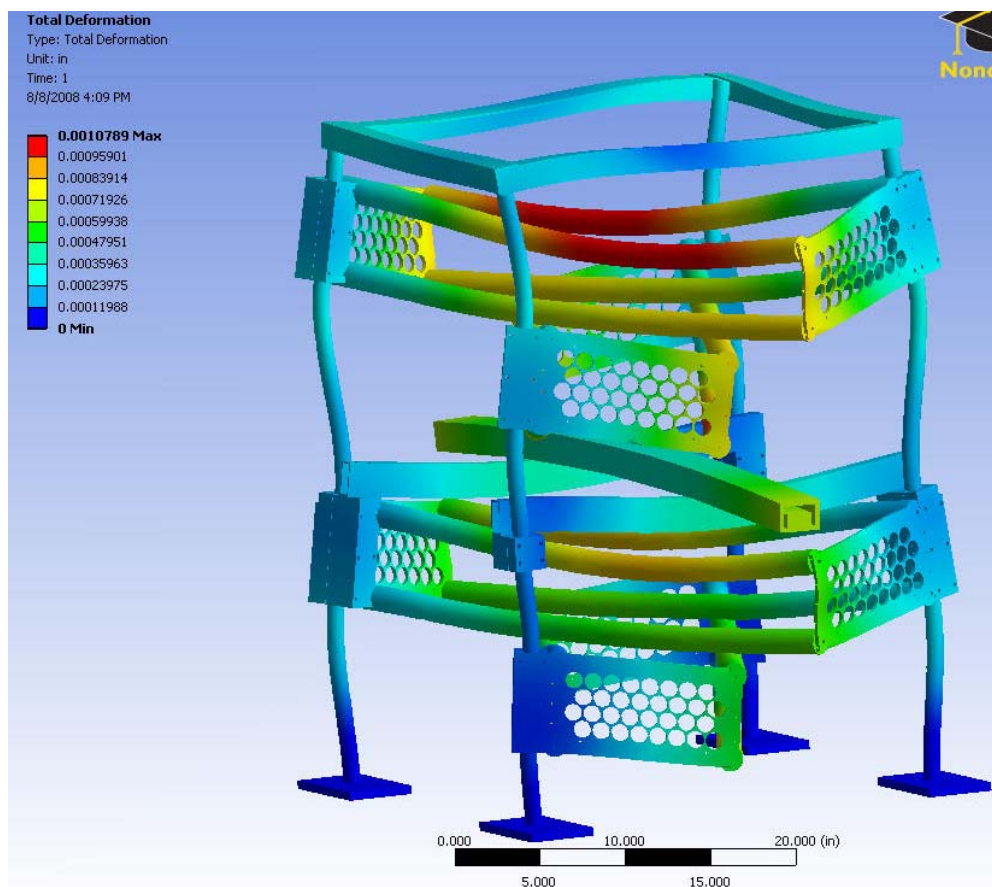
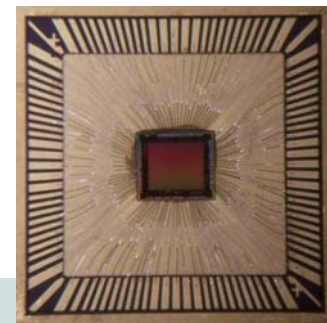
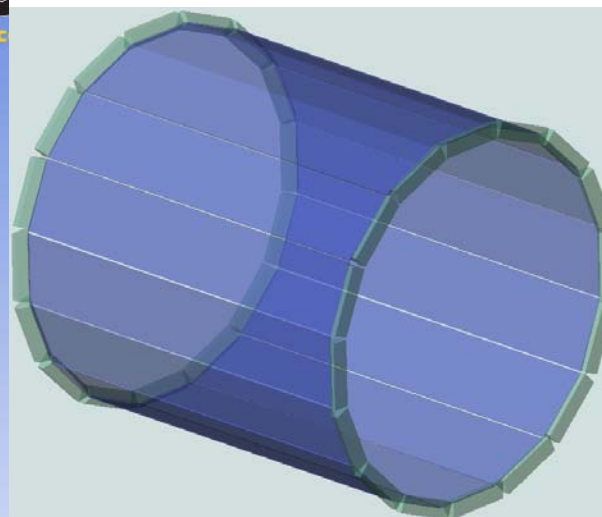


Updates from Hawaii



Larry Ruckman and Gary S. Varner



Tom Browder, Bryce Jacobson, James Kennedy, Kurtis Nishimura, Marc Rosen, Wen Yen, Andrew Wong

UNIVERSITY OF HAWAII AT
MANOA

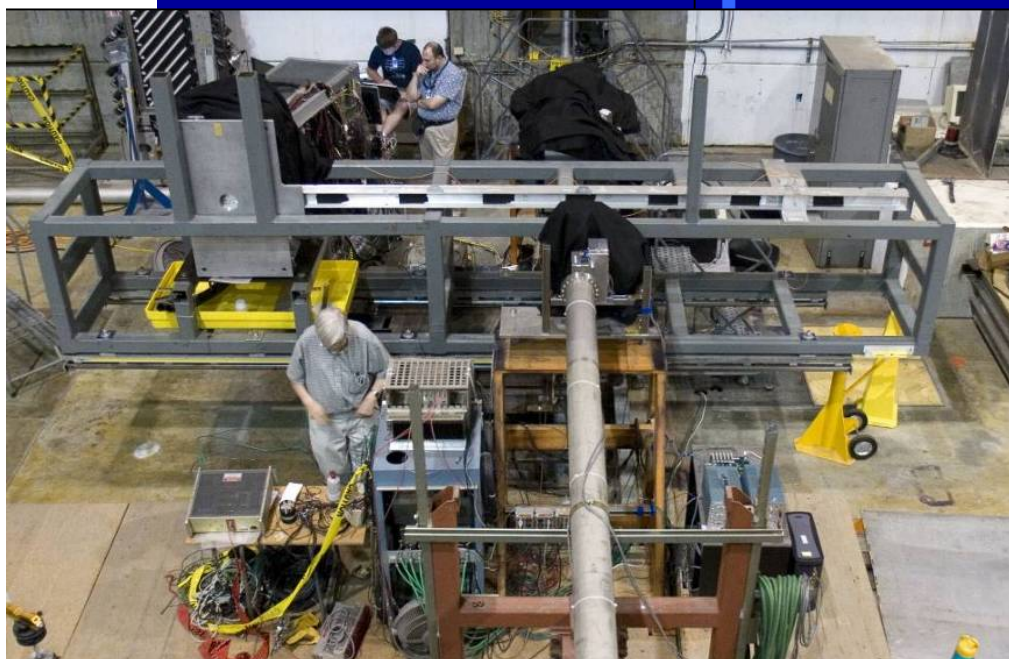
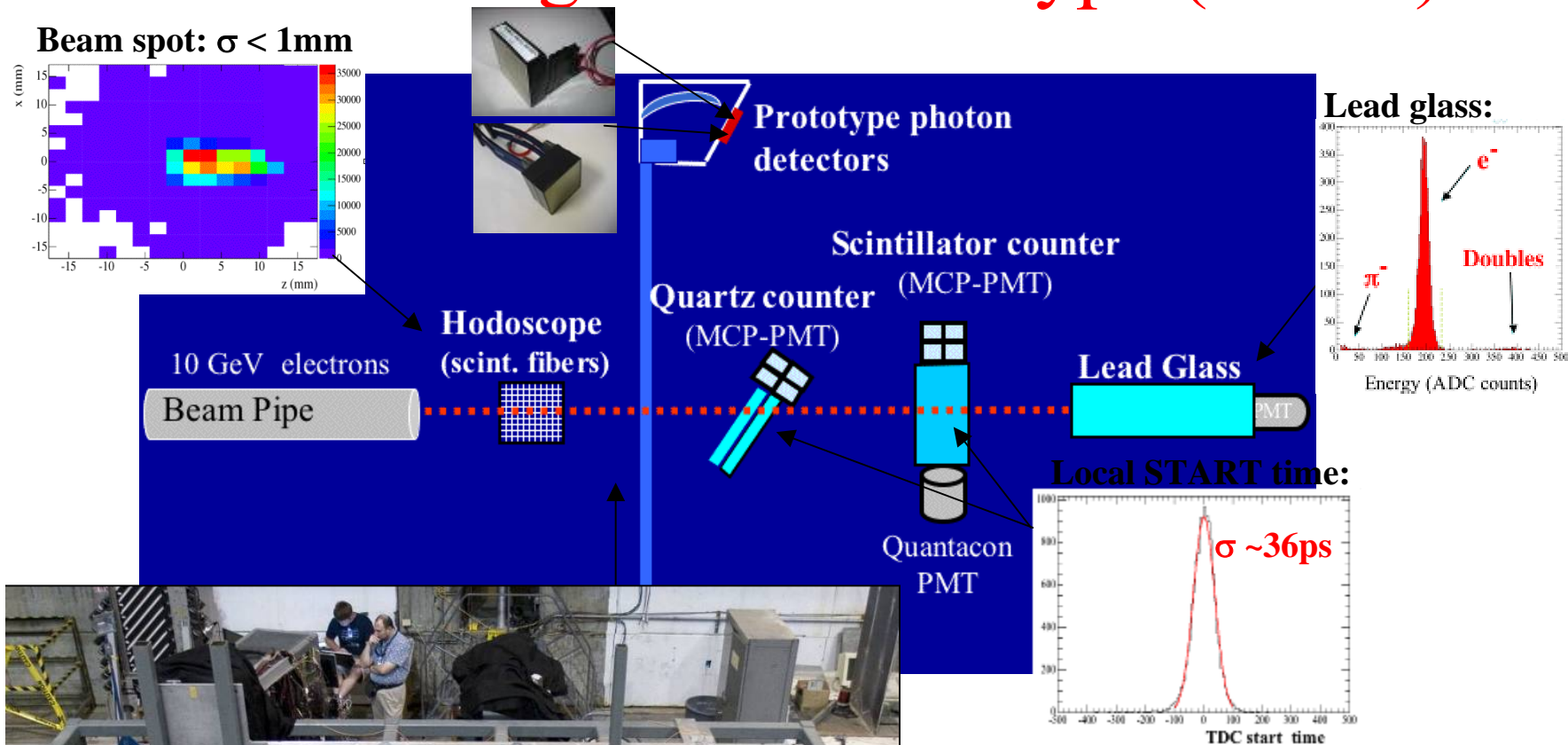


Super-B PID meeting 17-SEP-08

Today's Topics

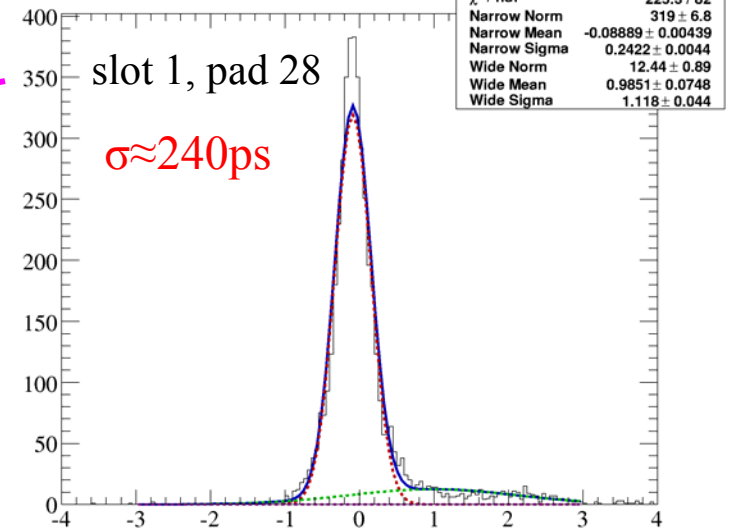
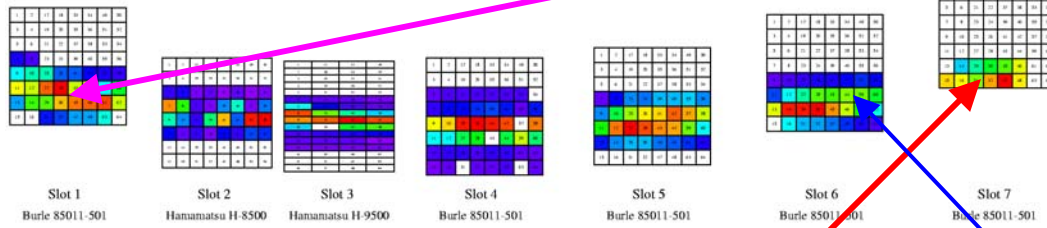
- Further studies with Fast Focusing-DIRC
 - Enhanced (waveform sampling) readout
 - Cosmic Ray running with new readout
- Readout ASIC Update
 - Initial BLAB2 Performance
 - Plans for full system
- Quartz bar $t, \{x, y\}$ studies (Hawaii cosmic)
 - Cosmic test stand (close GEANT4 loop)
 - Dense imaging plane concept
- Schedule and Plans

Focusing DIRC Prototype (T-492)

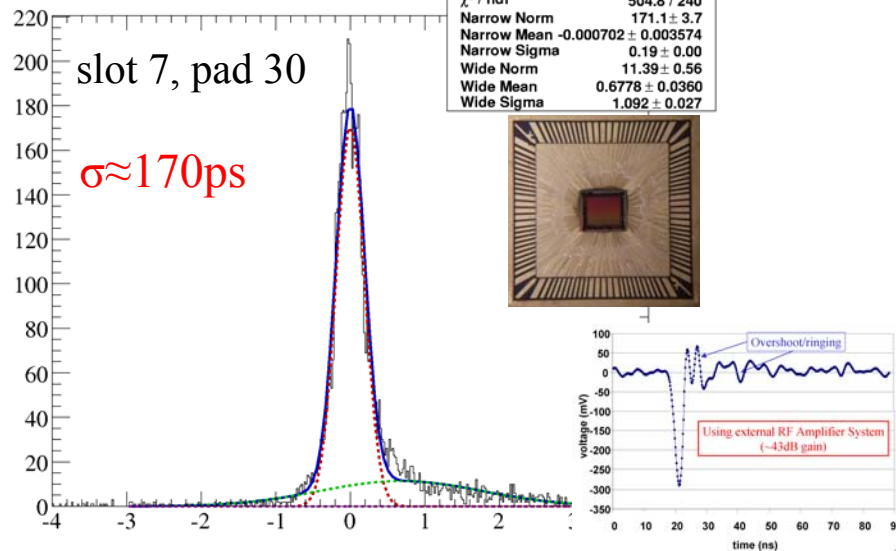


Test results: timing slot 7, pad 15
to Philips slot 1&6
for run 27, pos 1, direct photons

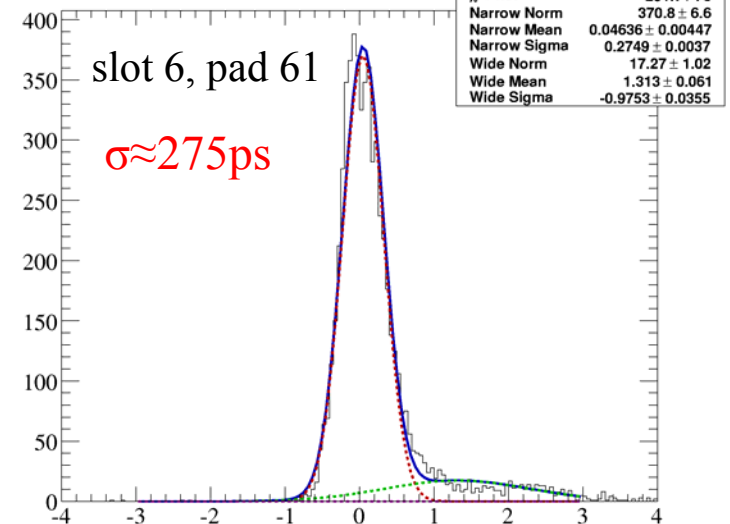
(symmetry partner in hit plane)



New BLAB-based Readout



(close neighbor in hit plane)

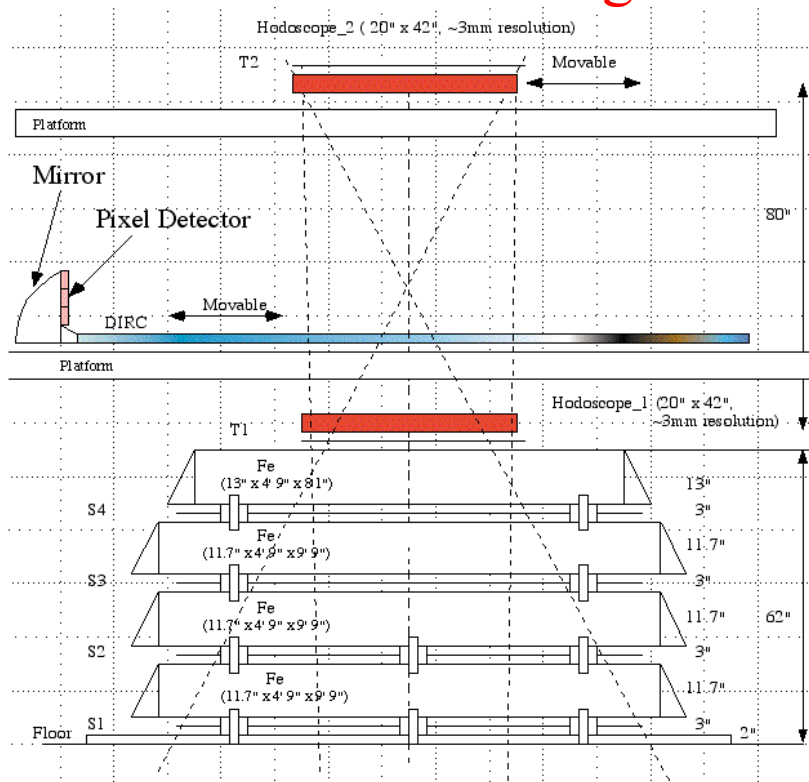


Decided to upgrade all channels to new BLAB electronics

delta(time) (ns)

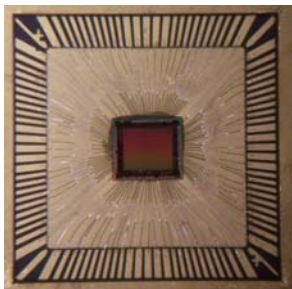
Moving to Bldg. 121

- LCLS Operations
 - Parasitic running possible, but
 - Rad safety system in ESA
- Move to nice cosmic stand
 - 1 mrad resolution
 - Precision timing and further studies w/ new electronics



1.6GeV/c
P_min
through
range
stack

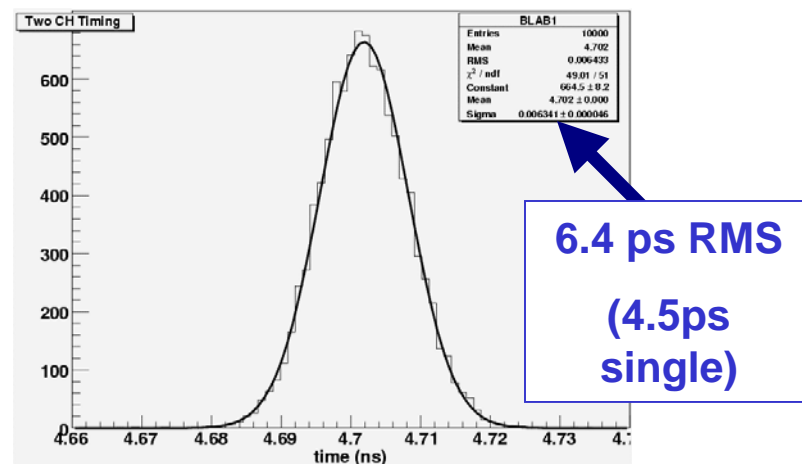
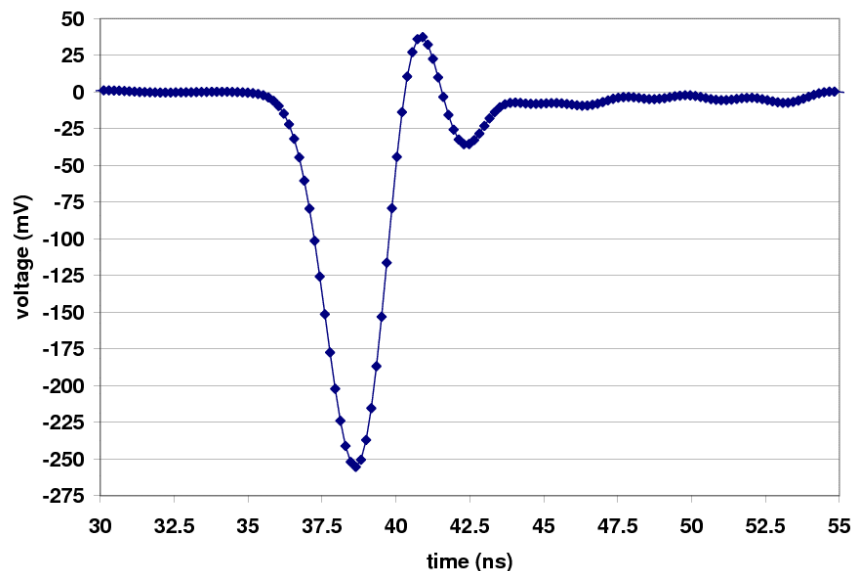
BLAB ASIC further studies



