

International Symposium on Multiparticle Dynamics Bologna, September 8-12, 2014



# $\beta \beta$ (0 $\nu$ ) Search with CUORE-0 and CUORE

Niccolò Moggi University of Bologna and INFN on behalf of the CUORE collaboration

# Outline

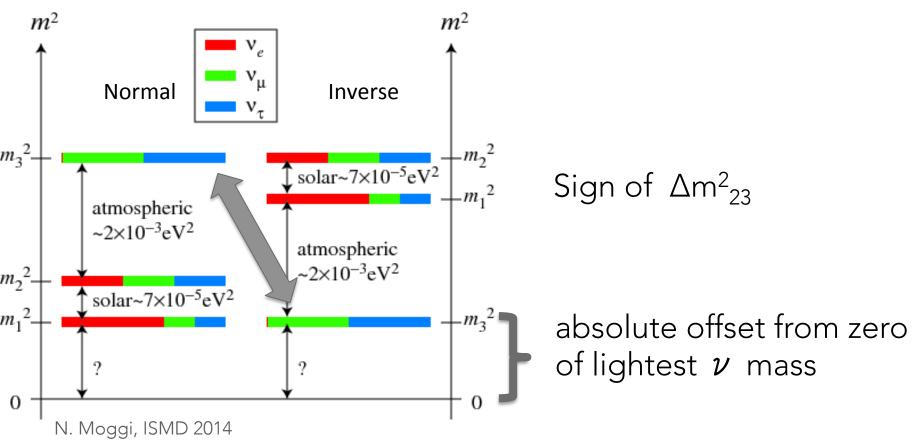
- Why search for  $\beta\beta(0\nu)$
- Physics signal
- The experimental parameters
- A phased search program:
   Cuoricino → CUORE-0 → CUORE
- Preliminary CUORE-0 results
- CUORE status and projected sensitivity
- Conclusions

# Some Neutrino Open Questions

1. Nature: Dirac or Majorana particles ?

 $\nu = \overline{\nu}$  ?  $\rightarrow$  Lepton Number Violation

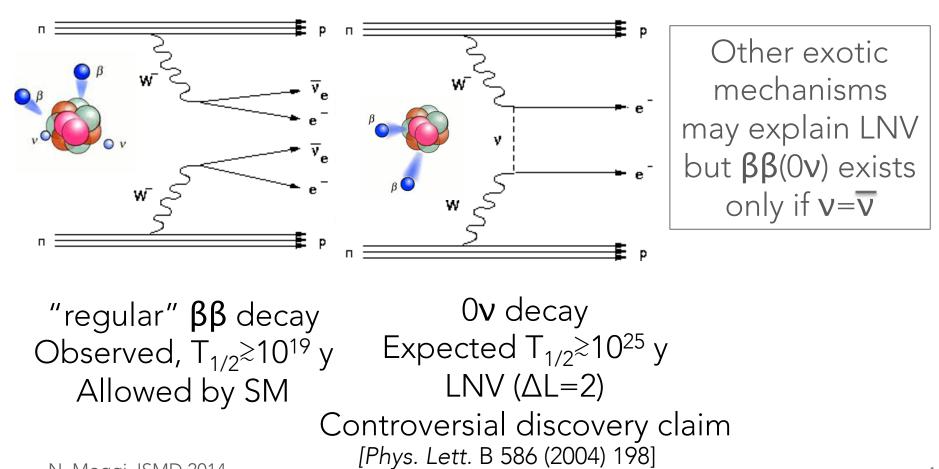
2. Mass: absolute scale, hierarchy of mass states



3

# $\beta \beta (0\nu)$

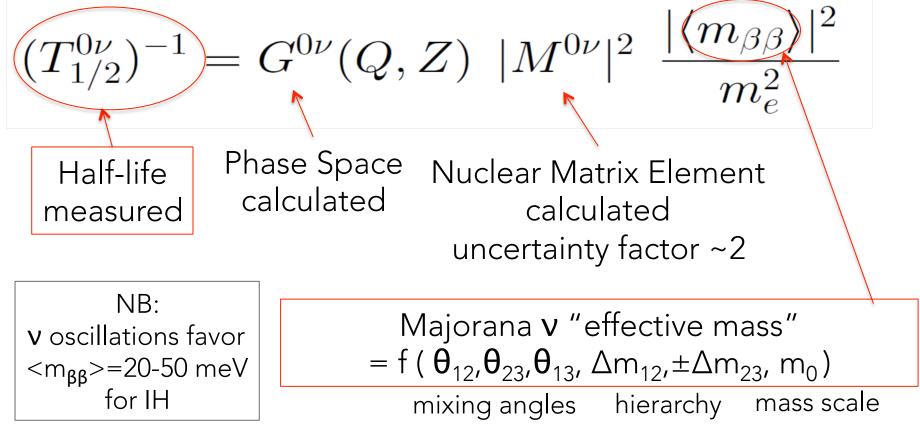
- Most sensitive method to investigate  $\,
  u\,$  nature
- (Z, A)  $\rightarrow$  (Z+2, A) +2e<sup>-</sup>(+2 $\nu$ )



