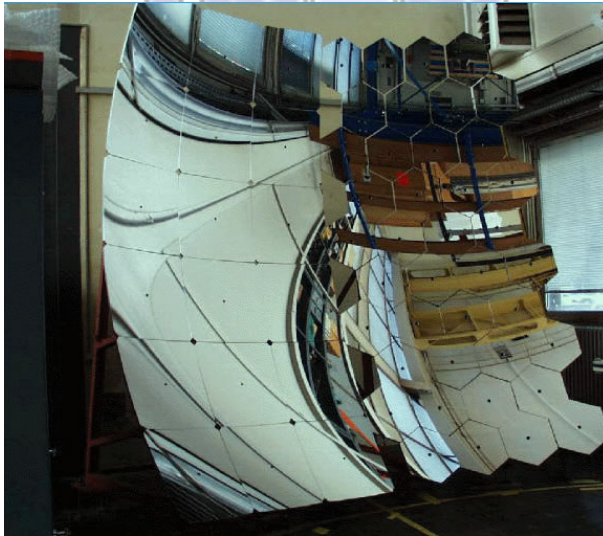


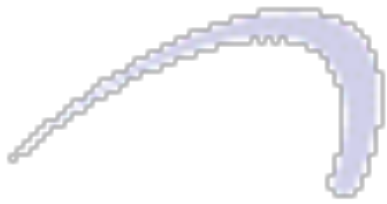


# Ground- and space-based experiments



INM

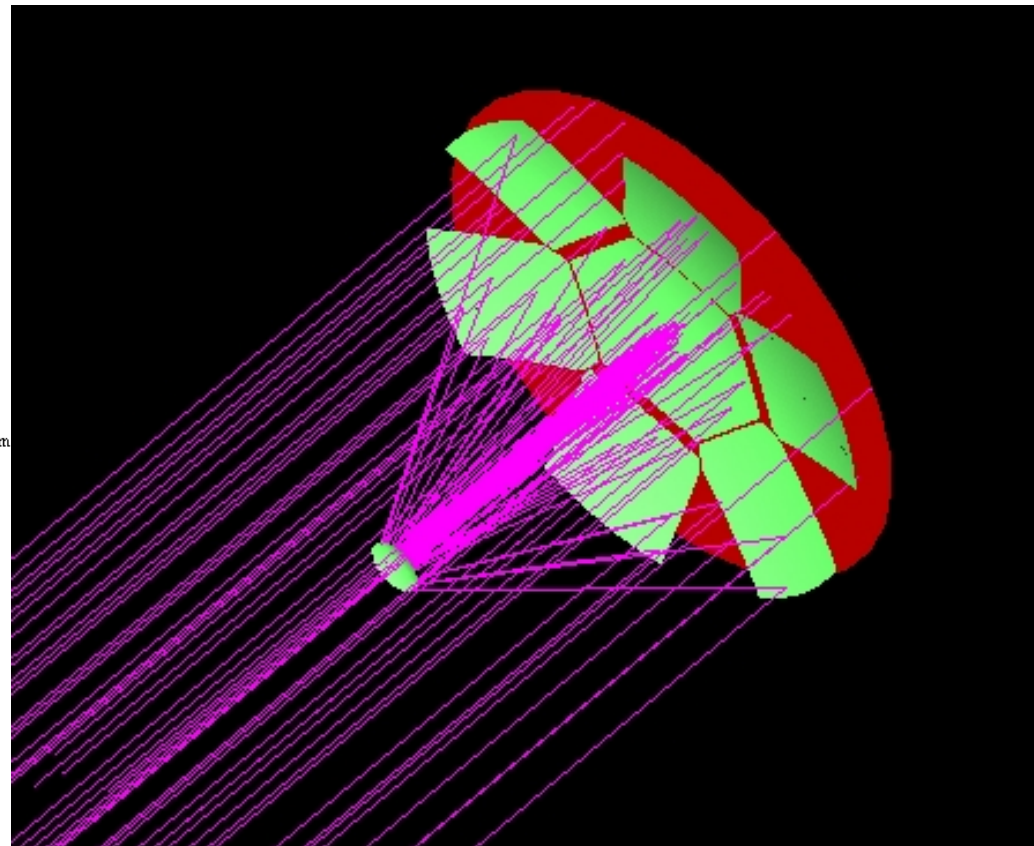
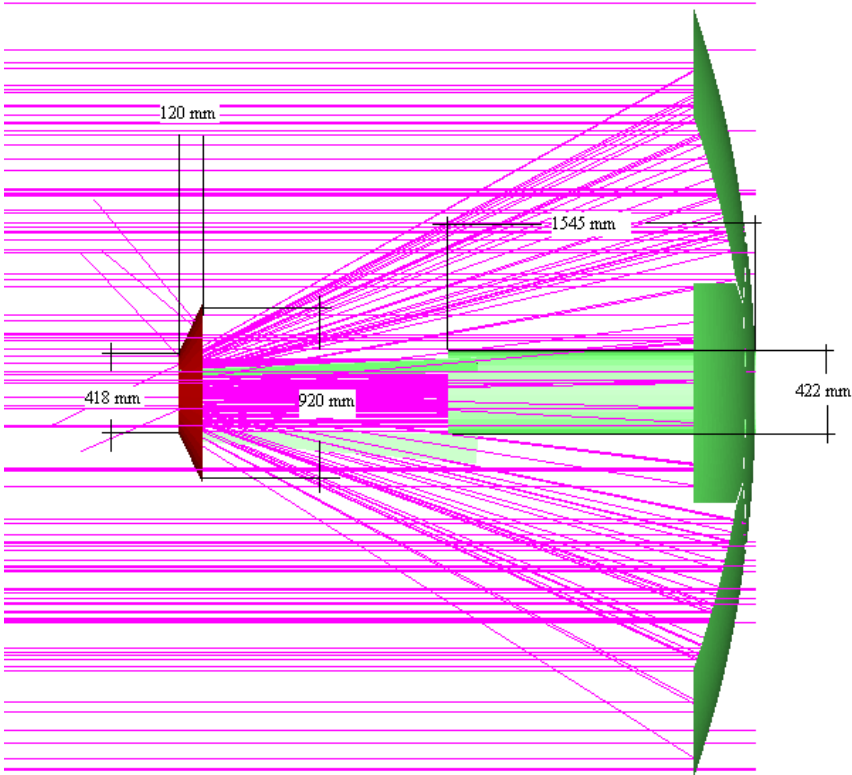




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# Deployable large mirrors

European Space Agency project



CNR/INO - Arcetri



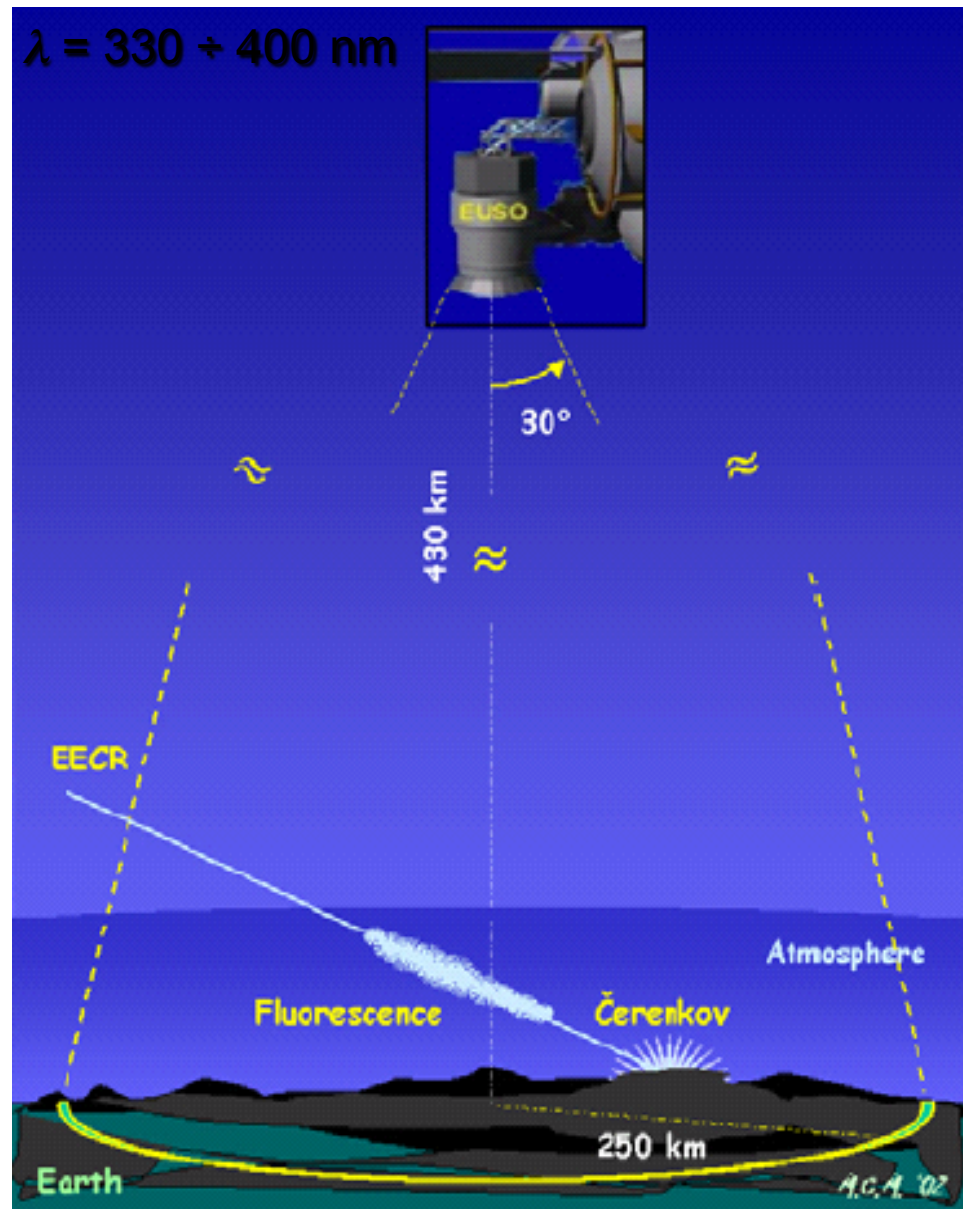
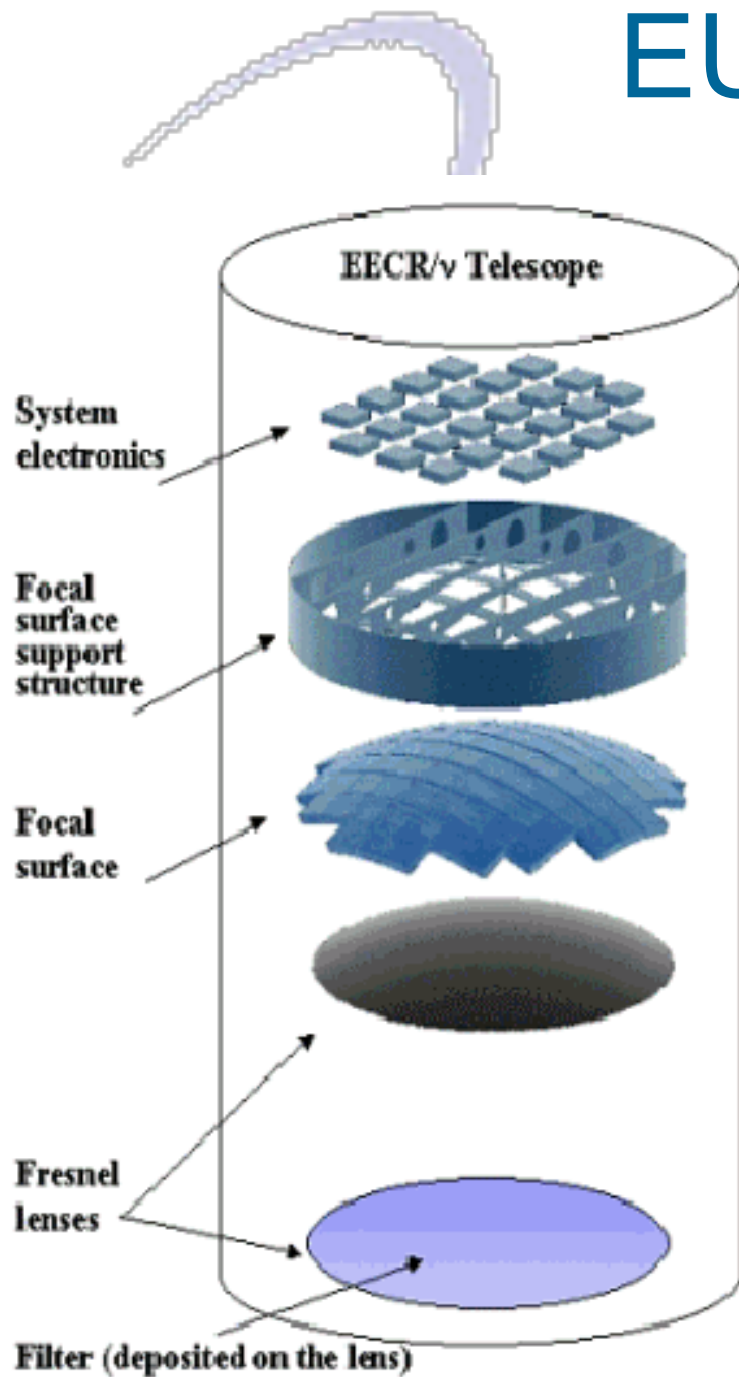
## Calibration Experiments for ultrafast UV transients

The UHECR detection through the UV fluorescence emission is mainly a technological.

- \* SOURCE is a support equipment facing the development of highly-sensitive optical and detection innovative technologies (mirror size up to 4 m).
- \* SOURCE is a facility that allow for the first time the real simulation of fluorescence UV tracks against a background in order to test the performance of UHECR ground- and space-based instrumentation in fully operative conditions.

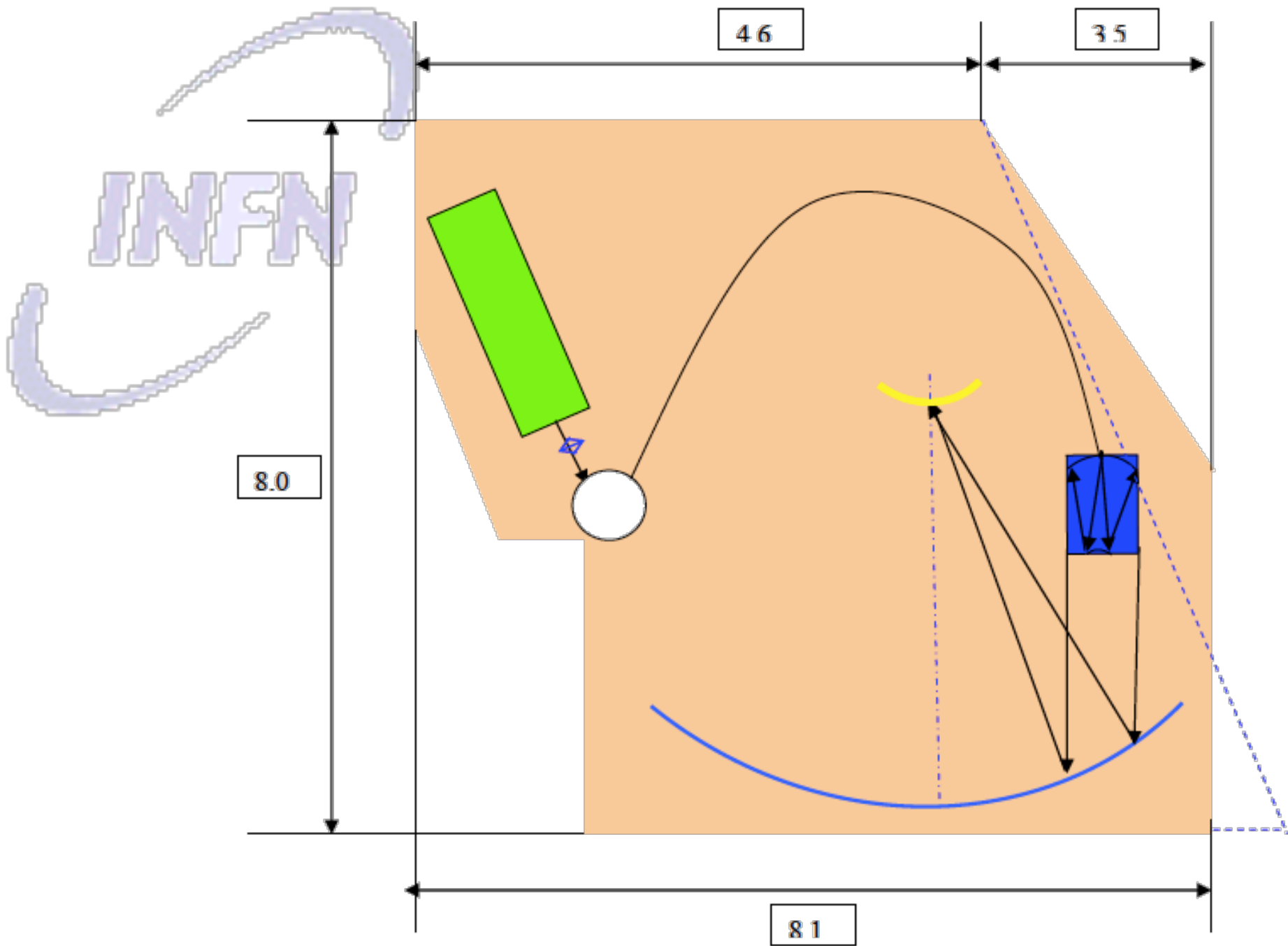


# EUSO





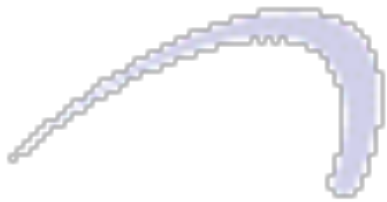




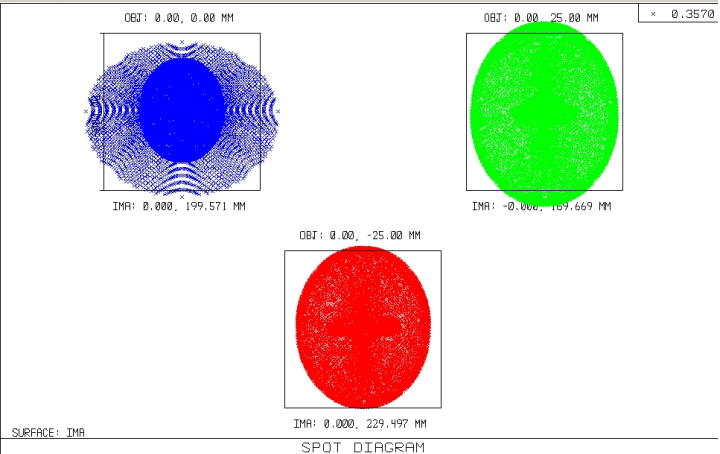


# Specifiche per il layout ottico

- Produrre tracce con sviluppo temporale minimo di 300  $\mu$ s
- Larghezza di banda 300-400 nm
- Produrre un background di luce diffusa
- Larghezza di banda 200-650 nm
- Flessibilità per misurare specchi con differenti caratteristiche ottiche
- Possibilità di misurare ottiche in trasmissione



# Schmidt da 4 metri



LENS HAS NO TITLE.  
 MON SEP 25 2006 UNITS ARE MICRONS.  
 FIELD : 1 2 3  
 RMS RADIUS : 1592.86 2272.76 1943.68  
 GEO RADIUS : 3686.72 3572.32 3130.48  
 BOX WIDTH : 6000 REFERENCE : CENTROID

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 CONFIGURATION 1 OF 1

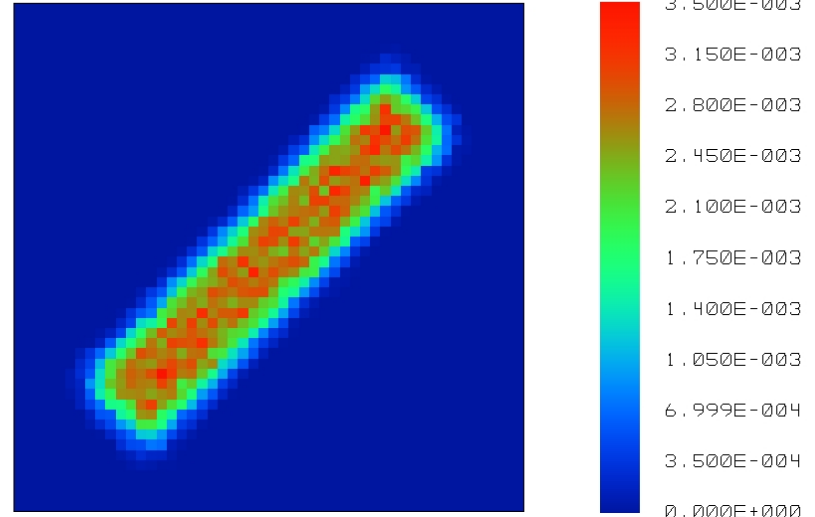


IMAGE DIAGRAM

LENS HAS NO TITLE.  
 MON SEP 25 2006  
 IMAGE WIDTH = 50.0000 MILLIMETERS, 50 X 50 PIXELS  
 FIELD POSITION: 0.0000, 0.0000 MM  
 PERCENT EFFICIENCY: 100.000%, 1.000E+000 WATTS  
 SURFACE: 4. UNITS ARE WATTS PER MILLIMETERS SQUARED.