BLAB1 -- NIM
A591 (2008) 534



- Comparable performance to best CFD + HPTDC
- MUCH lower power, no need for huge cable plant!
- Using full samples significantly reduces the impact of noise
- Photodetector limited



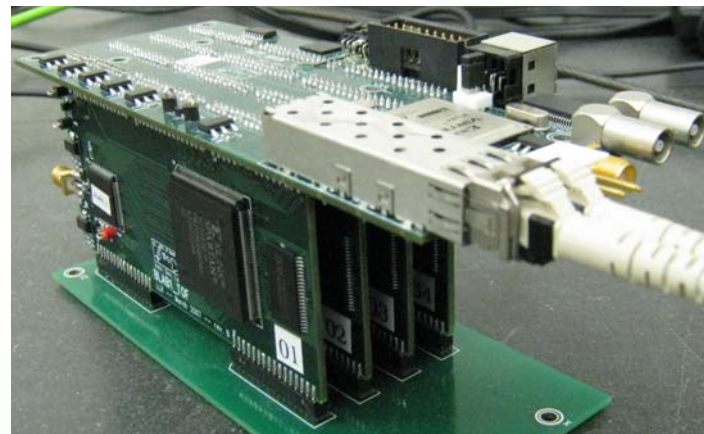
Submitted NIM, arXiv:0805.2225 6

Highly Integrated Readout

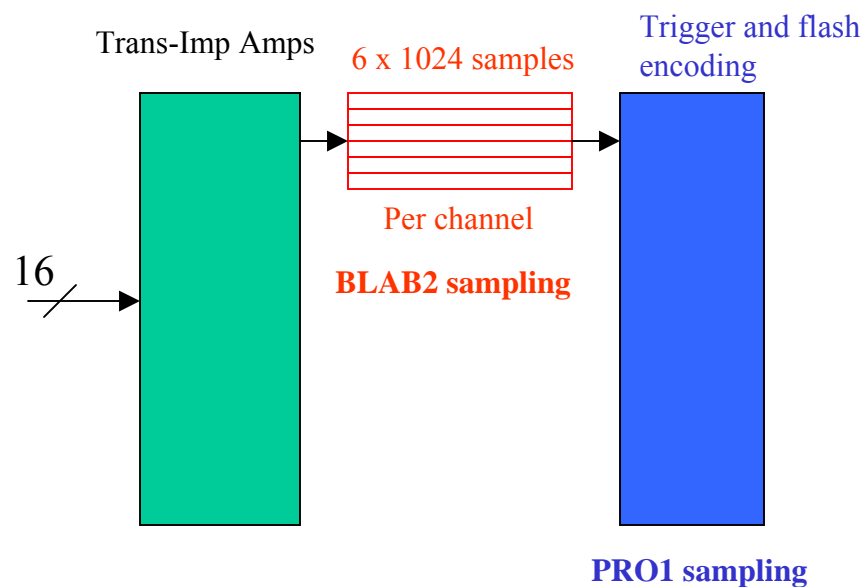
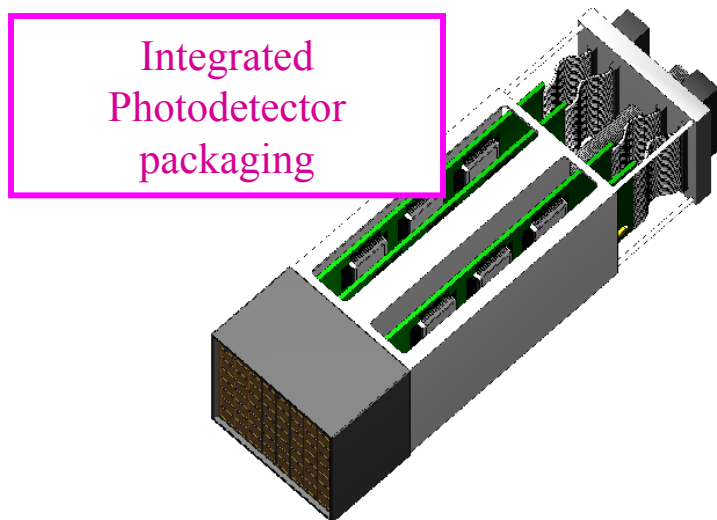
- **Buffered LABRADOR**

TABLE II: *BLAB2 ASIC Specifications.*

Item	Value
Photodetector Input Channels	16
Linear sampling arrays/channel	2 6
Storage cells/linear array	512 1024
Sampling speed (Giga-samples/s)	2.0 - 10.0
Outputs (Wilkinson)	32

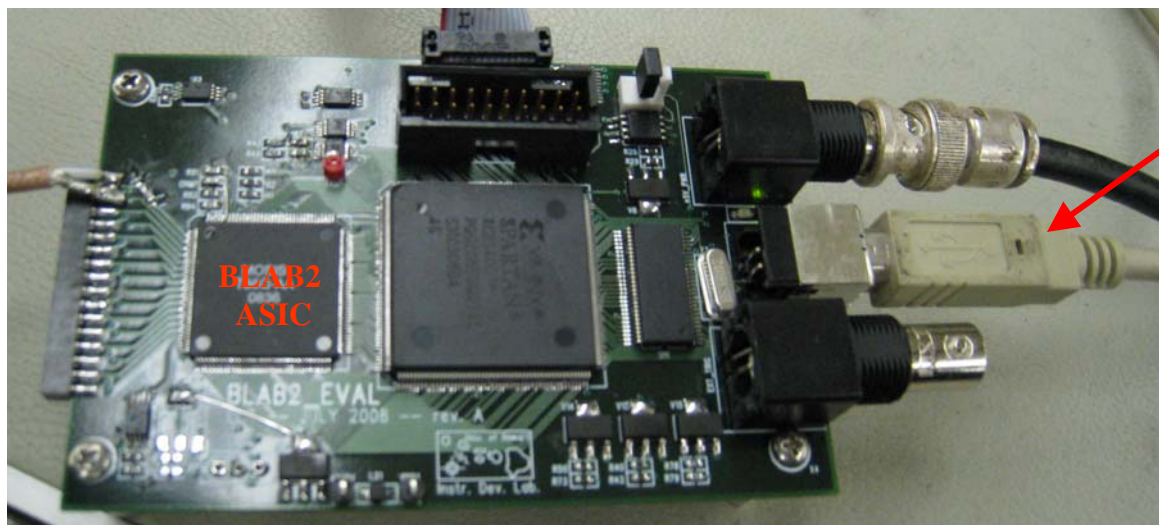
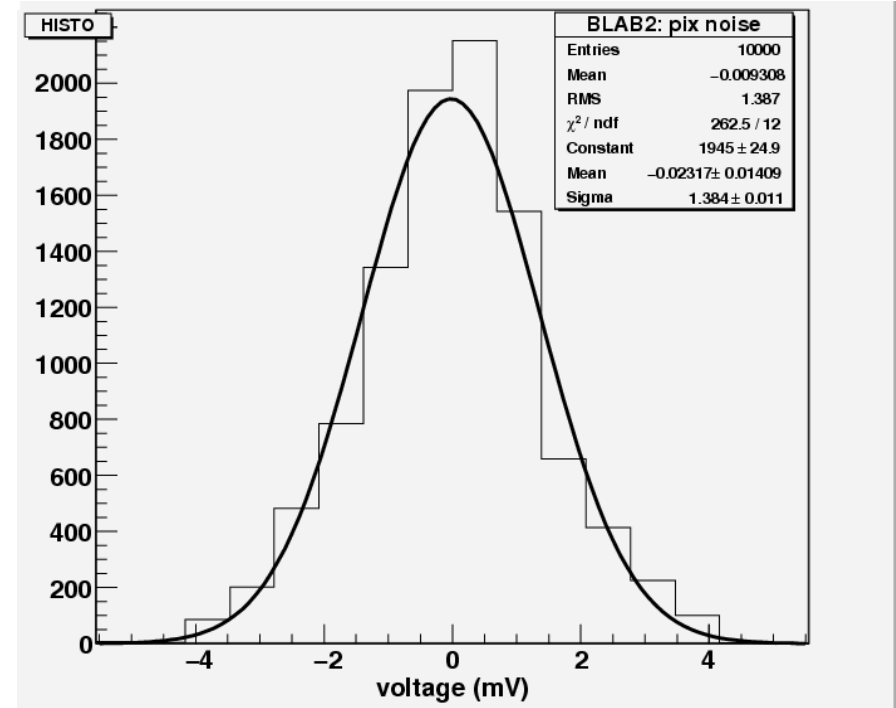
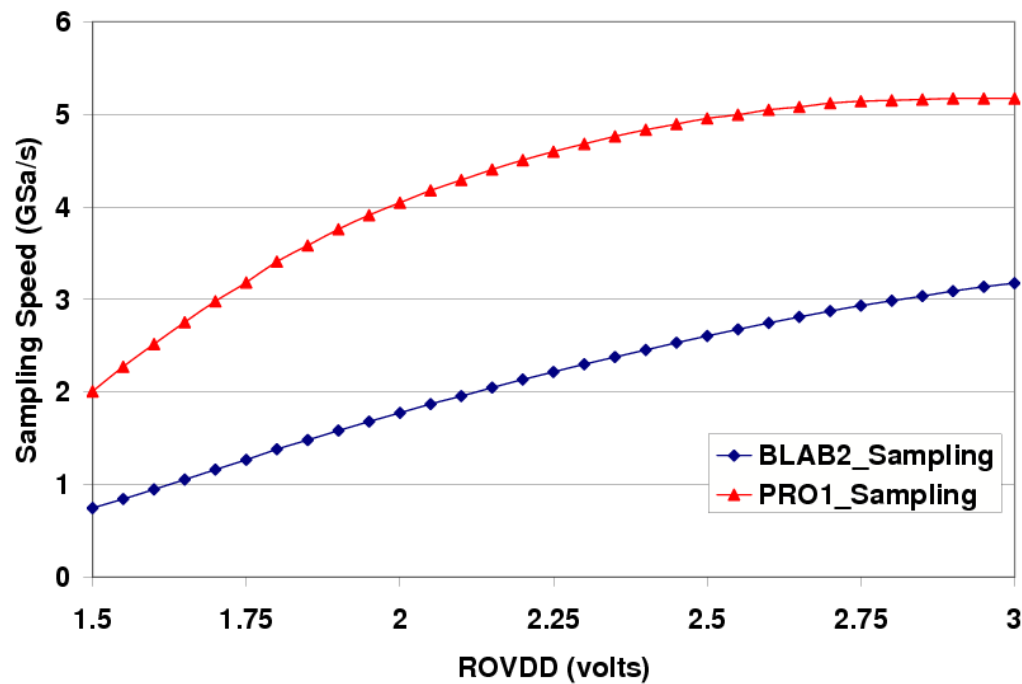


BLAB2 ASIC



BLAB2 submit 6/23, ASICs received last Monday

Sampling Speed and noise



Eval board USB2

- Win XP

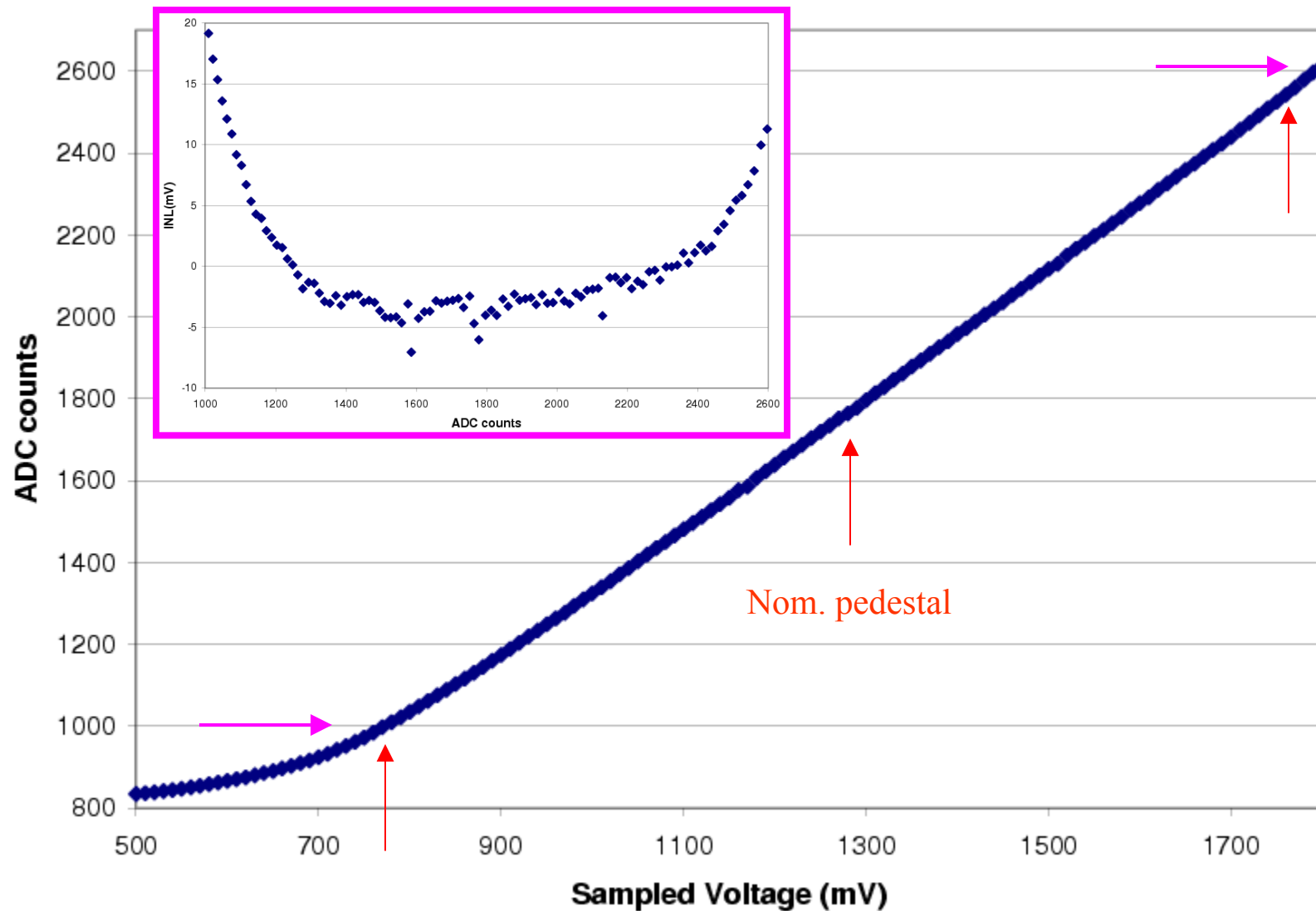
- Linux

- Mac OS-X

80 ASICs total
(1280 channels)
[448 needed f-fDIRC]

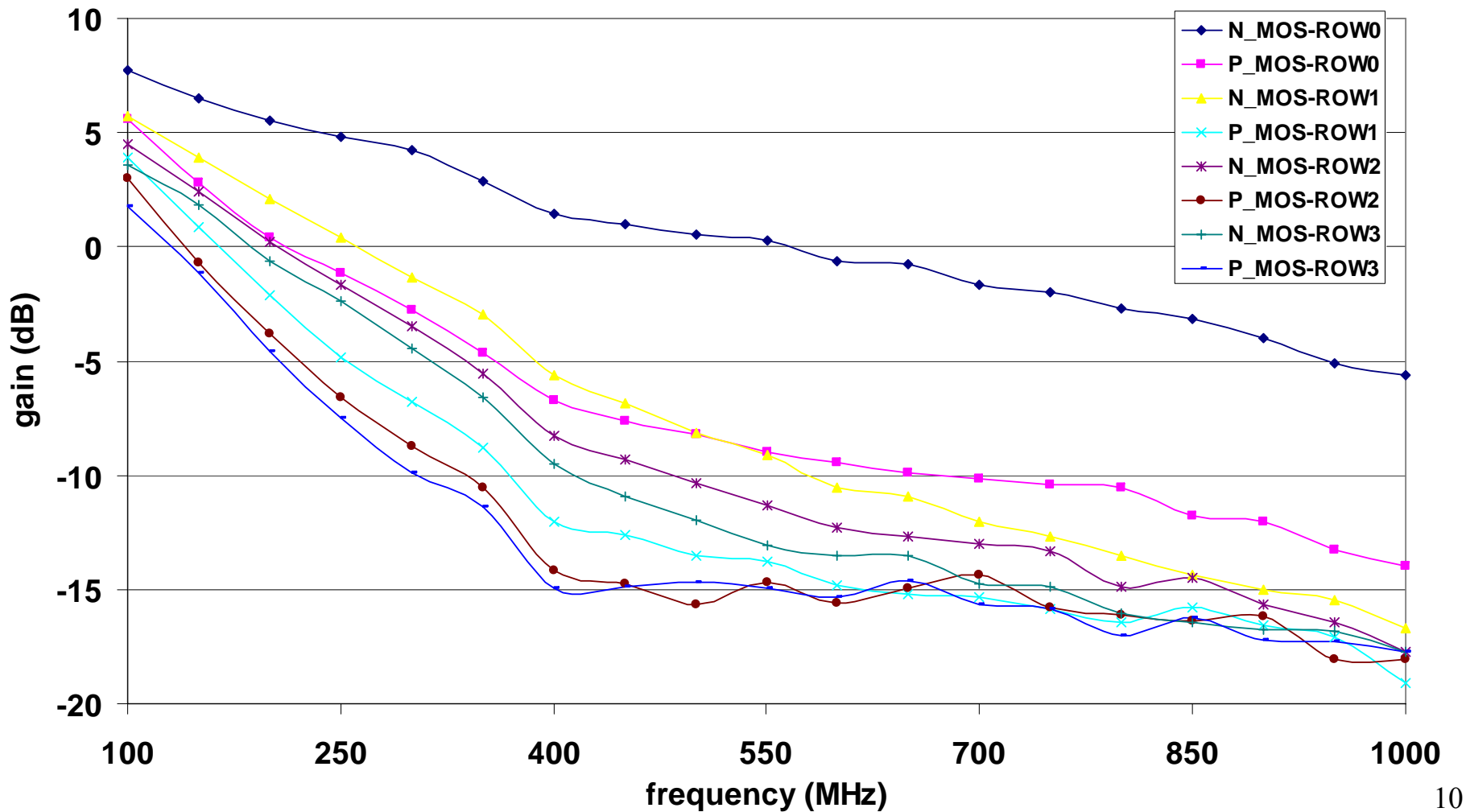
Linearity

In practice, correct with LUT so INL ~ 1 -2 mV



Frequency Response (with TIA)

