

The Darkside Program

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on behalf of Darkside-50 collaboration

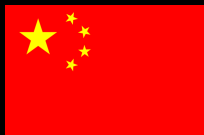
RICAP 2014
The Roma International Conference on
Astroparticle Physics

Noto, Italy

30 September 2014



The Darkside Collaboration



IHEP - China

Universite de Strasbourg - France

APC Universite Paris 7 Diderot - France

INFN Laboratori Nazionali del Gran Sasso - Italy

Gran Sasso Science Institute - Italy

INFN and Universita degli Studi Cagliari - Italy

INFN and Universita degli Studi Genova - Italy

INFN and Universita degli Studi Milano - Italy

INFN and Universita degli Studi Napoli - Italy

INFN and Universita degli Studi Perugia - Italy

INFN and Universita degli Studi Roma 3 - Italy

Jagiellonian University - Krakow, Poland

Joint Institute for Nuclear Research - Dubna, Russia

Lomonosov Moscow State University - Moscow, Russia

National Research Center Kurchatov Institute - Moscow, Russia

St. Petersburg Nuclear Physics Institute - Gatchina, Russia

KINR, NAS Ukraine - Keiv, Ukraine

Augustana College - USA

Black Hills State University - USA

University of Chicago - USA

University of Hawaii - US

University of Houston - USA

University of Massachusetts - USA

Princeton University - USA

Temple University - USA

UC Davis - USA

UCL A- USA

Virginia Tech - USA

FNAL, LANL, LLNL, PNNL, SLAC



The Darkside program @ LNGS

- A scalable technology for direct detection search for WIMP dark matter
 - Based on a two-phase argon time projection chamber (TPC)

DarkSide-10



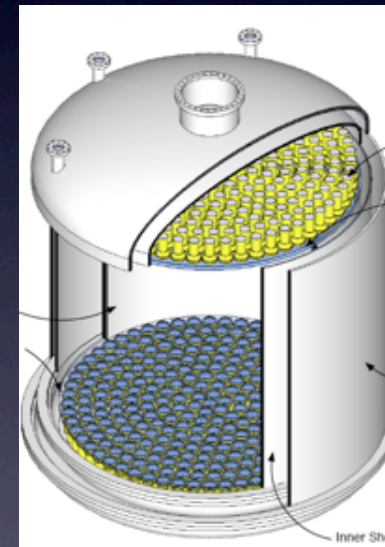
technical prototype
no DM goal

DarkSide-50



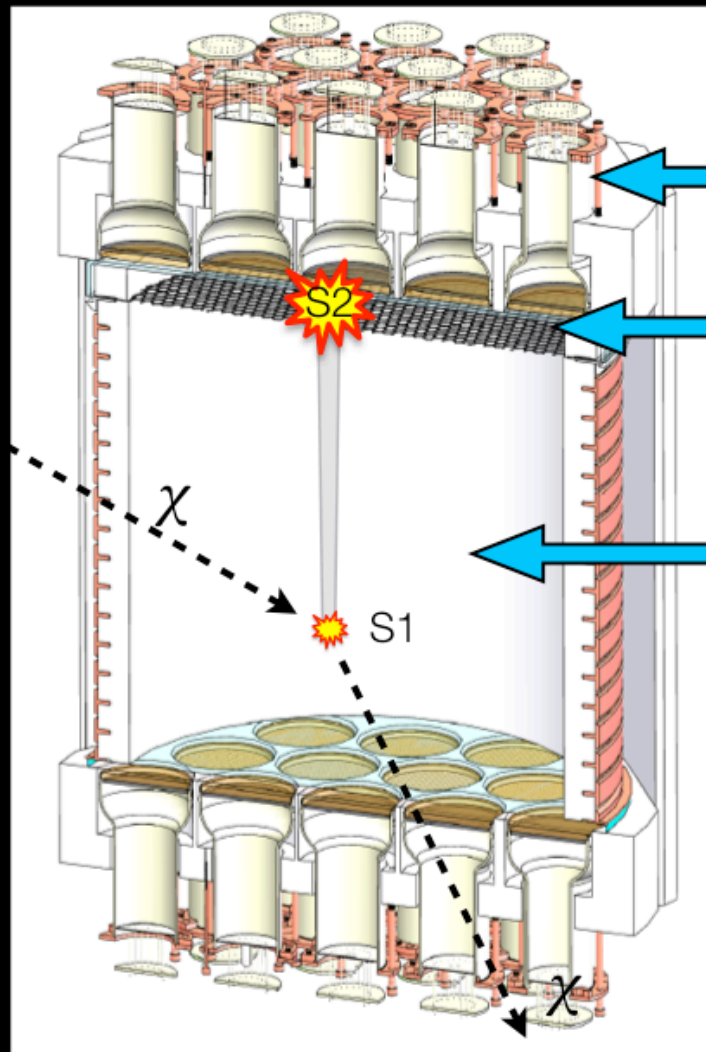
sensitivity
 10^{-45} cm^2

DarkSide-G2



sensitivity
 10^{-47} cm^2

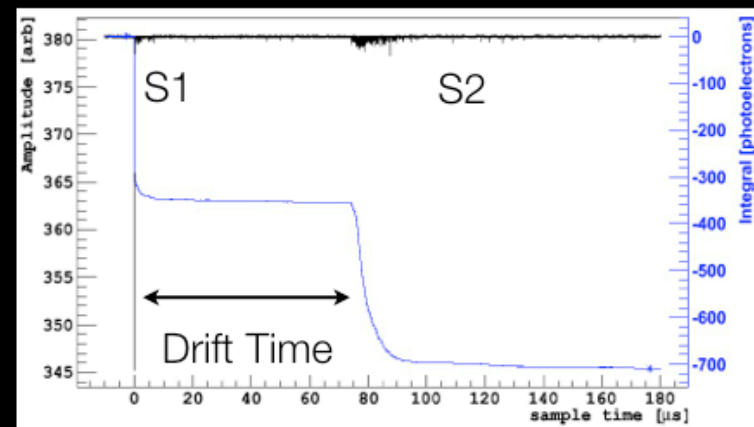
Detecting WIMPs



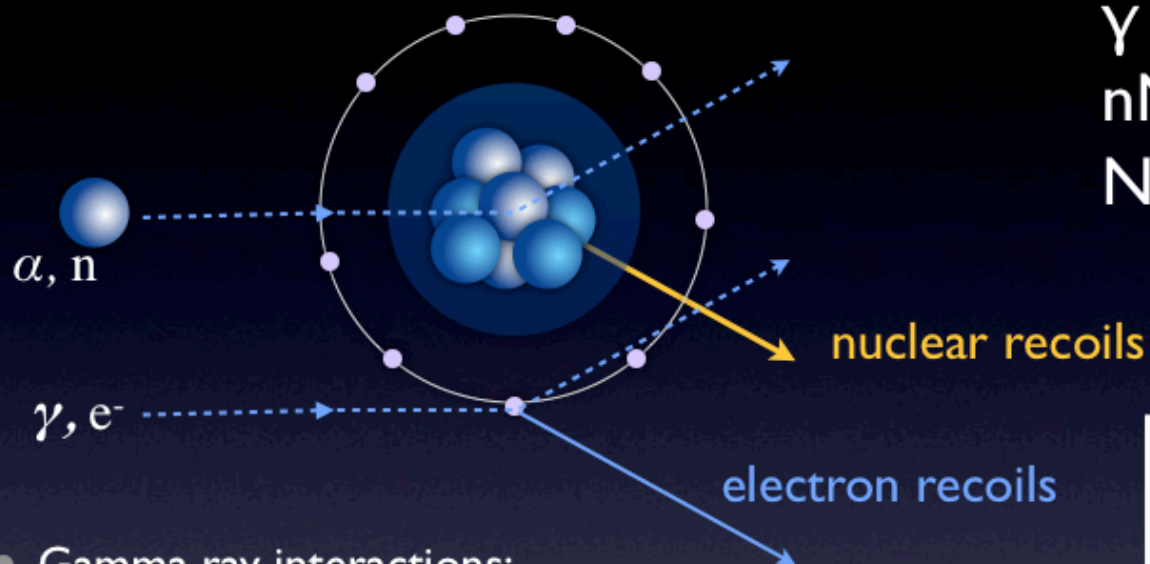
19 3" Photomultiplier Tubes (Top & Bottom)

Gas Ar ($E_{lum} \sim 4200$ V/cm)

Liquid Ar ($E_{drift} \sim 200$ V/cm)



