



The EXO-200 Detector

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Frontier Detectors for Frontier Physics

11th Pisa meeting on advanced detectors

24-30 May 2009

Double Beta Decay ...

- Rare nuclear transition between same mass nuclei
 - Energetically allowed for even-even nuclei

$$\beta\beta 2\nu : \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + \bar{\nu}_1 + e_2^- + \bar{\nu}_2$$

Allowed in SM and
already observed!

$$\left[T_{1/2}^{2\nu}(0^+ \rightarrow 0^+) \right]^{-1} = G^{2\nu}(Q_{\beta\beta}, Z) \left| M^{2\nu} \right|^2$$

$$\beta\beta 0\nu : \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + e_2^-$$

Neutrinos are Majorana particles!

$$\Delta L = 2 \quad (Z, A) \rightarrow (Z + 2, A) + e_1^- + e_2^- + \chi$$

$$\nu \equiv \bar{\nu} \quad m_\nu \neq 0$$

$$\left[T_{1/2}^{0\nu}(0^+ \rightarrow 0^+) \right]^{-1} = G^{0\nu}(Q_{\beta\beta}, Z) \left| M^{0\nu} \right|^2 \langle m_{\beta\beta} \rangle^2 \quad \langle m_{\beta\beta} \rangle^2 = \left| \sum_k m_k U_{ek}^2 \right|^2$$

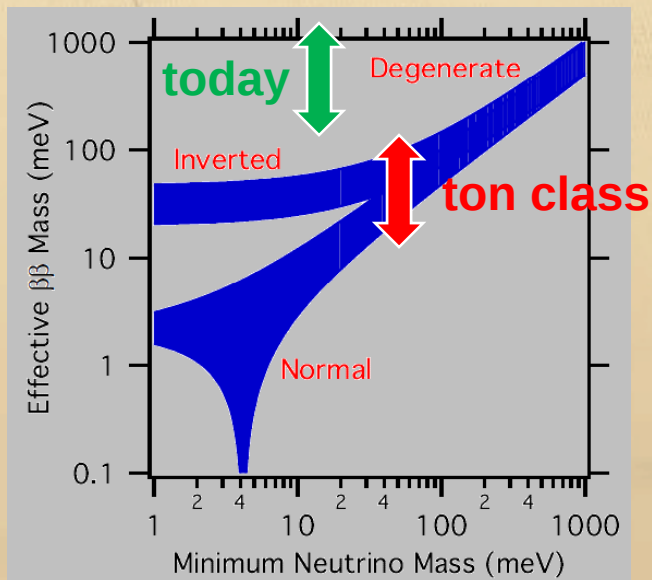
... and its Role in Neutrino Physics

Double beta decay experiments are part of a massive effort to determine the nature and properties of neutrinos!

$$\begin{aligned}
 & \langle m_\beta \rangle < 2eV \quad \text{Beta Decay Endpoint} \quad \Delta m_{23}^2 = (2.4^{+0.6}_{-0.5}) \times 10^{-3} eV \quad \text{Cosmological Constraints } \Sigma < 2eV \\
 & \langle m_\beta \rangle^2 = \sum_k m_k^2 |U_{ek}|^2 \quad \theta_{23} \approx 45^\circ \quad \text{Neutrino Oscillations} \quad \Sigma = \sum_k m_k = 92.5eV \times (\Omega_\nu h^2) \\
 & \quad \quad \quad \text{Atmospheric and Reactor} \quad \theta_{13} < 7^\circ \quad \text{Neutrino Oscillations} \quad \text{CP violation} \\
 & U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{-i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \dots \\
 & \theta_{12} \approx 34^\circ \quad \text{Neutrino Oscillations} \quad \text{Solar and Reactor} \quad \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\alpha/2} & 0 & 0 \\ 0 & e^{i\beta/2} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \text{Neutrinoless Double Beta Decay Only!} \\
 & \Delta m_{12}^2 = (8.0^{+0.4}_{-0.3}) \times 10^{-5} eV \quad \langle m_{\beta\beta} \rangle < 0.7eV
 \end{aligned}$$

Experimental Requirements

Effective neutrino mass as a function of the smallest neutrino mass for various scenarios



- **Large Mass:** at least 100 kg of source isotope
 - Scanning the quasi degenerate region
 - Ton scale for the inverted hierarchy region
 - Enrichment helps minimizing volume and improves source purity
- **Very Low Background:** 1 count per ton per year range
 - Survey, selection and purification of materials and components
 - Cleanroom assembly and detector operation
 - Deep underground installation and muon veto required
- **Very Good Energy Resolution:** in the 1% range
 - Limits the allowed double beta decay background
 - Increases signal to radioactive background ratio
- **And ...**
 - Large $Q_{\beta\beta}$ to have the signal out of the region densely populated by radioactive background for natural chains
 - Tagging the daughter isotope would eliminate most radioactive background or event topology and advanced kinematics details

EXO Project & EXO-200 Phase

- EXO project searches for double beta decay using ^{136}Xe
 - Ton scale implementation either as liquid or gas phase TPC
 - Relatively large Q value and straight forward enrichment technique
 - ^{136}Ba daughter tagging either in-situ or in external RF cage

$$\langle m_{\beta\beta} \rangle \propto \left(\frac{1}{Nt} \right)^{1/4} \text{👾}$$

$$\text{No Background! } \langle m_{\beta\beta} \rangle \propto \sqrt{\frac{1}{Nt}} \text{👾}$$