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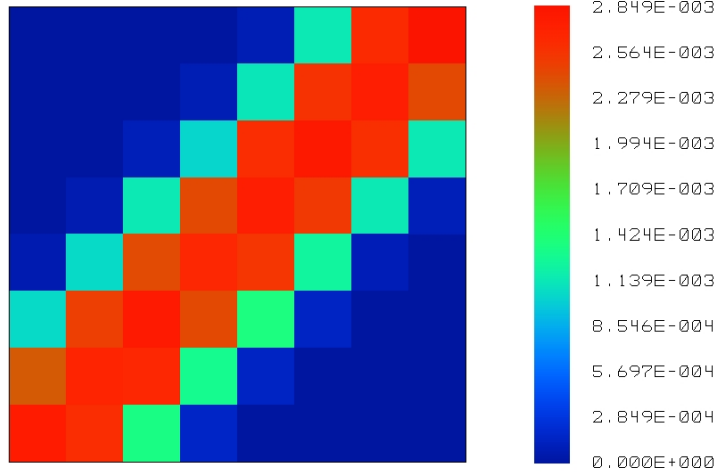
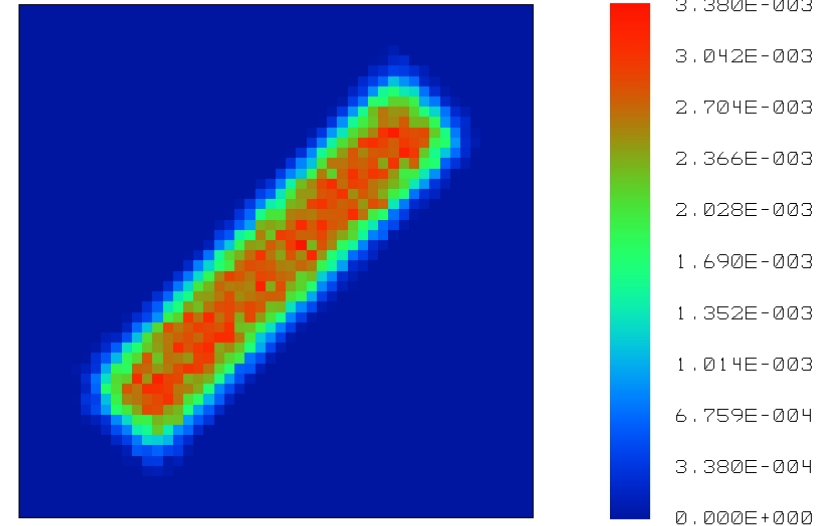
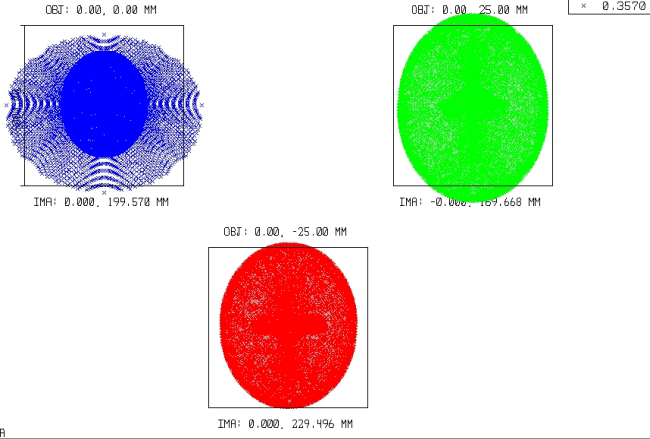


IMAGE DIAGRAM

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 FIELD POSITION: 0.0000, 0.0000 MM  
 PERCENT EFFICIENCY: 74.935%, 7.493E-001 WATTS  
 SURFACE: 4. UNITS ARE WATTS PER MILLIMETERS SQUARED.

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 CONFIGURATION 1 OF 1

# Schmidt ALC da 1 metro



SURFACE: IMR

SPOT DIAGRAM

LENS HAS NO TITLE.  
 MON SEP 25 2006 UNITS ARE MICRONS.  
 FIELD : 1 2 3  
 RMS RADIUS : 1592.87 2272.74 1943.66  
 GEO RADIUS : 3686.75 3572.30 3130.45  
 BOX WIDTH : 6000 REFERENCE : CENTROID

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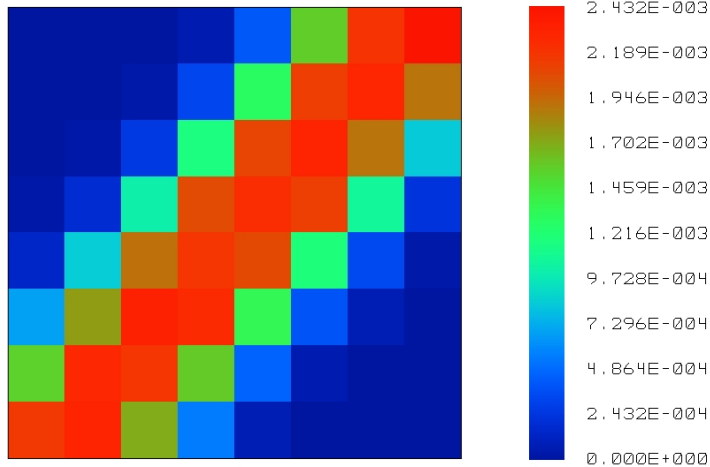


IMAGE DIAGRAM

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 FIELD POSITION: 0.0000, 0.0000 MM  
 PERCENT EFFICIENCY: 100.000%, 1.000E+000 WATTS  
 SURFACE: 4. UNITS ARE WATTS PER MILLIMETERS SQUARED.

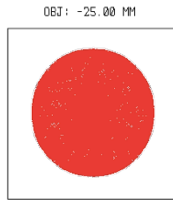
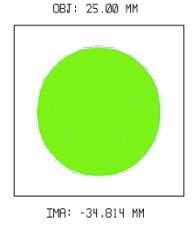
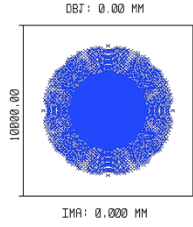
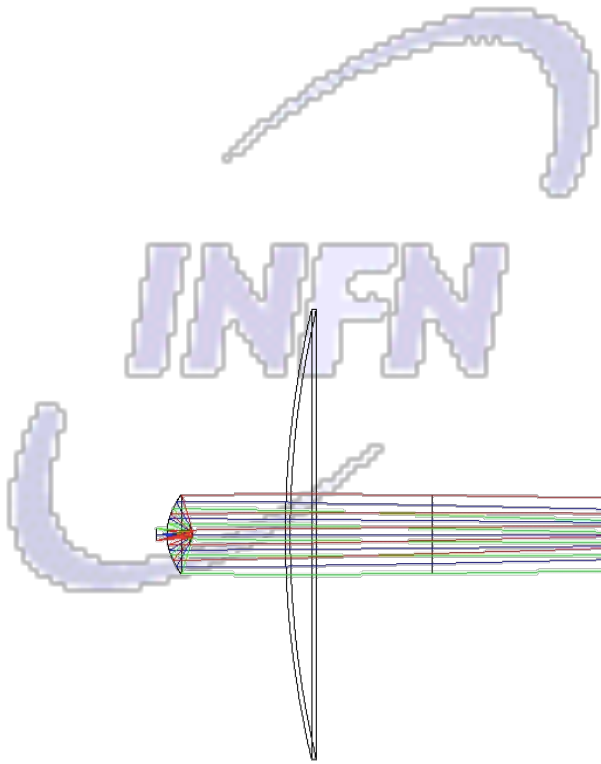
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IMAGE DIAGRAM

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 IMAGE WIDTH = 25.7000 MILLIMETERS, 8 X 8 PIXELS  
 FIELD POSITION: 0.00, 0.00 MM  
 PERCENT EFFICIENCY: 66.685%, 6.669E-001 WATTS  
 SURFACE: 4. UNITS ARE WATTS PER MILLIMETERS SQUARED.

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 CONFIGURATION 1 OF 1

# EUSO: Lenti di Fresnel



SURFACE: IMA

SPOT DIAGRAM

LENS HAS NO TITLE.  
MON SEP 25 2006 UNITS ARE MICRONS.

FIELD :	1	2	3
RMS RADIUS :	1830.54	2514.22	2514.22
GEO RADIUS :	3806.44	3718.29	3718.29
BOX WIDTH :	1E+004		

REFERENCE : CENTROID

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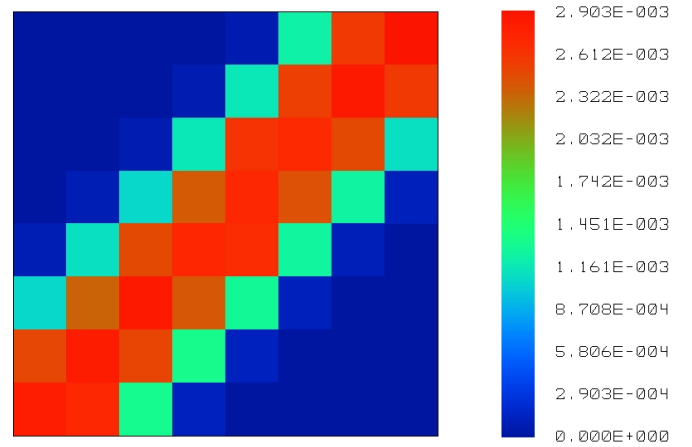
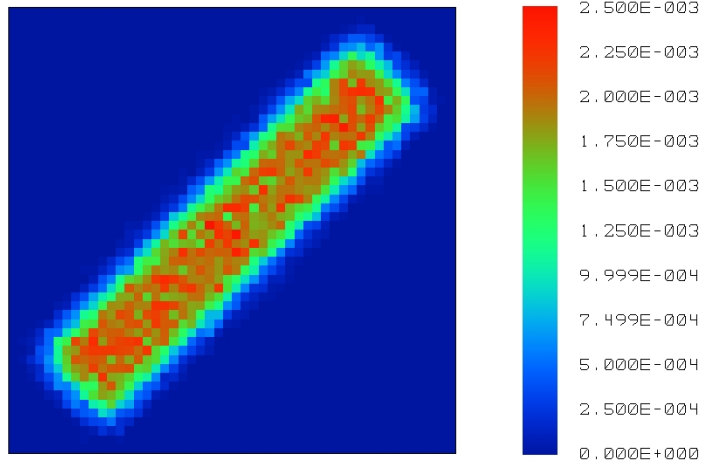


IMAGE DIAGRAM

LENS HAS NO TITLE.  
MON SEP 25 2006  
IMAGE WIDTH = 50.0000 MILLIMETERS, 50 X 50 PIXELS  
FIELD POSITION: 0.00 MM  
PERCENT EFFICIENCY: 100.000%, 1.000E+000 WATTS  
SURFACE: 8. UNITS ARE WATTS PER MILLIMETERS SQUARED.

IMAGE DIAGRAM

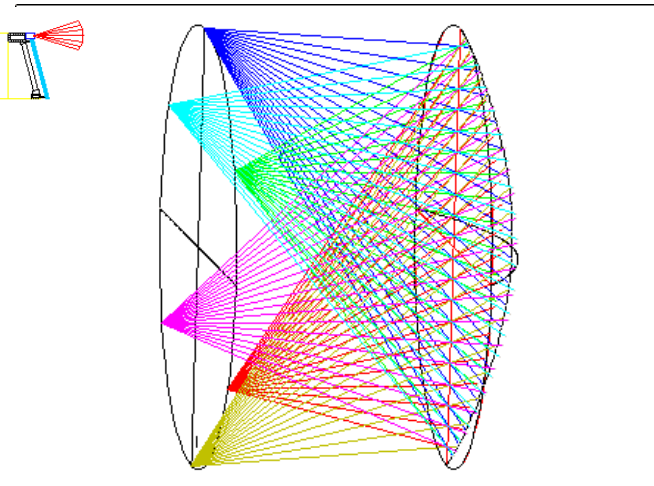
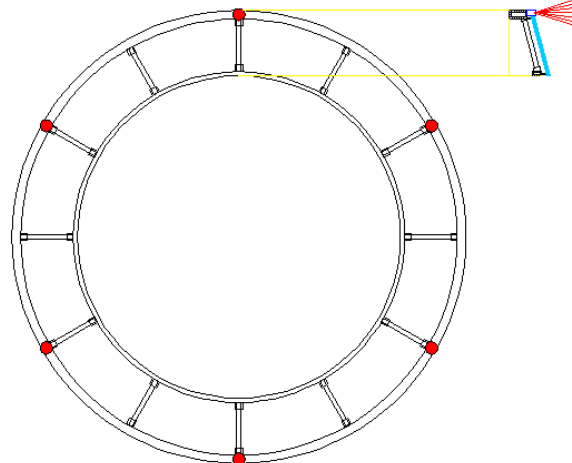
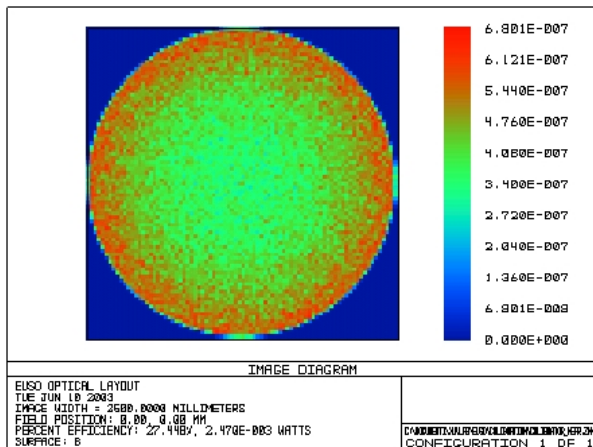
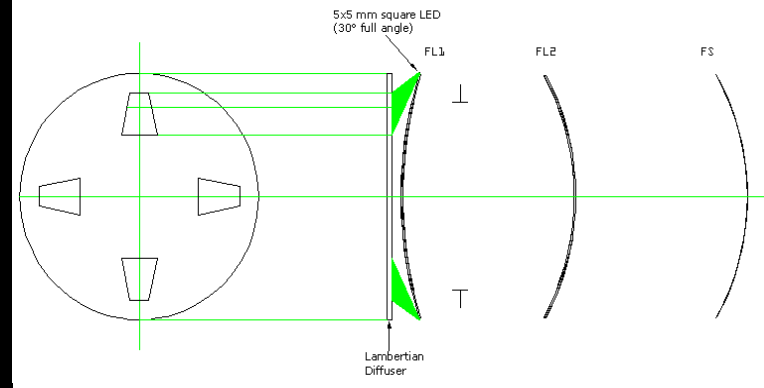
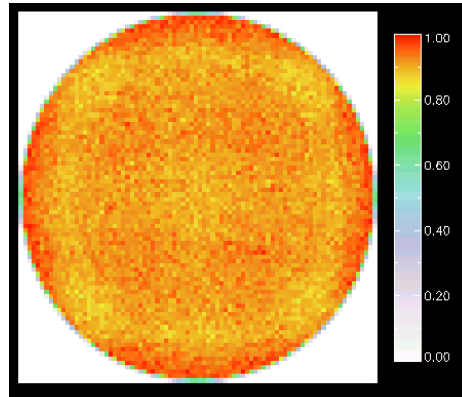
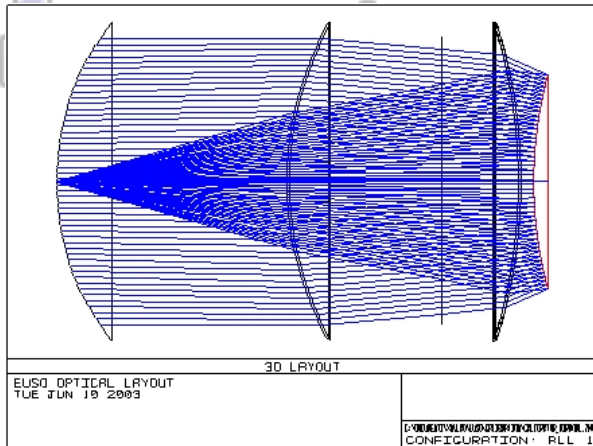
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MON SEP 25 2006  
IMAGE WIDTH = 25.7000 MILLIMETERS, 8 X 8 PIXELS  
FIELD POSITION: 0.0000, 0.0000 MM  
PERCENT EFFICIENCY: 75.276%, 7.528E-001 WATTS  
SURFACE: 4. UNITS ARE WATTS PER MILLIMETERS SQUARED.

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CONFIGURATION 1 OF 1

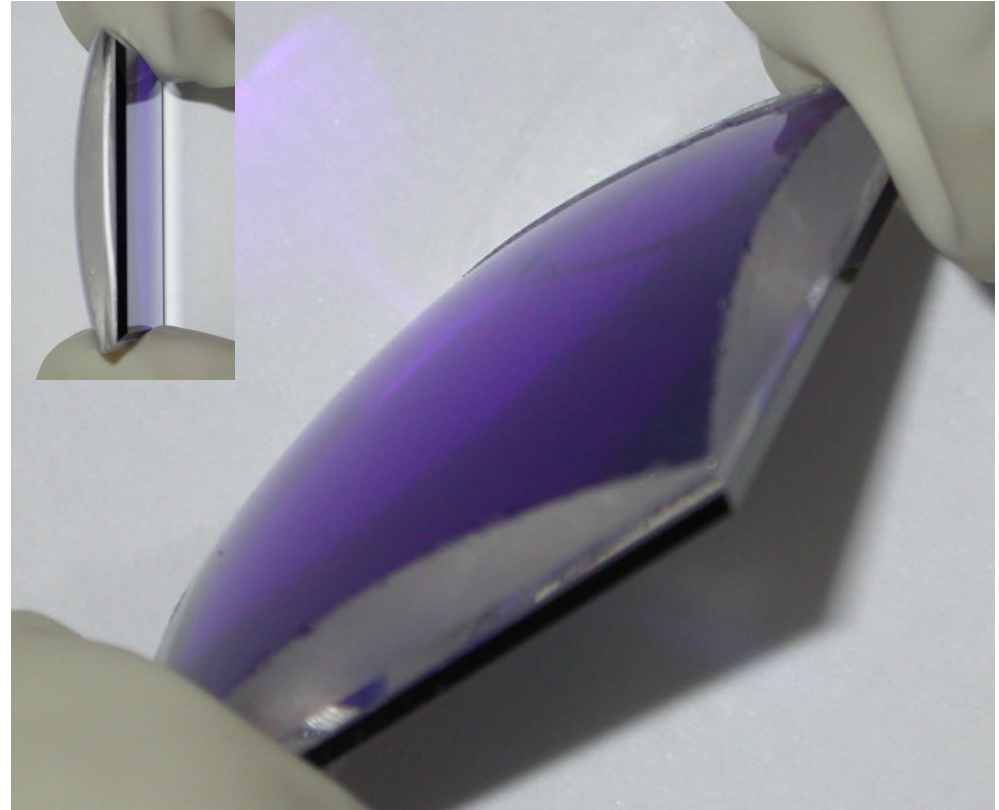
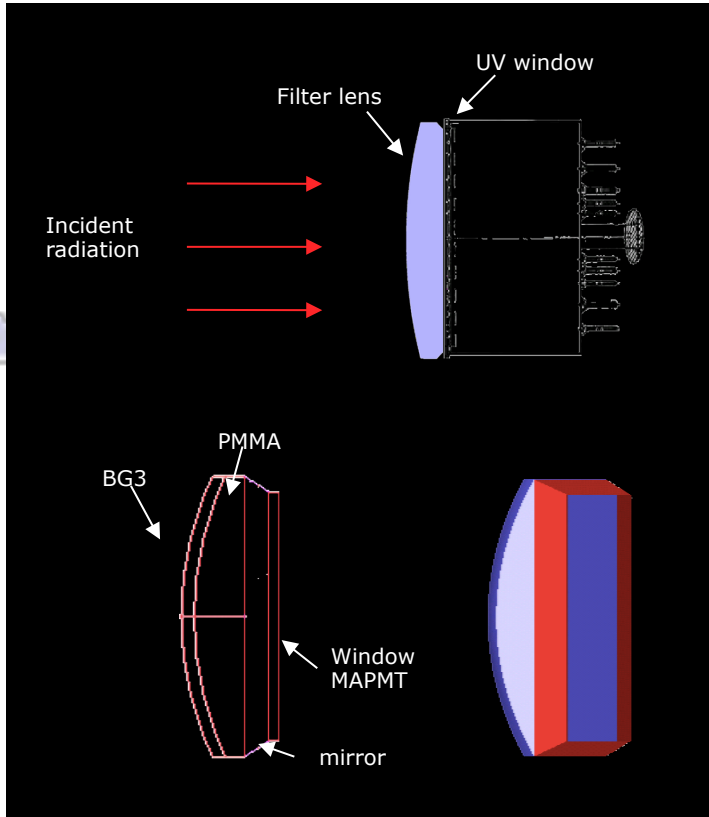
# In-flight calibration