Background



from natural radioactivity:

$$\gamma e^- \rightarrow \gamma e^-$$

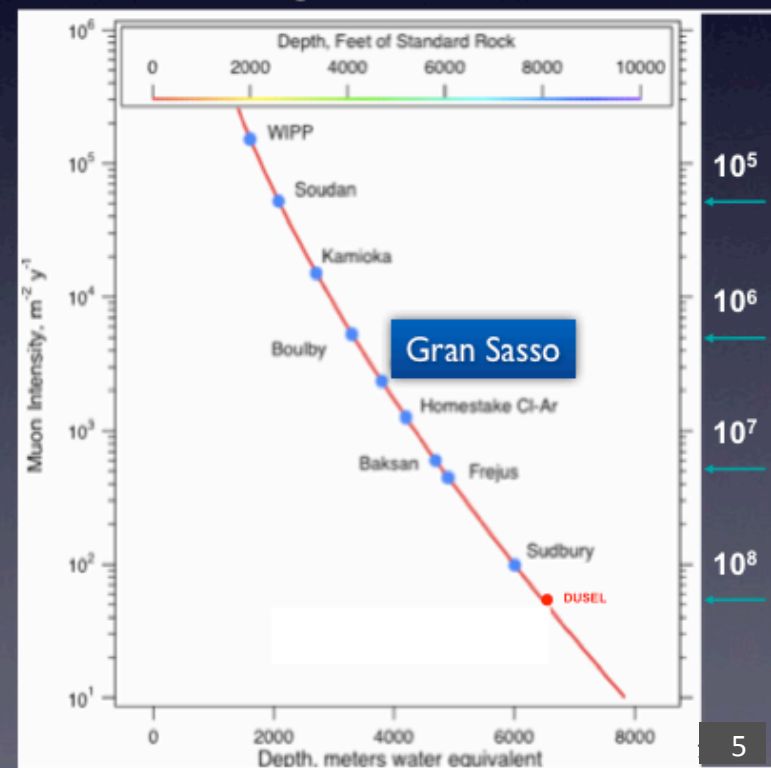
$$nN \rightarrow nN$$

$$N \rightarrow N' + \alpha, e^-$$

reduction
of muon
flux by:

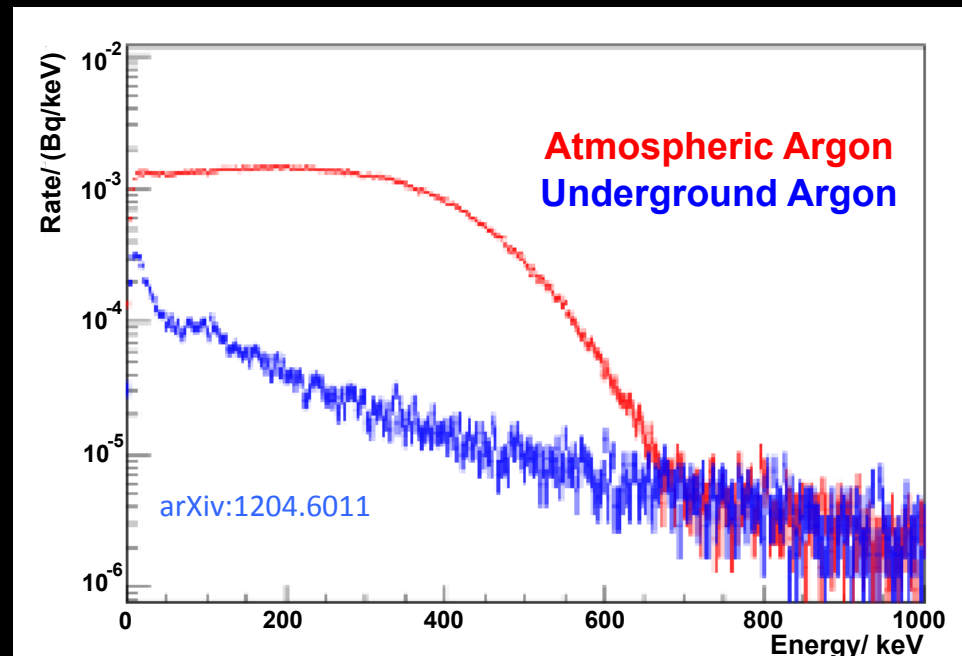
Underground labs

- Gamma ray interactions:
mis-identified electrons mimic nuclear recoil signals
- Neutrons:
(α, n), U,Th fission, cosmogenic spallation
- Contamination:
 ^{238}U and ^{232}Th decays, recoiling progeny mimic nuclear recoils



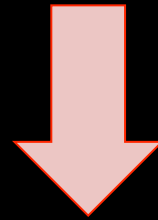
^{39}Ar reduction

- Dominant source of background: ^{39}Ar
- ^{39}Ar radioactivity present in atmospheric argon (1 Bq/kg)
- ^{39}Ar is a cosmogenic isotope, and the activity in argon from underground sources can be significantly reduced compared to atmospheric argon
- Source of **underground argon** measured to have **> 150 times** lower rate of ^{39}Ar , compared to atmospheric argon



Darkside-50

Design philosophy based on having very low background levels that can be further reduced through **active** suppression, for background-free operation



Liquid argon TPC,
within a neutron veto,
within a muon veto,
under a mountain

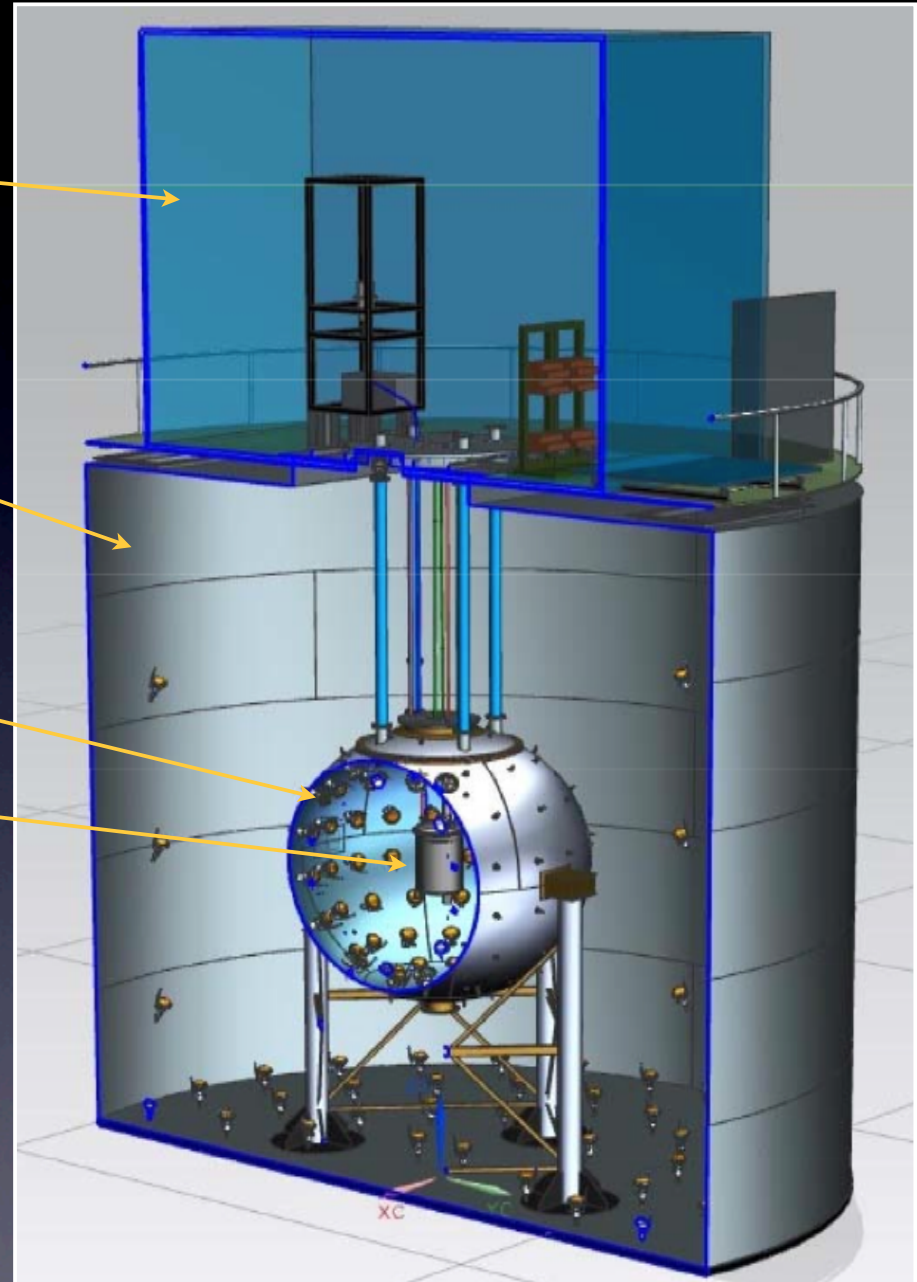
Radon-free clean assembly room
 $\leq 30 \text{ mBq/m}^3$ in $>100 \text{ m}^3$
(CRH)

μ veto and cosmogenic neutron passive shield
1000 ton water Cherenkov
(Borexino CTF)

Radiogenic neutron veto
30 ton borated liquid scintillator
(LSV)

WIMP LAr detector
150 kg of UAr $< 6.5 \text{ mBq/kg}$
(DS-50 TPC)

DarkSide design



Darkside-50 TPC

- 50 kg active mass of UAr
- 19 (top) + 19 (bottom) R11065 HQE Hamamatsu 3" PMTs
- ϕ 36 cm \times 36 cm drift
- Lateral walls made of high reflectivity polycrystalline PTFE
- All inner surfaces coated with TPB
- Fused silica diving bell (top) and window (bottom) in front of the PMT arrays coated with ITO.