- EXO-200 is the first phase using 200 kg of 80% enriched Xe
 - Major R&D effort precursory to the ton-scale experiment
 - Exploration of the quasi-degenerate region with ^{136}Xe
 - Allowed double beta decay never observed in xenon!
 - No tagging but massive progress for radioactive background reduction and energy resolution improvement (easily scalable to future detectors)

EXO Collaboration

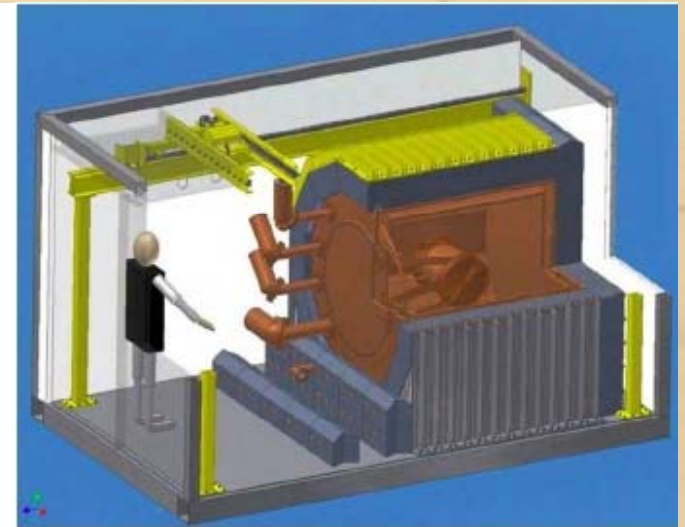
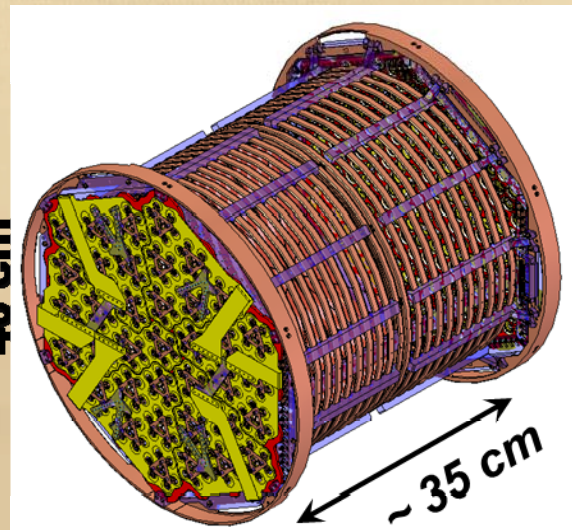
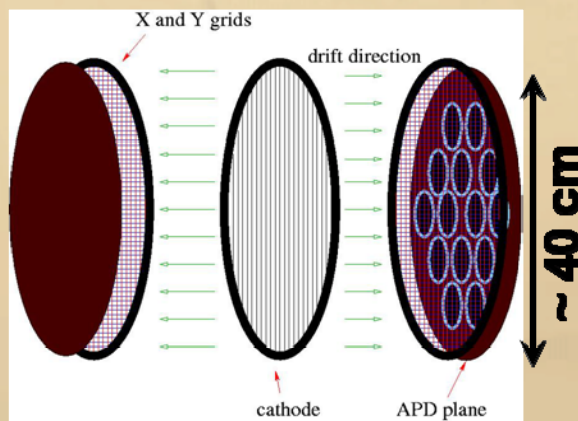
K.Barry, D.Leonard, E.Niner, A.Piepkne from **Physics Dept, University of Alabama, Tuscaloosa AL, USA**
P.Vogel from **Physics Dept Caltech, Pasadena CA, USA**
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L.Yang, S.Zalog from **SLAC, Menlo Park CA, USA**
J.Anthony, L.Bartoszek, R.DeVoe, P.Fierlinger, B.Flatt, G.Gratta, M.Green, S.Kolkowitz, F.LePort, M.Montero-Diez, R.Neilson, A. Rives,
K.O'Sullivan, K.Twilker from **Physics Dept, Stanford University, Stanford CA, USA**



EXO-200 Detector

- Liquid xenon TPC with two cylindrical drift volumes
 - Charge collection using 114 by 114 wire planes (at 60° pitch)
 - Scintillation light readout using 37 groups of 7 bare LAAPD (Large Area Avalanche Photodiodes) at both end caps
- High purity copper cryostat with external refrigeration-based cooling

