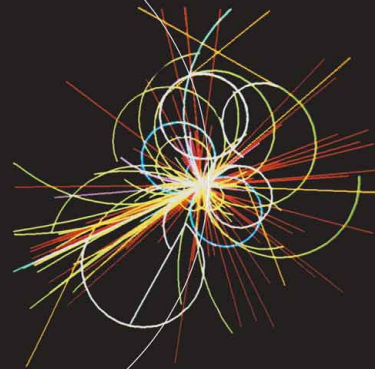


Frontier Detectors for Frontier Physics

11th Pisa meeting on
advanced detectors

La Biodola • Isola d'Elba • Italy
May 24 - 30, 2009



An antineutrino detector to monitor nuclear reactor's power and fuel composition

Marco Battaglieri

*Istituto Nazionale di Fisica Nucleare
Genova- Italy*

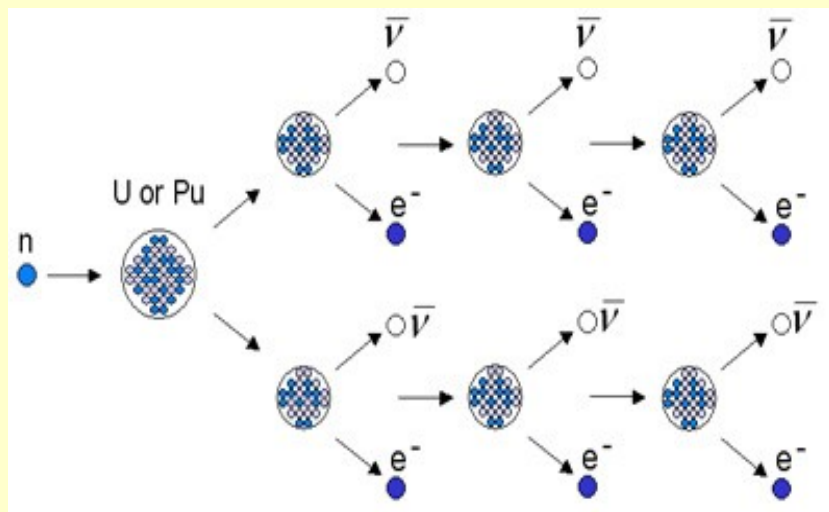
Motivations

How to detect $\bar{\nu}$

Detector set-up

R&D and prototyping

Power reactor core antineutrinos



Antineutrinos are directly produced in fission ($\sim 10^{20}$ $\bar{\nu}/s$)

Counts/day in organic scintillator detector (water-like)

$$10^6 \times V(m^3) \times P(GW) / D^2(m^2)$$

$$\sim 6000 \text{ ev/d} @ 1m^3 \text{ P}=3GW \text{ D}=20m$$

$\bar{\nu}$ flux proportional to power

↪ Integrated counts allow to monitor the reactor power

$\bar{\nu}$ energy spectrum related to isotope decay

↪ antineutrino energy spectrum reveals the reactor core isotopic composition (²³⁵U and ²³⁹Pu)

Why antineutrinos?

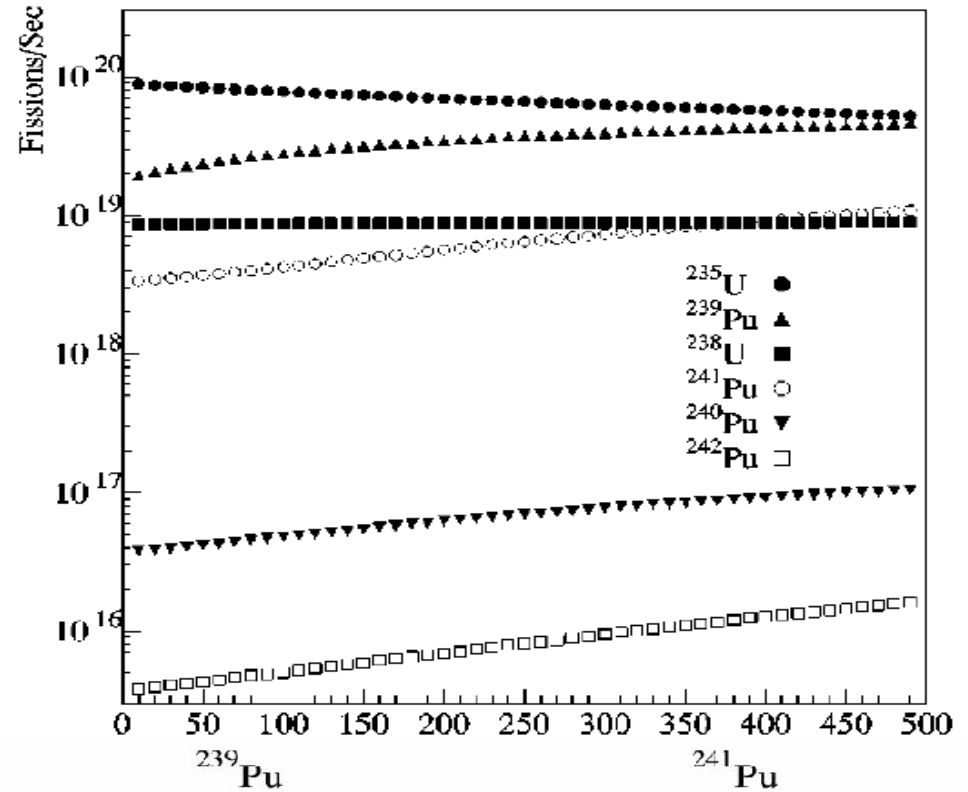
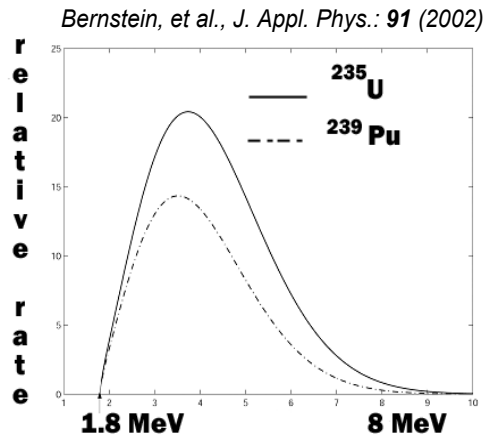
This is NOT the simplest way to monitor a reactor but ...

- **independent on-line monitor** of the reactor power
- the measurement is performed **outside the core** (safety issues) and does not affect the normal plant operations
- detector installation **does not need any engineering work** and can be done anytime in any plant
- antineutrino can not be shielded and are **only sensitive to fission material**
- direct measurement of core nuclear activity and **fuel burn-up**
- early detection of unauthorized **plutonium production** and subtraction (sensitivity down to 10th kg)

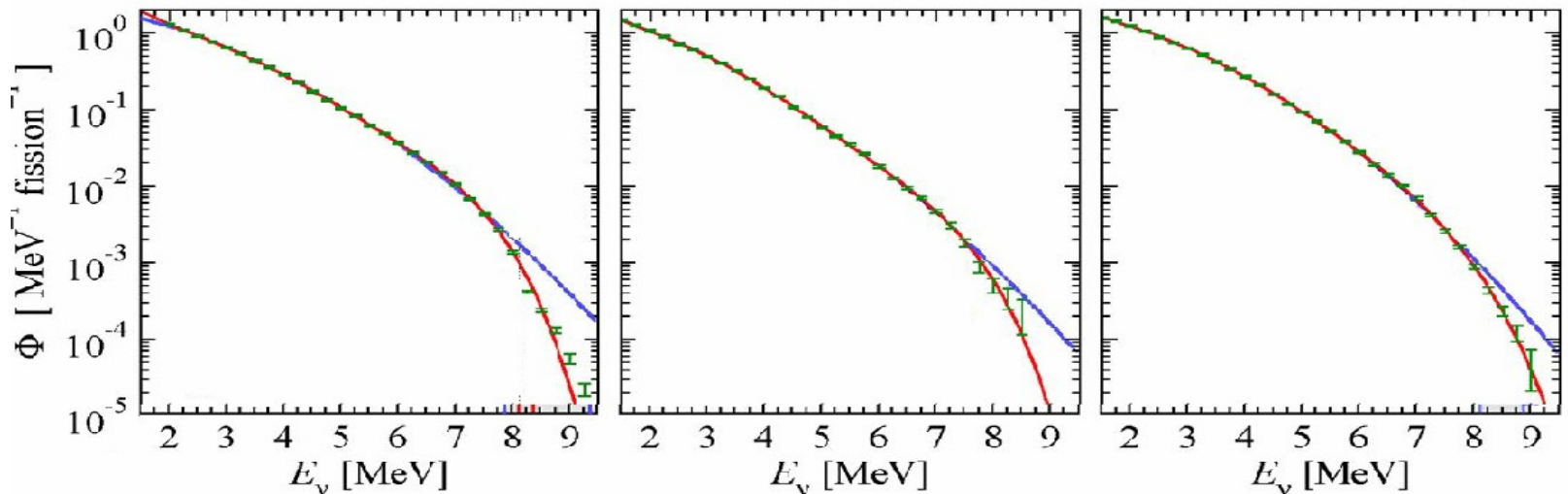
Core isotopic composition

★ Antineutrino rate varies with time and isotope

★ As reactor fuel burns the composition change



Antineutrino energy spectrum



How to detect antineutrinos?

both detected

Inverse beta decay: $\bar{\nu} p \rightarrow \beta^+ n$

$E_{\bar{\nu}}$

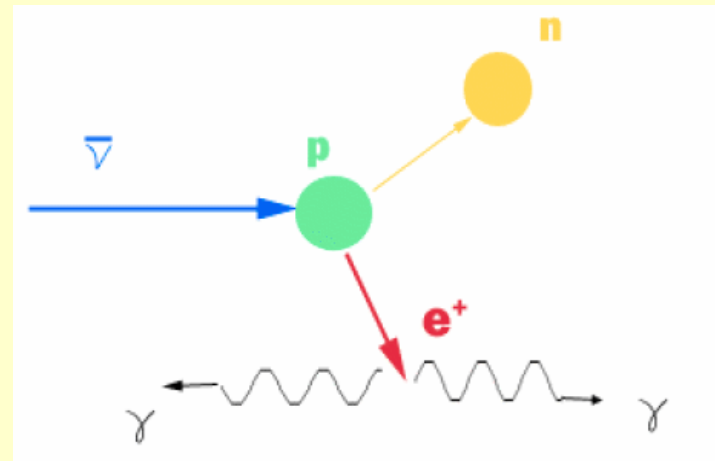
- **Reaction threshold $E_{\bar{\nu}} = 1.8 \text{ MeV}$**
→ reactor-off contribution (long-life isotopes emitting low E_n) is negligible!
- **$E_{\bar{\nu}}$ (1.8 - 7 MeV)**

β^+

- **$E_{\bar{\nu}} \sim E_{e^+}$**
- **few MeV positron: ionization**
- **$e^+ e^- \rightarrow \gamma\gamma$ ($2 \times 0.511 \text{ MeV}$)**
- **prompt signal ($t < 1 \mu\text{s}$)**

n

- **thermalization ($np \rightarrow np$)**
- **capture on nucleus A**
- **γ cascade ($E_{\text{tot}} = 2-8 \text{ MeV}$)**
- **delayed signal ($\tau \sim 5-100 \mu\text{s}$)**



Measured:

- ★ **energy deposition**
- ★ **time interval**

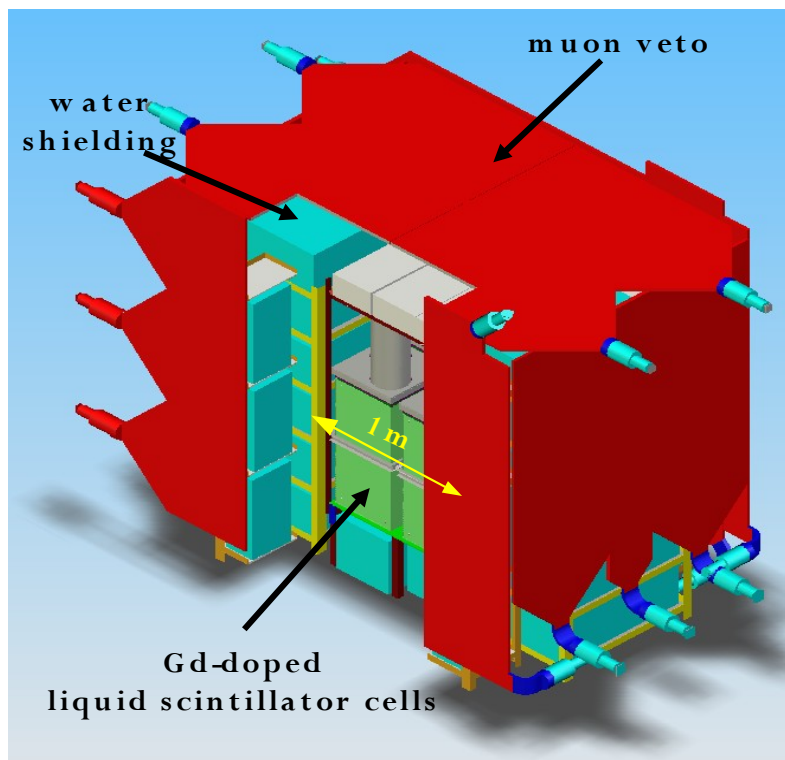
Detector requirements

Inverse beta decay: $\bar{\nu} p \rightarrow \beta^+ n$

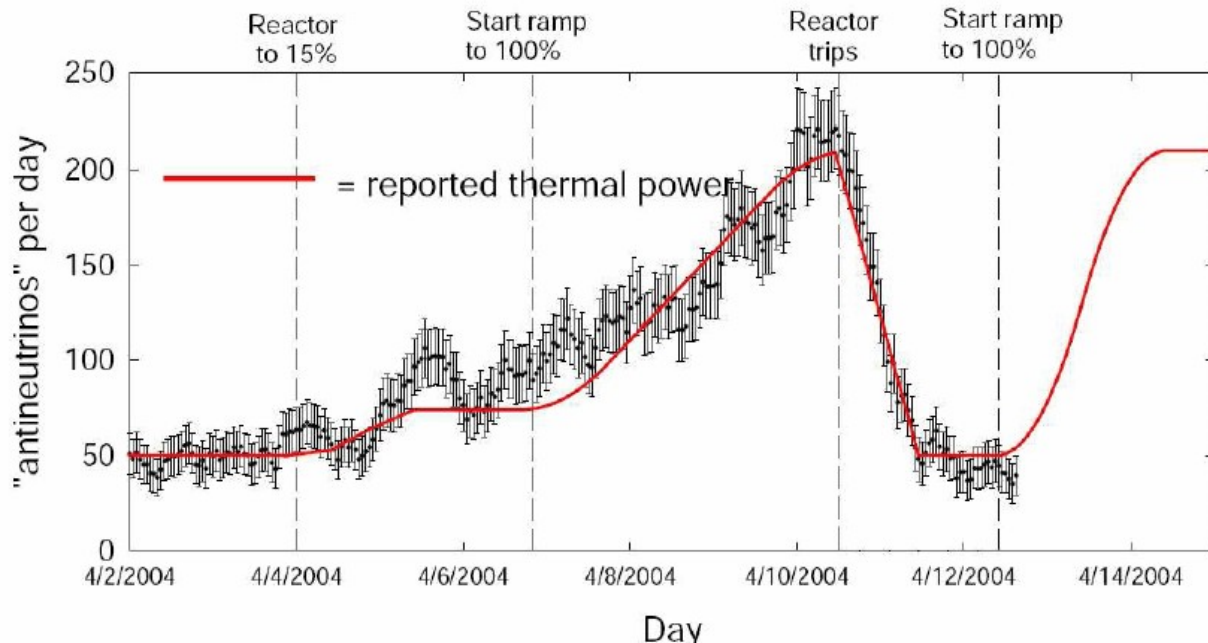
- proton rich target
- active target to detect e^+ ionization
- $\sim 1 \text{ m}^3$ target volume (low $\sigma_{\bar{\nu} p \rightarrow \beta^+ n} \sim 10^{-43} \text{ cm}^2$ @ $E_{\bar{\nu}} \sim \text{MeV}$)
- high segmentation for background rejection
- homogeneous target to have good e^+ energy resolution
- clear signature for neutron capture
- small n-capture time (τ_{cap}) to reduce e^+/n time coincidence window
- good cosmic rays background rejection

Does it work?