# INFN

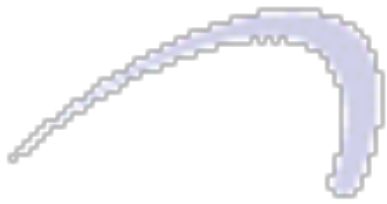


# Optical adapters



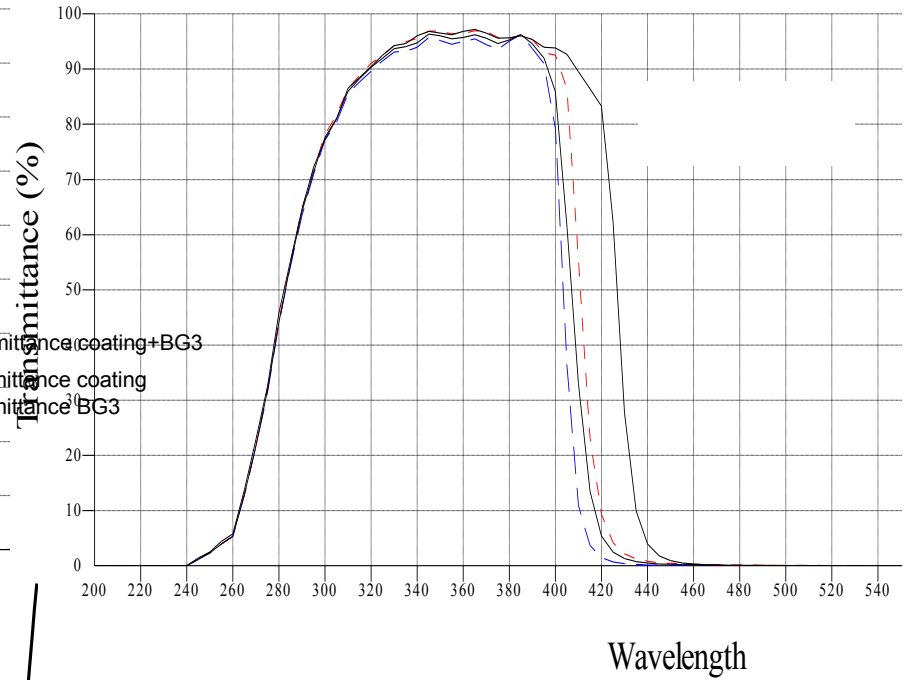
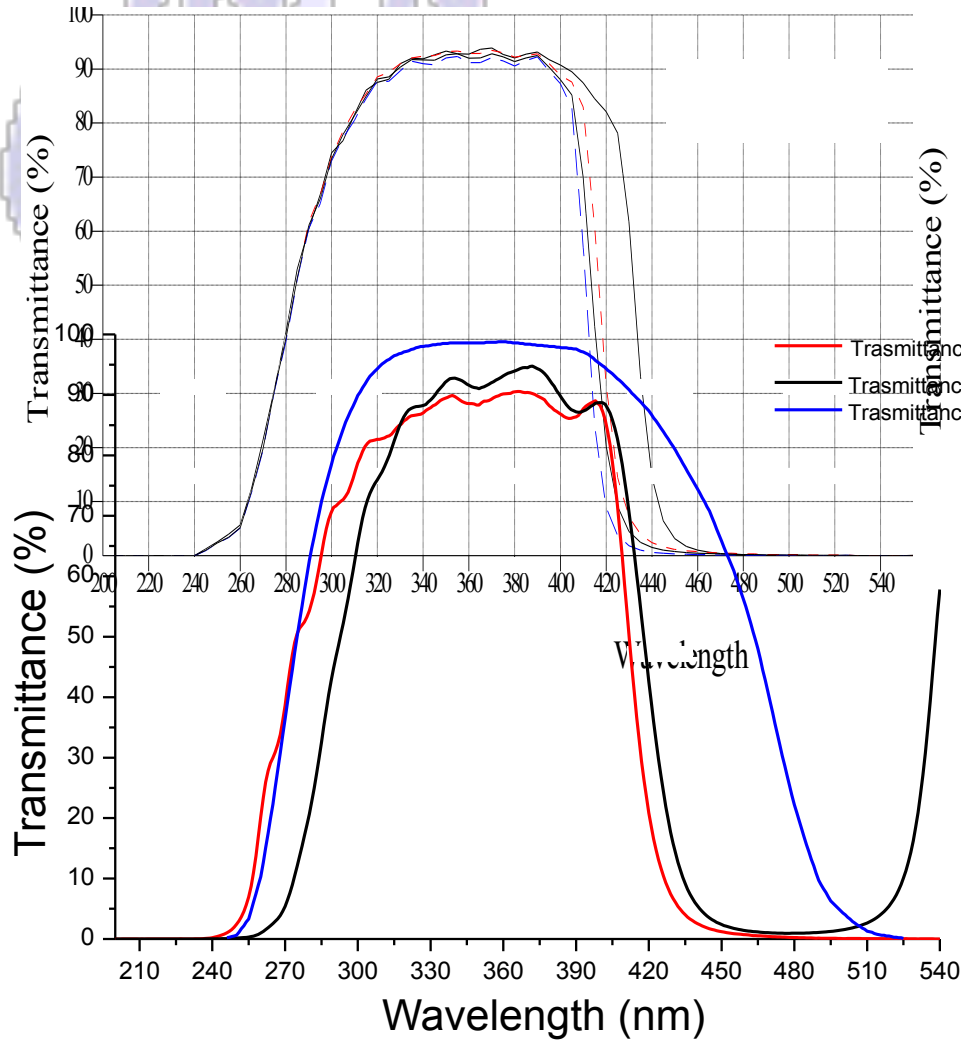
Weight of single lens	Total Weight
3.74 g	20.7 kg

$\lambda$ (nm)	Geometric Efficiency (%)	Radiometric Efficiency (%)
337	99.5	94.7
357	99.5	95.6
391	99.5	94.4

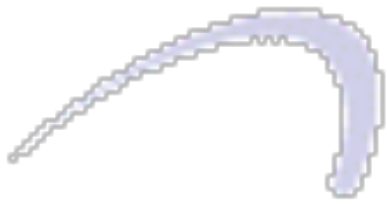


# Innovative filters and coatings

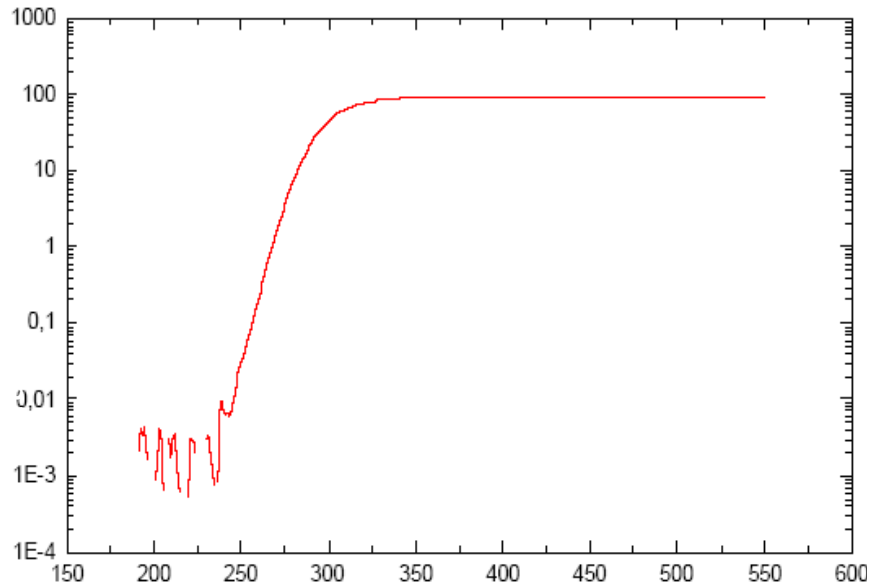
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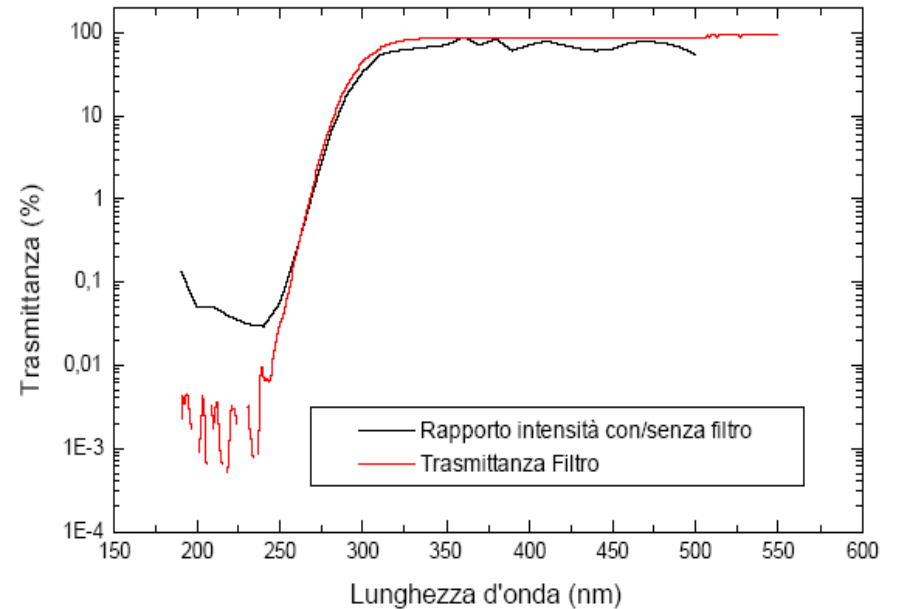




# Spectro-photometry



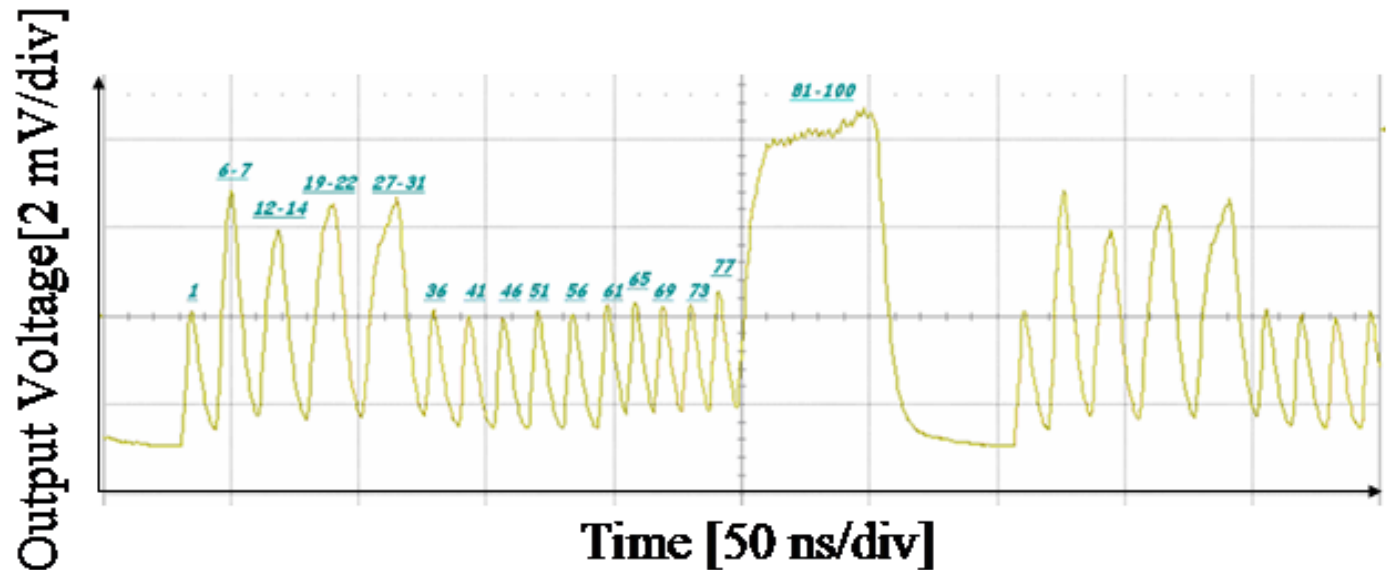
$$T_{\lambda} = \frac{I_{\lambda}}{I_{(\lambda_0)}}$$



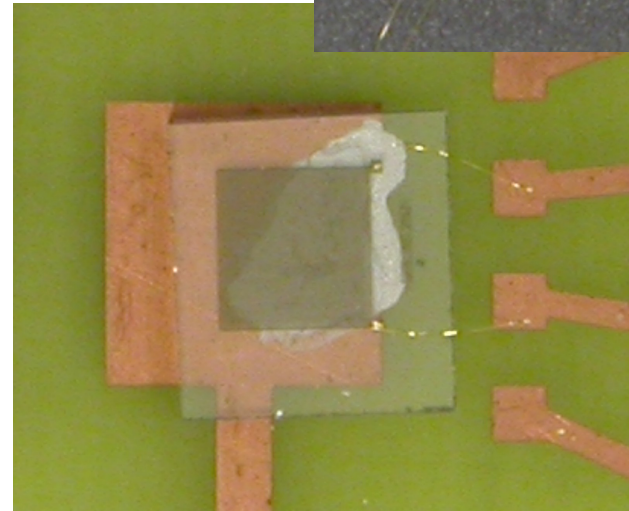
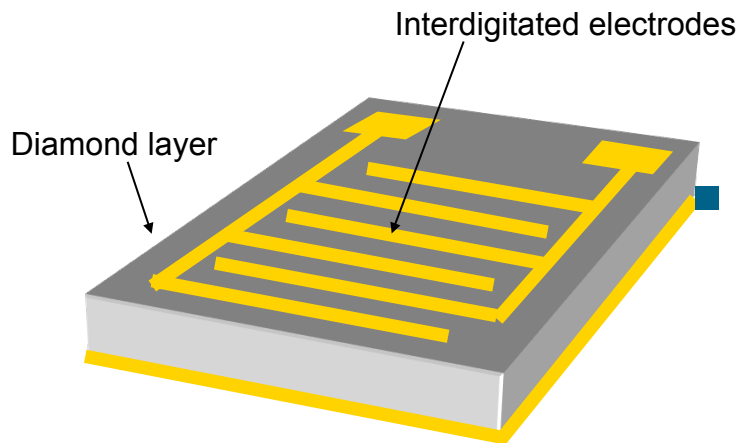
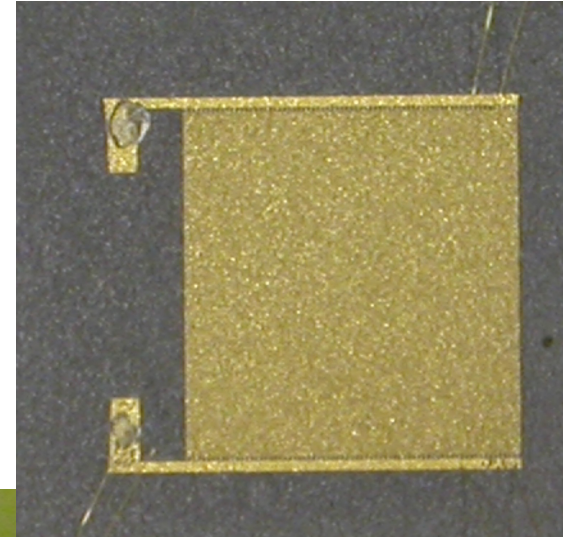
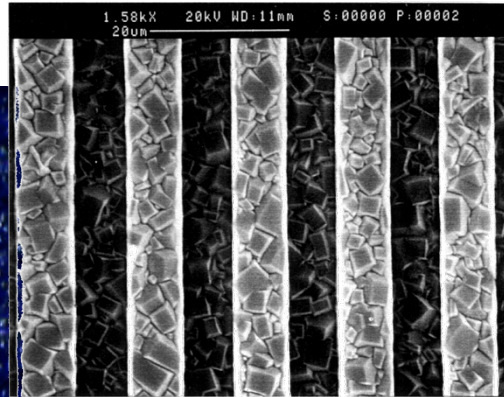
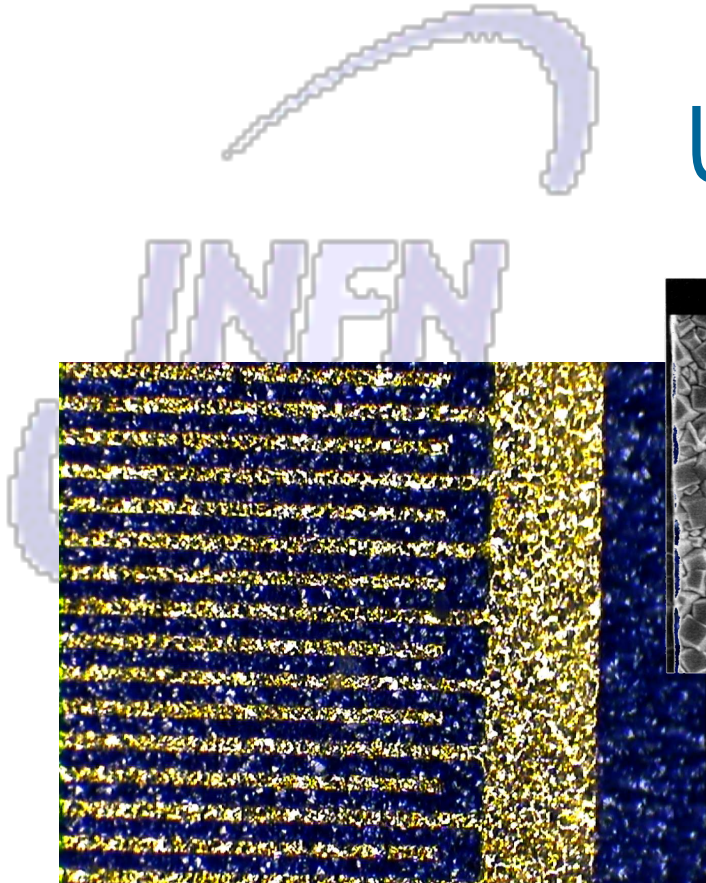


# Technique to measure response time

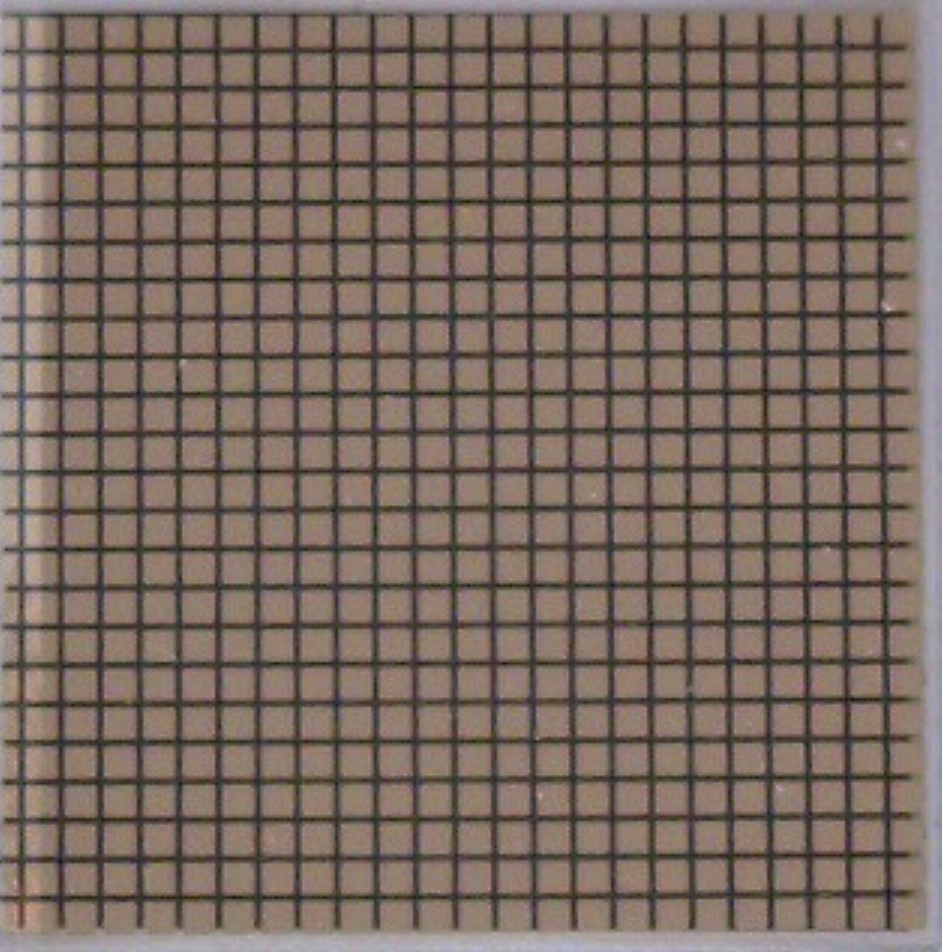
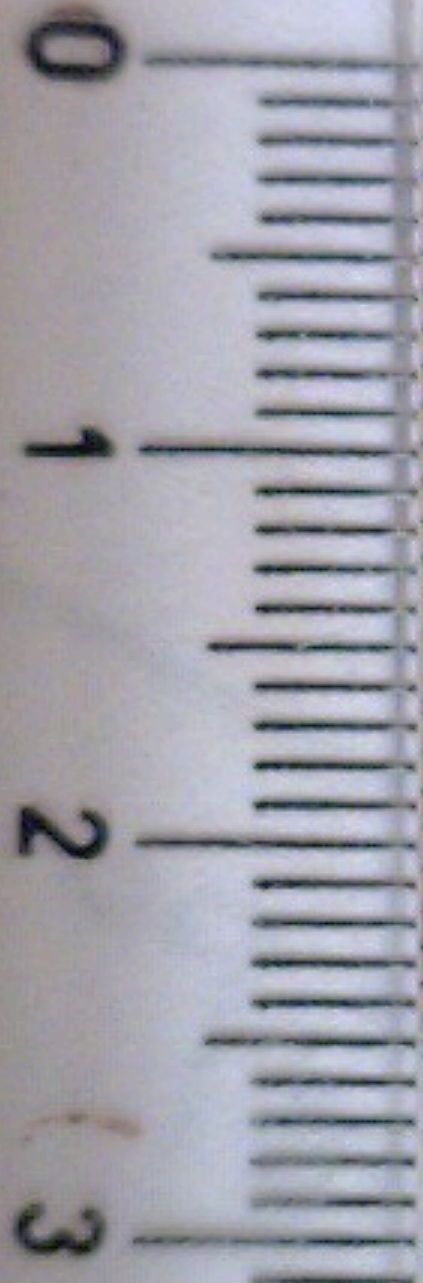
1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20
21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40
41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60
61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80
81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
101	102	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120



