

# Neutrino Interactions with Nucleons and Nuclei

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# Motivation and Contents

- Determination of neutrino oscillation parameters requires knowledge of neutrino energy
- Modern experiments use complicated nuclear targets: from Carbon to Argon
- Nuclear effects affect everything:
  - event identification
  - final state particles
  - reconstructed neutrino energy
  - event cross section measurements
  - neutrino oscillation parameters

# Neutrino Oscillations

- 2-Flavor Oscillation:

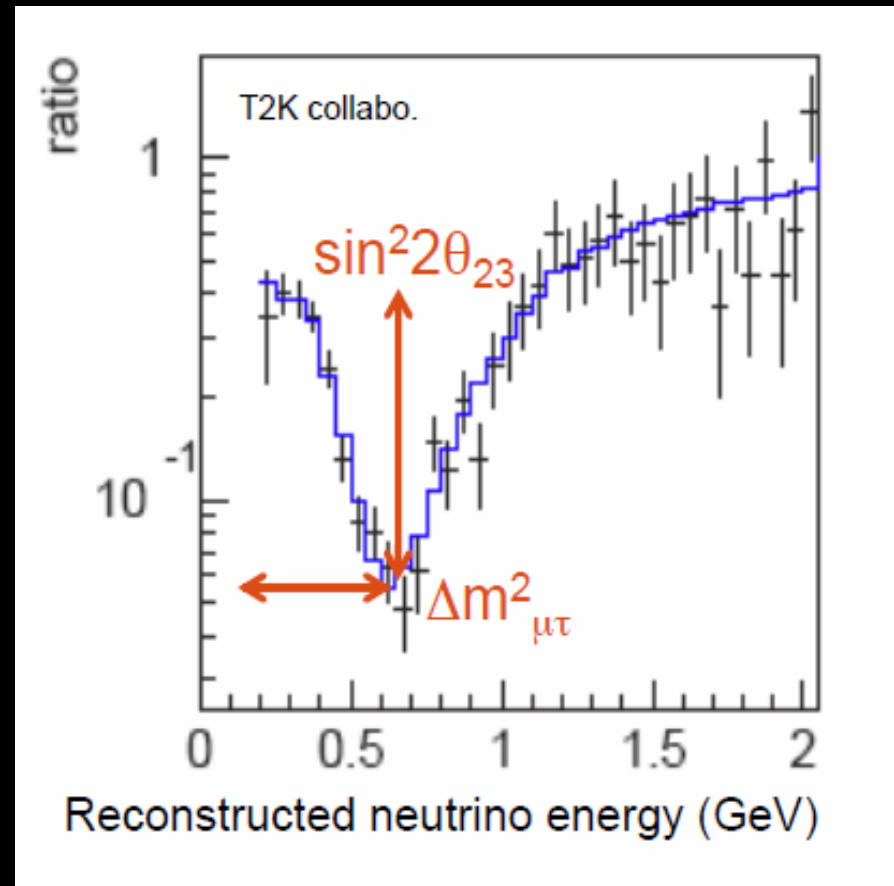
$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E_\nu} \right)$$

Know:  $L$ , need  $E_\nu$  to determine  $\Delta m^2, \theta$

- 3-Flavor Oscillation: allows for CP violation

# Observable Oscillation Parameters

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E_\nu} \right)$$



# Oscillation probability

## Long-Baseline Accelerator Appearance Experiments

- Oscillation probability complicated and dependent not only on  $\theta_{13}$  but also:

1. CP violation parameter ( $\delta$ )
2. Mass hierarchy (sign of  $\Delta m_{31}^2$ )
3. Size of  $\sin^2\theta_{23}$

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \frac{\Delta m_{31}^2 L}{4E} \times \left( 1 + \frac{2a}{\Delta m_{31}^2} (1 - 2S_{13}^2) \right) \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \sin \frac{\Delta m_{21}^2 L}{4E} \\
 & - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \sin \frac{\Delta m_{21}^2 L}{4E} \\
 & + 4S_{12}^2 C_{13}^2 \{ C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta \} \sin^2 \frac{\Delta m_{21}^2 L}{4E} \\
 & - 8C_{13}^2 S_{13}^2 S_{23}^2 \cos \frac{\Delta m_{32}^2 L}{4E} \sin \frac{\Delta m_{31}^2 L}{4E} \frac{aL}{4E} (1 - 2S_{13}^2)
 \end{aligned}$$

***⇒ These extra dependencies are both a “curse” and a “blessing”***

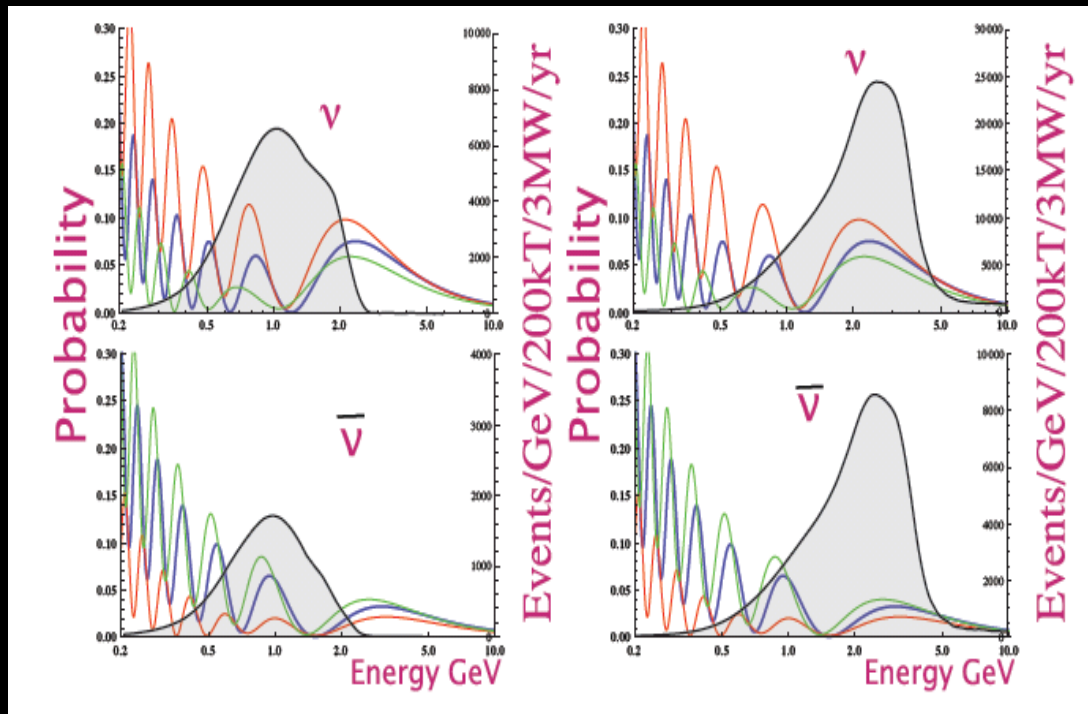
## Reactor Disappearance Experiments

- Reactor disappearance measurements provide a straight forward method to measure  $\theta_{13}$  with no dependence on matter effects and CP violation

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{13}^2 L}{4E} + \textit{small terms}$$

# LBNE, $\delta_{CP}$ Sensitivity

From: Bishai et al., hep-ex 12034090



8 GeV

60 GeV

proton energy

From:  
Bishai et al  
arXiv:1203.409

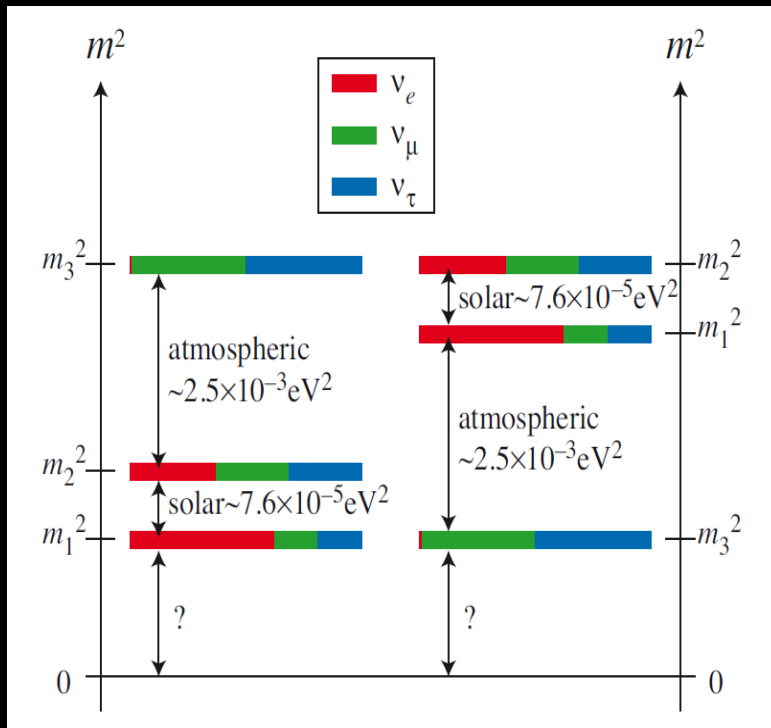
$$\delta_{CP} = 0$$

$$\delta_{CP} = \pi/2$$

$$\delta_{CP} = -\pi/2$$

Need energy to distinguish between different  $\delta_{CP}$

# Oscillation Signal Dependence on Hierarchy and Mixing Angle



Energy has to be known better than 50 MeV

Shape sensitive to hierarchy and sign of mixing angle

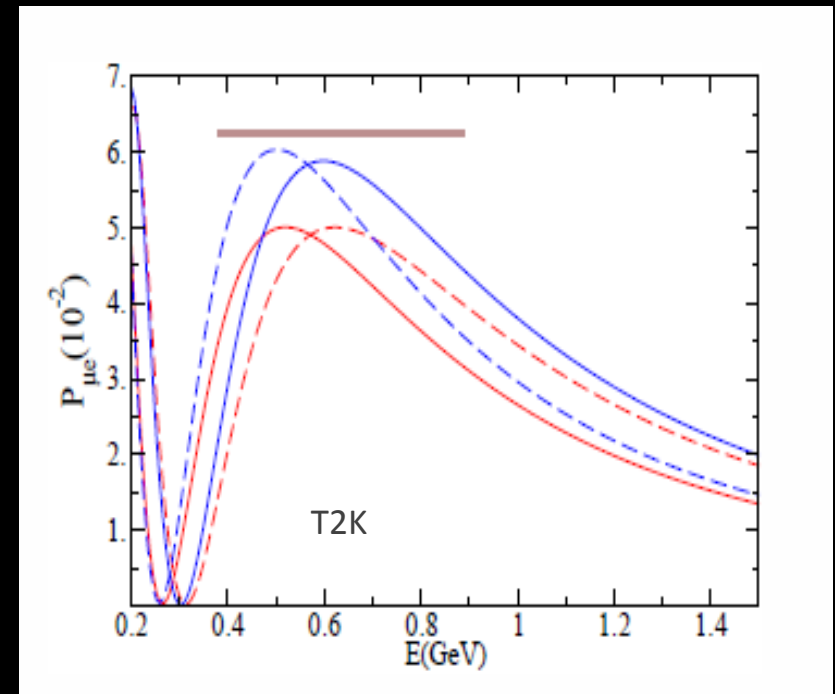


Fig. 2.  $\mathcal{P}_{\mu e}$  in matter versus neutrino energy for the T2K experiment. The blue curves depict the normal hierarchy, red the inverse hierarchy. Solid curves depict positive  $\theta_{13}$ , dashed curves negative  $\theta_{13}$

D.J. Ernst et al., arXiv:1303.4790 [nucl-th]

# Appearance experiment

- Near detector:

- Neutrino Flux
- Background
- Intrinsic  $\nu_e$
- Neutrino energy



- Far detector:

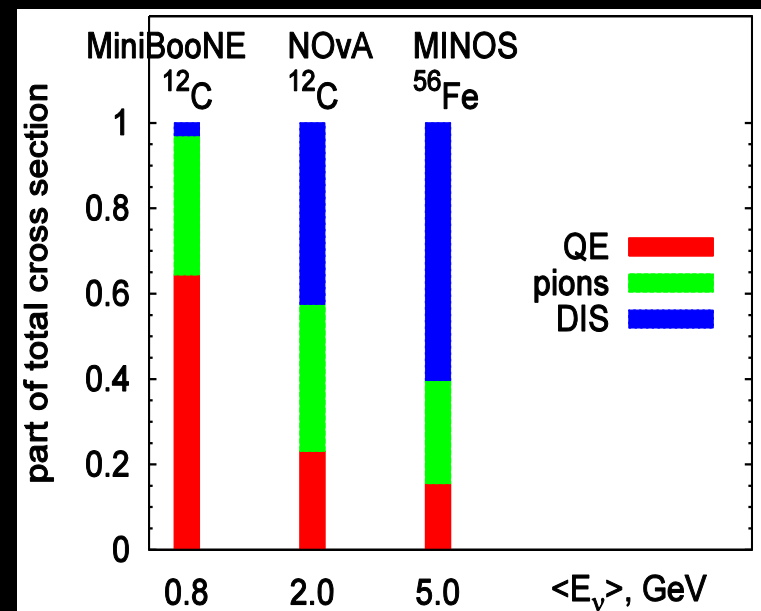
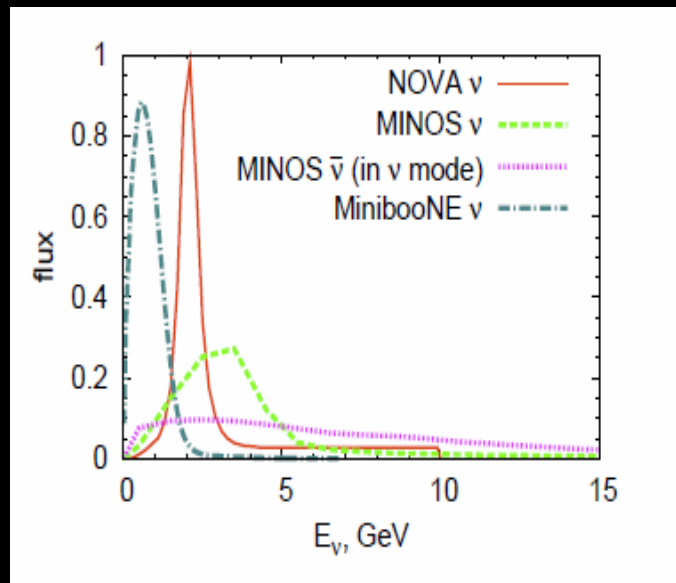
- Extrapolate Flux
- Background
- Neutrino energy

$$P(\nu_\mu \rightarrow \nu_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{1.27 \Delta m^2 L}{E_\nu} \right) + \text{other}$$



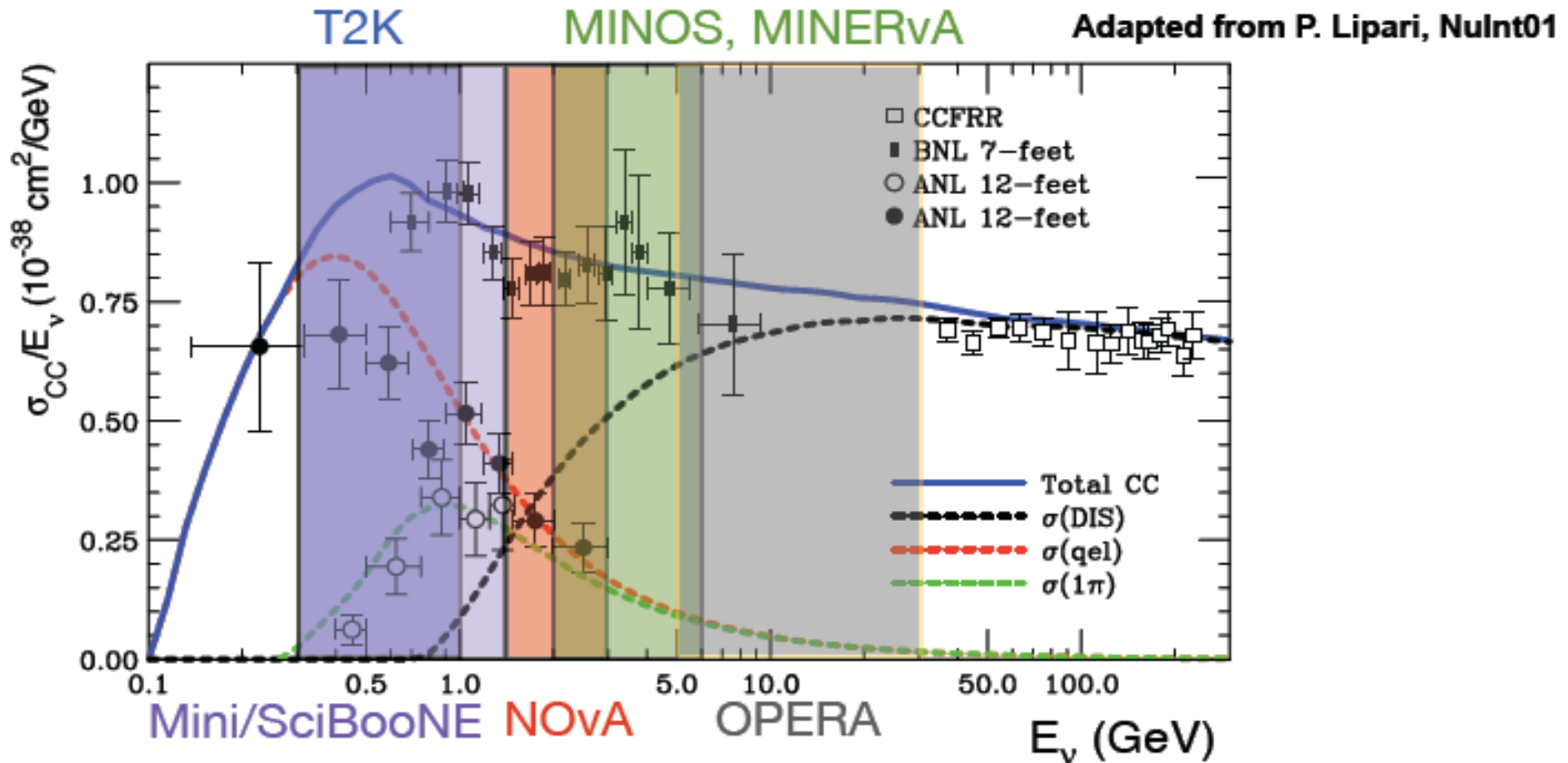
# Neutrino Beams

- Neutrinos do not have fixed energy nor just one reaction mechanism



Have to reconstruct energy from final state of reaction  
Different processes are entangled

# Neutrino Cross-Sections



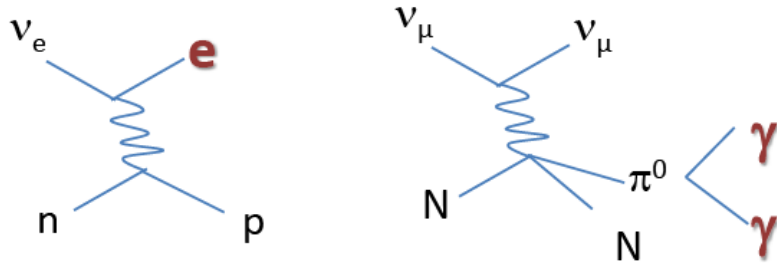
Upcoming experiments will continue to work in an “interesting” region:

- Large contributions from QE, Resonances and DIS regions
- Are these categories even sufficient ?

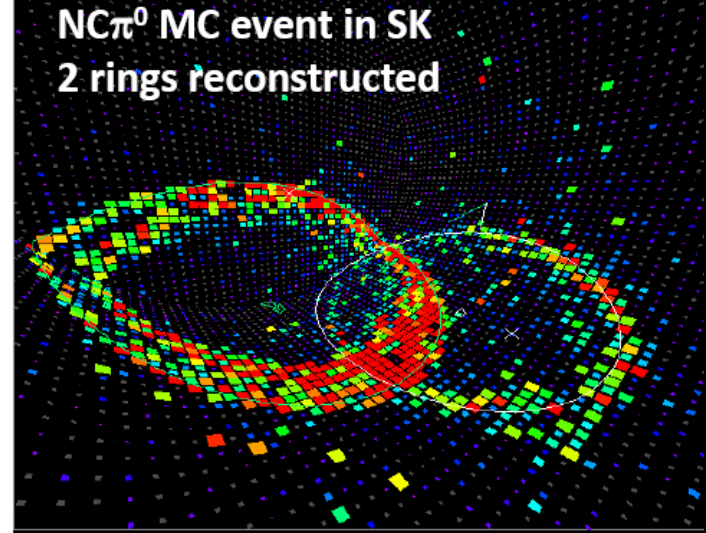
# Neutrino Interactions

## •for $\nu_e$ appearance

- beam  $\nu_e$
- NC $\pi^0$  events

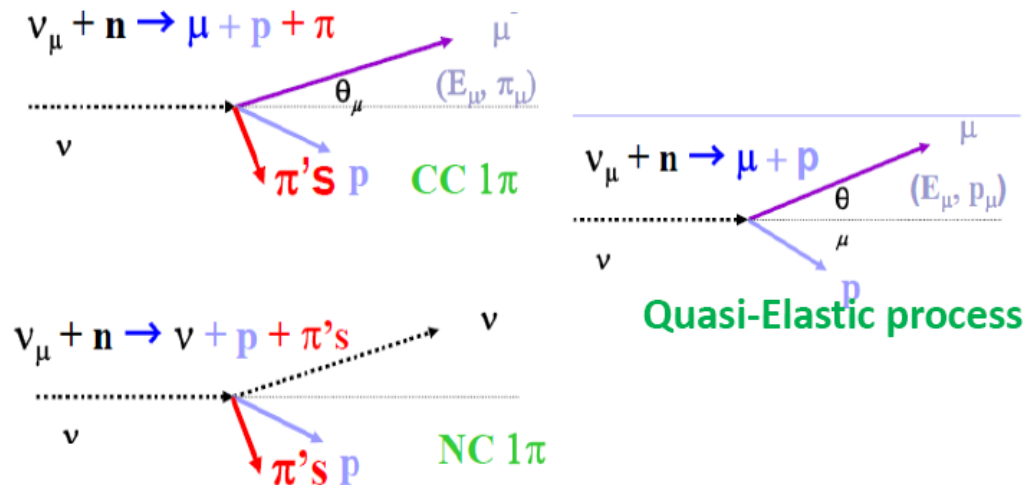


NC $\pi^0$  MC event in SK  
2 rings reconstructed

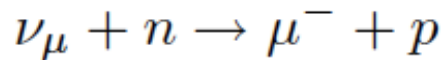


## •for $\nu_\mu$ disappearance (muon energy measurement)

- inelastic processes



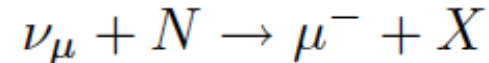
# Energy reconstruction



$$E_{\nu} = E_{\nu}(E_{\mu}, \theta_{\mu})$$

## Kinematic:

- Rely on underlying interaction to use relate outgoing lepton kinematics to neutrino energy
- Advantage:
  - don't need hadron reconstruction
- Disadvantages
  - energy is wrong if underlying interaction is wrong (i.e. not CCQE)
  - Nuclear effects smear resolution

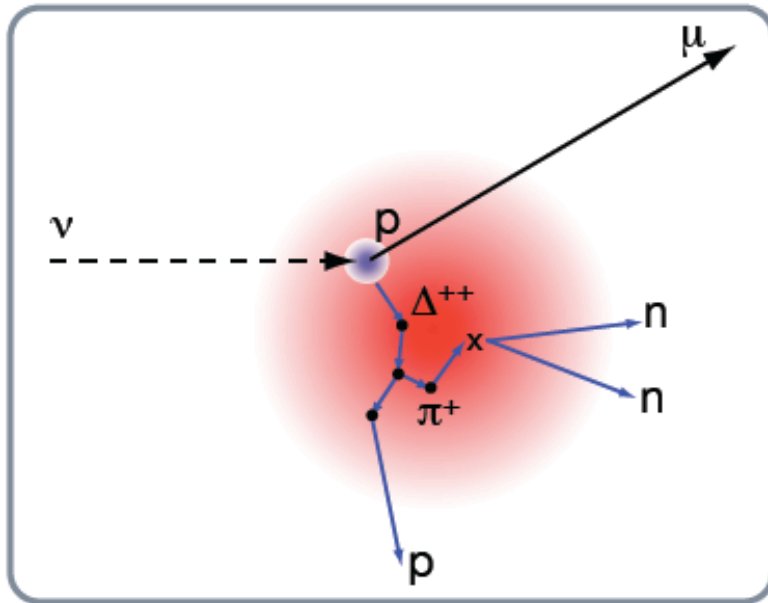


$$E_{\nu} = E_{\mu} + E_X$$

## Calorimetric

- Add up the energy from the leptonic and hadronic components
- Advantages
  - No *a priori* assumption about underlying interaction
- Disadvantages
  - Relies on hadron reconstruction

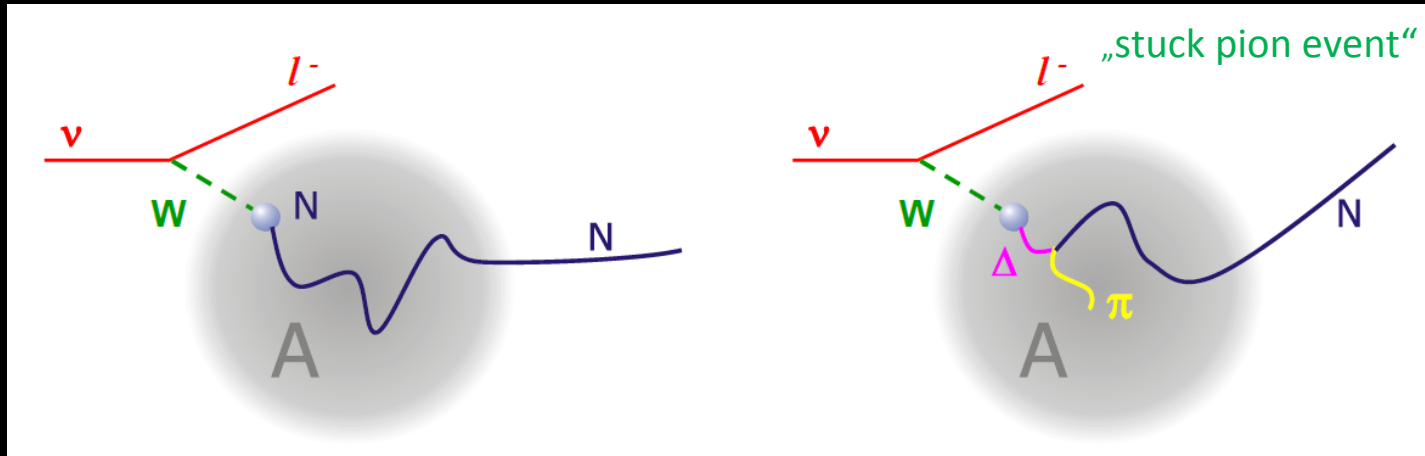
# Background: Nuclear re-interactions



- Lepton kinematics shifted/smeared
- Outgoing hadronic final state (“topology”) may differ from expectation from “underlying”  $\nu$ -nucleon interaction
- FSI effects may appear degenerate with hadronic interactions outside of the target nucleus.