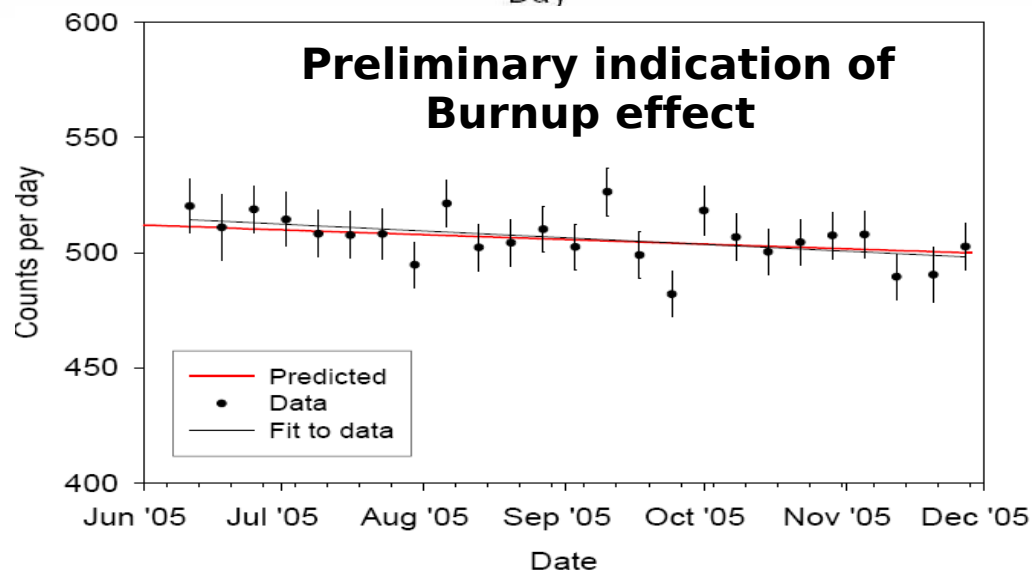
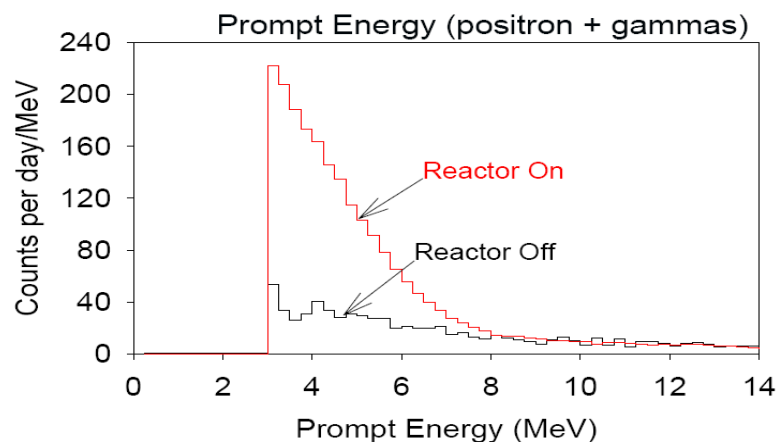
Sandia National Laboratory antineutrino detector (SONGS1) S.Onofre (Ca) 3.5 GW commercial power plant



★ Daily power monitoring using only antineutrinos



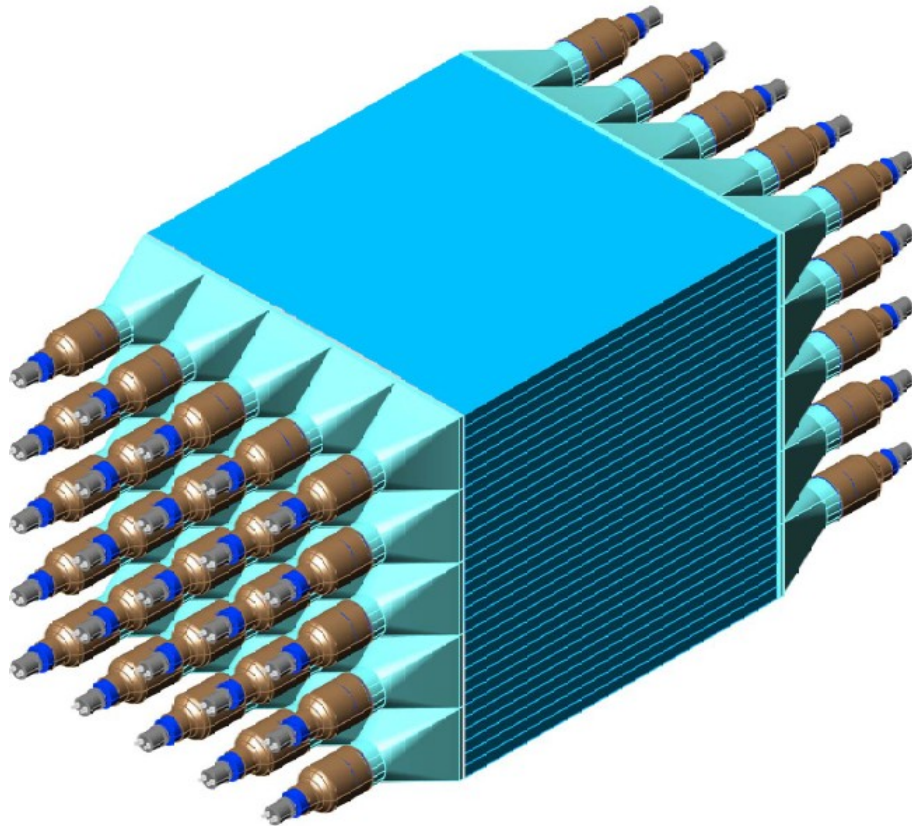
Detection efficiency $(10.7 \pm 1.5) \%$



CORMORAD

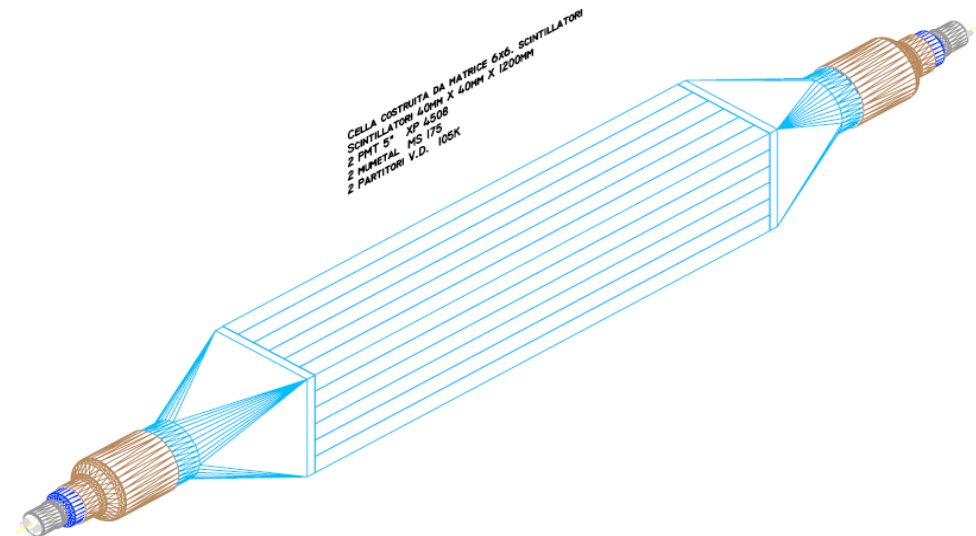
CORE MONITORING by ANTINEUTRINO DETECTION

★ Project part of the INFN-E strategic plan and in the INFN-Ansaldo agreement



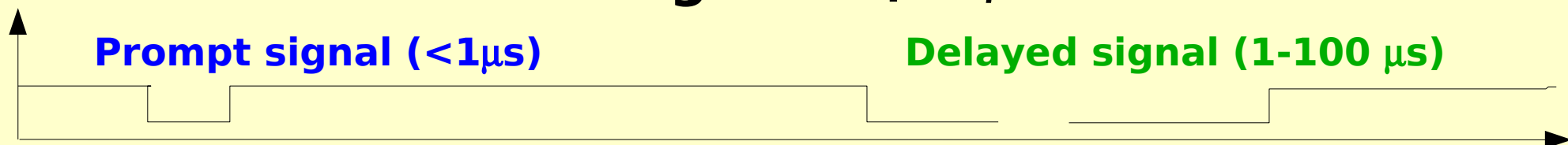
- 3" PMT on Each side
- 10-15% of produced light collected
- minimum light attenuation
a.l. of plastic scintillator
~300cm@425nm
- A total of 36+36 PMTs
- ADC and TDC read-out

- Active volume: $0.6 \times 0.6 \times 0.6 \text{ m}^3$
- $(3 \times 3) \times 4 \times 4 = 144$ plastic scintillator bars wrapped in $12.5\mu\text{m}$ Mylar-Gd foils
- Bar sizes: $5 \times 5 \times 60 \text{ cm}^3$



Reaction selection and bg rejection

Signal $\bar{\nu} p \rightarrow \beta^+ n$



- antineutrino interact with H in plastic
- positron releases energy within a cell
- Then annihilates
- two gammas mainly interact via Compton scattering
- $E_{\nu} \propto E_{\text{Prompt}} = E_{\text{ion}} + 2 \times 0.511 \text{ MeV}$
- $N_{\text{hit}} > 2$

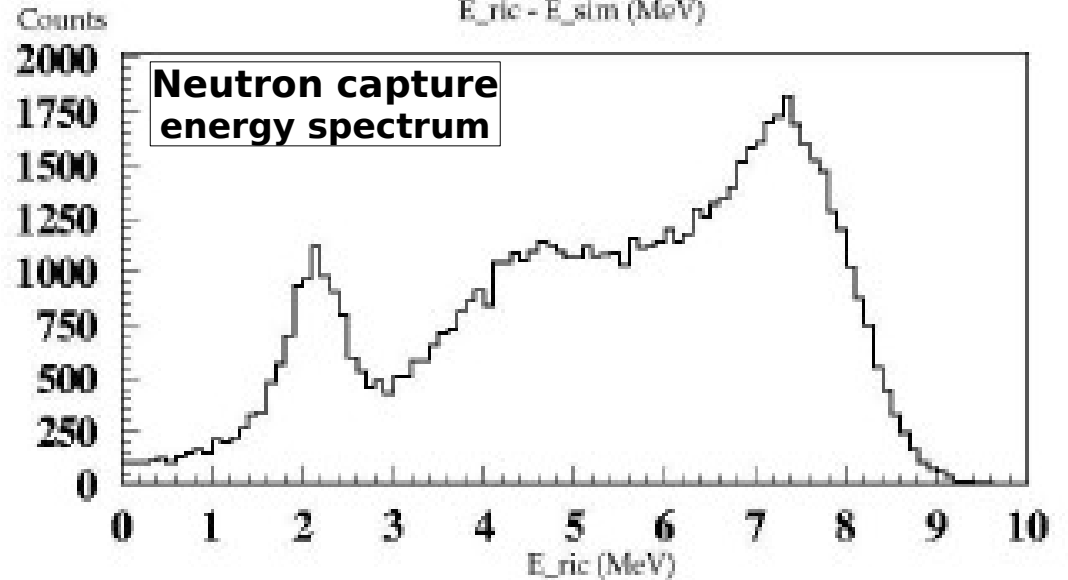
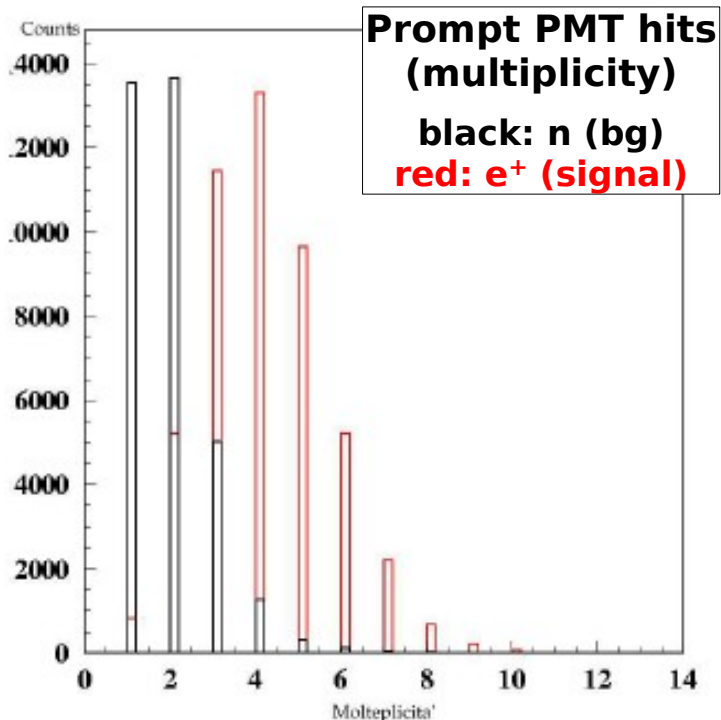
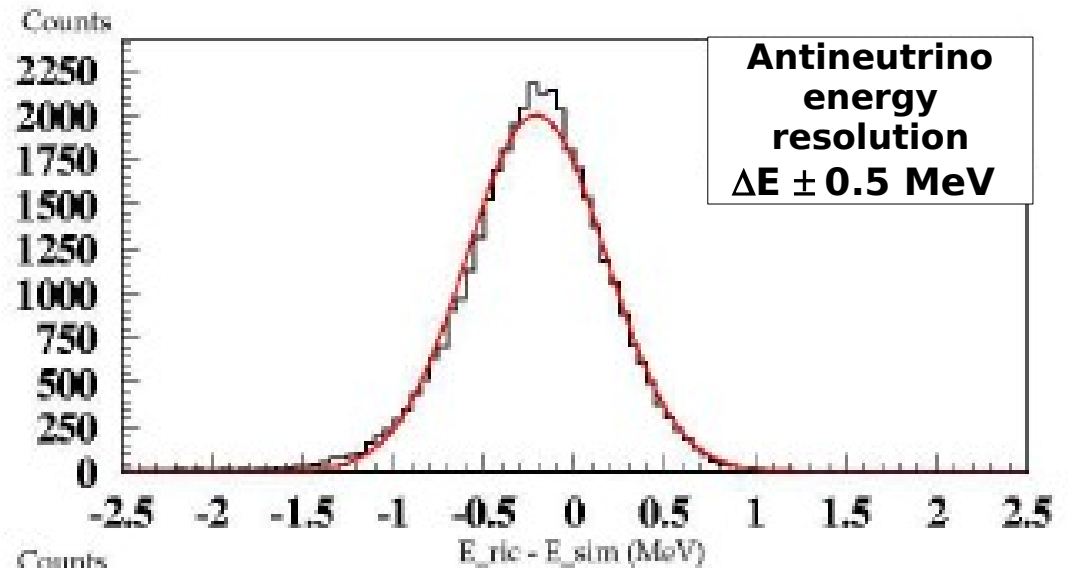
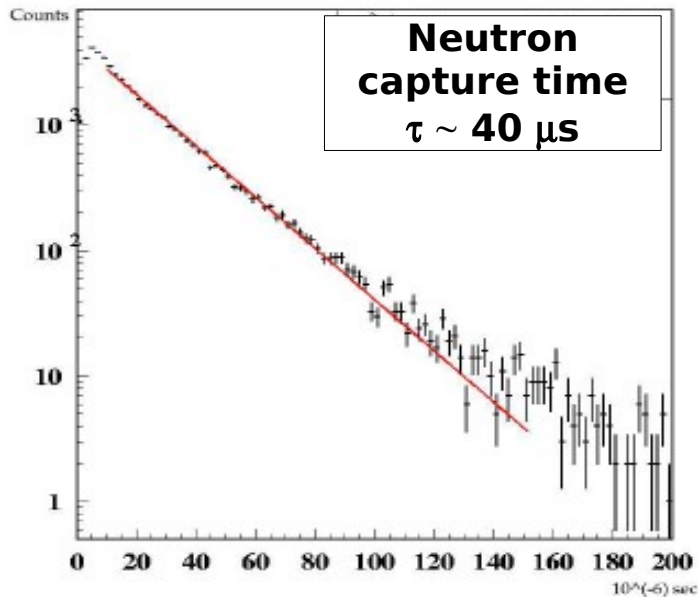
- n thermalizes ($np \rightarrow np$)
- tenth of cm ($>$ cell size)
- capture on A
- delayed signal ($\tau \propto 5-100 \mu\text{s}$)
- $E_{\text{Delayed}} = E_{\text{elastic}} + 7-8 \text{ MeV}$
- High hit multiplicity