# **Decay Rate**

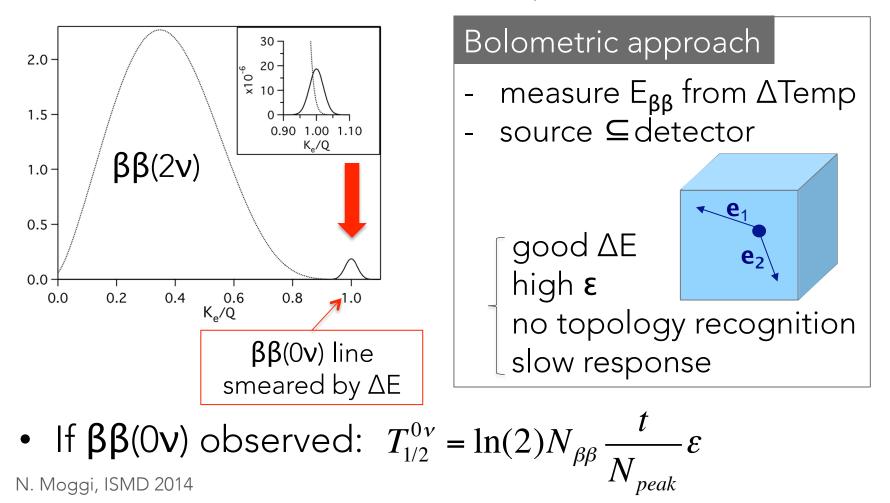
### If $\beta\beta(0\nu)$ is found:

- $\nu$  is Majorana && LNV
- Constraints on the absolute u mass scale

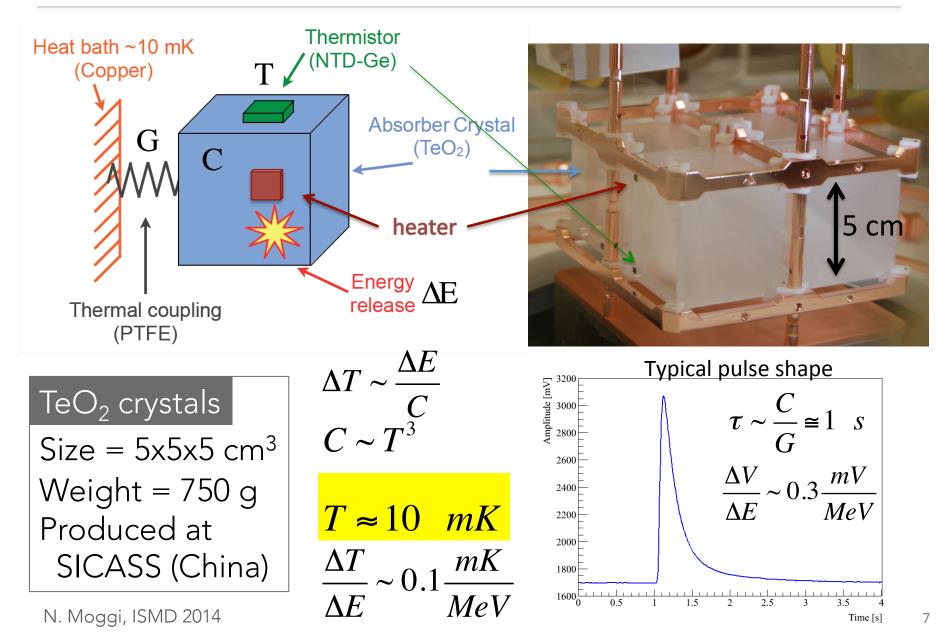


# **Experimental Signature**

- Measure  $E_{\beta\beta} = E_{\beta1} + E_{\beta2}$
- Signature: peak at  $Q=m_n m_p$



# **Bolometric Technique**

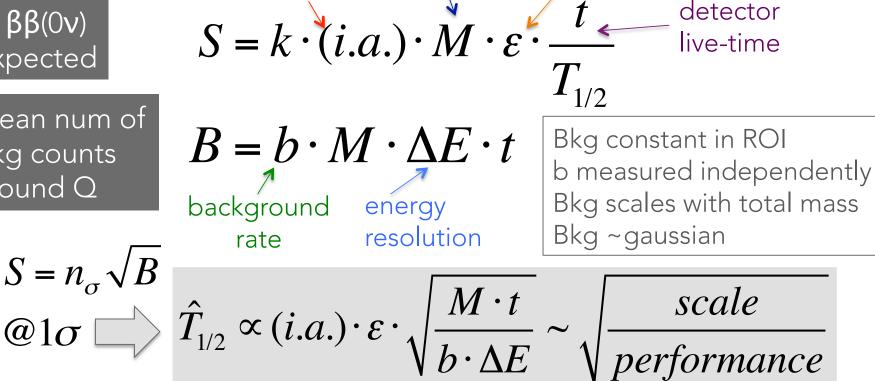


# Sensitivity

If no  $\beta\beta(0\nu)$ , we can express the sensitivity as the  $T_{1/2}$  that corresponds to a signal mimicked by a  $1\sigma$ fluctuation tot active efficiency isotopic mass abundance