~ 115 kg fiducial mass

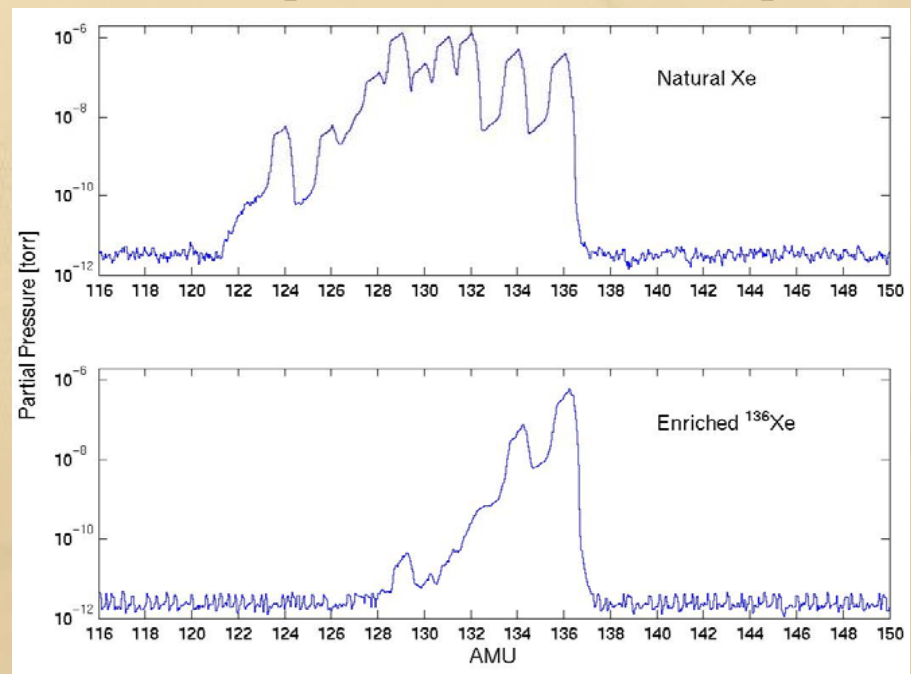


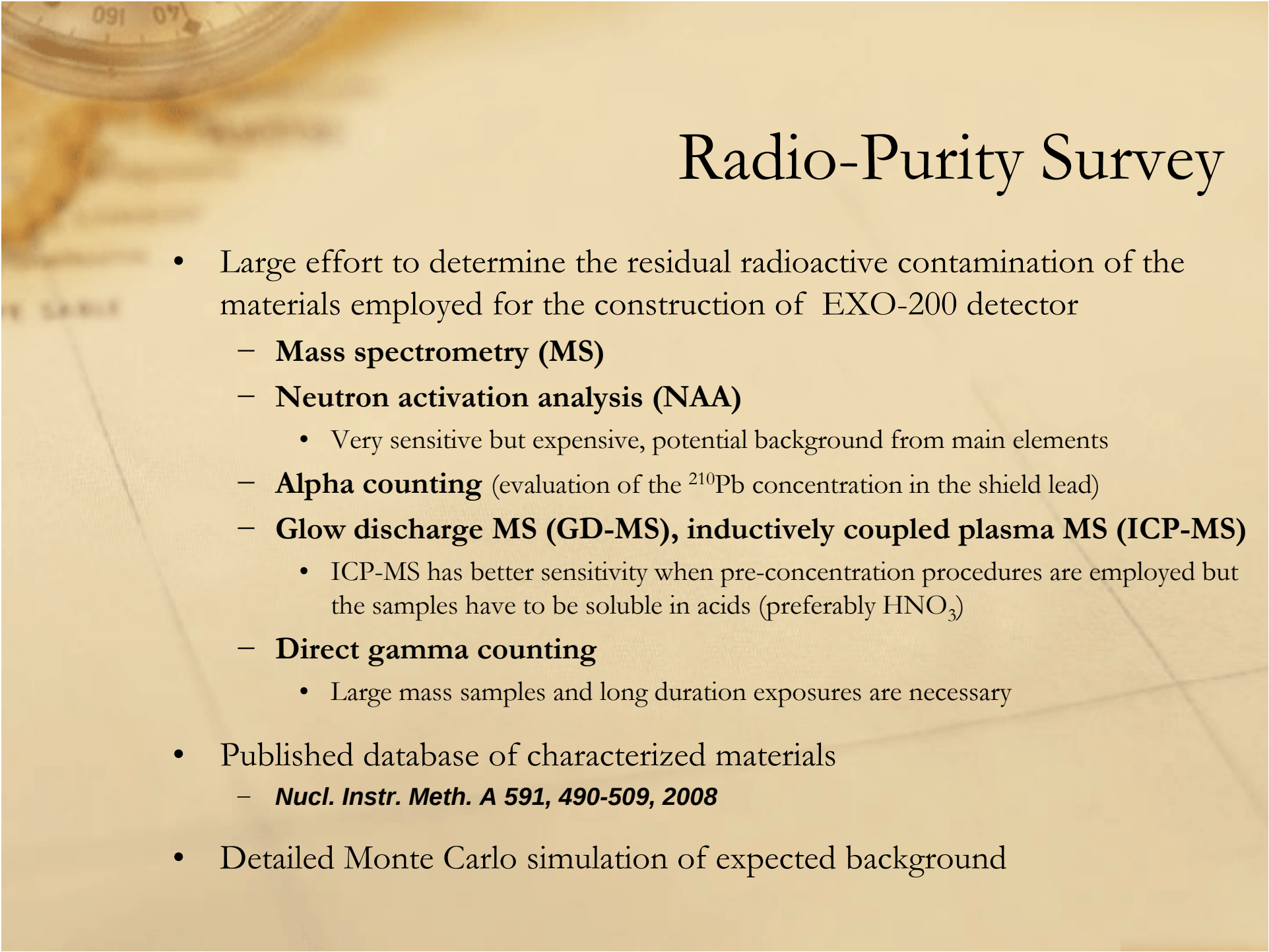
Successful Enrichment Program

200 kg of 80% enriched Xe delivered in 2003!

Used mass-separating centrifuges in Russia

The other isotopes can be returned to provider!