Background

- Atmospheric and solar → negligible
- Cosmic muons
random coincidence in a wide time-window → tracked and rejected in off-line analysis
- Residual radioactivity in PMTs and surrounding material → tracked and rejected in off-line analysis
- Spallation neutron
slowed down by elastic np scattering
p fast signal (similar to e^+)
n delayed signal (as in inverse β decay) → the real issue!
same time-coincidence signature
rejected by using hit multiplicity
need segmented detector !

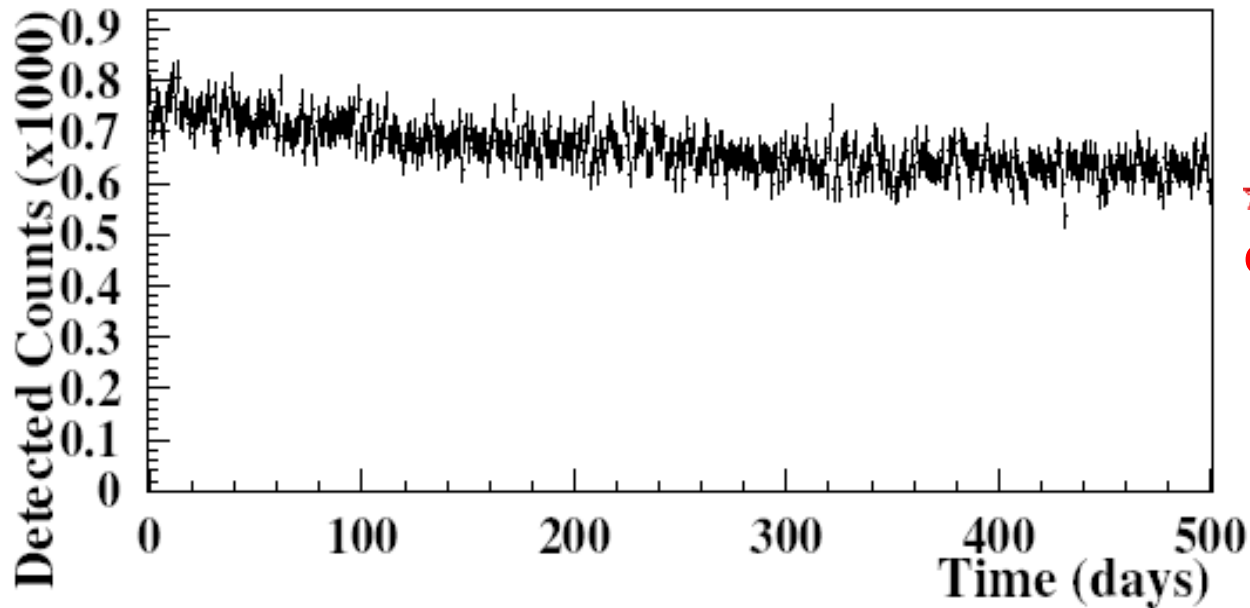
Expected performance and results (GEANT4)

R. De Vita

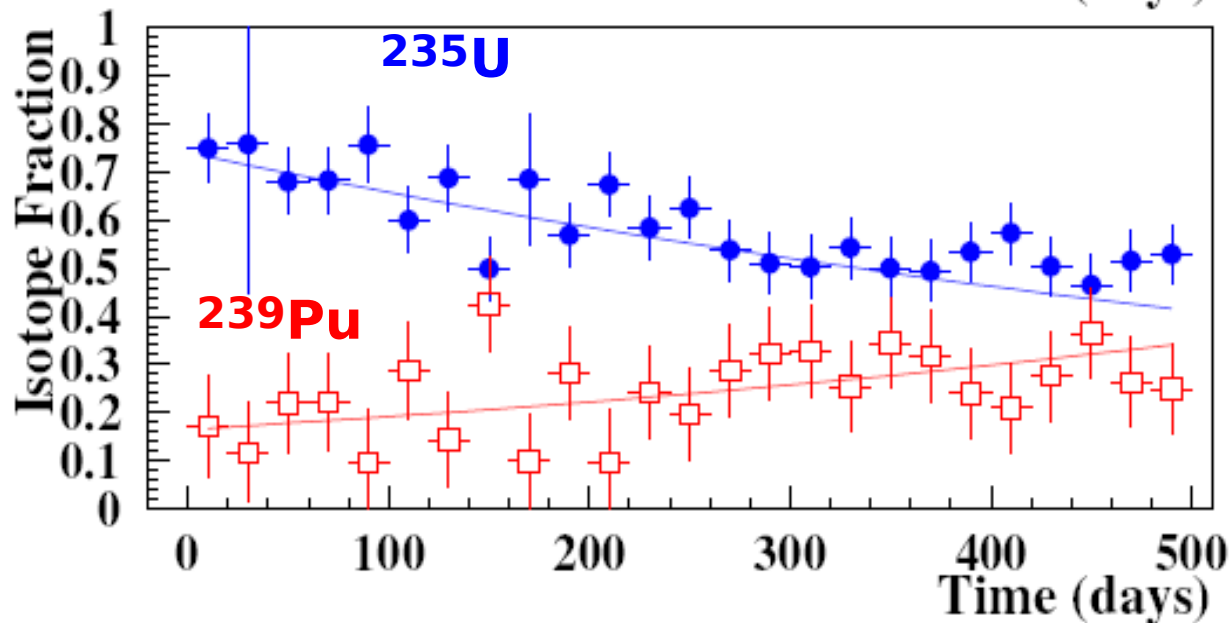


Detection Efficiency $\sim 40\%$

Expected performance and results (GEANT4)



★ Power plant monitor on daily base

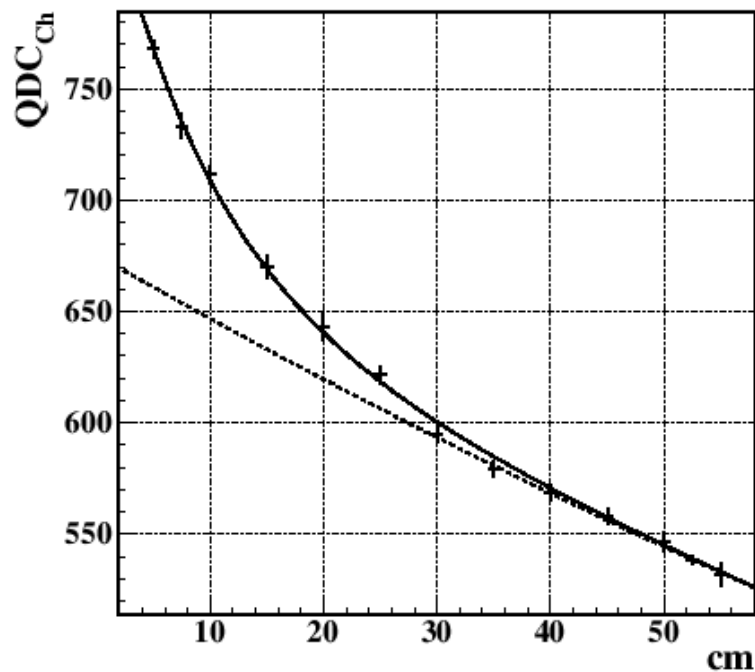


★ Fuel burn-up integrating over 10 days

R&D and prototyping

G. Firpo

★ Attenuation length (cm)



BC525: 26.3 ± 0.7 (HiFi)

32.1 ± 1.3 (3M)

BC408: 294 ± 93 (Mylar)

EJ200: 286 ± 45 (3M)

243 ± 45 (HiFi)

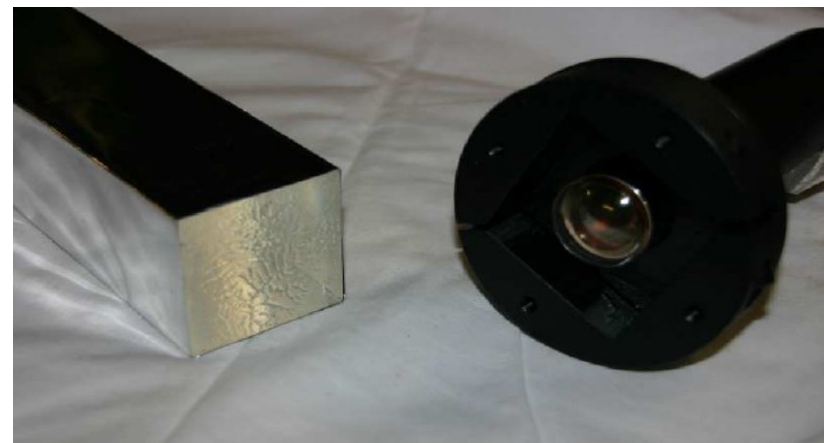
232 ± 39 (Mylar)

NE110: 106 ± 12 (Mylar)

JINR: 86 ± 14 (Mylar)

Liquid

Plastic



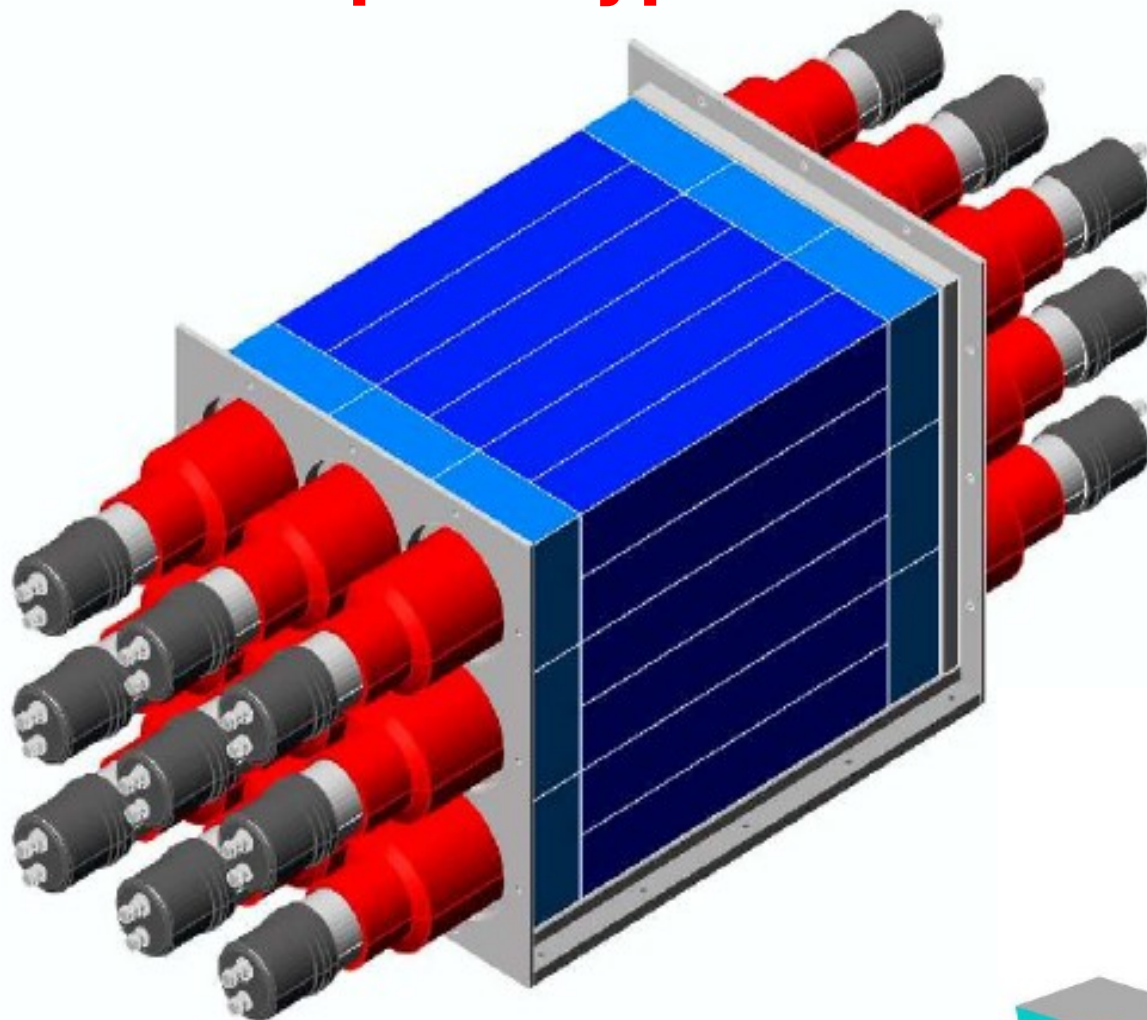
★ Energy resolution (e- @ 478 keV)