N ββ(0 $\nu$ ) expected

Mean num of <u>bkg co</u>unts around Q



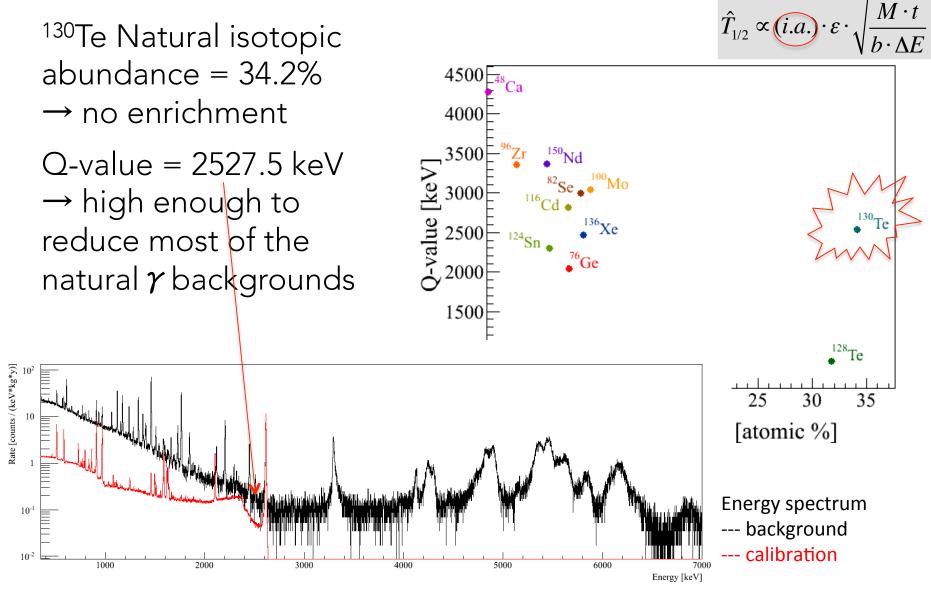
N. Moggi, ISMD 2014

# The Collaboration

Cryogenic Underground Observatory for Rare Events

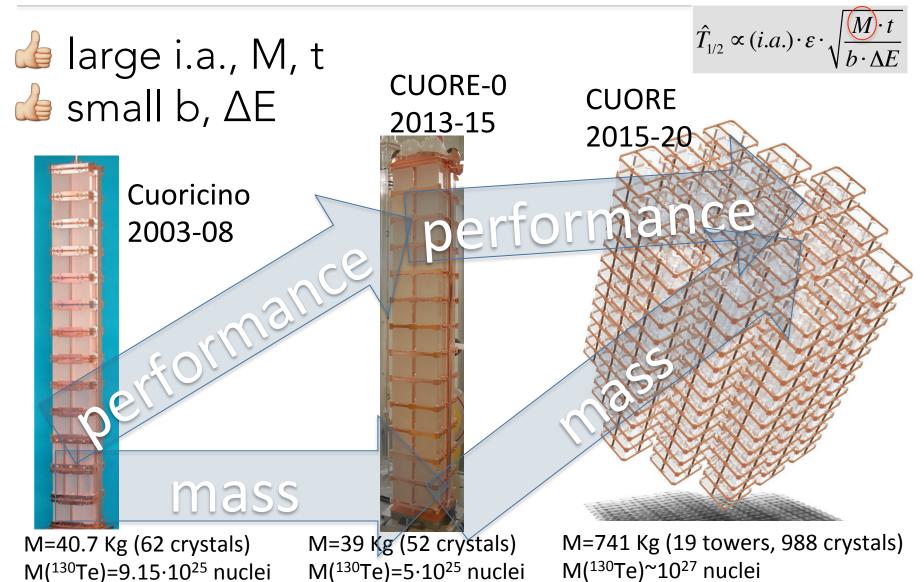


# Choice of Isotope



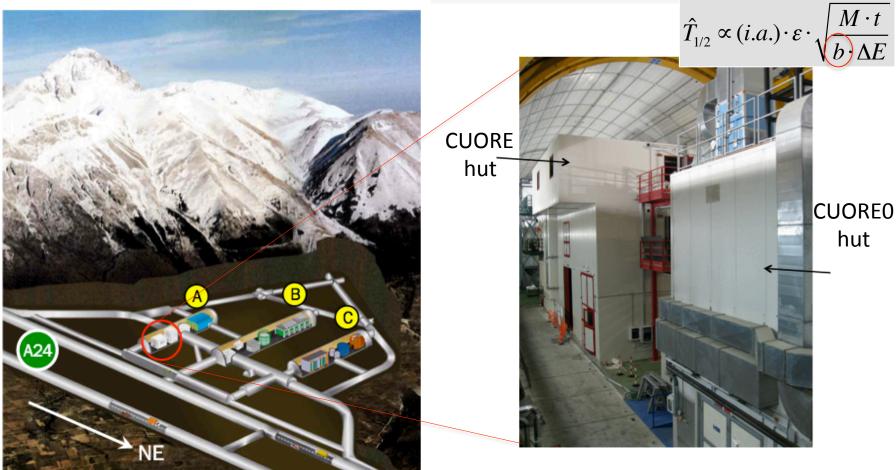
N. Moggi, ISMD 2014

# A Phased Search Program



N. Moggi, ISMD 2014

# **Experimental Site**



In A Hall of LNGS, Italy (Laboratori Nazionali del Gran Sasso)