# UV diamond photodetectors



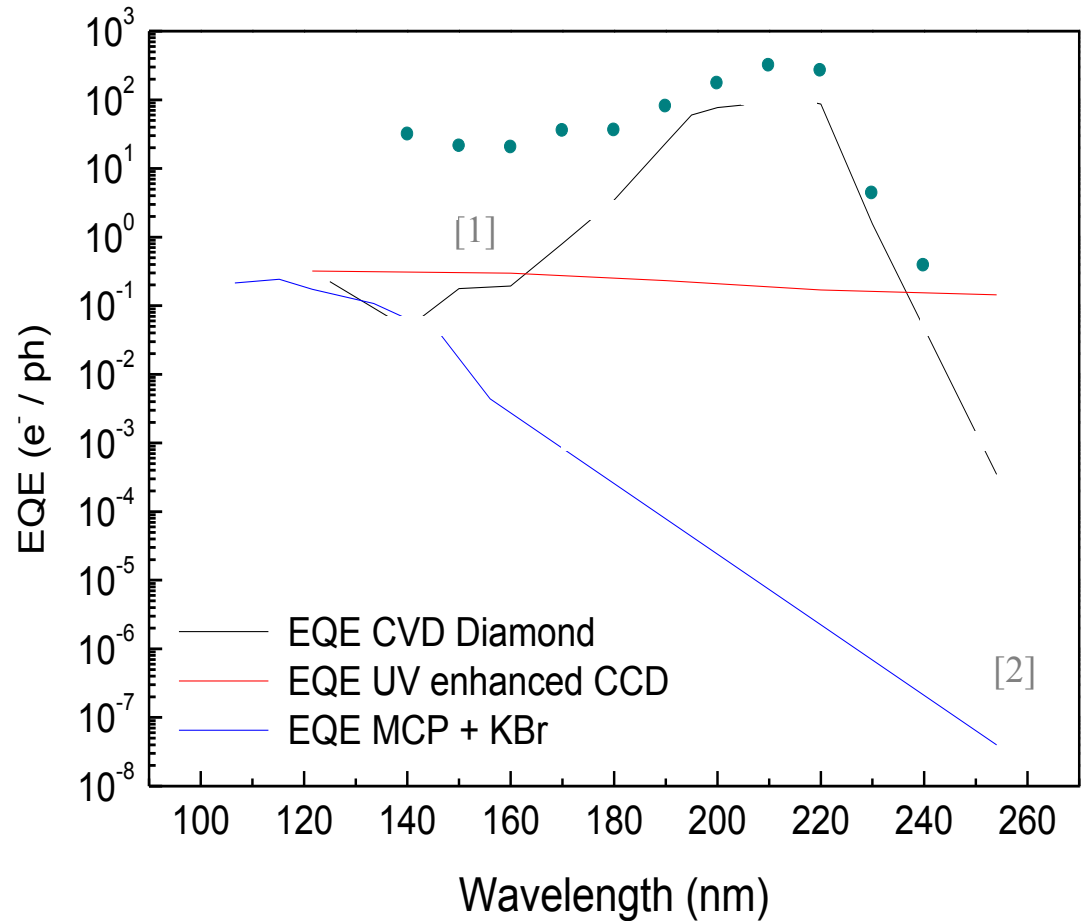






# Diamond detectors

## Confronto

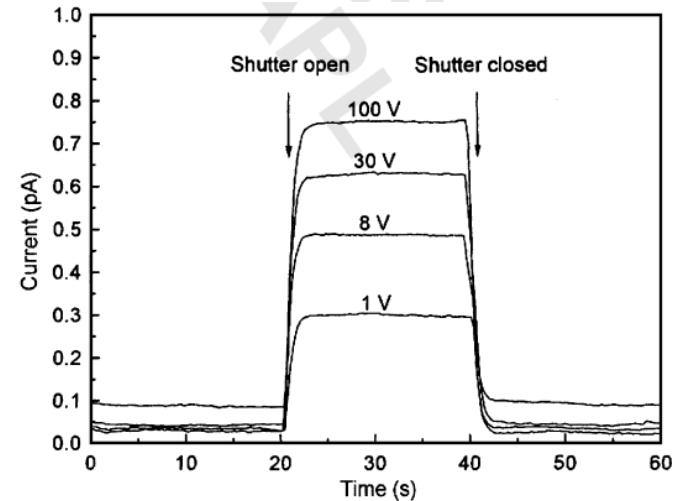
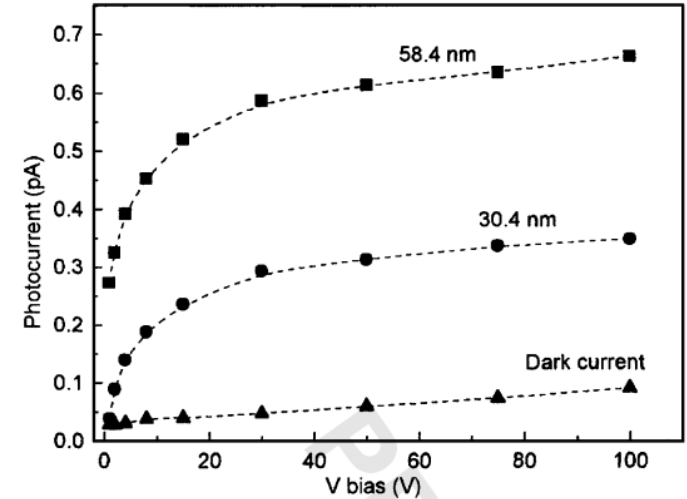
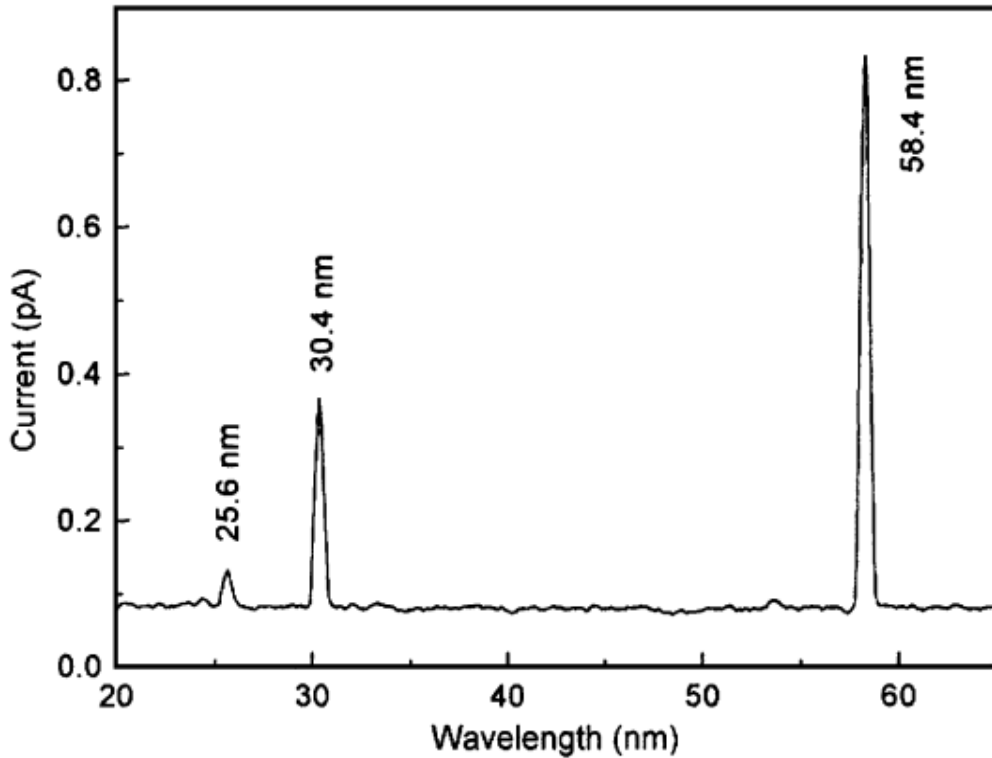


[1] Naletto, Pace et al, 1994

[2] Wilhelm et al., 1995

# Diamond detectors

## VUV response

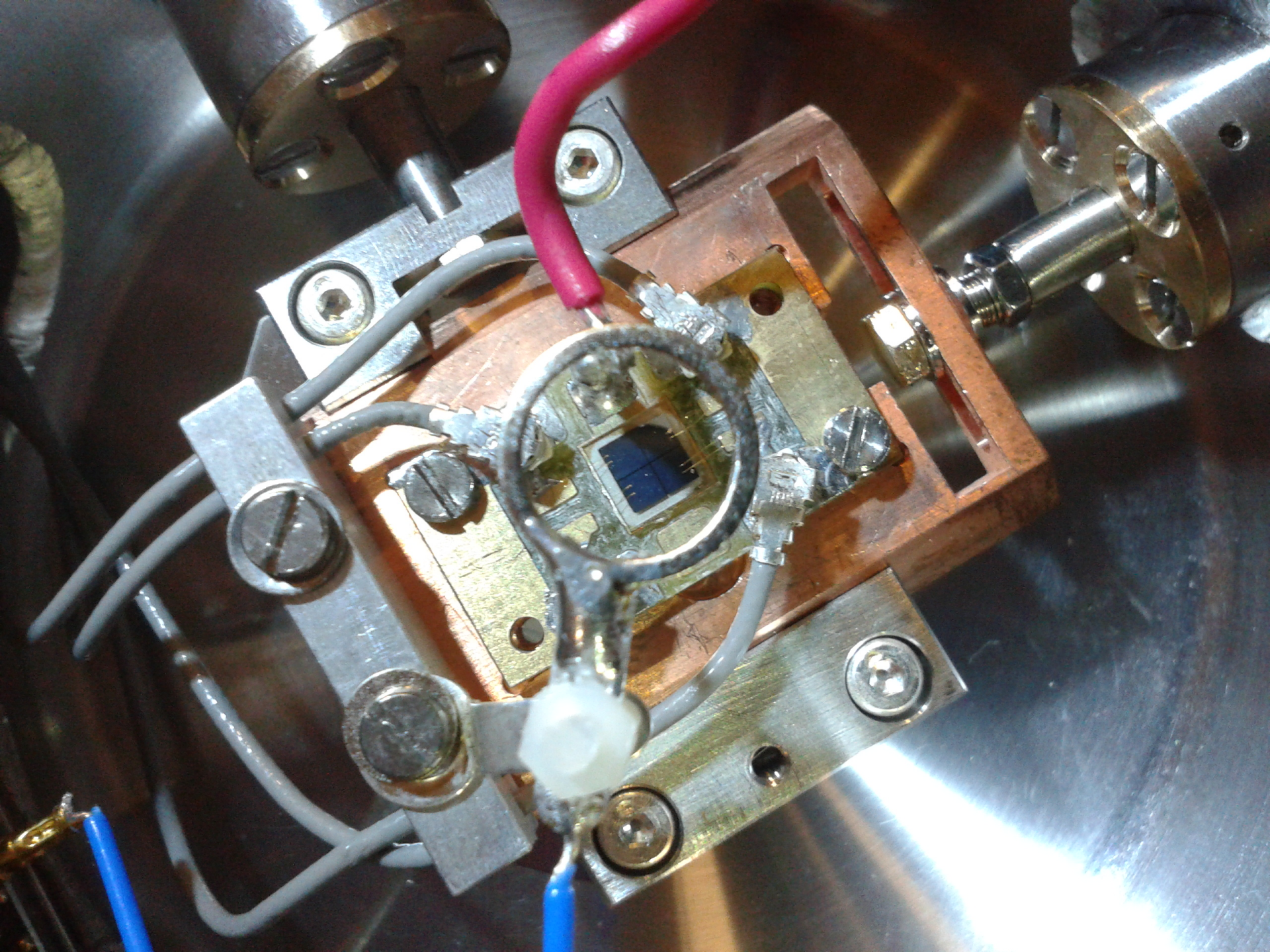




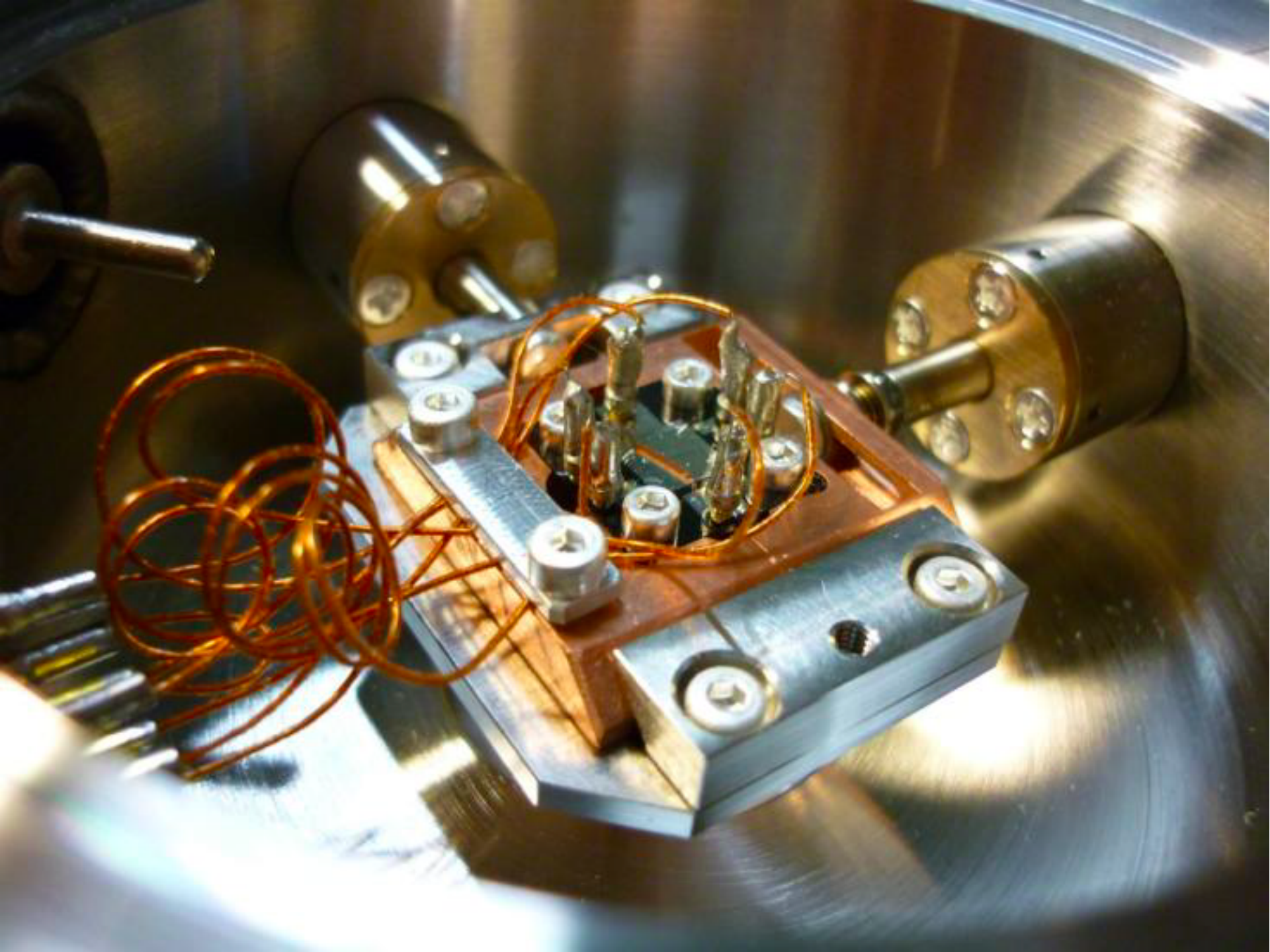
Scopo

Development of ultrafast X-UV diamond detectors for FEL and SR applications

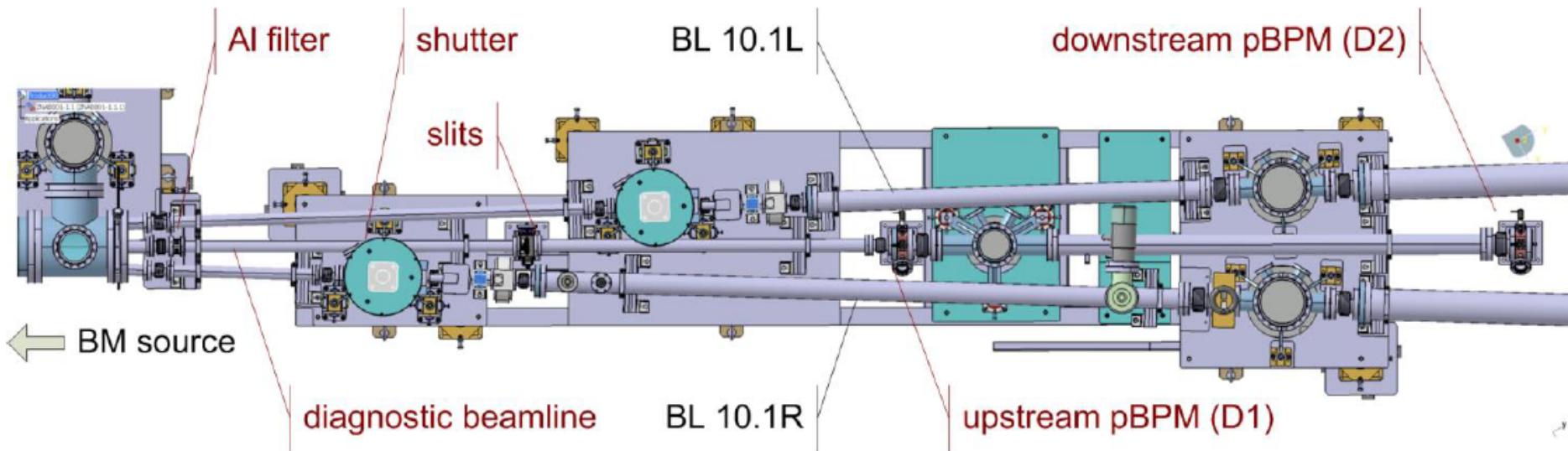




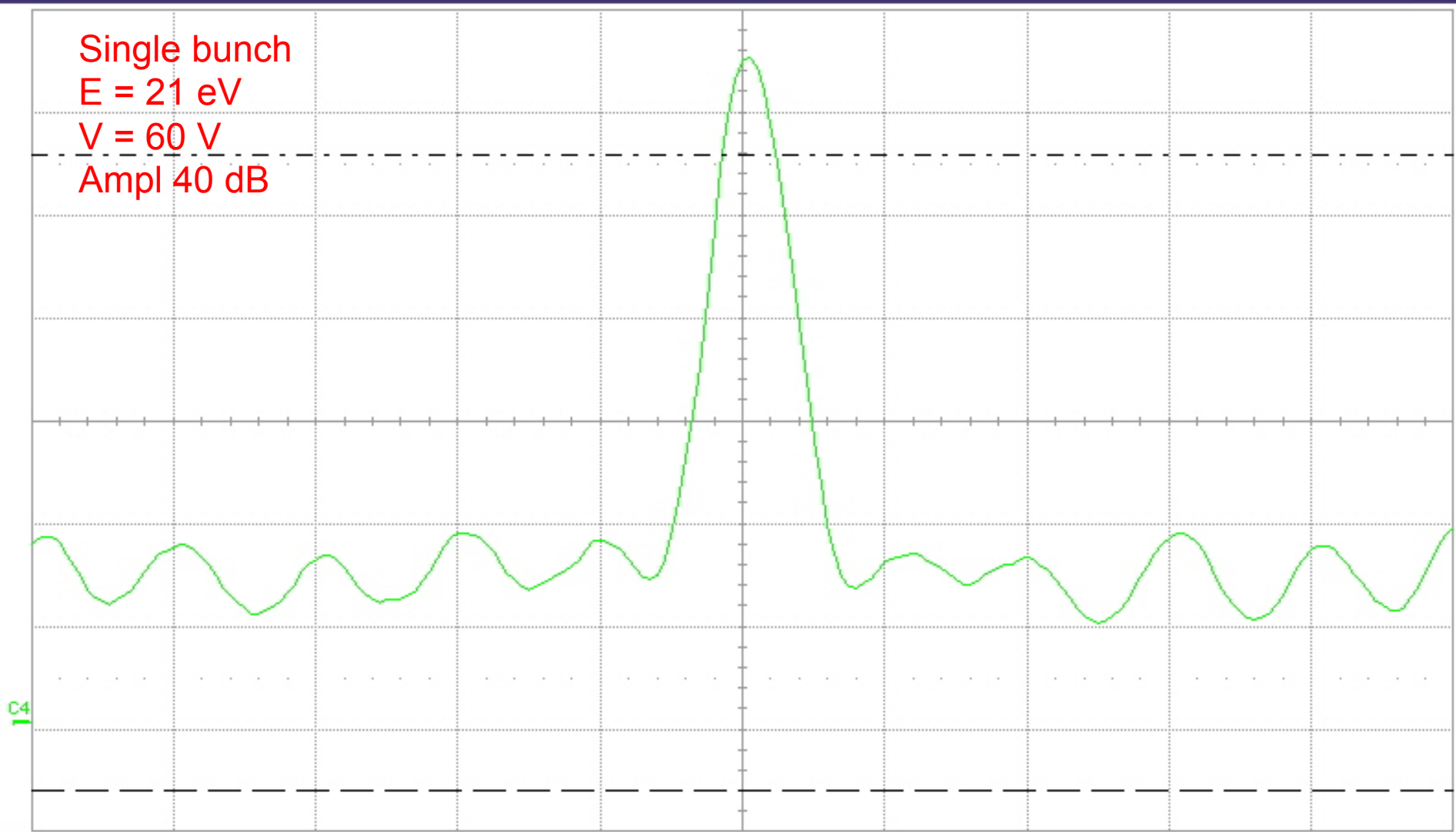




# pBPM

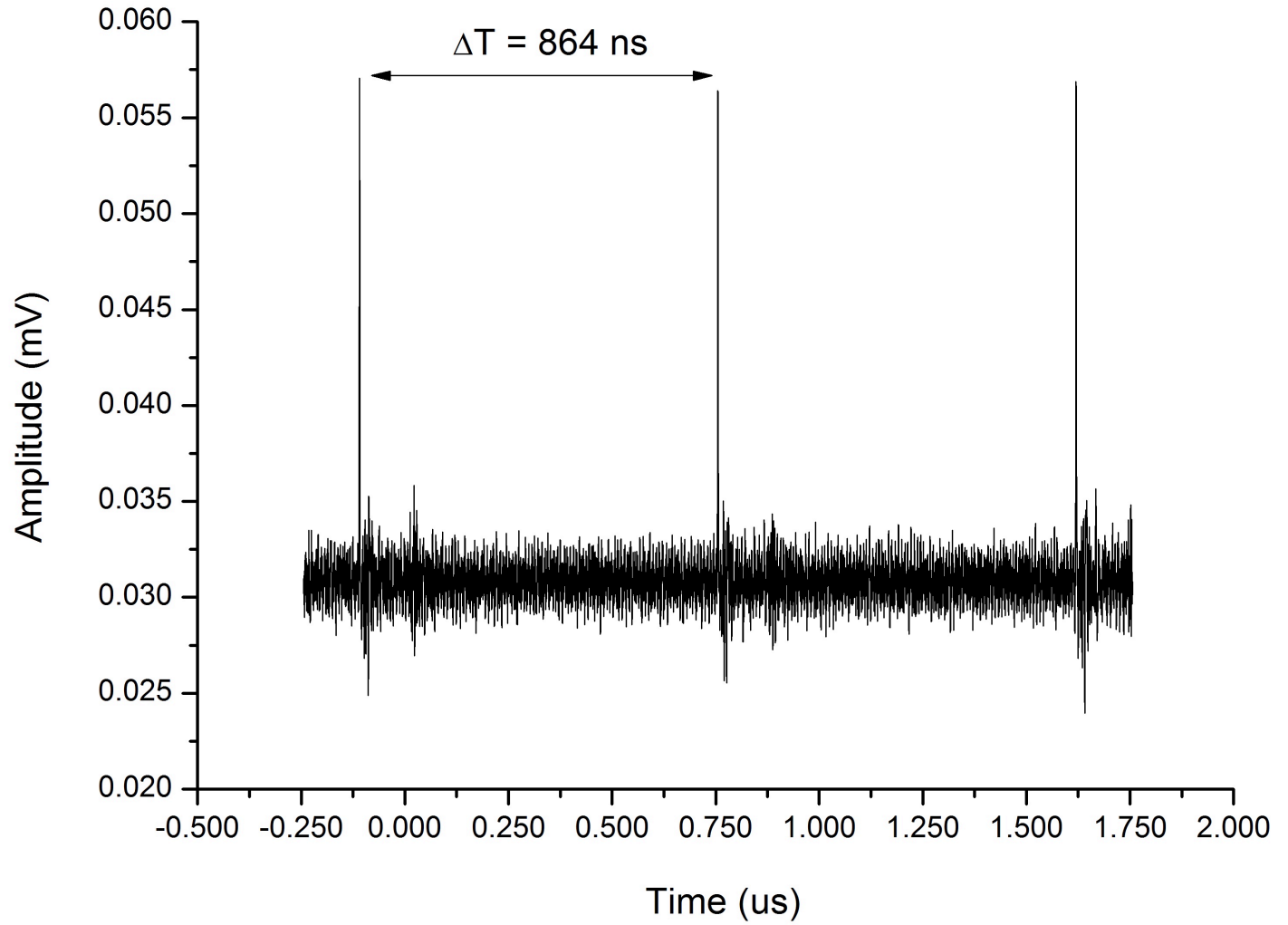
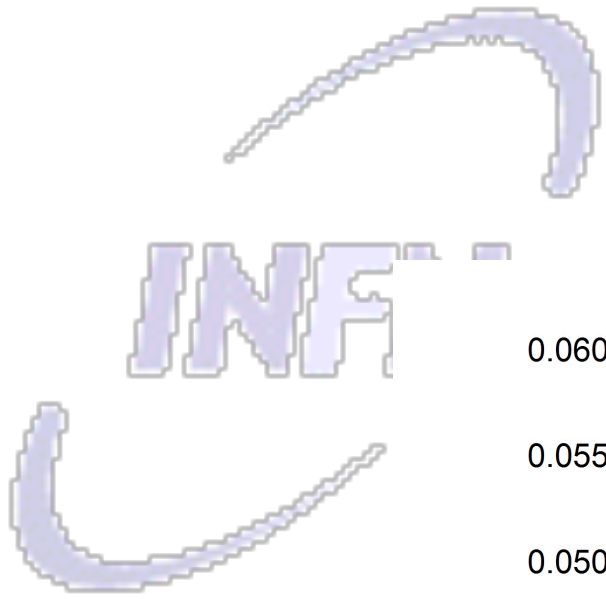