BC525	BC408	NE110	EJ200
30%	14%	18%	16%

★ Time resolution: $\sigma_t \sim 170$ ps

★ Space resolution: $\sigma_x \sim 8$ cm

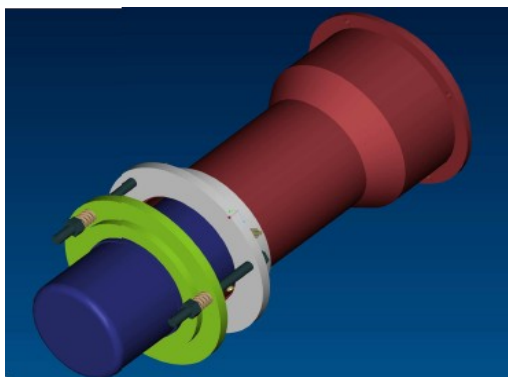
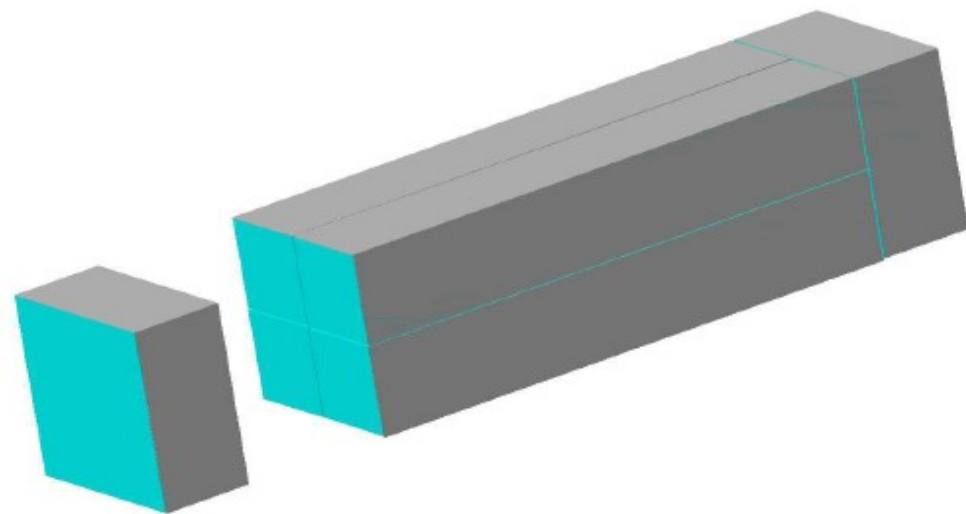
Scale 1:3 prototype



★ Size: 40 x 30 x 30 cm³

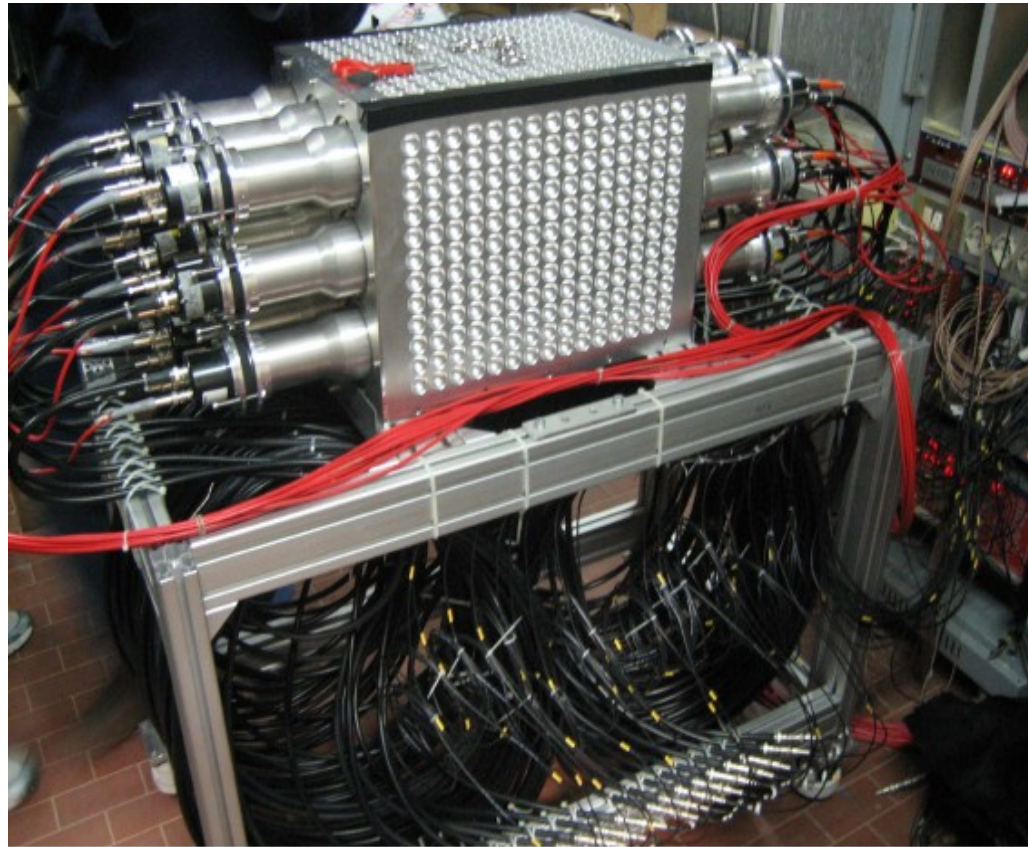
Prototype cell

- ★ 4 30x5x5 cm³ NE110 bars
- ★ 1 5x10x10 cm³ NE110 block
- ★ 12.5 μm Gd foils wrapping



★ Light read-out: 18 Photonis XP2312 3" PMTs

R&D and prototyping

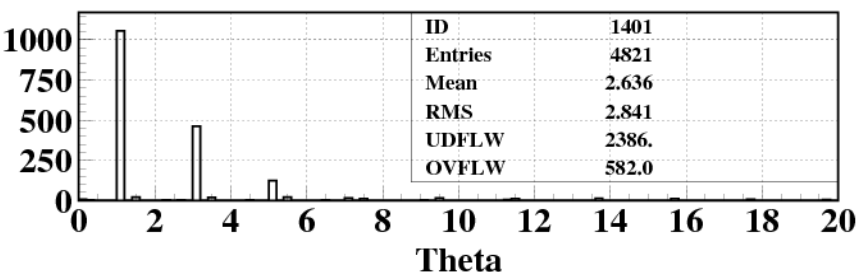
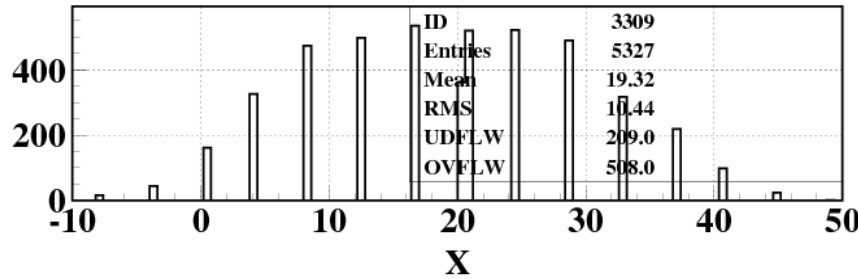
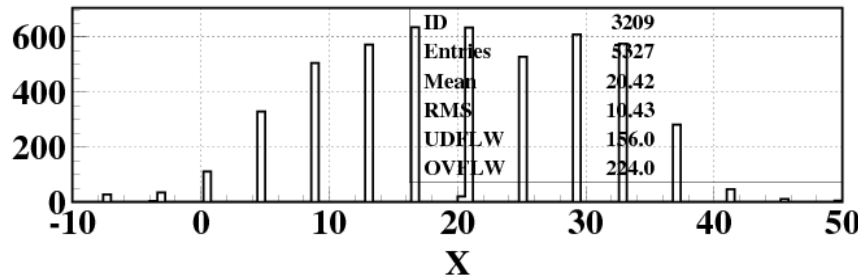
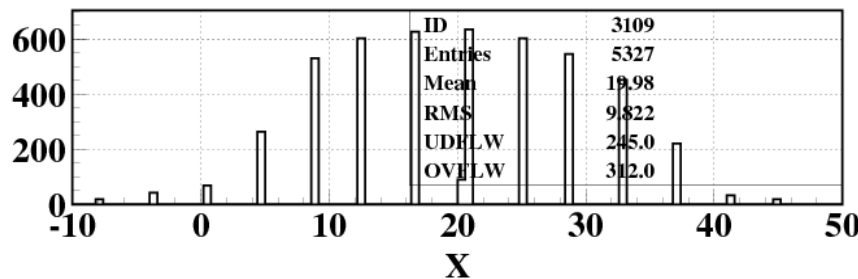
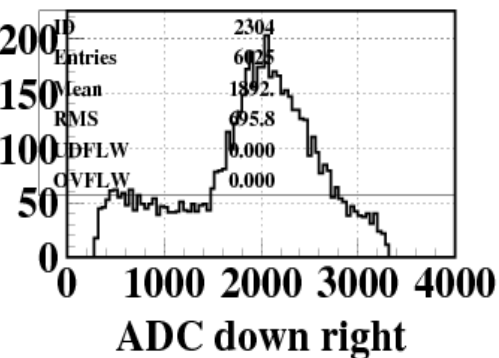
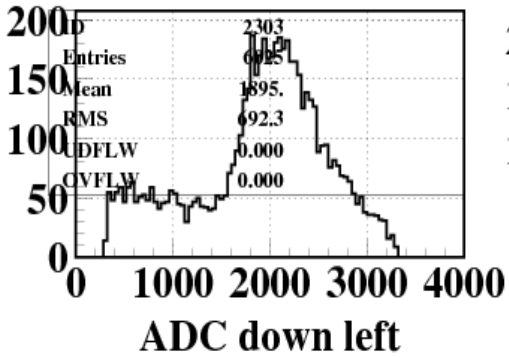
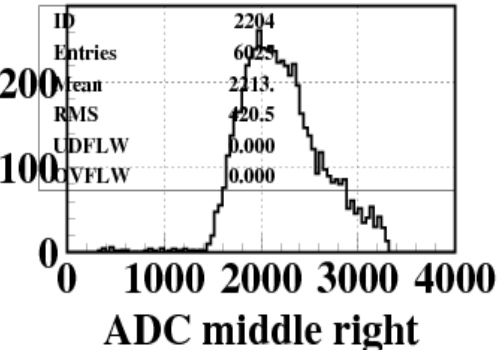
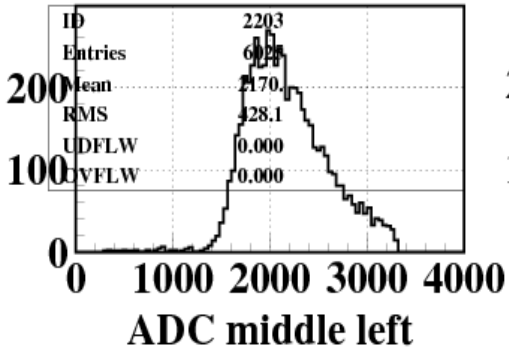
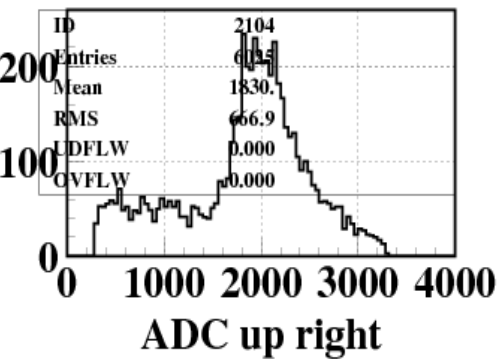
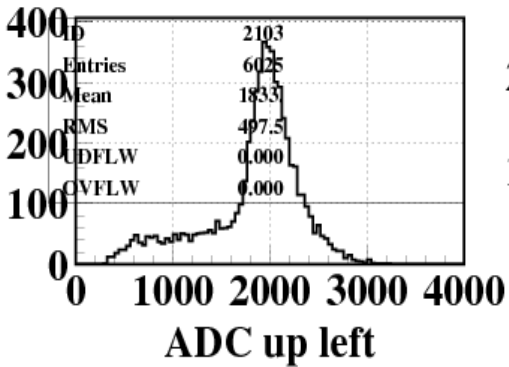


Energy and time calibration using cosmic rays

R. De Vita

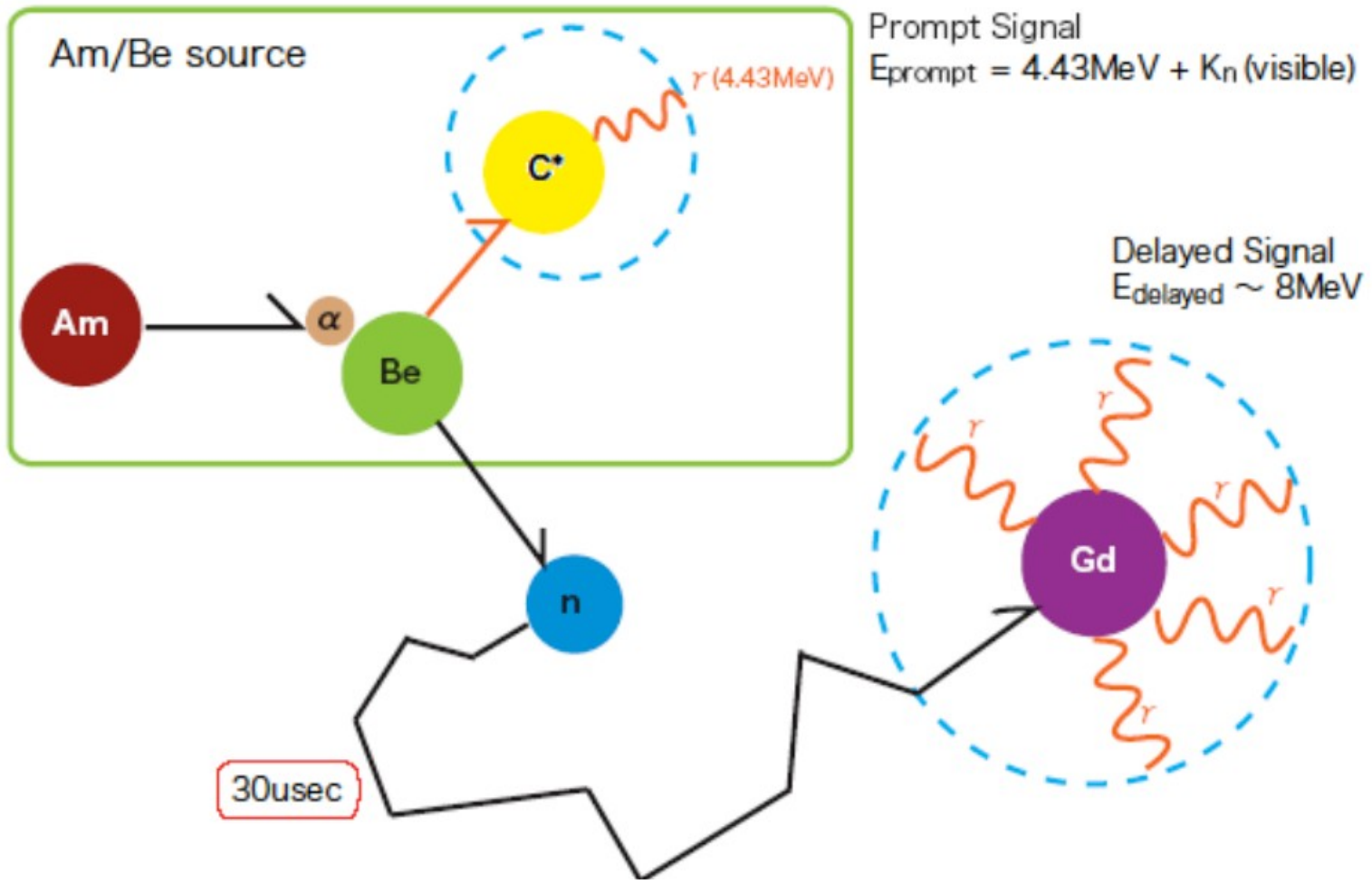
- Self calibration (Up/Middle/Down crossing)
- 12dB attenuation (x0.25) to avoid saturation
- All peaks at $ADC_{ch} = 2000$ ($1.35 \cdot 10^{-3}$ MeV/ch)