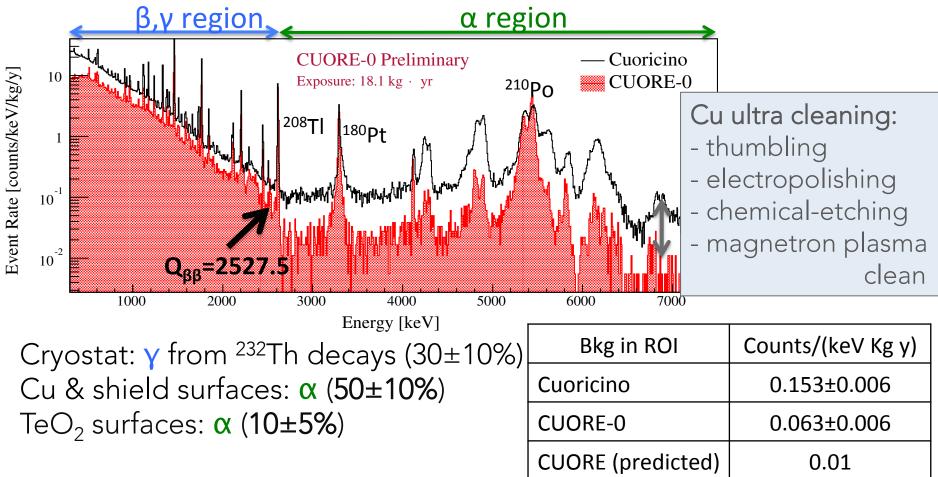
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Average depth ~ 3650 m.w.e.  $\mu$  flux: (2.58 ± 0.3)·10<sup>-8</sup>  $\mu$ /s/cm<sup>2</sup> n flux <10 MeV: 4·10<sup>-6</sup> n/s/cm<sup>2</sup>  $\gamma$  flux < 3 MeV: 0.73  $\gamma$ /s/cm<sup>2</sup> Sum negligible after shielding

# **CUORE-0** Background

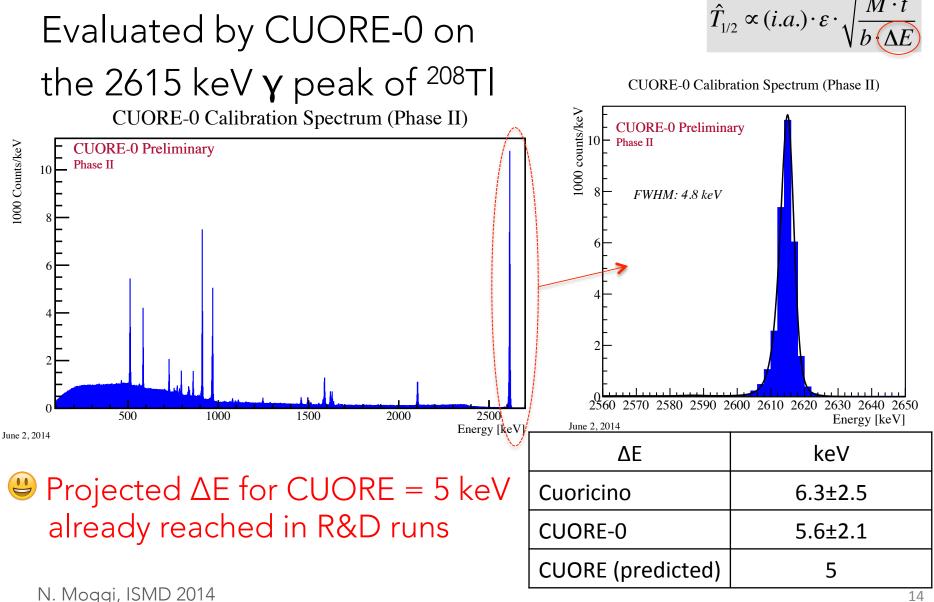
Bkg consistent with Cuoricino model...