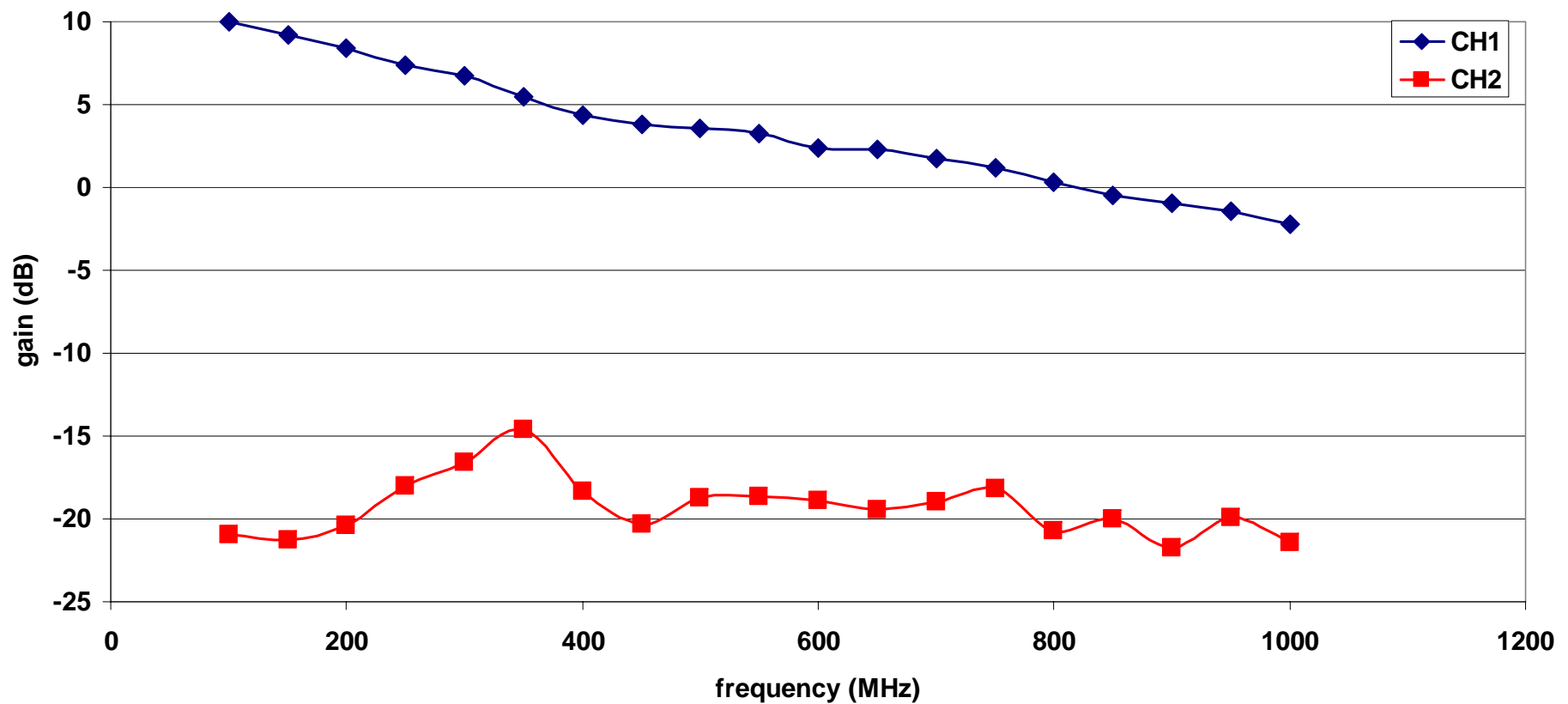
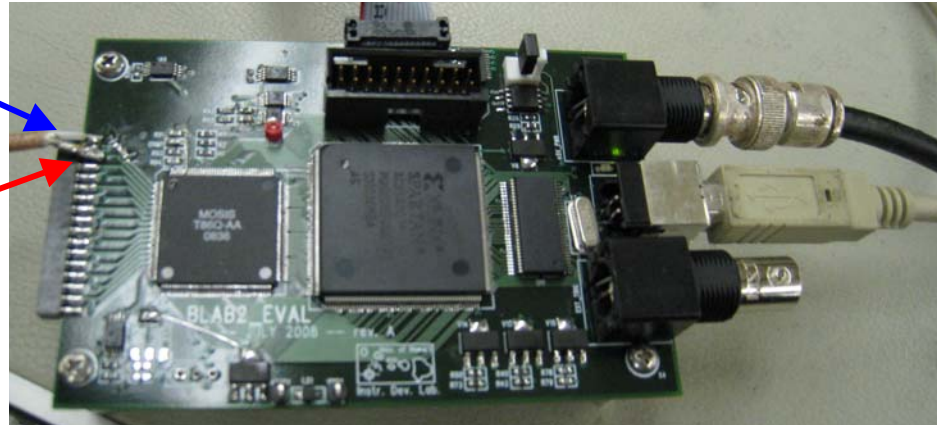
Each channel has 6 storage rows of 1024 samples
(512 NMOS and 512 PMOS interleaved)



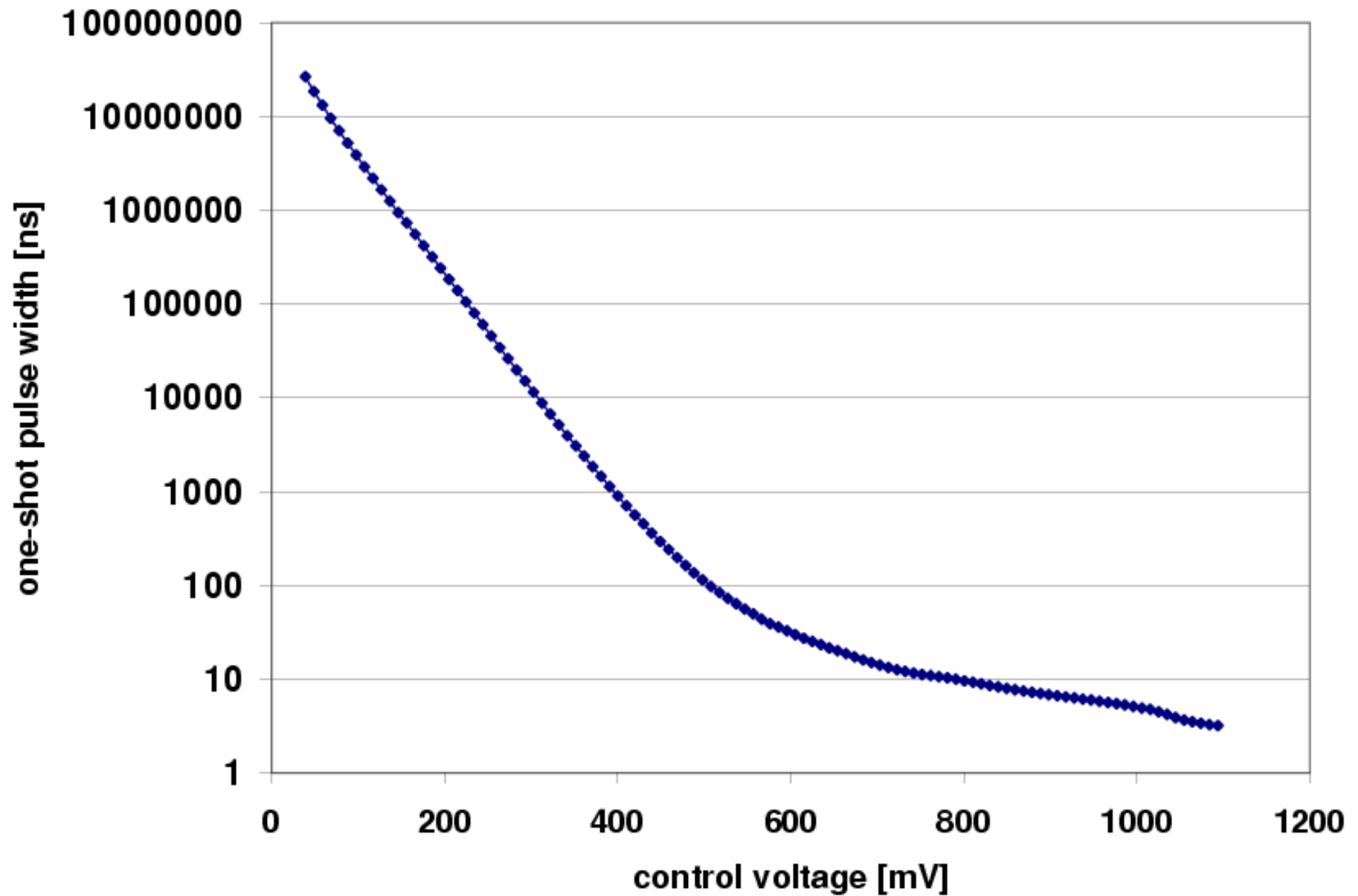
Measured cross-talk

Channel 1

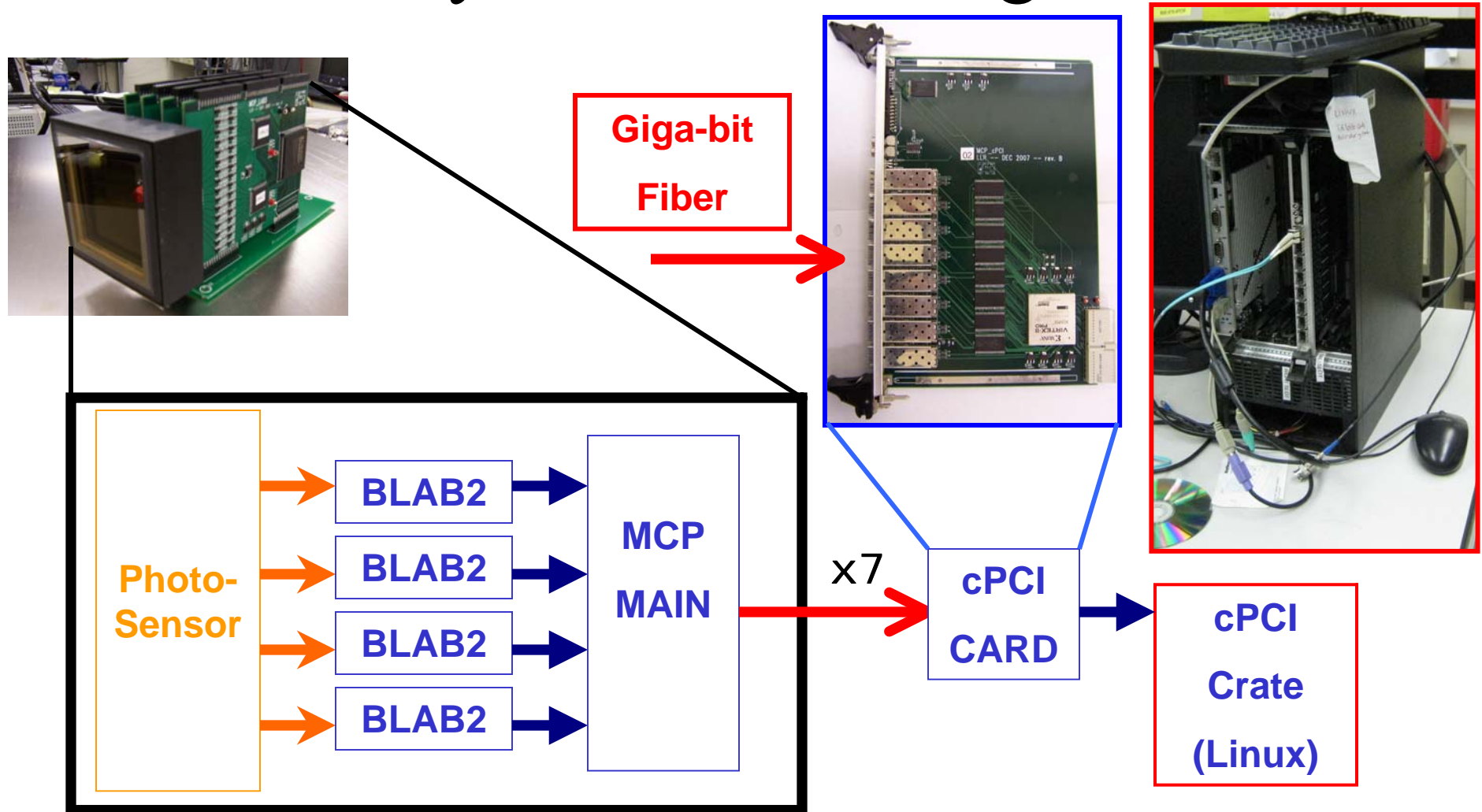
Channel 2



Trigger Performance



Readout System Block Diagram



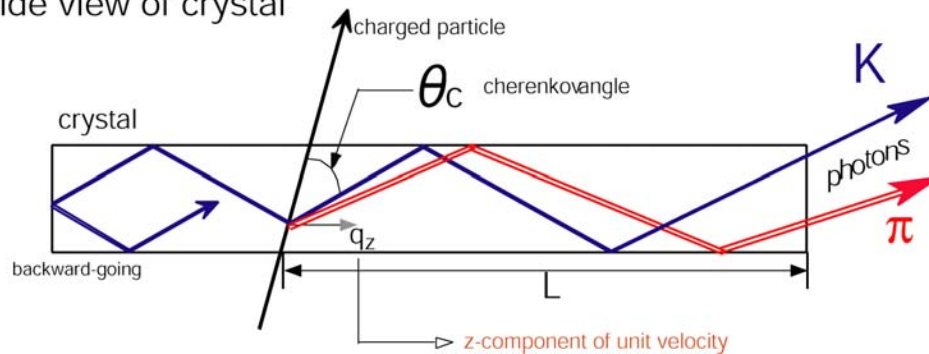
- Up to 7x64 channels per cPCI card
- Very portable DAQ
- Up to 32,256 channels/cPCI crate

Very cost effective, board hardware already exists, firmware/software dev.

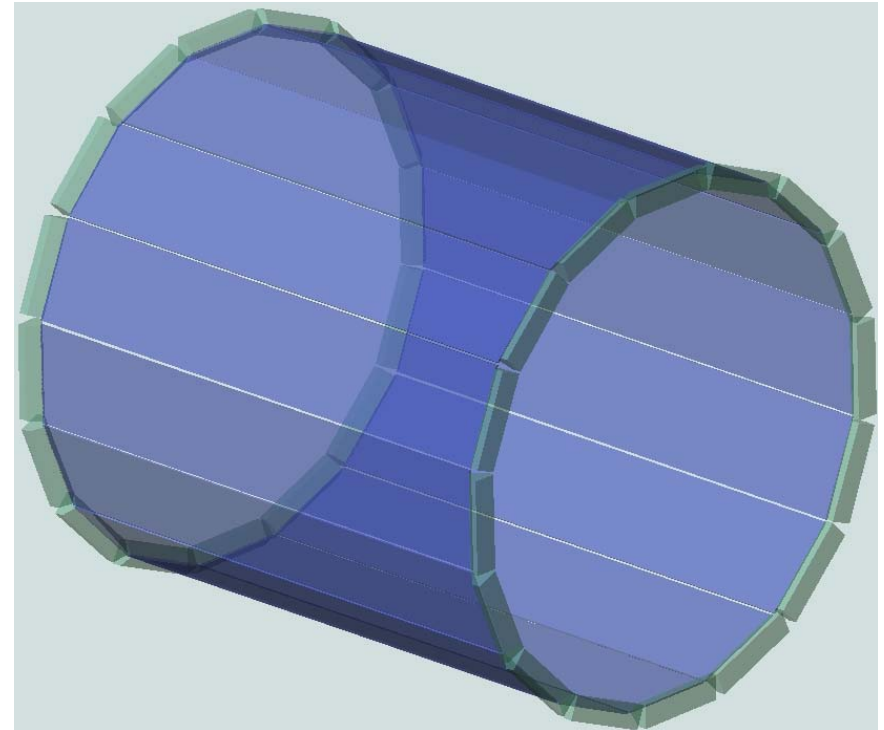
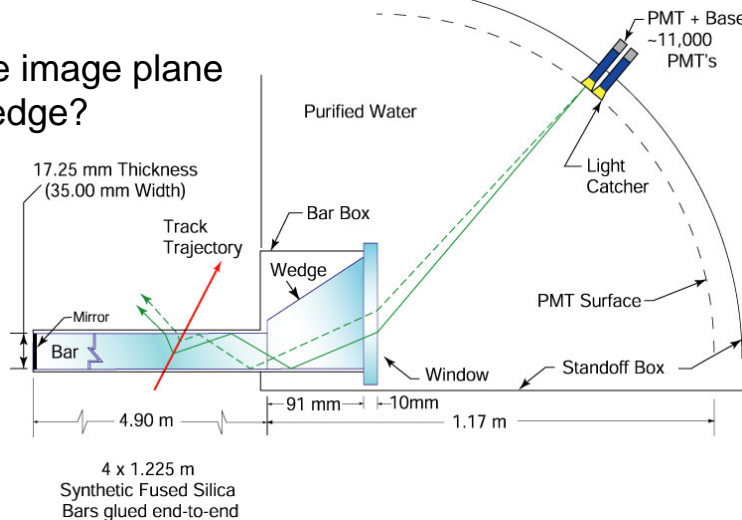
Improved readout timing

Concept: Use best of both ToF + TOP (timing) and DIRC and fit inside detector (background)

Side view of crystal

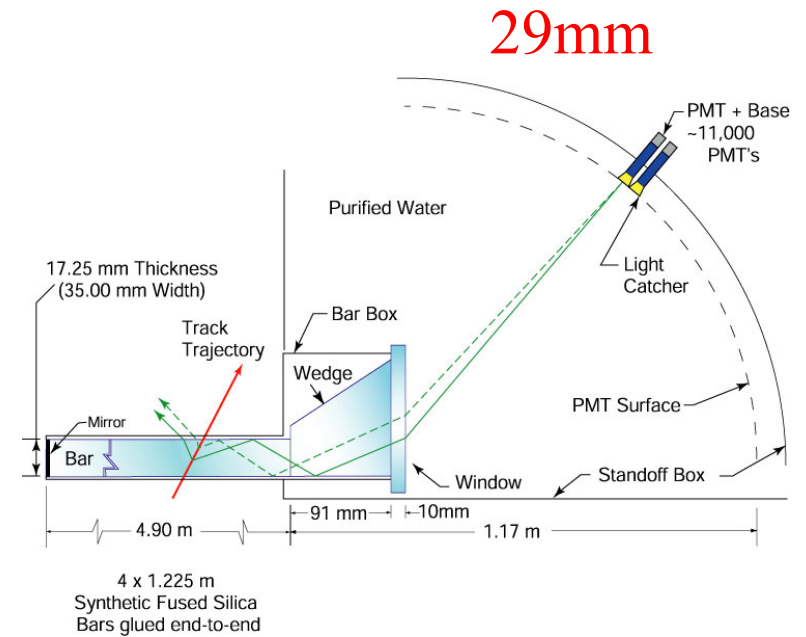
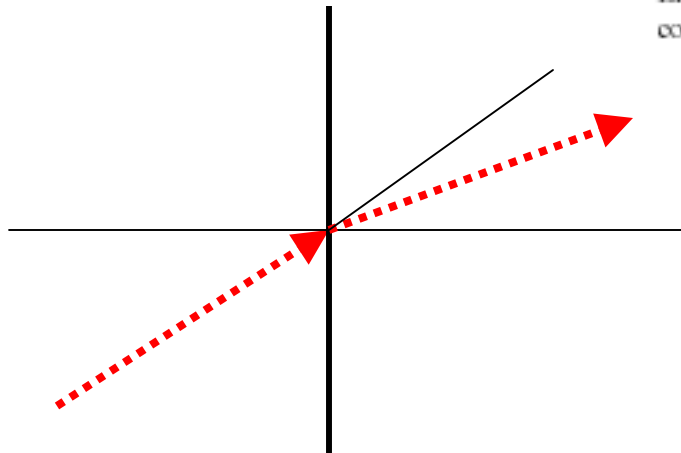
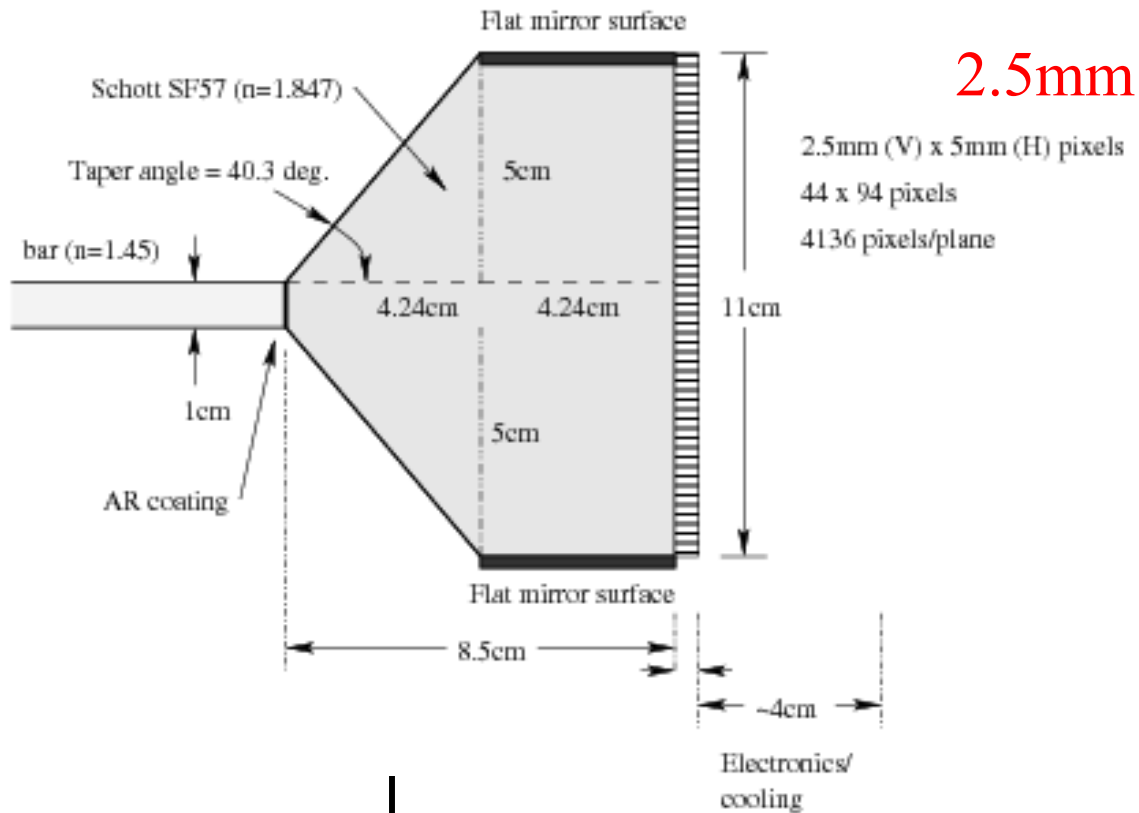


Scale image plane to wedge?