- Position = $(V \times \Delta T_{LR} + 40)/2$ $V=13$ cm/ns
- TDC res = 625ps $\rightarrow \Delta x \sim 5$ cm



Calibrations: AmBe source

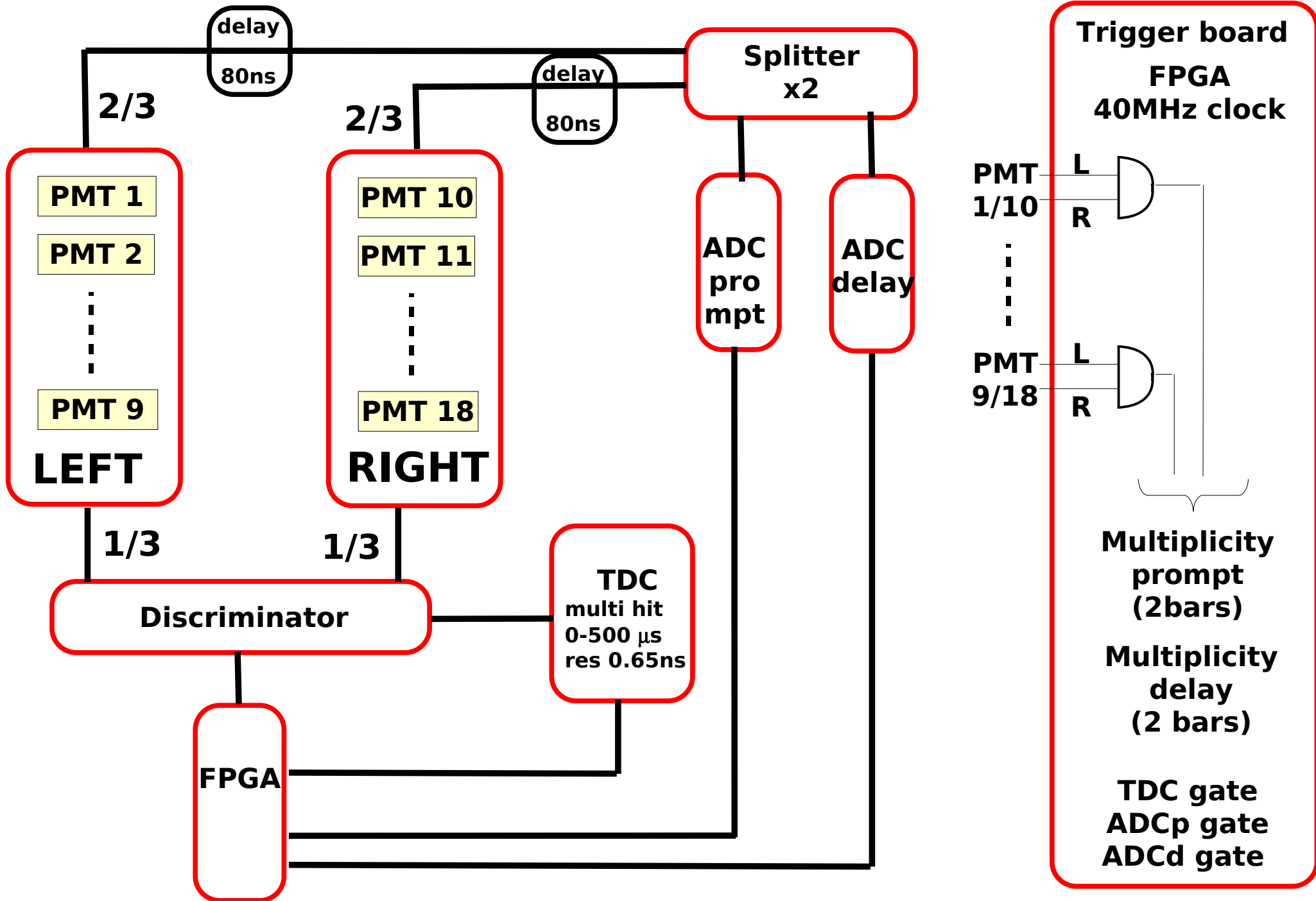
The prompt (gamma) and delayed n capture mimics the anti-neutrino signal



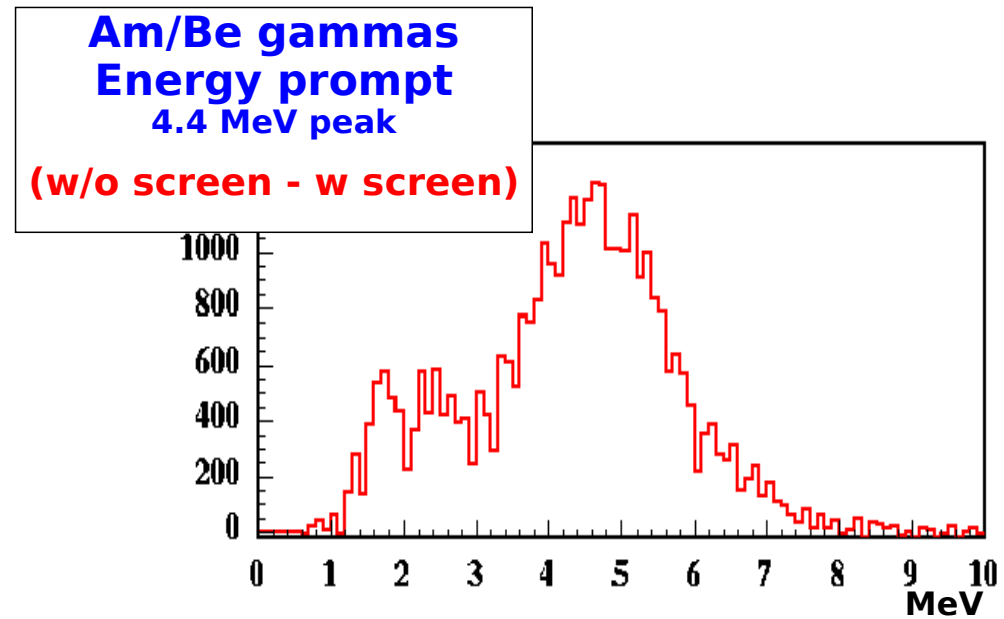
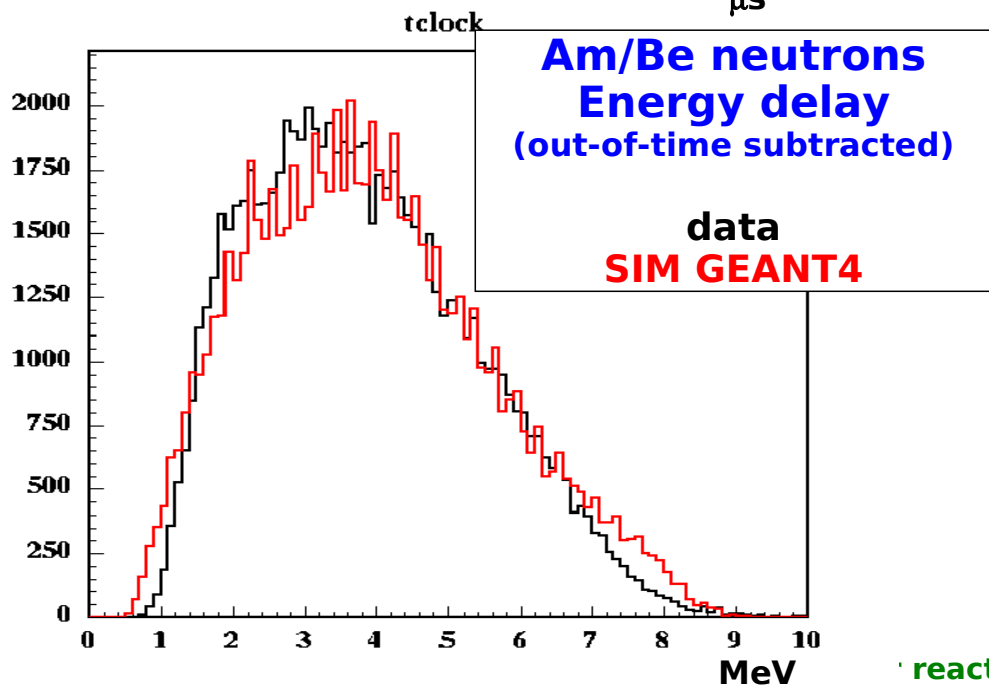
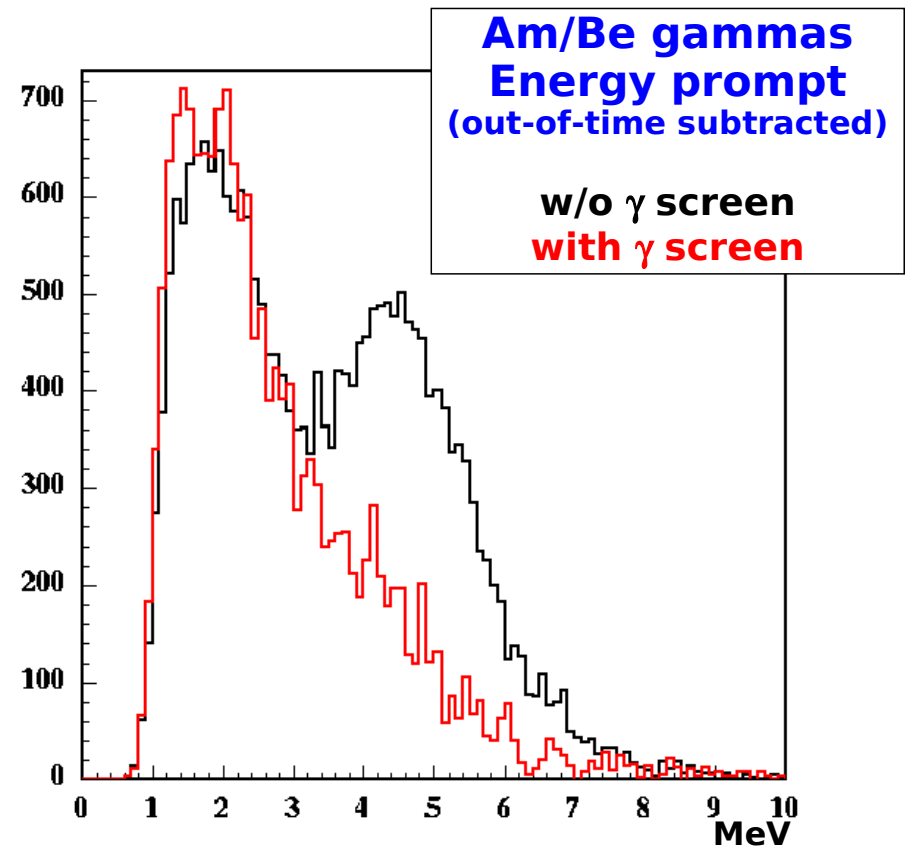
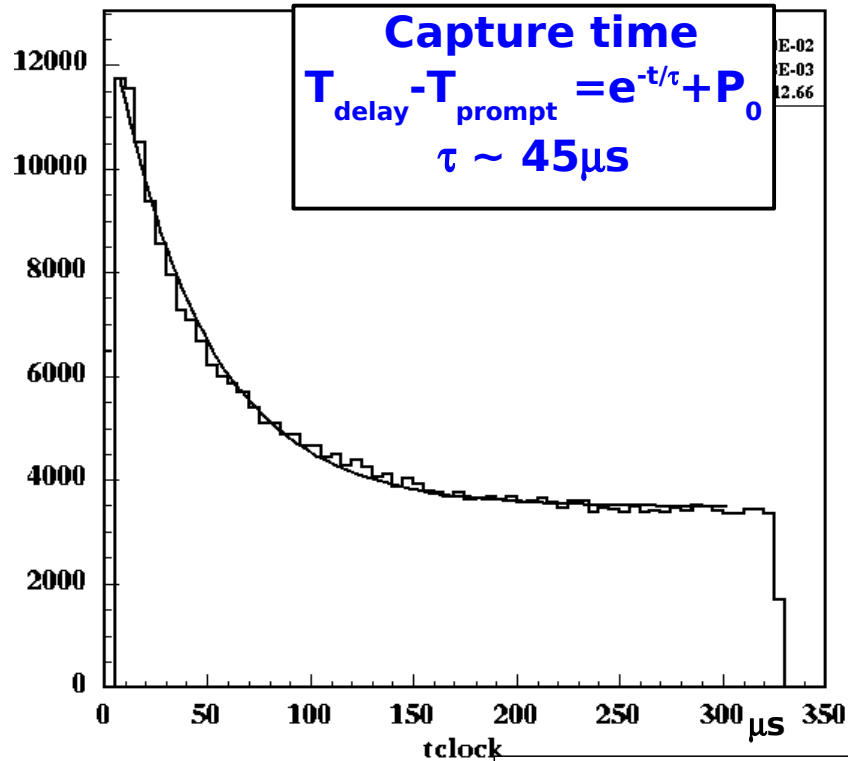
The real double-hit trigger is needed!

DAQ scheme (VME based)