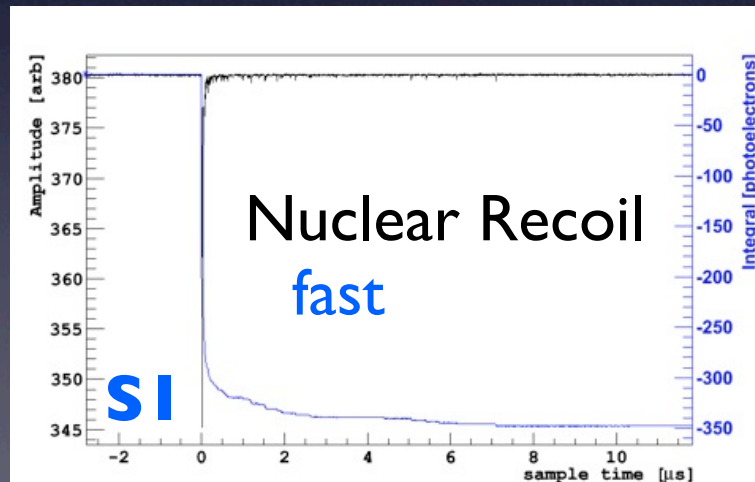
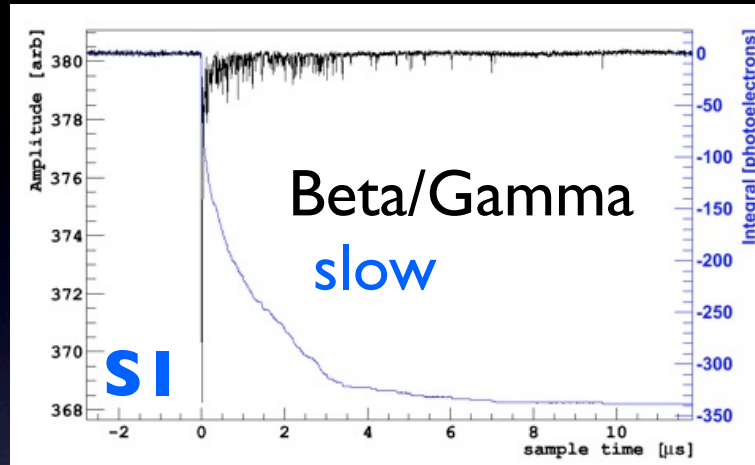
Designed to provide an extremely
high light yield
DATA TAKING IS ON GOING



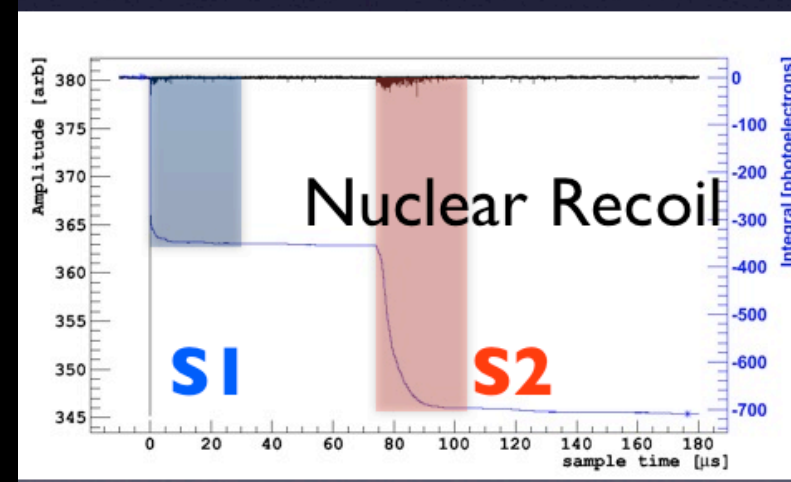
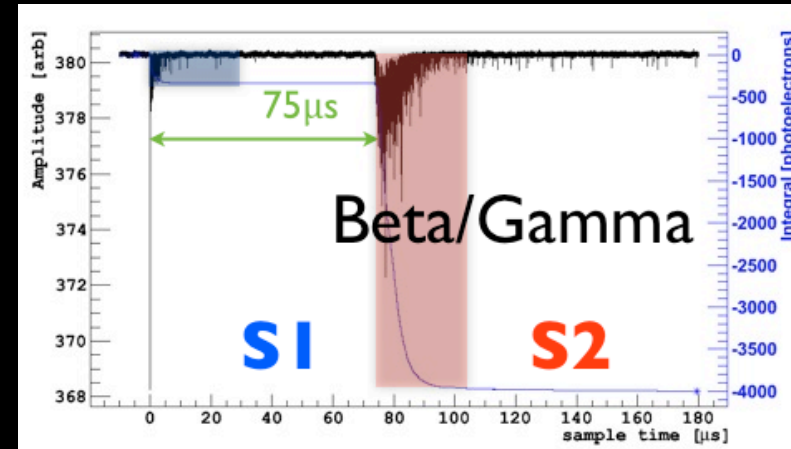
Background discrimination

Electron recoil
discrimination

Pulse shape discrimination

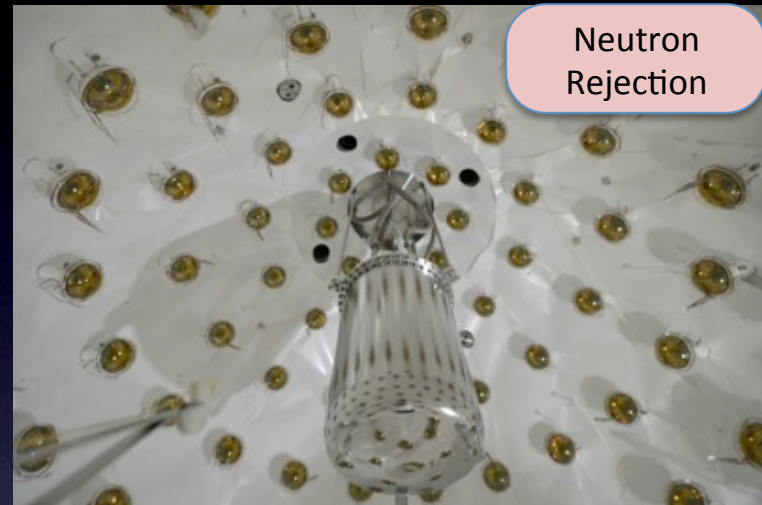


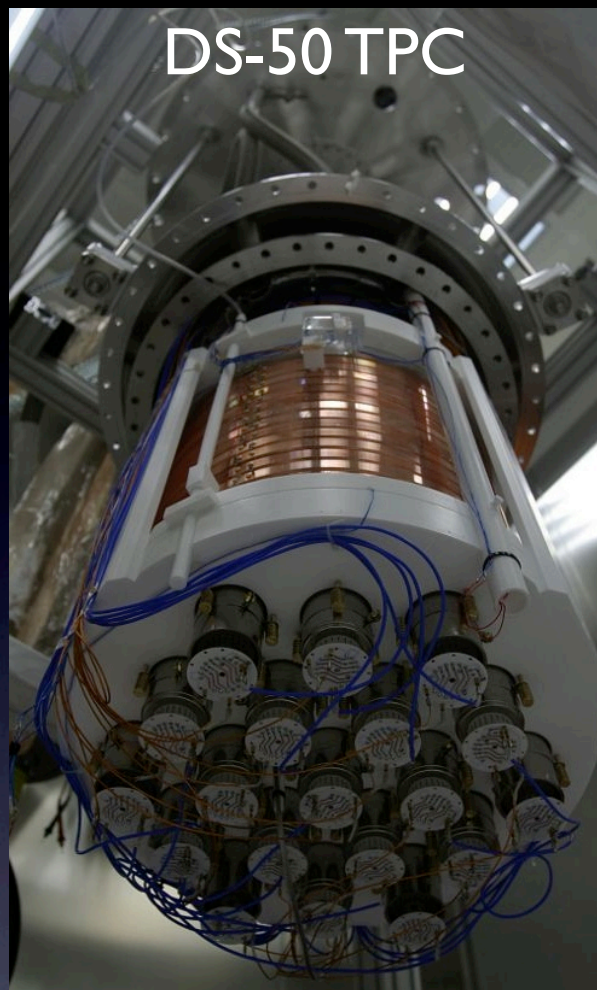
Charge-to-light ratio



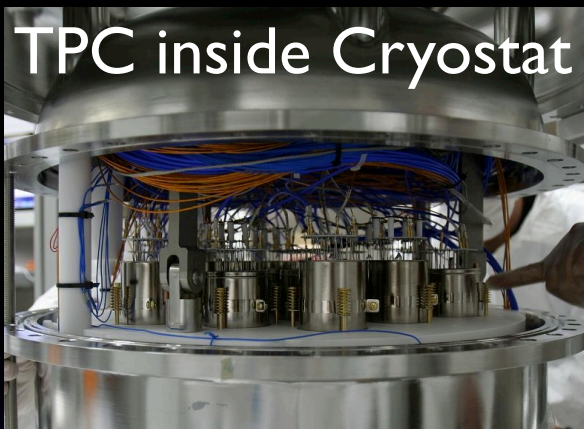
Neutron veto and muon veto

- The TPC is surrounded by a 30 ton **boron-loaded liquid scintillator** spherical veto, 4m diameter, instrumented with 110 low background 8" PMTs
- neutrons which escape the inner detector are detected via (n,α) reaction on ^{10}B
- >99.5% efficiency for radiogenic neutron detection, >95% for cosmogenic neutron detection [A.Wright et. al, NIMA 644, 18 \(2011\)](#)
- The LSV is installed inside a **Water Cherenkov** detector (Borexino CTF), 10 m height, 11 m diameter, filled with 1000 ton ultra-pure water, observed by 80 upward facing PMTs
- muon veto and passive shielding against external neutrons and gammas