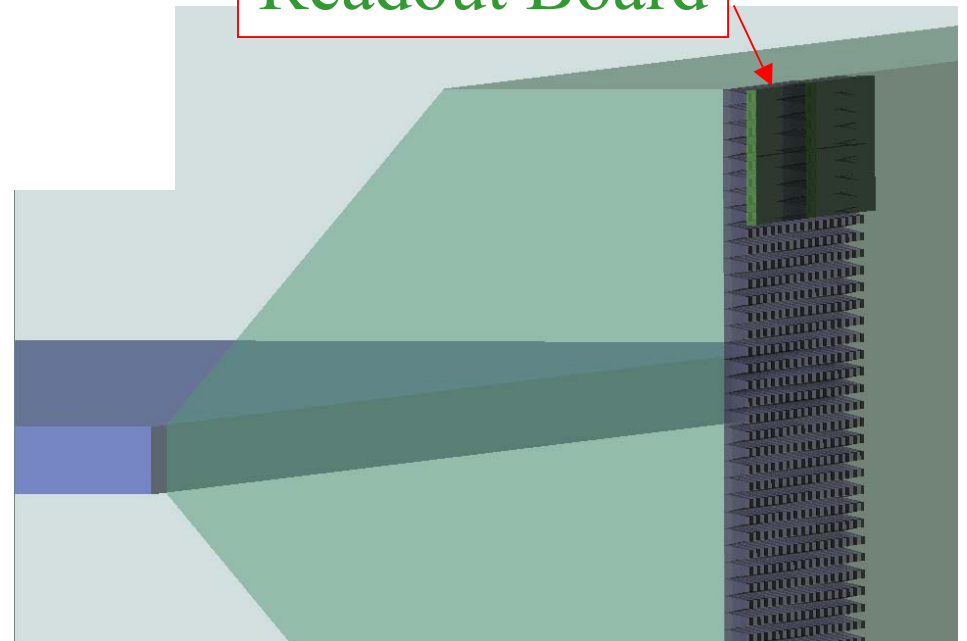


- Use new, compact solid-state photon detectors, new high-density electronics
- Use simultaneous T , θ_c [measured-predicted] for maximum K/π separation
- Keep pixel scale comparable to DIRC

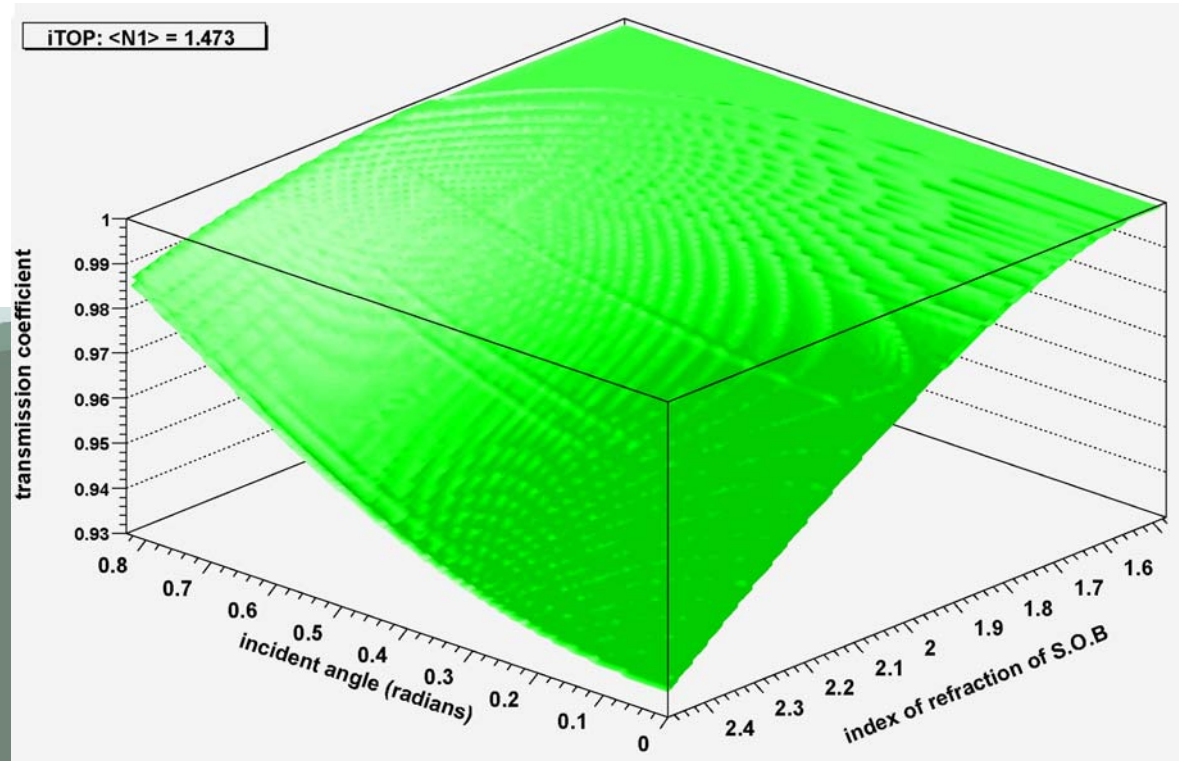
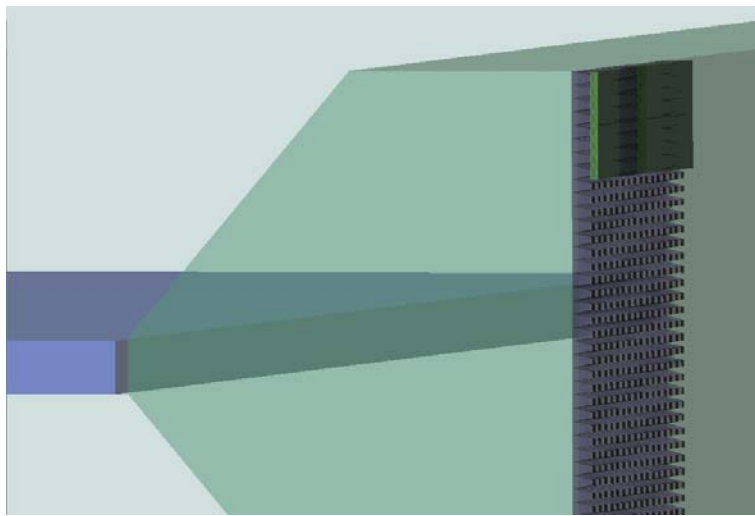
(too) Simple initial idea



Readout Board



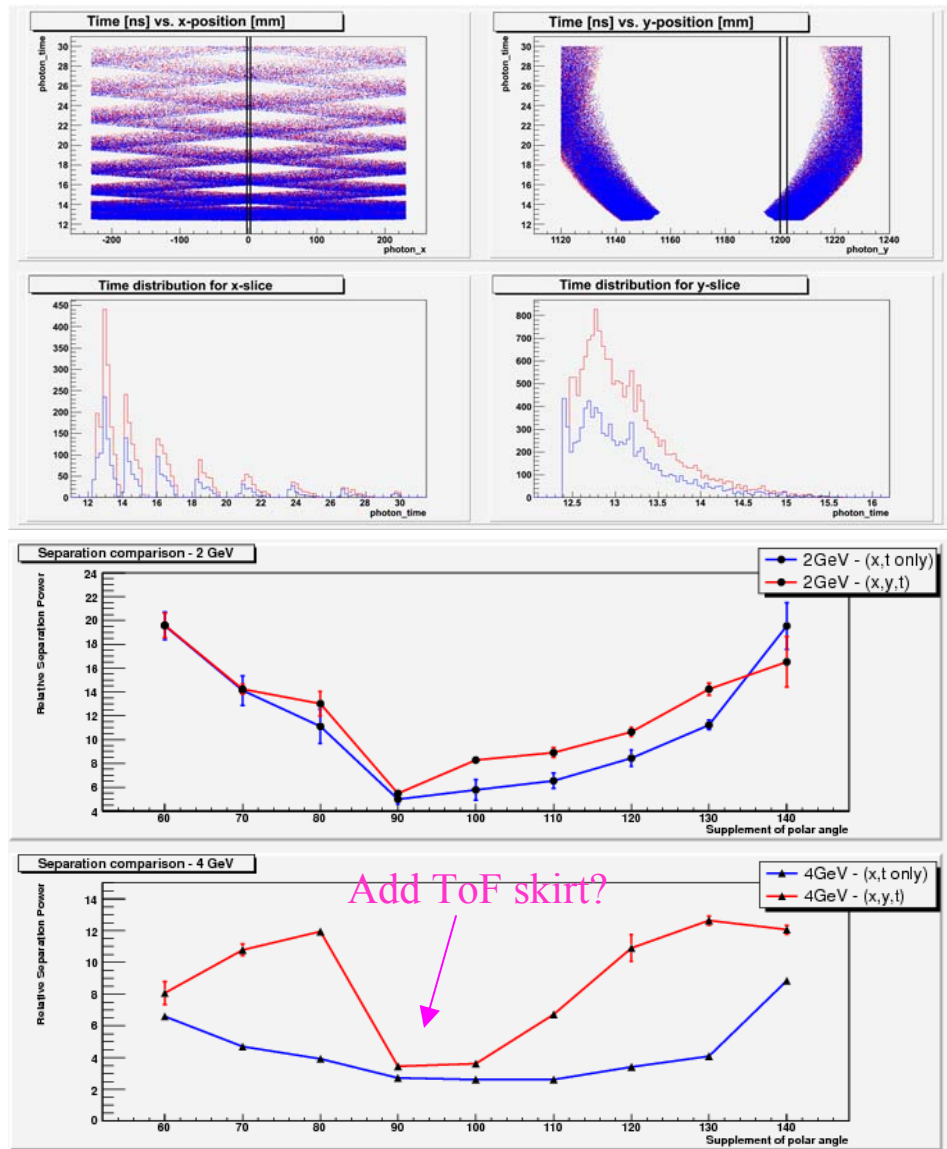
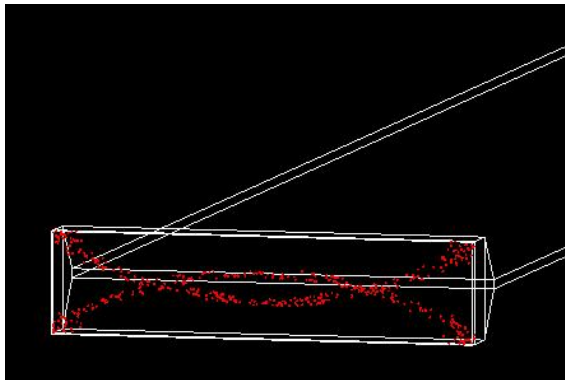
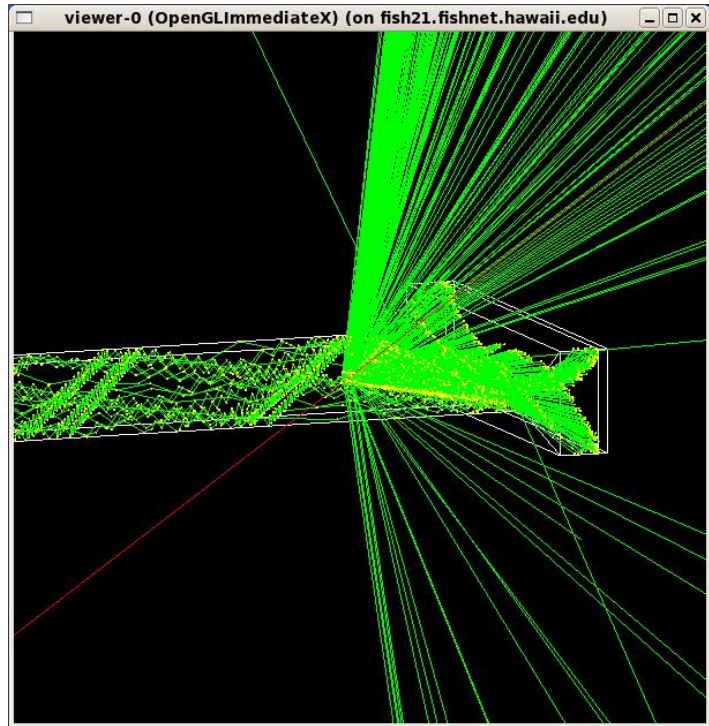
Stand-Off Block (SOB) Coupling



44 x 92 pix/plane = 4048 channels
16 bars x 2 ends x 4048 = ~130k channels

Problem is, once get in, hard to get out...
Exploring a number of concepts

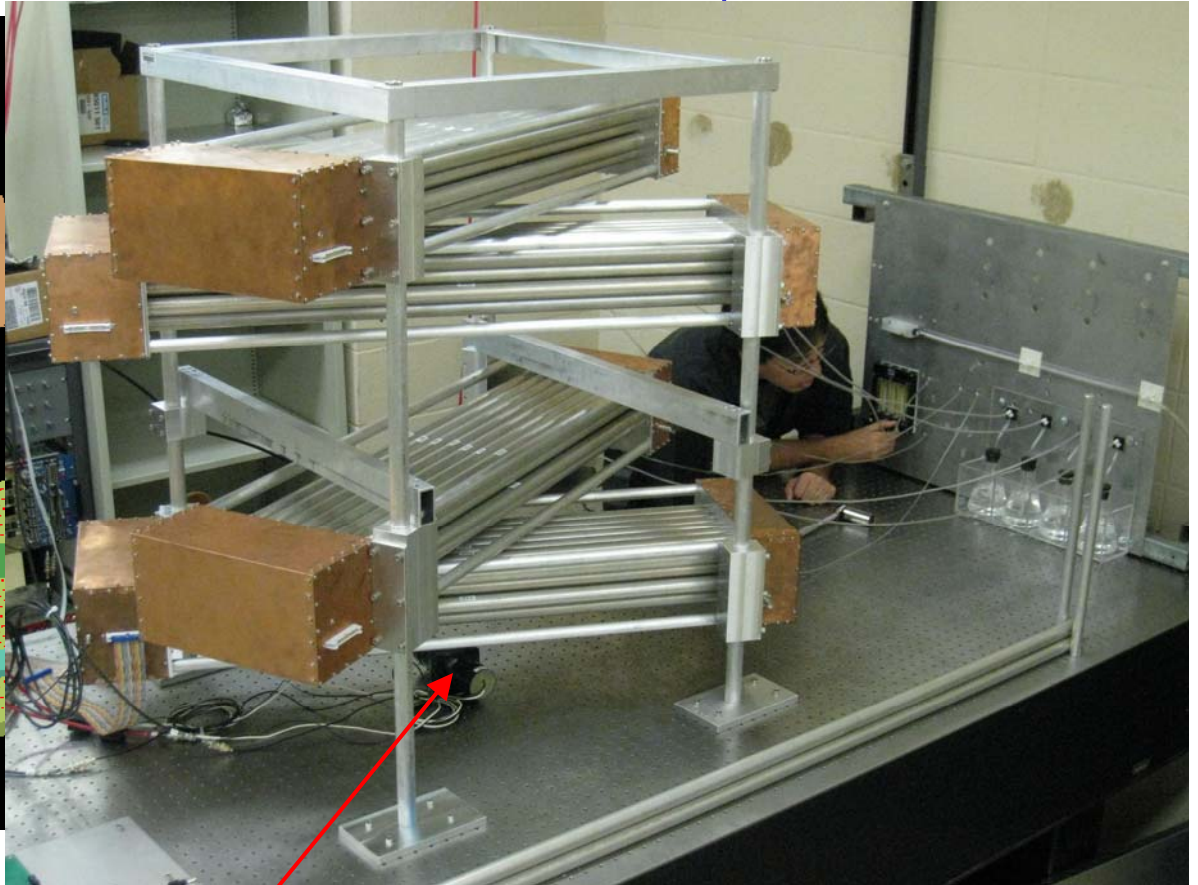
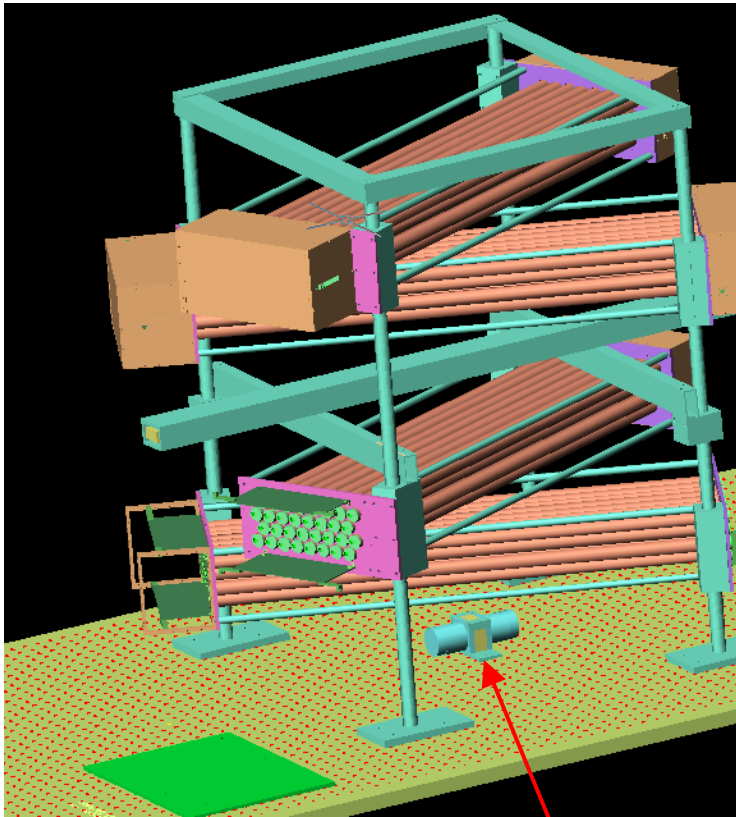
GEANT4 Simulation Kurtis Nishimura/Larry Ruckman



Looks promising, need to test

Cosmic Test Bench

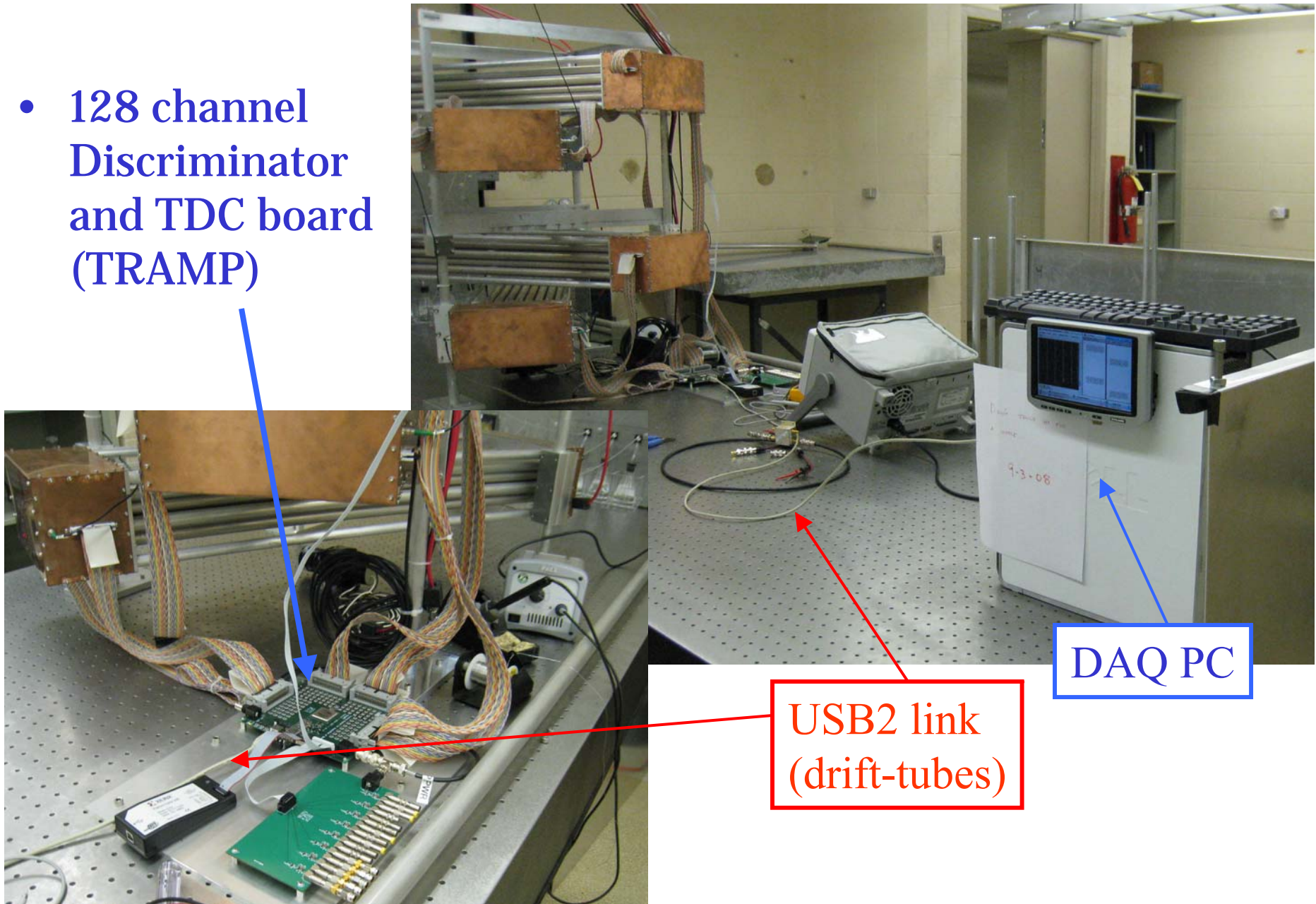
- Quartz bar test bed – 128 drift tube array