## Modeling $\nu$ interactions in nucleus

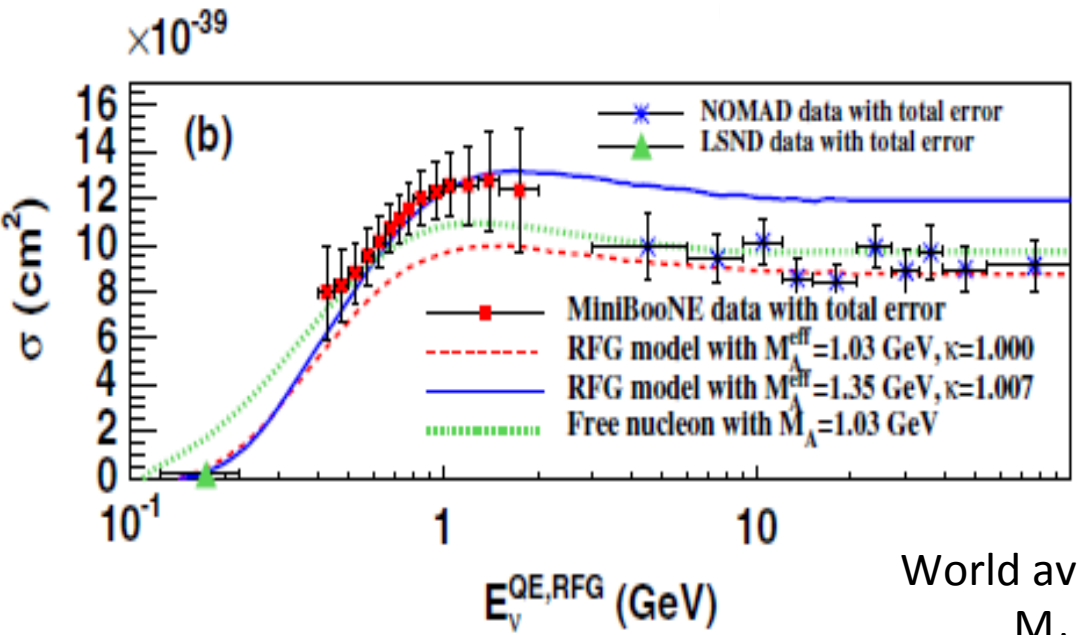
- Underlying  $\nu$ -nucleon/quark interaction
  - Mode (CCQE, resonance, etc.)
  - Determine “final” state of interaction
- Initial state nucleon/quark
  - Fermi motion, binding energy
- Final state effects
  - Pauli blocking
  - Propagate hadrons within nucleus
    - Absorption, scattering, CEX, etc.



Complication to identify QE, entangled with  $\pi$  production

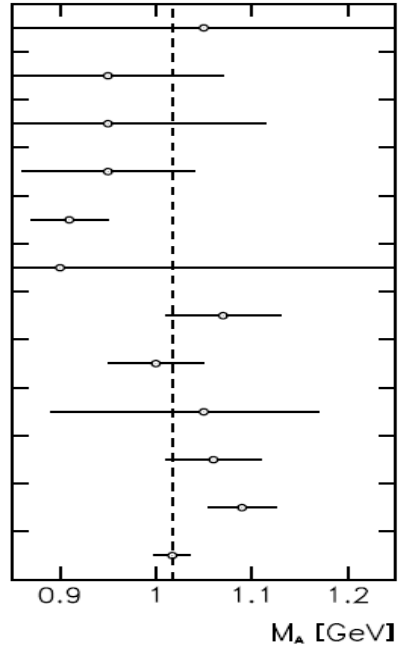
Nuclear Targets (K2K, MiniBooNE, T2K, MINOS, Minerva, ...)

# MiniBooNE QE puzzle



World average axial mass:  
 $M_A = 1.03$  GeV

- Argonne (1969)
- Argonne (1973)
- CERN (1977)
- Argonne (1977)
- CERN (1979)
- BNL (1980)
- BNL (1981)
- Argonne (1982)
- Fermilab (1983)
- BNL (1986)
- BNL (1987)



MiniBooNE use mineral oil (Cerenkov rings): identifies QE by muon and zero pion, corrects for 'stuck pions'

Can the nuclear effects be responsible for a higher axial mass value ?

# New nuclear model in neutrino interaction generators



# GENIE 2.8.0 + $\nu T$

- Based on GENIE 2.8.0 plus few packages developed at VT with inputs from other experts: C. Jen, A. Ankowski, O. Benhar, L. Kalousis, CM
- Implementation of SF following nuclear theory recipe. Implemented for C, O, Ca and Ar.
- Very good comparison with electron scattering data, few examples in next slides.
- Corrected the  $Q^2$  selection in GENIE 2.8.0, now the code conserve both energy and momentum.

# Spectral function

- RFG model

$$P(|\vec{p}|, E) = \frac{3}{4\pi k_F^3} \theta(k_F - |\vec{p}|) \delta(\sqrt{m^2 + |\vec{p}|^2} - m - E_B + E)$$

step function (no structure) energy-momentum conservation

- SF model (1p-1h : the simplest case)

$$P(|\vec{p}|, E) = \left| \langle {}^{40}\text{Ar} | {}^{39}\text{Ar}, p \rangle \right|^2 \delta(E_{{}^{40}\text{Ar}} - E_{{}^{39}\text{Ar}} - m_n + E)$$

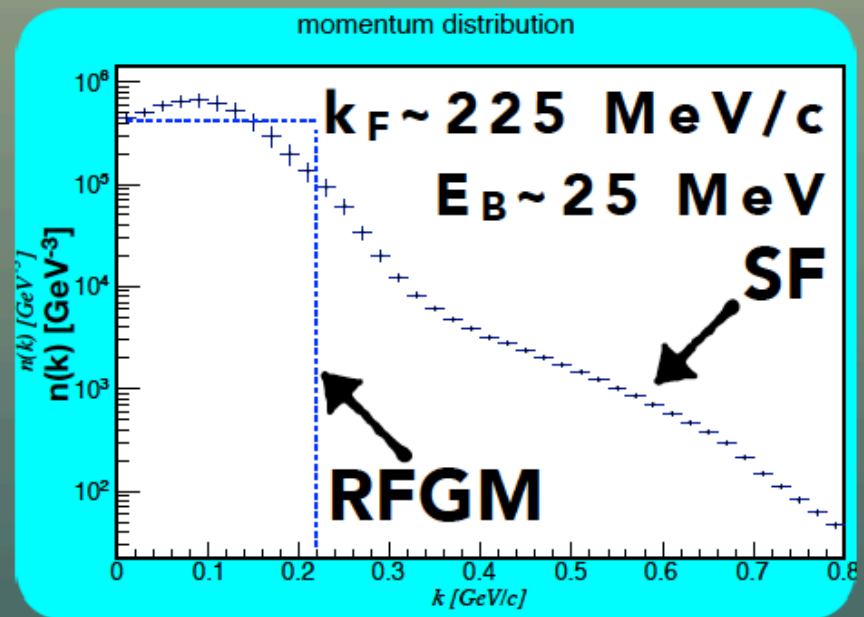
probability amplitude of knocking out one nucleon energy-momentum conservation

**Math Form of Spectral Functions**  
telling you the “probability” of knocking out one nucleon  
with one specific quantum number

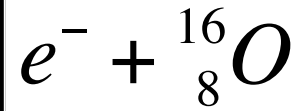
# Structure function from electron scattering data

- initial-state nucleon's momentum distribution,  $n_n$ , is assumed to be flat (no structure) for all shells below the Fermi sea ( $k_F$ )
- the spectral function theory considers the difference in the structure function for each shell and thus is capable of yielding a much better prediction while being compared to the measured electron cross-section than RFGM

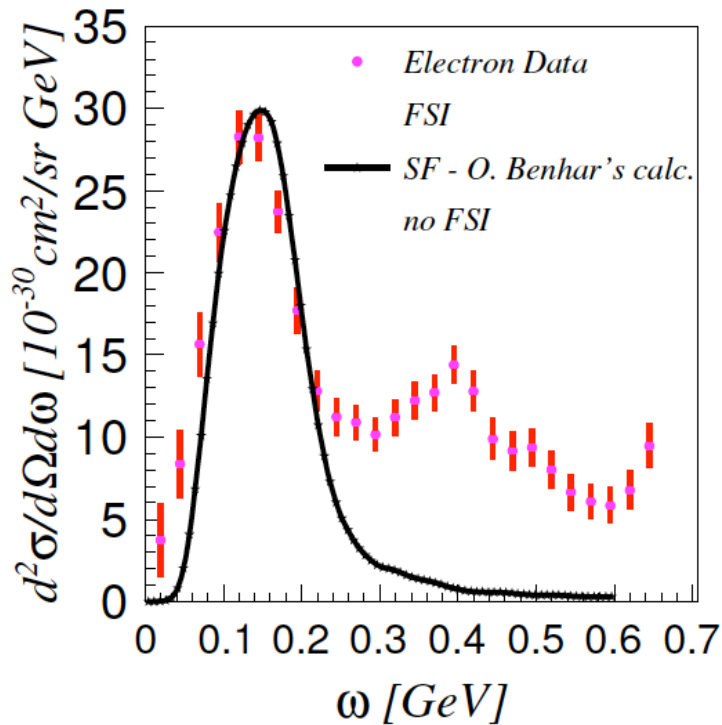
$$n^D = \sigma_{(e,e'p)} / K \sigma_{cc1} \quad \text{DWIA}$$



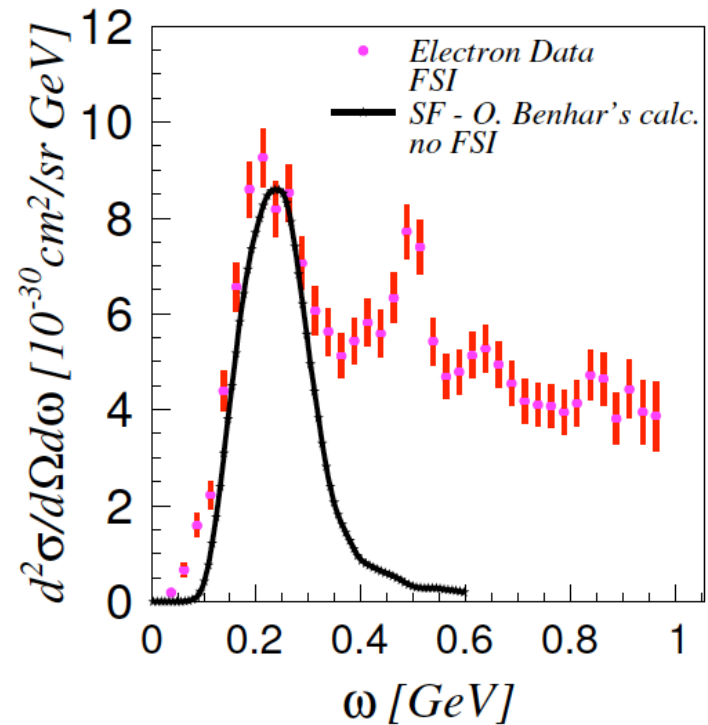
# GENIE 2.8.0 + $\nu T$



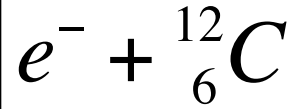
$E_e = 0.88$  GeV and  $\theta_e = 32$  deg