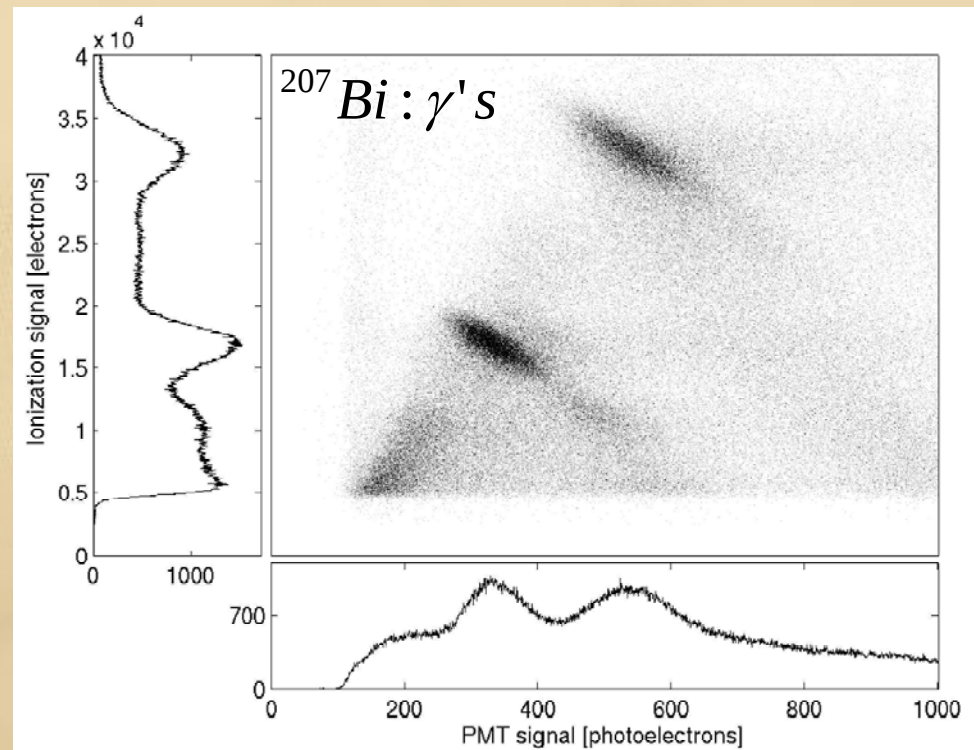
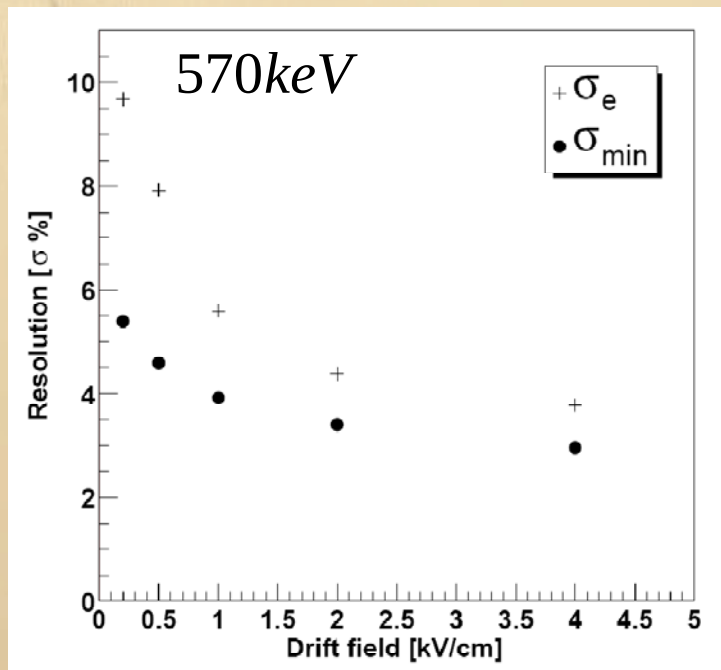
Radio-Purity Survey