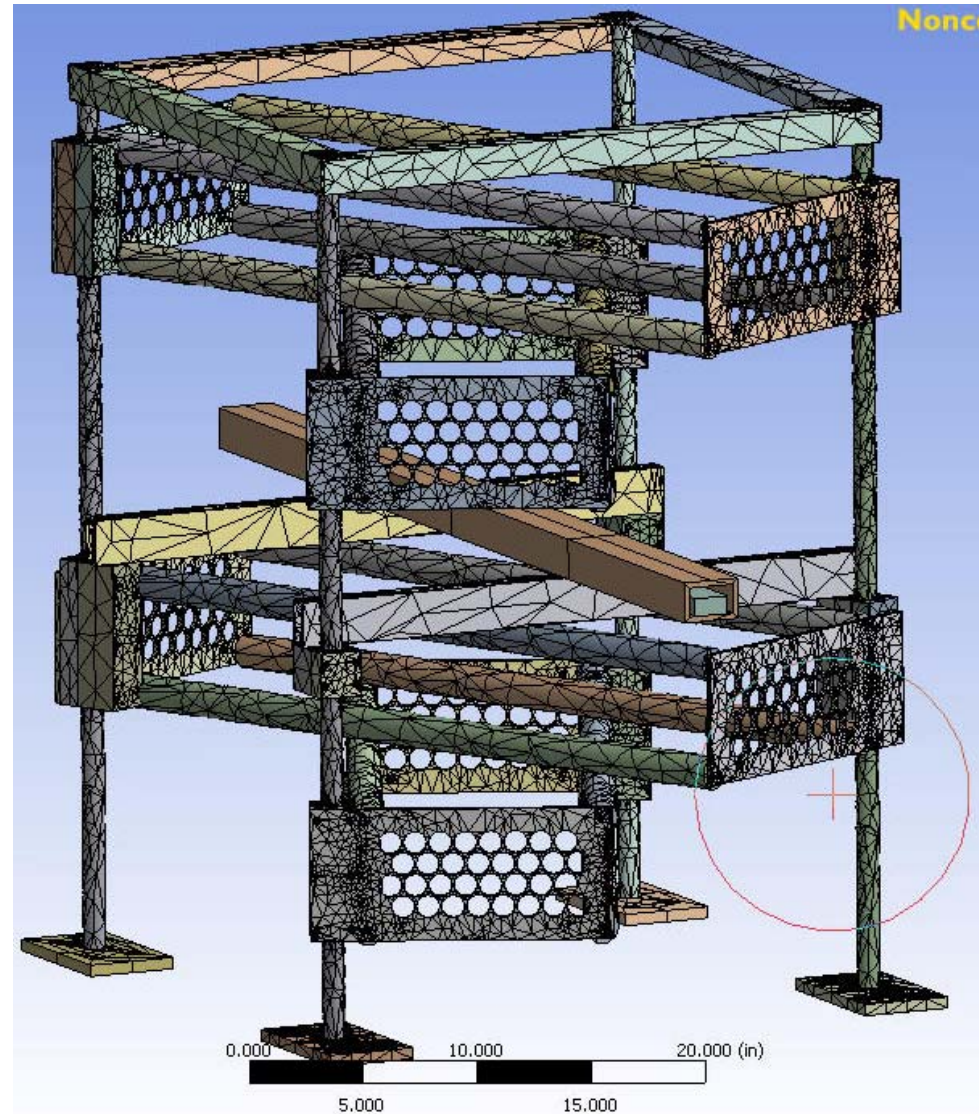
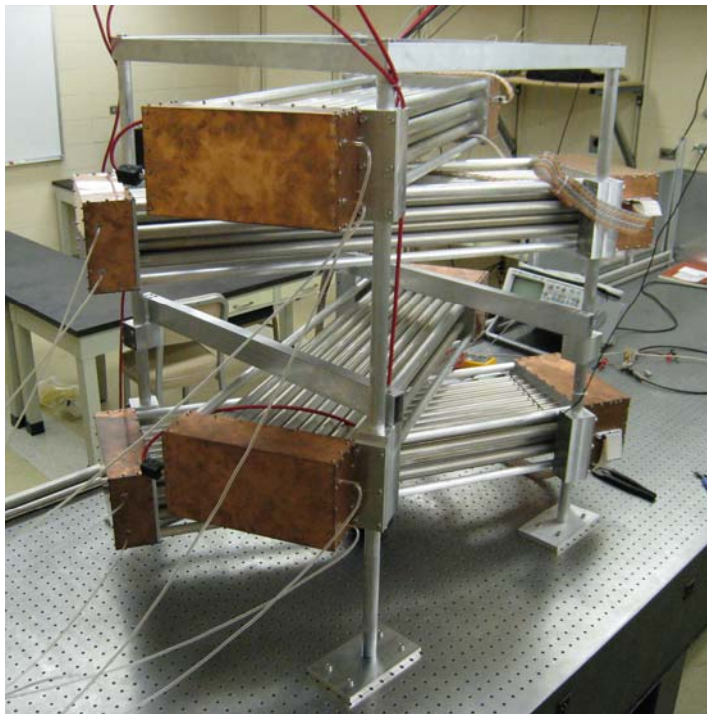
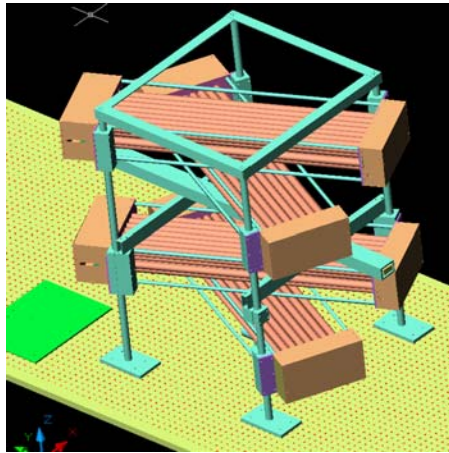
Precision Timing Block ($T=0$)
Radiator viewed by 2x fine-mesh PMTs

Cosmic Test Bench - Infrastructure

- 128 channel Discriminator and TDC board (TRAMP)

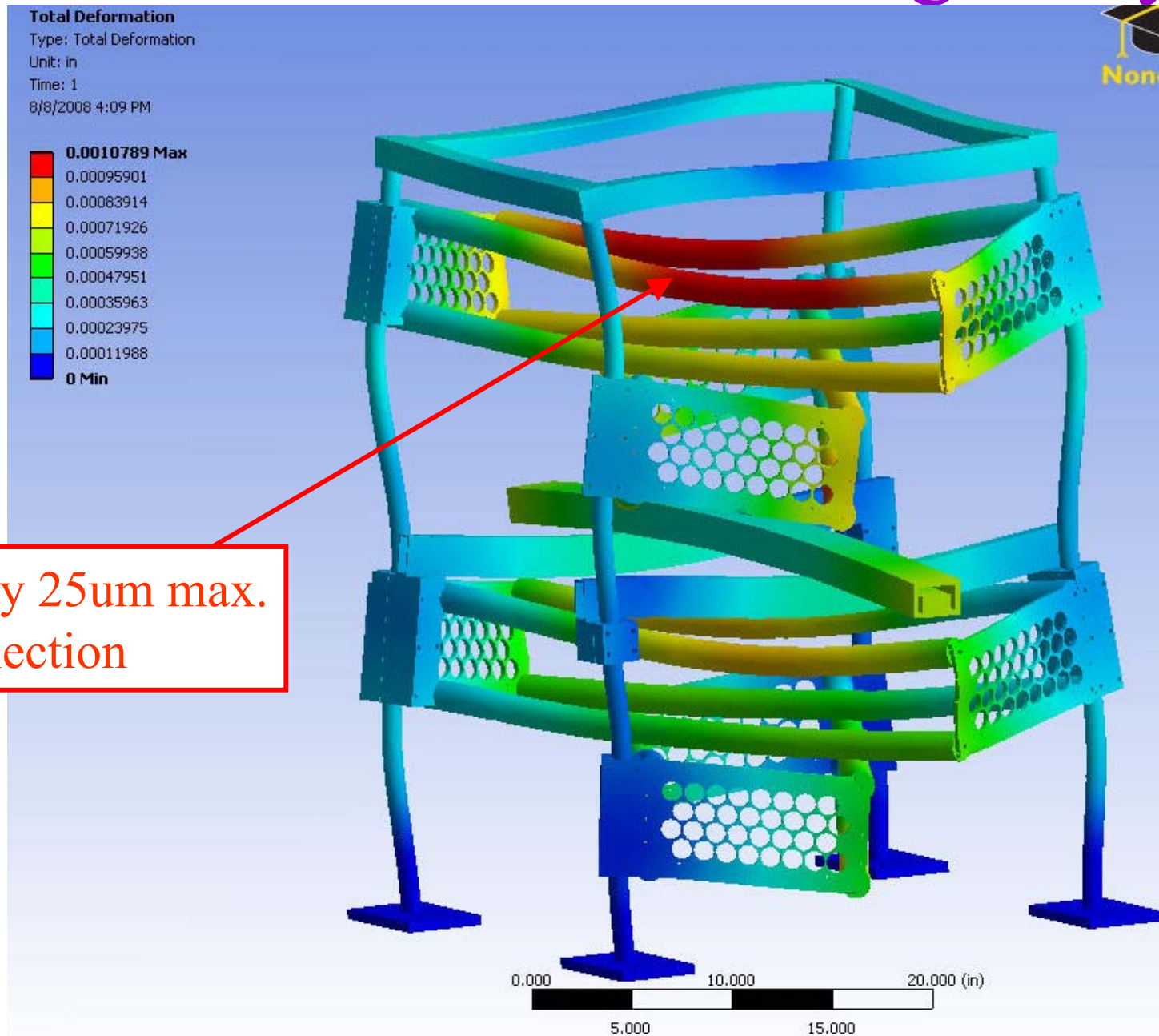


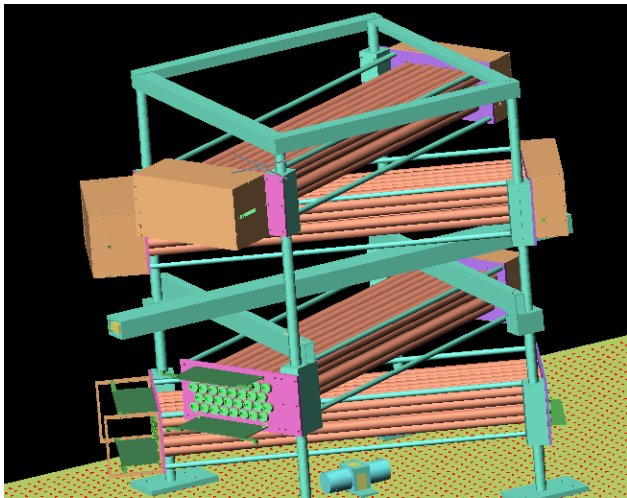
Cosmic Test Bench – mech. Stability



Rigidity and stability are excellent

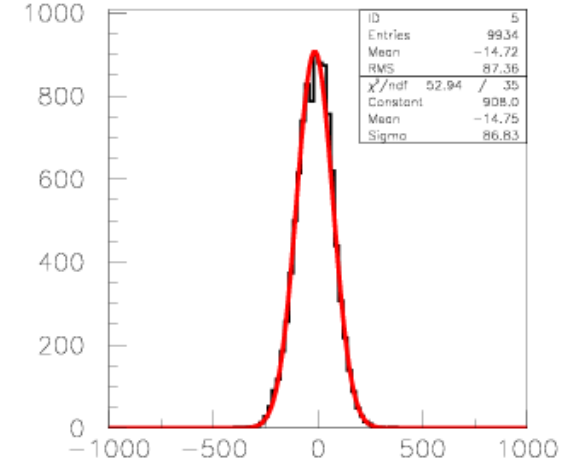
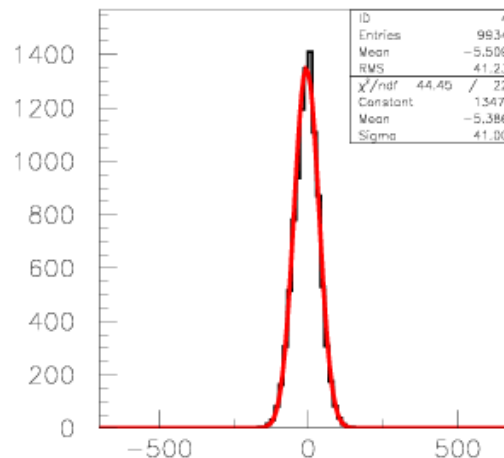
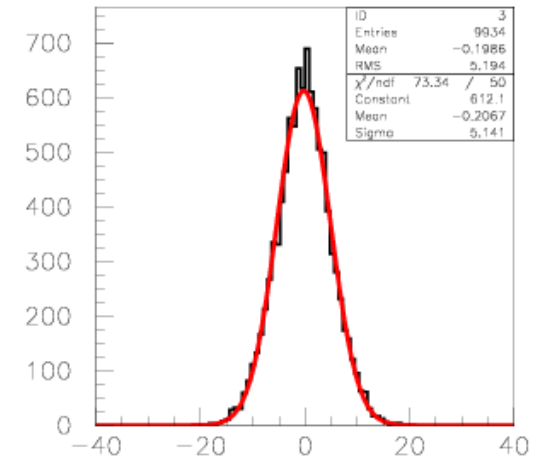
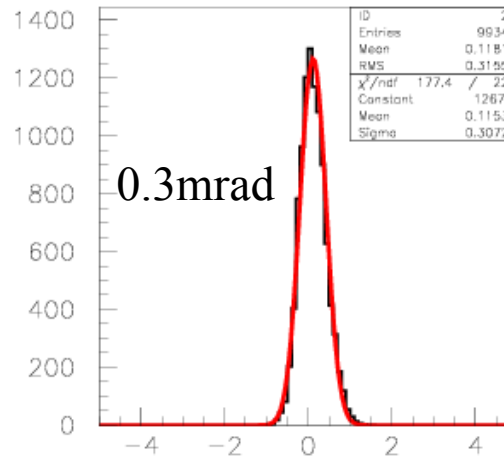
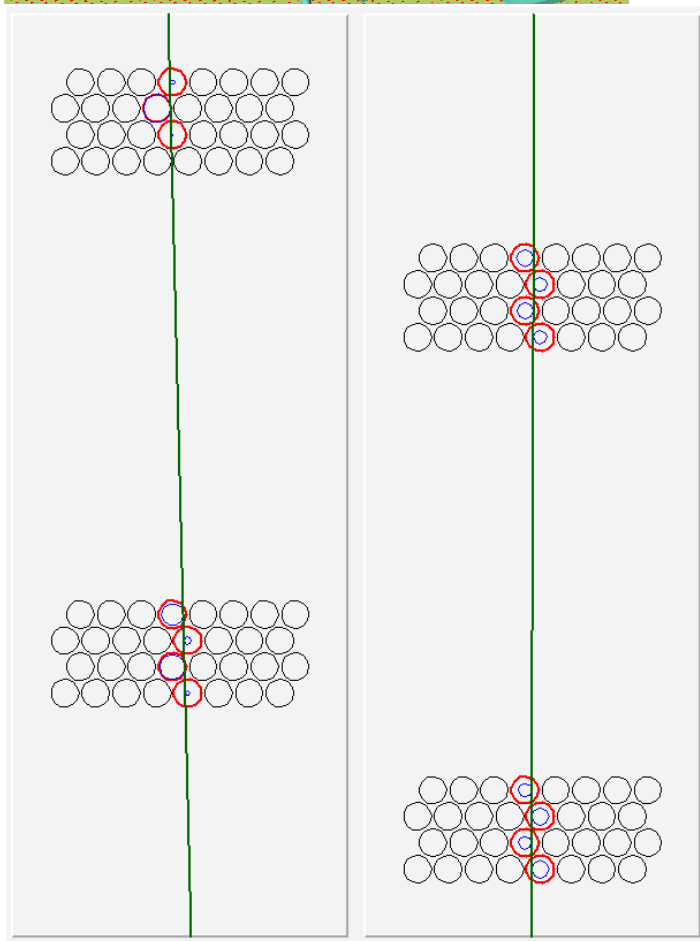
Cosmic Test Bench – Sag analysis





Tracking Resolution Sims.

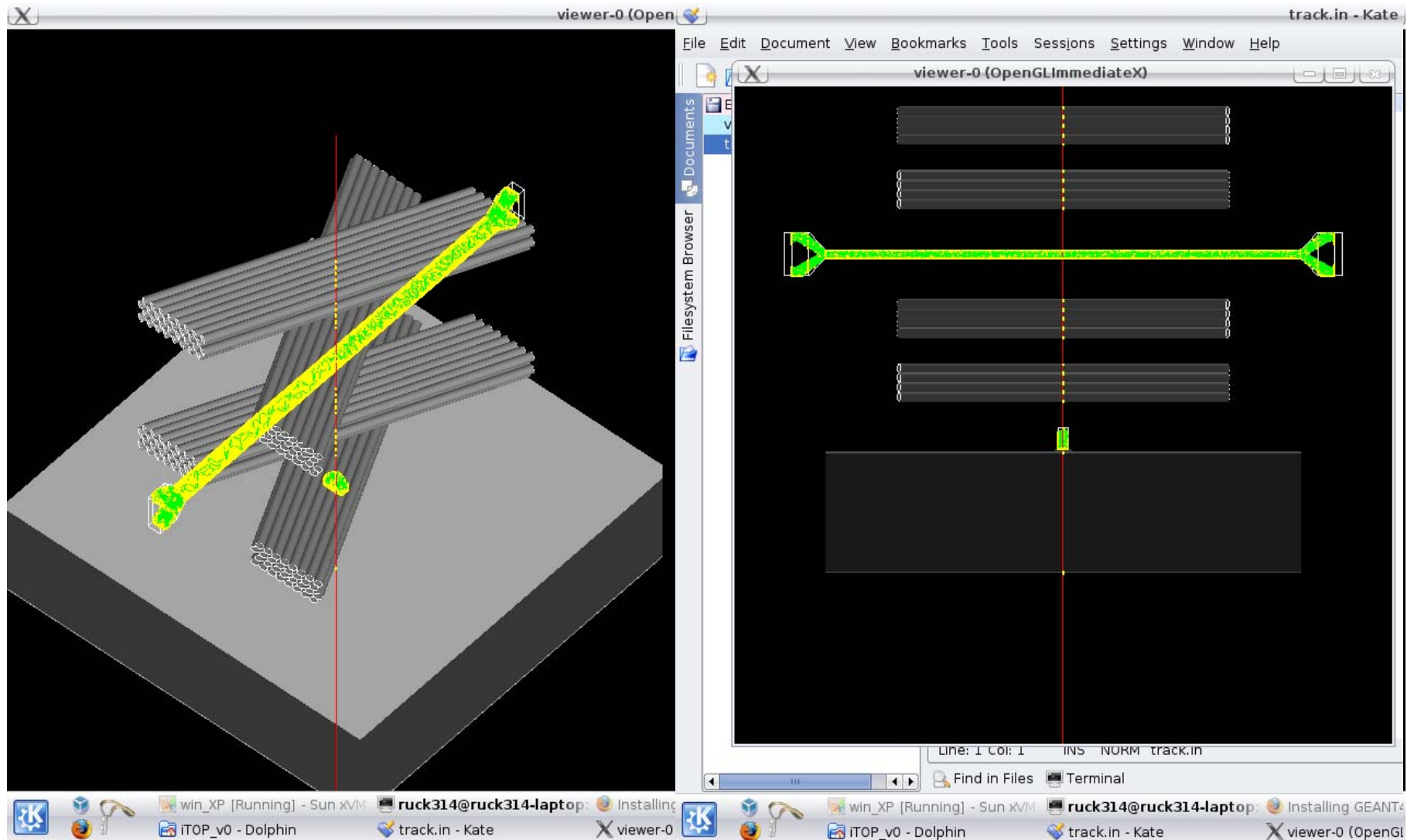
$\theta = 2 \text{ deg. } \phi = 70 \text{ deg.}$



$\sim 100\mu\text{m}$

G4 Simulation

- Studying optimizations for image planes



- 100% avail. beam, test different concepts – GEANT confirm

Photon detector options

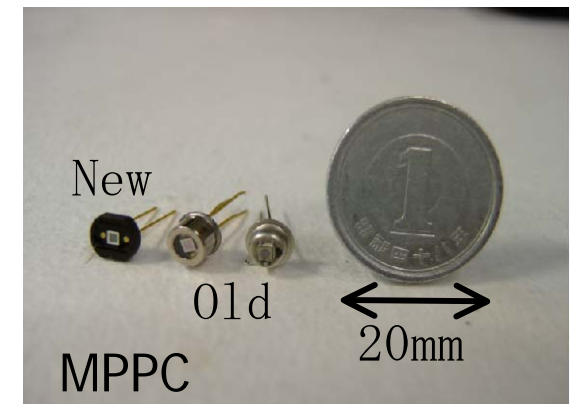
- **MCP-PMT (H-8500)**

- Will use for initial testing
 - Study realistic time resolution
 - Use BLAB gain to extend lifetime
- Demonstrate timing at lower gain



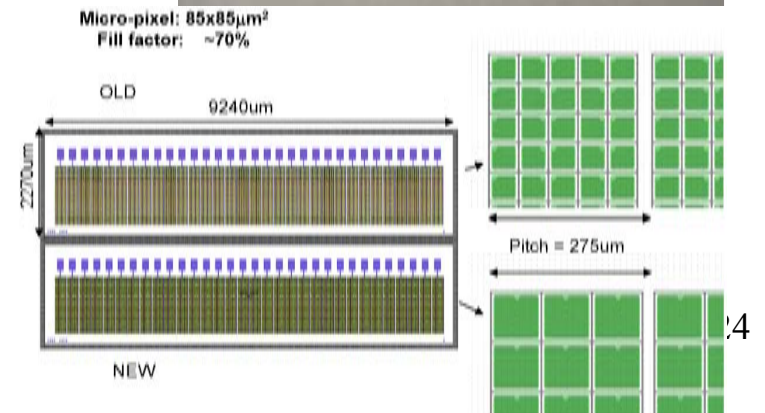
- **SiPM/MPPC**

- Good stability, 100ps TTS (N=100 p.e. \rightarrow 10ps)
- Need light guide to make work
- Radiation hardness?