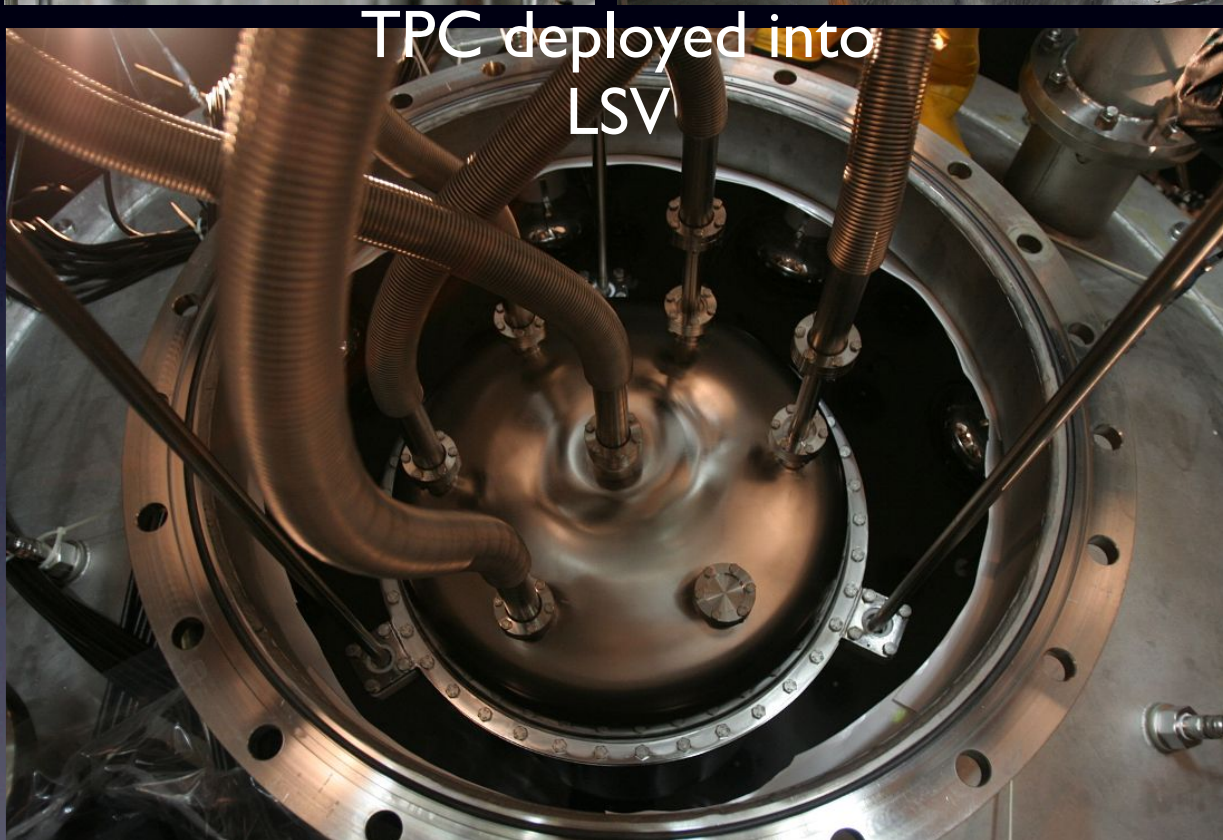
DS-50 TPC



TPC inside Cryostat



Cryostat and vacuum vessel



TPC deployed into LSV

DS50-TPC
Assembled,
Deployed

TPC hanging in LSV

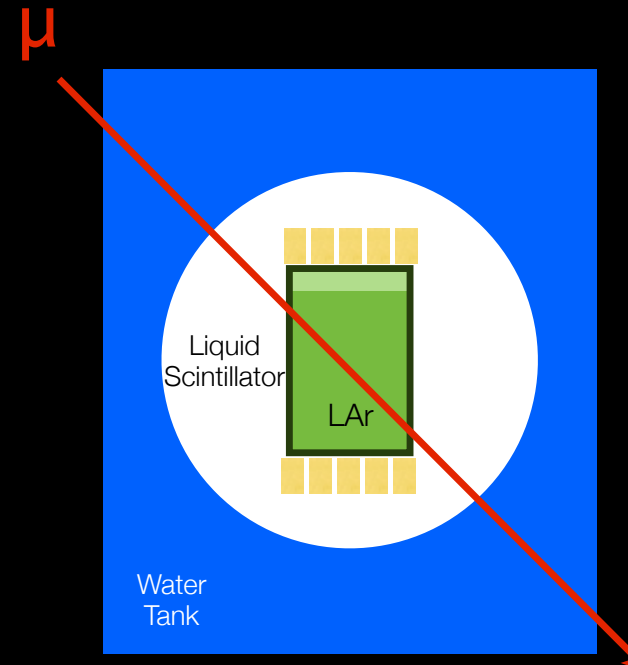
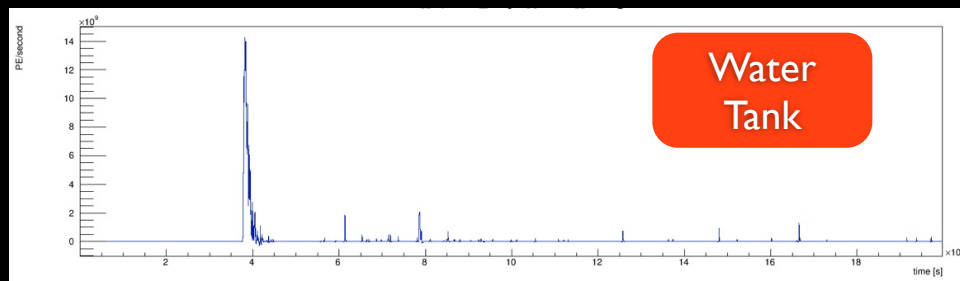
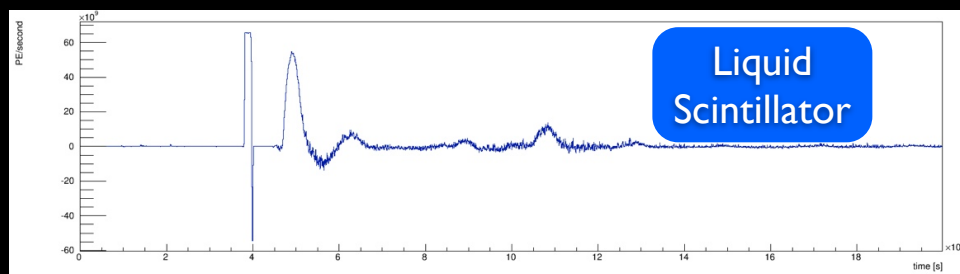
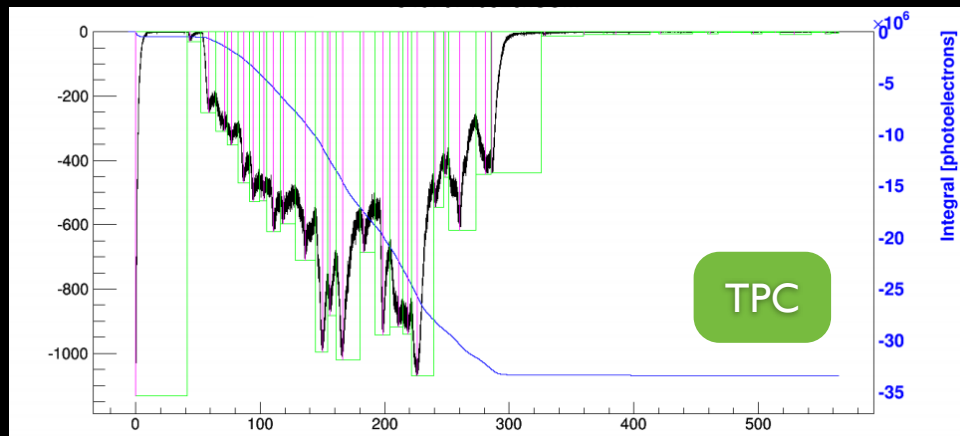