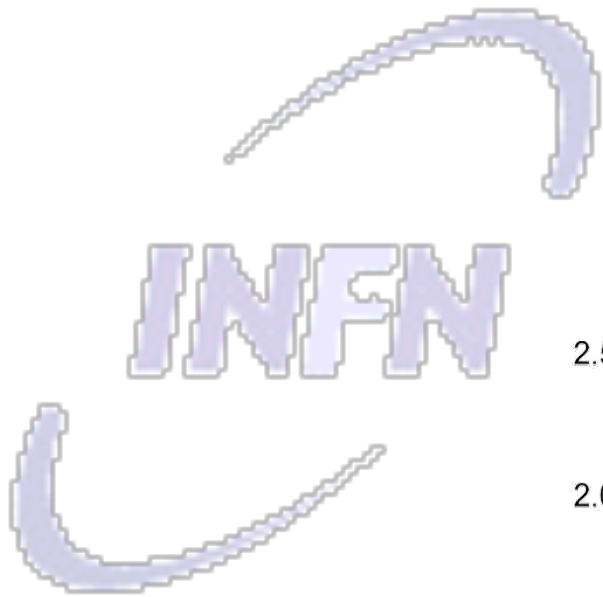
Single bunch  
E = 21 eV  
V = 60 V  
Ampl 40 dB



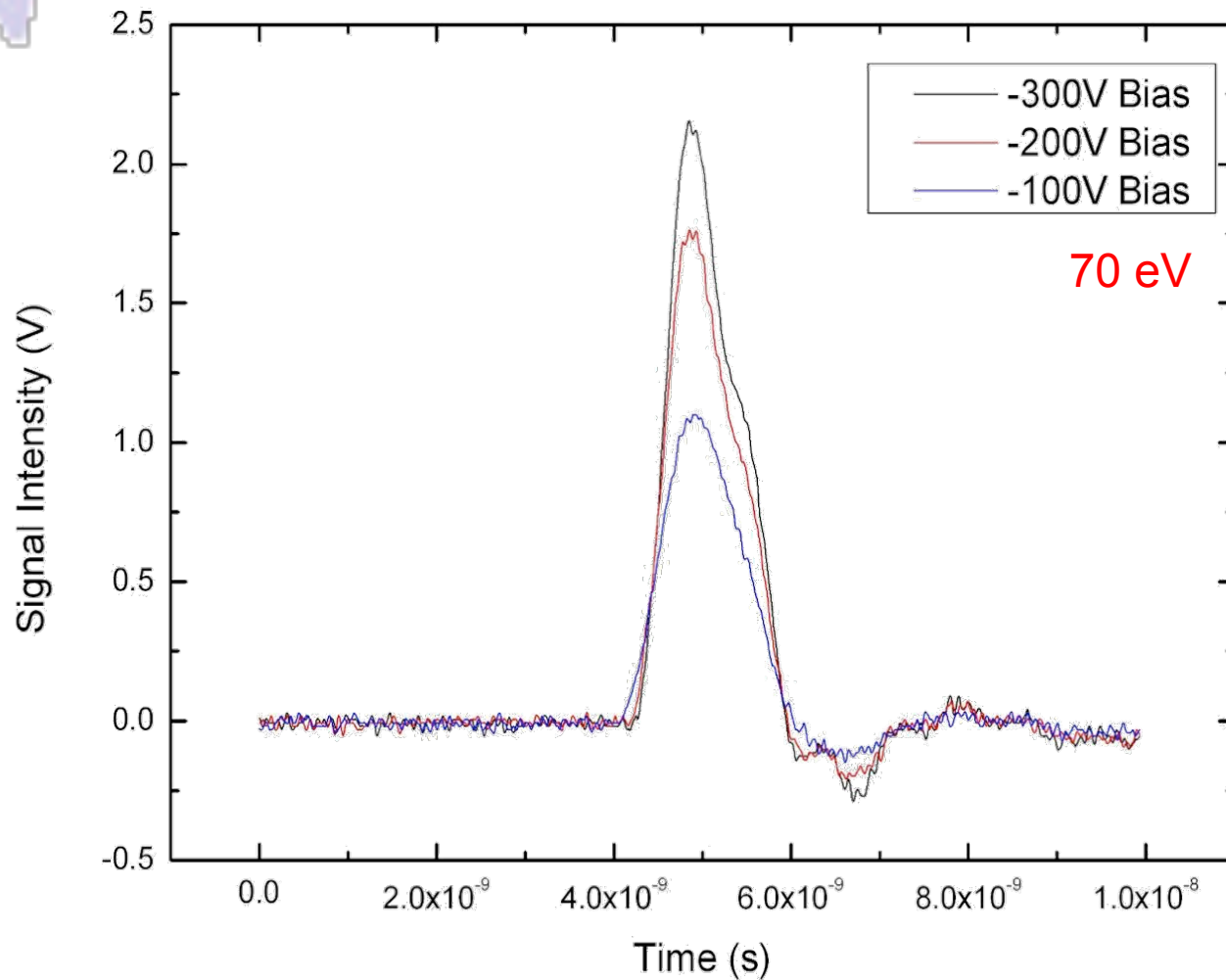
C4  
A1 DC50  
10.0 mV/div  
-29.400 mV  
3.523 k#  
---- 55.2 mV  
..... -6.5 mV

Tbase	-755.28 ns	Trigger	Ext
	2.00 ns/div	Normal	410 mV
200 S	10 GS/s	Edge	Negative

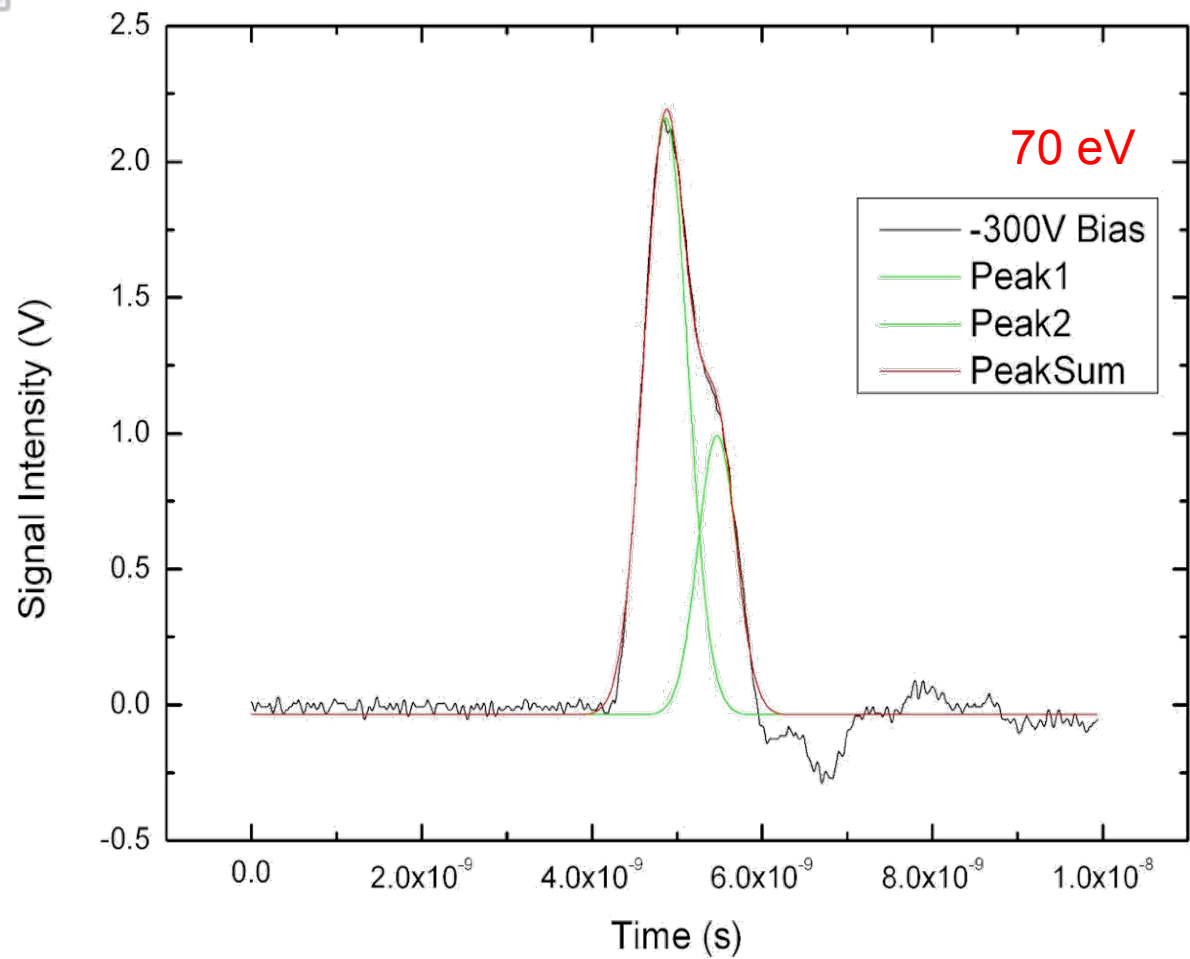
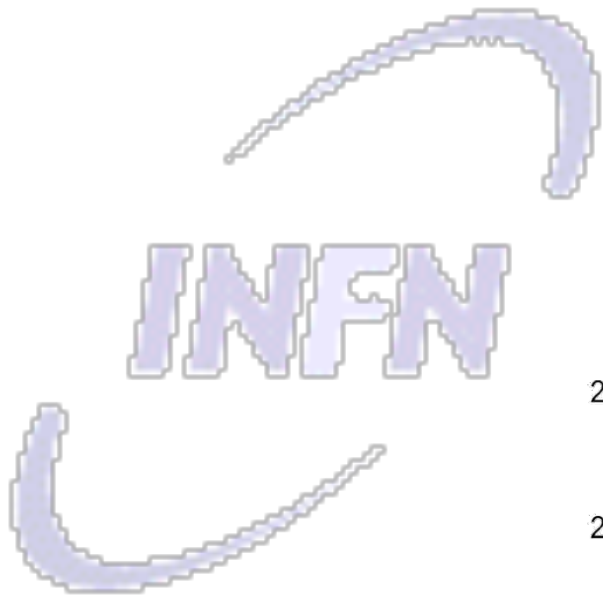