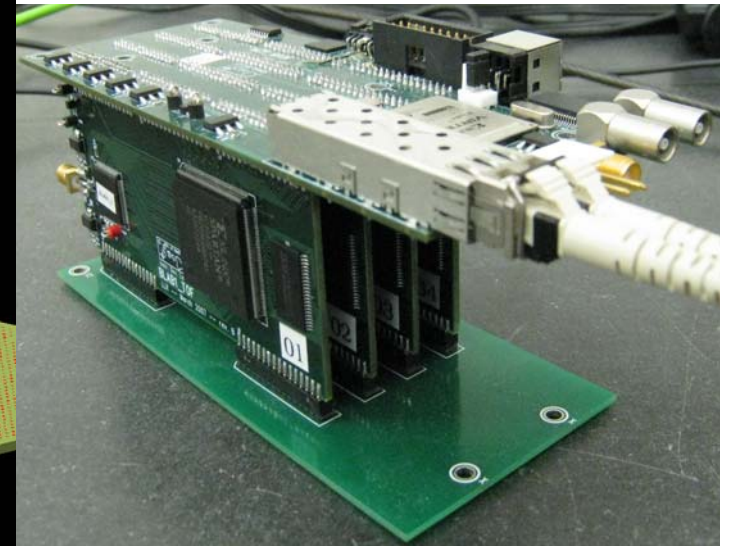
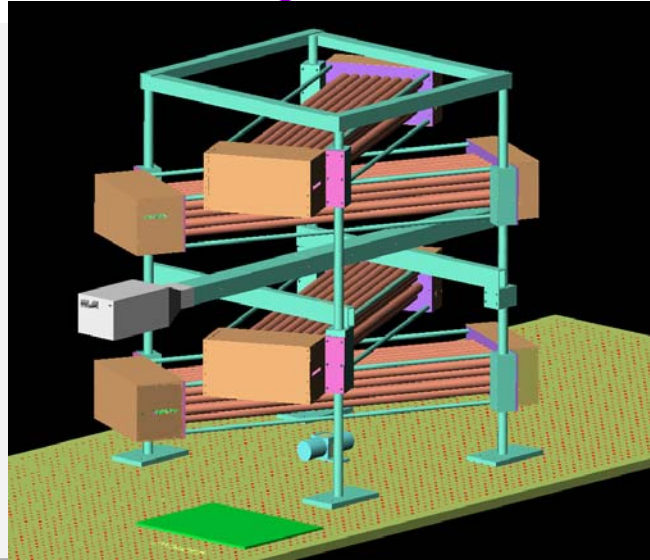
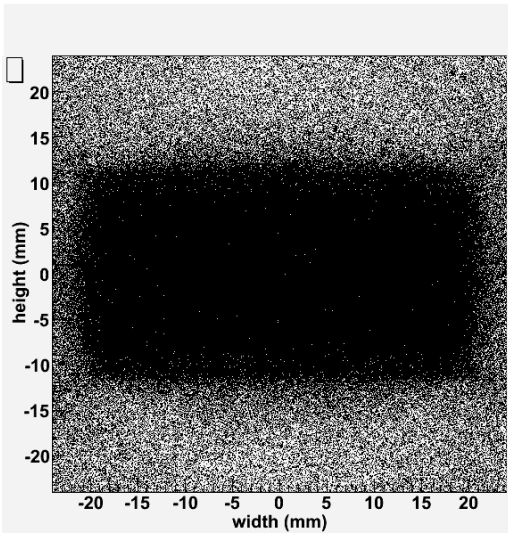


- **Linear arrays**

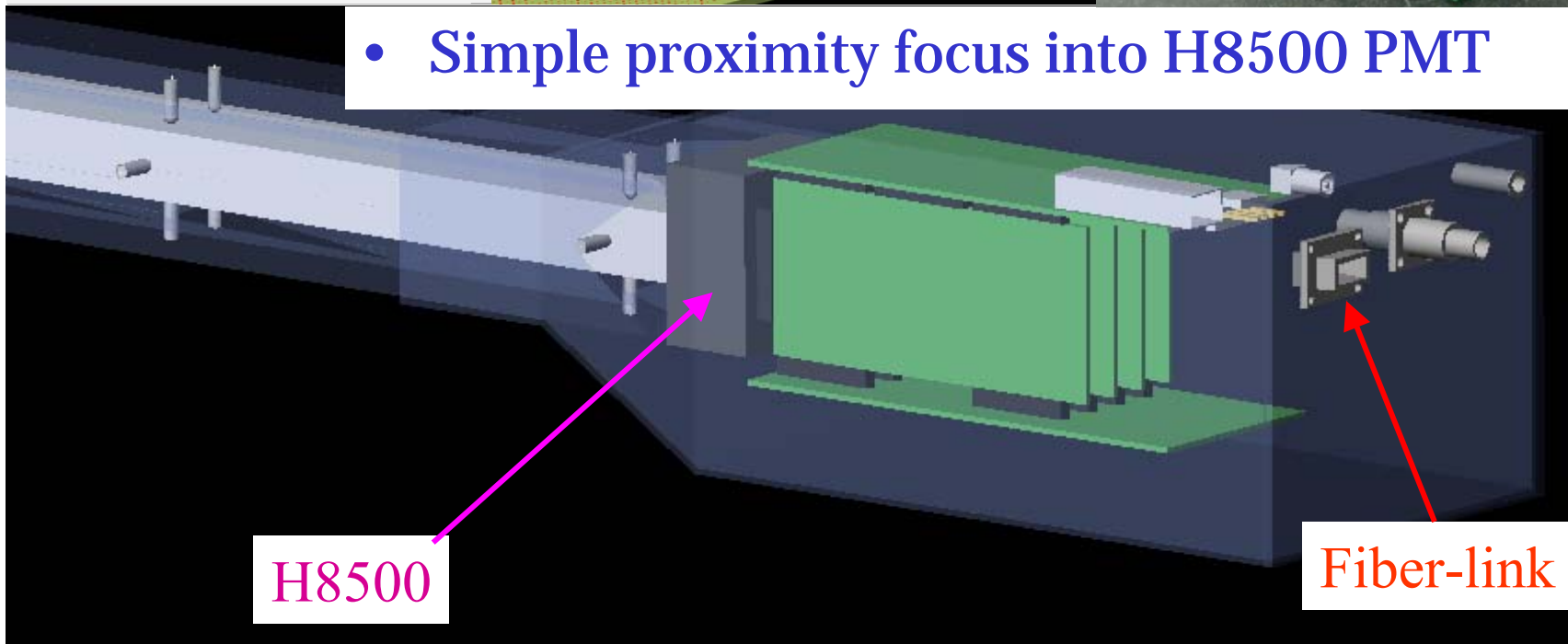
- Started evaluating (good vertical resolution)
- Packaging considerations



Initial Quartz bar test

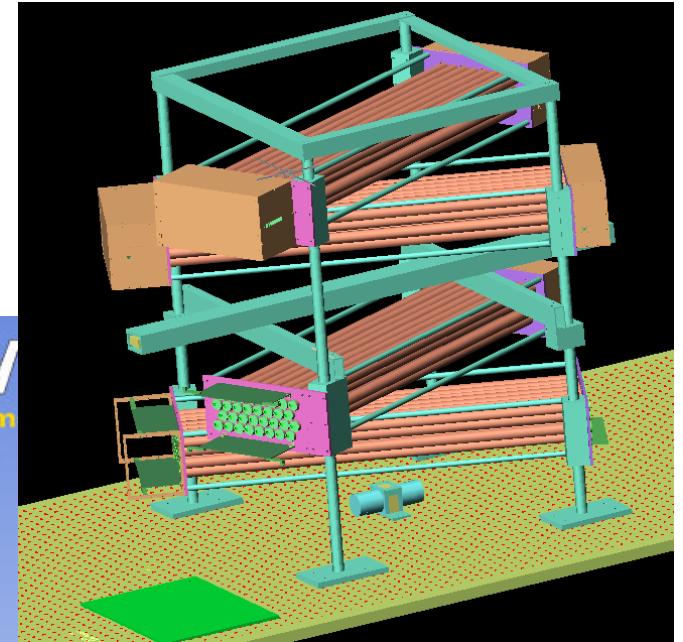
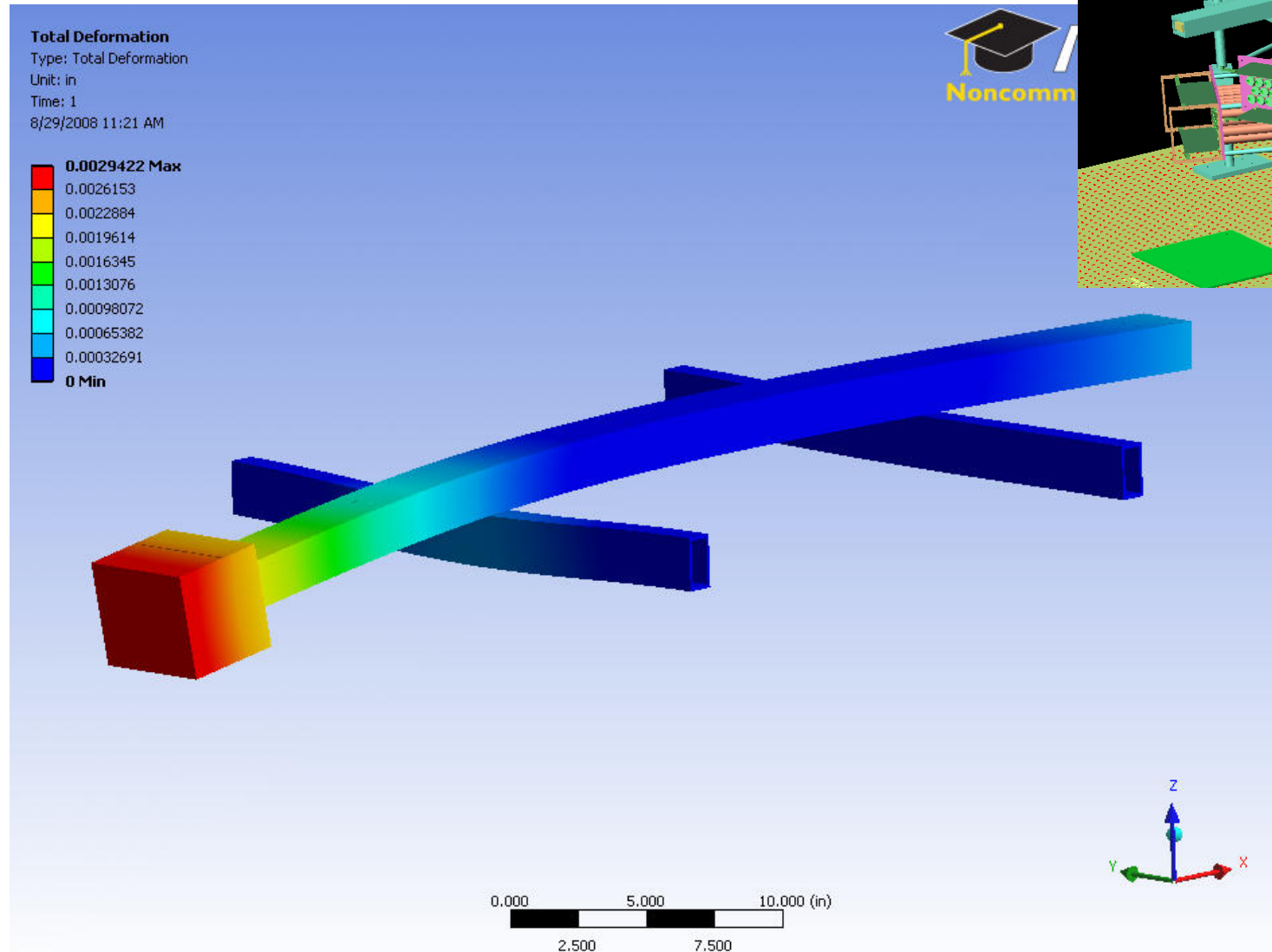


- Simple proximity focus into H8500 PMT



Cantilevered Loading

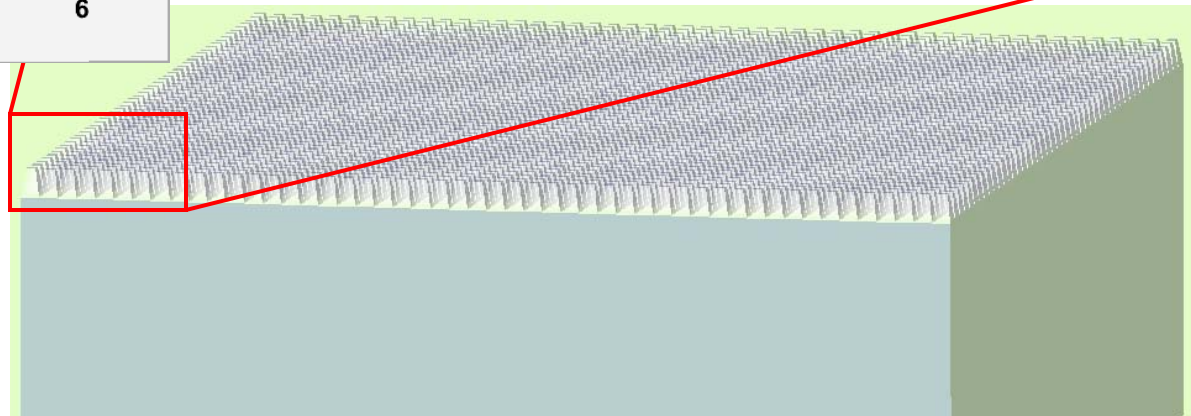
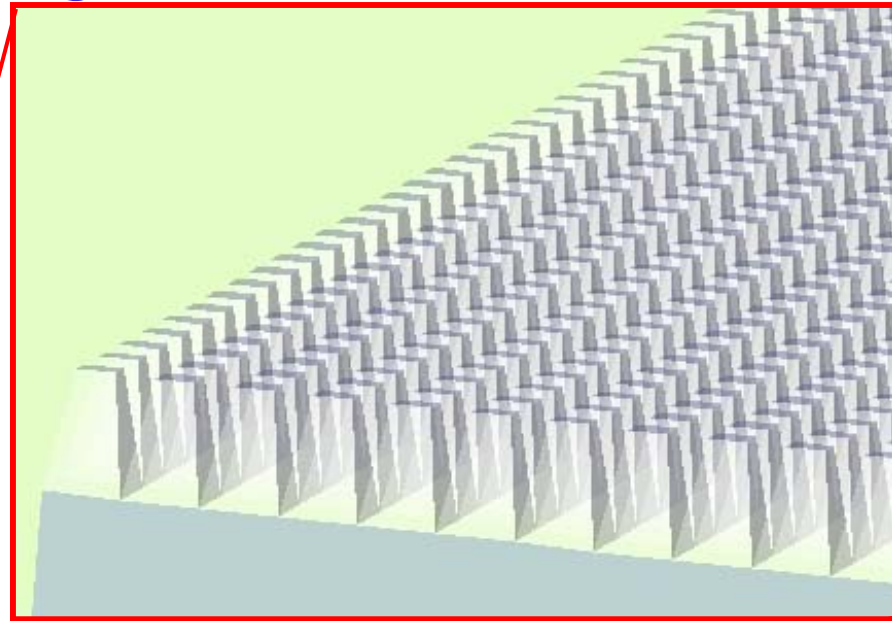
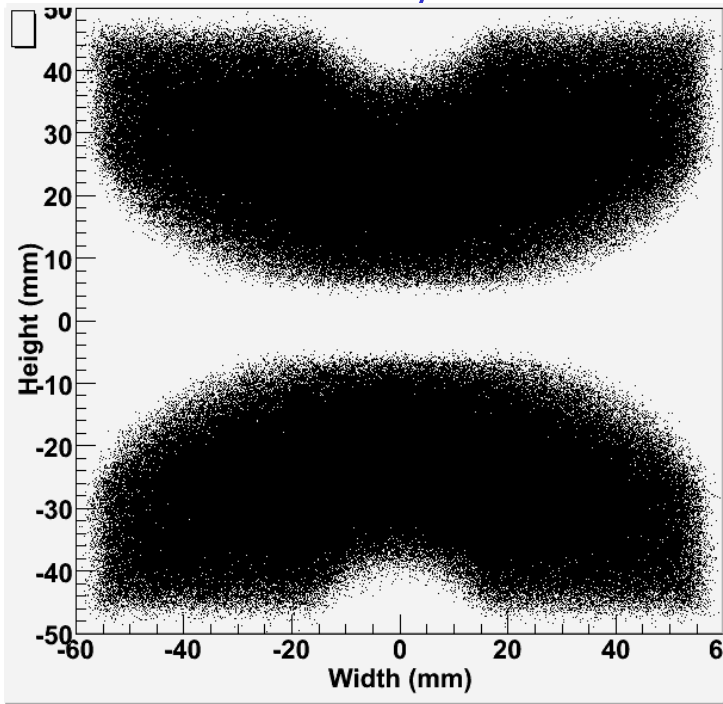
~75um max deflection (~3kg load)



Add
additional
support
(outrigger)

Full Image plane test (40mm SOB)