View inside the Water Cherenkov



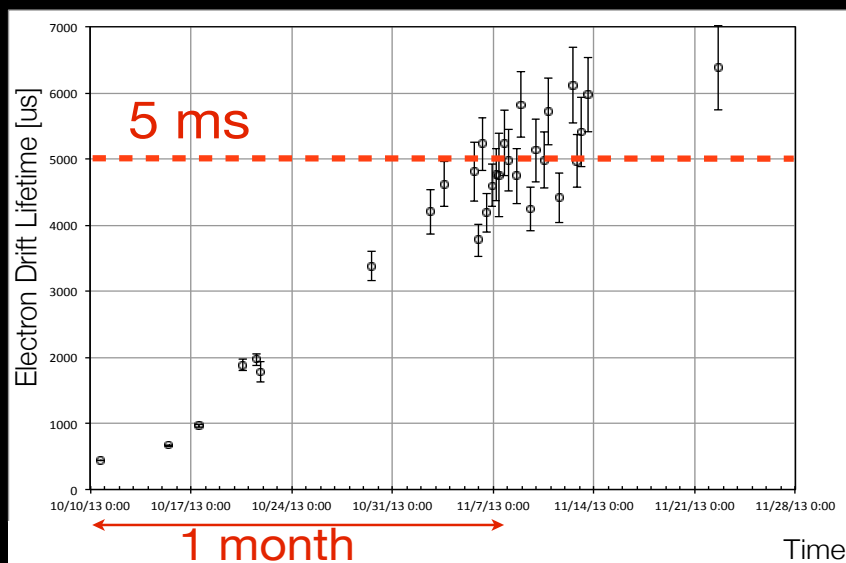
DS-50

All three detectors are currently operating

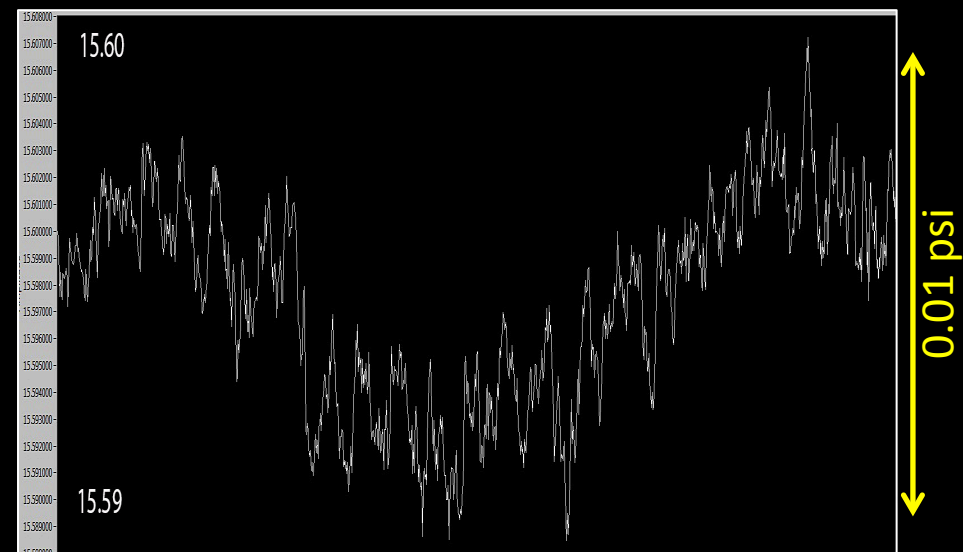


TPC commissioning

- Closed loop Ar recirculation (~30 slpm)
- Gaseous phase purification using commercial getter
- Cryogenic charcoal trap to remove Rn contamination
- Cryogenic system working in an extremely stable condition
- Max cooling power 300 W
- Max recirculation speed 75 kg/day



Electron drift lifetime > 5 ms, compared to max. drift time of ~ 375 μ s



LAr cryostat pressure showed very high stability

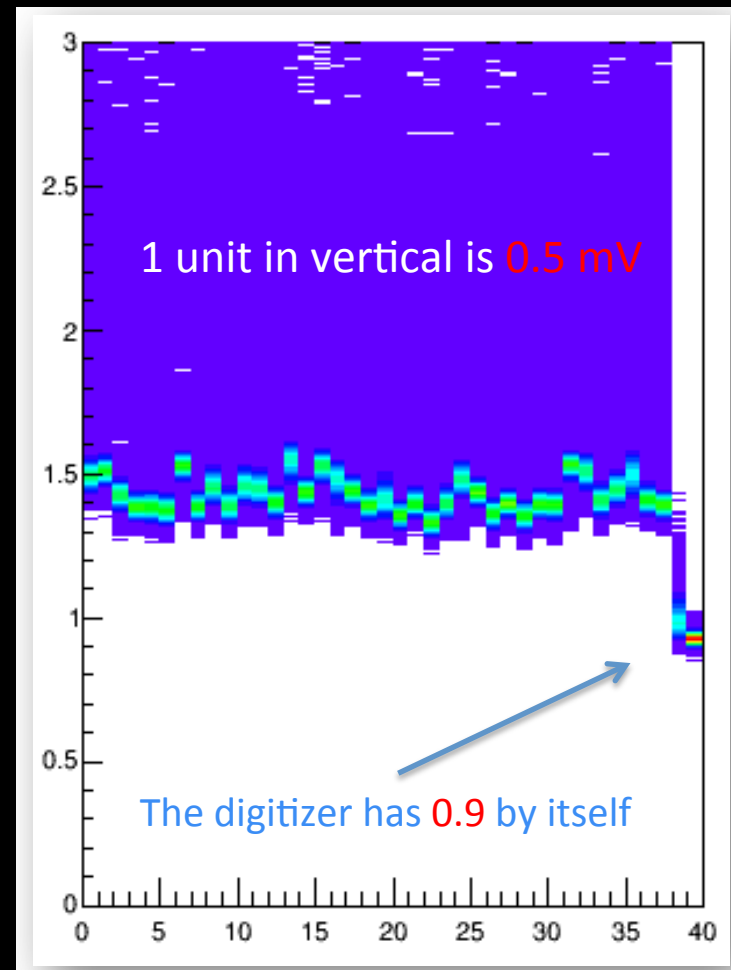
TPC commissioning: residual noise

19 R11065 PMTs working at a gain of 3×10^5 equipped with very low noise cold Amplifiers



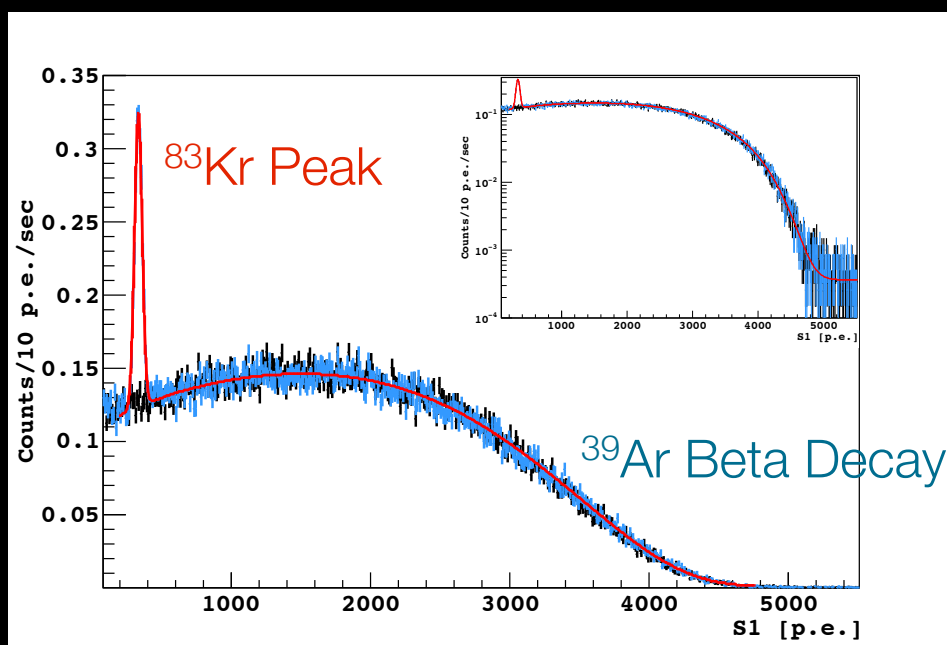
- Electronic noise about $700 \mu\text{V}$
- Contribution of the digitizers $450 \mu\text{V}$

The amplification of 240 V/V provides the same order of noise of the digitizer



TPC commissioning: calibration

- TPC currently filled with atmospheric argon (1 Bq/kg)
→ ^{39}Ar β decay spectrum
- $^{83\text{m}}\text{Kr}$ gas deployed into detector
41.5 keVee and half-time=1.83 hr



ER calibration at null field

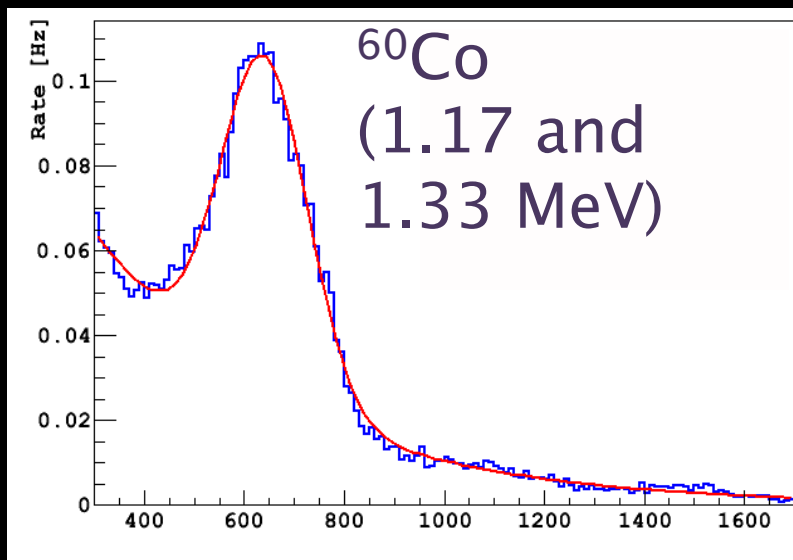
- dominated by ^{39}Ar
- $^{83\text{m}}\text{Kr}$ decays
- Fit to the ^{39}Ar and $^{83\text{m}}\text{Kr}$

Light Yield
8 PE/keV_{ee}

Measured light yield exceeds previous projections of 6 PE/keVee

Neutron Veto commissioning

- Neutron veto setup to trigger on events in the liquid Ar TPC
- Use high energy coincident ^{60}Co events from cryostat stainless steel to evaluate light yield in scintillator