..but reduction factor ~2 in ROI (~6 in  $\alpha$  region)



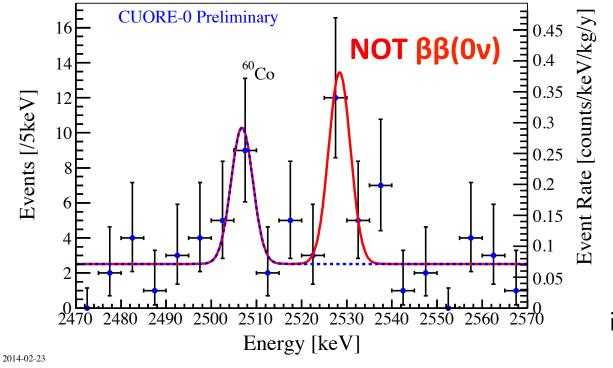
 $\hat{T}_{1/2} \propto (i.a.) \cdot \varepsilon \cdot j$ 

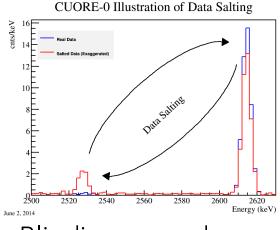
# **CUORE-0 Energy Resolution**



# **CUORE-0 Preliminary Results**

- Cuoricino:  $T_{1/2}^{0v} > 2.8 \cdot 10^{24} y$  (90% C.L.) best ever
- CUORE-0: EPJC 74, 2956 (2014)
  - blinded (unblinding early 2015 ?)



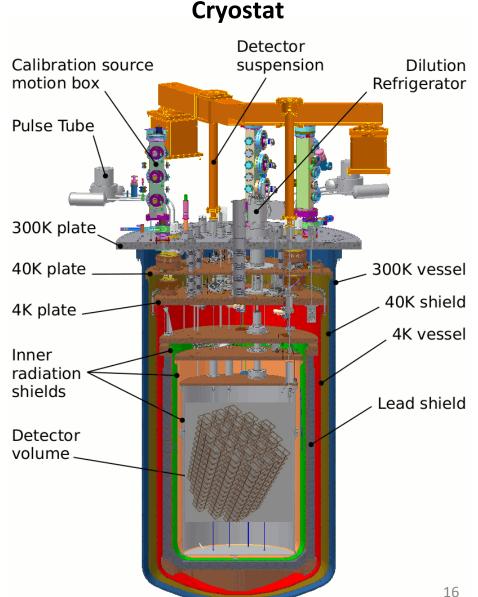


Blinding procedure: small (blinded) % of events in ±10 keV under <sup>208</sup>Tl peak exchanged with events ±10 keV around Q-value

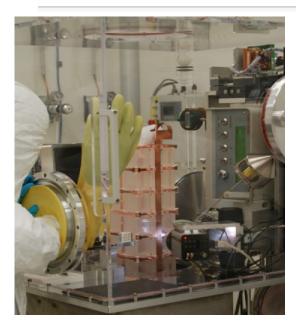
# **CUORE Background**

- More bolometers (selfshielding, coincidences, less crystals facing shields)
- Roman lead shield
- New cryostat
- T = 13 → 10 mK
- Improved operating conditions
- BKG = 0.01 c/(keV Kg y) within reach !





# **CUORE Status**





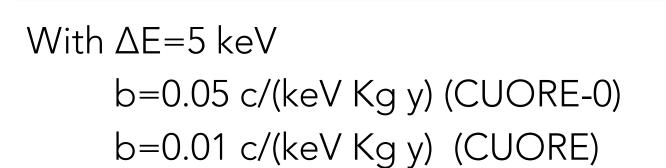


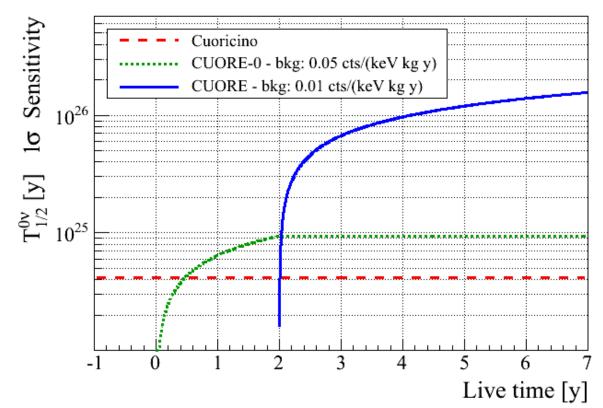
- All 19 towers now ready
- Moving to detector integration
- Now: new cryostat commissioning, working on DAQ, data analysis tools, slow-control, Farady cage...
  - still copper cleaning