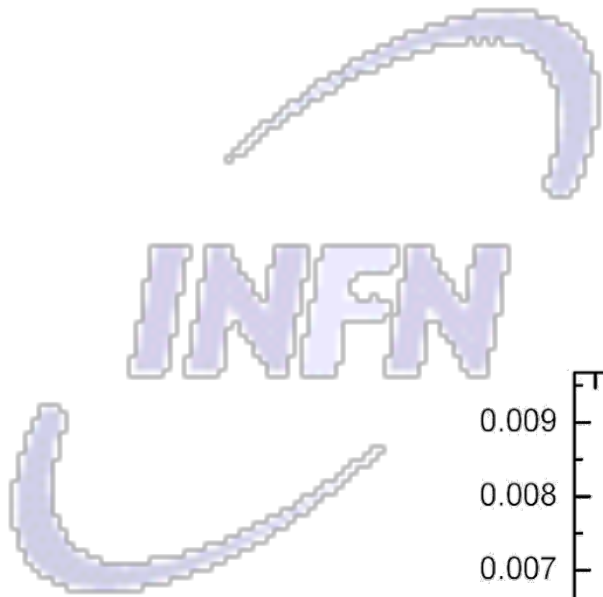


SCD/Cr-Au E6

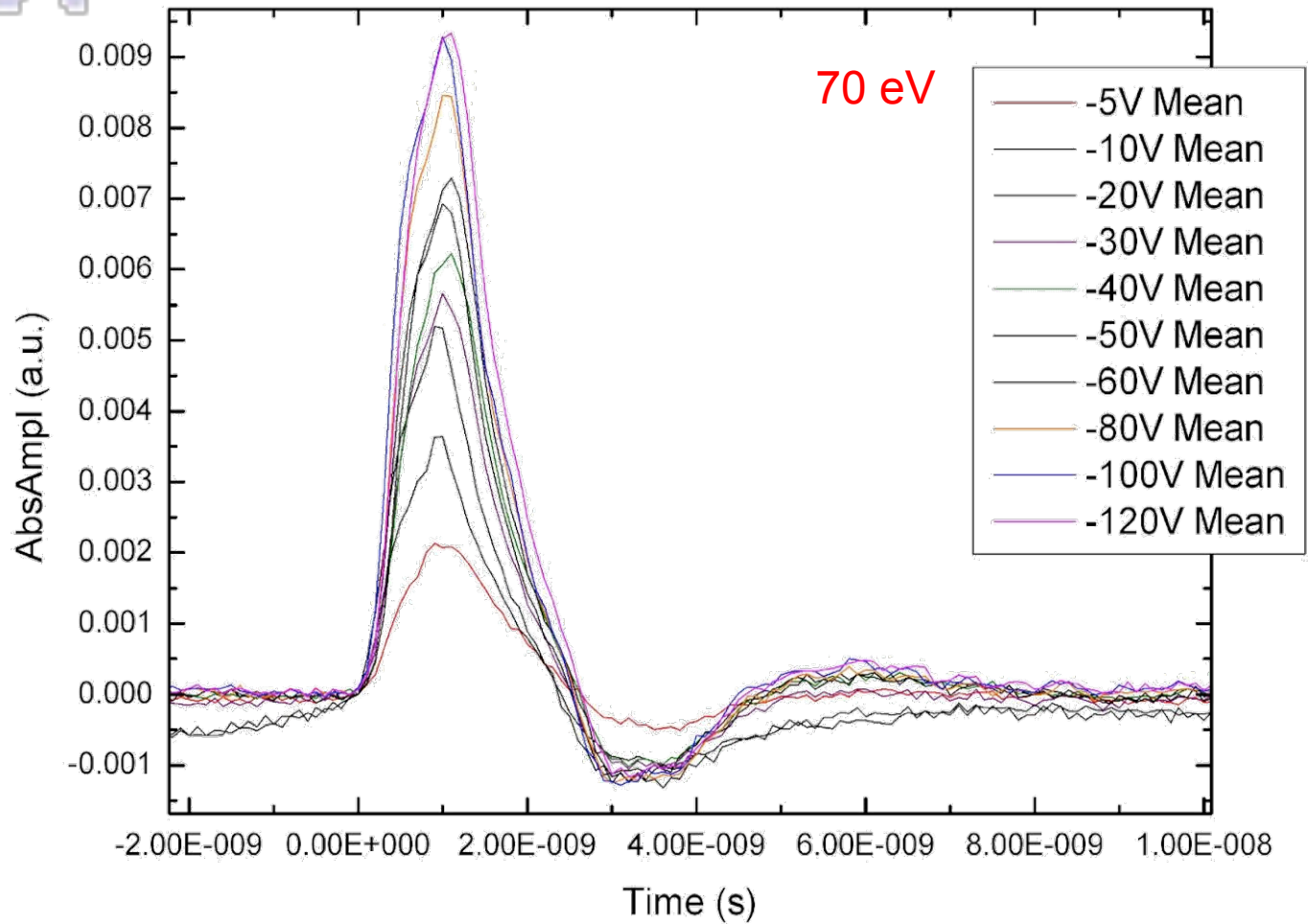








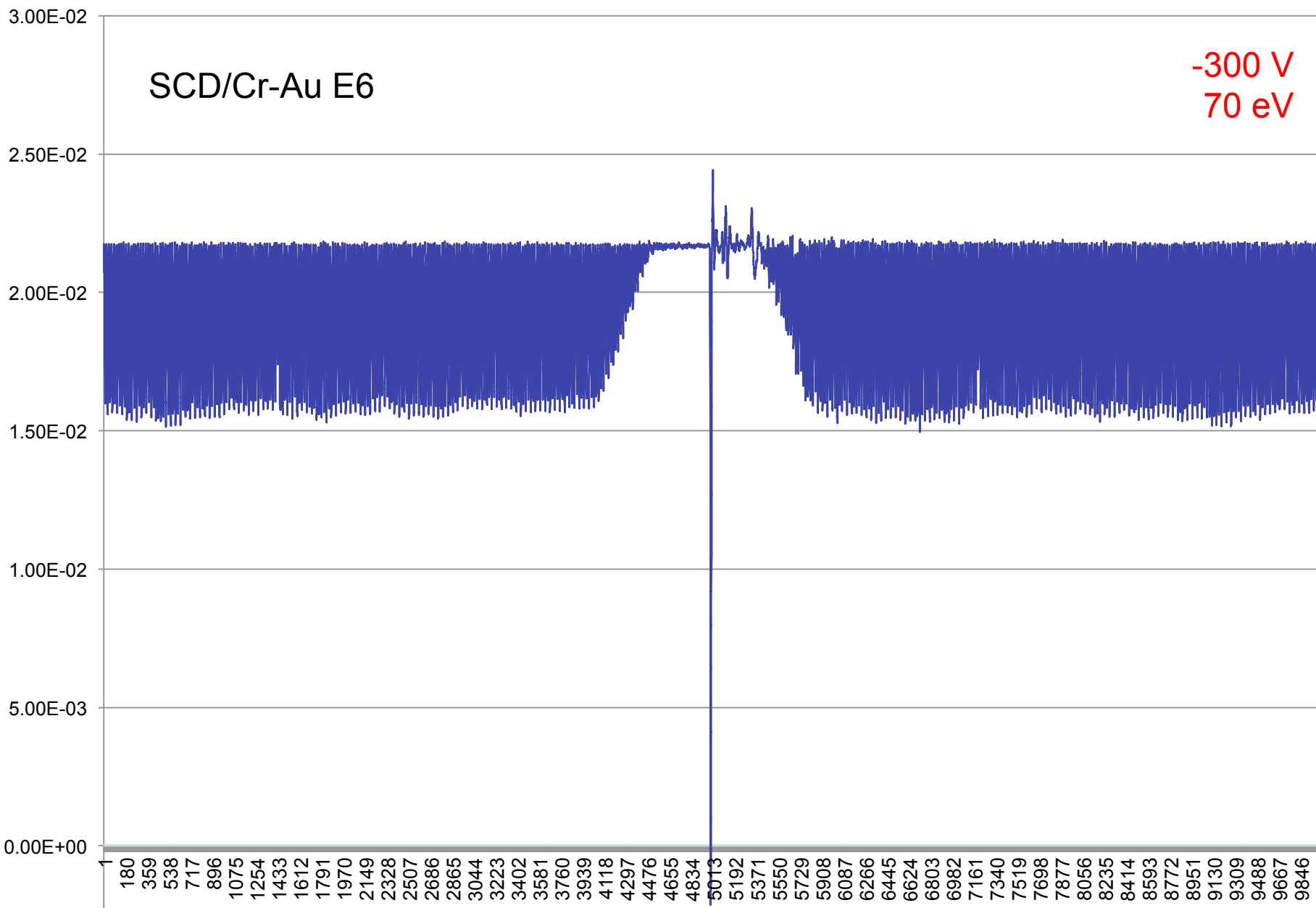
# SCD/AI interdigitated UTOV



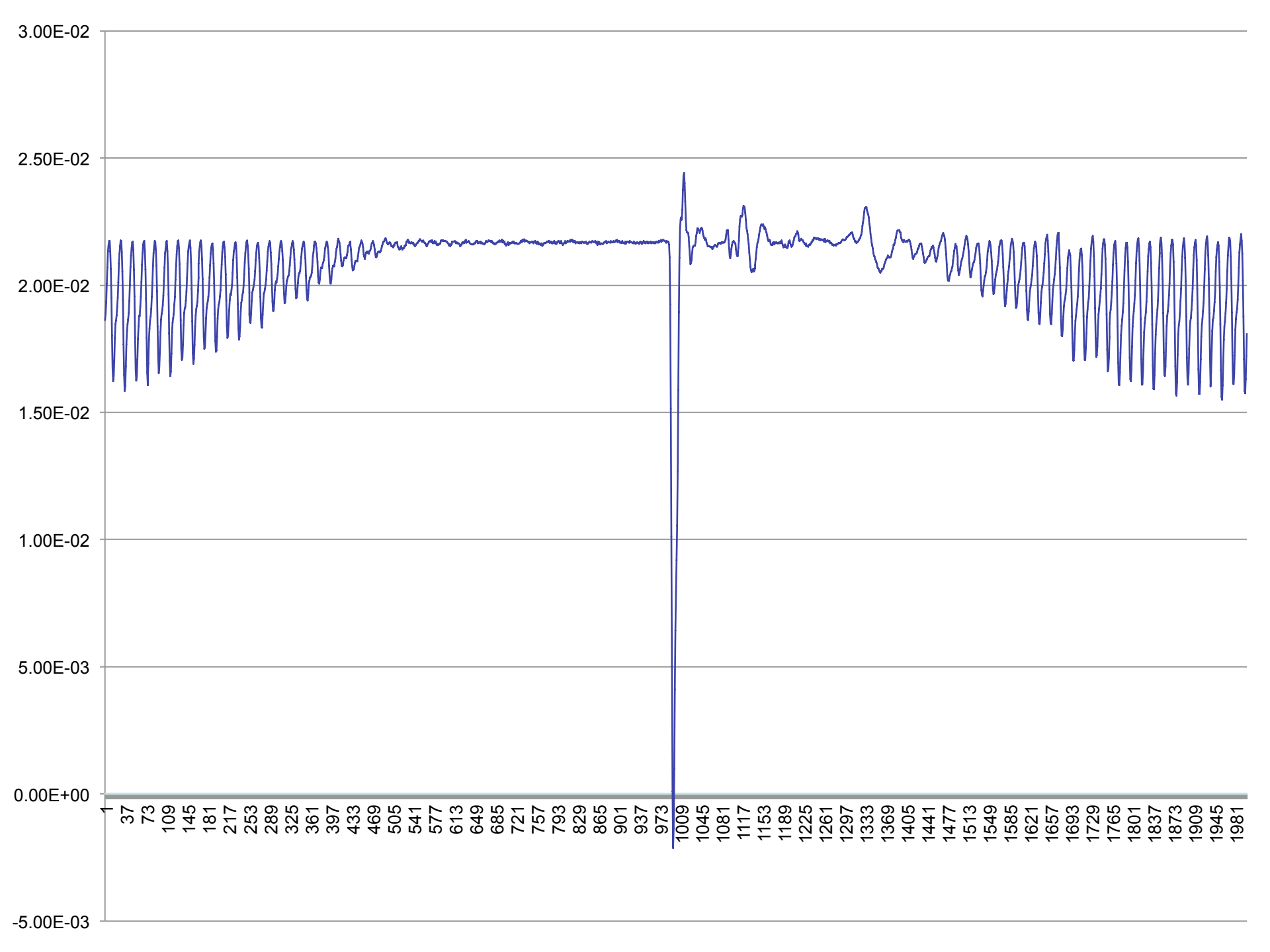


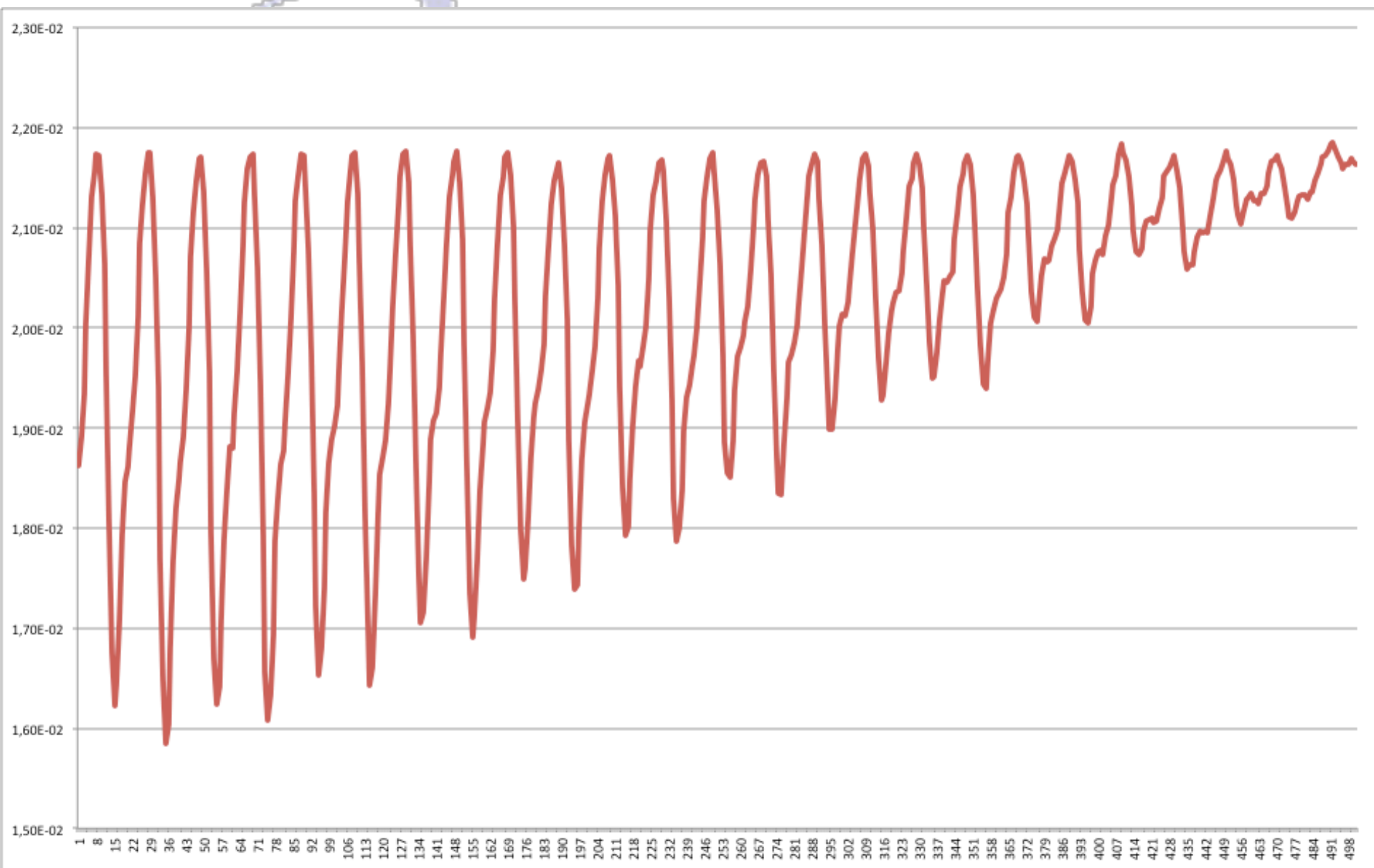
SCD/Cr-Au E6

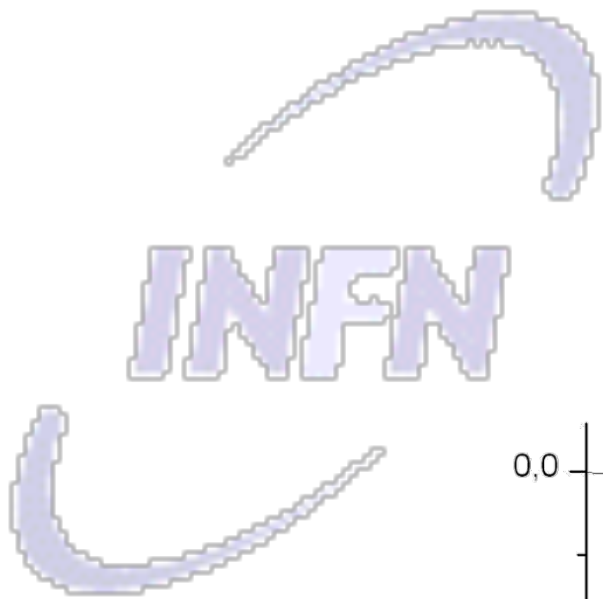
-300 V  
70 eV



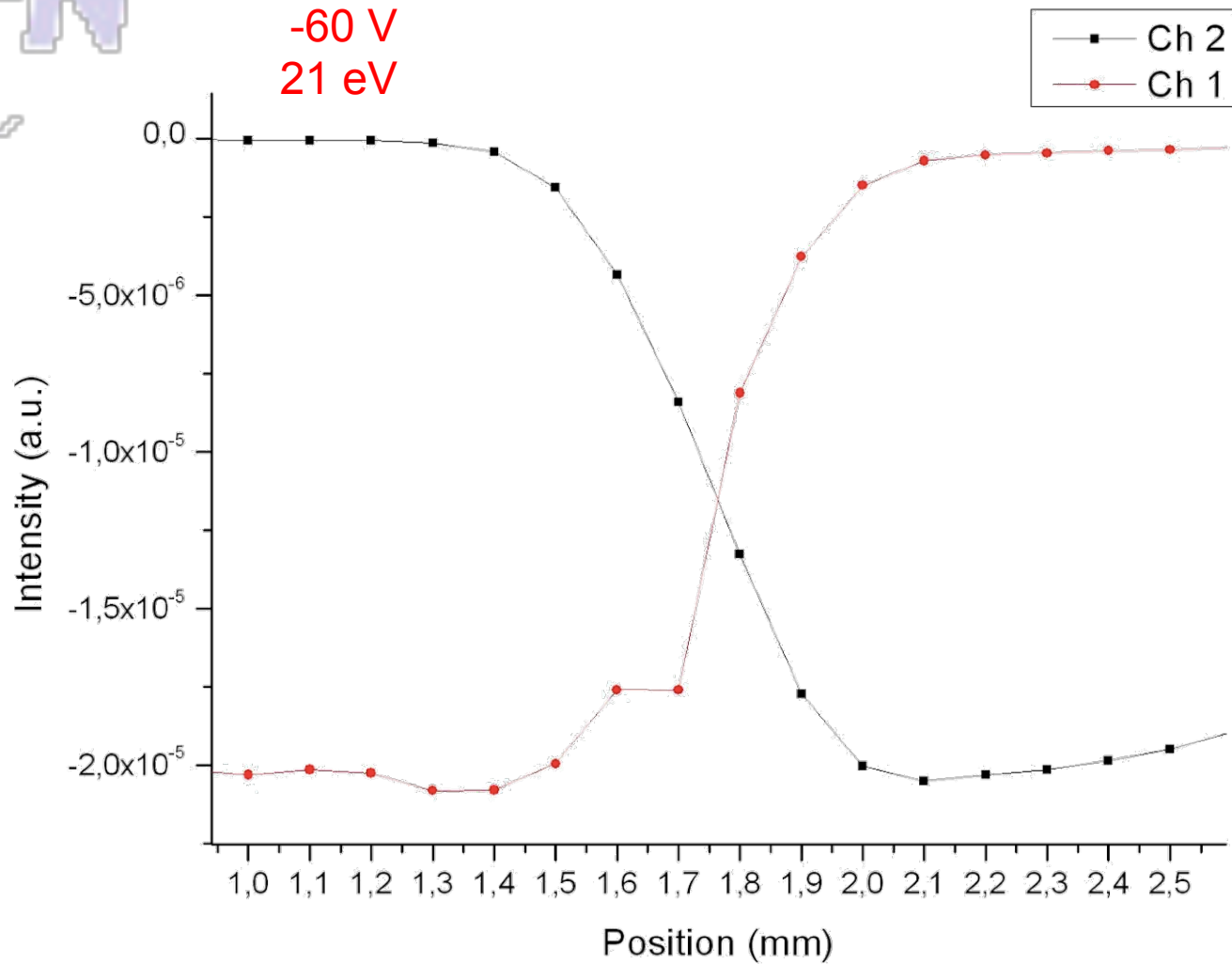
1 180 359 538 717 896 1075 1254 1433 1612 1791 1970 2149 2328 2507 2686 2865 3044 3223 3402 3581 3760 3939 4118 4297 4476 4655 4834 5013 5192 5371 5550 5729 5908 6087 6266 6445 6624 6803 6982 7161 7340 7519 7698 7877 8056 8235 8414 8593 8772 8951 9130 9309 9488 9667 9846

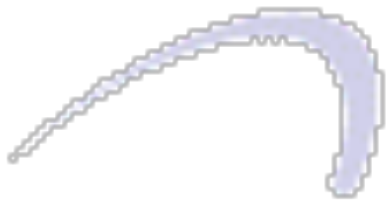






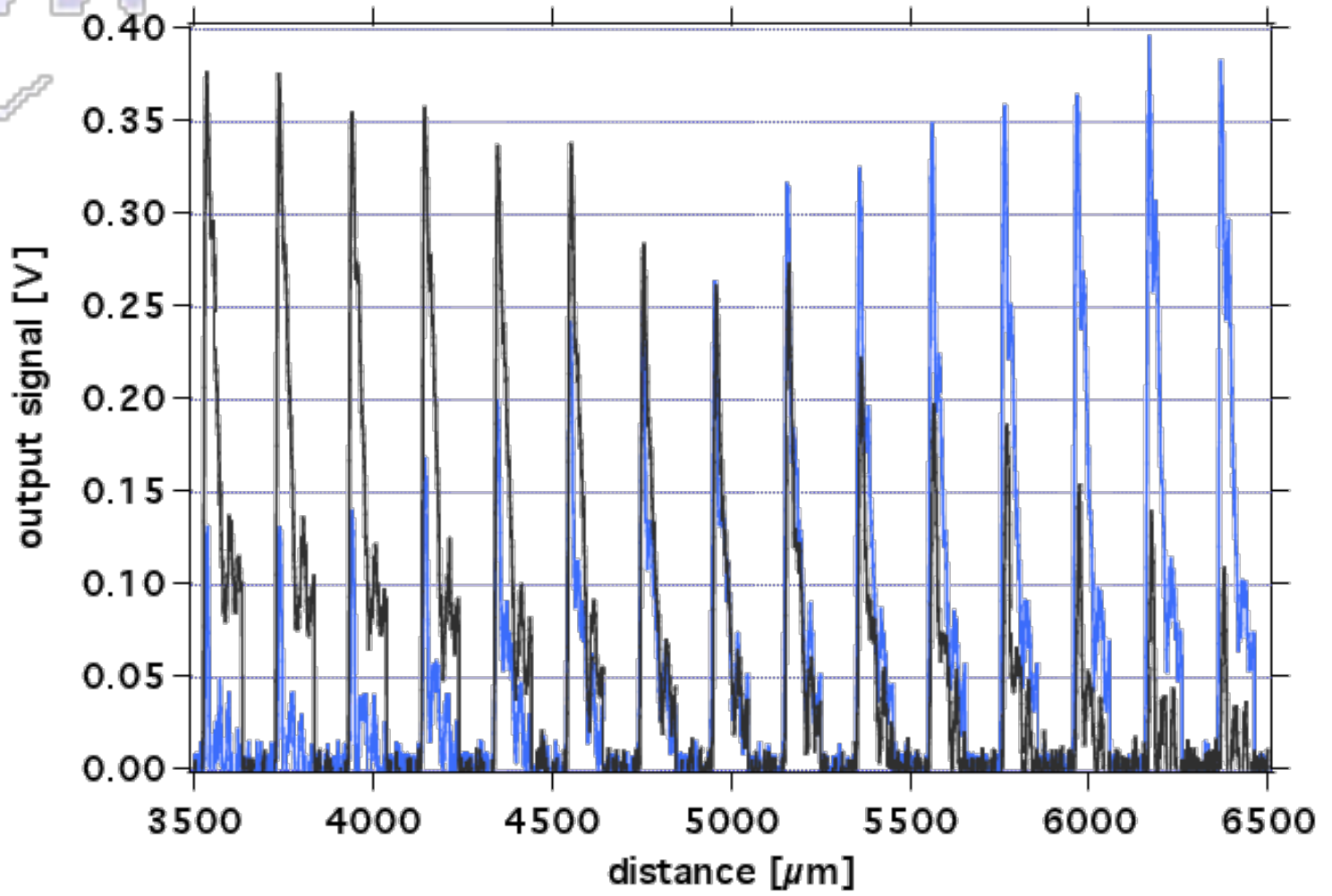
-60 V  
21 eV

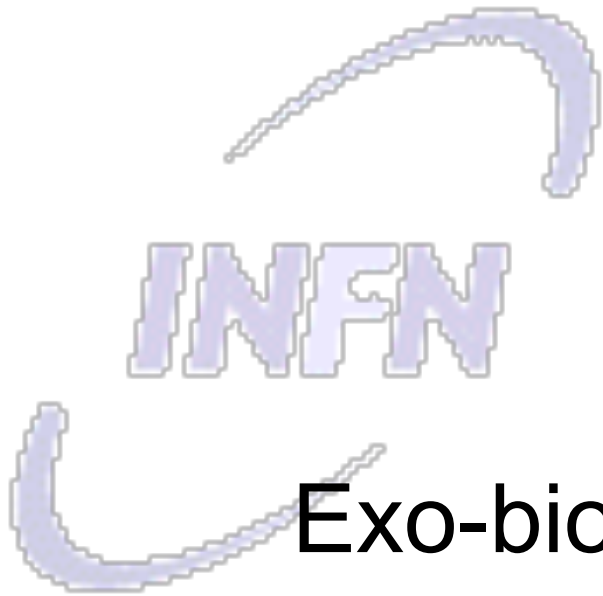




INFN

-300 V  
70 eV





# Exo-biosphere evolution and biosignature characterization by FT-IR spectroscopy

Debora Schierano<sup>1,2</sup>, Andrea Serra<sup>1,3</sup>, Mariangela Cestelli Guidi<sup>1</sup>, Emanuele Pace<sup>1,2</sup>

<sup>1</sup>INFN-Laboratori Nazionali di Frascati

<sup>2</sup>Università degli studi di Firenze

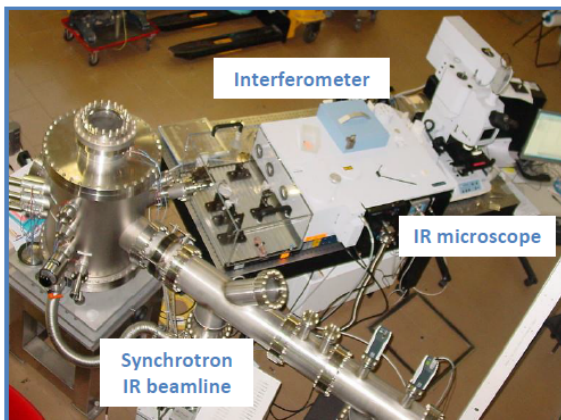
<sup>3</sup>Università degli studi di Roma 'La Sapienza'



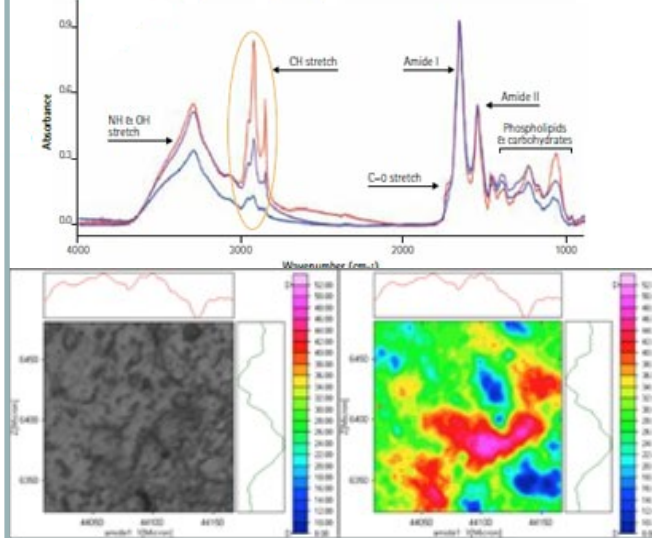
# INFN DaΦne-Ligh synchrotron facility

A synchrotron facility operating with synchrotron and standard sources in the Infrared and UV-VIS energy range is open to external users.