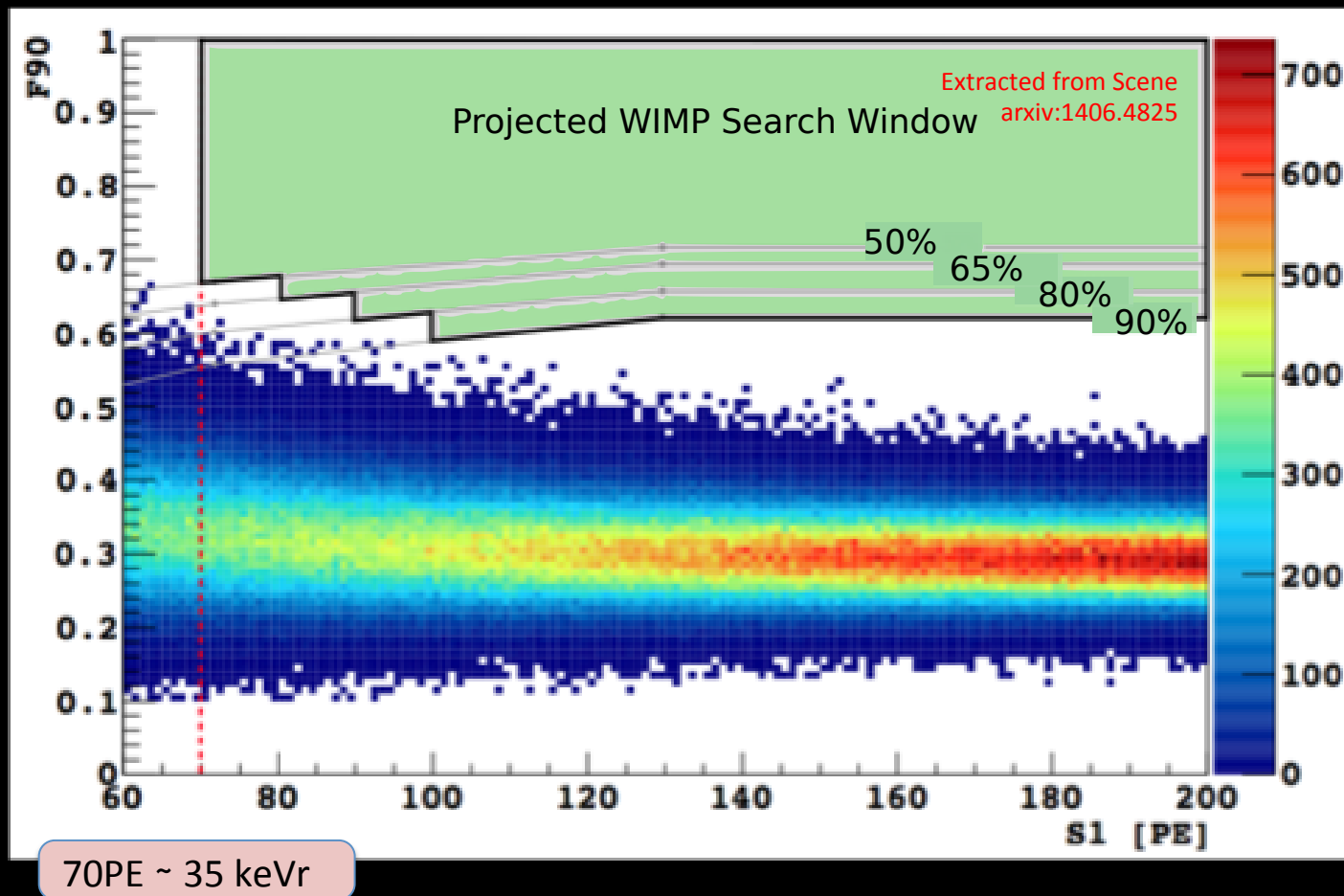


Light yield ~ 0.5 PE/keV
sufficient to detect ~ 50 keV α
from neutron capture

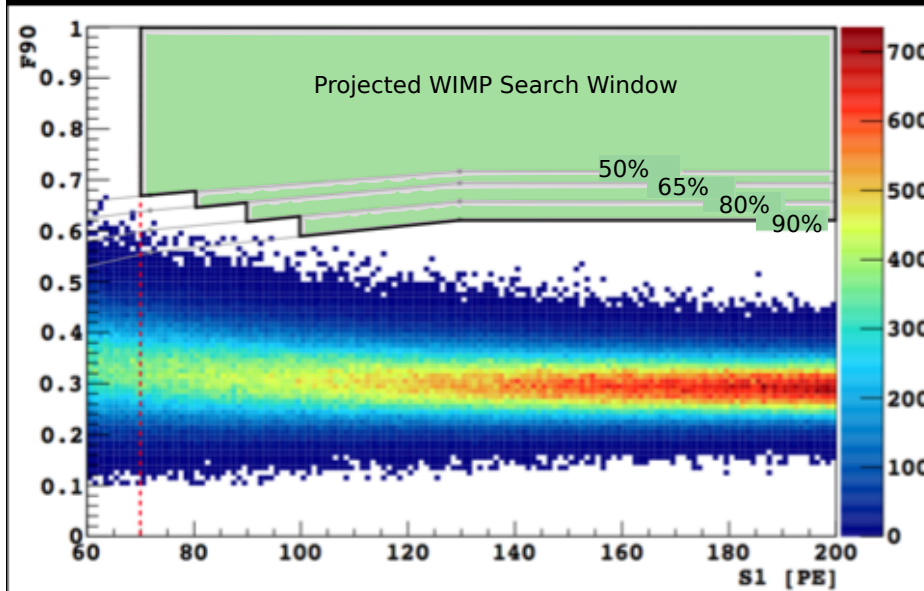
- Found high rate at low energies due to intrinsic ^{14}C in (biogenic) TMB
- Removed TMB from Scintillator (achieved a reduction: 50% \rightarrow 0.16%)
- Identified new batch of low- ^{14}C (underground) TMB

Initial exposure (280 kg-days)

We have **PROVEN** that S1 PSD at 200 V/cm (+ z fiducialization from S2) can efficiently suppress the dominant ER background that we expect in **2.6 years** of DS-50 UAr run, while maintaining high acceptance for WIMPs.

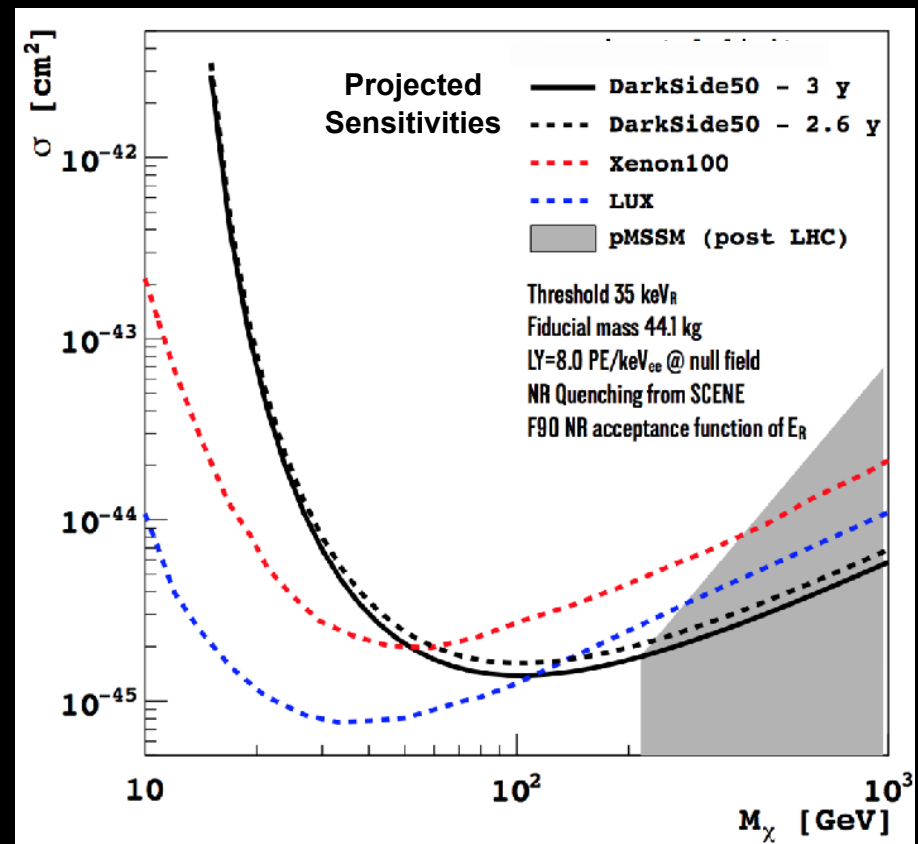


DS-50 projected sensitivity



Assumptions

- PSD as demonstrated
- No S2/S1 rejection
- Fiducial mass 44kg (z-cut only)
- NR energy & pulse shape taken from SCENE