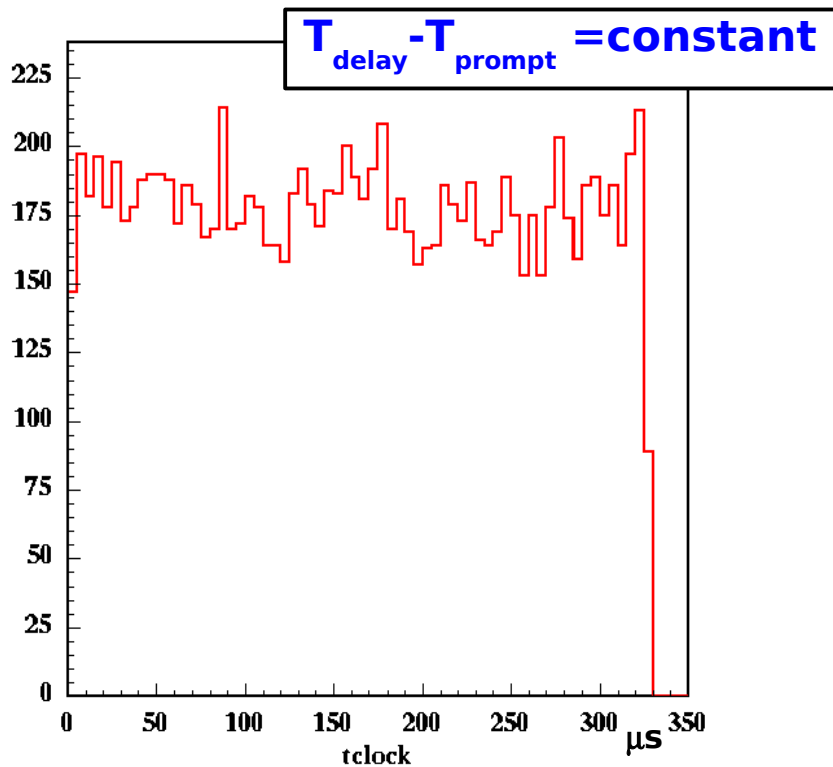
D. Piombo



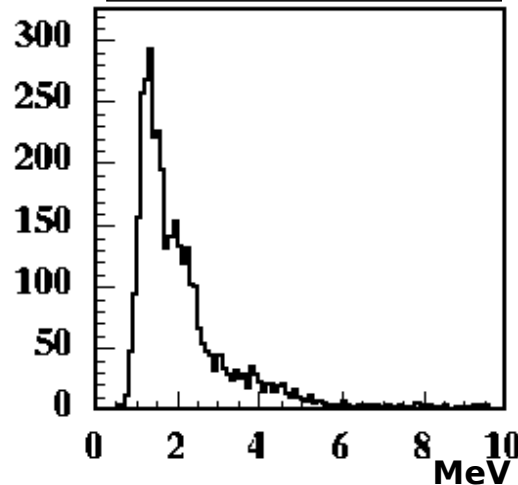
AmBe source



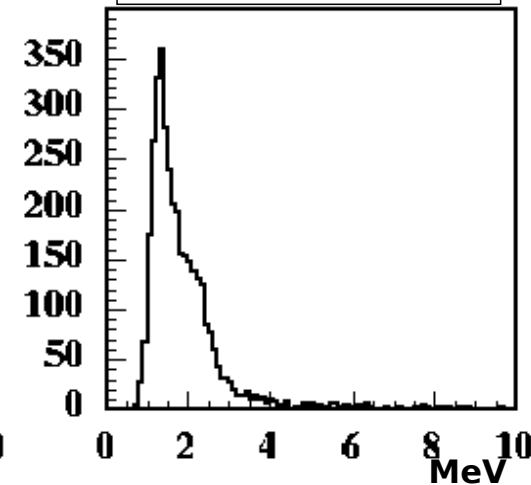
Background: cosmic rays



Cosmic rays
Energy prompt



Cosmic rays
Energy delay

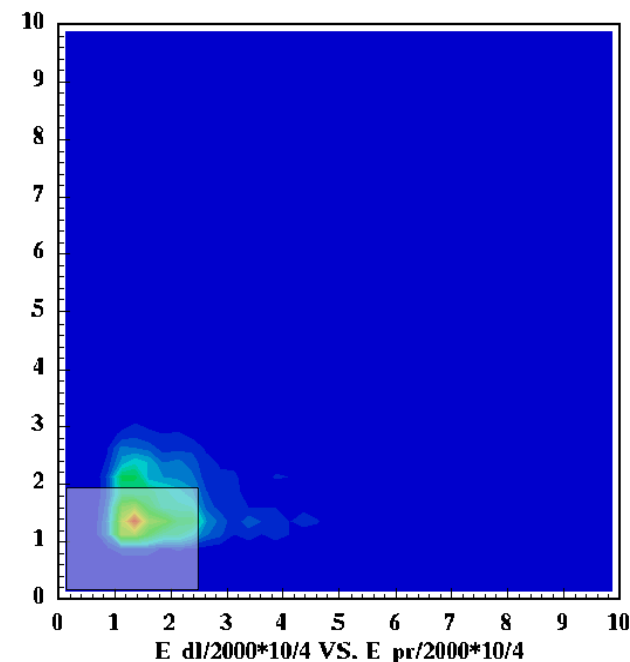
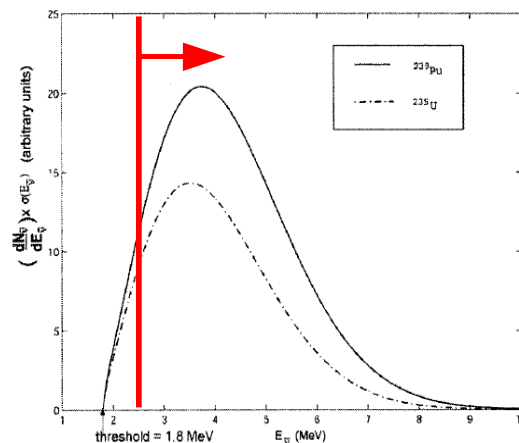


Cuts

- ★ $T < 100 \mu\text{s}$
- ★ $E_p > 2.5 \text{ MeV}$
- ★ $E_d > 2.0 \text{ MeV}$
- ★ Multiplicity $P \geq 2$
- ★ Multiplicity $D > 2$

Rejection factor:
~500

Bg	anti- ν
0.2%	20%



Need more detailed simulations to reduce bg (work in progress)

Measurement@ Cernavoda (Ro)



Test run in Cernavoda (Ro) SNN nuclear power plant during outage/restart in May 09

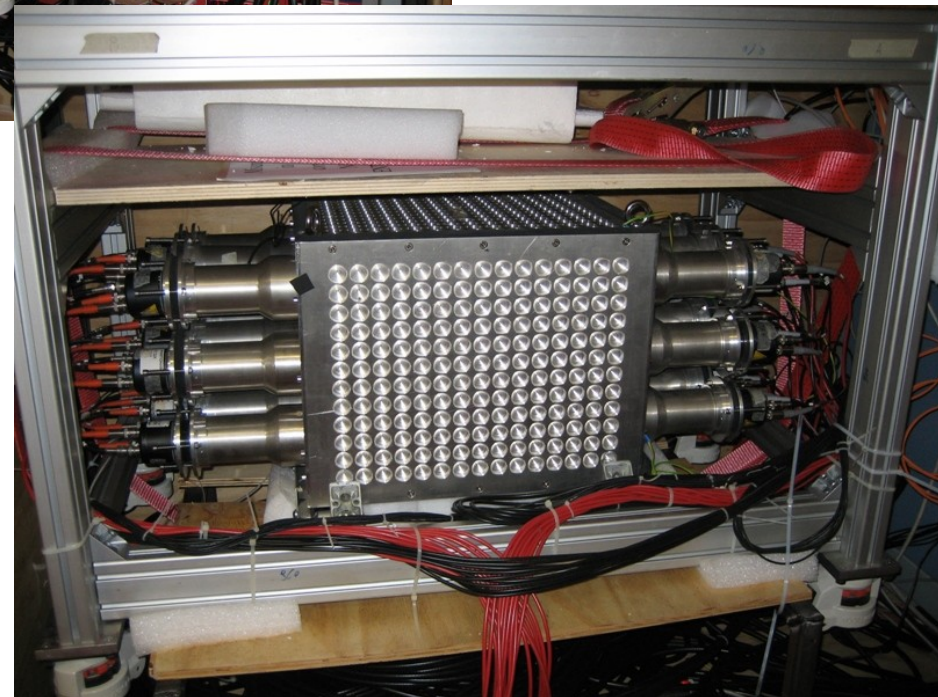
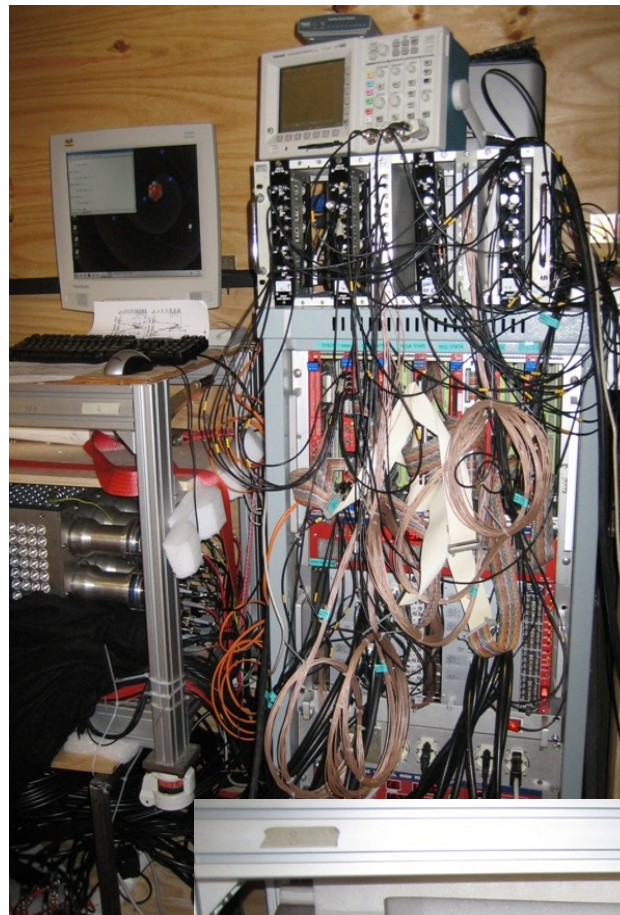
2 GW reactor off/on for maintenance
~20m from the reactor core
In-truck movable detector

Test Goals

In-situ measurement
Experimental set-up optimization
Background rates
Reactor On/Off change
MC validation
Data analysis optimization

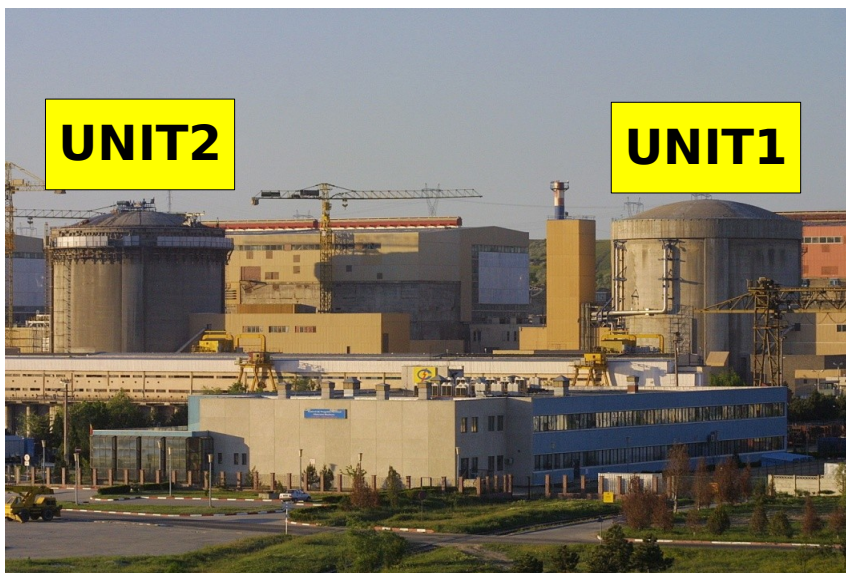
Expected rates/yields

antineutrino	cosmic	other
2000 ev/d/mq/GW	100 ev/d	?? ev/d
↓ 2GW 0.036mq Detection eff~20%	↓ based on home measurement off-line analysis rejection ~ 500	↓ spallation n residual n flux residual γ flux
30 ev/d x 25d = 750 ev	x 25d = 2500 ev	x 25d = ??? ev



21)

An antineutrino detector to monitor nuclear reactor's power and fuel composition - M.Battaglieri - INFN Genova



★ **Detector and electronics mounted in a van and moved to Cernavoda (Ro)**

★ **The van has been parked in front of Unit2 where will remain up to June 25th**

★ **After 2 days of commissioning it is now taking data with the reactor off**

★ **The reactor will be resumed on June 5th**

★ **We expect to collect data**
- 10 days reactor off
- 20 days reactor on

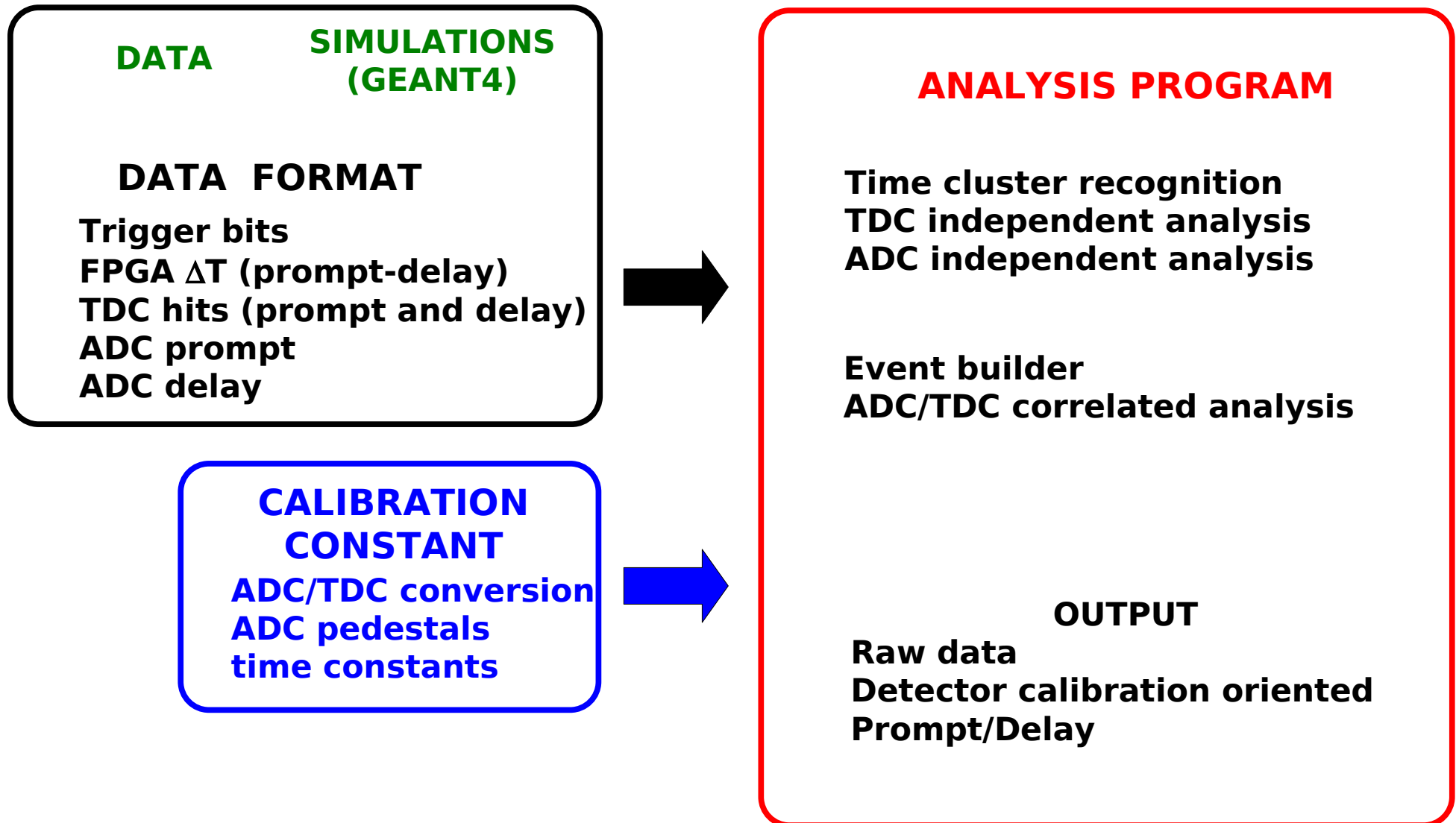


Conclusions

- **Antineutrinos can be used to monitor power and isotopic composition of a power nuclear reactor**
- **Antineutrino are detected via inverse-beta decay measuring the (fast) positron and the (delayed) neutron**
- **The 1 mq detector we proposed is made by plastic scintillator bars wrapped in thin gadolinium foils**
- **Segmentation (individual bar read-out) helps in reducing the background**
- **Extensive GEANT4 simulations shows we expect a 40% efficiency (power monitoring on day base and isotopic composition integrating 15 days of data)**
- **A prototype, scale 1:3, has been built in Genova and after lab tests, it has been installed in the Cernavoda (Ro) NPP where is taking data up to the end of June**