Cuore data taking expected in summer 2015

N. Moggi, ISMD 2014

# **Projected Sensitivity**

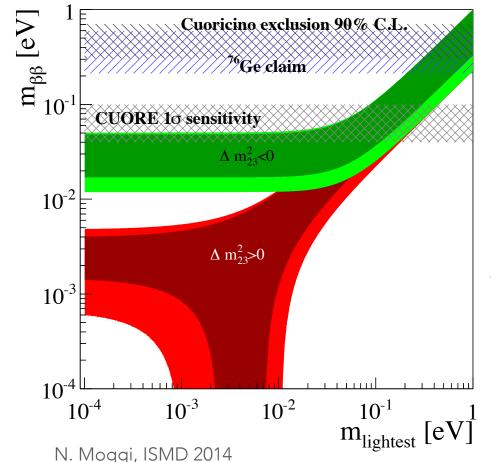




 $\hat{T}_{1/2} \propto (i.a.) \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{h \cdot \Lambda E}}$ 

# **Projected Sensitivity**

Sensitivity on  $T_{1/2}$  may be translated into sensitivity on  $\langle m_{\beta\beta} \rangle$  as:  $\sqrt{h \cdot \Lambda}$ 



$$< m_{\beta\beta} > \propto \sqrt[4]{rac{b \cdot \Delta E}{M \cdot t}}$$

≈ 40-100 meV @ 1σdepending on NME

CUORE will likely just reach the  $< m_{\beta\beta} >$  region favored by oscillation results for IH

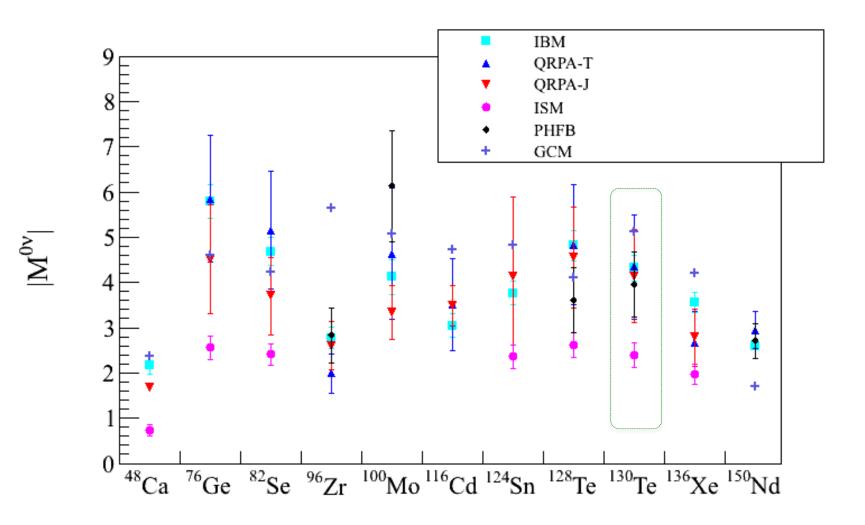
# Conclusions

- CUORE-0 is taking data
   confirms Cuoricino bkg model
  - demonstrates that  $\Delta E \lesssim 5$  keV is achievable
- CUORE expected to start in summer 2015 – observe  $\beta\beta(0\nu) \rightarrow \nu$  is Majorana && LNV – no observation  $\rightarrow T_{1/2}$  limit best ever
- Many R&D focused on bkg reduction
  - next generation experiments aim to reach IH  $<\!m_{\beta\beta}\!>$  region

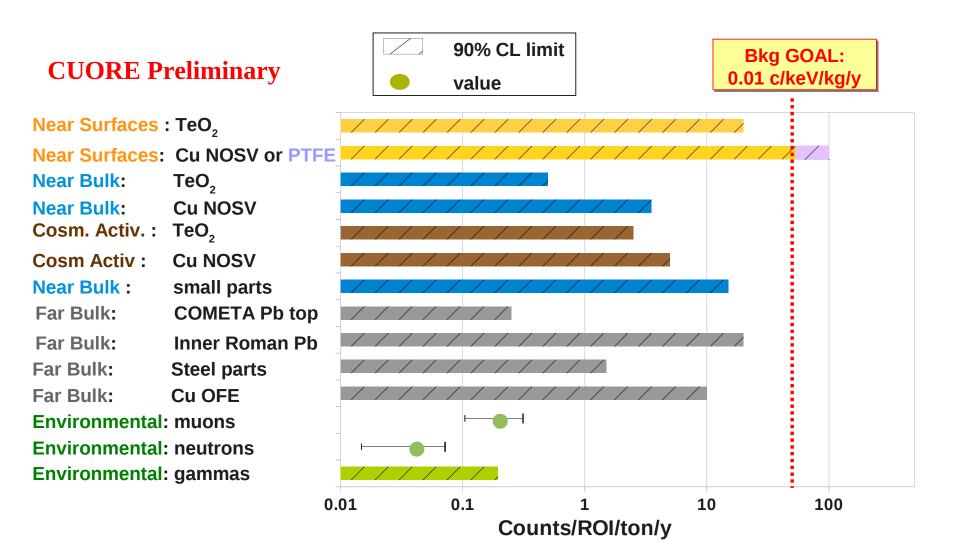
# Thank you !