DS-50 summary and plan

- **Detector running since Oct. 2013 with atmospheric Ar**
- **LAr TPC exceeded desired light yield**
- **Demonstrated excellent PSD performance background from ^{39}Ar : 2.6 years of DS-50**
- **Currently acquired and analyzed 2500 kg-days (55 days)**
 - Data being used to improve understanding of S2 signal, x-y position reconstruction, S1-S2 correlations
- **Neutron veto: High rate of ^{14}C found in TMB**
 - During last 2.5 months performed operations to remove TMB
 - **TMB removed** – achieved **600x** reduction - Will replace with low- ^{14}C TMB
- **Source calibration**
 - Plan to acquire gamma and neutron data very soon
 - Source insertion system installation on going
- **Underground argon**
 - Switch to using underground argon foreseen at the beginning of 2015

First physics results to be published soon

Darkside future prospects

G2 ($s \approx 10^{-45} \text{cm}^2$) **US NSF-DOE:**

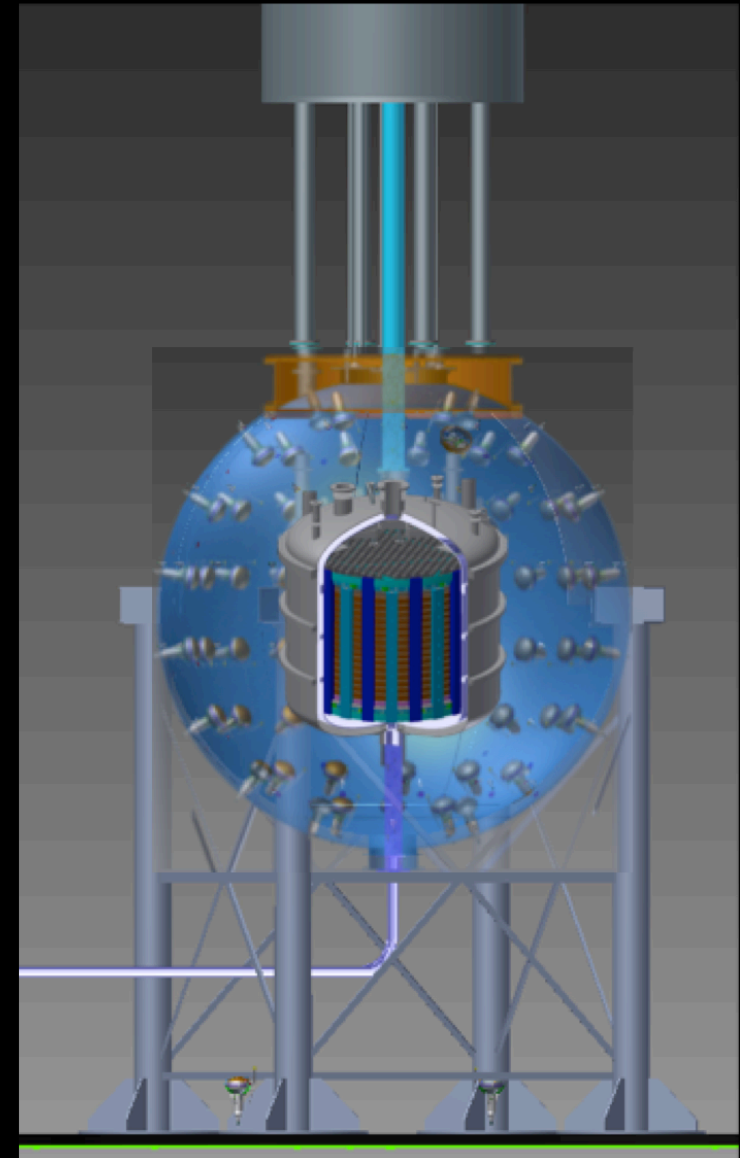
- DarkSide-G2 proposal declined (LZ, SuperCDMS, ADMX funded)
- The program will include R&D to test and develop technologies for future experiments, consistent with the recent P5 recommendations.

G3 ($s \approx 10^{-47} \text{cm}^2$) **EU H2020 INFRADEV-1-2014** (Research Infrastructures)

- DARWIN proposal (Xenon, DarkSide, ArDM groups) for multi-ton noble liquid detector submitted on Sep. 2

R&D activities:

- **Underground Argon provision and purification**
- Cryostat design materials for low radioactivity (Stability of pressure in gas region, Thermodynamics for smooth liquid-gas interface)
- **Photosensors**
- **Front-End & Digitizer (cables, noise, radioactivity)**

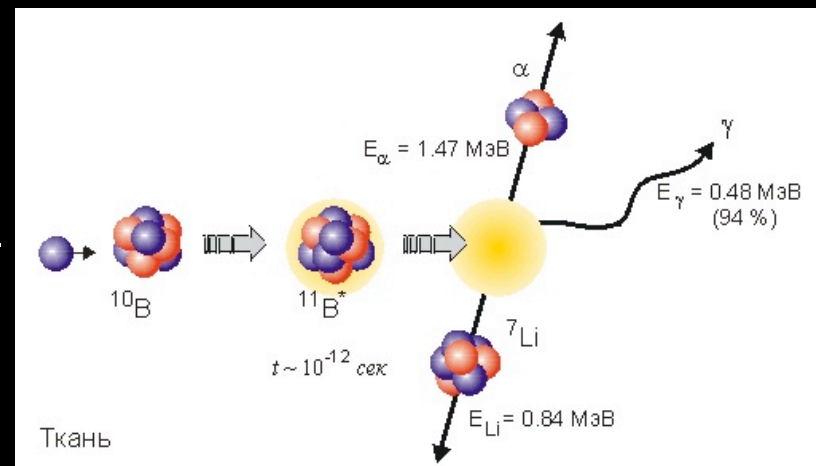


Thank you



Borated Liquid Scintillator

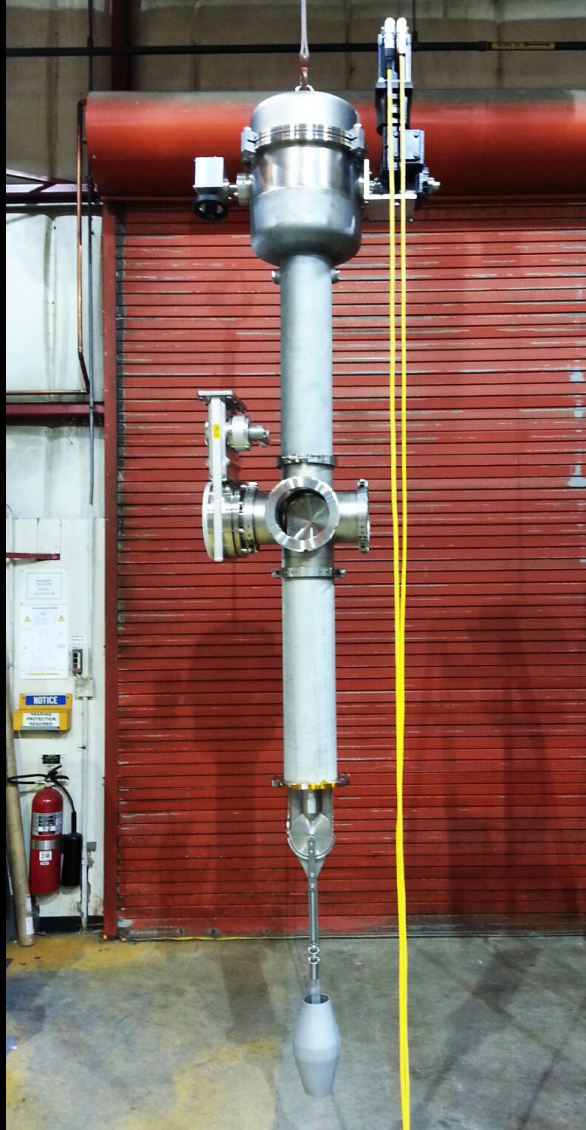
- High neutron capture cross section on boron allows for compact veto size
- Capture results in 1.47 MeV α particle - detected with high efficiency
- Short capture time ($2.3 \mu\text{s}$) reduces dead time loss