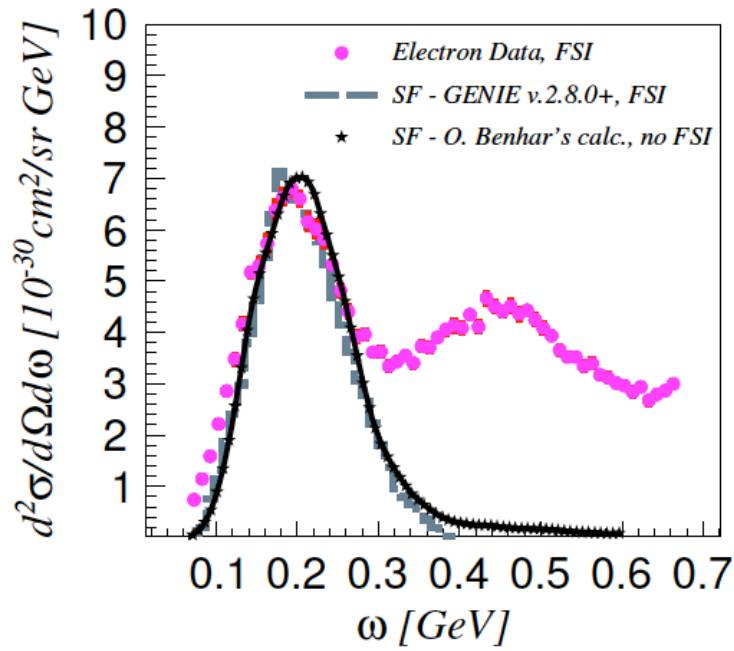
$E_e = 1.2$  GeV and  $\theta_e = 32$  deg



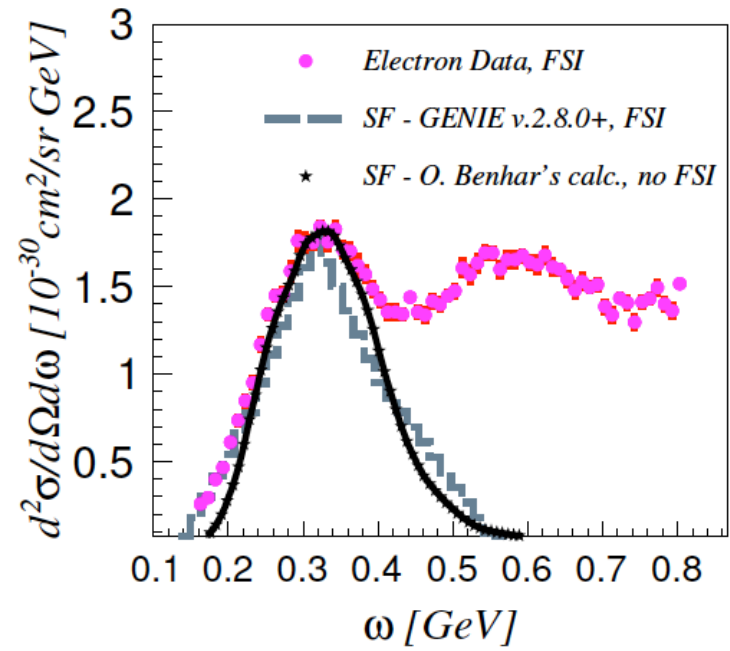
# GENIE 2.8.0 + $\nu T$



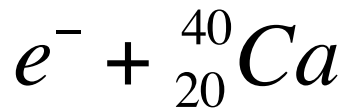
$E_e = 0.961$  GeV and  $\theta_e = 37.5$  deg



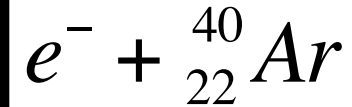
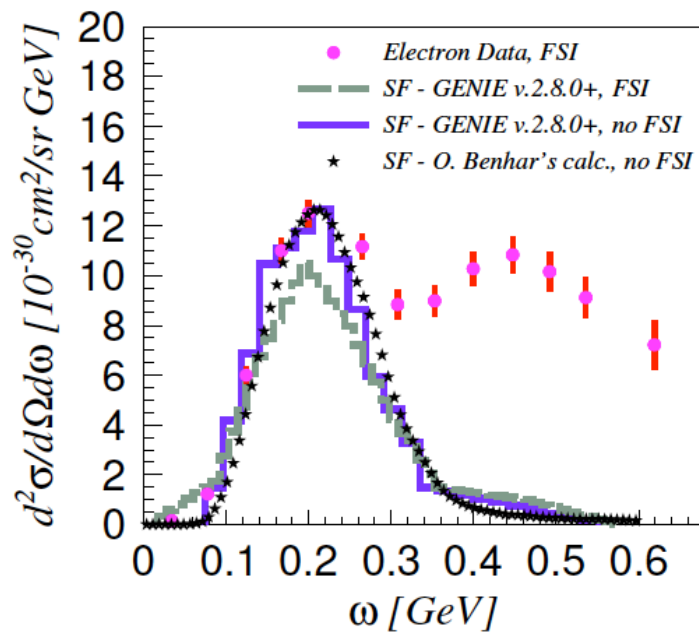
$E_e = 1.299$  GeV and  $\theta_e = 37.5$  deg



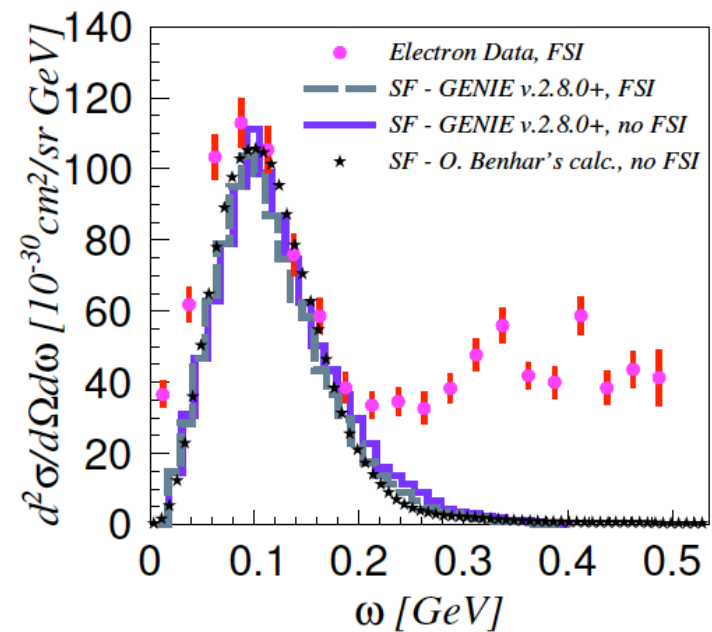
# GENIE 2.8.0 + $\nu T$



$E_e = 0.841$  GeV and  $\theta_e = 45.5$  deg



$E_e = 0.7$  GeV and  $\theta_e = 32$  deg



# Consider only events with no pion in final state: Cerenkov Experiments

- Experimental oscillation analyses requires QE identification (QE-like) with no pions in final state. This is why we care about QE.
- 0-pion events can involve pion production with subsequent pion absorption  $\rightarrow$  ,stuck pion events‘
- Experiments remove the contribution of pionless events due to absorption according to MC models
- Definition of QE cannot distinguish between true QE (1p-1h),  $N^*$  and 2p-2h interactions

Oscillation analysis  
and  
Energy Reconstruction  
in an ideal Long Baseline Experiment



# Experimental Setup

- Ideal and perfect near detector ( $^{12}\text{C}$  or  $^{16}\text{O}$ ), 1 km, 1kton
- Far detector at 295 km, 22.5 kton
  - Oxigen
  - Carbon (RFG and SF)
- Use T2K flux, peak at 0.6 GeV, 750kW, 5 years running
- Use SK reconstruction efficiency as function of energy
- Use migration matrices produced by GiBUU(1.6), GENIE(2.8.0) and GENIE 2.8.0+ $\nu\text{T}$
- Muon neutrino disappearance only -> fit to atmospheric parameters

# Go beyond simple case (arxiv:1311.4506,1402.6651)

- In a real experiment the “real” effects from data will be used in the oscillation analysis together with “some” simulation of nuclear effects
- Use one neutrino generator (GiBUU) to simulate the nuclear effects and use another neutrino generator (GENIE) to extract the oscillation parameters
- Neutrino generators are “enough” different to help understanding what will be the effect of different nuclear models on neutrino oscillation analyses
- Use different nuclear models - RFG and SF - and evaluate effects on neutrino oscillation parameters

- Neglecting all FSI and multinucleon contributions, we can compute the number of events as:

$$N_i^{QE} = \sigma_{QE}(E_i)\phi(E_i)P_{\mu\mu}(E_i)$$

$P_{\mu\mu}$  is the  $\nu_\mu$  oscillation probability - depends on neutrino energy ( $E_i$ )

$\phi$  is the neutrino flux - depends on neutrino energy ( $E_i$ )

$\sigma$  is the cross section for QE event - depends on neutrino energy ( $E_i$ )

- Neglecting all FSI and multinucleon contributions, we can compute the number of events as:

$$N_i^{QE} = \sigma_{QE}(E_i)\phi(E_i)P_{\mu\mu}(E_i)$$

- However, in practice we will observe a different distribution at the detector, given by:

$$N_i^{QE-like} = \sum_j M_{ij}^{QE} N_j^{QE} + \sum_{non-QE} \sum_j M_{ij}^{non-QE} N_j^{non-QE}$$

Where  $M^{QE}$  is the migration matrix for the QE events. Matrix is mostly diagonal and therefore just adds some smearing over the event sample

- Neglecting all FSI and multinucleon contributions, we can compute the number of events as:

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- However, an intermediate situation would most likely take place:

$$N_i^{test}(\alpha) = \alpha N_i^{QE} + (1 - \alpha) N_i^{QE-like}$$

Coloma and Huber, 1307.1243 [hep-ph]

# Nuclear effect factor

$$N_i^{test}(\alpha) = \alpha N_i^{QE} + (1 - \alpha) N_i^{QE-like}$$

$\alpha$  determine the amount of nuclear effect included:

$\alpha=0$  no nuclear effect or nuclear effect perfectly known in the fit

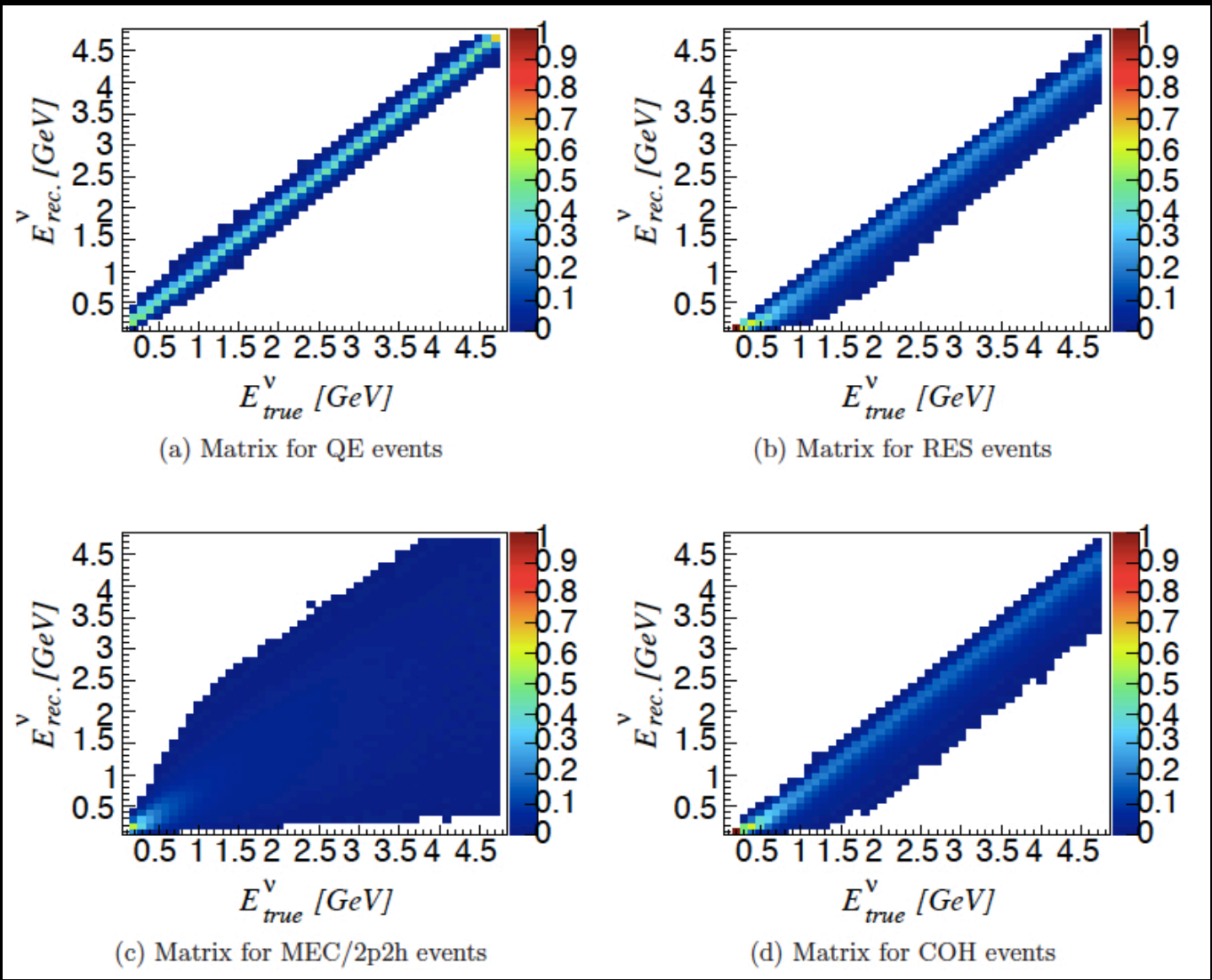
$\alpha=1$  maximum nuclear effect or nuclear effect not known in the fit

# Energy shift

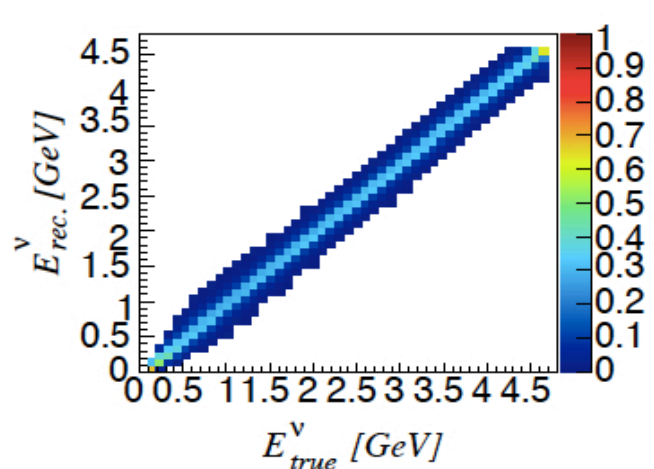
$$N[E] \rightarrow N[(1 + a) E]$$

Modify the number of events as function of energy introducing a calibration error “a” and additional pull term is added to the  $\chi^2$  of the fit.