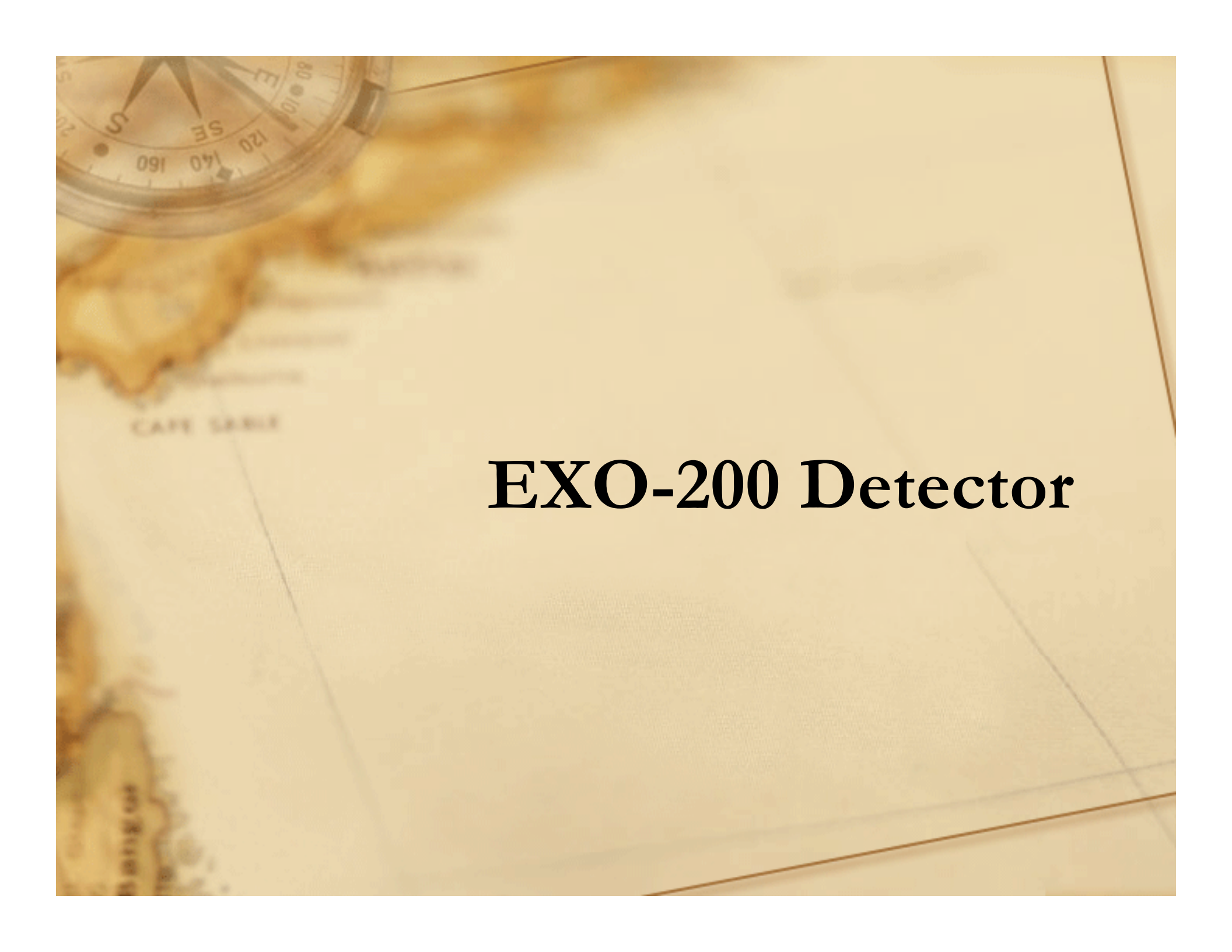
- Large effort to determine the residual radioactive contamination of the materials employed for the construction of EXO-200 detector
 - **Mass spectrometry (MS)**
 - **Neutron activation analysis (NAA)**
 - Very sensitive but expensive, potential background from main elements
 - **Alpha counting** (evaluation of the ^{210}Pb concentration in the shield lead)
 - **Glow discharge MS (GD-MS), inductively coupled plasma MS (ICP-MS)**
 - ICP-MS has better sensitivity when pre-concentration procedures are employed but the samples have to be soluble in acids (preferably HNO_3)
 - **Direct gamma counting**
 - Large mass samples and long duration exposures are necessary
- Published database of characterized materials
 - ***Nucl. Instr. Meth. A 591, 490-509, 2008***
- Detailed Monte Carlo simulation of expected background

Improving the Energy Resolution

Strong anti-correlation between ionization and scintillation signals in liquid xenon has been observed!

$$\frac{\Delta E}{E} = 1.4\% @ Q_{\beta\beta} = 2479\text{keV}$$

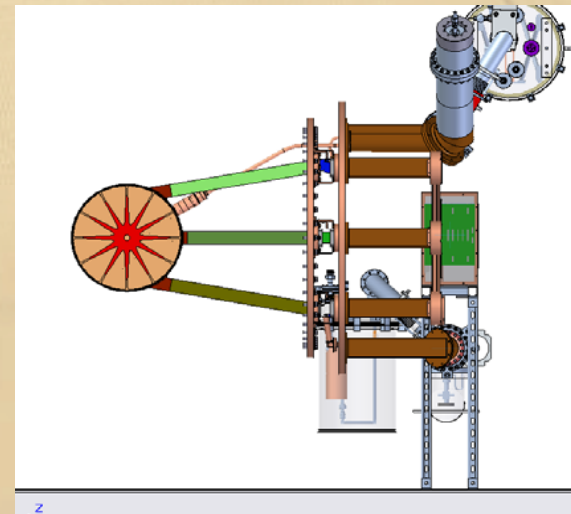
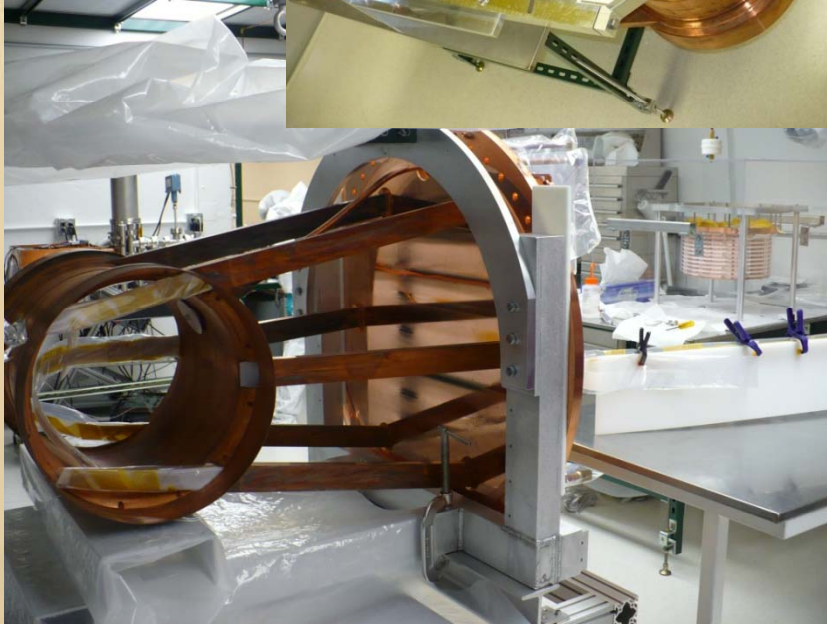
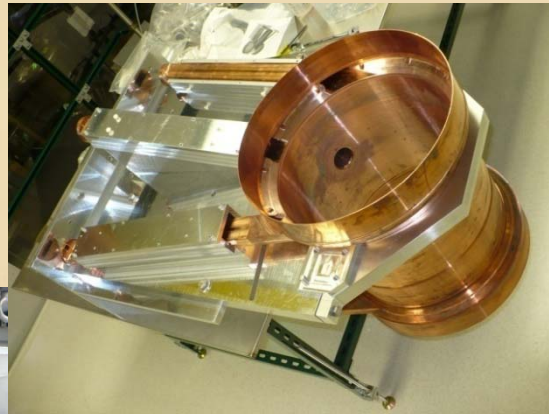