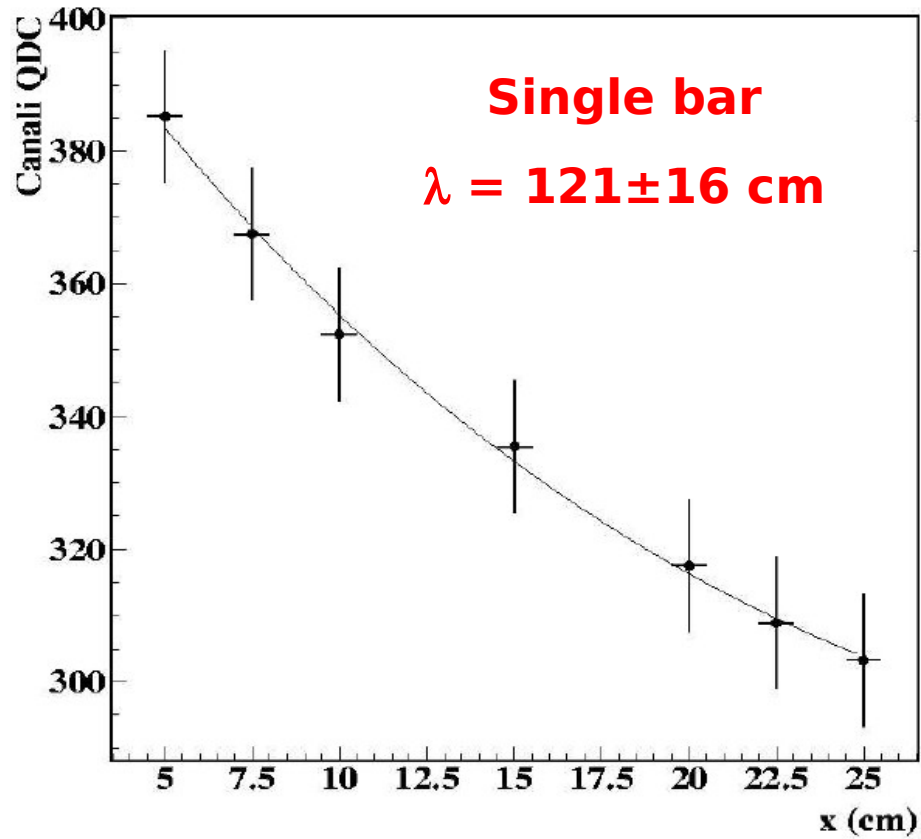
Back-up slides

Data analysis

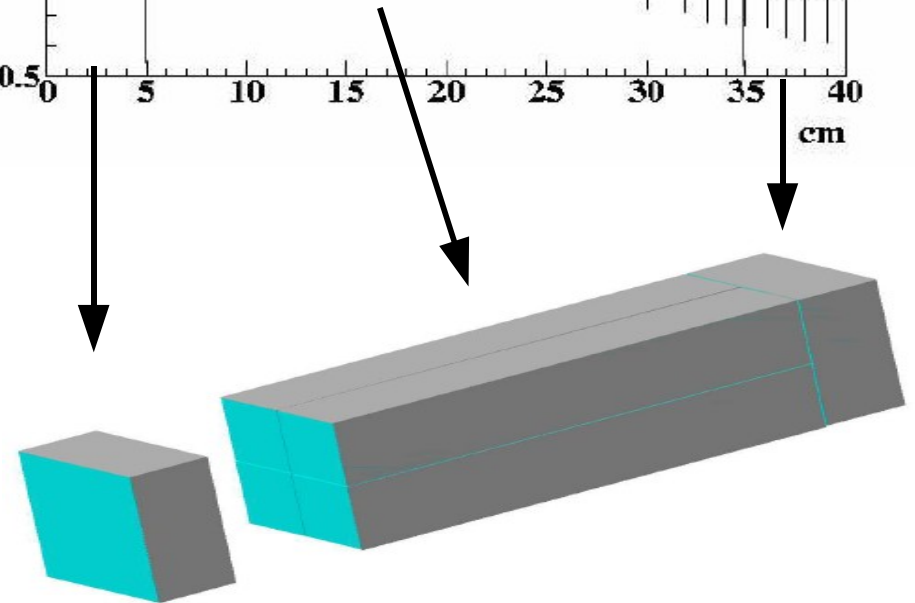
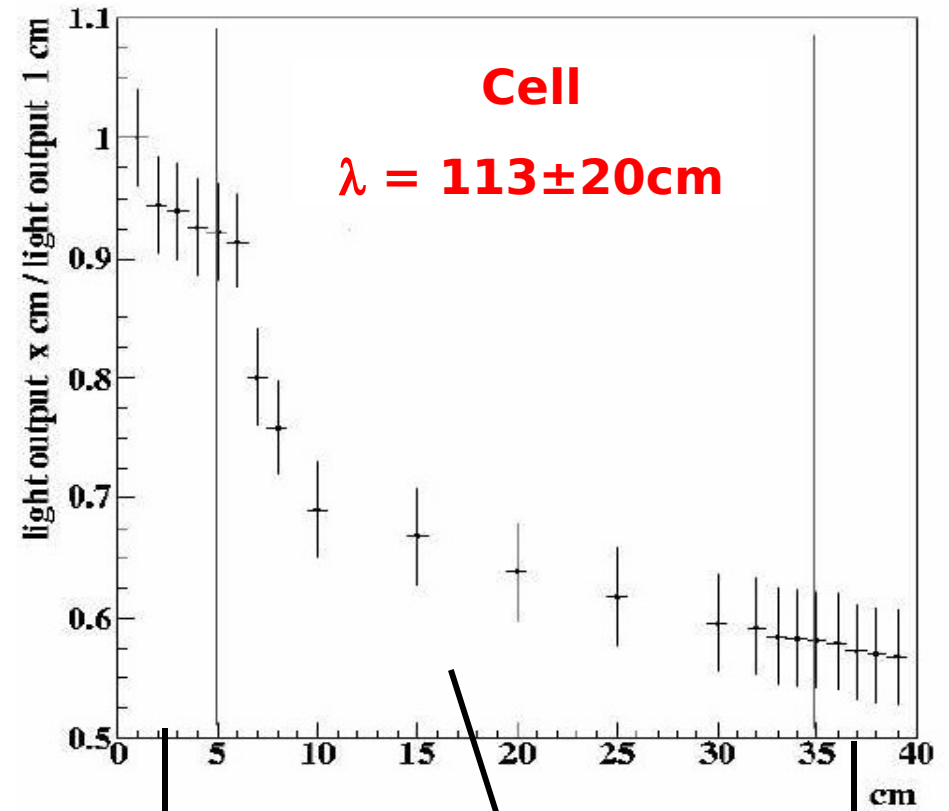


Attenuation length

G. Firpo

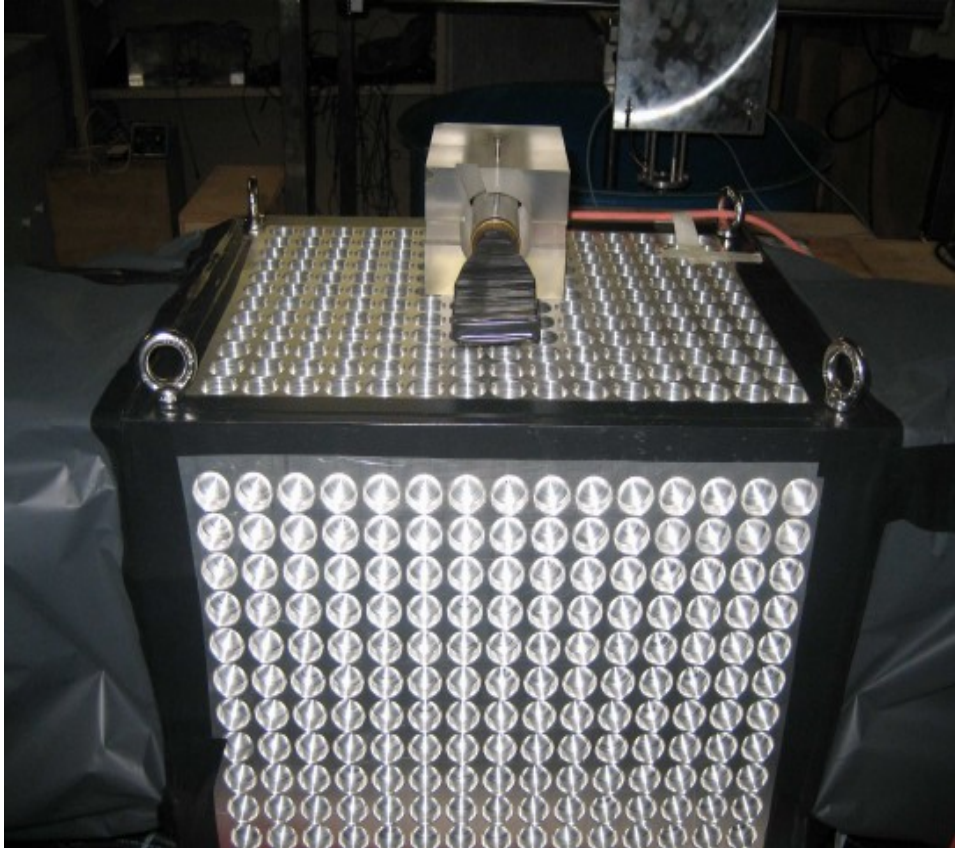


10% loss at the interface

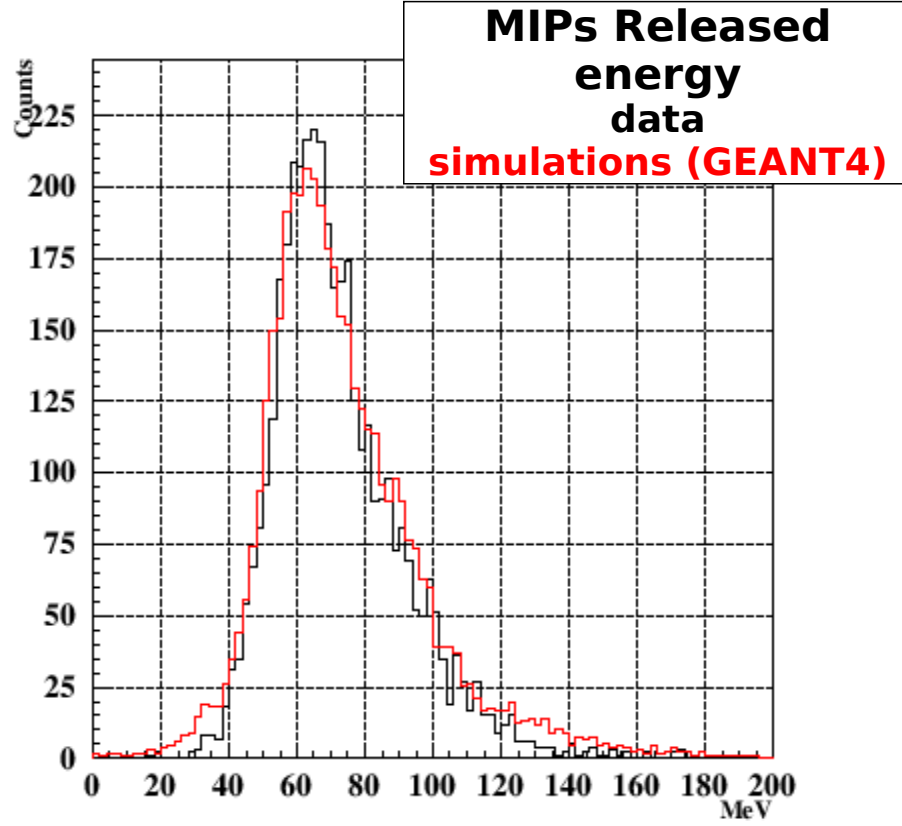
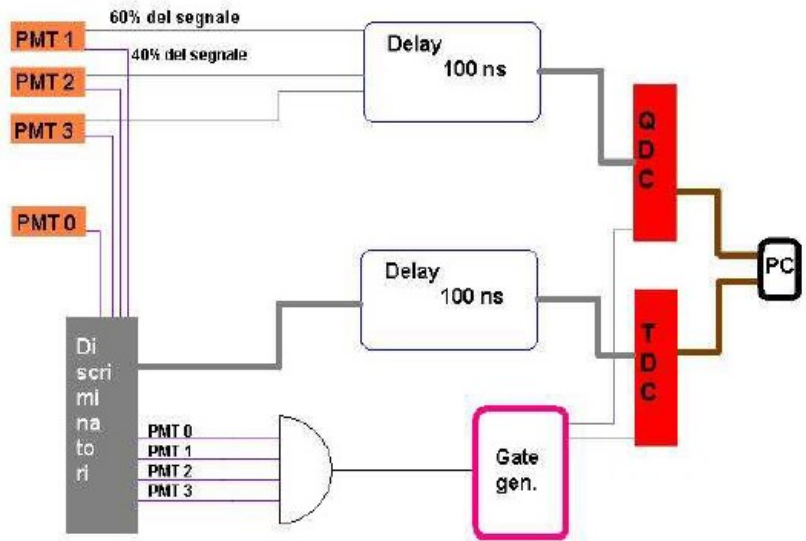


Calibrations: Cosmic rays

G. Firpo



Single hit trigger



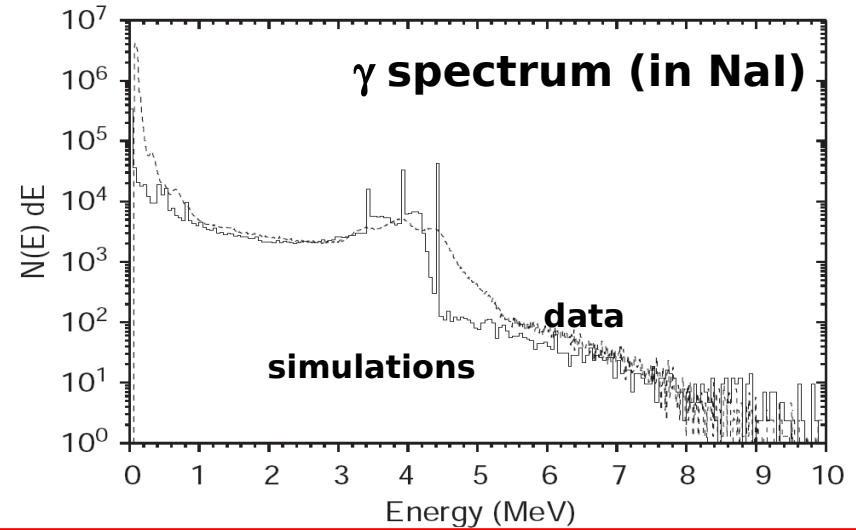
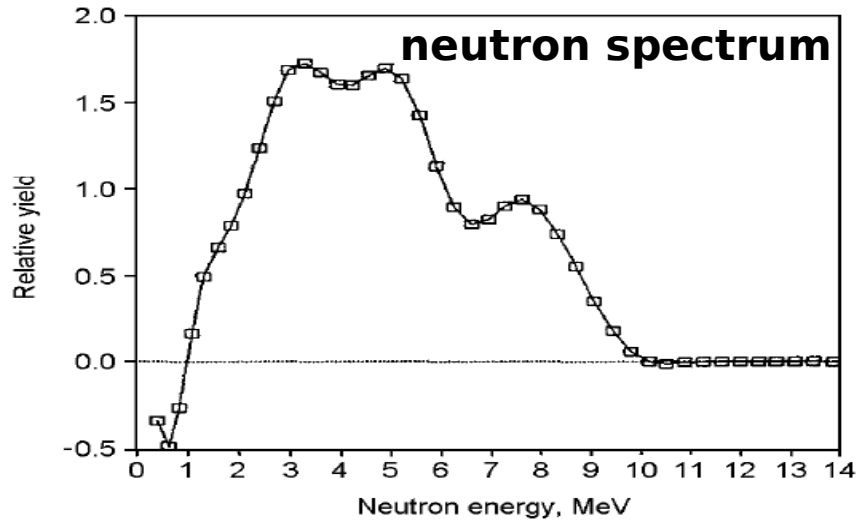
Calibrations: AmBe source