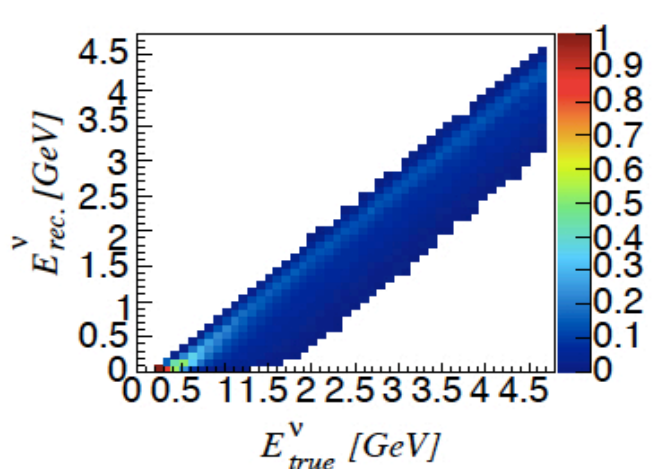
# Migration matrices: GiBUU ( $^{16}\text{O}$ )



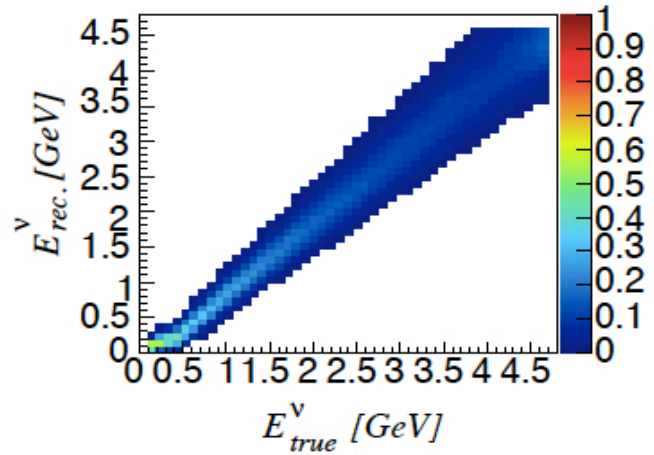
# Migration matrices: GENIE ( $^{16}\text{O}$ )



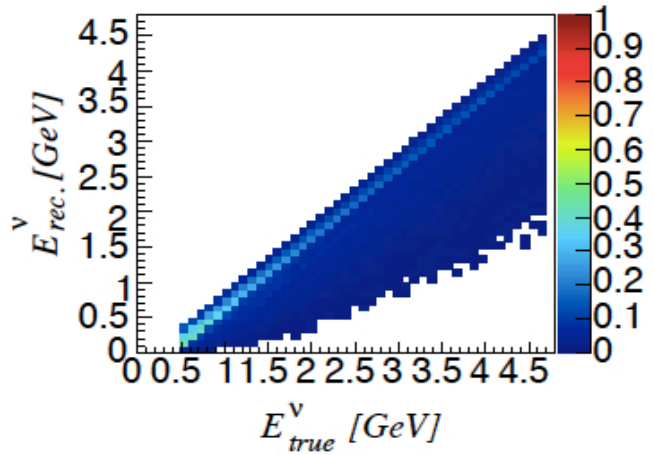
(a) Matrix for QE events



(b) Matrix for RES events



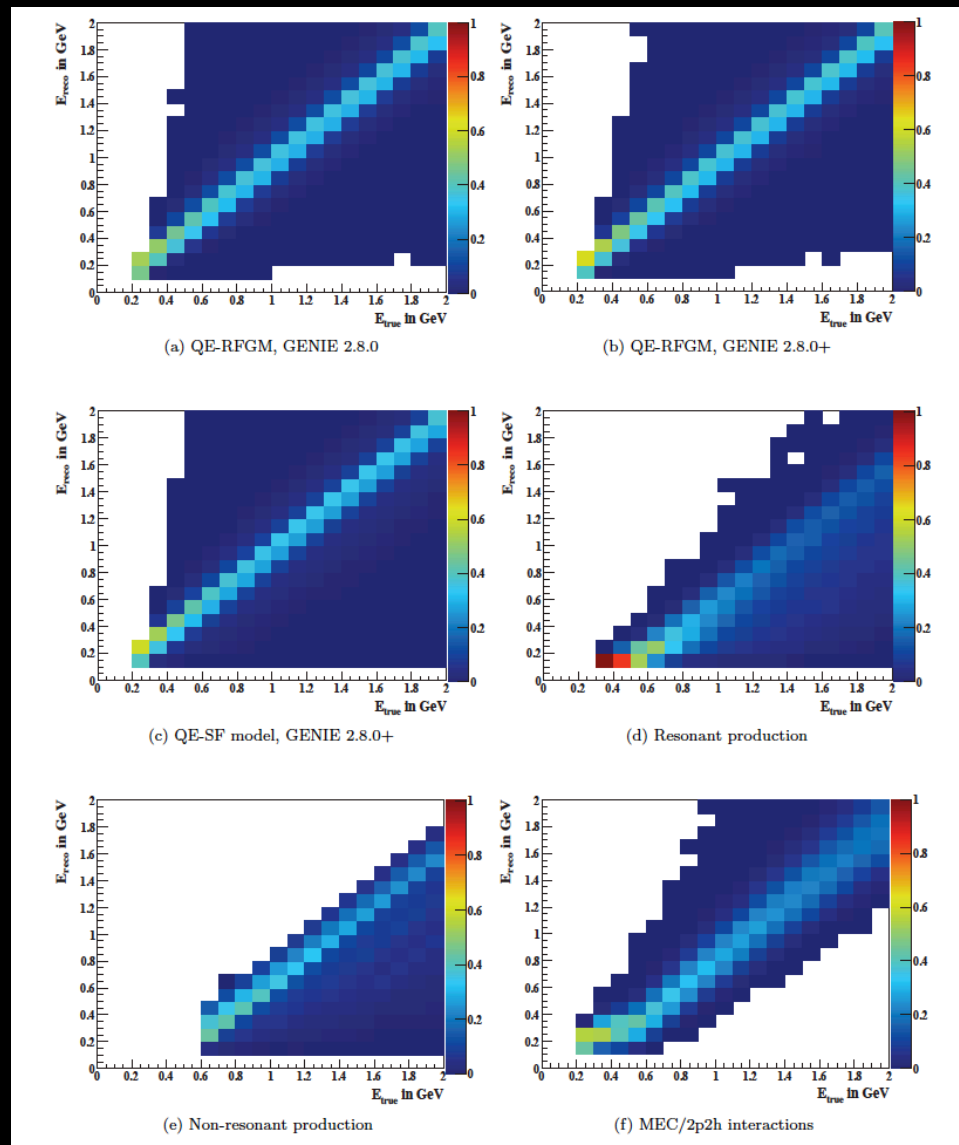
(c) Matrix for MEC/2p2h events



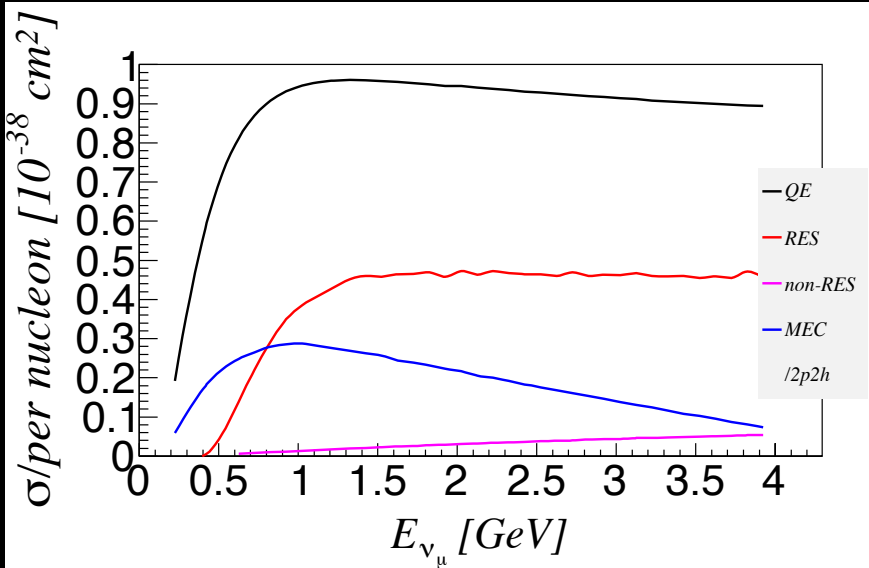
(d) Matrix for COH events



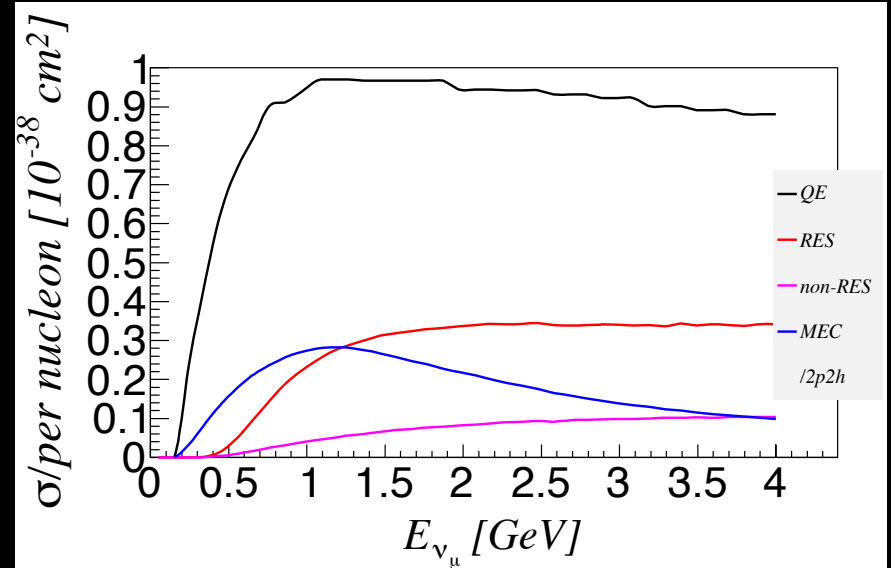
# Migration matrices: GENIE+ $\nu$ T ( $^{12}\text{C}$ )