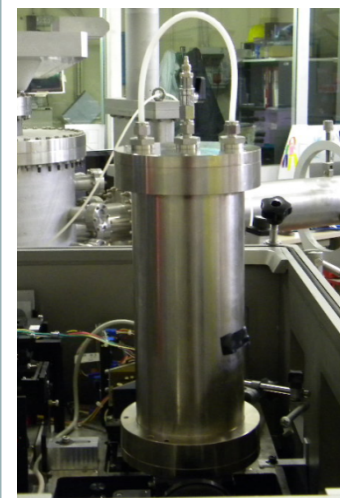
## Sinbad IR beamline (1 mm – 0.5 $\mu$ m)



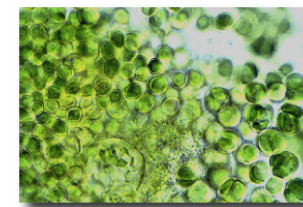
## FPA detector for chemical imaging



## Long path gas cell



- Extended IR range (from Far-IR to NIR-VIS)
- Chemical microimaging of materials
- High temperature (1200°C) / high pressure (20 GPa) setup
- Real time study of photoageing processes
- Exo-biosphere evolution and biosignature characterization

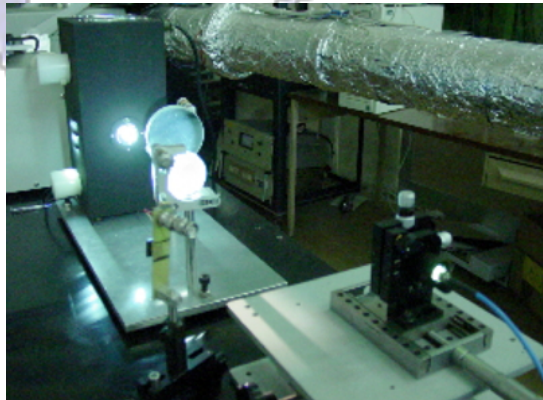






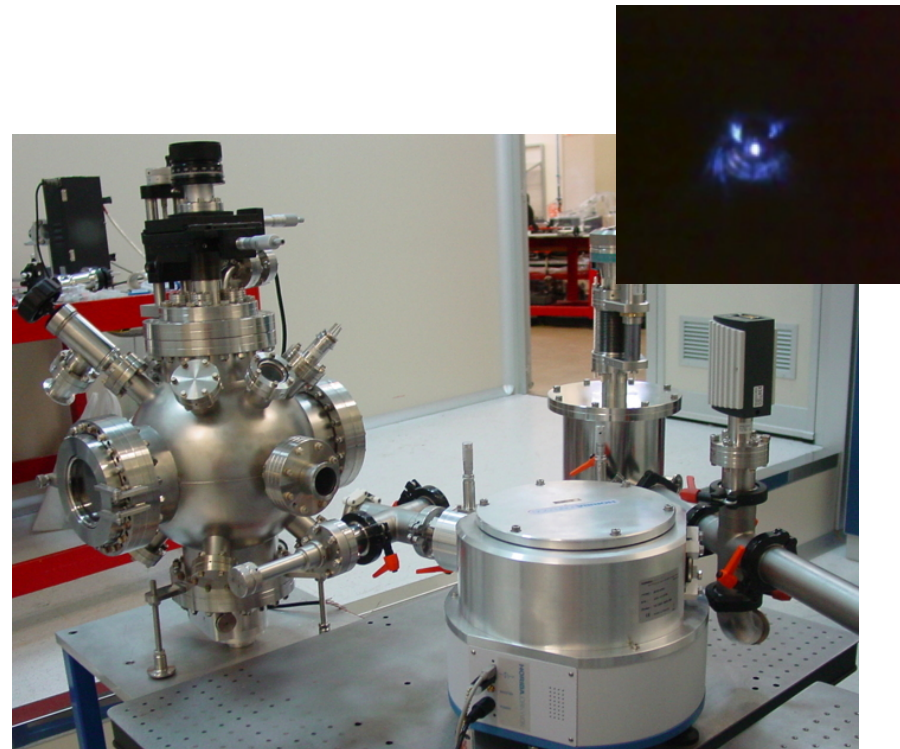
# UV-VIS new setup: first light

First light on 02.02.2011. The optical design of the branch line was revised to improve the expected performance of the whole system. The previous design was based on a focusing optical system producing 1 mm spot at 2 m distance from the last folding mirror at the entrance mirror camera. To improve aperture matching – thus spectral resolution and beam intensity – with the monochromators, the design was changed in a collimated beam from the folding mirror. The collimated beam is focused with the proper aperture at the entrance slit of each of the two monochromators and of the optical fiber feeding the white beam channel. A 35 mm collimated beam is now available.



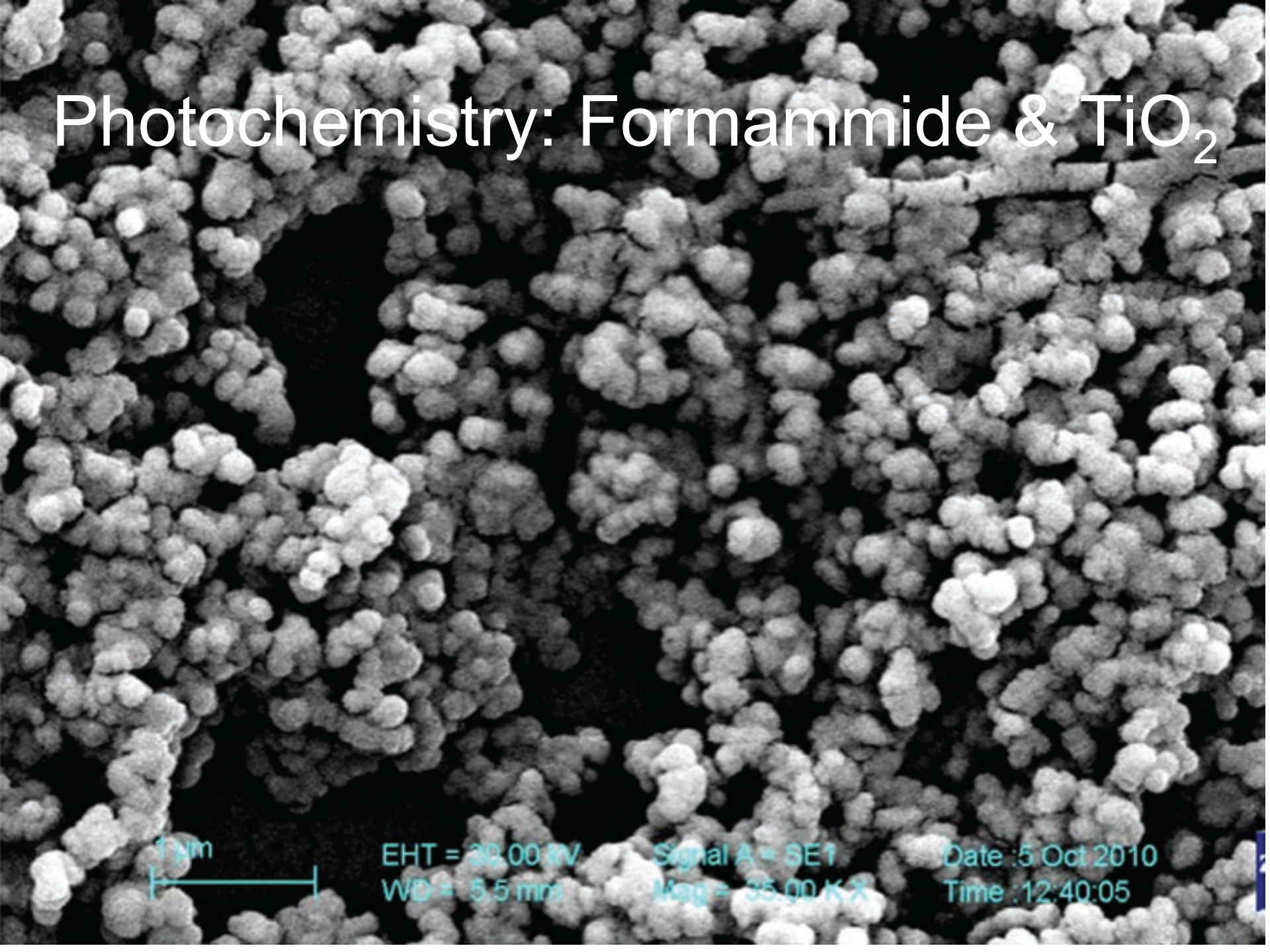
UV radiation transferred through solarized fiber optics

Instrumentation has been upgraded with a new VUV monochromator (UVXL200 by Jobin Yvon) operating in the 120-250 nm spectral range. It will be put on the SR beam line after focusing the collimated beam on the entrance slit with the appropriate aperture. The previous monochromator will be used coupled to a high power Deuterium lamp. This solution will improve the performance of the VUV SR channel and will offer the opportunity to double the VUV light sources available for experiments, such as UV irradiation, UV ageing, photobiology, etc.





# Photochemistry: Formamide & TiO<sub>2</sub>



1  $\mu$ m



EHT = 20.00 kV

WD = 5.5 mm

Signal A = SE1

Mag = 35.00 K X

Date : 5 Oct 2010

Time : 12:40:05

# Products



$$\lambda = (387.5 \pm 2.14 \text{ nm})$$

TABLE 2. YIELD OF BIOGENIC CARBOXYLIC ACIDS 5–12 IN THE PRESENCE OF LARGE (ENTRIES 1–8) AND SMALL (ENTRIES 9–16) ANATASE GRAINS

Entry	Product	Yield <sup>a</sup>
1	Glycolic acid 5	2.2
2	Lactic acid 6	8.0
3	Pyruvic acid 7	6.1
4	Succinic acid 8	0.2
5	Malic acid 9	0.5
6	Oxaloacetic acid 10	n.d.
7	$\alpha$ -Ketoglutaric acid 11	6.8
8	Fumaric acid 12	0.05
9 <sup>b</sup>	Glycolic acid 5	2.5
10 <sup>b</sup>	Lactic acid 6	8.3
11 <sup>b</sup>	Pyruvic acid 7	6.2
12 <sup>b</sup>	Succinic acid 8	0.1
13 <sup>b</sup>	Malic acid 9	0.8
14 <sup>b</sup>	Oxaloacetic acid 10	5.1
15 <sup>b</sup>	$\alpha$ -Ketoglutaric acid 11	6.9
16 <sup>b</sup>	Fumaric acid 12	0.1

<sup>a</sup>Product yield is expressed as milligrams of product with respect to initial amount in grams of formamide.

<sup>b</sup>Data obtained with TiO<sub>2</sub> anatase ground in an agate mortar.  
n.d., not determined.

TABLE 3. YIELD OF CARBOXYLIC ACIDS 13–19 AND HETEROCYCLES 20–21 IN THE PRESENCE OF LARGE (ENTRIES 1–9) AND SMALL (ENTRIES 9–18) ANATASE GRAINS

Entry	Product	Yield <sup>a</sup>
1	Ethanimidic acid 13	2.42
2	Propanoic acid 14	2.11
3	2,3-Dihydroxypropanoic acid 15	0.41
4	2,3-Dihydroxybutanoic acid 16	n.d.
5	2,4-Dihydroxybutanoic acid 17	n.d.
6	Oxalic acid 18	1.54
7	Malonic acid 19	2.46
8	3-Hydroxypyridine 20	0.02
9	4(3H)-Pyrimidinone 21	0.06
10 <sup>b</sup>	Ethanimidic acid 13	n.d.
11 <sup>b</sup>	Propanoic acid 14	n.d.
12 <sup>b</sup>	2,3-Dihydroxypropanoic acid 15	1.10
13 <sup>b</sup>	2,3-Dihydroxybutanoic acid 16	0.64
14 <sup>b</sup>	3,4-Dihydroxybutanoic acid 17	0.35
15 <sup>b</sup>	Oxalic acid 18	n.d.
16 <sup>b</sup>	Malonic acid 19	n.d.
17 <sup>b</sup>	3-Hydroxypyridine 20	0.07
18 <sup>b</sup>	4(3H)-Pyrimidinone 21	0.10

<sup>a</sup>Product yield is expressed as milligrams of product with respect to initial amount in grams of formamide.

<sup>b</sup>Data obtained with TiO<sub>2</sub> anatase ground in an agate mortar.  
n.d., not determined.



# UV irradiation of aminoacids

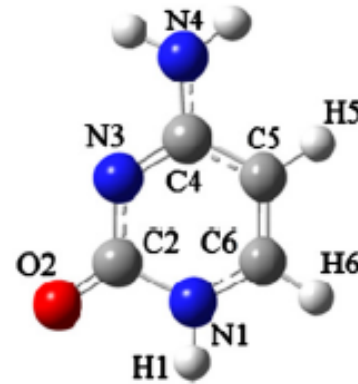
Analysis: FTIR spectroscopy



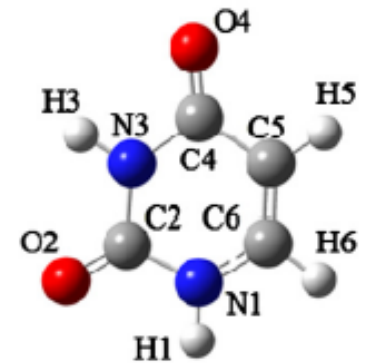
**Adenine**



**Hypoxanthine**



**Cytosine**



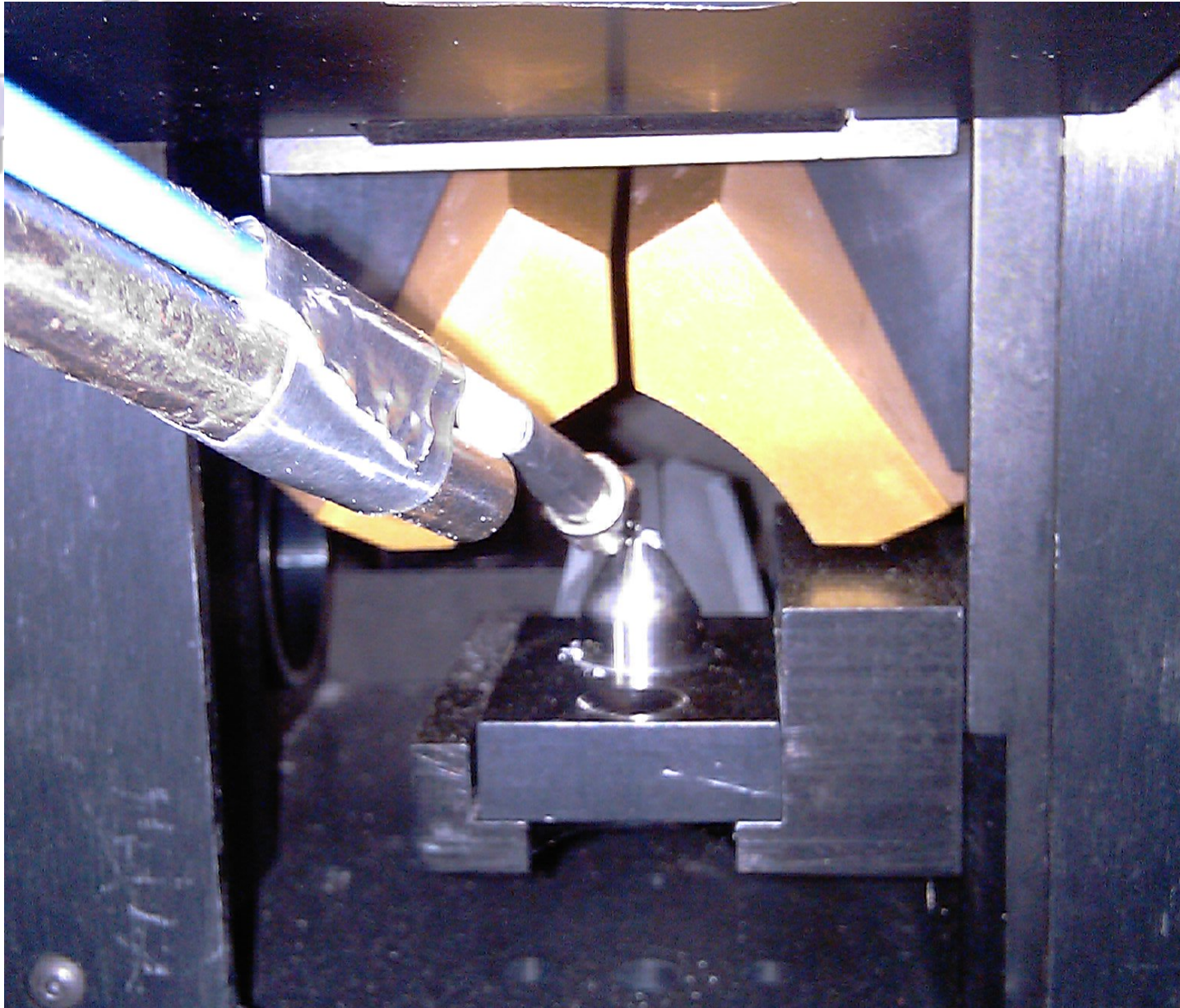
**Uracil**



# Exp Apparatus







INFI

# FTIR spectra

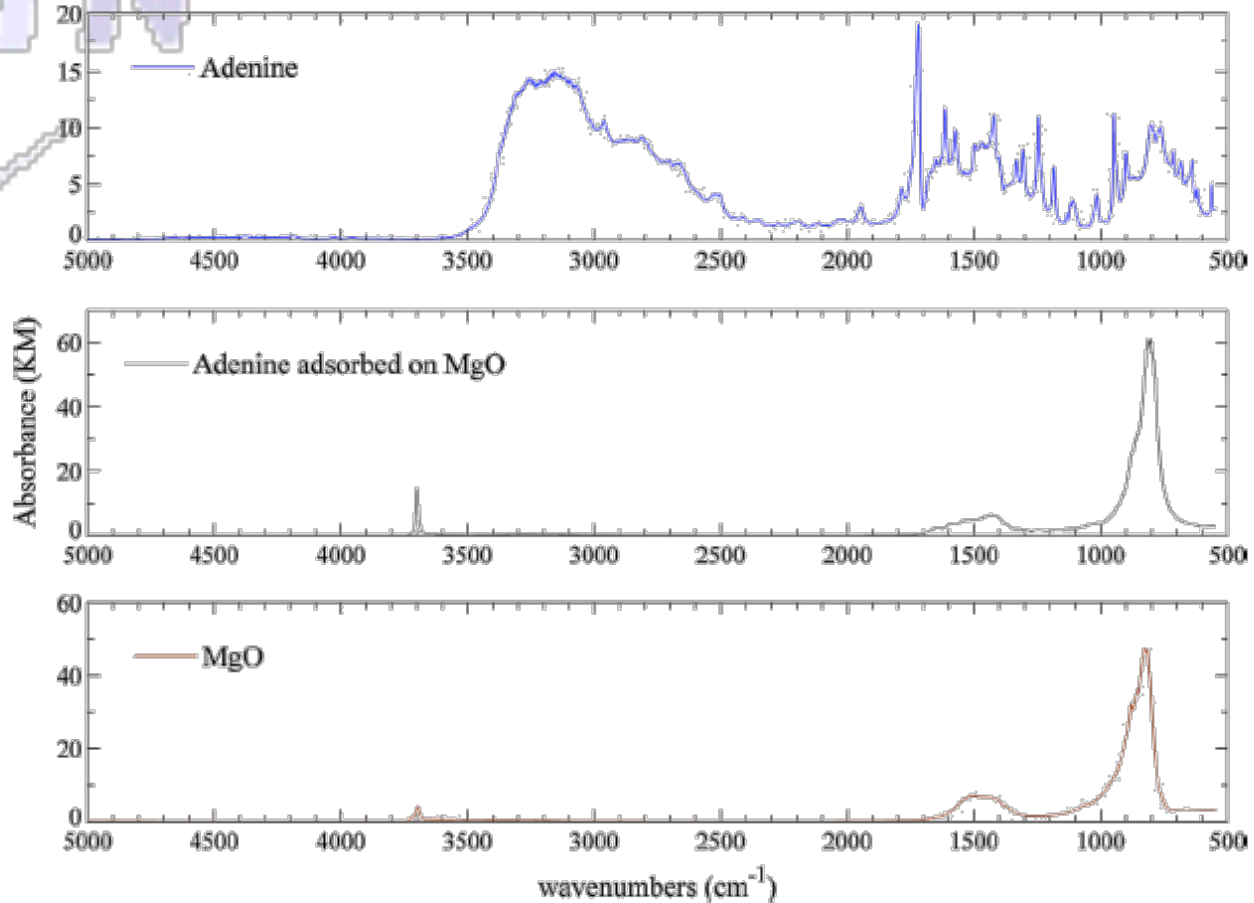


Fig. 2. IR spectra of pure adenine, adenine adsorbed on MgO and pure MgO, represented in Kubelka–Munk units (KM) vs wavenumbers ( $\text{cm}^{-1}$ ).