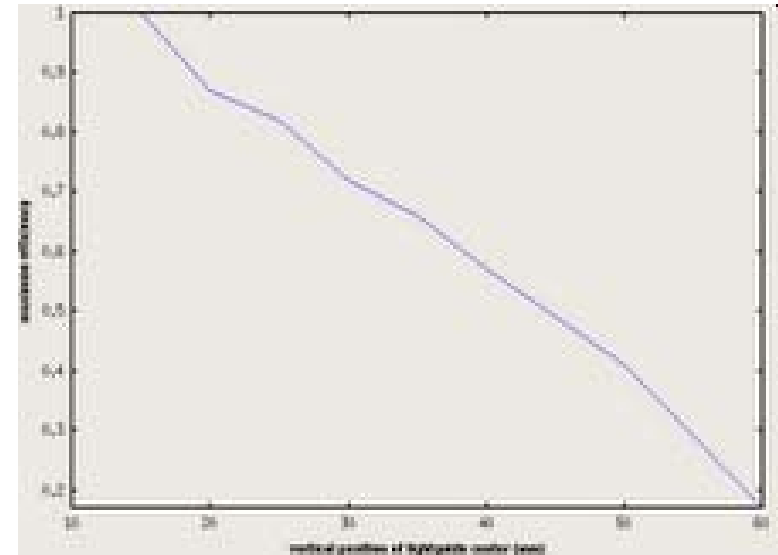
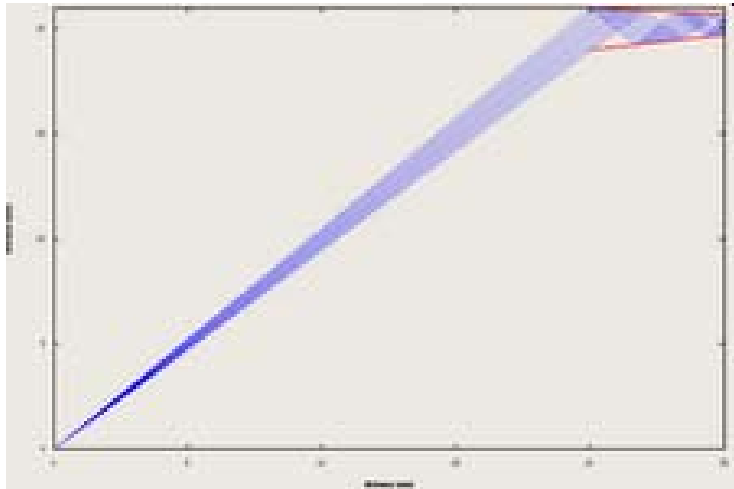
- Proximity focus initial target (100mm iTOP thickness)



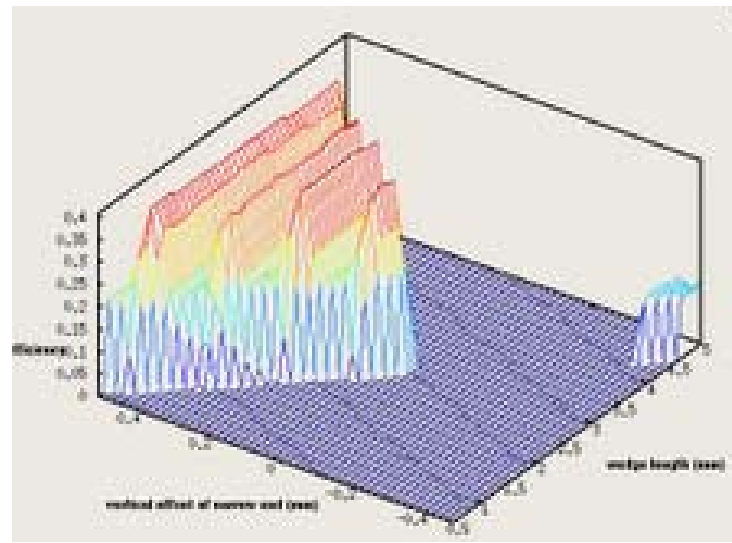
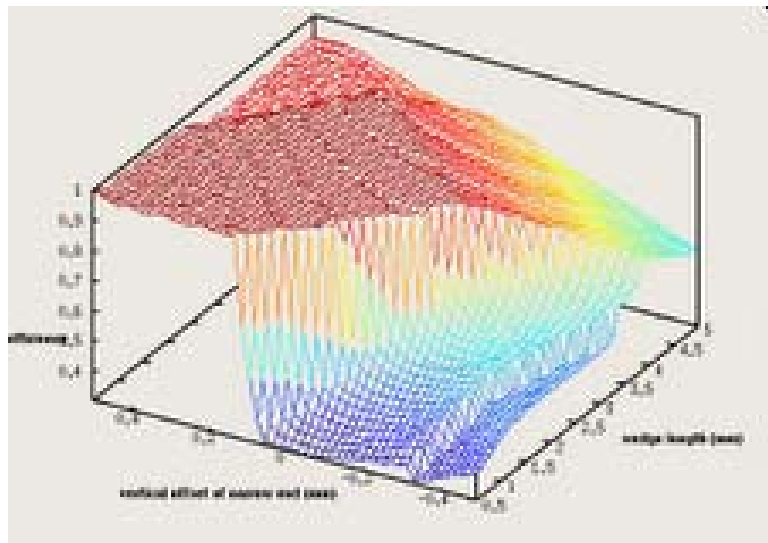
120mm (W)
x 100mm (H)
x 40mm (T)
2.4k pixels shown

- For good coupling, proposed to taper versus position, but...

Off-axis coupling issues...

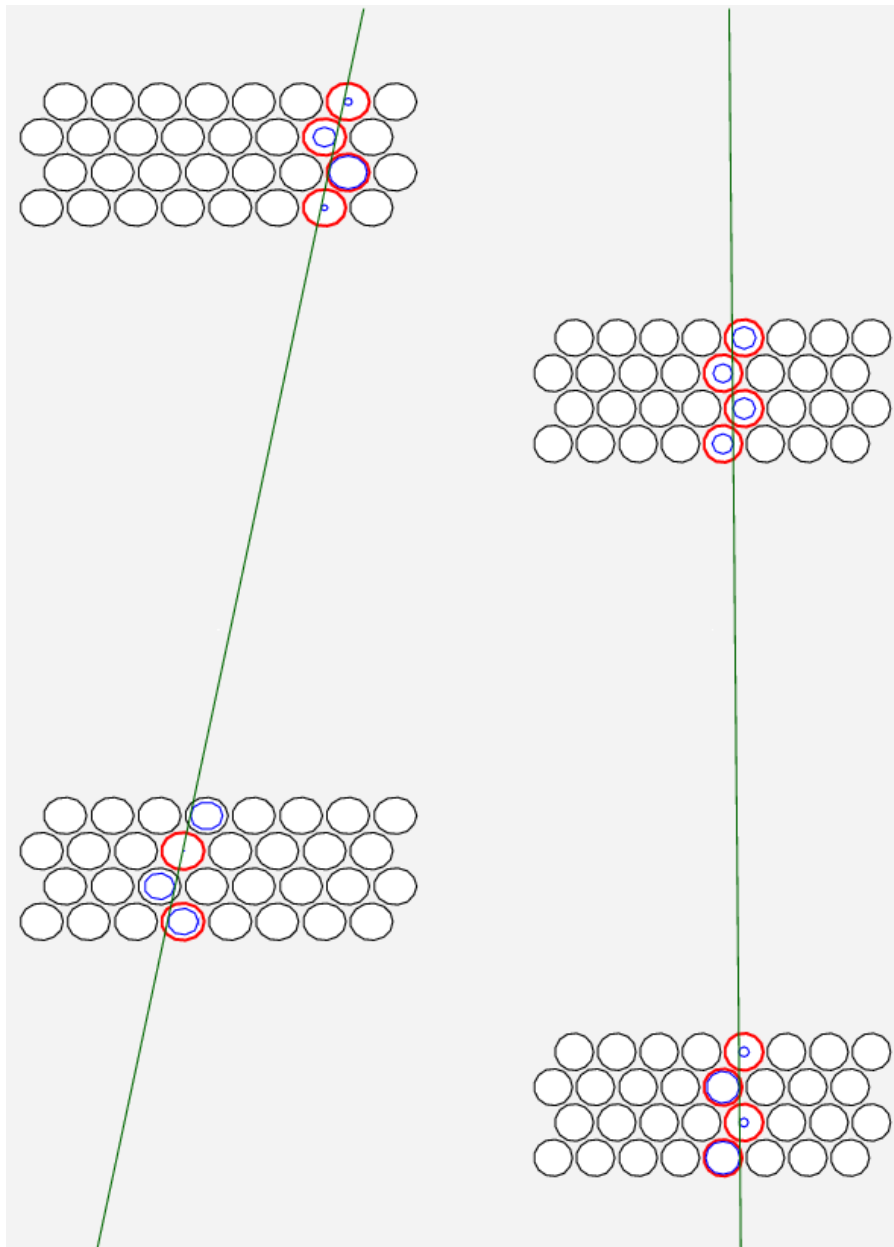


(blame octave for image quality...)



Test bench will allow study of planar arrays

First Fitted Track



Fit status was 1

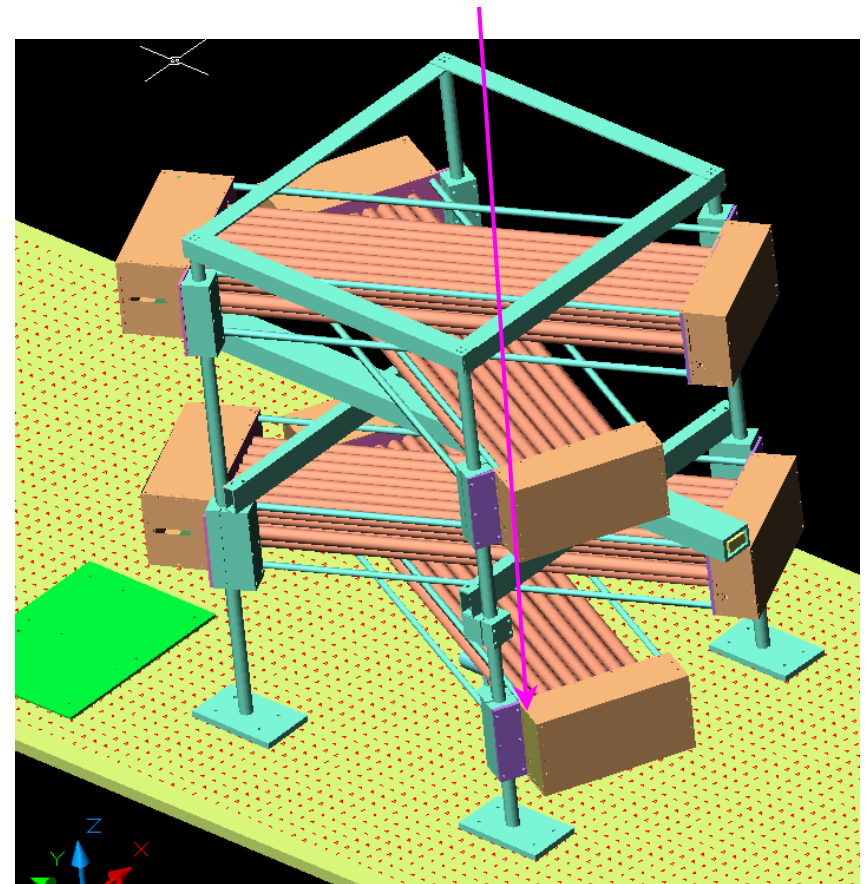
Theta = 14.1181 deg

Phi = 247.083 deg

x0 = 1.2125 cm

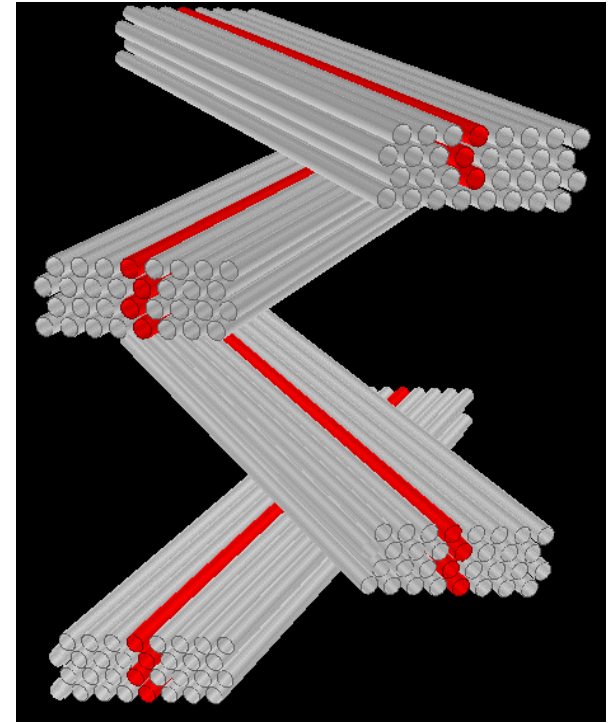
y0 = -0.943005 cm

t0 = 177.919 ns



Test Bench Status

Cosmic Data taking for tube alignment in Progress



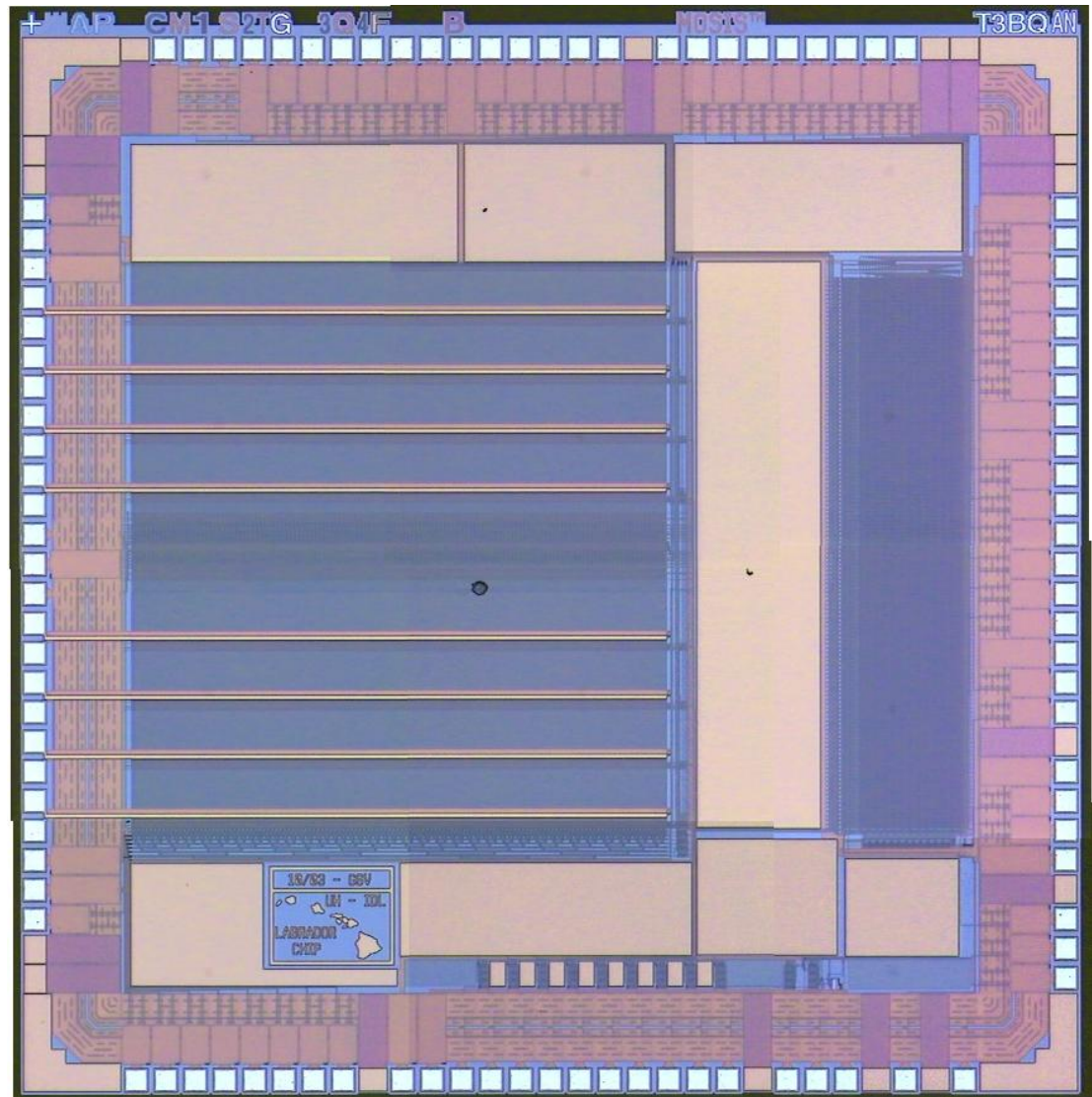
- Machining of Bar Support in progress
- First tests with H8500 thereafter (debug for f-DIRC)
- **Strategy** is to match full GEANT4 simulations to results seen (close the loop)
- In parallel develop larger image array (~2k channels)
- Add a 2nd Quartz bar
- Add momentum selector (DSSDs and magnet)

Summary

Building toward large-scale prototypes

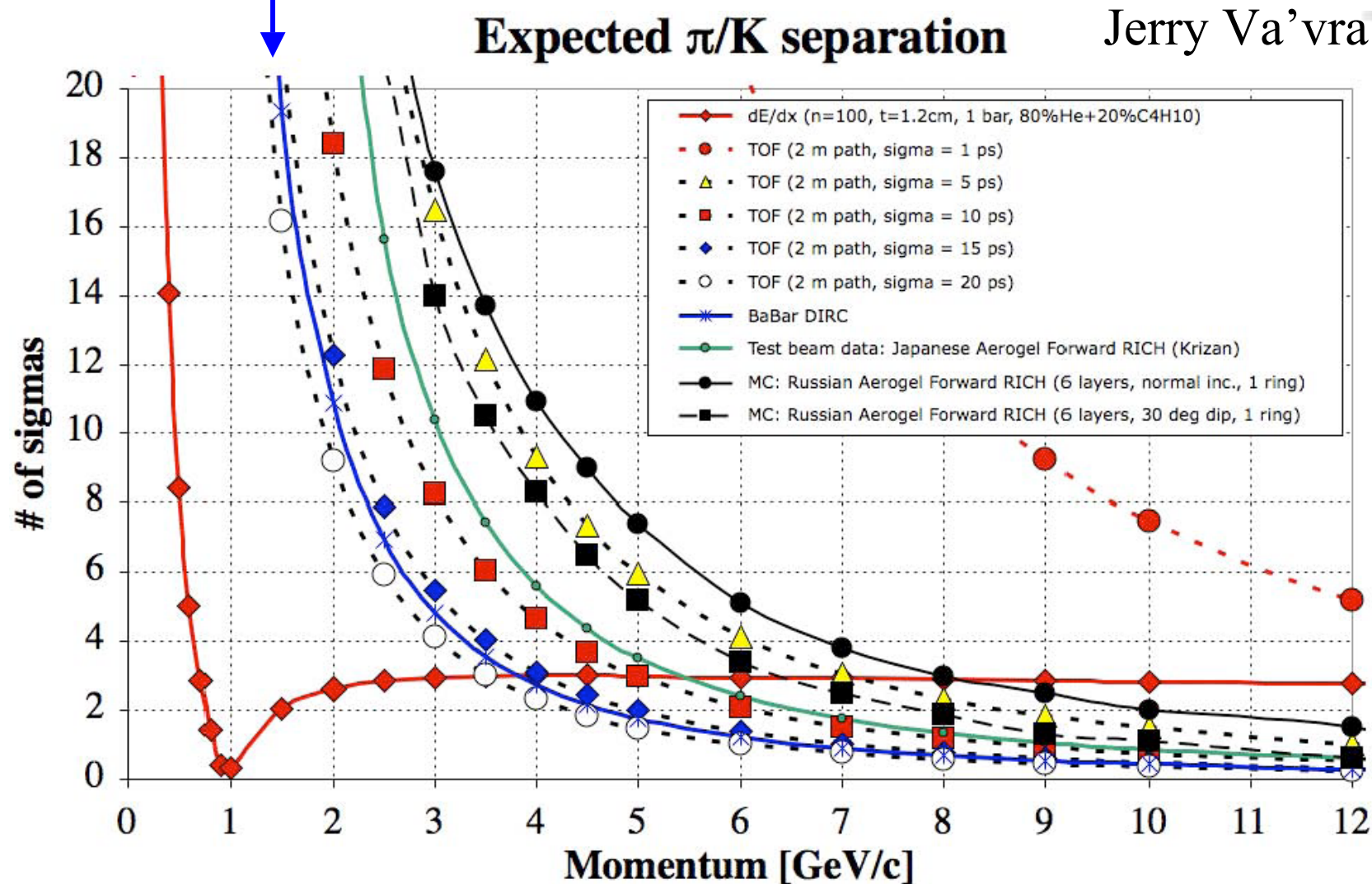
- December 2008 – 448 channel BLAB2-based readout system in Bldg 121 f-DIRC test stand
- Timing tests of semi-conductor G-APDs (continuing this autumn)
- Prototype k-Channel imaging plane (autumn)
 - Performance, cost, integration
 - Higher pixellation readout
- Close loop between GEANT4 and prototype (by year's end?)
- Define what needed for TDR input (2009)