# Cross-sections

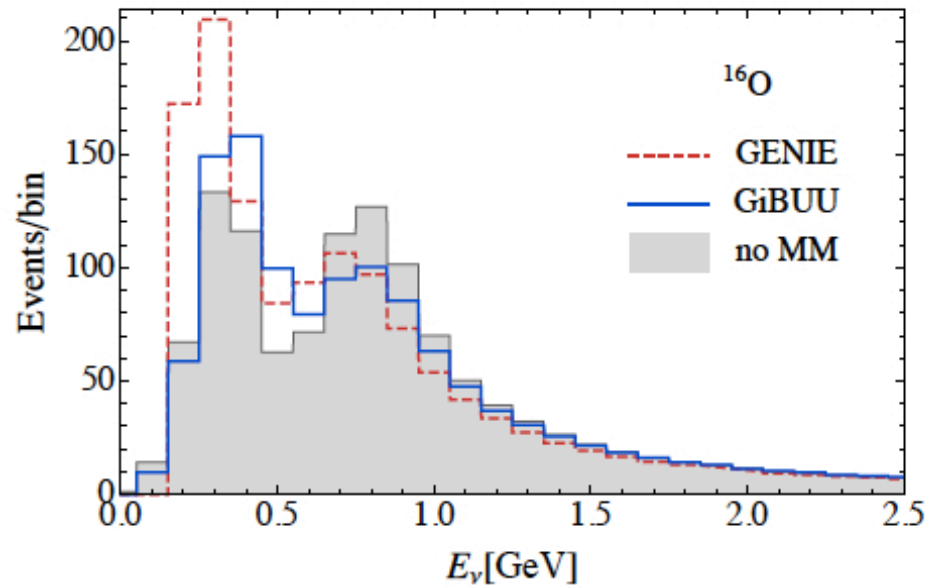


Genie 2.8.0

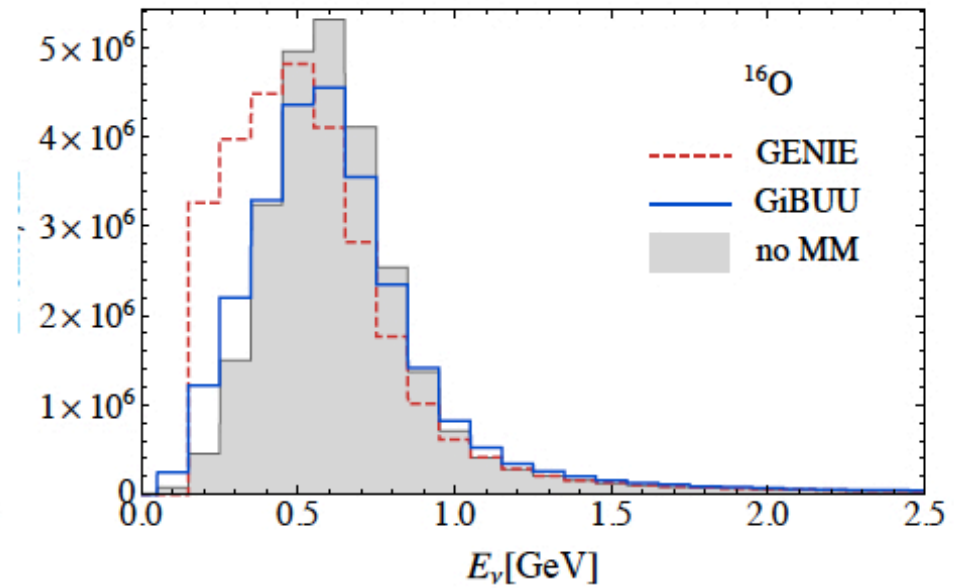


GiBUU 1.6

# Event distributions

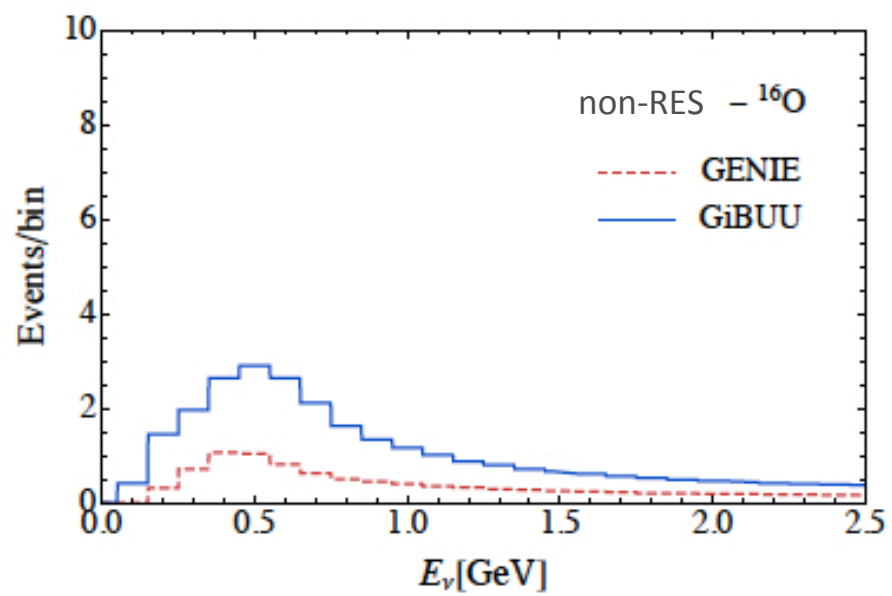
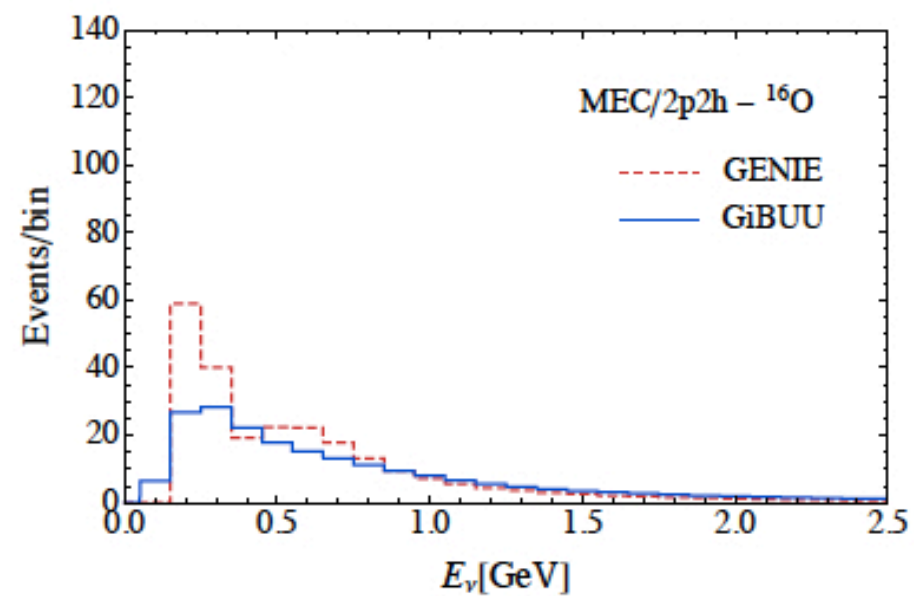
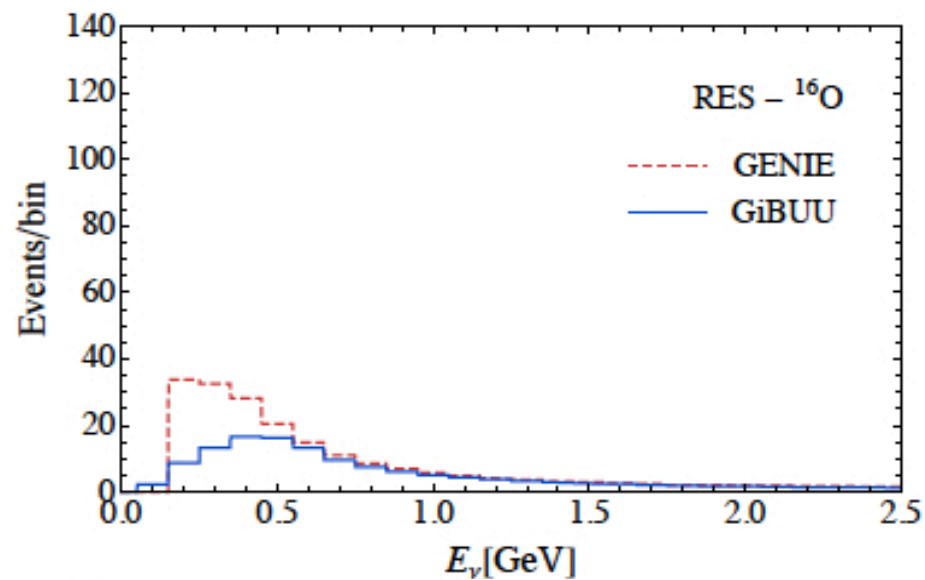
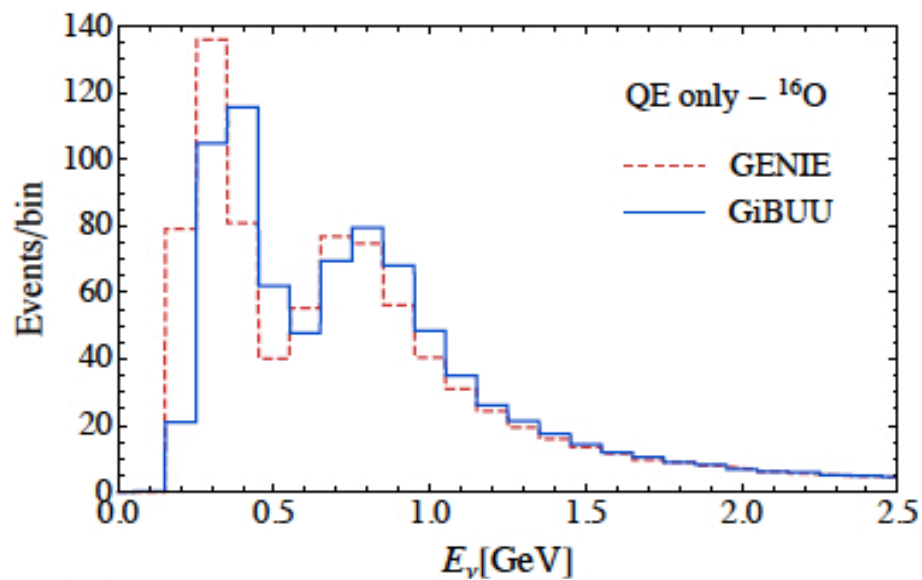


(a) Expected events at the far detector

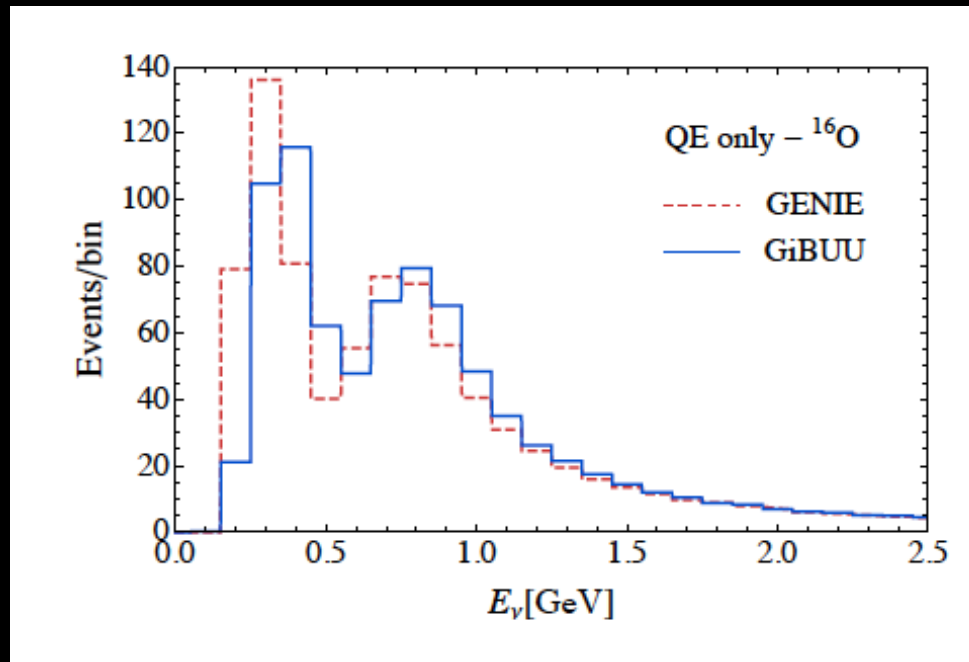


(b) Expected events at the near detector

	QE	RES	non-RES	MEC/2p2h	Total
GiBUU	870	152	32	214	1268
GENIE	877	221	11	249	1358



# A surprise ...

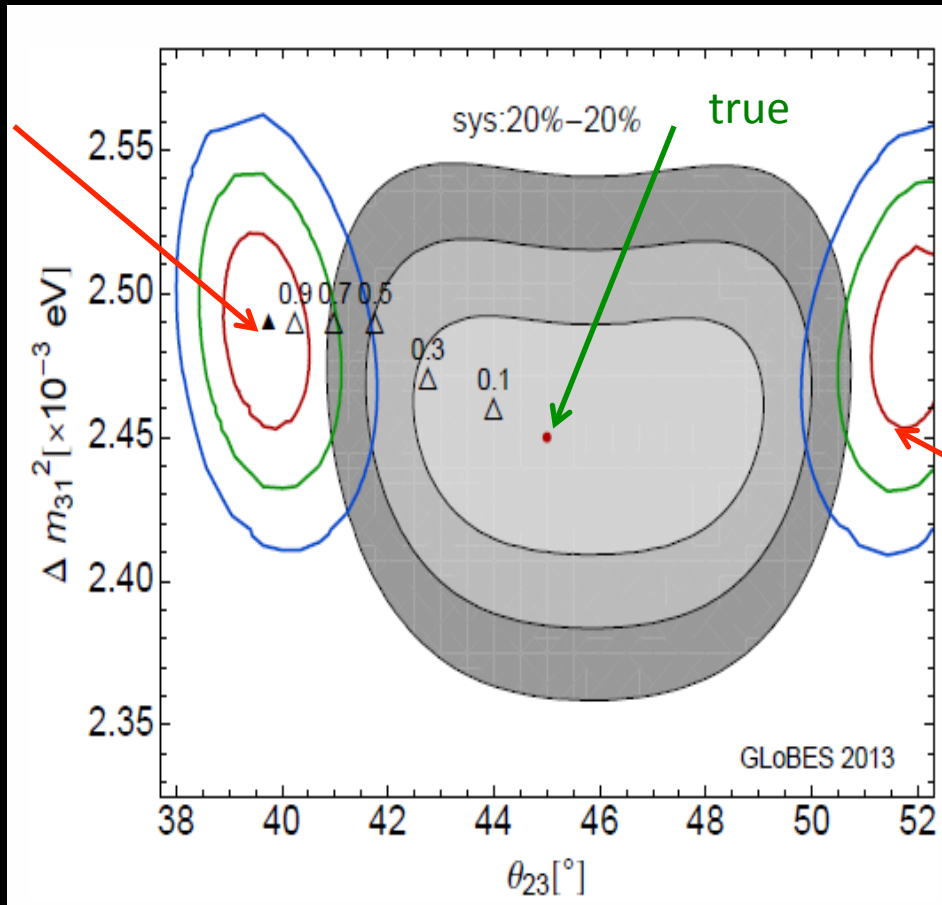


Number of events predicted as function of reconstructed neutrino energy shifted by 10% for pure QE and 17% for all the QE-like events