# FTIR spectra

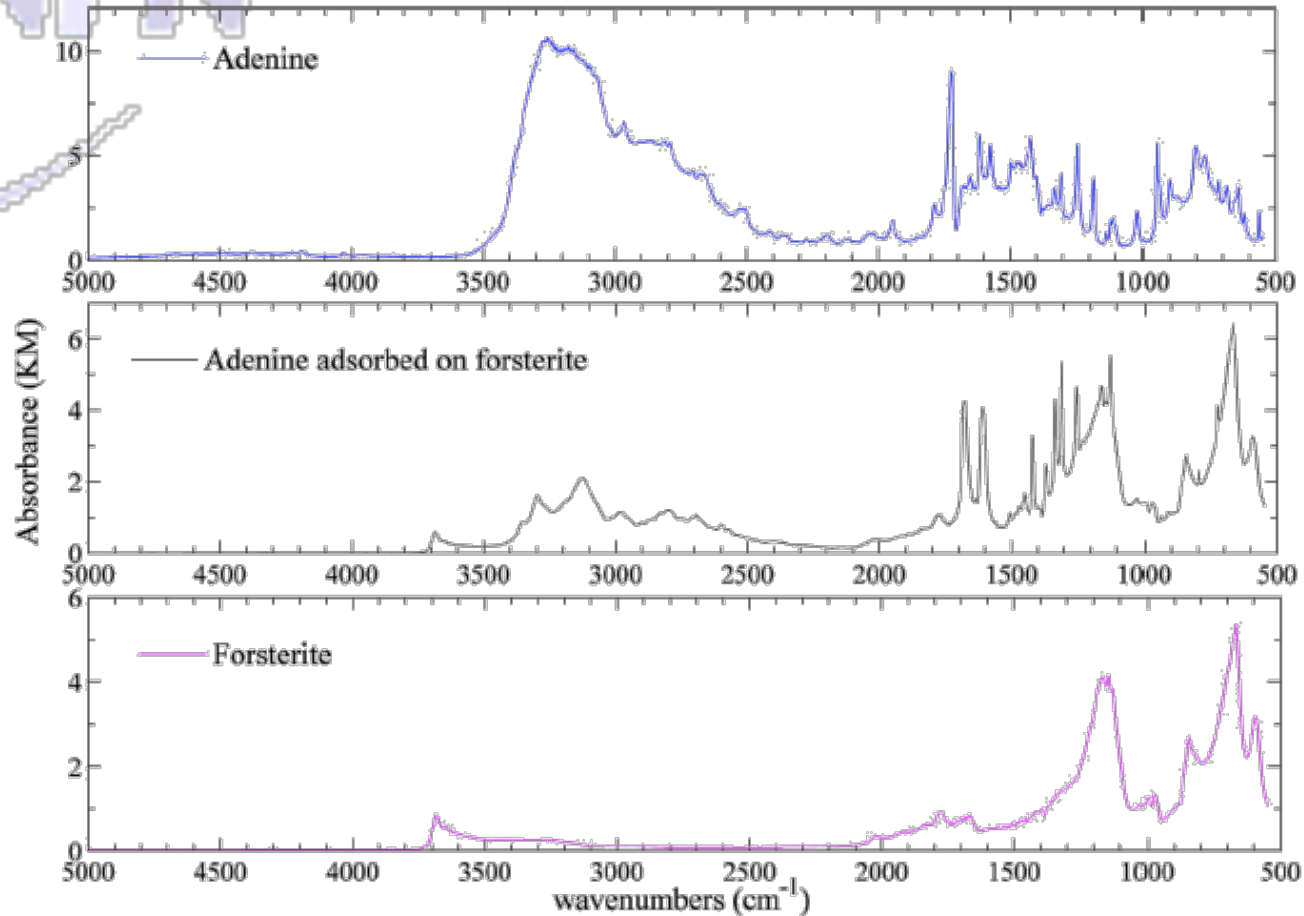
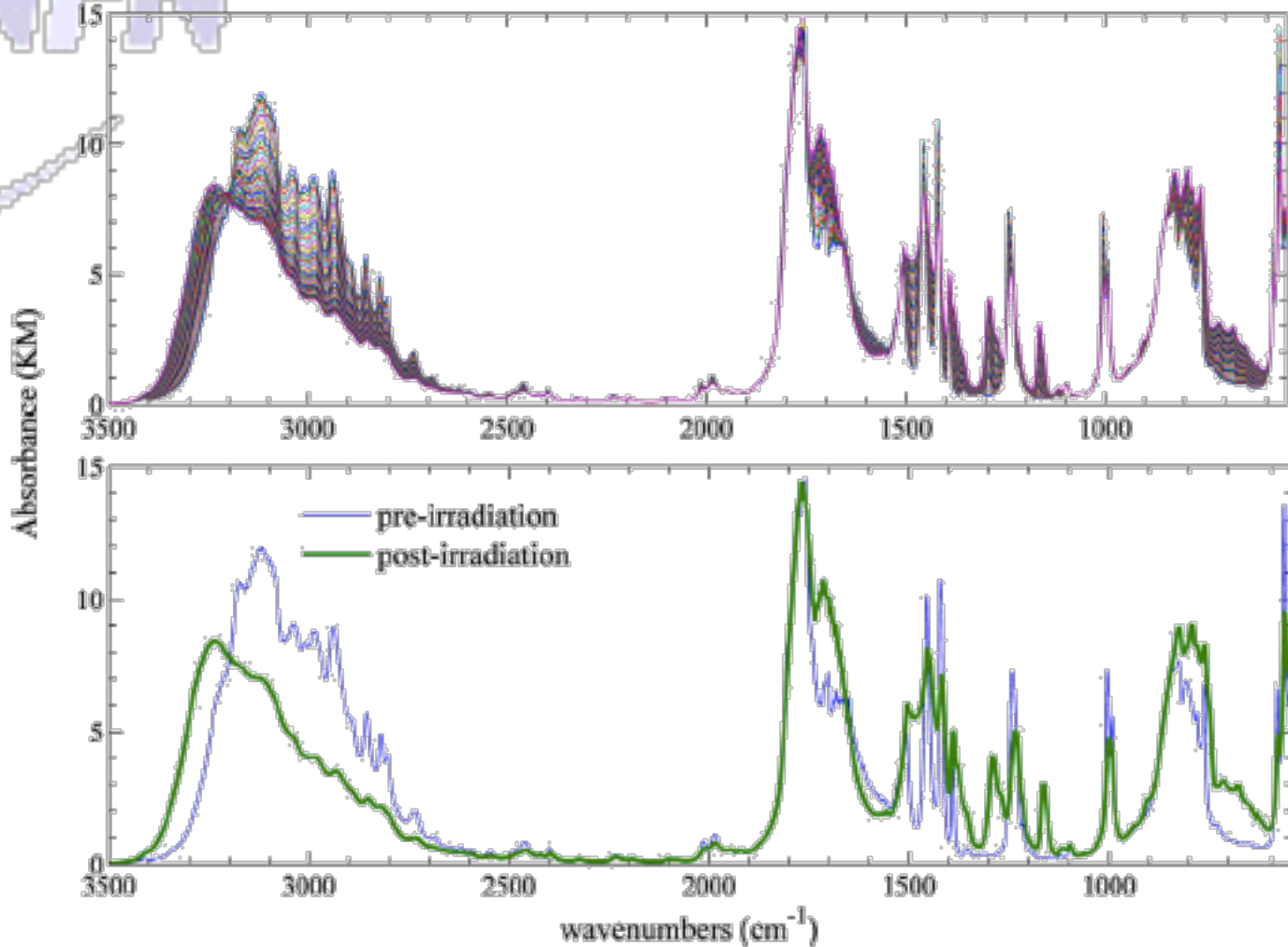


Fig. 3. IR spectra of pure adenine, adenine adsorbed on forsterite and pure forsterite, represented in Kubelka–Munk units (KM) vs wavenumbers ( $\text{cm}^{-1}$ ).

# FTIR spectra





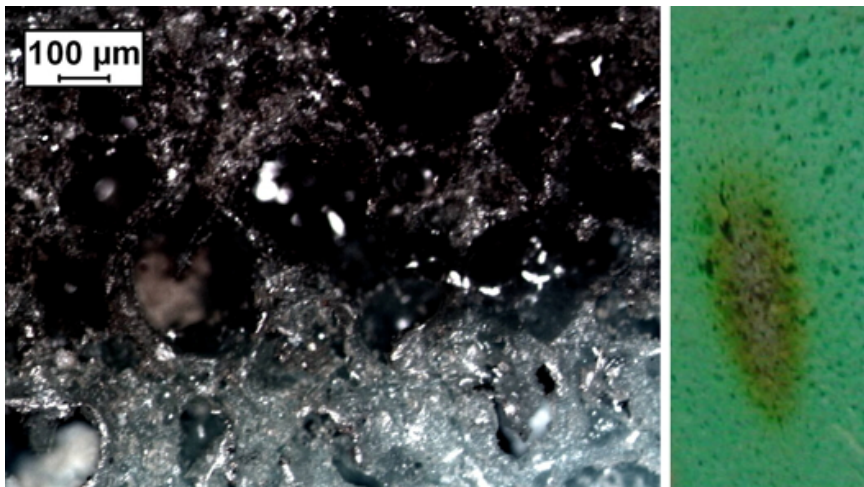
# Ageing of materials and coatings

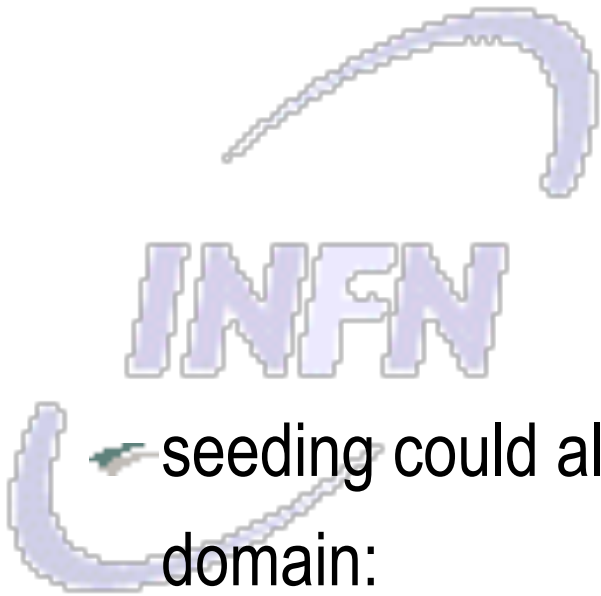
## Ageing of FOAM

FOAM is a memory form material that has been developed for space application. Memory form materials are used to recover the original form of the material after being properly modified for the application; so, it can be stretched, pressed, turned and it can recover the original form by applying thermal energy warming it at around 100°C. UV ageing is one of the main issues, because there is a lot of UV radiation in outer space and this material might undergo degradation after prolonged exposure to UV radiation.

Samples of this material were irradiated for several hours (from 1 up to 30 hours) with a  $10^{13}$  ph/s white beam. The samples were put firstly on the focus and then on the defocused beam to avoid any effect of thermal degradation. After irradiation the samples show brownish regions corresponding to the beam. Irradiated materials were characterized by the Differential Scanning Calorimetry (DSC) technique that shows no variation highlighting that UV ageing effects had no effects on in the amount of heat required to increase the sample temperature. To assess this result the sample was irradiated with IR radiation to have a direct comparison between photo ageing and thermal ageing.

Irradiated FOAM sample after 8 hours at the focus (right) as it appears at a zoomed observation (left)





# FEL VUV Sorgente

➤ seeding could allow to work with high harmonics in the UV domain: **250-90(~ 40!) nm @ 155 MeV**

➤ repetition rate: **1-10 Hz**

➤ radiation spectrum slightly **tunable**

➤ higher harmonics **linearly polarized** sources of high temporal and spatial coherence

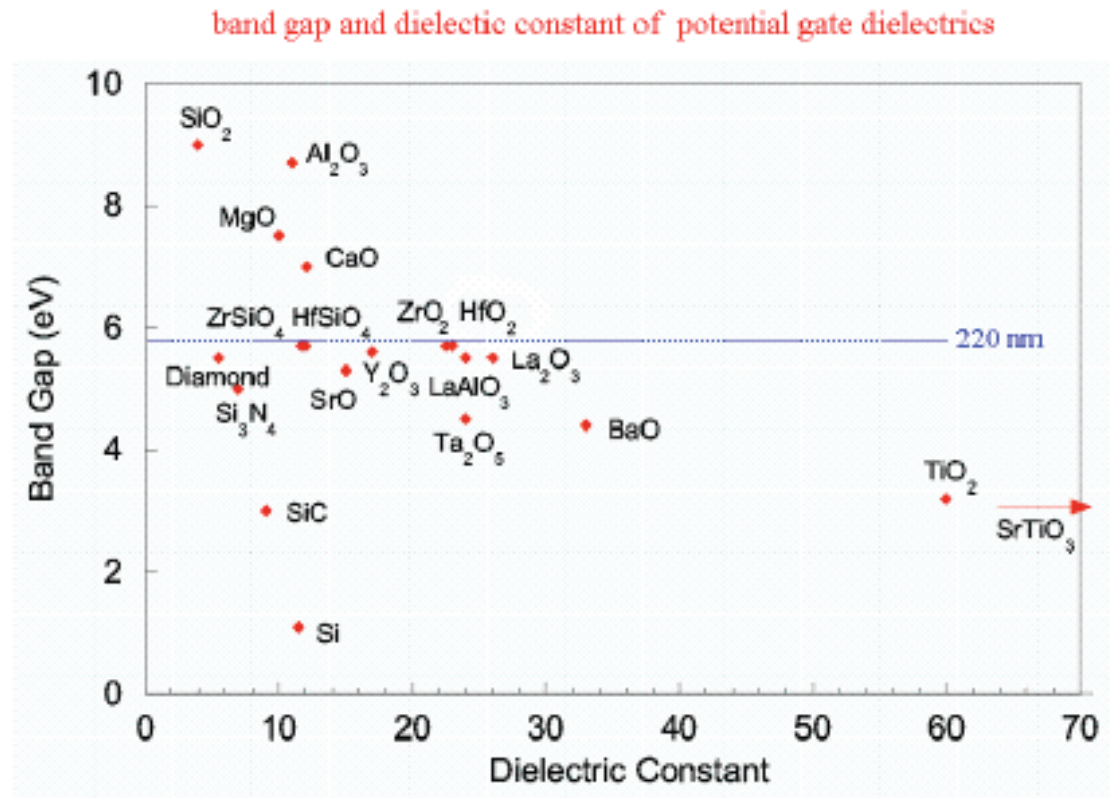
➤ pulse duration: **100 fs**

# FEL VUV Applicazioni

## Raman & PL spectroscopy on isolated building blocks of nanostructured films

**Raman spectroscopy** is among most popular methods for characterization of nano-structured systems. The ability of the VUV-FEL to shift the wavelength of scattered light from visible into deep UV will allow to probe new electronic transitions well within the **7-10 eV** range for classes of cluster materials such as nano-carbons and potential gap dielectrics from metal oxides

Raman Imaging...!



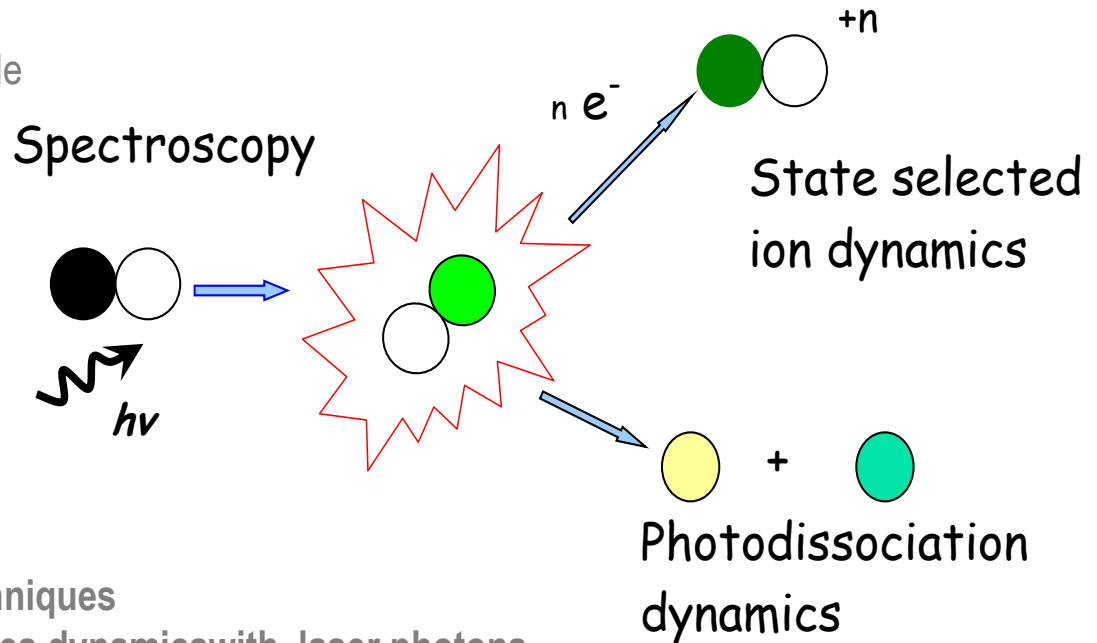
# FEL & Atomic and Molecular Physics

## SPECTROSCOPY OF “EXOTIC SPECIES”

High flux and Resolution allow thorough investigations of the electronic structure of low density matter: metastable and reactive species, i.e., high temperature species, atmospheric chemistry, combustion chemistry, isolated Molecules and Clusters in Supersonic Molecular Beam (condensation and nucleation studies, coupling to laser techniques for state selected excitation and photoemission studies)

## MULTIPHOTON PROCESSES

Access to electronic states, transparent to single photon techniques ruled by dipole interaction



## EXPERIMENTAL ADVANTAGES

Simple use of TOF and time-correlation techniques  
“Pump & Probe” techniques for excited states dynamics with laser photons

Direct measurement of lifetime; accurate analysis of intra- and inter-molecular energy transfer; interaction of photo-emitted electrons from gaseous samples with laser; electromagnetic field exploited at the same time as a diagnostic for FEL pulses (“side bands experiments”) and for the investigation of photo-ionization dynamics

# Electronic properties of biomolecules and their clusters

Solvent-free environment :

- easier modeling of different intermolecular forces
- results comparable with theoretical calculations

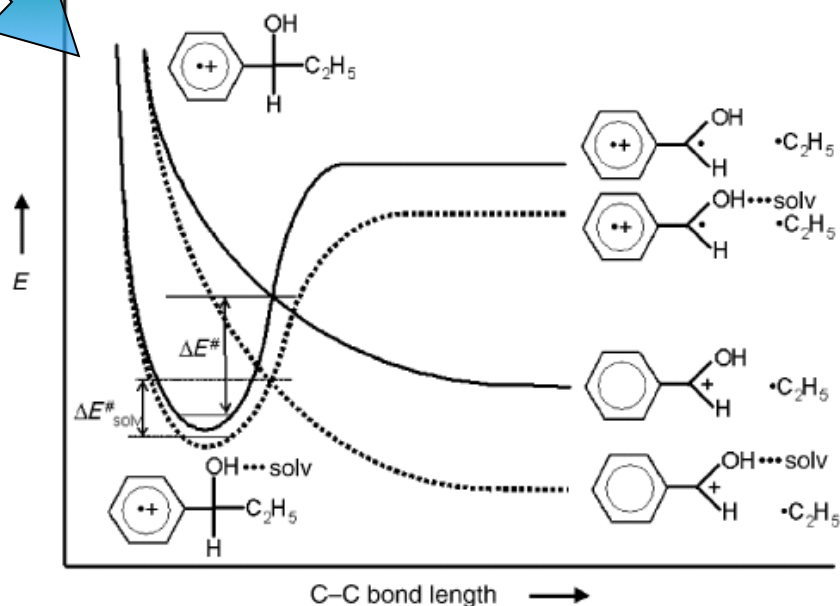
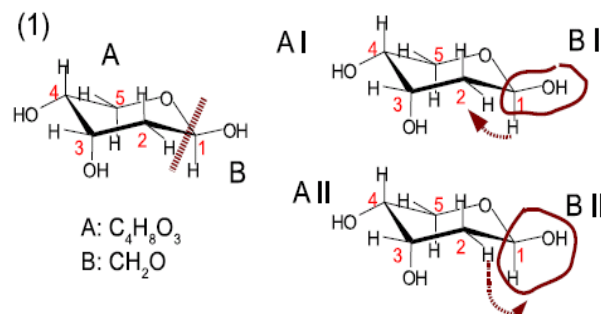
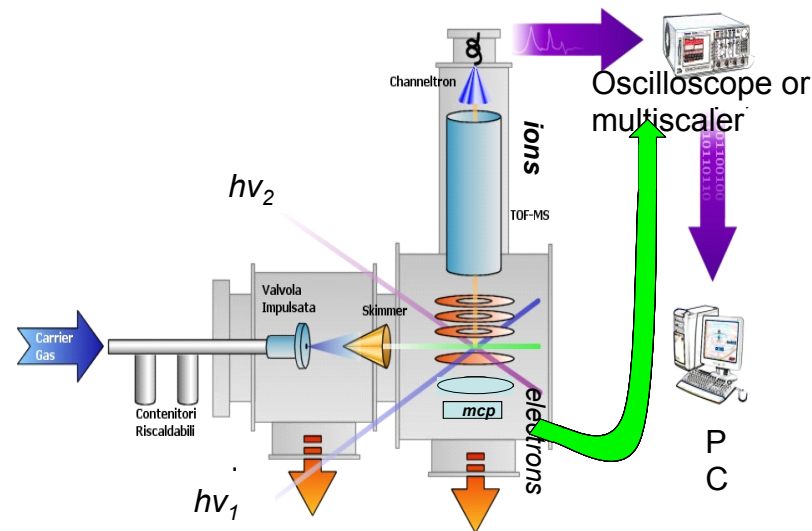


Figure 1. Pictorial crossing between the two lowest-energy electronic states of  $[(BZC_2H_5)_R]^+$  (full lines) and  $[(BZC_2H_5)_R]^+ \cdot \text{solv}$  (broken lines).

TOF set-up @ CNR-IMIP & University Tor Vergata



variable polarization and  $\lambda < 150 \text{ nm}$  from SPARC will extend investigation to s-bonded systems like sugars and non aromatic aminoacids and implement valence photoemission techniques for the study of chiral recognition



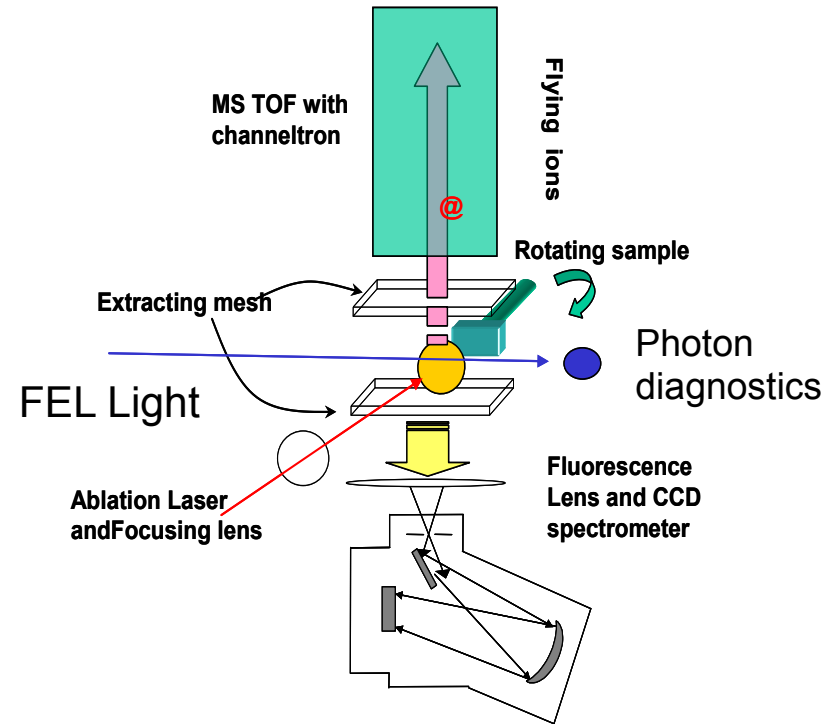
# Surface modification and ablation studies

Stefano Orlando (CNR-IMIP, Rome) & Jesus Alvarez Ruiz (CSIC, Madrid)

Many important elementary processes such as electron/hole recombination, excitation relaxation, etc., often occur on a much shorter time scale and only a time resolved spectroscopy is able to elucidate the dynamics of charge and energy transfer processes. Ultra-short laser pulses limit the secondary ionization and photo-fragmentation, and exclude the laser/plume interaction. Therefore, only within such excitation regime, time- and space-resolved optical spectroscopy of the generated plasma provides a direct investigation of laser-target interaction and ultimately of particle emission.

Photoemission and diffraction studies of the ablated species condensed on a substrate as a function of the laser pulses will be used for the analysis of the ablation products and optimization of the process with a view to application to thin film deposition.

# FEL VUV Applicazioni



**1.2 ps laser pulses**    **200 fs laser pulses**