EXO-200 Detector

EXO-200 Chamber

Ultra low radioactivity copper!
Shielded surface transport and storage
Only 1.5 mm thickness to reduce mass
PLC-based real-time pressure control
e-beam welded components
TIG welding for the final assembly



Charge and Light Readout

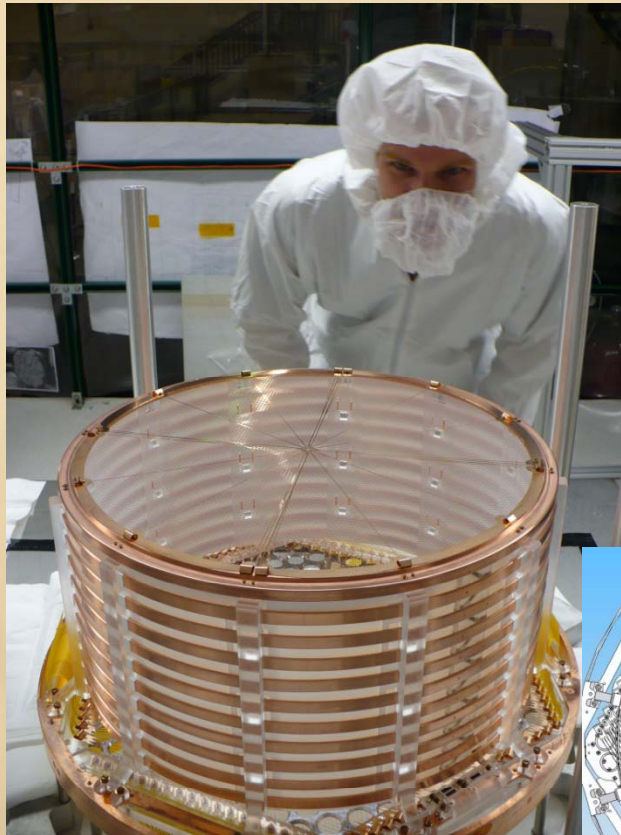
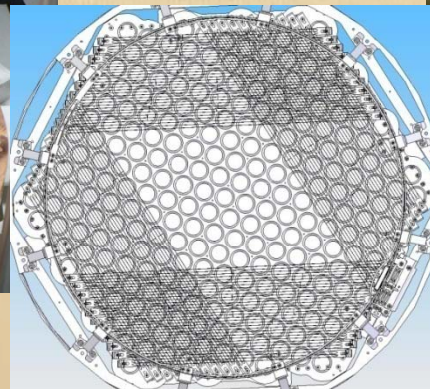


Photo-etched phosphor-bronze cathode
Induction & charge collection wire grids
259 LAAPD (37 groups of 7) per plane

- 1.6 cm active diameter
- very clean and low mass
- $QE > 1 @ 174 \text{ nm}$
- gain $100\times$ to $150\times @ \sim 1500\text{V}$