Back-up slides

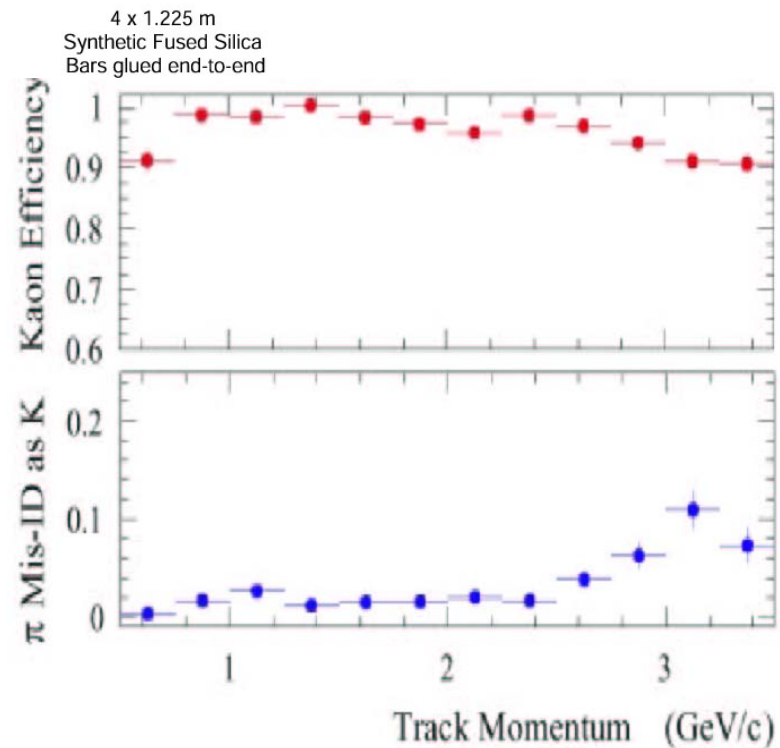
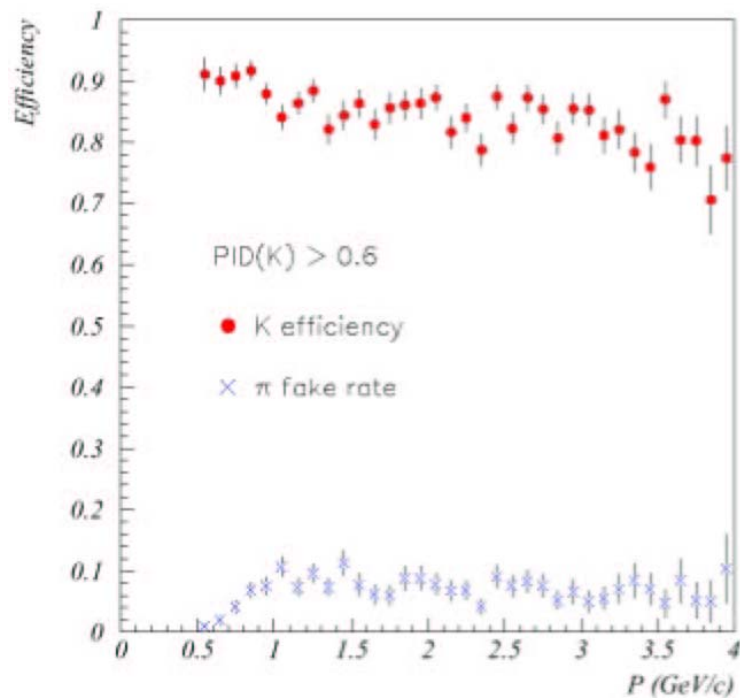
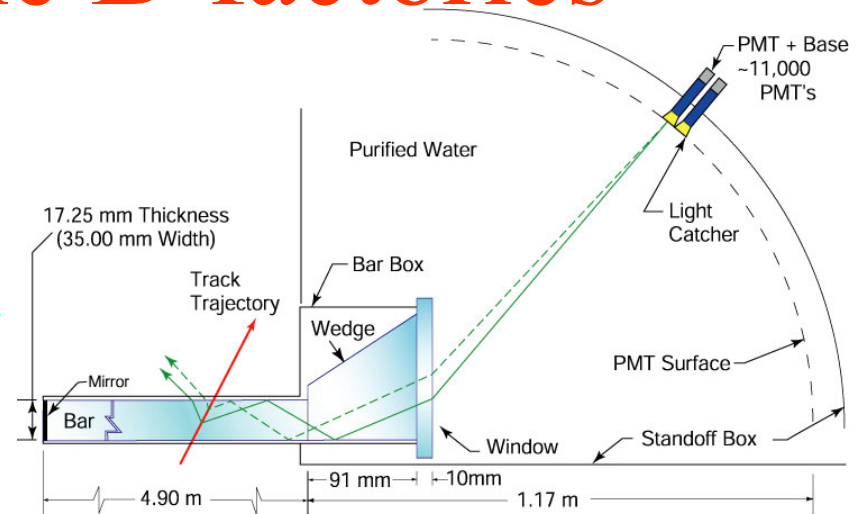
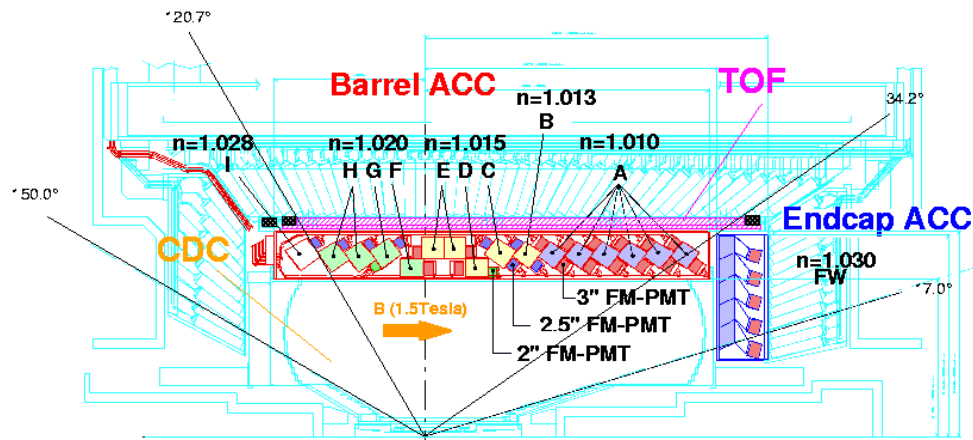


Particle ID Techniques

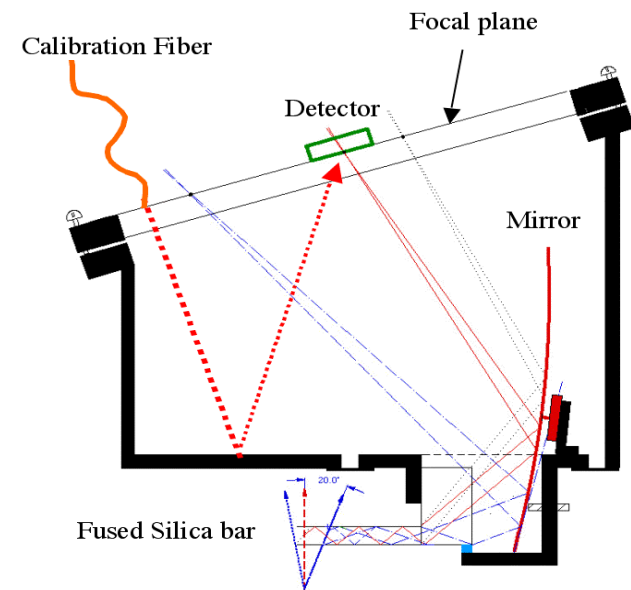
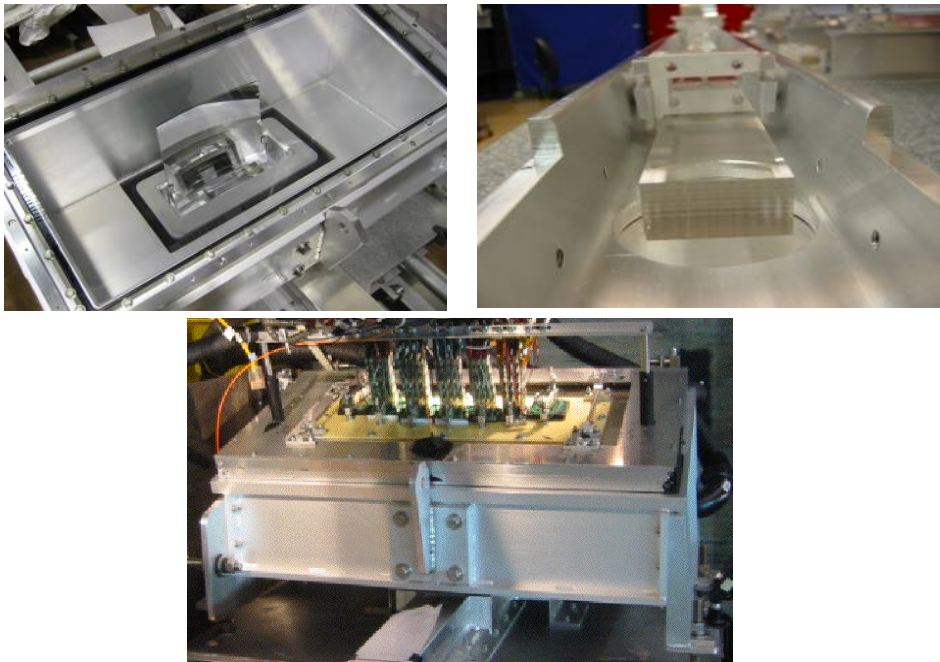
- BaBar DIRC is the starting place



Particle ID at the B-factories



Focusing DIRC Prototype Optics



- **Radiator:**
 - 1.7 cm thick, 3.5 cm wide, 3.7 m long fused silica bar (spares from BABAR DIRC).
- **Optical expansion region:**
 - filled with a mineral oil to match the fused silica refraction index (KamLand oil).
 - include optical fiber for the electronics calibration (PiLas laser diode).
- **Focusing optics:**
 - a spherical mirror with 49cm focal length focuses photons onto a detector plane.

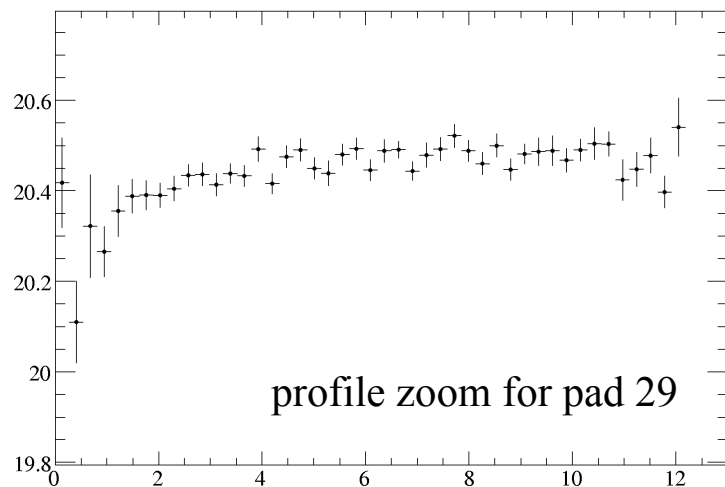
TDC vs. ADC for signal in run 27

Larry's offline correction method seems to come close to correcting time walk.

Some over-correction, some under-correction., more can be done offline with charge info.

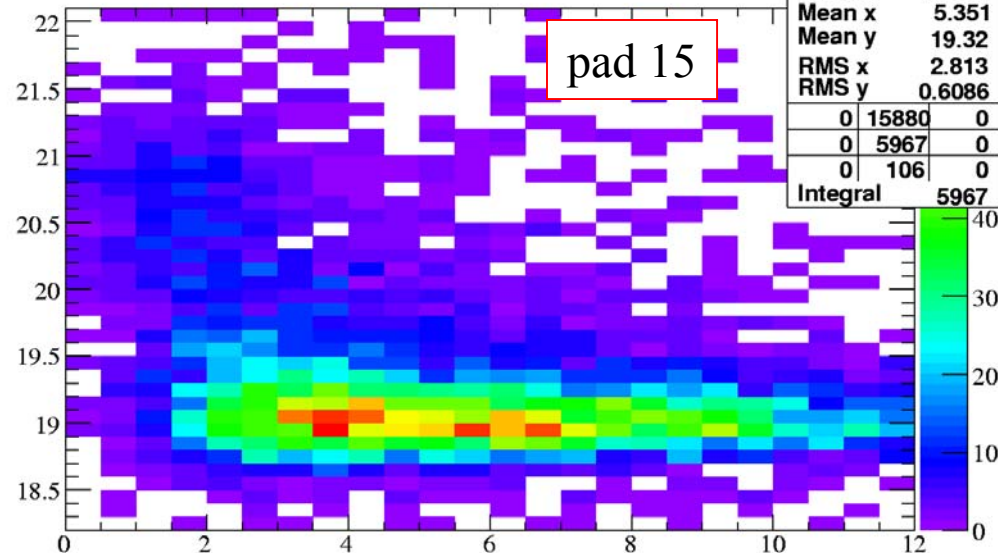
Jochen Schwiening
analysis (preliminary)

tdc_5:adc_5 {tdc_5>19.5&&tdc_5<21.5&&adc_5<12}

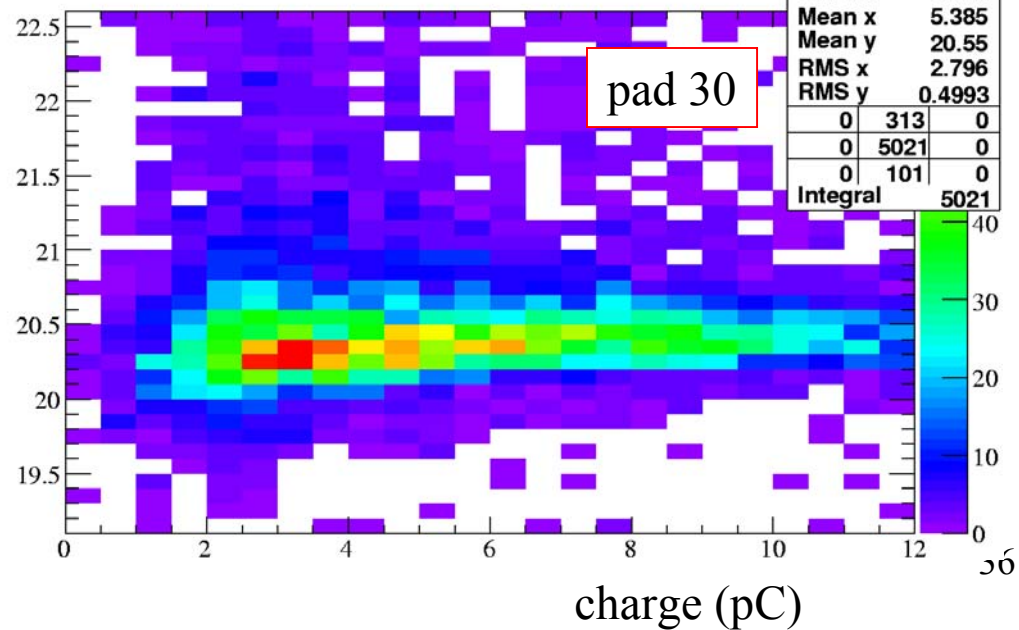


time (ns)

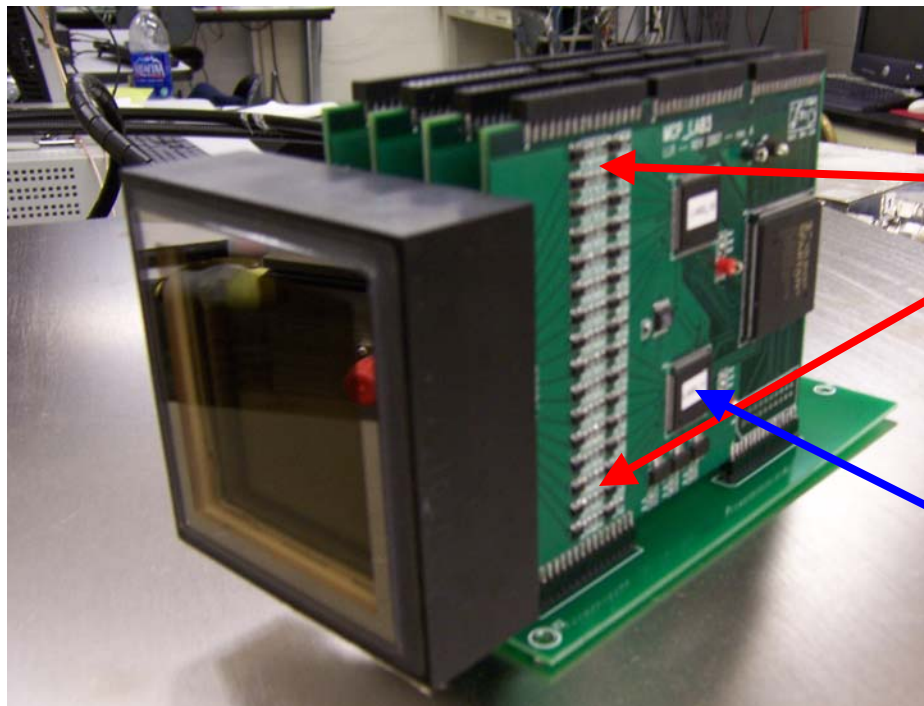
tdc_2:adc_2 {tdc_2>15&&tdc_2<35&&adc_2<12}



tdc_5:adc_5 {tdc_5>15&&tdc_5<35&&adc_5<12}



Gain Needed



Amplifiers dominate
board space

Readout ASIC tiny
(14x14mm for 16
channels)

- What gain needed?
 - At 10^6 gain, each p.e. = 160 fC
 - At 2×10^5 gain (better for aging), each p.e. = 32 fC
 - In typical ~ 5 ns pulse, $V_{\text{peak}} = dQ/dt * R = 32 \mu\text{A} * R = 32 \text{mV} * R [\text{k}\Omega]$ (6.4mV)

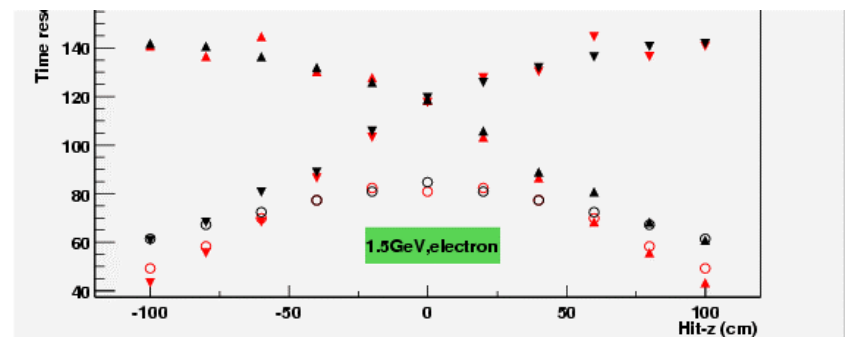
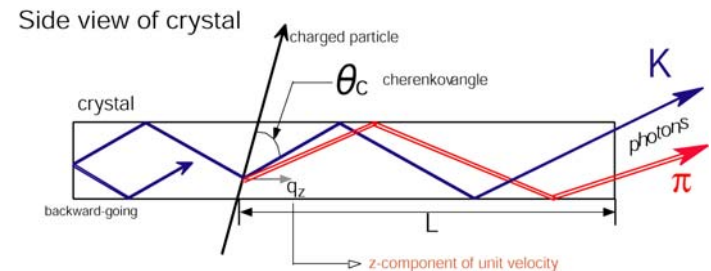
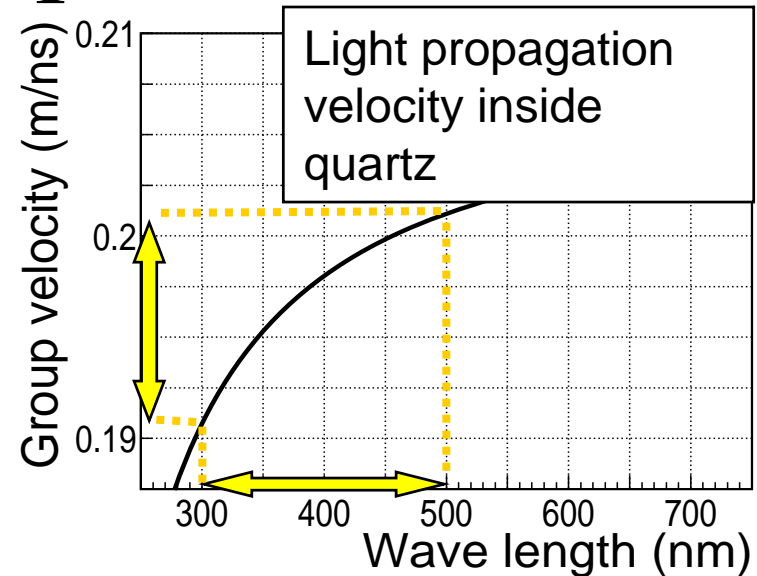
Gain Estimate	
Rterm	1 p.e. peak
50	1mV
1k	20mV
20k	400mV

Chromatic dispersion

Variation of propagation velocity depending on the wavelength of Cherenkov photons

- Due to wavelength spread of detected photons
- → propagation time dispersion
- Longer propagation length
→ Improves ring image difference
But, decrease time resolution.

Techniques complement each other!



Separation Concept using $\text{Log } \mathcal{L}$

90 degrees (normal incidence)