- Due to FSI – difference is in the migration matrices
- Intrinsic model differences between GENIE and GiBUU
- Intrinsic differences in the model implementations

# How to read the plots

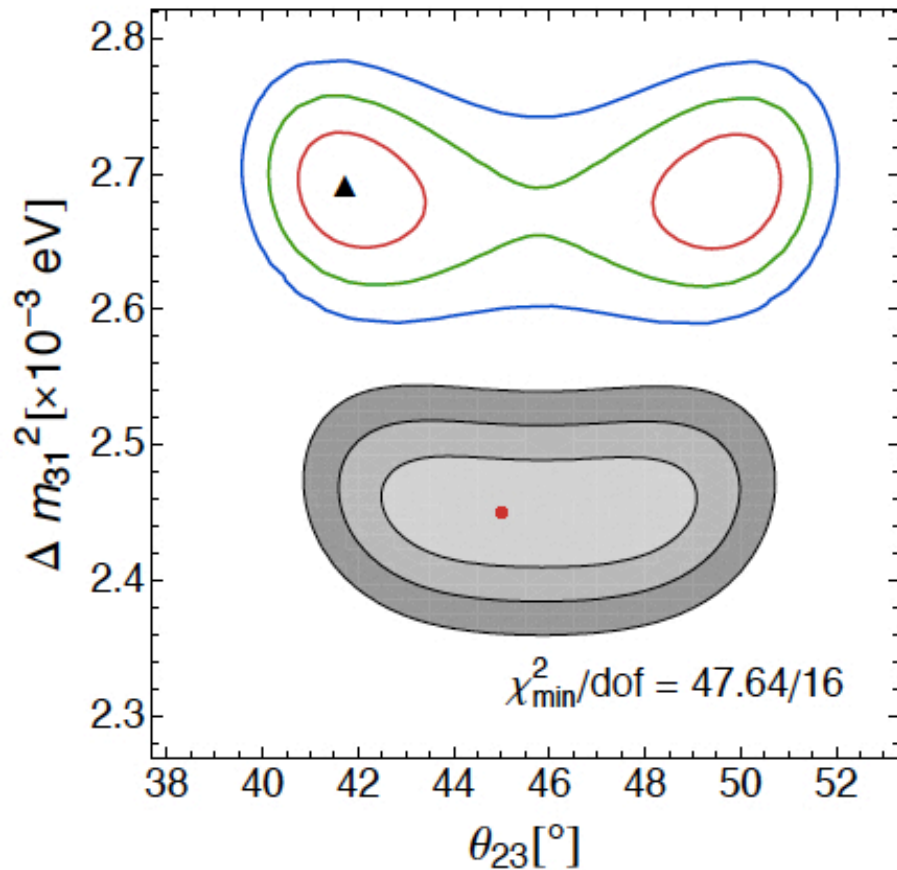
reconstructed  
from naive  
QE dynamics



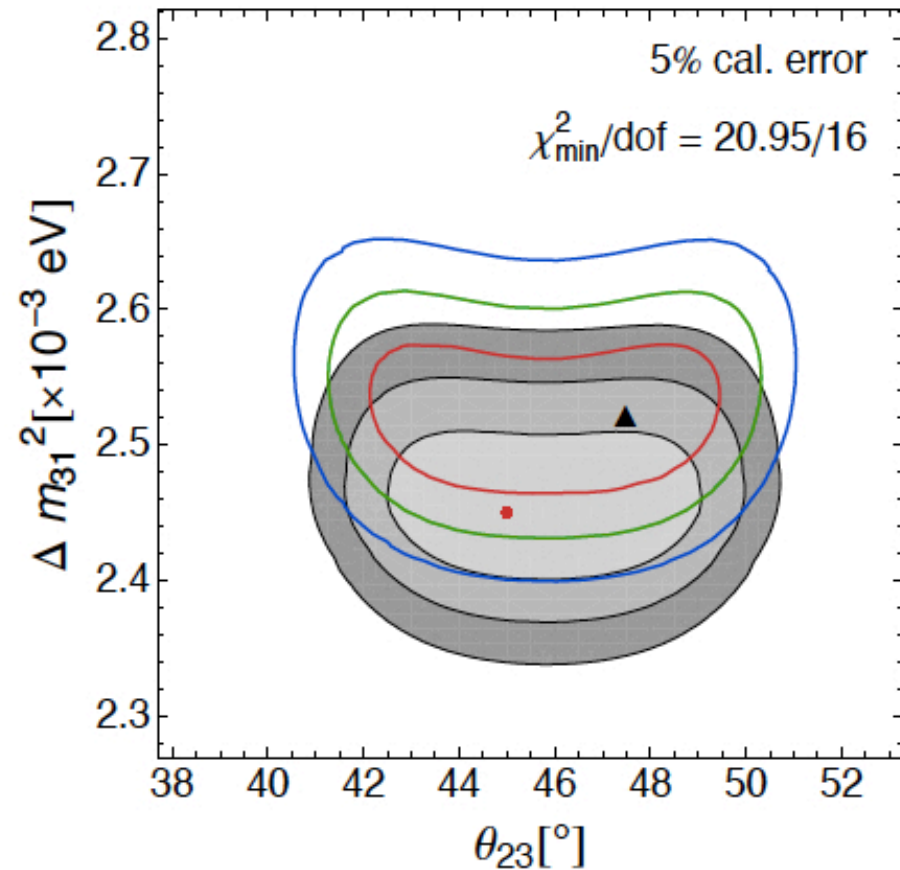
P. Coloma, P. Huber,  
arXiv:1307.1243, July 2013  
Analysis based on GiBUU