### BACKUP SLIDES

# NME

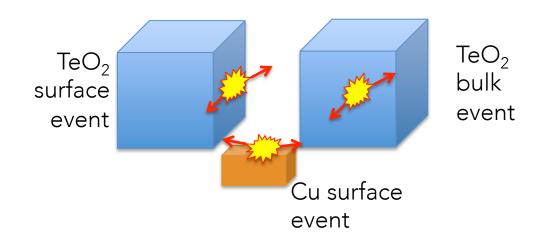


# Background Budget



# Signal Selection

- Only two signatures may distingush events from bkgs:
  - Energy release
  - Single hit :  $\beta\beta(0\nu)$  signal is confined in one single crystal  $\rightarrow$  multi-hit events are very likely background

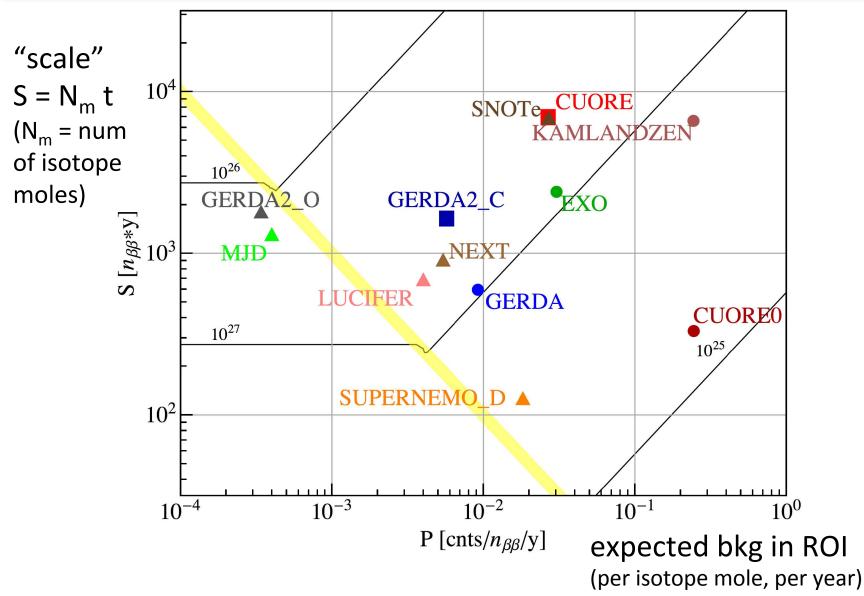


# The Past Experiments

Isotope	$T_{1/2}^{2\nu}$	$T_{1/2}^{0\nu}$	$ \langle m_{ u}  angle $
	$(10^{19} \text{ yr})$	$(10^{24}  m yr)$	(eV)
<sup>48</sup> Ca	$(4.4^{+0.6}_{-0.5})$ [80, 81, 82]	> 0.058[101]	< 19 - 36
$^{76}$ Ge	$(150 \pm 10)$ [83, 84, 85, 86]	$22.3^{+4.4}_{-3.1}[\overline{78}]$	
		> 19[70]	< 0.17 - 0.29
		> 15.7 71	< 0.19 - 0.32
$^{82}$ Se	$(9.2\pm0.7)$ 87, 88	> 0.36 [102]	< 1.23 - 1.88
$^{96}\mathrm{Zr}$	$(2.3 \pm 0.2)$ 89, 90	> 0.0092[90]	< 5.24 - 10.83
$^{100}Mo$	$(0.71 \pm 0.04)$ 91, 92	> 1.1 102	< 0.71 - 1.05
$^{116}Cd$	$(2.8 \pm 0.2)$ 82, 93, 94, 95	> 0.1794	< 1.64 - 2.69
$^{130}\mathrm{Te}$	$(70^{+9}_{-11})[96, 97]$	> 2.8[103]	< 0.45 - 0.70
$^{136}$ Xe	$(217 \pm 6)$ [98]	> 1.699	< 2.10 - 3.37
$^{150}$ Nd	$(0.82 \pm 0.09)[92], 100]$	> 0.018[100]	< 9.01 - 16.07

List of best reported results on  $\beta\beta(0\nu)$  process. Limits are at 90% C.L. [arXiv.1310.4692v1]

# The Competitors



# **Next Generation**

Experiment	Isotope	M <sub>ββ</sub>	Technique	Location	Start date
		(kg)			
<sup>130</sup> Te	CUORE0/CUORE	11/206	Bolometric	LNGS	2012/2015
<sup>76</sup> Ge	GERDA I/II	11/30	Ionization	LNGS	2012/2014
<sup>82</sup> Se	LUCIFER	9	Bolometric	LNGS	2014
	MJD	26	Ionization	SUSEL	2014
<sup>130</sup> Te	SNO+	163	Scintillation	SNOlab	2014
<sup>82</sup> Se or <sup>150</sup> Nd	SND/SuperNEMO	6/100	Tracko-calo	LSM	2014/2015
<sup>136</sup> Xe	EXO-200	79	Liquid TPC	WIPP	2012
<sup>136</sup> Xe	KamLAND-ZEN	179	Scintillation	Kamioka	2012
<sup>136</sup> Xe	<b>NEXT-100</b>	90	Gas TPC	Canfranc	2014