^{241}Am is an intense α emitter \rightarrow $^9\text{Be} (\alpha, n) ^{12}\text{C}$

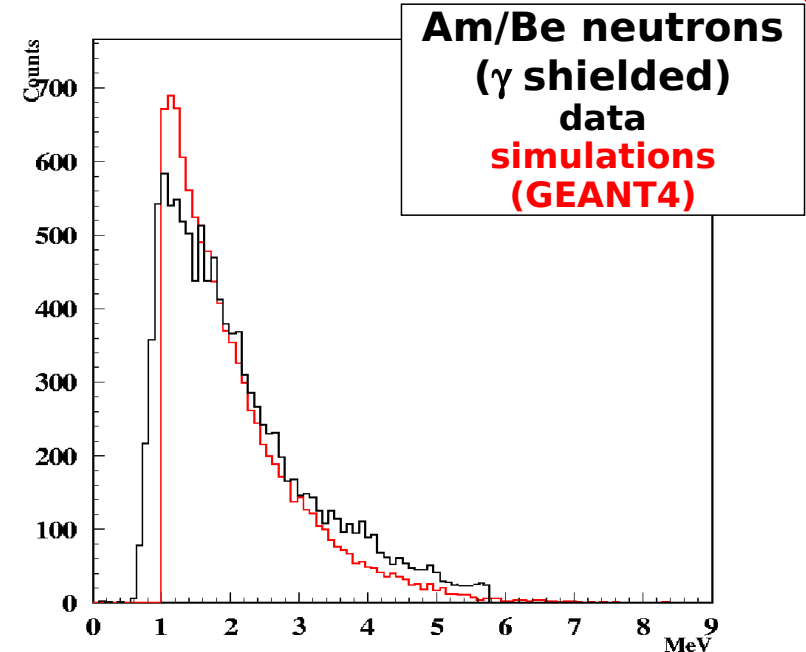
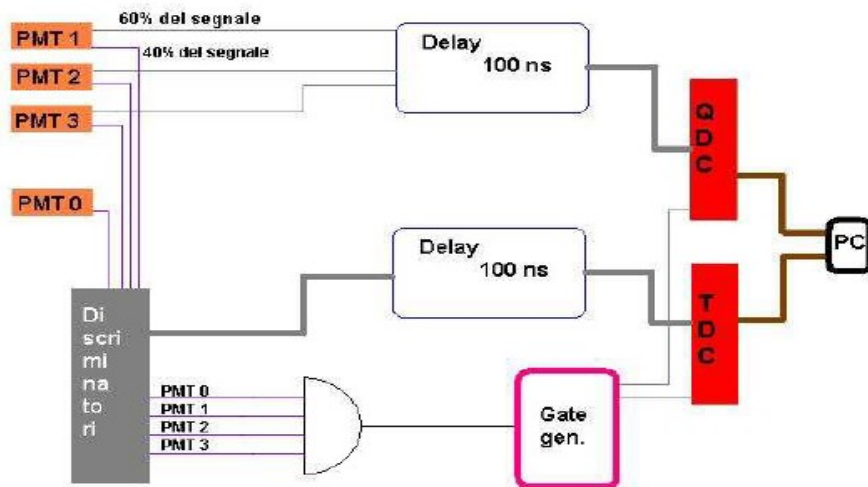
60% of n are emitted with a 4.4 MeV γ (first excited 2+ state of ^{12}C)

Encapsulated source has a more complex spectrum

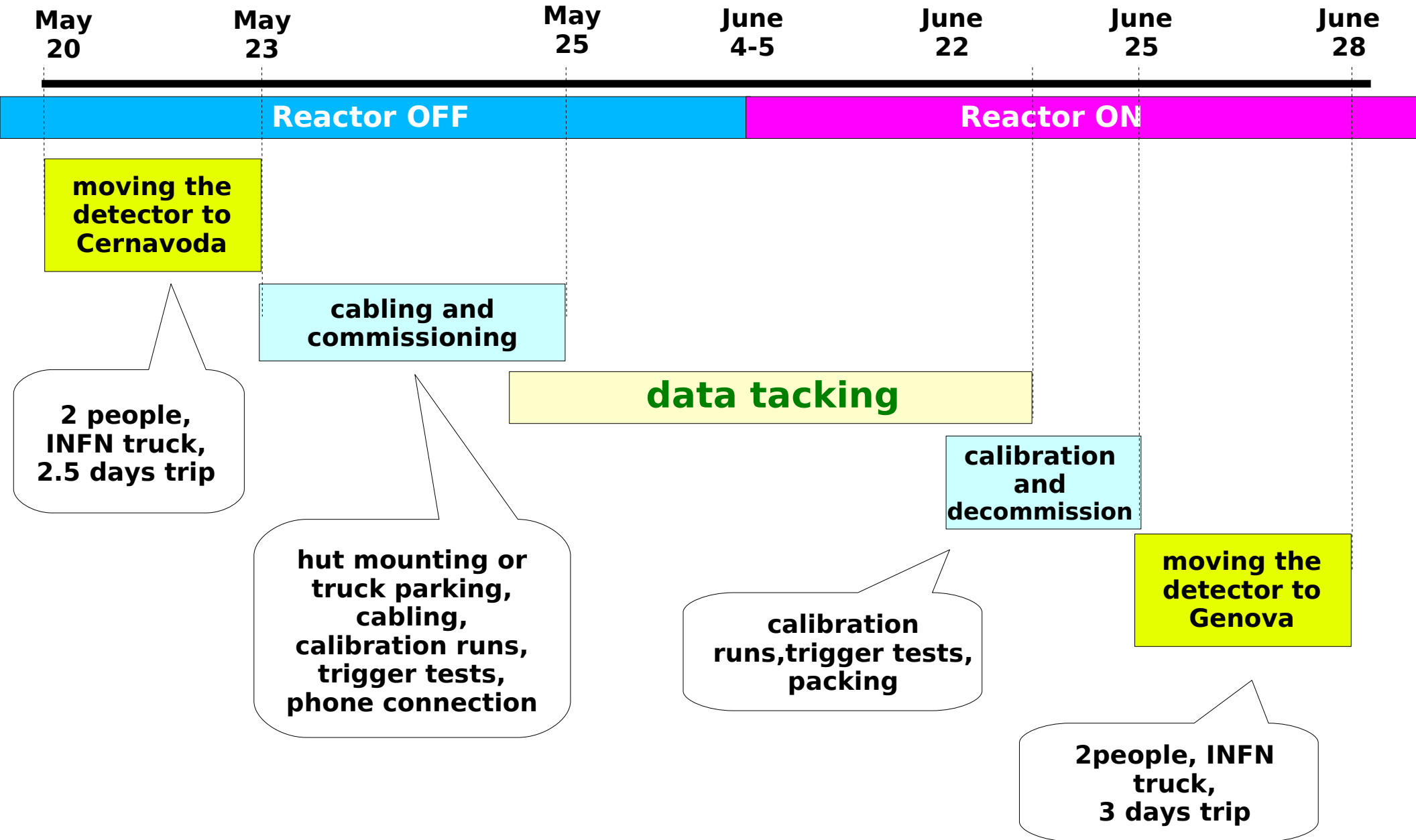
Neutron rate: $5 \cdot 10^3$ n/s



Single hit trigger: neutron detection

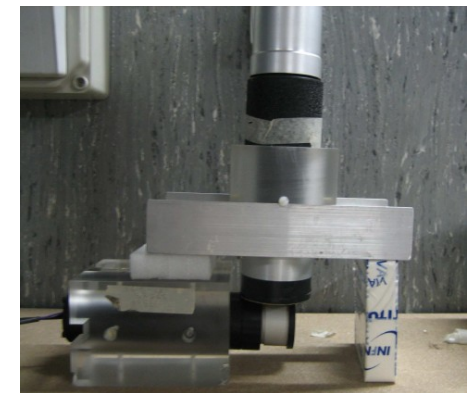
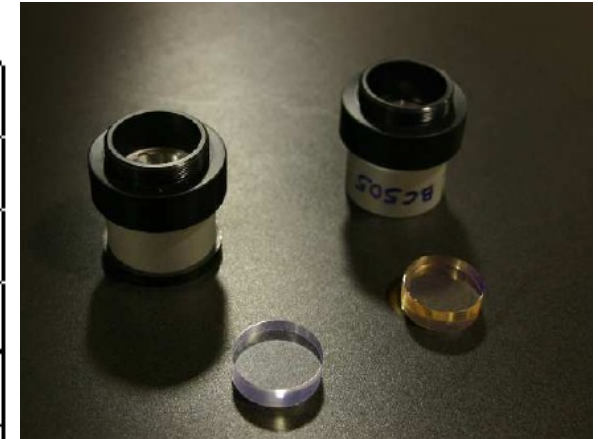


Tests in Cernavoda (Ro) Workplan



Gd- and B-doped plastic scintillators

Scintillatore	^{137}Cs (canali QDC)	Risol. en. (%)	^{60}Co (canali QDC)	Risol. en. (%)
BC505	1077 ± 9	11	1959 ± 8	7
BC517H	766 ± 5	19	1455 ± 7	9
BC525 (Gd)	736 ± 7	20	1424 ± 5	10
BC531	876 ± 5	19	1678 ± 4	10
B 0%	1153 ± 7	9	1850 ± 18	8
B 0.38%	1084 ± 7	11	1768 ± 6	7
B 5%	1044 ± 3	11	1716 ± 7	9
Gd 0%	508 ± 7	25	976 ± 8	18
Gd 1%	410 ± 7	35	885 ± 6	27
Gd 3%	302 ± 6	33	609 ± 8	26
Pol 1	990 ± 3	10	1584 ± 6	7
Pol 2	989 ± 3	10	1577 ± 7	7



Can we do it better?

Previous doped liquid scintillator showed aging problems
Need testing commercial available samples (BC-525)

Plastic scintillator advantages:
chemical stability longer in time
easier assembling
null chemical hazard

Different possible solutions:

1) standard plastic scintillator + Cd foils or Gd /foils-paint

Explored, G4 simulations done, reasonable results, easy mechanic assembly
but a high segmentation is required to not degrade energy resolution and light collection (higher costs?)

2) doped plastic

Commercial available: B-loaded plastic scintillator Saint-Gobain BC-454 (5%) but
10 x 10 x 100cm = 200k euro

Research group: Laboratori Nazionali di Legnaro (Italy)

Research group: Oak Ridge National Lab (USA)

Research group: Dubna (Russia)

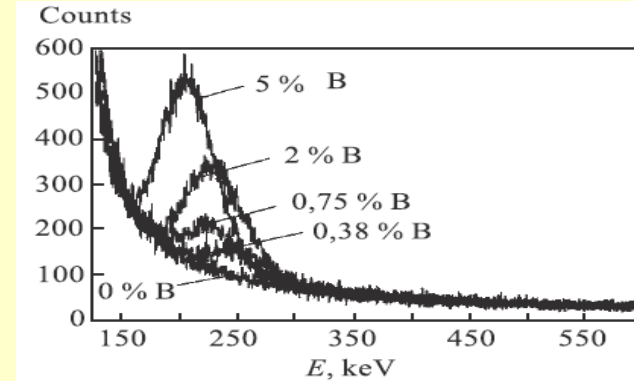
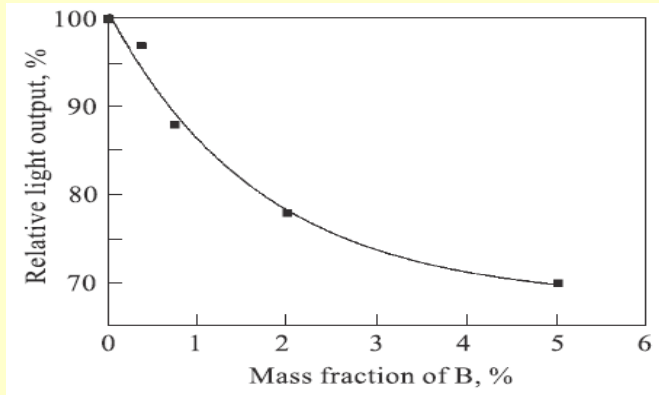
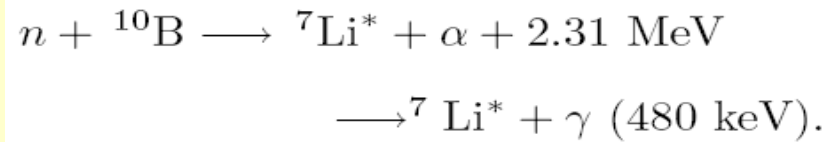
J INR-Dubna Group neutron detectors

1) B-loaded plastic scintillator (0,5% - 5%)

Polistyrene + p-Terphenil +POPOP

high light output (70% of unloaded with 5% B)

small sample 3 x 1 cm (reduced light emission of 1.5 MeV α particle)

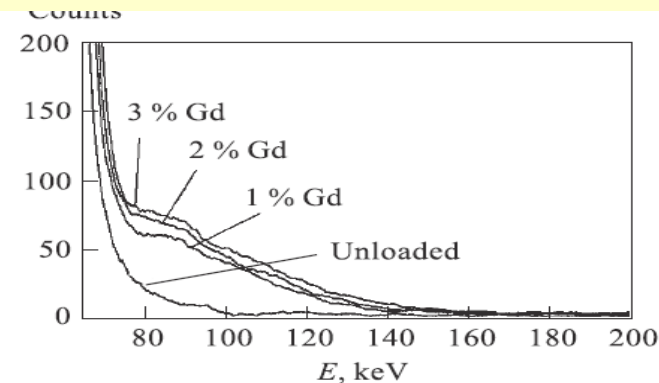
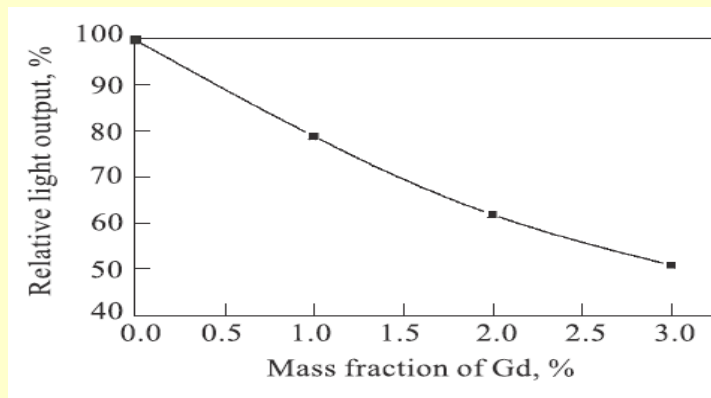


2) Gd-loaded plastic scintillator (0,5% - 3%)

PMMA Naphtalene +PPO +POPOP

high light output (60% of unloaded with 2% Gd)

small sample (3 x 1 cm) does not absorb the whole gamma cascade
claimed to be able to produce long bars (10 x 10 x 100 cm)



Oak Ridge neutron detectors



Two neutron sensors:

1) Silicone rubber Gd-loaded (1%)

- good radiation hardness
- high T operation (up to 200°C)
- small samples (~1cm)
- not cheap (1\$/g ?)
- small attenuation length (10-20 cm ?)
- need a container

2) Doped plastic (PS and PVT +PPO + Gd 1-1,5%)

- small samples
- good light transmission
- long stability (up to 6 years)

B. Zane from ORNL was contacted:
'Cheaper plastic can be home-made
manufactured but needs more R&D'