Radial Teflon UV light reflectors

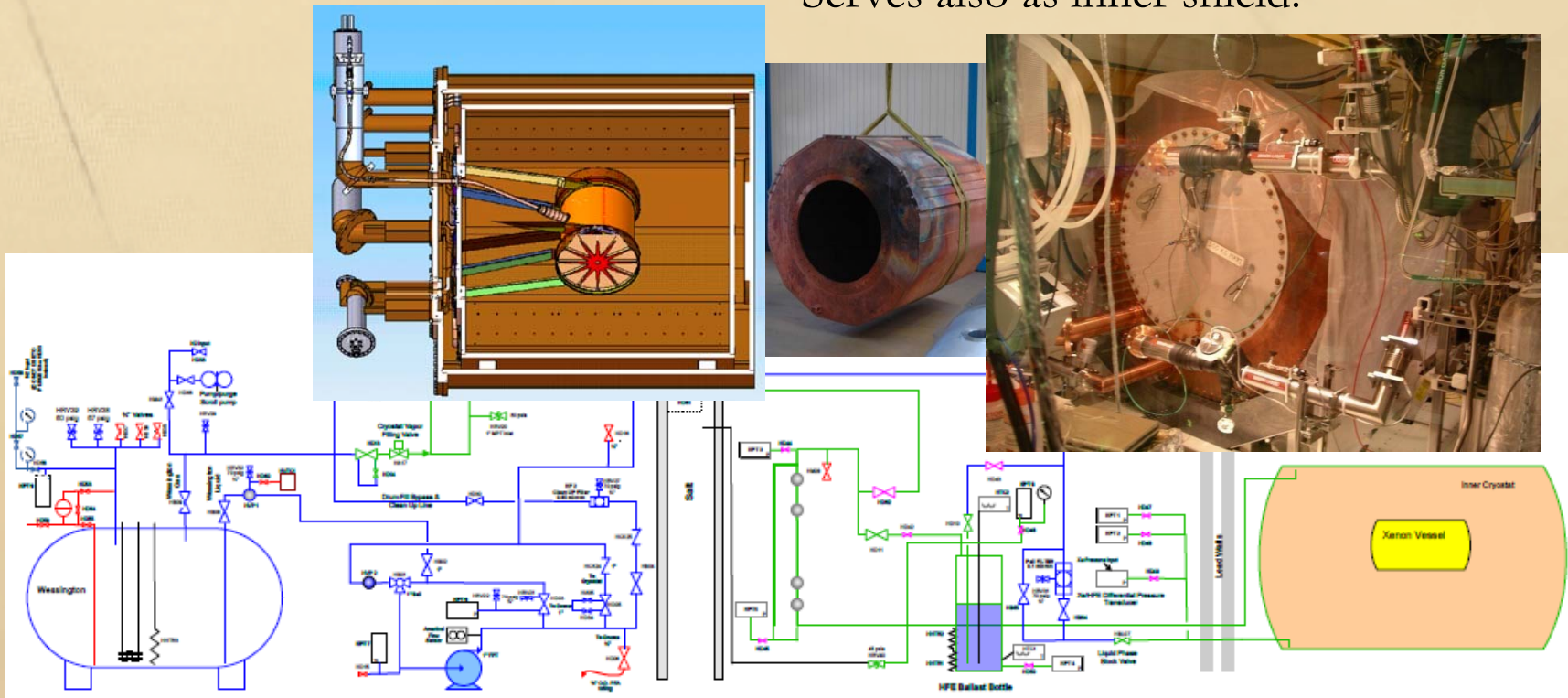


Chamber Assembly



Cryostat and Cooling System

Refrigeration based cooling (3×1500 W PolyCold units)
4.2 tons of high purity heat transfer fluid (3M HFE-7000)
Serves also as inner shield!



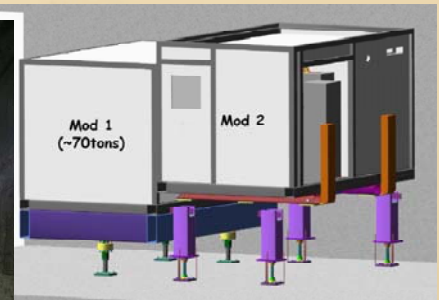
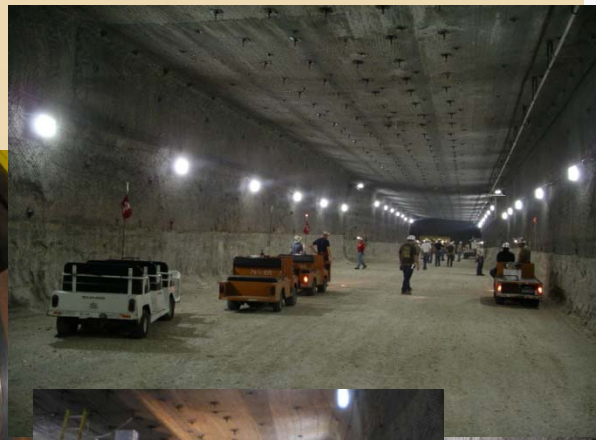


WIPP Installation

Experimental Area

Waste Isolation Pilot Plant, Carlsbad, New Mexico
~ 1600 m.w.e. (muon flux reduction by $\sim 10\times$)

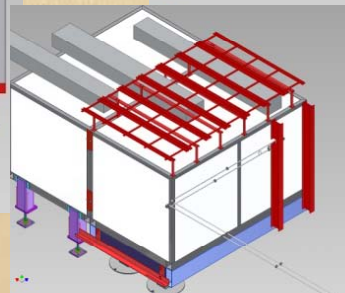
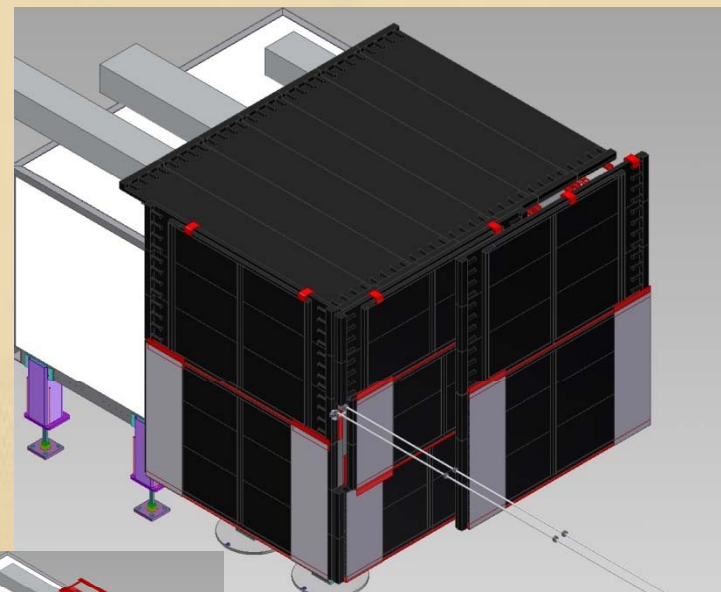
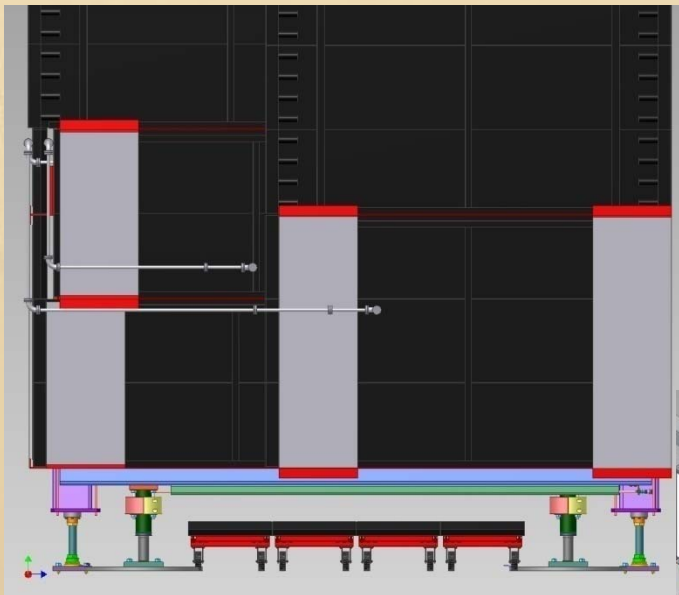
Large and wide remote experimental area available!