[O. Cremonesi, POS (EPS-HEP 2013) 146]

# **Next Generation**

	Isotope	Biso	FWHM (keV)	Status	$F_{68\% C.L.}^{0v}$ (5 yr)	$ \langle m_{\rm ee} \rangle $
CUORE0	<sup>130</sup> Te	266	5.6	R	1.5	224
CUORE	<sup>130</sup> Te	36	5	С	21	60
GERDA I	<sup>76</sup> Ge	21	4.8	R	9.4	165
GERDA II	<sup>76</sup> Ge	20/1.1	3.2	С	22/60*	107/65*
LUCIFER	<sup>82</sup> Se	1.9	13	D	16*	76*
MJD	<sup>76</sup> Ge	0.9	4	С	44*	77*
SNO+	<sup>130</sup> Te	0.9	240	D	20	62
EXO	<sup>136</sup> Xe	1.9	96	R	12	97
SND	<sup>82</sup> Se	0.6	120	D	3.3	166
SuperNEMO	<sup>82</sup> Se	0.6	130	D	13	85
KamLAND-Zen	<sup>136</sup> Xe	7.4	243	R	6.9	127
NEXT	<sup>136</sup> Xe	0.8	13	D	16	82

[O. Cremonesi, POS (EPS-HEP 2013) 146]

# If Next Generation Fails

- if (IH region is reached && no ββ(0ν) is observed ) then {we can still reach conclusions with some input from other experiments} :
  - input: **v** is a Majorana particle  $\rightarrow$  hierarchy is NH
    - will have to wait even longer
  - input: hierarchy is IH  $\rightarrow v$  is a Dirac particle
- if v is a Dirac particle we have no hope of  $\beta\beta(0v)$  observation

# The $\beta\beta(0\nu)$ observation

- Claim and confirmations (subset of HDM Coll.):
  - H.V Klapdor-Kleingrothaus et al., Phys. Lett. B 586 (2004) 198
  - H.V Klapdor-Kleingrothaus, International Journal of Modern Physics E 17 (2008) 505-517

 $T_{1/2}^{0\nu} = 2.23_{-0.31}^{+0.44} \times 10^{25} \text{ yr } @6\sigma \text{ in } {}^{76}\text{Ge}$ 

- More recent null results:
  - M. Auger et al. (EXO Coll.) Phys.Rev.Lett 109, 032505 (2012)
  - A. Gando et al. (KamLAND-Zen Coll.) Phys.Rev.Lett. 110 062502 (2013)
  - M. Agostini et al. (GERDA Coll.) Phys.Rev.Lett 111 22503 (2013)

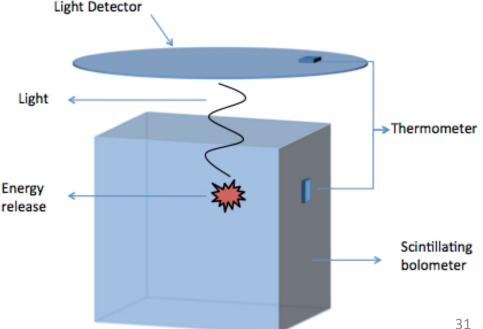
 $T_{1/2}^{0\nu} > 3.0 \times 10^{25}$  yr @90%C.L. (combination)

# Scintillating Bolometers

- Scintillating crystal + bolometric light detector
   → "heat" + "light"
- heat signal is ~ the same for all particles
- light signal depends on particle mass  $\rightarrow$  PID
- Separate  $\alpha$  (bkg. only) from  $\beta/\gamma$  (bkg. and signal)  $\rightarrow$  remove all  $\alpha$  bkg.

#### New technique

For scintillating bolometers the time shape of heat depends on particle type (light emission) → signal shape may be used without light detector



# Zero background

In the 0-background limit, the sensitivity formula becomes:

$$T_{1/2}^{0\nu} \propto (i.a.) \cdot \varepsilon \cdot M \cdot t$$

- does not depend on background and energy resolution
- scales linearly with mass and live-time

# Sensitivity to Dark Matter

- Low energy analisys ≤ 30 keV
- Dedicated trigger and pulseshape ID to lower the energy threshold
- Bolometers unable to distinguish nuclear recoils from  $\beta,\gamma$
- In principle CUORE could look for annual modulation: B+S(June+3)/B+S(Dec+3)
- Simulation of sensitivity assuming all DM is made of WIMP with spin-independent interactions