1, 2 and  $3\sigma$  allowed regions

# Simulating with GiBUU and extracting oscillations with GENIE 2.8.0: with and without calibration error

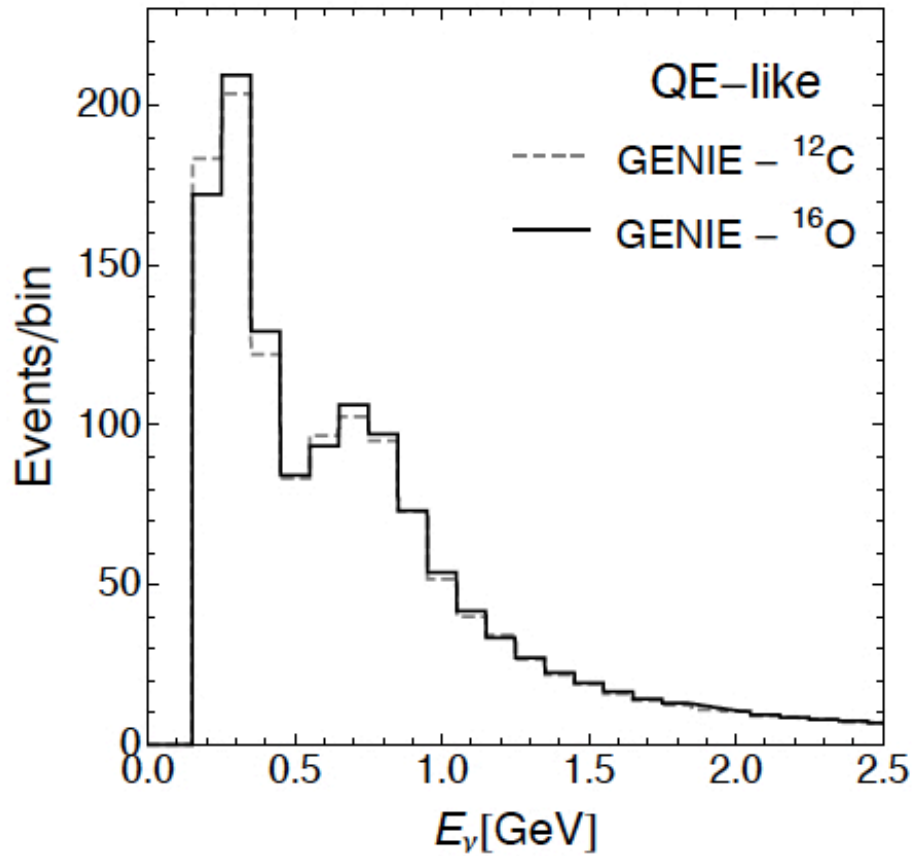


(a) No calibration error

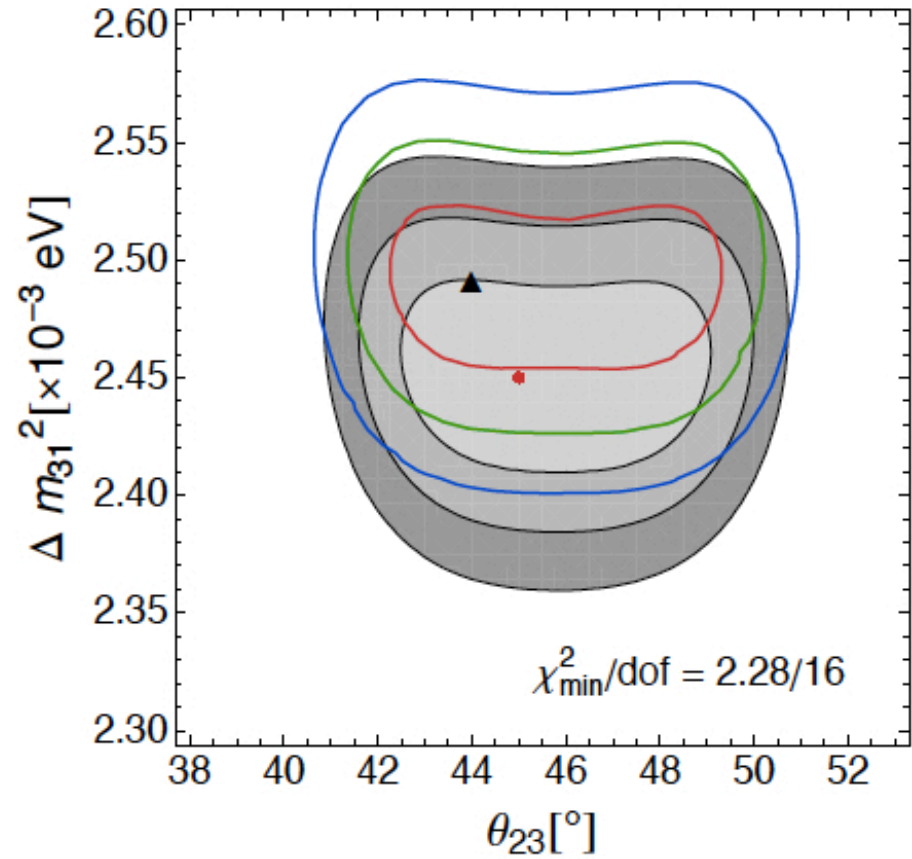


(b) 5% calibration error

# Carbon vs Oxygen



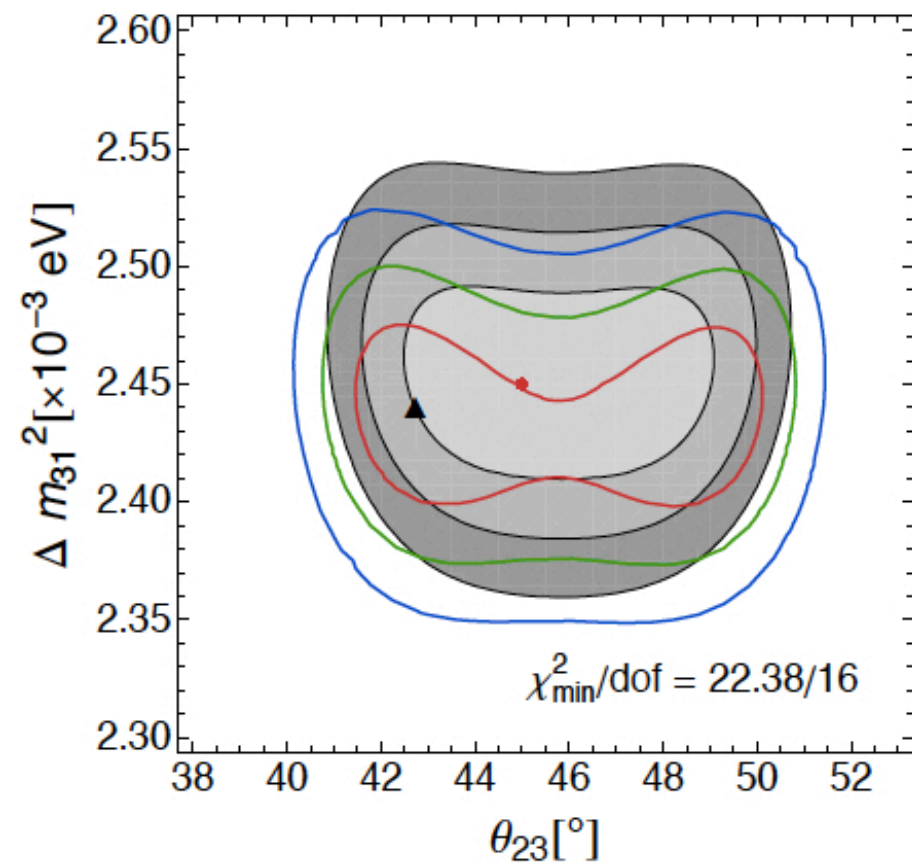
(a) QE-like event distributions



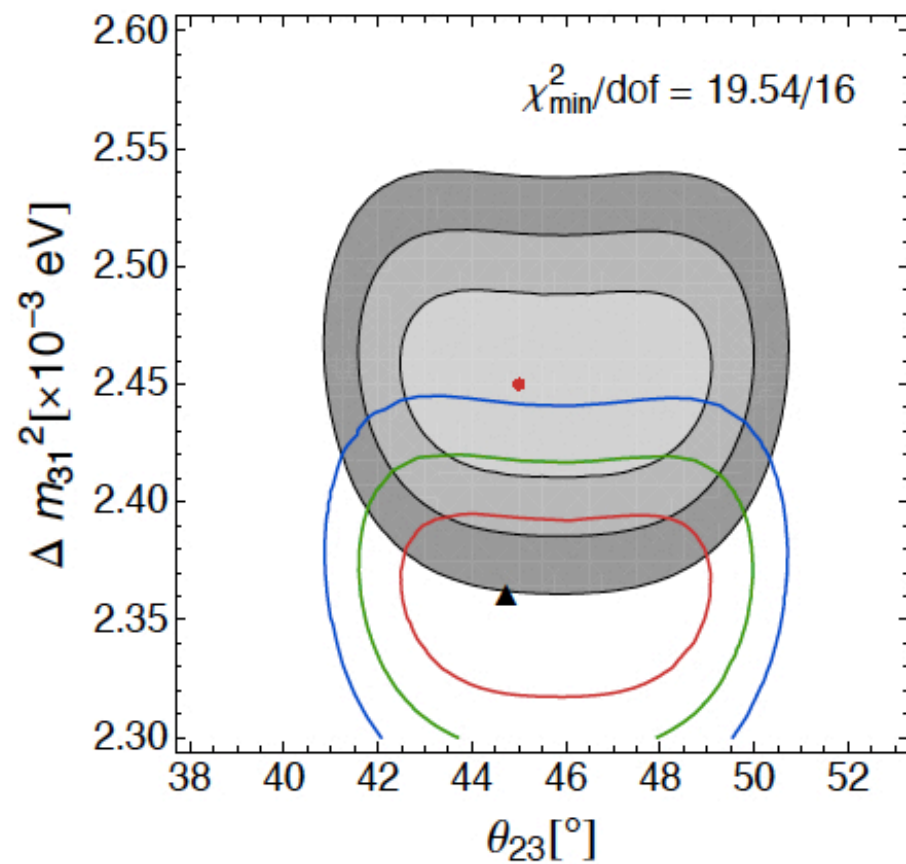
(b) Confidence regions



# With and without MEC/2p2h

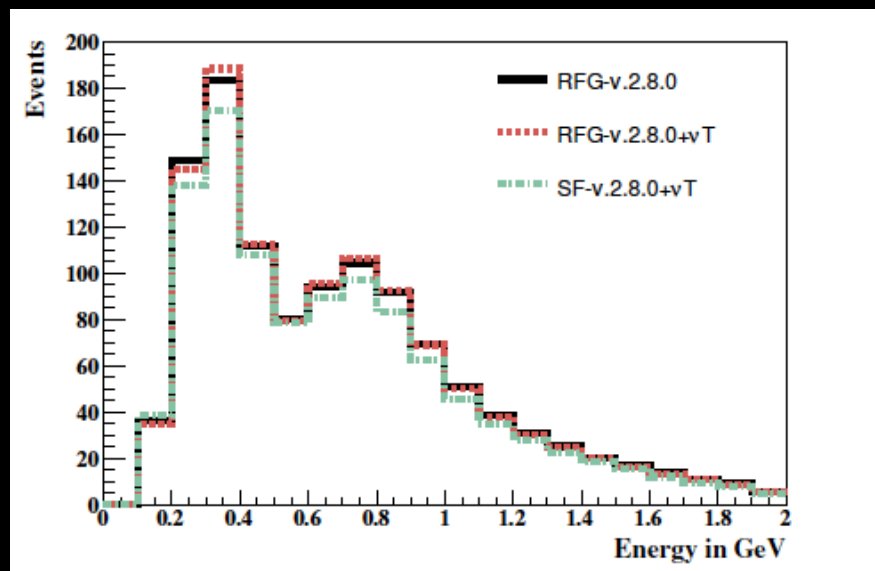
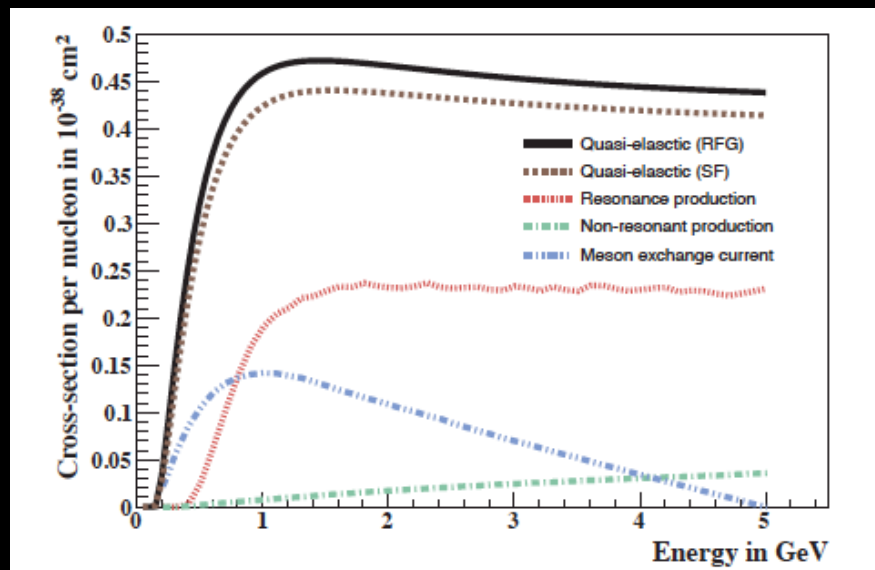
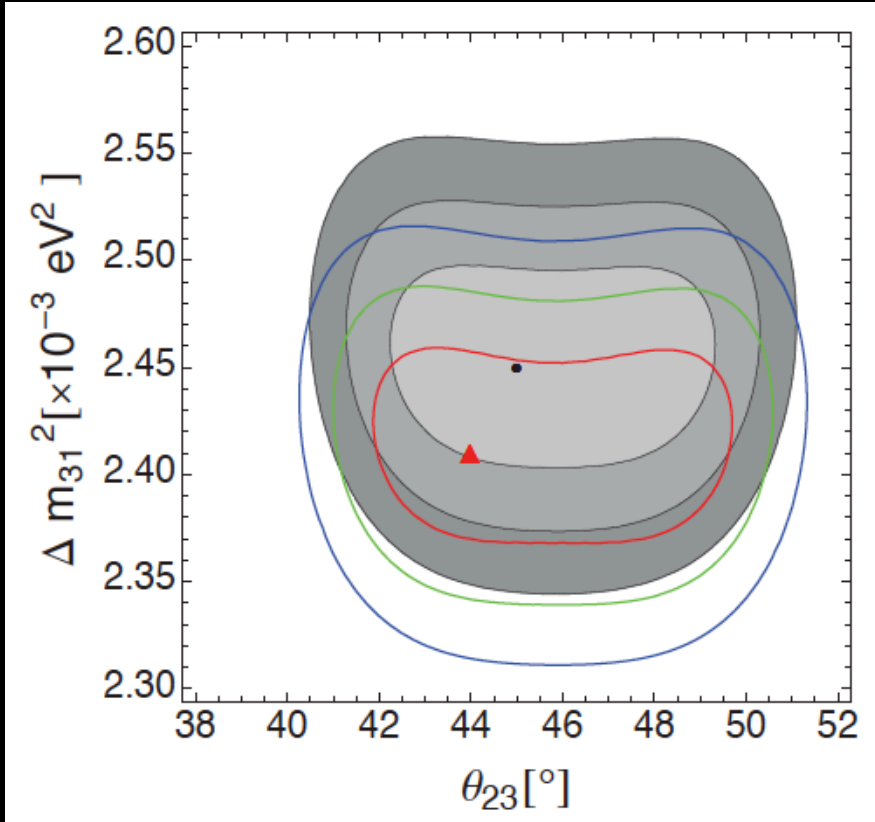


(a) Results using GiBUU matrices



(b) Results using GENIE matrices

# RFG vs SF



# Summary of results

## Input “true” Values

$$\begin{aligned}\theta_{12} &= 33.2^\circ & \Delta m_{21}^2 &= 7.64 \times 10^{-5} \text{ eV}^2 \\ \theta_{13} &= 9^\circ & \Delta m_{31}^2 &= 2.45 \times 10^{-3} \text{ eV}^2 \\ \theta_{23} &= 45^\circ & \delta &= 0^\circ\end{aligned}$$

## Fitted Values

True	Fitted	$\theta_{23,min}$	$\Delta m_{31,min}^2 [\text{eV}^2]$	$\chi_{min}^2$	$\sigma_a$
GENIE ( $^{16}\text{O}$ )	GENIE ( $^{12}\text{C}$ )	$44^\circ$	$2.49 \times 10^{-3}$	2.28	–
GiBUU ( $^{16}\text{O}$ )	GENIE ( $^{16}\text{O}$ )	$41.75^\circ$	$2.69 \times 10^{-3}$	47.64	–
		$47^\circ$	$2.55 \times 10^{-3}$	20.95	5%
GiBUU ( $^{16}\text{O}$ )	GiBUU ( $^{16}\text{O}$ ) w/o MEC	$42.5^\circ$	$2.44 \times 10^{-3}$	22.38	–
GENIE ( $^{16}\text{O}$ )	GENIE ( $^{16}\text{O}$ ) w/o MEC	$44.5^\circ$	$2.36 \times 10^{-3}$	19.54	–

# Summary of results cont'd

## Input “true” Values

$$\begin{aligned}\theta_{12} &= 33.2^\circ & \Delta m_{21}^2 &= 7.64 \times 10^{-5} \text{ eV}^2 \\ \theta_{13} &= 9^\circ & \Delta m_{31}^2 &= 2.45 \times 10^{-3} \text{ eV}^2 \\ \theta_{23} &= 45^\circ & \delta &= 0^\circ\end{aligned}$$

## Fitted Values

True	Fitted	$\theta_{23,min}$	$\Delta m_{31,min}^2 [\text{eV}^2]$
RFGM <sub>2.8.0+<math>\nu T</math></sub>	RFGM <sub>2.8.0</sub>	45.7 deg	$2.45 \times 10^{-3}$
SF <sub>2.8.0+<math>\nu T</math></sub>	RFGM <sub>2.8.0+<math>\nu T</math></sub>	44 deg	$2.41 \times 10^{-3}$
SF <sub>2.8.0+<math>\nu T</math></sub>	RFGM <sub>2.8.0</sub>	44.5 deg	$2.41 \times 10^{-3}$

# Conclusions

- Energy reconstruction essential for precision determination of neutrino oscillation parameters and neutrino-hadron cross sections
- Impact on neutrino oscillation experiments due to nuclear models, what they are and how they are implemented is not negligible (order 10%)
  - comparing systematically generators is important
  - neutrino event generators use almost same data set so there are correlations that are non-negligible
  - using wrong models affect neutrino oscillation parameters determination

- In future extend the case to CP violation:
  - neutrino vs anti-neutrino cross-section, : do we have reliable event generators for anti-neutrino ?
- MEC/2p2h: heavily tuned to MiniBooNE data in both GiBUU and GENIE.
  - the contribution of MEC/2p2h is the same between Carbon and Oxygen, should it be ?
- Energy reconstruction requires reliable event generators, of same quality as experimental equipment
- Precision era of neutrino physics requires much more sophisticated generators and a dedicated effort in theory
- Theorists-phenomenologists and experimentalists need to work together: NuSTEC

Generators are a crucial part  
of any experiment!  
Must be of same quality as the  
experimental equipment itself!  
Needed resources are relatively  
small, but still not available

# Thank You !