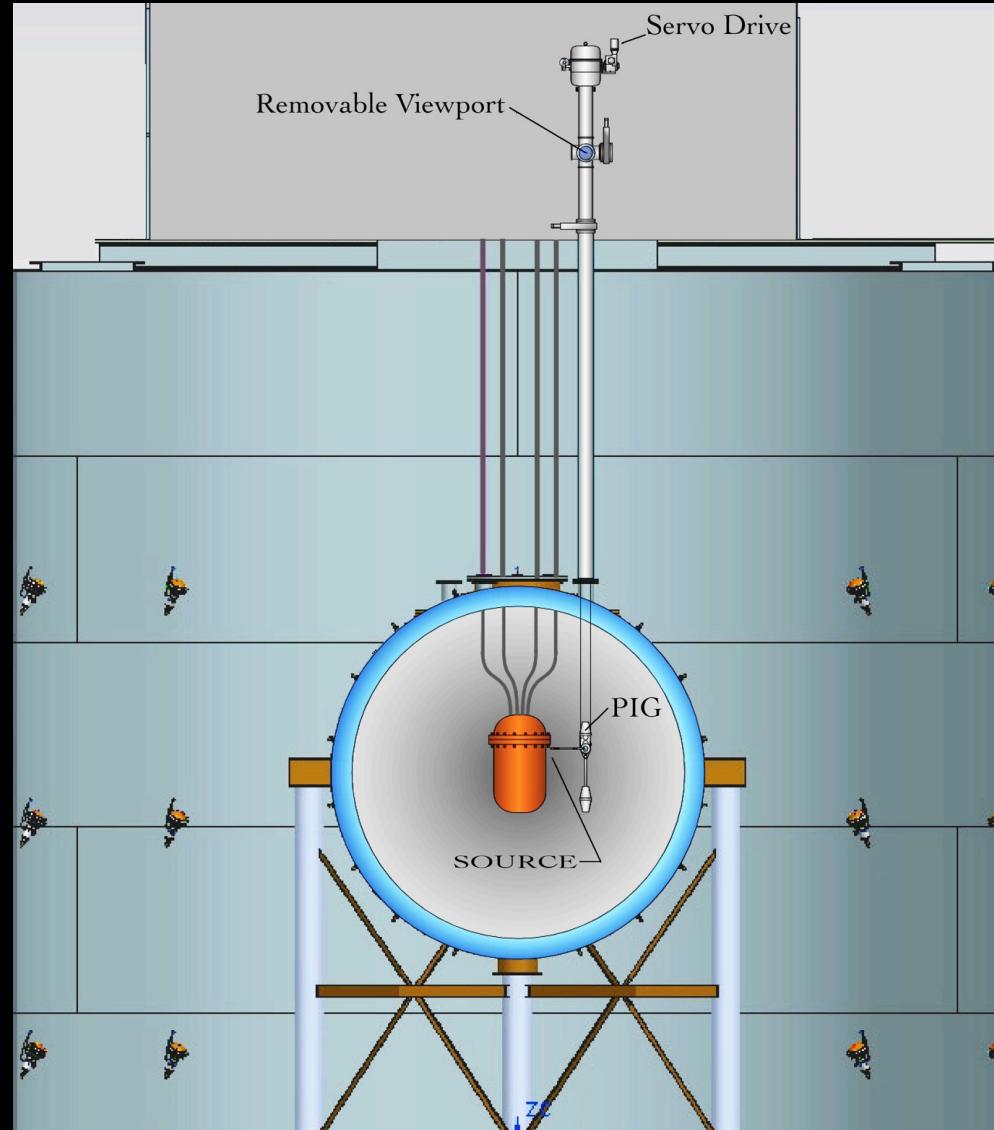
	Veto Efficiency (MC)
Radiogenic Neutrons	$> 99\%^*$
Cosmogenic Neutrons	$> 95\%$

Nuclear Instruments and Methods A 644, 18 (2011)

Calibration source system



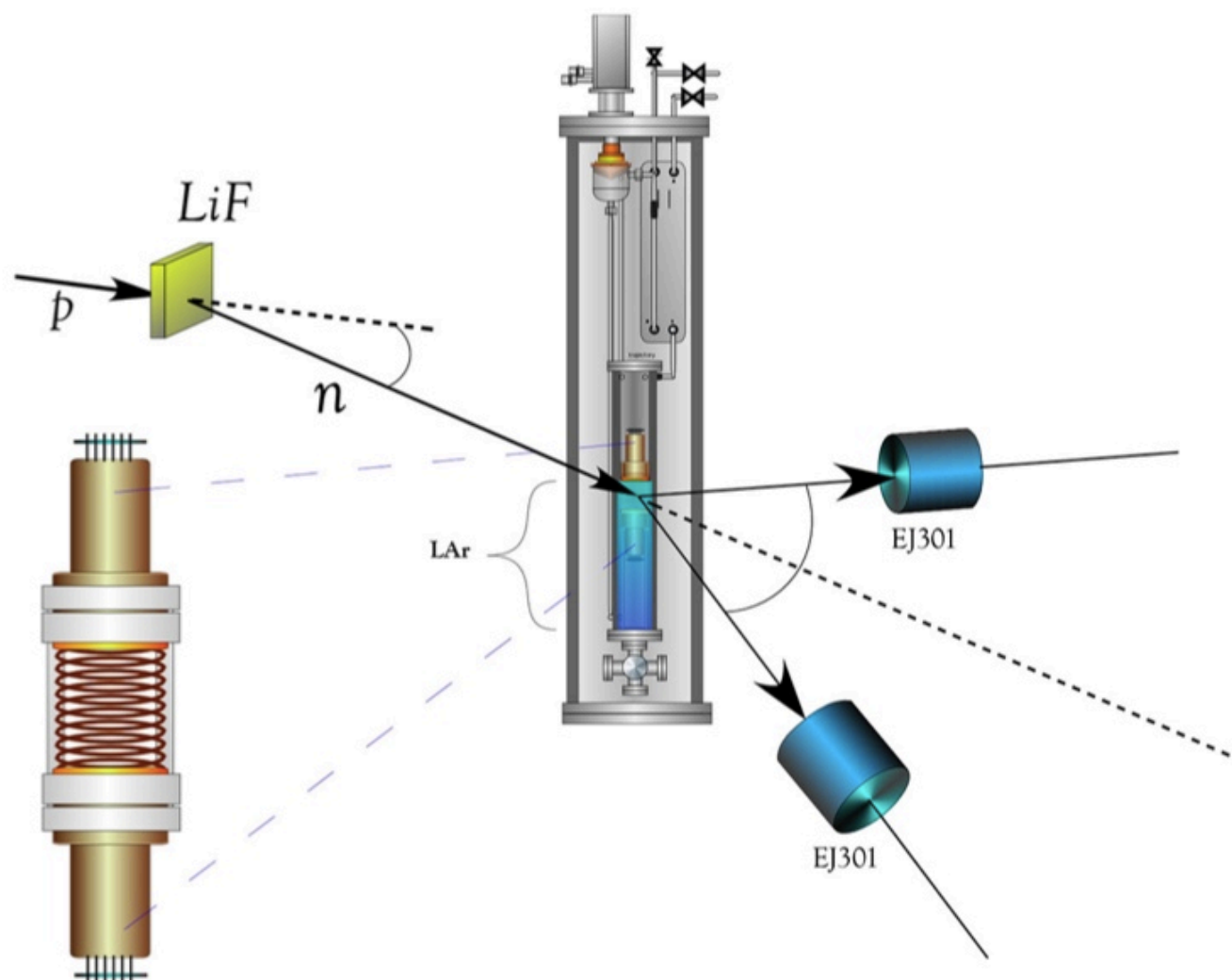
B. Rossi - 29 September 2014



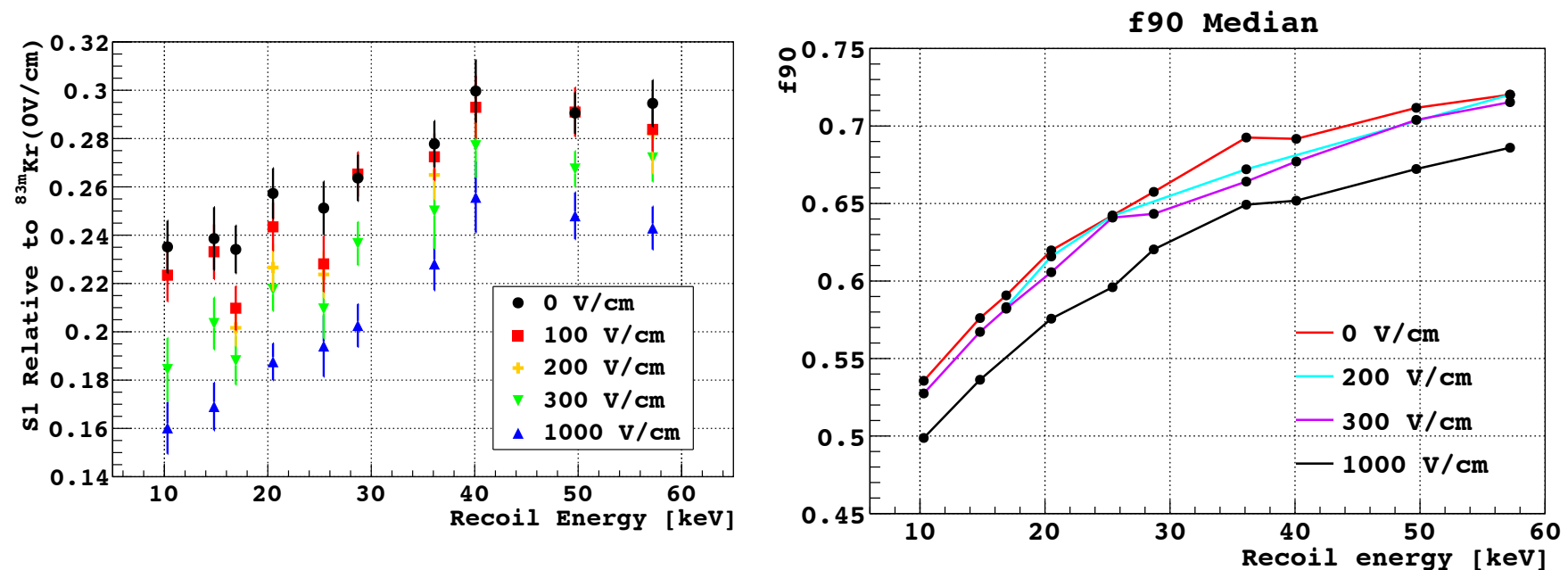
Ricap 2014 - Noto (SR) - Italy

SCENE Experimental Scheme

(Scintillation and Ionization Efficiency of Noble Elements)



Nuclear Recoil Scintillation Yield and Pulse Shape vs Drift Field



Paper on most recent results:

H. Cao et. al.

arXiv:1406.4825