Part of the EXO-200 infrastructure at WIPP, NM

Muon Veto

20× muon induced background
reduction (99,7% detection efficiency)



Extensive Monte Carlo
simulations to optimize
the veto configuration!

Expected Performance

- Very low radioactive background
 - Careful selection of materials, optimized custom design
 - Manufacturing, handling and installation in cleanrooms
- Very good energy resolution $\frac{S}{B} = \frac{m_e}{7Q_{\beta\beta}} \left(\frac{E}{\Delta E} \right)^6 \frac{T_{1/2}^{2\nu}}{T_{1/2}^{0\nu}}$

Chamber underground installation in august 2009

Physics runs starting in 2010, 2 years run time!

$$T_{1/2}^{2\nu} > 1.2 \times 10^{24} \text{ y @ 90\% C.L.}$$

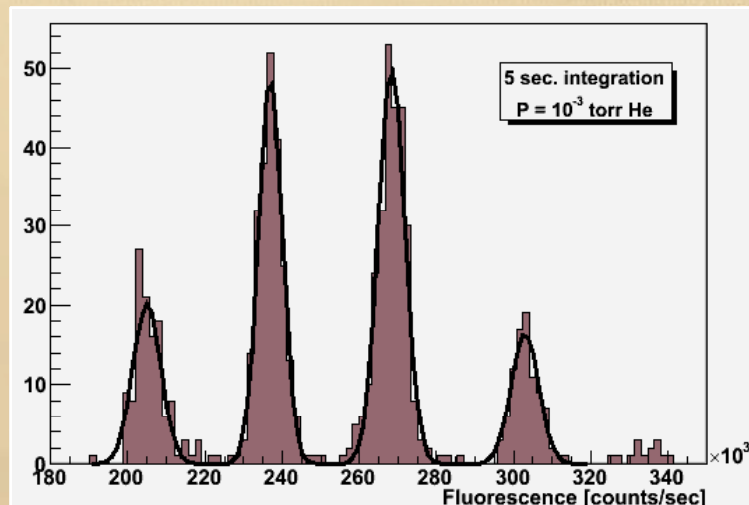
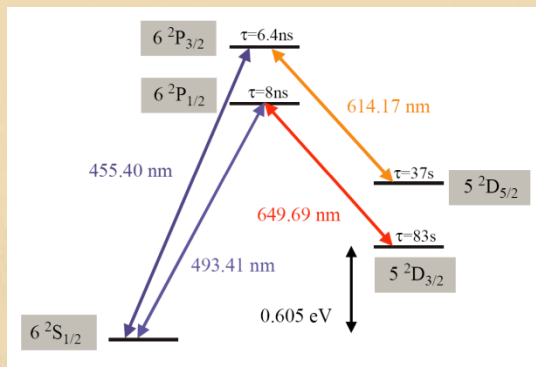
Case	Mass (ton)	Eff. (%)	Run Time (yr)	σ_E/E @ 2.5 MeV (%)	Radioactive Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (meV)	
							QRPA ¹	NSM ²
EXO-200	0.2	70	2	1.6	40	$6.4 * 10^{25}$	133	186

1) Rodin et. al., Nucl. Phys. A 793 (2007) 213

2) Caurier et. al., arXiv:0709.2137v1



Ba⁺ Tagging



RF cage with low pressure buffer gas

- Ba⁺⁺ → Ba⁺ conversion expected
 - Ionization potentials:
 - Xe⁺ = 12.13 eV vs. Ba⁺ = 5.21 eV
 - Xe⁺⁺ = 21.21 eV vs. Ba⁺⁺ = 10.00 eV
 - Solid Xe band gap (*Phys. Rev. B10 4464 1974*)
 - E_G = 9.22 +/- 0.01 eV
 - “Liquid Xe ionization potential” close to E_G (*J. Phys. C: Solid State Phys. Vol. 7 1974*)
 - 9.28 to 9.49 eV range
 - Use of additives for gas based detectors

Conclusion

- EXO-200 detector soon operational!
- The largest neutrino-less double beta decay detector!
- Successful large scale xenon enrichment proven!

